

New Directions for an Experimental Approach to the Chemistry of the Origin of Life

Grégoire Danger, L. Le Sergeant D'hendecourt, V. Vinogradoff, R. Pascal

▶ To cite this version:

Grégoire Danger, L. Le Sergeant D'hendecourt, V. Vinogradoff, R. Pascal. New Directions for an Experimental Approach to the Chemistry of the Origin of Life. Michele Fiore. Prebiotic Chemistry and Life's Origin, Royal Society of Chemistry, pp.424-455, 2022, Chemical Biology, 978-1-78801-749-7. 10.1039/9781839164798-00424 . hal-03858230

HAL Id: hal-03858230 https://hal.science/hal-03858230

Submitted on 8 Dec 2022

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers. L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.



1	
2	
3	X New Directions for an Experimental Approach to the Chemistry of the Origin
4	of Life
5	
6	
7	
8	
9	
10	
11	G. Danger ^{a,b,c} , L. Le Sergeant d'Hendecourt ^{a,c} , V. Vinogradoff ^{a,c} , R. Pascal ^{a*} .
12	^a Laboratoire de Physique des Interactions Ioniques et Moléculaires, UMR 7345, Aix-
13	Marseille Université, CNRS, Marseille, France, ^b Institut Universitaire de France
14	(IUF), ^c Laboratoire d'Astrophysique de Marseille, UMR 7326, Aix-Marseille
15	Université, CNRS, Marseille, France
16	*corresponding email address: robert.pascal@univ-amu.fr
17	
18	ABSTRACT
19	Organic matter is formed in molecular clouds from which planetary systems arise as
20	well as in planetary environments. Most of these locations are not likely to harbour
21	life, questioning the genuine prebiotic character of the corresponding chemistry and



its relevance to the origin of life. The formation of organic matter does therefore not 22 necessarily constitute a systemic trend towards life in our Universe. However, its 23 24 sluggish reactivity at low temperatures is likely to have allowed the delivery of volatile elements essential for life at the surface of planets like the Earth. Kinetic selection 25 may also have played a role for sorting specific active species or complex catalytic 26 27 processes during the self-organization preceding life owing to the kinetic barriers hindering the reactions of covalent bonds and holding the system in a far-from-28 equilibrium state. As a result of both processes, early environments on telluric 29 30 planets that exhibit other essential factors, namely, liquid water and light as an energy source, may then become favourable to the origin of life. In both processes, 31 reactivity matters as much, if not more, than structures. Next to focusing on the 32 synthesis of building blocks, emphasis must be made on reactivity for identifying 33 networks involving autocatalysis, replication, or positive feedback. 34



35 X. 1. Introduction

The topic of this book – Prebiotic Chemistry and Life's Origin – is connected with the 36 hypothesis, expressed almost one century ago, that life could have emerged from 37 abiotically formed organic matter^{1,2}. This hypothesis is reminiscent of the fact that the 38 designation of organic matter itself derives from its intimate association with life. The 39 report of the famous Miller experiment³, carried out a few decades later, apparently 40 provided support to this hypothesis. It suggested that the issue of the origin of life 41 could have been solved soon after it was demonstrated that biochemical "building 42 blocks of life" are abiotically formed under somehow arbitrary conditions judged at 43 that time compatible with the presumed environment of the early Earth. In spite of 44 the fact that vitalist beliefs had been refuted with the synthesis of urea from abiotic 45 and inorganic chemical compounds since the beginning of the 19th century⁴, the 46 enthusiasm raised by the Miller's experiment can still be considered as a remain of 47 the belief that organic matter and life on Earth and by extension elsewhere, are 48 intimately connected. Moreover, more than 50 years later, at a workshop organized 49 by Pier Luigi Luisi and Kepa Ruiz-Mirazo, the following guestion was set forth⁵: 50 "Would you agree with the statement, that from the conceptual point of view the field 51 has not progressed much since the early experiments of Stanley Miller? And why do 52 you think/not think so?" The fact that the relationship between abiotically formed 53 organic matter and the origin of life still seems questionable put under scrutiny the 54 relationship between these two topics. In other words, no definite evidence has been 55 provided by the scientific community in support of a more or less deterministic 56 development of chemistry, which could have yielded systems capable of evolving in 57 a way similar to living organisms. We therefore have a limited understanding of the 58 pathways through which such a hypothetical deterministic process could have 59 Royal Society of Chemistry – Book Chapter Template



happened. No comprehensive answer has been given to the question of how could 60 the chemical evolution of matter have been converted into the process of natural 61 62 selection. This question can be reformulated more precisely as how chemical changes, driven by chemical affinities, can give rise to a process in which occasional 63 replication errors can be selected as a source of self-organization and subsequent 64 evolution. Any change towards a more organized state associated with a decrease in 65 entropy contradicts the directives of the Second Law of thermodynamics. Proposing 66 that some unknown chemical determinism played a role at the early stages of the 67 self-organization of life requires that the process acted as a driving force to 68 overcome the constraints of the Second Law. It also implies that this process 69 disappeared without leaving any easily recognizable remnant after natural selection 70 became prevailing. 71

Alternatively, contingency cannot be excluded as a factor responsible for the self-72 73 organization of life, even though estimates of the probability of a single transition for the emergence of an organism from scratch imply that this event is virtually 74 impossible^{6,7,8}. Any transition process based on a highly improbable event 75 constitutes a violation of the Second Law of thermodynamics. In spite of this 76 drawback, it has been suggested that low-probability events may have been 77 responsible for the transition, such as the emergence of specific sequences having 78 the ability of catalysing their own replication in a population of ribonucleic acid 79 (RNA), polymers playing a role of information carriers⁹. However, this appealing 80 proposal constituting the 'RNA world' hypothesis, in which replication results from the 81 activity of ribozymes, is highly improbable. The evolution of the 'RNA World' would 82 indeed have required the occurrence of a process able of sustaining an environment 83



rich in a full set of activated RNA monomers, which constitutes a "prebiotic chemist's
nightmare"¹⁰.

From the dilemma between a "piece of statistical luck" or "an evolutionary process"⁸ 86 many scientists chose the latter^{7,11,12}. No sharp transition marked by a highly 87 improbable event is required in such an evolutionary process that may proceed 88 through stages separated by less improbable transitions. The onset of fully effective 89 natural selection may on the contrary constitute the conclusion in a fuzzy transition 90 towards life^{13,14}. The living world grew as a result of a driving force associated with 91 the occurrence of reproduction with modifications and natural selection, as 92 developed by the theory of evolution¹⁵. This evolutionary force is therefore mainly 93 related to kinetics. Its kinetic nature suggests that rudimentary selection based on 94 chemical kinetics, could have preceded its full implementation as the prevailing 95 evolutionary driving force. It could have led the system towards higher degrees of 96 97 organisation before natural selection is considered established and therefore provides grounds for a continuous process for the self-organization of life. This 98 process may be already considered as reminiscent of natural selection even before 99 actual biological life. 100

The whole process also involves a transition into a state presenting an alternative 101 102 form of stability, namely Dynamic Kinetic Stability (DKS), able to replace thermodynamic stability as a source of persistence^{16–22}. The direction of evolution is 103 not constrained by the Second Law since DKS can only be observed from open 104 systems maintained out of equilibrium. This is consistent with the thermodynamic 105 analysis of the development of self-organization and the conclusion that living 106 organisms must behave as open thermodynamic systems to avoid a violation of the 107 Second law²³. Kinetic selection requires the system to be held in a far from 108 Royal Society of Chemistry – Book Chapter Template



equilibrium situation. It must therefore be fed with energy and open to the 109 environment. All this body of evidence suggests that chemical evolution preceding 110 the transition should have been driven towards higher degrees of functional 111 organization by a factor orientating its direction of evolution that is of kinetic nature. 112 But kinetic selection may have also played a more general role, diverging from 113 natural selection, for instance in the segregation of a diversity of chemical 114 environments that allowed only some of them to become favourable to the origin of 115 life. 116

This chapter is aimed at demonstrating that life is governed by physical principles 117 and that these principles are sufficient to account for the main features of life and. 118 most importantly, of its origin. It is built on the basis of previous work that 119 emphasized on the importance of kinetics in the process^{16,24,25}. One of our main 120 tenets is based on the role of kinetic selection during the whole process. There is 121 122 indeed no reason to consider that this kind of selection was only present just before the onset of natural selection and that of the living world. Kinetic selection could also 123 play a role as early as in the differentiation of cosmic materials in astrochemistry and 124 continued over the whole process of self-organization based on the multiplication 125 and selection of material entities that led to biological life. Kinetic selection should 126 indeed be considered as a tool available for sorting essential building blocks from the 127 huge chemical diversity generally found in abiotic organic materials and most 128 prebiotic synthesis experiments^{26,27,28} before it reached its utter role in the 129 evolutionary biology of living organisms. Experimental studies in the origin of life 130 research should take these considerations into account and focus on processes by 131 which kinetic selection associated with the multiplication of chemical entities might 132 allow change to occur in the direction of increased functional organization¹⁶. 133 Royal Society of Chemistry – Book Chapter Template



134 X. 2. What Makes Organic Matter so Specific?

135 X. 2. 1. The Historical Status of Organic Matter as Regards the Origins of Life

To which extent is organic matter necessary and sufficient to the emergence of life? 136 The answer to this question is intimately associated with the process that has driven 137 the self-organization of matter to life. According to the vitalist belief, prevailing before 138 the nineteenth century, organic matter was only associated with living organisms. 139 Though devoid of scientific grounds, this description was considered to account for 140 141 its difference from mineral matter. Remains of this belief are so deeply anchored in our culture that inorganic and organic chemistries are still considered as different 142 fields of science in spite of the fact that their interface has become larger. A similar 143 comment could apply to the identification of organic matter in a space science 144 context that is often directly associated with the occurrence of life in popular science 145 media. A recent example was the glycine detection in the comet 67P/Churyumov-146 Gerasimenko²⁹, targeted by the Rosetta mission. The implicit idea that is suggested 147 by this description^{30,31} is that, as soon as ingredients are present and delivered in a 148 favourable environment, then life is bound to emerge. Hence, the sole detection of 149 organic matter in extraterrestrial environments would constitute a first indication for 150 the possibility of the presence of life in any favourable environment. However, these 151 favourable environments are quite far from being clearly defined at the present time. 152 Similar misconceptions with respect to the status of organic matter have also 153 hindered the development of science as regards its detection and characterization in 154 meteorites during the nineteen century and persisted all over the first half of the next 155 one³². This chapter is aimed at describing the factors that could explain why organic 156 chemistry is associated with life as we know it and that it is not the precise nature of 157 organic matter that is the most important but its reactivity. Our aim is not to suggest 158 Royal Society of Chemistry – Book Chapter Template



that life can emerge in the absence of organic matter, as one of our conclusions is
indeed that organic chemistry is necessary but not sufficient for the origin of life
when considering liquid aqueous environments. *The issues of the abiotic formation*of organic matter and that of the origin of life most probably constitute two different
processes even though the latter may not be considered independently of the
former.

