# Supplementary material: Interaction quenches in nonzero temperature fermionic condensates 

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## LINEARISATION OF THE HFB EQUATIONS

We give here additional details on our analytical solution of the BCS equations (1)-(2) in the regime of weak perturbations, leading to Eq. (6). We linearize the evolution around the initial equilibrium state, rather than around a virtual final equilibrium state as in Ref. [7]. At nonzero temperature, where $\Delta(t)$ never reaches $\Delta_{f}$, this is far more intuitive.

We introduce the fluctuations $\delta c_{\mathbf{k}}=c_{\mathbf{k}}-c_{\mathbf{k}, i}, \delta n_{\mathbf{k}}=$ $n_{\mathbf{k}}-n_{\mathbf{k}, i}$, and $\delta \Delta=\Delta-\Delta_{i}$ and linearize the BCS system (1)-(2):

$$
\begin{align*}
\mathrm{i} \partial_{t} \delta c_{\mathbf{k}} & =2 \xi_{\mathbf{k}} \delta c_{\mathbf{k}}-2 \Delta_{i} \delta n_{\mathbf{k}}+\frac{\xi_{\mathbf{k}}}{\epsilon_{\mathbf{k}}} F\left(\epsilon_{\mathbf{k}}\right) \delta \Delta  \tag{S1}\\
\mathrm{i} \partial_{t} \delta n_{\mathbf{k}} & =-\Delta_{i}\left(\delta c_{\mathbf{k}}-\delta c_{\mathbf{k}}^{*}\right)-\frac{\Delta_{i}}{2 \epsilon_{\mathbf{k}}} F\left(\epsilon_{\mathbf{k}}\right)\left(\delta \Delta-\delta \Delta^{*}\right) \tag{S2}
\end{align*}
$$

Although the quench scenario corresponds to $\delta n_{\mathbf{k}}=$ $\delta c_{\mathbf{k}}=0$ at $t=0^{-}$, we make so far no assumption on the initial state. In this more general case, the fluctuation of $\Delta$ has a time-dependent part caused by the $\delta c^{\prime}$ s, and a constant part $\delta \Delta_{0}=\frac{g_{f}-g_{i}}{V} \sum_{\mathbf{k}} c_{\mathbf{k}, i}$ caused by the quench on $g$ :

$$
\begin{equation*}
\delta \Delta(t)=\frac{g_{i}}{V} \sum_{\mathbf{k}} \delta c_{\mathbf{k}}(t)+\delta \Delta_{0} \tag{S3}
\end{equation*}
$$

In the spirit of Ref. [? ], we now move to the quasiparticle basis:

$$
\begin{align*}
& \alpha_{\mathbf{k}}^{+}=\frac{\xi_{\mathbf{k}}}{\epsilon_{\mathbf{k}}}\left(\delta c_{\mathbf{k}}+\delta c_{\mathbf{k}}^{*}\right)-\frac{2 \Delta_{i}}{\epsilon_{\mathbf{k}}} \delta n_{\mathbf{k}} \\
& \alpha_{\mathbf{k}}^{-}=\delta c_{\mathbf{k}}-\delta c_{\mathbf{k}}^{*}  \tag{S4}\\
& m_{\mathbf{k}}=\frac{2 \xi_{\mathbf{k}}}{\epsilon_{\mathbf{k}}} \delta n_{\mathbf{k}}+\frac{\Delta_{i}}{\epsilon_{\mathbf{k}}}\left(\delta c_{\mathbf{k}}+\delta c_{\mathbf{k}}^{*}\right) \tag{S5}
\end{align*}
$$

so as to diagonalise the individual parts of Eqs. (??)(??):

$$
\begin{align*}
& \mathrm{i} \partial_{t} \alpha_{\mathbf{k}}^{+}=2 \epsilon_{\mathbf{k}} \alpha_{\mathbf{k}}^{-}+F\left(\epsilon_{\mathbf{k}}\right)\left(\delta \Delta-\delta \Delta^{*}\right)  \tag{S6}\\
& \mathrm{i} \partial_{t} \alpha_{\mathbf{k}}^{-}=2 \epsilon_{\mathbf{k}} \alpha_{\mathbf{k}}^{+}+\frac{\xi_{\mathbf{k}}}{\epsilon_{\mathbf{k}}} F\left(\epsilon_{\mathbf{k}}\right)\left(\delta \Delta+\delta \Delta^{*}\right)  \tag{S7}\\
& \mathrm{i} \partial_{t} m_{\mathbf{k}}=0 \tag{S8}
\end{align*}
$$

