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1 **Comparison of ^{14}C and ^{210}Pb - ^{137}Cs - ^{241}Am dating methods of a recent bat**
2 **guano deposit (Lot, SW France).**

3

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10

11 **Abstract:**

12 Bat guano deposits are increasingly used as records of past environmental changes, an
13 approach that requires a precise chronology of the guano layers. This paper presents a
14 comparison between the well-established ^{14}C dating method and methods based on natural
15 ^{210}Pb excesses, $(^{210}\text{Pb})_{\text{ex}}$, and artificial radionuclides ^{137}Cs and ^{241}Am . The studied example is
16 a bat guano deposit from a cave in SW France (the Brantites III cave), which is currently
17 investigated for paleo-environmental reconstructions using stable isotopes. ^{14}C data show that
18 the upper part of the guano deposit accumulated during the last 150 years with a marked
19 increase in accumulation rates after around 1960 AD. While the incorporation of atmospheric
20 ^{14}C in guano is a well-understood process, the origin of ^{210}Pb excesses is more complex.
21 Based on consideration of ^{137}Cs and $(^{210}\text{Pb})_{\text{ex}}$ inventories [recorded in undisturbed soils in](#)
22 [France](#), and the measured inventories in the guano deposit, we suggest that most of the ^{210}Pb
23 excess is produced by ^{222}Rn decay in the cave air and then adsorbed onto the guano. As
24 Radon concentrations in caves can vary significantly on both short and long-term timescales,
25 one needs to be cautious before applying the often-used CRS (constant rate of supply of ^{210}Pb
26 excess) model to guano dating. Our $(^{210}\text{Pb})_{\text{ex}}$ data are best interpreted by two successive
27 periods of roughly constant, but widely different accumulation rates (0.3 m/y and 2.6 cm/y
28 before and after 1960, respectively) and $(^{210}\text{Pb})_{\text{ex}}$ fluxes. We suggest that these relatively
29 abrupt variations result from a change in cave ventilation leading to a more favourable shelter
30 for bats after 1960. The upper 40 cm of the deposit shows evidence of ^{210}Pb mobility, adding
31 a further complexity to the interpretation of $(^{210}\text{Pb})_{\text{ex}}$ profiles in guano deposits. However, the
32 existence of well-defined ^{137}Cs and ^{241}Am peaks allows a precise identification of the year of

33 maximum atmospheric fallouts (~ 1963-1964). When the ages provided by artificial
34 radionuclides are combined with the ^{210}Pb -derived accumulation rates, an age model can be
35 built, which is in good agreement with the ^{14}C age model. This example shows that the
36 $(^{210}\text{Pb})_{\text{ex}}$ method, when associated with ^{137}Cs (^{241}Am) data, can be used to date recent guano
37 deposit, although its application is not as straightforward as the ^{14}C method.

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40 **Keywords:** cave guano deposits; ^{14}C ages; ^{210}Pb , ^{137}Cs , ^{241}Am dating; age models; ^{210}Pb flux
41 and cave ventilation.

42 **1 Introduction**

43 Reconstruction of paleoenvironments and paleoclimates are based on many types of natural
44 archives, such as, for example, lacustrine or marine sediments, loess deposits, ice cores, or
45 speleothems. After the pioneering work of DesMarais et al. (1980), several studies have
46 emphasized the interest of bat guano deposits found in a number of caves as records of past
47 climatic and environmental changes, (e.g. Wurster et al., 2008; Johnston et al., 2010; Ferenc
48 et al., 2015). In order to date these changes and compare various records, a precise
49 chronology for such deposits is required. Wurster et al. (2009) showed that bat organic rich
50 guano material can be successfully dated through the ^{14}C method. Another possible
51 chronometer for young deposits (less than 150 y old) is ^{210}Pb . While the ^{210}Pb dating method
52 is widely used to infer sedimentation rates and ages of young sediments, few works tried to
53 apply this method to guano deposits. Among them are those of Wirmann et al. (2014) and
54 Gallant et al. (2020), whose results are further discussed by Bogdanovicz et al. (2020).
55 Wirmann et al. (2014) analysed bat and salangane guano from New Caledonia and found a
56 simple $(^{210}\text{Pb})_{\text{ex}}$ vs depth profile, allowing them to infer an accumulation rate of 3.5 cm/y. In
57 contrast Gallant et al. (2020) obtained a more complex profile in a guano deposit from
58 Jamaica, and the dates derived from ^{210}Pb are sometimes in conflict with those given by the
59 ^{14}C method. In a recent study, McFarlane and Lundberg (2021) showed that ^{210}Pb (and ^{137}Cs)
60 can be mobile in bat guano, a process that can seriously affect the chronology based on these
61 elements.

62 In this paper, we assess the potential of ^{210}Pb and ^{137}Cs , ^{241}Am as dating tools for recent bat
63 guano deposits and compare the results with those of the well-established ^{14}C method. The
64 studied example is a bat guano deposit found in a cave from SW France (Brantites III cave,
65 Lot), which is currently being investigated for paleo-environmental reconstructions using
66 stable isotopes.

67 **2 Site description and sampling**

68 Numerous caves and vertical pits (avens or “igues”) are found in the Causse du Quercy area,
69 an upper-Jurassic calcareous plateau, west of the French Massif Central, and part of the
70 sedimentary Aquitaine Basin. The Brantites I and II caves are located near the village of
71 Caniac-du-Causse at an altitude of 259 m a.s.l. Meteorological data are available from the
72 Gourdon station, located about 25 km to the NW. The Brantites III cave is connected to the

73 [larger](#) Brantites I cave, with an opening in the vertical Brantites I pit, a small un-explored
74 gallery at the opposite end, and a small aperture at the surface of the Causse (see map of Fig.
75 1). The Brantites I and II caves were first explored and described by the founder of speleology
76 E.A. Martel (Martel, 1894).

77 Observations made during the last decades show that the Brantites III cave is home to 1000 to
78 3000 insectivorous bats (3 rhinoloph species) during the [birthing](#) period (mostly from June to
79 September). Over the years a guano heap around 1.75 m high has accumulated (Fig. 2). The
80 guano deposit is located around 7m below the [land-surface](#). The top of the guano heap was
81 probably deposited in 2011 or 2012, since no bats were observed in the cave in 2013 and 2014,
82 although they were back in the cavity in 2015.

83 A vertical section was cut in the guano deposit with a shovel in March 2014. In the upper part,
84 alternating layers of light and dark brown colours can be easily seen, whereas the layering
85 tends to disappear below 80 cm depth. The lowermost part of the guano heap (below 120 cm)
86 has a dark brown colour and appears more compact (Fig. 2a). Four cores (C1 to C4) were
87 sampled using PVC half-pipes of 9 cm internal diameter. Their staggered arrangement with a
88 20 cm overlap between each core was intended to avoid gaps in the record. However this
89 choice resulted in some difficulties in assessing the depth of a given layer along the axis of
90 the guano heap, where the layers are roughly horizontal. Thus approximate corrections have
91 been applied to take into account the dip and thickness variations of the guano layers. This
92 recalculated axial depth is reported in the following [tables and diagrams](#).

93 **3 Sample preparation and analytical methods**

94 In the laboratory the guano cores were cut into 1cm-thick slices, put in plastic bags and stored
95 in a [freezer](#). Eight samples of about 1 cm³ each were subsampled and sent to Poznań Laboratory
96 for Radiocarbon dating (Goslar et al., 2004). The guano is a mixture of digested organic
97 matter, remains of insect parts (essentially made of chitine), and bat's hair. As all complex
98 organic materials, guano samples require a pre-treatment before AMS analyses. [The method](#)
99 [used is adapted from that described by Brock et al. \(2010\), and briefly summarised by](#)
100 [Johnston et al. \(2010\). The procedure involves treatments with acid and base solutions,](#)
101 [preceded by treatment with DCM/alcohol to get the fraction of SEG \(solvent-extracted guano\)](#)
102 [free of saturated hydrocarbons \(Wurster et al. 2009\). Guano samples are then combusted to](#)
103 [CO₂, and the CO₂ is finally reduced with H₂, using Fe powder as a catalyst, the final product](#)

104 to be analysed being a mixture of carbon and iron ($C/Fe = 1/2$) pressed into a 1 mm hole in an
105 aluminium cathode holder. The carbon isotope composition is determined on a “Compact
106 Carbon AMS”, manufactured by National Electrostatics Corporation, USA (detailed
107 procedures for AMS analyses are given by Goslar et al. (2004).

108 For ^{210}Pb and ^{137}Cs , ^{241}Am analyses, the samples were dried in an oven, for about one day at
109 60°C . The dried samples were then weighted, in order to estimate the “dry bulk density” DBD
110 (i.e. the dry weight divided by the volume of 1 cm-thick layer). The dry weights of the layers
111 were highly variable, from 2.8 to 6.5 g, with an average value of 4.4 g, i.e. an average DBD of
112 0.138 g/cm^3 . Weight variations are partly due to the difficulty of cutting regular slices in this
113 heterogeneous guano material and to variable proportions of its components. However, there
114 is no systematic increase of the DBD with depth, suggesting that compaction does play a
115 minor role in the studied part of the guano deposit.

