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# Analyzing nitrogen dioxide to nitrogen oxide scaling factors for data-driven satellite-based emission estimation methods: A case study of Matimba/Medupi power stations in South Africa

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#### ABSTRACT

In this paper, we propose improved nitrogen dioxide ( $NO_2$ ) to nitrogen oxide ( $NO_x$ ) scaling factors for several data-driven methods that are used for the estimation of  $NO_x$  power plant emissions from satellite observations of  $NO_2$ . The scaling factors are deduced from high-resolution simulations of power plant plumes with the MicroHH large-eddy simulation model with a simplified chemistry and then applied to Sentinel-5 Precursor (S5P) TROPOSpheric Monitoring Instrument (TROPOMI)  $NO_2$  satellite observations over the Matimba/Medupi power stations in South Africa. We show that due to the non-linear chemistry the optimal  $NO_2$  to  $NO_x$  scaling factors depend on both the method employed and the specific segments of the plume from which emission estimate is derived. The scaling factors derived from the MicroHH simulations in this study are substantially (more than 50%) higher than the typical values used in the literature with actual  $NO_2$  observations. The results highlight the challenge in appropriately accounting for the conversion from  $NO_2$  to  $NO_x$  when estimating point source emissions from satellite  $NO_2$  observations.

# 1. Introduction

Satellite observations of nitrogen dioxide ( $NO_2$ ) in the troposphere have been available since the launch of Global Ozone Monitoring Experiment (GOME) in 1995. More recent satellite instruments, such as the Ozone Monitoring Instrument (OMI, Levelt et al., 2018), have been providing observations at much higher resolution allowing to better characterize signals from local anthropogenic sources (e.g., cities and power stations). Several different methods have been proposed to infer emissions of nitrogen oxide ( $NO_x = NO_2 + NO$ ) from satellite  $NO_2$  observations (e.g., de Foy et al., 2015; Beirle et al., 2011, 2019). As  $NO_x$  is often co-emitted with carbon dioxide ( $CO_2$ ), space-based nitrogen dioxide ( $NO_2$ ) observations have been used as indicators for fossil fuel combustion and  $CO_2$  emissions at local scale (Reuter et al., 2014, 2019; Hakkarainen et al., 2021, 2023a). Furthermore, these observations have

been used to study the effects of environmental policies and economic changes (e.g., Castellanos and Boersma, 2012; de Foy et al., 2016).

Nitrogen oxides are mostly (e.g., 95%) emitted in the form of nitrogen monoxide (NO), which is not observed by satellites. A typical assumption is that NO is quickly oxidized to  $NO_2$  by reaction with ozone and the steady state  $NO:NO_2$  ratio is reached within a few minutes (Seinfeld and Pandis, 2006). Often a constant  $NO_x:NO_2$  ratio is assumed to infer the  $NO_x$  emissions from space-based  $NO_2$  observations. Many studies (e.g., Beirle et al., 2011, 2019; de Foy and Schauer, 2022; Merlaud et al., 2020; Shaiganfar et al., 2017; Ionov et al., 2022; Potts et al., 2022; Hakkarainen et al., 2021) use the steady-state noontime molar concentration ratio under typical urban conditions of 1.32 based on Seinfeld and Pandis (2006). Recently, model-based concentration ratios have also been calculated using simulations from Copernicus Atmospheric Monitoring Service (CAMS, Lorente et al.,

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2019; Rey-Pommier et al., 2022) and Comprehensive Air Quality Model with Extensions (CAMx, Goldberg et al., 2022). Beirle et al. (2021) calculated  $\mathrm{NO}_x$ : $\mathrm{NO}_2$  ratios according to the photo-stationary steady state. In general, these studies show small deviations from the value 1.32 (e.g., 1.16–1.83), but acknowledge that values near the point sources are likely to be higher. CHIMERE model simulations (Shaiganfar et al., 2017) further indicate that in large circles around Paris, the partitioning ratios are smaller during summer (1.32) than in winter (1.51), due to the higher ozone mixing ratios in summer. In contrast to model-based analyses, the Dutch aircraft measurements of in-plume  $\mathrm{NO}_x/\mathrm{NO}_2$  ratios from power stations (e.g., Janssen, 1988; Vilà-Guerau de Arellano et al., 1990; Bange et al., 1991; Hanrahan, 1999) often showed values higher than 10 near the source and values between 2 and 10 up to 15 km from the source.

Two classes of data-driven approaches have been used to infer local NO<sub>x</sub> emissions from satellites. The first approach is based on averaging NO<sub>2</sub> satellite data over long periods of time and the second is based on inferring emissions from instantaneous NO2 observations. In both cases the emission estimates are related to the local satellite overpass time. Both methods do not integrate full chemistry models but use simple assumptions to account for the NO2 lifetime and the conversion from NO<sub>2</sub>-derived emissions to NO<sub>x</sub> emissions. With the previous generation of NO2 observing satellites (e.g., OMI), averaging (via oversampling) was often necessary to improve the precision and the spatial resolution (Beirle et al., 2011). Emission estimation methods based on the analysis of individual emission plumes have been mostly applied to long-lived greenhouse gases such as CO2 and methane (CH4), which do not require adjustment due to the lifetime (Varon et al., 2019). Since the launch of the TROPOspheric Monitoring Instrument (TROPOMI) on board the Copernicus Sentinel-5 Precursor (S5P) satellite in 2017, it has been possible to derive NO<sub>x</sub> emissions from anthropogenic sources based on NO2 observations from individual orbits (e.g., Lorente et al., 2019; Rey-Pommier et al., 2022; Zhang et al., 2023), thanks to the improved spatial resolution (currently 5.5 km by 3.5 km at nadir for  $NO_2$ ).

