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The absorption spectrum of nitrous oxide between 7250 and 7653 cm⁻¹

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Abstract

45
46
47 We revisit the weak absorption spectrum of natural nitrous oxide between 7250 and
48 7653 cm^{-1} for which rovibrational assignments were previously reported in Lu et al.
49 (doi:10.1016/j.jqsrt.2012.03.005). Our main goal is to provide intensity information for the
50 high number of lines detected by cavity ring down spectroscopy in the region. Line parame-
51 ters of about 4180 transitions were retrieved from a line-by-line profile adjustment. They
52 belong to 69 bands of the $^{14}\text{N}_2^{16}\text{O}$, $^{14}\text{N}^{15}\text{N}^{16}\text{O}$, $^{15}\text{N}^{14}\text{N}^{16}\text{O}$, and $^{14}\text{N}_2^{18}\text{O}$ isotopologues, present
53 in natural isotopic abundance in the sample. Eight bands are newly reported. The band-by-
54 band analysis allowed for the determination of refined values of the rovibrational parame-
55 ters of a total of 68 upper vibrational states. Eight $^{14}\text{N}_2^{16}\text{O}$ bands were found affected by lo-
56 cal rovibrational perturbations and a total of 31 extra lines due to an intensity transfer could
57 be identified. The comparison with the recent NOSL-296 and Ames line lists is discussed in
58 detail. The overall agreement is very satisfactory. As concerns line intensities, the NOSL and
59 Ames lists show both specific advantages and drawbacks which are discussed.

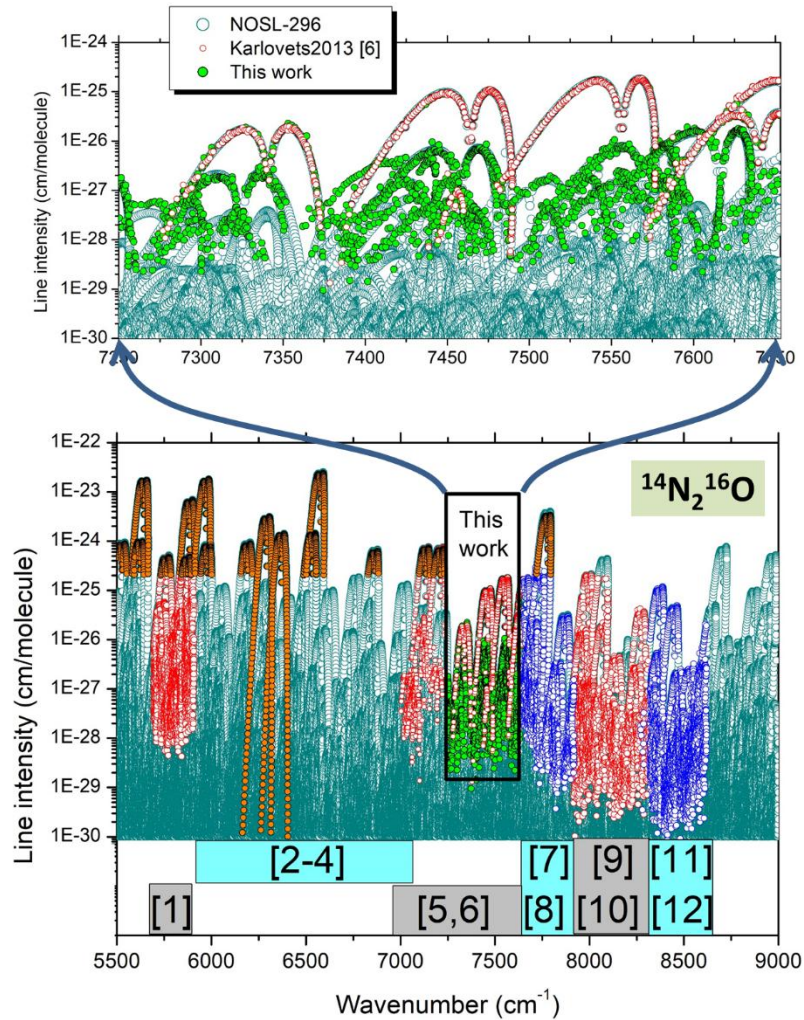
60 **1. Introduction**

61 The present study is part of a long-term project aiming at characterizing the high sensitivity
62 absorption spectrum of nitrous oxide in the near-infrared by cavity ring down spectroscopy (CRDS).
63 **Fig. 1** presents an overview of our preceding studies [1-12] which cover continuously the 5696-8630
64 cm^{-1} range. The HITRAN2020 list [13] and the very recent Nitrous Oxide Spectroscopic line list (NOSL-
65 296) (NOSL) [14] of $^{14}\text{N}_2^{16}\text{O}$ are included in the plot. The present version of the HITRAN list is mostly
66 based on the SISAM.N2O line list constructed by Toth from spectra measured by Fourier transform
67 spectroscopy (FTS) [15]. Most of the SISAM line positions and line intensities are calculated values
68 obtained using spectroscopic parameters obtained from a band-by-band fit of line positions and line
69 intensities. The SISAM list covers the 500-7800 cm^{-1} spectral range and includes transitions with line
70 strength larger than about 2×10^{-25} $\text{cm}/\text{molecule}$ at 296 K. The very recent NOSL-296 list [14] is a
71 calculated line list with positions obtained using an effective Hamiltonian (EH) model with
72 parameters fitted on experimental positions and intensities computed from effective dipole
73 moments (EDM) with parameters fitted on measured line intensities. The NOSL-296 intensity cut off
74 is 1×10^{-30} $\text{cm}/\text{molecule}$ at 296 K, about five orders of magnitude lower than the HITRAN intensity
75 threshold. Our previous CRDS studies [1-12] provided a large number of new line positions which
76 were valuable to refine the parameters of the EH model. As concerns line intensity calculations, they
77 require a set of EDM parameters for each ΔP value ($P = 2V_1 + V_2 + 4V_3$ is the polyad number where $V_{i=1-3}$
78 are the vibrational quantum numbers). The sensitivity of our CRDS setup allows for the detection of
79 lines with intensity on the order of 1×10^{-29} $\text{cm}/\text{molecule}$ at 296 K which were valuable to fit the
80 necessary EDM parameters [15]. We have highlighted in **Fig. 1** (red and blue open circles), the
81 transitions for which line intensities were previously retrieved from the CRDS spectra. As illustrated
82 by this figure, line intensities were reported for only part of the recorded CRDS spectra which makes
83 some sets of EDM parameters inaccurate or incomplete.

84 In the present work, we reconsider the 7250 - 7653 cm^{-1} interval, first assigned by Lu et al. [5]
85 in 2012. Intensity information was retrieved in Ref. [6] for about 740 lines of six strong bands
86 measured at 2 Torr and 10 Torr in the region, which allowed fitting the dominant EDM parameters.
87 Note that due to the relative weakness of the absorption in the region, (line intensities are smaller
88 than 2×10^{-25} $\text{cm}/\text{molecule}$), no transitions are included in the HITRAN database in the considered
89 interval. Here, using the same spectra as in Refs. [5,6], line intensities were systematically retrieved
90 using a multiline fitting program for about 4180 transitions presented on the upper panel of **Fig. 1**.
91 The resulting experimental list combined with improved predictions of the N_2O absorption spectrum
92 will allow for:

- 93 i. the rovibrational assignment of a few additional bands not reported in Ref. [5],

- 94 ii. detailed analysis of some local rovibrational interactions leading to the detection of
 95 extra lines resulting from an intensity transfer,
 96 iii. validation tests of the NOSL-296 list of $^{14}\text{N}_2^{16}\text{O}$ and the Ames line list for all the
 97 isotopologues, in particular, for the intensities of the weak bands. Deviations
 98 compared to NOSL intensities will help to refine the set of EDM parameters.



99

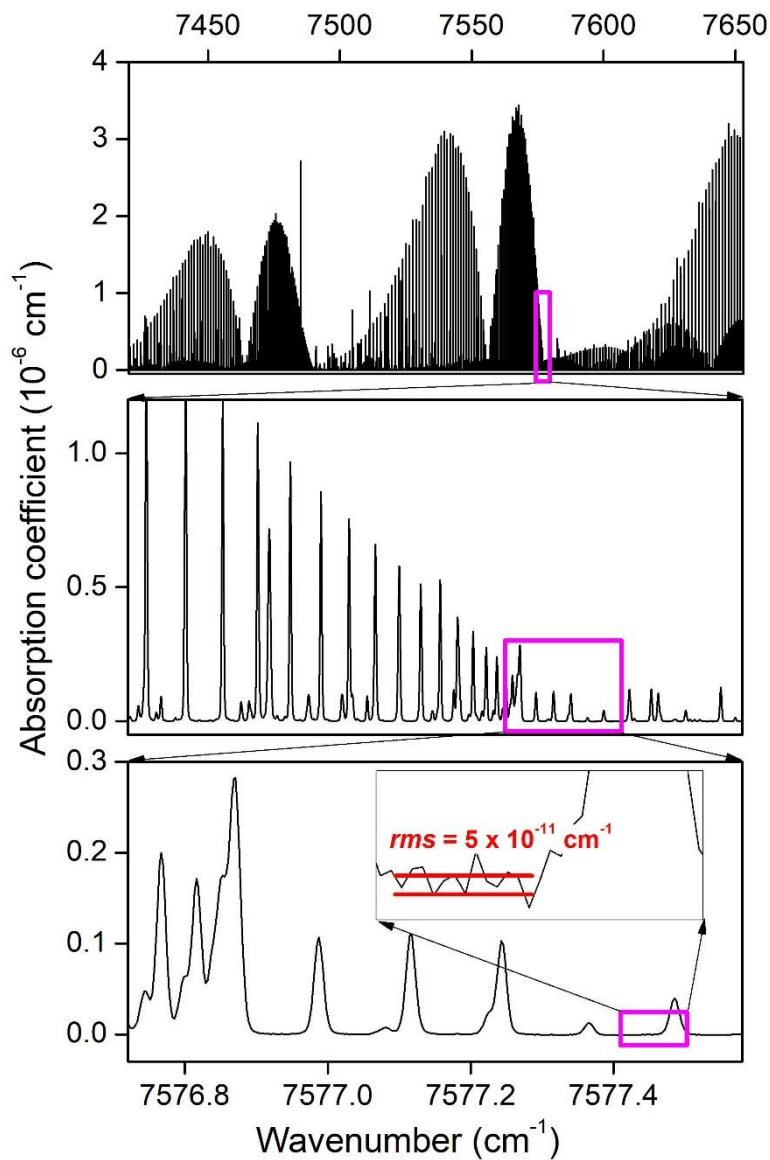
100 **Fig. 1.** Overview of various line lists of $^{14}\text{N}_2^{16}\text{O}$ between 5500 and 9000 cm^{-1} .
 101 *Lower panel:* NOSL-296 [14] and HITRAN2020 [13] line lists (cyan and orange symbols, respectively).
 102 The range of previous CRDS investigations is indicated: 5696-5910 cm^{-1} [1], 5905-7066 cm^{-1} [2-4],
 103 6950-7653 cm^{-1} [5,6], 7647-7918 cm^{-1} [7,8], 7915-8334 cm^{-1} [9,10], 8320-8622 cm^{-1} [11,12], and this
 104 work. Red and blue circles correspond to CRDS intensities previously reported.
 105 *Upper panel:* Enlargement corresponding to the 7250 - 7653 cm^{-1} range of the present study. The line
 106 intensities of a few bands were previously reported in Ref. [6] (red circles). About 4180 line
 107 intensities are retrieved in the present study (light green symbols).
 108

109 The rest of this paper is organized as follows. The experimental setup and the construction of
 110 the measured lists are briefly recalled in Section 2. In Section 3, we present the spectra analysis, the
 111 refined band-by-band fitting of the G_v , B_v , D_v , and H_v band parameters and the identification of

112 resonance perturbations. Finally, in Section 4, the comparison of the experimental data to the NOSL
113 and Ames line lists will allow for discussing their respective advantages and drawbacks.