In the quasiparticle basis, the fluctuations of $\Delta$ take the form:

$$
\begin{equation*}
\delta \Delta=\delta \Delta_{0}+\frac{g_{i}}{2} \int \frac{\mathrm{~d}^{3} k}{(2 \pi)^{3}}\left[\alpha_{\mathbf{k}}^{-}+\frac{\xi_{\mathbf{k}}}{\epsilon_{\mathbf{k}}} \alpha_{\mathbf{k}}^{+}+\frac{\Delta_{i}}{\epsilon_{\mathbf{k}}} m_{\mathbf{k}}\right] \tag{S9}
\end{equation*}
$$

To solve the time-dependent system, we introduce the Laplace transform of the variables:

$$
\begin{equation*}
A_{\mathbf{k}}^{ \pm}(\omega)=\int_{0^{-}}^{+\infty} \mathrm{e}^{\mathrm{i} \omega t} \alpha_{\mathbf{k}}^{ \pm}(t) \mathrm{d} t \tag{S10}
\end{equation*}
$$

and similarly for $M_{\mathbf{k}}(\omega)$ and $\delta_{ \pm}(\omega)$ the transform of $m_{\mathbf{k}}(t)$ and $\delta \Delta(t) \pm \delta \Delta^{*}(t)$ respectively. This allows us to express the microscopic variables in terms of the fluctuations of $\Delta$,

$$
\begin{align*}
& A_{\mathbf{k}}^{+}(\omega)=\frac{\mathrm{i} \omega \alpha_{\mathbf{k}}^{+}\left(0^{-}\right)+2 \mathrm{i} \epsilon_{\mathbf{k}} \alpha_{\mathbf{k}}^{-}\left(0^{-}\right)}{\omega^{2}-4 \epsilon_{\mathbf{k}}^{2}}+\frac{2 \xi_{\mathbf{k}} F\left(\epsilon_{\mathbf{k}}\right)}{\omega^{2}-4 \epsilon_{\mathbf{k}}^{2}} \delta_{+}(\omega)+\frac{\omega F\left(\epsilon_{\mathbf{k}}\right)}{\omega^{2}-4 \epsilon_{\mathbf{k}}^{2}} \delta_{-}(\omega)  \tag{S11}\\
& A_{\mathbf{k}}^{-}(\omega)=\frac{\mathrm{i} \omega \alpha_{\mathbf{k}}^{-}\left(0^{-}\right)+2 \mathrm{i} \epsilon_{\mathbf{k}} \alpha_{\mathbf{k}}^{+}\left(0^{-}\right)}{\omega^{2}-4 \epsilon_{\mathbf{k}}^{2}}+\frac{2 \epsilon_{\mathbf{k}} F\left(\epsilon_{\mathbf{k}}\right)}{\omega^{2}-4 \epsilon_{\mathbf{k}}^{2}} \delta_{-}(\omega)+\frac{\omega \xi_{\mathbf{k}} F\left(\epsilon_{\mathbf{k}}\right)}{\epsilon_{\mathbf{k}}\left(\omega^{2}-4 \epsilon_{\mathbf{k}}^{2}\right)} \delta_{+}(\omega)  \tag{S12}\\
& M_{\mathbf{k}}(\omega)=\frac{\mathrm{i} m_{\mathbf{k}}\left(0^{-}\right)}{\omega} \tag{S13}
\end{align*}
$$

and finally to eliminate them using the resummation

Eq. (??), yielding a closed system of equations on $\delta_{ \pm}$:

$$
\begin{equation*}
M\binom{\delta-\delta^{*}}{\delta+\delta^{*}}=-\mathrm{i}\binom{S_{-}}{S_{+}} \tag{S14}
\end{equation*}
$$

