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From hydrostatic to non-hydrostatic models in fluid mechanics: modeling, mathematical and numerical analysis, and computational fluid dynamics

Mehmet Ersoy

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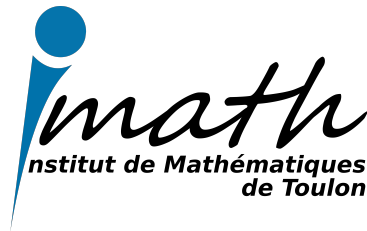
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HABILITATION À DIRIGER DES RECHERCHES

FROM HYDROSTATIC TO NON-HYDROSTATIC
MODELS IN FLUID MECHANICS: MODELING,
MATHEMATICAL AND NUMERICAL ANALYSIS,
AND COMPUTATIONAL FLUID DYNAMICS

présentée par

MEHMET ERSOY



FROM HYDROSTATIC TO NON-HYDROSTATIC MODELS
IN FLUID MECHANICS: MODELING, MATHEMATICAL
AND NUMERICAL ANALYSIS, AND COMPUTATIONAL
FLUID DYNAMICS

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par

MEHMET ERSOY

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La vie est loin d'être tranquille,
Elle est parsemée d'embruns et de méandres,
Fort heureusement, il y a la famille,
Elle m'a soutenue pour renaître de mes cendres.

— **Mehmet Ersoy**

Je dédie ce mémoire
à ma Famille,
à mes Parents,
à mon épouse ROSE,
à mes Enfants KAAAN, ENES et ALI

Mehmet Ersoy, La Garde, le 3 Décembre, 2020

RÉSUMÉ

Ce mémoire est le fruit de mes dix dernières années de recherche à l'IMATH. Il concerne l'analyse mathématique, numérique des équations compressibles et incompressibles dans lequel j'aborde les équations hydrostatiques et les équations non-hydrostatiques.

Je présente et je développe des outils numériques de résolution adaptative sur des maillages non structurés en dimension quelconque pour des méthodes volumes finis et Galerkin discontinus.

Je propose également la réduction de modèle, *i. e.* une approximation asymptotique, dans le cadre des ondes longues pour dériver ou revisiter des modèles non-linéaires hydrostatiques et non-hydrostatiques. En particulier, je propose la première généralisation des équations de Serre-Green-Naghdi monodimensionnel en géométrie quelconque pour l'hydraulique fluviale. Celle-ci est obtenue par une réduction 3D vers 1D.

ABSTRACT

This dissertation is the fruit of my last ten years of research at IMATH. It concerns the mathematical, numerical analysis of compressible and incompressible equations in which I tackle hydrostatic and non-hydrostatic equations.

I present and develop numerical tools for adaptive mesh refinement on unstructured meshes in any dimension for finite volume and discontinuous Galerkin methods.

I also propose model reduction, *i. e.* asymptotic approximation, in the long wave framework to derive or revisit hydrostatic and non-hydrostatic nonlinear models. In particular, I propose the first generalisation of the one-dimensional Serre-Green-Naghdi equations for arbitrary geometry for river hydraulics. This is achieved by a 3D to 1D reduction.

Garde en mémoire que l'avenir te sera plus doux si tu emploies bien le temps présent.

— Les proverbes et pensées grecques (1812)

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philosophe. Un jour un de ses étudiants avait écrit pour lui *Chaque fois qu'un homme étudie, il élève le niveau général de l'humanité* — **Pascal**. Enfin, c'est une merveille. Merci Tonton Jacques comme dirait Kaan.

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Bien entendu, je pourrai écrire autant de page que ce manuscrit sur chacun de mes collègues ... Mais je vais m'arrêter là... Enfin, pour ceux de l'IMATH... Donc sans oublier **Yves**, **Khaled**, **Guy**, **Thierry** (Mon "Nesterov" à moi ;-)), Professeur **LSD**, **Valérie**, **Fabien**, **Catherine**, **Nicooooo**, **Christian**, **Antonin**, **Jean-Marc**, **Pierre**, **JP**, **Thomas**, **Kévin**, etc. Et surtout, surtout **MERCI** à **Christelle** qui fait un travail remarquable.

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Mehmet Ersoy, La Garde, le 3 Décembre, 2020

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ACRONYMS

| | |
|----------|----------------------------------------------|
| MFN | Mécanique des Fluides Numériques |
| CFD | Computational Fluid Dynamics |
| DG | Discontinuous Galerkin |
| PFS | Pressurised and Free Surface |
| AMR | Adaptive Mesh Refinement |
| AMM | Adaptive Moving Mesh |
| BB-AMR | Block-Based Adaptive Mesh Refinement |
| CPU | Central Processing Unit |
| CFL | Courant, Friedrichs and Lewy |
| OS | Osher-Sanders |
| TW | Tang-Warnecke |
| MUSCL | Monotone Upwind Scheme for Conservation Laws |
| RK2 | second order Runge-Kutta scheme |
| SWASH | Simulation Waves till SHores |
| FS-model | Free Surface model |
| SGN | Serre Green Naghdi |

Part I

INTRODUCTION

*L'alchimie est à mi-chemin
entre la poésie et les mathématiques,
entre le monde du symbole et celui du nombre.*

— **Marie-Madeleine Davy**

PRÉAMBULE

Le rêve de tout scientifique n'est-il pas de comprendre et de prévoir les réactions de la nature qui l'entoure? Depuis toujours, connus sous différents noms : savant, "cheikh" ou encore chercheur, ces hommes ont établi des lois reliant des grandeurs observables à des grandeurs non ou moins observables, en se basant simplement à leur intuition.

La seule vraie chose précieuse est l'intuition.

— **Albert Einstein**

L'idéologie de ces savants était basée sur les quatre éléments : la terre, l'eau, l'air et le feu. Cette idéologie, bien que simpliste, avait une signification profonde, qui dans les faits n'est pas tellement éloignée de nos connaissances à ce jour. Cette tétralogie se traduit par :

- la terre désignant l'état solide,
- l'eau désignant l'état liquide,
- l'air désignant l'état gazeux,
- le feu désignant l'énergie.

Entre terre et ciel, ces éléments sont omniprésents et en particulier ces "fluides" sont en nous, et, leurs rôles est essentiel puisqu'ils sont les supports de toutes vies.

Contrairement à l'eau, l'air est omniprésent et fort heureusement. On la trouve en abondance dans certaines régions alors que dans d'autres elle est une denrée rare. C'est ainsi que la recherche de ce liquide met l'hydraulique au premier rang des sciences les plus anciennes. Il est facile de s'en convaincre. De la plus lointaine civilisation à celle de nos jours, l'homme a laissé une trace de son action sur l'eau. Les canaux, les aqueducs des mésopotamiens aux romains en sont des exemples, mais pas seulement, l'homme antique fit entrer ce "fluide" dans le panthéon de ces divinités. Outre le caractère vital de l'eau, ce "fluide" a toujours été un élément de fascination parfois religieuse et parfois scientifique.

La notion de "fluide" est bien-sûr bien plus générale que celui de l'eau, bien que celui-ci reste dans tous les esprits comme étant l'exemple typique. L'air et plus généralement tous les gaz sont des exemples plus subtils. On peut étendre à d'autres corps dont les métaux en fusion, la lave des volcans, les roches, la matière stellaire constituant une nébuleuse, ou encore une foule d'individus s'écoulant dans une rue ou encore le trafic routier *etc.* permettant de classer les fluides en fluide newtonien, un fluide qui garde sa fluidité même si on la remue contrairement au fluide non-newtonien. Le fluide non-newtonien permet de marcher dessus et sa viscosité augmente lorsque la contrainte mécanique augmente.

Tout au long des siècles, presque exclusivement entre les mains des artisans, comme l'est le travail du fer au forgeron, la mécanique des fluides d'Archimède à Vinci, a été et est une contemplation esthétique et un sujet de recherche vaste et fascinant dans tous les domaines scientifiques. De l'invention rudimentaire à l'invention moderne, de la roue hydraulique à l'éolienne en passant par les moulins à vent, à eau, et les premières "machines volantes", de nombreuses créations ont permis d'exploiter le fluide non pas comme une ressource vitale, mais comme un élément utile à rendre notre quotidien meilleur.

Par cette intuition encore et toujours, l'artisan a créé et innové. Il a fallu attendre plusieurs siècles pour voir le mariage entre le savant et l'artisan. On peut citer à titre d'exemple, les travaux de Belidor, ingénieur militaire et professeur de mathématiques, qui en 1737 publia les quatre tomes de son "Architecture hydraulique" ou "L'art de conduire, d'élever, et de ménager les eaux, pour les différents besoins de la vie", Bernouilli, membre d'une célèbre famille de mathématiciens de Bâle, qui formule l'expression de la conservation de l'énergie où figurent les trois termes de position, de pression et de vitesse (Traité d'hydrodynamique de 1738), Toricelli, en 1644 avait précédé et limité son travail aux jets d'eau des villas italiennes, Euler en 1741 donne l'expression dite "des moments cinétiques" qui est à la base des calculs de toutes les turbomachines, ou encore Lagrange qui éclairera la seconde moitié de ce siècle.

Si Vinci fut un exemple, Euler en est un autre. On oublie généralement qu'il est parti de la théologie, pour le conduire ensuite aux mathématiques, à la topologie, à la mécanique des fluides *etc.* tout en restant le philosophe de son siècle. Cette polyvalence s'oppose à la spécialisation des chercheurs modernes. Mais cette universalité a une limite. Si Euler était un éminent philosophe et savant, il n'était pas

ingénieur pour autant. Il faudra attendre quelques années avant que les théories de génie voient leurs réalisations se concrétiser par Fourneyron, Girard, Kaplan *etc.* Les réflexions de ces savants ne sont plus seulement des formules aujourd'hui, elles sont chaque jour utilisées et concrétisées, c'est la magie de la théorie que je décrirais sous la forme suivante :

*La théorie est folle !
Elle n'est pas assez précise pour être juste,
Mais assez pour la mettre en pratique.*

— **Mehmet Ersoy**

Loin d'un fluide réel, les fluides théoriques dits "fluides parfaits" ignorent toutes les difficultés que présente l'autre. Le phénomène essentiel qui caractérise le fluide réel est la combinaison des effets cinétiques d'inertie et de la viscosité génératrice de forces tangentes aux lignes de courant, forces qui vont engendrer des gradients de vitesse perpendiculairement aux parois limitant l'écoulement. Les premières études de cette la viscosité sont dues à un médecin et physicien, Poiseuille qui portera le nom de "Loi de Poiseuille" (1840). Cette loi communément appelée loi de Hagen-Poiseuille a été découverte en parallèle par Hagen. Ce dernier était un scientifique de renom connu pour ses travaux dans de multiples domaines comme le génie civil et maritime, les sciences hydrauliques et la mécanique des fluides.

Deux notions très importantes des fluides réels se déduisent de l'influence de la viscosité. C'est ici que se situent les travaux de Reynolds et en particulier son célèbre nombre sans dimension qui sépare les écoulements dits laminaires où la viscosité joue un rôle stabilisant, des écoulements dits turbulents où les forces d'inertie l'emportent et créent une instabilité interne. L'influence prépondérante de la viscosité au voisinage des parois est à l'origine de la couche limite à l'étude de laquelle s'illustrèrent entre autres Prandtl et Von Karman. Toutes les questions de sillage et décollement se déduisent de ces considérations. La mécanique des fluides "modernes" initiée par Euler et d'Alembert, le stade des fluides parfaits, fait des progrès avec Henri Navier et George Gabriel Stokes pour s'attaquer aux fluides incompressibles puis compressibles.

De la représentation concrète de l'intuition, aux pensées des grands philosophes de l'époque, tout mène indéniablement à la description par des équations, à savoir, à la modélisation mathématique, les mathématiques appliquées et le calcul scientifique dont Borel, Lagrange,

Lions, *etc.* furent les précurseurs indéniables de la modélisation mathématique moderne.

Borel a consacré une cinquantaine d'articles à la philosophie et à l'histoire des sciences, à la psychologie, à la pédagogie et à l'économie politique. Fervent défenseur de la conception intuitionniste des mathématiques, il a toujours insisté sur la nécessité de ne jamais perdre de vue le "réel" : pour lui,

Les mathématiques ne sont pas un jeu purement abstrait de l'esprit, mais sont, Au contraire, en étroite connexion avec la réalité concrète.

— **Émile Borel**

Il tenait à rappeler, dans les raisonnements les plus abstraits de l'analyse, comment ces raisonnements expriment des réalités, et par suite nous aident à comprendre la réalité physique. C'est pourquoi Borel a toujours préféré aux définitions descriptives les définitions constructives ; ces dernières sont indispensables pour passer de la théorie aux applications et permettent auparavant de s'assurer de l'existence et éventuellement de l'unicité de l'élément dont on parle. Jacques-Louis Lions qui était sans aucun doute le mathématicien d'exception, précurseur de biens de choses en analyse mathématique appliquée moderne et notamment en analyse numérique. Sa philosophie laisse rêveur, et comme le citait Antoine de Saint-Exupéry, "*Les rêves fournissent la nourriture de l'âme. Celui qui ne cherche pas à réaliser son rêve suicide son âme.*" :

*Ce que j'aime dans les mathématiques appliquées,
C'est qu'elles ont pour ambition de donner du monde des systèmes
une représentation qui permette de comprendre et d'agir.
Et, de toutes les représentations, la représentation mathématique,
Lorsqu'elle est possible, est celle qui est la plus souple et la meilleure.
Du coup, ce qui m'intéresse, c'est de savoir jusqu'où
On peut aller dans ce domaine de la modélisation des systèmes, c'est
d'atteindre les limites.*

— **Jacques-Louis Lions**

On ne peut pas parler de mathématiques appliquées sans parler d'algorithme et sans faire allusion au père de l'algèbre. Fondateur du premier observatoire permanent au monde, pionniers de l'algèbre avec son traité, *Kitab al jabr w'al muqabalah*, fondateur de la maison de la sagesse, Al-Khwarizmi traduit des textes scientifiques et philosophiques du monde entier. Al-Khwarizmi était philosophe et un

excellent passeur de connaissance. Il introduit dans son aire culturelle les connaissances mathématiques indiennes, notamment le système décimal de numération. La traduction latine de son ouvrage *Algorithmi de numero indorum* permit la transmission de ces connaissances jusque dans l'Occident du XIIe siècle. D'ailleurs, le mot *Algorithmi*, traduction latine d'Al-Khwarizmi, a donné naissance au mot "algorithme" et fait de lui l'irréfutable "père fondateur" de nos ordinateurs. À la fin du Xème siècle et au début du XIème siècle, le mathématicien persan Al-Karaji franchit une nouvelle étape dans l'histoire de l'algèbre : il la détache de la géométrie — ce que ni Al-Khwârisimî ni Abu Kamil n'avaient pu faire — dans ses ouvrages intitulés *Gloire de l'algèbre et de la muqabala et Merveilles du calcul*. Ensuite, la grande étape franchie par l'algèbre musulmane est la résolution de l'équation cubique dont le célèbre poète et mathématicien persan Omar Khayyâm (1048–1131), bercé et inspiré par la poésie

*Ceux qui ont embrassé science et littérature
ont récité leur fable et se sont endormis.*

— Omar Khayyâm

Les bases de l'analyse numérique et du calcul numérique "moderne" furent, sans aucun doute, initiées par Newton (1669) sur la recherche de zéro de fonction. Dix ans plus tard, c'est à Raphson, qu'on doit l'algorithme de Newton-Raphson qui met en évidence une formule de récurrence. Un siècle plus tard, Mouraille et Lagrange étudient la convergence des approximations successives en fonction des conditions initiales par une approche géométrique. Cinquante ans plus tard, Fourier et Cauchy s'occupe de la rapidité de la convergence *etc.*

Du traité d'Al-Khwarizmi à la naissance de la première machine de calcul, le calcul scientifique puis le calcul haute performance a vu le jour. Frontière entre les mathématiques appliquées et la mécanique, la Mécanique des Fluides Numériques ([MFN](#)) ou par le terme anglo-saxon Computational Fluid Dynamics ([CFD](#)), a grandi d'une curiosité mathématique pour devenir aujourd'hui un outil essentiel voire incontournable dans pratiquement toutes les branches scientifiques dont la dynamique des fluides et notamment en hydrodynamique côtière et celle des vagues.

De l'université de Savoie à l'Université de Toulon, sous la direction de Christian Bourdarias et Stéphane Gerbi, puis sous celle d'Enrique Zuazua lors de mon post-doc au Basque Center for Applied Mathematics, ainsi, j'ai l'immense plaisir de vous présenter sans aucune

prétention mes travaux de recherche, fruit de nombreuses années de passion.

Ce mémoire est dédié à *l'étude, l'approximation et la simulation des équations aux dérivées partielles issues de la mécanique des fluides compressibles ou incompressibles* – hydrostatique et non hydrostatique dans les domaines de l'hydrauliques et de l'hydrodynamique.

En amont, mes travaux comportent une part importante de *modélisation* — approximation de systèmes hyperboliques de systèmes de type "hydrostatique", modèles primitives compressibles, modèles de type Saint-Venant non homogène, modèles multi-fluide, modèles à faible Mach — dont les applications sont par exemples, les écoulements à surface libre et/ou en charge, les problèmes d'entraînement d'air et plus généralement les écoulements multi-fluides, la sédimentation, la dynamique de l'atmosphère, les problèmes de déferlements, les problèmes d'inondations, les écoulements non-hydrostatique . . .

En aval, mes travaux abordent un certain nombre de *problèmes théoriques* — existence et unicité des solutions faibles, stabilité des solutions faibles, contrôle optimal, problèmes de Riemann (systèmes autonomes et non autonomes), analyse asymptotique en temps long (propriétés des solutions stationnaires vs solutions non stationnaires) — dont le but principal est d'aboutir à certaines caractérisations nécessaires lors de l'approximation numérique de ces équations aux dérivées partielles. Mes travaux explorent donc *l'analyse numérique*, la construction de méthodes numériques Volumes Finis en multi-dimension et le calcul scientifique — solveur de Godunov, solveur cinétique, schémas équilibres, schémas anti-diffusif, techniques de raffinement de maillage, critère de raffinement — pour la simulation numérique des problèmes de modélisation mentionnés précédemment.

Mots clefs :

Fluide compressible, Fluide incompressible, Couplage compressible-incompressible, Modèle hydrostatique, Modèle non-hydrostatique, Modèle dispersif, Modèle à faible Mach, Existence et stabilité de solutions faibles, Analyse asymptotique, Contrôle, Problème de Riemann, Volumes Finis, Schéma cinétique, Schéma équilibre, Raffinement de maillage adaptatif h et hp .

TRAVAUX DE THÈSE

Le mémoire de thèse "MODÉLISATION, ANALYSE MATHÉMATIQUE ET NUMÉRIQUE DE DIVERS ÉCOULEMENTS COMPRESSIBLES OU INCOMPRESSIBLES EN COUCHE MINCE" est composé de 253 pages réparties en une préface, de trois parties totalisant cinq chapitres, d'une annexe et d'un index. Il s'articule autour du plan suivant :

PARTIE I Modélisation des équations en conduite fermée

CHAPITRE 1 Les équations Pressurised and Free Surface (PFS)

CHAPITRE 2 Approximation des équations PFS par Volumes Finitis

PARTIE II Autour des équations Primitives Compressibles

CHAPITRE 3 Existence et stabilités de solutions faibles

PARTIE III Travaux en cours et perspectives

CHAPITRE 4 Vers la cavitation

CHAPITRE 5 Modélisation de la sédimentation

Cette thèse est consacrée à la mise en oeuvre de nouveaux modèles et méthodes numériques pour des écoulements compressibles ou incompressibles en domaine "couche mince".

S'il est possible de calculer la dynamique d'un écoulement (par exemple en conduite fermée) à partir de la résolution des équations de Navier-Stokes, il est nécessaire d'effectuer la réduction de modèle afin de s'intéresser à la modélisation hydrodynamique aux échelles qui nous intéressent. Les modèles obtenus sont souvent plus exploitables d'un point de vue numérique en termes de gain de temps et introduit des nouveaux problèmes. En l'état actuel, les modèles concernés ne sont pas entièrement satisfaisants tant d'un point de vue théorique que numérique notamment en domaine "couches minces".

Ce travail apporte un nouveau regard à des problèmes existants par des dérivations de modèles originales, des preuves mathématiques astucieuses ou des schémas numériques alliant précision et simplicité.

2.1 ÉCOULEMENTS D'EAU MIXTES EN CONDUITES FERMÉES.

MOTIVATIONS.

Le *couplage* entre un *écoulement en charge* et un *écoulement à surface libre* (écoulements mixtes transitoires en charge/à surface libre) est, comme le montre l'expérience, un problème délicat compte tenu de l'instabilité du phénomène. Celui-ci est en outre peu accessible aux mesures directes. Le suivi en temps réel, à l'aide d'une simulation numérique, est d'un intérêt majeur pour les ingénieurs en tant qu'outil de validation et de prévision. Il existe en effet des risques très importants liés à des surpressions, compromettant l'intégrité de la structure. Ce type de problème se rencontre, par exemple, dans la modélisation des écoulements dans des réseaux d'assainissement lors d'orages violents, dans des canaux d'amenée en amont des usines hydroélectriques, etc.

L'ÉTAT DE L'ART.

Nous avons proposé un nouveau modèle pour les écoulements mixtes à *géométrie variable*, nommé **PFS** (voir [19] pour une dérivation complète ou [54, Chapitre 1]). Ce système est donné par les équations suivantes :

$$\begin{aligned} \partial_t A + \partial_x Q &= 0, \\ \partial_t Q + \partial_x \left(\frac{Q^2}{A} + p(x, A, E) \right) &= -gA \frac{dZ}{dx} \\ &\quad + Pr(x, A, E) - G(x, A, E) \\ &\quad - gK(x, A, E) \frac{Q|Q|}{A} \end{aligned}$$

où t est la variable de temps, x l'abscisse curviligne (suivant l'axe d'écoulement), $\frac{dZ(x)}{dx}$ le terme de pente, $Pr(x, A, E)$ le terme source de pression, $G(x, A, E)$ le terme de courbure et $K(x, A, E)$ le terme de frottement donné par la formule de Strickler (voir [19] ou [54, Chapitre 1] pour le détail des termes). L'inconnue $A(t, x)$, homogène à une aire, est la variable "mixte" car elle permet de calculer tantôt le niveau d'eau à surface libre, tantôt la pression en charge. L'inconnue $Q(t, x)$ correspond au débit. La donnée de l'aire mixte A n'est en effet suffisante ni pour déterminer la loi de pression ni les termes sources. Il faut suivre l'historique du point courant en termes d'état E (surface libre si $E = 0$ ou charge si $E = 1$). Le terme de pression $p(x, A, E)$, conforme à la physique du problème, est égal à la pression à surface libre lorsque la variable d'état E est égale à 0 (*i. e.* $p(x, A, 0) =$

$gI_1(x, A) \cos \theta$ la pression hydrostatique, c.f. Figure 2.1), ou bien de type mixte hydrostatique-acoustique en cas de charge si $E = 1$ (i.e. $p(x, A, 1) = gI_1(x, S) \cos \theta + c^2(A - S)$, c.f. Figure 2.1).

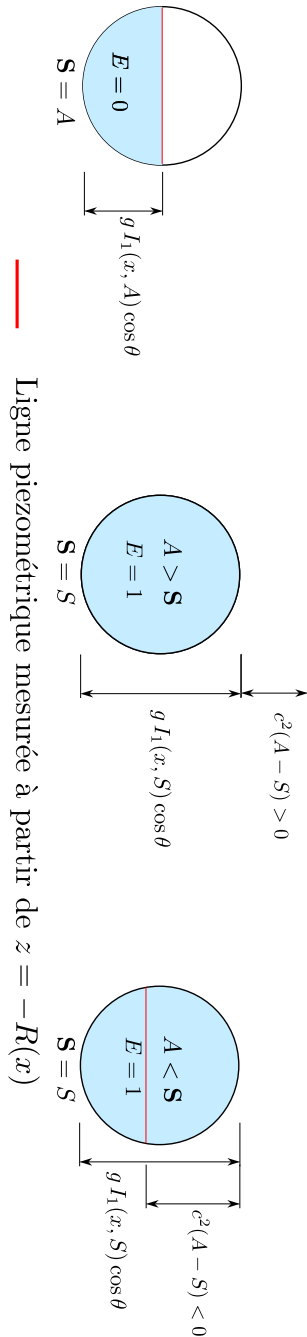


Figure 2.1: état à surface libre $p(x, A, 0)$ (gauche), état en charge avec surpression $p(x, A, 1) > 0$ (milieu), état en charge avec dépression $p(x, A, 1) < 0$ (droite) où S section d'eau physique, S section de la conduite, R rayon de la conduite.

ORIGINALITÉ : LE MODÈLE.

Ce nouveau modèle permet de prendre en compte la *compressibilité de l'eau lors d'un écoulement en charge* et les *points de transition* (points de passage d'un écoulement en charge vers un écoulement à surface libre et vice et versa) en géométrie quelconque. Contrairement aux méthodes basées sur l'artifice de la fente de Preissmann, il permet de simuler correctement les écoulements sub-atmosphériques (considérés jusqu'alors comme un passage d'un écoulement en charge vers un écoulement à surface libre).

Ce modèle est construit à partir des *équations d'Euler compressibles tridimensionnelles* pour la partie en charge avec une *loi de pression linéarisée* et les *équations d'Euler incompressibles tridimensionnelles* pour la partie à surface libre. Les effets de courbure locale sont pris en compte en écrivant les équations dans un *repère local mobile* (suivant l'axe d'écoulement moyen). Les deux modèles obtenus sont de type *Saint-Venant* et sont formellement proche dans l'écriture. Ils sont couplés pour obtenir un système d'équations *hyperboliques* : le modèle [PFS](#).

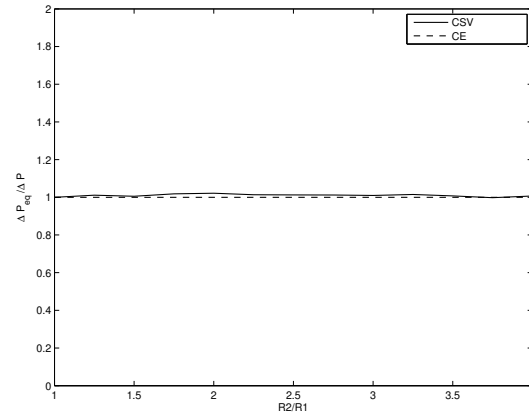
ORIGINALITÉ : ASPECTS NUMÉRIQUES.

L'approximation numérique naturelle utilisée est la méthode des *Volumes finis* de par la nature hyperbolique du système [PFS](#). La présence des termes sources "complexes" rendent la construction d'un schéma Volumes Finis consistant, stable et entropique difficile. C'est d'ailleurs la difficulté majeure de la discrétisation de ces équations. En combinant les techniques de *décentrement classique* pour les termes conservatifs, en utilisant la *théorie des produits non conservatifs* et en introduisant la notion de *pente dynamique* (la friction est intégrée au terme Z), nous avons obtenu des résultats très satisfaisants avec des solveurs *VFRoe* et *cinétique* (voir par exemple [17, 18, 21] ou [54, Chapitre 2] et [Figure 2.2a](#), par exemple).

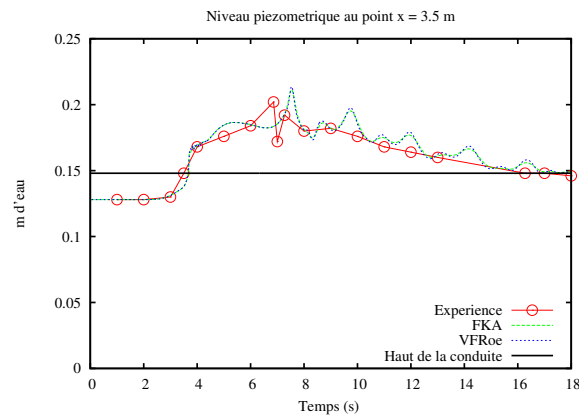
ORIGINALITÉ : TRAITEMENT DES POINTS DE TRANSITION ET CONDITIONS AUX LIMITES.

Les points de transition entre les deux types d'écoulement sont traités via la technique des *ondes fantômes*, i.e l'interface est considérée comme une frontière libre, correspondant à une discontinuité du gradient de pression. Les états de part et d'autre d'une telle interface sont obtenus via la résolution d'un problème de Riemann linéaire à matrice discontinue, la ligne de discontinuité coïncidant avec la trajectoire du point de transition. La généralisation de ces travaux en présence

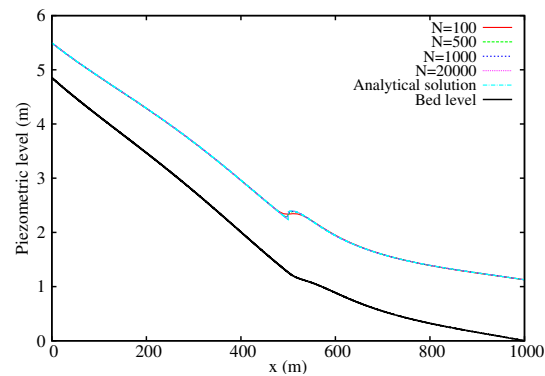
de termes sources complexes est présentée dans [17] (voir aussi [54, Chapitre 2.2] et Figure 2.2b).



(a) Comparaison de notre modèle (CSV, schéma cinétique) avec le modèle uniforme à section équivalente (CE).



(b) Cas test de Wiggert.



(c) Cas test de Baines (schéma cinétique).

Figure 2.2: Cas tests numériques.

Nous avons proposé une nouvelle méthode pour le traitement des points de transition dans [20], *alternative à la méthode des ondes fantômes* jusqu'alors utilisée (c.f. [17]), en adéquation avec le schéma cinétique (voir Figure 2.2c). Enfin, ces études théoriques et numériques ont contribué aux développements des codes EDF-CIH Roemix (solveur VFRoe) et Flowmix (solveur cinétique).

2.2 LA DYNAMIQUE DE L'ATMOSPHÈRE.

MOTIVATIONS.

Les équations primitives sont en pratique les équations les plus utilisées pour l'étude de la *circulation atmosphérique et océanique* à grande échelle et en particulier par les modèles de prévision météorologique. Elles sont issues de la dynamique des fluides géophysiques en domaine "couche mince". Les *équations primitives compressibles* dérivent de l'*approximation hydrostatique* des équations de Navier-Stokes compressibles.

L'ÉTAT DE L'ART.

Les premiers résultats mathématiques concernant les équations primitives à *viscosité constante* (celles utilisées en météorologie par exemple) ont été établis par Lions *et al.* [102] (formulation mathématique des équations et existence de solutions faibles globales en temps).

Dans la littérature, le modèle que nous appellerons *équations primitives compressibles simplifiées* (car les termes de diffusion de chaleur émanant du soleil et la quantité d'eau dans l'air sont négligés), est introduit la première fois par Kochin [89] en 1936. Curieusement, les premiers résultats mathématiques sont de Gatapov et Kazhikhov [69] en 2005 et qui plus est, pour un modèle similaire à celui de Kochin [89] que nous appellerons *modèle intermédiaire*.

ORIGINALITÉ.

Dans un premier temps, nous nous sommes intéressés à la dérivation des équations primitives compressibles introduites par Kochin [89]. Pour ce faire, nous avons proposé une approximation hydrostatique des équations bidimensionnelles et tridimensionnelles de Navier-Stokes compressibles en utilisant un *tenseur de viscosité anisotrope non constant* (voir [63] et [54, Chapitre 3.2]). Cette analyse asymptotique

permet de justifier les modèles en dimensions deux et trois. Le modèle en dimension deux s'écrit :

$$\begin{cases} \partial_t \rho + \partial_x(\rho u) + \partial_y(\rho v) & = 0, \\ \partial_t(\rho u) + \partial_x(\rho u^2) + \partial_y(\rho uv) + \partial_x p & = \partial_x(v_1 \partial_x u) + \partial_y(v_2 \partial_y u), \\ \partial_y p & = -\rho g \end{cases}$$

où x, y dénote la variable spatiale horizontale et verticale, ρ est la densité, (u, v) est le vecteur vitesse avec u (resp. v) la composante horizontale (resp. verticale), p est une loi de pression donnée par la loi d'état $p(\rho) = c^2 \rho$. La constante c est définie par $c^2 = \mathcal{R}\mathcal{T}$ où \mathcal{R} constante universelle des gaz parfaits de l'air et \mathcal{T} la température, supposée constante. Le profil de viscosité est choisi sous la forme stratifiée suivante $v_1(t, x, y) = v_0 e^{-g/c^2 y}$ où $v_0 \in \mathbb{R}$ et v_2 est une fonction donnée.

Les résultats dans [62, 63] sont obtenus en exploitant les informations données par l'équation hydrostatique $\partial_y p(\rho) = -\rho g$. En effet, la solution de cette équation s'écrit sous la forme (stratifiée) $\rho = \xi(t, x) e^{-g/c^2 y}$ où on montre que ξ est solution du modèle intermédiaire via le changement de variable $z = 1 - e^{-g/c^2 y}$. Le modèle intermédiaire est beaucoup simple à étudier et, en dimension deux, il s'écrit :

$$\begin{cases} \partial_t \xi + \partial_x(\xi u) + \partial_z(\xi w) & = 0 \\ \partial_t(\xi u) + \partial_x(\xi u^2) + \partial_z(\xi uw) + \partial_x \xi & = \partial_x(\partial_x u) + \partial_z(\partial_z u) \\ \partial_z \xi & = 0 \end{cases}$$

où $w(t, x, z) = e^{-y} v(t, x, y)$ est la vitesse verticale. Ce procédé astucieux permet d'établir une équivalence entre les équations primitives compressibles simplifiées et le modèle intermédiaire (voir [62, 63, 69]). Ainsi, nous avons établi l'existence de solutions faible globale en temps non élucidés depuis 1936 (voir [62] ou [54, Chapitre 3.3]) pour le modèle en dimension deux. Les solutions sont obtenues au sens des distributions et vérifient les estimations $\rho \in L^\infty(0, T; W^{1,2}(\Omega))$, $\partial_t \rho \in L^2(0, T; L^2(\Omega))$ et $u \in L^2(0, T; W^{2,2}(\Omega)) \cap W^{1,2}(0, T; L^2(\Omega))$, $v \in L^2(0, T; L^2(\Omega))$. Pour le modèle en dimension trois, nous avons obtenu un résultat de stabilité de solutions faibles (voir [63] et [54, Chapitre 3.4]).

2.3 LA SÉDIMENTATION.

MOTIVATIONS.

Sous l'action de la *sédimentation*, il existe une interaction dynamique

entre l'écoulement et la morphologie du fond, par exemple, le lit d'une rivière. Le transport de sédiments modifie la dynamique de l'écoulement et donc le transport de sédiments lui-même. Les deux écoulements sont donc interdépendants. La composante hydrodynamique de l'écoulement est régie par les équations classiques de *Saint-Venant* où la topographie évolue en fonction d'une équation de transport, l'équation d'*Exner*. Les équations de *Saint-Venant-Exner* sont données par les équations suivantes :

$$\begin{cases} \partial_t h + \operatorname{div}(q) = 0, \\ \partial_t q + \operatorname{div}\left(\frac{q \otimes q}{h}\right) + \nabla\left(g \frac{h^2}{2}\right) = -gh \nabla b \\ \partial_t b + \zeta \operatorname{div}(q_b(h, q)) = 0 \end{cases}$$

où h représente la hauteur d'eau, $q = hu$ est le débit d'eau, q_b est le débit de sédiments (ou le flux de transport solide) et $\zeta = 1/(1 - \psi)$ est un coefficient de porosité associé à la couche de sédiments. Il existe principalement trois modes de transport : par *suspension*, par *charriage* et/ou par *saltation*. Chaque mode est régi par des lois empiriques valable pour une plage de donnée. Par conséquent, il est difficile d'établir une loi générale décrivant les trois modes de transport. La réduction d'un modèle avec une *loi générale* (issue d'une description microscopique, par exemple) est un des enjeux majeurs pour la modélisation de la sédimentation.

L'ÉTAT DE L'ART ET ORIGINALITÉ.

Nous avons proposé une nouvelle manière d'aborder ce problème en décrivant la sédimentation par une description microscopique couplée à des équations de la mécanique des fluides.

Plus précisément, la dynamique des sédiments est décrite par l'équation de *Vlasov* :

$$\partial_t f + \operatorname{div}_x(vf) + \operatorname{div}_v((F + \mathbf{g})f) = r \Delta_v f$$

où f est la densité de particules, $v \in \mathbb{R}^3$ la vitesse cinétique d'une particule soumise à la force de gravité $\mathbf{g} = (0, 0, -g)^t$ et F la force de friction de Stokes. Le terme de diffusion $r \Delta_v f$ décrit le mouvement Brownien des particules où $r > 0$ est la vitesse de diffusion donnée par la formule d'Einstein. Cette équation est couplée au système de *Navier-Stokes compressible* :

$$\begin{aligned} \partial_t \rho_w + \operatorname{div}(\rho_w u) &= 0, \\ \partial_t(\rho_w u) + \operatorname{div}(\rho_w u \otimes u) + \nabla p(t, x) &= \operatorname{div} \sigma(\rho_w, u) + \mathfrak{F}, \end{aligned}$$

où $u(t, x) = (u_1, u_2, u_3)(t, x)$ est le champ de vitesses et $\rho_w(t, x)$ la densité. Le terme $p(t, x)$ est une *pression barotrope non homogène* tenant en compte l'effet des sédiments dans le fluide. Le terme $\sigma(\rho_w, u)$ est le tenseur de contrainte $-pI_3 + 2\Sigma(\rho_w) : D(u) + \lambda(\rho_w)\text{div}(u) I_3$ où I_3 correspond à la matrice identité et $\Sigma(\rho_w)$ est un tenseur de viscosité anisotrope.

L'interaction fluide/sédiments est décrite au niveau du terme de pression et du terme source. Le terme de pression s'écrit $p(t, x) = k(t, x_1, x_2)\rho(t, x)^2$ où $k(t, x_1, x_2) = \frac{gh(t, x_1, x_2)}{4\rho_f}$, $\rho := \rho_w + \rho_s$ est une "densité de mélange" avec ρ_s la densité macroscopique de sédiments, $\rho_s = \int_{\mathbb{R}^3} f dv$ et ρ_f une densité caractéristique du fluide. Le second couplage est décrit par le terme $-\int_{\mathbb{R}^3} Ff dv$ présent dans $\mathfrak{F} - \int_{\mathbb{R}^3} Ff dv + \rho_w \mathfrak{g}$.

Outre l'originalité d'introduire une loi de pression barotrope non-homogène (voir aussi [24, 25]), la réduction de ces équations vers un modèle de Saint-Venant-Exner a été obtenue par deux analyses asymptotiques successives. La première analyse asymptotique est la *limite hydrodynamique* de l'équation de Vlasov. En intégrant cette équation contre 1 et v , nous obtenons un système macroscopique que nous couplons avec les équations de Navier-Stokes compressibles pour aboutir à un modèle de type mélange (similaire à un modèle à deux phases). Nous procédons ensuite à la réduction de ce modèle par une analyse asymptotique en *couche mince* et nous obtenons les équations de Saint-Venant-Exner. Les détails de ces travaux sont disponibles dans [54, Chapitre 5].

TRAVAUX DE RECHERCHE : PÉRIODE POSTDOCTORALE

Vous trouverez ci-dessous le résumé de mes travaux de recherche : contrôle, analyse numérique, mécanique numérique et applications, et modèles asymptotiques hydrostatique et non-hydrostatique.

3.1 ANALYSE DE SENSIBILITÉ

MOTIVATIONS.

Les applications du contrôle optimal sont nombreuses et variées, par exemple, pour le contrôle des flux routiers, ferroviaires, aériens, fluviaux, barrages EDF, *etc.* Dans certains de ces exemples, les équations sont des lois de conservation hyperbolique (non-linéaires) pour lesquelles l'utilisation des techniques standard basées sur la linéarisation est délicate, en raison de la présence de discontinuités. Le problème qui a été abordé ici, à notre connaissance n'a pas encore été traité. Il s'agit d'étudier l'*analyse de sensibilité* pour des lois de conservation scalaire (stationnaire) avec terme source. Cette analyse est pertinente dans les applications aux problèmes de contrôle optimal et d'identification des paramètres (voir [Part ii](#)).

L'ÉTAT DE L'ART ET ORIGINALITÉ.

Dans ce travail [58], on considère un modèle simplifié de loi de conservation scalaire

$$\partial_x f(v(x)) + v(x) = g(x), \quad x \in \mathbb{R}, \quad (3.1)$$

avec des "conditions aux limites" $\lim_{x \rightarrow \pm\infty} v(x) = 0$. L'objectif principal est d'effectuer une analyse de sensibilité rigoureuse des solutions v du problème (3.1) par rapport aux perturbations du terme source (forçage) g . En notant $v_\varepsilon = v + \varepsilon(\delta v)_\varepsilon$ une solution du problème perturbé (3.1) avec $g = g + \varepsilon\delta g$ nous identifions la limite de la quantité

$$\frac{v_\varepsilon - v}{\varepsilon} = (\delta v)_\varepsilon \rightarrow \delta v \text{ faiblement-}^* \text{ dans } \mathcal{M}(\mathbb{R}) \text{ as } \varepsilon \rightarrow 0,$$

où (δv) est l'unique *solution de dualité* (c.f. [16]) du problème linéaire

$$\partial_x (f'(v)\delta v) + \delta v = \delta g.$$

Dans le cadre fonctionnel de Bressan and Marson [27], on peut aussi établir ce résultat à partir du problème non-stationnaire $\partial_t u(t, x) + \partial_x f(u(t, x)) + u(t, x) = g(x)$, $u(0, x) = u^0$ (voir [58] et Figure 3.1).

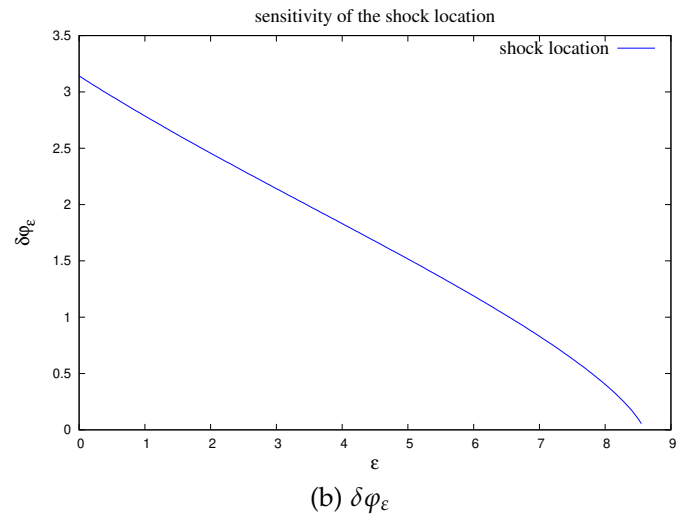
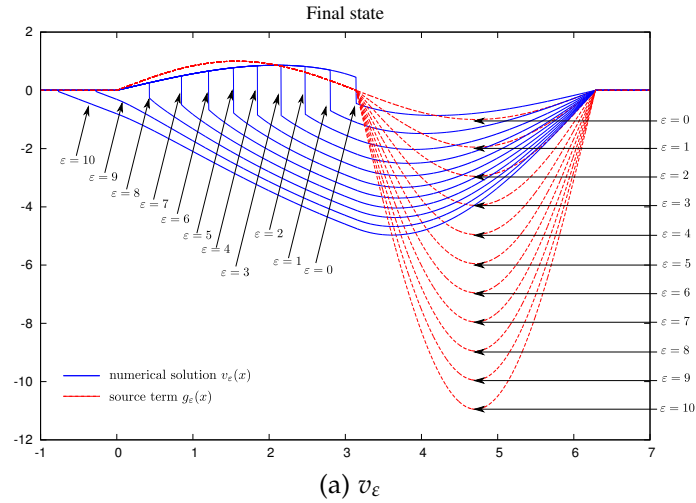


Figure 3.1: Illustration numérique de l'analyse de sensibilité de $(v_\epsilon, \delta\varphi_\epsilon)$ par rapport à g pour le problème avec $f(u) = \frac{u^2}{2}$.

