



Exploring the architectures of nearby exoplanetary systems using high-precision space astrometry

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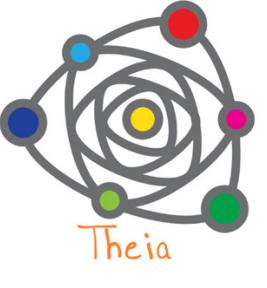
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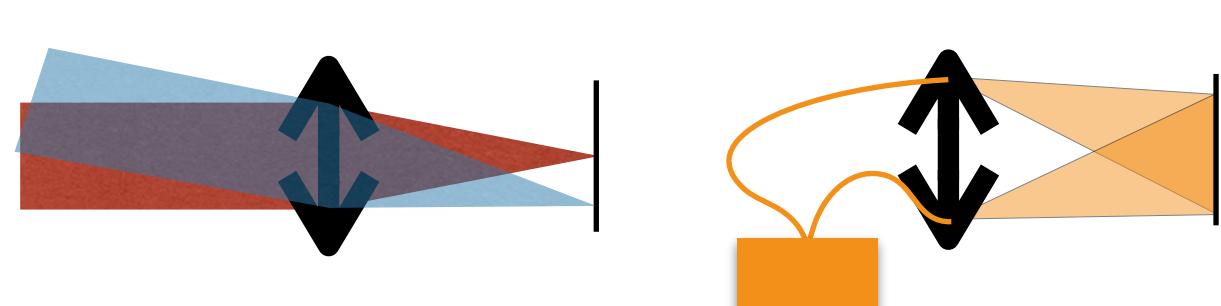
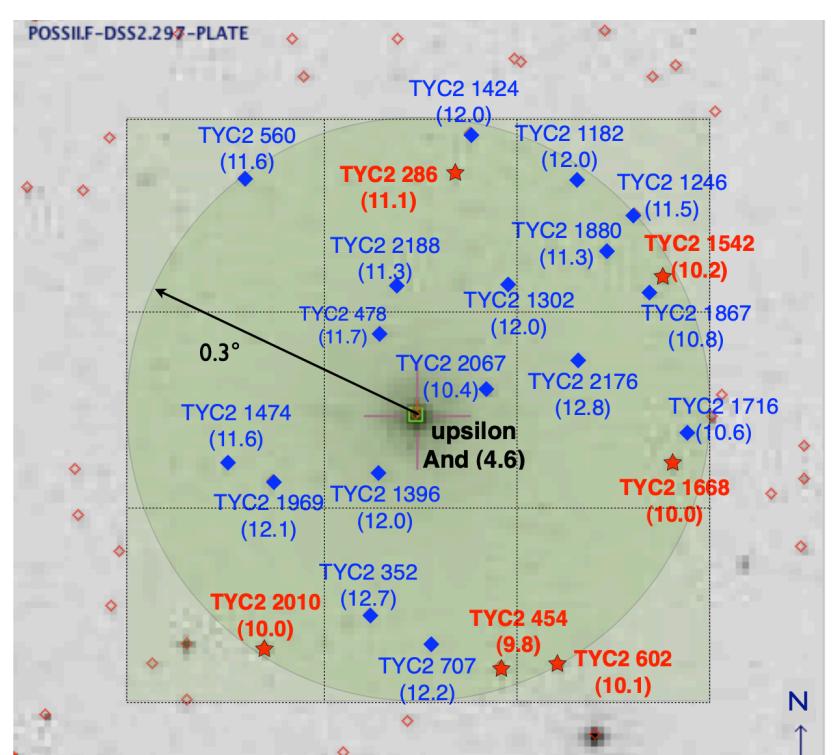


Exploring the architectures of nearby exoplanetary systems using high-precision space astrometry

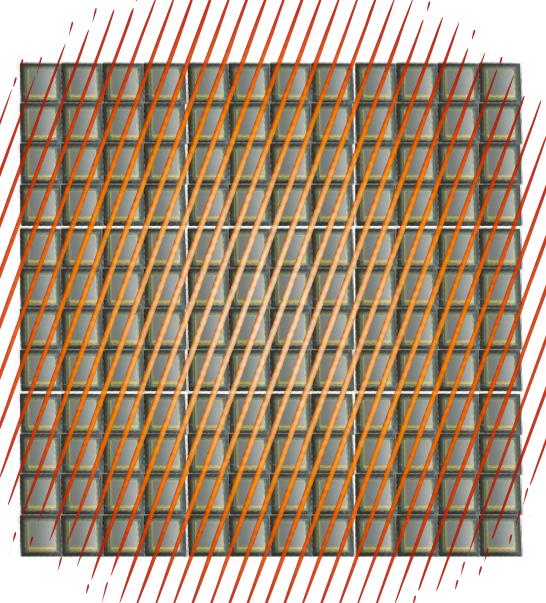
F. Malbet (1), M. Lizzana (1,2), A. Léger (3), G. Mamon (4), L. Labadie (5), A. Sozzetti (6), T. Lépine (7), R. Goullioud (8), and Theia collaboration (9)