114 2. Experimental details

115 The description of the CRDS recordings has been presented in Refs. [5,6] and is not repeated
116 here. The cavity ring down spectrometer has been described in detail in Refs. [16-18]. The ring down
117 cell (142 cm long) was filled with nitrous oxide in natural isotopic abundance (Air Liquide, Alphagaz,
118 purity $\geq 99.99\%$). The analyzed spectra were recorded at room temperature (294 ± 0.3 K) with a
119 sample pressure value of 10 Torr. The wavenumber scale of the spectra was calibrated with the help
120 of a lambdameter (Burleigh WA1650) and accurate reference line positions of H_2O (present as an
121 impurity) taken from the HITRAN database [19]. The typical value of the noise equivalent absorption
122 coefficient was $\alpha_{min} \sim 5 \times 10^{-11} \text{ cm}^{-1}$ as illustrated on **Fig. 2**



123
124 **Fig. 2.** CRDS spectrum of natural nitrous oxide recorded at a pressure of 10.0 Torr.

125

126 The overview of the N₂O spectrum at 10 Torr between 7250 and 7653 cm⁻¹ is presented in **Fig.**
127 **2.** The noise equivalent absorption on the order of $\alpha_{min} \approx 5 \times 10^{-11}$ cm⁻¹. It led to the detection of the
128 large number of lines with intensity as small as 10⁻²⁹ cm/molecule involving many hot bands of the
129 main ¹⁴N₂¹⁶O isotopologue and the contribution of minor ¹⁴N¹⁵N¹⁶O, ¹⁵N¹⁴N¹⁶O, and ¹⁴N₂¹⁸O
130 isotopologues in natural isotopic abundance.

131 A LabVIEW interactive least squares multi-lines fitting program was used to get the line centers
132 and intensities. The line profile was assumed to be of Voigt type with the Gaussian component
133 constrained to the calculated value of the Doppler profile. The fitting process produced the line
134 position, integrated absorbance, Lorentzian widths, and the corresponding local baseline (assumed
135 to be a linear function of the wavenumber).

136 The complete line list retrieved from the 10 Torr spectra consist of 6656 lines between 7250
137 and 7653 cm⁻¹ with line intensities ranging between 10⁻³⁰ and 2×10⁻²⁵ cm/molecule. Water vapor is
138 present as an impurity in our sample. Due to strong lines located in the region (line intensity up to
139 1×10⁻²⁰ cm/molecule), 1475 water lines were identified by comparison to the HITRAN spectroscopic
140 database [13], and removed from the list. A set of about 5180 lines was thus left for the rovibrational
141 assignment.

142 **3. Rovibrational analysis**

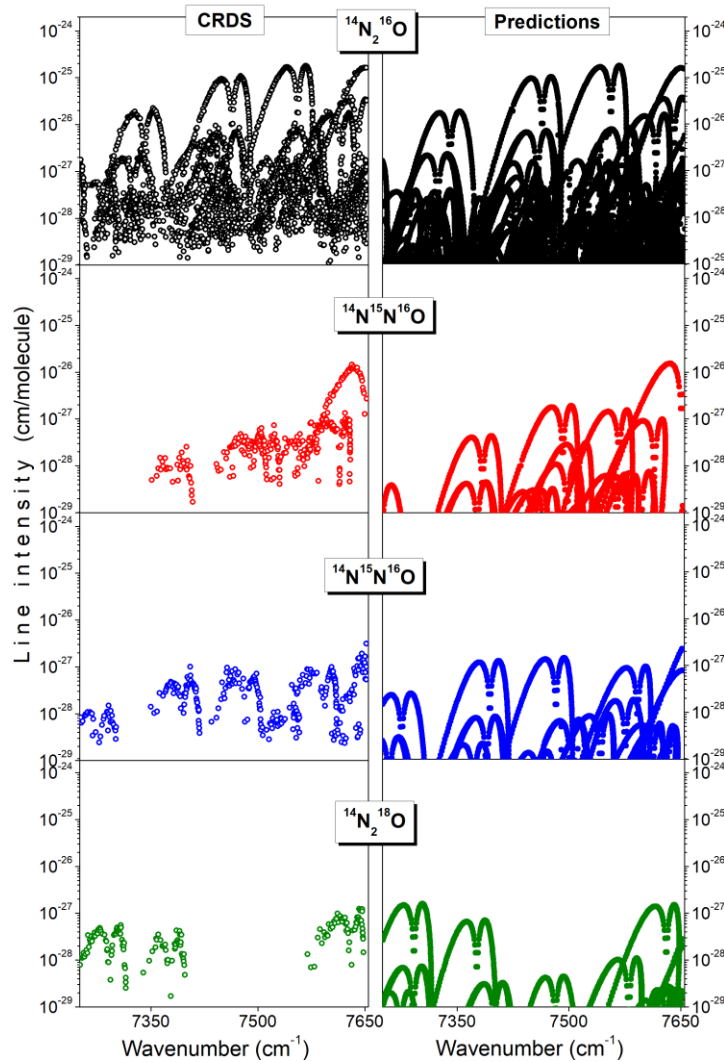
143 *3.1. Rovibrational assignments*

144 A total of 4186 lines belonging to 69 bands of ¹⁴N₂¹⁶O, ¹⁴N¹⁵N¹⁶O, ¹⁵N¹⁴N¹⁶O, and ¹⁴N₂¹⁸O were
145 assigned in the 7250-7653 cm⁻¹ region. The observed bands correspond to the ΔP= 12-14 series of
146 transitions. In the following, we will label the bands both (i) with the traditional V₁V₂l₂V₃ normal
147 mode notation where V₁V₂l₂V₃ corresponds to the maximum value of the modulo of the expansion
148 coefficients of the eigenfunctions in the normal mode basis and (ii) with the cluster labeling notation
149 of the upper state: (P l₂ i) where P= 2V₁+V₂+4V₃, l₂ is the quantum number of the angular momentum
150 associated to the vibrational bending mode and i is the order number within the considered polyad
151 increasing with the energy. Due to strong vibrational mixing, the normal mode labeling can be am-
152 biguous while it is not the case for the cluster labeling notation.

153 As in the previous analysis [5], the assignments were carried out by comparison with the
154 predictions of the effective operator models. In the case of the ¹⁴N₂¹⁶O isotopologue, we used the
155 recent NOSL-296 line list [14]. The NOSL list covers the 0.02–13378 cm⁻¹ spectral range and contains
156 almost 900,000 lines with line intensities larger than 10⁻³⁰ cm/molecule. This list is based on a non-
157 polyad model of the effective Hamiltonian [20] and effective dipole moment operators [14] with
158 parameters fitted against experimental data. In the studied region, the NOSL line list contains about

159 16860 $^{14}\text{N}_2^{16}\text{O}$ transitions above an intensity cut-off of 10^{-30} cm/molecule. As the line position values
160 of Ref. [5] were used as input data for the fit of the NOSL EH parameters, the agreement between
161 the NOSL and presently measured position values is expected to be very good (see below).

162 For minor isotopologues, the $^{14}\text{N}^{15}\text{N}^{16}\text{O}$ and $^{15}\text{N}^{14}\text{N}^{16}\text{O}$ line positions were calculated using the
163 polyad models of the effective Hamiltonian reported in Ref. [21]. The parameters of these models
164 were refitted to the experimental position input files of Ref. [21] supplemented by recently published
165 measurements from Refs. [11,12,22,23]. The same procedure for the EH model of $^{14}\text{N}_2^{18}\text{O}$ published
166 in Ref. [10] was used. The dipole moment parameters of $^{14}\text{N}_2^{16}\text{O}$ [14] were used to fit the measured
167 line intensities of $^{14}\text{N}^{15}\text{N}^{16}\text{O}$ and $^{15}\text{N}^{14}\text{N}^{16}\text{O}$ [9,11,12,15,23-25], and $^{14}\text{N}_2^{18}\text{O}$ [1,10-12,15,25]. Only those
168 parameters that affect the measured intensities were fitted. The remaining parameters were fixed at
169 the corresponding values of the $^{14}\text{N}_2^{16}\text{O}$ EDM parameters. The eigenfunctions of the fitted effective
170 Hamiltonian polyad models were used for the intensity calculations.



171

172

173 **Fig.3.** Overview comparison of the measured and predicted line lists of the four nitrous oxide
174 isotopologues contributing to the absorption spectrum between 7250 and 7653 cm^{-1} : $^{14}\text{N}_2^{16}\text{O}$ (446),

175 $^{14}\text{N}^{15}\text{N}^{16}\text{O}$ (456), $^{15}\text{N}^{14}\text{N}^{16}\text{O}$ (546), and $^{14}\text{N}_2^{18}\text{O}$ (448). Note that the line intensities are plotted in
 176 logarithmic scale and include the isotopic abundance factor.

177 **Fig. 3** shows the overview comparison of the CRDS observations of the four N_2O
 178 isotopologues contributing to the analyzed spectrum, to the predictions of their respective effective
 179 operator models. About 990 very weak lines (intensity less than 10^{-29} cm/molecule) were left unas-
 180 signed at the final stage of the assignment process.

181 **Table 1** summarizes the number of transitions and bands measured in this work. Overall,
 182 seven $^{14}\text{N}_2^{16}\text{O}$ bands and one $^{15}\text{N}^{14}\text{N}^{16}\text{O}$ band are newly reported. Their vibrational assignment is
 183 marked with bold character in **Table 2**. Both the normal mode and (P / i) cluster labeling assignments
 184 are given for each band in **Table 2**. Note that we correct here a ten of erroneous values of the i rank-
 185 ing number given in Ref. [5],

186
 187 **Table 1.** Number of transitions and bands of the different N_2O isotopologues measured in the 7250–
 188 7653 cm^{-1} region.

