Geodynamic significance of the Raspas Metamorphic Complex (SW Ecuador): geochemical and isotopic constraints

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Received 2 May 2000; received in revised form 30 January 2001; accepted 7 February 2001

Abstract

The Raspas Metamorphic Complex of southwestern Ecuador is regarded as the southernmost remnant of oceanic and continental terranes accreted in the latest Jurassic –Early Cretaceous. It consists of variably metamorphosed rock types. (1) Mafic and ultramafic rocks metamorphosed under high-pressure (HP) conditions (eclogite facies) show oceanic plateau affinities with flat REE chondrite-normalized patterns, $\epsilon Nd_{150 \text{ Ma}}$ ranging from +4.6 to 9.8 and initial Pb isotopic ratios intermediate between MORB and OIB. (2) Sedimentary rocks metamorphosed under eclogitic conditions exhibit LREE enriched patterns, strong negative Eu anomalies, Rb, Nb, U, Th, Pb enrichments, low $\epsilon Nd_{150 \text{ Ma}}$ values (from -6.4 to -9.5), and high initial ${}^{87}Sr/{}^{86}Sr$ and ${}^{206,207,208}Pb/{}^{204}Pb$ isotopic ratios suggesting they were originally sediments derived from the erosion of an old continental crust. (3) Epidote-bearing amphibolites show N-MORB affinities with LREE depleted patterns, LILE, Zr, Hf and Th depletion, high $\epsilon Nd_{150 \text{ Ma}}$ (>+10) and low initial Pb isotopic ratios. The present-day well defined internal structure of the Raspas Metamorphic Complex seems to be inconsistent with the formerly proposed interpretation of a "tectonic mélange". The association of oceanic plateau rocks and continent-derived sediments both metamorphosed in HP conditions suggests that the thin edge of the oceanic plateau first entered the subduction zone and dragged sediments downward of the accretionary wedge along the Wadatti-Benioff zone. Subsequently, when its thickest part arrived into the subduction zone, the oceanic plateau jammed the subduction processes, due to its high buoyancy. In Ecuador and Colombia, the latest Jurassic –Early Cretaceous suture involves HP oceanic plateau rocks and N-MORB rocks metamorphosed under lower grades, suggesting a composite or polyphase nature for the latest Jurassic –Early Cretaceous accretionary event. \oslash 2002 Elsevier Science B.V. All rights reserved.

Keywords: Raspas Metamorphic Complex; Geochemical constraints; Isotopic constraints; Eclogite; Subduction processes

1. Introduction

The western margin of South America is a classical example of an ocean –continent convergence, which eventually led to the Andean orogeny in Tertiary times. One puzzling problem concerns the significance of the scarce outcrops of high-pressure (HP) metamorphic rocks along the Andean margin. There is no consensus on their tectonic and geodynamic significance and they have been interpreted as related either to subduction, accretion, or collision processes (e.g. Feininger, 1980; Aspden and McCourt, 1986; Bourgois et al., 1987; Aspden et al., 1995; Lucassen et al., 1999).

The western part of northwestern South America (Ecuador and Colombia) is made up of exotic terranes mainly of oceanic origin (Fig. 1), accreted between latest Jurassic and Eocene times (Toussaint and Restrepo, 1994; Litherland et al., 1994; Reynaud et al., 1999; Lapierre et al., 2000; Hughes and Pilatasig, 2001). The eastern boundary of the accreted terranes is marked by the occurrence of tectonic slices of HP metamorphic rocks, which yielded Early Cretaceous cooling ages (\approx 132–110 Ma) and are thought to represent the suture zone of oceanic terranes accreted during the latest Jurassic –Early Cretaceous times (Feininger, 1982; Aspden and McCourt, 1986; Toussaint and Restrepo, 1994; Litherland et al., 1994, Fig.

Fig. 1. Structural sketch of northwestern South America showing the exotic terranes accreted in Mesozoic and Cenozoic times and location of the studied area (inset).

1). In Colombia (Kerr et al., 1996, 1997) and SW Ecuador (Arculus et al., 1999; Malfere et al., 1999), a part of the HP metamorphic rocks possesses oceanic plateau affinities.

This paper presents the results of a geochemical study carried out on rocks of the metamorphic Raspas ''Ophiolitic'' Complex of southwestern Ecuador (Aspden et al., 1995). Since the Raspas outcrops do not show clear evidences of a typical ophiolitic suite and in order to avoid interpretative ambiguities, we shall use in this work the term of ''Raspas Metamorphic Complex''. The present study has been undertaken in order to determine the nature and geodynamic significance of the terranes involved in the HP metamorphism, and to test the hypothesis of an accretionary complex.

2. Geological setting

In Western Ecuador, the latest Jurassic –Early Cretaceous accretion episode involved both continental and oceanic units (Litherland et al., 1994). The corresponding suture is located at the western edge of the Eastern Cordillera of Ecuador (Peltetec suture, Litherland et al., 1994). It is thought to extend into southwesternmost Ecuador in the El Oro metamorphic complex (Aspden et al., 1995). The latter belongs to the displaced Amotape–Tahuin block (Mourier et al., 1988) and comprises various lithologies and assemblages with ages ranging from Palaeozoic to Cretaceous (Aspden et al., 1995). The El Oro metamorphic complex is crosscut by several $E-W$ trending dextral fault systems that subdivide the El Oro metamorphic complex into various smaller units (Fig. 2).

Among these units, the $E-W$ trending Raspas Metamorphic Complex is bounded to the north and south by the "La Palma-El Guayabo" and the ''Tahuin Dam'' faults, respectively, which separate the Raspas Metamorphic Complex from rocks belonging to the Andean continental margin (Aspden et al., 1995; Fig. 2). The Raspas Metamorphic Complex consists of greenschists, serpentinised peridotites, pelitic schists, blueschists and eclogites, first described in detail by Feininger (1980). Phengite K-Ar data (132 ± 5) Ma; Feininger and Silberman, 1982) points to a lowermost Cretaceous age for the HP metamorphism (Duque, 1993), thus suggesting that it is related to the latest Jurassic –Early Cretaceous accretion episode (Aspden et al., 1995).