165 X. 2. 2. From Prebiotic Chemistry to the Origin of Life

166 The first conceptual approaches to the origin of life made by Oparin, assumed that life could evolve from abiotically formed organic matter (the so-called heterotrophic 167 hypothesis)¹. Actually, the main originality of the Oparin's hypothesis¹ relied on the 168 fact that he indeed considered the formation of organic matter and the origin of life 169 as constituting two separate processes. The Oparin hypothesis found experimental 170 support through the result of the Miller experiment³. This experiment, proving the 171 formation of α -amino acids from simple inorganic precursors, constituted a milestone 172 in the origin of life studies. It testified that, as soon as dioxygen is absent from the 173 environment, the chemistry of carbon at a reduced state of oxidation becomes 174 ubiquitous. The lack of knowledge on the precise conditions prevailing on the early 175 Earth, and especially on the composition and reduced state of the atmosphere 176 during the early Archean and Hadean periods (ca. 3.5-4.5 Gy before present), led 177 many scientists to consider that organic matter was not abundant on the Earth's 178 surface before the origin of life. The chemo-autotrophic hypothesis, stating that life 179 emerged in connexion with a carbon metabolism built from inorganic sources of 180 181 carbon and mineral reducing agents, was then proposed as a solution to circumvent a presumed limited availability of abiotic organic matter^{33,34}. In line with the above-182 developed ideas, the autotrophic hypothesis might be considered as step backwards 183 Royal Society of Chemistry – Book Chapter Template



to a conceptual stage in which the origin of life and the formation of organic matter 184 were not considered as separate processes. This assertion is worth mentioning in 185 spite of the popularity of a chemo-autotrophic origin of life in the scientific 186 community. The main tenets of this approach have been debated on the grounds 187 that the process clearly lacks experimental support³⁵. Other criticism stated that 188 biochemistry can hardly recapitulate prebiotic processes because the emergence of 189 metabolic pathways is unlikely in prebiotic chemistry³⁶. Additionally, a surface 190 metabolism would require multiple catalytic steps to work at the same time and an 191 appropriate arrangement of catalysts that seems even more unlikely³⁷. Among these 192 issues, the last one is probably the most crucial. The hypothesis of an autotrophic 193 origin of life indeed presumes that carbon fixation was not independent from the 194 origin of life process. Organics should therefore be both synthesized from inorganic 195 sources of carbon (CO₂ for instance) and used as food by the first living organisms. 196 One of its most achieved versions involves a catalytic contribution of inorganic 197 surfaces on which reactants could be adsorbed. However, the existence of a 198 metabolism requires that several chemical steps are efficiently proceeding at the 199 same time. Therefore, a collection of different catalysts must have been present and 200 the intermediates must have been able to move between the different surfaces 201 playing a catalytic role. This motion requires metabolites to be desorbed from the 202 surface and to be released in solution. This hypothesis is therefore highly unlikely 203 unless both the catalysts and the metabolites are sequestrated within a single 204 compartment. Otherwise the formation of organic matter and its metabolic use 205 should be considered as two distinct processes and the corresponding entities as 206 heterotrophs. We are thus conduced to the need of compartments possessing a 207 collection of catalytic surfaces between which metabolites could be able to migrate 208



without being released in the environment. This hypothesis implies that 209 compartments having a wide range of catalytic sites are available, which makes 210 211 them improbable unless they can be multiplied into numerous copies. The issue of the reproduction of entities based on organic matter would therefore be changed into 212 that of compartments containing a full set of homogeneous or heterogeneous 213 catalysts acting on the metabolism. These observations do not mean that studies on 214 prebiotic reaction networks that could predate metabolism are useless^{38–42}. On the 215 contrary such studies may help understanding why, at a given (but not necessarily 216 early) stage, these processes could have been harvested by systems able to 217 reproduce themselves and evolve on the pathway from non-life to the living state⁴³ 218 (Figure X.1). 219

220

[Insert Figure X.1 here]

In addition to the above-mentioned drawbacks of hypotheses involving contributions 221 of mineral surfaces, it is worth noting that they put the emphasis on the role of 222 catalysis that is generally understood as a mean to avoid kinetic barriers. This 223 emphasis appears to be in contradiction with the role independently assumed for 224 kinetic barriers needed to hold the system in a far from equilibrium state as a 225 prerequisite for self-organization^{44–47}. Otherwise, the system would rapidly reach the 226 227 equilibrium state, which prevents any possibility of self-organization associated with disequilibrium. The equilibrium state indeed corresponds to the statistically more 228 probable one. In other words, most components of a system capable of self-229 230 organizing far from equilibrium should preferably show sluggish reactivity rather than rapidly undergoing relaxation towards equilibrium. Any catalysis should only be 231 expressed on very specific chemical steps rather than on a wide range of processes. 232



In this way, fuelling the systems with energy would favour specifically the chemical
steps associated with the development of self-organization.

235 X. 2. 3. Extraterrestrial Chemistry

Extraterrestrial chemistry pertains to the chemistry, elements and molecules, that 236 can be observed in various extraterrestrial environments, evolving in a repeating 237 cycle that includes stars, the diffuse interstellar medium (ISM), molecular clouds, 238 planetary systems down to comets and asteroids and finally to meteorites recovered 239 240 on the Earth's surface. All these environments are prone to a rich chemistry where, in most of them, organic chemistry is at work. There, very diverse organic molecules 241 form and evolve. They are observed in the gas phase as well as in numerous solid 242 phases. The complex interplay between these different sites and processes are often 243 discussed in the astrochemistry community, which addresses the global chemical 244 evolution of the Galax y^{48} . 245

X. 2. 3. 1. Cosmic abundances, depletions and their significance for carbon based life

Cosmic abundances constitute a given characteristic of our "local" neighbourhood. It
is important to realize that these abundances, measured for example in the Sun's
photosphere but also in CI chondrites⁴⁹ are expected to be the same in the
interstellar medium from which the stars continually form. Therefore, the resulting
ISM gas chemistry and its subsequent solid phase chemistry are highly influenced by
the abundances of the elements, relative to H, initially present in the photospheres of
the stars.

Three groups of elements, decreasing in abundances, may be considered. The first
 one consists in non-condensable elements such as H and He (~90% and ~9% by
 Royal Society of Chemistry – Book Chapter Template



numbers respectively) by far the most abundant ones. Then next to it, the group of 257 "volatiles" mainly consists of O, C, N, and S (and the less abundant P) whose total 258 abundances respective to H amounts to $\sim 10^{-3}$. The third and guite substantial group, 259 as far of telluric planets are concerned, forms the rock-forming elements (Mg, Si, Fe, 260 Al, Ca, Ni, Ti) whose total cosmic abundances barely gets up to $\sim 10^{-4}$ that of H, an 261 order of magnitude less than the volatiles. Rock-forming elements are then almost 262 totally absent as atoms in the gas of the ISM, since they form the mineral 263 composition of the dust grains and thus reflects the refractory nature of the minerals, 264 265 kinetically frozen in the ejected and cooling atmospheres of asymptotic giant branch (AGB) stars ^{50,51}. 266

One of the salient point of all living systems, mainly the elemental composition of key 267 molecules of biochemistry (such as amino acids, proteins and the ATP molecule) 268 including the carriers of genetic information (nucleic acids), is that these molecules 269 appear to be strictly restricted only to very few elements, the so-called CHNOPS, a 270 terminology much in use in the astrobiology community. The availability of the 271 CHNOPS, for the initial astrochemical and subsequently astrobiological processes, 272 relates to the general astrophysical considerations discussed here. These 273 CHNO(P)S elements (with the notable exception of P that may be evidence for a 274 planetary process) are not only the most abundant ones at the cosmic level but their 275 volatility makes them natural partners for reacting together and with H and the 276 abundant C, which afterwards generates a rich organic chemistry in various 277 environments such as the surface of interstellar grains. The CHNOPS are poorly 278 depleted because, being more abundant than all the others, they cannot be locked 279 into mineral grains whose elements are already fully condensed at higher 280 281 temperatures.



Thus, the general view of the evolution of matter in the Galaxy (see Ref. 52 for a 282 comprehensive description of this evolution) from nucleosynthesis in stars to the 283 284 formation of refractory solids from AGB stars, the main sources of interstellar grains, provides a factual explanation for the abundance of the CHNOPS. Interstellar abiotic 285 organics may therefore be considered as the result of a generic trend at work at a 286 much larger scale than that of the solar system. Furthermore, at a more elementary 287 level, it also explains why the observed extraterrestrial chemistry in the gas phase is 288 largely dominated by volatile species (H₂O, CO...), which also holds true in the solid 289 290 phase, for interstellar ices for instance and their subsequent evolution into organic matter. Heavier elements, the rock-forming ones, are trapped in non-reactive grains. 291 In the gas phase, CHNOPS do react between themselves (and in particular with H) 292 because they cannot react with already depleted and much less abundant elements. 293 This dichotomy, imposed by the kinetics of solid formation in the mass loss of the 294 295 outer envelopes of stars, in the elemental distribution between volatiles (ice-forming) and refractory (rock-forming) elements crucially contributes to the rise of the organic 296 chemical world observed in extraterrestrial environments in addition to chemical 297 reactivity. 298

299 X. 2. 3. 2. Molecular astrochemistry

Molecular astrochemistry starts in the gas phase of the ISM with the elemental abundances described above. Up to 200 molecules, mostly organics, including radicals and ions, have been identified in the gas phase by radio astronomy. A nonexhaustive list of detected molecules in the gas phase in different settings (molecular clouds, protoplanetary disks and stars' envelopes) is given in Table X. 1. In this table, molecules (and radicals) are arranged by number of atoms in the species.