189

Molecule	HITRAN notation	Abundance [13]	Number of transitions	Number of bands ^a
$^{14}\text{N}_2^{16}\text{O}$	446	9.903×10^{-1}	3328	51(7)
$^{14}\text{N}^{15}\text{N}^{16}\text{O}$	456	3.641×10^{-3}	359	6(0)
$^{15}\text{N}^{14}\text{N}^{16}\text{O}$	546	3.641×10^{-3}	325	8(1)
$^{14}\text{N}_2^{18}\text{O}$	448	1.986×10^{-3}	174	4(0)
Total			4186	69(8)

190 **Note:**

191 ^a The number within parentheses correspond to newly assigned bands.

192

193 The N_2O assigned line list provided as Supplementary Material, includes the isotopologue
 194 identification, rovibrational assignment, measured line parameters, and corresponding values
 195 predicted by the effective operator models. For comparison purpose, the Ames line parameters are
 196 also included (see **Section 4**).

197 3.2. Band-by-band analysis

198 As usual, we used the following equation of the vibration-rotational energy levels to fit the
 199 upper state spectroscopic parameters to the observed wavenumbers:

$$200 F_v(J) = G_v + B_v J(J + 1) - D_v J^2(J + 1)^2 + H_v J^3(J + 1)^3 \quad (1)$$

201 where G_v is the vibrational term value, B_v is the rotational constant, D_v and H_v are the centrifugal
 202 distortion constants, J is the angular momentum quantum number. In the case of hot bands, the e
 203 and f sub-bands were considered independently. The results of the band-by-band fit of the
 204 spectroscopic parameters are given as a Supplementary Material. The lower state constants were
 205 constrained to their literature values from Refs. [15,26] and are included in the Supplementary

206 Material. A total of 68 bands were fitted. The typical *rms* value of the fits is about $6 \times 10^{-4} \text{ cm}^{-1}$,
 207 which is consistent with our claimed accuracy on the line positions ($\sim 10^{-3} \text{ cm}^{-1}$).

208 The retrieved constants ordered according to the band centers are listed in **Table 2**, for the
 209 four isotopologues. In the case of bands located near the borders of the investigated region, the
 210 input data set was supplemented with measurements from Refs. [5,8]. All these lines are indicated
 211 with their source in the Supplementary Material.

212 3.3. Local rovibrational perturbations

213 Eight $^{14}\text{N}_2^{16}\text{O}$ bands were found affected by local Coriolis or anharmonic interactions. Five of
 214 these interactions are evidenced for the first time. For some previously analyzed perturbations, we
 215 could newly assign some extra lines. The corresponding perturbed transitions were excluded from
 216 the fit of the spectroscopic constants. The perturbations were identified on the basis of the NOSL-
 217 296 line list [14]. **Table 3** lists for each of the perturbed bands, their center, the coupling mechanism,
 218 the *J* values corresponding to the energy crossing of the interacting states (J_{cross}), and the number of
 219 extra lines assigned in this work. When available, the references where these perturbations were
 220 firstly reported is given in the last column of **Table 3**.

221
 222 **Table 3.** Observed local rovibrational perturbations of $^{14}\text{N}_2^{16}\text{O}$ bands between 7250 and 7653
 223 cm^{-1} .
 224

Affected band	Center (cm^{-1})	Interacting states ^a	Interaction mechanism	J_{cross}	Number of extra lines	Ref. ^b
2112e-0000e 2112f-0000e	7443.008	(13 1 7) \leftrightarrow (13 1 6)	Intrapolyad anharmonic	17	6 3	[4,5]
6000e-0000e	7463.985	(12 0 14) \leftrightarrow (12 3 5)	Intrapolyad Coriolis	36	2	TW
6200e-1000e	7525.861	(14 0 19) \leftrightarrow (16 0 3)	Interpolyad anharmonic	24	11	[5]
6110e-0110e 6110f-0110f	7570.895	(13 1 15) \leftrightarrow (14 2 9)	Interpolyad Coriolis	<i>e</i> : 17	6	[5]
		(13 1 15) \leftrightarrow (14 2 9)	Interpolyad Coriolis	<i>f</i> : 21	6	[5]
		(13 1 15) \leftrightarrow (14 2 6)	Interpolyad Coriolis	<i>e</i> : 27	19	[5]
		(13 1 15) \leftrightarrow (14 0 11)	Interpolyad Coriolis	<i>e</i> : 37	0	TW
2113e-1110e 2113f-1110f	7657.001	(17 1 7) \leftrightarrow (15 1 11)	Interpolyad anharmonic	31	0	TW
1004e-0001e	7664.811	(18 0 4) \leftrightarrow (18 0 3)	Intrapolyad anharmonic	13	0	TW
		(18 0 4) \leftrightarrow (17 1 10)	Interpolyad Coriolis	35	0	TW
2003e-1000e	7691.585	(16 0 7) \leftrightarrow (16 0 6)	Intrapolyad anharmonic	32	1	[4,5,8]
1223e-0220e 1223f-0220f	7710.775	(16 2 3) \leftrightarrow (15 1 13)	Interpolyad Coriolis	<i>e</i> :32	0	[4,5,8
				<i>f</i> : 32	0]

225
 226 *Notes*

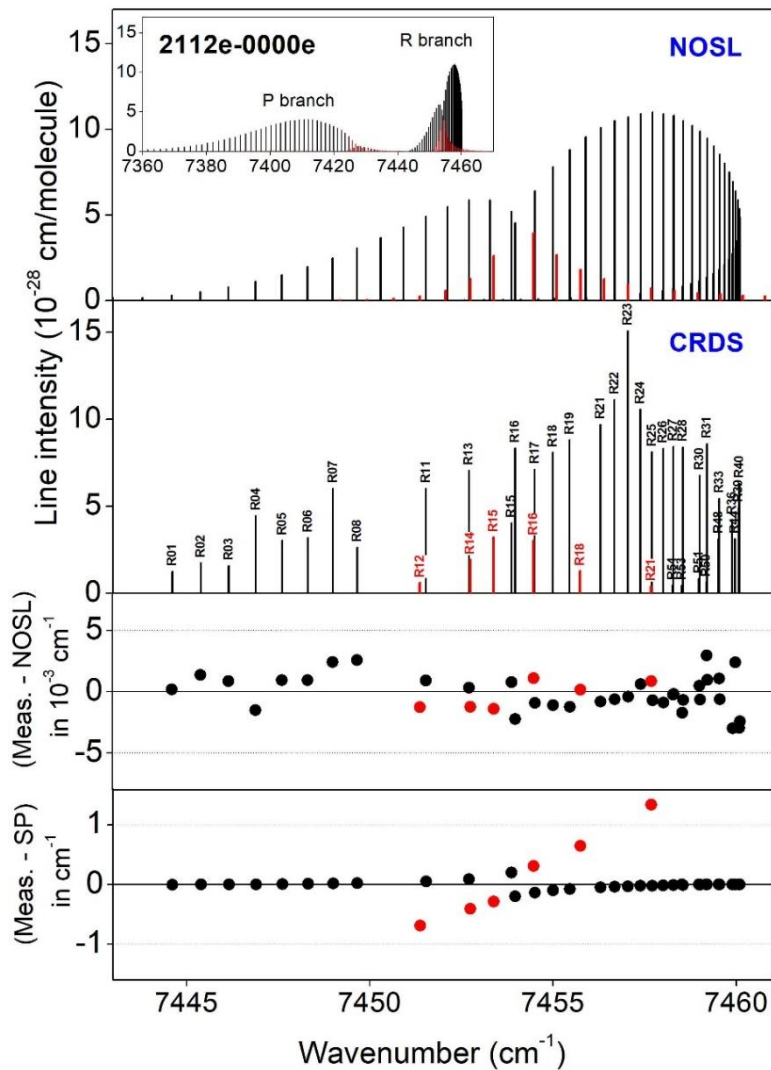
227 ^a Cluster labeling notation.

228 ^b First report of the considered perturbation. (TW: This work).

229

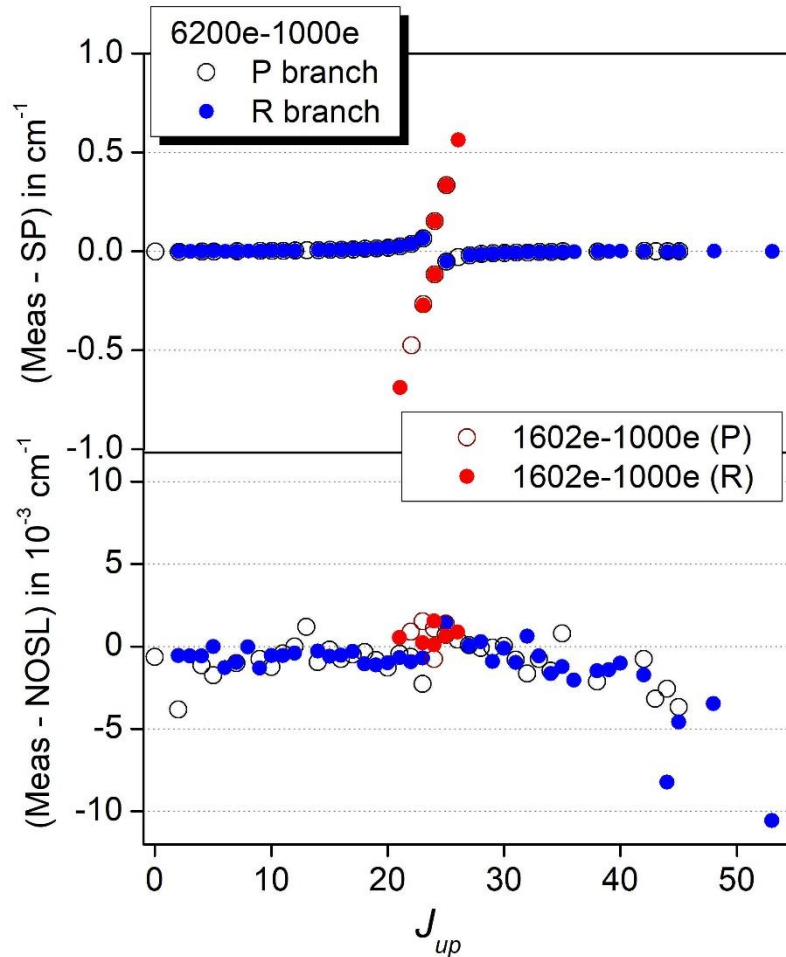
230 **Fig. 4** shows the effect of the anharmonic resonance interaction between the (13 1 7) and (13
 231 1 6) states on the rotational structure of the R branch of the 2112e-0000e band [the cluster labeling

232 notation of the upper state of this band is (13 1 7)]. The differences between the measured line posi-
 233 tions of the 2112e-0000e band and those provided by NOSL-296 [14] or calculated using the spectro-
 234 scopic parameters (SP) of **Table 2** are plotted on the lower panels. A total of six extra lines (R12, R14,
 235 R15, R16, R18, R21) belonging to the 0911e upper state [(13 1 6) in cluster labeling notation] could be
 236 identified around the energy crossing $J_{cross} = 17$. We also assigned three extra lines (Q16, Q17, and
 237 Q18) in the *Q* branch of the 2112f-0000e sub-band. The extra lines are provided in the Supplemen-
 238 tary Materials. The comparison of the two lowest panels (plotted with different Y-axis scales) illus-
 239 trates the quality of the NOSL-296 line positions in the present case of an intrapolyad interaction.
 240 The measured line positions (including extra lines) are reproduced within typically 10^{-3} cm^{-1} while the
 241 position perturbation of the R15 and R16 transitions of the bright 2112e-0000e band exceeds 0.1 cm^{-1}
 242 ¹.