Recent laser Ar/Ar datings have been performed on different samples from the present studied area (Monie et al., in preparation). The analyzed samples include the 97Ce1 eclogite and the 97Ce3 and 97Ce5 metapelites. Ar/Ar ages obtained on both phengite and amphibole range from ca. 122 to 130 Ma. This supports a lower Cretaceous age for the HP event recorded in the Raspas Metamorphic Complex rocks in good agreement with the previous dates (Feininger and Silberman, 1982).

The samples analyzed for geochemistry were collected from the various tectono-metamorphic units defined in the Raspas Metamorphic Complex: the La Chilca, El Toro and Río Panupali units (Aspden et al., 1995, Fig. 2).

3. Lithologies

From a petrographic point of view, the analyzed samples can be subdivided into two main groups, namely the metabasites and the metapelites. Table 1 summarizes the paragenesis of the analyzed metamorphic rocks.

3.1. The metabasites

The metabasites include eclogites, serpentinised peridotites and amphibolites of different metamorphic facies. The high-pressure rocks belong to the La Chilca Unit while the amphibolites and greenschists are the main components of the Río Panapuli Unit (Fig. 2).

The eclogites (97Ce1, 98RR4, 98RR6) and garnetamphibolites (97Ce4A) crop out in the Río Raspas (Fig. 2, UTM-WGS84: 17/620500E-9602100N). The eclogites consist of poikiloblastic garnet, quartz, omphacite, epidote, amphibole (barroisite) \pm phengite. Their contents in phengite and quartz vary significantly from one facies to another. Garnet commonly includes quartz, rutile and epidote. Zones rich in amphibole or omphacite can lead to a strongly layered texture. The garnet-amphibolites are associated with the eclogites and are characterized by the occurrence of blue-green amphibole.

According to Aspden et al. (1995), the serpentinised peridotites pinched within the La Palma –El Guayabo

Fig. 2. Structural sketch of the El Oro Metamorphic complex (adapted from Aspden et al., 1995) and location of the studied samples.

Sample	97Ce1 [eclogite]	98RR4 [eclogite]	98RR6 [eclogite]	97Ce4A [Grt amphibolite]	97Ce13 [amphibolite (metagabbro)]	97Ce18 [Ep-bearing] amphibolite]	98TH13 [Ep-bearing] amphibolite]	97Ce12 [Ep-bearing] amphibolite]	97Ce14 [serpentinised peridotite]	97Ce15 [serpentinised peridotite]
Qtz	\times	\times	\times	\times	\times	\times	\times	\times		
Ms		\times	\times							
Grt	\times	\times	\times	\times						
Omp	\times	\times	\times							
Hbl				\times	\times					
Act						\times	\times	\times		
Tr								\times	\times	\times
Bar	\times	\times	\times							
Ol									\times	\times
Srp									\times	\times
Czo	\times	\times	\times	\times						
Ep						\times	\times	\times		
Chl				\times	\times	\times	\times		\times	\times
Ab				\times		\times	\times	\times		
P ₁					\times					
Rt	\times	\times	\times	\times	\times					
Ttn	\times	\times	\times	\times		\times	\times	\times		
Cal/Dol					\times				\times	\times
Opaques	\times	\times	\times	\times				\times	\times	\times
Met. Facies	eclogite	eclogite	eclogite	epidote- amphibolite	amphibolite	greenschist	greenschist	greenschist	amphibolite	amphibolite

Table 1Mineral assemblages of the studied metabasites

Mineral assemblages of the studied metapelites

Abbreviations are after Kretz (1983).

Table 2 (continued)