Organic molecules, built from the CHNOPS elements, discernably dominate this list. 306 Note that rare molecules containing heavier atoms (Si, Al, Fe) are observed in 307 308 circumstellar shells, indeed prior to their condensation in solids. The molecular complexity in the gas phase appears rather limited. Molecular abundances decrease 309 roughly by a factor around 3 for each additional atom in the detected species. This 310 reflects a basic fact of interstellar chemistry: reactions are always two-body ones due 311 to the rather low density of the ISM. As a result, destruction routes prevail in the gas 312 on the contrary to what happens on grains surfaces that offer both a high local 313 density for further reactions and a protection able to mitigate the damaging effect of 314 energetic processes on molecular assemblies of increasing size and complexity. 315

316

[Insert Table X.1 here]

Interestingly, the initially larger abundances and availability of the CHNOPS in the 317 gas phase will determine the initial chemistry in dense molecular clouds and lead to 318 the abundant simple ices widely observed in regions of star formation. As 319 conjectured by Oort and Van de Hulst as far back as 1947⁵⁴, these atoms passing 320 from diffuse to denser clouds, in the presence of cold surfaces (10–30 K) and in such 321 a reducing H rich environment will stick onto the grains and easily react to form 322 hydrides like H₂O, CH₄ and NH₃. H₂ is formed, thanks to the grain surfaces, 323 324 throughout this accretion phase, during which specific molecules such as CO, HCO, HCO⁺ also form via ion-molecule reactions in the gas phase, then also stick onto the 325 same grains where they join the ices as CO, CO₂, H₂CO, and finally the abundant 326 CH₃OH resulting from the hydrogenation of CO. 327 The main interstellar ice molecules include by order of decreasing abundance H_2O_1 , 328

329 CH₃OH, CO, CO₂, NH₃, CH₄ and H₂CO^{55, 56}.



Energetic processing of interstellar ices from UV photons and cosmic ray particles 330 that are always present to some extent in molecular clouds, has triggered 331 332 simulations through irradiation experiments on laboratory analogues since the mid 80's. Ice analogues can be produced and fully studied in these experimental 333 simulations. Initial ice analogues are irradiated by hard ultraviolet photons or ionizing 334 particles. Molecules break-up and randomly recombine, generating a rich radical 335 chemistry since these atoms and radicals diffuse in the ice, particularly when the 336 temperature of the sample is slowly raised up to room temperature. The goal of 337 these experiments is to simulate the fate of icy materials that originate in molecular 338 clouds, are irradiated by various energetic particles during their journey to 339 protoplanetary disks, then are incorporated into small bodies of the solar system that 340 may finally deliver these on the Earth where some are recovered even nowadays as 341 meteorites. 342

X. 2. 3. 3. Organics in small bodies of the solar system and their relevance to the origin of life

If cold ices in the line of sight to protostellar objects reveal simple molecules and 345 some evidence for energetic and thermal processing of them, observational 346 complications arise from the fact that cold foreground ices mask efficiently warmer 347 regions where sublimation takes place. The expected formation of an organic 348 residue, as routinely achieved in laboratory simulations⁵⁷, is thus not easily 349 observable. The residue characteristics must then rely on these empirical 350 simulations approaches that provide analogues of organics to be compared with the 351 ones present in small bodies of the solar system and further on in carbonaceous 352 meteorites. The recovered residue samples can be studied by diverse and 353 complementary methods of analytical chemistry that involve usually infrared in-situ 354 Royal Society of Chemistry – Book Chapter Template



spectroscopy, mass spectrometry coupled to gas and liquid chromatography⁵⁸ as 355 well as more recent and elaborated methods using ultra-high resolution mass 356 spectrometry from FT-Orbitrap-MS and FT-ICR-MS analytical apparatuses. Infrared 357 in-situ spectroscopies at 300 K of the residues reveal that they contain numerous 358 and highly functionalized organic molecules containing side groups and chemicals 359 functions such as methyl and methylene groups, alcohols, esters, ethers, carboxylic 360 acids, cyanides, amines, amides, aromatics⁵⁹. A very peculiar molecule 361 hexamethylenetetramine (HMT) was identified by infrared spectroscopy in these 362 residues as early as 1995 by Bernstein et al.⁶⁰. More specific and targeted searches 363 for molecules believed to be of prebiotic interest were focused on amino acids, 364 sugars or nucleobases within these organic residues. A high number of racemic 365 amino acids (up to 18) were found by Munoz-Caro et al. in 2002⁶¹ as well as di-366 amino acids and even some dipeptides⁶². In 2016, Meinert et al.⁶³ evidenced a suite 367 of many sugars up to aldopentoses was recorded in comparable simulations. Ribose 368 was found in the residues, followed just after by the detection of deoxyribose⁶⁴. On 369 the other hand, non-targeted analyses using very high resolution mass spectrometry 370 revealed a very large molecular diversity that is also observed in extracts from 371 chondrites²⁶. 372

An important issue related to this experimental approach is the possible connection between these organic residues with the organic content of natural objects, namely with debris of the solar system formation such as comets or asteroids.

 All these identified molecules as well as less defined organic materials from
 laboratory residues have indeed been detected and characterized in carbonaceous
 meteorites originating from the fragmentation of parent bodies. FT-ICR-MS analyses
 of Murchison extracts²⁶ revealed the presence of tens of thousands of organic Royal Society of Chemistry – Book Chapter Template



compounds in meteorites, similarly to what is observed in the synthetized organic residues in the laboratory^{27,65}. Targeted analyses on the soluble organic fraction of some meteorites have shown the presence of nucleobases as well as up to 80 amino acids. Furthermore, the quite abundant HMT molecule in residues was recently detected in the Murchison meteorite⁶⁶, confirming that the photo- and thermo-chemistry of dense molecular ices⁶⁷ provides a reasonable explanation for the origin of a fraction, at least, of the organic matter of meteorites.

An important feature of this extra-terrestrial organic matter lies in its reactivity. It is 387 worth emphasizing the role of the kinetic barriers associated with covalent bonds that 388 are on the one hand involved for the long-term stability of organic matter in solid 389 bodies in space and, on the other hand, have been considered as key parameters in 390 the further self-organization process leading to life^{44–47}. This importance of kinetic 391 barriers with respect to these two unrelated processes, namely the formation of 392 393 organic solids in space and the self-organization of life, makes an unexpected connection between the chemistry of interstellar media and the habitability of the 394 surface of the Earth. The development of life could have opportunistically taken 395 advantage of the presence of diverse and abundant organic matter delivered from 396 space in conjunction with liquid water though, by itself, it is not the only determining 397 factor triggering the origin of life, as mentioned earlier. Neither the fall of meteorites 398 or comets, or alternative processes for the local production of organic matter (in the 399 atmosphere or in hydrothermal environments), are likely sufficient to induce the self-400 organization of matter. Dealing with these issues requires a deeper analysis of how 401 self-organization of life could occur. 402

In summary, since stars and planetary systems originate from the formation and
 subsequent evolution of molecular clouds, the organic matter found in minor bodies
 Royal Society of Chemistry – Book Chapter Template



resulting from the described process constitutes, at least partially, a remnant of the 405 one present before the accretion of the interstellar grains to form, in a sequence, 406 407 interplanetary dust disks, then planetesimals in the protoplanetary disk that will transform into planets but leaving various debris, the so-called minor bodies. The 408 whole process eventually leads to the meteorites analysed after they reached the 409 Earth's surface. For telluric planets in the inner solar system like the Earth, volatile 410 elements have most probably been lost⁶⁸. The delivery of water and organic matter 411 through the fall of these small bodies after the accretion process might turn out to be 412 a decisive parameter of their future ability to sustain the processes leading to life. 413

414 X. 3. The Self-organization of Life

415 X. 3. 1. Diversity, Complexity and Organization

In the extant world, actually nothing can be considered at equilibrium. Even the 416 elemental composition of nebulae and other objects in the Universe does not 417 correspond to equilibrium mixtures. The strong barriers that inhibit reactions between 418 nuclei actually thwarted the evolution of mixtures of gas that were initially mainly 419 constituted by hydrogen and helium. Interestingly in the context of the study of self-420 organization and as already pointed out by Eigen and Schuster²⁵, some of the 421 processes that rule the evolution of elements in hot stars actually correspond to 422 catalytic cycles as the Bethe and von Weizäcker cycle. Kinetic features are likely to 423 play a role in the self-organization of organic matter as well. Analyses of abiotic 424 organic matter found in meteorites²⁶ or formed under simulated conditions^{27,65} 425 demonstrate a wide molecular diversity exceeding the one observed in living 426 systems^{69,70}. For instance, its diversity in terms of isomer compositions should 427 correspond to the maximal degree allowed by stochastic mechanisms of formation 428



(including regio- and stereo-isomerism or chirality) rather than the more reduced set 429 driven by cellular activities. Biochemical systems have to be considered as complex 430 in term of functional organization, while abiotic organic matter is complex in term of 431 molecular diversity. Therefore, the transition from abiotic to living corresponds to 432 processes that drove the chemistry to a higher degree of organizational complexity 433 and a reduction of structural diversity. It has been suggested that high molecular 434 diversity in complex mixtures might be the source of functional properties⁷¹. 435 However, the processes through which such features of self-organization might be 436 437 brought about remain unknown. By contrast, the high degree of diversity in abiotic organic matter suggests that any pathway for self-organization must involve 438 selection. This requirement for selection is in agreement with our above-made 439 observation that the mere presence of the elements necessary for life at the surface 440 of the Earth may be considered as a result of their ability to be involved in the 441 formation of solid organic matter and thus to overcome the volatility of their lower 442 mass combinations. The late delivery of these elements as meteorite components 443 may constitute a way to overcome the loss of CHNOPS components, which are 444 essential for life, during the accretion process. Selection processes may actually 445 have played a much broader part in astrochemistry. It has been for instance 446 demonstrated that chiral imbalance can be observed under the influence of circularly 447 polarized light in simulated interstellar ices⁷². The observation of enantiomeric 448 excesses exceeding the outcome of these processes⁷³ also suggests that, in their 449 parent bodies, the organic matter of meteorites should have endured chemical 450 transformations leading to chiral amplification. Such processes, initially suggested 451 from the study of mathematical models^{74,75} and supported by experiments in the 452 laboratory (though not in a context compatible with life as we know it) by Soai et al.⁷⁶, 453



454 can obviously be classified as kinetic selection acting by driving the composition of
455 organic matter away from the statistical distribution expected from equilibrium
456 thermodynamics.

