

SimPINNs: Simulation-Driven Physics-Informed Neural Networks for Enhanced Performance in Nonlinear Inverse Problems

Sidney Besnard, Frédéric Jurie, Jalal M. Fadili

▶ To cite this version:

Sidney Besnard, Frédéric Jurie, Jalal M. Fadili. SimPINNs: Simulation-Driven Physics-Informed Neural Networks for Enhanced Performance in Nonlinear Inverse Problems. 2023. hal-04216669v1

HAL Id: hal-04216669 https://hal.science/hal-04216669v1

Preprint submitted on 25 Sep 2023 (v1), last revised 2 Oct 2023 (v2)

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

SIMPINNS: SIMULATION-DRIVEN PHYSICS-INFORMED NEURAL NETWORKS FOR ENHANCED PERFORMANCE IN NONLINEAR INVERSE PROBLEMS

Sidney Besnard^{1,2}, Frederic Jurie¹, Jalal Fadili¹

¹Univ. Caen Normandie, ENSICAEN, CNRS

²Safran Data Systems

ABSTRACT

This paper introduces a novel approach to solve inverse problems by leveraging deep learning techniques. The objective is to infer unknown parameters that govern a physical system based on observed data. We focus on scenarios where the underlying forward model demonstrates pronounced nonlinear behaviour, and where the dimensionality of the unknown parameter space is substantially smaller than that of the observations. Our proposed method builds upon physics-informed neural networks (PINNs) trained with a hybrid loss function that combines observed data with simulated data generated by a known (approximate) physical model. Experimental results on an orbit restitution problem demonstrate that our approach surpasses the performance of standard PINNs, providing improved accuracy and robustness.

Index Terms— Inverse problems, Neural Networks, Physics-Informed, Simulation

1. INTRODUCTION

Inverse problems play a crucial role in science by allowing to unravel the hidden properties and processes behind observed data. They allow scientists to infer and understand phenomena that are otherwise difficult or impossible to observe or measure directly. These problems involve determining the parameters of a system from some available measurements. Inverse problems have far-reaching applications spanning a wide spectrum ranging from medical imaging to non-destructive control or astronomical imaging, as we will see.

Machine and deep learning have recently emerged as powerful alternatives to variational models for solving inverse problems. These methods include supervised and unsupervised methods, such as Deep Inverse Prior (DIP) [1], Unrolling [2, 3], Plug-and-play (PnP) [4, 5], and generative models [6]. For a review see [?]. Unrolling and PnP rely on neural networks to learn the regularization from the data.

However, these approaches only make sense if the output parameter space can be equipped with a suitable notion of regularity. This is certainly the case if the input parameters are in the form of a structured signal, but is not always the case as in our setting (think of inferring a few parameters that are not structured on a grid). A naive technique would

then be to train a neural network using a dataset consisting of input-output pairs, where the input is the observed data and the output is the sought-after vector of parameters [7, 8, 9]. Clearly, the neural network learns to invert the forward model (i.e. the mapping between the observed data and parameters), with the hope that it would predict the unknown parameters for new observations. This approach leverages the ability of neural networks to capture intricate patterns and non-linear relationships in the data. Unfortunately, this type of technique is only applicable when a large set of training pairs is available, which is barely the case in most practical situations. Moreover, such approaches are completely agnostic to the forward model which would produce unrealistic solutions and may not generalize well.

Physics-informed neural networks (PINNs) were primarily proposed to solve partial differential equations (PDE) [10, 11, 12, 13]. Their core idea is to supplement the neural network training with information stemming from the measurement formation model, e.g. the PDE model. In turn, this allows to restrict the space of solutions by enforcing the output of the trained neural network to comply with the physical model as described by the PDE. In turn, these methods are expected to be trained with a smaller dataset.

An aspect to consider regarding PINNs is their reliance solely on the reconstruction error, which represents constraints imposed by the partial differential equation (PDE). However, in numerous cases, achieving a low reconstruction error does not guarantee an accurate prediction of the parameters (i.e., a low parameter error). Thus, it is essential to emphasize the network's requirement for regularization of the solution space during training.