La Chilca Unit: sedimentary rocks							
	Metapelite	Metapelite whole Rock	Phengite	Grt with inclusions	Metapelite whole Rock	Metapelite whole rock	Metapelite whole rock
Sample number	97CE2	97CE3	97CE3	97CE3	97CE4	97CE5	97CE6
SiO ₂	72.14^{a}	74.3	nd	nd	70.05	76.03	$76.86^{\rm a}$
TiO ₂	0.46	0.52	nd	nd	0.73	0.82	0.88
Al ₂ O ₃	12.9	11.87	$^{\rm nd}$	nd	13.86	13.05	12.45
Fe ₂ O ₃	5.26	5.71	nd	nd	6.56	4.15	4.68
MnO	0.14	$0.07\,$	nd	nd	$0.08\,$	bdl	0.03
MgO	1.26	1.28	nd	nd	1.46	0.45	0.54
CaO	0.51	0.69	nd	nd	$0.81\,$	0.18	0.28
Na ₂ O	0.84	0.63	nd	nd	0.88	0.68	0.58
K_2O	2.22	1.88	nd	nd	1.94	1.31	1.27
P_2O_5	0.18	0.09	nd	nd	0.13	$0.08\,$	0.1
LOI	nd	2.84	nd	nd	3.4	3.11	$^{\rm nd}$
Total	96.17	99.88	nd	nd	99.9	99.86	97.68
Sc	24	nd	nd	nd	nd	nd	$18\,$
$\ensuremath{\mathbf{V}}$	278	nd	nd	nd	120	71	71
Cr	122.8	nd	nd	nd	94	62	61.3
Ni	8.6	nd	nd	nd	25	12	7.8
Rb	82.8	78.78	309.3	4.58	40.14	18.74	49.9
$\rm Sr$	97.7	71.7	178.77	3.06	70.08	15.57	74.9
Ba	1547	607.98	2469.96	28.69	362.93	56.35	334.3
U	1.802	1.16	0.6	0.42	1.47	1.10	1.35
Th	9.804	5.79	3.24	1.17	6.52	1.96	4.87
Pb	16.19	14.01	43.91	1.12	14.32	12.25	12.73
Hf	4.93	0.25	0.07	1.36	1.31	3.53	8.67
Zr	159.8	9.55	2.09	53.01	51.3	137.43	318.3
Ta	0.752	0.53	$0.08\,$	0.25	0.9	1.17	0.95
Nb	10.18	8.16	1.83	3.26	14.01	17.43	14.02
Y	34.2	27.41	2.18	167.26	31.66	9.58	17.3
La	21.08	20.68	15.05	3.53	27.13	9.07	16.89
Ce	38.11	42.26	30.88	6.69	61.38	28.8	29.07
Pr	nd	5.3	3.85	0.9	$7.2\,$	2.65	nd
Nd Sm	20.95 4.32	19.97 4.05	14.23 2.73	3.55 1.88	27.81	9.71 1.95	13.57 2.73
$\mathop{\mathrm{Eu}}\nolimits$	0.96	0.96	0.73	$\mathbf{1}$	5.61	0.42	0.65
Gd	4.16	4.08	1.96	8.31	1.26 5.22	1.66	3.06
Tb	nd	0.65	0.18	2.68	$\rm 0.8$	0.27	nd
Dy	5.4	4.08	0.62	22.23	4.81	1.66	3.45
Ho	nd	0.88	$0.08\,$	5.53	1.03	0.34	nd
Er	3.47	2.63	0.17	17.66	3.03	1.06	1.93
Tm	nd	nd	nd	nd	nd	nd	nd
Yb	3.54	2.62	0.13	19.58	3.11	1.24	2.08
Lu	0.55	0.42	0.02	3.14	0.49	0.2	0.32
${}^{87}Sr/{}^{86}Sr$	nd	0.722593 ± 16	nd	nd	nd	0.721221 ± 16	nd
$({}^{87}Sr/{}^{86}Sr)i$ ^c		0.71581				0.71379	
143 Nd/ 144 Nd	nd	0.512240 ± 11	nd	nd	0.512126 ± 05	0.512079 ± 08	nd
$(^{143}Nd/^{144}Nd)i^c$		0.51212			0.51201	0.51196	
εNd_i		-6.35			-8.56	-9.47	
$(^{208}Pb/^{204}Pb)i^c$	nd	38.774	nd	39.653	38.780	38.890	38.910
$(207Pb)^{204}Pb)i^c$		15.700		15.688	15.700	15.710	15.708
$(206Pb)^{204}Pb)i^c$		18.925		19.657	18.907	18.972	18.957

(continued on next page)

	El Toro Unit: ultramafic rocks			Rio Panupili unit: mafic rocks					
	Serpentinite	Duplicate	Harzburgite	Amphibolite metabasalt	Amphibolite metagabbro	Amphibolite metabasalt	Amphibolite metabasalt		
	97CE14	97CE14	97CE15	97CE18	97CE13	98TH13	97CE12		
SiO ₂	40.08 ^d		40.97	50.59 ^d	49.92 ^d	48.76	48.95		
TiO ₂	0.09		0.08	1.42	0.77	1.16	1.41		
Al_2O_3	2.81		2.98	13.96	17.42	14.64	14.91		
Fe ₂ O ₃	8.25		8.27	10.31	7.35	10.25	10.62		
MnO	0.1		0.11	0.17	0.11	0.16	0.19		
MgO	37.04		36.67	7.51	7.65	8.19	8.02		
CaO	2.5		2.38	10.5	10.07	11.9	10.22		
Na ₂ O	0.05		0.5	3.01	3.28	3.01	2.89		
K_2O	0.01		bdl	0.1	0.18	0.1	0.13		
P_2O_5	$0.01\,$		0.07	0.14	$\rm 0.08$	0.1	0.13		
LOI	8.65		8.22	1.6	3.01	1.23	3.01		
Total	99.78		99.8	99.41	99.49	99.48	99.49		
Sc	11.32	nd	nd	40.26	27.07	nd	27.07		
V	81.31	nd	$^{\rm nd}$	355.65	210.52	nd	210.51		
Cr	2614.01	nd	nd	199.34	396.10	nd	48.95		
Ni	7112.94	nd	nd	nd	nd	nd	nd		
Rb	0.35	0.38	0.69	1.13	2.50	1.89	1.07		
Sr	12.46	12.8	10.58	124.90	200.70	116.94	119.99		
Ba	nd	6.16	$0.01\,$	10.83	8.31	29.82	10.36		
U	0.013^{b}	bdl	0.01	0.04^{b}	nd^b	0.12	0.048		
Th	0.026	0.02	0.03	0.12	0.011	$\rm 0.08$	0.11		
Pb	2.26	nd	0.53	2.72	3.19		66.63		
	0.074	0.11	0.09	32.05	17.28	1.14 1.82	0.82		
Hf									
Zr	2.17	2.97	1.54	1.03	0.69	52.89	22.31		
Ta	0.005	0.01	bdl	0.17	0.023	0.16	0.1		
Nb	0.43	0.66	1.65	2.50	0.38	2.31	1.01		
Y	2.27	2.39	2.68	31.18	14.53	42.18	3.03		
La	0.53	0.51	0.25	3.74	1.19	3.58	4.65		
Ce	0.89	0.87	0.42	12.30	3.87	11.19	12.89		
Pr	0.10	0.1	0.05	2.05	0.68	2.06	2.09		
Nd	0.40	0.43	0.28	10.38	3.97	11.26	10.08		
Sm	0.14	0.15	0.15	3.45	1.37	4.01	3.27		
Eu	0.11	0.11	0.13	1.37	0.56	1.36	1.27		
Gd	0.23	0.26	0.26	4.33	1.84	5.11	4.16		
Tb	0.05	0.05	0.06	0.79	0.34	0.99	0.76		
Dy	0.33	0.35	0.4	5.15	2.31	6.91	5.08		
Ho	0.077	$\rm 0.08$	$0.1\,$	1.12	0.50	1.49	1.09		
Er	0.24	0.26	0.29	3.17	1.47	4.04	3.15		
Tm	nd	nd	nd	2.28	2.21	nd	2.93		
Yb	0.25	0.25	0.31	3.00	1.40	4.10	3.03		
Lu	0.04	0.04	0.05	0.44	0.21	0.63	0.44		
${}^{87}Sr/{}^{86}Sr$	0.704242 ± 13	nd	0.704183 ± 12	0.704019 ± 07	0.703137 ± 06	nd	0.70335 ± 6		
$({}^{87}Sr/{}^{86}Sr)i$ ^c	0.70407		0.70378	0.70396	0.70306		0.703280		
143Nd/144Nd	nd	nd	0.513175 ± 19	0.513160 ± 07	0.513170 ± 13	0.513198 ± 07	0.513191 ± 6		
$(^{143}Nd/^{144}Nd)i^c$			0.51286	0.51296	0.51296	0.51299	0.512998		
εNd_i			$+8.04$	$+10.11$	$+10.14$	$+10.57$	$+10.8$		
$(^{208}Pb/^{204}Pb)i^c$	37.804	nd	38.457	nd	37.705	37.689	nd		
$(207Pb)^{204}Pb)i^c$	15.442		15.524		15.477	15.480			
$(^{206}Pb/^{204}Pb)i^c$	18.126		18.674		18.058	17.941			

