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► **To cite this version:**

Thomas Fix, Slimane Raissi, Dominique Muller, Corinne Bouillet, Daniele Preziosi, et al.. Insights on CaTiS<sub>3</sub> films grown by pulsed laser deposition. *Journal of Alloys and Compounds*, 2023, 964, pp.171272. 10.1016/j.jallcom.2023.171272 . hal-04163977

**HAL Id: hal-04163977**

**<https://hal.science/hal-04163977>**

Submitted on 18 Jul 2023

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1 **Insights on CaTiS<sub>3</sub> films grown by pulsed laser deposition**

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10

11 **Abstract**

12 We attempt to grow perovskite CaTiS<sub>3</sub> by pulsed laser deposition (PLD) on various substrates using a  
13 CaS:TiS<sub>2</sub> target. Using Al<sub>2</sub>O<sub>3</sub> (0001) substrates at 600°C deposition temperature in vacuum,  
14 Ca<sub>1.0</sub>Ti<sub>1.05</sub>S<sub>2.48</sub>O<sub>y</sub> stoichiometry is measured by Energy-dispersive X-ray spectroscopy (EDS) and  
15 Ca<sub>1.0</sub>Ti<sub>1.0</sub>S<sub>2.0</sub>O<sub>0.7</sub> by Rutherford Backscattering Spectrometry (RBS). This indicates that when the films  
16 are grown in vacuum at moderate temperature, a high amount of S can be transferred from the target  
17 to the film. While no phase segregation of CaS and TiS<sub>2</sub> could be observed, no perovskite CaTiS<sub>3</sub> phase  
18 could be obtained, but rather a phase governed by van der Waals interactions. The films show a highly  
19 doped n-type semiconductor behavior with absorption coefficient in the 10<sup>5</sup> cm<sup>-1</sup> range at 350-2500  
20 nm but no surface photovoltage signal. This work should stimulate further experimental efforts in PLD  
21 growth of chalcogenides and chalcogenide perovskites.

22

23 **Keywords:** thin films; sulphide materials

## 24 1. Introduction

25 In the world of inorganic materials, while so many materials have already been discovered and  
26 reported, the class of chalcogenide perovskites and related  $AB_2S_3$  sulfides forms an exceptional  
27 reservoir of emerging materials. In  $ABX_3$  structures, more covalent B-X bonds correspond to a lower  
28 difference in electronegativity between B and X which allows better charge transport in the material  
29 and lower bandgap [1]. As an example, the experimental bandgap of  $BaZrO_3$  is 2.87 eV while it falls to  
30 1.75 eV for  $BaZrS_3$  [1] because S has lower electronegativity than O. While halide perovskites are  
31 outstanding materials for photovoltaic and optoelectronic applications, partly due to the low  
32 electronegativity difference between Pb and I in  $MAPbI_3$  (MA=methylammonium), oxide perovskites  
33 seem far behind despite their exceptional multifunctional properties among to name a few  
34 superconductivity, multiferroicity and high spin polarization, combined with high stability [2]. Sulfur  
35 and selenium are important elements for thin film photovoltaics because of their presence in the  
36  $CuIn_xGa_{1-x}(S,Se)_2$  (CIGS) absorber. When included in chalcogenide perovskites, they can form  
37 compounds rather stable in air and water as reported for  $BaZrS_3$  [3, 4], as opposed to  $MAPbI_3$ . As well,  
38 oxidation in air starts only in the 500-600 °C range [5]. Unfortunately, not all  $ABO_3$  starting material  
39 can be envisaged as  $AB_2S_3$  perovskites due to phase formation issues. Still, 26 different perovskite  
40 chalcogenides are reported in the review of Sopiha *et al.* [4].

41 Distorted perovskite  $CaTiS_3$  (CTS) has been reported in 2015 as a promising candidate for photovoltaics  
42 with a bandgap of about 0.9 eV [6]. Since then there has been no experimental report on CTS material  
43 while a direct bandgap of 0.90 eV was confirmed using first-principles calculations in an orthorhombic  
44 structure [7]. The stability of putative chalcogenide perovskites is well discussed in [4]. In the case of  
45 CTS, while the Goldschmidt tolerance factor  $t$  is about 0.92, which is in the range of  $0.7 < t < 1.1$  where  
46 stable perovskites can form, the octahedral factor  $\mu$  is about 0.33, which is lower than the minimum  
47 value of 0.41 for stable perovskite structure. As well, the computed value of formation energy of  
48  $GdFeO_3$ -type  $CaTiS_3$  perovskite with respect to binaries ( $H_f$ ) is reported to be unfavorable for perovskite

49 formation while the  $\text{CaTiS}_3$  perovskite ground state is low enough, only 40 meV f.u.<sup>-1</sup> higher than the  
50 needle-like phase which is calculated to be the most stable phase [4, 6, 8].

51 While  $\text{CaTiS}_3$  has not been reported experimentally, a related compound has been more extensively  
52 studied,  $\text{SnTiS}_3$ . It has been reported in the form of a crystal as a self-assembled van der Waals  
53 heterostructure of  $\text{SnS}$  and  $\text{TiS}_2$ , referred to as a misfit layer compound because the alternating layers  
54 form a heterostructure with a less symmetric crystal structure [9-11].

