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Rapid detection of co-seismic ionospheric disturbances associated with the 2015 Illapel, the 2014 Iquique and the 2011 Sanriku-Oki earthquakes

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Key Points:

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8	•	A new methodology to rapidly estimate the time and the altitude of detection of
9		co-seismic ionospheric disturbances (CID),
10	•	The methodology uses data from near-epicenter seismic stations to calculate the
11		seismic peak time as an alternative to the earthquake onset,
12	•	First report on detection of CID as soon as 400 seconds after the earthquake on-
13		set and 250-430 seconds after the seismic peak time,

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14 Abstract

Co-seismic Ionospheric disturbances (CID, or "ionoquakes") are disturbances in the elec-15 tron density or total electron content (TEC) of the ionosphere, produced by the ground 16 motion due to earthquakes. Usually, ionoquakes are detected in the near-epicentral re-17 gion within 8-10 minutes after an earthquake onset time. In this work, we present a new 18 methodology that allows to estimate the CID arrival time based on determining the CID 19 peak time in TEC measurements with respect to the peak time of seismic waves regis-20 tered by the nearest seismic station. Our methodology also allows to understand the al-21 titude of GNSS detection that otherwise remains ambiguous. We apply the newly de-22 veloped techniques to detect CID signatures associated with three large earthquakes: the 23 2015 Illapel, the 2014 Iquique, and the 2011 Sanriku-Oki. We show that for these events, 24 the CID arrive 250-430 seconds after the time of the seismic wave peak, or 350-700 s af-25 ter the earthquake onset time. Our analysis show that the first CID are detected at the 26 altitudes of 150-180 km (the Sanriku earthquake) and of 200-300 km (the Illapel and the 27 Iquique earthquakes). The disturbances represent high-frequency acoustic oscillations 28 that propagate with a horizontal speed faster than 0.75 km/s. 29

30 1 Introduction

It is known that large earthquakes can generate acoustic and gravity waves that 31 further produce disturbances in the ionosphere (Calais & Minster, 1995; Heki & Ping, 32 2005; Liu et al., 2006; Lognonné, 2009; Astafyeva et al., 2009; Chum et al., 2012; Occhip-33 inti et al., 2013; Afraimovich, E. L. et al., 2013; Cahyadi & Heki, 2015; Thomas et al., 34 2018; Astafyeva & Shults, 2019; Sanchez et al., 2022; Bravo et al., 2022). Such distur-35 bances are referred to as Coseismic Ionospheric Disturbances (CIDs) or ionoquakes. They 36 usually propagate at horizontal speeds between 600 m/s and 3 km/s (Astafyeva et al., 37 2009; Rolland et al., 2011; Astafyeva et al., 2014), and their oscillation frequency lies in 38 the range 1-10 mHz (Bagiya et al., 2019; Manta et al., 2020; Sanchez et al., 2022). 39

The CIDs propagating at the acoustic wave speed (i.e., 600-1000 m/s) are usually
detected in the near-epicentral area only 8-10 min after an earthquake (Heki & Ping, 2005;
Kherani et al., 2012; Cahyadi & Heki, 2015; Thomas et al., 2018; Astafyeva & Shults,
2019). Such a short timing opens a possibility of use CIDs to enhance the capability of
early tsunami warning by using ionospheric data Occhipinti (2015); Astafyeva (2019, 2020).

In recent years, a few studies have reported early and rapid detection of CIDs, i.e. 45 less than the "nominal" 480-600 seconds after an earthquake. For instance, the first CIDs 46 due to the 2011 Tohoku-oki earthquake were detected as early as 420-464 seconds after 47 the earthquake onset time Astafyeva et al. (2011, 2013); Bagiya et al. (2020); Chum et 48 al. (2016). Early CID arrivals (440-480 sec) have also been reported for the Mw7.4 March 49 9, 2011 Sanriku-oki earthquake in Japan Thomas et al. (2018); Astafyeva and Shults (2019). 50 The phenomenon of early CID arrivals is not well understood yet. In the case of the Tohoku-51 oki earthquake, they were explained by the generation of large shock-waves Astafyeva 52 et al. (2011), and/or by low LOS elevation angles Astafyeva et al. (2013). 53

In this work, we present a new methodology allowing to 1) rapidly detect CID in total electron content (TEC) data times series and to estimate the CID arrival time; 2) estimate the altitude of ionospheric detection. We further apply this technique to analyze co-seismic ionospheric signatures due to the Mw8.3 Illapel earthquake of September 16, 2015, the Mw8.2 Iquique earthquake of April 01, 2014, and the Mw7.3 Sanrikuoki earthquake of March 09, 2011.