X. 3. 2. On the Possible Occurrence of Chemical Determinism as a Driving
 Force for Evolution

The development of living organisms, as that of embryos or the growth of 459 populations of cells, corresponds to processes in which the system increases its 460 461 overall degree of organization, in contradiction with the direction of evolution signified by the Second Law. These processes must then be coupled to the dissipation of 462 energy so that the overall entropy does not decrease. The systems must therefore 463 be open to exchanges of matter and energy and receive low-entropy forms of 464 energy²³. Understanding the processes through which the system achieves self-465 organization in a way coupled to energy dissipation is therefore essential to identify 466 the pathways through which life could have emerged. In evolutionary biology, this 467 process has been identified as an improved ability to give rise to a numerous 468 offspring in population of variants submitted to natural selection¹⁵. A logical 469 consequence would be to consider that natural selection was present and effective 470 at very early stages in the process. However as suggested by Pross¹⁶, there is a 471 472 logical issue in using concepts defined at a higher hierarchical level of complexity (within the context of biology for instance) to explain phenomena occurring at a lower 473 hierarchical level of integration of science (in chemistry for instance). Any circular 474 reasoning of this kind should be avoided when trying to explain how life proceeds. 475 However, it remains that the occurrence of more or less efficient kinetic selection 476 processes might be a first clue for the identification of self-organization processes. 477



The question of the driving force that could have preceded natural selection, before 478 life was established, has not received a definitive answer, yet. In order to answer this 479 question, the kind of evolutionary force, which can direct the systems in a direction 480 opposite to the imperative of the Second Law has to be identified. Since the variety 481 of biomolecules found in living organisms is limited compared to that produced by 482 abiotic synthesis, it is very likely that self-organization processes leading to life 483 involved selection of a kinetic kind, to avoid the combinatorial explosion leading to 484 the wide diversity that abiotic molecular chemistry can statistically produce. This 485 possibility is fully consistent with a principle of evolutionary continuity since by 486 essence natural selection is based on kinetics¹⁶. This observation suggests that the 487 driving force for the self-organization of life would rather be revealed by focusing on 488 the kinetics according to which change is produced in the corresponding system, 489 rather than by searching for any other chemical properties of biomolecules. This 490 491 emphasis made on kinetics constitutes the most important point that will be developed in our approach. The search for deviations to statistical distributions 492 between or within families of organics should therefore constitute a criterion for the 493 identification of these processes. 494

Nevertheless, other possibilities of driving forces have been considered in order to fill the gap between usual chemical systems in which evolution is ruled by the second law and biological ones in which natural selection is responsible for organization. No clear evidence in favour of a possibility of an independent driving force can be found in the historical record of the biological world. For instance, it can be suggested that an evolutionary process towards the equilibrium state might have led self-

organization at early stages. Hypotheses corresponding to that possibility have been
 proposed. One of them argues that some selectivity may be induced by the presence
 Royal Society of Chemistry – Book Chapter Template



of minerals in a given environment. This is the case for instance for the formation of 503 ribose, a low yield product of the formose reaction while it constitutes an unstable 504 intermediate in the formation of valuable nucleic acid building-blocks⁷⁷. Borates have 505 been proposed to solve this issue through the formation of covalent adducts⁷⁸. 506 However, though not impossible, the eventuality that the adduct corresponds to a 507 state in which a compound like ribose can be both protected from breakdown and 508 reactive in selected pathways is improbable. The occurrence of similar processes at 509 the same time and in a same place for different components of a protometabolism 510 actually corresponds to an almost zero probability. 511

Another hypothesis involves self-organization processes reminiscent to 512 crystallization in which the entropy loss associated with an increase in organization is 513 compensated by the formation of numerous weak but favourable interactions. For 514 instance, the replication of layers in clays⁷⁹ has been considered as a way of 515 reproducing information. How the arrangements formed in this way could have 516 evolved further into a system based on the polymerization of chemical building 517 blocks remains, however, merely speculative. We therefore have a limited 518 understanding of the pathways through which a deterministic process of evolution of 519 matter driven by chemical affinities could have been converted into the process of 520 natural selection, in which occasional replication errors can be selected as a source 521 of self-organization. 522

Other proposals consider that systems in a far from equilibrium state may undergo
 evolution towards self-organization as a result of processes like dissipative
 adaptation or of a tendency to maximize energy dissipation or entropy production^{80–}
 ⁸⁶. These hypotheses are consistent with the observation that life is associated with
 processes occurring in far from equilibrium situations. Living organisms must
 Royal Society of Chemistry – Book Chapter Template



therefore behave as open systems fed with low entropy forms of energy to avoid a 528 violation of the second law²³. Energy dissipation is therefore associated with these 529 processes. Living organisms indeed correspond to the definition of dissipative 530 structures and their presence induces an increase in dissipation as they exchange 531 energy and matter within an environment characterized by a strong disequilibrium. 532 Much work has been carried out to identify a principle of maximization of energy 533 dissipation as a driving force for the self-organization of living organisms as well as 534 that of the whole biosphere⁸²⁻⁸⁶. These theoretical approaches point towards a 535 536 maximization of the amount of heat or entropy released by the system, which characterizes changes in macroscopic state functions. They are in contradiction with 537 the view developed subsequently that understanding the process requires analysing 538 what takes place at the microscopic scale. The important parameter would therefore 539 rather lie in the potential of the source of energy able to drive the system than in the 540 scalar amount of heat dissipated^{87,88}. This property is just related to the fact that 541 energy is quantified when considering events occurring at the scale of particles, 542 atoms or molecules. Even if macroscopic thermodynamic state functions can be 543 helpful in predicting the evolution of systems near equilibrium, the role of kinetics 544 becomes essential at a larger distance from equilibrium⁸⁹. It must be emphasized 545 that there is no relationship between the thermodynamics and the kinetics of 546 chemical processes involving species bound by covalent bonds. In other words, the 547 kinetic barriers between reactants and products are not related to the relative 548 energies of the initial and final state of a reaction. Therefore, macroscopic 549 thermodynamic state functions could hardly govern the evolution of a far from 550 equilibrium system towards self-organization⁹⁰. 551



Independently of this issue, laws effective at the macroscopic scale could hardly be 552 helpful in understanding processes occurring at the microscopic scale like variation 553 in the structure of metabolites and catalysts as well as that of polymers that gain a 554 property of information carriers. Apart from determinism, the conditions present on 555 the surface of a planet are likely to be mostly the result of contingent events. These 556 events may lead to specific arrangements of matter or specific conditions (e.g. the 557 presence of both liquid water and organic matter at the surface of a telluric planet). 558 As mentioned in the preceding sections, these specific arrangements may be the 559 560 result of phenomena occurring at different stages (like the interstellar formation of ice grains and their further accretion into larger bodies). Though the relationship 561 between the two processes can have an explanatory power to account for these 562 contingencies, considering the two processes as a linear progress of a deterministic 563 kind corresponds to teleological views that must be firmly avoided in modern 564 science, the very essence of the Copernican revolution¹⁶. 565

566 X. 3. 3. Identifying the Actual Drivers for Self-organization

Finding a proper definition of life is a difficult issue and much work has been devoted 567 to this end⁹¹. Moreover, such a goal does not necessarily rests on well-defined 568 physical grounds and any definition would rather be considered as arbitrary than 569 570 objective. It is very likely that the long process starting from simple self-reproducing chemical systems and leading to a fully equipped cell involved many intermediary 571 stages. Selecting a given step in that process, as corresponding to the emergence of 572 573 the first living organism, may certainly not be straightforward and would most probably lead to a lengthy debate. It seems therefore not mandatory to address this 574 issue in an experimental origin of life perspective. Focusing on the driving forces that 575



rule the process would most likely be more important than on the specific nature of 576 the molecular or supramolecular structures through which the process ensued. 577 By contrast, when considering evolutionary biology, the driving force is clearly 578 identified as natural selection operating within populations of variants. Natural 579 selection is of essentially kinetic essence. It would be conform to the continuity 580 principle that the emergence of natural selection was the outcome of an earlier 581 process of selection based on kinetics. The wide diversity of abiotic organic matter 582 makes most of its components useless for the self-organization of life. The 583 occurrence of kinetically favourable processes, orientating synthesis towards that of 584 selected useful species rather than to that of a wide diversity (as a result of catalysis 585 for instance), should be highly favourable to its development. Kinetic features, 586 including catalysis, can indeed be efficiently selected just because they take place 587 rapidly (Figure X.2). 588

589

[Insert Figure X.2 here]

The idea that natural selection evolved from an earlier selection process of kinetic 590 nature was one of the main bases of the proposal of Dynamic Kinetic Stability (DKS) 591 in replicating systems by Addy Pross^{16,92}. This essential feature of systems able to 592 reproduce themselves, whatever their nature – chemical replicators or autocatalytic 593 networks of reactions – lead the most efficient of them to be selected among variants 594 just because they produce more of themselves. The idea that systems able to 595 multiply themselves express a specific kind of stability has been developed by Pross 596 since 2003^{16–19}. The essential property of these systems is that, provided that 597 conditions remain favourable to their survival, they can only evolve in the direction of 598 an increased kinetic stability. The concept of DKS in replicating system was 599



additionally proposed to account for the emergence of functions in the living world as 600 well as its teleonomic character through which changes leading to an improvement 601 of these functions appear as goal-directed^{93,94}, even though it is only the result of 602 chance and selection. It must however be expressed clearly that nothing in the 603 above description is in contradiction with the Second Law since a condition for the 604 expression of DKS is the presence of a disequilibrium such that reproduction can 605 take place according to an exponential law. Within a specific set of conditions 606 involving mainly a far from equilibrium situation, the evolution of systems able to 607 608 reproduce themselves do not follow a path leading to thermodynamic equilibrium characterized by an increased probability²⁰. On the contrary, provided that they are 609 fuelled to maintain the far from equilibrium situation, they can lead to the kinetic 610 selection of most efficient replicators than can be considered as the most kinetically 611 stable²⁰. This behaviour is closely associated with a general notion of stability, 612 different from its more specific thermodynamic meaning in physical sciences, which 613 can be understood as time stability or persistence²⁰. Tautologically, this definition of 614 stability implies that spontaneous processes can only lead the system to evolve in 615 one direction, that of more stable (more persistent) states, until they reach a situation 616 where they can no longer evolve. 617

The important consequence for the origin of life is that DKS in replicating systems
constitutes the fundamental driving force for their evolution. It can lead reproducible
systems to evolve in the direction of an improved ability to make more of themselves,
provided that possibilities of variations are present in the systems. The intimate
nature of the system is not important when considering this kinetic feature.
Autocatalytic reaction networks or template replicators are likely to behave similarly.
The only condition is that their growth takes place through an exponential law^{25,95}.
Royal Society of Chemistry – Book Chapter Template



Deviations to exponential growth have indeed been identified as limitations to this behaviour as in the template replication of polymers^{96–98}.