243
 244 **Fig. 4.** The perturbed *R(J)* branch of 2112e-0000e band of $^{14}\text{N}_2^{16}\text{O}$ at 7443.008 cm^{-1} .
 245 *Upper panels:* Overview of the NOSL-296 and experimental stick spectra of the *R(J)* branch of 2112e-
 246 0000e band. The (13 1 7) upper state is perturbed by an anharmonic interaction with the (13 1 6)

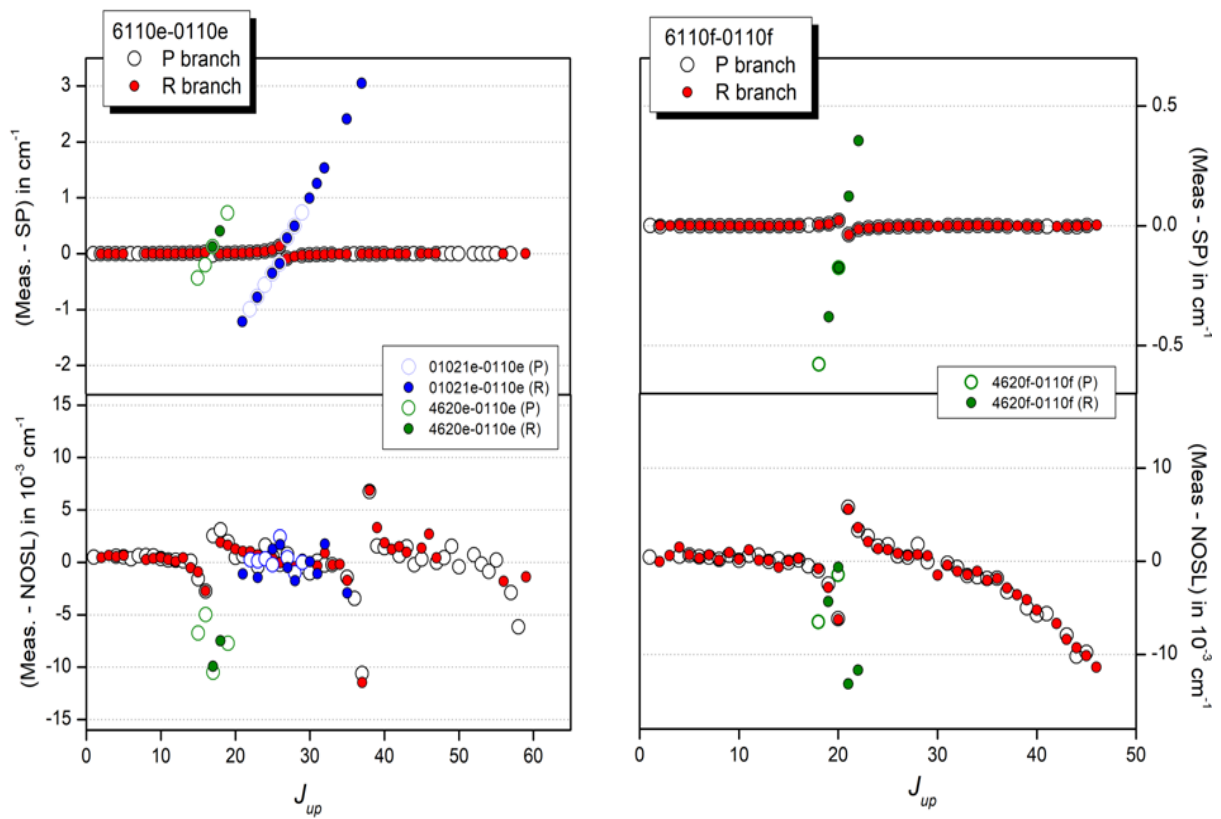
247 state (energy crossing at $J_{cross} = 17$). The extra lines (R12, R14, R15, R16, R18, R21) belong to the
 248 0911e-0000e band and are marked in red.
 249 *Lower panels:* Differences between the CRDS line positions and those provided by NOSL-296 [14] and
 250 the values calculated using the spectroscopic constants (SP) of **Table 2**.
 251 Note the different Y-axis scales of the two lower panels.
 252



253
 254 **Fig. 5.** Perturbation of the line positions of the 6200e-1000e band of $^{14}\text{N}_2^{16}\text{O}$ centered at 7525.861
 255 cm^{-1} . The residuals between the measured line positions and the values calculated using the spectro-
 256 scopic constants of **Table 2** and provided by NOSL-296 [14] are plotted *versus* the J rotational quan-
 257 tum number of the upper energy level. An interpolyad anharmonic resonance interaction between
 258 the (14 0 19) and (16 0 3) vibrational states is responsible of the observed position perturbations
 259 around the J crossing value of 24. A total of 11 extra lines belonging to the 1602e-1000e band could
 260 be detected (*P branch*: brown empty circles, *R branch*: red full circles).
 261

262 The 6200e-1000e band at 7525.861 cm^{-1} reaching the (14 0 19) upper state is perturbed by an
 263 anharmonic resonance interaction with the (16 0 3) perturber state (energy crossing at $J_{cross}= 24$).
 264 This interpolyad perturbation was firstly analyzed in Ref. [5]. Here, we could newly identify eleven
 265 extra lines (*P23-P26* and *R22-R25*) due to an intensity transfer towards the 1602e-1000e band. The
 266 residuals between the measured line positions of the 6200e-1000e band and those calculated by
 267 using the spectroscopic constants of **Table 2** and provided by NOSL-296 [14] are presented in **Fig. 5**.

268 On this figure, the deviations are plotted *versus* the rotational quantum number of the upper state of
 269 the transition, J_{up} . As a given upper state J_{up} level can be accessed through both P- and R-transitions,
 270 coincident deviation values for $P(J_{up}+1)$ and $R(J_{up}-1)$ insure that lower state combination difference
 271 relations are fulfilled. This agreement was particularly valuable to validate the assignment of the
 272 extra lines. As in the previous example, the NOSL predictions agree with the observations within
 273 about the experimental uncertainty on the line position (about $1 \times 10^{-3} \text{ cm}^{-1}$). Nevertheless, here the
 274 interaction mechanism involves states belonging to different polyads ($P=14$ and 16) and only recent-
 275 ly a non-polyad model of the effective Hamiltonian has been developed for $^{14}\text{N}_2^{16}\text{O}$ to account for
 276 interpolyad interactions [20]. This model is the one used to generate the NOSL-296 position values.



277
 278 **Fig. 6.** Differences between the measured line positions of the 6110-0110 hot band at 7570.895 cm^{-1}
 279 and the corresponding values calculated using the spectroscopic constants (SP) of **Table 2** (upper
 280 panels) and provided by NOSL-296 [14] (lower panels) for the *e* and *f* sublevels, respectively.

281 *Left hand:* The (13 1 15) upper state is perturbed by interpolyad Coriolis interactions with (i) (14 2 9)
 282 around $J_{cross} = 17$; (ii) (14 2 6) around $J_{cross} = 27$, and (iii) (14 0 11) around $J_{cross} = 37$.

283 *Right hand:* The (13 1 15) upper state is perturbed by interpolyad Coriolis interaction with (14 2 9)
 284 around $J_{cross} = 21$.

285
 286 The (13 1 15) state is the upper state of the 6110-0110 hot band at 7570.895 cm^{-1} which shows
 287 several local perturbations, all due to interpolyad Coriolis interactions with: (i) the (14 2 9) state (J_{cross}
 288 = 17 and 21 for *e* and *f* sublevels, respectively), (ii) the (14 2 6) state ($J_{cross} = 27$), and (iii) (14 0 11)
 289 state ($J_{cross} = 37$). Most of these perturbations were first evidenced in Ref. [5]. Here, we could newly

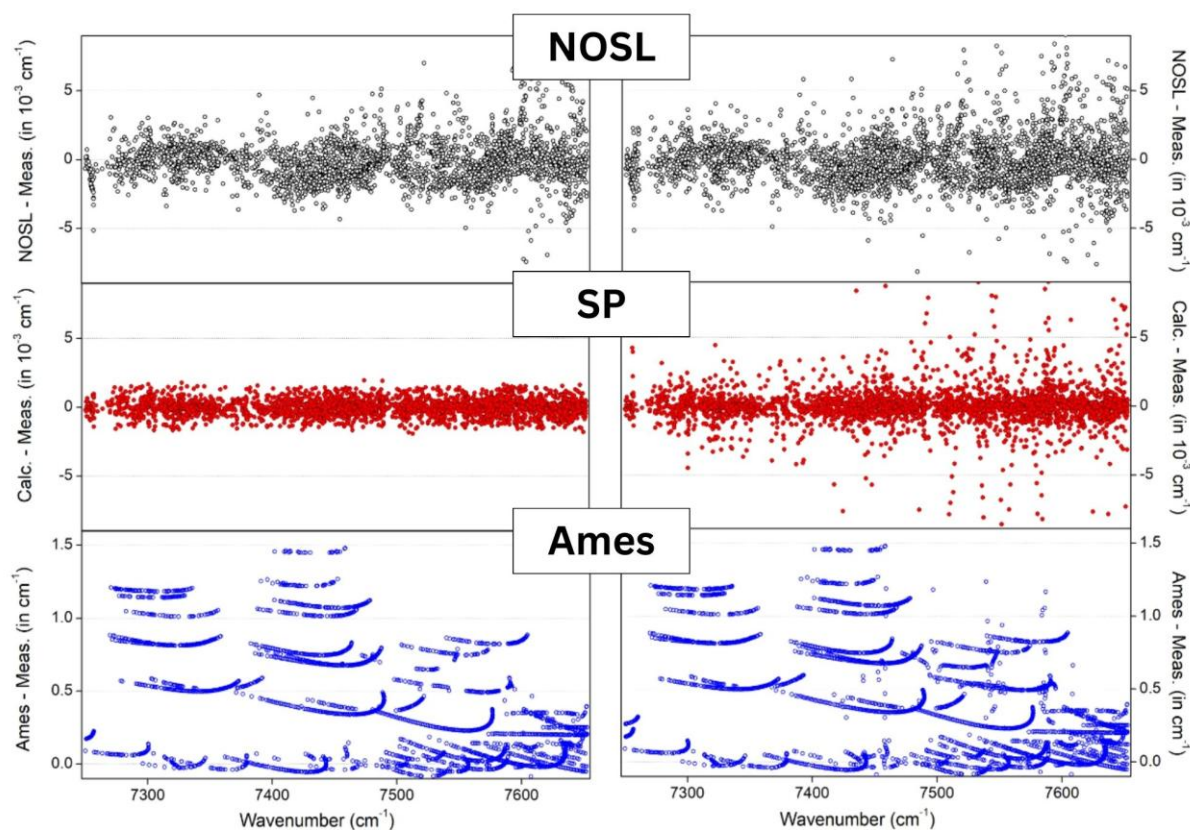
290 evidence the interpolyad Coriolis resonance around the energy crossing at $J = 37$ and a total of 31
291 extra lines were identified: twelve extra lines belonging to the 4620-0110 band [(14 2 9) upper state
292 in cluster labeling notation] could be assigned around the J crossing values 17 and 21 for e and f
293 sublevels, respectively and 19 extra lines were assigned around $J_{cross} = 27$. These lines are due to an
294 intensity transfer towards the 0(10)21-0110 band [(14 2 6) upper state in cluster labeling notation].
295 The residuals between the measured line positions of the 6110-0110 band and those calculated by
296 using the spectroscopic constants of **Table 2** and provided by NOSL-296 [14] are presented in **Fig. 6**
297 (e - e and f - f sub bands are presented on the left and right panels, respectively). The NOSL predictions
298 allow accounting satisfactorily of the different perturbations. Nevertheless, some significant residuals
299 remain near $J_{cross} = 17$ and 37 and $J_{cross} = 21$ for the e - e and f - f sub bands, respectively. It may indicate
300 that the EH parameters of the non-polyad model should be slightly refined to reproduce the line
301 positions within the experimental uncertainty.

