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Magnetic Thomas-Fermi theory for 2D abelian anyons

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Two-dimensional abelian anyons are, in the magnetic gauge picture, represented as fermions coupled to magnetic flux tubes. For the ground state of such a system in a trapping potential, we theoretically and numerically investigate a Hartree approximate model, obtained by restricting trial states to Slater determinants and introducing a self-consistent magnetic field, locally proportional to matter density. This leads to a fermionic variant of the Chern-Simons-Schrödinger system. We find that for dense systems, a semi-classical approximation yields qualitatively good results. Namely, we derive a density functional theory of magnetic Thomas-Fermi type, which correctly captures the trends of our numerical results. In particular, we explore the subtle dependence of the ground state with respect to the fraction of magnetic flux units attached to particles.

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The recognition that “anyons” with exotic exchange statistics (different from those of bosons and fermions) were a theoretical possibility in a 2D world historically came together with the observation that they would be equivalent to the attachment of magnetic flux to ordinary particles [31, 32, 41, 64, 65]. Nowadays, artificial magnetic flux attachment is the chief proposal to implement or simulate the behavior of 2D anyons, starting with the case of quasi-particles in the fractional quantum Hall effect [1, 39, 50] and continuing with more recent proposals using artificial gauge fields in a cold atoms context [9, 21, 63, 67, 68]. Experimental evidence was found in the quantum Hall context [6, 53] for 2D anyons and with cold atoms for 1D anyons [19, 24] (the latter being conceptually distinct from their 2D counterparts; see [12, 59, 60, 62]).

An immediate difficulty is that 2D anyons, being equivalent to quite peculiar composite fermions or bosons, are very difficult to study analytically even in the simplest of situations, with no other interactions than those coming from the flux attachment (“free” anyons). It is therefore important to be able to rely on good approximate models. Our purpose here is to propose a simple effective density functional theory for the ground state of 2D anyons in a trapping potential, inspired by the typical set-up of cold atoms experiments. This was studied before for “almost-bosonic” anyons [16, 17, 25, 27, 49, 63], where the flux attached to bosons is assumed small in the limit of large particle numbers, and for “almost-fermionic” anyons [13, 28, 29] for a similar situation but with fermion-based anyons instead. Here we aim at a description valid for more general flux, away from the almost-bosonic and almost-fermionic limits.

We start from a fermion-based magnetic gauge picture for N 2D anyons, i.e. from N fermions coupled to magnetic

flux tubes of intensity $2\pi\alpha \in [0, 2\pi[$, with α the exchange statistics parameter (counted from the fermionic end). To approximate ground-state properties at large N and fixed α , we perform a mean-field approximation. Available states are reduced to Slater determinants/quasi-free states, determined by a one-particle reduced density matrix. We solve for an approximate ground state by minimizing a Hartree functional with a self-consistent magnetic field, leading to Chern-Simons-Schrödinger-like variational equations. This general approach is akin to that of [35], where a Kohn-Sham method leads to a related model. We however fully take into account the contribution of the functional derivative of the Chern-Simons field to the variational equations. In the FQHE context, bosonic Chern-Simons-Schrödinger (CSS) was proposed in [70] and fermionic CSS in [46].

We numerically solve the fermionic Chern-Simons-Schrödinger variational problem and compare the results with those derived from a semi-classical/local density approximation. The latter is an instance of a magnetic Thomas-Fermi (mTF) theory, but with a self-consistent magnetic field.

The ground-state energy and density profiles in the local density approximation depend on α through the expression

$$\begin{aligned} c(\alpha) &:= \alpha + 2\alpha \lfloor \alpha^{-1} \rfloor - (\alpha \lfloor \alpha^{-1} \rfloor)^2 - \alpha^2 \lfloor \alpha^{-1} \rfloor \\ &= 1 + \alpha^2 (1 - \{ \alpha^{-1} \}) \{ \alpha^{-1} \}, \end{aligned}$$

where $\lfloor \alpha^{-1} \rfloor$ denotes the integer part of α^{-1} and

$$\{ \alpha^{-1} \} := \alpha^{-1} - \lfloor \alpha^{-1} \rfloor \in [0, 1[$$

its fractional part; see Figure 1 and compare with our numerical result in Figure 6.

Observe that for the more traditionally studied [11, 18, 22, 35, 40] cases of $\alpha = 1/n$ with integer n there is hence no dependence on n : $c(\alpha) = 1$. For more general values of α , which do not seem to have been thoroughly considered before, the dependence on α of ground-state energies and density profiles is quite small. Capturing it numerically is quite challenging:

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we tested systems of up to $N = 100$ particles, observing a satisfying qualitative agreement with theory, improving with increasing N . The dependence of direct-space densities on α stays quite small, leading us to suggest to look for signatures of anyon statistics in momentum-space densities instead, as put forward e.g. in [28] in the almost-fermionic limit $\alpha \rightarrow 0$. Such densities could be e.g. accessed by time-of-flight measurements in cold atomic gases [7, Section III.A].

We proceed to define the microscopic anyon and effective Hartree and CSS models more precisely in Sections 1 and 2, followed by a brief recollection of the almost-fermionic limit in Section 3. In Section 4, we introduce the mTF model, and in Section 5, we study it numerically. Finally, in Section 6, we provide some concluding remarks and outlooks for future research.

1. MAGNETIC GAUGE PICTURE FOR 2D ANYONS

We refer to the literature [38, 39, 47, 52, 56, 66] for general background, and in particular the link between (multi-valued) quantum wave functions with exchange phases $e^{i\pi(\alpha+1)}$ and the following model. We consider a gas of N fermions moving non-relativistically in a plane \mathbb{R}^2 , each coupled to a magnetic flux tube of intensity $2\pi\alpha$, and trapped in a confining (for example harmonic) external potential

$$V : \mathbb{R}^2 \mapsto \mathbb{R}^+.$$

Thus, the many-body Hamiltonian (defined as the Friedrichs extension of the associated quadratic form, and in units $\hbar = c = 2m = 1$)

$$H_N^\alpha := \sum_{j=1}^N \left[\left(-i\nabla_{\mathbf{x}_j} + \alpha \mathbf{A}(\mathbf{x}_j) \right)^2 + NV(\mathbf{x}_j) \right] \quad (1.1)$$

acts on the Hilbert space $L_{\text{asym}}^2(\mathbb{R}^{2N})$ of square-integrable complex-valued functions $\Psi_N(\mathbf{x}_1, \dots, \mathbf{x}_N)$ that are antisymmetric w.r.t. particle exchange, and we have defined the magnetic vector potential felt by particle j

$$\mathbf{A}(\mathbf{x}_j) := \sum_{k \neq j} \frac{(\mathbf{x}_j - \mathbf{x}_k)^\perp}{|\mathbf{x}_j - \mathbf{x}_k|^2} \quad (1.2)$$

corresponding to the magnetic field

$$B(\mathbf{x}_j) := \text{curl}_{\mathbf{x}_j} \mathbf{A}(\mathbf{x}_j) = 2\pi \sum_{k \neq j} \delta(\mathbf{x}_j - \mathbf{x}_k) \quad (1.3)$$

that we treat as a scalar field since it always points perpendicular to the plane. Note the convention to multiply V by the particle number N in (1.1). Effectively, this will set the length scale of the gas as $O(1)$ for large N . Indeed, the first kinetic term in (1.1) typically weighs $O(N^2)$ for N fermions in a bounded region (e.g. filling the Fermi disk, for $\alpha = 0$).

We study the ground-state problem

$$E(\alpha, N) := \inf \left\{ \langle \Psi_N | H_N^\alpha | \Psi_N \rangle_{L_{\text{asym}}^2(\mathbb{R}^{2N})} : \Psi_N \in L_{\text{asym}}^2(\mathbb{R}^{2N}), \|\Psi_N\|_{L^2} = 1 \right\}, \quad (1.4)$$

for which we propose effective models, aiming at approximations for macroscopic systems $N \rightarrow \infty$, at fixed parameter α (which we may restrict to the interval $[0, 1[$ by periodicity and complex conjugation). This contrasts with previous studies [13, 28, 29] valid in the almost-fermionic $\alpha \propto N^{-1/2}$ or almost-bosonic limits. The latter corresponds formally to $\alpha \rightarrow 1$ but is better understood [5, 25, 27, 48, 49] by studying the action of (1.1) on bosonic wave functions and letting $N \rightarrow \infty$, with $\alpha \propto N^{-1}$.