See poster 1180 (M. Lizzana)
for instrumental results

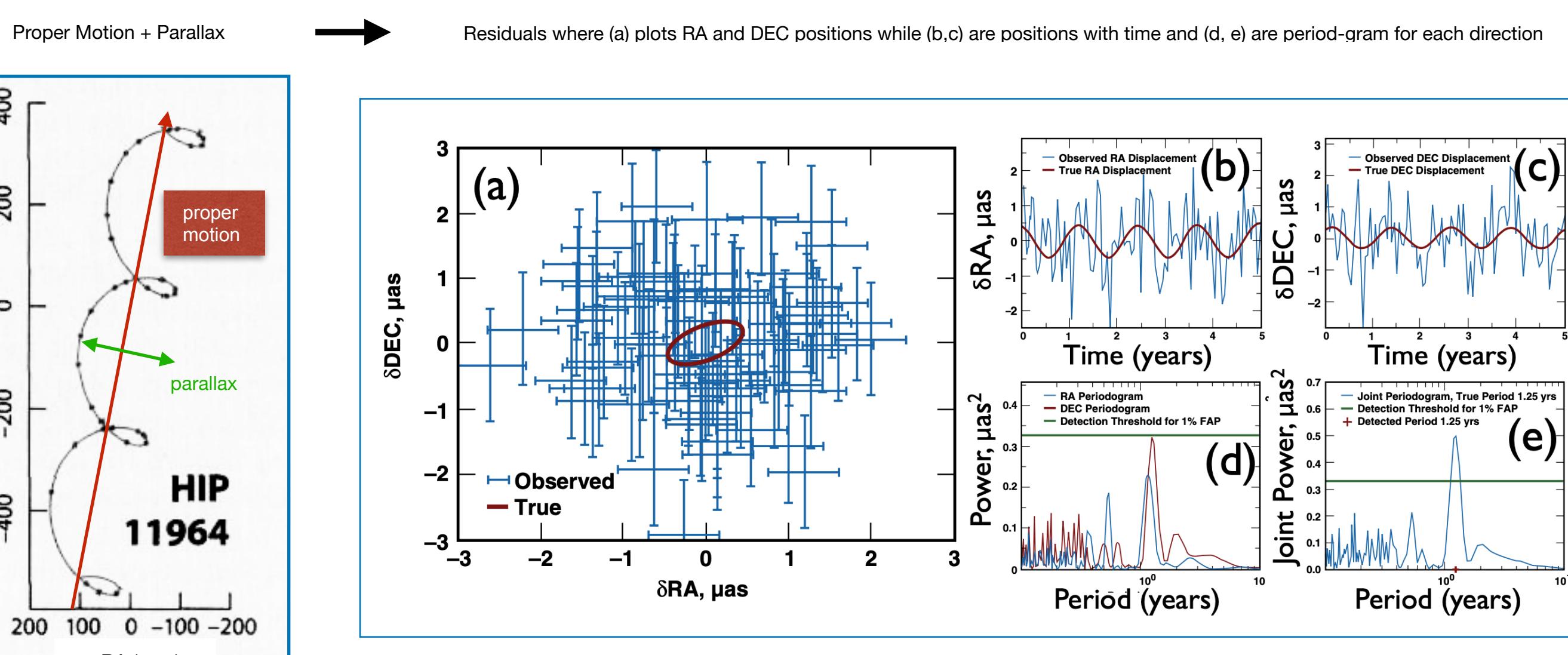
Principle of Theia : relative and precise differential astrometry



- The science target is at the center of the FOV
- We measure the offset angles between the target object (in black) and the reference stars (other stars in red and blue)
- The single measurement is repeated several times over the mission lifetime to reach the required accuracy on the target motion
- The detector calibration is done using interferometric modulated Young's fringes