The fluctuation matrix $M$ is introduced below equation (7) of the main text, and the sums encoding the initial conditions on $\Delta$ are given by

$$
\begin{array}{r}
S_{+}=\int \frac{\mathrm{d}^{3} \mathbf{k}}{(2 \pi)^{3}}\left[\frac{\omega \xi_{\mathbf{k}} \alpha_{\mathbf{k}}^{+}\left(0^{-}\right)+2 \xi_{\mathbf{k}} \epsilon_{\mathbf{k}} \alpha_{\mathbf{k}}^{-}\left(0^{-}\right)}{\epsilon_{\mathbf{k}}\left(\omega^{2}-4 \epsilon_{\mathbf{k}}^{2}\right)}+\frac{\Delta_{i}}{\epsilon_{\mathbf{k}}} \frac{m_{\mathbf{k}}\left(0^{-}\right)}{\omega}\right] \\
+\frac{\delta \Delta_{0}+\delta \Delta_{0}^{*}}{g_{i} \omega} \quad(\mathrm{~S} 15) \\
S_{-}=\int \frac{\mathrm{d}^{3} \mathbf{k}}{(2 \pi)^{3}} \frac{\omega \alpha_{\mathbf{k}}^{-}\left(0^{-}\right)+2 \epsilon_{\mathbf{k}} \alpha_{\mathbf{k}}^{+}\left(0^{-}\right)}{\omega^{2}-4 \epsilon_{\mathbf{k}}^{2}}+\frac{\delta \Delta_{0}-\delta \Delta_{0}^{*}}{g_{i} \omega} \tag{S16}
\end{array}
$$

Inverting Eq. (??) and switching back to the time domain [using the inverse Laplace transformation $f(t)=$ $\left.-\frac{1}{2 \pi} \int_{+\infty+\mathrm{i} \eta}^{-\infty+\mathrm{i} \eta} \mathrm{d} z \mathrm{e}^{-\mathrm{i} z t} F(z)\right]$, yields, for the time-evolution of $\Delta$ :

$$
\begin{equation*}
\binom{\delta \Delta(t)-\delta \Delta^{*}(t)}{\delta \Delta(t)+\delta \Delta^{*}(t)}=-\int_{+\infty+\mathrm{i} \eta}^{-\infty+\mathrm{i} \eta} \frac{\mathrm{~d} z}{2 \mathrm{i} \pi} \mathrm{e}^{-\mathrm{i} z t} M^{-1}\binom{S_{-}}{S_{+}} \tag{S17}
\end{equation*}
$$

We now input the initial condition corresponding to the interaction quench, that is (as explained above Eq. (3) of the main text) $\alpha_{\mathbf{k}}^{ \pm}=m_{\mathbf{k}}=0$, and $\delta \Delta_{0}=\epsilon g_{i} / \Delta_{i}$, which converts into $S_{-}=0$ and $S_{+}=2 \epsilon / \omega \Delta_{i}$. Finally, we
derive Eq. (6) of the main text by closing the integration contour in Eq. (??) around the branch cuts of $M$ (see Fig. 1 in Ref. [7]), and by remarking that the modulusmodulus response function (Eq. (7)) is related to $M$ by $f(z)=-\left(M^{-1}\right)_{22} / \Delta_{i}$.

## CALCULATION OF $M$

We numerically evaluate the elements of $M$ at nonzero temperature using the Kramers-Kronig relation

$$
\begin{align*}
\frac{M_{11}}{\omega_{0}^{2}}=\frac{M_{22}}{\omega_{0}^{2}-4 \Delta^{2}} & =\int_{-\infty}^{+\infty} \frac{\rho_{f}(\omega) \mathrm{d} \omega}{\omega_{0}-\omega}  \tag{S18}\\
M_{12} & =\int_{-\infty}^{+\infty} \frac{\rho_{g}(\omega) \mathrm{d} \omega}{\omega_{0}-\omega} \tag{S19}
\end{align*}
$$