Although in the idealized model (1.3) the charge-flux composites are pointlike, realistic emergent anyons do have some extent set by the underlying parameters, and we will in the following consider their local averaged/smeared distributions (this is well known as the ‘‘average-field approximation’’ [66]). Thus, we observe that, as per (1.3), we are for fixed α and large N effectively in a regime with magnetic field $B \propto N$. For usual 2D fermions in a constant magnetic field, this is precisely the regime where magnetic Thomas-Fermi theory emerges as an effective model [23, 43–45, 57, 58, 69]. This motivates us to adapt to our situation the semi-classical approximations of the aforementioned references, although we will not pursue their mathematically rigorous justification here.

2. HARTREE APPROXIMATION AND CHERN-SIMONS-SCHRÖDINGER EQUATIONS

If we ignore for a moment the many-body character of the vector potential (1.2) and magnetic field (1.3), the energy functional to be minimized takes the form

$$\langle \Psi_N | H_N^\alpha | \Psi_N \rangle \simeq \text{Tr} [(-i\nabla + \alpha \mathbf{A})^2 \gamma] + N \int_{\mathbb{R}^2} V \rho_\gamma \quad (2.1)$$

in terms of the one-particle density matrix, obtained via a partial trace,

$$\gamma := N \text{Tr}_{2 \rightarrow N} |\Psi_N\rangle\langle\Psi_N|,$$

and the associated one-particle density

$$\rho_\gamma(\mathbf{x}) = \gamma(\mathbf{x}; \mathbf{x}),$$

where we identify γ with its integral kernel. Minimizing the above quadratic functional amongst fermionic wave functions leads to a Slater determinant

$$\Psi_N(\mathbf{x}_1, \dots, \mathbf{x}_N) = \frac{1}{\sqrt{N!}} \det (u_i(\mathbf{x}_j))_{1 \leq i, j \leq N}, \quad (2.2)$$

where u_1, \dots, u_N are orthonormal orbitals in the one-body Hilbert space $L^2(\mathbb{R}^2)$. Such a quasi-free state is entirely characterized by the associated density matrix which takes the form

$$\gamma(\mathbf{x}; \mathbf{y}) = \sum_{j=1}^N u_j(\mathbf{x}) \overline{u_j(\mathbf{y})}, \quad \text{i.e.} \quad \gamma = \sum_{j=1}^N |u_j\rangle\langle u_j|. \quad (2.3)$$

We propose to approximate the original N -body problem by reinstating the self-consistency of the vector potential (and

thus magnetic field) in (2.1), namely by setting

$$\mathbf{A}(\mathbf{x}) = \mathbf{A}[\rho_\gamma](\mathbf{x}) := \int_{\mathbb{R}^2} \frac{(\mathbf{x} - \mathbf{y})^\perp}{|\mathbf{x} - \mathbf{y}|^2} \rho_\gamma(\mathbf{y}) \, d\mathbf{y}, \quad (2.4)$$

so that the magnetic field be proportional to matter density, as in (1.3):

$$\operatorname{curl} \mathbf{A}[\rho_\gamma](\mathbf{x}) = 2\pi \rho_\gamma(\mathbf{x}). \quad (2.5)$$

This way we obtain the Hartree-like functional

$$\mathcal{E}_\alpha^H[\gamma] := \operatorname{Tr} \left((-i\nabla + \alpha \mathbf{A}[\rho_\gamma])^2 \gamma \right) + N \int_{\mathbb{R}^2} V \rho_\gamma \quad (2.6)$$

that we should minimize amongst all fermionic one-particle density matrices γ , leading to the energy

$$E^H(N, \alpha) := \inf \left\{ \mathcal{E}_\alpha^H[\gamma] : \gamma \in \mathfrak{S}^1(L^2(\mathbb{R}^2)), \right. \\ \left. 0 \leq \gamma \leq \mathbb{1}, \operatorname{Tr} \gamma = N, \gamma \text{ of rank } N \right\}, \quad (2.7)$$

where $\mathfrak{S}^1(L^2(\mathbb{R}^2))$ denotes the trace-class operators over $L^2(\mathbb{R}^2)$, and the operator constraint $\gamma \leq \mathbb{1} = \text{Identity Operator}$ reflects the Pauli principle. The constraints in (2.7) impose that γ is a projector, $\gamma^2 = \gamma$.

The first variation of the energy (2.6) with respect to an infinitesimal variation $\gamma = \gamma^H + \delta\gamma$ is $\operatorname{Tr}(H_{\gamma^H} \delta\gamma)$ with

$$H_{\gamma^H} := (-i\nabla + \alpha \mathbf{A}[\rho_{\gamma^H}])^2 - 2\alpha \nabla^\perp w * \mathbf{J}[\gamma^H] + NV, \quad (2.8)$$

where we denote

$$w(\mathbf{x}) = \log |\mathbf{x}|, \quad \nabla^\perp w(\mathbf{x}) = \frac{\mathbf{x}^\perp}{|\mathbf{x}|^2},$$

and

$$\mathbf{J}[\gamma^H] = \frac{1}{2} \sum_{j=1}^N (\bar{u}_j (-i\nabla + \alpha \mathbf{A}[\rho_{\gamma^H}]) u_j + \text{h.c.}) \quad (2.9)$$

the current operator associated to γ^H . The second term in (2.8) (where $*$ denotes a convolution product) is perhaps the less usual. It stems from the fact that in (2.6) the kinetic energy operator itself depends on the density.

For a minimizer γ^H of (2.6), the optimality condition stemming from the first variation of the energy leads to a self-consistent equation of the form

$$[\gamma^H, H_{\gamma^H}] = 0 \quad (2.10)$$

In other words, in terms of the orbitals u_1, \dots, u_N of γ^H we find

$$H_{\gamma^H} u_j = \lambda_j u_j$$

for eigenvalues $\lambda_j, j = 1 \dots N$ of the mean-field Hamiltonian H_{γ^H} . This is a generalization to several orbitals of the Chern-Simons-Schrödinger system. Indeed, restricting to a single orbital $N = 1$, we find an equation for a complex classical field $u: \mathbb{R}^2 \rightarrow \mathbb{C}$,

$$(-i\nabla + \alpha \mathbf{A}[|u|^2])^2 u - 2\alpha \nabla^\perp w * \mathbf{J}[u] u + V u = \lambda u,$$

with the self-generated couplings

$$\operatorname{curl} \mathbf{A}[|u|^2] = 2\pi |u|^2, \\ \mathbf{J}[u] = \frac{1}{2} \bar{u} (-i\nabla + \alpha \mathbf{A}[|u|^2]) u + \text{h.c.}$$

Moving to a time-dependent setting (i.e. replacing λ with $i\partial_t$) we may recognize the equations of motion for a Schrödinger field u coupled to a Chern-Simons gauge field \mathbf{A} in Coulomb gauge; see e.g. [20, 36] or [27, Remark 1.1]. We further mention [2–5, 48, 54] and references therein for recent literature and connections with anyons.

3. THE ALMOST-FERMIONIC LIMIT

For later comparison, we recall from [26, 28] the behavior of the Hartree minimization problem (2.7) in the almost-fermionic limit

$$N \rightarrow \infty, \quad \alpha = \frac{\beta}{\sqrt{N}}, \quad \beta \text{ fixed.} \quad (3.1)$$

For N 2D fermions in a region of size $O(1)$ one should keep in mind the heuristic $\mathbf{p} = -i\nabla \propto \sqrt{N}$, so that the above scaling of parameters formally sets all terms of (2.6) of the same order of magnitude $O(N^2)$. In this regime one finds that the following semi-classical version of (2.6) on the “position \times momentum” phase space becomes accurate:

$$\mathcal{E}^{\text{cla}}[m] = \frac{1}{(2\pi)^2} \iint_{\mathbb{R}^4} (|\mathbf{p} + \alpha \mathbf{A}[m](\mathbf{x})|^2 + NV(\mathbf{x})) m(\mathbf{x}, \mathbf{p}) \, d\mathbf{x} d\mathbf{p}, \quad (3.2)$$

where the argument is a positive measure $m(\mathbf{x}, \mathbf{p})$ on phase space, properly normalized and satisfying a form of the Pauli principle:

$$0 \leq m(\mathbf{x}, \mathbf{p}) \leq 1, \quad \iint_{\mathbb{R}^4} m = (2\pi)^2 N. \quad (3.3)$$

Further, we have defined its magnetic potential

$$\mathbf{A}[m](\mathbf{x}) := \frac{1}{(2\pi)^2} \iint_{\mathbb{R}^4} \frac{(\mathbf{x} - \mathbf{y})^\perp}{|\mathbf{x} - \mathbf{y}|^2} m(\mathbf{y}, \mathbf{p}) \, d\mathbf{y} d\mathbf{p}.$$

Modulo a regularization of the basic Hamiltonian (1.1) (extended anyons), this approximate model has been rigorously derived in [28, 29]. The results can hence serve as benchmarks for our numerical method (see Section 5 B below).

