

## $\label{eq:2.1} ZnWO4\ grains\ characterisation\ for\ GRAiNITA-a\\ new-generation\ calorimeter$

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## **ZnWO<sub>4</sub> grains characterisation for GRAINITA – a new-generation calorimeter**

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**Introduction:** The electromagnetic calorimeters for high-energy physics experiments are often based on the use of scintillator materials and, in order to contain the cost of the detectors, the sampling technique is privileged with respect to the use of uniform inorganic crystals. With **GRAINITA**, we propose a new version of the Shashlik detector, for possible application in FCC e+e- experiments. The idea is to use, in the same volume, a mixture of high-Z scintillator grains and a high-density transparent liquid as absorber. Due to the stochastic confinement of the light, as in the LiquidO detection technique [1], the scintillation light remains localized in the zone of production and it can be collected on SiPMs, by means of wavelength shifting fibres, uniformly distributed in the detection volume. This new detector will provide extremely fine sampling of the electromagnetic shower and for this reason a very good photon energy resolution (R~2%/ $\sqrt{E}$ ) is expected with respect to conventional Shashlik detectors (typically R~10%/ $\sqrt{E}$ , [2]).

The GRAINITA detector

**Opaque scintillator technique for the development of a new** Shashlik calorimeter with extremely fine sampling

**Candidates scintillators: ZnWO<sub>4</sub> and BGO Candidate absorber: CH**<sub>2</sub>I<sub>2</sub>



#### Shashlik calorimeter: alternating layers of scintillator and absorber







- $\rightarrow$  For improved high energy and very good sampling
- $\rightarrow$  For costs reduction costs: grains production is less expensive than volumetric crystals production
- $\rightarrow$  For good energy measurements of the jets: study of the possibility of measuring the fraction of energy in the shower via Pulse Shape Discrimination.

#### ZnWO<sub>4</sub> can be produced directly in grain-shape via spontaneous crystallization X-ray excited emission spectrum

BGO can be reduced in the form of grains with the

• LY= 10 kph/MeV

- $\tau = 20 \ \mu s$
- $Z_{eff} = 61$
- $\lambda_{max} = 490 \text{ nm}$ • Crystal price/grains price = 2.5

• LY= 8-10 kph/MeV

•  $\tau = 300 \text{ ns}$ 



X-ray excited emission spectrum

	ZnWO <sub>4</sub>	$CH_2I_2$	Mix
Density [g/cm <sup>3</sup> ]	7.62	3.32	5.47
Refractive index	2.1	1.74	
X <sub>0</sub> [cm]	1.2	2.67	1.65
Molière radius[cm]	1.98	4.58	2.77

#### BGO $CH_2I_2$ Mix Density [g/cm<sup>3</sup>] 7.13 3.32 5.47 **Refractive index** 1.74 2.15 1.12 X<sub>0</sub> [cm] 2.67 1.50 Molière radius [cm] 2.26 4.58 2.84

## **Crystal grains production**

ISMA developed the technique to grow ZnWO<sub>4</sub> single crystal granules using the flux method. Several methods have been investigated:

- **NaCl as a solvent + ZnWO<sub>4</sub> salt**: ZnWO<sub>4</sub> grains yield  $\approx$  10 wt %; 44% of grains with size > 1mm, 33% of grains with size between 0.5 and 1 mm.
- Na<sub>2</sub>WO<sub>4</sub> as a solvent + ZnO and WO<sub>3</sub> : ZnWO<sub>4</sub> grains yield  $\approx 25$  wt % but brownish color
- Na<sub>2</sub>WO<sub>4</sub> as a solvent + ZnWO<sub>4</sub> salt: ZnWO<sub>4</sub> grains yield  $\approx$  85 wt %, colorless and transparent, average