116 The samples were then crushed in an agate mortar and homogenized. The powdered samples
117 were put in PE tubes of 14 mm internal diameter filled to a height of 39 mm for their analysis
118 by gamma spectrometry in a low-background Ge-well detector manufactured by
119 CANBERRA™. An average 2.7 g (from 2.1 to 3.7 g) was analysed in this way. ^{210}Pb
120 activities in guano samples were determined using its 46.5 keV gamma-ray by comparison of
121 their count rates with that of the ^{210}Pb peak of an in-house standard (volcanic rock) in which
122 activities of the U-series nuclides are well known. The efficiency calibration curve (absolute
123 efficiency vs gamma energy) obtained from this rock standard was used to calculate the
124 activities of ^{241}Am (at 59.5 keV) and ^{137}Cs (at 661.7 keV). Auto-absorption corrections can be
125 significant for low energy peaks such as those of ^{210}Pb and ^{241}Am . The low energy gamma
126 rays are much less absorbed in the guano samples, being essentially composed of light
127 elements (C, H, O, N) and having a low density, than in the volcanic rock standard, and thus a
128 correction has to be applied. The absorption coefficients (μ/ρ) were experimentally
129 determined on a few guano samples by measuring the attenuation of the 46.5 keV (^{210}Pb) and
130 63.2 keV (^{234}Th) gamma-rays emitted by a small pitchblende (U_3O_8) source, and passing
131 through a box containing the guano samples, using a broad-energy Ge detector BEGe 3825
132 made by CANBERRA™. The absorption coefficient for the 59.5 keV peak of ^{241}Am is
133 obtained by interpolation from the above measurements. Calculated auto-absorption
134 correction factors (i.e. the factors by which to multiply the activity/g of a guano sample

135 derived from the ratio of count rates/g of sample and standard) are about 0.70 and 0.80 for
136 ^{210}Pb and ^{241}Am respectively.

137 In order to calculate the ^{210}Pb excess (“unsupported” ^{210}Pb), $(^{210}\text{Pb})_{\text{ex}}$, ^{226}Ra must also be
138 evaluated. Because of the very low ^{226}Ra activity in most samples, ^{226}Ra was measured using
139 the peak at 352 keV of its progeny ^{214}Pb , and, when possible, also the other peaks of ^{214}Pb
140 (295 keV) and ^{214}Bi (609 keV). As discussed later in section 5.1, this calculated $(^{210}\text{Pb})_{\text{ex}}$
141 does not necessarily represent the true ^{210}Pb adsorbed on the surface of guano. Th content was
142 also estimated using the 238.6 keV peak of its progeny ^{212}Pb , assuming radioactive
143 equilibrium in the ^{232}Th -decay series.

144 All $^{210}\text{Pb}_{\text{ex}}$ ($T_{1/2} = 22.3$ y), ^{137}Cs ($T_{1/2} = 30$ y) and ^{241}Am ($T_{1/2} = 432$ y) data, obtained in 2020
145 and 2021, are decay-corrected to the date of sampling (8 March 2014).

146

147 **4 Results**

148 **4.1 ^{14}C data**

149 ^{14}C data and ages obtained on the studied part of the guano deposit are reported in Table 1.
150 Data are expressed in conventional ^{14}C ages BP for samples deposited before
151 1950, and in pMC (percentage of Modern Carbon) for samples deposited later
152 than 1950. The oldest ^{14}C date is 130 ± 30 BP at a depth of 138 cm. A calibrated date
153 around 1871 ± 72 AD is calculated using the OxCal program (Bronk Ramsey, 2009) in its 4.4
154 version (2021), with the IntCal20 calibration curve (Reimer et al., 2020), as the mean age of
155 the highest probability interval (1799-1942 AD). All the other layers, above 102 cm were
156 deposited during the period affected by atmospheric nuclear tests, with $^{14}\text{C}/^{12}\text{C}$ ratios higher
157 than the reference value of modern (1950 AD) carbon. Expressed in pMC, the values range
158 from 105.9 at 7 cm to 166.6 at 102 cm depth. The pMC values regularly decrease from 97 to 7
159 cm depth, indicating that these guano layers were deposited after 1964, year of the maximum
160 bomb ^{14}C content in the atmosphere. The dates reported in Table 1 are calculated using the
161 OxCal program and the Bomb21 NH1 calibration curve (Reimer et al., 2020; Hua et al., 2021).
162 The ^{14}C date for the layer at 102 cm depth (166.6 pMC) could correspond to 1966-1967 AD,
163 but our ^{137}Cs data below suggest a date of 1963.3, just before the ^{14}C bomb peak (1963.7).
164 Thus, according to ^{14}C data, the upper 1.50 m of guano accumulated for around 140 years,

165 from 1870 to 2011, although the accumulation rate was much lower in the bottom part of the
166 guano deposit (36 cm between 1870 and 1963, compared to 102 cm between 1963 and 2011,
167 with an average sedimentation rate of 2.4 cm/y). The age range covered by the ^{14}C dates
168 closely corresponds to the dating interval accessible through the ^{210}Pb method.

169 **4.2 ^{210}Pb , ^{137}Cs and ^{241}Am data**

170 These data are reported in Table 2. Note that, throughout the paper, parentheses denote
171 activities/g. $(^{210}\text{Pb})_{\text{ex}}$ activities/g of guano sample are reported vs depth in Fig. 3. Obviously
172 the $(^{210}\text{Pb})_{\text{ex}}$ profile does not conform to a simple model assuming a constant initial $(^{210}\text{Pb})_{0,\text{ex}}$
173 activity/g and a constant accumulation rate, that should result in a straight line with a negative
174 slope, when $(^{210}\text{Pb})_{\text{ex}}$ is plotted on a logarithmic scale. $(^{210}\text{Pb})_{\text{ex}}$ first increase over the depth
175 range 0 to 43 cm, from 59 to 114.5 mBq/g, then shows a roughly linear decrease with a ^{210}Pb
176 of 36 mBq/g at 104.5 cm. This general tendency is however rather irregular. Then, over a few
177 cm (104.5 to 108 cm), $(^{210}\text{Pb})_{\text{ex}}$ increases from 36 to 56.5 mBq/g. Finally, in the deepest
178 analysed part of the guano deposit (108 to 138 cm), $(^{210}\text{Pb})_{\text{ex}}$ displays a linear decrease until it
179 reaches 4.5 mBq/g, a near zero value within 2σ uncertainty.

180 ^{137}Cs activities are hardly detectable in the upper 45 cm of the guano sequence (≤ 1 mBq/g).
181 They then increase from 63.5 cm (1.17 mBq/g) to reach a peak value of 17.2 mBq/g at 100.5
182 cm, decreasing thereafter to values lower than 1 mBq/g below 110 cm (Fig. 4).

183 ^{241}Am , a longer-lived ($T_{1/2} = 432$ y) decay product of ^{241}Pu , produced by atmospheric nuclear
184 bomb testing, follows the same trend as ^{137}Cs , although with much lower activities, reaching
185 its maximum of 2.76 mBq/g at 101.5 cm, only 1 cm below the ^{137}Cs peak (Fig. 4).

186 **5 Discussion**

187 **5.1 The origin of ^{210}Pb excess**

188 ^{210}Pb is continuously produced by ^{222}Rn decay in the atmosphere, adsorbed onto aerosol and
189 dust particles and ultimately deposited on the soil surface by rain or snow, or by dust settling
190 (“dry deposition”). In lakes or marine waters, additional adsorbed ^{210}Pb is provided to settling
191 solid particles by decay of ^{222}Rn dissolved in the water column. Except in the case of ^{210}Pb
192 dating of ice or snow accumulation, where ^{210}Pb adsorbed onto detrital particles is negligible,
193 calculation of adsorbed, “excess ^{210}Pb ”, $(^{210}\text{Pb})_{\text{ex}}$, requires a correction based on the so-called

194 “supported ^{210}Pb ”, i.e. ^{210}Pb assumed to be in radioactive equilibrium with ^{226}Ra inside the
195 detrital component [$(^{210}\text{Pb})_{\text{ex}} = (^{210}\text{Pb})_{\text{measured}} - (^{226}\text{Ra})_{\text{measured}}$].

196 In guano, however, most of the material is made of organics with few detrital particles, as
197 shown by the very low Th contents of the analysed samples (0.4 ppm on average), except in
198 the deepest layers where Th increases up to 3.2 ppm at 138 cm depth. Thus, ^{226}Ra might not
199 originate from a detrital component in secular radioactive equilibrium, but rather be
200 incorporated during the plant-insect-bat cycle, with an unknown ($^{210}\text{Pb}/^{226}\text{Ra}$) ratio. On
201 average, (^{226}Ra) amounts to 10% of the measured (^{210}Pb) in the layers mostly devoid of
202 significant detrital minerals ($z < 108$ cm). Thus, the uncertainty linked to the “supported ^{210}Pb ”
203 correction should not significantly alter the interpretation of the ^{210}Pb data. It might however
204 contribute partly to [the dispersion of the data](#), as observed in Figure 3.

205 Of more concern is the fact that ^{210}Pb adsorbed on guano might not only come from
206 atmospheric ^{210}Pb fallout, a fraction of which is incorporated during the plant-insect-bat cycle,
207 but also from ^{210}Pb produced by ^{222}Rn decay in the air inside the cave. This can be
208 particularly important in caves where bats only stay for a few months in the year, as in the
209 Brantites III cave. For most of the year, the guano is preserved from direct atmospheric input
210 but can adsorb ^{210}Pb produced by ^{222}Rn decay in the cave air.

211 The relative contributions of these two sources of adsorbed ^{210}Pb could potentially be
212 estimated through the comparison between ^{210}Pb and ^{137}Cs (or ^{241}Am) inventories. The only
213 origin of ^{137}Cs in guano is from atmospheric ^{137}Cs fallout. If we assume that ^{137}Cs and
214 atmospheric (^{210}Pb)_{ex} are fractionated in a similar way during the plant-insect-bat-guano cycle,
215 i.e. if both nuclides have similar “transfer efficiencies” as defined by Robbins and Edgington
216 (1978), then the ratio of their atmospheric inventories, as recorded in soils for example,
217 should be the same as that recorded in the guano, if no other source of adsorbed ^{210}Pb was
218 implied.