Minimizing the functional (3.2) leads to a minimizer of the form

$$m^{\text{cla}}(\mathbf{x}, \mathbf{p}) = \mathbb{1} \left(\left| \mathbf{p} + \alpha \mathbf{A}^{\text{TF}}(\mathbf{x}) \right|^2 \leq 4\pi \rho^{\text{TF}}(\mathbf{x}) \right), \quad (3.4)$$

where ρ^{TF} minimizes a bona-fide Thomas-Fermi (TF) energy:

$$\mathcal{E}^{\text{TF}}[\rho] := \int_{\mathbb{R}^2} (2\pi \rho^2 + NV\rho), \\ \mathcal{E}^{\text{TF}}[\rho^{\text{TF}}] = E^{\text{TF}} := \inf \left\{ \mathcal{E}^{\text{TF}}[\rho] : \rho \geq 0, \int_{\mathbb{R}^2} \rho = N \right\},$$

and its associated magnetic potential

$$\mathbf{A}^{\text{TF}}(\mathbf{x}) := \mathbf{A}[\rho^{\text{TF}}] = \int_{\mathbb{R}^2} \frac{(\mathbf{x} - \mathbf{y})^\perp}{|\mathbf{x} - \mathbf{y}|^2} \rho^{\text{TF}}(\mathbf{y}) d\mathbf{y}. \quad (3.5)$$

Hence,

$$\rho^{\text{TF}} = \frac{N}{4\pi} (\lambda^{\text{TF}} - V)_+ \quad (3.6)$$

for a Lagrange multiplier λ^{TF} that one determines from the mass constraint. Given V , one can insert (3.6) in (3.5) to numerically compute \mathbf{A}^{TF} and then the momentum-space density

$$\begin{aligned} t^{\text{TF}}(\mathbf{p}) &:= \frac{1}{(2\pi)^2} \int_{\mathbb{R}^2} m^{\text{cla}}(\mathbf{x}, \mathbf{p}) d\mathbf{x} \\ &= \frac{1}{(2\pi)^2} \int_{\mathbb{R}^2} \mathbb{1} \left(\left| \mathbf{p} + \alpha \mathbf{A}^{\text{TF}}(\mathbf{x}) \right|^2 \leq 4\pi \rho^{\text{TF}}(\mathbf{x}) \right) d\mathbf{x}. \end{aligned} \quad (3.7)$$

Note that, with a radial trap, the TF density is also radial and Newton's theorem gives

$$\mathbf{A}^{\text{TF}}(\mathbf{x}) = \frac{2\pi \mathbf{x}^\perp}{|\mathbf{x}|^2} \int_{r=0}^{|\mathbf{x}|} \rho^{\text{TF}}(r) r dr. \quad (3.8)$$

For the particular choice $V(\mathbf{x}) = |\mathbf{x}|^s$, $s > 0$ one finds

$$\lambda^{\text{TF}} = \left(\frac{4(s+2)}{s} \right)^{\frac{s}{s+2}},$$

and thus

$$\mathbf{A}^{\text{TF}}(\mathbf{x}) = \begin{cases} \frac{N\mathbf{x}^\perp}{4} \left(\lambda^{\text{TF}} - \frac{2}{s+2} |\mathbf{x}|^s \right) & \text{for } |\mathbf{x}|^s \leq \lambda^{\text{TF}}, \\ \frac{N\mathbf{x}^\perp}{|\mathbf{x}|^2} \lambda^{\text{TF}} & \text{for } |\mathbf{x}|^s \geq \lambda^{\text{TF}}. \end{cases} \quad (3.9)$$

As is usual with ordinary Thomas-Fermi theory (see also [13]), the position-space density (3.6) is independent of the magnetic field, and thus of α . Instead, the momentum-space density (3.7) should be used to look for signatures of anyon statistics, i.e. of the Chern-Simons-like gauge field. Indeed,

$$\tilde{t}^{\text{TF}}(\mathbf{p}) := N^{-1} t^{\text{TF}}(\sqrt{N}\mathbf{p}) \quad (3.10)$$

converges in the regime (3.1) to a non-trivial profile depending on the parameter β .

4. MAGNETIC THOMAS-FERMI APPROXIMATION

We turn to an approximation that should be valid more generally for the trapped anyon gas, namely the high-density limit

$$N \rightarrow \infty, \quad \alpha \text{ fixed}, \quad (4.1)$$

corresponding to $\beta \sim \sqrt{N} \rightarrow \infty$ in (3.1). This will be of magnetic Thomas-Fermi type, and just as usual fermionic Thomas-Fermi can be recovered as a limit case of mTF [43–45], the

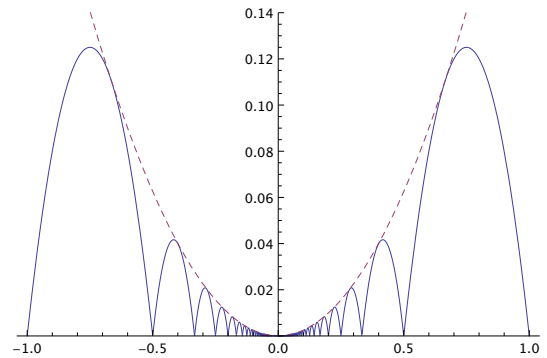


FIG. 1: Plot $\alpha \mapsto c(\alpha) - 1$ of the anyonic magnetic Thomas-Fermi constant (4.3), relative to the free fermion case, i.e. fermions at $\alpha = 0$ and bosons at $\alpha = 1$. Also shown (dashed) is the upper bound $\alpha^2/4$.

almost-fermionic description of the preceding section is a limit case of the considerations below.

We first describe the limit model briefly before presenting two parallel derivations in Sections 4 A and 4 B below. After eliminating momentum degrees of freedom similarly as above, the effective energy functional becomes

$$\mathcal{E}_\alpha^{\text{mTF}}[\rho] := \int_{\mathbb{R}^2} (2\pi c(\alpha) \rho(\mathbf{x})^2 + NV(\mathbf{x})\rho(\mathbf{x})) d\mathbf{x}, \quad (4.2)$$

to be minimized under the constraints

$$\rho \geq 0, \quad \int_{\mathbb{R}^2} \rho = N.$$

Here

$$\begin{aligned} c(\alpha) &= \alpha + 2\alpha[\alpha^{-1}] - (\alpha[\alpha^{-1}])^2 - \alpha^2[\alpha^{-1}] \\ &= 1 + \alpha^2 (1 - \{\alpha^{-1}\}) \{\alpha^{-1}\} \end{aligned} \quad (4.3)$$

can be interpreted as $(2\pi)^{-1} \rho^{-2}$ times the energy per particle of the homogeneous anyon gas; see (4.9) below.

We plot $c(\alpha) - 1$ in Figure 1, highlighting the quite non-trivial dependence on α and the reduction to the value for free fermions $c(\alpha) = 1$ for $\alpha = n^{-1}$, n integer or when $\alpha \rightarrow 0$ (this is the almost-fermionic limit of Section 3).

