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2 **Landslides induced by the 2017 Mw7.3 Sarpol-Zahab earthquake**  
3 **(Iran)**

4  
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7  
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16 **Abstract.** Landslides are the main secondary effects of earthquakes in mountainous areas. The spatial distribution of these  
17 landslides is controlled by the seismic ground motion and the local slope stability. While gravitational instabilities in arid  
18 and semi-arid environments are understudied, we document the landslides triggered by the Sarpol-Zahab earthquake  
19 (November 12, 2017, Mw7.3, Iran/Iraq border), the largest event ever recorded in the semi-arid Zagros Mountains. An  
20 original earthquake-induced landslide inventory was derived, encompassing landslides of various sizes and velocities  
21 (from rapid disrupted rockfalls to slow-moving coherent landslides). This inventory confirms the low level of triggered  
22 landslides in semi-arid environments. It also displays clear differences in the spatial and volumetric distributions of  
23 earthquake-induced landslides, having a 386 rockfalls of limited size triggered around the epicenter, and 9 giant (areas of  
24 ca. 10<sup>6</sup> m<sup>2</sup>) active and ancient rockslides coseismically accelerated at locations up to 180 km from the epicenter. This

25 unusual distant triggering is discussed and interpreted as an interaction between the earthquake source properties and the  
26 local geological conditions, emphasizing the key role of seismic ground motion variability at short spatial scales in  
27 triggering landslides. Finally, the study documents the kinematics of slow-moving ancient landslides accelerated by  
28 earthquakes, and opens up new perspectives for studying landslide triggering over different time-scales.

## 29 **1 Introduction**

30 Slope failures are one of the main secondary effects of earthquakes (Marano et al., 2010), with a large part of co- and  
31 post-seismic damage caused by landslides in mountainous areas (Fan et al., 2019; Keefer, 2002; Marc et al., 2015). The  
32 spatial distributions of earthquake-induced landslides highlight trigger parameters related to the earthquake parameters  
33 such as the moment magnitude and proximity to the fault (Rodríguez et al. 1999; Keefer 2002), the position in the  
34 hanging/footwall of faults (Gorum et al., 2011), or the fault rupture kinematics (Gorum et al., 2014; Specht et al., 2019).  
35 Trigger parameters are also related to the site conditions such as the geology (Bozzano et al., 2008; Roback et al., 2018),  
36 the groundwater (Watkinson and Hall, 2019), or the topography (Meunier et al., 2008; Lacroix et al., 2013). Complex  
37 interactions between these different factors can generate variability in seismic ground motions across a range of spatial  
38 scales, which is a key factor influencing landslides spatial distribution in mountainous areas. Subsequent questions arise  
39 on the characteristics of the seismic ground motion at the origin of the triggered landslide in relation to the corresponding  
40 mechanical behavior of soils and rocks under dynamic loading (e.g. frequency content, amplitudes, duration, Harp and  
41 Wilson, 1995; Meunier et al., 2007; Tataro et al., 2010; Harp et al., 2014; Specht et al., 2019; Valagussa et al., 2019).  
42 Landslides may also be induced at unexpectedly large distances from the epicenter, in relation to slope susceptibility, site  
43 effects, and/or antecedent rain (Delgado et al., 2011). These complexities make it difficult to quantify the controlling  
44 factors of earthquake-triggered landslides, which still remains an open question.

45 In the literature, a recent successful strategy to address this issue has been to compile and analyze all available coseismic  
46 landslide inventories (Tanyas et al., 2017; Marc et al., 2016; Tanyas et al., 2019), albeit with a clear under-representation  
47 of arid and semi-arid mountainous regions (e.g. Central Andes, Zagros, Tien Shan, Africa, Central Mountain Ranges),  
48 even though most of these areas have high seismic activity. This bias can be explained by the low density of population  
49 in these regions and the lower level of landslide triggering compared to other tropical or temperate areas for similar  
50 earthquake magnitudes (Lacroix et al., 2013, Barlow et al., 2015). Indeed, while shallow disrupted slides in soils and  
51 colluvial deposits highly predominate during earthquakes (Harp and Jibson, 1995; Wang et al., 2002; Keefer, 2002) the  
52 scarcity of rainfall in arid and semi-arid regions limits weathering and tends to generate hard, thin and dry soils with low  
53 organic matter and clay content (Singer., 1991; Dregne, 1976). These properties make soils less sensitive to seismic

54 shaking, generating fewer shallow disrupted slides in soils and colluvial deposits, which usually predominate during  
55 earthquakes in wet conditions (Harp and Jibson, 1995; Wang et al., 2002; Keefer, 2002). Arid and semi-arid regions are,  
56 however, of great interest for the study of factors controlling earthquake-triggered landslides, because the interfering  
57 effect of rainfall occurs little or not at all and site effects should be limited.

58 This study analyzes the distribution of landslides triggered by the Sarpol-Zahab earthquake (Mw7.3, Iran) that occurred  
59 on November 12th 2017 in the semi-arid (mean precipitation of 230 mm/yr) northwest region of the Zagros Mountains  
60 (Fig. 1). This major earthquake, associated with the rupture of a low angle blind thrust fault at depths of 15-20 km  
61 (Gombert et al., 2019), occurred at the end of the dry season in an area that encompasses many giant paleo-landslides, of  
62 volumes up to 30 km<sup>3</sup> (Ghazipour and Simpson, 2016). Following this earthquake, a few coseismic landslides of various  
63 types (debris fall, boulder/rock fall) were reported near the epicenter (Miyajima et al., 2018; Vajedian et al., 2018). The  
64 earthquake also reactivated the giant Mela-Kabod landslide (4-km-long, 1-km-wide) located ~40 km south of the  
65 epicenter (Vajedian et al., 2018; Goorabi, 2020). To our knowledge, no reactivation was reported at other giant landslides  
66 mapped by Ghazipour and Simpson (2016) in the region.

67 Accordingly, this study, which is based on an exhaustive inventory of the induced landslides, aims at understanding the  
68 factors controlling the landslide triggering during this earthquake. Landslide identification relies on a combination of  
69 remote sensing techniques, with the objective of detecting landslides of various types and sizes. Following Keefer (1984),  
70 we distinguish two main categories: coherent and disrupted slides. A disrupted landslide consists of many semi-  
71 independent units or granular flow. Co-seismic disrupted landslides generate scars and/or deposits that are mapped by  
72 change detection methods using pre/post- earthquake optical images (Guzzetti et al., 2012). The detection of coherent  
73 slow-moving landslides, activated or accelerated by the earthquake, was based on the complementary use of optical image  
74 correlation and InSAR time-series analysis.

75 This study reveals an original distribution pattern of the co-seismic landslides in a semi-arid region, which sheds new  
76 light on (1) the earthquake-induced effect on gravitational movements at large distances, (2) the role of seismic  
77 amplification, and (3) the co-seismic and post-seismic landslide deformation.

78

## 79 2 Study site

### 80 2.1 Geology and landslides of the NW Zagros Mountains

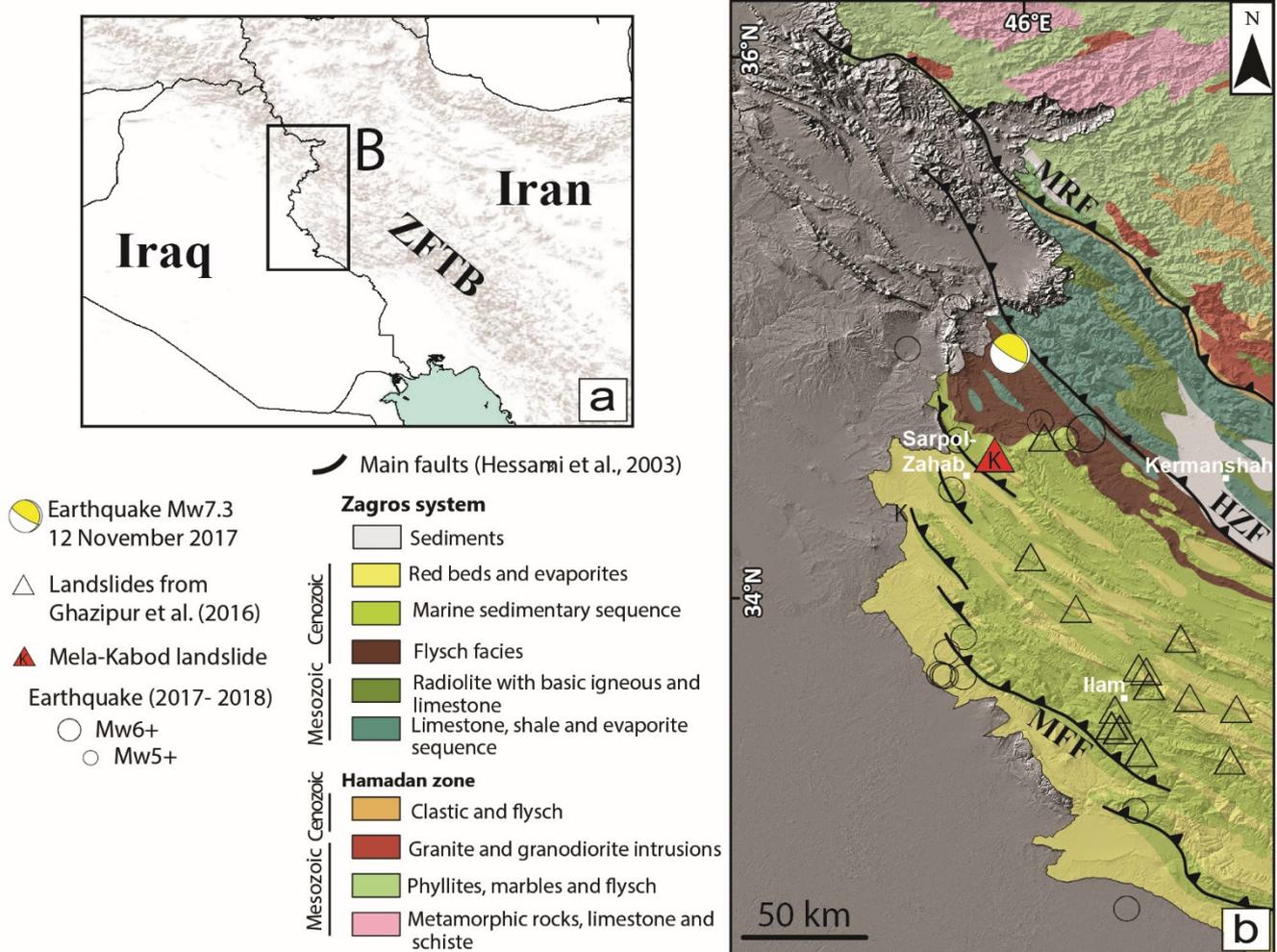
81 The study area is located in Western Iran, close to the Iraq border, in the local provinces of Kermanshah and Ilam, at the  
82 northwest extremity of the Zagros Mountains (Fig. 1a). The Zagros fold and thrust belt (ZFTB) formed in response to the  
83 collision between the Arabian and Eurasian plates, which initiated at ~35 Ma (McQuarrie and al., 2003) and continues  
84 today with a convergence rate of 8-23 mm/yr (Masson et al., 2014). This ~N-S convergence produced a zone of thin-  
85 skinned deformation with a succession of asymmetrical, NW-trending, inverted folds affecting a 7-12 km thick pile of  
86 Palaeozoic and Mesozoic to Cenozoic sedimentary rocks, comprising limestone, siltstone, conglomerate, shale and salt  
87 (Fig. 1b). For a detailed geological log, the reader can refer to Alavi (2007). The shortening was accommodated along the  
88 northwest-trending major thrust faults, which include the Main Recent Fault (MRF; bounding the Zagros to the northeast),  
89 the High Zagros Fault (HZF; in the central Zagros), and the Mountain Front Fault (MFF; bounding the belt to the SW) -  
90 see Fig. 1b. Regional convergence also implied thick-skin shortening, as most of these faults initiate in the crystalline  
91 Arabian basement through the reactivation of older normal faults (Barnhart et al., 2018). The northern part of the ZFTB  
92 hosts moderate seismicity ( $M=5-6$ ) with shallow epicenter depths ranging from 4 to 20 km (Talebian and Jackson 2004;  
93 Gombert et al. 2019). The last two significant historical earthquakes in the region ( $M_w$  5.9 and 6.4) occurred a thousand  
94 years ago (Ambraseys and Melville 2005). The 2017 Sarpol-Zahab earthquake ( $M_w$  7.3) is the largest event ever recorded  
95 in the Zagros mountains.