In this paper, we analyze NO and NO $_2$  plumes originating from the Matimba/Medupi power stations in South Africa based on high resolution NO $_x$  simulations and real NO $_2$  satellite data. We apply four NO $_x$  emission estimation methods based on instantaneous NO $_2$  scenes and two methods that infer NO $_x$  emissions from averaged NO $_2$  data. We use MicroHH (van Heerwaarden et al., 2017) large-eddy simulations (LES) with condensed and efficient plume chemistry to study the optimal NO $_2$  to NO $_x$  scaling factor f for each emission estimation method. Furthermore, we use S5P/TROPOMI NO $_2$  observations to infer the NO $_x$  emissions and analyze the sensitivity of the results to the assumptions made regarding the NO $_2$  to NO $_x$  scaling factor. We compare the emission estimates to the reported information.

# 2. Materials and methods

# 2.1. Matimba/Medupi power stations

We analyze the NO<sub>x</sub> emission estimates from the Matimba and Medupi coal-fired power stations previously studied by Hakkarainen et al. (2021). These two power stations operated by Eskom are located about 6 km apart in the Highveld region in South Africa (23.67°S, 27.61°E) and are isolated from other strong emission sources in the region. Eskom reports total monthly NO<sub>x</sub> emissions (using standards BS EN 14181:2004 - Quality Assurance of Automated Measuring Systems and ESKOM internal standard 240-56242363 Emissions Monitoring and Reporting) as well as daily energy production (see https://www.eskom. co.za/dataportal/emissions/ael/ for monthly reports and more details). We disaggregate the reported monthly emissions to daily emissions based on the daily energy consumption (as a fraction of total monthly energy consumption). The daily and monthly emissions for the year 2021 are given in Fig. 1. The total annual NO<sub>x</sub> emission of the two power stations is 101 ± 12 kton/year. As usual, NO<sub>x</sub> emissions are reported in terms of NO2 equivalent.

#### 2.2. MicroHH model

To calculate method-dependent  $\mathrm{NO}_2$  to  $\mathrm{NO}_x$  scaling factors, we use the MicroHH (van Heerwaarden et al., 2017) large-eddy simulations (LES) generated for the Matimba power station within the H2020 project CoCO2 (see Krol et al., 2024, for more details). The simulation period is 48 hours starting from 24 July 2020 00 UTC. Following the CoCO2 simulation protocol, based on the Community Emissions Data System (CEMS) for the year 2014, the simulation uses constant  $\mathrm{NO}_x$  emission of 1.9676 kg s<sup>-1</sup> (62 kton/year, in terms of  $\mathrm{NO}_2$  equivalent) of which 95% is NO and 5% is  $\mathrm{NO}_2$ .

The MicroHH LES simulations were initialized and driven by ERA5 (meteorology) and CAMS (non-reactive and reactive gases), with a coupling as described by Neggers et al. (2012) and van Stratum et al. (2023). The horizontal resolution was 100 m×100 m and the vertical resolution 25 m. The chemistry scheme is a condensed version of the scheme implemented in the European Centre for Medium-Range Weather Forecasts (ECMWF) Integrated Forecasting System (IFS, Huijnen et al., 2016) and focuses on correct calculations of the equilibrium between NO<sub>x</sub> and ozone (photo-stationary state) and NO<sub>x</sub> lifetime. Photolysis rates are calculated at 500 m above the surface for clear sky conditions with the TUV module (Madronich and Flocke, 1999). The simulated species are: HNO<sub>3</sub>, NO<sub>3</sub>, N<sub>2</sub>O<sub>5</sub>, H<sub>2</sub>O<sub>2</sub>, CO, HCHO, ROOH, C<sub>3</sub>H<sub>6</sub> (generic hydrocarbon), RO<sub>2</sub>, HO<sub>2</sub>, NO, NO<sub>2</sub>, OH (and fixed CH<sub>4</sub> and H<sub>2</sub>). The Rosenbrock solver of the kinetic rate equations was generated automatically by the Kinetic Pre Processor (KPP, Damian et al., 2002; Sandu et al., 2022).

The chemistry used in the MicroHH model (Krol et al., 2024) is simplified, and does not contain PAN and alkyl-nitrates. However, the model has been calibrated against a more complex chemistry implemented in the IFS. This version, which includes PAN and alkyl-nitrates, is used for the CAMS reanalysis (Inness et al., 2019) and serves as chemical boundary conditions for the MicroHH simulations. The simplified chemistry introduces additional uncertainty, but the main chemical processes that determine the  $\mathrm{NO}_x/\mathrm{NO}_2$  ratio and  $\mathrm{NO}_2$  lifetime are adequately modeled. MicroHH is one of the few Large Eddy Simulation models that resolves atmospheric chemistry (Krol et al., 2024). Plume dispersion in MicroHH was tested and evaluated against wind tunnel experiments in Razňjević (2023).

For the practical calculation we use the "Library of Plumes" dataset created in the H2020 project CoCO2 (Koene and Brunner, 2022, 2023). In the dataset, the vertical column densities (VCD) of NO,  $NO_2$  and  $CO_2$  are available with hourly resolution on a 2 km×2 km grid to mimic the upcoming CO2M satellite (Meijer et al., 2020).