If, for concreteness, the external potential is set to be

$$V(\mathbf{x}) = |\mathbf{x}|^s, \quad s > 0, \quad (4.4)$$

then explicit minimization of (4.2) as above leads to the α -dependent densities and energies

$$\begin{aligned} \rho_\alpha^{\text{mTF}}(\mathbf{x}) &= \frac{N}{4\pi c(\alpha)} (\lambda_\alpha - |\mathbf{x}|^s)_+, \\ \lambda_\alpha &= \left(4c(\alpha) \frac{(s+2)}{s} \right)^{\frac{s}{s+2}}, \\ E_\alpha^{\text{mTF}} &= N^2 \frac{s}{8c(\alpha)(s+1)} \lambda_\alpha^{\frac{2s+2}{s}} \\ &= N^2 \frac{s}{8(s+1)} \left(4 \frac{s+2}{s} \right)^{\frac{2s+2}{s+2}} c(\alpha)^{\frac{s}{s+2}}. \end{aligned} \quad (4.5)$$

We compare these expressions to numerical simulations in Section 5 C. We first derive them using the local density approximation, and then return to the question of the distribution in momentum space using magnetic semi-classical measures. Thus, in brief, the present section aims at supporting the magnetic Thomas-Fermi approximation in the form of the following claim:

$$\boxed{\text{For fixed } \alpha, E^H(N, \alpha) \underset{N \rightarrow \infty}{\simeq} E_\alpha^{\text{mTF}}.} \quad (4.6)$$

Moreover, with γ^H a minimizer for (2.7), the spatial density

$$\rho^H(\mathbf{x}) := \gamma^H(\mathbf{x}, \mathbf{x})$$

should satisfy

$$\boxed{\rho^H \underset{N \rightarrow \infty}{\simeq} \rho_\alpha^{\text{mTF}}.} \quad (4.7)$$

A. Local density approximation

The most direct way to argue for (4.6) is via a local density approximation (LDA), which we expect to be valid because the magnetic field in the Hartree model is a local function of the density. This was previously discussed briefly in [47, Section III.D].

Consider a bounded reference domain $\Lambda \subset \mathbb{R}^2$ of unit area, and the thermodynamic limit of the energy density for the corresponding homogeneous problem at fixed average density $\rho = N/L^2$:

$$e(\alpha, \rho) = \lim_{N, L \rightarrow \infty, NL^2 = \rho} \frac{1}{L^2} E^H(N, \alpha, L\Lambda), \quad (4.8)$$

where for any subset $\Omega \subset \mathbb{R}^2$ we define the local Hartree energy analogously to (2.6)-(2.7):

$$E^H(N, \alpha, \Omega) := \inf \left\{ \mathcal{E}_{\alpha, \Omega}^H[\gamma] : \gamma \in \mathfrak{C}^1(L^2(\Omega)), \right. \\ \left. 0 \leq \gamma \leq 1, \text{Tr } \gamma = N, \gamma \text{ of rank } N \right\},$$

with

$$\mathcal{E}_{\alpha, \Omega}^H[\gamma] := \text{Tr}_{L^2(\Omega)} \left((-i\nabla + \alpha \mathbf{A}[\rho_\gamma])^2 \gamma \right).$$

Clearly, by scaling extensive quantities, we will have

$$e(\alpha, \rho) \propto \rho^2,$$

whence the point of scaling V in (4.12) below, to impose a strong density limit.

We shall argue that, with $c(\alpha)$ as in (4.3), we have

$$\boxed{e(\alpha, \rho) \simeq 2\pi c(\alpha) \rho^2.} \quad (4.9)$$

Our claim (4.6) can now be rephrased as

$$E^H(N, \alpha) \simeq E^{\text{LDA}}(N, \alpha), \quad (4.10)$$

with

$$E^{\text{LDA}}(N, \alpha) := \inf \left\{ \mathcal{E}^{\text{LDA}}[\rho] : \rho \geq 0, \int_{\mathbb{R}^2} \rho = N \right\} \quad (4.11)$$

and

$$\mathcal{E}^{\text{LDA}}[\rho] := \int_{\mathbb{R}^2} (e(\alpha, \rho(\mathbf{x})) + NV(\mathbf{x})\rho(\mathbf{x})) \, d\mathbf{x}. \quad (4.12)$$

Our discussion in favor of (4.6) therefore now relies entirely on arguing for (4.9), which we now proceed to do.

For simplicity, we discuss a fixed box of side length = 1 with $N = \rho \gg 1$ particles inside. We can reduce to this case by scaling length units. We assume that the density will be essentially homogeneous:

$$\rho(\mathbf{x}) \simeq \rho \equiv \text{constant},$$

so that the magnetic field becomes

$$B = \alpha \text{curl } \mathbf{A}[\rho] = 2\pi\alpha\rho \equiv \text{constant} \gg 1. \quad (4.13)$$

Thus, we think of the problem at hand as equivalent to that of minimizing the energy of ρ free fermions in Landau's Hamiltonian

$$H_B = (-i\nabla + \mathbf{A})^2 \text{ with } \text{curl } \mathbf{A} = B \equiv \text{constant}. \quad (4.14)$$

The energy levels of the above are essentially the Landau levels

$$E_k = 2B \left(k + \frac{1}{2} \right), \quad k = 0, 1, 2, \dots \quad (4.15)$$

(recall that $\hbar = e = c = 1$ but the mass $m = 1/2$). These have degeneracy proportional to area times

$$\frac{B}{2\pi} = \alpha\rho. \quad (4.16)$$

Indeed, the magnetic length of the field is, in our units,

$$\ell_B := \frac{1}{\sqrt{B}} \ll 1 = \text{size of the box}, \quad (4.17)$$

which will ensure that we may safely approximate the energy levels by those of the full plane model, modulo taking the degeneracy (4.16) into account. Reviews of these facts are found e.g. in [57, Sections 2 and 3], [37, Section 3] or [61].

We set $n := \lfloor \alpha^{-1} \rfloor$ and write

$$\rho = n\alpha\rho + \rho^{\text{top}}, \quad \rho^{\text{top}} := (1 - n\alpha)\rho = \{\alpha^{-1}\} \alpha\rho,$$

to see that, taking the degeneracy (4.16) into account, we minimize the energy by completely filling the n lowest Landau levels (with $\alpha\rho$ particles per level) and putting the remaining $\rho^{\text{top}} < \alpha\rho$ fermions in a last, partially filled, level. Hence, an integer number n of levels are precisely filled if $\rho^{\text{top}} = 0$, i.e. if and only if the fractional part $\{\alpha^{-1}\} = 0$.

Now, summing the level energies (4.15) times the number of fermions occupying them, we find

$$\begin{aligned}
e(\alpha, \rho) &= \sum_{k=0}^{n-1} 2B \left(k + \frac{1}{2} \right) \alpha \rho + 2B \left(n + \frac{1}{2} \right) (1 - n\alpha) \rho \\
&= \sum_{k=0}^{n-1} 4\pi\alpha^2 \rho^2 \left(k + \frac{1}{2} \right) + 4\pi\alpha \rho^2 \left(n + \frac{1}{2} \right) (1 - n\alpha) \\
&= 2\pi\alpha \rho^2 + 4\pi\alpha^2 \rho^2 \sum_{k=0}^{n-1} k + 4\pi\alpha \rho^2 (1 - n\alpha) n \\
&= 2\pi \rho^2 \left(\alpha + n(n-1)\alpha^2 + 2\alpha n - 2\alpha^2 n^2 \right),
\end{aligned}$$

and thus the claim (4.9) follows.

B. Semi-classical measures

We turn to a more semi-classical point of view on our Claim (4.6). This will explain in which sense magnetic Thomas-Fermi theory should be expected to predict the behavior of Hartree minimizers. Indeed, implicit in the local density approximation above, is a magnetic semi-classical phase space [23, 43–45, 57, 58, 69] of the form “position of cyclotron orbit \times Landau level index”, instead of the more standard “position \times momentum” of Section 3 above. We first recall this in the context of free fermions in a large, constant, magnetic field, before turning to adaptations to the present context of 2D anyons. The general picture is derived from the well-known [10, 30, 61] classical motion in a large magnetic field, split into a cyclotron motion (whose radius, quantum mechanically, is quantized proportionally to \sqrt{n} , with n integer) and a drift of the orbit center \mathbf{R} .