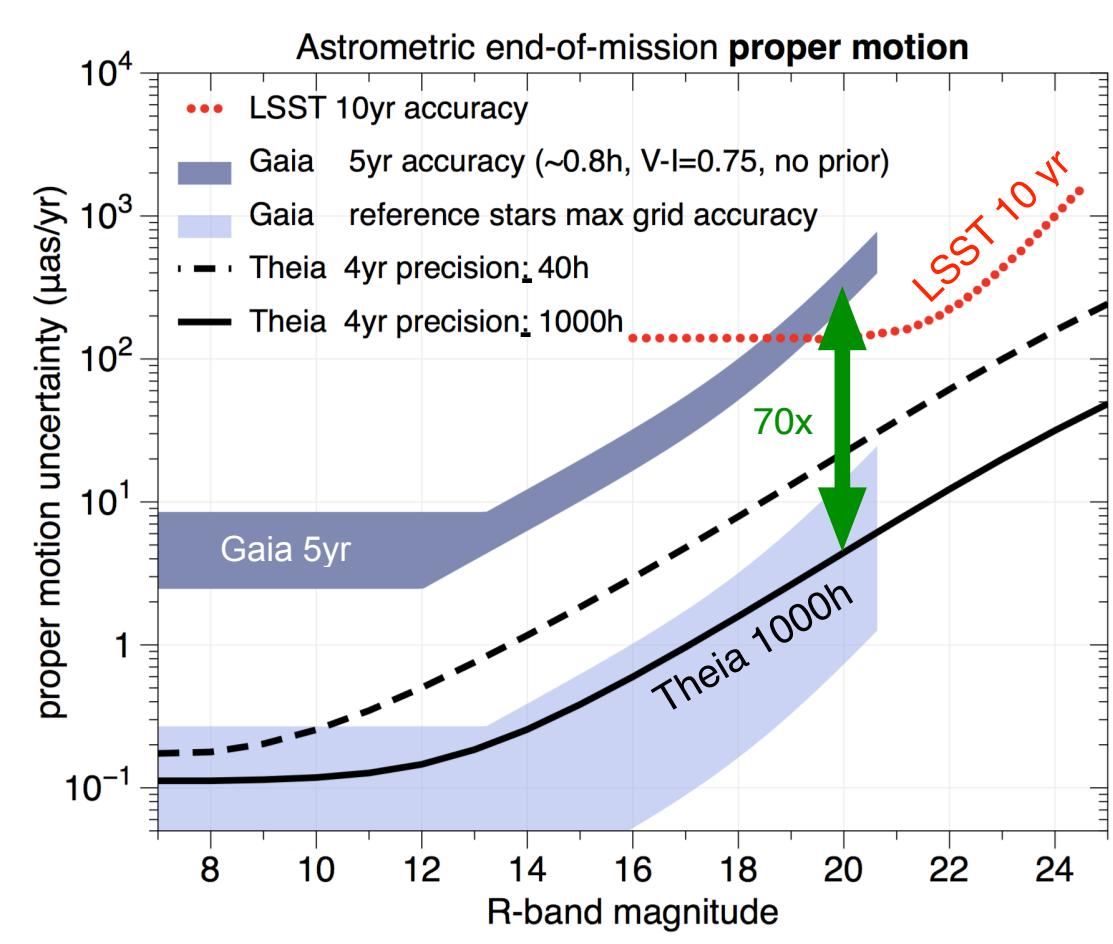
Astrometry: parallax, proper motion and reflex motion



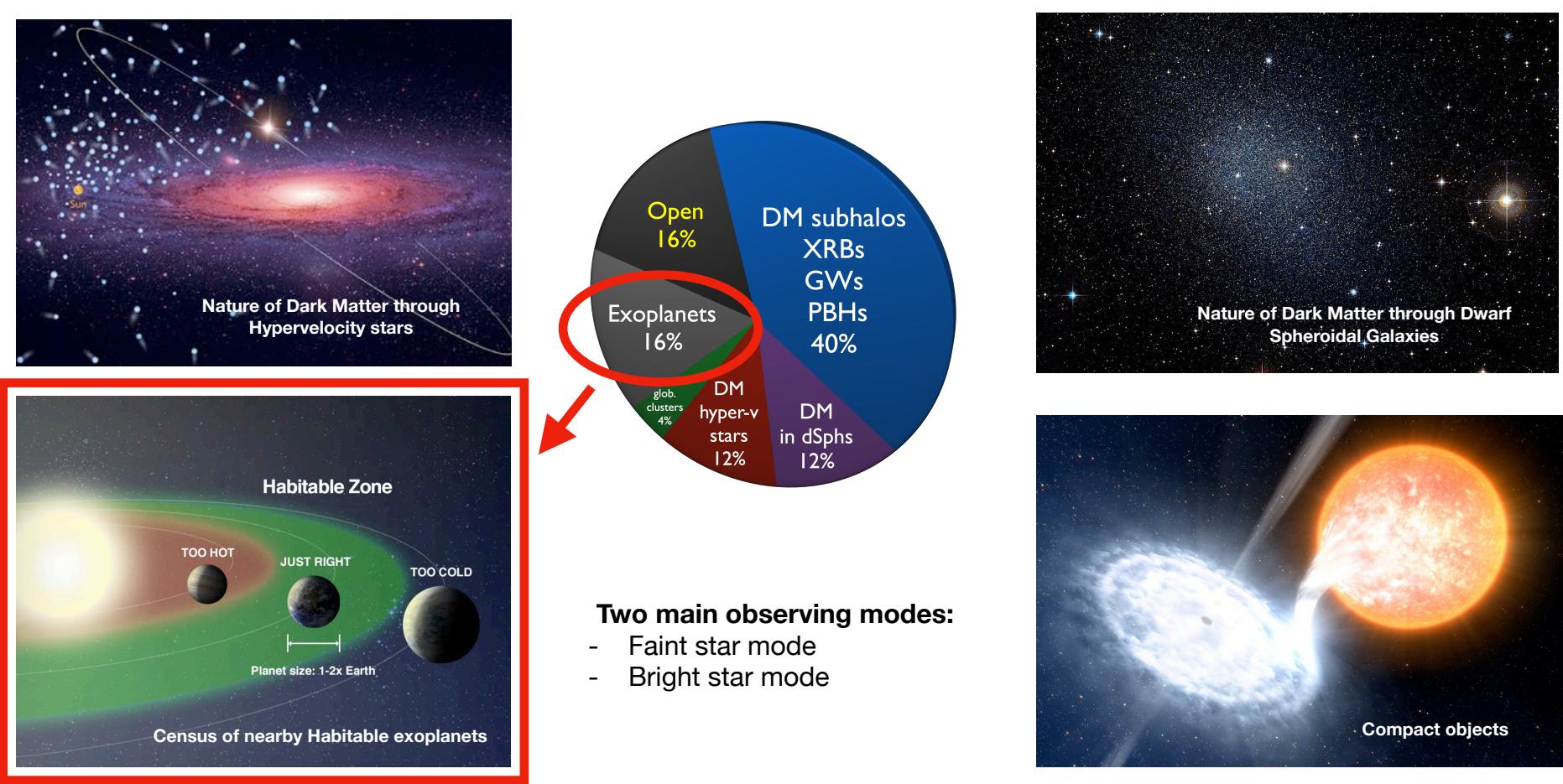
Unwin et al. (2008); Malbet et al. (2012)