⁶⁰ 2 Earthquake events and Seismic and TEC Data

The Illapel and Iquique earthquakes of Mw8.3 and 8.2 respectively were triggered 61 near the Chile subduction zone, in South America. Seismicity in the Chilean region is 62 defined by the subduction of the oceanic Nazca plate under the South American plate. 63 This subduction zone is well known for harboring large EQs (Carrasco et al., 2019). The 64 Illapel earthquake occurred on September 16, 2015, at 22:54:32 UT. According to the 65 U.S. Geological Survey (USGS), this earthquake was generated by thrust faulting, with 66 the epicenter located at latitude = 31.57° S and longitude = 71.67° W, at a depth of 22.4 67 km (https://earthquake.usgs.gov/earthquakes/eventpage/us20003k7a/technical). The Iquique earthquake occurred on April 01, 2014, at 23:46:47 UT as a result of thrust 69 faulting, with the epicenter located at latitude = 19.61° S and longitude = 70.77° W, at 70 a depth of 25.5 km, (https://earthquake.usgs.gov/earthquakes/eventpage/us20003k7a/ 71 technical). The Sanriku-Oki earthquake of Mw7.4, occurred on March 9, 2011 at 02:45:20 72 UT at a depth of 32 km, with an epicenter at 38.435°N, 142.842°E (https://earthquake 73 .usgs.gov/earthquakes/eventpage/usp000hvhj/executive). 74

Figure 1 illustrates the epicentral location and the shake map of the peak ground
 velocity for the Illapel, Iquique, and Sanriku-Oki earthquakes (source: https://earthquake
 .usgs.gov/earthquakes). The present study employs the seismic data derived from seis mometers and TEC data derived from GNSS receivers. Figure 1 also illustrates the lo cations of seismic stations and GNSS receivers for the three earthquake events.

2.1 vTEC estimation from GNSS receivers

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For Illapel and Iquique earthquakes, we examine the TEC data with 15 seconds sampling rate, retrieved from GNSS receivers of the permanent ground-based network UN-AVCO (http://www.unavco.org) and CSN (http://gps.csn.uchile.cl/). For the analysis of the Sanriku-Oki earthquake, the original TEC data is with a 1-second sampling rate, retrieved from the GNSS Earth Observation Network System (GEONET). However, we lower the sampling rate from 1 second to 15 seconds in order to have identical data cadences for all events.

⁸⁸ During the earthquakes, multiple GNSS satellites were visible by ground-based GNSS ⁸⁹ receivers. However, here we focus on PRNs = 12 and 24 for the Illapel earthquake, PRNs ⁹⁰ = 01, 20, and 23 for the Iquique earthquake, and PRN = 07 for the Sanriku-Oki earth-⁹¹ quake. We have selected these PRNs since their projected locations at the ionospheric ⁹² heights i.e., the Sub-Ionospheric Points (SIP) are within ~450 km of the epicentral dis-⁹³ tance, and the elevation angle is more than 42° , except for the Sanriku-Oki earthquake, ⁹⁴ where the elevation angle was ~ 30° .

The slant TEC (sTEC) and vertical TEC (vTEC) are estimated by using the phase and code measurements from ground-based GNSS-receivers (Hofmann-Wellenhof et al., 2008), based on Equation 1,

$$sTEC = \frac{1}{40.308} \frac{f_1^2 f_2^2}{f_1^2 - f_2^2} (L_1 \lambda_1 - L_2 \lambda_2 + const + nL),$$
(1)

where $f_1=1575.42$ MHz and $f_2=1227.60$ MHz are the carrier-wave frequencies, $\lambda_1 = c/f_1$ and $\lambda_2 = c/f_2$ are the corresponding wavelengths in meters, c is the speed of light, L_1 and L_2 are the carrier phases, *const* is the unknown initial phase ambiguity, and nLis the error in determining the phase path. The TEC is measured in TEC units (TECU) with 1 TECU = $10^{16} el/m^2$. The exactitude of TEC estimation from phase measurements is about 0.01–0.02 TECU (Coster et al., 2013).

The vTEC is derived from sTEC using the following conversion Equation 2, (Klobuchar, 105 1987).

$$vTEC = sTEC.cos \left[arcsin\left(\frac{r_e}{r_e + H_{\rm ION}}cos\theta\right) \right],\tag{2}$$

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where r_e is the Earth's radius, H_{ION} is the altitude of the ionospheric thin layer and θ is the satellite elevation angle.