For free fermions in a constant magnetic field, the one-body Hamiltonian (4.14) in the symmetric gauge is

$$H_B = (-i\nabla_{\mathbf{x}} + \mathbf{A}(\mathbf{x}))^2, \quad \mathbf{A}(\mathbf{x}) = \frac{B}{2} \mathbf{x}^\perp. \quad (4.18)$$

One can define a coherent operator by its integral kernel

$$\Pi_{n,\mathbf{R}}^B(\mathbf{x}, \mathbf{y}) := g_\varepsilon(\mathbf{x} - \mathbf{R}) \Pi_n^B(\mathbf{x}, \mathbf{y}) g_\varepsilon(\mathbf{y} - \mathbf{R}), \quad (4.19)$$

where

$$\begin{aligned}
\Pi_n^B(\mathbf{x}, \mathbf{y}) &= \frac{B}{2\pi} \exp \left(iB(x_1 y_2 - x_2 y_1) - B \frac{|\mathbf{x} - \mathbf{y}|^2}{4} \right) \\
&\quad \times L_n \left(B \frac{|\mathbf{x} - \mathbf{y}|^2}{2} \right), \quad (4.20)
\end{aligned}$$

in terms of Laguerre polynomials

$$L_n \left(B \frac{|\mathbf{x} - \mathbf{y}|^2}{2} \right) = \sum_{k=0}^n \binom{n}{k} \frac{(-1)^k}{k!} \left(B \frac{|\mathbf{x} - \mathbf{y}|^2}{2} \right)^k,$$

is the integral kernel of the orthogonal projector on the n -th Landau level (with energy (4.15)), and g_ε is a localization function normalized in L^2 , e.g.

$$g_\varepsilon(\mathbf{x}) = c_\varepsilon e^{-|\mathbf{x}|^2/\varepsilon^2}. \quad (4.21)$$

One typically asks

$$B^{-1/2} \ll \varepsilon \ll 1, \quad (4.22)$$

i.e. ε is larger than the magnetic length but smaller than the typical system size. One should think of the operator $\Pi_{n,\mathbf{R}}$ as localizing close to the n -th Landau level and close to the position \mathbf{R} (i.e., classically, close to a cyclotron orbit of radius $\propto \sqrt{n}$ around the guiding center \mathbf{R}).

Alternatively, one can define

$$\Pi_{n,\mathbf{R}}^B(\mathbf{x}, \mathbf{y}) := \frac{B}{2\pi} \psi_{n,\mathbf{R}}^B(\mathbf{x}) \overline{\psi_{n,\mathbf{R}}^B(\mathbf{y})}, \quad (4.23)$$

with the vortex coherent state (z and Z are complex numbers corresponding to \mathbf{x} and \mathbf{R} respectively)

$$\psi_{n,\mathbf{R}}^B(\mathbf{x}) := \frac{i^n \sqrt{B}}{\sqrt{2\pi n!}} \left(\sqrt{\frac{B}{2}} (z - Z) \right)^n e^{-B \frac{|\mathbf{x} - \mathbf{R}|^2}{4} - \frac{Bi}{2} \mathbf{R}^\perp \cdot \mathbf{x}}, \quad (4.24)$$

but we will mostly rely on (4.19) in the sequel.

Given a one-particle density matrix for N fermions

$$0 \leq \gamma \leq \mathbb{1}, \quad \text{Tr } \gamma = N, \quad (4.25)$$

we can form the associated Husimi function

$$m(n, \mathbf{R}) := \text{Tr} \left[\Pi_{n,\mathbf{R}}^B \gamma \right], \quad (4.26)$$

with $\Pi_{n,\mathbf{R}}^B$ set as either (4.19) or (4.23). Since

$$\Pi_n^B(\mathbf{x}, \mathbf{x}) \equiv \frac{B}{2\pi},$$

one deduces from (4.25) a semi-classical version of the Pauli principle:

$$0 \leq m(n, \mathbf{R}) \leq \frac{B}{2\pi}, \quad (4.27)$$

and since

$$\sum_{n \geq 0} \Pi_n^B(\mathbf{x}, \mathbf{y}) = \delta(\mathbf{x} - \mathbf{y}),$$

it follows that

$$\sum_{n \geq 0} \int_{\mathbb{R}^2} m(n, \mathbf{R}) d\mathbf{R} = N. \quad (4.28)$$

In view of the localization properties of $\Pi_{n,\mathbf{R}}^B$ on the $(n, \mathbf{R}) \in \mathbb{N} \times \mathbb{R}^2$ phase space, one approximates the quantum energy by the semiclassical one in the manner

$$\begin{aligned}
\mathcal{E}^q[\gamma] &:= \text{Tr} \left((H_B + NV) \gamma \right) \\
&\simeq \sum_{n \geq 0} \int_{\mathbb{R}^2} \left(2B \left(n + \frac{1}{2} \right) + NV(\mathbf{R}) \right) m(n, \mathbf{R}) d\mathbf{R} =: \mathcal{E}^c[m]. \quad (4.29)
\end{aligned}$$

In minimizing the right-hand side of the above, one can eliminate the Landau level variable n by setting

$$\varrho(\mathbf{R}) := \sum_{n \geq 0} m(n, \mathbf{R}) \simeq \gamma(\mathbf{R}, \mathbf{R}). \quad (4.30)$$

Solving next for ϱ as previously, and defining

$$n_\star(\mathbf{R}) := \left\lfloor \frac{2\pi\varrho(\mathbf{R})}{B} \right\rfloor$$

the number of filled levels, one finds

$$m(n, \mathbf{R}) = \begin{cases} \frac{B}{2\pi}, & \text{for } 0 \leq n \leq n_\star(\mathbf{R}) - 1, \\ \frac{B}{2\pi} \left(\frac{2\pi\varrho(\mathbf{R})}{B} - n_\star(\mathbf{R}) \right), & \text{for } n = n_\star(\mathbf{R}), \\ 0, & \text{for } n > n_\star(\mathbf{R}). \end{cases} \quad (4.31)$$

In the context of 2D anyons in the Hartree approximation of Section 2, we follow the previous steps mutatis mutandis. We only replace any occurrence of the fixed, constant, magnetic field B by

$$B(\mathbf{R}) := 2\pi\alpha\rho_\gamma(\mathbf{R}), \quad (4.32)$$

with ρ_γ the system's spatial density, in accordance with (2.5).

We start from the Hartree energy

$$\mathcal{E}_\alpha^H[\gamma] := \text{Tr} \left(\left(\tilde{H}_\gamma + NV \right) \gamma \right)$$

with

$$\tilde{H}_\gamma := (-i\nabla + \alpha\mathbf{A}[\rho_\gamma])^2$$

and a minimizer γ^H thereof, with density ρ^H . We form the Husimi function in the manner

$$m^H(n, \mathbf{R}) := \text{Tr} \left[\Pi_{n, \mathbf{R}}^{B(\mathbf{R})} \gamma^H \right], \quad (4.33)$$

with $\Pi_{n, \mathbf{R}}^{B(\mathbf{R})}$ as in (4.19) or (4.23), and

$$B(\mathbf{R}) \simeq 2\pi\alpha \sum_{n \geq 0} m^H(n, \mathbf{R}),$$

in accordance with (4.32). We then expect $\mathcal{E}_\alpha^H[\gamma]$ to be approximated by

$$\begin{aligned} \mathcal{E}_\alpha^c[m] &:= \sum_{n \geq 0} \int_{\mathbb{R}^2} \left(4\pi\alpha\rho_\gamma(\mathbf{R}) \left(n + \frac{1}{2} \right) + NV(\mathbf{R}) \right) m(n, \mathbf{R}) d\mathbf{R}, \\ \rho_\gamma(\mathbf{R}) &= \sum_{n \geq 0} m(n, \mathbf{R}), \end{aligned} \quad (4.34)$$

which should be minimized under the constraints

$$m(n, \mathbf{R}) \leq \alpha\rho_\gamma(\mathbf{R}), \quad \sum_{n \geq 0} \int_{\mathbb{R}^2} m(n, \mathbf{R}) d\mathbf{R} = N.$$

An explicit solution leads to a literal translation of (4.31):

$$m^H(n, \mathbf{R}) \simeq \rho^H(\mathbf{R}) \times \begin{cases} \alpha, & \text{for } 0 \leq n \leq \lfloor \alpha^{-1} \rfloor - 1, \\ \alpha \{ \alpha^{-1} \}, & \text{for } n = \lfloor \alpha^{-1} \rfloor, \\ 0, & \text{for } n > \lfloor \alpha^{-1} \rfloor. \end{cases} \quad (4.35)$$

The dependence on n in the above should be thought of as the distribution within local Landau levels leading to the energy (4.2). Approximating ρ^H by ρ_α^{TF} gives an explicit formula, and the final proposed approximation for the full ground-state density matrix is

$$\gamma^H \simeq \sum_{n \geq 0} \int_{\mathbb{R}^2} m^H(n, \mathbf{R}) \Pi_{n, \mathbf{R}}^{B(\mathbf{R})} d\mathbf{R}, \quad (4.36)$$

from which one can extract relevant physical quantities.