The requirement for exponential growth is also an essential feature of living 627 organisms. This is for instance observed in the irreversible character of reproduction, 628 meaning that the reproduction cycle cannot unfold in the reverse direction (Figure 629 X.3)⁸⁸. Transitions in the cell cycle or in the development of living organisms indeed 630 correspond to irreversible processes without any possibility of reversion into earlier 631 stages. This feature constitutes a drastic difference with the usual approaches. 632 through which physics describe changes in material systems that are commonly 633 simplified to ideal reversible processes before any analysis⁸⁸. For instance, the 634 motion of particles is governed by the principle of microscopic reversibility meaning 635 that it should unfold in the reverse direction upon time reversal. Irreversibility in 636 physics is usually missing at the microscopic scale but classically associated with the 637 statistical description of macroscopic systems as a consequence of Boltzmann's 638 analyses²⁰. Making a process irreversible at the microscopic scale requires 639 additional features to be present⁹⁹. To proceed irreversibly, an autocatalytic loop 640 must be fed with energy (Figure X.3) and the reversal of the intermediate I into 641 reactants must be avoided by kinetic barriers^{100,101}. An estimate of the height of the 642 kinetic barrier can be made by considering that the reaction reverting the reactants 643 from intermediate I is slow at the time scale of the reproduction process (generation 644 time). Under these conditions a favourable reproducible variation in the structure of 645 catalyst C would become predominant in a population of replicating systems 646 independently of its initial abundance, provided that the replication process leads to 647 exponential growth⁹⁵. 648

[Insert Figure X.3 here]



This analysis shows that the occurrence of a process mimicking natural selection in 650 simple chemical autocatalytic systems has to meet requirements in terms of distance 651 from equilibrium, kinetic barriers of the reverse process^{47,100,101}, as well as in terms of 652 the quanta of energy provided to the system to surpass these barriers⁸⁸. As soon as 653 these conditions are met and that possibilities of variation are available, the system 654 can only evolve in the direction of an increased kinetic stability, which may include 655 features of organization¹⁰². In other words, the maximization of DKS becomes a 656 driver for evolution in the direction of improved systems that may involve emergent 657 possibilities of self-maintenance and other features of self-organization. At a certain 658 stage these systems might be considered as living according to different arbitrary 659 criterions based for instance on the presence of a metabolism and of a sub-system 660 allowing the storage of information and its use in the reproduction cycle. 661

K. 3. 4. The Importance of Kinetic Barriers for Kinetic Selection

It therefore turns out that a condition for a system to undergo a selection process 663 working in a way analogous to natural selection rather lies in a limited spontaneous 664 reactivity of activated intermediates than in the presence of catalysts merely making 665 reactions faster. In other words, kinetic selection requires the system to evolve 666 slowly toward the equilibrium state unless no improvement can be made. In biology, 667 this is the case for instance of energy carriers like ATP that are spontaneously 668 almost unreactive at the time scale of usual biological processes but that can 669 undergo fast reactions as a result of enzyme catalysts selected and optimized by 670 evolution¹⁰³. This condition has been clearly stated by Albert Eschenmoser 671 repeatedly^{44–46}. He considered the prerequisite of "a chemical environment held in a 672 non-equilibrium state by kinetic barriers" ⁴⁴, as a condition for self-organization in 673 chemical systems. As soon as this condition is met, processes that can circumvent 674 Royal Society of Chemistry – Book Chapter Template



the kinetic barriers (catalysis or autocatalysis) can develop as features of self-675 organization¹⁰². It turns out that covalent bonds, stronger than many other weak 676 chemical interactions, are associated with significant kinetic barriers. Indeed, any 677 possibility of reaction requires the covalent bonds to be cleaved to a certain degree 678 before a new bond can be formed. Other types of chemical bonds like electrostatic 679 interactions or weak bonds like hydrogen bonds of Van der Waals forces can usually 680 be expressed without necessitating a cleavage of other interactions at a reacting 681 centre. Kinetic barriers, associated with covalent bonds, were considered in previous 682 sections of this review as essential for the stability of organic matter in interstellar 683 space. They may have allowed the delivery of the elements from which organic 684 matter is made at the surface of a telluric planet like the Earth, which may otherwise 685 be missing due to their volatile character. 686

Kinetic barriers are similarly found as critical in the self-organization of life as well. 687 This unanticipated influence of kinetic barriers on two otherwise non-related 688 processes occurring in the evolution of matter in the Universe provides some 689 grounds to the relationship between the abiotic formation of organic matter in space 690 or in the environment of planetary bodies and the development of life. Strong kinetic 691 barriers associated with covalent bonds are needed in both processes as 692 determining factors allowing, on the one hand, for the kinetic selection of stable 693 chemical species in space and, on the other hand, for that of stable features of self-694 organization. 695

The identification of a relationship between the generation time of a self-reproducing chemical system and the possibility of initiation of a process similar to natural selection leads to that of a quantitative threshold allowing for the emergence of kinetically based self-organization processes. The occurrence of a cycle working Royal Society of Chemistry – Book Chapter Template



irreversibly as in Figure X.3 (right) is dependent on the barrier of the reverse 700 reaction, which requires being high enough. The dissipation of the free energy 701 702 content of an energy carrier, i.e. its conversion into heat, to which the cycle is coupled, ensures the necessary irreversibility. The barriers can be assessed using 703 the Eyring equation relating temperature, kinetic barriers and rates constants. The 704 values of these three parameters are therefore bound to each other for systems in 705 706 which selection of the more kinetically stable variants is effective, i.e. evolving according to a maximization of replicating DKS. An assessment was made using the 707 following reasonable hypotheses: 708

Considering the temperature at which water is in its liquid state, which is
 essential for the development of its specific properties of hydrophobic
 interactions¹⁰⁴.

- Considering generation times of seconds to 100 years

This assessment resulted in an order of magnitude of the kinetic barriers of ca. 100 713 kJ mol⁻¹ (Refs. 100,101), close to 0.96 eV, expressed using a unit referring to 714 715 microscopic events. Taking into account the need of additional energy for the system to perform chemical work a value of ca. 150 kJ mol⁻¹ (1.45 eV) has been assessed 716 for a chemical reproduction cycle based on simple autocatalysis. Feeding an 717 autocatalytic cycle with several guanta of energy having a lower potential could not 718 be considered as equivalent⁸⁸. A cycle fed by such degraded forms of energy cannot 719 proceed irreversibly unless a very specific process is available for upgrading these 720 quanta into a more potent one and thus for compensating for the associated entropy 721 loss. This compensation is required to meet the Second Law and can only result 722 from the operation of a kind of thermodynamic engine⁸⁸. 723



The emphasis made on the importance of kinetic barriers as a prerequisite for self-724 organization was above noticed as an indication that, in principle, catalysis should 725 not be crucial for an evolutionary process ruled by an increase in DKS. This 726 statement seems in contradiction with the usual description of biochemical 727 processes that assigns an enzyme catalyst to almost every unitary step. Moreover, 728 729 the evolution of enzyme structures constitutes an ideal example in which the efficiency of genetically encoded biochemical processes can be monitored and 730 optimized^{105,106}. Indeed, enzyme efficiency is classically described by the 731 complementarity of the enzyme structure with that of the transition state of the 732 reaction stabilizing the later^{107,108}. Autocatalysis can indeed arise when a genetically 733 encoded product or intermediate of a reaction constitutes a catalyst for its own 734 formation. Evolution can therefore make a selection among a set of different variants 735 acting on this process (Figure X.4, left panel). The presence of a well-identified 736 737 catalytic step is not needed for observing autocatalysis (Figure X.4, right panel). There are others and actually simpler means through which autocatalysis can be 738 brought about. Actually, the autocatalytic property of a chemical reaction network is 739 rather related to a stoichiometric behaviour than to a catalytic one. The requirement 740 for an autocatalytic cycle lies in the outcome of one round of the cycle that results in 741 the formation of more than a stoichiometric amount of a given component (Figure 742 X.4, right). As a result, the total amount of each component of the cycle should 743 increase as the reaction proceeds, which can be observed from a variety of network 744 architectures^{109,110} without any need for a direct catalytic step. 745

746

[Insert Figure X.4 here]

The mechanism of the formose reaction¹¹¹, converting formaldehyde into

carbohydrates, indeed constitutes a classical example of network in which
 Royal Society of Chemistry – Book Chapter Template



autocatalysis does not result from a clearly identified catalytic step (Figure X.5). 749 However, the condition that autocatalysis can be observed provided more than one 750 entity of a component of the network is produced at every round of the cycle should 751 take into account side-reactions that have been omitted on Figure X.5. This 752 drawback introduces a limit on the complexity of the network but may also be 753 compensated by cross-catalysis or other additional features of complexity. In relation 754 with this observation, it is worth mentioning that catalysis can be brought about by 755 the activity of organic compounds^{112,113} and is not limited to metallic ions or surfaces 756 that can hardly be considered as components of a network of organic reactions. 757

758

[Insert Figure X.5 here]

The example of the formose reaction illustrates how the occurrence of catalysis by 759 organics, organocatalysis^{112,113}, may have played a specific role in the selection of 760 positive feedback before the emergence of encoded catalysts (Figure 3). Some of 761 these processes can be highly efficient. For instance, intramolecular reactions, in 762 which neighbouring groups are located in a proper position to react, can benefit from 763 a substantial part of the rate enhancement provided by enzymes¹¹⁴, which has been 764 assigned to an entropic origin¹¹⁵. Certain organocatalytic processes, involving a 765 transient formation of a covalent adduct, can take advantage of a similar situation. 766 This process of induced intramolecularity^{116–119} can result in rate increases 767 representing several orders of magnitude, which may be easily available in a 768 prebiotic context¹¹⁶. Though having a scope limited to compounds presenting a 769 favourable structure, the advantage of this kind of catalysis in the origin of life 770 context is its simplicity that does not require the occurrence of multiple weak 771 interactions with transition states in a properly defined three-dimensional structure 772 that is responsible for enzyme efficiency¹²⁰. It may also have orientated chemical 773 Royal Society of Chemistry – Book Chapter Template



evolution in the direction of multifunctional compounds like amino acids or
nucleotides for which some reactions can take place through 5- or 6-membered
cyclic transition states. However, the disadvantage of these systems as well as other
metabolic pathways is their limited evolvability, which should have been overcome
by the recourse to polymers of defined sequence to allow for a wider range of
variations^{121,122} and eventually open-ended evolution¹²³.