2.2 $V_{\rm SISM}$ estimation from seismometers

To obtain the vertical velocity (V_{SISM}) associated with the ground vibration from 109 the seismometers, we employ the Python library: obspyDMT. ObspyDMT is an open-110 source toolbox for querying, retrieving, processing, and managing seismological data sets 111 (Hosseini & Sigloch, 2017). The library downloads the data in count format, estimates 112 the ground vibration, and minimizes the instrumental response contributions associated 113 with the frequency response, amplifier, analog and digital filters, and digitization. We 114 used a bandpass filter on the seismograms before deconvolution, with a tuple defining 115 the four corner frequencies (0.02, 0.12, 10, 20) (Hosseini & Sigloch, 2017). 116

In the present study, V_{SISM} corresponds to the seismic station CO03 for the Illapel earthquake, PSGCX for the Iquique earthquake, and KSN for the Sanriku-Oki earthquake (Figure 1). To match the TEC sampling rate and to achieve identical spectral conditions between the two datasets, we reduce the seismic data sampling rate from 0.05 seconds to 15 seconds. The IRIS network (http://ds.iris.edu/wilber3/find_event) and National Research Institute for Earth Science and Disaster Resilience (NIED)-F-net (https:// www.fnet.bosai.go.jp/) administer the seismic data.

¹²⁴ 3 Methodology

In contrast to previous studies that rely on the onset time of an earthquake, here 125 we suggest to use the seismic peak time of the V_{SISM} for the estimation of the ionoquake 126 detection time ($t_{detection}$). The seismic peak time represents the time of the maximum 127 seismic oscillations rather than the beginning time of the oscillations. We note, however, 128 that the identification of the oscillation peak in the time series is comparatively unam-129 biguous. The analysis subjects the V_{SISM} and TEC data to the spectral analysis pro-130 cedure with identical spectral conditions, namely, the equal data length of 2 hours and 131 a sampling rate of 15 seconds. Then, we examine the spectrogram in a frequency range 132 of 0.13 mHz-33 mHz, and, for each frequency, we search for new oscillations in TEC start-133 ing from the corresponding peak onset time of V_{SISM} and in the vicinity of the epicen-134 ter. 135

We test two independent criteria to identify the peak time. In TEST-1, the peak time corresponds to the time of the first peak in the V_{SISM} and TEC oscillation. In TEST-2, the peak time corresponds to the time of the maximum in the V_{SISM} and TEC oscillations, i.e., corresponds to the time of the maximum amplitude. The peaks and the maxima are estimated by using the Python module "find-peaks". If the peak time of V_{SISM} and ionoquake are t_{SISM} and t_{TEC} respectively, then the ionoquake detection time can be defined as follows:

$$t_{detection} = t_{TEC} - t_{SISM} \tag{3}$$

Since both t_{SISM} and t_{TEC} suffer identical time shift from spectral analysis, the ionoquake detection time $t_{detection}$ remains unaffected by the time shift.

¹⁴⁵ 4 Results and Discussion

We first apply our methodology to the 2015 Illapel earthquake case. The time se-146 ries of the seismic V_{SISM} and the ionospheric TEC data are shown in Figure 2. The seis-147 mic data are from the seismic station CO03, which is the closest to the epicenter, and 148 TEC data corresponds to the SIP of Line-Of-Sight between the GPS station LSCH and 149 G12 satellite (Figure 1). The SIP of the trajectory of PRN 12 passes over the seismic 150 fault region and remains close to the epicenter (Figure 1). Figure 3 demonstrates the re-151 sults of TEST-1 and TEST-2 in the left and right panels, respectively. We note the fol-152 153 lowing characteristics:

154	- panels (A1, B1) reveal that the peak time of $V_{\rm SISM}$ depends on the frequency and
155	occurs between 22:54 UT and 23:03 UT,

- panels (A2, B2) reveal that the peak time of ionoquakes also depends on the frequency,
- panels (A3, B3) reveal that the ionoquake detection time ranges between 250-550 seconds from the peak time of V_{SISM} , depending on the frequency and testing criteria,
 - In both TEST-1 and TEST-2, the detection time of ionoquakes predominantly ranges between 250-400 seconds for the frequency range of 2-10 mHz,

Since both TEST-1 and TEST-2 confirm the ionoquake detection time range of 250-400
 seconds, it can be considered as the valid ionoquake detection time in this frequency range,
 invariant of the criteria.

To validate the methodology of the detection time estimation, we apply it to the 166 other SIP from various GPS stations within ± 400 km CO03-SIP distance that detected 167 ionoquakes, and we further examine the relationship between the detection time and the 168 CO03-SIP distance. Figure 4(A, B) demonstrates the relationship for the frequency of 169 3.7 mHz and a frequency range of 3.2-10 mHz, respectively. We note the expected in-170 crease of the detection time with the distance. According to Figure 4(B), the minimum 171 detection time is about 200 seconds at the distance of about -125 km, corresponding to 172 the CO03-SIP distance of LSCH-G12. Since detection time prolongs symmetrically across 173 the minimum ionoquake detection time location, the analysis identifies this detection as 174 the "earliest ionoquake". Noticeably, it is located very close to the epicenter and fault 175 region (denoted as red star and purple square in Figure 4). Therefore, our methodology 176 not only estimates the detection time vs. distance characteristics but also, locates the 177 earliest ionoquake close to the epicenter, based on the minimum detection time of the 178 ionoquake. 179

In supplementary Figure S1, we present time snapshots of the TEC map during -300-1400 seconds where 0 seconds correspond to the peak onset time of V_{SISM} at the frequency of 3.7 mHz. We note that the TEC response becomes noticeable approximately 240-300 seconds after the earthquake onset, in confirmation of findings in Figures 3-4.