Contributions. In this paper, we propose a novel hybrid approach that combines both physics-informed and data-driven methods by using simulated data to regularise the solution space. Given the difficulty in obtaining real training pairs (observations, parameters) for many real-world problems, simulations offer a convenient means to provide the missing complementary information. Our aim is to demonstrate the effectiveness of neural networks in dealing with non-linear problems, while presenting a method that implicitly infers the forward operator within the neural network parameters by minimising the reconstruction error.

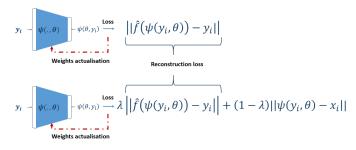


Fig. 1. Illustrations of the architectures used in PINNs and SimPINNs for inverse problems. The top part depicts the PINNs approach, where a neural network is trained to learn the inverse function of f, enabling the reconstruction of accurate x values from observed y values solely based on the observations and the underlying physics (unsupervised learning). In the bottom part, the SimPINNs (supervised) approach is shown, which utilizes the physics-based simulations to complement the training set with 'annotated' simulated data and regularize the solution.

2. PROPOSED METHOD

Let $\mathcal{X} \subset \mathbb{R}^n$ be the space of parameters of the (physical) model, and $\mathcal{Y} \subset \mathbb{R}^m$ be the space of observations. An inverse problem consists in reliably recovering the parameters $x \in \mathcal{X}$ from noisy indirect observations

$$y = f(x) + \epsilon, \tag{1}$$

where $f: \mathcal{X} \to \mathcal{Y}$ is the forward operator, and ϵ stands for some additive noise that captures the measurement and possibly the modeling error. Throughout, we assume that f is smooth enough (at least continuously differentiable).

In the sequel, for a neural network with parameters (weights and biases) $\theta \in \Theta$, $\psi : (y, \theta) \in \mathcal{Y} \times \Theta \mapsto \psi(y, \theta)$ denotes its output.

Finally, we also assume that we have an approximate explicit model of f, sometimes referred to as a digital twin. This model, \hat{f} , is obtained by modelling the physical phenomena involved in the observations. It will be used later to generate simulated data.

2.1. PINNs for non-linear inverse problems

The key idea of PINNs is to incorporate the physical model into the cost function during the training process. For a neural network ψ and training samples $\{y_i: i=1,\ldots,n\}$ with n samples, this amounts to solving the following minimization problem with the empirical loss:

$$\min_{\theta \in \Theta} \frac{1}{n} \sum_{i=1}^{n} ||y_i - \hat{f}(\psi(y_i, \theta))||^2$$
 (2)

This loss function leverages the information provided by the physical forward model directly into the training loss. It is also a non-supervised method that relies solely on observations, without any knowledge of the parameter vector x_i corresponding to each y_i . Unfortunately, as was observed previously in the literature (e.g. in [14]), when f is not iniective, there are infinitely many solutions $\psi(\cdot,\theta)$ which attain zero training error. This is because the forward model f may map multiple input vectors to the same output vector. For example, in the linear case, the action of f is invariant along its null space. This suggests that training a reconstruction network as (2) only from the observed data, without any additional assumptions or constraints, is not viable. Possible workarounds include explicitly constraining the output of the reconstruction network through regularization (and we are back to the variational world), or introducing invariances such as in [14]. In the forthcoming section, we will describe an alternative based on exploiting the forward model to simulate input-output pairs. This approach will help to regularize the training process and make it more robust to the non-injectivity of f.

2.2. SimPINNs: Simulation aided PINNs

In many areas of science, obtaining pairs of input (parameters)-output (observations) training data, can be a significant challenge. This can be due to various reasons, including that data are difficult or expensive to acquire. This is for instance the case in large instruments in physics. Furthermore, even if such pairs of data can be acquired, they are available only in limited quantity, which often impedes the use of data-intensive machine learning approaches.