780 X. 3. 5. Favourable Environmental Conditions for the Origin of Life

781 Liquid water is commonly considered as a prerequisite for life. There is substantial support to this assumption that includes its large abundance in the Universe as 782 justified by cosmic abundance constraints given above, the wide temperature range 783 in which it remains liquid, its very specific properties in inducing hydrophobic 784 interactions, or its ability in establishing hydrogen bonds and allowing proton 785 transfers¹⁰⁴. The requirement for liquid water defines a temperature range for the 786 occurrence of life based on organic matter. However, life is not compatible with the 787 whole range at which water is present as a liquid (below 647 K, the value of its 788 critical point). No life form has ever been observed to grow above 122 °C (395 K) 789 and most living beings are adapted to temperatures below 50°C (323 K), with the 790 exception of thermophilic and hyperthermophilic microorganisms¹²⁴. Only the lower 791 792 range in which water is liquid is therefore favourable to life as we know it and very probably to its origin. The requirement for kinetic barriers sufficient to hold the 793 system in a far from equilibrium state^{44–46} may account for this observation. Indeed, 794 raising the temperature from 300 to 400 K requires the barriers to be increased by 795 ca. 40 kJ mol⁻¹ for the system to have a similar lifetime^{100,101}. The rates of most 796 chemical reactions of covalent bonds increase in a non-selective way at high 797 temperatures rendering more difficult to meet the condition on kinetic barriers. 798 Royal Society of Chemistry – Book Chapter Template



Breakdown processes leading to the equilibrium state can indeed become 799 predominant at the expense of very specific metabolic pathways that allow living 800 801 organisms to avoid the relaxation towards the equilibrium state. The lower range of temperatures at which water remains in the liquid state (below or close to 300 K) 802 therefore constitutes the *most favourable* conditions for the origin of life. Under these 803 804 conditions, the relationship between the height of the kinetic barriers holding the chemical environment at distance from equilibrium and the potential energy feeding 805 the system led to consider a potential equivalent to that of visible or UV light as a 806 requirement. Under these conditions, reproducing chemical entities made of covalent 807 bonds with generation time equivalent to that of most living beings on Earth^{100,101} can 808 undergo a process mimicking natural selection. The set of parameters values 809 (temperature, energy and timescale) deduced from the above analysis is matching 810 the conditions that were likely present at the surface of the early Earth. Even though 811 812 an origin of life under less favourable conditions cannot be ruled out, this correspondence must not be considered as fortuitous and is on the contrary an 813 indication of the relevance of our analysis. It is worth noting, once again, that the 814 kinetic barriers associated with covalent bonds are also responsible for the stability 815 of organic matter in interstellar space, a matter which is easily formed by irradiation 816 of icy grains. That process was indeed considered above as essential for trapping 817 otherwise volatile elements like CHNOPS into interstellar ice grains for periods 818 exceeding that of the timescale of planetary systems formation and later for their 819 delivery on the surface of rocky planets. Kinetic barriers associated with covalent 820 bonds are then involved in processes as different as the origin of life and the 821 differentiation of environments like that of the surface of the early Earth. Hence, one 822 of the most important features of organic matter is related to its kinetic behaviour. 823



The low reactivity of covalent bonds at low and moderate temperatures, on the one hand, allows the delivery at the surface of rocky planets of otherwise volatile elements under some form rich in chemical components of prebiotic potential. On the other hand, this low reactivity avoids chemical systems to rapidly evolve towards the equilibrium state opening the possibility of complex reaction pathways that may include non-linear kinetic behaviours.

830 X. 3. 1. New Directions for Research

831 It turns out that DKS in chemical systems can express its power of directing the evolution of the system towards a higher degree of stability by selecting the 832 kinetically more efficient reproducing entities when the system is held at sufficient 833 distance from equilibrium. Such a distance implies that the reproduction cycle 834 proceeds irreversibly. Experimental research in the field of the origin of life should 835 therefore concentrate on processes more or less closely related to autocatalysis or 836 including positive feedback that can express a non-linear behaviour under far from 837 equilibrium conditions when the transformation in the reverse direction is negligible. 838 However, in many cases identifying autocatalysis, or similar kinetic behaviours, turns 839 out to be difficult. Autocatalysis could be defined as the kinetic outcome of a reaction 840 cycle or network in which more than one entity could be produced after one round. 841 Its definition is therefore stoichiometric. It is worth emphasizing the need of accuracy 842 in kinetic measurements since a second order kinetic law might escape a rough 843 analysis and could be overlooked by using non-accurate analytical methods or a 844 limited range of experimental data¹²⁵. Another issue is related to the fact that no 845 theoretical rule limits the size of the chemical network required for autocatalysis. 846 Furthermore, this kinetic behaviour appears as an emergent property of the whole 847 network that requires to be analysed as such. It has even been considered that the 848 Royal Society of Chemistry – Book Chapter Template



probabilities of autocatalysis may grow faster that the size of the network¹²⁶.
However, an increased size requires higher yields to achieve the condition that the
self-replicating entity is produced in more than 1 stoichiometric equivalent at every
round. That threshold could additionally depend on various experimental parameters
including temperature, concentrations, pH, etc. Identifying an authentic *emergent*autocatalytic network therefore needs a wide range of conditions to be analysed.

Moreover, the behaviour of complex systems involving several reactants are more 855 difficult to figure out than straightforward chemical processes. This is the source of 856 one of the most important barriers hindering the identification of autocatalytic 857 processes. It lies in the fact that the whole network of reactions must be performed 858 for the autocatalytic behaviour to be expressed. Indeed, autocatalysis is a property of 859 the network not that of a single reaction, whereas the usual custom of chemists 860 consists in isolating reaction steps, usually for increasing specific yields in a 861 synthetic approach, rather than considering systems of reactions. This custom, 862 though usually highly efficient, is likely to fail when looking for autocatalytic 863 properties of networks. These remarks constitute a strong indication in favour of 864 using a *Systems Chemistry*^{127–129} approach to the emergence of self-organization. 865 This approach means that chemical systems made of different components, 866 reactions, or even sub-systems should be considered to tackle self-organization with 867 the intuition that the most important features should emerge from the combination of 868 parts of the systems rather than from a fully reductionist approach. 869

With respect to an autocatalytic behaviour, the fact that it often results from a
collective behaviour of a network means that studying the whole system could be
compulsory for its detection instead of analysing its parts separately. Several
examples involving the use of original methods for the detection of autocatalytic
Royal Society of Chemistry – Book Chapter Template



behaviour have been reported for processes that are related to the origin of life. 874 Weber¹³⁰ indeed identified an autocatalytic property of the product in the reaction of 875 trioses with ammonia by iterating reactions in the presence of the crude product of 876 preceding experiments. This procedure allows the detection of the catalytic activity of 877 a product or side-product without needing the prior knowledge of its nature and is 878 therefore adapted to the identification of autocatalysis as a systems property^{131,132}. It 879 is reminiscent of the usual way through which strains of microorganisms can be 880 isolated in biology or of serial transfer experiments in which replicable biopolymers 881 can be evolved in test tubes¹³³. A similar procedure has also been checked to 882 identify solid surface environments able to select kinetic products out of mixtures by 883 iteratively seeding the product mixture in a fresh reaction medium¹³⁴. 884

885 X. 4. Conclusions

The ideas gathered in the present chapter do not constitute an attempt to define life 886 since it is an arbitrary task. However, they provide new insights on the process 887 through which self-organization may have occurred by favouring kinetically efficient 888 pathways. The transition of kinetic selection into natural selection might therefore not 889 be seen as a sharp transition but as a fuzzy one^{13,14}. Its identification should rather 890 turn out to be arbitrary rather than sharp and clear-cut. Accounting for the whole 891 process would be easier by looking for continuity than by trying to identify sharp 892 transitions. This approach provides an explanation for the specific ability of organic 893 matter in the chemistry of life. It is not directly related to the physical properties of 894 895 organic compounds as such but to the conditions ruling their transformations or breakdowns. It turns out that a connection can be made between the formation of 896 organic matter in interstellar settings and the origin of life that is not necessarily 897 related to a very specific nature of organic matter per se not to a fully deterministic 898 Royal Society of Chemistry – Book Chapter Template



process. Conversely, it must be based on the ability or inability of organic 899 compounds to be transformed. This observation would not constitute a surprise 900 901 when considering the fact that most constituents of cells are constantly regenerated. The stability of living organisms is less related to the conservation of their 902 constituents than to their continuous exchange with the environment. It is therefore 903 related to organic reactivity, a very specific domain of organic chemistry in which 904 905 properties emerge from chemical transformations within biochemicals having structures based on covalent bonds¹³⁵. It is not the organic character of the 906 907 extraterrestrial matter that confers it a prebiotic character but it is its ability to be transformed in favourable environments and to give rise to reaction networks in 908 which complex processes associated with non-linear kinetics can occur. The range 909 of kinetic barriers allowing these properties in aqueous solution, at temperatures 910 compatible with the existence of liquid water, matches the one corresponding to 911 912 species stable for billions of years at the low temperatures found in interstellar media. This stability may be responsible for the presence of CHNOPS elements at 913 the surface of a rocky planets at the same time as liquid water and therefore of the 914 constitution of environments favourable to the origin of life. Under these conditions, 915 the kinetic selection of systems capable of undergoing exponential growth may 916 become efficient. The non-limited diversity of organic molecules allowed possibilities 917 of non-limited variation in self-reproducing systems reaching eventually systems 918 made of biopolymers allowing an evolution based on natural selection. Emphasizing 919 the kinetic features of the process may be more helpful in the search for systems 920 presenting rudimentary traits of self-organization than looking at reactions able to 921 produce biochemical building blocks. Systems investigations aimed at identifying 922 autocatalysis or at least positive feedback behaviours should therefore be favoured. 923