3.2 MÉCANIQUE DES FLUIDES NUMÉRIQUES ET APPLICATIONS

MOTIVATIONS.

La mécanique des fluides numérique (MFN) ou par le terme anglais computational fluid dynamics (CFD), a grandi d'une curiosité mathématique pour devenir un outil essentiel voire incontournable dans

pratiquement toutes les branches de la dynamique des fluides et notamment en hydrodynamique côtière dont celle des vagues. Dans ce contexte, nous nous intéressons aux défis de la modélisation numérique (voir [Part iii](#)) qui consiste à simuler avec précision ces processus sur de très grandes échelles spatiales à l'aide de raffinement adaptatif de maillage (Adaptive Mesh Refinement ([AMR](#))) basé sur des blocs (Block-Based-AMR ou [BB-AMR](#)).

L'ÉTAT DE L'ART ET ORIGINALITÉ.

De manière générale, nous nous intéressons à l'approximation numérique par Volumes Finis de systèmes hyperboliques non linéaires multidimensionnels, écrit ici sous la forme 1d,

$$\begin{cases} \frac{\partial w}{\partial t} + \frac{\partial f(w)}{\partial x} = 0, & (t, x) \in \mathbb{R}^+ \times \mathbb{R} \\ w(0, x) = w_0(x), & x \in \mathbb{R}. \end{cases}$$

où $w : \mathbb{R}^+ \times \mathbb{R} \rightarrow \mathbb{R}^d$ représente le vecteur d'état inconnu et $f : \mathbb{R}^d \rightarrow \mathbb{R}^d$ désigne le flux. Nous avons développé un code numérique multidimensionnel pour l'approximation numérique de ces équations par Volumes Finis qui intègre un algorithme de pas de temps local, un schéma d'ordre 1 et 2 en espace-temps (MUSCL, AB2, RK2) basé sur un solveur cinétique et le solveur de Godunov dans un cadre de raffinement de maillage adaptatif [[1](#), [59](#), [60](#), [73](#), [74](#), [121](#), [122](#), [166](#), [167](#)] type décomposition de domaine ([BB-AMR](#)).

On rappelle succinctement le principe des méthodes [AMR](#) dans le cadre de la dimension 1. En notant $S_{k_b}^n$ un critère de raffinement positif sur une cellule C_{k_b} à l'instant t_n du maillage, on compare ce dernier à un seuil α , par exemple la moyenne,

$$\alpha = S_m = \frac{1}{|\Omega|} \sum_{k_b} S_{k_b}^n.$$

On définit ensuite deux coefficients $0 < \beta_{\min} \leq \beta_{\max}$, qui déterminent le pourcentage des cellules à raffiner ou à déraffiner. Ainsi, pour chaque cellule C_{k_b} :

- si $S_{k_b}^n > \alpha_{\max} = \beta_{\max}\alpha$, alors la cellule est raffinée et divisée en deux cellules $C_{k_{b0}}$ et $C_{k_{b1}}$,
- si $S_{k_{b0}}^n < \alpha_{\min} = \beta_{\min}\alpha$ et $S_{k_{b1}}^n < \alpha_{\min}$, alors les deux cellules sont fusionnées pour former une cellule C_{k_b} .

Dans le cadre multidimensionnel, nous avons développé une méthode de type décomposition de domaine *Block-Based AMR* ([BB-AMR](#)) qui

permet un contrôle efficace de la mémoire partagée, ce qui permet d'obtenir un *temps de calcul* bien mieux équilibré entre les cœurs de l'ordinateur (voir [Chapter 7](#)).

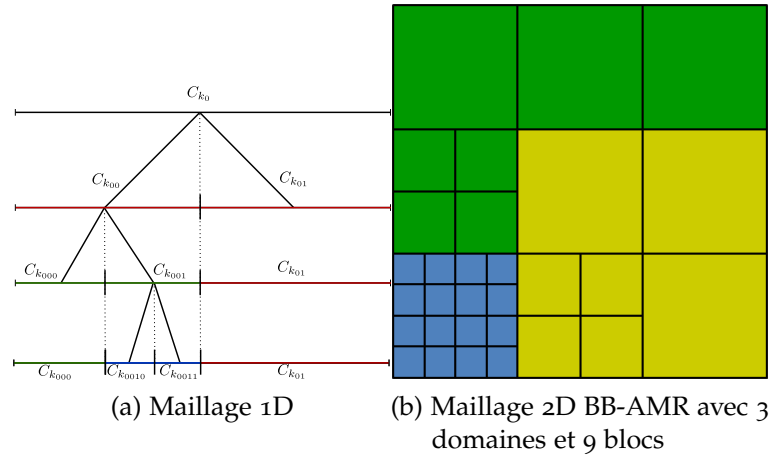


Figure 3.2: Exemple 1D et 2D.

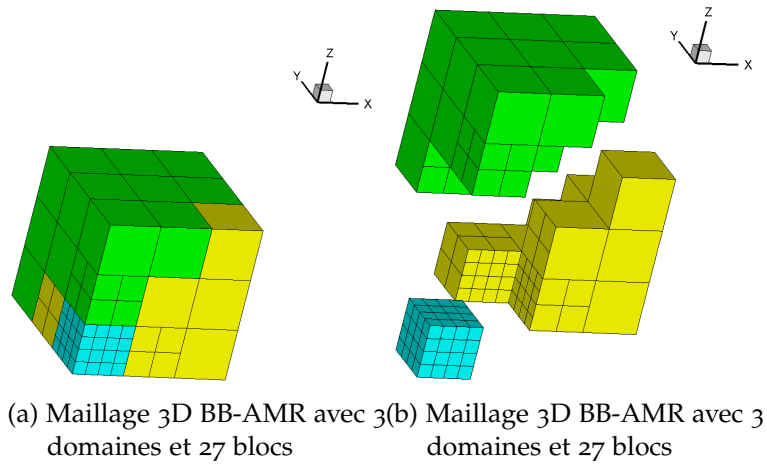
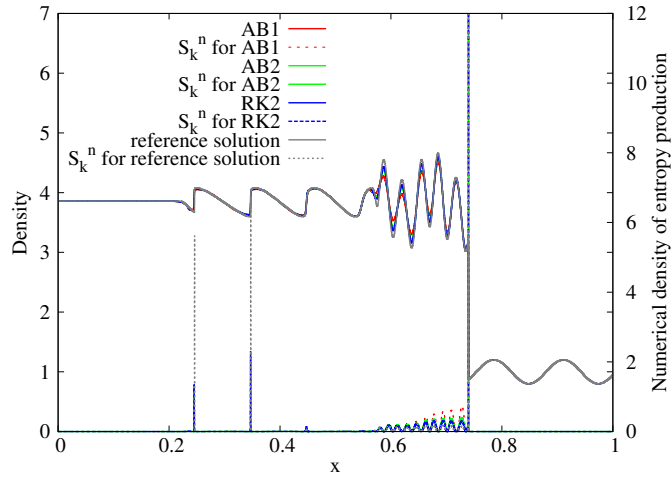
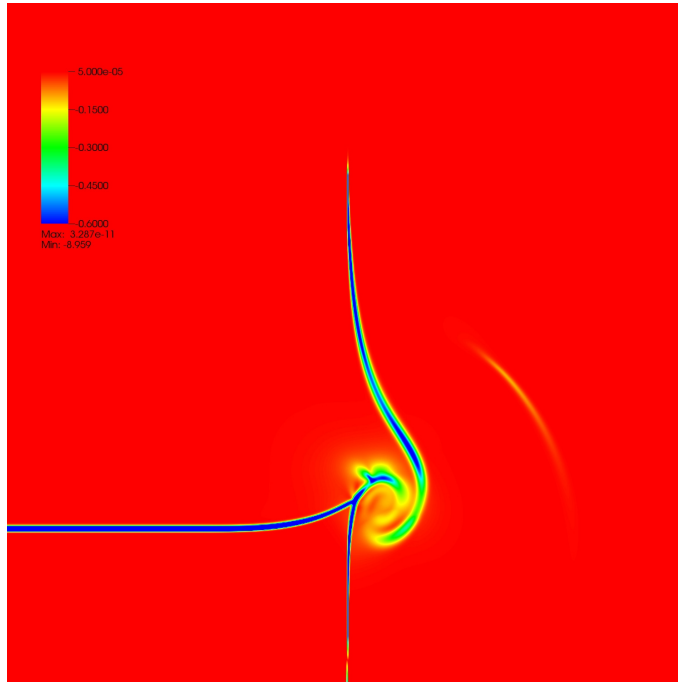


Figure 3.3: Exemple 3D.

Enfin, afin d'améliorer la méthode [AMR](#), nous avons proposé une méthode permettant de calculer le *seuil* $\alpha = \alpha_{PE}$ *automatiquement* de manière à "optimiser" les zones à raffiner (jusqu'alors le seuil α était calibré manuellement). Ce dernier est construit à partir du réarrangement décroissant du critère de raffinement (voir [Chapter 8](#)).

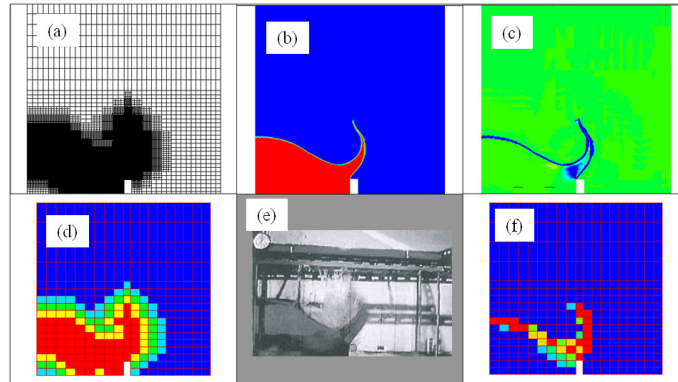


(a) Cas test Shu-Osher (Eqs d'Euler 1d (gaz parfait))

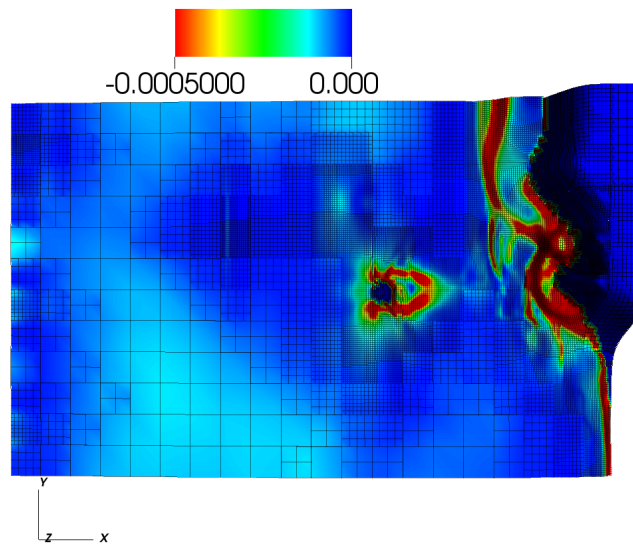


(b) Problème de Riemann (Eqs de Saint-Venant)

Figure 3.4: Exemples de résultats de simulation (1).



(a) Rupture de barrage 2D (Eqs Euler bi-fluide) : (a) Mail-
lage; (b) Densité (air-bleu, eau-rouge); (c) Densité de
production numérique d'entropie (valeurs vert-zero,
bleu-négatif); (d) Niveau de raffinement par bloc (1 à 5);
(e) Expérience Koshizuka (95); (f) Critère de raffinement
par bloc.



(b) Tsunami (vallée de la Monai) : temps $t = 16s$ (Eqs Saint-
Venant). Les couleurs correspondent aux valeurs de la
densité de production numérique.

Figure 3.5: Exemples de résultats de simulation (2).

3.3 DYNAMIQUE DES ÉCOULEMENTS DANS LES PLAGES SABLEUSES

MOTIVATIONS.

La compréhension de la dynamique des écoulements souterrains dans les zones côtières est d'un intérêt majeur dans les domaines de l'ingénierie côtière et du développement durable, notamment dans la région sud-est de la France. Une meilleure compréhension de la circulation de ces écoulements est d'une importance primordiale pour analyser et prévoir un ensemble de processus physiques et biogéochimiques des zones littorales, tels que le transport des sédiments (stabilité du lit), la diffusion des matières dissoutes tels que les polluants ou les éléments nutritifs ou encore le mélange entre les eaux continentales (douces) et marines (salées). Chacune de ces questions requiert une attention particulière dans le contexte du changement global (élévation du niveau de la mer, événements de submersion, salinisation des terres arables) et de la pression anthropique croissante (urbanisation côtière).

L'ÉTAT DE L'ART ET ORIGINALITÉ.

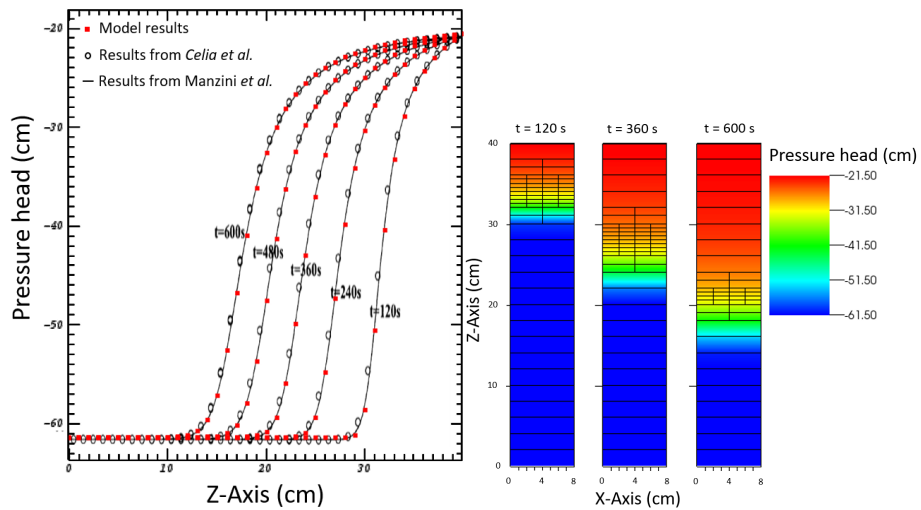
Ces travaux s'inscrivent dans un projet de recherche (voir [Part iv](#)) visant à simuler les *interactions entre les vagues et les eaux souterraines sur les plages de sable*. Des études récentes en laboratoire et sur le terrain [144] ont permis de bien comprendre expérimentalement la dynamique des eaux souterraines et ont montré un schéma de circulation sous la zone de swash. Jusqu'à présent, peu de modèles ont été proposés et leur portée reste limitée. De nouveaux développements numériques sont nécessaires pour capturer un maximum de processus physiques, en particulier pour résoudre tout au long de la phase d'onde la dynamique complexe des champs de saturation et de pression. La modélisation est assez difficile à réaliser car elle nécessite de résoudre à la fois les équations non linéaires pour l'écoulement des eaux souterraines et les équations des eaux peu profondes pour l'écoulement de surface. En outre, ces modèles doivent être couplés de manière appropriée afin de respecter la physique du problème. De ce fait, les infiltrations/exfiltrations de la zone de déferlement, les fluctuations de la saturation en sable, les multiples échelles de temps et les diverses échelles spatiales nécessitent une méthodologie robuste et précise tant pour les schémas numériques que pour les algorithmes, ainsi qu'un aperçu minutieux de la modélisation.

L'équation de Richards est une équation parabolique non linéaire classique pour décrire ces écoulements :

$$\partial_t(\theta(\psi)) - \nabla \cdot (\mathbb{K}(\psi)\nabla(\psi + z)) = 0$$

où ψ indique la pression (m), \mathbb{K} le tenseur de conductivité hydraulique (m/s), z l'élévation (m) et θ la teneur en eau. Ce système est fermé par deux lois constitutives sur la teneur en eau et la conductivité hydraulique.

La méthode *Galerkin discontinue* (ou Discontinuous Galerkin (DG) method) est utilisée car elle offre des avantages appropriés pour l'adaptation hp . Les méthodes discontinues de Galerkin peuvent être considérées comme une méthode partageant des propriétés à la fois des éléments finis et des volumes finis : elles sont basées sur une formulation variationnelle mais de façon élémentaire. Elles sont donc localement conservatrices, ce qui est crucial en dynamique des fluides ([35–39] et Figure 3.6).



(a) Profile de charge.

(b) Raffinement de maillage adaptatif autour du front.

Figure 3.6: Cas test 1d d'infiltration.

3.4 MODÈLES ASYMPTOTIQUES POUR LES ONDES LONGUES

MOTIVATIONS.

Par opposition à l'esprit "mécanique numérique", l'approche asymptotique permet de réduire la dimension du problème mathématique pour aboutir à des équations plus simples à résoudre dont le gain en

temps de calcul est significatif. C'est le cas par exemple, lorsque les écoulements sont en couche mince. On peut réduire ainsi le modèle mathématique 3d de départ en un modèle 2d, voire 1d (voir [Part v](#)).

Dans la continuité de ma thèse, j'ai justifié et développé quelques modèles réduits en couche mince (ou approximation onde longue) et notamment en introduisant le premier modèle réduit 3d vers un modèle 1d non-linéaire et faiblement dispersif pour l'hydraulique fluviale/urbaine. Les applications de ces modèles sont nombreuses et variées.

L'ÉTAT DE L'ART ET ORIGINALITÉ : MODÈLE HYDROSTATIQUE

Dans le cadre des *modèles à surface libre* dans une conduite fermée ou canal, par le biais de calcul asymptotique, j'ai déterminé l'expression du *terme de friction en fonction du périmètre mouillé* P_m pour une géométrie quelconque par la formule

$$K(x, u) = K_0(u) \frac{\int_{\Gamma_b(t,x)} ds}{A}$$

où u est la vitesse moyenne suivant x , $K_0(u) = C_l u + C_t |u| u$ est la friction de coefficient laminaire C_l et turbulent C_t , et $\Gamma_b(t, x)$ désigne la frontière mouillée (*i. e.* la partie de la paroi en contact avec le fluide). Ainsi, la quantité $\frac{A}{\int_{\Gamma_b(t,x)} ds}$ représente le rayon hydraulique qui tient compte de la géométrie de la section. Un résultat similaire a été obtenu pour le *modèle en charge*. Je vous invite à consulter ces articles [\[56, 57\]](#) qui ne sont pas détaillés dans ce mémoire.

L'ÉTAT DE L'ART ET ORIGINALITÉ : MODÈLE HYDROSTATIQUE INTÉGRANT DES TERMES DE PLUIE ET D'INFILTRATION

L'un des modèles les plus utilisés pour décrire l'*hydraulique pluviale* est le modèle de *Saint-Venant*. Nous proposons un nouveau modèle permettant non seulement de modéliser ce phénomène mais aussi de prendre en considération l'*influence de la trainée générée par l'apport de la pluie ou par l'infiltration*. Le modèle obtenu s'écrit comme suit

$$\begin{aligned} \partial_t h + \partial_x q &= S := R - I, \\ \partial_t q + \partial_x \left(\frac{q^2}{h} + g \frac{h^2}{2} \right) &= -gh \partial_x Z + S \frac{q}{h} \\ &\quad - \left(k_+(R) + k_- I + k_0 \left(\frac{q}{h} \right) \right) \frac{q}{h} \end{aligned}$$

où les inconnues $h(t, x)$ and $q(t, x) = h(t, x)u(t, x)$ désignent respectivement la hauteur d'eau et la vitesse moyenne de l'écoulement, g

étant la gravité et Z représente la topographie. Le terme source S représente l'apport par les eaux pluviales R ou la perte d'eau par l'infiltration I . Enfin les termes de la forme k_p représentent la trainée générée par l'apport de la pluie $p = +$ et l'infiltration $p = -$ et la topographie $p = 0$. Je vous invite à consulter cet article [61] qui n'est pas détaillé dans ce mémoire.

L'ÉTAT DE L'ART ET ORIGINALITÉ : EXTENSION AU CAS NON-HYDROSTATIQUE

La modélisation de l'hydrologie des bassins versants et des rivières occupe une place centrale dans les sciences de l'environnement, notamment en ce qui concerne la disponibilité de l'eau, les réseaux d'égouts urbains, les risques d'inondation, *etc.*

L'un des modèles les plus utilisés pour décrire le mouvement des cours d'eau est le modèle à surface libre moyenné par section (voir [19] ou [54, Chapitre 1]), à savoir, les modèles de type Free Surface model (*FS-model*). Cependant, pour certains régimes, ces modèles sont incapables de reproduire les trains d'ondes ou "ondes de choc dispersives" qui sont induits par une distribution de pression non hydrostatique. En procédant à un développement asymptotique à l'ordre deux, on peut ainsi construire un nouveau *modèle non-linéaire et faiblement dispersif* [46, 47] (voir [Part v](#)) qui généralise le *modèle FS-model* et les équations de *Serre-Green-Naghdi* :

$$\begin{cases} \partial_t A + \partial_x Q = 0 \\ \partial_t Q + \partial_x \left(\frac{Q^2}{A} + I_1(x, A) \right) + \mu_2 \partial_x (G(x, A) \mathcal{D}(u)) = \\ I_2(x, A) + \mu_2 \mathcal{G}(x, A, Q) + O(\mu_2^2) \end{cases}$$

où A est l'aire mouillée, $Q = Au$ est le débit moyen, u est la vitesse moyenne par section, I_1 (resp. I_2) est la pression hydrostatique (resp. terme source hydrostatique), $G(x, A)$ généralise le coefficient $\frac{h^3}{3}$ dans les équations de *Serre-Green-Naghdi*, $\mathcal{D}(v) = (\partial_x v)^2 - \partial_x \partial_t v - v \partial_{xx} v$ et $\mathcal{G}(x, A, Q)$ un terme source lié aux effets non-hydrostatique, fonction de la géométrie variable du domaine. Ce modèle est énergétiquement consistant avec les équations d'Euler incompressible en écoulement irrotationnel.

Part II

SENSITIVITY ANALYSIS

*Dans la nature, tout a toujours une raison.
Si tu comprends cette raison,
tu n'as plus besoin de l'expérience.*

— **Léonard de Vinci**

In this part, in collaboration with E. Feireisl and E. Zuazua, we analyze a 1d forced steady state scalar conservation laws [58].

In [Chapter 4](#), we first show the existence and uniqueness of entropy solutions as limits as $t \rightarrow \infty$ of the corresponding solutions of the scalar evolutionary hyperbolic conservation law.

In [Chapter 5](#), we then linearise the steady-state equation with respect to perturbations of the forcing term. This leads to a linear first order differential equation with, possibly, discontinuous coefficients. We show the existence and uniqueness of solutions in the context of duality solutions. We also show that this system corresponds to the steady state version of the linearised evolutionary hyperbolic conservation law. This analysis leads us to the study of the sensitivity of the shock location with respect to variations of the forcing term, an issue that is relevant in applications to optimal control and parameter identification problems.

Possible extensions and further discussions can be found in [Chapter 6](#).

INTRODUCTION

Optimal control of solutions to (non-linear) hyperbolic conservation laws is hampered by the presence of discontinuities (shock waves) making the use of the standard techniques based on linearization rather delicate. The same can be said about optimal design and parameter identification problems. For an analysis of these issues in the context of evolutionary hyperbolic conservation laws in 1d we refer to [29] and [30] and the references therein. The issue of the convergence of time-evolution controls towards the steady state ones as the time horizon tends to infinity is a subject that recently has attracted attention (see, for instance, [28] and [124]). But, as far as we know, this has not yet been addressed in the context of the scalar conservation laws considered here.

In this work, we consider a simplified model problem based on the 1d *scalar steady driven conservation law*

$$\partial_x f(v(x)) + v(x) = g(x), \quad x \in \mathbb{R}, \quad (4.1)$$

supplemented with a “far field” boundary condition

$$\lim_{x \rightarrow \pm\infty} v(x) = 0.$$

Note that, as we shall see in [Section 4.1](#), v can be viewed as the asymptotic limit for $t \rightarrow \infty$ of solutions to the associated evolutionary scalar hyperbolic conservation law

$$\partial_t u(t, x) + \partial_x f(u(t, x)) + u(t, x) = g(x), \quad u(0, x) = u^0. \quad (4.2)$$

Our main goal is to perform a rigorous *sensitivity analysis* of solutions v of the steady problem (4.1) with respect to perturbations of the forcing term g . In particular, denoting v_ε the solution of (4.1) with $g = g + \varepsilon \delta g$ we identify the limit of the quantity

$$\frac{v_\varepsilon - v}{\varepsilon} = h_\varepsilon \rightarrow h \text{ as } \varepsilon \rightarrow 0,$$

where v is the solution associated to g .

It turns out that h is in general a measure on \mathbb{R} , with a singular part sitting on the set of discontinuities of v , taking account of the sensitivity of the shock location of v , see [Chapter 5](#). Our approach is based on an adaptation of the concept of duality solutions introduced by Bouchut and James [15, 16] and Strömberg [146].

4.1 THE ASYMPTOTIC LIMIT AS $t \rightarrow \infty$

This chapter is devoted to analysis of the models under consideration from close, using the notion of entropy solutions. Of course, our analysis can be viewed as a particular instance of the general theory of nonlinear semigroups in $L^1(\mathbb{R})$ (see [44]). The steady problem and the evolution one are linked in the sense that the later can be solved in an unique manner by the semigroup of $L^1(\mathbb{R})$ -contractions generated by the accretive operator associated to the steady state problem.

Throughout this work we suppose that the flux function f is locally Lipschitz,

$$f \in W_{loc}^{1,\infty}(\mathbb{R}).$$

We start with the nowadays standard definition of *entropy solution* to the evolutionary problem (4.2) introduced by Kruřkov [92], [91] and [93].

Definition 4.1.1. Let $g \in L^\infty(\mathbb{R})$. A function $u = u(t, x)$,

$$u \in L^\infty((0, T) \times \mathbb{R}) \cap C([0, T]; L_{loc}^1(\mathbb{R})),$$

is an entropy solution to problem (4.2) if it solves the equation in the sense of distributions, and, moreover, fulfils the integral inequality

$$\begin{aligned} & \int_{\mathbb{R}} E(u(\tau, \cdot)) \varphi(\tau, \cdot) dx - \int_{\mathbb{R}} E(u^0) \varphi(0, \cdot) dx \\ & \leq \int_0^\tau \int_{\mathbb{R}} (E(u) \partial_t \varphi + F(u) \partial_x \varphi) dx dt \\ & + \int_0^\tau \int_{\mathbb{R}} -E'(u) u \varphi + E'(u) g \varphi dx dt, \end{aligned} \quad (4.3)$$

for any $\tau \in [0, T]$, any test function $\varphi \in C_c^\infty([0, T] \times \mathbb{R})$, $\varphi \geq 0$, and any convex entropy $E : \mathbb{R} \rightarrow \mathbb{R}$ with $F' \equiv E' f'$.

Note that the *existence* of global-in-time entropy solutions for $u^0 \in L^\infty(\mathbb{R})$ can be established by means of artificial viscosity approximations (see Kruřkov [93]).

It can also be shown that two entropy solutions u^1, u^2 emanating from the initial data u_0^1, u_0^2 satisfy

$$\int_{|x| \leq M} |u^1(\tau, \cdot) - u^2(\tau, \cdot)| dx$$

$$\begin{aligned}
 & + \int_0^\tau \int_{|x| \leq M + \lambda(\tau-t)} |u^1 - u^2| \, dx \, dt \\
 & \leq \int_{|x| \leq M + \lambda\tau} |u_0^1 - u_0^2| \, dx
 \end{aligned}$$

for any $\tau \geq 0$, $M > 0$, where

$$\begin{aligned}
 \lambda & = \sup \left\{ |f'(z)| \mid \right. \\
 & \left. |z| \leq (1 + \tau) \max \left\{ \|u_0^1\|_{L^\infty(\mathbb{R})}, \|u_0^2\|_{L^\infty(\mathbb{R})}, \|g\|_{L^\infty(\mathbb{R})} \right\} \right\}.
 \end{aligned}$$

Consequently,

$$\begin{aligned}
 & \int_{\mathbb{R}} |u^1(\tau, \cdot) - u^2(\tau, \cdot)| \, dx + \int_0^\tau \int_{\mathbb{R}} |u^1 - u^2| \, dx \, dt \\
 & \leq \int_{\mathbb{R}} |u_0^1 - u_0^2| \, dx.
 \end{aligned}$$

In particular, the entropy solutions are uniquely determined by the initial data.

As an immediate consequence of the previous estimate we deduce the exponential stability of entropy steady state solutions.

Corollary 4.1.1. *Let the flux f be a locally Lipschitz function on \mathbb{R} . Let $v \in L^\infty(\mathbb{R})$ be an entropy solution to the stationary problem (4.1) and u an entropy solution to the evolutionary problem (4.2), with $u(0, \cdot) = u^0 \in L^\infty(\mathbb{R})$, such that*

$$\|u^0 - v\|_{L^1(\mathbb{R})} < \infty.$$

Then

$$\|u(t, \cdot) - v\|_{L^1(\mathbb{R})} \leq \exp(-t) \|u^0 - v\|_{L^1(\mathbb{R})} \text{ for any } t \geq 0.$$

Remark 4.1.1. *In Corollary 4.1.1 the stationary solution v is arbitrary, in particular, it does not need to satisfy the far field condition provided $u^0 - v$ belongs to $L^1(\mathbb{R})$.*

The conclusion of Corollary 4.1.1 remains valid for solutions of general scalar conservation laws in \mathbb{R}^N provided the flux field is locally Lipschitz.

4.2 STATIONARY SOLUTIONS

For the sake of simplicity, we focus on the stationary solutions satisfying the far field condition. To this end, we restrict ourselves to the class of forcing terms

$$g \in L^1 \cap L^\infty(\mathbb{R}).$$

Under these circumstances, in view of (4.1), it is easy to see that the stationary problem (4.1) admits a unique entropy solution v in the same class, namely,

$$v \in L^1 \cap L^\infty(\mathbb{R}).$$

In addition, by the conservation of mass we have

$$\int_{\mathbb{R}} (v^1 - v^2) \, dx = \int_{\mathbb{R}} (g^1 - g^2) \, dx,$$

the comparison principle

$$\int_{\mathbb{R}} [v^1 - v^2]^+ \, dx \leq \int_{\mathbb{R}} [g^1 - g^2]^+ \, dx$$

and the L^1 -contraction property

$$\int_{\mathbb{R}} |v^1 - v^2| \, dx \leq \int_{\mathbb{R}} |g^1 - g^2| \, dx. \quad (4.4)$$

where v^1, v^2 are the entropy solutions of (4.1) with $g = g^1, g = g^2$, respectively.

The existence of stationary solutions out of (4.1) can be proved in several different ways. For instance, using the L^1 -distance as Lyapunov function and LaSalle's invariance principle (see [173]) one can show that there is a unique stationary solution and that all other solutions converge exponentially as $t \rightarrow \infty$ to it. One can also construct the stationary solutions as limits of time periodic solutions of time-period $\tau > 0$ with $\tau \rightarrow 0$. These periodic solutions can be built as fixed points (using Banach contraction principle) of the semigroup map associating the value of the solution at time $t = \tau$ to the initial datum. Furthermore, one can also build the stationary solutions as vanishing viscosity limits of elliptic equations. In fact, as mentioned above, these results are also a particular example of the classical theory of nonlinear contraction semigroups in $L^1(\mathbb{R})$.

The main result of this chapter is as follows:

Proposition 4.2.1. *Let the flux function f be continuously differentiable on \mathbb{R} and non-degenerate in the sense that the critical points $f'(y) = 0$ are isolated in \mathbb{R} . Let*

$$g \in L^1 \cap L^\infty \cap BV(\mathbb{R}).$$

Then the stationary problem (4.1) possesses an entropy solution v determined uniquely in the class

$$v \in L^1 \cap L^\infty(\mathbb{R})$$

such that v is continuous in \mathbb{R} with a possible exception of a countable set of points $\{s_i\}$, with

$$\lim_{x \rightarrow s_i^-} v(x) = v_-^i \neq v_+^i = \lim_{x \rightarrow s_i^+} v(x), \quad i = 1, \dots, N.$$

Each open interval with end points v_-^i, v_+^i contains at least one critical point y^i of f such that either

$$f(y^i) < f(v_-^i) = f(v_+^i) \text{ yielding } v_-^i > v_+^i$$

or

$$f(y^i) > f(v_-^i) = f(v_+^i) \text{ yielding } v_-^i < v_+^i.$$

Proof: Assuming, in addition to (4.2), that

$$g \in BV(\mathbb{R}),$$

(4.4) yields immediately

$$v \in BV(\mathbb{R}).$$

Identifying v with its Lebesgue means,

$$v(x) = \lim_{h \rightarrow 0} \frac{1}{2h} \int_{x-h}^{x+h} v(z) \, dz,$$

we deduce that v is continuous with a possible exception of countably many points at which the left and right limits exist.

Assume that $x_0 \in \mathbb{R}$ is a point of discontinuity of v , specifically,

$$\lim_{x \rightarrow x_0^-} v(x) = v_- \neq v_+ = \lim_{x \rightarrow x_0^+} v(x),$$

where, as $f(v)$ is Lipschitz continuous (since $\partial_x(f(v)) = g - v$),

$$f(v_-) = f(v_+).$$

Note that this null jump condition is the natural limit of the Rankine-Hugoniot condition for the damped evolutionary hyperbolic conservation law (4.2) or, in other words, the condition characterizing stationary shocks.

Since v is an entropy solution of (4.1) we deduce that

$$\partial_x [(f(v) - f(k)) \operatorname{sgn}(v - k)] \leq c \text{ in the sense of distributions,}$$

for some finite c . This can be easily seen by the vanishing viscosity argument, or directly from (4.3) in view of the fact that both g and v are bounded, for instance. In particular,

$$\begin{aligned} & \lim_{x \rightarrow x_0^-} (f(v) - f(k)) \operatorname{sgn}(v - k) \\ &= (f(v_-) - f(k)) \operatorname{sgn}(v_- - k) \geq (f(v_+) - f(k)) \operatorname{sgn}(v_+ - k) \\ &= \lim_{x \rightarrow x_0^+} (f(v) - f(k)) \operatorname{sgn}(v - k) \end{aligned}$$

provided k belongs to the open interval I with the end points v_- , v_+ .

We distinguish three complementary cases:

- $f(k) < f(v_-) = f(v_+)$ for some k in the interval linking v_- and v_+ . In this case, relation (4.2) implies

$$v_- > v_+.$$

- $f(k) > f(v_-) = f(v_+)$ for some k in the interval linking v_- and v_+ . Similarly, we deduce from (4.2) that

$$v_+ > v_-.$$

- The degenerate case in which $f(v_-) = f(k) = f(v_+)$ can be excluded by the main assumptions of the Proposition in which we impose the set of critical points of f to be isolated.

SENSITIVITY WITH RESPECT TO THE FORCING TERM

In this chapter, we study the sensitivity of solutions to the stationary problem (4.1), supplemented with the far field condition, with respect to perturbations of the right-hand side g . To this end, we introduce the concept of *duality solutions* in the spirit of Bouchut and James [16].

5.1 DUALITY SOLUTIONS FOR THE STATIONARY PROBLEM

We start with a prototype example of a scalar conservation law, where the flux function f is strictly convex with a (global) minimum attained in \mathbb{R} . By virtue of Proposition 4.2.1, solutions of problem (4.1) admit a countable (possibly empty) set of singularities where the solutions “jump down” across the shock, meaning they satisfy the so-called *Oleinik condition*. Note that, in particular, the degenerate case can be excluded because of the strict convexity assumption on f .

Furthermore, by comparison (the maximum principle), it can be shown that

$$\partial_x v \leq \bar{g} \text{ in } \mathcal{D}'(\mathbb{R}) \quad (5.1)$$

provided that

$$\partial_x g \leq \bar{g} \text{ in } \mathcal{D}'(\mathbb{R}). \quad (5.2)$$

We now study the effect of a perturbation

$$g_\varepsilon = g + \varepsilon(\delta g),$$

with

$$\delta g \in L^1 \cap L^\infty \cap BV(\mathbb{R})$$

$$\delta g \text{ of compact support in } \mathbb{R},$$

$$\partial_x(\delta g) \leq \overline{\delta g} \text{ in } \mathcal{D}'(\mathbb{R}).$$

This leads to an entropy solution $v_\varepsilon \in L^1 \cap L^\infty(\mathbb{R})$ of (4.1) that can be written in the form

$$v_\varepsilon = v + \varepsilon(\delta v)_\varepsilon,$$

more specifically,

$$\partial_x f([v + \varepsilon(\delta v)_\varepsilon]) + [v + \varepsilon(\delta v)_\varepsilon] = g + \varepsilon(\delta g).$$

Note that, since we deal with entropy solutions,

$$\|(\delta v)_\varepsilon\|_{L^1(\mathbb{R})} \leq \|\delta g\|_{L^1(\mathbb{R})}.$$

Furthermore, we have

$$\partial_x \left[\frac{1}{\varepsilon} \left(\int_0^\varepsilon f'(v + z(\delta v)_\varepsilon) \, dz \right) (\delta v)_\varepsilon \right] + (\delta v)_\varepsilon = \delta g;$$

whence the perturbation $h_\varepsilon = (\delta v)_\varepsilon$ solves the *linear problem*:

$$\partial_x (A_\varepsilon h_\varepsilon) + h_\varepsilon = \delta g,$$

where

$$A_\varepsilon = \frac{1}{\varepsilon} \left(\int_0^\varepsilon f'(v + z(\delta v)_\varepsilon) \, dz \right),$$

$$\|A_\varepsilon\|_{L^\infty(\mathbb{R})} \leq c \left(\|g\|_{L^\infty(\mathbb{R})}, \|(\delta g)\|_{L^\infty(\mathbb{R})} \right),$$

$$A_\varepsilon \rightarrow A \text{ weakly-}^* \text{ in } L^\infty(\mathbb{R}).$$

Furthermore, in accordance with (5.1), (5.2),

$$\begin{aligned} \partial_x A_\varepsilon &= \frac{1}{\varepsilon} \int_0^\varepsilon [f''(v + z(\delta v)_\varepsilon) \partial_x (v + z(\delta v)_\varepsilon)] \, dz \\ &\leq c(f)(\bar{g} + \overline{\delta g}), \end{aligned} \tag{5.3}$$

with

$$\partial_x g \leq \bar{g}, \quad \partial_x(\delta g) \leq \overline{\delta g}.$$

This yields local uniform *BV*-bounds on A_ε . These arguments, together with the continuity of f' allow us to deduce that

$$A = f'(v),$$

As a consequence we expect the limit measure h to be a solution of

$$\partial_x (f'(v)h) + h = \delta g. \tag{5.4}$$

Note however that $f'(v)$ may be discontinuous at the shock discontinuities of v . Thus, we need to introduce a suitable concept of weak solution in the class of measures in order to determine the solution h in a unique manner and to justify the limit process above. For doing that, the fact that the coefficients involved in the linearised equation fulfill a one-sided Lipschitz condition of the form (5.3) will play a key role.

Definition 5.1.1. *We say that h is a duality solution of*

$$\partial_x (Ah) + h = \delta g \quad (5.5)$$

if

$$\frac{d}{dt} \int_{\mathbb{R}} ph \, dx + \int_{\mathbb{R}} ph \, dx = - \int_{\mathbb{R}} (\delta g)p \, dx \quad (5.6)$$

for any reversible solution p of the evolutionary problem

$$\partial_t p + A\partial_x p = 0 \quad (5.7)$$

in the sense of Bouchut and James [16].

Remark 5.1.1.

- *In [16] reversible solutions of the adjoint equation (5.7) are built as limits of approximating sequences of solutions of a final value problem at $t = T$ for regularized potentials A . In [16] it is shown that, under the one-sided Lipschitz condition on A , this leads to a well identified unique solution, the so-called reversible one.*
- *Typically, as shown in [29], when linearizing scalar conservation laws along a solution with a shock discontinuity, the adjoint system of the linearised one has the structure (5.7) with the potential A being discontinuous along the shock. The corresponding solution p can be defined by characteristics away from the zone of influence of the shock. The reversible solution is that taking over this set the corresponding value at the shock location at the final time $t = T$. In this way, for locally Lipschitz continuous data at time $t = T$ reversible solutions are locally Lipschitz.*

Accordingly, the duality formula (5.6) makes sense for h a measure.

- *Note that duality solutions of (5.5) are uniquely determined by the right-hand side δg in the class of measures $h \in \mathcal{M}(\mathbb{R})$.*

Seeing that, by virtue of (5.3), the partial derivatives $\partial_x A_\varepsilon$ are bounded from above so that they fulfill uniformly the one-sided Lipschitz condition. Then we may use the abstract convergence result of Bouchut and James [16] to conclude that

$$h_\varepsilon \equiv (\delta v)_\varepsilon \rightarrow \delta v \text{ weakly-* in } \mathcal{M}(\mathbb{R}),$$

where $\delta v \in \mathcal{M}(\mathbb{R})$ is the unique duality solution of the linearised problem

$$\partial_x (f'(v)\delta v) + \delta v = \delta g.$$

Note that the compactness results in [16] refer to the evolution equation but, accordingly, they can also be applied to steady state solutions. This procedure yields a duality solution to the limit linearised steady-state problem (5.4).

The uniqueness of the duality solution for the steady problem (5.4) is easy to prove. In case there were two distinct solutions, h_1 and h_2 , then $h = h_1 - h_2$ would satisfy

$$\frac{d}{dt} \int_{\mathbb{R}} ph \, dx + \int_{\mathbb{R}} ph \, dx = 0$$

for all reversible solution of (5.7). Considering all possible reversible solutions of (5.7) associated to all Lipschitz continuous data $p_T = p_T(x)$ at $t = T$ we would conclude that

$$\int_{\mathbb{R}} p_T h \, dx = 0,$$

and this would yield $h \equiv 0$.

We have proved the following result.

Theorem 5.1.1. *Let $f : \mathbb{R} \mapsto \mathbb{R}$ be a strictly convex function in $C^2(\mathbb{R})$. Let*

$$g, \delta g \in L^1 \cap L^\infty \cap BV(\mathbb{R})$$

be given such that

$$\partial_x g \leq \bar{g}, \quad \partial_x(\delta g) \leq \overline{\delta g} \text{ in } \mathcal{D}'(\mathbb{R}).$$

Let

$$v_\varepsilon = v + \varepsilon(\delta v)_\varepsilon \in L^1 \cap L^\infty(\mathbb{R})$$

be the solution of the perturbed problem (5.1), where $v \in L^1 \cap L^\infty(\mathbb{R})$ is the unique solution of (4.1).