# 2.3. S5P/TROPOMI $NO_2$ observations

The TROPOspheric Monitoring Instrument (Veefkind et al., 2012) was launched onboard the Copernicus Sentinel-5P satellite on 13 October 2017. The satellite has an equatorial crossing time of 13:30 LT. The current spatial resolution is 5.5 km by 3.5 km at nadir and covers a swath of about 2600 km wide. Here we use the tropospheric NO $_2$  vertical columns with quality flag qa\_value  $\geq$  0.75 and the effective cloud fraction  $\leq$  0.3. Due to changes in the operational algorithm, we use the version v02.03.01 intermediate reprocessing on the S5P-PAL system https://data-portal.s5p-pal.com/, which provides a seamless connection with the operational version 2.3.1 data product. Technical details can be found from the readme file: https://data-portal.s5p-pal.com/product-docs/no2/PAL\_reprocessing\_NO2\_v02.03.01\_20211215.pdf.

We use the TROPOMI  $NO_2$  observations from the year 2021 to estimate the  $NO_x$  emission from Matimba/Medupi power stations. The S5P/TROPOMI overpass time is around 12 UTC over these stations. Following Beirle et al. (2019), we recalculate the air mass factors using  $NO_2$  profiles representative for the emission plume. Based on the MicroHH simulations, the  $NO_2$  inside the plume is rapidly mixed with background air in the planetary boundary layer with a mean



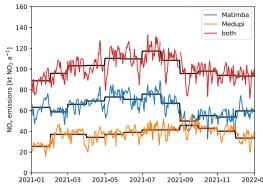


Fig. 1. (left) Matimba and Medupi power stations in South Africa about 6 km apart. (right) Monthly  $NO_x$  emissions converted to daily emissions using power generation reported by Eskom. Black lines indicate monthly emissions. © 2024 Google Earth

mole fraction of about 5 nmol mol $^{-1}$ . We therefore recomputed AMFs by updating NO $_2$  profiles inside the PBL height from ERA-5, which results in a multiplicative correction factor of 1.27 for TROPOMI NO $_2$  columns inside the plume. This value is in good agreement with the AMF correction factor of 1.35 found by Beirle et al. (2019) for South

# 2.4. NO<sub>x</sub> emission estimation methods

We use six emission estimation methods of which four (integrated mass enhancement, Gaussian plume, cross-sectional flux and light cross-sectional flux methods) are based on the analysis of instantaneous observations and two (exponentially modified Gaussian and divergence methods) are based on temporally averaged data. All the methods estimate the emission corresponding to the local TROPOMI overpass time. These computationally light methods were selected based on a benchmark analysis (Santaren et al., 2024) carried out in the H2020 project CoCO2 (Hakkarainen et al., 2023b) and are included in the ddeq Python library for data-driven emission quantification of hotspots. Kuhlmann et al. (2024) gives a comprehensive description of the ddeq implementation. The methods used in the analysis are briefly summarized below:

• In the Gaussian plume (GP) model method, the emission *E* is derived by fitting a 2D Gaussian plume function to the observations (e.g. Bovensmann et al., 2010; Nassar et al., 2017):

$$G(x,y) = \frac{E}{\sqrt{2\pi}\sigma(x)U} e^{-\frac{y^2}{2\sigma^2(x)}}$$
 (1)

where U is the effective wind speed and  $\sigma = \sqrt{2Kx^b/U}$  is the standard width depending on the eddy diffusion coefficient K and an additional exponential parameter b. With short-lived gases, an exponential decay that multiplies G(x,y) is also fitted:

$$D(x,\tau) = \exp\left(-\frac{x}{U\tau}\right) \tag{2}$$

where  $\tau$  is the exponential decay time. Furthermore, our implementation allows the Gaussian plume to be curved.

• For the cross-sectional flux methods, we use two different implementations: the cross-sectional flux (CSF) method described by Kuhlmann et al. (2020, 2021) and the light cross-sectional flux (LCSF) originally developed by Zheng et al. (2020) and Chevallier et al. (2022) for OCO-2 and adapted for TROPOMI in the CoCO2 project. In both versions, the emission is derived from the integrals of the cross-section of the plume perpendicular to the wind direction (i.e. line density) multiplied with the effective wind speed U. Gaussian curves are fitted for calculating the line densities (e.g., Reuter et al., 2019) to better constrain

noisy observations and to interpolate (possibly) missing data. The LCSF implementation uses information closer to the source (one hour downwind), whereas the CSF method analyzes the full detectable plume using sub-polygons every 5 km downwind of the source. The LCSF method assumes a NO $_{x}$  and NO $_{2}$  lifetime of 4 hours, while the CSF method computes the NO $_{x}$  decay time as an additional fit parameter.

• The integrated mass enhancement (IME) method calculates the emission from the total plume mass above the background. Following the approach by Frankenberg et al. (2016), for inert gases like  $\mathrm{CO}_2$  and  $\mathrm{CH}_4$ , the emission can be calculated as  $E=U\times\mathrm{IME}/L$ , where U is the wind speed, IME is the integrated mass enhancement (above the background) and L is the distance from the source up to where the plume is integrated. For  $\mathrm{NO}_x$ , a correction factor c can be derived from a Gaussian plume model multiplied with an exponential decay:

$$E = \frac{1}{c} \frac{U}{L} \text{IME} \quad \text{with} \quad c = \frac{U \tau}{L} \left( 1 - \exp\left(-\frac{L}{U\tau}\right) \right). \tag{3}$$

Here,  $\tau$  is the NO<sub>x</sub> lifetime, which is assumed to be 4 hours (see Kuhlmann et al., 2024, for details).