96 The topography of the northern ZFTB is strongly controlled by large wavelength folds affecting a succession of stiff  
97 calcareous units (Sarvak, Ilam and Asmari Formations), shales (Surgah, Gurpi Formations), or conglomerates (Kashkan  
98 Formation), forming ridges and valleys along the anticline and syncline axes, respectively (Alavi, 2007; Homke et al.,  
99 2009; Saura et al., 2011). Erosion and fluvial incision have carved this geological structure, revealing the contrast in  
100 competence between the stratigraphic units through morphological features such as typical cuesta landforms, breached  
101 anticlines, and secondary valleys that developed in less competent units.

102 This semi-arid zone receives about 230 mm/yr of annual rainfall, which occurs mostly between November and May.  
103 Despite the semi-arid conditions, several giant landslides were reported in the ZFTB, including the Seymareh rockslide-  
104 rock avalanche that was probably seismically triggered ten thousand years ago (Roberts and Evans 2013). The latter is  
105 considered to be the largest terrestrial slope failure (volume 44 km<sup>3</sup>) ever known on Earth (Delchiaro et al. 2019).  
106 Ghazipour and Simpson (2016) provide a general inventory of large landslides for the ZFTB based on a review of  
107 geological mapping, complemented by manual mapping with Google Earth and Digital Elevation Models (DEMs) and

108 were able to identify about three hundred large paleo landslides with areas ranging between  $10^3$  m<sup>2</sup> and  $10^8$  m<sup>2</sup>. The  
 109 landslides reported by Ghazipour and Simpson (2016) and located in our study area are plotted in Fig. 1b. Nonetheless,  
 110 Ghazipour & Simpson didn't give clear explanation for the existence of these slides or showed signs of recent activities.

111



112 **Fig. 1** (a) Study area location. (b) Geological map of the Northwestern Zagros adapted from the 1:2,500,000 tectonic map  
 113 of the National Iranian Oil Company (Huber, 1976). The yellow beachball is the moment tensor estimated by the US  
 114 Geological survey (USGS) for the 12 November 2017 earthquake. Earthquakes (Mw>5) are also reported from the USGS  
 115 for the period of 2017-2018 and presented with empty circles. The empty triangles correspond to the giant landslides  
 116 mapped by Ghazipur and Simpson (2016) while the red triangle shows the location of the Mela-Kabod landslide (Vajedian  
 117 et al., 2018; Goorabi, 2020). The main faults are from Hessami et al. (2003). MRF: Main Recent Fault; HZF: High Zagros  
 118 Fault; MFF: Main Frontal Fault.

119

## 120 **2.2 The 2017 Sarpol Zahab earthquake and coseismic induced landslides**

121 The Mw7.3 Sarpol-Zahab earthquake occurred on the 12/11/2017. The EQ epicenter was located at ca. 35°N – 46°E,  
122 about 60 km north to Sarpol-Zahab city, close to the lineament of the HZF fault (Fig. 1b). It killed 620 people and caused  
123 considerable damage, especially in the city of Sarpol-Zahab (Miyajima et al., 2018). Several recent studies (e.g. Barnhart  
124 et al., 2018; Chen et al., 2018; Nissen et al., 2019; Gombert et al., 2019) investigated the earthquake source properties,  
125 and fault rupture process, using a combination of radar interferometry (InSAR) and seismology (body waveform  
126 inversion), and discussed the tectonic context of this earthquake. These studies reached mainly the following conclusions:  
127 (1) the fault rupture did not reach the surface, the earthquake occurred along a blind thrust fault dipping (east) 10-20° and  
128 located at a depth of about 14-20 km in the crystalline basement; (2) afterslip may have propagated aseismically up-dip  
129 of the earthquake epicenter onto a nearly horizontal basal decollement of the thin-skinned ZFTB; (3) the fault rupture was  
130 about 50 km long and 30 km wide with a maximum coseismic slip of  $5.5 \pm 0.5$  m; (4) a high impulsive source with a  
131 southward rupture directivity is inferred, and produced the largest peak ground acceleration ( $PGA = 0.700 \text{ cm/s}^2$ ) in the  
132 city of Sarpol-Zahab (~60 km south of the epicenter); (5) high horizontal peak ground accelerations (~100  $\text{cm/s}^2$ ) were  
133 also recorded up to 150 km south of the source (Babaie Mahani and Kazemian, 2018).

134 Miyajima et al. (2018) reported several coseismic rockfalls and landslides that affected roads and houses around the  
135 Sarpol Zahab town, including the triggering of a massive landslide (4×1 km) named Mela-Kabod and located a few  
136 kilometers south-east of Sarpol Zahab (Fig. 1b). Coseismic interferometric measurements reveal decorrelation along or  
137 close to pre-existing faults, suggesting surface processes (Vajedian et al., 2018) that were not precisely mapped. Authors  
138 also estimated that the surface motions at the Mela-Kabod landslide reached over 30 and 10 meters in the horizontal and  
139 vertical directions, respectively (Vajedian et al., 2018; Goorabi, 2020). However, to our knowledge, no comprehensive  
140 inventory of Sarpol Zahab earthquake-induced landslides has been made, and the possible coseismic effects on the many  
141 large paleo landslides identified in this region have not been studied.

142

## 143 **3 Data and Methods**

144 The working strategy aimed at detecting the maximum possible number of earthquake-induced landslides in the study  
145 area, extending more than 200 kilometers along the Iran-Iraq border using different remote sensing methods (Fig. 2). First,  
146 the scars and deposits of disrupted coseismic landslides were mapped by a visual comparison of pre- and post-seismic  
147 Planet satellite images. Second, the slow-moving coherent landslides (velocities ranging from m/yr to mm/yr) were

148 detected by deriving the ground deformation from optical image correlation (Planet and Spot6/7 images) and from radar  
 149 satellite images (Sentinel-1 Interferometric Synthetic Aperture Radar) for the co-seismic period. Finally, once slow-  
 150 moving landslide sites were identified and mapped, their kinematics was studied by deriving time-series of landslide  
 151 displacements over a 20-month period surrounding the Sarpol Zahab earthquake from InSAR data (Table 1).