55 Because pulsed laser deposition (PLD) is well known to enable the formation of metastable phases, as  
56 in the case for example of h- $\text{TbMnO}_3$  less stable than o- $\text{TbMnO}_3$  [12], we attempt in this work to obtain  
57  $\text{CaTiS}_3$  perovskites. We use PLD for film growth, using a target mixing the two  $\text{CaS}$  and  $\text{TiS}_2$  precursors.  
58 Calcium sulfide  $\text{CaS}$  is an insulator with a bandgap from 2.5 to 4.5 eV, which provides high luminescence  
59 when doped with rare earth elements [13].  $\text{TiS}_2$  finds applications in batteries and is either described  
60 as a metal or semimetal, or a degenerate semiconductor, although a positive temperature dependence  
61 of resistivity is often observed [14-16]. In the following text, we will call the thin films deposited using  
62 the  $\text{CaS} : \text{TiS}_2$  target by the acronym CTS.

## 63 **2. Experimental procedure**

### 64 **PLD fabrication**

65 The target for PLD deposition was made by mixing the required quantities of  $\text{CaS}$  (purity 99.9%, Alfa  
66 Aesar) and  $\text{TiS}_2$  (purity 99.9%, Sigma Aldrich) and fabricating a 1'' pellet which was pressed isostatically  
67 with an actual pressure on the target of 460 MPa (Cold Isostatic Pressing Press from MIT Corporation  
68 YLJ-CIP-500M-30). The target was then mounted in the PLD chamber without any further treatment.  
69 The isostatic press is a key feature for obtaining a sulfur rich PLD target without the need of annealing  
70 under sulfur-rich atmosphere which would require special equipment and safety measures.

71 The substrates used for PLD deposition were  $\text{SrTiO}_3$  (001) (STO),  $\text{Al}_2\text{O}_3$  (0001) (ALO),  $\text{MgAl}_2\text{O}_4$  (001)  
72 (MAO),  $\text{TiO}_2$  (001) and (100) from Crystal GmbH, and fused quartz. The PLD process was performed

73 using a KrF laser (248 nm) with a 10 Hz repetition rate and a laser fluence of around 1-2 J/cm<sup>2</sup> (typically  
74 18000 pulses). The depositions were carried out at 600-900°C in vacuum (typically 5·10<sup>-6</sup> – 1·10<sup>-5</sup> mbar  
75 during deposition while at high temperature). The target-substrate distance was fixed at 55 mm.  
76 Around 15 samples were deposited by PLD for the study.

77

## 78 **Film characterization**

79 Films grown by PLD on crystal substrates were investigated using at room temperature a Rigaku  
80 SmartLab diffractometer equipped with a monochromatic source delivering a Cu K<sub>α1</sub> incident beam (45  
81 kV, 200 mA, 0.154056 nm). The same apparatus was used for x-ray diffraction (XRD) and x-ray  
82 reflectivity (XRR). Rutherford Backscattering Spectrometry (RBS) was used to determine atomic  
83 composition and thickness of the layers by using 1.5 MeV He<sup>+</sup> particles provided by the Van de Graaff  
84 accelerator (HVEE KN4000) facility of ICube. The analysis of the spectra was performed using the  
85 SIMNRA tool. The optical properties were obtained by measuring transmittance with a UV-vis-IR  
86 Perkin-Elmer Lambda 19 spectrophotometer. The optical properties and thickness of the films were  
87 also investigated by spectroscopic ellipsometry using a HORIBA Uvisel Lt M200 FGMS (210–880 nm)  
88 apparatus. The dispersion formula for CTS films is based on the double Tauc-Lorentz model [17]. The  
89 goodness-of-fit value (Chi Square  $\chi^2$ ) was 3.2 for a CTS film thickness of around 80 nm.  
90 Photoluminescence (PL) was carried out at room temperature using an excitation at 355 nm with a  
91 frequency-tripled neodymium-doped yttrium aluminum garnet (Nd-YAG) laser. The laser pulse  
92 duration, frequency and power were 10 ns, 20 kHz and up to 45 mW (7 W cm<sup>-2</sup>), respectively. The PL  
93 signal was collected by means of a concave mirror into an optical fiber and analyzed by a multichannel  
94 CCD. The PL signal was integrated over time for a duration of 600 ms.

95 The surface of the films was investigated using atomic force microscopy (AFM) with a NT-MDT Smena  
96 B AFM in tapping mode. Scanning electron microscopy (SEM) was performed with a JEOL JSM-IT200LA  
97 apparatus and the stoichiometry was checked by Energy-dispersive X-ray spectroscopy (EDS). Hall

98 effect was measured using an ECOPIA HMS-5000 system at room temperature. Resistance vs.  
99 temperature was measured with a Physical Properties Measurement System (Dynacool, Quantum  
100 Design) from 300 K to 5 K using a van der Pauw configuration. Kelvin Probe (KP) and Surface  
101 PhotoVoltage (SPV) measurements were performed with a single point Kelvin Probe system (KP  
102 Technology KP020) with a 2 mm diameter gold tip, and using a halogen source with calibrated  
103 illumination for white light SPV.

104 To further investigate the structural properties of the films, Transmission Electron Microscopy (TEM)  
105 was performed using a JEOL 2100 F microscope, operating at 200 kV, with a point to point resolution  
106 of 0.2 nm and a corrector on the condenser lenses, to perform high-resolution imaging. The sample  
107 was prepared, as a cross section in the STO substrate [100] zone axis, using the focused ion beam (FIB)  
108 technique.