There are however situations cases where even if such data are unavailable, it is possible to artificially generate pairs by generating the input-output pairs by leveraging knowledge of the forward model (Eq. (1)), even if the latter is only approximately known. This involves generating a parameter/input vector x sampled from the range of possible input values in the model or based on the known distribution of input data. We propose to compute the corresponding simulated observation $\hat{f}(x)$, i.e. without noise. It is important to note that the forward model serves anyway as an approximation of the underlying physical phenomenon f it represents, and the simulated observation can only be considered as a perturbed version of the unknown observation due to model imperfections.

Summarizing our discussion above, we propose to train a neural network ψ by replacing Eq. (2) with

$$\min_{\theta \in \Theta} \frac{1}{n} \sum_{i=1}^{n} \mathcal{L}_{\lambda}(x_i, y_i, \theta), \text{ where}$$
 (3)

 $\mathcal{L}_{\lambda}(x,y,\theta) = \lambda \|y - \hat{f}(\psi(y,\theta))\|^2 + (1-\lambda)\|\psi(y,\theta) - x\|^2.$ Here $\lambda \in]0,1[$ balances between the two terms: fidelity to the observation and reconstruction error. The determination of an

optimal value for λ can be a challenging task. In our study, we employed an empirical approach to estimate this value by performing cross-validation.

In the case of real observations, the value of x is unknown, making it impossible to calculate the second term in the loss function (Eq. (3)). Therefore, only the first term, which focuses on reconstruction fidelity, is utilized for such data.

The ratio between the number of real data, denoted as N_o (where only the observation is known), and the number of simulated data, denoted as N_s , plays a significant role in the analysis. Consequently, the influence of this ratio have to be thoroughly examined through experimental studies.

3. EXPERIMENTAL RESULTS

We validate the proposed method by applying it to an orbit restitution problem, where the dimensionality of \mathcal{X} (n=6) is relatively smaller than the dimensionality of \mathcal{Y} ($m=64^2$). This problem encompasses several intriguing aspects that make it particularly compelling. Firstly, the underlying physics and the involved forward operator exhibit nonlinearity, which is a primary focus in real-world research problems. Secondly, an orbit is defined by six orbital elements, while the received data exists in a significantly larger space, such as the image space in our case. Consequently, the forward operator maps from a smaller parameter space to a substantially larger image space. The third aspect of this problem pertains to the challenging nature of obtaining labelled data, as it requires integrating raw acquisition with non-trivial evaluations and determinations of orbit parameters.

It is important to note that despite initially appearing simple due to the presence of more equations than unknown variables, this problem poses additional difficulties. The involved physics operator is nonlinear, rendering the problem ill-posed with non-trivial equivalence classes. This complexity makes it more challenging than it may seem. In the subsequent sections, we present the details of the problem, including the experimental settings and the obtained results.

3.1. Problem and dataset

The objective is to derive an inverse operator for an orbit propagator using images obtained from a simulated sensor.

The forward operator performs the projection of orbits expressed in the Terrestrial Reference Frame (TRF) onto an image, similar to a ground track orbit projection. Consequently, the received data correspond to the projection of an orbit onto a sensor positioned on the Earth (see Figures 2 and 3).

For this problem, we opt to represent each orbit using 3 Keplerian parameters: inclination, eccentricity, and periapsis. This choice leads to a wide range of images and results, while avoiding trivial equivalence classes in the parameter set.

In this problem, the forward operator as follows:

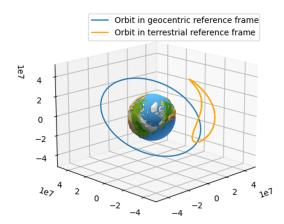


Fig. 2. Orbit Restitution: Visualization of an orbit in multiple reference frames. In this example, the depicted orbit has an inclination of 45 degrees, an eccentricity of 0.4, and an apogee of 42164×10^3 meters.