5. NUMERICAL SOLUTIONS

We turn to the direct numerical minimization of the Hartree energy functional (2.6)-(2.7) from Section 2, and the comparison of the results with the expectations derived in Sections 3 and 4. The theoretical results in the almost-fermionic limit should be considered much more solidly grounded than those for general α involving mTF. We hence think of their reproduction by the simulations in Section 5 B below as a benchmark for the numerical method, giving confidence for the more challenging simulations of Section 5 C.

A. Methods

We minimize the functional (2.6) as a function of the orbitals u_j , $j = 1, 2, \dots, N$ using a small adaptation of the code DFTK [34]. This code was originally designed to solve the Kohn-Sham equations for electrons, but has an architecture flexible enough to accommodate (2.6) easily. We discretize the orbitals using a spectral method in a large box with periodic boundary conditions. We use a standard Riemannian L-BFGS method on the Stiefel manifold to perform the minimization [8].

The numerically most sensitive part of the scheme is the solution of the system

$$\text{curl } \mathbf{A} = 2\pi\rho, \quad \text{div } \mathbf{A} = 0, \quad \rho = \sum_{j=1}^N |u_j|^2.$$

Solving this equation naively in the Fourier domain imposes an artificial periodicity of \mathbf{A} which is incompatible with its slow decay at infinity, leading to very slow convergence with respect to the box size. This problem is strictly analogous to that encountered in density functional theory simulations of electrons in charged cells (e.g. ions), where the Coulomb potential is long-ranged. To correct for this, we analytically solve the equation

$$\text{curl } \mathbf{A}_{\text{ref}} = 2\pi\rho_{\text{ref}}$$

for a gaussian ρ_{ref} of reasonable width and of total integral equal to that of ρ . Then, we solve for the correction

$$\text{curl } \mathbf{A}_{\text{diff}} = 2\pi(\rho - \rho_{\text{ref}})$$

by a Fourier method, assuming periodic boundary conditions for \mathbf{A}_{diff} , and reconstruct the total potential

$$\mathbf{A} = \mathbf{A}_{\text{ref}} + \mathbf{A}_{\text{diff}}$$

in real space. Since ρ is expected to be roughly spherical, \mathbf{A}_{diff} decays quickly at infinity, and this method is found to yield rapid convergence with respect to the box size in practice.

The Thomas-Fermi solution (3.6) can be used to provide rough guesses of the box and grid size. In particular, in the scaling we use, the width of the Thomas-Fermi solution does not scale with N , but its momentum-space width does. We use this to guide the selection of numerical parameters in our numerical experiments, by taking a box size equal to $2q_x$ times the Thomas-Fermi real-space width, and a grid size equal to $2\pi q_p$ over the momentum-space width (which scales with N). The parameters q_x and q_p are quality factors. Due to the spectral accuracy of our numerical scheme and the strong locality of ρ and \mathbf{A}_{diff} , the convergence with respect to both q_x and q_p is very fast; we found $q_x = 2$, $q_p = 6$ to give visually converged results for the values of N reported.

A further key technical point concerns the accuracy of the radially averaged densities we plot below. It is guaranteed by performing the averaging exactly, i.e. using the analytical expression for the average of a plane wave (the average over a circle of radius r of $e^{iq \cdot \mathbf{x}}$ is given in terms of Bessel functions as $J_0(|q|r)$).

B. Almost-fermionic regime

We first check the agreement between our numerical method and the theoretical findings recalled in Section 3. Some preliminary checks appeared in [26, Section 1.5.4]. Our new simulations confirm those.

We solve for N orthonormal orbitals u_1, \dots, u_N and form the associated density matrix

$$\gamma(\mathbf{x}, \mathbf{y}) = \sum_{j=1}^N u_j(\mathbf{x}) \overline{u_j(\mathbf{y})}.$$

In the almost-fermionic limit, the densities in position space

$$\rho(\mathbf{x}) = \gamma(\mathbf{x}, \mathbf{x}) = \sum_{j=1}^N |u_j(\mathbf{x})|^2 \quad (5.1)$$

and momentum space (with \hat{u} the Fourier transform of u)

$$t(\mathbf{p}) = \sum_{j=1}^N |\hat{u}_j(\mathbf{p})|^2 \quad (5.2)$$

should both be correctly approximated by the Thomas-Fermi description from Section 3. We plot their average over rotations after dividing by N , and rescaling \mathbf{p} by a factor $N^{-1/2}$ in momentum space (recall that (3.10) is expected to converge for large N). Comparisons are then made with the explicit (3.6) (insert $c(\alpha) = 1$ in (4.5)) and semi-explicit (3.7) (where the

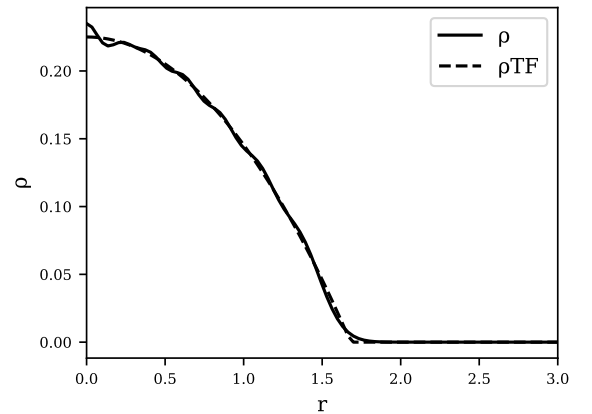


FIG. 2: Position density (normalized), $N = 100$, $\alpha = \frac{1.5}{\sqrt{N}} = 0.15$, $V(\mathbf{x}) = |\mathbf{x}|^2$.

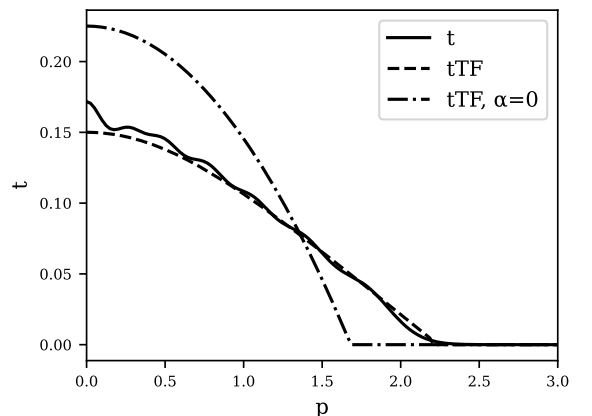


FIG. 3: Momentum density (rescaled), $N = 100$, $\alpha = \frac{1.5}{\sqrt{N}} = 0.15$, $V(\mathbf{x}) = |\mathbf{x}|^2$.

integrations involving the TF position density are performed numerically using a Monte Carlo method [33, 42]). Recall that the expected position density is independent of α in this regime. For reference we also plot the momentum density one would obtain from ordinary TF theory, i.e. at $\alpha = 0$.

We provide two typical sets of plots:

- Figures 2 and 3 correspond to $N = 100$ particles in a harmonic trap

$$V(\mathbf{x}) = |\mathbf{x}|^2 \quad (5.3)$$

with

$$\alpha = 1.5N^{-1/2} = 0.15.$$

Note that in a harmonic trap, the usual TF momentum and position densities (i.e. at $\alpha = 0$) are identical.

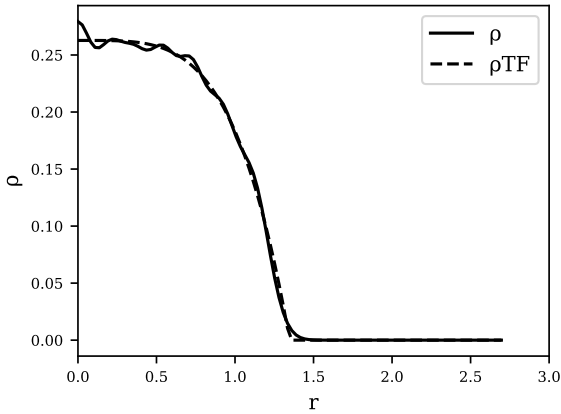


FIG. 4: Position density (normalized), $N = 100$, $\alpha = \frac{1.5}{\sqrt{N}} = 0.15$, $V(\mathbf{x}) = |\mathbf{x}|^4$.