### 109 **3. Results and discussion**

110 Films were grown first on STO (001) in vacuum at different deposition temperatures from 600°C to  
111 900°C using a target containing the CaS and TiS<sub>2</sub> precursors as described in the experimental  
112 procedure. Peaks of CaS and TiS<sub>2</sub> are identified as shown in Figure 1. This is confirmed by TEM  
113 observations described later in the text. As shown in Table 1, the films grown on STO (001) in vacuum  
114 at 600°C and 700°C contain an appropriate amount of S as determined by EDS. This means that the  
115 preservation of stoichiometry due to the so-called congruence of PLD is also valid for the S element  
116 under certain conditions. The films grown at 800°C and 900°C show a progressive decrease of S molar  
117 amount, which is caused by the re-evaporation of sulfur due to high vapor pressure. The Ti  
118 stoichiometry cannot be determined for these samples because of the presence of Ti in the STO  
119 substrate; the O stoichiometry cannot also be determined by EDS. However, the film grown on ALO  
120 (0001) indicates that the Ca and Ti elements are present in the right composition according to EDS:  
121 Ca<sub>1.0</sub>Ti<sub>1.05</sub>S<sub>2.48</sub>O<sub>y</sub>. This is confirmed by RBS that indicates a stoichiometry of Ca<sub>1.0</sub>Ti<sub>1.0</sub>S<sub>2.0</sub>O<sub>0.7</sub>, revealing  
122 the amount of oxygen in the film on ALO (0001) (Figure S1 of Supplementary Information (SI)). Coming

123 back to the 800°C and 900°C samples on STO, the decrease of S amount in the films is accompanied in  
124 the XRD shown in Figure 1 by the appearance of CaTiO<sub>3</sub> (00l) peaks. The CaTiO<sub>3</sub> (002) peak corresponds  
125 to a pseudocubic out-of-plane parameter of 0.381 ±0.001 nm, very close to the pseudocubic out-of-  
126 plane parameter of perovskite CaTiO<sub>3</sub> (0.3820 nm, ICDD 04-007-5451). This indicates that there is no  
127 incorporation of S in the CaTiO<sub>3</sub> phase that is formed. CaTiO<sub>3</sub> can form because at high temperature  
128 the base vacuum pressure is higher and oxygen is more reactive to form the CaTiO<sub>3</sub> phase. The CaS  
129 (002) and (004) peak positions are not moving when increasing the deposition temperature. This can  
130 indicate again no incorporation of S in the corresponding CaTiO<sub>3</sub> lattice. However, when looking at TiS<sub>2</sub>  
131 (001) there is a small progressive shift towards higher angles when going from 600°C to 800°C while  
132 the peak disappears at 900°C.

133 When now keeping the PLD growth at 600°C on STO (001) we introduce voluntarily an oxygen  
134 atmosphere at 10<sup>-3</sup> or 10<sup>-2</sup> mbar O<sub>2</sub>. Figure 2 shows that when oxygen is introduced the CTS phase  
135 disappears and the CaTiO<sub>3</sub> (001) and (002) peaks appear. This shows that even a low oxygen  
136 atmosphere is detrimental to the formation of CTS. This is confirmed by Table 1 showing the absence  
137 of S in the samples grown in O<sub>2</sub>.

138 In order to further understand the growth of the CTS phase, samples were grown on different crystal  
139 substrates such as MAO (001) and ALO (0001). Surprisingly, the same peaks of TiS<sub>2</sub> (001) and CaS (002)  
140 present with the STO (001) substrate can be observed with MAO and ALO substrates, and at the same  
141 angular position (Figure 3). This is unusual for the following reasons. STO is cubic and has a lattice  
142 parameter of 0.3905 nm while MAO is also cubic but with a lattice parameter of 0.8083 nm, i.e. two  
143 times 0.4015 nm. If there was epitaxy or texture of the CTS film, such a change in lattice parameters  
144 of the substrates would induce a change in the film out-of-plane parameter, unless fully relaxed.  
145 Concerning ALO, because this substrate has a R-3c space group it is not expected that CTS would grow  
146 in the same way as when grown of STO (001). As well, XRD of CTS on TiO<sub>2</sub> and fused quartz substrates  
147 provide the same texture of the films as with the other substrates. All these results indicate that the

148 CTS films grown do not obey to a traditional epitaxy as observed in many oxide systems using these  
149 crystal substrates. We will therefore use TEM to further investigate the films.

150 SEM observations indicated that the films grown on STO (001) are continuous, but with some splashing  
151 on the surface which is well known in the PLD process (Figure S2). Similar observations are made for  
152 films on ALO (0001) (Figure S3). When observing the sample in compositional mode, no phase  
153 segregation or elemental contrast can be observed apart from the previously mentioned splashing.  
154 This means that the surface of the films is homogeneous in composition at the scale achievable by the  
155 SEM. The surface quality is further investigated by AFM. Figure 4 shows an AFM topography image of  
156 a CTS sample on STO (001). The root mean square (RMS) roughness is 2.7 nm and the grain size is in  
157 the 100 nm range. This indicates that the films have a good surface quality for a film of around 80 nm  
158 thickness.

159 We now look into the optical properties of the films. The absorption coefficient of a CTS film is shown  
160 in Figure 5, and compared to a spectrum for diamond silicon. It is observed that CTS has a strong  
161 absorption coefficient in the  $10^5 \text{ cm}^{-1}$  range at 350-2500 nm. This is confirmed by spectroscopic  
162 ellipsometry and the refractive index and extinction coefficient could be determined (Figure S4 of SI).  
163 The Tauc plot in Figure 5 indicates a direct semiconductor with a bandgap of 1.59 eV.  
164 Photoluminescence measurements were also performed on CTS films grown on STO and ALO but no  
165 contribution other than the contribution of the substrates could be measured (Figure S5 of SI).