Then

$$(\delta v)_\varepsilon \rightarrow \delta v \text{ weakly-}^* \text{ in } \mathcal{M}(\mathbb{R}) \text{ as } \varepsilon \rightarrow 0,$$

where (δv) is the unique duality solution of the linear problem

$$\partial_x (f'(v)\delta v) + \delta v = \delta g \tag{5.8}$$

in the sense specified in Definition 5.1.1.

Remark 5.1.2. *In accordance with [16, Theorem 2.2], the function $f'(v)$ in (5.8) may be redefined on the set of zero measure in such a way that the resulting function is Borel and (5.8) holds in the sense of distributions. In particular, the singular part of the measure δv is supported by shocks of v .*

5.2 SENSITIVITY OF SHOCKS

As in the previous chapter, we assume that the flux function f is strictly convex with a (global) minimum attained in \mathbb{R} . Actually, to simplify the presentation we assume that f is even, as it is for instance the case for the Burgers equation where $f(u) = u^2/2$.

Assume that an entropy solution v exhibits a shock discontinuity at $x = 0 \in \text{supp } g$. Our goal is to justify the following relation:

$$\delta\varphi = \lim_{\varepsilon \rightarrow 0} \frac{\varphi_\varepsilon}{\varepsilon} = \frac{[f'(v)\delta v]_0}{[v]_0} \quad (5.9)$$

where $[u]_{x_0} := u(x_0^+) - u(x_0^-)$ stands for the jump of a function u across $x = x_0$, φ_ε stands for the shock location of the perturbed solution $v_\varepsilon = v + \varepsilon\delta v_\varepsilon$ under forcing $g \equiv g + \varepsilon\delta g_\varepsilon$. In particular, we have

$$\lim_{x \rightarrow 0^-} v(x) = v(0^-) > v(0^+) = \lim_{x \rightarrow 0^+} v(x) = -v(0^-).$$

Note that $v(0^+) = \lim_{x \rightarrow 0^+} v(x) = -v(0^-)$ since f is even and $f(v(0^+)) = f(v(0^-))$.

The function $x \mapsto f(v(x))$ is Lipschitz continuous (note that the Rankine-Hugoniot condition ensures in this case the continuity of $f(v)$ across the shock), in particular, there exists $a < 0$ and $b > 0$ such that the solution has the following structure

$$v(a) = 0, v(x) > 0 \text{ in } (a, 0], v \in C([a, 0]),$$

and

$$v(b) = 0, v(x) < 0 \text{ in } [0, b), v \in C([0, b]).$$

Accordingly, the solution may not have an infinite number of shocks accumulating at a point and shocks are isolated. Moreover, if g is as in Theorem 5.1.1, the function $f'(v)$ satisfies the one-sided Lipschitz condition (5.3) and the linearised equation

$$\partial_x(f'(v)\delta v) + \delta v = \delta g, \delta g \in L^1 \cap L^\infty,$$

admits a unique solution

$$\delta v \in L^1(a, 0) \text{ and } \delta v \in L^1(0, b).$$

Consider now the solutions $v_\varepsilon = v + \varepsilon\delta v_\varepsilon$ of the perturbed problem

$$\partial_x(f(v_\varepsilon)) + v_\varepsilon = g + \varepsilon\delta g.$$

Since v_ε, v are entropy solutions, we get

$$\|v_\varepsilon - v\|_{L^1(\mathbb{R})} \leq \varepsilon \|\delta g\|_{L^1(\mathbb{R})}$$

and, in particular,

$$\delta v_\varepsilon \rightarrow \delta v \text{ weakly-}^* \text{ in } \mathcal{M}(\mathbb{R}).$$

In accordance with Theorem 5.1.1, the limit measure δv is unique, the so-called duality solution of the limiting stationary problem. We have

$$f(v_\varepsilon) \rightarrow f(v) \text{ in } C([a, b]),$$

and we deduce that v_ε possesses a unique shock discontinuity at the point $x = \varphi_\varepsilon$ in any compact interval $[a + \delta, b - \delta]$ for any $\delta > 0$ provided $\varepsilon = \varepsilon(\delta)$ is small enough.

Moreover, $\varphi_\varepsilon \rightarrow 0$ as $\varepsilon \rightarrow 0$ and one can quantify the variation of the shock location with respect to the perturbation $g \rightarrow g + \varepsilon \delta g$. Indeed, I_ε being an open interval with the end points $0, \varphi_\varepsilon$. We have

$$\|\delta v_\varepsilon - \delta v\|_{C([a+\delta, b-\delta] \setminus I_\varepsilon)} \rightarrow 0 \text{ as } \varepsilon \rightarrow 0.$$

On the other hand,

$$\int_{I_\varepsilon} \delta v_\varepsilon dx = \frac{1}{\varepsilon} \int_{I_\varepsilon} (v_\varepsilon - v) dx = \frac{\varphi_\varepsilon}{\varepsilon} \frac{1}{|I_\varepsilon|} \int_{I_\varepsilon} |v_\varepsilon - v| dx,$$

where

$$\int_{I_\varepsilon} \delta v_\varepsilon dx \rightarrow [f'(v)\delta v](0+) - [f'(v)\delta v](0-)$$

while

$$\frac{1}{|I_\varepsilon|} \int_{I_\varepsilon} |v_\varepsilon - v| dx \rightarrow [v](0-) - [v](0+).$$

Thus we have proved (5.9).

5.3 STATIONARY PROBLEM SEEN AS AN EVOLUTIONARY ONE

In this chapter, we still assume that the flux function f is strictly convex with a (global) minimum attained in \mathbb{R} .

The sensitivity analysis of the steady equation (4.1) with respect to forcing can be also obtained as a formal limit as $t \rightarrow \infty$ of the unsteady one (4.2) in view of the exponential stability of those solutions (see Corollary 4.1.1).

First of all we consider the evolution problem (4.2) and, for the sake of simplicity, we assume that the solution develops at most one

shock. Let us assume that $u(t, x)$ is a Lipschitz continuous solution of equation (4.2) on Ω^\pm separated by a regular curve

$$\Sigma = \{(t, \varphi(t)), t > 0\},$$

where it satisfies the Rankine-Hugoniot condition

$$\varphi'(t) = \frac{[f(u)]_{\varphi(t)}}{[u]_{\varphi(t)}}$$

and the entropy condition $[u]_{\varphi(t)} \leq 0$.

In the presence of shocks, the state of equation (4.2) needs to be viewed as a pair $(u(t, \cdot), \varphi(\cdot))$ and problem (4.2) can be written as follows:

$$\begin{cases} \partial_t u + \partial_x(f(u)) + u = g(x), & t \geq 0, x \in \Omega(t), \\ \varphi'(t) = \frac{[f(u)]_{\varphi(t)}}{[u]_{\varphi(t)}}, & t \in (0, T), \\ u(0, x) = u^0(x), & x \in \Omega(0), \\ \varphi(0) = \varphi_0 \end{cases}$$

where $\Omega(t)$ stands for the union of the two regions to both sides of the shock:

$$\Omega(t) = \{x < \varphi(t)\} \cup \{x > \varphi(t)\}.$$

We then analyse the sensitivity of the pair (u, φ) with respect to perturbations δg of the source term g . The functional framework adopted here is based on the generalized tangent vectors introduced by Bressan and Marson [27] (see also [29]).

Let

$$g \in L^1 \cap L^\infty \cap BV(\mathbb{R}) \text{ with supp } g \text{ compact in } \mathbb{R}$$

be the source term. Let

$$u^0 \in L^1 \cap L^\infty(\mathbb{R}) \text{ with supp } u^0 \text{ compact in } \mathbb{R}$$

be the initial datum with a single discontinuity at

$$x = \varphi_0 \in \text{supp } g$$

and consider a generalized tangent vector

$$(\delta g, \delta u^0, \delta \varphi_0) \in (L^1 \cap L^\infty \cap BV(\mathbb{R})) \times (L^1 \cap L^\infty(\mathbb{R})) \times \mathbb{R}.$$

Then, for ε small enough, the perturbed solution $u_\varepsilon(t, \cdot) = u(t, \cdot) + \varepsilon \delta u(t, \cdot)$ of (4.2) with $g = g_\varepsilon (\equiv g + \varepsilon \delta g)$ is Lipschitz continuous developing a single discontinuity at $x = \varphi_\varepsilon(t)$ for $t > 0$.

Thus, $u_\varepsilon(t, \cdot)$ generates a generalized tangent vector

$$(\delta u(t, \cdot), \delta \varphi(t)) \in (L^1 \cap L^\infty(\mathbb{R})) \times \mathbb{R}.$$

One can prove (see for instance [26]) that it solves the following linearised system:

$$\begin{cases} \partial_t \delta u + \partial_x (f'(u) \delta u) + \delta u = \delta g, & (t, x) \in (0, T) \times \Omega(t), \\ \delta \varphi'(t) [u]_{\varphi(t)} + \delta \varphi(t) \left(\varphi'(t) [u_x]_{\varphi(t)} - [f'(u) u_x]_{\varphi(t)} \right) \\ + \varphi'(t) [\delta u]_{\varphi(t)} - [f'(u) \delta u]_{\varphi(t)} = 0, & t \in (0, T), \\ \delta u^0(x) = \delta u(0, x), & x \in \Omega(0), \\ \delta \varphi(0) = \delta \varphi_0. \end{cases} \quad (5.10)$$

This system has a unique solution which can be computed in two steps. Characteristics of (5.10) and (4.2) being the same, providing u solution of (4.2), the method of characteristics determines δu outside the shock curve. Thus the values of u and u_x to both sides of the shock curve allow to determine the jump relation through $x = \varphi(t)$. As a consequence all coefficients of the ODE are known and $\delta \varphi$ is then obtained by solving the second equation in (5.10).

Taking into account the exponential stability of u (see Corollary 4.1.1), the linear structure of equation (5.10), and passing formally to the limit as $t \rightarrow \infty$ we get the following linearised steady equation

$$\begin{cases} \partial_x (f'(v) \delta v) + \delta v = \delta g, & x \in \mathbb{R} / \{\varphi\}, \\ \delta \varphi [f'(v) v_x]_\varphi + [f'(v) \delta v]_\varphi = 0. \end{cases}$$

Note that this is a formal argument. To make it rigorous one should carefully prove that the limit as $t \rightarrow \infty$ of all terms actually exists what requires uniform regularity estimates on solutions.

In particular, formally, the second equation is obtained as limit as $t \rightarrow \infty$ of the linearised evolution shock condition in (5.10). Indeed, using, formally, that both $\varphi'(t)$ and $\delta \varphi'(t)$ tend to zero as $t \rightarrow \infty$ we get the linearised shock condition in (5.3).

Moreover it provides the sensitivity of the shock location (which can be also obtained easily by linearizing the static shock condition $[f(v)]_\varphi = 0$) and can be written as

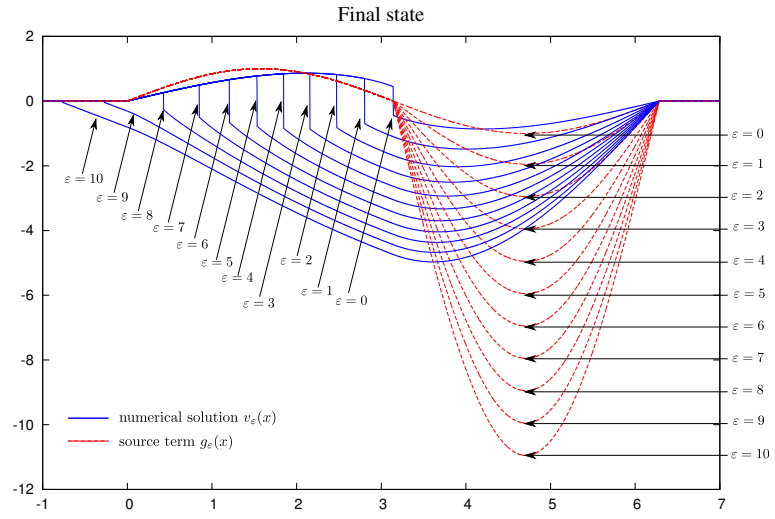
$$\delta \varphi = \frac{[f'(v) \delta v]_\varphi}{[v]_\varphi}$$

(see also equation (5.9)).

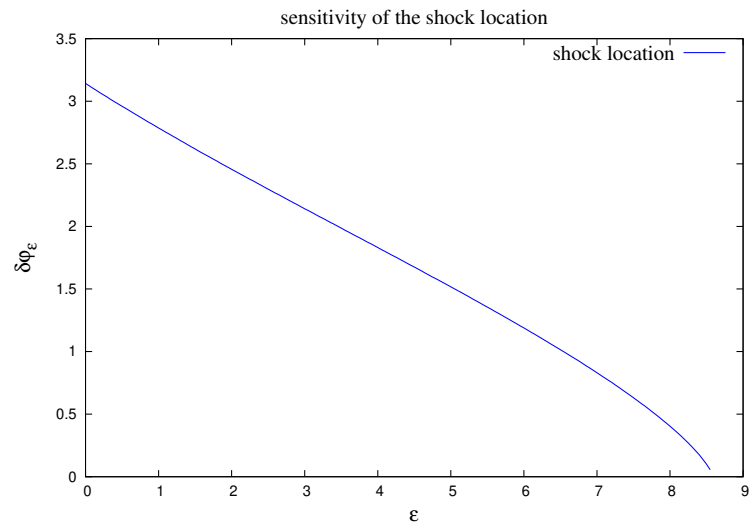
Let us illustrate the sensitivity of the shock location with respect to small perturbations of the source term for T large enough. We consider in Figure 5.1 $g_\varepsilon(x) = g(x) + \varepsilon\delta g(x)$ where

$$\delta g(x) = \begin{cases} 0 & \text{if } 0 \leq x \leq \pi \\ g(x) & \text{if } \pi < x \leq 2\pi \end{cases}$$

as a perturbation of $g(x) = \sin(x)$. In Figure 5.1a we show the behavior of the solution v_ε under the perturbation g_ε and in Figure 5.1b the sensitivity of the shock location. In particular, according to numerical simulations the shock seems to disappear for $\varepsilon \gtrsim 8.6$.



(a) v_ϵ (dashed line: source term, solid line: numerical solution)



(b) $\delta\phi_\epsilon$

Figure 5.1: Numerical illustration of the sensitivity of $(v_\epsilon, \delta\phi_\epsilon)$.

 EXAMPLES, EXTENSIONS, CONCLUDING REMARKS

In this chapter, throughout several examples, we describe some qualitative properties of steady solutions such as their support, shock location, *etc.* We also discuss a property that numerical simulations seem to indicate, according to which, solutions, at least in part, stabilize in finite time. This constitutes an interesting open problem.

For the sake of simplicity, let us again consider the Burgers equation (4.1) with $f(v) = v^2/2$ and $g \in L^1(\mathbb{R}) \cap L^\infty(\mathbb{R})$ of compact support.

6.1 STRUCTURE OF STEADY SOLUTIONS

In what follows, given a finite number of points $\{a_0, \dots, a_N\}$ we consider g as the sum of

$$g(x) = \sum_{i=0}^{N-1} g_i(x) \mathbb{1}_{(a_i, a_{i+1})}(x) \quad (6.1)$$

where $g_i(x) = g(x)|_{(a_i, a_{i+1})}$ is either positive or negative and so that sign changes from a subinterval to the other.

Following Mascia and Terracina [109], under such circumstances, one can easily prove the following:

Properties 6.1.1. *For a given $1 \leq i \leq N$, the stationary solution has at most one shock $s \in [a_{i-1}, a_i]$.*

Proof. Assume by contradiction that there exist two shocks, say, $(s_1, s_2) \in [a_{i-1}, a_i]^2$ with $s_1 < s_2$. Then there exist three C^1 solutions of problem (4.1) v^- , v^m and v^+ such that the solution on the whole interval $[a_{i-1}, a_i]$ is:

$$\begin{aligned} v(x) &= v^-(x) \mathbb{1}_{(a_{i-1} \leq x < s_1)}(x) \\ &+ v_m(x) \mathbb{1}_{(s_1 < x < s_2)}(x) + v^+(x) \mathbb{1}_{(s_2 < x \leq a_i)}(x) \end{aligned}$$

where

$$v^-(s_1^-) > 0 > v_m(s_1^+) \text{ and } v_m(s_2^-) > 0 > v^+(s_2^+).$$

On one hand, since we have $v_m(s_1^+) < 0 < v_m(s_2^-)$ then there exists

$$s_1 < s < s_2 \text{ such that } v_m(s) = 0.$$

On the other, $v_m(x)$ satisfies

$$\forall x \in (s_1, s_2), f'(v_m(x))v'_m(x) = g(x) - v_m(x)$$

with

$$f'(v_m(s)) = f'(0) = 0.$$

As a consequence, we have $g(s) = 0$ which is a contradiction with the decomposition (6.1) of g . \square

In what follows, we illustrate examples of steady solutions of equation (4.1) for $f(v) = v^2/2$ when considering continuous source terms (see Figure 6.1a and Figure 6.2-6.3),

$$g_c(x) = \sin(4x)\mathbb{1}_{(0,2\pi)}(x)$$

and discontinuous source terms (see Figure 6.1b),

$$\begin{aligned} g_{dc}(x) &= \mathbb{1}_{(0,1)}(x) - 0.8\mathbb{1}_{(1,2)}(x) + 0.4\mathbb{1}_{(2,3)}(x) \\ &\quad - 0.5\mathbb{1}_{(3,4)}(x) + 0.7\mathbb{1}_{(4,5)}(x) \\ &\quad - 2\mathbb{1}_{(5,6)}(x) + 2\mathbb{1}_{(6,7)}(x) - \mathbb{1}_{(7,8)}(x) + 3\mathbb{1}_{(8,9)}(x). \end{aligned}$$

The results are displayed on Figure 6.1 and, for both cases, there is only one shock per interval where g keeps the same sign.

Note also that, while in the first example, the support of the steady solution v is included in the one of the forcing term g_c , that is not the case in the second case. Indeed, in that one the support of v goes beyond that of g_{dc} to the right. Over there the solution exhibits a simple linear behavior in agreement with the equation fulfilled beyond the support of the forcing term ($\partial_x(v^2/2) + v = 0$).

6.2 EVOLUTION PROBLEM

Let us come back to the two previous examples. The inspection of the time evolution of several numerical simulations ¹ seems to indicate that there exists one part of the solution (the one localized on the support of g) that stabilizes in finite time (this can be observed, for instance in Figure 6.2), while the other one decays exponentially to zero (as displayed on Figure 6.3) beyond the support of g where the equation reads $\partial_t u + \partial_x f(u) + u = 0$.

¹ We have used the standard Godunov scheme to compute each numerical solution on a fine grid with $N = 10000$ and free boundary conditions.

More precisely, it seems that the following assertion holds:

Conjecture: For any u^0 and g with compact support there exists a finite time $T^* > 0$ such that

$$\forall t \geq T^*, \forall x \in \text{supp } g, \quad u(t, x) = v(x),$$

where v is the unique entropy steady state solution.

Note that rigorous analysis leads to the exponential convergence of the total mass but that the behavior inside the support of the forcing term g requires further work. indeed, let $m(t) = \int_{\mathbb{R}} u(t, x) dx$ be the mass of u over \mathbb{R} . It satisfies

$$m' + m = \int_{\mathbb{R}} g dx.$$

Thus, one has

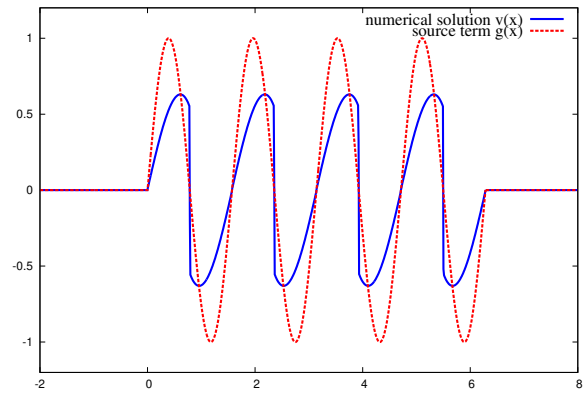
$$\begin{aligned} m(t) &= \int_{\mathbb{R}} g(x) dx + (m(0) - \int_{\mathbb{R}} g(x) dx) e^{-t} \\ &= \int_{\mathbb{R}} v(x) dx + (m(0) - \int_{\mathbb{R}} g(x) dx) e^{-t} \end{aligned}$$

since $\int_{\mathbb{R}} v(x) dx = \int_{\mathbb{R}} g(x) dx$. For a better understanding of the large time behavior and the possible stabilization in finite time a finer analysis of the characteristic of Equation (4.2).

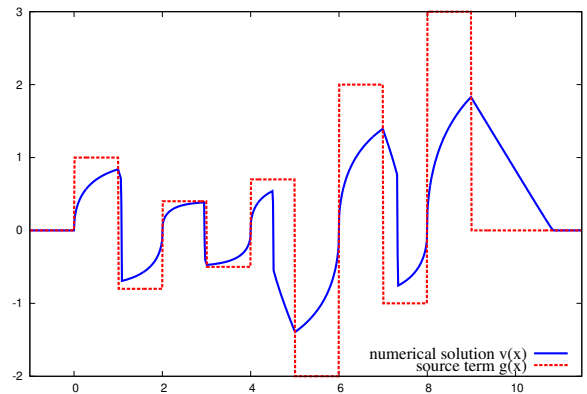
The numerical simulations seem to indicate the following:

- In [Figure 6.2](#), one has $m(0) = \int_{\mathbb{R}} g(x) dx = 0$ and thus $m(t) = 0$, for all $t > 0$. Moreover it seems that $\text{supp } u \subset \text{supp } g$, for all $t > 0$. Such a solution should stabilize in finite time.
- In [Figure 6.3](#), one has $m(0) = 0$, $\int_{\mathbb{R}} g(x) dx \neq 0$ and thus $m(t) = (1 - e^{-t}) \int_{\mathbb{R}} g(x) dx$, for all $t > 0$. Part of $u(t)$ escapes $\text{supp } g$, then it seems to converge to zero exponentially but not in finite time. It seems that, nevertheless, the part of the solution restricted to $\text{supp } g$ reaches the steady state in finite time.

As we said above these global or local finite time stabilization properties constitute interesting open problems to be rigorously analyzed.



(a) g_c



(b) g_{dc}

Figure 6.1: Multiple steady shocks (dashed line: source term, solid line: numerical solution).

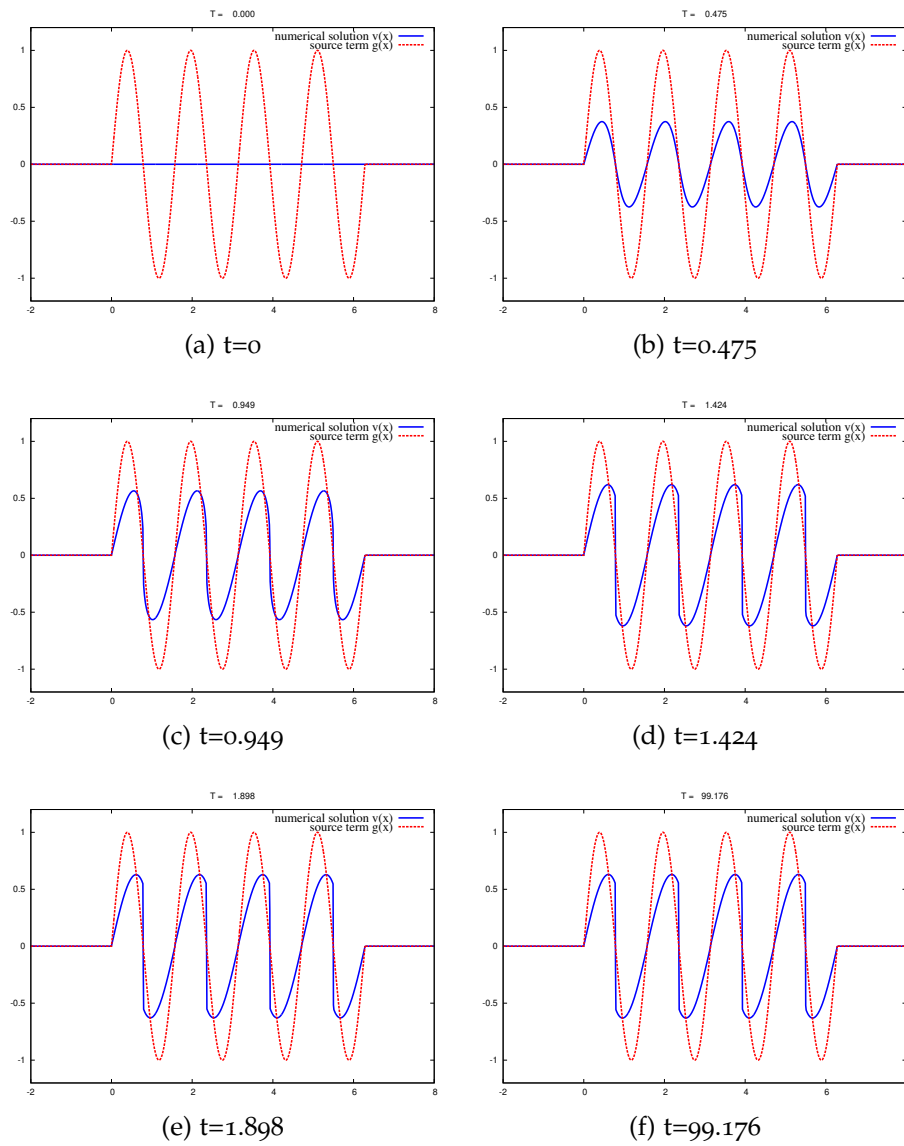


Figure 6.2: A steady solution emerging in finite time with g_c (dashed line: source term, solid line: numerical solution).

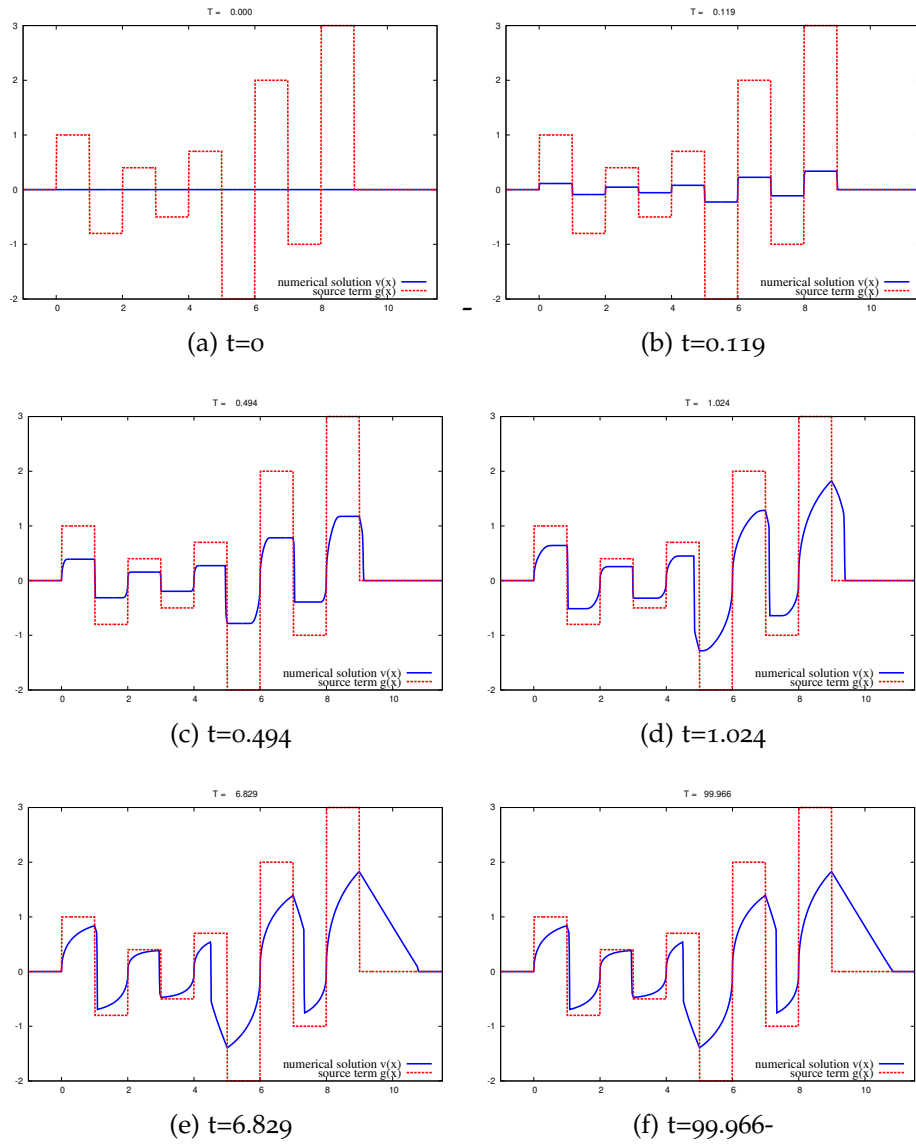


Figure 6.3: A local finite time emerging steady solution with g_{dc} (dashed line: source term, solid line: numerical solution).

Part III

CFD: FINITE VOLUME METHODS, h -AMR AND APPLICATIONS

*Lorsque deux forces sont jointes,
leur efficacité est double.*

— Isaac Newton

In this part, I present the work in collaboration with T. Altazin, F. Golay, R. Marcer, K. Pons, D. Sous, L. Yushchenko done through several papers [[1](#), [59](#), [60](#), [73](#), [74](#), [121](#), [122](#), [166](#), [167](#)].

In [Chapter 7](#), we present the finite volume method developed for non-linear hyperbolic systems in adaptive mesh refinement framework. The mesh refinement algorithm is driven by

- a mesh refinement indicator and,
- a mesh refinement threshold.

A mesh refinement indicator is based on an error estimation procedure or a feature detection technique. In practice, we use the well-known numerical density of entropy production as a mesh refinement indicator and we refer to [[1](#), [59](#), [60](#), [73](#), [74](#), [121](#), [122](#), [166](#), [167](#)] and the reference therein, where we show its efficiency through several numerical test cases in 1-d, 2-d and 3d cases.

As a well-known fact, the mesh refinement threshold is a problem-dependent parameter whose choice might be the critical weakness of mesh adaptation methods. This issue is addressed and an automatic selection of the threshold is proposed in [Chapter 8](#) for which one can show that whatever the mesh refinement indicator is, the obtained results is almost of the same order of accuracy.

PRINCIPLE OF ADAPTIVE MESH REFINEMENT

CFD has grown from a mathematical curiosity to become an essential tool in all branches of fluid dynamics, particularly in coastal hydrodynamics, including wave hydrodynamics. Boosted by the continuous improvement of computer technology, CFD simulations are expected to provide a unique insight, in terms of spatio-temporal resolution and controlled conditions, into such wave dynamics. Therefore, the numerical simulation of waves is both an attractive research topic and a challenging task for coastal, environmental and hydraulic engineering. The involved physical processes are quite complex and the scales are often large in general. A major issue of many modelling challenges is thus to accurately simulate those processes over very large ranges in spatial scale inexorably leading to heavy computational time on uniform meshes.

In principle, AMR [13] or Adaptive Moving Mesh (AMM) [4, 88] methods allow to solve in a reasonable Central Processing Unit (CPU) time these equations adjusting the computational effort locally to maintain a quite uniform level of accuracy. The first one relies on macro cells which can be refined (and then possibly coarsened) while the other method allows all mesh nodes to move to generate big or small cells (with a fixed nodes number for all the simulation). The zones where the mesh needs to be moved or refined is determined thanks to a mesh refinement criterion. Depending on the application, equations and numerical methods, a variety of different criteria might be used based on an error estimation procedure or a feature detection technique. Basically, the zones where a given threshold is exceeded are moved or refined. In particular, for AMR methods, the more the threshold is small, the more accurate is the result at the expense of the computational time. In this work, we focus on h -AMR method.

7.1 GOVERNING EQUATIONS AND FINITE VOLUME APPROXIMATION

In the most of the applications concerned in this work, wave dynamics are governed by non-linear hyperbolic systems:

$$\begin{cases} \frac{\partial \mathbf{w}(t)}{\partial t} + \nabla \cdot \mathbf{f}(t, \mathbf{w}) = \mathbf{G}, & (t, x) \in \mathbb{R}^+ \times \mathbb{R}^d \\ \mathbf{w}(0, x) = \mathbf{w}_0(x), & x \in \mathbb{R}^d \end{cases}$$

where \mathbf{w} , \mathbf{f} , \mathbf{G} stands respectively for conservative variables, flux and source.

In what follows, for the sake of simplicity, we present the numerical algorithms in 1d. We refer to the corresponding papers for the multi-dimensional case (see for instance [1]).

We recall here the well-known construction of a numerical approximation of the following general non-linear hyperbolic conservation laws

$$\begin{cases} \frac{\partial w}{\partial t} + \frac{\partial f(w)}{\partial x} = 0, & (t, x) \in \mathbb{R}^+ \times \Omega \\ \mathbf{w}(0, x) = \mathbf{w}_0(x), & x \in \mathbb{R}. \end{cases} \quad (7.1)$$

where $\Omega \subset \mathbb{R}$, $\mathbf{w} \in \mathbb{R}^d$ denotes the vector state (referred as the conservative variables) and \mathbf{f} denotes the flux governing the physical description of the flow.

The computational domain is split into control volumes $C_k =]x_{k-1/2}, x_{k+1/2}[$ of mesh size h_k with $x_{k\pm 1/2} = x_k \pm h_k/2$. The unknowns $\mathbf{w}(t, x)$ are approximated by their mean values on the cell C_k at time t :

$$\mathbf{w}_k(t) \simeq \frac{1}{h_k} \int_{C_k} \mathbf{w}(t, x) dx.$$

Integrating (7.1) over each cell and applying the Green's formula we obtain:

$$\frac{\partial \mathbf{w}_k}{\partial t} + \frac{\mathbf{F}_{k+1/2}(t) - \mathbf{F}_{k-1/2}(t)}{h_k} = 0, \quad (7.2)$$

where $\mathbf{F}_{k\pm 1/2}$ stands for the numerical flux at the interface $x_{k\pm 1/2}$ (see for instance [65, 71, 153]).

In our case, $\mathbf{F} = \mathbf{f}(\mathbf{R}(0^\pm, \mathbf{w}_L, \mathbf{w}_R))$ is determined from the exact solution $\mathbf{R}(0^\pm, \mathbf{w}_L, \mathbf{w}_R)$ of the Riemann problem associated with the

$$\text{left } \mathbf{w}_L \text{ and the right } \mathbf{w}_R \text{ state: } \mathbf{w}_0(x) = \begin{cases} \mathbf{w}_L & \text{if } x < 0, \\ \mathbf{w}_R & \text{if } x > 0. \end{cases}$$

Equation (7.1) is completed with an entropy inequality of the form:

$$S(\mathbf{w}) = \frac{\partial s(\mathbf{w})}{\partial t} + \frac{\partial \psi(\mathbf{w})}{\partial x} \leq 0, \quad (7.3)$$

where (s, ψ) , with $\psi'(\mathbf{w}) = s'(\mathbf{w})f'(\mathbf{w})$, stands for a convex entropy-entropy flux pair. S is called the *density of entropy production*.

Remark 7.1.1. *As a well-known result, the uniqueness of the (weak) solution of (7.1) is lost even if the initial data are smooth. It can be recovered by completing Equation (7.1) with Equation (7.3). Numerical approximation of Equations (7.1) with Equation (7.3) leads to the so-called numerical density of entropy production which is a measure of the amount of violation of the entropy equation (as a measure of the local residual as in [13, 83, 86, 87]). As a consequence, the numerical density of entropy production provides information on the need to locally refine the mesh (e. g. if the solution develops discontinuities) or to coarsen the mesh (e. g. if the solution is smooth and well-approximated) and can be useful for AMR methods (as already done in [127–129] and [1, 59, 60, 72–74, 121, 122, 166, 167]).*

The Equation (7.3) is discretized using the same scheme as Equation (7.1). Thus, the semi discrete formulation for the entropy inequality reads:

$$S_k(t) = \frac{\partial s(\mathbf{w}_k)}{\partial t} + \frac{\delta \psi_k(t)}{h_k} \quad (7.4)$$

where

$$\delta \psi_k(t) = \psi(\mathbf{R}(0^-, \mathbf{w}_k(t), \mathbf{w}_{k+1}(t))) - \psi(\mathbf{R}(0^+, \mathbf{w}_{k-1}(t), \mathbf{w}_k(t))) .$$

To also achieve a second order approximation in space, the Monotone Upwind Scheme for Conservation Laws (MUSCL) reconstruction can be used as done in [59].

The numerical time integration of Equations (7.2) and (7.4) can be performed through Runge-Kutta or Adams-Bashforth (we refer to [59] for details). Thus, for instance, the first order space-time scheme is :

$$\frac{\mathbf{w}_k^{n+1} - \mathbf{w}_k^n}{\delta t_n} + \frac{\mathbf{F}_{k+1/2}^n - \mathbf{F}_{k-1/2}^n}{h_k} = 0,$$

and

$$S_k^n := \frac{s(\mathbf{w}_k^{n+1}) - s(\mathbf{w}_k^n)}{\delta t_n} + \frac{\delta \psi_k^n}{h_k} \quad (7.5)$$

where δt_n is the time step,

$$F_{k+1/2}^n = f(\mathbf{R}(0^\pm, \mathbf{w}_k^n, \mathbf{w}_{k+1}^n))$$

and,

$$\delta\psi_k^n = \psi(\mathbf{R}(0^-, \mathbf{w}_k^n, \mathbf{w}_{k+1}^n)) - \psi(\mathbf{R}(0^+, \mathbf{w}_{k-1}^n, \mathbf{w}_k^n)) .$$

7.2 ADAPTIVE MESH REFINEMENT ALGORITHM

We recall here the adaptive mesh refinement algorithm using the numerical density of entropy production, for instance Equation (7.5), as a mesh refinement indicator.

7.2.1 Mesh refinement criterion based on the numerical density of entropy production

The construction of an efficient criterion remains a difficult subject, since it depends on physical phenomena and numerical methods. Multi-resolution analysis (see for instance, [80, 113]) provides some robust tools to estimate and control rigorously the error at each refinement level. The *a posteriori* error estimates can be also based on the measure of the local residual as done in [158] and in [87, 171] for instance. Based upon Richardson-type estimates of the local truncation error [13], one can also refine or coarsen to reach a given accuracy for a minimum amount of work. However, near discontinuous solutions, Richardson extrapolation is invalid [126]. It will work but it is no more reliable than heuristic methods and the extra cost cannot be justified. The mesh refinement criterion can be also constructed, as simple as possible, as the combination of gradients based on the underlying physical quantities to detect pertinent regions to refine or to coarsen [130] under the condition that the mesh refinement threshold is well-chosen.

But paradoxically, to our knowledge, very few works use a physical criterion as [84]: the numerical density of entropy production. More recently, Puppo [127, 128] has proved that the numerical density of entropy production can be used as a discontinuity indicator and a local error indicator. Indeed, the entropy inequality being related mathematically and physically to the system (7.1), the numerical density of entropy production (7.5) should be (and actually is, as we will see

in Section 8.3, see also [59, 74] for instance) an accurate and a useful mesh refinement criterion whenever the mesh refinement threshold is well-calibrated (we refer to Chapter 8 for a possible construction of an automatic "well-calibrated" mesh refinement threshold).

7.2.2 Mesh refinement process

We present here a local mesh refinement procedure driven by the numerical density of entropy production, for instance Equation (7.5). In order to reduce the time necessary to manage the refinement, we use "macro cells" which could be refined by generating hierarchical grids. Each cell can be split in two sub-cells. We thus produce a dyadic cells graph, whose numbering (in basis 2) allows a quick computing scan to determine the adjacent cells. For the convenience of the reader, we make use of the following notations: let k_b be the index which makes reference to the macro cell numbered k and b a binary number which contains the hierarchical information of a sub-cell. In particular, the level of a sub-cell C_{k_b} is defined as the $\text{length}(b) - 1$. For instance, a macro cell C_{k_0} of level 0 will be split into two sub-cells $C_{k_{00}}$ and $C_{k_{01}}$ of level 1. A mesh refinement example is proposed in Figure 7.1.

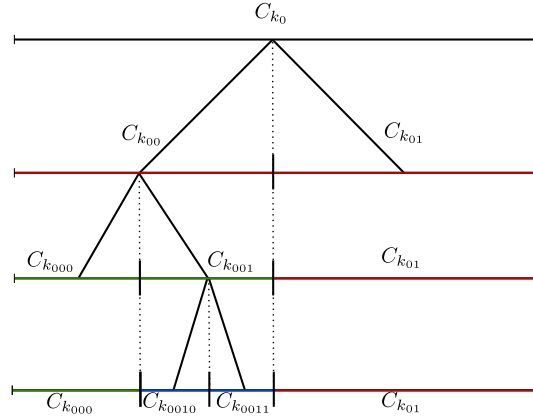


Figure 7.1: Example of hierarchical dyadic tree.

The mesh refinement procedure can be simply expressed as follows. We introduce a non negative mesh refinement criterion $S_{k_b}^n$, for instance Equation (7.5), computed on each cell C_{k_b} at time t_n and compared to a threshold α , for instance,

$$\alpha = S_m = \frac{1}{|\Omega|} \sum_{k_b} S_{k_b}^n$$

In this case, this approach is the so-called mean method (see [121] for other methods).

We then define two coefficients $0 < \beta_{\min} \leq \beta_{\max}$, which determine the ratio of the cells to be refined or coarsened. Thus, for each cell C_{k_b} :

AL 1.

- if $S_{k_b}^n > \alpha_{\max} = \beta_{\max}\alpha$, the cell is refined and split into two sub-cells $C_{k_{b0}}$ and $C_{k_{b1}}$.
- if $S_{k_{b0}}^n < \alpha_{\min} = \beta_{\min}\alpha$ and $S_{k_{b1}}^n < \alpha_{\min}$, the cell is coarsened into a cell C_{k_b} .

The adaptive algorithm is stopped when the level l of a cell C_{k_b} reaches the maximum level l_{\max} .

Remark 7.2.1.

- The threshold parameters β_{\min} and β_{\max} allow to set a percentage of mesh refinement and mesh coarsening with respect to the quantity S_m . The grid is not refined or coarsened for $S_{k_b}^n \in [\alpha_{\min}, \alpha_{\max}]$.
- It is not surprising that these settings can deteriorate or improve the accuracy of the numerical solution. For instance, the more β_{\min} and β_{\max} are small, the more accurate are the results at the expense of the computational time. Thus β_{\min} and β_{\max} are a problem-dependent parameters whose choice might be the critical weakness of the mesh adaptation method. This issue is addressed and an automatic selection of the threshold is proposed in [Chapter 8](#).