In dimensionless units $\left(\check{\rho}=(2 \pi)^{3} \Delta \rho / k_{\Delta}^{3}, \check{\omega}=\omega / 2 \Delta \check{\beta}=\right.$ $\beta \Delta)$, we have

$$
\begin{align*}
& \check{\rho}_{f}=\frac{\pi}{2 \check{\omega}} \frac{\operatorname{th}(\check{\beta} \check{\omega})}{\sqrt{\breve{\omega}^{2}-1}}\left(k_{1}(\check{\omega})+k_{2}(\check{\omega})\right)  \tag{S20}\\
& \check{\rho}_{g}=\frac{\pi}{2} \operatorname{th}(\check{\beta} \check{\omega})\left(k_{1}(\check{\omega})-k_{2}(\check{\omega})\right) \tag{S21}
\end{align*}
$$

with

$$
\left\{\begin{array} { l } 
{ k _ { 1 } ( \omega ) = \Theta ( \check { \omega } - 1 ) \sqrt { \check { \mu } + \sqrt { \check { \omega } ^ { 2 } - 1 } } }  \tag{S22}\\
{ k _ { 2 } ( \omega ) = \Theta ( \check { \omega } - 1 ) \Theta ( \sqrt { 1 + \check { \mu } ^ { 2 } } - \check { \omega } ) \sqrt { \check { \mu } - \sqrt { \check { \omega } ^ { 2 } - 1 } } \quad \text { if } \check { \mu } > 0 }
\end{array} \left\{\begin{array}{l}
k_{1}(\omega)=\Theta\left(\check{\omega}-\sqrt{1+\check{\mu}^{2}}\right) \sqrt{\check{\mu}+\sqrt{\check{\omega}^{2}-1}} \\
k_{2}(\omega)=0
\end{array} \quad \text { if } \check{\mu}<0\right.\right.
$$

The divergence of the integrals (??)-(??) in $\omega=\omega_{0}$ is easily compensated by adding/subtracting $\rho\left(\omega_{0}\right)$. Still, the evaluation can become difficult when $\check{\omega}_{0}$ is close to
the angular points 1 or $\sqrt{1+\check{\mu}^{2}}$. To overcome this, we subtract the nearly diverging behavior of the integrand. (For here on, all quantities are implicitly dimensionless).

- When $\omega_{0} \rightarrow 1^{+}$, the integrand nearly diverges around $\omega \rightarrow 1^{+}$. We set $e=(\omega-1) /\left(\omega_{0}-1\right)$ and derive:

$$
\begin{align*}
& \frac{\rho_{f}(\omega)-\rho_{f}\left(\omega_{0}\right)}{\omega_{0}-\omega}=\frac{\pi \operatorname{th} \beta}{\sqrt{2}} \frac{\sqrt{\mu}}{\left(\omega_{0}-1\right)^{3 / 2}} \frac{1}{\sqrt{e}+e}+O\left(\omega_{0}-1\right)^{-1 / 2}  \tag{S23}\\
& \frac{\rho_{g}(\omega)-\rho_{g}\left(\omega_{0}\right)}{\omega_{0}-\omega}=O\left(\omega_{0}-1\right)^{-1 / 2} \tag{S24}
\end{align*}
$$

- When $\omega_{0} \rightarrow \omega_{3}^{ \pm}$, the integrand nearly diverges around $\omega \rightarrow \omega_{3}^{\mp}$. We set $e=\left(\omega-\omega_{3}\right) /\left(\omega_{3}-\omega_{0}\right)$ and derive:

$$
\begin{align*}
& \frac{\rho_{f}}{\omega_{0}-\omega}=\frac{\pi \operatorname{th}\left(\beta \omega_{3}\right)}{\sqrt{2}} \frac{1}{\omega_{3} \sqrt{\mu}} \frac{1}{\omega_{0}-\omega_{3}} \frac{1}{1+e}+O(1)  \tag{S25}\\
& \frac{\rho_{g}}{\omega_{0}-\omega}=\frac{\pi \operatorname{th}\left(\beta \omega_{3}\right)}{\sqrt{2}} \frac{\sqrt{\mu}}{\omega_{0}-\omega_{3}} \frac{1}{1+e}+O(1) \tag{S26}
\end{align*}
$$

[7] V. Gurarie, Phys. Rev. Lett. 103, 075301 (2009), URL https://link.aps.org/doi/10.1103/PhysRevLett.
[| H. Kurkjian and J. Tempere, New Journal of Physics 19, 113045 (2017), URL http://stacks.iop.org/ $1367-2630 / 19 / i=11 / a=113045$.