Improvement compared to Gaia

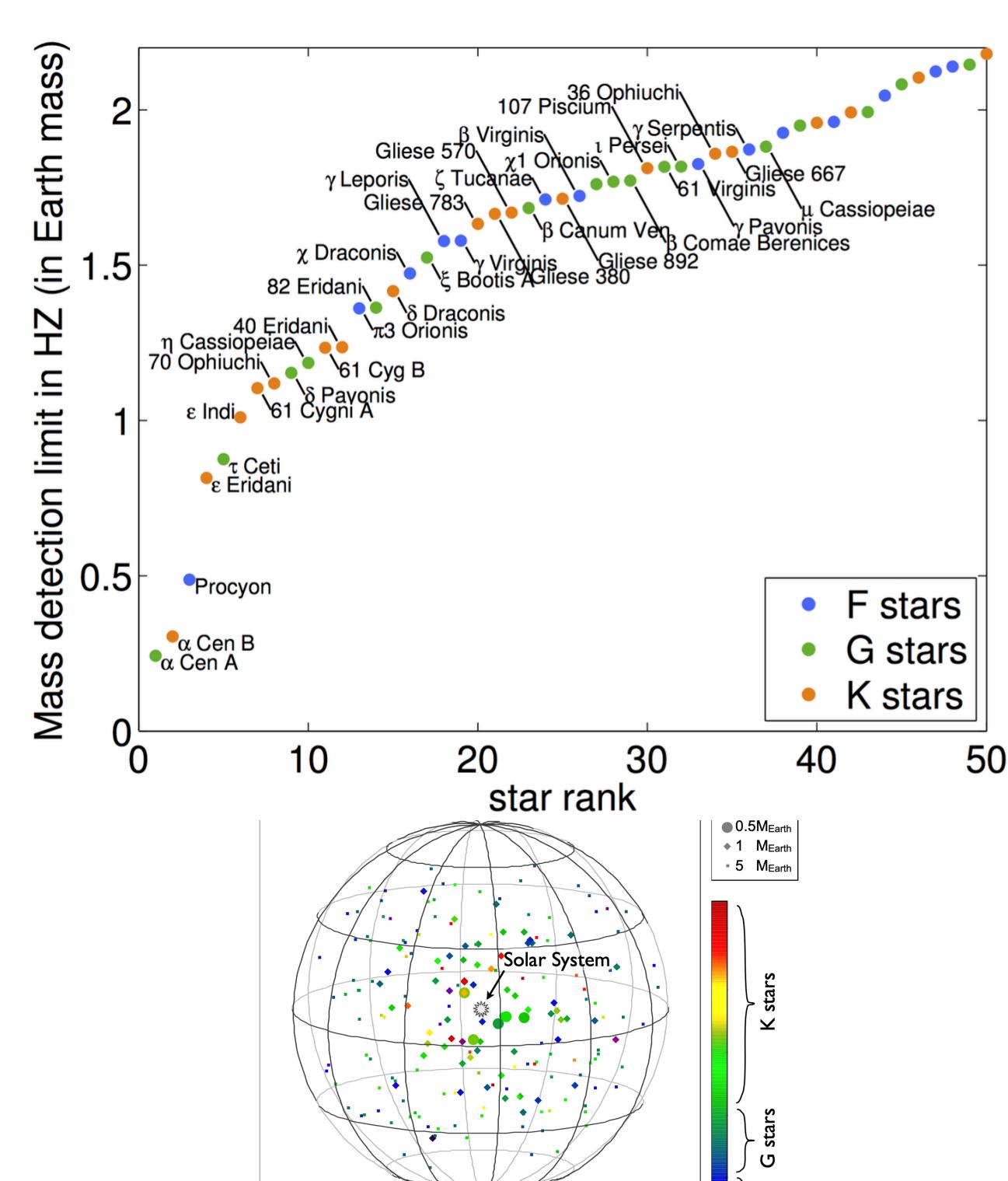
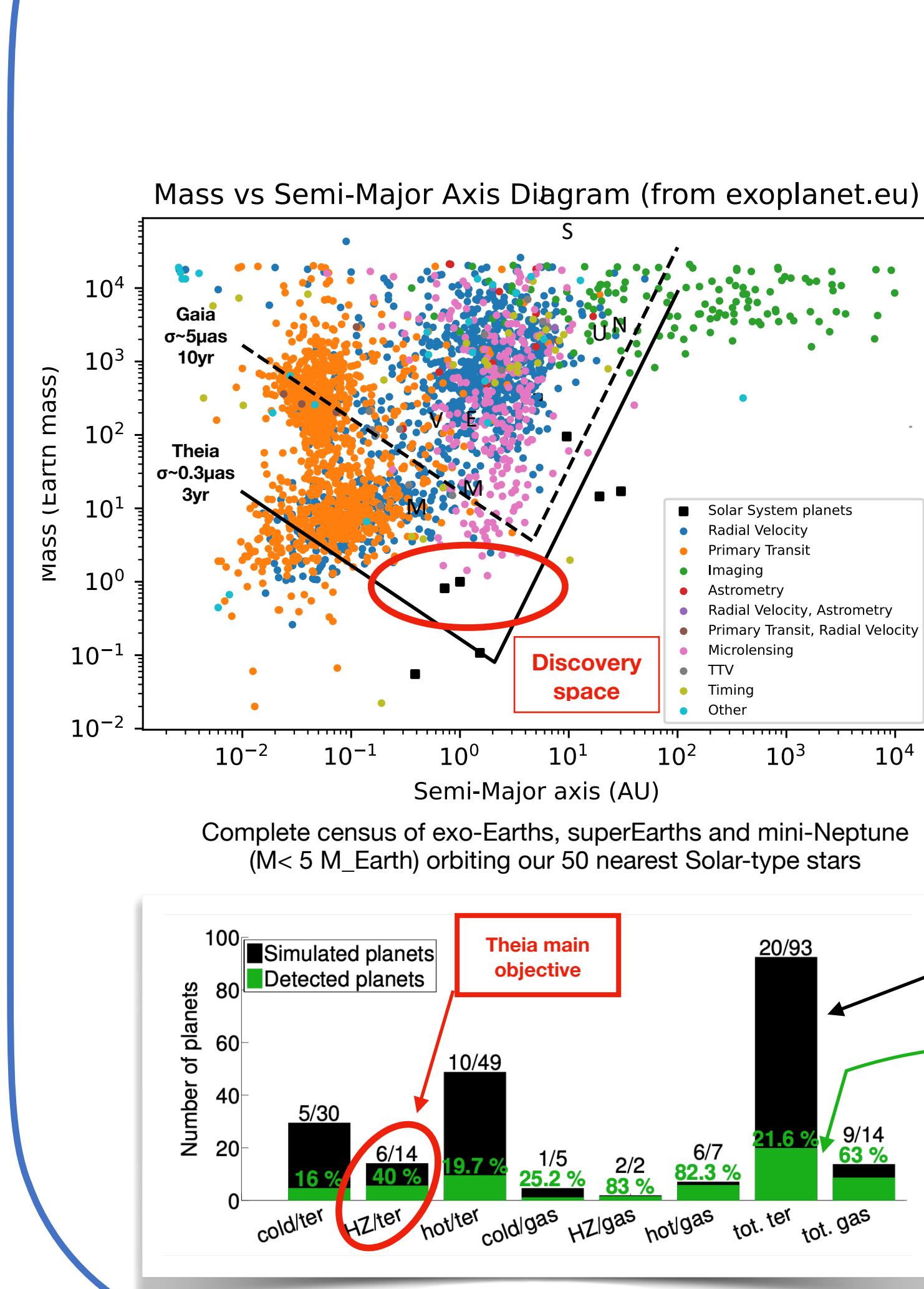
	Gaia post-launch	Theia	improvement
Telescope Aperture	1.49 x 0.54 = 0.80 m ²	0.8 → 0.50 m ²	0.62
Field of view		0.5 deg	
Coverage	Survey	Pointed	
Astrometry	Global	Differential	
Exposure time per field	76 x 9 (CCD) x 4.4 sec = 0.84 hr	25 x 40 hr = 1000 hr	1200
Additional factors	2D vs 1D astrometry: 1.4, Gaia Stray light: 1.9		
Proper motion accuracy G=10 star	3.5 μas/yr	0.11 μas/yr	30
Proper motion accuracy G=15 star	14 μas/yr	0.4 μas/yr	35
Proper motion accuracy G=20 star	330 μas/yr	5 μas/yr	66