On one hand, for the first order scheme, if a cell C_{k_b} is split into two sub-cells $C_{k_{b0}}$ and $C_{k_{b1}}$, averaged values at time t_n are projected on each sub-cell:

$$\mathbf{w}_{k_{b0}}^n = \mathbf{w}_{k_{b1}}^n = \mathbf{w}_{k_b}^n$$

and the numerical fluxes between the two sub-cells (see [Figure 7.2a](#)) are defined as:

$$\left\{ \begin{array}{l} \mathbf{F}_{k_{b0}-1/2}^n = \mathbf{F}_{k_b-1/2}^n, \\ \mathbf{F}_{k_{b0}+1/2}^n = \mathbf{F}_{k_{b1}-1/2}^n = \mathbf{f}(\mathbf{w}_{k_b}^n), \\ \mathbf{F}_{k_{b1}+1/2}^n = \mathbf{F}_{k_b+1/2}^n. \end{array} \right.$$

In the case of second order method, as the gradient is available, we also use:

$$\begin{cases} w_{k_{b0}}^n = w_{k_b}^n - \frac{h_{k_b}}{4} \frac{\partial w_{k_b}^n}{\partial x} , \\ w_{k_{b1}}^n = w_{k_b}^n + \frac{h_{k_b}}{4} \frac{\partial w_{k_b}^n}{\partial x} . \end{cases}$$

On the other hand, if two sub-cells $C_{k_{b0}}$ and $C_{k_{b1}}$ are coarsened, we initialize the new cell C_{k_b} as (see Figure 7.2b):

$$w_{k_b}^n = \frac{1}{2} (w_{k_{b0}}^n + w_{k_{b1}}^n) \text{ and } \begin{cases} F_{k_b-1/2}^n = F_{k_{b0}-1/2}^n , \\ F_{k_b+1/2}^n = F_{k_{b1}+1/2}^n . \end{cases}$$

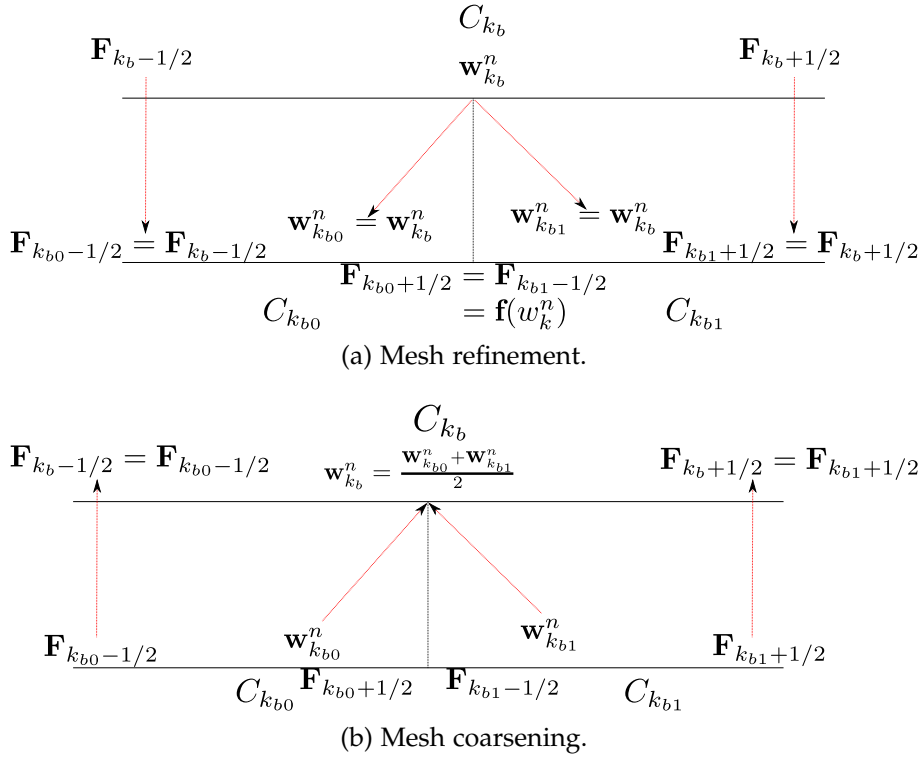


Figure 7.2: Mesh refinement and mesh coarsening: construction of the fluxes.

7.3 THE LOCAL TIME STEPPING METHOD

Unlike the Runge-Kutta method, the Adams-Bashforth time integration can be easily implemented in the *local time stepping* framework in

order to reduce the computational cost. The main approach to update averaged quantities will be recalled and the general algorithm will be presented.

7.3.1 The local time stepping algorithm

Explicit adaptive schemes are well-known to be time consuming due to a Courant, Friedrichs and Lewy (CFL) stability condition. The CPU-time increases rapidly as the mesh is refined since the CFL stability condition imposes an upper bound on $\frac{\delta t}{h}$ where δt is the time step and h the finest mesh size. Nevertheless, the CPU-time can be significantly reduced using the local time stepping algorithm (see *e. g.* [2, 112, 150]). In recent years, such numerical schemes have been widely developed for equations which arise in many fluid flows: traffic flows, multi-phase flows, multi-layer flows, *etc.* The local time stepping algorithm has been employed successfully for real two and three dimensional steady and unsteady problems (see for instance [85, 112, 115, 129, 149, 151]).

In this section, we combine the local time stepping algorithm (see for instance [2]) with the mesh refinement process presented above. For the sake of clarity, we do not use the binary subscript notation introduced in the previous section.

Let h_{\min} be the minimum diameter of the mesh and h_{\max} the maximum diameter of the mesh. All cells of the mesh are sorted in groups, called "level", based on their diameter. Thus, the level of refinement L_k of the k^{th} cell C_k is defined by:

$$2^{N-L_k}h_{\min} \leq h_k < 2^{N+1-L_k}h_{\min} \text{ and } N = \log_2 \left(\frac{h_{\max}}{h_{\min}} \right) + 1$$

where N stands for the maximum level. The coarsest cells are therefore of level 1, while the finest cells are of level N .

Remark 7.3.1. *For any mesh, the previous inequality holds and in the particular case of a dyadic mesh, we have*

$$h_k = 2^{N-L_k}h_{\min} .$$

Let δt_n be the minimum time step at time t_n according to the CFL condition associated with the smallest cell. The macro time step Δt_n is then defined by $\Delta t_n = 2^{N-1}\delta t_n$.

We define the level of interface of two adjacent cells $L_{k+1/2}$ by

$$L_{k+1/2} = \max(L_k, L_{k+1}) .$$

Assuming that the maximum level of refinement is N at the current time t_n , following [2], the local time stepping algorithm reads as follows:

```

foreach  $i \in \{1, 2^{N-1}\}$  do
  Let  $j$  be the biggest integer such that  $2^j$  divides  $i$ 
  foreach interface  $x_{k+1/2}$  such that  $L_{k+1/2} \geq N - j$  do
    1. compute the integral of  $F_{k+1/2}(t)$  on the time interval
        $2^{N-L_{k+1/2}}\delta t_n$ ,
    2. distribute  $F_{k+1/2}(t_n)$  to the two adjacent cells,
    3. update only the cells of level greater than  $N - j$ .
  end
end

```

Remark 7.3.2. *This scheme uses the Osher and Sanders [115] projection. This point is motivated in Section 7.3.2.*

Let us illustrate for a global time advancement the local time stepping algorithm for a first order scheme (the principle for a second order is essentially the same since we have to use the value of the flux stored in the two previous local time steps). We consider a mesh composed of 4 cells of levels 1, 2 and 3 respectively, i.e. here $N = 3$, as displayed in Figure 7.3, Figure 7.4, Figure 7.5 and Figure 7.6. Let us note w_m^n the state vector at time t_n . At the first stage (see Figure 7.3), $i = 1$, only the small cells are advanced with a step δt^n , i.e. $w_3^{n1} := w_3(t_n + \delta t^n)$ and $w_4^{n1} := w_4(t_n + \delta t^n)$ are computed. While $i = 2$ (see Figure 7.4) concerns the small cells of level L_N and the cells immediately larger at level L_{N-1} . The cells L_N travel with a time step δt_n , while the cells L_{N-1} travel with a time step $2\delta t_n$, etc.

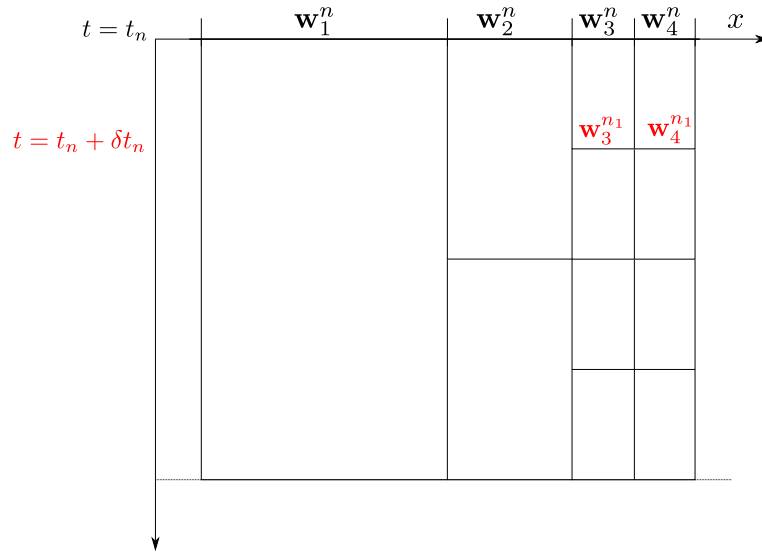


Figure 7.3: Illustration of a global time advancement with a first order scheme at time $t_n + \delta t_n$.

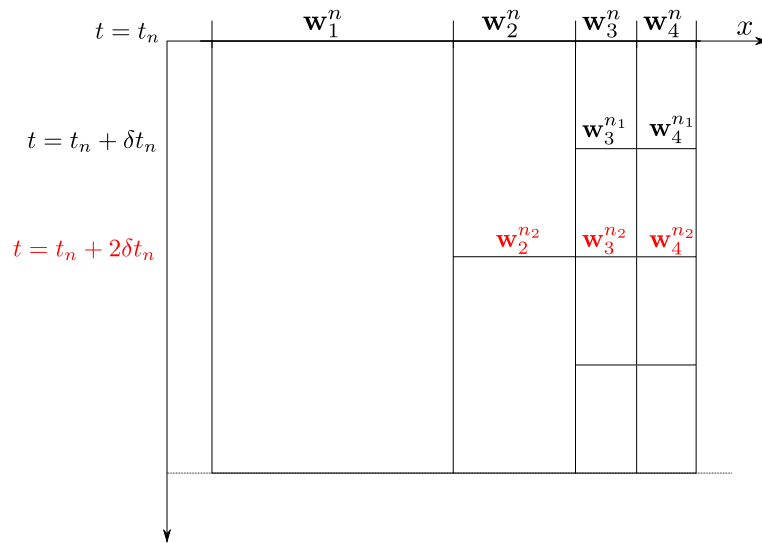


Figure 7.4: Illustration of a global time advancement with a first order scheme at time $t_n + 2\delta t_n$.

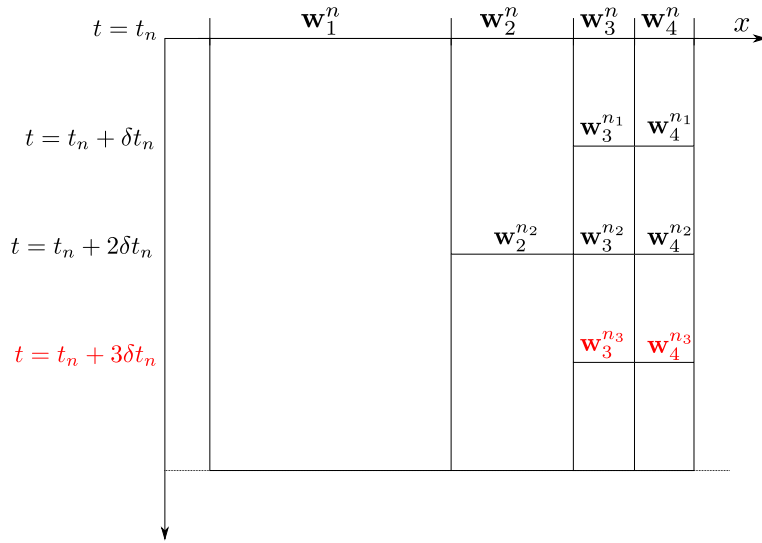


Figure 7.5: Illustration of a global time advancement with a first order scheme at time $t_n + 3\delta t_n$.

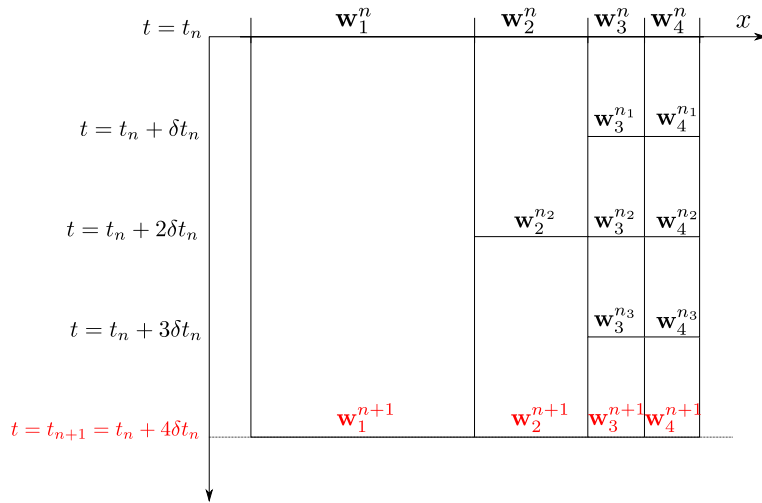
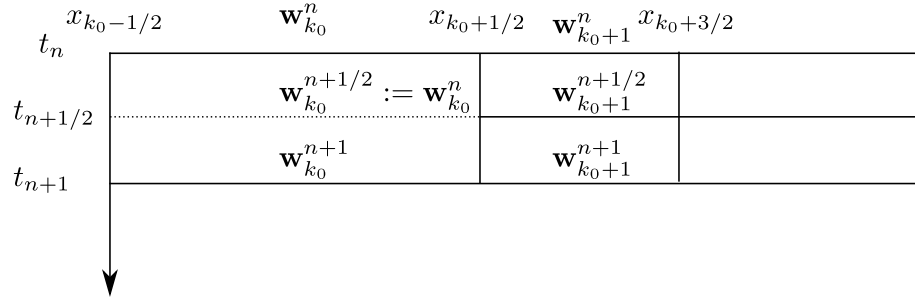


Figure 7.6: Illustration of a global time advancement with a first order scheme at time $t_n + 4\delta t_n$.

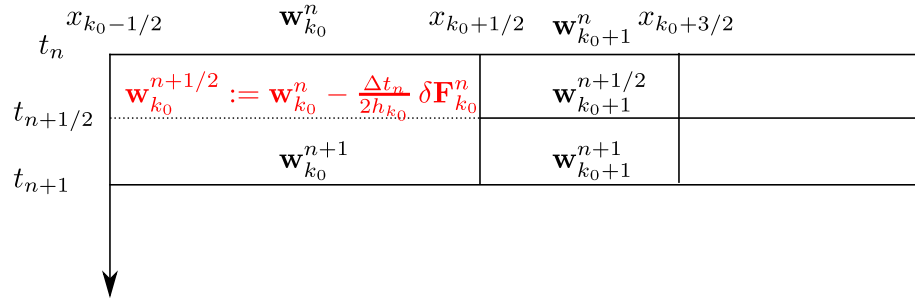
7.3.2 Computation of the flux

We recall the procedure introduced by Osher-Sanders (OS) [115] and, Tang-Warnecke (TW) [151] for the computation of the flux at the cell interface of two different levels. For the sake of simplicity, we assume here only two time increments which means that we consider cells

of only two different kinds, say $h(x) = h_{k_0} = h_0$ for $x < x_{k_0+1/2}$ and $h(x) = h_{k_0+1} = h_0/2$ for $x > x_{k_0+1/2}$. Thus, for every $k \leq k_0$, all cells C_k are of level L_{k_0} and for every $k > k_0$, all cells C_k are of level $L_{k_0} + 1$ (as displayed in Figure 7.7). Let $\Delta t_n := t_{n+1} - t_n = 2\delta t_n$ be the macro time step. To avoid boundary terms, we consider the example with an initial condition with compact support.



(a) Osher and Sanders projection.



(b) Tang and Warnecke projection.

Figure 7.7: Projection's methods.

7.3.2.1 A simple projection method

The OS scheme is defined as follows (see Figure 7.7a): $\forall k \in \mathbb{Z}$,

$$\begin{aligned} w_k^{n+1/2} &= \begin{cases} w_k^n & \text{if } k \leq k_0, \\ w_k^n - \frac{\Delta t_n}{2h_k} \delta F_k^n & \text{if } k > k_0 \end{cases} \\ w_k^{n+1} &= w_k^n - \frac{\Delta t_n}{2h_k} \left(\delta F_k^n + \delta F_k^{n+1/2} \right). \end{aligned} \quad (7.6)$$

It was pointed out by Tang and Warnecke [151] that the present scheme loses the conservative property and it is locally non consistent in the sense of the truncation error. Indeed, at the intermediate time step, one has

$$\sum_{k \in \mathbb{Z}} w_k^{n+1/2} = \sum_{k \in \mathbb{Z}} w_k^n + \frac{\Delta t_n}{2h_{k_0+1}} F_{k_0+1/2}^n(w_{k_0}^n, w_{k_0+1}^n)$$

which shows that this scheme is locally non conservative. At the next intermediate step, the conservativity is preserved, i.e.

$$\sum_{k \in \mathbb{Z}} w_k^{n+1} = \sum_{k \in \mathbb{Z}} w_k^{n+1/2}$$

but the global conservativity is lost.

The fact that the scheme is not consistent in the sense of truncation error can be observed in the simplest case of the one-dimensional transport equation (as done in [151]):

$$\partial_t w + c \partial_x w = 0$$

with a positive constant c . In that case, using the Godunov solver, the scheme (7.6) writes for $k = k_0$:

$$\frac{w_{k_0+1}^{n+1} - w_{k_0+1}^n}{\Delta t_n} - \frac{c}{2h_{k_0}} \left(w_{k_0+1}^n - w_{k_0}^n + \left(w_{k_0+1}^{n+1/2} - w_{k_0}^n \right) \right) = 0 .$$

Assuming w smooth enough, writing the Taylor expansion of w at point x_{k_0+1} at time t_{n+1} and $t_{n+1/2}$, we deduce the following equation

$$\partial_t w + c \partial_x w = -\frac{c}{4} \frac{\Delta t}{h} \partial_t w + \mathcal{O}(h, \Delta t).$$

Thus, if $\frac{\Delta t}{h}$ converge to a constant $d > 0$ then the previous equation does not converge to the initial transport equation.

7.3.2.2 A natural projection method

Contrary to the Osher and Sanders [115] approach, Tang and Warnecke [151] propose the following scheme: $\forall k \in \mathbb{Z}$,

$$\begin{aligned} w_k^{n+1/2} &= w_k^n - \frac{\Delta t_n}{2h_k} \delta F_k^n \\ w_k^{n+1} &= w_k^{n+1/2} - \frac{\Delta t_n}{2h_k} \begin{cases} \delta F_k^n & \text{if } k \leq k_0 \\ \delta F_k^{n+1/2} & \text{if } k > k_0 \end{cases} . \end{aligned}$$

One can easily check that the global conservativity is also lost and

$$\sum_{k \in \mathbb{Z}} w_k^{n+1} = \sum_{k \in \mathbb{Z}} w_k^n + \frac{\Delta t_n}{2h_{k_0+1}} F_{k_0+1/2}^{n+1/2}(w_{k_0}^n, w_{k_0+1}^n) .$$

Although the scheme is not conservative, they claim that it is consistent in the sense of the truncation error.

Even if the OS projection leads to non conservative and non consistent scheme at the space-time grid between two levels of refinement, it has the advantage to be less time consuming than the second one. In general, the apparent loss of consistency on the local truncation error seems not to affect the error of the scheme (see [41, 141] for interesting issues).

Moreover, one can "reinforce" the consistency and the stability of the OS projection by preventing two adjacent cells from having a level difference greater than two. This trick is referred as "*smoothing grid*" in the following.

Comparing numerically the OS and TW projection methods, in the case of the Sod's shock tube problem for the one-dimensional gas dynamics equations for ideal gas, we get more stable results with the smoothing grid technique and we save the computational time (see [59] for more details) with respect to the TW projection:

$$\begin{aligned} \text{CPU-time for TW projection} &= 288.21 \text{ (s)}, \\ \text{CPU-time for OS projection} &= 80.04 \text{ (s)}. \end{aligned}$$

7.4 DATA MANAGEMENT IN HIGH DIMENSION

The multi-dimensional method is almost the same as described in the previous Section 7.1, Section 7.2 and Section 7.3 to the exception of the data management.

Keeping in mind the following requirements: parallel treatment, fast computation and hierarchical grid (mesh level cannot exceed 2 for stability reasons as pointed above in Section 7.3.2 and in [59]), the main and challenging difficulty is to develop a suitable mesh refinement tool which is related on the design of software data structures and numerical scheme.

As pointed out by several authors, the extension toward three-dimensional case is not an easy task. Interesting works have been presented for 2d Cartesian grid or quad-tree [12, 111, 165, 169], octree for 3d simulations [68, 105], and anisotropic AMR [43, 79].

We aim to generalise the scheme proposed Section 7.1, Section 7.2 and Section 7.3 (see [59]) with a finite volume solver. This naturally leads to quad-tree meshing in two dimensional problem and octree

meshing in three dimensional problem.

In the most of the applications, the presence of a complex moving air-water interface or more generally wave interactions implies to re-mesh at each time step, which is clearly time consuming. Keeping in mind our willingness to find a relevant compromise between the contradictory aims of physical relevancy and numerical performance, we introduce a Cartesian block-based mesh approach (somehow like in [162, 172]). The computational domain is split in several blocks. In practice, each block corresponds to the initial unstructured mesh composed of hexahedral cells. According to the mesh refinement procedure (with respect to S_k^n , S_m , the mesh refinement level N), each block is sub-divided in a Cartesian way $(2^{N-1}n_x, 2^{N-1}n_y, 2^{N-1}n_z)$ where (n_x, n_y, n_z) stands for the initial block discretisation.

As done in Section 7.1, Section 7.2 and Section 7.3 (see [59]), for each refined cell (or block), averaged values at time t^n are projected on each sub-cell and each flux is computed as simply as possible to avoid heavy computation.

The interface defining two blocks is most of the time is a non conforming one as displayed in Figure 7.9b.

Next, in order to spare proportionally the balanced distribution of CPU load, the cells of each block are re-distributed in a fixed number of domains according to the Cuthill-McKee numbering. The number of domain being fixed, each domain are loaded in a given MPI process. The procedure is certainly not optimal but the synchronisation time between each domain are almost the same, hence overall more efficient. Finally, the re-numbering and re-meshing being expensive, the mesh is kept constant on a time interval, called AMR time-step, given by *the smallest block* (and not by the smallest cell) and the maximum velocity. This procedure provides a good compromise between computational time and accuracy (see for instance [1, 74, 167] and in Section 8.3).

Let us illustrate the algorithm to define unstructured meshes using the BB-AMR scheme in the two-dimensional case (*i.e.* $n_z = 1$ always). During the refinement process, the number of blocks and domains are kept always constant. As displayed in Figure 7.8a, we consider an initial mesh composed of 9 blocks, assumed to be of level 1 (the coarsest level with $n_x = n_y = 1$), and decomposed on 3 domains according to a Cuthill-McKee numbering of the blocks (3 + 3 + 3 internal cells). For example the first domain (blue) is composed of

three internal cells and three adjacent "ghost" cells in order to share to send/receive information between domains. After some computations, at the AMR-time T_1 , let us suppose that blocks 1 and 4 have to be refined. As displayed in Figure 7.8b, the blocks distribution into the domains has been modified (5 + 5 + 5 internal cells). One can observed that the new mesh is a non-conforming mesh. For example, the cell belonging to the block 2 is bordered by 7 faces. If, at the AMR-time T_2 , only the block 1 exceeds the mesh refinement criterion, it has to be refined. But, as the level of two adjacent blocks cannot exceeds 2 in order to avoid oscillations, the block 2 must also be refined. As displayed in Figure 7.8c, the blocks distribution has been modified (16 + 7 + 7 internal cells). The distribution is not well balanced because we distribute the blocks and not the cells for sake of simplicity. This drawback is reduced as the number of blocks is much bigger than the number of domains.

The three dimensional case is dealt in the same way. For illustration purpose, Figure 7.9 shows an example of 3d block-based mesh decomposition (64 + 39 + 39 internal cells) with 3 domains and 27 blocks, where the first block in the corner is of mesh refinement level 3.

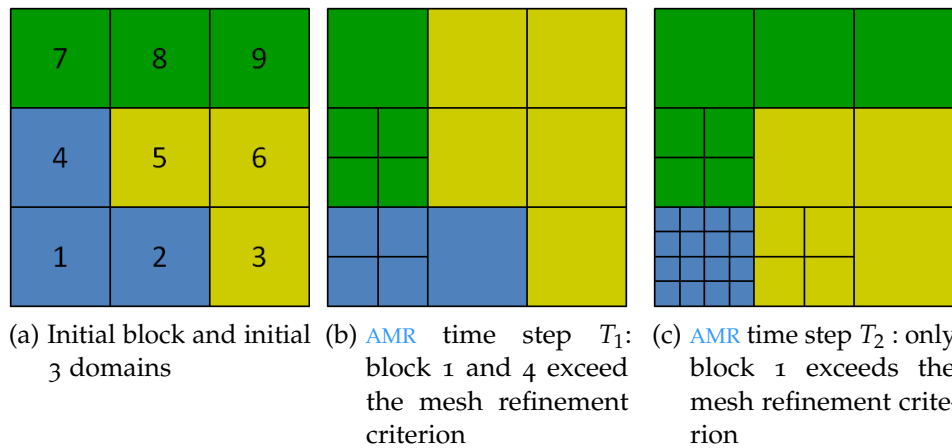
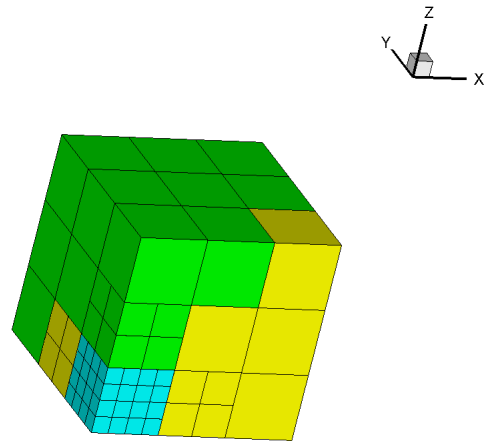
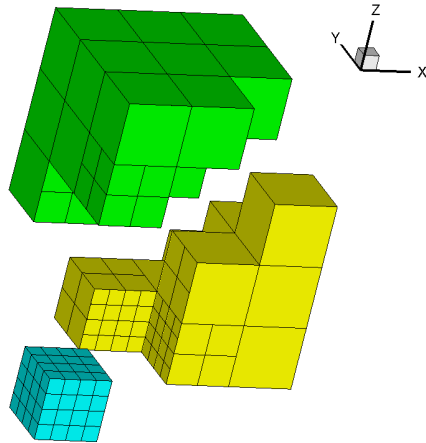


Figure 7.8: Example of two dimensional BB-AMR with 3 domains (blue-yellow-green) and 9 blocks.



(a) Block-based mesh



(b) Domain decomposition

Figure 7.9: Example of three dimensional block-based mesh with 3 domains and 27 blocks.

AUTOMATIC THRESHOLDING METHOD

In [Chapter 7](#), we have overlooked the mesh refinement threshold which is a problem-dependent parameter whose choice might be the critical weakness of mesh adaptation methods. In this chapter, we address a new construction of an automatic selection of a suitable mesh refinement threshold. It allows to detect almost all relevant regions to refine without any hand-calibration as usually done for mesh adaptation algorithms (see [AL 1](#) and [\[121\]](#)).

The construction of the automatic selection of the threshold is based on the decreasing rearrangement function of the mesh refinement criterion. If S denotes the mesh refinement criterion, the decreasing rearrangement is the Lebesgue measure of the set $\{S(x) > \alpha\}$. This function sorts the local maxima of S from the smallest to the largest. One can therefore set automatically the threshold small enough to detect the most relevant cells to refine. In practice, we are able to reach a good balance between the accuracy and the computational time without any hand-calibration as usually done in mesh adaptation algorithms as shown in [Section 8.3](#) (see also [\[122\]](#)).

In what follows, we define the piecewise mesh refinement criterion as follows :

$$S(x, t) = \sum_k S_k^n \mathbb{1}_{C_k \times [t_n, t_{n+1}[}(x) \quad (8.1)$$

$$\text{where } \mathbb{1}_A(x) = \begin{cases} 1 & \text{if } x \in A, \\ 0 & \text{otherwise.} \end{cases}$$

8.1 SETTINGS AND DEFINITION

Let us consider a smooth (at least twice differentiable) mesh refinement criterion $S(x, t) \in \mathbb{R}^+$, $x \in [0, L]$ and $t > 0$ where L is the length of the domain. The time t being fixed, we write in the sequel $S = S(x)$. Without loss of generality, we suppose that $S(0) = S(L) = S'(0) = S'(L) = 0$ and $0 < S_\infty = \max_{x \in (0, L)} S(x) < \infty$ (if $S_\infty = 0$ the numerical

solution is identically equal to zero). The function S can be regarded here as the regularisation of the discrete mesh refinement criterion (8.1).

In view of the above assumptions, the set

$$Z_\alpha = \{x \in (0, L); \varphi_\alpha(x) = S(x) - \alpha = 0 \text{ and } S'(x) \neq 0\}$$

is not empty. Indeed, for each $0 < \alpha < S_\infty$, since $S_\infty > 0$, S has at least one maximum. Then, there exists $p_\alpha \in \mathbb{N}_*$ such that the number of elements in the set Z_α is

$$\#Z_\alpha = 2p_\alpha.$$

Thus, for all $\alpha \in (0, S_\infty)$, one can describe the set Z_α as follows

$$Z_\alpha = \{x_0(\alpha) < x_1(\alpha) < \dots < x_{2p_\alpha-2}(\alpha) < x_{2p_\alpha-1}(\alpha)\}.$$

Let us assume that S has p local maxima. Then there exists an *increasing* sequence $(\alpha_k^*)_{1 \leq k \leq p}$ and a sequence $(x_k^*)_{1 \leq k \leq p}$ such that

$$\forall k = 1, \dots, p \quad S'(x_k^*) = 0, \quad S''(x_k^*) < 0 \text{ and } S(x_k^*) = \alpha_k^*.$$

By construction, the sequence $(\alpha_k^*)_{1 \leq k \leq p}$ represents the local maxima sorted from the smallest to the largest with $\alpha_p^* = S_\infty$. Let us also remark that $x_k^* \notin Z_\alpha, \forall k = 1, \dots, p, \forall \alpha \in (0, S_\infty)$.

With these settings, one can define the decreasing rearrangement function of S :

$$\alpha \in [0, S_\infty] \mapsto d(\alpha) := \begin{cases} L & \text{if } \alpha = 0, \\ \sum_{k=1}^{p_\alpha} x_{2k+1}(\alpha) - x_{2k}(\alpha) & \text{if } 0 < \alpha < S_\infty, \\ 0 & \text{if } \alpha = S_\infty. \end{cases} \quad (8.2)$$

We have

Theorem 8.1.1. *Let $l \in \mathbb{N}_*$, $S \in C^l([0, L], \mathbb{R}^+)$ be a smooth function such that $S(0) = S(L) = S'(0) = S'(L) = 0, 0 < S_\infty < \infty$ with p local maxima. Then*

1. $\int_0^{S_\infty} d(\alpha) d\alpha = \int_0^L S(x) dx.$
2. $d \in C^0([0, S_\infty], \mathbb{R}^+)$ and
 - a) $\forall \alpha \in [0, S_\infty], d'(\alpha) < 0$
 - b) $\forall k \in \llbracket 0, p \rrbracket, \lim_{\alpha \rightarrow \alpha_k^*} d'(\alpha) = -\infty$ with the convention $\alpha_0^* := 0$

3. $d \in C^l(D_*, \mathbb{R}^+)$ on the set $D^* := \bigcup_{k=0}^{p-1} (\alpha_k^*, \alpha_{k+1}^*)$.

Let us now illustrate the previous introduced notations using the given mesh refinement criterion displayed in Figure 8.1a. The decreasing rearrangement function d of S is represented in Figure 8.1b. The first and the second order derivatives of d are shown in Figure 8.1c and Figure 8.1d. In these figures, the distribution is normalised by L and we have $p = 2$, i.e. p local maxima.

In Figure 8.1a, for $\alpha = 0.5$, the set Z_α is composed of $4 = 2p_\alpha$ elements which are approximately

$$x_0(\alpha) \approx 0.8803, \quad x_1(\alpha) \approx 1.6196, \quad x_2(\alpha) \approx 3.379, \quad x_3(\alpha) \approx 4.121.$$

The distribution function for $\alpha = 0.5$ is given by

$$d(\alpha) = \frac{(x_1(\alpha) - x_0(\alpha)) + (x_3(\alpha) - x_2(\alpha))}{L} \approx 0.3.$$

In Figure 8.1b (see also Figure 8.1a), each vertical asymptote of d indicates a local maximum of $S(x)$. More precisely, we have

$$\alpha_1^* = S(x_1^* = 1.25) = 1 \quad \text{and} \quad \alpha_2^* = S(x_2^* = 3.75) = 2 = S_\infty$$

where $(x_k^*)_{1 \leq k \leq 2}$ are the local maxima of S . Thus, by construction, $(\alpha_k^*)_{1 \leq k \leq 2}$ is an increasing sequence.

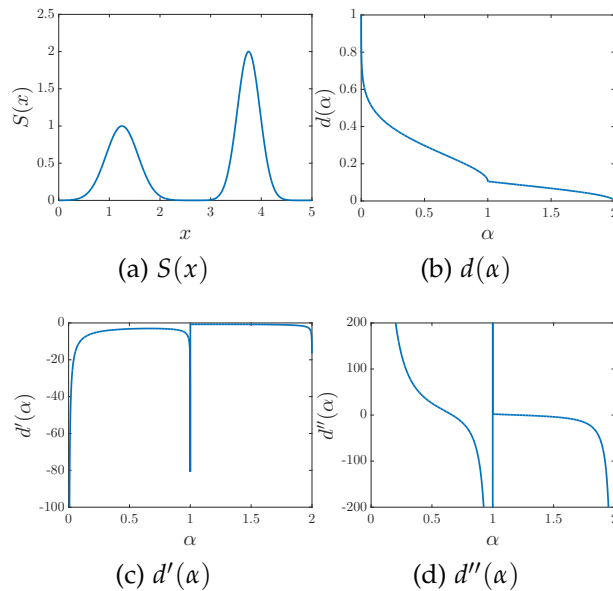


Figure 8.1: The distribution function d , the first order d' and the second order d'' derivatives for the mesh refinement criterion $S(x) = a \exp(-b(x - x_B)^2) + \exp(-5(x - x_A)^2)$ where $x_A = 1.25$, $x_B = 3.75$, $a = 2$ and $b = 10$.

8.2 CONSTRUCTION OF THE THRESHOLD

As pointed out before in [45, 125], it is difficult to define an automatic threshold from the decreasing rearrangement function (8.2) without using the first order or/and the second order derivatives of d to characterise the local maxima. Moreover, we emphasise in [121] that the use high order derivatives requires an accurate interpolation of d to be accurate in the localisation of this local maxima which can be quite costly. To this purpose, in order to avoid these disadvantages, we have proposed the following weight function (8.3). This function is built only from the decreasing rearrangement function (8.2):

$$f(\alpha) = \alpha d(\alpha) \quad (8.3)$$

and has the following useful localisation properties:

Corollary 8.2.1. *Assume that S is twice differentiable and has p local maxima. Then, by virtue of Theorem 8.1.1, the function d (8.2) has p inflexion points*

$$\forall k = 0 \dots p - 1, \alpha_{k+1}^{**} \in (\alpha_k^*, \alpha_{k+1}^*) \text{ such that } d''(\alpha_{k+1}^{**}) = 0 .$$

Moreover, the function f (8.3) has p local maxima $\bar{\alpha}_1, \dots, \bar{\alpha}_p$ such that

$$\forall k = 1 \dots p, \bar{\alpha}_k \in (\alpha_k^{**}, \alpha_k^*) .$$

As emphasised in [121], $\alpha = S_m := \frac{1}{|\Omega|} \int_{\Omega} S(x) dx$, i.e. $\beta_{\min} = \beta_{\max} = 1$, is in general relevant to detect cells to refine for smooth solutions and $\alpha < S_m$, i.e. $\beta < \beta_{\min} = \beta_{\max} = 1$, for discontinuous solutions (see AL 1).

Thus, by virtue of the Corollary 8.2.1 and the previous statement, it is natural to look for an automatic threshold as follows:

$$\alpha_{\text{PE}} \text{ such that } f(\alpha_{\text{PE}}) = \max_{0 < \alpha \leq S_m} f(\alpha) .$$

In order to illustrate the previous statements, we consider the function $S(x) = a \exp(-b(x - x_B)^2) + \exp(-5(x - x_A)^2)$ where $x_A = 1.25$ and $x_B = 3.75$. We represent in Figure 8.2 the shape of the function f (8.3) and in Figure 8.3 the mesh refinement criterion for a discontinuous and a smooth solution. As expected, we get

- for a discontinuous solution, see Figure 8.2a and 8.3a, $\alpha_{\text{PE}} < S_m$.

- for a smooth solution, see Figure 8.2b and 8.3b, $\alpha_{PE} = S_m$.

In both cases, the computed threshold is pertinent to detect the region to refine.

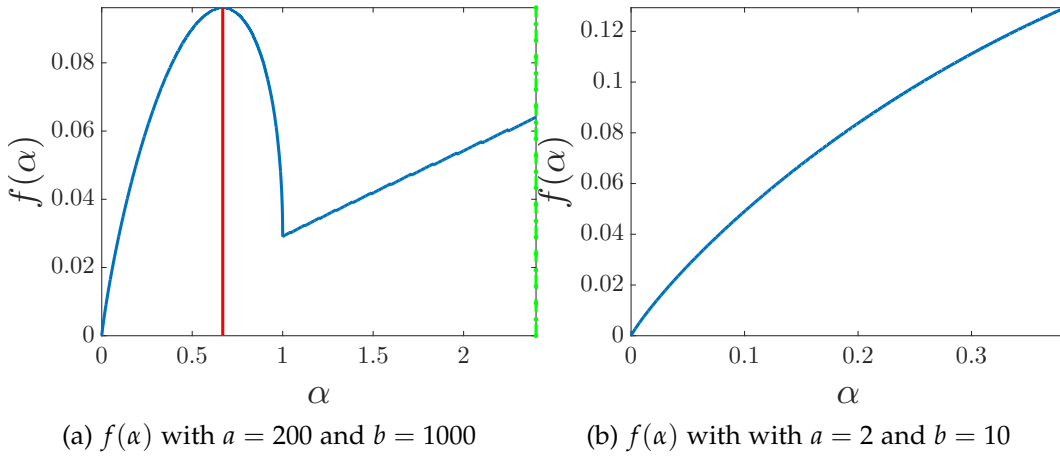


Figure 8.2: The function f for the mesh refinement criterion $S(x) = a \exp(-b(x - x_B)^2) + \exp(-5(x - x_A)^2)$ where $x_A = 1.25$ and $x_B = 3.75$.

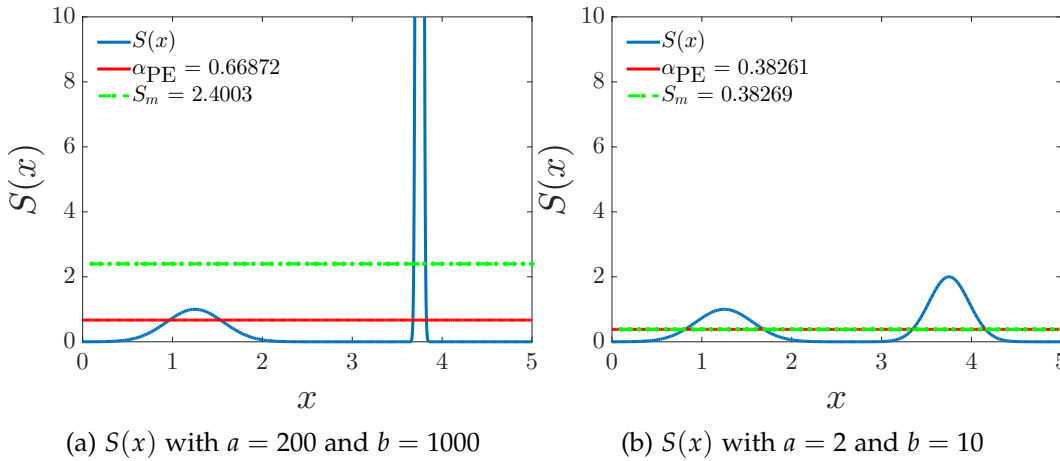


Figure 8.3: The mesh refinement criterion $S(x) = a \exp(-b(x - x_B)^2) + \exp(-5(x - x_A)^2)$ where $x_A = 1.25$ and $x_B = 3.75$.

Finally, the discretisation of the decreasing rearrangement function (8.2) is performed as follows

$$d(\alpha) = \frac{1}{N} \sum_{j=0}^{M-1} d_j \mathbb{1}_{(\alpha_j, \alpha_{j+1})}(\alpha) \text{ with } d_j = \#\{k; S_k^n > \alpha_j\}$$

where $\#$ is the number of elements in the set $\{S_k^n > \alpha_j\}$ and where α_j is defined by

$$\alpha_j = S_m \left(\frac{j}{M} \right), \quad 0 \leq j \leq M. \quad (8.4)$$

In [121], we show that for $M \geq N$, the threshold α_{PE} is quasi-insensitive and make our method totally *parameterless* and less expensive than the methods proposed in [45] and [125].

The algorithm **AL 1** is then defined by

$$\alpha_{\min} = \beta_{\min} \alpha_{\text{PE}} \text{ and } \alpha_{\max} = \beta_{\max} \alpha_{\text{PE}}$$

where $0 < \beta_{\min} \leq \beta_{\max}$.

8.3 NUMERICAL RESULTS

To illustrate the ability of the method to catch relevant information, we start first with simple waves in the case of a one dimensional Riemann problem in Section 8.3.1. Then, the case of wave interactions is dealt in Section 8.3.2 for the Monai-Walley tsunami run-up onto a complex three dimensional beach. We also show the efficiency of the **BB-AMR** presented in Section 7.4. For both test case here, we have considered the Saint-Venant equations :

$$\begin{aligned} \partial_t h + \operatorname{div}(hu) &= 0 \\ \partial_t(hu) + \operatorname{div}\left(hu \otimes u + \frac{g}{2}h^2I\right) &= -gh\nabla Z \end{aligned} \quad (8.5)$$

where the unknowns $h(x, t)$ and $\mathbf{u}(x, t) = (u_1(x, t), u_2(x, t))$ are respectively the height of the water and the depth-averaged velocity of the water at a space-time point (x, t) , $\mathbf{x} = (x_1, x_2) \in \mathbb{R}^2$ is the space coordinate, $t > 0$ is the time, g is the gravitational constant set to 9.81m/s^2 , $Z(x)$ is the bottom surface elevation and I is the identity matrix.