• In the exponentially modified Gaussian (EMG) method, the emission is calculated from the averaged data. Here, the data is rotated according to the wind direction so that all scenes have wind direction from west to east. The resulting rotated mean field is then integrated along the latitudinal dimension to derive the  $NO_2$  line densities. The emission E is obtained by fitting the averaged line densities with the EMG model M(x) (as in Beirle et al., 2011; de Foy et al., 2014) as follows:

$$M(x) = \frac{E}{U} \times (e * G)(x) + B. \tag{4}$$

Here (e \* G)(x) is the convolution between the exponential decay and the Gaussian function, and B is the background.

• The divergence (DIV) method is based on calculating the divergence D of the averaged fluxes that are derived from the vertical column densities and the wind components. The emission field E is obtained from the divergence D according to the continuity equation for steady state, D=E-S, where S indicates the sink term. The emissions from point sources are obtained from the emission field by using the Gaussian peak fitting approach (Beirle et al., 2019, 2021). The version used here is described in Hakkarainen et al. (2022). First, the fluxes are calculated using the tropospheric  $NO_2$  columns and effective wind speed. The divergence D is calculated from the averaged fluxes. The sinks are calculated as  $S=V/\tau$ , where V is the mean vertical column density and  $\tau$  the assumed lifetime. The emissions fields are calculated as E=D+S and fitted with the Gaussian peak fitting function.

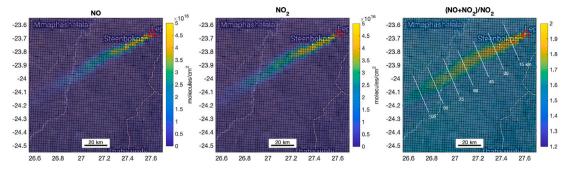


Fig. 2. MicroHH simulations of NO and NO<sub>2</sub> columns and the NO<sub>x</sub>:NO<sub>2</sub> molar concentration ratio. The maximum value is 8.7 in the grid cell closest to the Matimba power station marked with red triangle. The background value is about 1.5. White lines perpendicular to the plume indicate transects at different distance from the source (red triangle). The spatial resolution of the data is  $2 \text{ km} \times 2 \text{ km}$ .

The methods based on instantaneous images also require plume detection, background extraction and center line fitting. Here we use the plume detection described by Kuhlmann et al. (2021), which has the following four steps: (1) pixels significantly enhanced above the background are detected based on a statistical test; (2) detected pixels are grouped into regions (plumes); (3) a region of connected enhanced pixels is assigned to a point source; and (4) a centerline is fitted to the detected pixels for each plume as a two-dimensional curve. The algorithm is used by the GP, CSF and IME method. However, the LCSF methods identifies the centers and width of plume transects fitting a Gaussian function to the  $\mathrm{NO}_2$  enhancements in the across-wind direction, close and downwind to the source, using the wind direction provided by the model.

For the GP, CSF and IME method, the background is computed by masking the detected plume and applying a normalized convolution, while the LCSF method computes the background by fitting a linear term in addition to the Gaussian function to the  $\rm NO_2$  concentrations in the across-wind direction.

The  $\mathrm{NO}_x$  emission estimates are calculated from satellite-based  $\mathrm{NO}_2$  observations. We use a scaling factor f, to convert  $\mathrm{NO}_2$ -derived emissions to  $\mathrm{NO}_x$  emissions, i.e.,  $E_{\mathrm{NO}_x} = f \times E_{\mathrm{NO}_2}$ . We note that the widely used scaling factor 1.32 is a molar ratio and therefore  $E_{\mathrm{NO}_x}$  corresponds to the emission reported as  $\mathrm{NO}_2$  equivalents. Here we derive the method-specific scaling factors as  $f = E_{\mathrm{NO}_x}/E_{\mathrm{NO}_2}$ , where  $E_{\mathrm{NO}_x}$  and  $E_{\mathrm{NO}_2}$  are the emissions calculated from the  $\mathrm{NO}_x$  and  $\mathrm{NO}_2$  MicroHH LES simulations, respectively.

# 3. Results

# 3.1. Analysis of MicroHH simulations

Fig. 2 shows an example of NO and  $\mathrm{NO}_2$  vertical column densities from MicroHH large-eddy simulations (25 July 2020 12:00 UTC). As expected, the  $\mathrm{NO}_x$ :NO $_2$  molar concentration ratio is higher near the emission source, where values higher than 2 are observed until about 20 km downwind from the source. The maximum value is 8.7 in the grid cell closest to the source. In this case, the background ratio is about 1.5, which is in the typical range (e.g., Fig. 2 in Beirle et al., 2023), but slightly higher than 1.32. Due to chemistry, the molar concentration ratios are exponentially decreasing along the plume.

We use the MicroHH simulations to derive the emissions from the NO $_x$  VCDs (NO + NO $_2$ ) and from the NO $_2$  VCDs only. The latter option is the only one available with real satellite observations. The NO $_2$  to NO $_x$  scaling factors for each method are inferred from the mean emissions as  $f = E_{\rm NO} / E_{\rm NO}$ .