302 **4. Comparison with the NOSL-296 and Ames calculated line lists**

303 4.1. $^{14}\text{N}_2^{16}\text{O}$

304 We present in **Fig. 7** an overview comparison for the measured $^{14}\text{N}_2^{16}\text{O}$ line positions. From top
305 to bottom, the experimental data are compared to (i) the NOSL-296 values [14], (ii) the values
306 calculated using the spectroscopic constants of **Table 2**, and (iii) the Ames-296 values [27]. **Fig.7**
307 presents two columns: in the right panels, all the position measurements are considered for the
308 different comparisons while in the left panels, the comparisons are limited to the lines whose
309 positions were included in the fit of the band-by-band spectroscopic parameters listed in **Table 2** (let
310 us call these lines “SP lines”, hereafter). Thus, the positions of perturbed or poorly determined lines
311 are excluded from the left panels. SP lines have their positions well reproduced by the band-by-band
312 fits of the SP parameters ($|\text{meas.} - \text{calc.}|$ position difference smaller than 10^{-3} cm^{-1}) and the
313 excluded lines are clearly evidenced by the comparison of the middle panels.

314 Let us recall that the NOSL and Ames calculated lists were obtained with very different
315 theoretical approaches. While the NOSL list is based on effective operators with parameters fitted
316 against measured line positions and line intensities, the Ames list is based on an empirical potential
317 energy surface (PES) and a pure *ab initio* dipole moment surface. In the following comparison, we will
318 use the Ames-296 line list available at <http://huang.seti.org> with file name n2o.6iso.296K.1E-
319 31.15Kcm-1.Y02-A8.dmsC.dat.

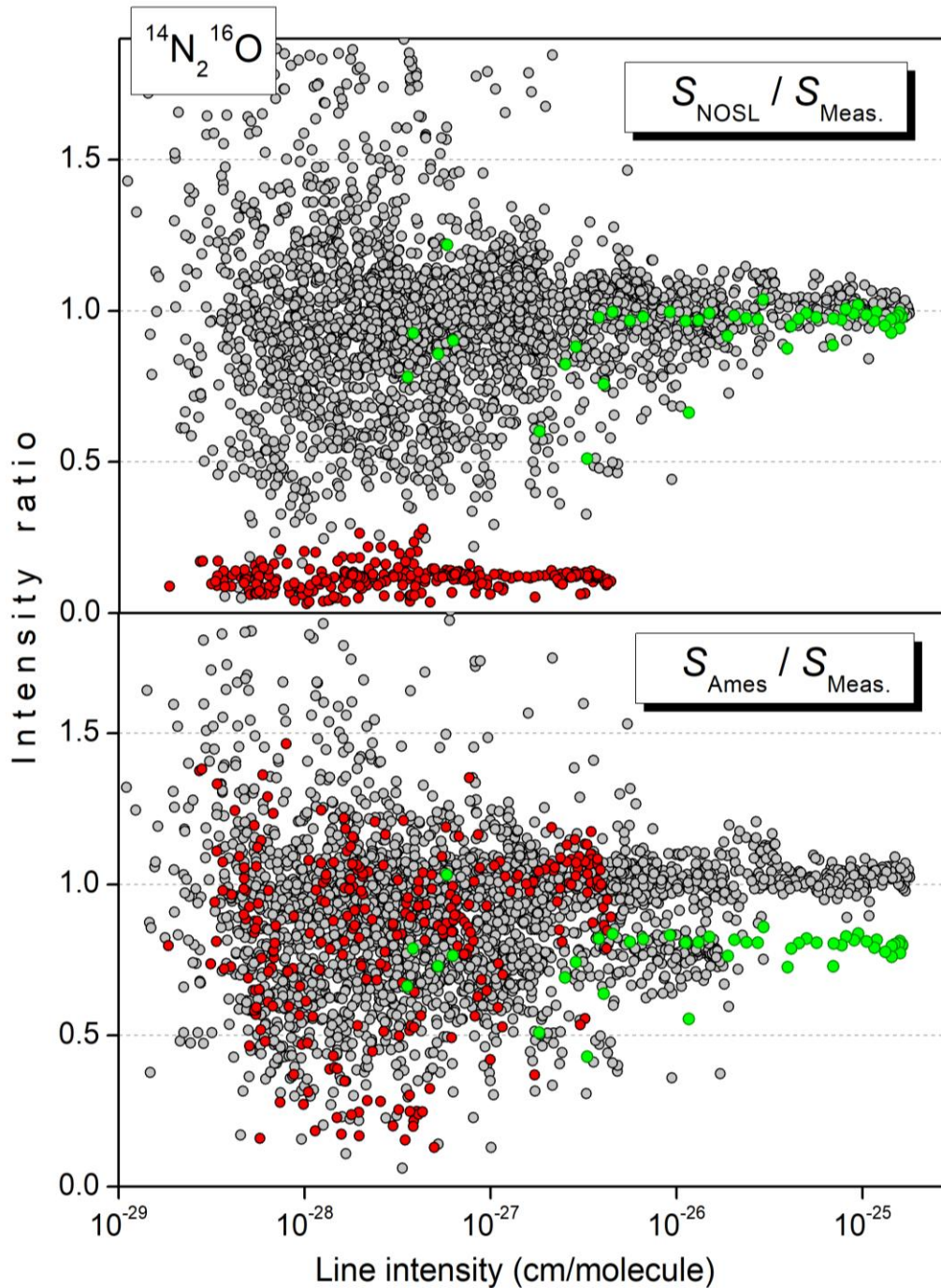


320
 321 **Fig. 7.** Position differences between the measured and the calculated line lists between 7250 and
 322 7653 cm^{-1} : (i) NOSL-296 [14] (upper panels), (ii) calculated using the band-by-band spectroscopic
 323 constants of Table 2 (SP) (middle panels), (iii) Ames-296 [27] (lower panels).
 324 *Right hand:* all the experimental lines are shown. *Left hand:* only the lines that are included in the fit
 325 of the spectroscopic constants are shown.
 326 Note the different Y-axis scales of the panels.

327
 328 As expected, the comparison of the lower and upper panels of **Fig. 7** illustrates the fact that
 329 the accuracy of the Ames positions cannot compete with the NOSL (and SP) position accuracy. The
 330 Ames deviations are mostly band dependent with a limited rotational dependence. On average,
 331 Ames term values of the upper vibrational states involved in the considered bands appear to be
 332 overestimated, by an amount up to 1.5 cm^{-1} . Although limited, the rotational dependence of the
 333 deviations is clear and systematic. The Ames rotational energies are overestimated by amounts
 334 increasing with J up to about 0.2 cm^{-1} .

335 As most of the presently measured line positions were previously reported in Ref. [5] and thus
 336 included in the transition dataset used to fit the EH parameters [20], the NOSL line positions are in
 337 close agreement to the observations even in case of interpolyad interaction (see above).
 338 Nevertheless, as concerns SP lines (for which SP calculated positions deviate by less than 10^{-3} cm^{-1}),
 339 the comparison of the two upper left panels of **Fig. 7** seems to indicate that the band-by-band SP

340 parameters allow for a slightly better reproduction of the measured positions than the (non-polyad)
 341 EH model, the NOSL position deviations of some SP lines near 7600 cm^{-1} being on the order of 5×10^{-3}
 342 cm^{-1} .



343
 344 **Fig. 8.** Variation of the ratio of the calculated and measured line intensities *versus* the measured in-
 345 tensity values of the $^{14}\text{N}_2\text{ }^{16}\text{O}$ transitions in the $7250\text{-}7653\text{ cm}^{-1}$ region. The 0203-0000 band is plotted
 346 with green full circles. Six $\Delta P = 14$ bands (0004-0200, 0004-1000, 0403-1000, 2112-0000 (Q branch),
 347 0114-0310, and 0114-1110) with strongly underestimated NOSL intensities are plotted with red full

348 circles. Upper and lower panels are relative to the NOSL-296 [14] and Ames [27] databases, respec-
349 tively.