I also refer to [1, 59, 60, 73, 74, 121, 122, 166, 167] for more test cases.

8.3.1 1d Riemann problem over a flat bottom

We consider the Riemann problem with the initial data

$$\forall x \in [0, 80], (h(x, 0), u(x, 0)) = \begin{cases} (5.64, 8) & \text{if } x < 20, \\ (0.6, 8) & \text{if } x > 20 \end{cases} \quad (8.6)$$

where the exact solution is composed of a left-going rarefaction and a right-going shock wave (see for instance [Figure 8.4a](#), black line).

The goal is to show that, whatever the mesh refinement criterion is, the threshold α_{PF} is selected automatically to adapt the mesh without hand-calibration. To this purpose, we propose :

CRITERION 0: Since the exact solution of the Riemann problem (8.5)-(8.6) is known, we consider the following criterion

$$S_k^n := |h_k^n - h_{\text{ex}}(x_k, t_n)|$$

where h_{ex} stands for the exact solution of (8.5)-(8.6).

CRITERION 1: We consider the numerical density of entropy production:

$$S_k^n = |\mathcal{E}_k^n|$$

where

$$\mathcal{E}_k^n = \frac{E(\mathbf{w}_k^{n+1}) - E(\mathbf{w}_k^n)}{\delta t_n} + \frac{(\mathcal{G}_{k+1/2}(\mathbf{w}_k^n, \mathbf{w}_{k+1}^n) - \mathcal{G}_{k-1/2}(\mathbf{w}_{k-1}^n, \mathbf{w}_k^n))}{\delta x_k}$$

with $E(\mathbf{w}) = \frac{q^2}{2h} + \frac{gh^2}{2}$, $q = hu$ and $G(\mathbf{w}) = \left(E(\mathbf{w}) + \frac{gh^2}{2}\right)u$. The numerical flux $\mathcal{G}_{k+1/2}$ is computed with the exact local solution of the 1D Riemann problem at the interface $x_{k+1/2}$ with the left state \mathbf{w}_k^n and the right state \mathbf{w}_{k+1}^n as done in [Chapter 7](#).

CRITERION 2: Finally, the last mesh refinement criterion is based on the gradient method. It can be constructed as a combination of gradients based on physical quantities to detect pertinent regions to refine or to coarsen [130]. To make it as simple as possible, we consider the gradient of h which can be an efficient mesh

refinement criterion if the threshold is well-chosen. Thus, for instance, we define

$$S_k^n = \left| \frac{h_{k+1}^n - h_k^n}{x_{k+1} - x_k} \right|.$$

Each numerical experiment starts initially with $N(0) = 100$ cells, the maximum level of refinement is set to $l_{\max} = 3$, $M = N(t)$ for all $t \geq 0$ (see Eq. (8.4)) and we prescribe free boundary conditions.

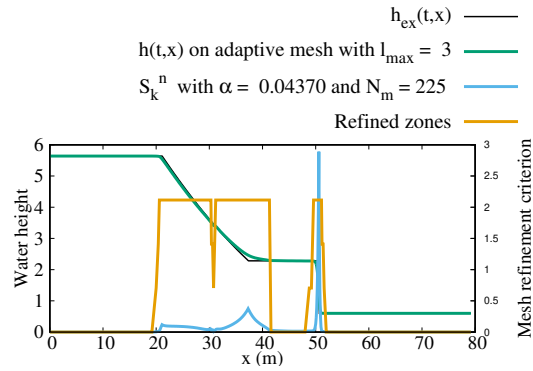
Even if we perform the following strategy for each cell C_k :

- if $S_k^n > \alpha_{\text{PE}}$, the mesh is refined and split,
- if $S_k^n < \alpha_{\text{PE}}$ the mesh is coarsened

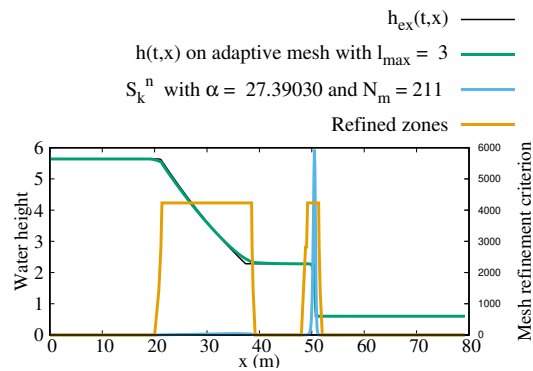
which is known to be a source of spurious oscillations, the proposed method does not generate spurious oscillations in the computation of the numerical solution (see [121] for further details) at least for simple waves as shown in Figure 8.4. Moreover, whatever the mesh refinement criterion, we obtain almost the same numerical solution with almost the same order of accuracy as presented in Table 8.1 (see also Figure 8.4). Finally, in comparison with the mean method, if the threshold is not well-calibrated, the region of interests are refined or too refined as displayed in Figure 8.5.

| | $N_m(l) \rightarrow \ h - h_{ex}\ _1$ | $N_m(l) \rightarrow \ u - u_{ex}\ _1$ |
|-------------|---------------------------------------|---------------------------------------|
| Criterion 0 | 2.0136 | 2.1342 |
| Criterion 1 | 2.0491 | 2.0901 |
| Criterion 2 | 2.1109 | 2.1789 |

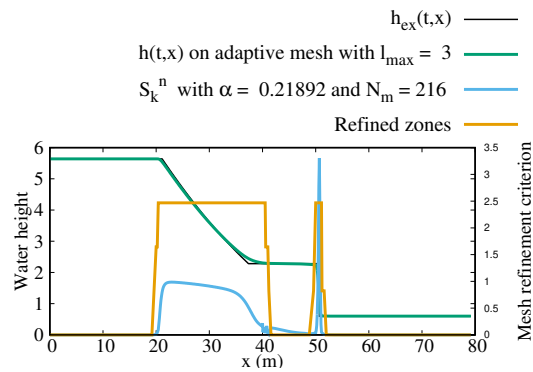
Table 8.1: Convergence rate of the L^1 discretisation error obtained with the refinement level $l = 1, 2, 3$ and 4.



(a) Criterion 0



(b) Criterion 1



(c) Criterion 2

Figure 8.4: Numerical results for the water height at time $t = 2$ s using the automatic threshold α_{PE} . N_m is the mean number of cells used during the simulation.

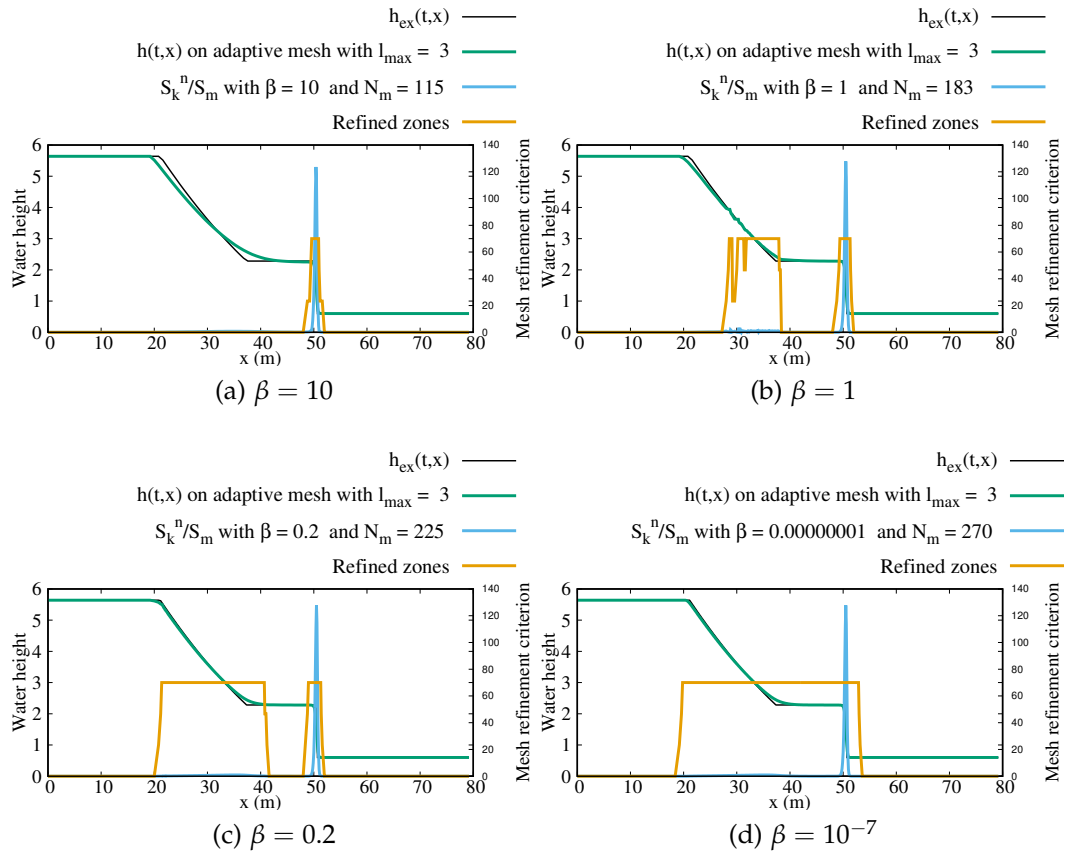


Figure 8.5: Numerical results for the water height at time $t = 2$ s using the criterion 1 with the mean method. N_m is the mean number of cells used during the simulation.

8.3.2 Tsunami runup onto a complex three dimensional Monai-Walley beach

We now focus on the overall performance of the **BB-AMR** (presented in [Section 7.4](#)) scheme with automatic thresholding compared to experimental and state-of-the-art numerical results¹. For each test case, the numerical simulations are carried out on adaptive grids and on uniform grids. The minimum adaptive cell size $\delta x_{adaptive}$ is equal to the uniform one $\delta x_{uniform}$. The comparisons of both simulations allow to show the reliability and the efficiency of the **BB-AMR** method. To simplify the analysis, we have considered only 1 domain, *i.e.*, 1 processor. Interested readers can find a detailed numerical study with several domains in [74]. For all computations,

- the **MUSCL** reconstruction and the second order Runge-Kutta scheme (**RK2**) method are used to achieve a second order scheme in space and time.
- to simplify the analysis and to focus on the automatic threshold, we have chosen $\beta_{min} = \beta_{max} = 1$.
- $M = N$ (see Equation (8.4)).
- the **CFL** number is set to 0.9,
- For the sake of simplicity, the re-meshing time step is fixed constant to 0.25 s. This choice roughly verifies the macro **CFL** condition.

This test case concerns the Hokkaido-Nansei-Oki tsunami, in 1993, that struck the Okushiri Island in Japan. The tsunami run-up records was about 30 m height and the currents speed of order 10-18 m/s for which relevant high-quality data were saved. ² The largest value run-up (32 m) was recorded near the Monai-Walley beach.

8.3.2.1 Experimental settings

To understand this complex run-up, a 1/400 scale laboratory model of Monai was realized in a large-scale tank of 205 m long, 6 m deep,

¹ The numerical soft used for the following test cases is the EOLENS code developed by the Institut de Mathématiques de Toulon (IMATH) and Principia. The code is written in Fortran 90, compiled with gfortran on Linux machine Intel(R) Core(TM) i5-2500 CPU @ 3.30GHz.

² Several sources and data can be found, see for instance http://nctr.pmel.noaa.gov/benchmark/Laboratory/Laboratory_MonaiValley/ or http://isec.nacse.org/workshop/2004_cornell/bmark2.html

and 3.5 m wide at the Central Research Institute for Electric Power Industry (CRIEPI) in Abiko (Japan). The offshore incident wave on a water depth $d = 13.5$ cm is prescribed. There are reflective side-walls at $y = 0$ and 3.5 m as for the x boundaries. The bathymetry as well as the coastal topography reproduced in the laboratory experiment are represented³ in [Figure 8.6b](#) and [Figure 8.6c](#).

The input wave at $x = 0$ m is a leading-depression height of -2.5 mm with a crest of 1.6 cm, as displayed in [Figure 8.6a](#). In the experiment, the waves are measured at thirteen locations and complete time histories are given at three locations, $(x, y) = (4.521, 1.196)$, $(4.521, 1.696)$, and $(4.521, 2.196)$ in meters, see also [\[104\]](#). In contrast with the previous experimental test cases, the dispersive phenomena can be fully neglected here.

8.3.2.2 Numerical vs experimental results.

The initial condition for this test case is a lake at rest and an imposed water height on the left side wall, see [Figure 8.6a](#). Wave gauges data⁴ experimental results are given up to time $t = 100$ s. However, the information concerning the input wave profile (see [Figure 8.6a](#)) at $x = 0$ m is only available up to time $t = 22$ s. For numerical purpose, we prescribe for $t > 22$ s, a constant free surface level $\eta = 0$ m on the left wall. In practice, due to this assumption, some artificial reflecting waves can occur on the wave-maker beyond $t > 22$ s and can reach the free surface gauges around $t = 30$ s. Finally, to avoid any wrong comparisons due to these artificial reflections, the comparisons between the experimental gauges and the numerical results will be only done over the first 30 s of the experiment⁵

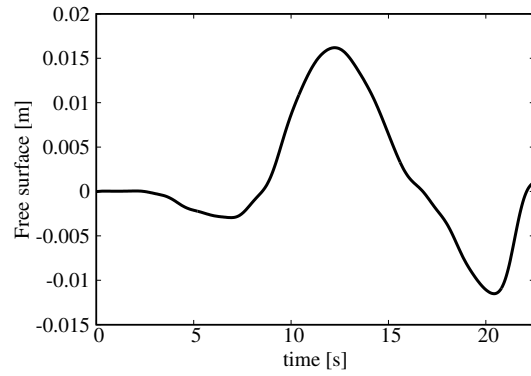
For computational purpose, we have considered 240 initial blocks composed of 8 000 (varying up to 40 000) cells for the initialisation of the adaptive simulation and 62 000 cells for uniform mesh simulation. The time simulation is 30s. Reflecting boundary conditions are prescribed to walls.

To illustrate this test case we show in [Figure 8.7](#)–[Figure 8.10](#), the propagation of the wave at time $t = 11.25$ s ([Figure 8.7](#)), $t = 13.25$ s

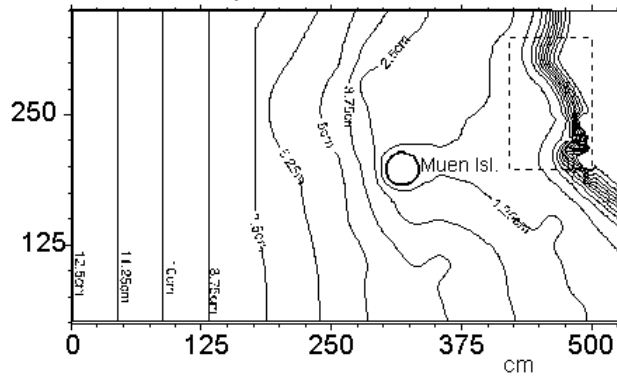
³ sources and pictures are available at http://isec.nacse.org/workshop/2004_cornell/bmark2.html

⁴ Wave gauges data are available at the address http://nctr.pmel.noaa.gov/benchmark/Laboratory/Laboratory_MonaiValley/

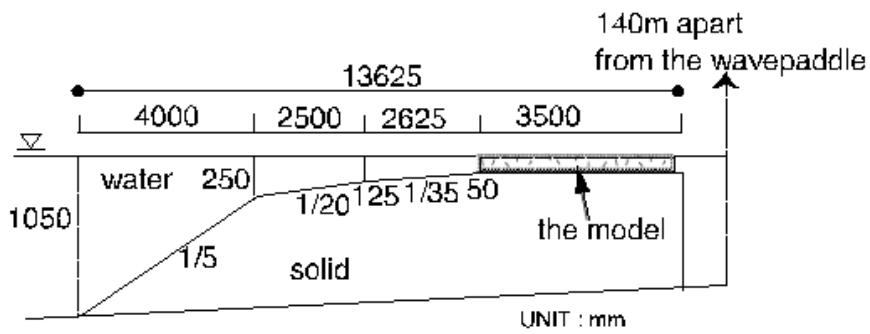
⁵ However, let us note that in [\[52, 53\]](#), the simulation is done over 30 s and can be a point of comparison for future work. At the time of writing [\[122\]](#), we are unaware of Dutykh' results [\[52, 53\]](#).



(a) Initial wave profile



(b) Top view



(c) Side view

Figure 8.6: Experimental settings.

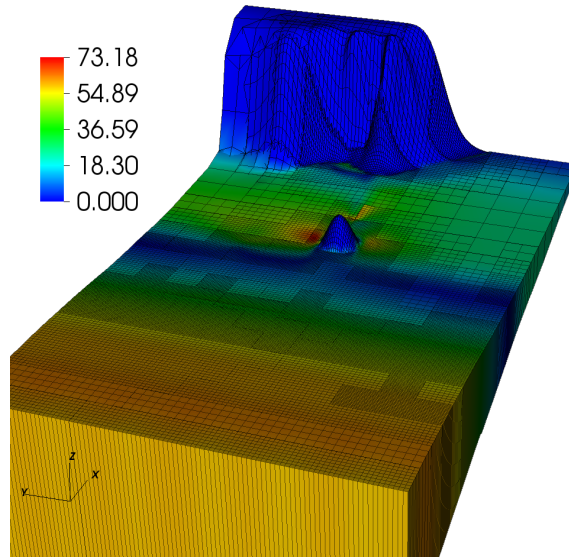


Figure 8.7: Numerical result for the water height at time $t = 11.25$ s. The colours correspond to the magnitude of the kinetic energy.

(Figure 8.8), $t = 16$ s (Figure 8.9) and $t = 17.5$ s (Figure 8.10). The Figure 8.7 represents the tsunami wave entering inside the computational domain. The Figure 8.8 shows the submersion of the conical island. Figure 8.9 and Figure 8.10 show respectively the large flooding and run-up of the wave and its reflection on the cliff.

The numerical solutions are now quantitatively compared to the experimental measures. The comparisons are shown in Figure 8.11 at the three waves gauges WG 1 (see Figure 8.11a), 2 (see Figure 8.11b) and 3 (see Figure 8.11c). These figures show that the wave amplitude and the wave shape are accurately predicted at three different location in the computed domain. As emphasised before, dispersive effects being negligible, the numerical results are accurate. However, it can be noticed in Figure 8.11 a small amplitude errors occurring at the beginning of the simulations. This error can be attributed to the lack of accuracy of the prescribed initial condition and boundary condition at $x = 0$. This problem was already encountered by Popinet [123].

8.3.2.3 AMR performance

To highlight the efficiency of the mesh refinement parameter and the automatic selection of the threshold, the numerical density of entropy production (without absolute value) and the associated mesh are displayed at several times in Figure 8.12–Figure 8.15, (time $t = 11.25$ s

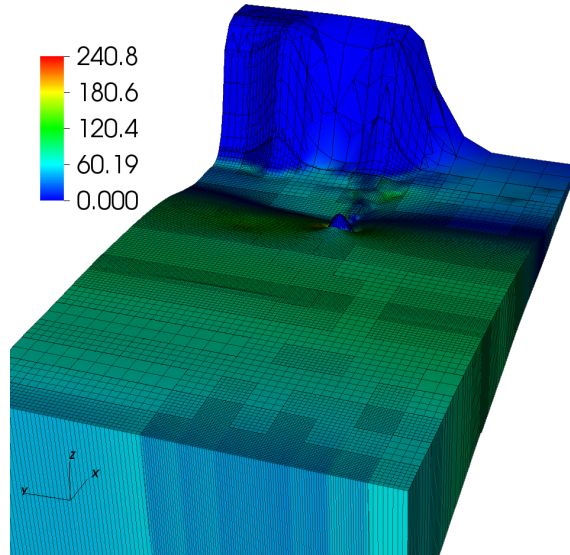


Figure 8.8: Numerical result for the water height at time $t = 13.25$ s. The colours correspond to the magnitude of the kinetic energy.

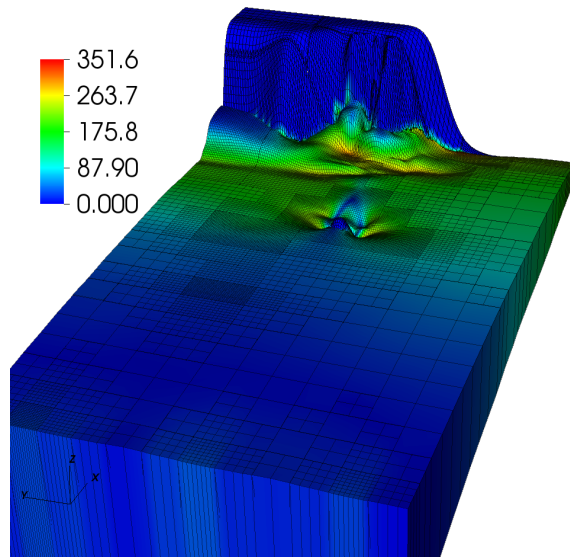


Figure 8.9: Numerical result for the water height at time $t = 16$ s. The colours correspond to the magnitude of the kinetic energy.

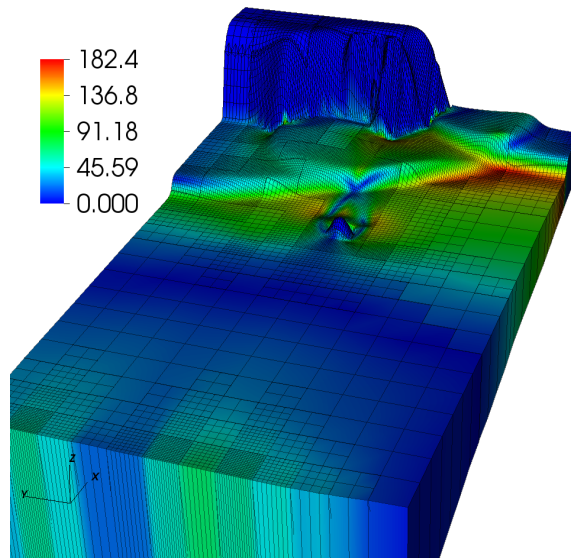
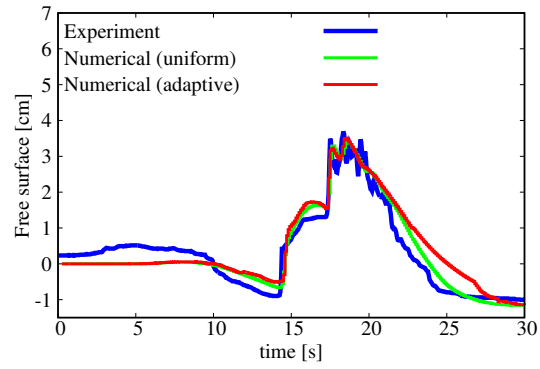


Figure 8.10: Numerical result for the water height at time $t = 17.5$ s. The colours correspond to the magnitude of the kinetic energy.

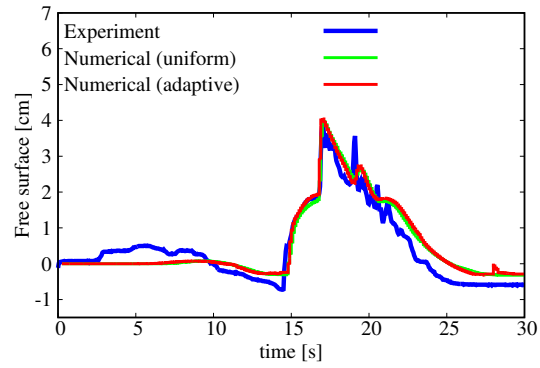
in Figure 8.12, for $t = 13.25$ s in Figure 8.13, for $t = 16$ s in Figure 8.14 and for $t = 17.5$ s in Figure 8.15).

We observe mainly that the wave front and subsequent waves are very well-captured by the criterion and the mesh refinement threshold is well chosen to refine in those areas (see for instance Figure 8.10 and Figure 8.15). Between 3 and 7 seconds, the tsunami propagates in the domain involving a grow-up of the number of cells. Then the wave passes through the island and reaches the reef around 15 seconds. After this, the wave is reflected and goes back to the wave maker with a second interaction with the island. The maximum number of cells is reached when the wave is passing on the island and begins to run-up the coast (see Figure 8.13), and when the wave goes back to the wave maker. These results are coherent with the physical process involved since it corresponds to the both times where multiple active flow regions are observed. The overall scales of the tsunami propagation with its reflections are very well captured as displayed in Figure 8.12, Figure 8.14 and Figure 8.15.

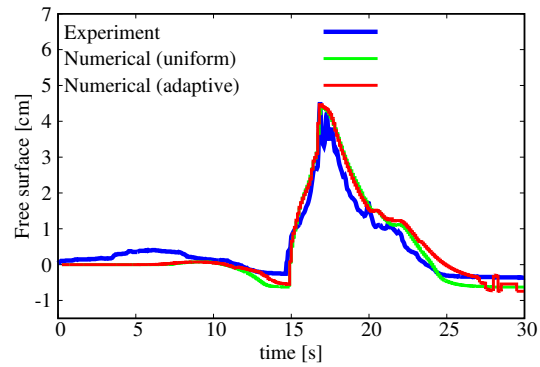
Compared to the computation on the uniform grid (see Figure 8.11), the adaptive mesh refinement method requires an average of almost 25000 cells against 62000 cells. For the same accuracy, the AMR method allows to save almost 60% of cells with respect to the uniform sim-



(a) Gauge 1



(b) Gauge 2



(c) Gauge 3

Figure 8.11: Free surface results at different positions : experimental data versus numerical simulation with and without mesh adaptivity.

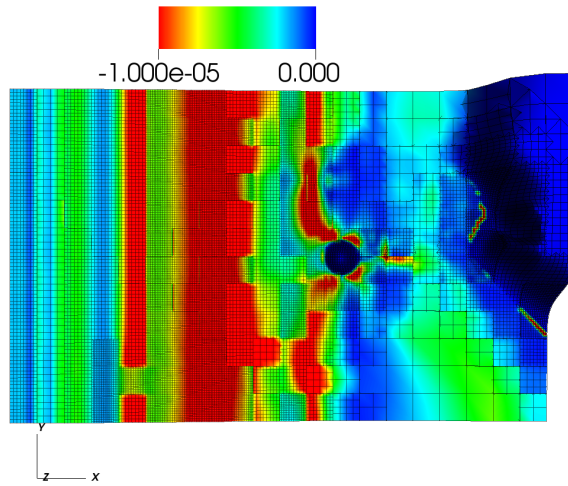


Figure 8.12: Adaptive mesh : time $t = 11.25$ s. The colours correspond to the numerical density of entropy production \mathcal{S} .

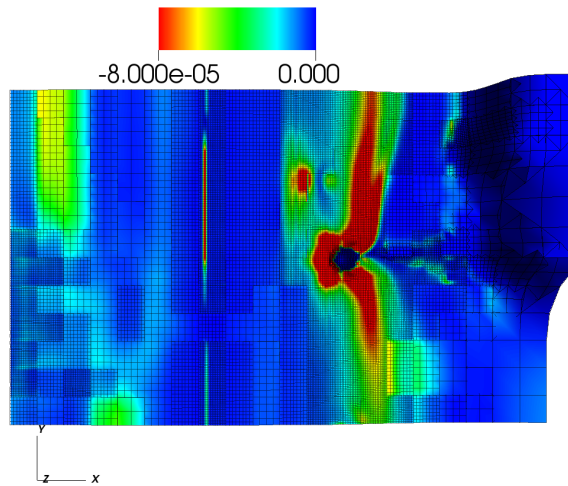


Figure 8.13: Adaptive mesh : time $t = 13.25$ s. The colours correspond to the numerical density of entropy production \mathcal{S} .

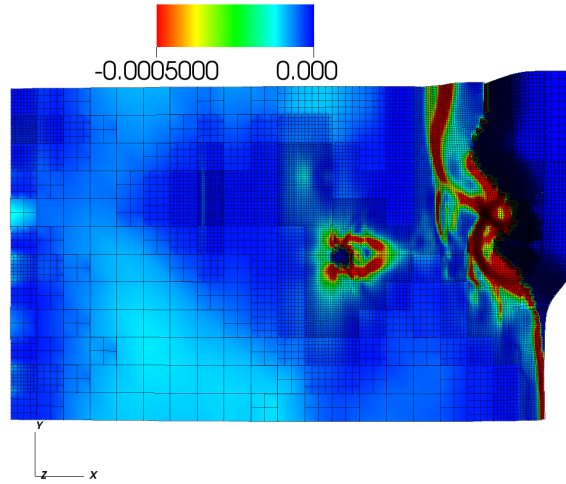


Figure 8.14: Adaptive mesh : time $t = 16$ s. The colours correspond to the numerical density of entropy production \mathcal{S} .

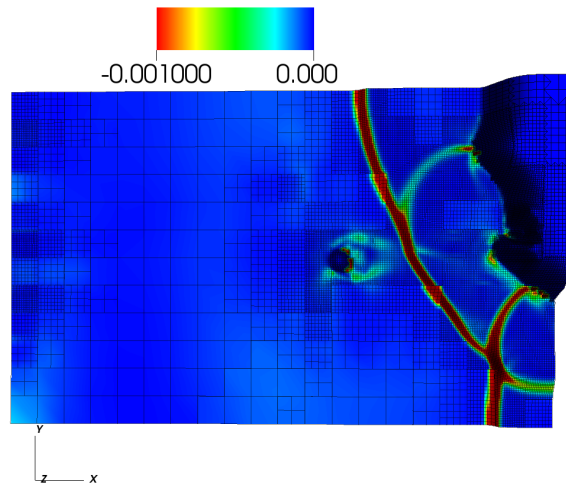


Figure 8.15: Adaptive mesh : time $t = 17.5$ s. The colours correspond to the numerical density of entropy production \mathcal{S} .

ulation. Thus, the [AMR](#) method allows to speed up the computation almost 3 times.

CONCLUDING REMARKS AND EXTENSIONS

In this work, a multidimensional finite volume solver is developed to solve non-linear hyperbolic equation system on unstructured meshes.

An [AMR](#) is introduced to improve both solution accuracy and [CPU](#) performance. It is based on a useful numerical criterion: the numerical density of entropy production. From a computational viewpoint, to make the local time step method more efficient in a parallel context, a new Block-Based [AMR](#) technique ([BB-AMR](#)) is applied. A series of 1D, 2D and 3D test cases have been performed to test and validate our approach, using several model configurations (uniform or adaptative grids, Adams-Bashforth or Runge-Kutta schemes, *etc.*). Moreover, the [BB-AMR](#) method is improved using an Automatic Mesh Refinement Threshold based on the decreasing rearrangement of the mesh refinement criterion function yielding to a robust parameterless [AMR](#) method.

From an application viewpoint, we have successfully applied and solved several problems [[74](#), [122](#), [167](#)]: dam-break, breaking waves, run-up, reflection and/or flooding, drying phenomena, *etc.*

We are now interested in the extension of the method to one-dimensional and two-dimensional dispersive equations for dispersive waves for tsunami propagation for instance.

Part IV

CFD: DG METHOD, hp -AMR AND APPLICATIONS TO WAVE-DRIVEN GROUNDWATER FLOWS

*Les vagues ne naissent pas
quand elles déferlent sur la plage.
Elles roulent longtemps en mer
avant de se casser sur un repli de sable.*

— **Geneviève Dormann**

In this part, I present the work in collaboration with J.B. Clément, F. Golay and D. Sous for wave-driven groundwater flows in sandy beaches. This work is a part of the PhD of J-B. Clément (see [35–39]).

Partially-saturated porous media are described by Richards' equation. Simulations of Richards' equation involve sharp wetting fronts which evolve both in space and in time as well as unsaturated/saturated regions of parabolic/elliptic natures which stands alongside. Getting numerical solutions for Richards' equation can be hard to achieve. In this work, a DG method is chosen since they share advantages both with finite elements methods and finite volumes methods (as developed in [Part iii](#)). It allows in principle to reach high-order accuracy schemes, h -adaptivity, p -adaptivity and more generally hp -adaptivity.

[Chapter 10](#) introduces Richards' equation to get the model problem. In [Chapter 11](#), Richards' equation discretization is presented through a DG framework. Then, several numerical experiments are discussed in [Chapter 12](#).

INTRODUCTION

This work enters into a long-term research project aiming to simulate the interactions between waves and groundwater in sandy beaches. Recent studies in laboratory [145, 155] and field studies [82, 144] provided a good experimental understanding of groundwater dynamics and shown a circulation pattern below the swash zone. Up to now, few models have been suggested and their scope remains limited [101, 107].

New numerical developments are necessary to capture a maximum of physical processes, in particular to resolve all along the wave phase the complex dynamics of saturation and pressure fields. The modelling is quite hard to achieve because it requires to solve both nonlinear equations for subsurface water flow and the shallow water equations for surface flow. Additionally, these models has to be coupled in a suitable way in order to respect the physics of the problem. Indeed, on one hand, the dynamics of incoming waves is very fast due to wave run-up and run-down which are covering and uncovering alternatively the beach. But, on the other hand, groundwater circulation is particularly slow [144] because of sediments acting as low-pass filter. Thereby, infiltrations/exfiltrations of the swash zone, sand saturation fluctuations, multiple timescales and various space-scales require a robust and accurate methodology for both numerical schemes and algorithms as well as a careful insight into the modelling.

In this work, we focus on the subsurface flow which is assumed to occur in a partially saturated porous medium and can be described by Richards equation. The latter takes into account the actions of gravity (advection) and capillarity (diffusion) but neglecting the flow of the non-wetting phase, namely the air [148]. Richards equation has been widely used to simulate water flow in unsaturated porous media [9, 66, 143]. Nevertheless, its derivation is generally rather roughly introduced, which leads us to first set out the main steps to underline some issues in regards to our model problem. Richards equation is a nonlinear parabolic equation but it degenerates into an elliptic equation under saturated conditions. Nonlinearities are due to constitutive laws and, consequently, the solution includes sharp moving fronts with dynamic

smooth regions which challenge numerical methods to obtain robust and efficient solutions of Richards equation [66]. Extensive efforts have been made to overcome these difficulties, in particular to find efficient nonlinear solvers for Richards equation [42, 66, 98, 99, 103, 110, 116, 134, 152]. Results of existence, regularity and uniqueness were found [161] and solution was assessed through diverse simulations involving different numerical methods.

Current research into Richards equation consist in advance spatial and temporal discretizations [66] allowing construction of adaptive approximations in size and/or in order for space and time. A primary focus on spatial discretization can be outlined [6, 66, 143] with the use of finite differences, finite volumes, finite elements, mixed finite elements and DG. We choose to use DG method because it offers suitable advantages for adaptivity and the treatment of flow dynamics [132]. Indeed, DG methods may be seen as a method sharing properties both from finite elements and finite volumes : they are based on a variational formulation but in an element-wise fashion. So they are locally conservative which is crucial in fluid dynamics and explains why the method is more and more used for problems in porous media. Moreover, it enables to change the degree of polynomial approximation and to use non-conforming mesh (hanging node). This is an important benefit since it is possible to handle high-order accuracy and hp -adaptation. Various forms of DG methods [3] can be used for Richards equation which can be discretized in a primal or mixed formulation. Here, we will focus on primal interior penalty DG methods.

10.1 RICHARDS EQUATION AND CONSTITUTIVE LAWS

Richards equation is a classic nonlinear parabolic equation to describe flow occurring in the unsaturated zone of an aquifer [66, 148]:

$$\partial_t (\theta (\psi)) - \nabla \cdot (\mathbb{K} (\psi) \nabla (\psi + z)) = 0 \quad (10.1)$$

where ψ denotes the pressure head (m), \mathbb{K} the hydraulic conductivity tensor (m/s), z the elevation (m) and θ the water content. Equation (10.1) is called the mixed form of Richards equation because it involves both variables, θ and ψ . This form was chosen because it is mass conservative and is applicable to a wide range of porous media [66]. Throughout this work, hydraulic head h (m) and flux q (m/s) are also used: $h = \psi + z$ and $q = -\mathbb{K} (\psi) \nabla h$. Equation (10.1) is rewritten as:

$$\partial_t (\theta (h - z)) - \nabla \cdot (\mathbb{K} (h - z) \nabla h) = 0 . \quad (10.2)$$

To close this system of equations, two constitutive laws are needed for water content and hydraulic conductivity. To describe these relations, it exists several models which depends on hydraulic properties of the porous medium. In this work, the model used for the simulations [81, 157] is:

$$(\mathbb{K}(\psi); \theta(\psi)) = \begin{cases} \left(\frac{AK_s}{A + |\psi|^B}; \frac{C(\theta_s - \theta_r)}{C + |\psi|^D} + \theta_r \right), \\ \text{if } \psi \leq 0 \text{ (unsaturated zone)} \\ (K_s; \theta_s), \\ \text{if } \psi > 0 \text{ (saturated zone)} \end{cases} \quad (10.3)$$

where A , B , C and D are empirical parameters, K_s is the hydraulic conductivity at saturation, θ_s is the saturated water content and θ_r is the residual water content.

10.2 SEEPAGE BOUNDARY CONDITION

The seepage boundary condition is specific to subsurface model [138]. It models the interface between porous medium and atmosphere. If the porous medium is saturated and outflow occurs, then water pours out at atmospheric pressure ($\psi = 0$). Otherwise, the interface acts as an impervious boundary. The difficulty is that the length of seepage face is unknown *a priori*. It can be seen as a non-linear Robin boundary condition:

$$a(h)h - b(h)\mathbb{K}(h-z)\nabla h \cdot \mathbf{n} = c(h) \quad \text{where} \quad \begin{cases} a(h) = \mathcal{H}(h) \\ b(h) = 1 - \mathcal{H}(h) \\ c(h) = \mathcal{H}(h)z \end{cases}$$

$$\text{with } \mathcal{H}(h) = \begin{cases} 1 \text{ if "seepage",} \\ 0 \text{ otherwise.} \end{cases}$$

Sometimes, the Heaviside function \mathcal{H} is regularized [50] to overcome convergence difficulties. The seepage boundary condition is also taken into account as a Signorini's type:

$$\psi \leq 0, \quad -\mathbb{K}(h-z)\nabla h \cdot \mathbf{n} \geq 0, \quad (-\mathbb{K}(h-z)\nabla(h) \cdot \mathbf{n})\psi = 0.$$

In this work, the seepage boundary condition is not regularized and is treated like in [138] by switching between Dirichlet and Neumann

boundary condition for each degree of freedom within the weak formulation of the next section. Therefore, if a seepage boundary condition lies on Γ_S , it is written:

$$\begin{cases} h = z & \text{if } \psi \geq 0 \text{ and } -\mathbb{K}(h - z) \nabla h \cdot \mathbf{n} > 0, \\ -\mathbb{K}(h - z) \nabla h \cdot \mathbf{n} = 0 & \text{otherwise.} \end{cases}$$

Paniconi and Putti [116] reported that seepage condition can cause oscillations leading to convergence problems if rapid changes of saturation state occur near the seepage face, no matter convergence requirements were relaxed or not.

DG METHOD

11.1 DISCRETIZATION

First of all, some definitions are set for the **DG** formulation. More careful developments can be found in Dolejší and Feistauer [49] or Rivière [132].

Let Ω be a space of \mathbb{R}^2 with $\partial\Omega = \Gamma_D \cup \Gamma_N$ if no seepage boundary condition is considered. From equation (10.2) and by adding initial and boundary conditions, the model problem is:

$$\left\{ \begin{array}{ll} \partial_t \theta (h - z) - \nabla \cdot (\mathbb{K} (h - z) \nabla h) = 0, & \text{in } \Omega \times [0, T] \\ h = h_D, & \text{on } \Gamma_D \times [0, T] \\ -\mathbb{K} (h - z) \nabla h \cdot \mathbf{n} = q_N, & \text{on } \Gamma_N \times [0, T] \\ h = h_0, & \text{in } \Omega \times \{0\} \end{array} \right. \quad (11.1)$$

with T the total duration and h_0 the initial condition. Ω is subdivided into elements E which form the partition denoted by \mathcal{E}_h . Elements are made of faces F . \mathcal{F}_h^I stands for the set of interior faces, \mathcal{F}_h^D for the set of Dirichlet boundary faces and \mathcal{F}_h^N for the set of Neumann boundary faces.

Let be two neighbouring elements E_l and E_r sharing one face F . There are two traces of a function u along F denoted by u_l and u_r . The jump across a face is denoted by $\llbracket u \rrbracket = u_l - u_r$, the mean by $\{u\} = \frac{1}{2} (u_l + u_r)$ and \mathbf{n} is the unit normal vector oriented from E_l to E_r .

Several **DG** methods are possible. Here, the incomplete interior penalty Galerkin (IIPG) method is chosen to avoid the addition of stabilization terms from symmetrization [49]. Penalization terms are added to enforce Dirichlet boundary condition and continuity constraint through respectively, the boundary penalty σ_b and the interface penalty σ_i . Problem (11.1) is multiplied by a test function v , integrated on each element E . After using Green's theorem, summing over all ele-

ments and performing some calculations, the space semi-discretization is obtained:

$$s > \frac{3}{2}, \text{ Find } h \in H^s(\mathcal{E}_h) \text{ such that } \forall v \in H^s(\mathcal{E}_h), \quad (11.2)$$

$$\sum_{E \in \mathcal{E}_h} \int_E \partial_t \theta(h-z)v \, dx + a_h(h, v) = l_h(v) \text{ with}$$

$$\begin{aligned} a_h(h, v) &= \sum_{E \in \mathcal{E}_h} \int_E \mathbb{K}(h-z) \nabla h \cdot \nabla v \, dx \\ &\quad - \sum_{F \in \mathcal{F}_h^I} \int_F \{ \mathbb{K}(h-z) \nabla h \cdot \mathbf{n}_F \} [[v]] \, ds \\ &\quad + \sum_{F \in \mathcal{F}_h^I} \int_F \frac{\sigma_i^F}{|F|} [[h]] [[v]] \, ds \\ &\quad - \sum_{F \in \mathcal{F}_h^D} \int_F \mathbb{K}(h-z) (\nabla h \cdot \mathbf{n}_F) v \, ds \\ &\quad + \sum_{F \in \mathcal{F}_h^D} \int_F \frac{\sigma_b^F}{|F|} h v \, ds \\ l_h(v) &= \sum_{F \in \mathcal{F}_h^D} \int_F \frac{\sigma_b^F}{|F|} h_D v \, ds - \sum_{F \in \mathcal{F}_h^N} \int_F q_N v \, ds. \end{aligned}$$

Contrary to the Finite Elements method, there is no continuity constraint between elements for the test functions. Therefore, any polynomial basis functions is suitable what makes DG method flexible. Even if a nodal basis can be chosen, this work is based on a modal basis, the monomial basis, because it is simple and can be extend easily to higher orders.