219 A ^{137}Cs inventory was measured by Monna et al. (2009) in [an undisturbed soil \(neoluvisol
220 developed in eolian loess deposits\)](#) at an experimental station in Versailles near Paris. These
221 authors found a value of 3200 Bq/m² in 1992 (corresponding to 1925 Bq/m² in 2014). Le
222 Roux et al. (2010) reported ^{137}Cs inventories in undisturbed soils from Montagne Noire (SW
223 of the French Massif Central), and demonstrated their relationship with the mean annual
224 precipitation (MAP): I_{Cs} (Bq/m², in 2007) = 3.04 * MAP (mm) - 110. The MAP at the Gourdon
225 meteorological station is around 900 mm for the period 1971-2010 (www.infoclimat.fr). With

226 this value, an inventory of 2626 Bq/m² is calculated in 2007, corresponding to 2234 Bq/m² in
227 2014. An approximate value of around 2000 Bq/m² can thus be estimated, in agreement with
228 the maps of ¹³⁷Cs soil inventories recently published by Meusburger et al. (2020).

229 Atmospheric ²¹⁰Pb fluxes are reported by Garcia-Orellana et al. (2006) for coastal areas from
230 West-Mediterranean Sea. They also show a good correlation with MAP. Taking into account
231 the MAP measured at Gourdon (900 mm), a flux F of 150 Bq/m²/y is calculated, resulting in
232 an inventory of 4839 Bq/m² ($I = F/\lambda_{Pb}$). A similar value around 5000 Bq/m² was obtained
233 from sediments deposited in the Thau lagoon (along the Gulf of Lions coast in France) by
234 Monna et al. (1996). A value slightly above 4000 Bq/m² can be derived from a global
235 correlation of ²¹⁰Pb inventories vs MAP in European inland sites (Appleby 2001; Le Roux,
236 2008). Taking into account the uncertainties associated with these data, a value of 4500 Bq/m²
237 seems a reasonable estimate.

238 Thus the ratio I_{Pb}/I_{Cs} is 2.25 (4500/2000). The inventories of ¹³⁷Cs and ²¹⁰Pb_{ex} in the guano
239 deposit can be estimated from our data. They are calculated according to the classical
240 equation:

$$241 \quad I \text{ (Bq/m}^2\text{)} = 10 \cdot \sum A(z) \cdot \rho(z) \cdot \Delta z ,$$

242 where $A(z)$ is the ¹³⁷Cs or ²¹⁰Pb_{ex} activity/g (in mBq/g) in the layer at depth z (cm), ρ is the
243 “dry bulk density” (g/cm³) at depth z and Δz (cm) the thickness of layer z . As all the guano
244 layers have not been analysed, activities have been interpolated between measured values, and
245 the average dry bulk density $\bar{\rho}$ used in the estimation, i.e. $I = 10 \cdot \bar{\rho} \cdot \int_0^\infty A(z) \cdot dz$.

246 The ¹³⁷Cs inventory (in 2014) in the guano deposit is around 290 Bq/m². If the ²¹⁰Pb_{ex} was
247 only derived from atmospheric deposition, as ¹³⁷Cs, with the same transfer efficiency, one
248 should expect a ²¹⁰Pb_{ex} inventory of 725 Bq/m². The value estimated from the curve (²¹⁰Pb)_{ex}
249 vs depth is around 11700 Bq/m², higher by a factor of 16 than the expected value. This
250 strongly suggests that most of the ²¹⁰Pb adsorbed onto the guano comes from ²²²Rn decay in
251 the cave air.

252 ²²²Rn activities/m³ in cave air are generally much higher than in the atmosphere, due to its
253 release from sediments, residual clays or alterites filling many caves, and to degassing of
254 flowing or dripping waters that were enriched in ²²²Rn during infiltration through the soil or
255 even through the limestone micro-fractures. While the atmosphere has ²²²Rn activities of a

256 few tens of Bq/m³ (e.g. Gamarini, 2006), ²²²Rn activities in caves can be a factor of 1000
257 higher (e.g. Hakl et al., 1997; Cigna, 2005). For example, Guibert et al. (2015) measured a
258 value of 7000 Bq/m³ for ²²²Rn in the air of the prehistoric Chauvet cave in France. These
259 authors discuss in detail the mechanisms leading to ²²²Rn accumulation in the cave. They
260 convincingly demonstrate that ²¹⁰Pb is enriched at the surface of rock fragments exposed to
261 the cave air, probably from direct adsorption of short-lived nuclides produced by ²²²Rn decay.

262 Radon monitoring in a number of caves has revealed large temporal variations at different
263 timescales: diurnal, seasonal, annual and even longer-term variations (e.g. Hakl et al., 1997).
264 It is well known for example that ²²²Rn activities are higher in the summer than in the winter
265 (Cigna, 2005). During winter, the outside air is generally colder and thus denser than the cave
266 air, and it can invade the cave, whereas during summer the warmer and lighter outside air
267 does not penetrate the cave, allowing ²²²Rn accumulation. The degree of ventilation of the
268 cave plays an essential role in its ²²²Rn content and the ²¹⁰Pb flux produced by ²²²Rn decay.
269 Ventilation will depend on external factors (temperature, wind...) as well as internal factors
270 linked to the morphology of the cave (e.g. natural or artificial opening or obstruction of
271 conduits modifying air flows). It is thus possible that the usual models of ²¹⁰Pb dating
272 invoking a constant ²¹⁰Pb flux (or rate of supply, like in the CF:CS or CRS models) do not
273 apply for the entire period of cave guano deposition.

274 ***5.2 Interpretation of (²¹⁰Pb)_{ex} vs depth profile***

275 As mentioned in section 4.2, we have divided the (²¹⁰Pb)_{ex} profile into 4 segments, according
276 to Fig. 3. The ²¹⁰Pb_{ex} data of segment 1 (between the surface and 42.5 cm) can not be
277 explained by any classical ²¹⁰Pb model. Indeed, the “constant rate of supply” (CRS = □)
278 model (Appleby and Oldfield, 1978) is sometimes advocated to explain reverse ²¹⁰Pb_{ex}
279 profiles as a consequence of an acceleration of the accumulation rate w , and thus a decrease in
280 ²¹⁰Pb_{0,ex} initial activity/g A_0 (since $A_0 = \square/\square$). However, such an abrupt increase in
281 accumulation rate is not supported by the ¹⁴C data, which rather suggest a relatively stable
282 rate in the upper part of the guano deposit (after 1963). A possible cause of this peculiar
283 profile would be a sudden decrease in the ²¹⁰Pb_{ex} flux, linked to a decrease in ²²²Rn
284 concentration in the cave as a consequence of an increased ventilation of the cave with
285 outside air. But such a change occurring during the last 20 years would probably have been
286 noticed by speleologists, who regularly visit the cave to survey the bat population. Another
287 possible explanation is that the adsorbed ²¹⁰Pb_{ex} has been leached from the upper part of the

288 guano heap. Indeed the upper layers of guano deposits are often characterised by low pH,
289 often below 4 that might result from infiltration of bat's urine and/or microbial decomposition
290 of guano (e.g. Audra et al., 2019; MacFarlane and Lundberg, 2021). In such conditions,
291 $^{210}\text{Pb}_{\text{ex}}$ might well be leached from the shallow layers and be either accumulated deeper
292 downwards or exported **beyond** the guano heap. In their recent paper, MacFarlane and
293 Lundberg (2021) argue, on the basis of both natural examples and experimental data, that
294 such a process is probably not uncommon and can affect both ^{210}Pb and ^{137}Cs . In agreement
295 with their conclusion, we suggest that the adsorbed $^{210}\text{Pb}_{\text{ex}}$ was mobile in the **acidic** shallower
296 part of the Brantites III guano deposit.

297 $^{210}\text{Pb}_{\text{ex}}$ data **from** segment 2 are compatible with an exponential decrease, despite a large
298 scatter around the best-fit line. They probably reflect a period of approximately constant
299 sedimentation rate, and constant $^{210}\text{Pb}_{\text{ex}}$ flux (like in the CF:CS model), implying a roughly
300 constant initial $(^{210}\text{Pb})_{0,\text{ex}}$ activity/g. The slope $(-\lambda/v)$ of the straight line in Fig. 3 indicates a
301 sedimentation rate v of 2.59 ± 0.28 cm/y. This value is very close to the average
302 sedimentation rate obtained from ^{14}C dates for the period 1963-2010 (2.4 cm/y). With such a
303 high sedimentation rate, the analysis of 1cm-thick layers will inevitably lead to a dispersion of
304 the data. Indeed, at an intra-annual scale, the sedimentation rate is not constant, since bats live
305 in the cave and produce guano for only a few months. For the rest of the year, the
306 sedimentation rate is near zero, and the $^{210}\text{Pb}_{\text{ex}}$ concentration thus much higher at the surface
307 of the fresh guano. In addition, the probable seasonal variation in Radon concentration in the
308 cave will also induce variation in $^{210}\text{Pb}_{\text{ex}}$ fluxes and concentrations.