166 Now that good light absorption properties have been demonstrated, it is fundamental to focus in the  
167 electrical properties of the films. Because CTS is grown in vacuum, STO substrates are not appropriate  
168 for electrical measurements of the films because they become conductive due to oxygen deficiency.  
169 Therefore, the electrical measurements were performed on the ALO samples. Hall effect using CTS on  
170 ALO (0001) indicates a n-type conduction with a carrier density of  $-3.4 \cdot 10^{22} \text{ cm}^{-3}$ , mobility of  $0.34 \text{ cm}^2 \text{V}^{-1} \text{ s}^{-1}$   
171 and resistivity of  $2.6 \cdot 10^{-3} \Omega \cdot \text{cm}$ . The value of carrier density is very high and the value of resistivity  
172 is very low which could suggest a metallic nature of the conduction of the film. To confirm this, we



173 performed resistance vs. temperature measurements of the film. Figure 6 shows a R-T measurement  
174 in the 300 K – 5 K range, showing that the sheet resistance increases when the temperature decreases,  
175 which is expected for a semiconductor.

176 As well, we measured the work function of the films using the Kelvin probe technique as detailed in  
177 the experimental section. The work function was measured to be in the 4.5-4.7 eV range. Using the  
178 same Kelvin probe apparatus while adding visible light on the sample during the measurement, surface  
179 photovoltage (SPV) can be measured. SPV is a widely used method to estimate the minority carrier  
180 diffusion length of semiconductors. It consists in measuring the potential of a semiconductor surface  
181 while generating electron-hole pairs with a light source. Typical highly efficient absorbers for solar cells  
182 provide a high SPV signal. Here, the photoactive response was negligible despite the high precision of  
183 the apparatus (Figure S6), indicating that the CTS films cannot be used as such as a photovoltaic  
184 absorber in a solar cell. This could be due to the high carrier density measured in the films.

185 Now that the basic properties of the films such as morphology, surface, optical, electrical properties  
186 are determined there is still some uncertainty about the structural arrangement of the CTS films. SEM  
187 has shown no phase segregation within the detection limits, tending to say that CaS and TiS<sub>2</sub> atomic  
188 planes are intricately combined within the structure. To do so, high resolution TEM is performed on a  
189 CTS sample on STO (001) as shown in Figure 7. It shows that the CTS film is crystalline but there are  
190 many Moiré fringes visible, of d-spacing of about 0.586 nm. Using the Moiré formula for TiS<sub>2</sub> d-spacing  
191 of 0.5705 nm and CaS d-spacing of 0.2811 nm yields a periodicity of 0.5541 nm quite close to the  
192 observed value. Moiré fringes are often observed in van der Waals superstructures that can be based  
193 on transition metal dichalcogenides [18]. The selected area electron diffraction (SAED) pattern is  
194 shown in Figure S7 of SI and the CTS film and substrate spots are identified, revealing the presence of  
195 d-spacing corresponding to CaS (002) and also TiS<sub>2</sub> (100) and (001), indicating that there is texture but  
196 multiple variants in plane for TiS<sub>2</sub> in the CTS film.

## 197 **Conclusions**

198 In conclusion, we attempted to grow perovskite  $\text{CaTiS}_3$  films by PLD in the hope to benefit from the  
199 non-thermodynamic equilibrium nature of PLD growth and from the concept of epitaxial stabilization  
200 using various crystalline substrates. After having grown about 15 films we conclude that the putative  
201 perovskite phase of  $\text{CaTiS}_3$  is not obtained experimentally. The PLD process is unexpectedly preserving  
202 relatively well the sulfur stoichiometry, if the deposition temperature is limited to 600-700°C and the  
203 films are deposited in vacuum. We found that the nature of the substrate has little influence on the  
204 diffraction patterns of the films, and therefore the growth of the films can be related to van der Waals  
205 interactions that are often encountered in chalcogenide materials. The films are of a highly-doped n-  
206 type semiconductor nature as observed by Hall effect and R-T measurements. TEM observations in  
207 cross section show that the films are crystalline but the observations in high resolution are disrupted  
208 by Moiré fringes. Surface photovoltage measurements indicate that the films are not photoactive,  
209 although the absorption coefficient is in the  $10^5 \text{ cm}^{-1}$  range and the direct bandgap is about 1.59 eV.  
210 This could be due to the high carrier density in the  $10^{22} \text{ cm}^{-3}$  range that can be a source of high Auger  
211 recombinations. The work should stimulate further experimental efforts into fabrication of emerging  
212 films of chalcogenide and chalcogenide perovskites by PLD.

### 213 **Author contributions**

214 TF conceptualization, methodology, investigation, writing – original draft, supervision; SR  
215 investigation, writing – review & editing; DM RBS measurements, writing – review & editing; CB TEM  
216 observations, writing – review & editing; DP investigation, writing – review & editing; AS writing –  
217 review & editing, project administration.