Theia's main science based on high precision astrometry

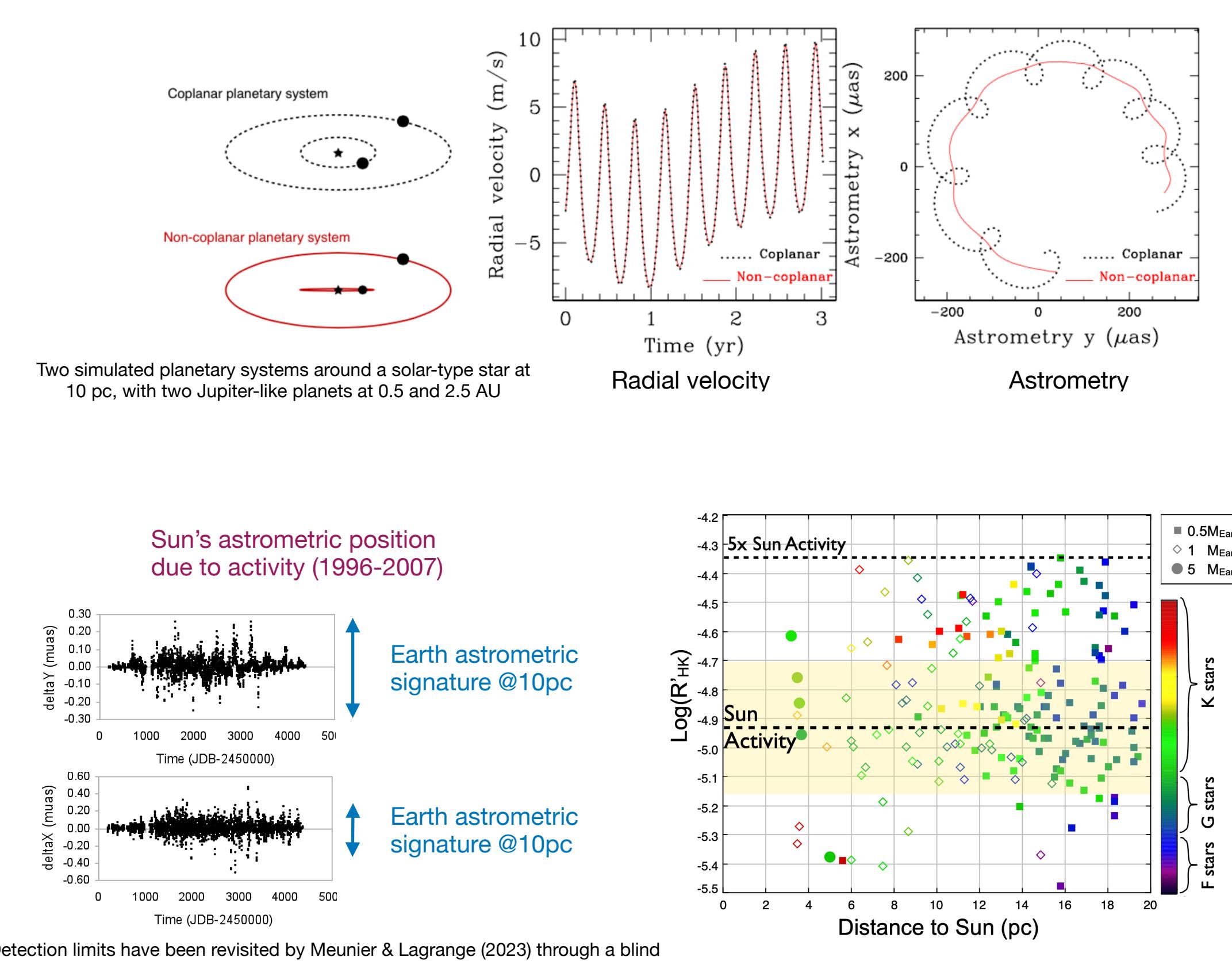


Exoplanets primary goal : Earth-mass planets in HZ in the vicinity of the Sun



Averaging 10000 random exoplanets distributions around the 50 target stars (Crouzier 2014).