Fig. 3 shows the time series of the estimated  $NO_x$  emissions using the LCSF, CSF, GP, and IME methods. Overall, across all the methods, the mean  $NO_x$  emissions (solid lines in Fig. 3) are in good agreement with the emission value used in the simulation (dashed lines in Fig. 3). Here, and throughout the rest of the manuscript we express

NO<sub>x</sub> emission as NO<sub>2</sub> equivalents. In the case of chemically inactive species like CO2, the emissions from each method are also very close to the emission value used in the simulation (see supplementary file, Figs. S1–S4). The NO<sub>2</sub>-derived emissions are multiplied with a scaling factor f computed as the ratio of the mean NO<sub>x</sub> and NO<sub>2</sub> emission estimates. The uncertainty was estimated from the standard deviation of the ratios. The estimated mean scaling factors  $f = E_{NO_x}/E_{NO_y}$  are  $1.71 \pm 0.29$ ,  $2.73 \pm 0.64$ ,  $2.17 \pm 0.48$  and  $1.83 \pm 0.58$  for the IME, GP, CSF and LCSF methods, respectively. These values are significantly higher than the values typically used in the literature (e.g., 1.32). The scaling factors f for each method are dependent on the part of the plume used for emission calculation, as discussed below. The daytime (9–15 UTC) scaling factors  $f_d$  are about 3%–10% larger (except for GP method 13% smaller), with the mean values of 1.83, 2.38, 2.24 and 2.02 for the corresponding methods. The daytime standard deviation is about 0.2 to 0.3 for all the methods.

The optimal scaling factor depends on the plume length used in the calculation as shown for the IME approach in Fig. 4. The factor is about 3.7 for a 15 km long plume and drops rapidly reaching values smaller than 1.7 for 50 km long plumes. Note that the scaling factor for the IME method is independent of the assumed decay time that impacts  $\mathrm{NO}_2$  and  $\mathrm{NO}_x$  in the same way.

The scaling factor can be used to understand where a method obtains information about the emissions. The IME method uses all measurements up to 50 km downwind of the source. Likewise, the LCSF method uses only measurements one hour downwind of the source, which ranges between 13 and 42 km depending on wind speed. Consequently, the LCSF method has a larger scaling factor than the IME method. The GP method has the highest scaling factor of 2.73 implying that it mostly uses VCDs at distances below 30 km where the signal-to-noise ratio is largest. In contrast, the CSF method computes fluxes at several distances downwind from which the emission rate and decay time are estimated, which apparently gives a higher weight to downwind values.

The decay time for the  $\mathrm{NO}_x$  columns was estimated as an additional fit parameter in the Gaussian plume and cross-sectional flux inversion methods (Figs. S2 and S3). The corresponding median values of 2.8 and 6.6 hours are consistent with the typical value of 4 hours used in previous studies and assumed here for the IME and LCSF methods. For the  $\mathrm{NO}_2$  columns, the estimated decay times are higher (9.0 h and 13.5 for GP and CSF method) than for  $\mathrm{NO}_x$ . This is likely caused by the increase of  $\mathrm{NO}_2$  columns downwind from the source due to the conversion of  $\mathrm{NO}$  to  $\mathrm{NO}_2$ . We note that lifetimes reported here are effective decay times, and in general, the lifetime as mass/loss is not constant along the plume.

The analyzed MicroHH simulation is 48 hours long. For the methods that use averaged data for emission estimation (EMG and divergence method), calculating the emissions from a 2 day simulation is challenging. Fig. 5 shows the line densities of the rotated data and the EMG fits for both  $NO_x$  and  $NO_2$  emission estimation. By comparing

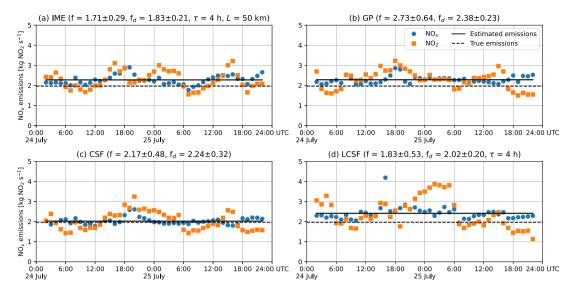


Fig. 3. Time series of estimated  $NO_x$  emissions from MicroHH simulations using (a) IME, (b) GP, (c) CSF and (d) LCSF method. Emissions were estimated from  $NO_2$  observations. The  $NO_2$ -derived emissions were multiplied with a method-specific scaling factor f. In addition, the scaling factor  $f_d$  was computed using only daytime data (9–15 UTC). The uncertainty was estimated from the standard deviation of the ratios. IME and LCSF assume a fixed decay time of four hours.

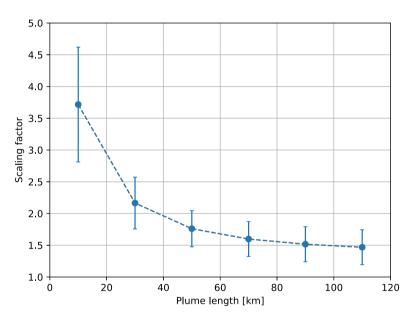


Fig. 4. Dependency of the scaling factor on plume length estimated from the IME method.

the emission estimates derived from the respective EMG fits, we obtain  $f=2.14\pm0.10$ . We note that the  $\mathrm{NO}_2$  peak is located further away from the source (as compared to  $\mathrm{NO}_x$ ) due to the conversion from NO to  $\mathrm{NO}_2$ . The photo-stationary state equilibrium under these conditions is reached only after about 1 h. The lifetime estimated by the EMG fit is 1.5 times higher for  $\mathrm{NO}_2$  than for  $\mathrm{NO}_x$  (3.1 h vs 2.1 h). Moreover, the lifetimes are not constant along the plume. As discussed in more detail in Krol et al. (2024), the plume lifetime is determined both by mixing (entrainment of background ozone into the plume) and the strength of the  $\mathrm{NO}_x$  emissions. When emissions are large, as in the Matimba case, the plume stays chemically intact longer, because the emitted NO (95% of  $\mathrm{NO}_x$ ) efficiently titrates the entrained ozone.