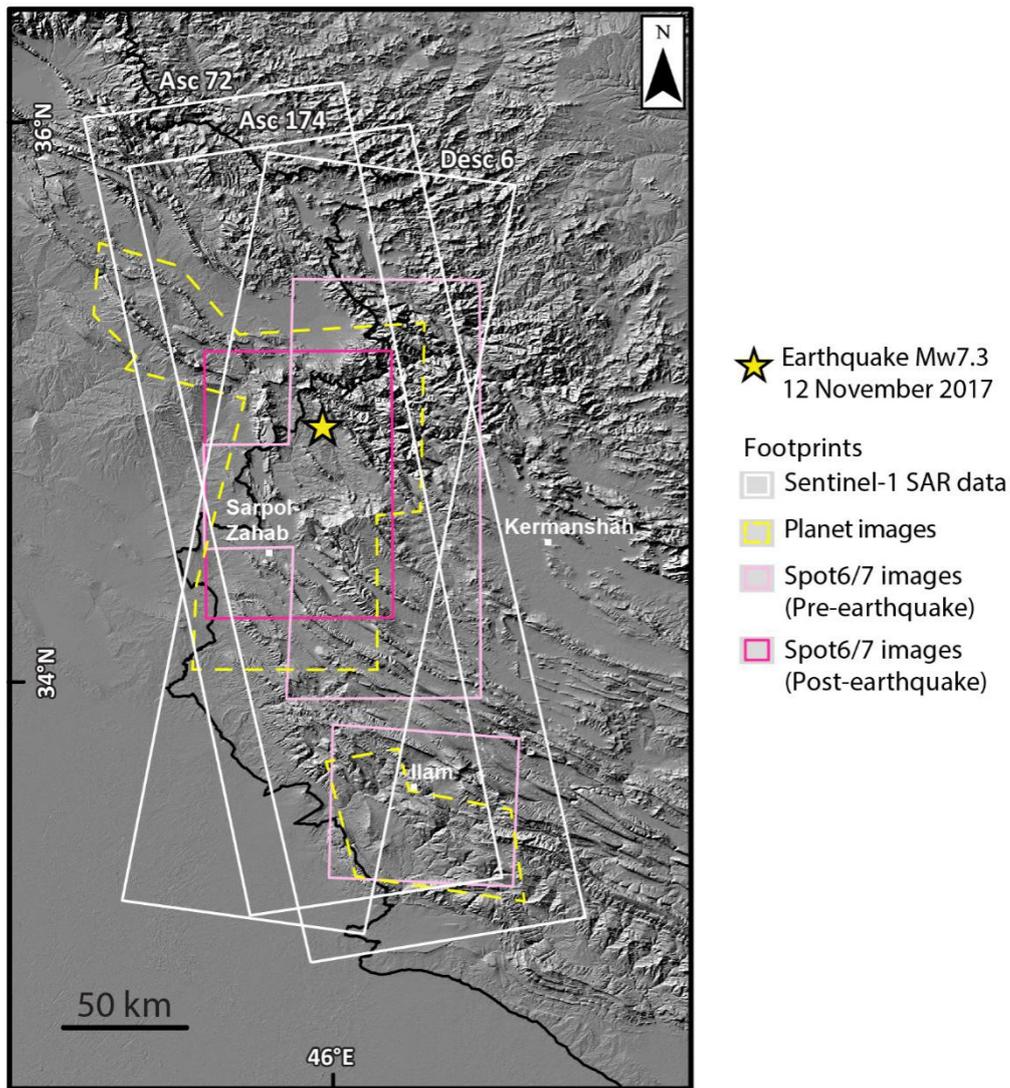
152

Data type / satellites (resolution)		Dates of Acquisition	
		Pre-earthquake	Post-earthquake
Optical	Planet (3-4 m)	19-Oct-2017 (N*) 07-Nov-2017 (S*)	13-Nov-2017 (N*) 17-Nov-2017 (S*)
	SPOT6/7 (1.5 m)	13-Oct-2013 (N*), 24-Apr-2014 (N*), 04-May-2014 (N*), 14-Aug-2014 (N*), 09-Nov-2014 (S*)	29-Nov-2017 (N*), 12-Dec-2017 (N*)
Radar	Sentinel-1 (~4x15 m)	10-Jan-2017 to 06-Nov-2017 (acquisition every 12 days)	12-Nov-2017 to 27-Aug-2018 (acquisition every 6 days)

153

154 **Table.1** Synthesis of satellite data used in this study and their characteristics. N\* and S\* correspond respectively to the  
 155 areas around the epicenter and in the Southern part.

156



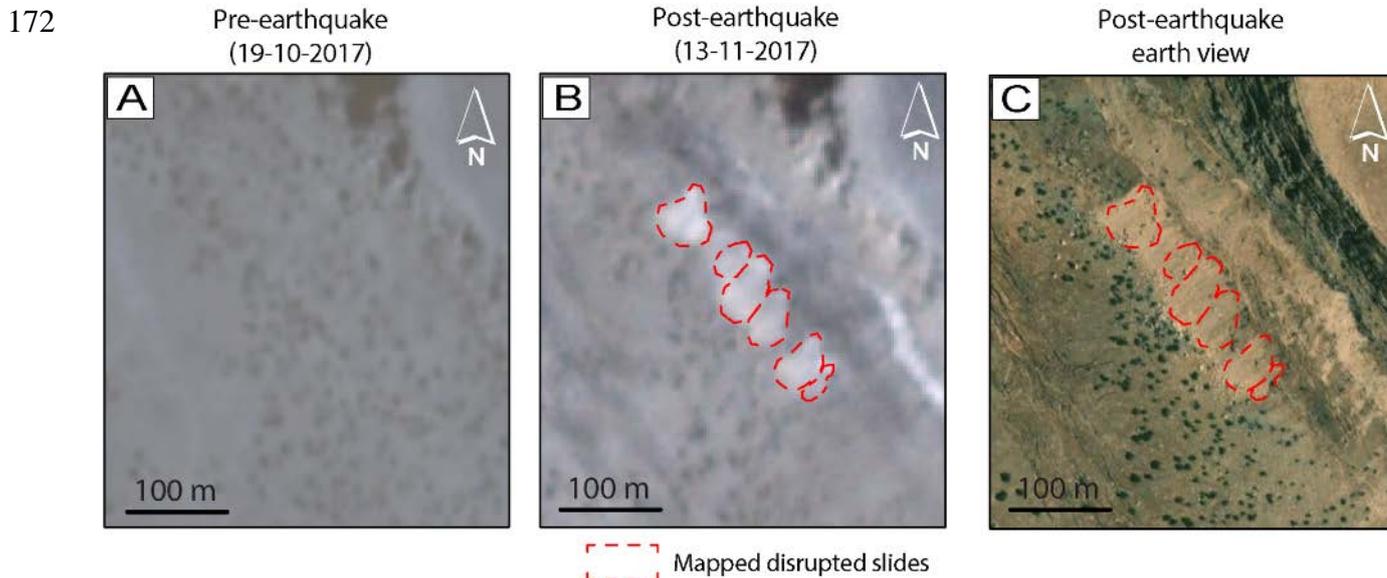
158 **Fig. 2** Footprints of the satellite data used in this study. See Table 1 for the data characteristics.

159

### 160 3.1 Inventory of disrupted landslides

161 A detailed inventory of coseismic disrupted landslides was conducted by the visual comparison of multispectral Planet  
 162 satellite images (3-4 m resolution) acquired before and after the earthquake (19/10/2017 and 13/11/2017) covering an  
 163 area of 12000 km<sup>2</sup> centered on the epicenter (see footprint on Fig. 2). The available pre- and post-earthquake orthorectified  
 164 Planet images were compared and analyzed in ArcGIS software using available toolsets, DEMs, and Earth view base  
 165 maps. The analysis was complemented by geological maps of unequal resolution: the 1:250,000 Ilam province geological  
 166 map from Llewellyn (1974) and the 1: 2,500,000 Iran tectonic map from the National Iranian Oil Company (Huber, 1976).  
 167 As shown, for instance, on Fig. 3, we typically detected and mapped new rockfall scars and their associated debris deposits

168 induced by the earthquake as polygons. To extract the slope of each detected disrupted landslide, a slope map was derived  
169 from the high-resolution post-earthquake DEM using arcmap slope (3D analysis) tool (see section 3.2.1). We ascribed to  
170 each landslide the steepest slope found within a buffer zone of 50 m around the highest altitude point (i.e. headscarp) of  
171 each mapped polygon.



173 **Fig. 3** Typical example of coseismic disrupted landslide deposits mapped from Planet imagery. (a) pre and (b) post Sarpol-  
174 Zahab earthquake Planet imagery showing the same area. (c) Post-earthquake Google Earth view.

175

## 176 3.2 Inventory of coherent slow-moving landslides

### 177 3.2.1 High resolution DEMs

178 Two 4 m resolution pre- and post-earthquake DEMs were generated in the region of the epicenter (Fig. 2) from tri-stereo  
179 pairs of SPOT6/7 images, acquired in 2014 and November-December 2017 (Table 1). Ames Stereo Pipeline (Beyer et al.,  
180 2018) was used for the DEM generation, following the iterative method proposed by Lacroix (2016). A 4 m resolution  
181 pre-earthquake DEM was similarly generated with Ames Stereo Pipeline using tri-stereo SPOT7 images (acquired in  
182 2014, Table 1) covering the southern part of the study area (Fig. 2). These DEMs were used both for geomorphological  
183 purposes and for precise orthorectification of optical images (see following section).

184

### 185 3.2.2 Correlation of optical images

186 The earthquake-induced horizontal ground displacement was calculated from the correlation of pre/post-earthquake  
187 orthorectified optical satellite images (Leprince et al., 2007), using SPOT6/7 images for the northern area, and Planet  
188 satellite images for the southern area where immediate post-earthquake SPOT6/7 images were not available (Table 1).  
189 The pre- and post-earthquake panchromatic bands of the SPOT6/7 images from the northern area (Fig. 2) were first  
190 orthorectified using the high resolution concurrent DEMs and then mosaicked. The COSI-Corr phase correlator was then  
191 used to measure the horizontal ground motion (Leprince et al., 2007). The correlation was conducted in the frequency  
192 domain using a sliding window of 64×64 pixels, which was chosen after several trials to maximize the signal to noise  
193 ratio on the Mela-Kabod landslide. In a similar manner, the Planet images were correlated (Fig. 2). For that, the green  
194 bands were utilized for correlation due to artifacts observed in the other bands.

195 The EW and NS displacement fields were post-processed to correct them from classical artifacts (e.g. Leprince, et al.,  
196 2007; Bontemps et al., 2018): a 2D linear ramp was removed from the data to detrend them and pixels of low correlation  
197 coefficient ( $CC < 0.65$  for SPOT6/7 and  $CC < 0.7$  for Planet images) were masked. Finally, to eliminate a potential constant  
198 bias of georeferencing, the median of the displacement in each direction was removed from the entire EW and NS  
199 displacement fields (e.g. Lacroix et al. 2018).

200 These displacement fields were then inspected visually to detect landslide-like patterns of displacement with magnitudes  
201 larger than the uncertainty of the displacement (calculated as the standard deviation of the EW and NS components over  
202 the whole area), oriented in the direction of the slope. The detected patterns are then plotted on the google-earth image to  
203 validate the landslide origin of this signal.

204

### 205 3.2.3 InSAR processing

206 InSAR data were used to detect and monitor slow coherent landslides. We derived ground motion time-series from InSAR  
207 using three Sentinel-1 SAR tracks spanning a 20-month time period centered on the earthquake (Table 1) with acquisitions  
208 of 12 and 6 days before and after the earthquake respectively. differential interferograms were generated using the NSBAS  
209 (New Small BAseline Subset) (Doin et al., 2011) processing chain based on the ROI\_PAC software (Rosen et al., 2004).  
210 Two ascending (tracks 174 and 72, subswath iw1 for both) and one descending (track 6, subswath iw2) tracks of Sentinel  
211 1A and 1B covering an area of 33500 km<sup>2</sup> (see footprint on Fig. 2 and Table 1) were used for this purpose. Initially, all  
212 secondary SLC (Single Look Complex) images in a single reference SLC geometry were resampled and co-registered to

213 reference using precise orbits and an ASTER digital elevation model (30 m resolution), combined with empirical offsets  
214 between secondary and reference images. Then, a small baseline subset was defined using temporal and perpendicular  
215 baseline constraints (Doin, et al., 2011). After calculating differential interferograms we corrected them from atmospheric  
216 delays using ERA-5 ECMWF (European Center for Median-Range Weather Forecast) reanalysis (Doin et al., 2009).  
217 Finally, we made an empirical correction for topographically-correlated atmospheric-delays (Doin et al., 2011).