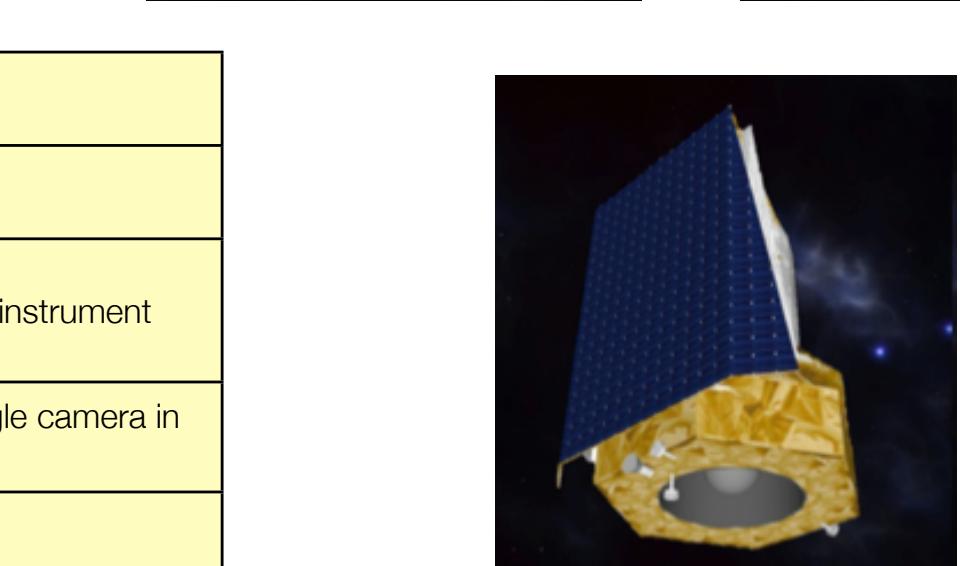
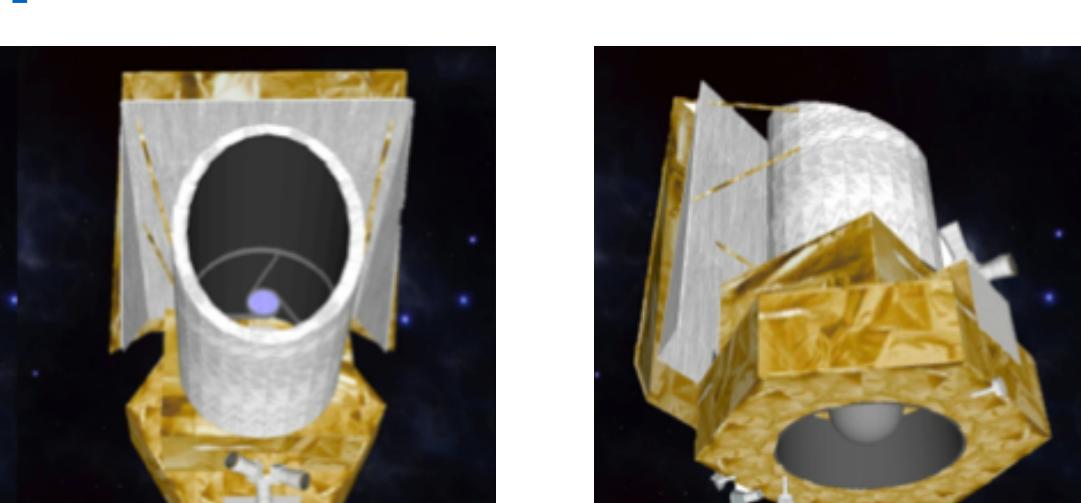
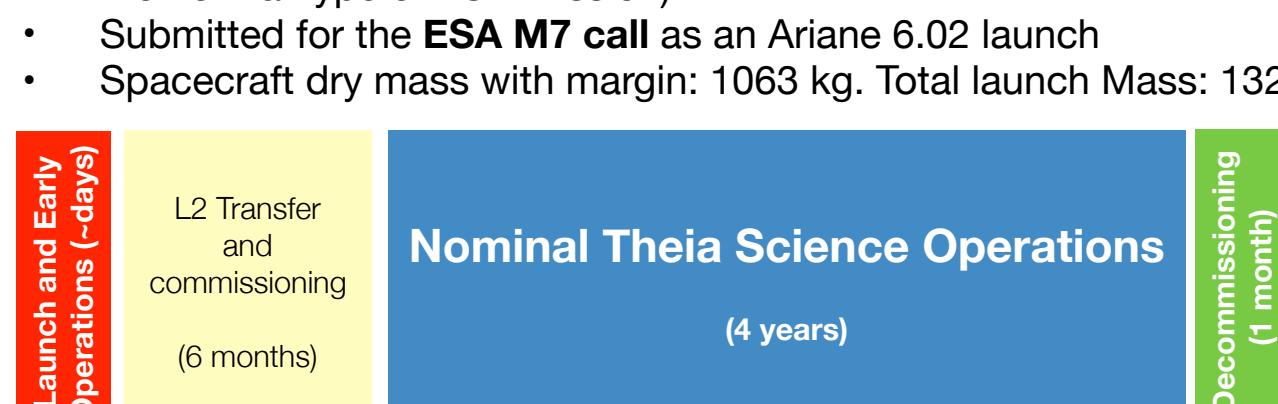
Advantages and limitations of high precision astrometry



Detection limits have been revisited by Maunier & Lagrange (2023) through a blind test method and found lower detection limits, except for a handful of stars, showing that it is probably possible to reach lower masses.

Theia mission profile

- ESA-led, ESA-operated mission with consortium funded payload (this is the normal type of ESA mission)
- Submitted for the **ESA M7 call** as an Ariane 6.02 launch
- Spacecraft dry mass with margin: 1063 kg. Total launch Mass: 1325 kg



- Univ. Grenoble Alpes CNRS IPAG, Grenoble, France
- CNES, Paris, France
- Univ. Paris Saclay CNRS IAS, Orsay, France
- Sorbonne Univ. CNRS IAP, Paris, France
- Univ. Cologne, Cologne, Germany
- INAF Oss. Di Torino, Turin, Italy
- Univ. Jean Monnet, CNRS, IOGS, Curien lab, Saint Etienne, France,
- JPL, Pasadena, USA
- <http://theia.phyip3.dur.ac.uk/>