Fig. 5 (inset) also shows the molar ratio of  $NO_x$  and  $NO_2$  line densities. Ratios of about 3.2 are found near the source and are rapidly decreasing to about 1.5 after 20 km. This also indicates that the methods that use information closer to the source need a higher scaling factor. We were not able to calculate the  $NO_2$  to  $NO_x$  scaling factor for the divergence method, but the peak fitting approach included in

this method primarily uses information close to the source as well. In this case, we would need a longer simulation and data across all wind directions to successfully apply the peak fitting approach.

Finally, we note that all the methods (GP, CSF and EMG) that estimate the exponential decay time  $\tau$  from the observations give higher scaling factors than the methods (IME and LCSF) that assume fixed NO $_x$  and NO $_2$  decay time of four hours (2.38, 2.24 and 2.14 vs. 1.83 and 2.02). Part of the reason is that estimated decay times for NO $_x$  are significantly lower (30%–70%) than the decay times obtained from the NO $_2$  observations, because the estimated NO $_2$  lifetime compensates for the conversion of NO to NO $_2$ , which increases the apparent lifetime of NO $_2$ . In general, lower estimated lifetimes correspond to higher emissions, which results in higher scaling factors as  $f = E_{NO_x}/E_{NO_2}$ . In particular, here the LCSF method gives a lower scaling factor than the CSF method, which would be unexpected (as LCSF uses information closer to the emission source) without considering the fact that with the CSF method the estimated decay times for NO $_x$  are about 50% lower than for NO $_2$ .

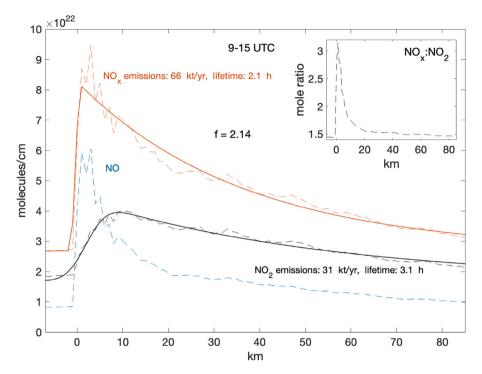


Fig. 5. MicroHH  $NO_x$  (red), NO (blue) and  $NO_2$  (black) line densities (dashed lines) as a function of distance from the source. The solid lines are fits from the EMG model. The inset shows the molar ratio of  $NO_x$  and  $NO_y$  line densities.

#### 3.2. Results with S5P/TROPOMI data

Fig. 6 illustrates an example of an individual  $NO_2$  plume emitted by the Matimba/Medupi power plants as observed by S5P/TROPOMI on 25 July 2021. As the distance between the power stations is only 6 km, they appear as a single source. The  $NO_x$  emission estimates based on the IME, GP, CSF and LCSF methods range between 78 kton/year and 99 kton/year, which is lower than the reported emission values of 130 kton/year. These estimates are derived by already applying the larger method-dependent  $NO_2$  to  $NO_x$  scaling factors f obtained from the MicroHH analysis and the AMF-related correction factor of 1.27. For all methods, we use the vertically averaged ERA5 winds based on a typical emission profile used in simulations as presented by Brunner et al. (2019). For the CSF and GP methods, the lifetimes are fitted from the observations. For the IME and LCSF methods, which do not provide an estimate of the lifetime, we assume a constant  $NO_x$  lifetime of 4 h.

Fig. 7 shows the  $NO_x$  emissions estimation time series based on the four methods compared to the reported Matimba/Medupi  $NO_x$  emissions. All methods, especially IME and CSF, show higher emission estimates during the austral winter months. During the austral summer, we generally obtain values lower than reported. The temporal variability is also quite large compared to the reported  $NO_x$  variability. There are also large differences between the methods in terms of temporal evolution and scatter. One must note that the scaling factors f were derived from 48h-LES simulations during austral winter (July 2020) and might not be representative of the conditions during the rest of the year (e.g., Shaiganfar et al., 2017).

Next, we analyze the Matimba/Medupi  $\mathrm{NO}_x$  emissions based on the methods using temporal averaging. Fig. 8 shows the  $\mathrm{NO}_2$  line densities and the fit derived using the EMG approach. We obtain a lifetime of 3.5 hours and a  $\mathrm{NO}_x$  emission of 80 kton/year assuming a  $\mathrm{NO}_2$  to  $\mathrm{NO}_x$  scaling factor of 1.32. Considering the AMF-related correction factor of 1.27 used here, the emission estimate of 80 kton/year is in agreement with those  $\mathrm{NO}_x$  emissions of about 60 kton/year reported in recent studies (Hakkarainen et al., 2021; Potts et al., 2022) that also used the EMG method, but were applied to different TROPOMI  $\mathrm{NO}_2$  data versions and time periods.