218 The coseismic interferograms were then inspected for landslide-like patterns characterized by InSAR fringes with sharp  
219 limits delineating spatially small anomalies located on slopes (e.g. Dini et al., 2019). Special attention was paid to the  
220 giant paleo-slides identified by Ghazipour and Simpson (2016) to detect potential signs of activity. When a landslide-like  
221 pattern is detected, we then inspected the Google Earth images and the generated DEMs to check for typical  
222 geomorphological signs of gravitational deformation (i.e. scarps, depletion zones, debris deposits, slope bulging, river  
223 deflection, etc.).

224 The displacement time-series were then retrieved for the detected landslide sites. As the coseismic signals on landslides  
225 were too large to be unwrapped, due to phase ambiguities across landslide boundaries, the pre- and post-earthquake time-  
226 series were analyzed independently to determine the pre- and post- seismic landslide kinematics. After dividing the  
227 interferograms into pre- and post-earthquake groups, phase delays of the unwrapped interferograms were inverted pixel  
228 by pixel in order to solve the total phase delay, relative to the first date (Doin et al., 2011). Time-series for each landslide  
229 were then constructed using the cumulative deformation maps obtained from the inversion (one map at each date of the  
230 Sentinel-1 images). The mean displacement ( $\mu$ ) was calculated over a selected window, of about 15x15 pixels on the  
231 landslide body at each date, relative to a mean displacement extracted from a surrounding buffer zone of several hundreds  
232 of meter width. The displacement deviation ( $\sigma$ ) was then estimated from the mean absolute deviation of the displacement  
233 in the reference area.

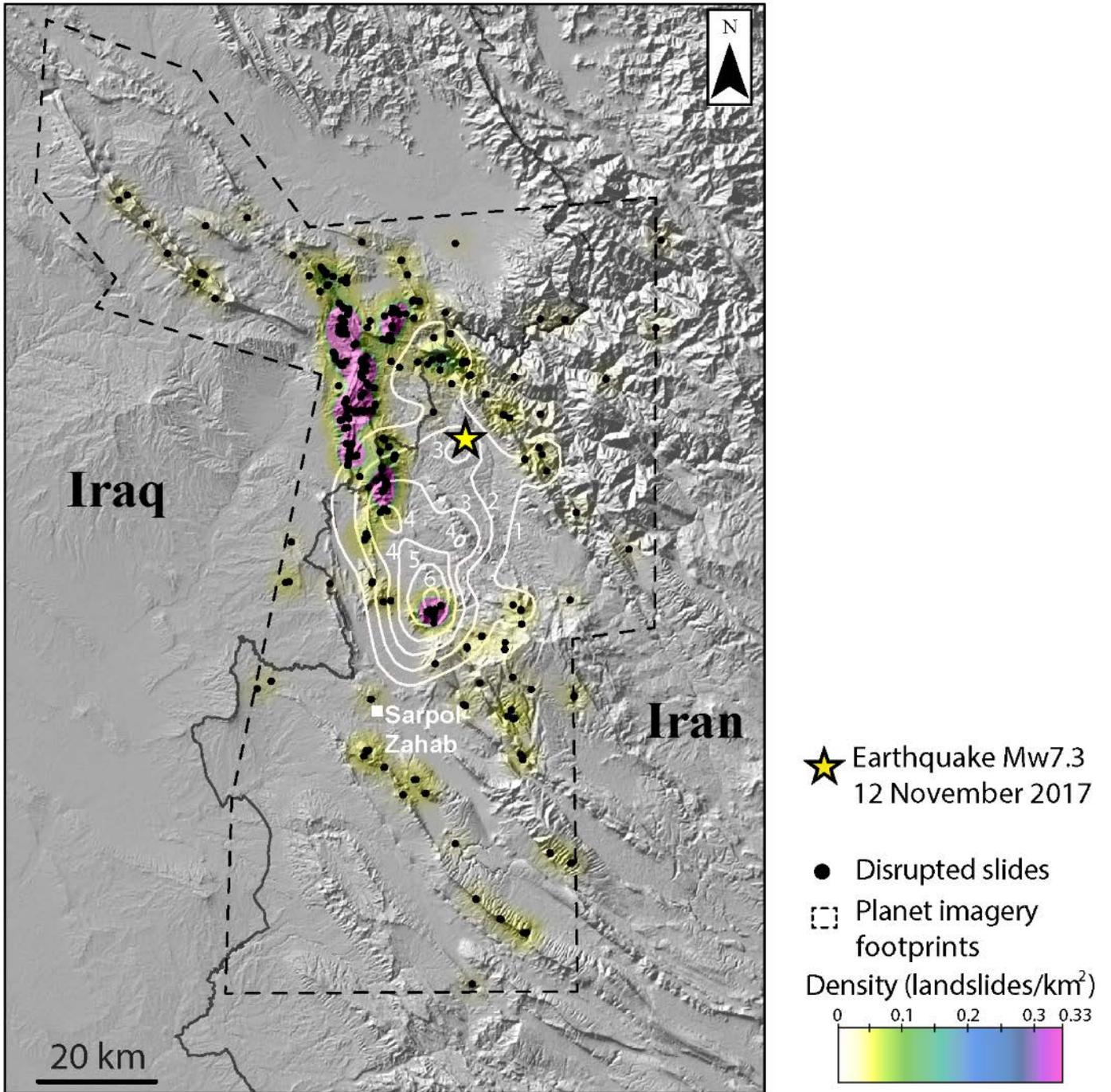
234 In a next step, the pre- and post-seismic landslide mean velocities and their associated uncertainties was calculated. Each  
235 point “ $i$ ” of the time-series was considered as a random variable of normal distribution ( $\mu_i, \sigma_i$ ). 10,000 realizations of this  
236 random variable were randomly picked at each point of the time series, and the associated 10,000 pre and post velocities  
237 were calculated by a linear regression of these time-series. The mean and standard deviation of these 10,000 velocities  
238 gave us an estimate of the mean velocity and its uncertainty.

239

240 **4. Results**

241 **4.1 Disrupted landslides**

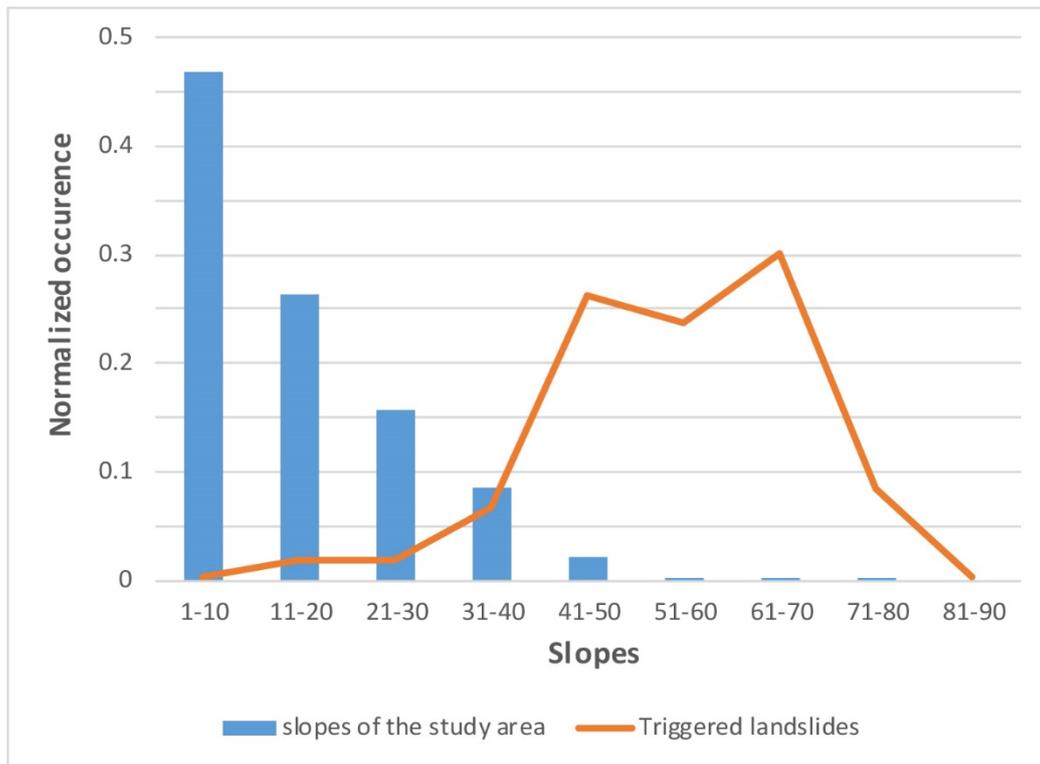
242 386 coseismic disrupted landslides were mapped (Fig. 4) with areas ranging from 200 to 20000 m<sup>2</sup>. These events are  
243 spread over 140 km along the fault in the mountain ranges, with the more distant one located 90 km south of the epicenter.  
244 They are mostly rockfalls that affect mainly limestone and flysch formations. Approximately 85% of these landslides are  
245 located within 40 km of the epicenter, with the highest concentration northwest of the epicenter (Fig. 4). This clustering  
246 does not correspond to the coseismic slip of the fault that occurred preferentially southward from the epicenter (Fig. 4).  
247 As the fault is blind with a low dip angle, the majority of the disrupted landslides are located some distance to the fault,  
248 ranging from 14 to 30 km. As seen in Fig. 4 and Fig. 5, landslide locations present a clear relationship with the topography,  
249 with most landslides occurring on slopes between 40° and 80°, which are significantly steeper than the mean ~18° slope  
250 of the area (Fig. 5).



252 **Fig. 4** Disrupted slides and falls inventory, and corresponding density map. White isolines indicate cumulative fault-slip  
 253 (in meters) 12 seconds after earthquake reported from Gombert et al. (2019).

254

255



256 **Fig. 5** Normalized number of landslides per slope category, and comparison with the slope distribution in the area of  
257 study.