924	
925	
926	References
927	1. A. I. Oparin, <i>Proiskhozhedenie Zhizni</i> , Mosckovskii Rabochii, Moscow, 1924.
928	Reprinted and translated in J. D. Bernal The Origin of Life, Weidenfeld and Nicolson,
929	London, 1967.
930	2. J. B. S. Haldane, <i>Rationalist Annual</i> , 1929. 148 , 3–10.
931	3. S. L. Miller, <i>Science</i> , 1953, 117 , 528–529.
932	4. F. Wöhler, <i>Ann. Phys.</i> , 1828, 88 , 253–256.
933	5. P. L. Luisi, K. Ruiz-Mirazo, Book of Abstracts, Workshop OQOL'09: Open
934	Questions on the Origins of Life, San Sebastian (Spain), May, 2009.
935	6. H. J. Morowitz, Beginnings of Cellular Life. Metabolism Recapitulates Biogenesis,
936	Yale University Press, New haven, 1992.
937	7. R. Pascal, A. Pross, and J. D. Sutherland, <i>Open Biol.</i> , 2013, 3 , 130156.
938	8. I. Fry, in Handbook of Astrobiology, ed V. M. Kolb, CRC Press, Boca Raton, 2019,

- 939 pp. 109–124.
- 9. R.F. Gesteland, T.R. Cech and J.F. Atkins, The RNA World, 2nd edn, CSHL 940

- 941 Press, Cold Spring harbour, 1999.
- 10. G. F. Joyce, and L. E. Orgel, in The RNA World, 2nd edn, eds R.F. Gesteland, 942
- T.R. Cech and J.F. Atkins, CSHL Press, Cold Spring harbour, 1999, pp. 49–77. 943
- 11. J. D. Sutherland, *Nat. Rev. Chem.* 2017, **1**, 0012. 944



- 12. J. Peretó, in *Handbook of Astrobiology*, ed. V. M. Kolb, CRC Press, Boca Raton,
 2019, pp. 219–233.
- 947 13. G. Bruylants, K. Bartik and J. Reisse, *Orig. Life Evol. Biosph.*, 2010, 40, 137–
 948 143.
- 14. G. Bruylants, K. Bartik and J. Reisse, C. R. Chim. 2011, 14, 388–391.
- 15. C. Darwin, *On the origins of species*, John Murray, London, 1859.
- 16. A. Pross *What is Life? How Chemistry Becomes Biology*, 2nd edition, Oxford
- University Press, Oxford, UK, 2016.
- 17. A. Pross and V. Khodorkovsky, J. Phys. Org. Chem. 2004, 17, 312–316.
- 18. A. Pross, *Chem. Eur. J.* 2009, **15**, 8374–8381.
- 955 19. A. Pross, J. Syst. Chem. 2011, 2, 1.
- 956 20. R. Pascal and A. Pross, *Chem. Commun.*, 2015, **51**, 16160–16165.
- 957 21. R. Pascal and A. Pross, *J. Syst. Chem.*, 2014, **5**, 3.
- 22. A. Pross and R. Pascal, *Open Biol.*, 2013, **3**, 120190.
- 23. E. Schrödinger, *What is Life?* Cambridge University Press, Cambridge, 1944.
- 960 24. M. Eigen, *Naturwissenschaften*, 1971, **58**, 465–523.
- 25. M. Eigen and P. Schuster, *Naturwissenschaften*, 1977, **64**, 541–565.
- 26. P. Schmitt-Kopplin, Z. Gabelica, R. D. Gougeon, A. Fekete, B. Kanawati, M.
- Harir, I. Gebefuegi, G. Eckel and N. Hertkorna, Proc. Natl. Acad. Sci. U.S.A., 2010,
- 964 **107**, 2763–2768.



- 27. G. Danger, F.-R. Orthous-Daunay, P. de Marcellus, P. Modica, V. Vuitton, F.
- 966 Duvernay, L. Flandinet, L. Le Sergeant d'Hendecourt, R. Thissen and T. Chiavassa,
- 967 Geochim. Cosmochim. Acta, 2013, **118**, 184–201.
- 28. E. Wollrab, S. Scherer, F. Aubriet, V. Carré, T. Carlomagno, L. Codutti and A. Ott
- 969 Orig. Life Evol. Biosph. 2016, **46**, 149–169.
- 29. K. Altwegg, H. Basiger, A. Bar-Nun et al. Sci Adv. 2016, 2, e1600285. doi:
- 971 10.1126/sciadv.1600285.
- 30. C. Gearing, New Scientist, 27 May, 2016,
- 973 https://www.newscientist.com/article/2090696-building-blocks-of-life-spotted-around-
- 974 <u>comet-for-the-first-time/#ixzz68jP2Hw8q</u>
- 975 31. Astrobiology Magazine 1 June, 2016. https://www.astrobio.net/also-in-
- 976 <u>news/rosettas-comet-contains-ingredients-life/</u>
- 32. F. Mullie and J. Reisse, *Top. Curr. Chem.*, 1987, **139**, 83–117.
- 978 33. G. Wächtershäuser, *Microbiol. Rev.*, 1988, **52**, 452–484.
- 979 34. G. Wächtershäuser, *Chem. Biodiversity*, 2007, **4**, 584–602.
- 980 35. C. De Duve and S. L. Miller, *Proc. Natl. Acad. Sci. U.S.A.*, 1991, **88**, 1014–1017.
- 36. A. Lazcano and S. L. Miller, *J. Mol. Evol.*, 1999, **49**, 424–431.
- 982 37. L. E. Orgel, *Proc. Natl. Acad. Sci. U.S.A.*, 2000, **97**, 12503–12507.
- 38. H. J. Morowitz, J. D. Kostelnik, J. Yang and G. D. Cody, Proc. Natl. Acad. Sci.
- 984 U.S.A., 2000, **97**, 7704–7708.
- 985 39. E. Smith and H. J. Morowitz, Proc. Natl. Acad. Sci. U.S.A., 2004, 101, 13168-
- 986 13173.



- 40. K. B. Muchowska, S. J. Varma, E. Chevallot-Beroux, L. Lethuillier-Karl, G. Li and
- 988 J. Moran, *Nat. Ecol. Evol.*, 2017, **1**, 1716–1721.
- 989 41. S. J. Varma, K. B. Muchowska, P. Chatelain and J. Moran, *Nat. Ecol. Evol.*,
 990 2018, **2**, 1019–1024.
- 42. K. B. Muchowska, S. J. Varma and J. Moran, *Nature*, 2019, **569**, 104–107.
- 992 43. R. Pascal, *Nature*, 2019, **569**, 47–49.
- 993 44. A. Eschenmoser, Orig. Life Evol. Biosph., 1994, **24**, 389–423.
- 45. A. Eschenmoser, Orig. Life Evol. Biosph., 2007, **37**, 309–314.
- 995 46. A. Eschenmoser, *Angew. Chem. Int. Ed.*, 2011, **50**, 12412–12472.
- 996 47. R. Pascal, *Israel J. Chem.*, 2015, **55**, 865–874.
- 48. L. Le Sergeant d'Hendecourt, *EPJ Web of Conferences* 2011, **18**, 06001. DOI:
- 998 10.1051/epjconf/20111806001
- 49. H. Palme and A. Jones, in *Treatise on Geochemistry, Vol 1 Meteorites, Comets*
- and Planets, ed. M. Davis, Elsevier Pergamon, Oxford, 2003, p. 41.
- 1001 50. J. M. Greenberg, *Astrophys. J.*, 1974, **189**, L81–L85.
- 1002 51. J. S. Lewis, *Science*, 1974, **186**, 440–443.
- 1003 52. J. N. Tolstikhin and D. Kramers, *The Evolution of Matter: From the Big Bang to*
- 1004 *the Present Day*, Cambridge University Press, Cambridge, 2008.
- 1005 53. V. Vinogradoff, PhD thesis, Aix-Marseille University, 2013.
- 1006 54. J. Oort and H. C. Van de Hulst, Recherches Astronomiques de l'Observatoire
- 1007 d'Utrecht, 1946, **11**, 2.



- 1008 55. E. L. Gibb, D. C. B. Whittet, A. C. A. Boogert and A. G. G. M. Tielens, *Astrophys.*
- 1009 *J. Suppl. Ser.*, 2004, **151**, 35–73.
- 1010 56. E. Dartois, *Space Sci. Rev.*, 2005, **119**, 293–310.
- 1011 57. G.M. Munoz-Caro and W. Schutte, *Astron. Astrophys.*, 2003, 412, 121–132.
- 1012 58. G. M. Munoz-Caro and W. Schutte, *Astron. Astrophys.*, 2003, **412**, 121–132.
- 1013 59. M. Nuevo, J. H. Bredehöft, U. J. Meierhenrich, L. d'Hendecourt and W. H.-P.
- 1014 Thiemann, *Astrobiology* , 2010, **10**, 245–256.
- 1015 60. M. P. Bernstein, S. Sandford, L. J. Allamandola and S. Chang, Astrophys. J.,
- 1016 1995, **454**, 327–344.
- 1017 61. G. M., Munoz-Caro, U. J. Meierhenrich and W. A. Schutte, *Nature*, 2002, 416,
 1018 403-407.
- 62. R. Kaiser, A. M. Stockton, Y. S. Kim, E. C. Jensen and R. A. Mathies, Astrophys.
 J., 2013, **765**, 111.
- 1021 63. C. Meinert, I. Myrgodoroska, P. de Marcellus, T. Buhse, L. Nahon, S. V.
- Hoffmann, L. Le Sergeant d'Hendecourt and U. J. Meierhenrich, *Science*, 2016, **352**,
 208–212.
- 1024 64. M. Nuevo, G. Cooper and S.A. Sandford, *Nat. Commun.*, 2018, **9**, 5276.
- 1025 65. G. Danger, A. Fresneau, N. Abou Mrad, P. de Marcellus, F.-R. Orthous-Daunay,
- 1026 F. Duvernay, V. Vuitton, L. Le Sergeant d'Hendecourt, R. Thissen and T. Chiavassa,
- 1027 Geochim. Cosmochim. Acta, 2016, **189**, 184–196.
- 1028 66. Y. Oba, Y. Takano, H. Naraoka, Y. Furukawa, D. P. Glavin, J. P. Dworkin and S.
- 1029 Tachibana, *Nat. Comm.*, 2020, **11**, 6243.