Table 2 (continued)

fault system belong to the El Toro unit. The studied samples (97Ce14, 97Ce15) were collected close to the village of Limón Playa (Fig. 2, UTM-WGS84: 17/ 618250E-9604200N). Most olivine porphyroblasts are transformed into serpentine aggregates. Preserved crystals correspond to Mg-rich olivine (Fo90). Fibrous aggregates of colourless amphibole underlie a rather well developed schistosity. Recent observations suggest that the harzburgites of the El Toro unit underwent HP metamorphic conditions (Gabriele and Piccardo, in preparation).

The epidote bearing amphibolites were sampled in the ''Quebrada de Damas'' (97Ce12, UTM-WGS84: 17/625500E-9597500N), 0.8 km W of the village of Zaracay (97Ce18), and in a quarry located 1.8 km W of Zaracay (98TH13, UTM-WGS84: 17/624748E-9596892N). They show an association of actinolite, chlorite and epidote typical of the greenschist facies metamorphism. All samples have a well defined schistosity. An amphibolite (metagabbro) (97Ce13) in which green hornblende porphyroblasts underline a poorly developed cleavage was also collected in the Quebrada de Damas North of the Zaracay village (Fig. 2).

3.2. The metapelites

The metapelites (97Ce2 to 97Ce6) were sampled in the mentioned Río Raspas, where they are closely associated with the eclogites (Fig. 2). The metapelites are medium-to coarse-grained graphite bearing garnetmicaschists, where the eclogitic paragenesis consists of quartz, white mica, garnet, chloritoid, kyanite and rutile. The white mica defines the main schistosity, within which garnet poikiloblasts developed. This latter includes quartz, kyanite, chloritoid and rutile inclusions. Peak metamorphic conditions for the metapelites reached approximately $600\degree$ C and 20 kbar (Gabriele et al., 1999).

4. Analytical procedures

Trace element abundances have been determined by solution- and selective laser ablation (LA)-ICP-MS for amphibole, omphacite, phengite and garnet separates and garnet in thick-section from a mafic eclogite and pelitic schists. LA – ICP –MS analyses were performed at School of Earth Sciences at the Australian National University. Solution trace element analyses were performed by inductively coupled plasma mass spectrometry (ICP-MS) at the Université Joseph Fourier (Grenoble) after acid digestion. Trace elements are spiked with pure Tm using procedures described by Barrat et al. (1996). Standards JB2, WSE, BIR-1, JR1, and UBN were analyzed together with unknown samples. Major element analyses were performed at the Centre de Recherche Pétrographiques et Géochimiques (CRPG, Nancy).

Isotopic data (Sr, Nd and Pb) have been obtained for almost all the lithologies. The $143Nd/144Nd$ and ${}^{87}Sr/{}^{86}Sr$ isotopic compositions were measured on a Finnigan MAT261 multicollector mass spectrometer at the Laboratoire de Géochimie de l'Université Paul Sabatier, using the analytical procedures of Lapierre et al. (1997). The chemical separation of lead was done at the Laboratoire de Géochimie de l'Université de Montpellier II following a procedure modified from Manhès et al. (1978). Total Pb blanks are less than 65 pg for a 100 mg sample. The Pb/Pb isotopic ratios were measured on the MC-ICP-MS P54 at the Ecole Nationale Supérieure de Lyon.

The complete isotopic data set has been corrected for in situ decay assuming an age of 150 Ma, based on the median age of the Amotape –Chaucha Terrane (ACT) lithologies (Litherland et al., 1994).

Analytical data for representative lithologies of the Raspas Metamorphic Complex are presented in Table 2.

Notes to Table 2:

 ε Nd_i calculated with actual $(^{143}Nd)^{144}Nd)_{CHUR} = 0.512638$ and $(^{147}Sm)^{144}$ _{CHUR} = 0.1967 (Wasserburg et al., 1981).

Pb isotopic ratios measured with external precision of ca. 100–150 ppm for the ^{206, 207, 208}Pb/²⁰⁴Pb ratios.

Abbreviations: nd, not determined; bdl, below detection limit.

^a Major and trace elements analyses performed at the Department of Geology and Research School of Earth Sciences, ANU.

^b Trace element analyses (ppm) performed at the Laboratoire des Chaînes Alpines, Univ. Joseph Fou

5. Geochemistry ofrocks of theRaspasMetamorphic Complex

5.1. Eclogites (97Ce1, 98RR4, 98RR6) and garnet amphibolite (97Ce4A)

Eclogites are very fresh rocks $(LOI < 1.4%)$ with homogeneous major and trace element compositions (Table 2). The MgO contents (from 6% to 9.1%) are similar to those of fractionated basalts while the $TiO₂$ and $Fe₂O₃$ contents as well as La/Nb (0.75–0.91) and La/Th $(12.68 - 13.42)$ ratios display enriched MORB signature.