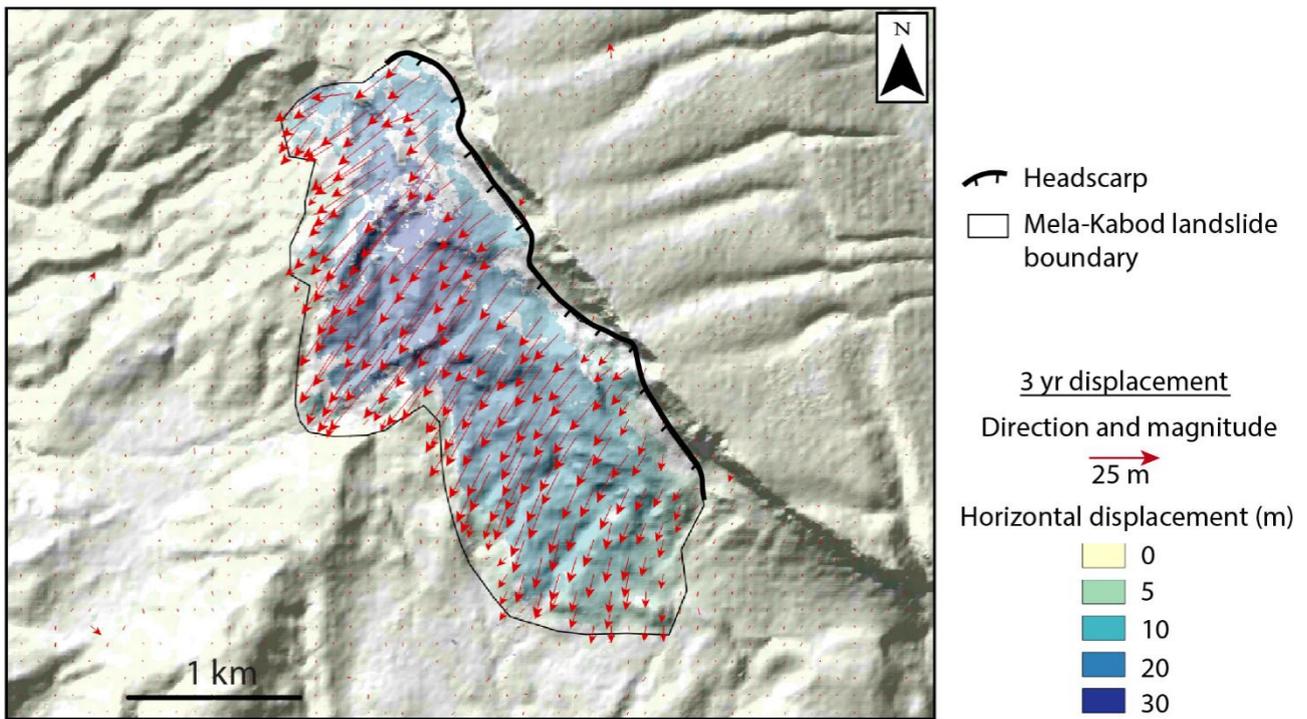
258

## 259 4.2 Coherent landslides

### 260 4.2.1 Detection and characterization

261 Nine coseismically coherent landslides were detected, their locations and properties are summarized in Table 2.

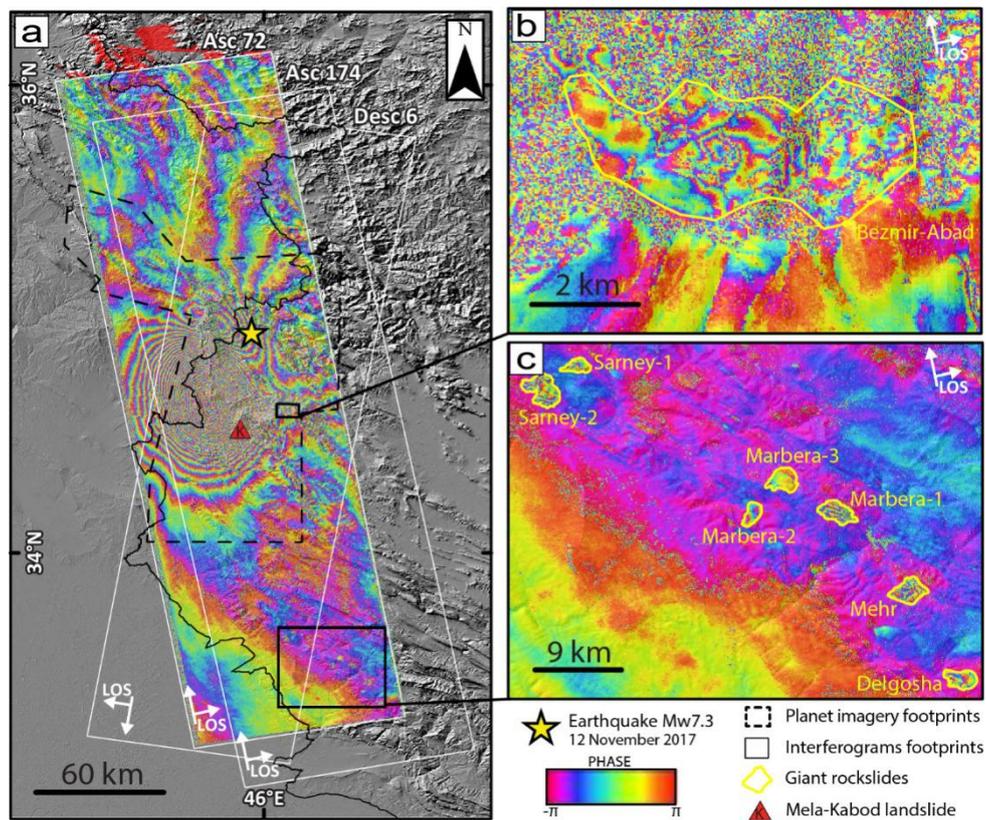
262 One site located ~40 km south of the epicenter was detected by optical image correlation: the Mela-Kabod landslide (Fig.  
263 6), already reported by Miyajima et al. (2018) and Vajedian et al. (2018). This massive landslide (area 4.5 km<sup>2</sup>) affects  
264 the red conglomerates of the Kashkan formation underlying the stiff limestone of the Asmari Formation (Valkaniotis et  
265 al., 2018). It showed southwestward movement with a maximum horizontal coseismic displacement of about 30 m, in  
266 agreement with the value found by Vajedian et al. (2018) using SAR interferometry. The detection threshold was found  
267 to be 0.5-1.6 and 1.5-2.2 for SPOT6/7 and Planet image correlation respectively. With this detection threshold, no other  
268 earthquake-induced landslides were found in the targeted area.



270 **Fig. 6** Spot6/7 image (Table S1) correlation results for the Mela-Kabod landslide obtained using the COSI-Corr tool,  
 271 showing the 3 yr cumulated displacement in meters including the coseismic period of the Sarpol-Zahab earthquake. The  
 272 black contour shows the limit of the horizontal displacement field while the displacement direction and magnitude are  
 273 presented by the red arrows.

274

275 Analyzing the coseismic SAR interferogram over the study area (Fig. 7a), eight landslides characterized by sharp phase  
 276 discontinuities and 3-4 fringes were detected (Fig. 7b and Fig. 7c; Table 2). One landslide; the Bezmir-Abad with an area  
 277 of 15 km<sup>2</sup> (Fig. 7b) is located 36 km south of the epicenter, close to the Mela Kabod landslide, and also affects the Kashkan  
 278 and Asmari formations. The other seven landslides are clustered between 140-180 km south of the epicenter (Fig. 7c,  
 279 Table 2), in a region where no disrupted landslides were detected with the Planet satellite imagery. The size of the affected  
 280 areas varies between 2 and 7.8 km<sup>2</sup> (see Table 2). Four of these sites (Mehr, Marbera-1, Marbera-3 and Bezmir-Abad)  
 281 coincide with the location of paleo landslides reported by Ghazipour and Simpson (2016), two are pointed on the  
 282 geological map (Marbera-1 and Delgosha; see Fig. S1 in supplementary information), while the other two (Sarney-1 and  
 283 Sarney-2) were previously unknown.



285 **Fig. 7** (a) Coseismic SAR interferogram computed over the study area from the ascending track 72 between 11/11/2017  
 286 and 17/11/2017. (b) and (c) close-ups show the 8 landslides detected with InSAR near the earthquake and in the southern  
 287 far-field, respectively. The location of the Mela-Kabod landslide, detected with optical images, is identified with a red  
 288 triangle.

289

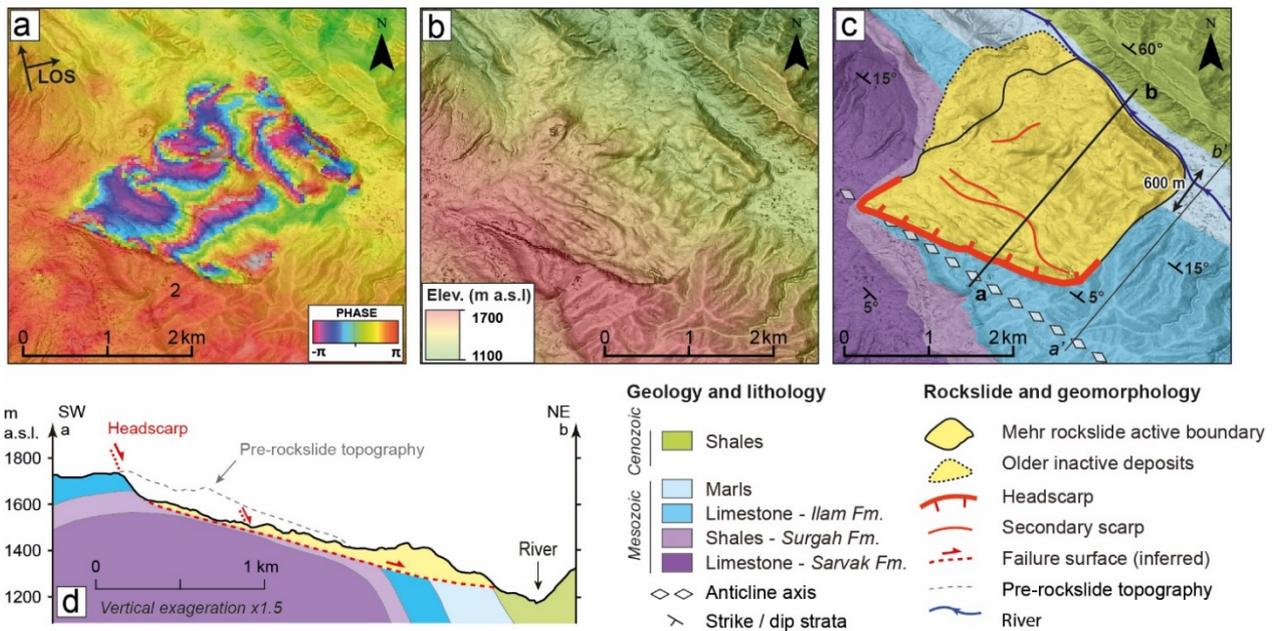
290 Assessment of the geomorphology and the geology of each site revealed clear evidence for pre-existing gravitational  
 291 slope deformations such as (1) headscarps of pluri decametric height bounding the top of the zones (Table2), (2) slope  
 292 bulging and river deflection in some cases, and (3) screes and old landslide deposits.