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4.1 Possible subjectivities of the new methodology and their impacts on the ionoquake detection time

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4.1.1 Seismic station Vs. Epicenter

Figure 4 shows the ionoquake detection time estimated with reference to a particular seismic station: CO03, instead of the epicenter location. In order to examine any possible subjectivity arising from the location of a seismic station, we carry out the ionoquake detection time estimation for various seismic stations (Figure S2). We note that the ionoquake detection time remains within 250-400 seconds near the epicenter, independently on a choice of a seismic station. Therefore, the results of the present method ology are not subjective to a choice of seismic station.

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4.2 Detection altitude for rapid ionoquakes

Figure 4(A, B) reveals the potential of the present methodology to identify the epi-195 central distance of the earliest ionoquake close to the epicenter. However, the location 196 of CID/ionoquakes depends on the altitude (H_{ION}) of the thin ionospheric layer that is 197 assumed to be $H_{ION} = 300$ km in Figure 4(A, B). Consequently, the horizontal distribu-198 tion of ionoquakes will also depend on H_{ION}. To have an unambiguous ionoquake de-199 tection time, we examine the horizontal distribution of ionoquake detection time with 200 varying $H_{ION}=200, 250, 300, and 350 \text{ km}$ (Figure 5). Interestingly, in Figure 5(C), for 201 $H_{ION}=300$ km, the epicenter and the locations of minimum detection time are the clos-202 est. Therefore, the altitude region around 300 km is the most favorable altitude for the 203 detection of the ionoquakes. 204

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4.2.1 Onset time of an earthquake Vs. Peak onset time of $V_{\rm SISM}$

In previous studies, the onset time of an earthquake was used as a reference for the 206 seismic source time (Astafveva et al., 2011, 2013; Astafveva & Shults, 2019; Thomas et 207 al., 2018). However, strictly speaking, the earthquake onset time does not represent the 208 source time. The theory considers the co-seismic crustal uplift to be the source of CID. 209 Consequently, it is the time of the co-seismic uplift that should be taken as the source 210 time. However, seismic ruptures take time to propagate and cause crustal uplifts. For 211 large earthquakes that are characterized by large-dimension faults, the delay between 212 the earthquake onset and the maximum uplift can reach up to 3 minutes. For smaller 213 events, the delay of 10 to 20 seconds is usually observed. While this time is the most cor-214 rect to use, it cannot be calculated rapidly and without numerical modeling of seismic 215 faults. Besides, different seismological and seismo-geodetic techniques provide different 216 solutions. 217

Our approach suggests using the peak onset time of V_{SISM} that is calculated from seismic stations. The main advantage of our method is the independence on the seismological models, and also in the fact that it allows to calculate the CID arrivals very rapidly, i.e. potentially it can be used in near-real-time.

According to the USGS solutions, the Illapel earthquake onset time is 22:54:32 UT, and the maximum uplift occurred at 22:55:22 UT on the north-east from the epicenter. The peak onset time of V_{SISM} varies between 22:54 and 23:03 UT (Figure 3(A1, B1)), i.e., it is delayed between 32 seconds and 9 minutes from the USGS onset time, and 87 seconds to nearly 10 min from the seismic uplift time.

To examine the effects of the delay and other subjectivities arising from the usage of the seismic peak onset time rather than the earthquake onset time and the uplift time, we carry out the ionoquake detection time estimation with the following conventional definitions:

$$t1_{detection} = t_{TEC} - 22:54:32, \quad t2_{detection} = t_{TEC} - 22:55:25$$
 (4)

Here, t_{TEC} is the peak onset time of ionoquake is the same as defined for TEST-1 in Equation (3). The $t_{2detection}$ in Equation (4) the maximum uplift time. Figure 4(C, D) shows the results for the $t_{1detection}$ and $t_{2detection}$. One can see that within 50 to 200 km epicentral distance, both the $t_{1detection}$ and $t_{2detection}$ are less than 400 seconds.