Chondrite-normalised (CN; Sun and McDonough, 1989) patterns for eclogites and mineral separates are presented in Fig. 3A and A'. The REE abundances of these rocks are 20 times the chondritic values and their REE patterns are flat $[0.99 < (La/Yb)_{CN} < 1.29$; $0.88 < (Ce/Sm)_{CN} < 1.13$], suggesting an Oceanic Plateau Basalt (OPB) affinity (Kerr et al., 1997). These eclogites have Eu negative anomalies [0.64 < Eu/ Eu * < 0.88]. 97Ce1 clinopyroxene and barroisite separates have lower REE abundances than the host rock. Clinopyroxene differs from the amphibole by a strong LREE enrichment $[(La/Yb)_{CN} = 5.92]$. Conversely, the amphibole is enriched in HREE, compared to the pyroxene $[(La/Yb)_{CN} = 0.41]$.

Primitive mantle-normalised multi-elements plots of the Raspas eclogites are flat with the exception of large ion lithophile elements (LILE). Pb and Th show positive or negative anomalies depending on the rock lithologies (Fig. 4A). All eclogites exhibit negative Sr anomalies. 97Ce1, which does not contain phengite minerals, is relatively depleted in Rb, Ba and Pb contrasting with the two other eclogites (98RR4 and 98RR6) that are enriched in Rb and Ba. In addition, 98RR6 differs from the other eclogites by higher Pb and $TiO₂$ contents. The absence or presence of phengites in the eclogites often determines the K, Ba and Rb abundances of the rock (Becker et al., 2000).

It is slightly different for Pb where the content of this element in the eclogite depends on more than

one mineral. Pb may substitute for K in phengite, it substitutes for Ca in epidote-group minerals (epidote, zoisite, clinozoisite) and may also replace Na in paragonite. Thus, the multi-element distribution in the eclogites can be substantially influenced by the nature of the protolith, by the mineral assemblage, as well as by the impact of element partitioning of phases with high LILE or HFSE abundances (such as rutile). Similarly, the Nb/U and Ce/Pb of the eclogites are highly variable. The Nb/U (51.1) and Ce/Pb (28.4) of the phengite-free eclogite (97Ce1) approach those of unaltered MORB or OIB (Hofmann et al., 1986). In contrast, these ratios are significantly lower in the phengite-bearing eclogites $[15.0 < Nb/U < 16.5; 3.4 < Ce/Pb < 18.2]$. The U and Nb contents of eclogites may be partly controlled by the occurrence and abundance of specific minerals such as garnet and omphacite (for U^{4+}) and rutile (for U^{6+} and Nb) (Brenan et al., 1994).

Bulk rock major element composition of the garnet-amphibolite (97Ce4A) is similar to that of a cumulate gabbro (Table 2). The REE pattern of this rock has a marked LREE depletion $[(La/Yb)_{CN}]$ $= 0.54$; (Ce/Sm)_{CN} = 0.56] and an Eu positive anomaly $[Eu/Eu^* = 1.5]$ suggesting that the protolith was formed by the accumulation of cpx and plagioclase (Fig. 3A). This is confirmed by the high Al_2O_3 (17.1%) and Cr (543 ppm) contents of this rock (Table 2). Primitive mantle-normalised multi-elements plot of this garnet amphibolite (Fig. 4A) supports this interpretation with a marked depletion in the most incompatible elements (Nb, Ta, Zr, Hf). This metamorphic rock is also depleted in Rb but enriched in Pb and Sr. Positive anomalies for these two latter elements are likely related to plagioclase accumulation, suggesting that high-pressure metamorphism had little effect on the primary chemical composition of the rock.

The Sr isotopic compositions of the garnet-amphibolite (97Ce4A) and eclogite (97Ce1) and their corresponding cpx and amphibole have been measured (Table 2). 97Ce1 whole-rock and cpx have the highest and lowest $({}^{87}Sr/{}^{86}Sr)i$ ratios (0.70649 and 0.70331,

Fig. 3. Chondrite-normalized (CN) rare element patterns of the minerals and host rocks from the Raspas metamorphic rocks. (A) Eclogites and garnet-amphibolite (97Ce4A); (A') mineral separates from the 97Ce1 eclogite; (B) serpentinised harzburgites; (C) greenschists and amphibolites; (D) high-pressure metasedimentary rocks; (D') mineral separates from the metapelites (97Ce3 and 97Ce6).

Fig. 4. Multi-trace-element patterns for the whole rocks from the Raspas metamorphic rocks. (A) Primitive Mantle (PM)-normalised plots for the eclogites and garnet – amphibolite (97Ce4A); (B) (PM) plots for the serpentinised harzburgites; (C) (PM) plots for the amphibolites; (D) Bulk Continental Crust (BCC)-normalised plots for the metapelites.

respectively), while that of the amphibole is slightly higher (0.70391) than the cpx value. $({}^{87}Sr/{}^{86}Sr)i$ of the garnet-amphibolite is significantly lower than that of the 97Ce1 eclogite (0.70519).