293 As an example, the geomorphological analysis of the Mehr landslide, located 170 km south of the epicenter, is shown in  
 294 Fig. 8. The coseismic displacement extent, clearly visible in the interferogram (Fig. 8a), delineates a 3 km long by 2.5 km  
 295 wide region. It is bounded to the SW by a ~160 m high headscarp and to the NE by the toe of debris deposits that propagate  
 296 ~600 m over the valley floor (see DEM in Fig. 8b). The geological map (Fig. 8c) and the cross-section (Fig. 8d) indicate  
 297 that the landslide initiated at the contact between Ilam limestones and Surgah shales, along the northern flank of a NW-  
 298 SE-striking anticline. The mechanism is a rockslide consisting of stiff limestone blocks sliding over a large distance on a  
 299 shale layer dipping 5 to 15° NE, and probably eroding it throughout the movement. From Google Earth images, the mass

300 in motion is destructured with a mixing of limestone blocks, entrained clayey material, and debris deposits probably  
 301 generated by secondary rockfalls. Comparison between a topographic profile extracted along the rockslide and one taken  
 302 along the non-disrupted slope suggests a maximum depth of 200 m for the slip surface (Fig. 8d), implying a rockslide  
 303 volume of  $\sim 0.5 \text{ km}^3$ .

304 We found similar geomorphological and geological characteristics for the other sites (Fig. S1 and Table 2) that can be  
 305 summarized as follows: (1) the detected interferometric patterns match the positions of pre-existing giant rockslides with  
 306 a volume range from  $0.16$  to  $2.2 \text{ km}^3$ , as estimated from the empirical surface/volume law from Ghazipur and Simpson  
 307 (2016); (2) all rockslides occurred in a geological structure with a limestone layer (Ilam or Asmari formations) overlying  
 308 a softer rock (shale, marl or conglomerate of the Surgah, Pabdeh and Kashkan formations, respectively); (3) for the seven  
 309 sites located furthest from the epicenter, the rockslides developed on the flanks of anticlines; (4) as revealed by the  
 310 headscarp height and the debris propagation, most rockslides exhibit a total displacement of several hundred meters.

311



312 **Fig. 8** The Mehr giant rockslide (see location in Fig. 7c) presented from (a) a coseismic interferogram computed along  
 313 the Sentinel-1 ascending track-72 between 11/11/2017 and 17/11/2017, (b) a pre-seismic SPOT6-7 hillshaded DEM (4 m  
 314 resolution, 09/09/2014), (c) a geological map adapted from Llewellyn (1974), (d) a cross-section built along the profile  
 315 “ab” (see location on Fig. 8c). The pre-rockslide topography is presented from the reference profile a’b’.

316

317 4.2.2 Co-seismic motion

318 The precise quantification of the coseismic motion with InSAR data is not possible due to (1) the sharp limits of the  
 319 patterns in the coseismic interferograms (Fig. 7), which prevent extraction of the phase value during the unwrapping  
 320 process, and (2) the absence of a displacement signal in the correlation of optical images (Fig. S1). Accordingly, the upper  
 321 and lower limits of the coseismic motion was bracket using the optical image correlation uncertainty and fringes numbers  
 322 of the interferogram respectively. The LOS coseismic displacement at the 8 sites previously affected by rock avalanches  
 323 was estimated to be between a few centimeters and a few tens of centimeters (Table 2).

324

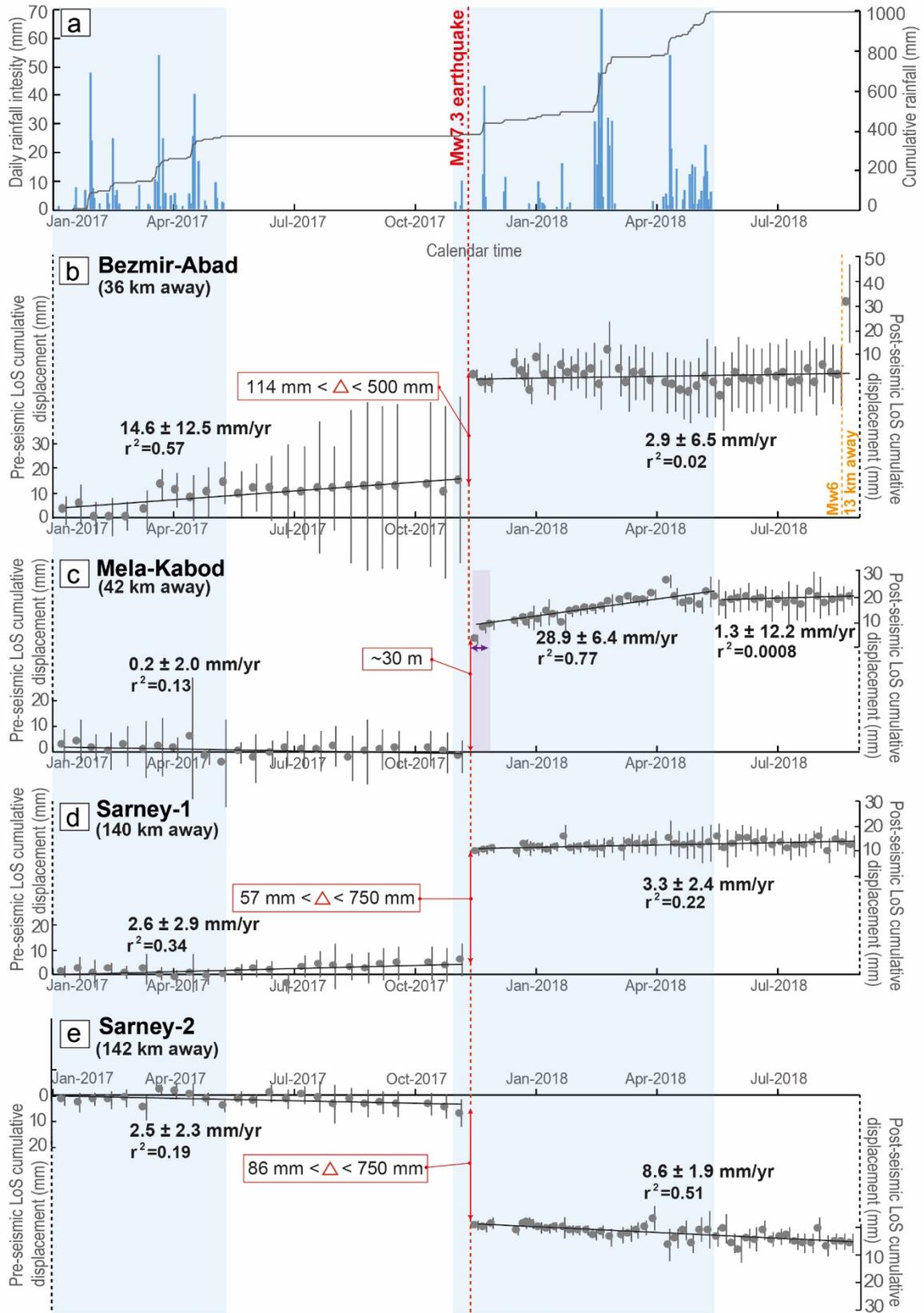
Landslide name		Mehr	Marbera-1	Marbera-2	Marbera-3	Sarney-1	Sarney2	Delgosha	Bezmir-Abad	Mela-Kabod
Location (° .dd)	Longitude	46.43909	46.36963	46.28953	46.32007	46.12247	46.0821	46.49005	46.10286	45.906
	Latitude	33.4379	33.5133	33.5103	33.5446	33.6532	33.6291	33.3525	34.6157	34.535
Ghazipur and Simpson (2016)	Name	L21	L22	NR	L23	NR	NR	NR	L56	NR
	Area (m <sup>2</sup> )	5.2x10 <sup>6</sup>	3.9x10 <sup>6</sup>	NR	6.3x10 <sup>5</sup>	NR	NR	NR	1.6x10 <sup>6</sup>	NR
Area (m <sup>2</sup> )		5.8x10 <sup>6</sup>	5.2x10 <sup>6</sup>	2.7x10 <sup>6</sup>	4.9x10 <sup>6</sup>	2.6x10 <sup>6</sup>	7.8x10 <sup>6</sup>	4.4x10 <sup>6</sup>	15x10 <sup>6</sup>	4.5x10 <sup>6</sup>
Width (m)		2500	3100	1000	2400	1600	3600	2000	6100	3800
Length (m)		3100	2000	1800	2300	2500	2500	2100	2500	1700
Volume (m <sup>3</sup> )		5.2x10 <sup>8</sup>	4.5x10 <sup>8</sup>	1.7x10 <sup>8</sup>	4.1x10 <sup>8</sup>	1.6x10 <sup>8</sup>	7.8x10 <sup>8</sup>	3.4x10 <sup>8</sup>	2.2x10 <sup>9</sup>	3.5x10 <sup>8</sup>
ΔH (m)		410	425	340	550	370	520	520	510	560
Maximum height of headscarp (m)		160	130	30	120	100	160	100	300	> 30
Slope average (°)		9	10	8	12	9	8	8	10	14
Landslide orientation		North-East	North-East	South-West	North-Northeast	North-East	South-West	South-West	North-East	South-West
Distance from epicenter (km)		169.7	159.7	158.8	155.5	140.1	142.4	179.6	35.8	41.9
LOS velocity (mm/yr)	Pre-earthquake period	25.4 ± 3.4	11.8 ± 3.9	0.3 ± 2.3	24.5 ± 4.2	2.6 ± 2.9	2.5 ± 2.3	23.2 ± 6.6	14.6 ± 12.5	0.2 ± 2.0
	Post-earthquake rain period	25.9 ± 8.5	38.1 ± 7.9	6.2 ± 5.6	46.5 ± 7.8	5.9 ± 6.1	9.5 ± 4.5	28.9 ± 6.7	9.1 ± 12.5	13.7 ± 2.0
	Post-earthquake dry period	16.7 ± 16.7	14.8 ± 17.4	2.6 ± 11.7	2.9 ± 16.2	2.5 ± 13.6	10.6 ± 10.6	34.3 ± 11.2	18.1 ± 32.1	28.9 ± 6.7
Range of coseismic displacement (cm)		9 to 40	11.4 to 36	5 to 42	2.9 to 42	5.7 to 75	8.6 to 75	5.7 to 67	11.4 to 50	100 to 3500