309 A similar CF:CS model can explain the data **from** segment 4 in Fig. 3. The steeper slope
310 indicates however a much lower sedimentation rate of 0.30 ± 0.04 cm/y. This value is close
311 to that deduced from the two oldest ^{14}C ages (0.39 cm/y). It should be noted that the $^{210}\text{Pb}_{\text{ex}}$
312 data of segment 4 are not in continuity with those of segment 2. It means that the initial
313 $^{210}\text{Pb}_{0,\text{ex}}$ concentration for this older period was different, and higher, compared to that of the
314 younger period. The transition between these two contrasting periods corresponds to segment
315 3 in Fig. 3, where $(^{210}\text{Pb})_{\text{ex}}$ increases with depth.

316 In summary, $^{210}\text{Pb}_{\text{ex}}$ data display two distinct stages of guano accumulation, an older one with
317 a low accumulation rate (0.3 cm/y) and a younger one with a much higher rate (2.6 cm/y).
318 Without ^{14}C ages, it would not be possible to assess either the age of this abrupt change, or the
319 deposition date of a particular guano layer. Indeed the only available “anchor point” date,

320 2011 at the top of the deposit, cannot be used because the mobility of $^{210}\text{Pb}_{\text{ex}}$ prevents
321 calculation of accumulation rate in the segment 1 zone. However, ^{137}Cs (and ^{241}Am) data
322 provide additional and essential information, as discussed below.

323

324 *5.3 The essential contribution of ^{137}Cs (and ^{241}Am) profiles*

325 The ^{137}Cs and ^{241}Am profiles displayed in Fig. 4 clearly show the typical profiles expected
326 from the curve of atmospheric fallout of these bomb-related radionuclides (e.g. Barker and
327 Bouisset, 2001). The peak values corresponding to the year 1963 occur between 100 and 102
328 cm depth for both nuclides and coincides with the ^{14}C date of 1963.4 obtained at 102 cm.
329 Thus ^{137}Cs and ^{241}Am data provide a robust anchor date. Another possible anchor date,
330 although less precisely defined, corresponds to the beginning of significant increase in ^{137}Cs
331 in the atmosphere after the first nuclear tests, i.e. the first years of the 1950-1960 decade.
332 Using these anchor dates and the two accumulation rates determined for stages 2 and 4, an
333 age model can be proposed from the combined ^{210}Pb and ^{137}Cs (^{241}Am) data. The dates
334 calculated from each of the analysed layers are reported in Table 2.

335 From the $(^{210}\text{Pb})_{\text{ex}}$ data and the age of each guano layer obtained from the interpolated ^{14}C
336 ages and the ^{210}Pb -derived accumulation rate for the stage 4, the initial $(^{210}\text{Pb})_{0,\text{ex}}$ activities/g,
337 at the time of deposition can be calculated. They are reported in Fig. 5. As expected from Fig.
338 3, $(^{210}\text{Pb})_{0,\text{ex}}$ was higher in the pre-1960 period (around 312 mBq/g) than in the post-1960
339 period (216 mBq/g). The difference however is small, compared to the large difference in the
340 accumulation rates (an increase by a factor of 8.6). This means that the flux of $^{210}\text{Pb}_{\text{ex}}$ should
341 have been much lower by a factor of 6 in the pre-1960 period. No major change in climatic
342 conditions is detectable around 1960 in the South of France (Lelièvre et al., 2011). Such a
343 large variation of the flux of $^{210}\text{Pb}_{\text{ex}}$ in the cave could be explained if there was an abrupt
344 change in the cave ventilation around 1960, i.e. if the cave was largely open to the atmosphere
345 with efficient air circulation and a low Radon concentration before 1960, and was much less
346 ventilated with more Radon accumulation after 1960. Such a change could result for example
347 from the partial closure of the gallery in the distal part of the cave (Fig. 1), as a consequence
348 of rock-fall or sediment discharge obstructing the gallery. Since no direct observation by
349 speleologists is available for this period, it is difficult to go further in this discussion. It should
350 be noted however that such an hypothetical change in cave ventilation might well explain the
351 low accumulation rate (probably resulting from a less numerous bat population), because bats

352 prefer more stable cave conditions without strong air flows (Zukal et al., 2017; Lüthi and
353 Blant, 2017).

354 Our data suggest that $^{210}\text{Pb}_{\text{ex}}$ flux in caves as recorded by guano deposits may vary
355 considerably over time. This implies that great care should be taken before applying the
356 “constant rate of supply” (CRS) model, as sometimes done in the literature (e.g. Bogdanovicz
357 et al., 2020). In order to show that the CRS model would give inaccurate dates in the present
358 case, we have calculated the age of the 1963 layer at 100-102 cm depth, using the classical
359 formula (Appleby and Oldfield, 1978):

$$360 \quad t = -\frac{1}{\lambda} \cdot \ln (I_t/I_0) ,$$

361 where I_t and I_0 are the $^{210}\text{Pb}_{\text{ex}}$ inventories below the layer z (age t), and the total inventory
362 respectively. The calculated age t is 70.6 y (before 2014), i.e. 1943, 20 years older than the
363 true age.

364 ***5.4 Comparison of $^{210}\text{Pb}_{\text{ex}}$, ^{137}Cs (^{241}Am) and ^{14}C age models***

365 In Fig.6 are shown the curves interpolated from ^{14}C dates and $^{210}\text{Pb}_{\text{ex}}$ and ^{137}Cs - ^{241}Am dates
366 (Table 1). There is a good agreement between the two age models, with a maximum
367 difference of about 5 y in the 1970-1980 AD range (Fig. 6). This shows that $^{210}\text{Pb}_{\text{ex}}$ data can
368 provide a useful means of dating guano deposits, but that $^{210}\text{Pb}_{\text{ex}}$ data should be
369 complemented by at least ^{137}Cs data. We thus agree with the recommendation made by Smith
370 (2001) that “ ^{210}Pb geochronology must be validated using at least one independent tracer”,
371 and with the assertion by McFarlane and Lundberg (2021) that the interpretation of $^{210}\text{Pb}_{\text{ex}}$
372 data alone will often prove difficult. And, obviously, in spite of its higher analytical cost, ^{14}C
373 remains the best method to date bat guano deposits.

374 **6 Conclusions**

375 The presence, in the Brantites III cave, in SW France, of a relatively thick guano deposit
376 allowed a detailed comparison of the ^{14}C and ^{210}Pb , ^{137}Cs , ^{241}Am dating methods.

377 The analysed upper part of the guano accumulation was deposited for the last 150 years as
378 indicated by a ^{14}C date obtained for the layer at 138 cm depth. All the other ^{14}C data display
379 $^{14}\text{C}/^{12}\text{C}$ ratios higher than the value of modern C, clearly indicating that their ages are younger
380 than the introduction in the atmosphere of anthropogenic ^{14}C derived from nuclear tests

381 (around 1945). Unambiguous ^{14}C ages in stratigraphic order are obtained by comparison with
382 the post-bomb atmospheric curve.

383 Whereas the origin of ^{14}C ultimately found in the guano is rather straightforward, we show
384 that the origin of ^{210}Pb excesses is more complex. A comparison of $^{210}\text{Pb}_{\text{ex}}$ and ^{137}Cs
385 inventories derived from atmospheric fallout, and those measured in the guano deposit
386 suggests that most $^{210}\text{Pb}_{\text{ex}}$ adsorbed onto guano is produced by ^{222}Rn decay inside the cave.
387 As ^{222}Rn concentration in caves might show both short-term (e.g. seasonal) and long-term
388 (several years or more) fluctuations, it can affect the adsorbed $^{210}\text{Pb}_{\text{ex}}$ flux, making the
389 assumption of a constant rate of supply (in the CRS model) questionable.

390 The profile of $^{210}\text{Pb}_{\text{ex}}$ vs depth reveals two contrasting periods of accumulation reflecting
391 large differences in guano accumulation rates (0.3 cm/y before 1960 and 2.6 cm/y after 1960),
392 which agree with the rates derived from ^{14}C ages. Calculation of initial $(^{210}\text{Pb})_{0,\text{ex}}$ activities/g
393 indicates a higher activity/g for the older period, but a much lower $^{210}\text{Pb}_{\text{ex}}$ flux. We tentatively
394 interpret the large difference in fluxes as the result of a relatively abrupt change in cave
395 ventilation from widely open (to the atmosphere) to more confined conditions. This could also
396 explain the smaller number of bats living in the cave before 1960, [as suggested by the lower](#)
397 [accumulation rate](#).

398 The shallower part (30 cm) of the deposit displays a reversed $^{210}\text{Pb}_{\text{ex}}$ profile that might result
399 from the partial leaching of ^{210}Pb in acidic conditions, as proposed by McFarlane and
400 Lundgren (2021).

401 Both ^{137}Cs and ^{241}Am activities/g vs depth display simple profiles with a clearly marked peak
402 value at a depth of 100-102 cm, corresponding to 1963, the year of maximum atmospheric
403 fallout. It also coincides with the ^{14}C date obtained for that layer. Using this age and the age
404 of beginning of ^{137}Cs increase (~1952) as anchor ages, and the $^{210}\text{Pb}_{\text{ex}}$ - derived accumulation
405 rates, a combined ^{210}Pb - ^{137}Cs - ^{241}Am age model is proposed, which is in good agreement
406 with the ^{14}C age model.

407 Although the origin of $^{210}\text{Pb}_{\text{ex}}$ adsorbed on guano deposits is more complex than usually
408 assumed, ^{210}Pb can still be used as a dating tool, when combined with other tracers like ^{137}Cs
409 or ^{241}Am , as demonstrated in this study. Moreover, as $^{210}\text{Pb}_{\text{ex}}$ is dependent on Radon
410 concentration in the cave air, it can offer insights into past variations in the cave ventilation.