Eclogites, mineral separates and garnet-amphibolite εNd_i range from $+4.58$ to $+9.79$ (Table 2). Reported in the $\epsilon N d_i$ versus $({}^{87}Sr)^{86}Sr)i$ (Fig. 5A), they fall in the N-MORB and OIB fields. One out of four of these high-pressure metamorphic rocks

(98RR6) exhibits a significantly lower εNd_i (+4.58). In these rocks, no clear correlation is observed between εNd_i and (La/Yb) ratio (Fig. 5B). In Fig. 5A, the lateral shift observed for the $87\text{Sr}/86\text{Sr}$ ratios of 97Ce1 and 97Ce4A whole rocks, without any decrease of the εNd_i , could be attributed to a hydrothermal alteration or could result from altered oceanic crust assimilation. Initial lead isotopic compositions of these samples are scattered (Table 2) and display

Fig. 5. (A) ϵ Nd_i versus (⁸⁷Sr/⁸⁶Sr)*i* diagram for rocks of the Raspas Metamorphic Complex (this study). The field of the Duarte Complex basalts and cumulates has been reported for comparison (Lapierre et al., 2000); (B) eNdⁱ versus (La/Yb) diagram. La/Yb ratios for E-MORB, N-MORB and Primitive Mantle (PM) are taken from Sun and McDonough (1989).

typical MORB and/or OIB signature. Two samples (98RR6 and 97Ce4A), however, exhibit significantly higher Pb/Pb values (Fig. 6). This feature could be due to contribution of a radiogenic component, which may be assimilated to continental crust or pelagic sediments or to a strong seawater hydrothermal alteration. The lowest lead isotopic composition has been obtained on cpx from the 97Ce1 eclogite, which plots in the N-MORB domain.

5.2. Serpentinised harzburgites (97Ce14, 97Ce15)

The bulk rock compositions of the serpentinised ultramafic rocks are typical to those of harzburgites with very low Al_2O_3 , CaO and TiO₂ contents and very high MgO amounts (Table 2). The high LOI $(> 8\%)$ testifies to the alteration degree of these rocks and to the abundance of serpentinite after olivine and orthopyroxene pseudomorphs. High Cr content (2641 ppm for 97Ce14) is due to the occurrence of preserved Cr-rich chromite spinels (picotite). Chondrite-normalised REE patterns of these ultramafic rocks are characterized by La and Ce enrichments $[1.76 < (La/Md)CN < 2.58]$ and positive Eu anomalies [1.78 < Eu / Eu * < 2.04] (Fig. 3B). The medium to heavy REE are within the range expected for residual ultramafic lithologies remaining after basalt extraction. The LREE and Eu enrichments may reflect a more complex origin but can also be related to late hydrothermal alteration. Ultramafic rocks (primitive mantle normalized) multi-element plots show trace element contents close to the primitive mantle abundances with the exception of Pb, which is slightly enriched in both samples (Fig. 4B). 97Ce15 also exhibits a significant negative Ba anomaly and more limited positive Nb and P anomalies relative to 97Ce14 and the primitive mantle. Both ultramafic rocks possess $({}^{87}\text{Sr})^{86}\text{Sr})$ *i* ratios which fall in the range of the Raspas mafic amphibolites (Table 2, Fig. 5A). εNd_i ratio of 97Ce15 $(+8.04)$ is similar to those of most of the Raspas eclogites but is significantly lower than those of abyssal harzburgites (Snow et al., 1994). In Fig. 5A,B, 97Ce15 (black triangle) plots close to the less hydrothermally altered eclogites. Reported in Fig. 6, the initial lead isotopic ratios of the serpentinised harzburgites plot on or close to the NHRL (Hart, 1984) and fall in the MORB field.

5.3. Greenschists to amphibolite metamorphic rocks (97Ce12, 97Ce13, 97Ce18, 98TH13)

Amphibolites bulk rock major and trace element compositions are similar to those of N-MORB. They are LREE depleted $[0.61 \leq (La/Yb)CN \leq 0.90;$ $0.70 < (Ce/Sm)CN < 0.89$] and relatively flat for medium to heavy REE $[0.70 < (Sm/Yb)CN < 1.85]$ (Fig. 3C). 97Ce13 differs from the metabasalts by lower REE and $TiO₂$ abundances, higher $Al₂O₃$ content and a very slight positive Eu anomaly (Eu/ $Eu^* = 1.09$, suggesting that the protolith was a plagioclase cumulate gabbro. Primitive mantle normal-

ised multi-element plots of these amphibolites confirm their N-MORB affinities (Fig. 4C) with large depletions in LILE, Th, Zr and Hf. Positive Pb, Sr and Eu anomalies in samples 97Ce12, 97Ce13 and 97Ce18 likely reflect plagioclase accumulation, suggesting that these elements were not intensely mobilised during metamorphism. $({}^{87}Sr)^{86}Sr)i$ ratios of these amphibolites do not significantly differ from those of the serpentinised harzburgites of the El Toro unit (Table 2). The metagabbro and metabasalts yield very similar εNd_i ratios $(+10.11 \le \varepsilon Nd_i \le +10.81)$, in the typical range of MORB (Fig. 5A). They are significantly more radiogenic than those of the eclogites and serpentinised harzburgite. In the $\epsilon N d_i$ -(La/Yb)CN diagram (Fig. 5B), these amphibolites exhibit the highest ε Nd_i and the most LREE-depleted patterns of all analyzed rocks from the Raspas Metamorphic Complex. The $({}^{206}Pb/{}^{204}Pb)i$, $({}^{207}Pb/{}^{204}Pb)i$ and $(208Pb/204Pb)i$ ratios of the 98TH13 and 97Ce13 amphibolites are very similar to the 97Ce14 serpentinite and are located in the more depleted part of the MORB domain (Fig. 6).