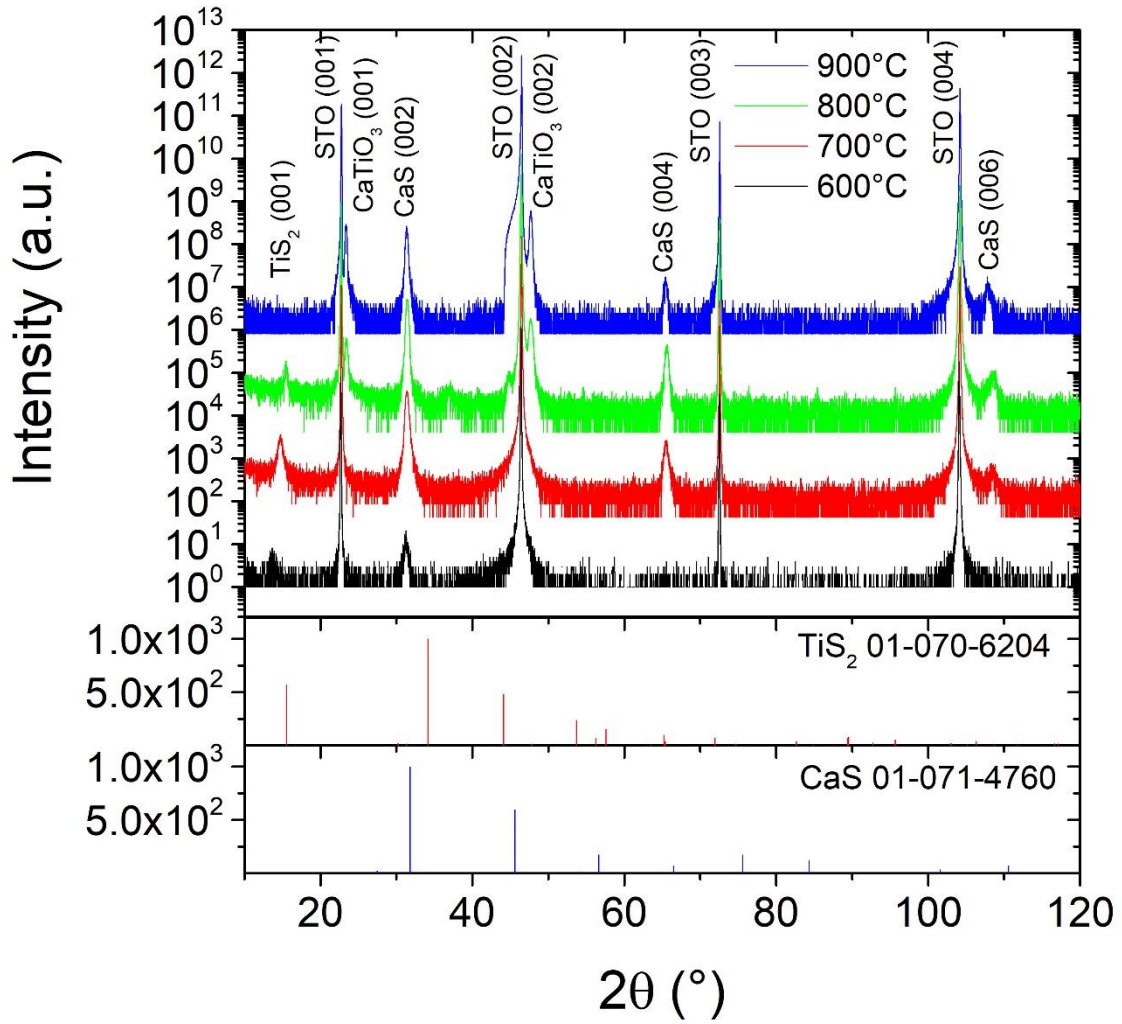
### 218 **Declaration of Competing Interest**

219 The authors declare that they have no known competing financial interests or personal relationships  
220 that could have appeared to influence the work reported in this paper.

### 221 **Acknowledgments**

222 The authors thank the XRD platform of IPCMS including M. Lenertz, and the staff of the C<sup>3</sup>Fab platform  
223 of ICube, including J. Bartringer for SEM and PL measurements and S. Roques. D. Troadec is greatly  
224 thanked for FIB lamella preparation.

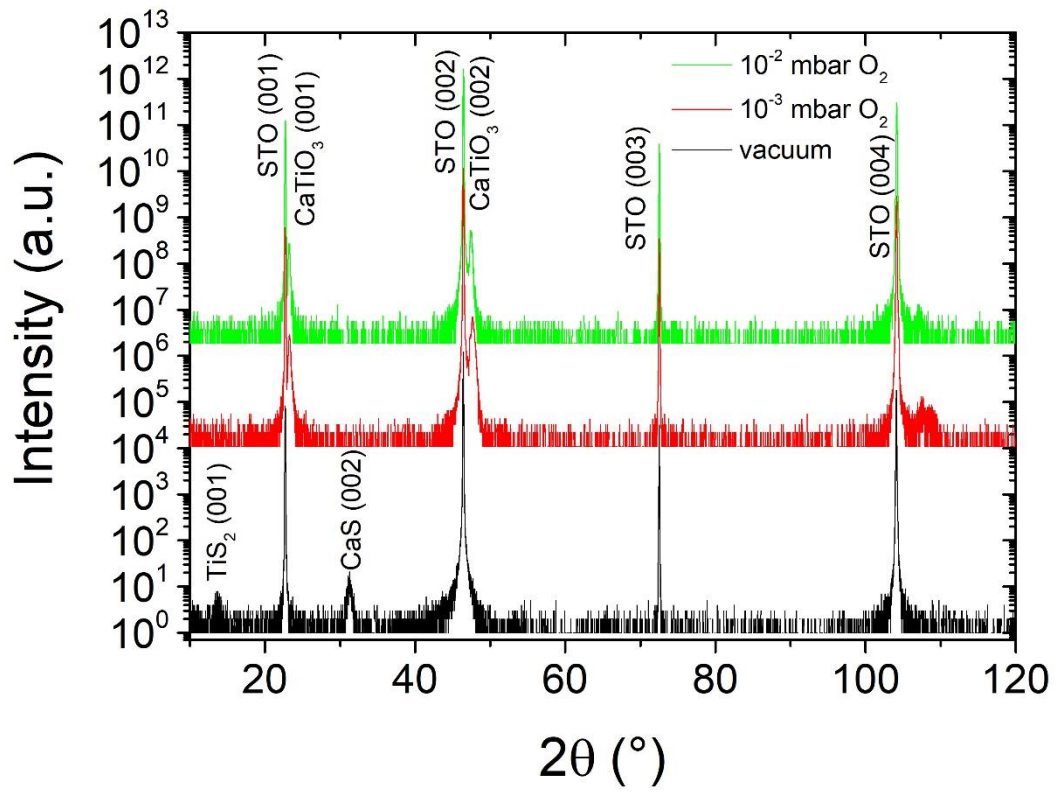
225 **Figure 1:** XRD  $\theta$ - $2\theta$  patterns of films grown by PLD on STO (001) at different deposition temperatures,  
226 and corresponding ICDD references for  $\text{TiS}_2$  and  $\text{CaS}$ .