To get the full discretization, the first term of the weak formulation (11.2) is approximated by a backward Euler scheme which gives:

$$\begin{aligned} &\text{Find } h \in H^s(\mathcal{E}_h) \text{ such that } \forall v \in H^s(\mathcal{E}_h), \\ &\sum_{E \in \mathcal{E}_h} \int_E \frac{\theta(h^{n+1}-z) - \theta(h^n-z)}{\delta t} v \, dx + a_h(h^{n+1}, v) - l_h(v; t^{n+1}) = 0 \quad (11.3) \\ &\iff R(h^{n+1}, v) = 0 \end{aligned}$$

where δt is the time step and n is the number of time iteration. The choice is quite natural for backward Euler scheme because its implicit nature provides natural stability properties. Good convergence behaviour for Richards equation was also observed in [116]. Together with its simple implementation, this scheme is widespread and use of higher order scheme is questionable.

11.2 NONLINEAR SOLVER WITH ADAPTIVE TIME STEPPING

Solving equation (11.3), which represents the residual R , can be challenging because of strong nonlinearities. Newton-Raphson methods converge quadratically but they often fail for parabolic degenerate problem where the Jacobian matrix might become singular and because of poor initializations [98, 103, 116]. That's why fixed-point iteration methods are robust and popular techniques for Richards equation even if they converge linearly¹.

Let m be the number of nonlinear iteration and J is either the Jacobian matrix computed from (11.3) or is the matrix $K(h)$ coming from the variational formulation (11.3) written under a matrix system $K(h)h = f$. These two schemes reads:

$$\begin{cases} J(h^{n+1,m}) \delta h = -R(h^{n+1,m}) \\ h^{n+1,m+1} = h^{n+1,m} + \delta h \end{cases}$$

In this study, the fixed-point iteration method is chosen and constructed like in [98] by avoiding the first derivative terms inside the Jacobian matrix. The stopping criterion is $\|R(h)\|_{L^2(\Omega)} < \varepsilon$ where ε is a user-defined tolerance.

Time adaptation is motivated by the convergence of the nonlinear solver which often failed under too strong boundary and initial conditions. On one hand, transient simulations have difficulties to converge if the time step is too large but, on the other hand, shorter time steps mean more time steps and so, a longer computational time. That is why time adaptation is very attractive and common for Richards equation. However, it exists different strategies to adjust the time step [66]. First ones are heuristic and mainly based on convergence performance of the nonlinear solver. Second ones are rational and based on error control. The latter are generally much more efficient than the former. Yet, heuristic methods may outperform them in regards of their simple implementation.

For this study, the time step is adjusted heuristically according to the previous number of iterations N_{it} from the nonlinear solver such

¹ Let us note that recent progress in [168] shows that the Jacobian rank deficiency is not an obstacle anymore for the convergence of Newton-Raphson methods. Even stronger, the quadratic convergence may be preserved for rank-deficient systems.

as [11, 152]. The simulations begins with a time step δt^0 . The next time step δt^{n+1} is calculated from the previous one δt^n according to the time-stepping scheme:

$$\left\{ \begin{array}{l} \delta t^{n+1} = \begin{cases} \lambda_{\text{amp}} \delta t^n & \text{if } N_{it} \leq \inf_{lt} \\ \delta t^n & \text{if } \inf_{lt} < N_{it} \leq \sup_{lt} \\ \lambda_{\text{red}} \delta t^n & \text{if } \sup_{lt} < N_{it} \leq \max_{lt} \end{cases} \\ \delta t^n = \lambda_{\text{red}} \delta t^n \text{ if } \max_{lt} < N_{it} \\ \text{or if the solver has failed (time step is started again)} \end{array} \right.$$

The factors $0 < \lambda_{\text{red}} < 1$ and $1 < \lambda_{\text{amp}}$ as well as the threshold values $0 < \inf_{lt} \leq \sup_{lt} < \max_{lt}$ are prescribed by the user. A minimum time step is defined to avoid excessive small time step. With this approach, the nonlinear solver is more robust and computational time is saved compared to a fixed time step. Indeed, whatever δt^0 , the time step is adjusted until the nonlinear solver succeeds to converge. Then, thanks to the coefficients, the time step is adjusted smoothly. Nevertheless, the method depends on many empirical parameters and does not provide an optimal time step because the amplification/reduction coefficients restrict the possible size of the new time step.

11.3 ADAPTIVE MESH REFINEMENT

Solution of Richards equation holds sharp wetting fronts which are moving dynamically through the domain among smooth regions[110]. They arise typically from boundary conditions and dissipate over time by coming across the saturated zone. Mesh adaptation is a technique which can handle these particularities to reach the desired accuracy [50, 99, 110]. Recent strategies involve both size and order adaptation, based on *a posteriori* errors [99] or interpolation error [50].

In this study, an *h*-adaptation based on a simple *a posteriori* error is performed like [110]. Yet, whereas they use changes in effective saturation [110], that's directly changes in the solution, the hydraulic head h , which serve as error indicator. The latter has the advantage of working both for the saturated and unsaturated regions.

Mesh adaptation is a 2D block-based technique adapted from [1] and Part iii. The process of mesh adaptation begins with the evaluation of the criterion defined by:

$$C_E = \frac{1}{\max_{E \in \mathcal{E}_h} \left(\frac{1}{|E|} \|\nabla h\|_{L^2(E)} \right)} \frac{1}{|E|} \|\nabla h\|_{L^2(E)}$$

The criterion is normalized so that $0 \leq C_E \leq 1$. The second step consist in deciding to refine, coarsen or do nothing thanks to user-defined threshold values $0 < \beta_{min} \leq \beta_{max} < 1$ for the criterion. Elements are merged or split according to the scheme described in Figure 11.1. To avoid excessive refinement and coarsening, minimum and maximum levels are set up. Then, the mesh is smoothed by checking that each neighbouring block of elements cannot have a difference of refinement level superior to one as done in Part iii.

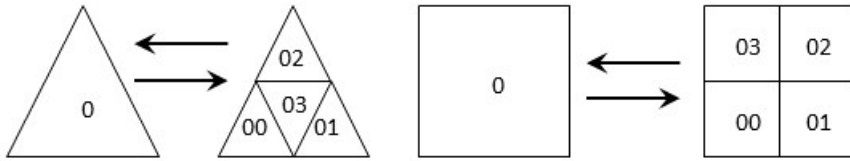


Figure 11.1: Refinement and coarsening process for one triangular and quadrangular element.

In the last step, the solution is projected from the previous mesh to the new one by solving a local DG problem. The element-wise weak formulation corresponding to $h_{E_{new}} = h_{E_{old}}$ reads:

$$\int_{E_{new}} h_{E_{new}} v \, dx = \int_{E_{old}} h_{E_{old}} v \, dx$$

This strategy of adaptation is very simple but it uses an *a posteriori* indicator and requires empiric parameters which balance computational performance and accuracy. This forces to calibrate each numerical simulation.

NUMERICAL RESULTS

12.1 ONE-DIMENSIONAL INFILTRATION NUMERICAL TEST CASE

This numerical test-case was firstly introduced by [31] on the basis of an experiment from [81]. It was taken afterwards by others [108] and corresponds to a class of popular test cases, for example, see [110].

This problem deals with an 1D-infiltration into a soil column. \mathbb{K} and θ are taken from (10.3) with $A = 1.175 \times 10^6$, $B = 4.74$, $C = 1.611 \times 10^6$, $D = 3.96$, $K_s = 0.00944 \text{ cm/s}$, $\theta_s = 0.287$ and $\theta_r = 0.075$. The initial condition is $h_0 = -61.5 + z \text{ cm}$. On the top and bottom, hydraulic head is kept fixed: $h_{top} = 19.3 \text{ cm}$, $h_{bottom} = -61.5 \text{ cm}$. Sides are impervious. Initial condition is: $h = -61.5 + z \text{ cm}$.

The solutions are in very good agreement to one another as seen for pressure head in the Figure 12.1a. Moreover, h -adaptation is capturing the wetting front efficiently as it is shown in Figure 12.1b.

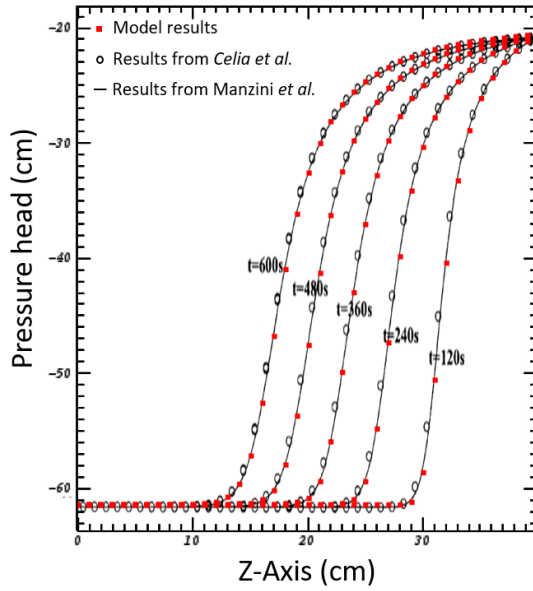
12.2 COMPARISON WITH WATER TABLE DRAINAGE EXPERIMENT

Vauclin *et al.* made a drainage experiment of a slab of soil in which experimental details and positions of water table can be found [157].

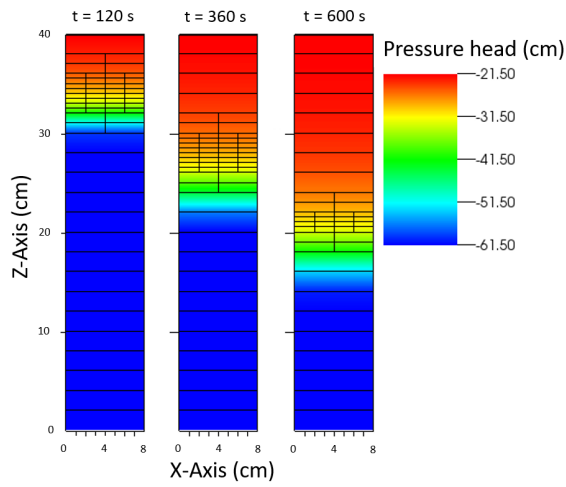
A water table depletion occurs after an instantaneous drop of the imposed hydraulic head. \mathbb{K} and θ are taken from (10.3) with $A = 359720.49$, $B = 4.5$, $C = 40000$, $D = 2.9$, $K_s = 40 \text{ cm/h}$, $\theta_s = 0.3$ and $\theta_r = 0$. This experiment is used as a test-case to show a transient simulation with seepage boundary condition. The domain is initially in hydrostatic equilibrium: $h_0 = 145 \text{ cm}$. Boundary conditions on the right, at the bottom, on the top and on the left for $z > 145 \text{ cm}$ are $\nabla h \cdot \mathbf{n} = 0$. On the left, the level is kept constant in the ditch for $z \leq 75 \text{ cm}$, $h = 75 \text{ cm}$, and, the seepage boundary condition is prescribed for $45 \text{ cm} < z \leq 145 \text{ cm}$.

Numerical simulation is performed with linear approximation $p = 1$, $\sigma_i = \sigma_b = 100$ on a mesh with $\Delta x = 10 \text{ cm}$ and $\Delta z = 5 \text{ cm}$. Water table location matches $\psi = 0$. Numerical positions of water table are

displayed in Figure 12.2 together with the experimental data from [157].



(a) Pressure head profiles adapted from [108]



(b) Adaptive mesh refinement around the wetting front

Figure 12.1: Results for the one-dimensional infiltration numerical test case.

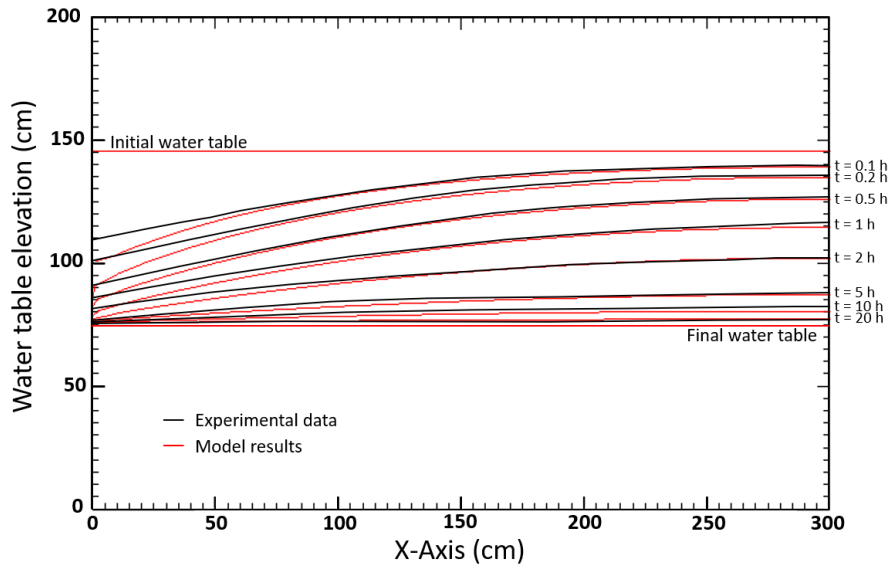


Figure 12.2: Numerical simulation of two-dimensional drainage with water table position compared to experimental data from [157].

Results presented in the Figure 12.2 show that the model is able to reproduce satisfactory the transient water table dynamics. Disagreements near the seepage boundary are found between numerical and experimental profiles. Three reasons can explain it. One is the lack of spatial resolution. Refining the grid would improve the computed solution and catch better the nonlinear seepage boundary condition. Another reason is in relation with the experimental set-up because water should not pour out completely freely through the experimental device in such a way that seepage boundary condition does not model it properly. Finally, the determination of hydraulic properties was done by correlation from measurements with some fairly important dispersion, due to natural heterogeneity, neglecting hysteresis [157]. This can explain some discrepancies. Other numerical studies based on this experiment are reported [40, 152]. They also struggle to match water table profiles near the seepage.

12.3 IDEALIZED BEACH TEST CASE

In this test case, the Simulation Waves till SHores (*SWASH*) code is used to simulate waves in coastal waters and to produce data which enforce Richards equation through suitable boundary conditions. This weak coupling can take into account seepage along the beach face. This idealized case is constructed to investigate the beach groundwater

response to the action of swash: infiltration/exfiltration, water table position, hydraulic head and saturation distribution as well as pore velocities are outlined.

12.3.1 Description of the idealized beach case

Figure 12.3 shows a schematic profile of the beach considered. Y-axis was multiplied by 20 to make it more visible. It's a straight slope of ratio 1:35 which spans over 105 meters. Hydraulic properties for constitutive laws of Equation (11.1) are written in the Table 12.1.

| Parameters | Hydraulic conductivity | Water content |
|---------------------|------------------------|---------------|
| A/C | 2.99×10^6 | 40 000 |
| B/D | 5.0 | 2.90 |
| Value at saturation | 35.0 | 0.30 |

Table 12.1: Parameters of constitutive laws for the idealized beach case [156].

Initial conditions are hydrostatic with water table level at same height of surface water: $h_0 = 1.5$ m. The bottom is an impervious boundary. The water table and head are kept fixed on the inland boundary: $h = 1.5$ m. The beach face is monitored by the SWASH code. In SWASH, a sinusoidal time-series of free surface elevation is imposed at the offshore boundary to simulate a long infragravity-like wave of period 50 s and amplitude 0.3 m. These types of waves have been observed to be a major driver of beach groundwater dynamics of the Rousty microtidal dissipative beach [144]. This results in two swash events on the beachface. Time simulated lasts 120 s.

Two waves are generated by SWASH and their characteristics are: wavelength, frequency. This results in two swash events on the beach face. Time simulated lasts 7.2 s. The mesh used for the simulation is displayed on the Figure 12.3.

12.3.2 Results

An overview of groundwater features during a swash cycle is presented. Then, results focus on uprush and backwash events. Some selected snapshots from the simulation show interesting observations for the idealized beach case. Surface water and water table are depicted

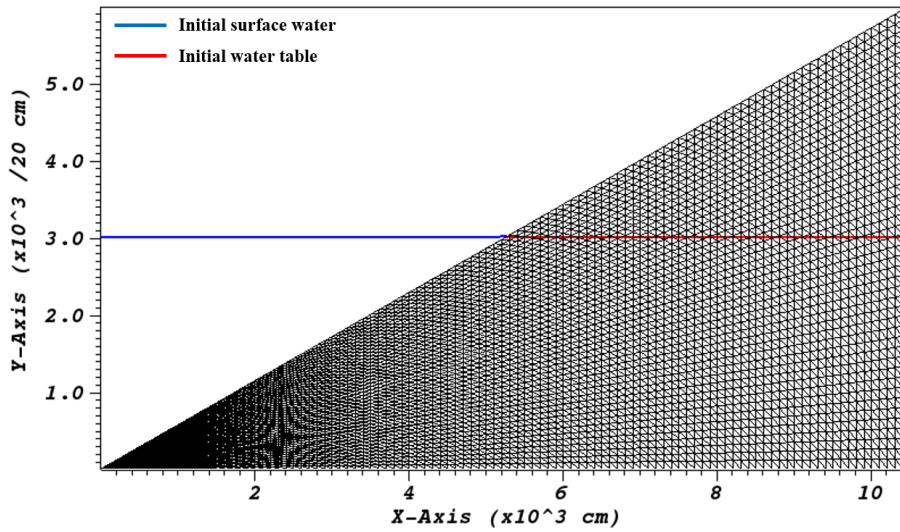


Figure 12.3: Topography of the idealized beach and mesh used for the simulation.

by blue and white solid lines respectively. Hydraulic head contours are displayed together with velocities vectors.

OVERALL PATTERN FOR SWASH CYCLE:

A first observation concerns wave propagation. When a wave is moving, groundwater flows in the saturated zone respond immediately. Hydraulic head spreads vertically quite sharply under the wave front. Infiltration/exfiltration occurs at the back/front of the bore respectively. It is particularly visible in the [Figure 12.4](#) even if this phase is completely ideal because it's about the first wave moving through hydrostatic conditions.

Throughout the simulation, a nearly hydrostatic state is observed in the upper beach area. Moreover, flow divides into seaward and landward components near the uprush limit. Velocities magnitude is rather small everywhere, about 1 cm/h, except for uprush infiltration at the top where it reaches up to 160 cm/h. Around exit point of water table, a local overpressure is observed. It is probably caused by numerical artefacts due to the lack of mesh resolution for the seepage boundary condition. An overall pattern can be depicted for swash cycles. Two recirculation zones are attached respectively to the back and the front of the swash tongue. These two groundwater cells are interacting together and are evolving during uprush and backwash,

growing or diminishing. Sometimes one vanishes completely, sometimes another one gets trapped by recharge induced by water table elevation.

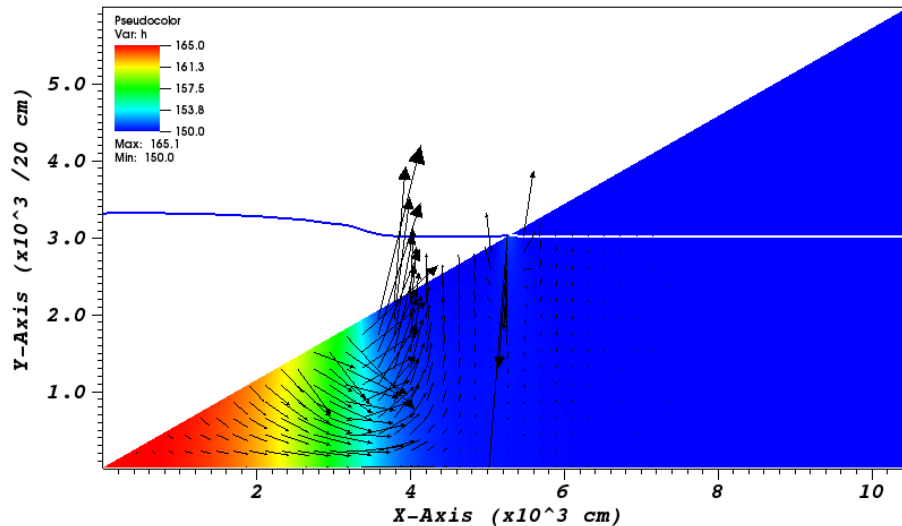


Figure 12.4: Snapshot of hydraulic head contours at $t = 12.5s$ during the first wave propagation on calm waters.

PATTERN DURING UPRUSH EVENT:

During uprush, there is downward infiltration while the swash tongue is covering the beach. It is mainly concentrated in the upper zone, through the boundary in contact with the fully saturated part of the beach. Infiltration induces a rise of the water table near the exit point, with, as seen in [Figure 12.5](#).

PATTERN DURING BACKWASH EVENT:

During backwash, a low hydraulic head zone emerges at the base of the swash zone where exfiltration occurs. Velocities are mainly seaward under the swash tongue moving back. Meanwhile, a small region of high hydraulic head persists in the corner of the saturated zone where water table have been raised by the precedent uprush. It results in recharge and, as water is percolating inside the porous medium, flow divides between seaward and landward direction. Moreover, above the unsaturated zone which has been covered by uprush, water is infiltrating downward in a sharp wetting front. While the swash toe retreating to the sea, a seepage face is appearing which drops, locally

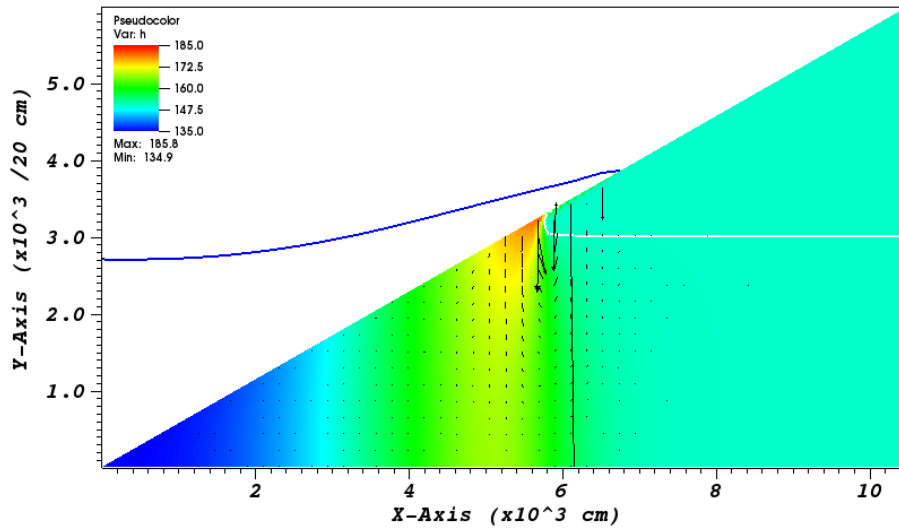


Figure 12.5: Snapshot of hydraulic head contours at $t = 37.5$ s for the end of the first uprush. Velocities range from about 0.3 to 0.6 cm/h in the beach. Maximum velocity is 35 cm/h for infiltration.

but quickly, the water table what gives form to a small hump-like structure. [Figure 12.6](#) presents these observations.

PATTERN BETWEEN BACKWASH AND UPRUSH EVENTS:

The next incoming wave arrives just after the backwash. It pushes the region of low hydraulic head up. As this zone swept away, it reaches the remaining small region of high hydraulic head induced by the water table elevation and makes it disappear temporarily. The incoming wave is bringing a new high hydraulic head front. This causes infiltration which arises water table again, ending seepage and erasing the hump-like structure of water table. Then, a new cycle begins. This intermediate state is shown in [Figure 12.7](#).

PATTERN BETWEEN BACKWASH AND UPRUSH EVENTS:

The next incoming wave arrives just after the backwash. It pushes the low hydraulic head region up. As it swept away, the zone concentrates more and more and reaches the remaining small high hydraulic head region around hump-like structure to make it disappear. The wave which is coming goes with a new high hydraulic head front. This causes infiltration which arises again water table, ending seepage and erasing the hum-like structure. The situation is shown in [Figure 12.7](#).

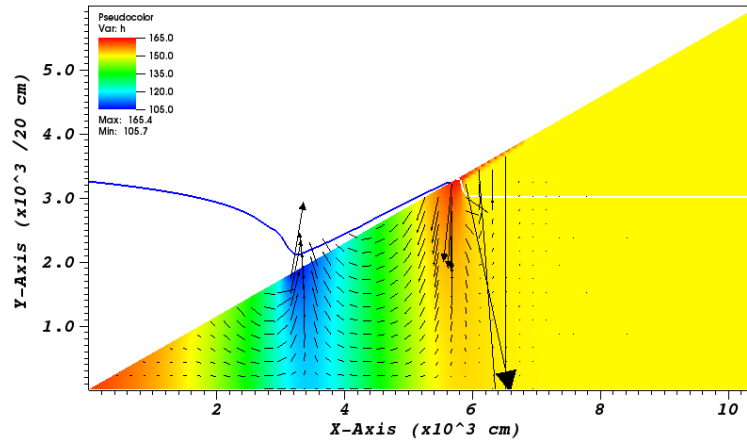


Figure 12.6: Snapshot of hydraulic head contours at $t = 57.5$ s for the end of the first backwash. Most velocities are less than 1 cm/h.

12.3.3 Discussion

In the simulation, the offshore water level is similar to the inland water table, therefore no steady pressure gradient is observed across the beach-face by contrast to field or laboratory studies where cross-beach gradients can be observed depending on boundary conditions [144, 155]. Within the beach, overall velocity magnitude is about 1 cm/h which is fairly comparable to measurements made in [144]. The only exception is for infiltration/exfiltration where maximum velocity order is about 100 cm/h. This seems to be less compared with the experiments of [155]. Further tests are now engaged in the exact configurations of the laboratory [155] and field [144] experiments to allow direct comparisons. As described by [119], the moving bore is carrying a sharp hydraulic front. Exfiltration occurs at its front and infiltration at its back when it meets the groundwater conditions of the precedent swash. A local super-elevation of the water table has also been revealed in these observational studies due to the water table charge by short or IG wave forcing under the swash zone. [119] showed a domed water table in their study. The present benchmark case will be extended to a longer duration to test if the model is able to represent such feature. Water table elevation induces charge in this zone as long as the low-pressure zone does not affect it with the ongoing uprush. As a consequence, infiltration with flow division occurs near the uprush limit. Laboratory experiments and numerical simulations [100, 155] both predict this feature but [144] do not observe it. During

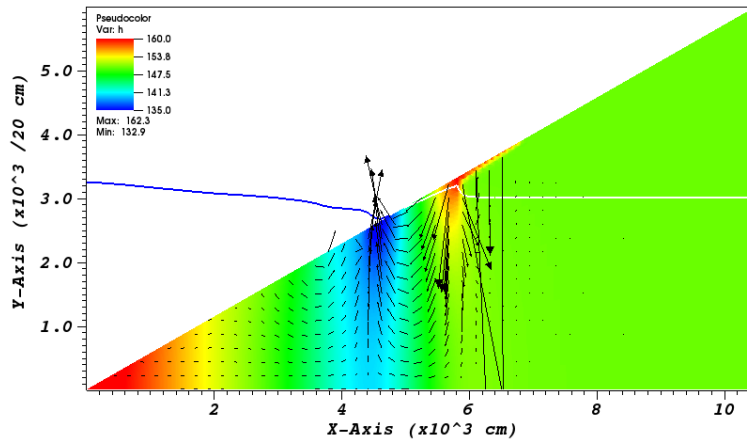


Figure 12.7: Snapshot of hydraulic head contours at $t = 67.5$ s for the next incoming uprush. Most velocities are less than 0.8 cm/h.

the simulation, seepage face has small length because water is able to pour out from sand quite fast following surface water. As a consequence, a small hump-like structure appears during backwash because of seepage along the boundary. This last feature is not described in any study found, but further tests are ongoing with more realistic configurations and longer wave forcings to better understand involved processes. The two circulation zones with infiltration/exfiltration in the upper/lower part of the beach are also identified in the simulation from [100] and experiments from [155]. Infiltration and exfiltration involve greater pore velocities than in other parts of the beach. Like it is depicted by [119], gradients brought by the wave are much larger than the average values in the beach due to larger differences in water depth across the wave.

CONCLUDING REMARKS AND EXTENSIONS

In this work, we have developed a full numerical code on unstructured meshes for Richards equation through a [DG](#).

This numerical code, associated with a backward Euler scheme, is able to simulate variably saturated flows in porous media for a wide range of problems steady and unsteady problems, including seepage boundary condition. Richards equation gives rise to some issues concerning non-linearities and convergence robustness. In order to provide a computationally simple but efficient algorithm in terms of solution accuracy and robustness, the nonlinear solver is based on a fixed-point iteration method with adaptive time stepping. Besides, an adaptive mesh refinement technique is employed to achieve accuracy and deal with steep gradients. Through several test cases ([\[35, 36, 39\]](#)) numerical results present the functioning of the method and its potential.

Further research is underway to extend adaptive mesh refinement to an efficient hp adaptation, assess more carefully the convergence behaviour of seepage boundary condition and make time step adaptation less heuristic.

Part V

DIMENSION REDUCTION: FROM HYDROSTATIC TO NON HYDROSTATIC MODELS

*L'art de réduire n'est pas une dépréciation
mais une volonté de compréhension.*

— **Mehmet Ersoy**

In this part, I focus on the model reduction in the shallow water limit for compressible and incompressible flows with hydrostatic and non-hydrostatic pressure distribution. Opposite to the [CFD](#) viewpoint (see [Part iii](#) and [Part iv](#)), the idea in this part is to reduce the dimension of the initial model preserving some of the main physical and geometrical features. This is one of the most challenging issues that we address with the obvious consequence to decrease the computational time.

In this part, in collaboration with M.A. Debyaoui, we present the first non-linear and weakly dispersive extension of the [FS-model](#) [\[54\]](#) (or the Serre Green Naghdi ([SGN](#)) equations) for natural rivers and channel hydraulics ([\[47\]](#)). This work is a part of the PhD of M.A. Debyaoui (see also [\[46, 47\]](#)).

In [Chapter 14](#), we first motivate our work. Then, in [Chapter 15](#), we present the geometrical settings and the Euler equations. [Chapter 16](#) is devoted to the asymptotic limit of the Euler equations. In [Chapter 17](#), we present the new model for arbitrary non rectangular channel/river. Finally, the rectangular case is developed in [Chapter 18](#) and a finite numerical scheme is proposed.

I also refer to [\[56, 57\]](#) for a rigorous model reduction including friction and to [\[61\]](#) for overland flows including rainfall and infiltration (or exfiltration). These works are not presented here.

INTRODUCTION

The modelling of hydrology of catchment basins and rivers holds a central place in environmental sciences, particularly in connection with water availability, urban sewer systems, flood risks and in particular for tsunamis. Indeed, rivers are known to be the tsunami highways. Waves penetrate through rivers much faster inland than the coastal inundation reaches over the ground, and may lead flooding in low-lying areas located several km away from the coastline [154]. This is important today in understanding and forecasting the impact of climate variability on the human and natural environment. Modelling these processes and predicting the motion of water is a difficult task for which substantial effort has been devoted [64, 76, 77, 133, 142, 160, 164, 170].

One of the most widely used models to describe the channel and river motion of watercourses is the *section-averaged free surface model* [19, 48, 56] which is a generalisation of the well-known *Saint-Venant system* (introduced by Adhémar Jean Claude Barré de Saint-Venant in the 19th Century [70, 135]):

$$\begin{cases} \partial_t A + \partial_x Q & = 0, \\ \partial_t Q + \partial_x \left(\frac{Q^2}{A} + I_1(x, A) \right) & = I_2(x, A). \end{cases} \quad (14.1)$$

In these equations, A is the wet area of fluid cross-section, Q is the water discharge, I_1 is the hydrostatic pressure and I_2 is the hydrostatic pressure source term which takes in account of the variation of the section. The model (14.1) reduces to the well-known one-dimensional Saint-Venant equations for uniform rectangular section. The free surface model is the first order shallow water approximation of the section-averaged Navier-Stokes or Euler equations under suitable assumptions on the horizontal and the vertical scales (see, *e. g.*, [19, 48, 56, 70, 75] and the reference therein).

Thanks to the hyperbolic structure of these equations, sharp transitions between two different flow states result in a discontinuous solution, both in the water surface and in the velocity. These discontinuous solutions (called shocks and also referred as bores) are well-suited to approximate breaking waves with turbulent rollers for large transitions

of the Froude's number. However, for small or moderate transitions, the advancing wave front can be followed by a train of free-surface undulations, sometimes called "whelps". This phenomenon, called undular bore (also often called dispersive shock waves), is induced by a non hydrostatic pressure distribution [118]. As a consequence, wave solutions spread out in space as they evolve in time, *i. e.* waves of different wavelengths travel with different speeds. This is the so-called dispersive effect. Consequently, undular bores are not reproducible with the non-dispersive free surface system and non hydrostatic pressure is required.

Dispersive equations were first introduced by Boussinesq [23] in 1872 to mathematically justify the existence of solitary waves observed by Russell's experiments in 1834. These equations enters in the framework of shallow water equations. They can be obtained as the second order asymptotic approximation in μ^2 , with $\mu = (\frac{H}{L})^2$, of the depth-averaged Euler equations where H represents a characteristic water depth and L is a characteristic horizontal scale.

The Boussinesq type equations, which are weakly non-linear and weakly dispersive, are defined by an additional small non-linearity parameter $\varepsilon = \frac{a}{H} = O(\mu) \ll 1$ where a is the order of the free surface amplitude. These equations for a flat bottom, for instance in 1d, are given by

$$\begin{cases} \partial_t \zeta + \partial_x(h\bar{u}) & = O(\mu^2) \\ \partial_t \bar{u} + \varepsilon \bar{u} \partial_x \bar{u} + \nabla \zeta + \mu \mathbb{D} & = O(\mu^2) \end{cases} \quad (14.2)$$

where h is the water depth, ζ is the free surface elevation and \bar{u} is the depth-averaged velocity. The term \mathbb{D} represents the dispersive term. In 1877, the KdV equation was discovered by Boussinesq [23] and was later derived by Korteweg and Gustav de Vries (KdV) [90]. This equation approximately describes the evolution of long, one-dimensional waves in many physical settings, including shallow-water waves with weakly non-linear behaviour. This equation can also be obtained in the case of unidirectional waves for which the theoretical framework allows to compute an analytical solution such as 1d solitary wave propagation on flat bottom (see e.g. Boussinesq [8, 23, 90]). We have also the Benjamin, Bona, and Mahony (BBM) equation which is an improvement of the KdV equation [10]. However, in practice and especially in coastal engineering applications, the nearshore wave dynamics being often varying-bottom dependent, dispersive and non-linear, the Boussinesq type equations are not appropriate.

In 1967, Peregrine [117] introduced the first weakly non-linear two dimensional Boussinesq type equations for non flat bottom. Witting [163] proposed a method based on Padé expansion to improve the frequency dispersion of the Boussinesq-type equations. From this method, several equations of order $O(\mu^2)$ with improved dispersion characteristic have been proposed, see for instance [106, 114, 136].

In 1953, a one-dimensional, fully non-linear ($\varepsilon = O(1)$) and weakly dispersive equations for flat bottom were derived by Serre [140], independently of Su and Gardner [147] with

$$\mathbb{D}(v) = \frac{1}{h} \partial_x \left(\frac{h^3}{3} \mathcal{D}(v) \right) \quad (14.3)$$

where

$$\mathcal{D}(v) = (\partial_x v)^2 - \partial_x \partial_t v - v \partial_{xx} v. \quad (14.4)$$

This model was then extended to non flat bottom by Seabra-Santos *et al* [139], see also [33]. Finally, Green and Naghdi [78] derived the two dimensional fully nonlinear dispersive equations for uneven bottom which are the extension of Serre equations. In the literature, this system is often called **SGN** equations. Further interesting extensions are also presented in [34, 95, 96, 120, 131, 159] based on either Euler equations, water waves equations or variational principles. All of the previous dispersive models are obtained either from 3d-2d or 2d-1d reduction, but, up to our knowledge, the 3d-1d reduction has never been done before. Thus, our main goal is to derive from the three-dimensional incompressible and irrotational Euler equations with suitable boundary conditions, a model akin to (14.2) – (14.4) via section averaging under the shallow water assumption. The section averaged model that we obtain extends the section-averaged free surface model (14.1) and the **SGN** equations (14.2) – (14.4):

$$\begin{cases} \partial_t A + \partial_x Q = 0 \\ \partial_t Q + \partial_x \left(\frac{Q^2}{A} + I_1(x, A) \right) + \mu_2 \partial_x (G(x, A) \mathcal{D}(u)) = \\ I_2(x, A) + \mathcal{G}(x, A, Q) + O(\mu_2^2) \end{cases}$$

where A is the wet area, $Q = Au$ is the water discharge, u is the section-averaged velocity, I_1 (resp. I_2) is the hydrostatic pressure (resp. source) term, $G(x, A)$ generalises $\frac{h^3}{3}$ in (14.3), the term $\mathcal{G}(x, A, Q)$ extends the uneven bottom source term, \mathcal{D} is the term given by (14.4) and therefore $\mathbb{D} = \partial_x (G(x, A) \mathcal{D}(u))$ represents the dispersive term.

THE THREE-DIMENSIONAL INCOMPRESSIBLE EULER EQUATIONS

We start in [Section 15.1](#) by reviewing the irrotational and incompressible Euler equations in the special geometric setting, describing the physics with a wet boundary on the bottom of the water course and a free surface on the top. Boundary conditions are presented in [Section 15.2](#).

15.1 GEOMETRIC SET-UP AND THE EULER EQUATIONS

Let $T > 0$ be an arbitrary time. We consider an incompressible and irrotational fluid moving in the time-space box $[0, T] \times \mathcal{C}$ with typical point $(t, (x, y, z))$ where \mathcal{C} is the geometrical definition of a convex (non-rectangular) channel/river

$$\mathcal{C} = \left\{ (x, y, z) \in \mathbb{R}^3; x \in [0, L_c], \alpha(x, z) \leq y \leq \beta(x, z) \right. \\ \left. \text{and } d(x, y) \leq z \right\} .$$

$L_c > 0$ is the horizontal length of the domain and for $z \geq d(x, y)$, $\alpha(x, z)$ (respectively $\beta(x, z)$) is the left (respectively right) boundary point at the elevation z as displayed in [Figure 15.1](#). The height of the surface of the water level and the boundary of the section are modelled, respectively, by the functions $\eta(t, x, y)$ and $d(x, y)$ with respect to a reference horizontal height $z = 0$. For all $x \in [0, L_c]$, we assume that the function $y \mapsto d(x, y)$, has a global minimum at

$$y^*(x) \text{ i.e. } d^*(x) = d(x, y^*(x)) = \min_y d(x, y)$$

where y^* describes the transversal variation of the channel with respect to the main channel/river direction.

We define the local height of the water by

$$h(t, x, y) := \eta(t, x, y) - d(x, y).$$

The boundary of the section can be also described as a function of y by

$$\forall x \in [0, L_c], \varphi(x, z) = \begin{cases} \alpha(x, z) & \text{if } \varphi(x, z) < y^*(x) \\ \beta(x, z) & \text{otherwise.} \end{cases} \quad (15.1)$$

The width of the section at the elevation z is given by an increasing function

$$z \mapsto \sigma(x, z) = \beta(x, z) - \alpha(x, z) .$$

The wet region is defined as the region in which the fluid resides at each time $t \in [0, T]$

$$\Omega(t) = \bigcup_{0 \leq x \leq L} \Omega(t, x)$$

with its global counterpart

$$\Omega := \bigcup_{0 \leq t \leq T} \Omega(t) . \quad (15.2)$$

$\Omega(t, x)$ is the cross-section of fluid at the position x of the channel \mathcal{C} :

$$\Omega(t, x) = \{ (y, z) \in \mathbb{R}^2; \alpha(x, z) \leq y \leq \beta(x, z) , \quad d(x, y) \leq z \leq \eta(t, x, y) \} \quad (15.3)$$

and $A(t, x) = |\Omega(t, x)|$ is the wet area of the cross-section of fluid:

$$\begin{aligned} A(t, x) &= \int_{\Omega(t, x)} d\omega \\ &= \int_{y^-(t, x)}^{y^+(t, x)} \eta(t, x, y) - d(x, y) dy \\ &= \int_{d^*(x)}^{\max_y \eta(t, x, y)} \int_{\alpha(x, z)}^{\beta(x, z)} \mathbb{1}_{\{d(x, y) \leq z \leq \eta(t, x, y)\}} dy dz \end{aligned} \quad (15.4)$$

where

$$y^-(t, x) := \min\{y \in \mathbb{R}; \eta(t, x, y) = d(x, y)\}$$

and

$$y^+(t, x) := \max\{y \in \mathbb{R}; \eta(t, x, y) = d(x, y)\}$$

stands for the left and the right transversal limit at $z = \eta$ as shown in [Figure 15.1](#).

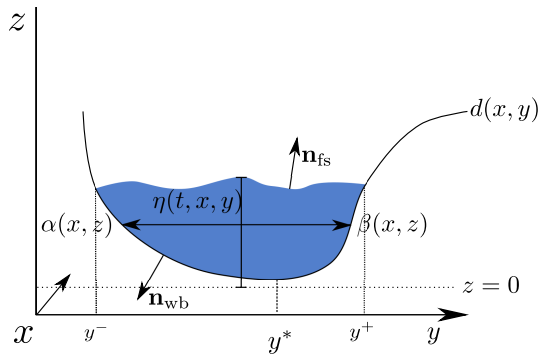


Figure 15.1: Geometric set-up.

We assume that the flow is governed, on the space-time domain Ω , by the three-dimensional irrotational and incompressible Euler equations

$$\begin{aligned} \operatorname{div} \rho_0 \mathbf{u} &= 0, \\ \frac{\partial}{\partial t}(\rho_0 \mathbf{u}) + \operatorname{div} \rho_0 \mathbf{u} \otimes \mathbf{u} + \nabla p - \rho_0 \mathbf{F} &= 0 \end{aligned} \quad (15.5)$$

where $\mathbf{u} = (u, v, w)$ is the velocity field, ρ_0 is the density of the fluid (taken to be constant), $\mathbf{F} = (0, 0, -g)$ is the external force of gravity with constant g and p is the pressure.

These equations are completed by the irrotational equation:

$$\operatorname{curl} [\mathbf{u}] = 0. \quad (15.6)$$

15.2 BOUNDARY CONDITIONS

For a given time $t \in [0, T]$ and $x \in [0, L_c]$, the boundary of the fluid cross-section $\Omega(t, x)$ is composed of a free surface $\Gamma_{\text{fs}}(t, x) = \{(y, z) \in \mathbb{R}^2; z = \eta(t, x, y)\}$ and a wet boundary (part of the boundary in contact with water) $\Gamma_{\text{wb}}(x) = \{(y, z) \in \mathbb{R}^2; z = d(x, y)\}$ such that

$$\partial\Omega(t, x) = \Gamma_{\text{fs}}(t, x) \cup \Gamma_{\text{wb}}(x).$$

We prescribe a kinematic boundary condition at the free surface boundary and a no-penetration condition at the wet boundary as described below.