5.4. Garnet-micaschists (97Ce2, 97Ce3, 97Ce4, 97Ce5, 97Ce6)

Metapelites have very homogeneous major element compositions, close to that of continental crust (Taylor and McLennan, 1985). Differences in $TiO₂$ and MgO contents are likely to be linked to the abundance of oxides. The REECN patterns of these metapelitic rocks are similar to that of the Bulk Continental Crust (BCC), (Taylor and McLennan, 1985; Fig. 3D) with a characteristic LREE-enrichment $[4.27 < (La/Yb)CN < 6.26]$ and a marked negative Eu anomaly $[0.46 < Eu/Eu * < 0.91]$. 97Ce5 differs from the other rocks by a positive Ce anomaly $[Ce/Ce^* = 1.44]$, which could be linked to seawater alteration. 97Ce3 phengite differs from its related host rock by a significantly higher (La/Yb)CN ratio (83.04) (Fig. $3D'$). The difference between the whole-rock and phengite pattern is attributed to the presence of garnet in the bulk rock, which balanced the strong HREE depletion of the phengite. Most of the trace element abundances of these metapelites are similar to those of BCC, as emphasized by the essentially nonfractionated Bulk Continental Crust normalized-patterns reported in Fig. 4D. However,

Fig. 6. Correlation plots of (A) ²⁰⁶Pb/²⁰⁴Pb versus ²⁰⁷Pb/²⁰⁴Pb and (B) ²⁰⁶Pb/²⁰⁴Pb versus ²⁰⁸Pb/²⁰⁴Pb for rocks of the Raspas Metamorphic Complex. Data sources for OIBs (Hart et al., 1986; Stille et al., 1986; White et al., 1993); St-Helena (Newson et al., 1986); Tubuaii (Vidal et al., 1984), Lesser Antilles (Davidson, 1987), Canary Islands (Hoernle and Tilton, 1991); MORB (White et al., 1987; Mahoney et al., 1989; Staudigel et al., 1991), Modern marine sediments (Ben Othman et al., 1989); NHRL, Northern Hemisphere Reference Line (Hart, 1984).

abundances of Ba, Th, Zr and Hf differ from one rock to another (Fig. 4D). Relative to BCC, 97Ce2 is enriched in Ba and Th while 97Ce5 and 97Ce3 are strongly depleted. 97Ce6 and 97Ce3 are enriched and depleted in Zr and Hf, respectively. Losses in Sr affect all the metapelites. The $({}^{87}\text{Sr})^{86}\text{Sr})i$, εNd_i and $(206, 207, 208Pb/204Pb)i$ ratios of the metapelites are homogeneous and typical of the continental crust (Figs. 5A and 6). The negative εNd_i values (< -6.3) associated with both high $(^{87}Sr)^{86}Sr)i$ (> 0.7137) and $(^{207}Pb/^{204}Pb)i$ ratios (> 15.68) indicate that these metapelites were derived from the erosion of an old crustal component and/or sediments.

6. Discussion

6.1. Main rock groups of the Raspas Metamorphic **Complex**

Field work, geochemical analyses and preliminary petrographic studies lead to distinguish three main rock groups within the Raspas Metamorphic Complex.

. Mafic HP metamorphic rocks include the eclogites and probably the serpentinised peridotites. Geochemical data indicate that these rocks likely represent remnants of a crustal section of an oceanic plateau (OPB) with the serpentinised peridotites corresponding to the residual part. No MORB-type rocks were found in this sample set.

. Metapelites, metamorphosed under HP conditions, were originally terrigenous sediments resulting from the erosion of an old continental crust. They are structurally associated with the eclogites and current petrographic studies suggest that both rock lithologies underwent comparable tectonic and $P-T$ evolutions.

. Epidote-bearing amphibolites were metamorphosed into greenschist to amphibolite facies (Río Panupali unit). These rocks display geochemical features indicating oceanic floor (MORB) affinities. No rocks with oceanic plateau characteristics occur in this unit.

Thus, the Raspas Metamorphic Complex exhibits a well defined internal structure, with independent tectono-metamorphic units, each one characterized by distinctive lithologies, metamorphic evolution and chemical affinities. This does not support the interpretation of a mere "tectonic mélange". Since these tectono-metamorphic units are bounded by tectonic contacts, the present-day structural organization of the Raspas Metamorphic Complex is subsequent to the peak metamorphic conditions, and, therefore, results from its exhumation history. Additionally, this postmetamorphic tectonic evolution obscured the primary relationships between the distinct units and makes it difficult to propose a geodynamic evolution at this stage of the study.

6.2. Possible accretion processes

Because of their overthickened crust (15–40 km), oceanic plateaus are considered to be too buoyant to subduct and are more easily accreted or even obducted onto the continental margin (Ben Avraham et al., 1981; Cloos, 1993; Abbott et al., 1997). In the Raspas Metamorphic Complex, oceanic plateau basalts underwent eclogite facies metamorphism, indicating that part of the oceanic plateau can be subducted down to 40 –60 km depth. However, we can reasonably consider that only a rather thin edge of the oceanic plateau, formed by basaltic flows and shallow level gabbroic

intrusions, has been subducted. When the thickest part of the oceanic plateau entered the subduction zone, it probably jammed the subduction and was accreted, thus leading to the westward jump or flip of the subduction zone.

The association of sediments of continental origin (metapelites) with remnants of oceanic plateau rocks (eclogites and peridotites), both metamorphosed under comparable high-pressure conditions, suggests that they underwent the same burial history. As the geochemical features of the metapelites clearly indicate that they were derived from the erosion of an old continental crust, they cannot constitute the submarine stratigraphic cover of the oceanic plateau. The subduction zone usually acts as a sediment trap that inhibits the bypass of products deriving from the continent onto the oceanic floor before the latter enters the subduction zone. On the other hand, the oceanic plateau is a submarine relief, on top of which no detrital sediments can be deposited, except—possibly—products of its own erosion.

Therefore, the continent derived metapelites were probably originally deposited at the foot of the Andean continental margin and structured as a classical sedimentary accretionary prism, such as those known in present-day accretionary margins (Von Huene and Scholl, 1991; Lallemand, 1999). As the edge of the oceanic plateau entered the subduction zone, it dragged down the accretionary wedge, which was buried along the Wadatti-Benioff zone, until the main body of the oceanic plateau jammed the subduction process.