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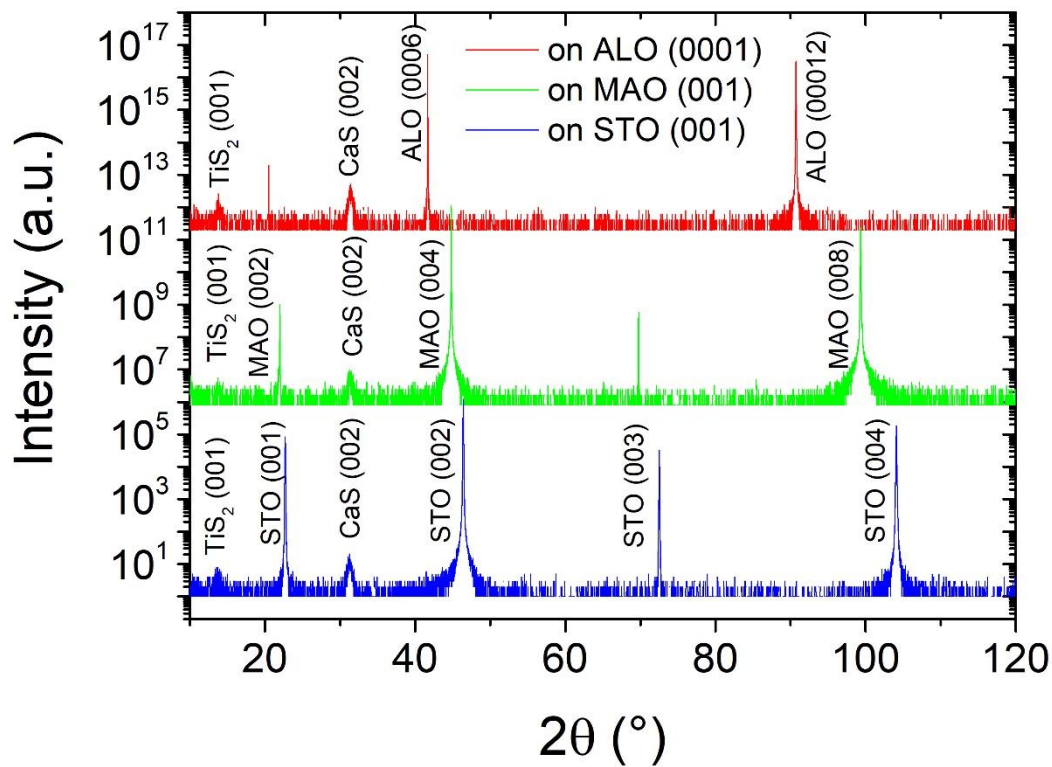
229 **Figure 2:** XRD  $\theta$ - $2\theta$  patterns of films grown by PLD on STO (001) at 600°C in different atmospheres.



230

231 **Figure 3:** XRD  $\theta$ - $2\theta$  patterns of films grown by PLD on different substrates at 600°C in vacuum.