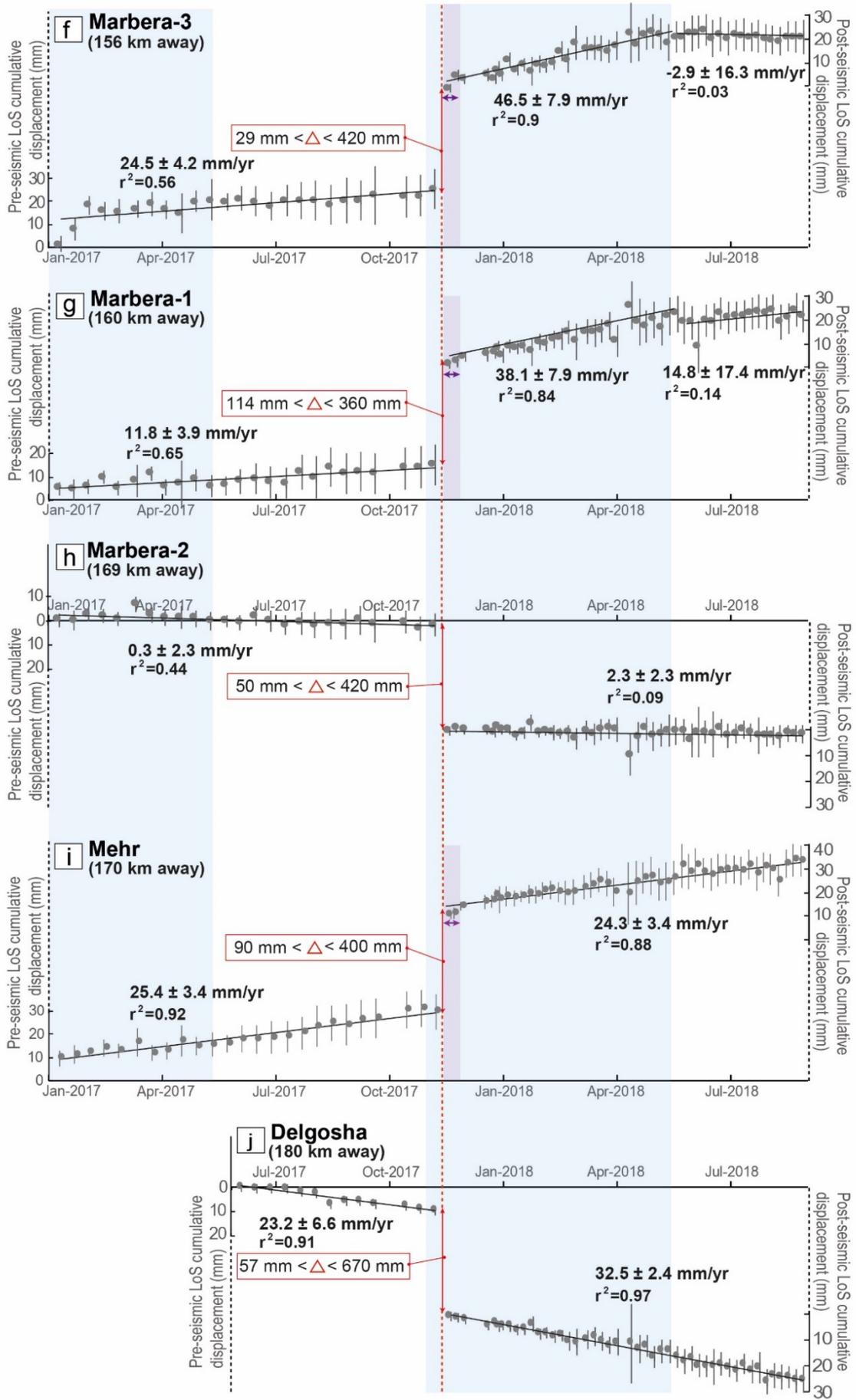
325 **Table 2:** Characteristics of the giant slow-moving rockslides detected from Sentinel-1 interferograms and optical image  
326 correlation. The area of the rockslides already identified by Ghazipur and Simpson (2016) is given for comparison. NR  
327 stands for landslides that were Not Reported in the database of Ghazipur and Simpson (2016). The area deduced from  
328 this study corresponds to the area affected by co-seismic displacement as seen in the interferograms. The volume is  
329 estimated from the empirical law proposed in the study of Ghazipur and Simpson (2016) for the Zagros region. The  
330 elevation difference  $\Delta H$ , average slope and length are calculated from the headscarp top to the landslide toe. Landslide  
331 orientation gives the direction toward which the material moves. Line-of-sight (LOS) velocities are calculated by linear  
332 interpolation of cumulative displacements from time series over several months (see §3.2.3 for details).

333

#### 334 4.2.3 Pre- and post-seismic time-series

335 The cumulative LOS displacement time-series show different kinematic patterns for the coherent landslides before and  
336 after the earthquake (Fig. 9). In the pre-seismic period, landslide velocities range between 0 to 25 mm/yr, with either  
337 dormant (almost zero velocity within the error limits: Fig. 9c, d, e, h) or slightly active landslides of constant velocities  
338 (Fig. 9b, f, g, i, j). After the coseismic motion, we observe a transient relaxation over 20 days of about 2-5 mm clearly  
339 seen at several sites: Mela-Kabod, Marbera-3, Marbera-1 and Mehr (Fig. 9c, f, g, i). Then, three different post-seismic  
340 patterns emerge with landslides showing: (1) a constant post-seismic velocity indistinguishable, within uncertainties, of  
341 the pre-seismic one (Fig. 9b, d, h, i), (2) a constant post-seismic velocity higher than the pre-seismic one (Fig. 9e, j), and  
342 (3) a transient increase in velocity of several months before returning to their pre-seismic rates (Fig. 9c, f, g). In this latter  
343 case, the transient increase and its subsequent lower velocity phase coincide with rainy and dry seasons, respectively, as  
344 shown by the comparison with rainfall time-series (Fig. 9a). Finally, the displacement observed for the last acquisition  
345 date of the Bezmir Abad time series (Fig. 9b) may correspond to the coseismic effect of a Mw 6.0 earthquake that occurred  
346 in the vicinity (~13 km) of this site (Fig. S4).





349 **Fig. 9** (a) Cumulative rainfall and rainfall intensity collected at the Ilam meteorological station (Fig. 1b; National Climatic  
350 Data Center). (b) to (j) Cumulative LOS displacement time-series with the error bars, computed for all the detected  
351 rockslides from InSAR over 20 months spanning the Mw 7.3 Sarpol-Zahab earthquake. The coseismic motion ( $\Delta$ ) is fixed  
352 for all time-series. The pre-seismic InSAR time-series for the Delgosha rockslide (j) extends back only 6 months before  
353 the earthquake due to unavailable data. The purple horizontal double arrow indicates the post-seismic relaxation period  
354 observed for some of the landslides.

355

## 356 **5. Discussion**

### 357 **5.1. Uncertainty on the detection of landslides**

358 The Sarpol-Zahab earthquake induced coseismic displacements for two types of landslides: 386 small disrupted slides  
359 clustered in a radius of a few tens of kilometers around the epicenter, and 9 giant rockslides, 7 of which are located in the  
360 far-field (up to 5 times the fault length). The validation of this inventory is complex due to the lack of accurate field  
361 observations, except for the well-detected Mela-Kabod giant landslide (Vajedian et al., 2018; Miyajima et al., 2018;  
362 Goorabi, 2020). The smallest disrupted slide detected has an area of 200 m<sup>2</sup> but smaller scale events certainly occurred,  
363 as suggested by field observations (Miyajima et al., 2018), which can therefore not be captured by our inventory. Also,  
364 the complex coseismic InSAR images makes the detection of landslide-like patterns difficult in the area affected by the  
365 strongest earthquake deformation (Fig. 7b). Thereof, there is a possibility that coherent landslides with small  
366 displacements ( $< 0.9$  m) in this area were not identified. On the other hand, landslides with large coseismic displacement  
367 such as the Mela Kabod landslide cannot be detected by InSAR due to phase decoupling (Goorabi, 2020). Finally, small  
368 coherent landslides (area less than  $\sim 10^4$  m<sup>2</sup>) with displacement less than one meter are not detected by our methodology  
369 due to the limited resolution of the Sentinel-1 SAR data ( $\sim 4 \times 15$  m original pixel size, 30x30m after multi-looking).

### 370 **5.2 Controlling factors of disrupted landslides**

371 Disrupted landslides are spread over ca. 140 km all along the fault plane with a higher concentration around the epicenter  
372 (Fig. 4), as commonly observed during earthquakes (e.g. Gorum et al., 2011; Gorum et al., 2014). The highest  
373 concentration is located northwest of the epicenter, and the farthest disrupted landslide was triggered 90 km south of the  
374 epicenter. The landslide concentration observed to the northwest is not consistent with the southward directivity of the  
375 source and the co-seismic slip displacement inferred from (Gombert et al. 2019), contrary to what was observed for other  
376 earthquakes (e.g. Gorum et al., 2011; Specht et al., 2019). Only few seismological stations were installed near the epicenter,

377 making it challenging to resolve the spatial distribution of the seismic ground motion in this area. Five seismic stations  
378 are available within 50 km of the epicenter  
379 (<https://earthquake.usgs.gov/earthquakes/eventpage/us2000bmcg/shakemap/pga>). In the area of high landslide  
380 concentration, 28 to 35 km north of the epicenter, three stations recorded PGA values between 0.27g and 0.57g, whereas  
381 the highest ground motion is recorded 40 km south of the epicenter in Sarpol-Zahab city (in an area with little topography  
382 and few landslides) with a peak ground acceleration of 0.76g. This strong asymmetry of the ground motion associated  
383 with the fault rupture directivity (Babaie Mahani and Kazemian, 2018; Gombert et al., 2019) is not reflected in the spatial  
384 distribution of landslides. Instead, the cluster of disrupted landslides northwest of the epicenter may be explained by the  
385 steep topography of the area (Fig. 4 and Fig. 5). Finally, although some small landslides may have been missed in our  
386 inventory (see discussion above), the number of recorded landslides (about 400) is still very low for a Mw7.3 earthquake,  
387 which would be expected to trigger a few thousand landslides in such a mountainous area (Keefer, 2002, Tanyas et al.,  
388 2017). This low number of landslides is, however, consistent with other observations from arid and semi-arid  
389 environments (Lacroix et al., 2013; Barlow et al., 2015). For this earthquake, we suggest that it can be explained by (1)  
390 the aridity of the region, which limits weathering and soil production (Singer, 1991; Dregne, 1976), and (2) relatively  
391 deep (and blind) thrust faulting, which induces lower ground motions than surface rupturing earthquakes (Aki, 1987).