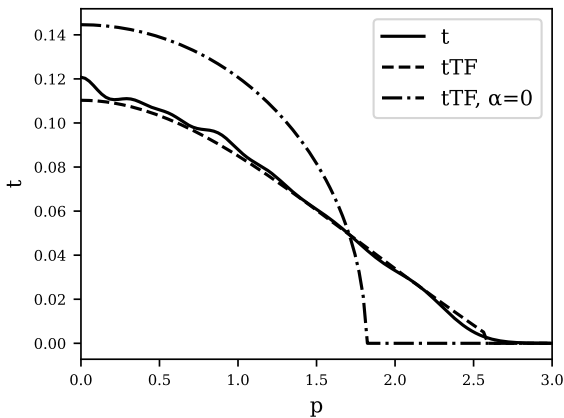


FIG. 5: Momentum density (rescaled), $N = 100$, $\alpha = \frac{1.5}{\sqrt{N}} = 0.15$, $V(\mathbf{x}) = |\mathbf{x}|^4$.

- Figures 4 and 5 correspond to $N = 100$ particles in a quartic trap

$$V(\mathbf{x}) = |\mathbf{x}|^4$$

with again $\alpha = 0.15$.

In both cases we find a rather good agreement between the theoretical expectations and the numerical simulations, in particular in view of the fact that the numerical densities result from averaging a sum of many rather oscillating profiles over rotations. The agreement found already occurs for a “moderately small” value of α , indicating a rapid convergence of the Hartree minimization problem to the almost-fermionic Thomas-Fermi one for small α . This could be guessed from the rapid convergence

$$c(\alpha) \xrightarrow{\alpha \rightarrow 0} 1$$

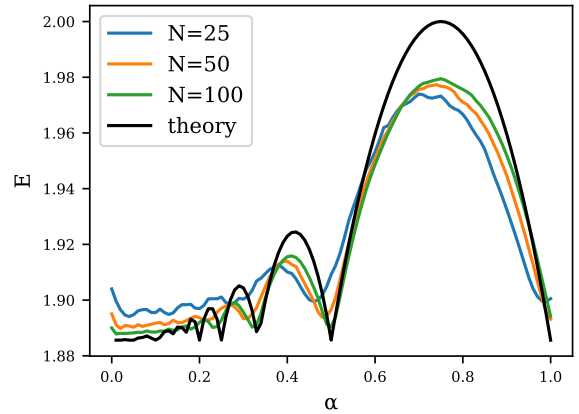


FIG. 6: Ground-state energy (normalized) in a harmonic potential as a function of the statistics parameter α (fermions for $\alpha = 0$), for different values of particle number N , and compared to the theoretical value $E_\alpha^{\text{mTF}}/N^2$ in (4.5).

apparent in Figure 1, and is in agreement with previous simulations [35]. As anticipated, the position-space density is indistinguishable from the purely fermionic case $\alpha = 0$, $c(\alpha) = 1$. The momentum-space density is much more interesting, and markedly differs from that expected for free fermions, thus showing a clear signature of anyonic statistics via the interaction with the Chern-Simons-like gauge field.

C. Magnetic Thomas-Fermi regime

We now move away from the almost-fermionic limit to check the prediction of mTF theory at general α . Inserting reasonable numbers in (4.5) immediately reveals that, as regards energies and position-space densities, we are looking for relatively small effects. Roughly (cf. again the relatively small variations in Figure 1) at the level of energies, one looks for variations in the 5% range. We thus have to perform rather challenging simulations (in terms of particle number, grid size and computation time). We estimate the precision of our numerical results as circa 1%.

First, in Figure 6 we plot the numerically evaluated energies for $N = 25$, $N = 50$ and $N = 100$ particles in a harmonic trap, as a function of α . We find a systematic numerical error compared to the theory (4.5), but the expected oscillations are correctly reproduced, in particular the absolute maximum at $\alpha = 3/4$ expected from (4.3) and Figure 1.

Note that

- the agreement becomes noticeably better for larger α , deeper in the mTF regime. This is also the most challenging situation, since the self-consistent magnetic field is larger.
- the agreement improves when increasing N , as it should. This occurs rather slowly, and hence we have

not pushed the numerics further than the already relatively costly $N = 100$.

Next, in Figures 7 and 8 we plot the position space and momentum space densities obtained for $N = 25$, $N = 50$ and $N = 100$ particles in a harmonic trap (5.3) at $\alpha = 3/4$. We again observe a correct match between numerical position space densities and mTF theory, but the oscillations of the former make it hard to argue for a clear signature of a $c(\alpha) \neq 1$. The trend with N is however encouraging: around $N = 100$, the mTF density (solid black line) starts to approximate the numerical plots better than the usual TF density (dashed black line).

The momentum space densities show, perhaps, our most interesting finding. There is a priori no obvious theory to compare them to, for mTF theory is not based on the usual “position \times momentum” phase space. We however observe that some extrapolation of the TF momentum density fits the numerical curves pretty well. Namely, this is the semi-explicit formula (3.7) where we replace ρ^{TF} by ρ_α^{mTF} (given by (4.5)–(4.3) at $\alpha = 3/4$). This leads to the expression

$$t_\alpha^{\text{mTF}}(\mathbf{p}) = \frac{1}{(2\pi)^2} \int_{\mathbb{R}^2} \mathbb{1} \left(\left| \mathbf{p} + \alpha \mathbf{A}_\alpha^{\text{mTF}}(\mathbf{x}) \right|^2 \leq 4\pi \rho_\alpha^{\text{mTF}}(\mathbf{x}) \right) d\mathbf{x} \quad (5.4)$$

with

$$\mathbf{A}_\alpha^{\text{mTF}}(\mathbf{x}) := \mathbf{A}[\rho_\alpha^{\text{mTF}}] = \int_{\mathbb{R}^2} \frac{(\mathbf{x} - \mathbf{y})^\perp}{|\mathbf{x} - \mathbf{y}|^2} \rho_\alpha^{\text{mTF}}(\mathbf{y}) d\mathbf{y}. \quad (5.5)$$

and thus, if $V(\mathbf{x}) = |\mathbf{x}|^s$,

$$\mathbf{A}_\alpha^{\text{mTF}}(\mathbf{x}) = \frac{N\mathbf{x}^\perp}{4c(\alpha)} \left(\left(4c(\alpha) \frac{(s+2)}{s} \right)^{\frac{s}{s+2}} - \frac{2}{s+2} |\mathbf{x}|^s \right) \quad (5.6)$$

on the support of ρ_α^{mTF} . Interestingly, we find that this use of ρ_α^{mTF} instead of ρ^{TF} leads to better fits of the numerically computed densities. These are shown in Figure 8. Note the non-monotonic shape of the curve, which marks the larger intensity of the effective magnetic field, and is a distinctive feature of the mTF regime.

Finally, in Figure 9, we plot the local filling of Landau levels at the center of the trap, where the density is largest and hence mTF theory should be the most accurate. Namely, in accordance with (4.33), we start from the numerically calculated density matrix γ^{H} to form the Husimi function

$$m^{\text{H}}(n, 0) := \text{Tr} \left[\Pi_{n,0}^{B(0)} \gamma^{\text{H}} \right], \quad (5.7)$$

with $\Pi_{n,0}^{B(0)}$ as in (4.19) (which seems to work better than (4.23) according to our numerical tests). To limit the influence of numerical errors we take the theoretical value

$$B(0) = 2\pi\alpha\rho_\alpha^{\text{mTF}}(0).$$

There remains to choose the length ε in (4.21). This is slightly tricky in practice, because at the values of N we performed calculations for, there is not much room between magnetic

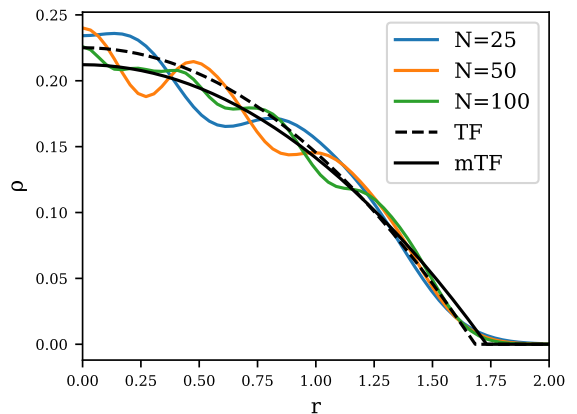


FIG. 7: Real/position space density $\rho(r)$ (normalized) in a harmonic trap for $\alpha = 0.75$. Different values of N , and compared to the theoretical ρ^{TF} in (3.6) (dashed) resp. ρ_α^{mTF} in (4.5) (solid).

