

## Depositional processes and environments of the Layered Sulfate-bearing unit, Gale crater, Mars

Amelie L. Roberts, Sanjeev Gupta, Aster Cowart, Lauren Edgar, William E. Dietrich, William Rapin, Edwin S. Kite, Steven G. Banham, Joel M. Davis, Claire A. Mondro, et al.

## ▶ To cite this version:

Amelie L. Roberts, Sanjeev Gupta, Aster Cowart, Lauren Edgar, William E. Dietrich, et al.. Depositional processes and environments of the Layered Sulfate-bearing unit, Gale crater, Mars. Tenth International Conference on Mars 2024, Lunar and Planetary Institute, Jul 2024, Pasadena (California), United States. pp.3515. hal-04650564

## HAL Id: hal-04650564 https://hal.science/hal-04650564v1

Submitted on 16 Jul2024

**HAL** is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers. L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

## Depositional processes and environments of the Layered Sulfate-bearing unit, Gale crater, Mars

Amelie L Roberts<sup>1\*</sup>, Sanjeev Gupta<sup>1</sup>, Aster Cowart<sup>2</sup>, Lauren A Edgar<sup>3</sup>, William E Dietrich<sup>4</sup>, William Rapin<sup>5</sup>, Edwin Kite<sup>6</sup>, Steven G Banham<sup>1</sup>, Joel M Davis<sup>1</sup>, Claire Mondro<sup>7</sup>, Bill Farrand<sup>8</sup>, Alexander B Bryk<sup>4</sup>, Tex Kubacki<sup>9</sup>, Natalie Moore<sup>9</sup>, Rachel Sheppard<sup>2</sup>, Linda Kah<sup>10</sup>, Gwénaël Caravaca<sup>5</sup>, Abigail Fraeman<sup>11</sup>, Ashwin Vasavada<sup>11</sup>.

<sup>1</sup>Imperial College London (\*<u>alr21@ic.ac.uk</u>), <sup>2</sup>Planetary Science Institute, <sup>3</sup>USGS, <sup>4</sup>UC Berkeley, <sup>5</sup>IRAP-CNRS, <sup>6</sup>University of Chicago, <sup>7</sup>Caltech, <sup>8</sup>Space Science Institute, <sup>9</sup>Malin Space Science Systems, <sup>10</sup>University of Tennessee, Knoxville, <sup>11</sup>NASA JPL.

**Introduction:** From orbit, a shift from clay-bearing to sulfate-bearing stratigraphy is recorded in the layered sedimentary rocks of the central mound of Gale crater [1]. This transition potentially records a shift from wet to dry conditions as seen from orbit elsewhere on Mars [2]. The Mars Science Laboratory *Curiosity* rover can test these hypotheses *in-situ*. Here, we describe the sed-imentary facies and architecture within the sulfate-bearing strata exposed in Marker Band valley and Gediz Vallis explored by Curiosity between Sols 3560 to 3917 (elevations -3880 to -3775 m) to interpret their depositional processes and environments.

**Stratigraphic context:** The clay-bearing stratigraphy, found in the Murray and Caroline Shoemaker formations, was predominantly deposited in aqueous palaeoenvironments [e.g. 3, 4]. So far, *Curiosity's* exploration of the overlying clay-sulfate transition, within the Mirador formation, reveals a transition to aeolian deposits [e.g. 5; 6] with intermittent aqueous activity [7].

Marker Band valley transitions into Gediz Vallis and is eroded into the sulfate-bearing strata. The stratigraphy here is subdivided into the (1) Contigo, (2) Catrimani, (3) Amapari, and (4) Chenapau members of the Mirador formation (Fig. 1).

Unit architecture: Here, we summarise the  $\sim$ 70 m of vertical stratigraphy exposed in Marker Band valley (Fig. 1) which we subdivide into 11 units labelled A through K (Fig. 1) which remain broadly, laterally consistent across the study area:  $\sim$ 1 km<sup>2</sup>.

The strata exposed at the base of Marker Band Valley comprises the upper part of the Contigo member. Here Unit A (>8 m thick) is characterized by m-scale sets of low-angle (apparent angles of 5-6°) cross-stratification with sharp, planar bounding surfaces. This is sharply overlain by the Catrimani member, Unit B (~5 m to ~15 m thick [2]), characterized by planar stratification. Between the Catrimani and Amapari members is a light-toned, mottled unit (Unit C; ~0.5 m thick) with a lower contact that penetrates into the underlying strata by a set of converging inclined fractures creating a sawtooth pattern [also see 8]. This unit is sharply overlain by the Amapari member (Unit D; ~1 m thick) containing a dark-toned, indurated band (~0.5 m thick) seen from orbit [9] and a recessive component (~0.5 thick). This is sharply overlain by the Chenapau member (>35 m thick) composed of dark- and light-toned stratigraphic sections which we have subdivided into 7 units, E through K. Units E (~3.6 m thick), G (2.7 m thick), I (~4.9 m thick), and K (>7.8 m thick) are slope-forming and appear mostly structureless from afar but occasionally preserve evidence of planar stratification. Units F (~4.2 m thick), H (~5.3 m thick), and J (~7.3 m thick) appear dark-toned and scarp forming. These characteristics are visible from orbit. Unit F contains localized corrugated textures and north-dipping bands. Units H and J are characterized by planar interstratified resistant bands with recessive strata sets which create shadows giving the dark-toned appearance. Truncation of lowangle sets are seen in Unit J.

Embedded within the stratigraphy of Unit E, there is an indurated, dark-toned asymmetric, lens ( $\sim$ 7 m x  $\sim$ 0.5 m) which pinches out laterally in a 'wing-shape'. Between units H and I, a set of nine distinct, broad, concave-up scour-and-fill structures ( $\sim$ 60 to  $\sim$ 40 m x  $\sim$ 1 to 4 m) are observed. The scours are simple, symmetric, with occasional planar bases, infilled by onlapping light-toned strata characteristic of Unit I. An additional scour truncates into Unit J and is overlain by Unit K.