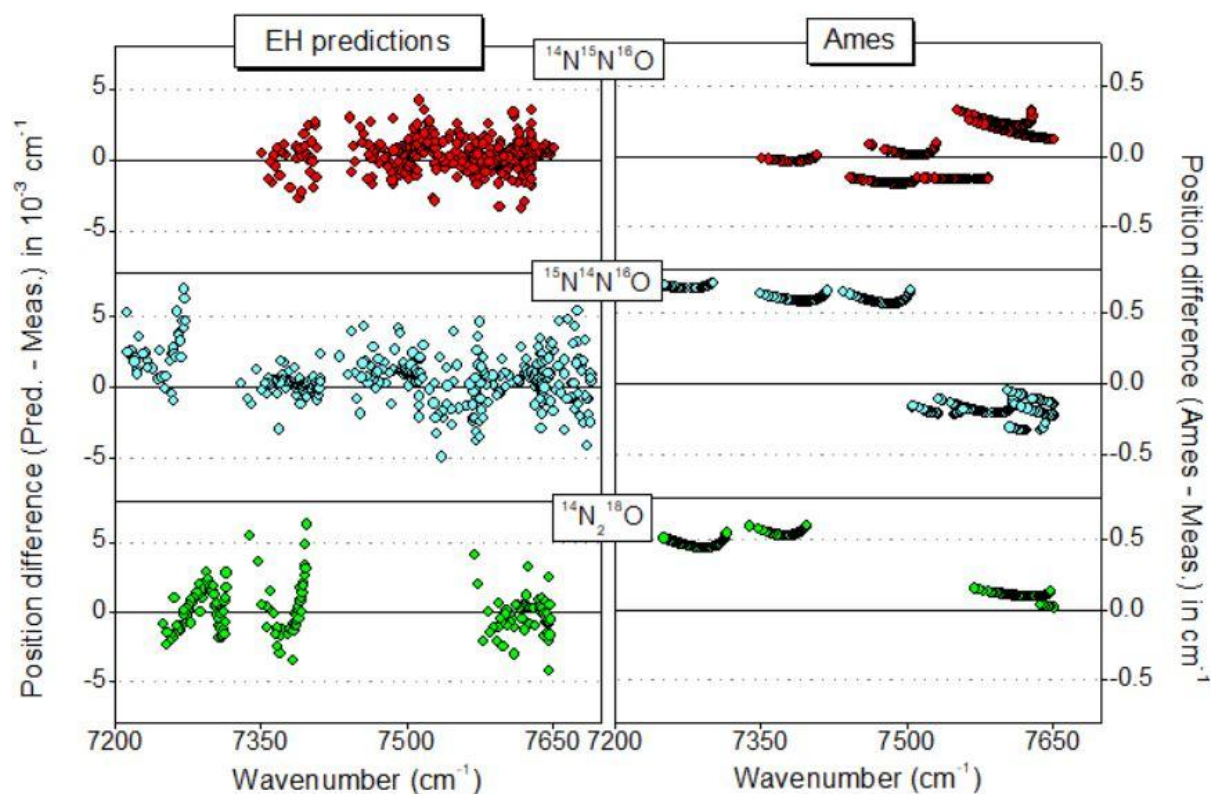
350

351 The comparison of the measured $^{14}\text{N}_2^{16}\text{O}$ line intensities to those provided by NOSL-296 [14]
352 and Ames-296 [27] line lists is presented in **Fig. 8**. The overall agreement is good for both calculated
353 lists but each of them shows important deviations for some specific bands. The most obvious prob-
354 lem concerns the NOSL intensity of a series of $\Delta P = 14$ bands, namely 0004-0200, 0004-1000, 0403-
355 1000, 2112-0000 (*Q* branch), 0114-0310, and 0114-1110 – which have their NOSL line intensities
356 underestimated by about one order of magnitude (red symbols in **Fig. 8**). The intensity of these
357 bands is mainly determined by the $M_{-1,0,4}^0$ EDM parameter which, due to the lack of intensity infor-
358 mation, could not be determined in the global fit of the EDM parameters used to generate NOSL
359 intensities. The present intensity measurements will allow correcting this issue. We note that, in
360 spite of a high dispersion of the intensity ratios, the Ames intensities do not show a similar problem
361 for the considered $\Delta P = 14$ bands (red dots in the lower panel of **Fig. 8**).

362 At the opposite, the NOSL line intensities of the 0203e-0000e band centered at 7665.273 cm^{-1}
363 agree with the measurements while the Ames intensities are systematically underestimated by about
364 20 % (green dots in **Fig. 8**). According to Dr. X. Huang from the Ames team, this band is “a sensitive
365 band which intensity may vary by 15-50% according to the dipole moment surface”. Note that the
366 agreement observed between the NOSL intensities and the measurements is not surprising because
367 this band is one of the bands for which intensity information was reported in Ref. [6] and used to
368 derive the EDM $\Delta P = 14$ parameters of the NOSL database.

369 4.2. $^{14}\text{N}^{15}\text{N}^{16}\text{O}$, $^{15}\text{N}^{14}\text{N}^{16}\text{O}$, and $^{14}\text{N}_2^{18}\text{O}$

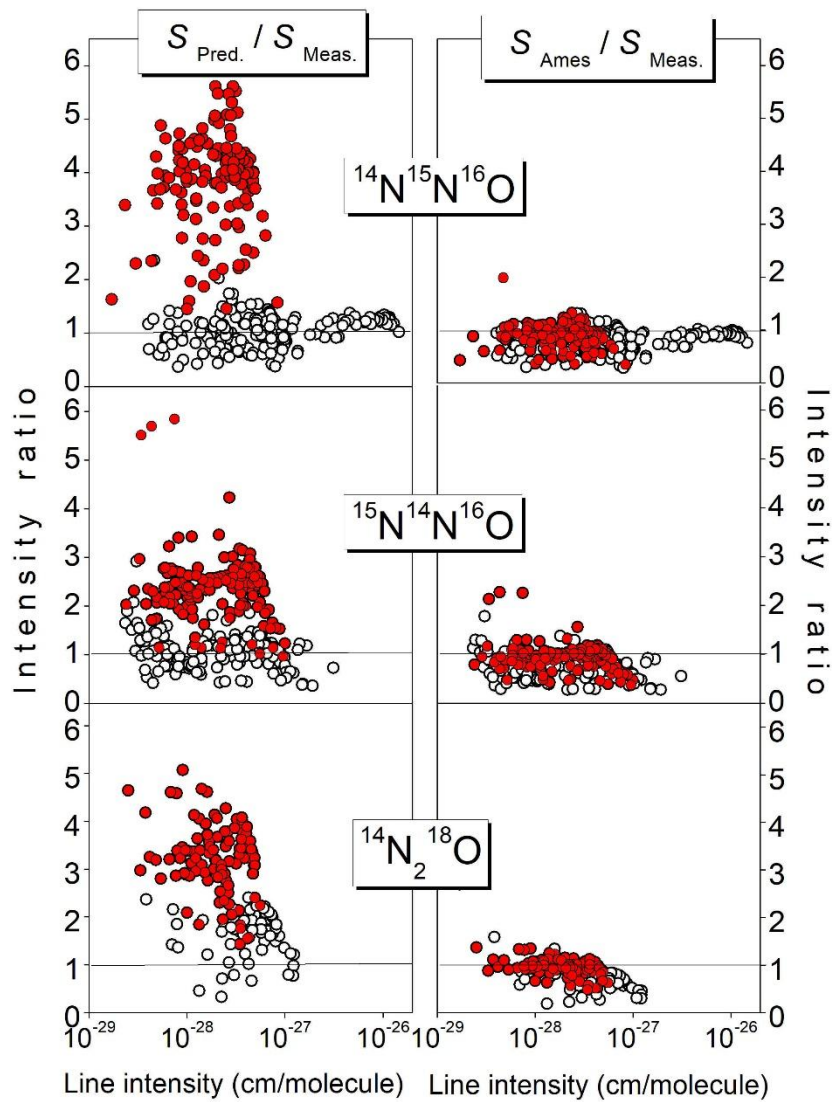
370 **Fig. 9** shows the overview comparison of the CRDS line positions for the $^{14}\text{N}^{15}\text{N}^{16}\text{O}$, $^{15}\text{N}^{14}\text{N}^{16}\text{O}$,
371 and $^{14}\text{N}_2^{18}\text{O}$ isotopologues to the effective Hamiltonian positions and the Ames-the NOSL list is lim-
372 ited to the main isotopologue and the origin of the EH parameters used for each of the considered
373 minor isotopologue has been detailed in section 3.1. The largest ($v_{EH} - v_{CRDS}$) deviations are about
374 $5 \times 10^{-3}\text{ cm}^{-1}$ for the three minor isotopologues whereas the Ames line positions show deviations with
375 similar general appearance as for $^{14}\text{N}_2^{16}\text{O}$ and an amplitude up to 0.7 cm^{-1} .



376
 377 **Fig. 9.** Differences between the line positions of the $^{14}\text{N}^{15}\text{N}^{16}\text{O}$, $^{15}\text{N}^{14}\text{N}^{16}\text{O}$, and $^{14}\text{N}_2^{18}\text{O}$ isotopologues
 378 assigned between 7250 and 7653 cm^{-1} and their values predicted by the effective Hamiltonian (left
 379 panels) and Ames database (right panels). Note the different Y-axis scales of the left-hand and right-
 380 hand panels.

381
 382 The ratios of the line intensities predicted by the effective operator model and provided in
 383 the Ames list to those measured in this work are presented in **Fig. 10**. Note that due to the low rela-
 384 tive abundance values (see **Table 1**), the bands observed for these minor species are all weak (inten-
 385 sity smaller than 2×10^{-26} $\text{cm}/\text{molecule}$). They belong to the $\Delta P = 12$ and $\Delta P = 14$ series. Overall, con-
 386 sidering the weakness of the considered lines, the agreement with the Ames intensity values is satis-
 387 factory for the three minor isotopologues. As concerns line intensities calculated using the effective
 388 operator approach, a satisfactory agreement is also observed for the intensities of the $\Delta P = 14$ bands
 389 but the line intensities of the 5200e-0000e and 6000e-0000e bands ($\Delta P = 12$) are largely overesti-
 390 mated (by a factor of 2 to 4). The line intensities of these bands were calculated using the corre-
 391 sponding dipole moment parameters of the main isotopologue, $^{14}\text{N}_2^{16}\text{O}$ [14]. As a result, the quality
 392 of predictions for these $\Delta P = 12$ bands is low. Although the Ames/meas. intensity ratios show signifi-
 393 cant dispersion (mostly due to experimental uncertainty related to the weakness of the considered
 394 lines), the overall agreement is satisfactory. This observation is of importance to insure that the iso-
 395 topic abundance in the used N_2O sample is close to its natural value. Indeed, while the chemical puri-
 396 ty of the sample is characterized, no information on the isotopic abundance is generally provided by

397 the gas supplier. The Ames intensity values allow validating the normal isotopic abundance which is
 398 generally assumed by default.



399
 400 **Fig. 10.** Ratios of the line intensities predicted by the effective operator model (left panels) and pro-
 401 vided by the Ames database (right panels) to the values measured for the $^{14}\text{N}^{15}\text{N}^{16}\text{O}$, $^{15}\text{N}^{14}\text{N}^{16}\text{O}$, and
 402 $^{14}\text{N}_2^{18}\text{O}$ isotopologues in the 7250 - 7653 cm^{-1} spectral region. All the bands belong to the $\Delta P = 14$
 403 series except for the 5200e-0000e and 6000e-0000e bands which belong to $\Delta P = 12$ series (red full
 404 circles).

405 **5. Conclusion**

406 The absorption spectrum of nitrous oxide has been reconsidered in the 7250 - 7653 cm^{-1} inter-
 407 val. The systematic retrieval of the line parameters using a multiline fit combined with improved pre-
 408 dictions of the N_2O absorption spectrum have allowed to identify eight additional bands not reported
 409 in Ref. [5], to identify or extend the analysis of rovibrational interaction affecting a few bands and to
 410 tests the NOSL-296 list of $^{14}\text{N}_2^{16}\text{O}$ and the Ames line list for all the isotopologues, in particular for line
 411 intensities.