Moreover, we note that in Equation (4), the t_{TEC} corresponds to the first peak of ionoquake, and it is subtracted from the earthquake onset time of 22:54:32. However,

in this scenario, the onset time of an ionoquake is more correct than the time of the first 237 peak of ionoquake. This will however decrease the $t1_{detection}$. For instance, at a frequency 238 of 4.3 mHz, the onset time of the ionoquake will occur about 116 seconds earlier than 239 the time of the peak of the ionoquake. Therefore, in Figure 4(C, D), the appropriate iono-240 quake detection time will shift by -116 seconds at the frequency of 4.3 mHz. Consequently, 241 in the vicinity of the epicenter the ionoquake detection time can be around 300 seconds, 242 which might seem too short knowing that the "nominal" propagation of CID is about 243 7-10 minutes. Such a short timing could be related to low elevation angles during the 244 detection of ionoquakes, that will lead to lower and much lower altitudes of detection, 245 or higher vertical and horizontal propagation speeds because of transformation of acous-246 tic waves into shock-acoustic waves due to non-linear effects. Below we discuss all these 247 possible explanations. 248

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4.3 Propagation speed and acoustic wave energetics

250 251 Relative to the location of the minimum detection time, the averaged propagation speed of the ionoquakes can be estimated as follows:

$$v = \frac{d - d_{\text{generation}}}{t - t_{\text{generation}}} \implies v1 = \frac{d - d_0}{t - t_0} \quad or \quad v2 = \frac{d - d_{eq}}{t - t_{eq}} \tag{5}$$

where (d, t) are the coordinates of the ionoquake in Figure 4(B), (d_{generation}, t_{generation}) 252 are the coordinates of the ionoquake at the time of generation, (d_0, t_0) are the coordi-253 nates of the ionoquake corresponding to the minimum detection time in Figure 4(B), and 254 (d_{eq}, t_{eq}) are the location of the epicenter and the onset time of the earthquake. The speed 255 v1 sets the upper limit for the actual speed since the ionoquakes are possibly generated 256 either at the minimum detection time or slightly earlier. The speed v^2 sets the lower limit 257 for the actual speed since ionoquakes are certainly generated after the onset time of earth-258 quake t_{eq} . Therefore, the actual speed resides in between the lower v2 and upper v1 lim-259 its. Figure S3 shows the distribution of v1 and v2 as a function of ionoquake detection 260 time and CO03-SIP distance. We note that they are in the range of 0.25-1.5 km/s such 261 that the early detected ionoquakes have predominantly large speeds. For instance, rapid 262 ionoquakes with a detection time of 250-400 seconds predominantly propagate faster than 263 0.75 km/s which is the acoustic speed range in the upper atmosphere. Previous stud-264 ies have found that acoustic-gravity waves resulting from ground vibration can efficiently 265 couple with the ionosphere and give rise to the ionoquakes (Rolland et al., 2013; Sanchez 266 et al., 2022). If a wave responsible for coupling propagates faster than 0.75 km/s, it ar-267 rives at 180 km or higher altitude in 250-400 seconds. Therefore, the detection time and 268 propagation speed of rapid ionoquakes suggest the altitude of detection of rapid iono-269 quakes to be above 180 km altitude, i.e., in the upper atmosphere where acoustic speed 270 is faster than 0.75 km/s. Therefore, the methodology of ionoquake detection time esti-271 mation and their propagation speed estimation validate each other. Moreover, it is cor-272 rect to say that the majority of rapid ionoquakes originate in the altitude range between 273 180 km and $H_{\rm ION}$ =300 km. Interestingly, simulation study by Chum et al. (2016) for 274 the Illapel earthquake finds significant air particle disturbance in the altitude range of 275 170-250 km, raising the possibility of the majority of ionoquakes to be in this altitude 276 range, as found in the present study. Also, the simulation study by Kherani et al. (2012) 277 demonstrated the acoustic-gravity wave with a vertical phase speed of more than 600 278 m/s to give rise to the coseismic TEC disturbances within (300) 360 seconds at the height 279 of 180 (250) km. 280

The lower limit v2 distribution in Figure S3 also attains the lowest of about 0.25 km/s for the early detected ionoquakes in the vicinity of the epicenter. This is due to the instantaneous generation $(t_{generation} = t_{eq})$ assumption in (5) which is not quite realistic. For the large earthquake of the 2011 Japan tsunami, the simulation of the Seismo-Atmosphere-Ionosphere (SAI) coupling revealed that the ionoquake can be developed within 360 seconds from the earthquake onset due to the fast-propagating Acoustic-Gravity wave energetics (Kherani et al., 2012). Therefore, the definition v1 in (5) and the corresponding distribution in Figure S3(A) represent a realistic scenario. Moreover, in the present study, the ionoquake detection time of 250-400 seconds can be associated with the SAI coupling mechanism energized by the acoustic-gravity wave.