Fig. 9 shows the results using the divergence method. The first panel shows the sinks calculated using the assumed lifetime of four hours, the second panel is the divergence, and the third panel is the emissions, i.e., E=S+D. Like for the EMG method, we use a  $\mathrm{NO}_2$  to  $\mathrm{NO}_x$  scaling factor of 1.32. The emissions from the Matimba/Medupi power stations are clearly visible and easy to fit. When a peak function is fitted to the average emission map (third panel), we obtain the  $\mathrm{NO}_x$  emissions of 52 kton/year. This is comparable to previous studies: Beirle et al. (2021) obtained emissions of 21 kton/year, whereas Beirle et al. (2019) obtained emissions of 54 kton/year (after applying correction for different factors such as AMF, lifetime, and winds, see Beirle et al. (2021) for more details). Beirle et al. (2023) obtained  $\mathrm{NO}_x$  emissions of 76 kton/year for the year 2021 with a significantly updated method.

Fig. 10 summarizes the annual emission estimates using the different methods. The emissions are presented using the scaling factor 1.32 (dark colors) as well as with the method-dependent scaling factors (light colors). With the scaling factor 1.32, all the methods systematically underestimate the reported Matimba/Medupi annual NO<sub>x</sub> emissions of about 100 kton/year. With the method-based scaling factors, the annual emissions are substantially closer to the reported emissions. Assuming that the reported values are correct and with the right temporal representation, these results indicate that methodspecific scaling factors are needed to better calculate the true NO<sub>x</sub> emissions for strong point sources like the Matimba/Medupi power station. As the MicroHH simulation was only 48 hours long, we could not calculate a specific scaling factor for the divergence method. To obtain NO<sub>x</sub> emission values with the divergence method compatible with the reported emissions, we would need a NO<sub>2</sub> to NO<sub>x</sub> scaling factor of about 2.5. This value is plausible as the emissions are inferred via peak fitting above the source. In this case, the standard deviations of the 2D Gaussian peak function were 7 km and 5 km in x and y directions, respectively.

# 4. Discussion and conclusions

In this paper, we analyzed optimal  $NO_2$  to  $NO_x$  scaling factors for several data-driven satellite-based emission estimation methods. The

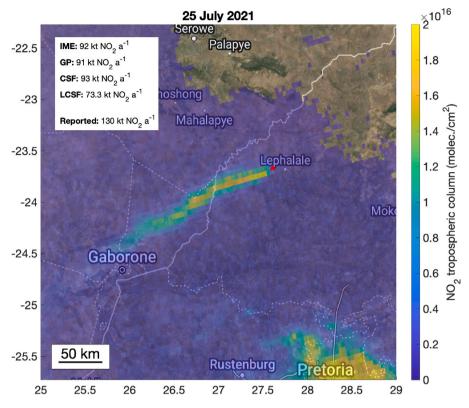


Fig. 6. Sentinel-5P/TROPOMI NO<sub>2</sub> plume from Matimba/Medupi (red triangle) observed on 25 July 2021 (on © Google Maps 2023 background). The NO<sub>x</sub> emission estimates obtained with the IME, GP, CSF, and LCSF methods as well as the reported emissions are shown in the top left.

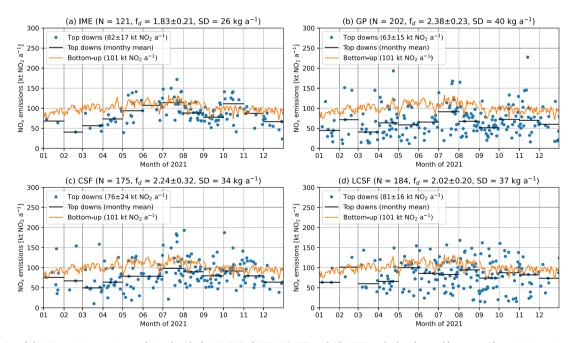


Fig. 7. Time series of the  $NO_x$  emission estimates obtained with the (a) IME, (b) GP, (c) CSF, and (d) LCSF method with monthly means. The emission estimates are compared against Eskom-reported emissions. The title shows the number of estimates N, the used scaling factor  $f_d$  and the scatter between top-down and bottom-up emission estimates.

results based on the MicroHH simulations show that the optimal scaling factors depend on the selected method because the different methods exploit information from different down-wind distances of the plume. In general, the scaling factors are higher if the information comes from near the point source and smaller if larger spatial areas are considered. This is because the  $NO_x/NO_2$  concentration ratios vary strongly within the plume, with the highest values at the source pixel, while the values downwind are closer to literature values. Depending on the method,

the scaling factors vary between 1.8 and 2.4, which is substantially higher than the values typically used in the literature in the case of Matimba/Medupi power plants. Aircraft measurements of plume  $\mathrm{NO}_x/\mathrm{NO}_2$  ratios (e.g., Janssen, 1988; Vilà-Guerau de Arellano et al., 1990; Bange et al., 1991; Hanrahan, 1999) also indicate much higher values near the source. These higher scaling factors could partly explain why some studies (e.g., Beirle et al., 2021) have reported too low  $\mathrm{NO}_x$  emission estimates. Also, the three-dimensional radiative transfer effects, reported

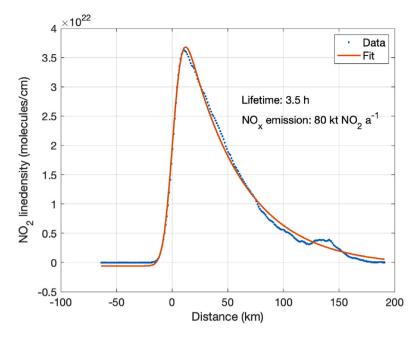


Fig. 8. Illustration of the exponentially modified Gaussian (EMG) method applied to the Matimba/Medupi power stations for 2021 based on Sentinel-5P/TROPOMI  $NO_2$  observations using f = 1.32.