412 The HITRAN database does not provide any line parameters in the region. The comparison val-
413 idates the quality of the NOSL list of $^{14}\text{N}_2^{16}\text{O}$ in the region. This can be partly explained by the fact
414 that line positions [5] and part of the line intensities [6] were already considered in the transition
415 database used to refine the effective operator model used to generate NOSL. The large set of 3328
416 intensities presently reported are overall in good agreement with NOSL intensities while only 741
417 lines were measured in Ref. [6], illustrating the good extrapolation capabilities of the effective opera-
418 tor approach. Nevertheless, six weak $\Delta P = 14$ bands have their NOSL line intensities underestimated
419 by about one order of magnitude compared to measurements because their intensities were not
420 previously measured and relevant $\Delta P = 14$ EDM parameters could not be determined. For the con-
421 sidered bands the Ames intensities are validated by the measurements. Unfortunately, while Ames
422 intensity calculations are found in general very good, the Ames intensities of the 0203-0000 band,
423 which is one of the strongest of the region, are underestimated by about 20 %. Ideally, the best da-
424 tabase for N_2O could be obtained by combining the advantages of the NOSL and Ames intensities. For
425 line positions, no doubt that the NOSL values should be preferred. As concerns line intensities, the
426 choice of the best source is less obvious and would require to systematically identify the NOSL bands
427 with unreliable intensities due to the lack of intensity formation and the Ames “sensitive” bands for
428 which intensity calculations may be less accurate. Let us not that the quality of the NOSL intensities
429 benefits from the large experimental information in the near infrared investigated by CRDS (see **Fig.**
430 **1**). The advantages of the Ames intensities against NOSL are expected to be more obvious at higher
431 energy where much less experimental data are available.

432 The above considerations relative to the main isotopologue, apply also to the minor species. In
433 case of experimental data too sparse to determine the EDM parameters of a ΔP series of transitions,
434 the Ames intensity values appear to be preferable.

435

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523
524**Table 2.** Spectroscopic parameters (in cm^{-1}) for the different bands of N_2O isotopologues assigned in the CRDS spectrum between 7250 and 7653 cm^{-1} .

Band ^a	$(P_l, i)^b$	ΔG_V^c	G_V	B_V	$D_V \times 10^7$	$H_V \times 10^{12}$	RMS^d	N_{fit}/N_{obs}^e	J_{max} $P/Q/R^f$	ΔP^g	Note ^h
¹⁴ N ₂ ¹⁶ O											
3421e-0220e	(14 2 12)	7237.06191(22)	8414.80658(22)	0.41121531(85)	1.3333(83)	-6.28(22)	0.65	75/92	P52/R50	12	1
3421f-0220f		7237.06060(23)	8414.80527(23)	0.41121686(81)	1.2218(77)	1.18(20)	0.70	86/105	P50/R51	12	1
2731e-0330e	(15 3 12)	7241.05404 ⁱ						/19	R40	12	
2732f-0330f		7241.05416 ⁱ						/19	R40	12	
3401e-0200e	(14 0 15)	7284.50398(12)	8452.63628(12)	0.40903512(65)	0.3789(79)	6.25(24)	0.54	93/110	P46/R48	12	1
0(14)00e-0200e	(14 0 16)	7307.59339(31)	8475.72569(31)	0.4154876(29)	8.218(67)	55.3(41)	0.66	37/39	P31/R32	12	
0(14)20e-0220e	(14 2 13)	7310.11337(39)	8487.85804(39)	0.4158003(28)	-2.999(48)	-69.2(22)	0.98	39/45	P40/R36	12	
0(14)20f-0220f		7310.11466(29)	8487.85933(29)	0.4158091(10)	3.0303(73)	-	0.73	30/37	P36/R38	12	
5310e-0110e	(13 1 13)	7324.35925(16)	7913.12712(16)	0.41380938(77)	3.1243(88)	7.19(27)	0.56	72/73	P48/R45	12	
5310f-0110f		7324.35825(16)	7913.12612(16)	0.41719121(34)	3.4113(13)	-	0.69	74/75	P54/R50	12	
0512e-0000e	(13 1 4)	7325.61195(29)	7325.61195(29)	0.41110450(95)	2.0622(57)	-	0.64	22/26	P28/R40	13	
0512f-0000e		7325.61265(27)	7325.61265(27)	0.41263612(89)	2.1256(53)	-	0.62	21/23	Q42	13	
6200e-1000e	(14 0 17)	7328.04505(29)	8612.94839(29)	0.4131284(14)	4.978(16)	19.33(51)	0.64	36/41	P44/R46	12	
5200e-0000e	(12 0 13)	7340.791714(98)	7340.791714(98)	0.41527767(37)	5.3702(31)	25.066(67)	0.45	94/97	P57/R56	12	
5220e-0000e	(12 2 10)	7355.03464(95)	7355.03464(95)	0.4157264(22)	-0.662(16)	-25.14(33)	0.76	39/46	P53/R55	12	
0114e-1110e	(17 1 1)	7365.81705(32)	9246.08279(32)	0.40553414(83)	1.7836(43)	-	0.61	21/33	P45/R40	14	
0114f-1110f		7365.81784(37)	9246.08358(37)	0.4062805(14)	1.8153(90)	-	0.80	17/24	P40/R39	14	
0004e-1000e	(16 0 1)	7429.23729(11)	8714.14063(11)	0.40518474(31)	1.7741(15)	-	0.53	81/85	P48/R47	14	
7110e-1110e	(15 1 18)	7430.65489(29)	9310.92063(29)	0.4098825(13)	2.266(11)	-	0.57	23/27	P36/R31	12	
7110f-1110f		7430.65576(46)	9310.92150(46)	0.4128955(22)	2.273(18)	-	0.74	14/16	P34/R33	12	
7110e-0310e	(15 1 17)	7436.63295(60)	9185.69818(60)	0.4115991(28)	2.897(26)	-	0.69	15/19	P32/R29	13	
7110f-0310f		7436.63275(73)	9185.69790(73)	0.4151885(29)	3.307(21)	-	0.72	9/11	P26/R34	13	
7000e-1000e	(14 0 18)	7440.19671(13)	8725.10005(13)	0.41049861(66)	3.0211(73)	8.74(21)	0.52	81/87	P49/R50	12	
2112e-0000e	(13 1 7)	7443.00813(34)	7443.00813(34)	0.4089097(16)	1.039(16)	-10.18(36)	0.65	18/53	P45/R53	13	2
2112f-0000e		7443.00360(86)	7443.00360(86)	0.4099529(18)	1.286(10)	-	0.67	8/26	Q42	13	
6200e-0200e	(14 0 17)	7444.81579(20)	8612.94809(20)	0.4131324(10)	5.035(13)	21.26(39)	0.60	56/60	P48/R45	12	
6220e-0220e	(14 2 14)	7453.38703(27)	8631.13170(27)	0.4138709(15)	-0.237(21)	-19.65(75)	0.75	49/51	P43/R43	12	
6220f-0220f		7453.38851(21)	8631.13318(21)	0.41386573(60)	2.2853(34)	-	0.75	48/58	P46/R41	12	
6220f-0220e								4/5	Q09	12	
0403e-1000e	(16 0 2)	7454.439982(54)	8739.30316(54)	0.4106908(28)	3.817(26)	-	0.95	17/20	P27/R32	14	

6110e-0110e	(13 1 14)	7457.58225(12)	8046.35012(12)	0.41192828(44)	2.4229(38)	3.262(87)	0.49	90/94	P57/R53	12	
6110e-0110f								4/9	Q16	12	
6110f-0110f		7457.58177(11)	8046.34964(11)	0.41473850(41)	2.5632(33)	2.439(67)	0.52	88/94	P59/R58	12	
6110f-0110e								7/7	Q21	12	
6000e-0000e	(12 0 14)	7463.98523(11)	7463.98523(11)	0.41279933(27)	3.2651(15)	7.080(22)	0.56	123/135	P71/R70	12	3
5220e-0000e	(12 2 11)	7488.62925(53)	7488.62925(53)	0.4140536(11)	0.8974(61)	-9.26(10)	0.65	54/65	P63/R61	12	
0114e-0310e	(17 1 1)	7497.01312(59)	9246.07835(59)	0.4055641(43)	2.146(53)	-	0.84	12/15	P27/R27	14	
0114f-0310f		7497.01677(64)	9246.08192(64)	0.4062965(38)	1.976(35)	-	0.82	14/18	P31/R30	14	
0533e-0330e	(17 3 3)	7522.09826(32)	9289.01065(32)	0.4114482(20)	1.516(25)	-	0.60	29/32	P28/R26	14	
0533f-0330f		7522.09834(31)	9289.01058(31)	0.4114513(20)	1.537(25)	-	0.59	29/32	P28/R26	14	
6200e-1000e	(14 0 19)	7525.86136(27)	8810.76470(27)	0.4094690(28)	0.627(25)	9.90(57)	0.78	24/87	P46/R52	12	4
0513e-0310e	(17 1 2)	7526.07994(44)	9275.14517(44)	0.4100991(20)	2.368(15)	-	0.70	18/24	P36/R26	14	
0513f-0310f		7526.07724(37)	9275.14239(37)	0.4121525(14)	2.737(11)	-	0.66	22/27	P33/R35	14	
0004e-0200e	(16 0 1)	7546.00813(17)	8714.14043(17)	0.40518536(51)	1.7769(28)	-	0.67	57/69	P45/R45	14	
6000e-0000e	(12 0 15)	7556.135163(88)	7556.135163(88)	0.41065810(31)	1.6107(25)	7.501(52)	0.42	113/114	P58/R57	12	
7000e-0200e	(14 0 18)	7556.96979(30)	8725.10209(30)	0.41048250(99)	2.7817(58)	-	0.89	37/45	P43/R35	12	
0423e-0220e	(16 2 1)	7568.84589(15)	8746.59056(15)	0.41087041(76)	0.1084(87)	-16.76(26)	0.59	58/63	P51/R40	14	
0423e-0220f								5/5	Q08	14	
0423f-0220f		7568.84639(13)	8746.59106(13)	0.41087048(36)	2.1331(17)	-	0.58	57/64	P50/R42	14	
0423f-0220e								5/5	Q08	14	
6110e-0110e	(13 1 15)	7570.89543(25)	8159.66330(25)	0.4106216(19)	1.566(14)	2.70(27)	0.78	31/119	P58/R58	12	5
6110f-0110f		7570.89374(16)	8159.66161(16)	0.4130766(14)	1.789(17)	34.81(55)	0.52	48/90	P46/R45	12	
0403e-0200e	(16 0 2)	7571.17235(18)	8739.30465(18)	0.41069182(81)	3.9995(84)	16.97(22)	0.62	65/75	P56/R52	14	
5420e-0220e	(14 2 15)	7582.34992(29)	8760.09459(29)	0.4125766(15)	1.274(20)	-10.17(66)	0.82	46/56	P36/R45	12	
5420e-0220f								3/3	Q06	12	
5420f-0220f		7582.34948(30)	8760.09415(30)	0.4125778(17)	1.632(22)	4.36(74)	0.84	39/50	P38/R45	12	
5420f-0220e								3/3	Q06	12	
0911e-0000e	(13 1 8)	7589.42023(80)	7589.42023(80)	0.4130956(48)	5.155(77)	8.19(35)	0.62	17/22	R36	13	
0911f-0000e		7589.4243(15)	7589.4243(15)	0.4155234(91)	5.29(10)	-	0.62	3/5	Q27	13	
1203e-1000e	(16 0 4)	7592.13856(29)	8877.04189(29)	0.4079641(12)	2.3436(98)	-	0.60	18/24	P26/R36	14	
3620e-0000e	(12 2 12)	7610.6169(28)	7610.6169(28)	0.4134111(56)	1.709(33)	-8.70(61)	0.57	20/25	P46/R53	12	
0313e-0110e	(15 1 1)	7616.50056(10)	8205.26843(10)	0.40960363(33)	2.0982(25)	1.129(48)	0.49	94/100	P63/R58	14	
0313e-0110f								10/10	Q13	14	
0313f-0110f		7616.500136(93)	8205.268006(93)	0.41102324(20)	2.19690(79)	-	0.52	86/92	P53/R55	14	