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4.4 Ionoquake detection time during the Mw8.2 Iquique earthquake of 01 April, 2014

Figures 6-7 show our results for the Iquique earthquake of 01 April 2014. The temporal-293 spectral characteristics of V_{SISM} and TEC data in Figure 6(A1-A2) and the estimation 294 of $t_{detection}$ in Figure 6(A3) reveal the detection of ionoquakes starting from 430 seconds 295 after the peak onset time of V_{SISM} for the frequency range of 2 mHz-10 mHz. The con-296 ventional detection time $t_{1}_{detection}$, represented by yellow circles in Figure 6(A3) is in 297 between 600-700 seconds which suffers from the time shift effects, as discussed in sec-298 tion 4.1.1. The relationship between $t_{detection}$ and seismic-SIP distance in Figure 7 shows 299 the two-direction propagating ionoquakes with the earliest ionoquake location near the 300 epicenter for $H_{\rm ION}=200$ km. We note the detection of several rapid ionoquakes in 400-301 430 seconds within 200 km seismic-SIP distance, i.e., within about 250 km epicentral dis-302 tance. The detection of rapid ionoquakes in the case of the Iquique earthquake is slightly 303 delayed, in comparison to the much earlier detection (lower than 400 seconds) in the case 304 of the Illapel earthquake. This suggests that not all strong earthquakes produce rapid-305 ionoquakes detectable within 400 seconds from the peak onset time of the ground vibra-306 tion. 307

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4.5 Ionoquake detection time during the Mw7.4 Sanriku-Oki earthquake of March 9, 2011

We applied our newly developed method to seismic and TEC data around the epi-310 central area of the Sanriku-oki earthquake (Figures 8-9). The temporal-spectral char-311 acteristics of V_{SISM} and TEC data in Figure 8(A1-A2) and the estimation of $t_{detection}$ 312 in Figure 8(A3) reveal the detection of ionoquakes between 240-400 seconds for the fre-313 quency range of 2.5 mHz-10 mHz. The conventional detection time $t_{1}_{detection} = t_{TEC}$ -314 02: 45: 20, represented by yellow circles in Figure 8(A3) also below 400 seconds for 315 frequencies below 5 mHz. Figure 9 reveals the detection of rapid ionoquakes, as close as 316 100 km of the epicentral distance. The relationship between the $t_{detection}$ and the seismic-317 SIP distance shows two-direction propagating ionoquakes from the epicentral region for 318 H_{ION}=150 km-180 km. 319

Previously, Thomas et al. (2018) and Astafyeva and Shults (2019) reported the de-320 tection of the first ionoquakes at 430 seconds and 470-480 seconds respectively at the 321 altitudes 150 km and 180-190 km, respectively. Our methodology applied for the same 322 LOS as in Thomas et al. (2018) and Astafyeva and Shults (2019) shows quite similar re-323 sults: The conventional detection time $t1_{detection}$, represented by yellow-circles in Fig-324 ure S4(A3-B3) is 430-520 seconds. However, based on the $t_{detection}$, the earliest ionoquakes 325 are detected at about 320 to 460 seconds, as shown by the gray circles in Figure S4(A3-326 B3). Moreover, $H_{\rm ION}$ =150 km-180 km of the present methodology confirms the altitude 327 of the earliest ionoquakes reported in the previous studies of Thomas et al. (2018) and 328 Astafyeva and Shults (2019). Therefore, our methodology estimates the true earliest ar-329 rivals of the ionoquakes/CID, in addition to the altitude of detection. 330

Supplementary Figure S5(A) shows the results for conventional detection time $t1_{detection}$ in the same format as Figure 9(B). For comparison, Figure 9(B) is re-drawn as Figure S5(B). The conventional method detects the earliest ionoquakes 350-400 seconds after the earthquake onset time though they are few compared to the number of earliest ionoquakes detected from the newly developed method in Figure S5(B). Therefore, both conventional and the newly developed methods detect the earliest ionoquakes in less than 400 seconds from the earthquake onset time and from the time of the peak seismic up-lift.

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5 Rapid Ionoquakes from Seismo-Atmosphere-Ionosphere Coupling

Past simulation studies (e.g., Kherani et al. (2012, 2016)) have found that the SAI 340 coupling energized by the Acoustic-Gravity waves (AGWs) generates rapid ionoquakes 341 in the upper thermosphere within 360 seconds from the earthquake onset. The rapid iono-342 quakes of the present study are likely to be associated with this coupling mechanism. How-343 ever, the simulation study of Kherani et al. (2016) was for the case of a tsunami. For an 344 earthquake, no simulation study is available to support the early arrival within 360 sec-345 onds from the mainshock onset of the present study. The simulation study of Chum et 346 al. (2016) for the Illapel earthquake demonstrates the onset time of air particle distur-347 bances only 530 seconds after the earthquake at 800 km epicentral distance. Consequently, 348 in the vicinity of the epicenter, the disturbance should be detected several hundreds of 349 seconds earlier. 350