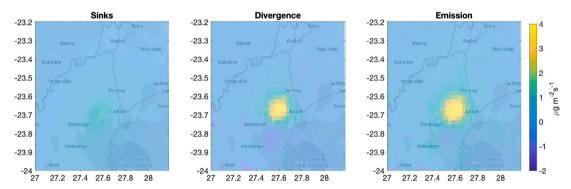


Fig. 9. Illustration of the divergence method for Matimba/Medupi power stations in 2021 based on Sentinel-5P/TROPOMI NO2 observations.

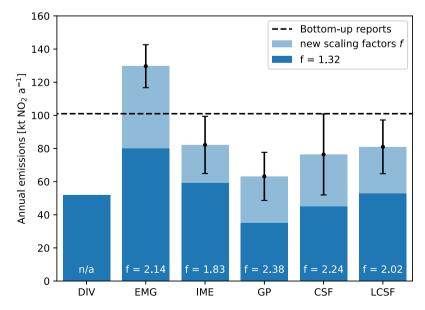


Fig. 10. Summary of the annual NO<sub>x</sub> emission estimates for Matimba and Medupi power stations using method-dependent scaling factors ( $f_d$ , light colors) and f = 1.32 (dark colors). The error bars show variability ( $1\sigma$ ) of the scaling factor. Other sources of uncertainty are not included in the error bar.

by Schwaerzel et al. (2020) and Wagner et al. (2023), can lead to underestimation of the true trace gas content. Here, we further verify that emission estimates obtained with method-specific scaling factors derived from a small set of MicroHH simulations are closer to the reported emissions in the Matimba/Medupi power station case study. The optimal scaling factors are likely not constant for a given method but additionally depend on the strength of the emission source, the meteorological conditions, background ozone concentration levels, and NO, lifetime. In general, the scaling factors are expected to be smaller during summer than in winter, due to the higher ozone mixing ratios in summer (e.g., Shaiganfar et al., 2017). We note that if the emission values used for the MicroHH simulations are substantially overestimated or underestimated the scaling factors could be overestimated or underestimated, respectively, due to the changes in the NO-NO<sub>2</sub>-O<sub>3</sub> chemistry and hence the partitioning between NO and NO2. When emissions are large the plume stays chemically intact longer (Krol et al., 2024). The method-dependent scaling factors obtained here should thus not be interpreted as being universally applicable for the respective

We used a specific 48 hour MicroHH large-eddy simulation to calculate optimal  $\mathrm{NO}_2$  to  $\mathrm{NO}_x$  scaling factors for the Matimba/Medupi power station. In general, calculating optimal scaling factors from large-eddy simulations can only be a partial solution when data-driven and computationally light emission inversion methods are used. High resolution model simulations with  $\mathrm{NO}_x$  chemistry for each source and time period can be computationally expensive as the MicroHH simulations here. However, more comprehensive simulations with systematic variation of all influencing factors could potentially lead to a formulation of method-specific scaling factors as a function of these factors. The methods that are based on data averaging, the divergence method in particular, would optimally need one year of simulations for each source.

The method-specific  $\mathrm{NO}_2$  to  $\mathrm{NO}_x$  scaling factors are needed as only  $\mathrm{NO}_2$  is observed from satellites. Another option could be to convert  $\mathrm{NO}_2$  columns to  $\mathrm{NO}_x$  columns using spatially explicit maps of molar concentration ratio and then use these fields to infer emissions without scaling factors. The explicit concentrations maps could be generated for small sets of plumes with high-resolution modeling, and in larger studies parametric solutions could be used. These options are left for future studies. All in all, the results presented here highlight the need to appropriately account for the conversion from  $\mathrm{NO}_2$  to  $\mathrm{NO}_2$ .

# CRediT authorship contribution statement

Janne Hakkarainen: Conceptualization, Formal analysis, Investigation, Methodology, Writing - original draft, Writing - review & editing. Gerrit Kuhlmann: Conceptualization, Formal analysis, Investigation, Methodology, Writing - original draft, Writing - review & editing. Erik Koene: Data curation, Formal analysis, Investigation, Methodology, Writing - review & editing. Diego Santaren: Formal analysis, Investigation, Methodology, Writing – review & editing. Sandro Meier: Formal analysis, Writing - review & editing. Maarten C. Krol: Data curation, Formal analysis, Investigation, Methodology, Resources, Writing - original draft, Writing - review & editing. Bart J.H. van Stratum: Data curation, Resources, Writing - review & editing. Iolanda Ialongo: Investigation, Writing - original draft, Writing - review & editing. Frédéric Chevallier: Methodology, Writing – review & editing. Johanna Tamminen: Supervision, Writing - review & editing. Dominik Brunner: Supervision, Writing – review & editing, Investigation. Grégoire Broquet: Supervision, Writing - review & editing.

# Declaration of competing interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

#### Data availability

MicroHH columns created in EU H2020 project CoCO2 are available online at https://doi.org/10.5281/zenodo.7448143 (Koene and Brunner, 2023). The Copernicus Sentinel-5P/TROPOMI data used in this study can be accessed from the S5P-PAL system https://data-portal.s5p-pal.com/.

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# Appendix A. Supplementary data

Supplementary material related to this article can be found online at https://doi.org/10.1016/j.apr.2024.102171.

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