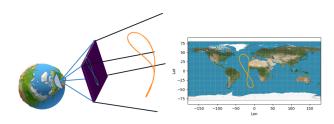


Fig. 3. Orbit restitution: Simulation of the orbit projection, as depicted in Figure 2. The left image exhibits the simulated sensor exhibiting the projected orbit, while the right image depicts the orbit projected onto the ground plane.

$$f : [0,1[\times[0,2\pi[\times[0,2\pi[\quad \to \quad \mathcal{I}_{64\times64}(\mathbb{R}) \quad e,i,\omega \quad \mapsto \quad .$$
 (4)

This operator simulates the entire system, encompassing the satellite position (achieved through an orbital propagator), the radiation pattern, and the projection to the final image (assumed to be 64×64 pixels). In this experiment, it is assumed that the approximated and actual physics models are identical, denoted as $\hat{f}=f$. To compute the satellite position while preserving the differentiability criterion of \hat{f} , an analytical Keplerian propagator is employed.

Finally, the neural network utilized for this problem consists of 5 dense layers, with each layer containing 784 neurons and employing a ReLU activation function. To assess the performance of the neural network, two previously defined loss functions are employed, as illustrated in Figure 1. Specifically, the Mean Squared Error (MSE) is utilized for both the reconstruction error and the parameter error evaluation.