411

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416

417 **References**

- 418 Appleby, P.G., Oldfield, F. (1978). The calculation of Lead-210 dates assuming a constant
419 rate of supply of unsupported ²¹⁰Pb to the sediment. *Catena*, 5, 1-8.
- 420 Appleby, P.G. (2001). Chronostratigraphic techniques in recent sediments. Tracking
421 environmental change using lake sediments. Basin analysis, coring, and chronological
422 techniques. *Developments in Paleoenvironmental Research* 1, 171–203.
- 423
- 424 Audra, P., De Waele, J., Bentaleb, I., Chroňáková, A., Křišťůfek V., D’Angeli, I.M., Carbone,
425 C., Madonia, G., Vattano, M., Scopelliti, G., Cailhol, D., Vanara, N., Temovski, M., Bigot,
426 J.Y., Nobécourt, J.C., Galli, E., Rull, F., Sanz-Arranz, A. (2019). Guano-related phosphate-
427 rich minerals in European caves. *Int. J. of Speleology*, 48 (1), 75-105.
- 428
- 429 Barker, E., Bouisset, P. (2001). Evolution de la radioactivité en France Métropole et en
430 Outre-Mer dans les aérosols et les retombées depuis la mise en place de l’observatoire
431 atmosphérique de l’IPSN. Rapport DPRE/SERNAT/2001-27, 51p.
- 432
- 433 Bogdanowicz, W., Worobiec, E., Grooms, C., Kimpe, L.E., Smol, J.P., Stewart, R.S.,
434 Suchecka, E., Pomorski, J.J. , Blais, J.M., Clare, E.L., Fenton, M.B. (2020). Pollen
435 assemblage and environmental DNA changes: A 4300 year-old bat guano deposit from
436 Jamaica. *Quaternary International*, 558, 47-58.
- 437
- 438 Brock, F., Higham, T., Ditchfield, P., Bronk Ramsey, C. (2010). Current pretreatment
439 methods for AMS radiocarbon dating at the Oxford Radiocarbon Accelerator Unit (ORAU).
440 *Radiocarbon*, 52 (1), 103-112.
- 441 Bronk Ramsey, C. (2009). Bayesian analysis of radiocarbon dates. *Radiocarbon*, 51(1), 337–
442 360.
- 443 Cigna, A, A. (2005). Radon in caves. *International Journal of Speleology*, 34 (1-2), 1-18.
- 444
- 445 Des Marais, D.J., Mitchell, J.M., Meinschein, W.G., Hayes, J.M. (1980). The carbon isotope
446 biogeochemistry of the individual hydrocarbons in bat guano and the ecology of the
447 insectivorous bats in the region of Carlsbad, New Mexico. *Geochimica et Cosmochimica*
448 *Acta*, 44(12), 2075-2086.

449

450 Ferenc, L.F., Onac, B.P., Tanțău, I., Wynn, J.G., Tămaș, T., Coroiu, I., Giurgiu, A.M. (2015).
451 A Late Holocene environmental history of a bat guano deposit from Romania: an isotopic,
452 pollen and microcharcoal study. *Quaternary Science Reviews*, 127, 141-154.

453

454 Gallant, L.R., Grooms, C., Kimpea, L.E., Smolb, J.P., Bogdanowicz, W., Stewart, R.S., Clare,
455 E.L., Fenton, M.B., Blaisa, J.M. (2020). A bat guano deposit in Jamaica recorded agricultural
456 changes and metal exposure over the last > 4300 years. *Palaeogeography, Palaeoclimatology,*
457 *Palaeoecology* 538, 109470.

458

459 Galmarini, S. (2006). One year of ^{222}Rn concentration in the atmospheric surface layer.
460 *Atmos. Chem. Phys.*, 6, 2865-2887.

461

462 Garcia-Orellana, J, Sanchez-Cabeza, J. A., Masqué, P., Avila, A., Costa, E., Loÿe-Pilot, M.D.,
463 Bruach-Menchén, J.M. (2006). Atmospheric fluxes of ^{210}Pb to the western Mediterranean Sea
464 and the Saharan dust influence. *J. Geophys. Res.*, 111, D15305, doi:10.1029/2005JD006660.

465

466 Goslar T., Czernik J., Goslar E. (2004). Low-energy ^{14}C AMS in Poznan Radiocarbon
467 Laboratory, Poland. *Nuclear Instruments and Methods in Physics Research B* 223-224: 5-11.

468

469 Guibert, P., Brodard, A., Quiles, A., Geneste, J.M., Baffier, D., Debard, E., Ferrier, C. (2015).
470 When were the walls of the Chauvet-Pont d'Arc Cave heated? A chronological approach by
471 thermoluminescence. *Quaternary Geochronology* 29, 36-47.

472

473 Hakl, J., Hunyadi, I., Csige, I., Géczy, G., Lénart, L., Varhegyi, A. (1997). Radon transport
474 phenomena studied in Karst caves. International experiences on Radon levels and exposures.
475 *Radiations Measurements*, 28, 1-6, 675-684.

476

477 Hua, Q., Turnbull, J. C., Santos, G. M., Rakowski, A. Z., Ancapichún, S., De Pol-Holz, R.,
478 Hammer, S., Lehman, S. J., Levin, I., Miller, J. B., Palmer, J. G., & Turney, C. S. M. (2021).
479 Atmospheric Radiocarbon For The Period 1950–2019. *Radiocarbon*, pp 1–23.
480 doi:10.1017/rdc.2021.95

481 Johnston, V.E, McDermott, F., Tămaș, T. (2010). A radiocarbon dated bat guano deposit from

482 N.W. Romania: Implications for the timing of the Little Ice Age and Medieval Climate
483 Anomaly. *Palaeogeography, Palaeoclimatology, Palaeoecology*, 291(3-4), 217-227.
484

485 Lelièvre, F., Sala, S., Ruget, F., Volaire, F. (2011). Evolution climatique du Sud de la France
486 1950-2009, Projet CLIMFOUREL PSDR3, Régions LOR, MIP, RIA. Série Les Focus PSDR3.
487 ([https://www1.montpellier.inra.fr/PSDR/doc/climfourel/FOCUS-PSDR3-](https://www1.montpellier.inra.fr/PSDR/doc/climfourel/FOCUS-PSDR3-CLIMFOUREL_Clim%20Chgt.pdf)
488 [CLIMFOUREL_Clim%20Chgt.pdf](https://www1.montpellier.inra.fr/PSDR/doc/climfourel/FOCUS-PSDR3-CLIMFOUREL_Clim%20Chgt.pdf))

489 Le Roux, G., Pourcelot, L., Masson, O., Duffa, C., Vray, F., Renaud, P. (2008). Aerosol
490 deposition and origin in French mountains estimated with soil inventories of ^{210}Pb and
491 artificial radionuclides. *Atmospheric Environment*, 42 (7), 1517-1524.
492

493 Le Roux, G., Duffa, C., Vray, F., Renaud, P. (2010). Deposition of artificial radionuclides
494 from atmospheric Nuclear Weapon Tests estimated by soil inventories in French areas low-
495 impacted by Chernobyl. *Journal of Environmental Radioactivity*, Elsevier, 101, 211-218.
496

497 Lüthi, C., Blant, M. (2017). Spéléologie et chauves-souris.
498 (https://www.isska.ch/pdf/Fr/Rd/Biospeleologie/Brochure_Chiro_V3_f.pdf)
499

500 Martel, E.A. (1894). *Les abîmes : les eaux souterraines, les cavernes, les sources, la*
501 *spéléologie : explorations souterraines effectuées de 1888 à 1893 en France, Belgique,*
502 *Autriche et Grèce, Paris, Librairie Ch. Delagrave, 578 p.*

503 McFarlane, D.A., Lundberg, J. (2021). Geochronological implications of ^{210}Pb and ^{137}Cs
504 mobility in cave guano deposits. *International Journal of Speleology*, 50(3), 239-248.
505 <https://doi.org/10.5038/1827-806X.50.3.2391>
506

507 Meusburger, K., Evrard, O., Alewell, C., Borrelli, P., Cinelli, G., Ketterer, M., Mabit, L.,
508 Panagos, P., Van Oost, K., Ballabio, C. (2020). Plutonium aided reconstruction of caesium
509 atmospheric fallout in European topsoils. *Nature*. <https://doi.org/10.1038/s41598-020-68736-2>
510

511 Monna, F., Lancelot, J., Bernat, M., Mercadier, H. (1996). Taux de sédimentation dans l'
512 étang de Thau à partir des données géochronologiques, géochimiques et des repères
513 stratigraphiques. *Oceanologica Acta*, 20, 4, 627-638.

514
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529
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531
532
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540
541
542
543
544
545
546

Monna, F., Van Oort, F., Hubert, P., Dominik, J., Bolte, J., Loizeau, J.L., Labanowski, J., Lamri, J., Petit, C., Le Roux, G., Château, C. (2009). Modeling of ¹³⁷Cs migration in soils using an 80-year soil archive: role of fertilizers and agricultural amendments. *Journal of Environmental Radioactivity*, 100, 9–16.