We examine this possibility by solving the governing equations of the ground uplift-351 Acoustic-Gravity wave-Ionospheric disturbances coupling mechanism, presented by Kherani 352 et al. (2012) for the Tohoku-Oki event and later employed by Sanchez et al. (2022) for 353 the Ridgecrest earthquake. In Figures 10-11, we demonstrate the simulation results for 354 the Illapel earthquake. We note in Figure 10 that the observed and simulated waveforms 355 of ionoquakes are fairly good in agreement. Moreover, in Figure 11, observed and sim-356 ulated ionoquake detection times are in the same time range of 260-400 seconds. There-357 fore, the physical mechanism responsible for the rapid ionoquakes Seismo-Atmosphere-358 Ionosphere coupling dynamics energized by the Acoustic-Gravity waves. 359

The supplementary Figure S6 demonstrates the vertical propagation of simulated 360 acoustic-gravity waves above the epicenter. We note that from the ground uplift, numer-361 ous waves with wavefronts of different slops i.e., of different phase speeds are launched 362 into the atmosphere. This is owing to the numerous scale heights and duct sizes present 363 in the atmosphere, that allow numerous wavelengths at a given frequency to be sustained 364 in the atmosphere. We note that the waves with significant amplitudes of about 10-20 365 m/s arrive at 160 km altitude at 22.975 hours from the mainshock onset time of 22.91 366 hours i.e., the waves arrive in about 240 seconds from the onset. These waves have wave-367 lengths comparable to the size of the longest atmospheric duct of about 150 km and at 368 the acoustic frequencies, they propagate with a phase speed of about 600 m/s or more. 369 Therefore, in the rapid development of ionoquakes, the long wavelength AGWs partic-370 ipate, as also found by Kherani et al. (2012). We note in Figure S6 that the though the 371 phase speed is about 600 m/s or more, the amplitude of the wave i.e., the fluid oscilla-372 tion is about 10-20 m/s, much slower than the average atmospheric sound speed. There-373 fore, such a fast propagating wave is not a shock acoustic wave, the wave which is char-374 acterized by a phase speed slower than the average atmospheric sound speed and an am-375 plitude comparable to the thermospheric sound speed (Zettergren et al., 2017). 376

6 Summary

We report early detections of co-seismic ionospheric disturbances (or ionoquakes) 378 associated with the 2015 Illapel, the 2014 Iquique, and the 2011 Sanriku-Oki earthquakes. 379 Using Total-Electron-Content (TEC) and seismic measurements, the study compares iono-380 quakes from our new and previous methods. The new method relies on applying the same 381 data processing procedures to the seismic and TEC data, and on estimating the time of 382 the peak of the seismic and TEC vibrations to obtain the ionoquake detection time. The 383 advantage of our method is its independence from the seismological models. The method 384 produces the expected prolongation in ionoquake detection time with increasing distance 385 from the epicenter. Moreover, the method allows locating the earliest detected ionoquakes 386

which turns out to be near the epicenter. The localization is more accurate for the thin 387 ionospheric laver centered around an altitude of 300 km for the Illapel, 200 km for the 388 Iquique, and 150-180 km altitude for the Sanriku-Oki earthquakes. The new method also 389 finds spectral and propagation characteristics of the earliest ionoquakes predominantly in the acoustic range. A comparative study with the conventional ionoquake detection 391 time method highlights a new result that the detection time of earliest ionoquakes is within 392 400 seconds from the earthquake onset time and from the time of peak seismic uplift, 393 for the Illapel and Sanriku-Oki earthquakes. For the Iquique earthquake, the new method 394 detects the ionoquakes as early as 430 seconds from the time of the peak seismic uplift. 395 Based on numerical simulation of the Seismo-Atmosphere-Ionosphere (SAI) coupling mech-396 anism, the study associates the rapid ionoquakes with the SAI coupling energized by the 397 acoustic-gravity waves. 398

³⁹⁹ Open research statement

The GNSS data for the earthquakes in Chile are available at the UNAVCO webservice (https://www.unavco.org/data/gps-gnss/data-access-methods/gnss-data -access-notebook/gnss-data-access-notebook.html) and CSN (http://gps.csn .uchile.cl/). The GNSS data for the 2011 Sanriku earthquake are available from the GeoSpatial Authority of Japan (GSI, terras.go.jp) via http://datahouse1.gsi.go.jp/ terras/terras_english.html). The seismic data can be downloaded from the IRIS (http:// ds.iris.edu/wilber3/find_event)

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Figure 1. Locations of the 2015 Illapel (A), the 2014 Iquique (B) and the 2011 Sanriku-oki (C) earthquakes. The epicenter of the earthquakes is represented by a red star and the blue triangles show the location of the seismic stations used in this study. The dots of different colors represent the SIPs at the time when the ionoquake occurred, each color corresponds to a particular PRN. The black squares depict the location of the GNSS stations. The colored contours represent the PGVs (Peak-Ground-Velocity) for each of the earthquakes. Beach ball shows the Global Centroid Moment Tensor



Figure 2. The 2015 Illapel earthquake: Vertical ground velocity recorded by seismic station CO03 (A), TEC time series recorded by the GNSS satellite PRN G12 and the LSCH ground receiver (B). Both the V_{SISM} and TEC time series have a resolution of 15 seconds.