15.2.0.1 Free surface boundary conditions

Assuming a kinematic boundary condition, we set for all points $\mathbf{m}(t, x, y) = (x, y, \eta(t, x, y)) \in \Gamma_{\text{fs}}(t, x)$,

$$\mathbf{u} \cdot \mathbf{n}_{\text{fs}} = \partial_t \mathbf{m} \cdot \mathbf{n}_{\text{fs}}$$

where

$$\mathbf{n}_{\text{fs}}(t, x, y) = \frac{1}{\sqrt{1 + (\partial_x \eta(t, x, y))^2 + (\partial_y \eta(t, x, y))^2}} \begin{pmatrix} -\partial_x \eta(t, x, y) \\ -\partial_y \eta(t, x, y) \\ 1 \end{pmatrix}$$

is the unit outward normal vector to the free surface. This leads to the following explicit form of the kinematic boundary condition

$$\partial_t \eta + u \partial_x \eta + v \partial_y \eta = w \text{ on } z = \eta. \quad (15.7)$$

We also assume that the pressure at the free surface level is equal to the atmospheric pressure p_0

$$p = p_0 \text{ on } z = \eta . \quad (15.8)$$

In the sequel, without loss of generality, we set $p_0 = 0$.

15.2.1 Wet boundary conditions

On the wet boundary, *i. e.*, the part of the boundary in contact with water, prescribing a no-penetration condition, we set for all points $\mathbf{m}(x, y) = (x, y, d(x, y)) \in \Gamma_{\text{wb}}(x)$,

$$\mathbf{u} \cdot \mathbf{n}_{\text{wb}} = 0$$

which leads to the following explicit form:

$$u\partial_x d + v\partial_y d = w \text{ on } z = d \quad (15.9)$$

where

$$\mathbf{n}_{\text{wb}}(x, y) = \frac{1}{\sqrt{1 + (\partial_x d(x, y))^2 + (\partial_y d(x, y))^2}} \begin{pmatrix} \partial_x d(x, y) \\ \partial_y d(x, y) \\ -1 \end{pmatrix}$$

is the unit outward normal vector to the wet boundary. In view of the definition of the function φ , see Equation (15.1), the no-penetration condition can be also expressed as a function of (x, z) by

$$u\partial_x \varphi + w\partial_z \varphi = v \text{ on } y = \varphi \quad (15.10)$$

where the unit outward normal vector to the wet boundary is

$$\mathbf{n}_{\text{wb}}(x, z) = \frac{1}{\sqrt{1 + (\partial_x \varphi(x, z))^2 + (\partial_z \varphi(x, z))^2}} \begin{pmatrix} \partial_x \varphi(x, z) \\ -1 \\ \partial_z \varphi(x, z) \end{pmatrix} .$$

In this work, we neglect some physical process arising in river flows: sedimentation, exchange between groundwater flows and subsurface flows, porosities, *etc.* However, these phenomena can be easily integrated to this work, see for instance [55, 61] and Part iv, by considering the following boundary condition $\mathbf{u} \cdot \mathbf{n}_{\text{wb}} = (\partial_t \mathbf{m} + I) \cdot \mathbf{n}_{\text{wb}}$ where $\partial_t \mathbf{m}$ models the evolution in time of the bed and I is the infiltration function. I models the amount of water that leaves ($I > 0$) or enters ($I < 0$) the flow per elementary boundary element.

WIDTH-AVERAGED AND DEPTH-AVERAGED ASYMPTOTIC EXPANSIONS

In this section, we present the strategy to derive the section-averaged non-linear dispersive model (see [Chapter 17](#)). To this end, we first introduce in [Section 16.1](#) the dimensionless problem by introducing the classical dispersive parameter μ_2 . We also introduce a dispersive parameter μ_1 but in the transversal direction. In this work, we assume that $\mu_1 < \mu_2$ to obtain the section-averaged dispersive model with respect to μ_2 as usually done. Due to the structure of the equations, we cannot obtain the model by a direct section-averaging. We need to develop first suitable asymptotic expansions in two steps, briefly summarized below:

- in [Section 16.2](#), the Euler equations (15.5)-(15.6) are width-averaged to get the following asymptotic expansion of the horizontal fluid velocity

$$u(t, x, y, z) = \langle u \rangle(t, x, z) + O(\mu_1)$$

where $\langle u \rangle$ is the width-averaged velocity.

- in [Section 16.3](#), the width-averaged Euler equations allows to obtain the asymptotic expansion of the horizontal width-averaged fluid velocity

$$\langle u \rangle(t, x, z) = \bar{u}(t, x) + \mu_2 f(\bar{u}(t, x), \Omega(t, x)) + O(\mu_2^2)$$

where \bar{u} is the section-averaged velocity for some function f given later on.

Thus, the asymptotic expansion of u up to order $O(\mu_2^2)$, the so-called "motion by slices" (see [70]), can be written as follows:

$$u(t, x, y, z) = \bar{u}(t, x) + \mu_2 f(\bar{u}(t, x), \Omega(t, x)) + O(\mu_2^2) .$$

Finally, using these asymptotic expansions, we are able in [Chapter 17](#) to section-average the Euler equations (15.5)-(15.6) to obtain the new one-dimensional non-linear dispersive equations.

16.1 DIMENSIONLESS EULER EQUATIONS

Let us consider the following scales involved in the wave motion: L a characteristic wave-length in the longitudinal direction, H_2 a characteristic water depth, H_1 a characteristic scale of the channel width and h_1 a characteristic wave-length in the transversal direction. We then define the classical dispersive parameter μ_2 (see *e.g.* [94])

$$\mu_2 = \frac{H_2^2}{L^2}$$

and $\mu_1 = \frac{h_1^2}{L^2}$ where μ_1 is also a dispersive parameter but in the transversal direction.

In the following, we consider the asymptotic regime:

$$h_1 < H_1 = H_2 \ll L$$

such that the following inequality holds

$$\mu_1 < \mu_2^2.$$

Under these assumptions, we get the following ordering:

$$\begin{aligned} \mu_1^2 < \frac{\mu_1^2}{\mu_2} < \min\left(\frac{\mu_1^2}{\mu_2^2}, \mu_1\mu_2\right) < \max\left(\frac{\mu_1^2}{\mu_2^2}, \mu_1\mu_2\right) < \mu_1 \\ < \min\left(\frac{\mu_1}{\mu_2}, \mu_2^2\right) < \max\left(\frac{\mu_1}{\mu_2}, \mu_2^2\right) < \mu_2. \end{aligned}$$

We also introduce $\mathbf{U} = (U, V = \sqrt{\mu_1}U, W = \sqrt{\mu_2}U)$ the scale of fluid velocity so that $V < W < U$. The time scale is $T = \frac{L}{U}$. We set $P = \frac{p}{\rho_0}$ and we define $\mathcal{P} = U^2$.

This allows us to introduce the dimensionless quantities of time \tilde{t} , space $(\tilde{x}, \tilde{y}, \tilde{z})$, pressure \tilde{P} , depth \tilde{d} , water elevation $\tilde{\eta}$ and velocity field $(\tilde{U}, \tilde{V}, \tilde{W})$, via the following scaling relation

$$\begin{aligned} \tilde{x} &= \frac{x}{L}, & \tilde{P} &= \frac{P}{\mathcal{P}}, & \tilde{\varphi} &= \frac{\varphi}{h_1}, & \tilde{y} &= \frac{y}{h_1}, & \tilde{u} &= \frac{u}{U}, & \tilde{d} &= \frac{d}{H_2}, \\ \tilde{z} &= \frac{z}{H_2}, & \tilde{v} &= \frac{v}{V}, & \tilde{\eta} &= \frac{\eta}{H_2}, & \tilde{t} &= \frac{t}{T}, & \tilde{w} &= \frac{w}{W}. \end{aligned} \tag{16.1}$$

Finally, we define the non-dimensional Froude's number

$$F_r = \frac{U}{\sqrt{gH_2}}.$$

For the sake of clarity and simplicity dropping $\tilde{\cdot}$, dividing the Euler equations (15.5) by ρ_0 , using the dimensionless variables (16.1), and reordering the terms with respect to the powers of μ_1 and μ_2 , the dimensionless incompressible Euler system (15.5) reads as follows:

$$\begin{aligned} \partial_x u + \partial_y v + \partial_z w &= 0 \\ \partial_t u + u\partial_x u + v\partial_y u + w\partial_z u + \partial_x P &= 0 \\ \mu_1 (\partial_t v + u\partial_x v + v\partial_y v + w\partial_z v) + \partial_y P &= 0 \\ \mu_2 (\partial_t w + u\partial_x w + v\partial_y w + w\partial_z w) + \partial_z P &= -\frac{1}{F_r^2} \end{aligned}$$

The boundary conditions (15.7), (15.8), (15.9) and (15.10) read

$$\begin{aligned} \partial_t \eta + u\partial_x \eta + v\partial_y \eta &= w \text{ and } P = 0 \text{ on } z = \eta \\ u\partial_x d + v\partial_y d &= w \text{ on } z = d \\ \text{or } u\partial_x \varphi + w\partial_z \varphi &= v \text{ on } y = \varphi \end{aligned}$$

and the irrotational equation (15.6) becomes

$$\partial_y u = \mu_1 \partial_x v, \quad \mu_1 \partial_z v = \mu_2 \partial_y w, \quad \partial_z u = \mu_2 \partial_x w. \quad (16.9)$$

16.2 3D-2D MODEL REDUCTION AND ASYMPTOTIC EXPANSIONS

In this first step, we focus on the width-averaging of System (16.2)–(16.5). In particular, we compute the asymptotic expansions of the velocity (u, v) and the pressure P of System (16.2)–(16.5) as a function of their width-averages. We obtain what we call "flat free surface approximation" property which means that the variations of the free surface following y can be neglected in the three-dimensional Euler system (16.2)–(16.5). Finally, the asymptotic approximation of the width-averaged Euler system is obtained. This averaged model is the starting point for the second step (see Section 16.3).

16.2.0.1 Asymptotic expansions of the fluid velocity

Given a function $(t, x, y, z) \mapsto X(t, x, y, z)$ and $(x, z) \mapsto q(x, z)$, we define X_q as follows

$$X_q(t, x, z) := X(t, x, q(x, z), z)$$

and we use the notations

$$\mathbf{w} = \begin{pmatrix} u \\ w \end{pmatrix}, \quad \operatorname{div}_{x,z} [\mathbf{w}] = \partial_x u + \partial_z w \quad \text{and} \quad \nabla_{x,z}(u) = \begin{pmatrix} \partial_x u \\ \partial_z u \end{pmatrix}$$

to represent the vector \mathbf{w} , the two dimensional divergence operator and the gradient operator with respect to the variable x and z .

Integrating the two first equations of the irrotational equations (16.9) for $s \in [\alpha(x, z), y]$, we get

$$u(t, x, y, z) = u_\alpha(t, x, z) + \mu_1 \int_{\alpha(x, z)}^y \partial_x v \, ds \quad (16.10)$$

and

$$w(t, x, y, z) = w_\alpha(t, x, z) + \frac{\mu_1}{\mu_2} \int_{\alpha(x, z)}^y \partial_z v \, ds. \quad (16.11)$$

Then, integrating the divergence equation (16.2) for $s \in [\alpha(x, z), y]$, we obtain

$$\int_{\alpha(x, z)}^y \partial_y v \, dy = - \int_{\alpha(x, z)}^y \operatorname{div}_{x,z} [\mathbf{w}] \, dy. \quad (16.12)$$

Thanks to Equations (16.10), (16.11) and (16.12) and the wet boundary condition (16.8) for $\varphi = \alpha$ the fluid velocity v is approximated by

$$\begin{aligned} v(t, x, y, z) &= -\operatorname{div}_{x,z} [\mathbf{w}_\alpha(t, x, z)(y - \alpha(x, z))] \\ &\quad + O\left(\frac{\mu_1}{\mu_2}\right). \end{aligned} \quad (16.13)$$

Coming back to Equations (16.10) and (16.11) together with Equation (16.13), the fluid velocity (u, w) can be written as

$$\begin{aligned} u(t, x, y, z) &= u_\alpha(t, x, z) - \frac{\mu_1}{2} \partial_x \operatorname{div}_{x,z} [\mathbf{w}_\alpha(t, x, z)(y - \alpha(x, z))^2] \\ &\quad + O\left(\frac{\mu_1^2}{\mu_2}\right) \end{aligned} \quad (16.14)$$

and

$$\begin{aligned} w(t, x, y, z) &= w_\alpha(t, x, z) - \frac{\mu_1}{2\mu_2} \partial_z \operatorname{div}_{x,z} [\mathbf{w}_\alpha(t, x, z)(y - \alpha(x, z))^2] \\ &\quad + O\left(\frac{\mu_1^2}{\mu_2^2}\right). \end{aligned} \quad (16.15)$$

16.2.1 *Width-averaged Euler equations*

Given a function $(t, x, y, z) \mapsto X(t, x, y, z)$, we define the width-average of X by the quantity

$$\langle X \rangle(t, x, z) := \frac{1}{\sigma(x, z)} \int_{\alpha(x, z)}^{\beta(x, z)} X(t, x, y, z) dy$$

where $\sigma(x, z) = \beta(x, z) - \alpha(x, z)$ is the width of the section at the elevation z .

We average the equations (16.2)–(16.5) for $y \in [\alpha(x, z), \beta(x, z)]$ using Leibniz integral rule, to get the width-averaged Euler system:

$$\left\{ \begin{array}{l} \frac{\partial}{\partial x} (\sigma \langle u \rangle) + \frac{\partial}{\partial z} (\sigma \langle w \rangle) = 0 \\ \frac{\partial}{\partial t} (\sigma \langle u \rangle) + \frac{\partial}{\partial x} (\sigma \langle u^2 \rangle) + \frac{\partial}{\partial z} (\sigma \langle uw \rangle) + \frac{\partial}{\partial x} (\sigma \langle P \rangle) = P_\beta \frac{\partial \beta}{\partial x} - P_\alpha \frac{\partial \alpha}{\partial x} \\ \mu_2 \left(\frac{\partial}{\partial t} (\sigma \langle w \rangle) + \frac{\partial}{\partial x} (\sigma \langle uw \rangle) + \frac{\partial}{\partial z} (\sigma \langle w^2 \rangle) \right) + \frac{\partial}{\partial z} (\sigma \langle P \rangle) = -\frac{\sigma}{F_r^2} \\ \quad + P_\beta \frac{\partial \beta}{\partial z} - P_\alpha \frac{\partial \alpha}{\partial z} \\ P_\beta = P_\alpha - \mu_1 \left\langle \frac{Dv}{Dt} \right\rangle \end{array} \right. \quad (16.16)$$

where $\frac{Dv}{Dt}$ stands for the material derivative

$$\frac{Dv}{Dt} := \partial_t v + u \partial_x v + v \partial_y v + w \partial_z v .$$

ASYMPTOTIC EXPANSIONS OF THE WIDTH-AVERAGED TERMS IN SYSTEM (16.16):

Thanks to the expressions (16.14) and (16.15), the average of the terms in System (16.16) can be written

$$\begin{aligned}
\sigma(x, z)\langle u \rangle(t, x, z) &= \sigma(x, z)u_\alpha(t, x, z) \\
&\quad - \frac{\mu_1}{6}\partial_x \operatorname{div}_{x,z} \left[\mathbf{w}_\alpha(t, x, z)\sigma(x, z)^3 \right] \\
&\quad + O\left(\frac{\mu_1^2}{\mu_2}\right), \\
\sigma(x, z)\langle w \rangle(t, x, z) &= \sigma(x, z)w_\alpha(t, x, z) \\
&\quad - \frac{\mu_1}{6\mu_2}\partial_z \operatorname{div}_{x,z} \left[\mathbf{w}_\alpha(t, x, z)\sigma(x, z)^3 \right] \\
&\quad + O\left(\frac{\mu_1^2}{\mu_2^2}\right), \\
\sigma(x, z)\langle u^2 \rangle(t, x, z) &= \sigma(x, z)u_\alpha^2(t, x, z) + O\left(\frac{\mu_1^2}{\mu_2}\right), \\
\sigma(x, z)\langle w^2 \rangle(t, x, z) &= \sigma(x, z)w_\alpha^2(t, x, z) + O\left(\frac{\mu_1^2}{\mu_2^2}\right), \\
\sigma(x, z)\langle uw \rangle(t, x, z) &= \sigma(x, z)u_\alpha(t, x, z)w_\alpha(t, x, z) \\
&\quad - u_\alpha(t, x, z)\frac{\mu_1}{6\mu_2}\partial_z \operatorname{div}_{x,z} \left[\mathbf{w}_\alpha(t, x, z)\sigma(x, z)^3 \right] \\
&\quad + O\left(\frac{\mu_1^2}{\mu_2^2}\right).
\end{aligned}$$

IRROTATIONALITY:

We lose the irrotational condition (16.9) by width-averaging, since we get

$$\frac{\partial}{\partial z}(\sigma\langle u \rangle) = \mu_2 \frac{\partial}{\partial x}(\sigma\langle w \rangle) + (u_\beta \partial_z \beta - w_\beta \partial_x \beta - u_\alpha \partial_z \alpha + w_\alpha \partial_x \alpha).$$

However, using the last equation in (16.9) together with Equations (16.14) and (16.15), we obtain

$$\frac{\partial u_\alpha}{\partial z} = \mu_2 \frac{\partial w_\alpha}{\partial x} + O(\mu_1). \quad (16.22)$$

The approximated irrotational condition can be also written as a function of $(\langle u \rangle, \langle w \rangle)$ by means of Equations (16.17) and (16.18).

ASYMPTOTIC EXPANSION OF THE PRESSURE:

The last equation in System (16.16) allows to write

$$P_\beta = P_\alpha + O(\mu_1). \quad (16.23)$$

Thanks to (16.23), on one hand, the terms in the right hand side of the second and third equations of System (16.16) can be simplified in

$$P_\beta \nabla_{x,z}(\beta) - P_\alpha \nabla_{x,z}(\alpha) = P_\alpha \nabla_{x,z}(\sigma) + O(\mu_1) .$$

On the other hand, coming back to System (16.2)–(16.5) together with Equation (16.23), integrating Equation (16.4) for $s \in [\alpha(x,z), y]$, we show that

$$\begin{aligned} P(t, x, y, z) &= P_\alpha(t, x, z) - \mu_1 \int_{\alpha(x,z)}^y \frac{Dv}{Dt} ds \\ &= P_\alpha(t, x, z) + O(\mu_1) . \end{aligned} \quad (16.24)$$

As a consequence, the width-averaged pressure $\langle P \rangle$ can be approximated by P_α at order $O(\mu_1)$, *i. e.*

$$\langle P \rangle(t, x, z) = P_\alpha(t, x, z) + O(\mu_1) .$$

ASYMPTOTIC EXPANSION OF THE FREE SURFACE:

Using Equations (16.15) and (16.24) in Equation (16.5), we can write the z -gradient of the pressure as

$$\frac{\partial}{\partial z} P_\alpha(t, x, z) = -\frac{1}{F_r^2} - \mu_2 \frac{D}{Dt} w_\alpha(t, x, z) + O(\mu_1) .$$

Next, integrating this equation for $s \in [z, \eta(t, x, y)]$ using the boundary condition (16.6) with Equation (16.24) lead to

$$P_\alpha(t, x, z) = \frac{\eta(t, x, y) - z}{F_r^2} + \mu_2 \int_z^{\eta(t,x,y)} \frac{D}{Dt} w_\alpha(t, x, s) ds + O(\mu_1) .$$

Thus, taking the y -derivative of this expression yields to

$$0 = \partial_y \eta \left(\frac{1}{F_r^2} + \mu_2 \frac{D}{Dt} w_\alpha|_{z=\eta} \right) + O(\mu_1) = -\partial_y \eta \partial_z P|_{z=\eta} + O(\mu_1) .$$

Consequently, since $\partial_z P|_{z=\eta} \neq 0$, we get $\partial_y \eta = O(\mu_1)$, *i. e.*, we obtain the flat surface approximation

$$\eta(t, x, y) = \eta_{\text{eq}}(t, x) + O(\mu_1) \quad (16.25)$$

for some function η_{eq} defined hereafter. It means that one can neglect the y -variations of the free surface of the three-dimensional model (16.2)–(16.5) (see Figure 16.1). In other words, the fluid cross-section $\Omega(t, x)$ (15.3) and the wet area $A(t, x)$ (15.4) can be simplified in

$$\begin{aligned} \Omega_{\text{eq}}(t, x) &= \{(y, z) \in \mathbb{R}^2; \alpha(x, z) \leq y \leq \beta(x, z) \\ &\quad \text{and } d^*(x) \leq z \leq \eta_{\text{eq}}(t, x)\} \end{aligned} \quad (16.26)$$

with

$$A_{\text{eq}} = |\Omega_{\text{eq}}(t, x)| \quad (16.27)$$

thanks to

$$\begin{aligned} A(t, x) &= \int_{\Omega(t, x)} dy dz \\ &= \int_{y^-(t, x)}^{y^+(t, x)} \eta(t, x, y) - d(x, y) dy \\ &= \int_{y^-(t, x)}^{y^+(t, x)} \eta_{\text{eq}}(t, x) - d(x, y) dy + O(\mu_1) \\ &= \int_{d^*(x)}^{\eta_{\text{eq}}(t, x)} \sigma(x, z) dz + O(\mu_1) \\ &= \int_{\Omega_{\text{eq}}(t, x)} dy dz + O(\mu_1) \\ &= A_{\text{eq}}(t, x) + O(\mu_1) \end{aligned} \quad (16.28)$$

Therefore, in the following, we consider the fluid cross-section Ω_{eq} instead of Ω for which computations are easier. In the following, we set η_{eq} as

$$\eta_{\text{eq}}(t, x) = \eta(t, x, y^*(x)).$$

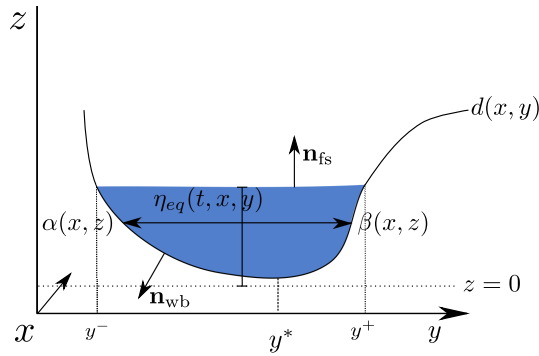


Figure 16.1: Equivalent geometric set-up.

WIDTH-AVERAGED EQUATIONS:

Using all the previous equations (16.17)–(16.21) and (16.24)–(16.25) in System (16.16), we obtain the width-averaged Euler equations expressed as the couple of unknown (w_α, P_α) :

$$\left\{ \begin{array}{l} \operatorname{div}_{x,z} [\sigma \mathbf{w}_\alpha] + O\left(\frac{\mu_1^2}{\mu_2^2}\right) = \frac{\mu_1}{6\mu_2} \frac{\partial}{\partial z} (\sigma \partial_z (\operatorname{div}_{x,z} [\mathbf{w}_\alpha \sigma^3])) \\ \frac{\partial}{\partial t} (\sigma u_\alpha) + \operatorname{div}_{x,z} [\sigma u_\alpha \mathbf{w}_\alpha] + \frac{\partial}{\partial x} (\sigma P_\alpha) + O\left(\frac{\mu_1^2}{\mu_2^2}\right) = P_\alpha \frac{\partial \sigma}{\partial x} \\ + \frac{\mu_1}{6\mu_2} \partial_x (u_\alpha \partial_z \operatorname{div}_{x,z} [\mathbf{w}_\alpha \sigma^3]) \\ \mu_2 \left(\frac{\partial}{\partial t} (\sigma w_\alpha) + \operatorname{div}_{x,z} [\sigma w_\alpha \mathbf{w}_\alpha] \right) + \frac{\partial}{\partial z} (\sigma P_\alpha) = -\frac{\sigma}{F_r^2} \\ + P_\alpha \frac{\partial \sigma}{\partial z} + O(\mu_1) \end{array} \right. \quad (16.29)$$

where the fluid domain is now defined (see Equations (16.25)-(16.26) and Figure 16.2) by

$$\langle \Omega \rangle (t, x) = \{z \in \mathbb{R}; d^*(x) \leq z \leq \eta_{\text{eq}}(t, x)\} .$$

For a given time $t \in [0, T]$ and $x \in [0, L_c]$, the boundary of the fluid cross-section $\partial \langle \Omega \rangle (t, x)$ is composed of a free surface $\langle \Gamma_{\text{fs}} \rangle (t, x) = \{z \in \mathbb{R}; z = \eta_{\text{eq}}(t, x)\}$ and a wet boundary $\langle \Gamma_{\text{wb}} \rangle (x) = \{z \in \mathbb{R}; z = d^*(x)\}$ such that

$$\partial \langle \Omega \rangle (t, x) = \langle \Gamma_{\text{fs}} \rangle (t, x) \cup \langle \Gamma_{\text{wb}} \rangle (x)$$

The unit outward normal vector to the free surface is

$$\langle \mathbf{n}_{\text{fs}} \rangle (t, x) = \frac{1}{\sqrt{1 + (\partial_x \eta_{\text{eq}}(t, x))^2}} \begin{pmatrix} -\partial_x \eta_{\text{eq}}(t, x) \\ 1 \end{pmatrix}$$

and the unit outward normal vector to the wet boundary is

$$\langle \mathbf{n}_{\text{wb}} \rangle (x) = \frac{1}{\sqrt{1 + (\partial_x d^*(x))^2}} \begin{pmatrix} \partial_x d^*(x) \\ -1 \end{pmatrix} .$$

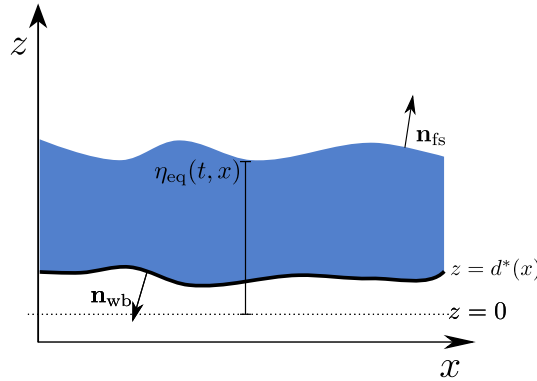


Figure 16.2: Width-averaged geometric set-up.

These equations are completed with the irrotational condition (16.22) and the following boundary conditions

$$\begin{aligned} \frac{\partial \eta_{\text{eq}}}{\partial t} + u_\alpha \frac{\partial \eta_{\text{eq}}}{\partial x} &= w_\alpha + O\left(\frac{\mu_1}{\mu_2}\right) \text{ and } P_\alpha = O(\mu_1) \text{ on } z = \eta_{\text{eq}} \\ u_\alpha \partial_x d &= w_\alpha + O\left(\frac{\mu_1}{\mu_2}\right) \text{ on } z = d \end{aligned}$$

which are obtained through Equations (16.6) and (16.7) using Equations (16.13), (16.14), (16.15).

These end the first step.

16.3 2D-1D LIKE MODEL REDUCTION AND ASYMPTOTIC EXPANSIONS

In what follows, we note $f_b(t, x) = f_\alpha(t, x, d^*(x))$ for a given function f and S

$$S(x, z) = \int_{d^*(x)}^z \sigma(x, s) ds$$

the wet area of water between $d^*(x)$ and z .

In the sequel, we consider the fluid cross-section (16.26) $\Omega_{\text{eq}}(t, x)$ instead of $\Omega(t, x)$ thanks to (16.28) (see also (15.3)). Thus, we define the local height of the water in $\Omega_{\text{eq}}(t, x)$ by

$$h_{\text{eq}}(t, x, y) = \eta_{\text{eq}}(t, x) - d(x, y).$$

In this new setting, the section-average \bar{f}_{eq} of a function f can be defined by

$$\bar{f}_{\text{eq}} = \frac{1}{A_{\text{eq}}(t, x)} \int_{d^*(x)}^{\eta_{\text{eq}}(t, x)} \int_{\alpha(x, z)}^{\beta(x, z)} f(t, x, y, z) dy dz.$$

16.3.1 Asymptotic expansion of the fluid velocity

Let us first integrate Equation (16.22) for $s \in [d^*(x), z]$, to get

$$u_\alpha(t, x, z) = u_b(t, x) + \mu_2 \int_{d^*(x)}^z \frac{\partial w_\alpha}{\partial x} ds + O(\mu_1). \quad (16.32)$$

Then, dropping all the lower terms in $O\left(\frac{\mu_1}{\mu_2}\right)$ in the first equation of System (16.29):

$$\frac{\partial}{\partial z}(\sigma w_\alpha) = -\frac{\partial}{\partial x}(\sigma u_\alpha) + O\left(\frac{\mu_1}{\mu_2}\right)$$

and integrating the above equation for $s \in [d^*(x), z]$, keeping in mind the boundary condition (16.31), we obtain

$$\sigma(x, z)w_\alpha(t, x, z) = -\frac{\partial}{\partial x} \left(\int_{d^*(x)}^z \sigma u_\alpha ds \right) + O\left(\frac{\mu_1}{\mu_2}\right).$$

By means of Equation (16.32), the previous equation becomes:

$$w_\alpha(t, x, z) = -\frac{1}{\sigma(x, z)} \frac{\partial}{\partial x} (u_b(t, x)S(x, z)) + O(\mu_2).$$

By injecting Equation (16.33) in (16.32), we get

$$u_\alpha(t, x, z) = u_b(t, x) - \mu_2 \int_{d^*(x)}^z \partial_x \mathcal{S}(u_b, x, s) ds + O(\mu_2^2) \quad (16.34)$$

where

$$\mathcal{S}(u, x, z) = \frac{1}{\sigma(x, z)} \frac{\partial}{\partial x} (uS(x, z)). \quad (16.35)$$

As a consequence, we have

$$\begin{aligned} \bar{u}_{\text{eq}} &= \frac{1}{A_{\text{eq}}(t, x)} \int_{d^*(x)}^{\eta_{\text{eq}}(t, x)} \int_{\alpha(x, z)}^{\beta(x, z)} u(t, x, y, z) dy dz \\ &= \frac{1}{A_{\text{eq}}(t, x)} \int_{d^*(x)}^{\eta_{\text{eq}}(t, x)} \sigma(x, z) \langle u \rangle(t, x, z) dz \\ &= \frac{1}{A_{\text{eq}}(t, x)} \int_{d^*(x)}^{\eta_{\text{eq}}(t, x)} \sigma(x, z) u_\alpha(t, x, z) dz + O(\mu_1) \end{aligned}$$

and through Equation (16.34), we obtain

$$\begin{aligned} u_b(t, x) &= \bar{u}_{\text{eq}}(t, x) \\ &+ \frac{\mu_2}{A_{\text{eq}}(t, x)} \int_{d^*(x)}^{\eta_{\text{eq}}(t, x)} \sigma(x, z) \left(\int_{d^*(x)}^z \partial_x \mathcal{S}(u_b, x, s) ds \right) dz + O(\mu_2^2). \end{aligned} \quad (16.36)$$

Using the fact that

$$u_b(t, x) = \bar{u}_{\text{eq}}(t, x) + O(\mu_2)$$

in Equation (16.36) allows to write $u_b(t, x)$ as a function of \bar{u}_{eq} :

$$\begin{aligned} u_b(t, x) &= \bar{u}_{\text{eq}}(t, x) \\ &+ \frac{\mu_2}{A_{\text{eq}}(t, x)} \int_{d^*(x)}^{\eta_{\text{eq}}(t, x)} \sigma(x, z) \left(\int_{d^*(x)}^z \partial_x \mathcal{S}(\bar{u}_{\text{eq}}, x, s) ds \right) dz + O(\mu_2^2). \end{aligned}$$

$$(16.37)$$

Finally, thanks to (16.10), (16.34), (16.37), by linearity of \mathcal{S} (16.35), we obtain

$$\begin{aligned} u(t, x, y, z) &= u_\alpha(t, x, z) + O(\mu_1) \\ &= u_b(t, x) - \mu_2 \int_{d^*(x)}^z \partial_x \mathcal{S}(u_b, x, s) ds + O(\mu_2^2) \\ &= \bar{u}_{\text{eq}}(t, x) + \mu_2 B_0(\bar{u}_{\text{eq}}, x, z) + O(\mu_2^2) \end{aligned} \quad (16.38)$$

where

$$\begin{aligned} B_0(\bar{u}_{\text{eq}}, x, z) &= \frac{1}{A_{\text{eq}}(t, x)} \int_{d^*(x)}^{\eta_{\text{eq}}(t, x)} \left(\sigma(x, z) \int_{d^*(x)}^z \partial_x \mathcal{S}(\bar{u}_{\text{eq}}, x, s) ds \right) dz \\ &\quad - \int_{d^*(x)}^z \partial_x \mathcal{S}(\bar{u}_{\text{eq}}, x, s) ds. \end{aligned} \quad (16.39)$$

Repeating the computations with Equations (16.15), (16.33) and (16.38), we find

$$w(t, x, y, z) = -\mathcal{S}(\bar{u}_{\text{eq}}, x, z) + O\left(\frac{\mu_1}{\mu_2}\right). \quad (16.40)$$

16.3.2 Pressure decomposition

Integrating for $s \in [z, \eta_{\text{eq}}]$, the third equation of System (16.29), we get

$$\begin{aligned} P_\alpha(t, x, z) &= \frac{(\eta_{\text{eq}}(t, x) - z)}{F_r^2} \\ &\quad + \int_z^{\eta_{\text{eq}}(t, x)} \frac{\mu_2}{\sigma} \left(\frac{\partial}{\partial t}(\sigma w_\alpha) + \text{div}_{x,z}[\sigma w_\alpha w_\alpha] \right) ds \\ &\quad + O(\mu_1). \end{aligned}$$

Thanks to the free surface boundary condition (16.30), Equation (16.38) and Equation (16.40), we obtain the asymptotic expansion of the pressure P at order $O(\mu_2^2)$

$$P(t, x, y, z) = P_h(t, x, z) + \mu_2 P_{\text{nh}}(t, x, z) + O(\mu_2^2)$$

where

$$P_h(t, x, z) = \frac{(\eta_{\text{eq}}(t, x) - z)}{F_r^2} \quad (16.41)$$

is the usual hydrostatic pressure and

$$\begin{aligned}
P_{\text{nh}}(t, x, z) = & \int_z^{\eta_{\text{eq}}(t, x)} \frac{1}{2\sigma(x, s)^2} \partial_z \left((\sigma(x, s) \mathcal{S}(\bar{u}_{\text{eq}}, x, s))^2 \right) ds \\
& - \int_z^{\eta_{\text{eq}}(t, x)} \partial_t \mathcal{S}(\bar{u}_{\text{eq}}, x, s) \\
& + \frac{\bar{u}_{\text{eq}}(t, x)}{\sigma(x, s)} \partial_x (\sigma(x, s) \mathcal{S}(\bar{u}_{\text{eq}}, x, s)) ds
\end{aligned}$$

is a non-hydrostatic pressure.

Handling the terms in the non-hydrostatic pressure differently, one can write

$$P_{\text{nh}}(t, x, z) = \mathcal{D}(\bar{u}_{\text{eq}}) \int_z^{\eta_{\text{eq}}} \frac{S(x, s)}{\sigma(x, s)} ds + \int_z^{\eta_{\text{eq}}} \mathcal{Q}(\bar{u}_{\text{eq}}, S, \sigma) ds \quad (16.42)$$

where

$$\mathcal{D}(u) = (\partial_x u)^2 - \partial_t \partial_x u - u \partial_{xx} u, \quad (16.43)$$

and

$$\begin{aligned}
\mathcal{Q}(u, S, \sigma) = & \frac{u^2}{\sigma(x, s)} \left(\frac{\partial_x S(x, s) \partial_x \sigma(x, s)}{\sigma(x, s)} - \partial_{xx} S(x, s) \right) \\
& + \partial_x \left(\frac{u^2}{2} \right) \frac{S(x, s) \partial_x \sigma(x, s)}{\sigma(x, s)^2} - (\partial_t u + u \partial_x u) \frac{\partial_x S(x, s)}{\sigma(x, s)}. \quad (16.44)
\end{aligned}$$

The non-hydrostatic pressure (16.42) is more suitable to have a formulation akin to the [SGN](#) equations (14.2)-(14.4).

A NEW NON-LINEAR DISPERSIVE MODEL

In this section, we assume that the flow is governed, on the space-time domain $\Omega_{\text{eq}} = \bigcup_{0 \leq t \leq 1} \Omega_{\text{eq}}(t)$, instead of Ω (see Def. (15.2)), by the three-dimensional incompressible and irrotational Euler equations (16.2)–(16.5), (16.9) where the wet region is defined as the region in which the fluid resides at each time $t \in [0, 1]$, $\Omega_{\text{eq}}(t) = \bigcup_{0 \leq x \leq 1} \Omega_{\text{eq}}(t, x)$ with $\Omega_{\text{eq}}(t, x)$ the fluid cross-section defined by (16.26).

These equations (16.2)–(16.5) are completed with the kinematic boundary condition at the free surface (16.6) $\Gamma_{\text{fs}}(t, x) = \{z \in \mathbb{R}; z = \eta_{\text{eq}}(t, x)\}$, and a no-penetration condition on the wet boundary (16.7) $\Gamma_{\text{wb}}(x) = \{(y, z) \in \mathbb{R}^2; z = d(x, y)\}$ where

$$\partial\Omega_{\text{eq}}(t, x) = \Gamma_{\text{fs}}(t, x) \cup \Gamma_{\text{wb}}(x) .$$

These boundary conditions can be written under the following form:

$$\int_{\partial\Omega_{\text{eq}}(t, x)} (\partial_t \mathbf{m} + u \partial_x \mathbf{m} - \mathbf{v}) \cdot \mathbf{n} \, ds = 0 \text{ on } \partial\Omega_{\text{eq}}(t, x) \quad (17.1)$$

where $\mathbf{m}(t, x, y, z) = \begin{cases} (y, \eta_{\text{eq}}(t, x)) & \text{if } \mathbf{m} \in \Gamma_{\text{fs}}(t, x) \\ (y, d(x, y)) & \text{if } \mathbf{m} \in \Gamma_{\text{wb}}(t, x) \end{cases}$ is a boundary point,

$\mathbf{n} = \begin{cases} \mathbf{n}_{\text{fs}} & \text{if } \mathbf{m} \in \Gamma_{\text{fs}}(t, x) \\ \mathbf{n}_{\text{wb}} & \text{if } \mathbf{m} \in \Gamma_{\text{wb}}(t, x) \end{cases}$ stands for the outward unit normal

vector in the $\Omega_{\text{eq}}(t, x)$ -plane (see Figure 16.1) and $\mathbf{v} = \begin{pmatrix} v \\ w \end{pmatrix}$.

To work with the wet region, we introduce its indicator function

$$\Phi(t, x, y, z) := \mathbb{1}_{\Omega_{\text{eq}}(t, x)} \text{ for all } t, x, y, z \in \mathbb{R}$$

where

$$\mathbb{1}_P := \begin{cases} 1 & \text{if } P \text{ is true,} \\ 0 & \text{if } P \text{ is false.} \end{cases}$$

The function Φ is advected by the flow so its material derivative, with respect to the flow \mathbf{u} , must therefore be zero. Moreover, thanks to

the incompressibility condition, Φ satisfies the following indicator transport equation

$$\partial_t \Phi + \partial_x(\Phi u) + \text{div}_{y,z} [\Phi \mathbf{v}] = 0 \text{ on } \Omega_{\text{eq}}(t) \quad (17.2)$$

where $\text{div}_{y,z} [\mathbf{v}] = \partial_y v + \partial_z w$ with $\mathbf{v} = (v, w)$.

17.1 EQUATION OF THE CONSERVATION OF THE MASS

Integrating Equation (17.2) over the section $\Omega_{\text{eq}}(t, x)$ and using Leibniz integral rule, we get the following conservation of the mass equation

$$\begin{aligned} & \int_{\Omega_{\text{eq}}(t,x)} \partial_t \Phi + \partial_x(\Phi u) + \text{div}_{y,z} [\Phi \mathbf{v}] \, dy \, dz \\ & = \partial_t A_{\text{eq}} + \partial_x Q_{\text{eq}} = 0 \end{aligned} \quad (17.3)$$

where Q_{eq} is the water discharge

$$Q_{\text{eq}} = A_{\text{eq}} \bar{u}_{\text{eq}} \quad (17.4)$$

and \bar{u}_{eq} is $\Omega_{\text{eq}}(t, x)$ section-averaged velocity given by

$$\bar{u}_{\text{eq}} = \frac{1}{A_{\text{eq}}(t, x)} \int_{\Omega_{\text{eq}}(t,x)} u(t, x, y, z) \, dy \, dz .$$

17.2 EQUATION OF THE CONSERVATION OF THE MOMENTUM

In order to get the momentum equation of the section-averaged free surface model, we integrate each terms of (16.3) along the section $\Omega_{\text{eq}}(t, x)$ as follows:

$$\int_{\Omega_{\text{eq}}(t,x)} \underbrace{\partial_t(u)}_{a_1} + \underbrace{\partial_x(u^2)}_{a_2} + \underbrace{\text{div}_{y,z} [u\mathbf{v}]}_{a_3} + \underbrace{\partial_x P}_{a_4} \, dy \, dz = 0 .$$

17.2.1 Computation of the term $\int_{\Omega_{\text{eq}}(t,x)} a_1 \, dy \, dz$

We have

$$\int_{\Omega_{\text{eq}}(t,x)} \partial_t(u) \, dy \, dz = \partial_t \int_{\Omega_{\text{eq}}(t,x)} u \, dy \, dz - \int_{\partial\Omega_{\text{eq}}(t,x)} u \partial_t \mathbf{m} \cdot \mathbf{n} \, ds . \quad (17.5)$$

17.2.2 Computation of the term $\int_{\Omega_{\text{eq}}(t,x)} a_2 \, dy \, dz$

Thanks to the asymptotic approximation of the term u (16.38) up to order $O(\mu_2^2)$, the non-linear term u^2 can be written as:

$$u^2(t, x, y, z) = \bar{u}_{\text{eq}}(t, x)^2 + 2\mu_2 \bar{u}_{\text{eq}} B_0(\bar{u}_{\text{eq}}, x, z) + O(\mu_2^2) .$$

Therefore, we have

$$\int_{\Omega_{\text{eq}}(t,x)} \partial_x(u^2) \, dy \, dz = \partial_x \int_{\Omega_{\text{eq}}(t,x)} u^2 \, dy \, dz - \int_{\partial\Omega_{\text{eq}}(t,x)} u^2 \partial_x \mathbf{m} \cdot \mathbf{n} \, ds$$

i. e.

$$\begin{aligned} \int_{\Omega_{\text{eq}}(t,x)} \partial_x(u^2) \, dy \, dz &= \partial_x(Q_{\text{eq}}^2 / A_{\text{eq}}) + \mu_2 \partial_x \left(\frac{Q_{\text{eq}} B(x, Q_{\text{eq}} / A_{\text{eq}})}{A_{\text{eq}}} \right) \\ &\quad - \int_{\partial\Omega_{\text{eq}}(t,x)} u^2 \partial_x \mathbf{m} \cdot \mathbf{n} \, ds + O(\mu_2^2) \end{aligned} \quad (17.6)$$

where B stands for the quadratic part of the fluid velocity, defined via B_0 (16.39), and

$$\begin{aligned} B(x, Q_{\text{eq}} / A_{\text{eq}}) &= \int_{\Omega_{\text{eq}}(t,x)} 2B_0(Q_{\text{eq}} / A_{\text{eq}}, x, z) \, dy \, dz \\ &= 2 \int_{d^*(x)}^{\eta_{\text{eq}}(t,x)} \sigma(x, z) B_0(Q_{\text{eq}} / A_{\text{eq}}, x, z) \, dz \\ &= 0 . \end{aligned}$$

17.2.2.1 Computation of the term $\int_{\Omega_{\text{eq}}(t,x)} a_3 \, dy \, dz$

We have

$$\int_{\Omega_{\text{eq}}(t,x)} \text{div}_{y,z} [u\mathbf{v}] \, dy \, dz = \int_{\partial\Omega_{\text{eq}}(t,x)} u\mathbf{v} \cdot \mathbf{n} \, ds . \quad (17.7)$$

17.2.3 Computation of the term $\int_{\Omega_{\text{eq}}(t,x)} a_1 + a_2 + a_3 \, dy \, dz$

Gathering the results of the computations (17.5), (17.6) and (17.7), using the boundary conditions (17.1), we get

$$\begin{aligned} \int_{\Omega_{\text{eq}}(t,x)} a_1 + a_2 + a_3 \, dy \, dz &= \partial_t Q_{\text{eq}} + \partial_x(Q_{\text{eq}}^2 / A_{\text{eq}}) \\ &\quad + \mu_2 \partial_x \left(\frac{Q_{\text{eq}} B(x, Q_{\text{eq}} / A_{\text{eq}})}{A_{\text{eq}}} \right) \\ &\quad - \int_{\partial\Omega_{\text{eq}}(t,x)} u (\partial_t \mathbf{m} + u \partial_x \mathbf{m} - \mathbf{v}) \cdot \mathbf{n} \, ds \\ &\quad + O(\mu_2^2) \\ &= \partial_t Q_{\text{eq}} + \partial_x(Q_{\text{eq}}^2 / A_{\text{eq}}) + O(\mu_2^2) . \end{aligned}$$

(17.8)

17.2.4 Computation of the term $\int_{\Omega_{\text{eq}}(t,x)} a_4 dy dz$

We apply the Leibniz rule to the gradient of the pressure and we obtain

$$\int_{\Omega_{\text{eq}}(t,x)} \partial_x P dy dz = \partial_x \int_{\Omega_{\text{eq}}(t,x)} P dy dz - \int_{\partial\Omega_{\text{eq}}(t,x)} P \partial_x \mathbf{m} \cdot \mathbf{n} ds \quad (17.9)$$

where

$$\int_{\partial\Omega_{\text{eq}}(t,x)} P \partial_x \mathbf{m} \cdot \mathbf{n} ds = \int_{\Gamma_{\text{wb}}(x)} P \partial_x \mathbf{m} \cdot \mathbf{n} ds$$

thanks to the free surface boundary condition (15.8).