Reimer, P., Austin, W., Bard, E., Bayliss, A., Blackwell, P., Bronk Ramsey, C., Butzin, M., Cheng, H., Edwards, R., Friedrich, M., Grootes, P., Guilderson, T., Hajdas, I., Heaton, T., Hogg, A., Hughen, K., Kromer, B., Manning, S., Muscheler, R., Palmer, J., Pearson, C., van der Plicht, J., Reimer, R., Richards, D., Scott, E., Southon, J., Turney, C., Wacker, L., Adolphi, F., Büntgen, U., Capano, M., Fahrni, S., Fogtmann-Schulz, A., Friedrich, R., Köhler, P., Kudsk, S., Miyake, F., Olsen, J., Reinig, F., Sakamoto, M., Sookdeo, A., & Talamo, S. (2020). The IntCal20 Northern Hemisphere radiocarbon age calibration curve (0–55 cal kBP). *Radiocarbon*, 62. (4), 725-757. doi :10.1017/RDC.2020.41

Robbins, J.A., Edgington, D.N. (1975). Determination of recent sedimentation rates in Lake Michigan using Pb-210 and Cs-137. *Geochim. Cosmochim. Acta*, 39, 285-304.

Smith, J.N. (2001). Why should we believe ²¹⁰Pb sediment geochronologies? *Journal of Environmental Radioactivity*, 55, 121–123.

Vermeesch, P. (2018). IsoplotR: A free and open toolbox for geochronology. *Geosci. Front.*, 9, 1479–1493.

Wirrmann, D., Sémah, A.M., Mendez-Millan, M., Schmidt, S., Boissenin, M., Boucher, H., Bouloubassi, I., Brescia, F., Cetin, F., Djouarev, I., Klein, V. (2017). Signification environnementale de guano de salanganes et de chiroptères de Nouvelle-Calédonie. Premiers résultats. *Quaternaire*, 28(3), 401-412.

Wurster, C.M., Patterson, W.P., McFarlane, D.A., Wassenaar, L.I., Hobson, K.A., Athfield, N.B., Bird, M.I. (2008). Stable carbon and hydrogen isotopes from bat guano in the Grand Canyon, USA, reveal Younger Dryas and 8.2 ka events. *Geology*, 36(9), 683-686.

- 547 Wurster, C.M., Bird, M.I., Bull, I.D., Bryant, C., Ascough, P. (2009). A protocol for
548 radiocarbon dating tropical subfossil cave guano. *Radiocarbon*, 51(3), 977-986.
549
- 550 Zúkal, J., Berková, H., Bandouchová, H., Kováčová, V., Píkula, J. (2017). Bats and caves :
551 Activity and ecology of bats wintering in caves. <http://dx.doi.org/10.5772/intechopen.69267>

552 **Table and Figure captions :**

553

554 Table 1 : ^{14}C results. Data are reported in « percent Modern Carbon » (pMC) for the period
555 affected by atmospheric nuclear tests. Ages are calculated from the OxCal program (see text
556 for details).

557

558 Table 2 : (^{210}Pb), (^{137}Cs) and (^{241}Am) activities/g in the guano samples, calculated in 2014,
559 with their 2σ uncertainties. ^{14}C ages are reported together with the ages derived from Pb, Cs
560 and Am data. ^{14}C ages in bold characters correspond to the measurements reported in Table 1;
561 the ages in italic are obtained by linear interpolation between two successive ^{14}C dates. The
562 dates in bold characters in the “Pb, Cs, Am dates” column represent anchor dates for the
563 derivation of the age model.

564

565 Fig.1: Map of the Brantites III cave. Inset: location of the cave in the SW France.

566

567 Fig.2: a) vertical cross-section of the guano deposit. The guano heap is 1.75 m high.

568 b) Detail showing the staggered arrangement of the cores from C1 at the top to C3 at
569 the bottom. Open black and yellow circles show examples of corresponding layers in C1-C2
570 and C2-C3 cores, respectively. [colour pictures].

571

572 Fig. 3: (^{210}Pb)_{ex} activities/g vs depth. The whole distribution is divided into 4 segments S1 to
573 S4. Only segments 2 and 4 are interpreted as reflecting a roughly constant accumulation rate
574 (2.59 ± 0.28 cm/y and 0.30 ± 0.04 cm/y, respectively). Uncertainties on slopes and derived
575 accumulation rates are calculated using the IsoplotR program (Vermeesch, 2018). Data for S1
576 are assumed affected by ^{210}Pb mobility (see text for further discussion).

577

578 Fig. 4: (^{137}Cs) and (^{241}Am) activities/g vs depth. For clarity, only the average 2σ uncertainties
579 have been reported.

580

581 Fig. 5: Initial (^{210}Pb)_{0,ex} at the time of guano deposition vs age. Note the higher values before
582 1960. The data for the post-1990 layers are not reported since they are affected by ^{210}Pb
583 mobility (see text for a detailed discussion).

584

585 Fig. 6: ^{14}C and ^{210}Pb - ^{137}Cs - ^{241}Am age models for the guano deposit.

586

587

¹⁴C lab. n°	Sample layer	Depth (cm)	pMC values or age BP	calibrated age AD
Poz-62924	C1 (5-6)	7	105.85 ± 0.33	2005 ± 2
Poz-62926	C2 (5-6)	34.5	113.03 ± 0.31	1993.5 ± 1.5
Poz-134057	C1 (28-29)	37.5	113.77 ± 0.35	1991.5 ± 1.5
Poz-62927	C3 (10-11)	63.5	125.23 ± 0.33	1981 ± 1
Poz-134058	C2 (47-48)	76.5	130.46 ± 0.32	1978.5 ± 0.5
Poz-62928	C4 (10-11)	97	156.37 ± 0.35	1969 ± 1
Poz-134059	C3 (46-47)	102	166.57 ± 0.44	1963.4 ± 0.5
Poz-62929	C4 (50-51)	138	130 ± 30 BP	1871 ± 72

588

589 Table 1

Samples	Depth z (cm)	(²¹⁰ Pb) _m mBq/g	(²²⁶ Ra) _m mBq/g	(²¹⁰ Pb) _{ex,2014} mBq/g	(¹³⁷ Cs) ₂₀₁₄ mBq/g	(²⁴¹ Am) ₂₀₁₄ mBq/g	¹⁴ C dates (cal AD)	Pb,Cs,Am dates AD	(²¹⁰ Pb) _{0,ex} mBq/g
Top layer	0.0	51.8	4.3	58.9 ± 6.6	<DL	<DL	2011	2011	
C1 (0-1)	0.5	41.7	3.2	47.3 ± 5.7	<DL	<DL	2010		
C1 (5-6)	7.0	53.5	2.8	62.2 ± 6.5	0.9 ± 0.5	<DL	2005.0 ± 1.5		
C1 (11-12)	15.0	64.3	1.5	77.1 ± 6.1	0.2 ± 0.4	<DL	2001.7		
C1 (21-22)	28.0	75.8	0.8	92.0 ± 7.8	1.0 ± 0.5	<DL	1996.2		
C2 (5-6)	34.5	nd	nd	nd	<DL	<DL	1993.5 ± 1.5		
C1 (28-29)	37.5	78.5	1.5	94.6 ± 6.6	<DL	<DL	1991.5 ± 1.5		
C2 (13-14)	42.5	95.8	2.8	114.5 ± 7.4	<DL	<DL	1989.4	1985.9 ± 5.0	14.71 ± 0.95
C2 (14-15)	43.5	85.8	4.3	101.0 ± 8.7	0.8 ± 0.6	<DL	1989.0	1985 ± 4.9	13.13 ± 0.93
C2 (15-16)	44.5	67.2	3.3	78.6 ± 7.4	<DL	<DL	1988.6	1985.1 ± 4.8	10.36 ± 0.76
C2 (16-17)	45.5	69.0	2.5	81.9 ± 7.4	<DL	<DL	1988.2	1984.7 ± 4.7	10.93 ± 0.79
C3 (10-11)	63.5	62.8	4.8	71.4 ± 5.7	1.2 ± 0.4	<DL	1981.0 ± 1.0	1977.7 ± 3.3	11.91 ± 0.61
C2 (47-78)	76.5	63.7	5.8	71.1 ± 5.3	1.1 ± 0.4	<DL	1978.5 ± 0.5	1972.7 ± 2.3	12.83 ± 0.52
C3 (30-31)	85.0	57.2	3.0	67.1 ± 5.4	2.3 ± 0.5	<DL	1974.7	1969.5 ± 1.7	13.64 ± 0.73
C3 (39-40)	94.5	51.3	5.3	55.1 ± 5.8	3.3 ± 0.6	<DL	1970.2	1965.9 ± 1.1	12.84 ± 0.81
C3 (40-41)	95.5	54.2	10.0	57.4 ± 6.2	3.6 ± 0.6	0.56 ± 0.30	1969.7	1965.4 ± 1.1	13.57 ± 0.88
C3 (41-42)	96.5	57.5	11.5	56.3 ± 6.2	4.1 ± 0.7	<DL	1969.3	1965.0 ± 1.1	13.52 ± 0.89
C4 (10-11)	97.0	52.3	7.2	58.2 ± 5.1	4.2 ± 0.6	<DL	1969.0 ± 1.0	1964.8 ± 1.0	14.09 ± 0.76
C3 (42-43)	97.5	52.8	5.0	59.3 ± 5.8	6.9 ± 0.7	0.86 ± 0.50	1968.3	1964.6 ± 1.0	14.71 ± 0.88
C3 (43-44)	98.5	39.5	3.2	45.2 ± 5.0	7.5 ± 0.6	1.44 ± 0.40	1967.1	1964.2 ± 1.0	11.78 ± 0.77
C3 (44-45)	99.5	47.5	2.7	55.5 ± 5.8	10.4 ± 0.7	1.66 ± 0.42	1965.9	1963.8 ± 1.0	15.20 ± 0.95
C3 (45-46)	100.5	42.3	3.2	48.4 ± 5.8	17.2 ± 1.0	2.07 ± 0.44	1964.6	1963.4 ± 0.5	14.25 ± 0.88
C3 (46-47)	101.5	39.5	7.2	39.8 ± 5.3	16.5 ± 0.9	2.73 ± 0.50	1963.4 ± 0.2	1963.0 ± 1.0	12.01 ± 1.31
C3 (47-48)	102.5	41.7	9.2	40.0 ± 4.1	15.5 ± 0.7	2.18 ± 0.32	1960	1962.6 ± 1.0	12.33 ± 1.23
C4 (16-17)	103.0	45.5	16.5	44.3 ± 5.1	12.8 ± 0.8	1.66 ± 0.38	1959	1962.5 ± 1.0	13.72 ± 1.42
C3 (48-50)	103.5	40.2	5.2	36.0 ± 5.4	9.0 ± 0.7	1.54 ± 0.42	1958	1962.3 ± 1.0	11.29 ± 1.28
C4 (17-18)	104.0	37.3	5.8	39.7 ± 4.2	6.4 ± 0.6	1.25 ± 0.30	1956	1961.0 ± 5.0	12.57 ± 1.27
C4 (18-19)	105.0	39.8	4.5	44.8 ± 4.6	3.5 ± 0.6	0.86 ± 0.32	1954	1960.0 ± 4.0	14.50 ± 1.46
C4 (19-20)	106.0	42.3	2.3	49.7 ± 5.0	2.5 ± 0.5	0.94 ± 0.38	1951	1959.0 ± 3.0	16.46 ± 1.64
C4 (20-21)	107.0	49.8	4.3	56.5 ± 6.2	3.0 ± 0.7	0.94 ± 0.38	1949	1958.2 ± 2.7	19.13 ± 1.92
C4 (21-22)	108.0	43.3	2.8	50.4 ± 4.6	1.7 ± 0.4	0.91 ± 0.36	1946	1954.8 ± 2.2	18.93 ± 1.55
C4 (22-23)	109.0	36.0	3.0	41.0 ± 5.4	1.1 ± 0.5	0.19 ± 0.20	1943	1951.5 ± 2.0	17.10 ± 1.55
C4 (23-24)	110.0	40.3	3.0	46.5 ± 5.0	0.6 ± 0.3	0.31 ± 0.30	1941	1948.2 ± 2.2	21.47 ± 1.87
C4 (24-25)	111.0	33.7	6.7	33.5 ± 3.7	0.6 ± 0.4	0.22 ± 0.18	1938	1944.8 ± 2.7	17.17 ± 1.73
C4 (39-40)	126.0	20.5	14.3	7.7 ± 4.1	<DL	<DL	1899	1895 ± 15	18 ± 10
C4 (50-51)	137.0	30.7	27.0	4.5 ± 7.0	<DL	<DL	1871 ± 72	1858 ± 25	34 ± 38