Figure 3. The 2015 Illapel earthquake: Results of TEST-1 (left) and TEST-2 (right). Panels A1-B1 show the V_{SISM} time series recorded by the CO03 seismic station for different frequencies. Panels A2 and B2 show the Δ TEC time series obtained at different frequencies for the LSCH PRN G12 station. The filled circles in (A1,A2) represent the time of the peak in the first seismic oscillation and the first Δ TEC oscillation, respectively. The filled circles in (B1, B2) show the peak in the entire series of seismic and Δ TEC oscillations, respectively. The filled circles in panels (A3, B3) represent the detection time of ionoquakes, t_{detection}, derived from the definition (3).



Figure 4. The 2015 Illapel earthquake: Ionoquake detection time vs. distance diagram where distance is in between seismic station (C003) and Sub-Ionospheric-Point (SIP) of Line-of-sight (LOS) of LSCH-G12 pair. Panels (A-B) correspond to the new definition $t_{detection}$ in (3) for a frequency of 3.7 mHz, and a frequency range 3.2-10 mHz, respectively. Panels (C-D) correspond to the conventional definitions $t_{detection}$ and $t_{2detection}$ in (4), respectively for a frequency range 3.2-10 mHz. The distance of 0 km indicates the location of the CO03 seismic station. In (A-B), $t_{detection} = 0$ in the x-axis corresponds to the time of peak seismic uplift at each corresponding frequency. In (C-D), $t_{1detection} = 0$ in the x-axis corresponds to the time of earthquake onset.



Figure 5. The 2015 Illapel earthquake: $t_{detection}$ - distance diagram where distance is in between the seismic station (C003) and SIPs of LOSs of numerous GPS receiver-satellite pairs. The frequency range is 3.2-10 mHz. Four panels correspond to four values of altitude of detection (H_{ION}). The distance of 0 km indicates the location of the CO03 seismic station. $t_{detection} = 0$ in the x-axis corresponds to the time of peak seismic uplift at each corresponding frequency.



Figure 6. The 2014 Iquique earthquake with TEST-1. Panel A1 shows the V_{SISM} time series recorded by the PSGCX seismic station for different frequencies. Panel A2 shows the Δ TEC time series obtained at different frequencies for the TRTA PRN G20 station. In (A1-A2), the filled circles represent the time of the peak in the first seismic oscillation and the first Δ TEC oscillation, respectively. In panel A3, filled circles in grey and yellow correspond to the new definition, $t_{detection}$ in (3), and the conventional definition, $t_{1detection} = t_{TEC} - 23 : 46 : 47$, respectively.



Figure 7. The 2014 Iquique earthquake: $t_{detection}$ - distance diagram where distance is in between the seismic station (PSGCX) and SIPs of LOSs of numerous GPS receiver-satellite pairs. The frequency range is 3.2-10 mHz. Four panels correspond to four values of altitude of detection (H_{ION}). The distance of 0 km indicates the location of the PSGCX seismic station. $t_{detection} = 0$ in the x-axis corresponds to the time of peak seismic uplift at each corresponding frequency.



Figure 8. The 2011 Sanriku-Oki earthquake with TEST-1. Panel A1 shows the V_{SISM} time series recorded by the KSN seismic station for different frequencies. Panel A2 shows the ΔTEC time series obtained at different frequencies for the 0560 PRN G07 station. In (A1-A2), the filled circles represent the time of the peak in the first seismic oscillation and the first ΔTEC oscillation, respectively. In panel A3, filled circles in grey and yellow correspond to the new definition, t_{detection} in (3), and the conventional definition, t_{1detection} = t_{TEC} - 02 : 45 : 20, respectively.



Figure 9. The 2011 Sanriku-Oki earthquake: $t_{detection}$ - distance diagram where distance is in between the seismic station (KSN) and SIPs of LOSs of numerous GPS receiver-satellite pairs. The frequency range is 3.2-10 mHz. Four panels correspond to four values of altitude of detection (H_{ION}). The distance of 0 km indicates the location of the KSN seismic station. $t_{detection} = 0$ in the x-axis corresponds to the time of peak seismic uplift at each corresponding frequency.



Figure 10. Δ TEC comparison between observation (blue) and simulation (yellow), for the Illapel earthquake.



Figure 11. The 2015 Illapel earthquake. Comparison of Ionoquakes detection time between simulation and observation.