## **ZnWO<sub>4</sub> scintillation properties: grains vs crystals**

## The experimental set-up:

- ZnWO<sub>4</sub> crystals:  $2 \times 2 \times 4 \text{ cm}^3$  and  $1 \times 1 \times 1 \text{ cm}^3$
- ZnWO<sub>4</sub> plates (cut from the same single crystal) : 2 cm x 2 cm, different thickness: 0.85 mm, 1.03 mm, 2.14 mm, 3.14 mm, 4.25 mm
- ZnWO<sub>4</sub> grains, approximate later size: 0.5÷1.5 mm
- High gain PMT Hamamastu R2083, silicon grease, teflon tape
- Integrator circuit:  $\tau_{\rm RC} = 100 \ \mu s$
- Digital oscilloscope: LeCroy WavePro 715Zi (acquisition at 100MS/s)
- <sup>241</sup>Am (60keV, completely absorbed in one grain), <sup>137</sup>Cs, <sup>22</sup>Na, <sup>133</sup>Ba sources



### • $Z_{eff} = 73$ • $\lambda_{\rm max} = 480 \, \rm nm$ • Crystal price/grains price > 1.1

« Planetary Ball Milling » technique:

## size of grains $\approx 1$ mm lateral size.



1<sup>st</sup> batch: 40 g Size ~ 1÷3 mm

2<sup>nd</sup> batch: 170 g 3<sup>rd</sup> batch: 1380 g Size ~ 1 mm Size ~ 1÷2 mm

For **GRAINITA**, ZnWO<sub>4</sub> crystalline grains were grown with the flux method from thoroughly mixed ZnWO<sub>4</sub> (5N) and Na<sub>2</sub>WO<sub>4</sub> (4N) starting materials, as this resulted to be the best effective solution to obtain sub-millimetric grains. Starting materials were charged into Pt crucible, placed into the resistive heating furnace, heated up to T=1250°C and then kept for 10 hours at this temperature to assure complete homogenization of the melt. According to he required size of the grains, the melt was cooled down to 750°C and then it was allowed to spontaneous cooling down to room temperature. Obtained ZnWO<sub>4</sub> crystalline grains were rinsed in distilled water at 100°C and dried at T= 400°C.

Spectrophotometer: PerkinElmer Lambda 850 UV/Vis



The light yield and energy resolution of a 1cm<sup>3</sup> ZnWO<sub>4</sub> crystal are comparable with that of a same-sized BGO. At 662 keV we measured an energy resolution of 12.9% for ZnWO<sub>4</sub> and 13% for BGO.







The second production of grains shows a much smaller variance in the amplitude of the <sup>241</sup>Am peak at 60keV, thus indicating a better homogeneity in the grains light yield.



	Max Amp	Sigma	
Crystals	[mV]	[mV]	
l grain 1st batch	257.1	36.1	
l grain 1st batch	307.7	34.9	PMT
I grain 2nd batch	302.5	32.9	
I grain 2nd batch	290.2	36.6	
Grains 1st batch	260.7	59.2	
Grains 2nd bacth	292.7	43.1	PMT
).85 mm plate	298.2	36.9	
1.03 mm plate	300.3	34.0	
2.14 mm plate	284.2	36.6	PMT
3.14 mm plate	266.9	35.1	
1.25 mm plate	272.8	34.5	
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L cm <sup>3</sup> crystal	181.4	26.7	PMT
2x2x4 cm <sup>3</sup> crystal	125.9	23.2	

ΡΜΤ

# **Transmittance spectra for ZnWO**<sub>4</sub> samples of different thickness

## **Conclusions and future work**

Crystalline grains of ZnWO<sub>4</sub>, grown with the flux method, proved to have scintillation properties comparable with that of volumetric optics. The latest batch of grains produced by ISMA (delivered at the end of October) is enough to assemble a medium-size prototype with approximate size of 5 cm x 5 cm x 10 cm and equipped with 16 WLS fibers, and SiPMs for the signal read-out. At present, we are working on the selection of the best suited optical fiber to be coupled to ZnWO<sub>4</sub> and BGO grains, respectively, as well as on the selection of the best suited liquid to be used as absorber. In the next future, we intent to develop two medium-size GRAiNITA prototypes, employing BGO and ZnWO<sub>4</sub> grains respectively, and to characterize the detectors response with cosmic rays and during an in-beam campaign.

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