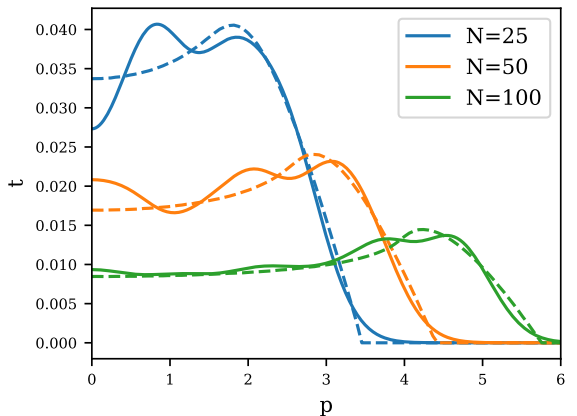


FIG. 8: Momentum space density $t(p)$ for $\alpha = 3/4$, with p rescaled by $N^{-1/2}$. Dashed lines are the semi-explicit theoretical guesses from (5.4)–(5.6).

length and system size to satisfy (4.22) (on top of that, the magnetic length seen as the radius of cyclotron orbits actually depends on the Landau level). A typical plot is shown in Figure 9, with comparison to the theoretical finding (4.35).

Observe that

$$m^{\text{H}}(n, 0) \leq \alpha \rho_{\gamma^{\text{H}}}(0), \quad \sum_{n \geq 0} m^{\text{H}}(n, 0) = g_\varepsilon * \rho_{\gamma^{\text{H}}}(0),$$

so that it is fairly natural to normalize plots by the value

$$\rho_\alpha^{\text{mTF}}(0) \simeq g_\varepsilon * \rho_{\gamma^{\text{H}}}(0)$$

and this is what we elected to do. Hence what is plotted is really the filling relative to the total density.

We plot the (relative) filling of the lowest four Landau levels $n = 0, 1, 2, 3$. The fit with theory (as calculated from (4.35))

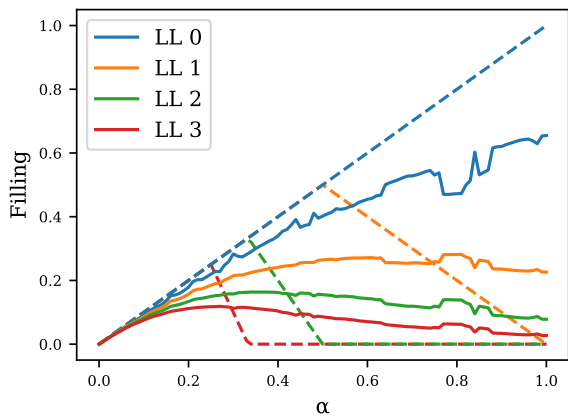


FIG. 9: Relative filling $m^H(n, 0)$ of local Landau levels at the origin, as a function of α , normalized by $g_\varepsilon * \rho_{\gamma^H}(0)$, compared to its theoretical value from (4.35) (dashed line). Parameters are $N = 100$, $\varepsilon = 0.1$, $V(\mathbf{x}) = |\mathbf{x}|^2$.

is certainly not perfect, which can be related to discrepancies in energies seen on Figure 6. We do not observe the expected fact that any Landau level stays full until the next one is entirely empty. The noteworthy feature is however that starting around $\alpha \simeq 1/2$, a very limited number of Landau levels are required to account for the whole density. We interpret this as a transition towards mTF theory, whose main feature is indeed that only few Landau levels are taken into account. The residual Landau level mixing we observe may be due to finite size effects, variations of the magnetic length and the (chosen by hand) ε length scale, etc. We have however observed a positive trend in N . For $N = 100$ (shown in Figure 9), although the theoretical value in each level is not reproduced, we find that, for $\alpha \gtrsim 1/4$, at least 80% of the total density is accounted for by the expected lowest $\lfloor \alpha^{-1} \rfloor + 1$ levels.

6. CONCLUDING REMARKS

We have simulated a mean-field system of fermions coupled to magnetic flux tubes, effectively via a Chern-Simons self-generating gauge field, and compared the results to semiclassical approximations, based either on the usual “position \times momentum” phase space or on the “position \times quantized cyclotron radius” phase space. Being better suited for systems in large magnetic fields, we found that the latter approximation gives a better match to the numerics for general values of the magnetic flux, reproducing correctly ground-state energies and densities. A theoretical guess for the momentum density also fits numerics rather well, giving predictions on an observable which shows more marked dependence on the magnetic flux, and could thus serve as a better signature in putative experiments. Such a momentum density could indeed be obtained from time-of-flight measurements if the system under consideration is realized with cold atoms coupled to density dependent artificial gauge fields. This is one of the proposed

realizations of anyons, quantum particles with fractional exchange statistics.

We note that our energy functional (2.6) contains only the most salient feature one would expect from anyons or particles coupled to a Chern-Simons field, namely the self-consistent magnetic field (2.5). It could be desirable to go beyond that by taking e.g. the following considerations into account:

- If one directly inserts a Slater determinant trial state in the many-body energy functional (1.4), one finds as energy the functional (2.6) that we studied thoroughly, but with the addition of extra terms. Besides exchange terms that would be of lower order for large N one finds a two-body term [28, Section 2.1]

$$\alpha^2 \iint_{\mathbb{R}^4} \frac{1}{|\mathbf{x} - \mathbf{y}|^2} (\gamma(\mathbf{x}, \mathbf{x})\gamma(\mathbf{y}, \mathbf{y}) - |\gamma(\mathbf{x}, \mathbf{y})|^2) d\mathbf{x}d\mathbf{y}. \quad (6.1)$$

Here γ is, as in the main text, the one-particle density matrix of the Slater determinant, and the two terms inside the parenthesis are respectively the direct and exchange terms obtained from applying the Wick rule to compute the two-particle density matrix. A possible approximation of the above could be the “direct only”

$$\alpha^2 \iint_{\mathbb{R}^4} W(\mathbf{x} - \mathbf{y})\gamma(\mathbf{x}, \mathbf{x})\gamma(\mathbf{y}, \mathbf{y}) d\mathbf{x}d\mathbf{y}, \quad (6.2)$$

where the effect of the exchange term (and thus the Pauli principle) is only taken into account by replacing $|\mathbf{x} - \mathbf{y}|^{-2}$ with a $W(\mathbf{x} - \mathbf{y})$ regularized at short distances to make (6.2) finite.

We do not expect that the addition of terms of this type would alter our qualitative findings, while possibly blurring the main features somewhat, which is why we elected to stick with the pure Chern-Simons-Schrödinger formalism of Section 2.

- Using more sophisticated trial states, incorporating repulsive Jastrow factors, would lead [4, 14, 15, 51, 55] to the replacement of (6.1)-(6.2) by a purely local term

$$f(\alpha) \int_{\mathbb{R}^2} \gamma(\mathbf{x}, \mathbf{x})^2 d\mathbf{x} \quad (6.3)$$

for some positive function $f(\alpha)$, see [4]. Such a term could have the interpretation of a Pauli term coupling the effective, density-dependent, magnetic field, to some spin of the particles. It could be incorporated with no extra difficulty in our basic scheme. With attractive Jastrow factors instead one could imagine a negative value of $f(\alpha)$.

- Anyons realized as quasi-particles of condensed matter systems would have a non-zero extension, leading to a smearing of the magnetic flux over some, typically small, length scale. We have not considered this so far, but we do not expect this to influence our main conclusions dramatically, at least not at the level of a mean-field approximation. If anything, this would tend to smoothen the two-particle density of the system, making the mean-field approximation even more efficient.
- An additional, constant, external magnetic field could be added to the model, to better match the fractional quantum Hall effect context, and/or check for signs of superfluidity/superconductivity; see e.g. [35] and references therein.

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