392

### 393 **5.3 Far field forcing of coherent landslides**

394 The striking feature of this earthquake resides in the coseismic motion of the giant paleo-rockslides located at epicentral  
395 distances of up to 180 km (Fig. 7 and S1). For such a high magnitude earthquake (Mw 7.3), this distance corresponds to  
396 the upper bound, as can be inferred from worldwide data base of seismically induced landslides (Keefer, 2002; Delgado  
397 et al. 2011). The concentration and the huge size (volumes up to 2 km<sup>3</sup>) of the rockslides triggered by this earthquake is  
398 however intriguing and, to our knowledge, has never been reported so far from the epicenter (4-5 times the fault length).

399 All these giant rockslides were pre-existing features with very large slide displacements accumulated in the past (a few  
400 hundred meters), and for most of them still active before the earthquake (Fig. 8, Table 2) highlighting their unstable  
401 character. All these rockslides are situated south of the epicenter in the direction of the rupture. As emphasized by several  
402 authors (Chen et al., 2018; Gombert et al., 2019) the directivity of the rupture toward the south led to stronger ground  
403 motion south compared to north of the rupture (Babaie Mahani and Kazemian, 2018). The recorded PGA closest to these  
404 rockslides reached 0.13g at the Ilam station (146 km far from the epicenter) and is significantly higher than the PGA  
405 recorded north of the epicenter at approximately the same distance (0.03g in Mahabad or Bukan stations).

406 The ground shaking can also be amplified by both local topographic and geologic effects (e.g. Murphy, 2015), more  
 407 specifically as several rockslides developed on flanks of anticline ridges and in mechanically heterogeneous lithologies.  
 408 Indeed, the five more distant giant landslides affected the same 200 m thick Ilam formation (limestone) overlying a 100  
 409 m thick shale layer (Surgah formation). This two-layer structure constitutes a dynamic oscillator on the thick underlying  
 410 carbonate of the Sarvak formation, generating resonance and favoring rock sliding initiation along the weak shale layer.  
 411 During the long-distance slide of the rockslides, the limestone blocks progressively disintegrated and dragged part of the  
 412 Surgah formation, creating a highly destructive deposit with a maximum thickness ( $t$ ) of around 150 m, consisting mainly  
 413 of a mixture of shale and limestone. The amplification of the seismic waves during the earthquake is due to the seismic  
 414 impedance contrast (product of the density  $\rho$  and shear wave velocity  $V_s$ ) between this deposit and the underlying  
 415 calcareous substratum. For vertically incident SH waves in a 1D elastic structure, the resonance frequency  $f_0$  and the  
 416 corresponding amplification  $Af_0$  are given by (Kramer, 1996):

$$417 \quad f_0 = \frac{V_{SD}}{4t} \quad (1)$$

$$418 \quad Af_0 = \frac{\rho_B V_{SB}}{\rho_D V_{SD}} \quad (2)$$

419 where  $V_{SD}$  and  $V_{SB}$  are the shear wave velocities of the rockslide deposit and the bedrock, respectively, and  $\rho_D$  and  $\rho_B$  are  
 420 the corresponding densities.

421 Rockslide deposit and bedrock  $V_s$  values at these sites are not available, but plausible values can be taken from a similar  
 422 rockslide comprising limestone and marls (e.g. Socco et al., 2010), and bedrock (e.g. Telford et al., 1990):  $V_{SD} = 600$  m/s;  
 423  $V_{SB} = 3000$  m/s;  $\rho_D = 2$ ;  $\rho_B = 2.5$ . Considering these values, a resonance frequency  $f_0 \approx 1$  Hz with a 1D amplification over  
 424 6 has been obtained. This resonance frequency, around 1Hz, is compatible with the frequency content of the earthquake  
 425 source at such large distances (Babaie Mahani and Kazemian, 2018), suggesting a potential amplification effect that can  
 426 thus favor landslide triggering.

427 Interestingly, several other ancient landslides previously mapped in this region (Ghazipour and Simpson, 2016) were not  
 428 reactivated during the Sarpol-Zahab earthquake (Fig. 1b). Some explanations can be given for this phenomenon: (1) these  
 429 landslides were inactive during the whole period of monitoring, thus indicating their higher stability; (2) they mostly  
 430 occurred in different lithologies (Oligocene and Eocene units) with a structure (absence of shale layer) less sensitive to  
 431 sliding and seismic amplification; (3) these non-reactivated landslides are mostly northwest oriented while the reactivated

432 ones are oriented northeast or southwest. Stronger seismic ground motions have been recorded perpendicular to the fault  
433 for long-period waves (Babaie Mahani and Kazemian, 2018), favoring the initiation of NE-SW oriented landslides.

434

#### 435 **5.4 Post-seismic motion of coherent landslides**

436 The 9 giant rockslides present a co-seismic motion of at least 10 cm, 4 of them (Fig. 9c, f, g, i) being affected by a transient  
437 increase of the post-seismic velocity of a few mm in the 20 days following the earthquake. This co-seismic and relaxation  
438 process is consistent with previous observations (Lacroix et al., 2014), although the ratio between co- and post-seismic  
439 motion is far greater here, with most of the motion released during the co-seismic phase (against 3 to 10 times more post-  
440 seismic than co-seismic motion released on the Maca landslide, Peru - see Lacroix et al., 2014; Bontemps et al., 2020).  
441 Five of the landslides do not even present a post-seismic relaxation phase. These observations suggest that the rheology  
442 of the sliding surface does not follow strain-weakening friction laws (Lacroix et al., 2014), a rheology which can be  
443 caused by undrained loading of the landslide sliding surface during the shaking (Wang and Sassa, 2002) and progressive  
444 release of excess pore-water out of the loaded sliding surface.

445 Time-series of landslide kinematics also reveal an increase of the post-seismic annual velocity compared to the pre-  
446 seismic annual velocity for at least 5 of the studied rockslides (Fig. 9c, e, f, g, j). For 3 of the rockslides (Mela-Kabod,  
447 Marbera-1, Marbera-3) we document a higher velocity during the 2018 rainy season (Fig. 9c, f, g) followed by a  
448 subsequent decrease in velocity, eventually returning to the pre-earthquake velocity during the dry season. This  
449 observation may suggest a seasonal motion and threshold rainfall effects on the landslide kinematics (Zerathe et al., 2016).  
450 Furthermore, moderate earthquakes (Mw5-6) that occurred close to the landslides (Fig. 1b) may have contributed to the  
451 transient velocity increase by damaging rocks, thus promoting water infiltration (Bontemps et al., 2020) up to the  
452 impermeable formation (e.g. Surgah). However, exploring these issues are beyond the scope of this study, as this will  
453 require longer time-series and detailed field measurements of hydrologic and kinematic parameters (e.g. Schulz et al.,  
454 2009).

455

#### 456 **6. Conclusions**

457 In this paper we provide an inventory of landslides induced by the Mw 7.3 Sarpol-Zahab earthquake which struck the  
458 northern Zagros Mountains on the 12th November 2017. To detect a maximum of triggered landslides of various sizes,  
459 typologies, and co-seismic displacements, we performed: (1) a visual change detection comparing orthorectified Planet

460 optical satellite images (3-4 m resolution) spanning the earthquake, (2) optical image correlations using SPOT6/7 (1.5 m  
461 resolution) and Planet images spanning the earthquake, and (3) InSAR detection and monitoring using Sentinel-1 images.  
462 With this strategy, 395 triggered landslides of different sizes and kinematics were detected, including 386 rockfalls  
463 concentrated within a radius of 40 km around the epicenter and 9 pre-existing giant rockslides (area range 3-30 km<sup>2</sup>)  
464 distributed at distances between ~40 km and up to 180 km from the epicenter. Results revealed the exceptionally low  
465 number of landslides recorded during this Mw 7.3 earthquake, which is attributed to the semi-arid conditions of the Zagros  
466 Mountains that limit weathering and soil production. On the other hand, the striking element revealed by this earthquake-  
467 induced landslide inventory was the co-seismic displacement of 9 pre-existing giant rockslides, with minimum  
468 displacements ranging between 10 and 90 cm. Seven of these coherent landslides are clustered in the far field at distances  
469 of 140-180 km from the epicenter (4-5 times the fault length), and affect old deposits generated by the slide of a thick  
470 limestone layer over a shale formation on limbs of anticlinal structures striking NW-SE. We propose that the co-seismic  
471 forcing of these rockslides may be related to a combined effect of the southward directivity of the source, the NE-SW  
472 polarization of the ground motions, their sensitivity to low-frequencies (~1 Hz), the site effect due to the seismic  
473 impedance contrast on the flanks of the anticlines, and their pre-existing activity. From the InSAR time-series  
474 documenting the rockslide displacements (covering 20 months centered on the earthquake date), we observed that half of  
475 those giant rockslides were already slowly moving before the earthquake (velocity < 25 mm/yr).

476 Finally, most of these giant coherent rockslides have certainly been initiated several millennia ago, as shown by the dating  
477 of the giant Seymareh landslide (Roberts and Evans, 2013). The aridity of the area helped to preserve their geomorphology,  
478 and open the possibility to date the scarp evolution (e.g. LeRoux et al., 2009) for a long-term study of the landslide  
479 kinematics. The recent activity documented here shows the interest of these objects for landslide forcing studies over both  
480 short and long term of arid environments.

481

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