$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	${\rm ArgPer}(\times 10^{-5})$		5 \	Number of real observations (N_o)							
$ \begin{array}{ c c c c c c }\hline 1000 & 2.0546 & 2.3412 & 1.1546 & 0.9731 & 1.1742 \\ \hline & 10000 & 1.1470 & 2.1942 & 0.8591 & 0.7854 & 1.0644 \\ \hline & 20000 & 0.8401 & 0.7721 & 0.6924 & 0.6571 & 0.6431 \\ \hline & 40000 & 0.4541 & 0.4412 & 0.4201 & 0.4121 & 0.4212 \\ \hline & & & & & & & & & & & & & & & & & &$, –	0		1000		10000	20000	40000	
$ \begin{array}{ c c c c c c c c } \hline >^{*} & 10000 & 1.1470 & 2.1942 & 0.8591 & 0.7854 & 1.0644 \\ \hline 20000 & 0.8401 & 0.7721 & 0.6924 & 0.6571 & 0.6431 \\ \hline 40000 & 0.4541 & 0.4412 & 0.4201 & 0.4121 & 0.4212 \\ \hline \hline Eccentricity(\times 10^{-5}) & Number of real observations (N_o) & 0 & 1000 & 10000 & 20000 & 40000 \\ \hline 0 & - & 1.5587 & 1.4424 & 1.1232 & 0.8638 \\ \hline 1000 & 1.3545 & 1.5225 & 0.7542 & 1.1036 & 0.8452 \\ \hline 20000 & 1.4684 & 1.0845 & 1.2781 & 0.6306 & 0.6251 \\ \hline 40000 & 0.3895 & 0.3856 & 0.3776 & 0.3435 & 0.2861 \\ \hline Inclination & Number of real observations (N_o) & 0 & 1000 & 10000 & 20000 & 40000 \\ \hline 0 & - & 0.0132 & 0.0021 & 0.0009 & 0.0004 \\ \hline 1000 & 0.0134 & 0.0114 & 0.0017 & 0.0012 & 0.0003 \\ \hline 20000 & 0.0014 & 0.0007 & 0.0007 & 0.0007 & 0.0005 \\ \hline 40000 & 0.0002 & 0.0002 & 0.0002 & 0.0001 & 0.0001 \\ \hline Reconstruction & Number of real observations (N_o) & 0 & 1000 & 10000 & 20000 & 40000 \\ \hline 0 & - & 0.3433 & 1.1667 & 0.8742 & 0.3023 \\ \hline 10000 & 1.4581 & 0.7064 & 0.1314 & 0.0989 & 0.0852 \\ \hline >^{*} & 10000 & 0.2467 & 0.1510 & 0.1394 & 0.0920 & 0.0955 \\ \hline \end{array}$		0		-		2.4541		1.2970	0.9947	0.9081	
$ \begin{array}{ c c c c c c c c } \hline 20000 & 0.8401 & 0.7721 & 0.6924 & 0.6571 & 0.6431 \\ \hline 40000 & 0.4541 & 0.4412 & 0.4201 & 0.4121 & 0.4212 \\ \hline \hline Eccentricity(\times 10^{-5}) & Number of real observations (N_o) & 0 & 1000 & 10000 & 20000 & 40000 \\ \hline \hline 0 & - & 1.5587 & 1.4424 & 1.1232 & 0.8638 & 10000 & 1.6556 & 1.5245 & 0.7054 & 1.1023 & 0.7214 & 20000 & 1.4684 & 1.0845 & 1.2781 & 0.6306 & 0.6251 &$	N_s	1000		2.0546		2.3412		1.1546	0.9731	1.1742	
$ \begin{array}{ c c c c c c c c } \hline 20000 & 0.8401 & 0.7721 & 0.6924 & 0.6571 & 0.6431 \\ \hline 40000 & 0.4541 & 0.4412 & 0.4201 & 0.4121 & 0.4212 \\ \hline \hline Eccentricity(\times 10^{-5}) & Number of real observations (N_o) & 0 & 1000 & 10000 & 20000 & 40000 \\ \hline \hline 0 & - & 1.5587 & 1.4424 & 1.1232 & 0.8638 & 10000 & 1.6556 & 1.5245 & 0.7054 & 1.1023 & 0.7214 & 20000 & 1.4684 & 1.0845 & 1.2781 & 0.6306 & 0.6251 &$		10000		1.1470		2.1942		0.8591	0.7854	1.0644	
		20000		0.8401		0.7721		0.6924	0.6571	0.6431	
Constructive Color Colo		40000		0.4541		0.4412		0.4201	0.4121	0.4212	
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	Eccentricity(>10-5		7-5	Number of real observations (N_o)							
$ \begin{array}{ c c c c c c c c c } \hline 1000 & 1.3545 & 1.5225 & 0.7542 & 1.1036 & 0.8452 \\ \hline & 10000 & 1.6556 & 1.5245 & 0.7054 & 1.1023 & 0.7214 \\ \hline & 20000 & 1.4684 & 1.0845 & 1.2781 & 0.6306 & 0.6251 \\ \hline & 40000 & 0.3895 & 0.3856 & 0.3776 & 0.3435 & 0.2861 \\ \hline \\ \hline & & & & & & & & & & & & & & & &$	Eccentricity(×10			0		1000		10000	20000	40000	
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$				-		1.5587		1.4424	1.1232	0.8638	
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	N_s	1000		1.3545		1.5225					
$ \begin{array}{ c c c c c c c c c }\hline & 40000 & 0.3895 & 0.3856 & 0.3776 & 0.3435 & 0.2861\\\hline \hline & & Number of real observations (N_o)\\\hline & & & & 1000 & 10000 & 20000 & 40000\\\hline & & & & & 1000 & 10000 & 20000 & 40000\\\hline & & & & & & 1000 & 10001 & 0.0009 & 0.0004\\\hline & & & & & 1000 & 0.0134 & 0.0114 & 0.0017 & 0.0012 & 0.0003\\\hline & & & & & 10000 & 0.0022 & 0.0028 & 0.0024 & 0.0008 & 0.0007\\\hline & & & & & & 10000 & 0.0002 & 0.0002 & 0.0001 & 0.0005\\\hline & & & & & & & & & & & & & & & \\\hline & & & &$				1.6556		1.5245		0.7054	1.1023	0.7214	
				1.4684		1.0845					
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$		40000		0.3895		0.3856		0.3776	0.3435	0.2861	
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	Inclination			Number of real observations (N_o)							
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$				0		1000		10000	20000	40000	
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$	N_s	0		-		0.0132		0.0021	0.0009	0.0004	
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$		1000	0.	0.0134		0.0114		0.0017	0.0012	0.0003	
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$		10000	0.	0.0022		0.0028		0.0024	0.0008	0.0007	
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$		20000	0.	0.0014		0.0007		0.0007	0.0007	0.0005	
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$		40000	0.	0.0002		0.0002		0.0002	0.0001	0.0001	
$ \begin{array}{ c c c c c c c c c c c c c c c c c c c$				Number of real observations (N_o)							
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	Reconstruction			0		1000		10000	20000	40000	
≥° 10000 0.2467 0.1510 0.1394 0.0920 0.0955	N_s	0		-	1.3243			1.1667	0.8742	0.3023	
2 ^{\infty} \frac{10000}{20000} \frac{0.2467}{0.0966} \frac{0.1510}{0.0831} \frac{0.1394}{0.0968} \frac{0.0920}{0.0750} \frac{0.0955}{0.0701}		1000	1.4581		0.7064			0.1314	0.0989	0.0852	
20000 0.0966 0.0831 0.0968 0.0750 0.0701		10000	(0.2467		0.1510		0.1394	0.0920	0.0955	
		20000	(0.0966		0.0831		0.0968	0.0750	0.0701	
40000 0.0304 0.0305 0.0298 0.0271 0.0251		40000	(0.0304		0.0305		0.0298	0.0271	0.0251	