In the Raspas Metamorphic Complex, rocks of N-MORB affinities metamorphosed under greenschist to amphibolite facies conditions are presently tectonically juxtaposed with oceanic plateau rocks showing HP assemblages. Such a tectonic framework seems to characterize the latest Jurassic –Early Cretaceous suture of northwestern South America. The Peltetec suture of the eastern Cordillera of Ecuador contains Jurassic lavas and greywackes of N-MORB and arc affinities, metamorphosed under greenschist facies during the latest Jurassic –Early Cretaceous (Litherland et al., 1994). In the suture zone of Colombia, most blueschist to amphibolite facies metamorphic rocks, dated between 129 and 110 Ma (Aspden and McCourt, 1986; Toussaint and Restrepo, 1994), are interpreted as remnants of oceanic plateau (Kerr et al., 1997).

If these two rock types had undergone the same geodynamic history, then we would expect an HP metamorphic overprint in the N-MORB rocks and a comparable or lower metamorphic grade in the oceanic plateau rocks, since the latter are theoretically less easily subducted than the former. This situation is not observed in the presently known outcrops ascribed to the latest Jurassic –Early Cretaceous suture. Therefore, the present-day Raspas-Peltetec-Colombian suture may represent a composite and/or polyphase feature involving at least two distinct large-scale tectono-metamorphic units, reflecting distinct geodynamic evolutions and significances.

6.3. Comparison with the northern Andes and Caribbean

Due to the northward component of the oceanic convergence in Late Cretaceous and Tertiary times (Pardo-Casas and Molnar, 1987) and to the evidences of large-scale dextral strike-slip movements of Tertiary age in northwestern South America (Freymuller et al., 1993), remnants of Jurassic oceanic terranes, possibly equivalent to the Raspas mafic rocks, must be sought in the northern Andes and around the Caribbean plate.

In the suture zone of Colombia, remnants of oceanic plateau locally underwent HP metamorphism in the Early Cretaceous and, therefore, are of Jurassic age (Aspden and McCourt, 1986; Toussaint and Restrepo, 1994; Nivia, 1996; Kerr et al., 1997).

In Venezuela, the Siquisique ophiolites of Middle Jurassic age (Bartok et al., 1985) are considered petrologically comparable to the oceanic plateau assemblages of the Curac ao island (Donnelly et al., 1990), while Beets et al. (1984) postulate that the Margarita HP metabasic rocks (Sisson et al., 1997) are tholeiites of Late Jurassic age, possibly of MORB affinity. The Aptian –Albian La Rinconada unit (Margarita island, Venezuela) comprises HP gneisses and eclogites of MORB chemistry metamorphosed at ca. 500 $^{\circ}$ C and 12-14 kbar (Stockhert et al., 1995). The tectonic evolution of this complex requires a Pacific origin for the Caribbean plate. In the Lower Unit of the Santa Elena Peninsula (Costa Rica), Early Jurassic to Middle Cretaceous alkaline rocks (De Wever et al., 1985) exhibit OIB-like geochemical characteristics (Unit I, Hauff et al., 2000) and are associated with an Early Cretaceous island arc suite (Astorga, 1997; Hauff et al., 2000).

In Hispaniola and Puerto Rico, pillow basalts are interbedded with ribbon radiolarian cherts of Late Jurassic –Early Cretaceous age and exhibit a Pacific affinity (Mattson and Pessagno, 1979; Montgomery et al., 1994). In Hispaniola and possibly in Puerto Rico, these MORB-type basalts probably emplaced near a hotspot (Donnelly et al., 1990; Lapierre et al., 1999).

The Raspas Metamorphic Complex may represent the southernmost outcrop of these Jurassic oceanic thickened crustal fragments scattered around the Cretaceous Caribbean Plate, since some of them present well documented oceanic plateau affinities or plume activity related signatures (Colombia, Costa Rica, Hispaniola).

7. Conclusions

. Geochemical and petrographic study of the Raspas Metamorphic Complex allowed us to identify several distinct litho-tectonic units: (1) HP mafic and ultramafic rocks with oceanic plateau affinities, (2) HP metapelites derived from an old continental crust, and (3) greenschist –amphibolite basalts and gabbros of N-MORB origin.

. Therefore, the Raspas Metamorphic Complex exhibits tectono-metamorphic units characterized by distinct lithologies, geochemical affinities and metamorphic conditions. Such a well defined internal structure does not support the interpretation of an accretionary wedge or a "tectonic mélange".

. The evolution of the HP metamorphic rocks of the Raspas Metamorphic Complex may be explained by the dragging down of continent-derived sediments accumulated into an accretionary wedge as the edge of the oceanic plateau entered the subduction zone. Subsequently, the main body of the oceanic plateau stopped the subduction processes. Scattered Jurassic oceanic remnants, locally with plume-source affinity, known from northern South America and the Caribbean, may represent parts of this disrupted oceanic terrane.

. As for most Late Cretaceous –Paleogene accretions in Ecuador (Jaillard et al., 1997; Reynaud et al.,

1999), the accretion of the Raspas oceanic plateau occurred without obduction. In this model, underthrusting of parts of accreted oceanic terranes might contribute to the crustal thickening of the upper plate in this part of the Andes, as already suggested by Megard (1987, 1989) in Ecuador and further documented in Colombia by Weber et al. (2002).

. Along the latest Jurassic –Early Cretaceous suture of northwestern South America, oceanic plateau exhibit HP assemblages, whereas arc or N-MORB rocks show lower metamorphic grades. This suggests that this suture involves at least two distinct oceanic terranes characterized by distinct geodynamic evolution.

Acknowledgements

This work was funded by a CNRS-INSU IT Programm, contribution no. 263, by the IRD, CSS1, and by the SNSF grant no. 20-50812.97. We would like to thank P. Telouk and the ''Service Commun National du MC-ICPMS de l'Ecole Normale de Lyon'', as well as P. Brunet (UMR-5563), who carried out all the Sr– Nd TIMS analyses. The authors are grateful to M. Ballèvre (Univ. Rennes) and to G. Martinotti (Univ. Turin) for their invaluable help in petrographic study and field work, respectively. Luis Aguirre and Andrew Kerr are thanked for their constructive and helpful reviews of the manuscript.

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