0313f-0110e								12/16	Q22	14	
4600e-1000e	(14 0 20)	7621.71108(30)	8906.61442(30)	0.4113021(13)	-1.267(11)	-	0.83	31/34	P33/R36	14	
0623e-0420e	(18 2 4)	7634.70063(47)	9965.82214(47)	0.4090960(25)	0.826(22)	-	0.67	14/17	P20/R31	14	
1443e-0440e	(18 4 12)	7637.13944(50)	9993.39186(50)	0.4098035(19)	1.653(15)	-	0.65	17/20	P27/R33	14	
3600e-0000e	(12 0 16)	7640.47386(11)	7640.47386(11)	0.41185011(39)	-0.0798(32)	8.863(69)	0.50	112/114	P56/R57	12	6
2113e-1110e	(17 1 7)	7657.00168(53)	9537.26742(53)	0.4055794(36)	1.711(20)	-	0.88	14/40	P42/R27	14	6,7
2113f-1110f		7657.00042(67)	9537.26416(67)	0.4066022(84)	1.12(22)	-	0.75	11/30	P38/23	14	
1004e-0001e	(18 0 4)	7664.81080(48)	9888.56757(48)	0.4033296(35)	1.461(57)	-10.4(19)	0.69	20/40	P47/R28	14	6,8
0203e-0000e	(14 0 1)	7665.273348(86)	7665.273348(86)	0.40978360(26)	2.5061(17)	3.448(31)	0.44	122/125	P64/R63	14	6
0223e-0000e	(14 2 1)	7673.63709(64)	7673.63709(64)	0.4099805(15)	1.1548(99)	-3.54(19)	0.76	56/63	P59/R52	14	6
1333e-0330e	(17 3 3)	7674.09722(19)	9441.00960(19)	0.40909436(43)	1.5884(18)	-	0.63	53/58	P51/R44	14	6
1333f-0330f		7674.09763(34)	9441.00987(34)	0.4090959(12)	1.625(12)	3.53(32)	0.65	45/48	P51/R41	14	6
3710e-0110e	(13 1 16)	7677.51544(20)	8266.28331(20)	0.41080657(98)	1.024(12)	6.42(36)	0.65	67/76	P49/R41	12	6
3710f-0110f		7677.51602(19)	8266.28389(19)	0.41338623(99)	0.371(13)	2.67(44)	0.64	71/78	P46/R44	12	6
2003e-1000e	(16 0 7)	7691.58474(18)	8976.48808(18)	0.40513676(94)	1.6902(75)	2.99(16)	0.73	57/98	P59/R51	14	6,9
1203e-0200e	(16 0 4)	7708.90881(12)	8877.04111(12)	0.40796932(44)	2.4140(35)	2.953(76)	0.53	94/101	P58/R56	14	6
1223e-0220e	(16 2 3)	7710.77485(15)	8888.51952(15)	0.40837749(72)	1.2477(45)	-2.095(75)	0.56	44/94	P53/R62	14	6,10
1223f-0220f		7710.77513(11)	8888.51980(11)	0.40837379(36)	1.79191(88)	-	0.42	41/98	P57/R62	14	
1113e-0110e	(15 1 3)	7747.02838(12)	8335.79625(12)	0.40719885(27)	1.7219(14)	0.222(19)	0.51	107/119	P74/R65	14	6
1113f-0110f		7747.02832(11)	8335.79619(11)	0.40804222(26)	1.7036(13)	0.353(18)	0.48	113/126	P73/R68	14	6
¹⁴N¹⁵N¹⁶O											
5200e-0000e	(12 0 14)	7384.07735(36)	7384.07735(36)	0.4129979(18)	4.404(22)	19.42(69)	0.70	26/33	P32/R46	12	
6000e-0000e	(12 0 15)	7492.12185(17)	7492.12185(17)	0.41043674(59)	2.3639(38)	-	0.67	56/63	P42/R41	12	
0203e-0000e	(14 0 1)	7511.63701(21)	7511.63701(21)	0.4101381(14)	2.842(20)	6.63(75)	0.60	43/52	P42/R43	14	
5200e-0000e	(12 0 16)	7567.42932(24)	7567.42932(24)	0.40922107(89)	0.6800(64)	-	0.72	37/43	P40/R35	12	
1113e-0110e	(15 1 3)	7614.99808(16)	8190.43173(16)	0.40740129(50)	1.7501(27)	-	0.58	46/54	P46/R43	14	
1113e-0110f								2/4	Q06	14	
1113f-0110f		7614.99849(19)	8190.43214(19)	0.40822660(58)	1.7286(32)	-	0.68	47/57	P46/R41	14	
1113f-0110e								2/2	Q03	14	
1003e-0000e	(14 0 3)	7650.753947(89)	7650.753947(89)	0.40710109(19)	1.72924(73)	-	0.48	93/100	P57/R51	14	6
¹⁵N¹⁴N¹⁶O											
5200e-0000e	(12 0 13)	7275.89874(28)	7275.89874(28)	0.4013148(20)	4.375(35)	6.16(16)	0.60	29/37	P28/R38	12	
6000e-0000e	(12 0 14)	7394.27113(14)	7394.27113(14)	0.39887412(75)	2.7847(88)	5.50(28)	0.51	60/63	P42/R47	12	
6000e-0000e	(12 0 15)	7482.85458(19)	7482.85458(19)	0.39745044(98)	1.075(12)	7.32(41)	0.53	54/65	P43/R45	12	
0313e-0110e	(15 1 1)	7543.00758(76)	8128.31970(76)	0.3956707(38)	1.776(34)	-	0.71	9/16	P33/R28	14	

0313f-0110f		7543.00305(64)	8128.31517(64)	0.3970185(28)	2.184(26)	-	0.73	16/20	P33/R26	14	
0203e-0000e	(14 0 1)	7592.48553(20)	7592.48553(20)	0.39584295(90)	2.2379(97)	2.18(28)	0.63	58/65	P48/R49	14	
1203e-0200e	(16 0 4)	7629.70073(67)	8789.67244(67)	0.3941395(29)	2.109(21)	-	0.69	11/16	P22/R33	14	
1113e-0110e	(15 1 3)	7667.80648(36)	8253.11860(36)	0.3934939(12)	1.5783(65)	-	0.82	28/45	P44/R27	14	6
1113f-0110f		7667.80474(35)	8253.11686(35)	0.3943209(26)	1.923(45)	14.1(21)	0.74	29/49	P40/R28	14	6
1003e-0000e	(14 0 3)	7702.49927(11)	7702.49927(11)	0.39310930(37)	1.5910(26)	0.660(48)	0.53	95/101	P64/R59	14	6
¹⁴ N ₂ ¹⁶ O											
6000e-0000e	(12 0 14)	7292.37397(14)	7292.37397(14)	0.38920233(72)	2.2164(85)	6.11(26)	0.47	59/63	P40/R47	12	
5200e-0000e	(12 0 15)	7376.57120(22)	7376.57120(22)	0.38938671(94)	0.1517(70)	-	0.69	39/42	P37/R37	12	
0203e-0000e	(14 0 1)	7630.84269(17)	7630.84269(17)	0.38684019(45)	2.1569(22)	-	0.66	55/63	P50/R40	14	
1003e-0000e	(14 0 3)	7718.94756(11)	7718.94756(11)	0.38434089(38)	1.4574(30)	0.536(61)	0.47	90/99	P57/R58	14	6

525 Notes

526 The confidence interval (1SD) is in the units of the last quoted digit.

527 ^a $V_1V_2I_2V_3$ correspond to the maximum value of the modulo of the expansion coefficients of the eigenfunction. V_2 is given between parenthesis when it is larger than 10. Bold characters are used for the newly assigned bands,

529 ^b Cluster labeling notation: $(P=2V_1+V_2+4V_3, I_2, i)$, i is the order number within the cluster increasing with the energy. We correct here a ten of erroneous values of the i ranking number given in Ref. [5],

531 ^c $G_v=G_v'-G_v''$ is the band center,

532 ^d Root Mean Square of the (Obs.-Calc.) differences of the position values (in 10^{-3} cm^{-1} unit),

533 ^e N_{fit} is the total number of transitions included in the fit; N_{obs} is the total number of measured transitions of the considered band,

534 ^f Observed branch with the maximum value of the total angular momentum quantum number of the input data,

535 ^g $P=P'-P''$, where P' and P'' are the upper and lower polyads.

536 ^h Notes

537 1 – Line positions from Ref. [5] were included in the fit;

538 2 – Anharmonic resonance (13 1 7) \leftrightarrow (13 1 6) ($J_{cross}=17$);

539 3 – Intrapolyad Coriolis resonance (12 0 14) \leftrightarrow (12 3 5) ($J_{cross}=36$);

540 4 – Interpolyad anharmonic (14 0 19) \leftrightarrow (16 0 3) ($J_{cross}=24$);

541 5 – Three interpolyad Coriolis resonances (13 1 15) \leftrightarrow (14 2 6) ($J_{cross}=27$); (13 1 15) \leftrightarrow (14 2 9) ($J_{cross}=e: 17$ and $f: 21$) and (13 1 15) \leftrightarrow (14 0 11) ($J_{cross}=37$);

542 6 – Line positions from Ref. [8] were included in the fit;

543 7 – Interpolyad anharmonic resonance (17 1 7) \leftrightarrow (15 1 11) ($J_{cross}=31$);

544 8 – Intrapolyad anharmonic resonance (18 0 4) \leftrightarrow (18 0 3) ($J_{cross}=13$) and Interpolyad Coriolis resonance (18 0 4) \leftrightarrow (17 1 10) ($J_{cross}=35$);

545 9 - Intrapolyad anharmonic (16 0 7) \leftrightarrow (16 0 6) ($J_{cross}=32$);

546 10 - Interpolyad Coriolis resonance (16 2 3) \leftrightarrow (15 1 13) ($J_{cross}=e: 32, J_{cross}=f: 32$);

547 ⁱ Band center provided by the global effective Hamiltonian.