Table 1. Error evaluation across various parameters, using a test set comprising solely of observations.

3.2. Results and analysis

As shown in Table 1, the SimPINNs method effectively utilizes and leverages information from both real observations and generated data. It outperforms the PINNs method in terms of parameter reconstruction; PINNs is the row with $N_s=0$. The most favorable results are achieved when employing the largest number of real and generated data points $(N_o=N_s=40k)$, with an average improvement of a factor of 2 for the parameter error compared to the unsupervised PINNs method. Additionally, the SimPINNs method demonstrates a significant advantage, reducing the reconstruction error by a factor of 10.

The impact of incorporating orbital physics in the forward operator becomes evident in the significant benefit it provides for image reconstruction in this particular use case. Due to the influence of orbital dynamics, even small changes in the parameters can have a substantial impact on the satellite's orbit and drastically alter the resulting projection on the observed image. As a result, two images with nearly identical parameters can exhibit significant differences. In this context, the reconstruction loss plays a crucial role in assisting the network in handling these high-gradient values that may not be adequately captured by the supervised loss alone.

By combining both the generated and real data, the model can benefit from regularization through \mathcal{L}_{λ} applied to the real dataset. This approach allows the model to leverage

the strengths of both data types and converge towards an improved solution with reduced parameter error.

Figure 4 presents a selection of projected orbits along with their corresponding reconstructions.

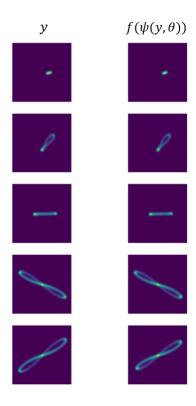


Fig. 4. Reconstruction Results using the SimPINNs Method.

4. CONCLUSIONS

This article explores an NN approach to solve non-linear inverse problems, focusing on the promising SimPINNs method. SimPINNs leverages the physics model to generate labeled training data and effectively regularize the neural network. The study demonstrates that simulation-aided training provides more information compared to conventional PINNs or vanilla neural network training. By utilizing the approximated physics operator, the model achieves improved learning and generalization over non-labeled datasets. SimPINNs shows potential for addressing challenging inverse problems with limited labeled data, offering insights into training neural networks with physics-based knowledge.

Acknowledgements

The research presented in this paper is, in part, funded by the French National Research Agency (ANR) through the grant ANR-19-CHIA-0017-01-DEEP-VISION.