To compute the above boundary integral, we consider the parametrisation

$$z \in [d^*(x), \eta_{\text{eq}}(t, x)] \mapsto \mathbf{m}(x, z) = \begin{pmatrix} \varphi(x, z) \\ z \end{pmatrix}$$

where φ is given by (15.1). The unit speed curve parameterisation is

$$\|\partial_y \mathbf{m}(x, z)\| = \sqrt{1 + (\partial_y \varphi(x, z))^2}.$$

The outward unit normal is given by

$$\mathbf{n} = \begin{cases} \frac{1}{\|\mathbf{n}\|} \begin{pmatrix} -1 \\ \partial_z \alpha \end{pmatrix} & \text{if } \varphi(x, z) < y^*(x) \\ \frac{1}{\|\mathbf{n}\|} \begin{pmatrix} 1 \\ \partial_z \beta \end{pmatrix} & \text{otherwise} \end{cases}$$

where

$$\|\mathbf{n}(x, z)\| = \|\partial_y \mathbf{m}(x, z)\|.$$

Thus, we obtain

$$\int_{\Gamma_{\text{wb}}(x)} P \partial_x \mathbf{m} \cdot \mathbf{n} ds = \int_{d^*(x)}^{\eta_{\text{eq}}(t,x)} P(t, x, z) \partial_x \sigma(x, z) dz.$$

HYDROSTATIC PART: (SEE EQUATION (16.41)) Using $P = P_h$ in Equation (17.9), we get:

$$\int_{\Omega_{\text{eq}}(t,x)} \partial_x P_h(t, x, z) dy dz = \partial_x (I_1(t, x)) - I_2(t, x) \quad (17.10)$$

where

$$\begin{aligned} I_1(t, x) &= \int_{\Omega_{\text{eq}}(t, x)} P_h(t, x, z) dy dz \\ &= \int_{d^*(x)}^{\eta_{\text{eq}}(t, x)} \frac{(\eta_{\text{eq}}(t, x) - z)}{F_r^2} \sigma(x, z) dz \end{aligned} \quad (17.11)$$

and

$$I_2(t, x) = \int_{d^*(x)}^{\eta_{\text{eq}}(t, x)} \frac{(\eta_{\text{eq}}(t, x) - z)}{F_r^2} \partial_x \sigma(x, z) dz \quad (17.12)$$

are respectively the hydrostatic pressure term and the hydrostatic pressure source term.

NON-HYDROSTATIC PART: (SEE EQUATION (16.42)) Using $P = P_{\text{nh}}$ in Equation (17.9), we get

$$\int_{\Omega_{\text{eq}}(t, x)} \partial_x P_{\text{nh}}(t, x, z) dy dz = \partial_x (DI_1(t, x)) - DI_2(t, x) \quad (17.13)$$

where

$$\begin{aligned} DI_1(t, x) &= \int_{d^*(x)}^{\eta_{\text{eq}}(t, x)} P_{\text{nh}}(t, x, z) \sigma(x, z) dz \\ &= G(x, A_{\text{eq}}(t, x)) \mathcal{D}(\bar{u}_{\text{eq}}(t, x)) \\ &\quad + \int_{d^*(x)}^{\eta_{\text{eq}}} \sigma(x, z) \int_z^{\eta_{\text{eq}}} \mathcal{Q}(\bar{u}_{\text{eq}}(t, x), S(x, s), \sigma) ds dz \end{aligned}$$

where

$$G(x, A_{\text{eq}}) = \int_{d^*(x)}^{\eta_{\text{eq}}} \sigma(x, z) \int_z^{\eta_{\text{eq}}} \frac{S(x, s)}{\sigma(x, s)} ds dz . \quad (17.14)$$

The quantities \mathcal{D} and \mathcal{Q} are given by (16.43), (16.44), and

$$DI_2(t, x) = \int_{d^*(x)}^{\eta_{\text{eq}}(t, x)} P_{\text{nh}}(t, x, z) \partial_x \sigma(x, z) dz . \quad (17.15)$$

Each term DI_j for $j = 1, 2$ is the non-hydrostatic counterpart of the term I_j for $j = 1, 2$.

Remark 17.2.1. The terms $I_j(t, x)$ and $DI_j(t, x)$, respectively, can be written $I_j(x, A_{\text{eq}}(t, x))$ and $DI_j(x, A_{\text{eq}}(t, x), Q_{\text{eq}}(t, x))$ for $j = 1$ and $j = 2$. In what follows, we make use of these notations.

Finally, gathering Equation (17.10) and Equation (17.13), noting

$$\begin{aligned} \mathcal{G}(x, A_{\text{eq}}, Q_{\text{eq}}) &= -\partial_x \left(\int_{d^*(x)}^{\eta_{\text{eq}}} \sigma(x, z) \int_z^{\eta_{\text{eq}}} \mathcal{Q} \left(\frac{Q_{\text{eq}}}{A_{\text{eq}}}, S, \sigma \right) ds dz \right) \\ &\quad + DI_2, \end{aligned} \quad (17.16)$$

we obtain

$$\begin{aligned} \int_{\Omega_{\text{eq}}(t,x)} \partial_x P dy dz &= \partial_x (I_1(x, A_{\text{eq}})) - I_2(x, A_{\text{eq}}) \\ &\quad + \mu_2 \partial_x (G(x, A_{\text{eq}}) \mathcal{D}(\bar{u}_{\text{eq}})) \\ &\quad - \mu_2 \mathcal{G}(x, A_{\text{eq}}, Q_{\text{eq}}) \end{aligned} \quad (17.17)$$

Gathering results (17.8) and (17.17), we get the equation of the conservation of the momentum.

17.3 THE DISPERSIVE MODEL FOR ARBITRARY NON RECTANGULAR CHANNEL/RIVER

From now on, we omit the notations X_{eq} and \bar{X} for the sake of readability.

Gathering Equations (17.3), (17.8) and (17.17), we present the new one-dimensional dispersive model for open channel/river flows

$$\begin{cases} \partial_t A + \partial_x Q = 0 \\ \partial_t Q + \partial_x \left(\frac{Q^2}{A} + I_1(x, A) \right) + \mu_2 \partial_x (G(x, A) \mathcal{D}(u)) = I_2(x, A) \\ + \mu_2 \mathcal{G}(x, A, Q) + O(\mu_2^2) \end{cases} \quad (17.18)$$

where A is the wet area (16.27), Q is the water discharge (17.4), I_1 (resp. I_2) is the hydrostatic pressure (resp. source) terms (17.11) (resp. (17.12)), $G(x, A)$ (17.14) generalises $\frac{h^3}{3}$ in the classical SGN equations (see (14.3)), $\mathcal{G}(x, A, Q)$ (17.16) extends the uneven bottom source term in the classical SGN equations and $\partial_x (G(x, A) \mathcal{D}(u))$ is the dispersive term where \mathcal{D} is given by (16.43).

The new section-averaged model extends the section-averaged free surface model for open channel flows [19, 55] by taking $\mu_2 = 0$ and extends the SGN equations for uneven bottom [7, 78, 139, 140] to arbitrary channel/river section.

THE DISPERSIVE MODEL FOR RECTANGULAR SECTION

For the specific case of the rectangular section, almost all the previous computations are the same. The changes are mainly in the geometrical definition and the boundaries of the channel.

We consider the motion of an incompressible and irrotational fluid with constant density $\rho_0 > 0$ in a three dimensional domain (see [Figure 18.1](#))

$$\Omega(t) = \{(x, y, z) \in \mathbb{R}^3; x \in [0, L_c], \\ \alpha(x) \leq y - y^*(x) \leq \beta(x), d(x) \leq z \leq \eta(t, x)\}$$

where $y^*(x)$ describes the transversal variation of the channel with respect to the main channel direction, defined by

$$y^*(x) = \frac{\alpha(x) + \beta(x)}{2}$$

where α and β are the transversal limit of the channel. Here the bottom d is now a function of x only and $d^*(x) = d(x)$ by definition. Thus, in view of the definition of the fluid domain, since $\sigma = \sigma(x) = \beta(x) - \alpha(x)$ and $y^-(x) = \alpha(x)$ and $y^+(x) = \beta(x)$, the wet-area can be simply defined by

$$A(t, x) = \sigma(x) \langle h(t, x) \rangle$$

where $\langle h(t, x) \rangle$ is width-averaged of the local height of the water $h(t, x, y) := \eta(t, x, y) - d(x)$, *i.e.*, $\langle h(t, x) \rangle = \langle \eta(t, x) \rangle - d(x)$.

The boundary of the domain $\Omega(t)$ is defined by $\partial\Omega(t)$ and is decomposed into four parts: the free surface $\Gamma_{fs}(t)$, the wet boundary $\Gamma_{wb}(t)$, the inflow boundary $\Gamma_i(t)$ and the outflow boundary $\Gamma_o(t)$. The wet boundary can be decomposed itself in three parts: the bottom $\Gamma_b(t)$, the left lateral boundary $\Gamma_{lb}(t)$ and the right one $\Gamma_{rb}(t)$.

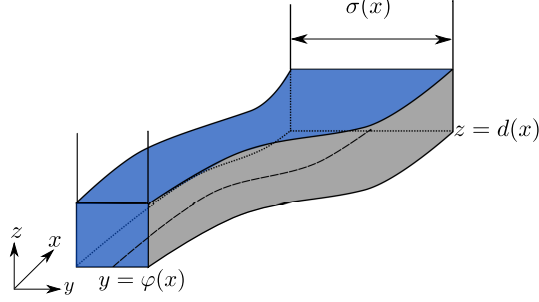


Figure 18.1: Geometric set-up.

The kinematic free surface condition and the wet boundary condition are the same, except the definition of the outward unit normal vector which is now

$$\mathbf{n}_{\text{wb}} = \begin{cases} \frac{1}{\sqrt{1 + (\partial_x d)^2}} (\partial_x d, 0, -1)^T & \text{if } \mathbf{n}_{\text{wb}} = \mathbf{n}_b \\ \frac{1}{\sqrt{1 + (\partial_x \alpha)^2}} (\partial_x \alpha, -1, 0)^T & \text{if } \mathbf{n}_{\text{wb}} = \mathbf{n}_{\text{lb}} \\ \frac{1}{\sqrt{1 + (\partial_x \beta)^2}} (\partial_x \beta, 1, 0)^T & \text{if } \mathbf{n}_{\text{wb}} = \mathbf{n}_{\text{rb}} \end{cases}$$

Thus, we have now

$$\begin{aligned} u \partial_x d - w &= 0 & \text{on } \Gamma_b(t), \\ u \partial_x \alpha - v &= 0 & \text{on } \Gamma_{\text{lb}}(t), \\ u \partial_x \beta + v &= 0 & \text{on } \Gamma_{\text{rb}}(t). \end{aligned}$$

Almost all the computations are the same to the exception of the computation of the $\int_{\Omega_{\text{eq}}(t,x)} a_4 dy dz$ which yields to different expression of the hydrostatic pressure source term (17.12) and the non hydrostatic pressure source term (17.15). Indeed, since we have now $d^*(x) = d(x)$ and $\sigma = \sigma(x)$, thus $A_{\text{eq}}(t, x) = \sigma(x) h_{\text{eq}}(t, x)$. Applying successively the one-dimensional Leibniz integral rule in the y -direction and in the z -direction, we obtain

$$\begin{aligned} \int_{\Omega_{\text{eq}}(t,x)} \partial_x P dy dz &= \partial_x \left(\int_{d(x)}^{\eta_{\text{eq}}(t,x)} \sigma(x) P(t, x, z) dz \right) \\ &- \int_{d(x)}^{\eta_{\text{eq}}(t,x)} P(t, x, z) \sigma'(x) dz + \sigma(x) P(t, x, d(x)) d'(x). \end{aligned}$$

Thus, in contrast with the non-rectangular case, the hydrostatic (17.12) and the non-hydrostatic (17.15) pressure source terms read

$$I_2(t, x) = \sigma'(x) \frac{h_{\text{eq}}(t, x)^2}{2F_r^2} - \sigma(x) \frac{h_{\text{eq}}(t, x)}{F_r^2} d'(x) \quad (18.1)$$

and

$$DI_2(t, x) = \int_{d(x)}^{\eta_{\text{eq}}(t, x)} P_{\text{nh}}(t, x, z) \sigma'(x) dz - \sigma(x) P_{\text{nh}}(t, x, d(x)) d'(x) . \quad (18.2)$$

Thus, the model (17.18) is the same as before but with the formula (18.1) and (18.2) instead of (17.12) and (17.15).

18.1 REFORMULATED MODEL FOR RECTANGULAR CASE

The main drawback of Equations (17.18) is that it has third order terms in space which may lead to instabilities at the numerical level. Therefore, we first propose a more stable formulation of Equations (17.18) before to present its numerical approximation.

Skipping the technical details, defining a linear operator \mathbb{L} (where \mathcal{L} is defined below)

$$\mathbb{L}[A, d, \sigma](u) = A \mathcal{L}[A, d, \sigma] \left(\frac{u}{A} \right) ,$$

one can show that System (17.18) can be written:

$$\begin{cases} \partial_t A + \partial_x(Au) = 0 \\ (I_d - \mu_2 \mathbb{L}[A, d, \sigma]) (\partial_t(Au) + \partial_x(Au^2)) \\ + \partial_x I_1(x, A) + \mu_2 A \mathcal{Q}[A, d, \sigma](u) = I_2(x, A) + O(\mu_2^2) \end{cases} \quad (18.3)$$

where I_d stands for the identity operator, \mathcal{L} is a linear operator

$$\begin{aligned} \mathcal{L}[A, d, \sigma](u) &= \frac{1}{A} [\partial_x(\overline{\mathcal{T}}[A, d, \sigma](u, \sigma)) - \overline{\mathcal{T}}[A, d, \sigma](u, \partial_x \sigma)] \\ &\quad + \frac{1}{A} \sigma(x) d'(x) \mathcal{T}[A, d, \sigma, z = d(x)](u) \end{aligned}$$

and \mathcal{Q} is a quadratic operator

$$\begin{aligned} \mathcal{Q}[A, d, \sigma](u) &= \frac{1}{A} [\partial_x(\overline{\mathcal{G}}[A, d, \sigma](u, \sigma)) - \overline{\mathcal{G}}[A, d, \sigma](u, \partial_x \sigma)] \\ &\quad + \frac{1}{A} \sigma(x) d'(x) \mathcal{G}[A, d, \sigma, z = d(x)](u) \end{aligned}$$

with \mathcal{T}, \mathcal{G} are given by

$$\mathcal{T}[A, d, \sigma, z](u) = \partial_x(u) \int_z^\eta \frac{S(x, s)}{\sigma(x)} ds + u \int_z^\eta \frac{1}{\sigma(x)} \partial_x S(x, s) ds ,$$

and

$$\begin{aligned} \mathcal{G}[A, d, \sigma, z](u) &= \int_z^\eta 2 (\partial_x u)^2 \frac{S(x, s)}{\sigma(x)} + \frac{u^2}{\sigma(x)} \left(\frac{\partial_x S(x, s) \partial_x \sigma(x)}{\sigma(x)} - \partial_x \partial_x S(x, s) \right) \\ &\quad + \partial_x \left(\frac{u^2}{2} \right) \frac{S(x, s) \partial_x \sigma(x)}{\sigma(x)^2} ds \end{aligned}$$

with

$$\bar{\mathcal{X}}[A, d, \sigma](u, \psi) = \int_{d(x)}^\eta \psi \mathcal{X}[A, d, \sigma, z](u) dz .$$

In particular, one can explicitly compute those operators:

- if $\sigma \in \mathbb{R}_*^+$ and $d \in \mathbb{R}$ are constant then we recover the standard one dimensional [SGN](#) equations (see for instance [\[95–97\]](#)) over flat bottom with

$$\mathcal{L}[A, d, \sigma](u) = \mathcal{L}_0[A, \sigma](u) = \frac{1}{\sigma h} \partial_x \left(\frac{\sigma h^3}{3} \partial_x u \right)$$

and

$$\mathcal{Q}[A, d, \sigma](u) = \mathcal{Q}_0[A, \sigma](u) = \frac{1}{\sigma h} \partial_x \left(\frac{2}{3} \sigma h^3 (\partial_x u)^2 \right) .$$

- if $\sigma \in \mathbb{R}_*^+$ is constant and $d = d(x)$ then we recover the standard one dimensional [SGN](#) equations (see for instance [\[95–97\]](#)) over uneven bottom with

$$\begin{aligned} \mathcal{L}[A, d, \sigma](u) &= \mathcal{L}_1[A, d, \sigma](u) = \mathcal{L}_0[A, \sigma](u) \\ &\quad - \frac{1}{\sigma h} \partial_x \left(\frac{\sigma h^2}{2} u d'(x) \right) + \frac{h}{2} \partial_x u d'(x) - u (d'(x))^2 \end{aligned}$$

and

$$\begin{aligned} \mathcal{Q}[A, d, \sigma](u) &= \mathcal{Q}_1[A, d](u) = \mathcal{Q}_0[A, \sigma](u) \\ &\quad + \frac{1}{\sigma h} \partial_x \left(\sigma \frac{h^2}{2} u^2 d''(x) \right) + h (\partial_x u)^2 d'(x) \\ &\quad + u^2 d'(x) d''(x) . \end{aligned}$$

- if $\sigma = \sigma(x)$ and $d = d(x)$ then we get the generalised one dimensional [SGN](#) equations for non uniform rectangular channel over uneven bottom with

$$\begin{aligned} \mathcal{L}[A, d, \sigma](u) &= \mathcal{L}_1[A, d, \sigma](u) + \frac{1}{\sigma h} \partial_x \left(\sigma'(x) \frac{h^3}{3} u \right) \\ &\quad - \frac{\sigma'(x)}{\sigma} \left(\partial_x u \frac{h^2}{3} + u \frac{h^2}{3} \frac{\sigma'(x)}{\sigma} - u \frac{h}{2} d'(x) \right) \end{aligned}$$

and

$$\begin{aligned} \mathcal{Q}[A, d, \sigma](u) &= \mathcal{Q}_1[A, d, \sigma](u) + \frac{1}{\sigma h} \partial_x \left((\sigma'(x))^2 \frac{u^2 h^3}{3} \right) \\ &\quad + \frac{1}{\sigma h} \partial_x \left(d'(x) \sigma'(x) u^2 \frac{h^2}{2} \right) - \frac{1}{\sigma h} \partial_x \left(\sigma'(x) u^2 \frac{h^3}{3} \right) \\ &\quad + \partial_x \left(\partial_x \left(\frac{u^2}{2} \right) \sigma'(x) \frac{h^3}{3} \right) - \frac{1}{\sigma h} \sigma'(x) \mathcal{R}[A, d, \sigma](u) \end{aligned}$$

with

$$\begin{aligned} \mathcal{R}[A, d, \sigma](u) &= (\partial_x u)^2 \frac{h^3}{3} + u^2 \left(\frac{\sigma'(x)}{\sigma} \right)^2 \frac{h^3}{3} + u^2 \left(\frac{\sigma'(x)}{\sigma} \right) d'(x) \frac{h^2}{2} \\ &\quad - u^2 \left(\frac{\sigma''(x)}{\sigma} \right)^2 \frac{h^3}{3} + u^2 d''(x) \frac{h^2}{2} \\ &\quad + \partial_x \left(\frac{u^2}{2} \right) \frac{\sigma'(x)}{\sigma} \frac{h^3}{3} - u^2 d'(x) \frac{\sigma'(x)}{\sigma} \frac{h^2}{2} \\ &\quad - u^2 \sigma'(x) (d'(x))^2 h + u^2 \sigma''(x) d'(x) \frac{h^2}{2} \\ &\quad - \partial_x \left(\frac{u^2}{2} \right) \sigma'(x) d'(x) \frac{h^2}{2}. \end{aligned}$$

It is known that third order derivatives involved in the initial model (17.18) may create high frequencies instabilities, but the presence of the $(I_d - \mu_2 \mathbb{L}[A, d, \sigma])^{-1}$ in the second equation of (18.3) stabilises the equations with respect to these perturbations. Therefore, in the following, we construct a numerical scheme for Equations (18.3) instead of Equations (17.18).

18.2 A WELL-BALANCED FINITE VOLUME APPROXIMATION

This section is devoted to the numerical method to solve the reformulated dispersive model (18.3). It is rather natural to split the hyperbolic part to the dispersive one as done by several authors (see for instance [22, 32, 33]).

Let $N \in \mathbb{N}^*$. Let us consider the following uniform mesh on $[0, L_c]$. Cells are denoted for every $i \in [0, N + 1]$, by $m_i = (x_{i-1/2}, x_{i+1/2})$ with $x_i = \frac{x_{i-1/2} + x_{i+1/2}}{2}$ the cell center and $\delta x = x_{i+1/2} - x_{i-1/2}$ the space mesh. The interfaces $x_{1/2} = 0$ and $x = x_{N+1/2}$ denote the upstream and the downstream ends. We also consider a time discretisation t_n defined by $t_{n+1} = t_n + \delta t_n$ where the time step δt_n is computed through a CFL condition related to the hyperbolic part.

Let us first highlight that the still water steady state for Equations (18.3) is independent of μ_2 . Indeed, one has $\forall \mu_2 > 0$, the still water steady state equation reads

$$u = 0, \frac{A}{\sigma} + d = h_0$$

for some positive h_0 . As a consequence, the construction of a well-balanced scheme can be easily achieved considering only the hyperbolic part of Equations (18.3), for instance, by the use of the hydrostatic reconstruction (see for instance [5]).

Let us define $d_{i+1/2} = \max(d_i, d_{i+1})$ where $d_i = \frac{1}{\delta x} \int_{m_i} d(x) dx$, $\sigma_{i+1/2} = \max(\sigma_i, \sigma_{i+1})$ where $\sigma_i = \frac{1}{\delta x} \int_{m_i} \sigma(x) dx$ and let us define the reconstructed states

$$A_{i+1/2}^- = \sigma_{i+1/2} \left(\frac{A_i}{\sigma_i} + d_i - d_{i+1/2} \right),$$

$$A_{i+1/2}^+ = \sigma_{i+1/2} \left(\frac{A_{i+1}}{\sigma_{i+1}} + d_{i+1} - d_{i+1/2} \right)$$

with

$$U_{i+1/2}^- = (A_{i+1/2}^-, A_{i+1/2}^- u_i), \quad U_{i+1/2}^+ = (A_{i+1/2}^+, A_{i+1/2}^+ u_{i+1})$$

where $U_i = (A_i, A_i u_i)^T \approx \frac{1}{\delta x} \int_{m_i} (A, Au)^T dx$.

Let us introduce the flux

$$F_1(U) = Q, \quad F_2(U) = Q^2 / A,$$

$$F_3(x, U) = I_1(x, A) + \mu_2 \bar{\mathcal{G}}[A, d, \sigma](u, \sigma)$$

and

$$\begin{aligned} \mathfrak{S}(x, U) &= I_2 + \mu_2 \bar{\mathcal{G}}[A, d, \sigma](u, \partial_x \sigma) \\ &\quad - \mu_2 \sigma(x) d'(x) \mathcal{G}[A, d, \sigma, z = d(x)](u). \end{aligned}$$

Then, one can write System (18.3) as follows:

$$\begin{aligned} \partial_t A + \partial_x F_1(U) &= 0 \\ (I_d - \mu_2 \mathbb{L}[A, d, \sigma]) (\partial_t Q + \partial_x F_2(U)) + \partial_x F_3(x, U) - \mathfrak{S}(x, U) &= 0 \end{aligned}$$

With these settings, we define the following numerical scheme:

$$\begin{aligned} A_i^{n+1} &= A_i^n - \frac{\delta t_n}{\delta x} \left(\mathcal{F}_1 \left(U_{i+1/2}^{-,n}, U_{i+1/2}^{+,n} \right) - \mathcal{F}_1 \left(U_{i-1/2}^{-,n}, U_{i-1/2}^{+,n} \right) \right) \\ Q_i^* &= Q_i^n - \frac{\delta t_n}{\delta x} \left(\mathcal{F}_2 \left(U_{i+1/2}^{-,n}, U_{i+1/2}^{+,n} \right) - \mathcal{F}_2 \left(U_{i-1/2}^{-,n}, U_{i-1/2}^{+,n} \right) \right) \\ Q_i^{n+1} &= Q_i^* - \frac{\delta t_n}{\delta x} (Y^n)_i \end{aligned}$$

where

$$\begin{aligned} A^n Y^n &= \left(\mathcal{F}_3 \left(x_{i+1/2}, U_{i+1/2}^{-,n}, U_{i+1/2}^{+,n} \right) \right. \\ &\quad \left. - \mathcal{F}_3 \left(x_{i-1/2}, U_{i-1/2}^{-,n}, U_{i-1/2}^{+,n} \right) + \mu_2 \mathcal{N}_i^n \right)_{1 \leq i \leq N}. \end{aligned}$$

The matrix \mathcal{A}^n is the cell-centered approximation of the linear operator $(I_d - \mu_2 \mathbb{L}[A, d, \sigma])$ and \mathcal{N}_i^n is the cell-centered approximation of $-\overline{\mathcal{G}}[A, d, \sigma](u, \partial_x \sigma) + \sigma(x) d'(x) \mathcal{G}[A, d, \sigma, z = d(x)](u)$.

The numerical fluxes are defined by

$$\begin{aligned} \mathcal{F}_1 \left(U_{i+1/2}^{-,n}, U_{i+1/2}^{+,n} \right) &= \frac{F_1(U_{i+1/2}^{-,n}) + F_1(U_{i+1/2}^{+,n})}{2} \\ &\quad - s_{i+1/2}^n (A_{i+1/2}^{+,n} - A_{i+1/2}^{-,n}) \\ \mathcal{F}_2 \left(U_{i+1/2}^{-,n}, U_{i+1/2}^{+,n} \right) &= \frac{F_2(U_{i+1/2}^{-,n}) + F_2(U_{i+1/2}^{+,n})}{2} \\ &\quad - s_{i+1/2}^n (Q_{i+1/2}^{+,n} - Q_{i+1/2}^{-,n}) \\ \mathcal{F}_3 \left(x_{i+1/2}, U_{i+1/2}^{-,n}, U_{i+1/2}^{+,n} \right) &= \frac{F_3(x_{i+1/2}, U_{i+1/2}^{-,n}) + F_3(x_{i+1/2}, U_{i+1/2}^{+,n})}{2} \\ &\quad + \left(\frac{A_i^{n2}}{2\sigma_i F_r^2} - \frac{A_{i+1/2}^{-,n2}}{2\sigma_{i+1/2} F_r^2} \right) \\ \mathcal{F}_3 \left(x_{i+1/2}, U_{i+1/2}^{-,n}, U_{i+1/2}^{+,n} \right) &= \frac{F_3(x_{i+1/2}, U_{i+1/2}^{-,n}) + F_3(x_{i+1/2}, U_{i+1/2}^{+,n})}{2} \\ &\quad + \left(\frac{A_{i+1}^{n2}}{2\sigma_{i+1} F_r^2} - \frac{A_{i+1/2}^{+,n2}}{2\sigma_{i+1/2} F_r^2} \right) \end{aligned}$$

such that whenever $\mu_2 = 0$, we recover the classical numerical scheme¹ for the hyperbolic part

$$\begin{aligned} U_i^{n+1} &= U_i^n - \frac{\delta t_n}{\delta x} \\ &\quad \times \left(\mathcal{F}(x_{i+1/2}, U_{i+1/2}^{-,n}, U_{i+1/2}^{+,n}) - \mathcal{F}(x_{i-1/2}, U_{i-1/2}^{-,n}, U_{i-1/2}^{+,n}) \right) \end{aligned}$$

with $\mathcal{F}(x, U, V) = (\mathcal{F}_1(U, V), \mathcal{F}_2(U, V) + \mathcal{F}_3(x, U, V))$.

In these expressions,

$$s_{i+1/2} = \max_{j=1,2} \left| \lambda_j(x_{i+1/2}, U_{i+1/2}^{-,n}) \right|, \left| \lambda_j(x_{i+1/2}, U_{i+1/2}^{+,n}) \right|$$

where $\lambda_j(x, U) = Q/A + (-1)^j \sqrt{\frac{A}{\sigma(x) F_r^2}}$, $j = 1, 2$ are the eigenvalues of the Jacobian matrix of $(F_1, F_2 + F_3)^T$.

The numerical scheme is consistent and stable under the CFL condition

$$\max_{1 \leq i \leq N} (|\lambda_1(x_i, U_i^n)|, |\lambda_2(x_i, U_i^n)|) \frac{\delta t_n}{\delta x} \leq 1.$$

¹ For the sake of simplicity and clarity, we have presented the finite volume method using the Rusanov solver but the method is not limited to this one.

18.3 NUMERICAL TEST CASE

In this section, we test the accuracy of the method and we show numerically the influence of the section variation in the case of the propagation of a solitary wave. For this purpose, we consider the exact solitary wave solutions of the Green-Naghdi equations in the one-dimensional setting over a flat bottom (see [96]), given in variables with dimensions, by

$$\begin{aligned} \eta(t, x) &= a \operatorname{sech}^2(k(x - ct)), & u(t, x) &= c \left(\frac{\eta(t, x)}{\eta(t, x) + z_0} \right) \\ \text{with } k &= \frac{\sqrt{3a}}{2z_0\sqrt{z_0+a}} \text{ and } c = \sqrt{g(z_0 + a)} \end{aligned} \quad (18.4)$$

where z_0 is the depth of the fluid and a is the relative amplitude.

ACCURACY: The propagation of the solitary wave (18.4) is initially centered at $x_0 = 10$ m with a relative amplitude $a = 0.2$ m over a constant water depth $z_0 = 1$ m. The computational domain is $L_c = 100$ m and it is discretized with N cells. The single solitary wave propagates from left to right. In this test, since the solitary wave is initially far from boundaries, the boundary conditions do not affect the computation, thus we choose to impose free boundary conditions at the downstream and upstream ends. The exact solution is computed in a channel of width $\sigma = 1$.

In what follows, we quantify the numerical accuracy of our numerical scheme by computing the numerical solution for this particular test case for an increasing number of cells N over a duration $T = 20$ s. Starting with $N = 100$ number of cells, we successively multiply the number of cells by two. For all n , we compare, in Figure 18.2, $M^n := \max_{0 \leq i \leq N+2} (h_i^n)$ of our numerical solution provided by Equations (18.3) with the exact one $M(t_n) := \max_{x \in [0, L_c]} h(t_n, x) = 2.2$ given by (18.4). One can easily remark that the first order discretisation is not accurate for long time simulation due to the numerical dissipation. However, to limit the numerical dissipation of the first order numerical scheme, one can either limit the simulation time or consider a very large number of cells. However, it is better to increase the order of the numerical scheme but this is left to a future work. Therefore, in what follows, we consider a shorter simulation time and a large number of cells, just to illustrate the influence of the variation of the channel.

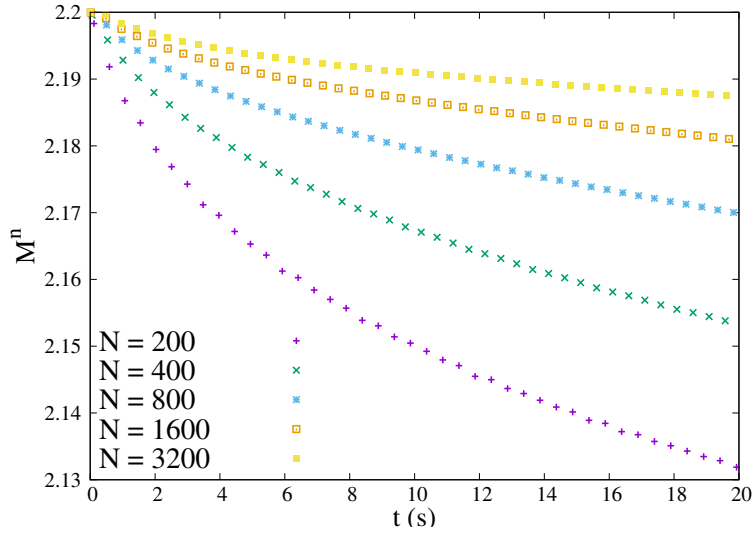


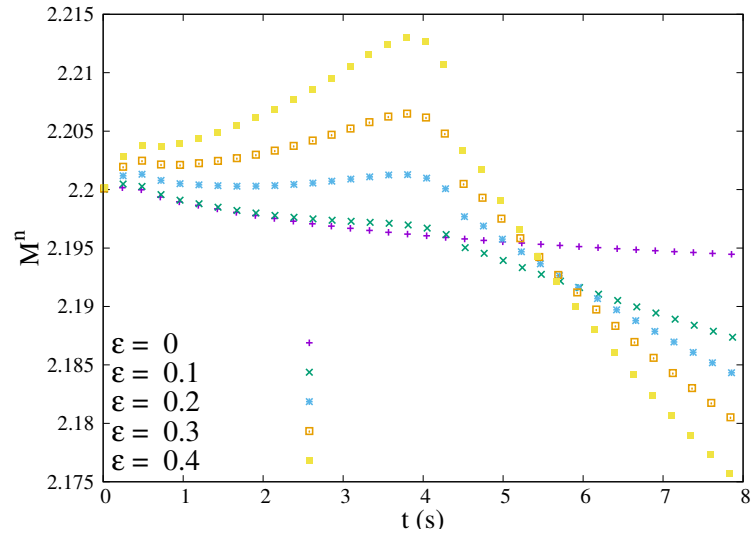
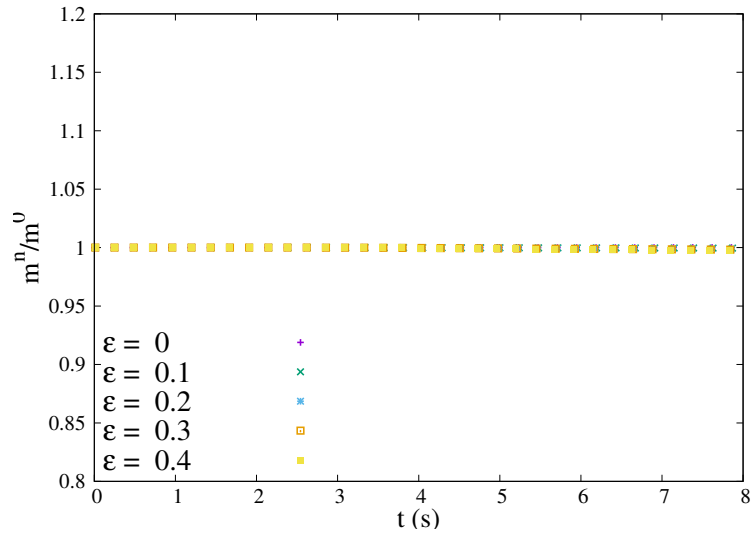
Figure 18.2: $M^n := \max_{0 \leq i \leq N+2}(h_i^n)$.

INFLUENCE OF THE SECTION VARIATION: We consider again the propagation of a solitary wave initially centered at $x_0 = 10$ m of relative amplitude $a = 0.2$ m, over a constant water depth $z_0 = 1$ m onto a computational domain of $L_c = 50$ m and discretized with $N = 5000$ cells. Initially starting with $(\eta(0, x), u(0, x))$ (see Equations (18.4)), we compute the numerical simulation for the channels defined by

$$\sigma(x; \varepsilon) = \beta(x; \varepsilon) - \alpha(x; \varepsilon)$$

$$\text{with } \beta = \frac{1}{2} - \frac{\varepsilon}{2} \exp\left(-\varepsilon^2 \left(x - L/2\right)^2\right) \text{ and } \alpha = -\beta$$

with $\varepsilon = 0$, $\varepsilon = 0.1$, $\varepsilon = 0.2$, $\varepsilon = 0.3$ and $\varepsilon = 0.4$. The obtained results are presented in Figure 18.3. In Figure 18.3a, for each geometry, we show the evolution of the maximum of the water level $M^n := \max_{0 \leq i \leq N+2}(h_i^n)$. As expected, since the first part for $x \leq 25$ is linearly converging, the water level increases while for $x > 25$, the channel is linearly diverging and therefore, the amplitude of the water level decreases. Moreover, in all numerical simulations, the mass is conserved. Indeed, for each value of ε , we have displayed in Figure 18.3b, the ratio of $\frac{m^n}{m^0}$ where $m^n := \frac{1}{N+2} \sum_{i=0}^{N+1} A_i^n$ is the mass of water at time t_n . The ratio $\frac{m^n}{m^0}$ is almost equal to 1, up to the order of accuracy of the numerical scheme.

(a) $M^n := \max_{x \in [0, L_c]} (h_i^n)$ (b) $\frac{m^n}{m^0}$ with $m^n = \frac{1}{N+2} \sum_{i=0}^{N+1} A_i^n$ Figure 18.3: Influence of σ .

In what follows, we quantify the numerical accuracy of our numerical scheme. Starting with $N = 100$ number of cells, we successively multiply the number of cells by two. The errors on the water surface deformation are presented in Table 18.1 and in Table 18.2. These errors are computed at $t = 8$ s using the L^2 :

$$\|\eta_{\text{num}} - \eta_{\text{ref}}\|_2 = \sqrt{\delta x \sum_i |\eta_{\text{num}_i}(t = 8) - \eta_{\text{ref}}(t = 8, x_i)|^2}$$

and the L^∞ norms where η_{ref} is the exact solution in the case $\varepsilon = 0$ and is a reference one computed with 10 000 cells for $\varepsilon = 0.4$. Since the results are almost the same whatever ε is, we have decided to present only the results for $\varepsilon = 0$ and $\varepsilon = 0.4$ in Table 18.1 and in Table 18.2.

As expected, the obtain numerical order is slow because of the numerical dissipation of the solitary wave (as already pointed out in several works, see for instance [22] for which we obtain almost the same order of convergence in the case of uniform section). Moreover, as expected, cross-sectional variations have no influence on the convergence rate. Let us just emphasise that the convergence rates are slightly better in the case of non-uniform section because we are comparing our results to a reference solution and not to the exact one.

| N | $\ \eta_{\text{num}} - \eta_{\text{exact}}\ _2$ | $\ \eta_{\text{num}} - \eta_{\text{exact}}\ _\infty$ |
|-------|-------------------------------------------------|------------------------------------------------------|
| 100 | 0.0789 | 0.0449 |
| 200 | 0.0497 | 0.0288 |
| 400 | 0.0304 | 0.0180 |
| 800 | 0.0198 | 0.0116 |
| 1600 | 0.0153 | 0.0081 |
| 3200 | 0.0138 | 0.0062 |
| Order | 0.53 | 0.58 |

Table 18.1: Convergence rate of the L^2 error for $\varepsilon = 0$. The order is computed through the first order interpolation polynomial

| N | $\ \eta_{\text{num}} - \eta_{\text{ref}}\ _2$ | $\ \eta_{\text{num}} - \eta_{\text{ref}}\ _\infty$ |
|-------|-----------------------------------------------|----------------------------------------------------|
| 100 | 0.05212 | 0.02533 |
| 200 | 0.02096 | 0.01082 |
| 400 | 0.01079 | 0.00554 |
| 800 | 0.00748 | 0.00503 |
| 1600 | 0.00635 | 0.00412 |
| 3200 | 0.00505 | 0.00300 |
| Order | 0.64 | 0.56 |

Table 18.2: Convergence rate of the L^2 error for $\varepsilon = 0.4$. The reference solution is computed with 10 000 cells. The order is computed through the first order interpolation polynomial

CONCLUDING REMARKS AND EXTENSIONS

In this part, we have presented the first non-linear and weakly dispersive model for natural rivers and channel hydraulics. The obtained model is a generalisation of the well-known one dimensional Serre-Green-Naghdi equation.

In the specific case of rectangular channels, we have proposed a stable reformulation of the model for numerical purpose. To this end, we have constructed a first order well-balanced finite volume numerical scheme for which we have proposed some simple computations to study the influence of the section variation. Our numerical scheme is stable under the classical CFL condition of the hyperbolic part of the equations. However, it is clear that, unless to use fine grids, the numerical scheme is too dissipative to capture correctly the dispersive waves.

On one hand, the natural extension is to increase the order of this numerical scheme, not only for the rectangular sections but for arbitrary ones. The accuracy can be improved by using Finite Volume adaptive methods as done in [Part iii](#) or Continuous Galerkin methods¹, or even better Discontinuous Galerkin methods (see [Part iv](#)), especially in the context of Dissipative SGN models, see for instance [[14](#), [67](#)].

On the other hand, at least for the performed numerical test case, it seems, that the operator is invertible. Theoretical mathematical and numerical framework are addressed in a forthcoming works.

¹ [[51](#)] demonstrated all the advantages and accuracy of the continuous Galerkin methods for the nonlinear dispersive wave propagation.

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