Table 2



Figure 1

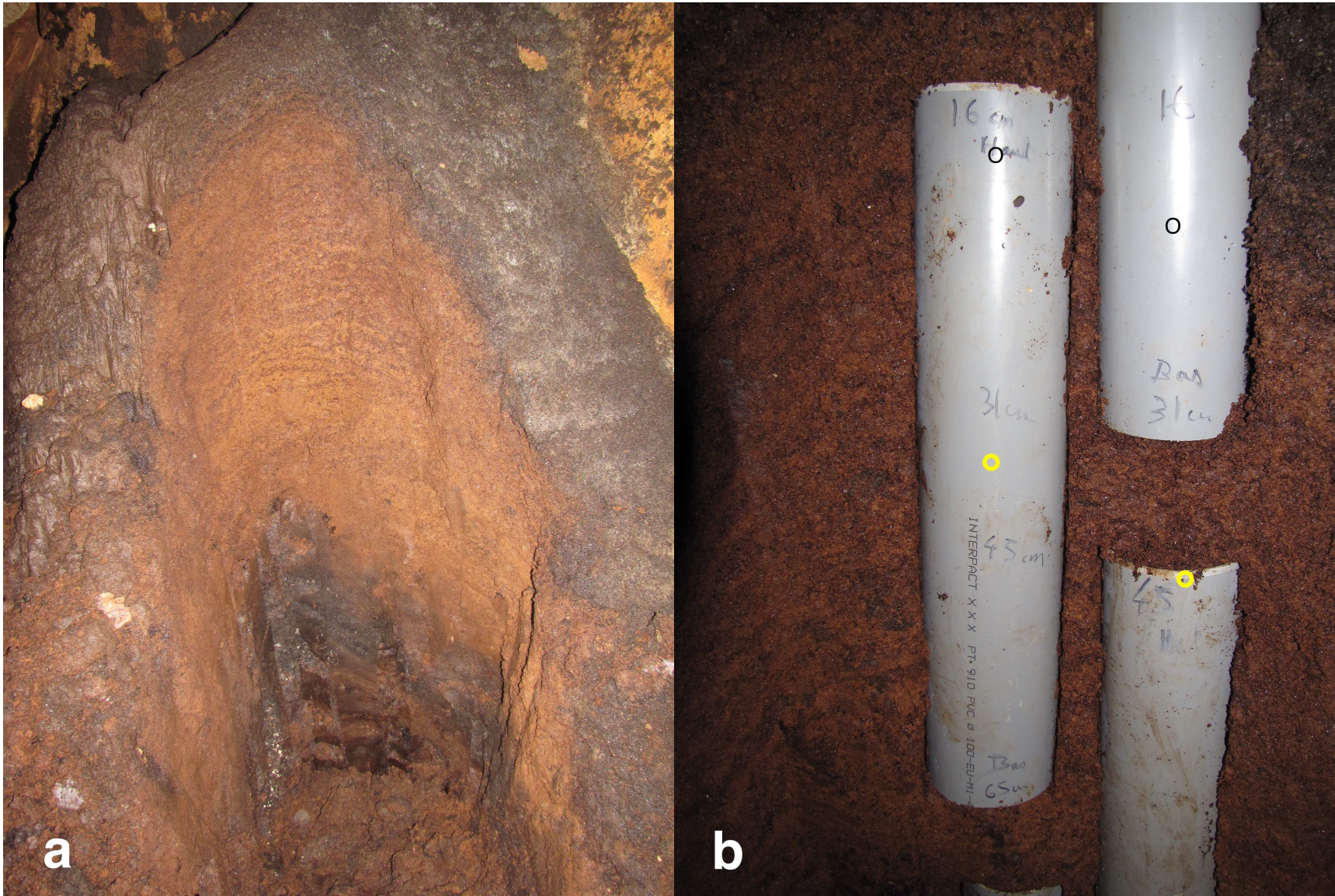


Figure 2

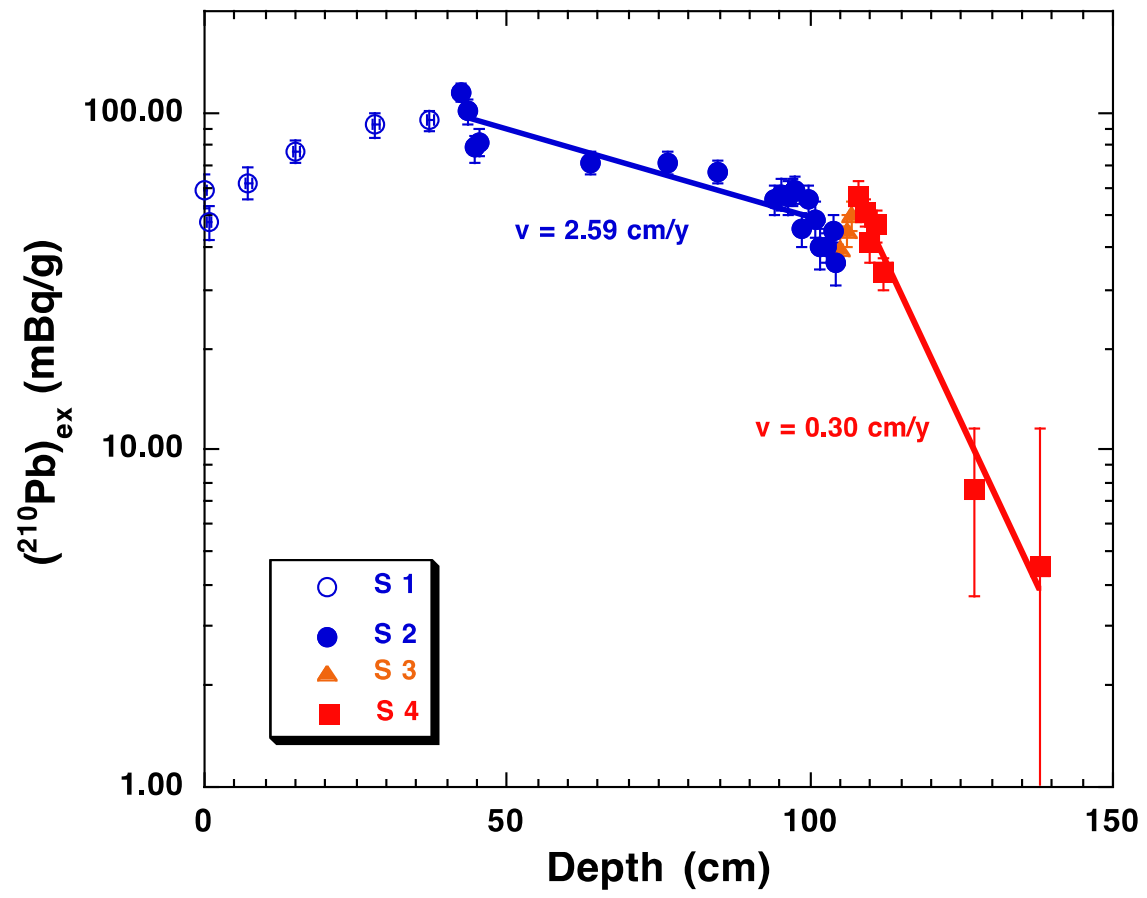


Figure 3