5. REFERENCES

- [1] Dmitry Ulyanov, Andrea Vedaldi, and Victor Lempitsky, "Deep image prior," <u>International Journal of Computer Vision</u>, vol. 128, no. 7, pp. 1867–1888, mar 2020.
- [2] Pei Peng, Shirin Jalali, and Xin Yuan, "Solving inverse problems via auto-encoders," <u>IEEE J. Sel. Areas Inf. Theory</u>, vol. 1, no. 1, pp. 312–323, 2020.
- [3] Karol Gregor and Yann LeCun, "Learning fast approximations of sparse coding," in <u>ICML</u>. 2010, pp. 399–406, Omnipress.
- [4] Singanallur V. Venkatakrishnan, Charles A. Bouman, and Brendt Wohlberg, "Plug-and-play priors for model based reconstruction," in 2013 IEEE Global Conference on Signal and Information Processing, 2013, pp. 945– 948.
- [5] Zihui Wu, Yu Sun, Alex Matlock, Jiaming Liu, Lei Tian, and Ulugbek S. Kamilov, "SIMBA: Scalable inversion in optical tomography using deep denoising priors," <u>IEEE Journal of Selected Topics in Signal Processing</u>, vol. 14, no. 6, pp. 1163–1175, oct 2020.
- [6] Ashish Bora, Ajil Jalal, Eric Price, and Alexandros G. Dimakis, "Compressed sensing using generative models," in Proceedings of the 34th International Conference on Machine Learning, ICML 2017, Sydney, NSW, Australia, 6-11 August 2017, Doina Precup and Yee Whye Teh, Eds. 2017, vol. 70 of Proceedings of Machine Learning Research, pp. 537–546, PMLR.
- [7] Christopher A. Metzler and Gordon Wetzstein, "Deep s3pr: Simultaneous source separation and phase retrieval using deep generative models," in <u>ICASSP 2021 2021 IEEE International Conference on Acoustics, Speech and Signal Processing (ICASSP)</u>, 2021, pp. 1370–1374.
- [8] Christopher Metzler, Phillip Schniter, Ashok Veeraraghavan, and Richard Baraniuk, "prDeep: Robust phase retrieval with a flexible deep network," in Proceedings of the 35th International Conference on Machine Learning, Jennifer Dy and Andreas Krause, Eds. 10–15 Jul 2018, vol. 80 of Proceedings of Machine Learning Research, pp. 3501–3510, PMLR.
- [9] Yair Rivenson, Yibo Zhang, Harun Günaydın, Da Teng, and Aydogan Ozcan, "Phase recovery and holographic image reconstruction using deep learning in neural networks," <u>Light: Science and Applications</u>, vol. 7, no. 2, pp. 17141–17141, oct 2017.
- [10] Maziar Raissi, Paris Perdikaris, and George Em Karniadakis, "Physics-informed neural networks: A deep

- learning framework for solving forward and inverse problems involving nonlinear partial differential equations," <u>Journal of Computational Physics</u>, vol. 378, pp. 686–707, 2019.
- [11] Maziar Raissi, "Deep hidden physics models: Deep learning of nonlinear partial differential equations," <u>The Journal of Machine Learning Research</u>, vol. 19, no. 1, pp. 932–955, 2018.
- [12] Ziya Uddin, Sai Ganga, Rishi Asthana, and Wubshet Ibrahim, "Wavelets based physics informed neural networks to solve non-linear differential equations," Scientific Reports, vol. 13, no. 1, pp. 2882, Feb 2023.
- [13] Shengze Cai, Zhicheng Wang, Sifan Wang, Paris Perdikaris, and George Em Karniadakis, "Physics-Informed Neural Networks for Heat Transfer Problems," <u>Journal of Heat Transfer</u>, vol. 143, no. 6, pp. 060801, 04 2021.
- [14] Dongdong Chen, Julian Tachella, and Mike Davies, "Equivariant imaging: learning beyond the range space," in ICCV. 2021, pp. 4379–4388, IEEE/CVF.