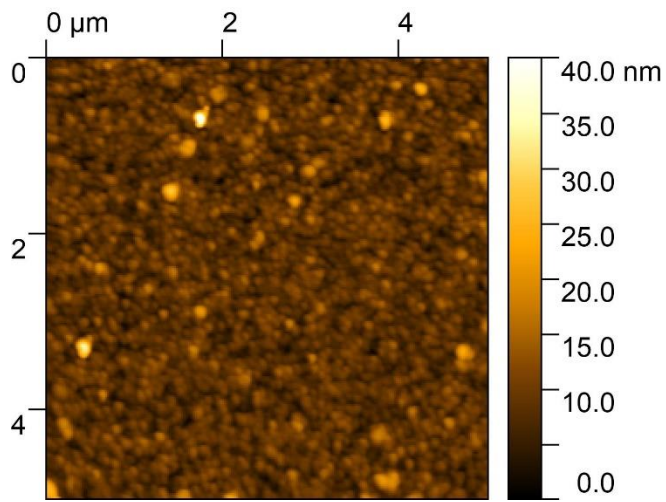
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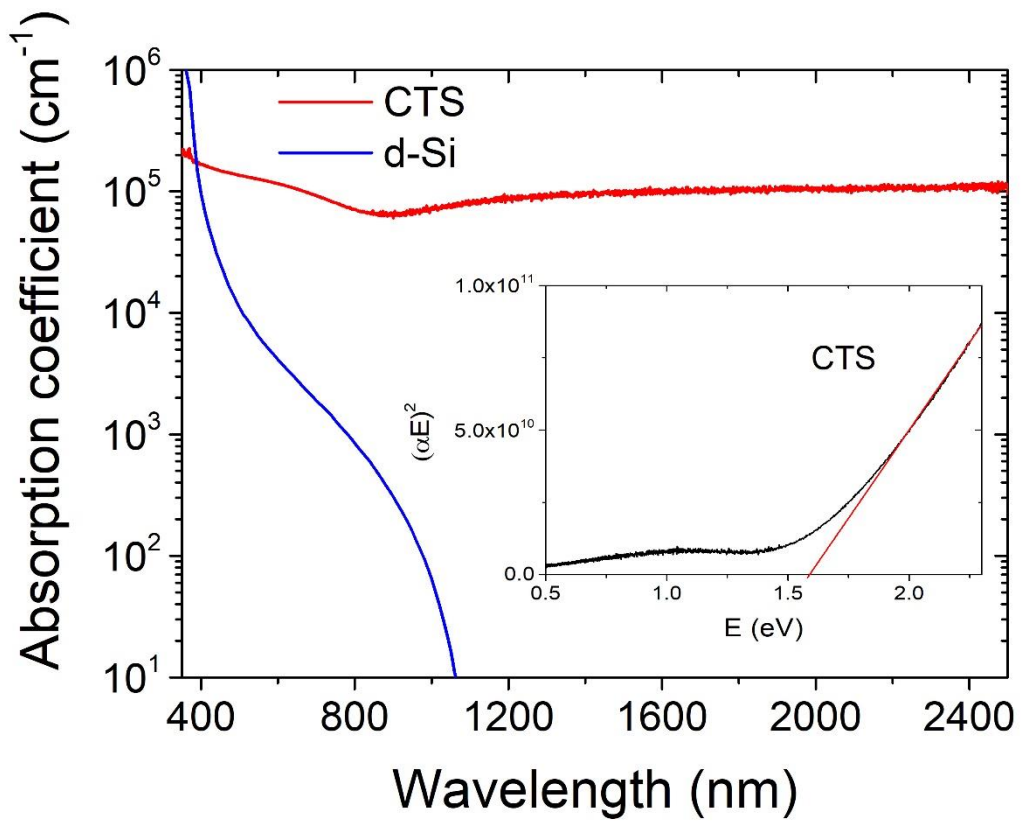
235 **Figure 4:** AFM topography image of a CTS sample grown on STO (001) at 700°C in vacuum.



236

237

238 **Figure 5:** Absorption coefficient determined by UV-visible spectroscopy for a CTS sample grown on  
239 quartz at 600°C in vacuum. The reference for diamond silicon is taken from [19]. Inset: Tauc plot of  
240 CTS for a direct bandgap material.  
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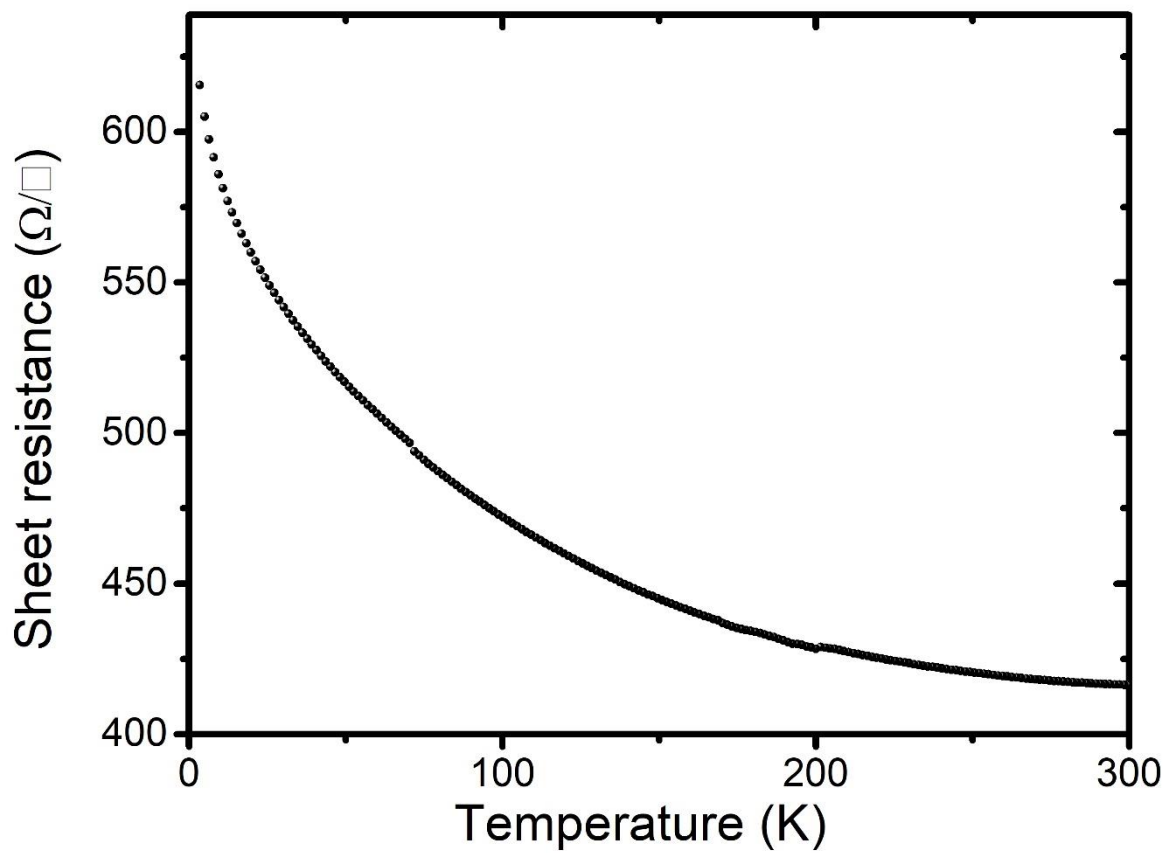