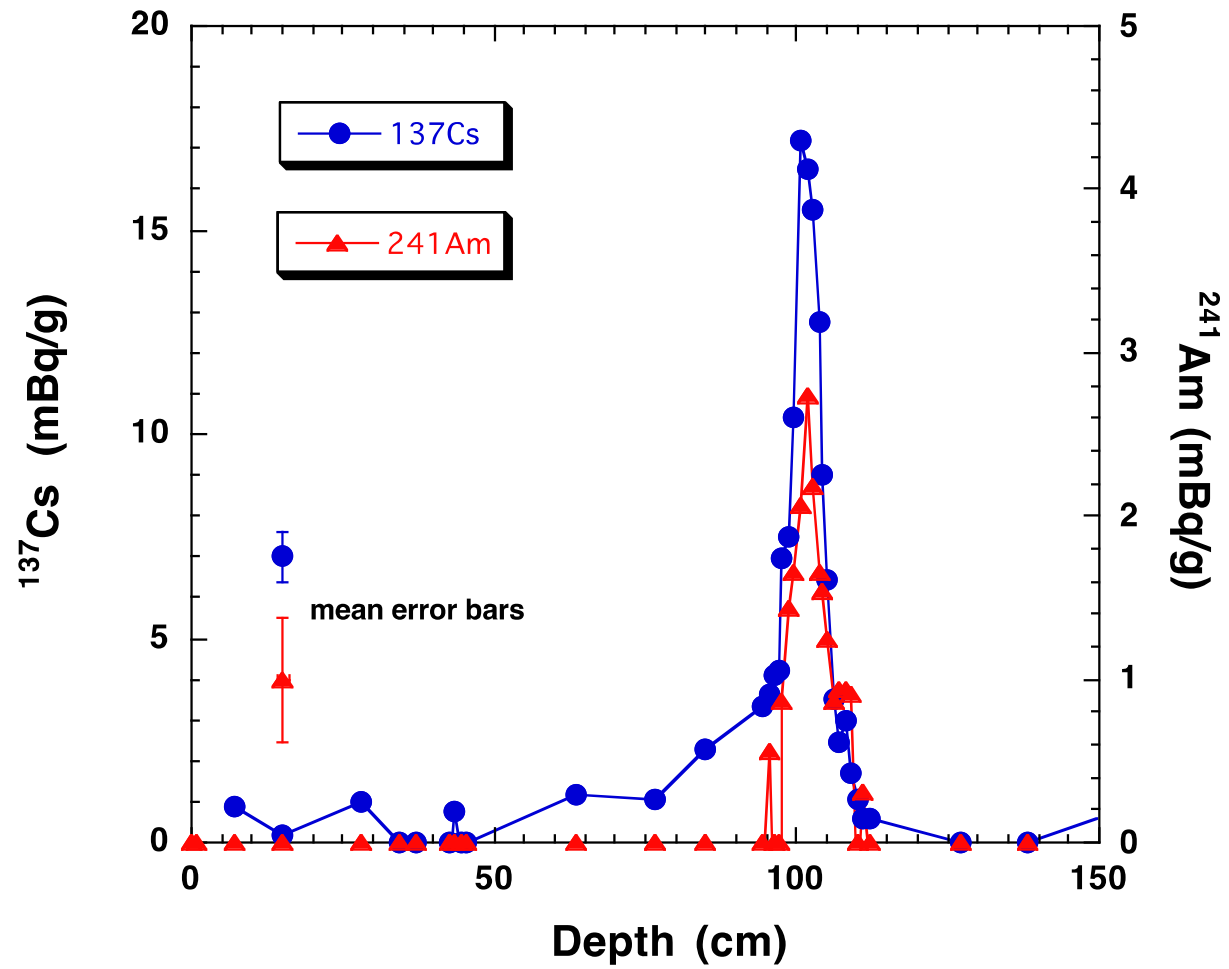


Fig. 4

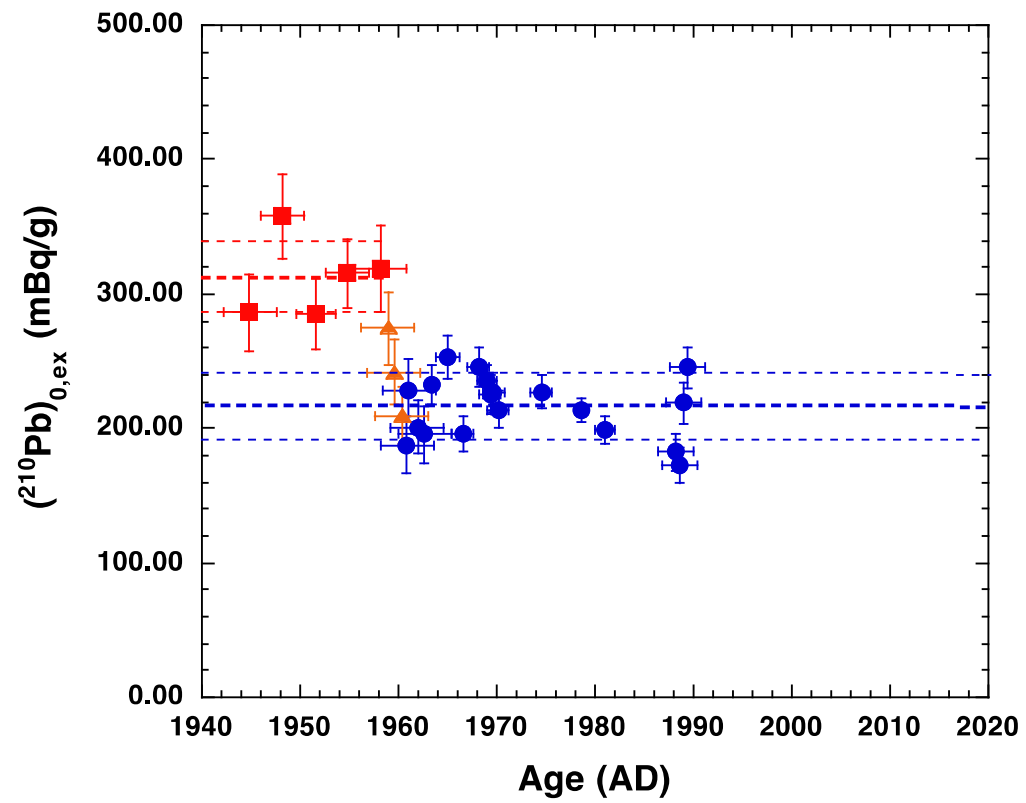


Fig. 5

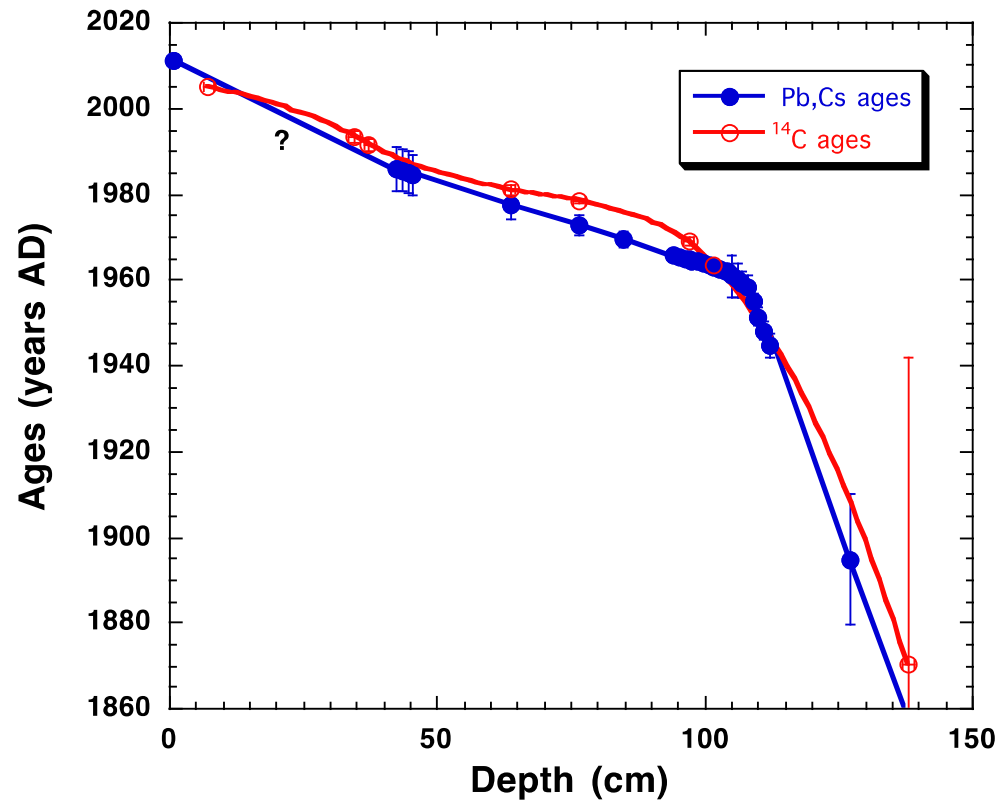


Figure 6