**Description of nearfield facies observations:** Unit A is characterized by dark-toned bands separating recessive laminae consistent with the thicknesses of the inclined cross-strata. Bounding surfaces with associated asymptotic cross-laminae were identified along the traverse. Unit B is characterized by a facies with light-toned, planar, uniform, laterally extensive laminae (~0.74 mm thick). Both Units A and B are overprinted by <1 mm thick nodules obscuring grain size. In Unit C, a DRT underlying the Marker Band indicates a softer texture with mottled red, white, and green coloring. Unit D consists of a symmetric rippled and a thickly laminated component [6, 9, 10, 11].

The base of Unit E is sharp, overlain by strata with undulations (~50 cm in wavelength) and an abrupt change into a massive-appearing horizon composed of irregularly-shaped nodules that become planar up-section. The infill of the dark-toned lens within unit E consists of amalgamated bedsets of centimeter-scale tabular cross-stratification dipping to the north-west. Unit F is characterized by sets (~5-14 cm thick) of corrugated

3515.pdf

textures composed of wavy laminae which resemble climbing, asymmetric ripples. These have amplitudes of <1 mm and vary in wavelength by 4-8 cm. Units G through K have repeating facies associations. Units G, I, and K are all characterized by uniform, planar, laterally extensive, ~0.76 mm thick laminae. Each individual laminae preserves either a light-toned and dark-toned component (Fig. 2) creating a pin-stripe pattern. Units H and J are characterized by the gradual introduction of dark-toned, indurated, laterally discontinuous laminae which break up the uniform, planar, light-toned laminae into 12 - 1 cm thick sets. These dark-toned laminae have a smooth surficial texture and protrude out of the rock as elongated, parallel fins, spaced ~4-5 cm apart. Units H and J are associated with abundant, small disc-shaped features and more angular nodules.

Stratigraphic synthesis and implications: The transition recorded from Units A through D is interpreted to represent a shift in environment from the predominantly aeolian dune setting recorded in the claysulfate transition region. The cross-strata of Unit A is the preserved expression of migrating, simple dunes. The dark-toned bands are interpreted as preferential diagenesis from changes in grain size and sorting associated with dune depositional processes [12]. The extent, uniformity, and planar expression of the laminae are characteristic of the migration of wind ripples in an aeolian sand-sheet environment where dune formation was inhibited by near-surface water [e.g. 13]. The mottled texture, red/white coloring and fracture morphology of Unit C are consistent with terrestrial exposure surfaces [7, 14]. The facies of Unit D are consistent with lacustrine environments [e.g 6, 9, 11, 12].

The sharp contact of Unit E represent a transition into a distinct environment. The dark-toned lens within Unit E with the asymmetric central body, wing-shaped features, and cm-scale amalgamated bedsets of tabular cross-stratification are consistent with fluvial erosion and infill [15, 16]. Fluvial scouring indicates Unit E must have occurred in a subaerial environment.

Units F through K represent a return to a sand-sheet environment. Even though both aqueous and aeolian wind-ripples climb [12, 17, 18], the wide-spacing and small amplitudes of the corrugations in Unit F are consistent with wind-ripples [e.g. 17] potentially indicating a comparatively higher sedimentation rate and lower wind energy [e.g. 18]. The clear pinstripe pattern seen in Units G, I, and K caused by changes in tone [e.g 19] and cementation [e.g 20] are distinct to wind-ripple migration in a sand-sheet environment (Fig. 2). The darktoned bands in units H and J may represent preferential cementation or diagenesis from a change in grain-size, pore size, or pore space.

The geometry of the scours with broad cross-

sectional concave profiles, flat central portions, and similar facies outside of and within the scour-and-fill structures suggest that the scours likely formed by aeolian processes. We interpret that the scours likely represent saucer-shaped aeolian 'blowout' structures [e.g 18] formed by deflation from intense winds.

Overall, the sedimentary structures found in the layered sulphate-bearing strata in marker band valley represent a predominantly 'dry' aeolian environment with intermittent high winds causing blow-out structures and a thick succession of sand-sheet strata. The sand-sheets show that there was likely sustained, near-surface water during the accumulation of the layered sulfates.

References: [1] Fraeman et al. (2016) JGR: Planets. [2] Bibring et al. (2006) Science, [3] Grotzinger et al. (2014), [4] Fedo et al. (2022), [5] Edgar et al. (2024) LPSC, [6] Rapin et al. (2021) Geology. [7] Gupta et al., (2024) AGU, [8] Farrand et al (2024) Mars 10., [9] Weitz et al. (2022) JGR : Planets, [10] Weitz et al. (2023) LPSC, [11] Dietrich et al. (2023) LPSC, [12] Kocurek (1996), [13] Kocurek & Nielson (1986) Sedimentology, [14] Marriott & Wright (1996) J Geol Soc London, [15] Priddy and Clarke (2020), [16] Gibling, (2006) J. Sediment Res, [17] Hunter (1977) Sedimentology, [18] Fernandes et al (2009), [19] Ferronatto et al. (2019), [20] Fryberger et al. (1988).



Fig. 1: The stratigraphic units as seen in one of the buttes in Marker Band Valley featuring one of the broad, concave-up scours. Image credit: NASA/JPL-Caltech/MSSS.



Fig. 2: A) Pinstripe pattern in laminae taken by the Mars Hand Lens Imager (MAHLI) on sol 3866. Image credit: NASA/JPL-Caltech/MSSS. B) Terrestrial pinstripe laminae from wind ripple [from 19].