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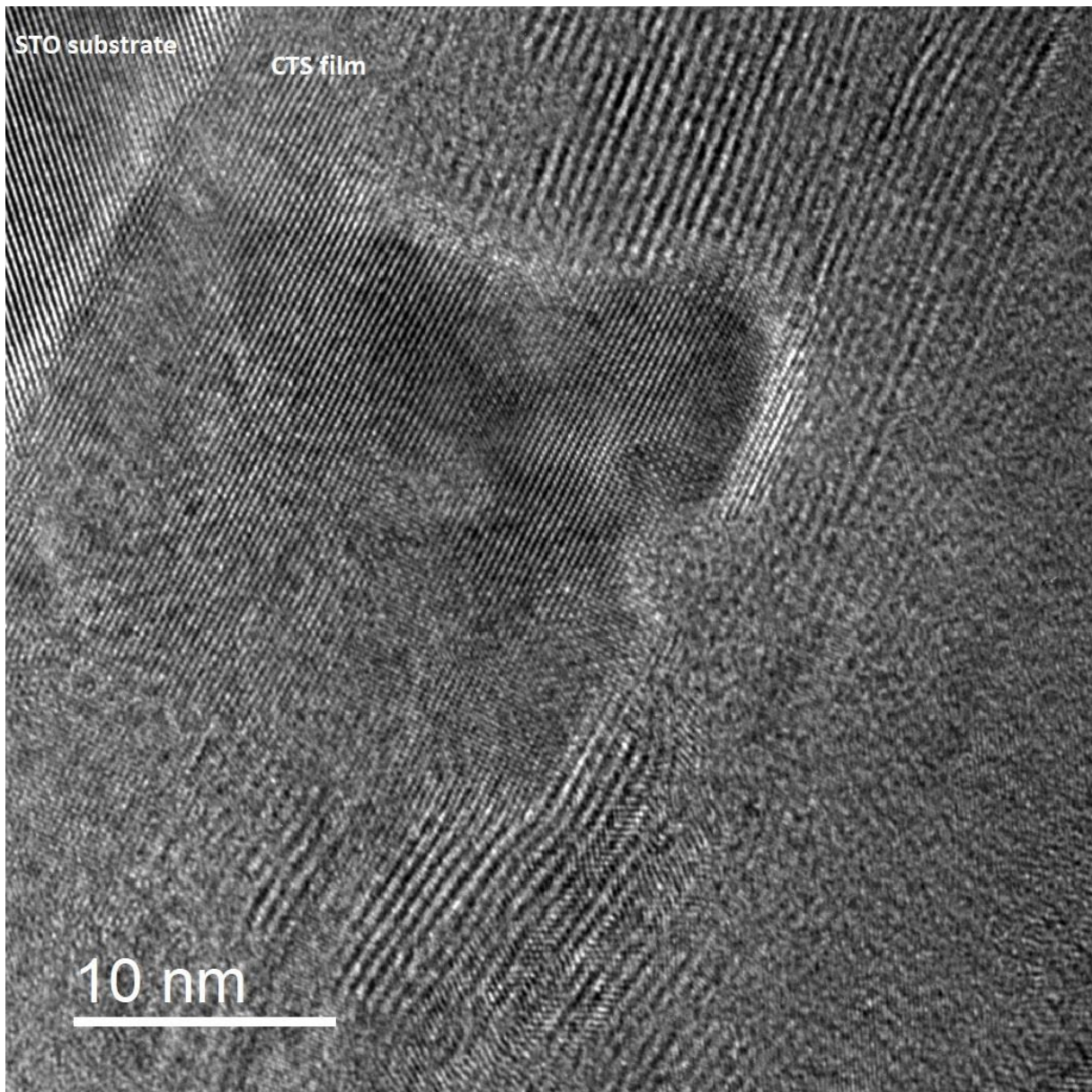
244 **Figure 6:** Sheet resistance vs. temperature of a CTS film grown on ALO (0001).



245

246

247 **Figure 7:** High resolution TEM image in the STO [100] zone axis.  
248



249  
250

251 **Table 1:** Stoichiometry determined by EDS of the samples grown under different conditions such as  
252 substrate, deposition temperature and atmosphere.

253

Substrate	Deposition temperature (°C)	Deposition atmosphere	Stoichiometry (EDS)
STO (001)	600	vacuum	$\text{Ca}_1\text{Ti}_x\text{S}_{3.16}\text{O}_y$
STO (001)	700	vacuum	$\text{Ca}_1\text{Ti}_x\text{S}_{3.11}\text{O}_y$
STO (001)	800	vacuum	$\text{Ca}_1\text{Ti}_x\text{S}_{1.78}\text{O}_y$
STO (001)	900	vacuum	$\text{Ca}_1\text{Ti}_x\text{S}_{0.71}\text{O}_y$
STO (001)	600	$10^{-3}$ mbar $\text{O}_2$	$\text{Ca}_1\text{Ti}_x\text{O}_y$ (S=0)
STO (001)	600	$10^{-2}$ mbar $\text{O}_2$	$\text{Ca}_1\text{Ti}_x\text{O}_y$ (S=0)
ALO (0001)	600	vacuum	$\text{Ca}_1\text{Ti}_{1.05}\text{S}_{2.48}\text{O}_y$
MAO (001)	600	vacuum	$\text{Ca}_1\text{Ti}_{0.92}\text{S}_{2.53}\text{O}_y$

254

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308