Theia instrumental concept

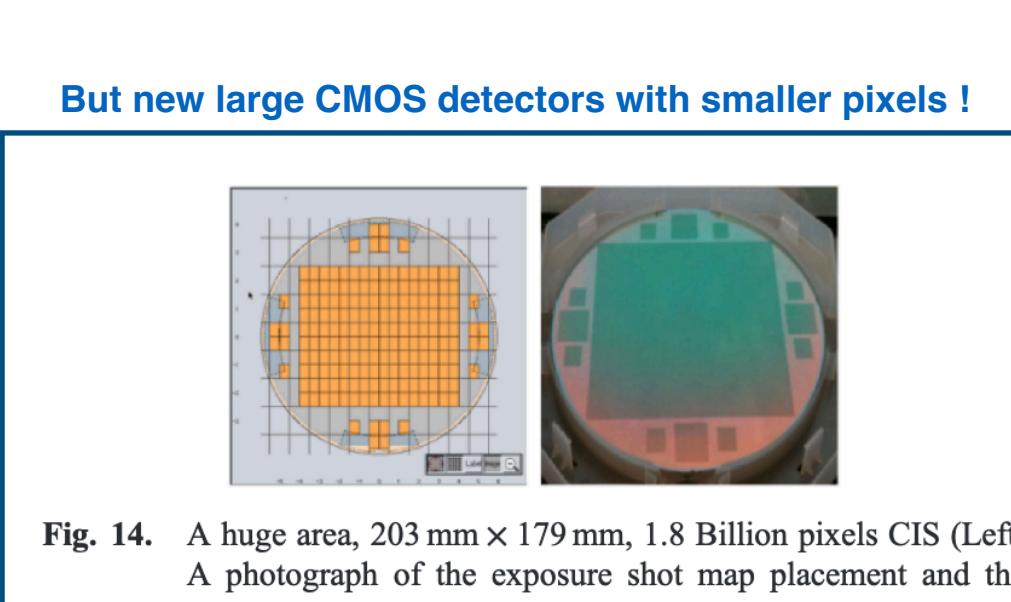
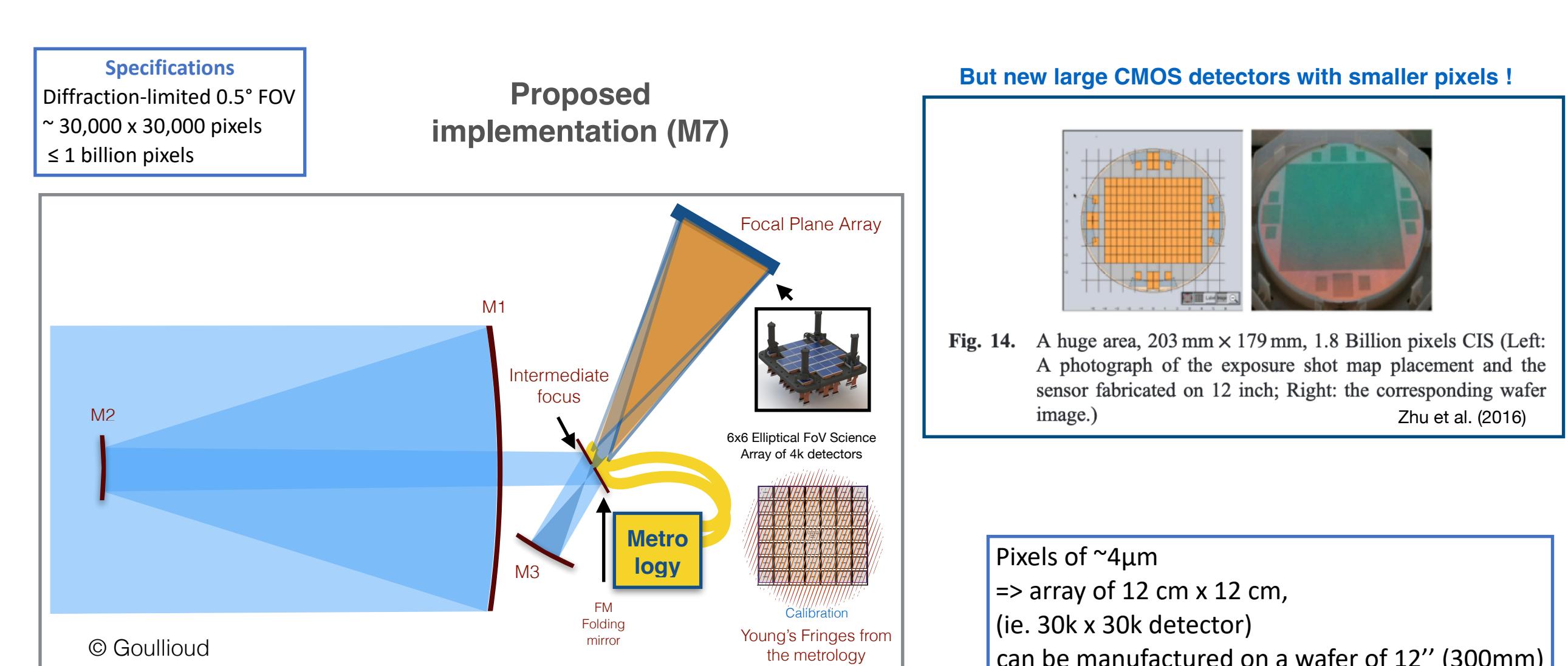


Fig. 14. A huge area, 203 mm x 179 mm, 1.8 Billion pixels CIS (Left: A photograph of the exposure shot map placement and the sensor fabricated on 12 inch; Right: the corresponding wafer image.) Zhu et al. (2016)

Pixels of ~4μm
=> array of 12 cm x 12 cm,
(i.e. 30k x 30k detector)
can be manufactured on a wafer of 12" (300mm)

References

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