- 1030 67. V. Vinogradoff, S. Bernard, C. Le Guillou and L. Remusat, *Icarus*, 2018, **305**,
 1031 358–370.
- 1032 68. D. C. Catling and J. F. Kastings, Atmospheric evolution on inhabited and lifeless
- 1033 *planets*, Cambridge University Press, Cambridge, 2017.
- 1034 69. P. Schmitt-Kopplin, M. Harir, B. Kanawati, R. Gougeon, F. Moritz, N. Hertkorn, S.
- 1035 Clary, I. Gebefügi and Z. Gabelica, in *Astrobiology. An Evolutionary Approach y*, ed.
- 1036 V. M. Kolb, CRC Press, Boca Raton, 2014, pp. 63–82.
- 1037 70. P. Schmitt-Kopplin, D. Hemmler, F. Moritz, R. D. Gougeon, M. Lucio, M.
- 1038 Meringer, C. Müller, M. Harir and N. Hertkorn, *Faraday Discuss.*, 2019, **218**, 9–28.
- 1039 71. N. Guttenberg, N. Virgo, K. Chandru, C. Scharf and I. Mamajanov, *Phil. Trans. R.*
- 1040 Soc. A, 2017, **375**, 20160347.
- 1041 72. P. de Marcellus, C. Meinert, M. Nuevo, J.-J. Filippi, G. Danger, D. Deboffle, L.
- 1042 Nahon, L. Le Sergeant d'Hendecourt and U. J. Meierhenrich, Astrophys. J. Lett.
- 1043 2011, **727**, L27.
- 1044 73. J. E. Elsila, J. C. Aponte, D. G. Blackmond, A. S. Burton, J. P. Dworkin and D. P.
 1045 Glavin, *ACS Cent. Sci.*, 2016, **2**, 370–379.
- 1046 74. F. C. Frank, *Biochim. Biophys. Acta*, 1953, **11**, 459–463.
- 1047 75. R. Plasson, H. Bersini, and A. Commeyras, *Proc. Natl. Acad. Sci. U.S.A.*, 2004,
 1048 **101**, 16733–16738.
- 1049 76. K. Soai, T. Shibata, H. Morioka and K. Choji, *Nature* 1995, **378**, 767–768.
- 1050 77. R. Larralde, M. P. Robertson and S. L. Miller, Proc. Natl. Acad. Sci. U.S.A., 1995,
- 1051 **92**, 8158–8160.



- 1052 78. A. Ricardo, M. A. Carrigan, A. N. Olcott and S. A. Benner, *Science*, 2004, **303**,
- 1053 **196–196**.
- 1054 79. A. G. Cairns-Smith, *Chem. Eur. J.*, 2008, **14**, 3830–3839.
- 1055 80. J. L. England, J. Chem. Phys., 2013, **139**, 121923.
- 1056 81. J. L. England, *Nat. Nanotechnol.*, 2015, **10**, 919–23.
- 1057 82. R. Swenson, Syst. Res., 1989, 6, 187–197.
- 1058 83. L. M. Martyushev and V. D. Seleznev, *Phys. Rep.*, 2006, **426**, 1–45.
- 1059 84. A. Kleidon, *Phys. Life Rev.*, 2010, 7, 424–460.
- 1060 85. A. Belkin, A. Hubler and A. Bezryadin, *Sci. Rep.*, 2015, **5**, 8323.
- 1061 86. K. R. Skene, *Entropy*, 2015, *17*, 5522–5548.
- 1062 87. A. Pross and R. Pascal, *Beilstein J. Org. Chem.*, 2017, **13**, 665–674.
- 1063 88. G. Danger, L. d' Hendecourt and R. Pascal, *Nat. Rev. Chem.*, 2020, **4**, 102–109.
- 1064 89. I. Prigogine, *Science*, 1978, **201**, 777–785.
- 1065 90. J. Ross, A. D. Corlan, S. C. Müller, *J. Phys. Chem. B*, 2012, **116**, 7858–7865.
- 1066 91. R. Popa, Between Necessity and Probability: Searching for the Definition and
- 1067 Origin of Life, Springer-Verlag, Berlin Heidelberg, 2004.
- 1068 92. R. Pascal and A. Pross, J. Syst. Chem., 2019, 7, 1–8.
- 1069 93. A. Pross, Orig. Life Evol. Biosph., 2005, **35**, 383–394.
- 1070 94. A. Pross, J. Phys. Org. Chem., 2008, **21**, 724–730.
- 1071 95. E. Szathmáry and I. Gladkih, *J. Theor. Biol.*, 1989, **138**, 55–58.



- 1072 96. G. von Kiedrowski, *Angew. Chem. Int. Ed.*, 1986, **25**, 932–935.
- 1073 97. D. Sievers and G. von Kiedrowski, *Nature*, 1994, **369**, 221–224.
- 1074 98. D. Sievers and G. von Kiedrowski, *Chem. Eur. J.*, 1998, **4**, 629–641.
- 1075 99. D. G. Blackmond, *Angew. Chem. Int. Ed.*, 2009, **48**, 2648–2654.
- 1076 100. R. Pascal, J. Syst. Chem., 2012, **3**, 3.
- 1077 101. R. Pascal, in Astrochemistry and Astrobiology: Physical Chemistry in Action,
- Eds. I. W. L Smith, C. S. Cockell and S. Leach, Springer, Berlin, 2013, pp. 243–269.
- 1079 102. A. Pross, J. Mol. Evol., 2013, 76, 185–191.
- 1080 103. F. H. Westheimer, *Science*, 1987, **235**, 1173–1178.
- 1081 104. H. Cottin, J. Kotler, K. Bartik, J. Cleaves, C. Cockell, J.-P. de Vera, P.
- 1082 Ehrenfreund, S. Leuko, I. Ten Kate, Z. Martins, R. Pascal, R. Quinn, P. Rettberg and
- 1083 F. Westall, *Space Sci. Rev.*, 2017, **209**, 1–42.
- 1084 105. F. H. Arnold, Acc. Chem. Res., 1998, **31**, 125–131.
- 1085 106. D. R. Liu, P. G. Schultz, *Angew. Chem., Int. Ed.*, 1999, **38**, 36–54.
- 1086 107. L. Pauling, *Chem. Eng. News*, 1946, **24**, 1375–1377.
- 1087 108. L. Pauling, *Nature*, 1948, **161**, 707–709.
- 1088 109. R. Plasson, A. Brandenburg, L. Jullien and H. Bersini, J. Phys. Chem. A, 2011,
- 1089 **115**, 8073–8085.
- 1090 110. A. Blokhuis, D. Lacoste and P. Nghe, *Proc. Natl. Acad. Sci. U.S.A.*, 2020, **117**,
 1091 25230–25236.
- 1092 111. R. Breslow, *Tetrahedron Lett.*, 1959, **1**, 22–26.



- 1093 112. D. W. C. MacMillan, *Nature*, 2008, **455**, 304–308.
- 1094 113. C. F. Barbas III, *Angew. Chem. Int. Ed.*, 2008, **47**, 42–47.
- 1095 114. A. J. Kirby, *Adv. Phys. Org. Chem.*, 1980, **17**, 183–278.
- 1096 115. M. I. Page and W. P. Jencks, *Proc. Natl. Acad. Sci. U.S.A.*, 1971, 68, 1678-
- 1097 1683.
- 1098 116. R. Pascal, *Eur. J. Org. Chem.*, 2003, **2003**, 1813–1824.
- 1099 117. A. M. Beauchemin, *Nat. Chem.*, 2013, **5**, 731–732.
- 1100 118. M. P. Jamshidi, M. J. MacDonald and A. M. M. Beauchemin, Orig. Life Evol.
- 1101 *Biosph.*, 2017, **47**, 405–412.
- 1102 119. K. L. Tan, *Nat. Chem.*, 2012, **4**, 253–254.
- 1103 120. W. P. Jencks, *Adv. Enzymol. Relat. Areas Mol. Biol.*, 1975, **43**, 219–410.
- 1104 121. V. Vasas, E. Szathmáry and M. Santos, *Proc. Natl Acad. Sci. U.S.A.*, 2010,
- 1105 **107**, 1470–1475.
- 1106 122. V. Vasas, C. Fernando, M. Santos, S. Kauffman and E. Szathmáry, *Biol. Direct*1107 2012, **7**, 1.
- 1108 123. K. Ruiz-Mirazo, J. Pereto and A. Moreno, *Orig. Life Evol. Biosph.*, 2004, 34,
 1109 323–346.
- 1110 124. M. T. Madigan, J. M. Martinko and J. Parker, *Brock biology of microorganisms*,
- 1111 Prentice Hall, Upper Saddle River, NJ, 2003.
- 1112 125. R. Pascal, J. Taillades, A. Commeyras, *Tetrahedron*, 1980, **36**, 2999–3008.
- 1113 126. S. A. Kauffman, *Investigations*, Oxford University Press, Oxford, 2000.



- 1114 127. M. Kindermann, I. Stahl, M. Reimold, W. M. Pankau and G. von Kiedrowski,
- 1115 Angew. Chem. Int. Ed., 2005, **44**, 6750–6755.
- 1116 128. R. F. Ludlow and S. Otto, *Chem. Soc. Rev.*, 2008, **37**, 101–108.
- 1117 129. K. Ruiz-Mirazo, C. Briones and A. de la Escosura, Chem. Rev., 2014, 114,
- 1118 285–366.
- 1119 130. A. L. Weber, *Orig. Life Evol. Biosph.*, 2007, **37**, 105–111.
- 1120 131. R. Clairmont, A. Bommarius and A. Weber, *J. Mol. Evol.*, 2015, **81**, 1–9.
- 1121 132. A.L. Weber and A.C. Rios, *Orig. Life Evol. Biosph.*, 2019, **49**, 199–211.
- 1122 133. D. R. Mills, R. L. Peterson and S. Spiegelman, *Proc. Natl. Acad. Sci. U.S.A.*,
- 1123 1967, **58**, 217–224.
- 1124 134. S. Colón-Santos, G. J. T. Cooper and L. Cronin, *ChemSystemsChem*, 2019, 1,
 1125 e1900014.
- 135. W. P. Jencks, *Catalysis in Chemistry and Enzymology*, McGraw-Hill, New York,
 1127 1969.



1129 **FIGURE CAPTIONS**

1130

1131	Figure X. 1 An abiotic reaction network predating metabolism may have been
1132	involved in the self-organization of life at different stages e.g. by feeding a prebiotic
1133	network of reactions with essential components. Notwithstanding the role of the
1134	environment in fuelling dissipative processes, organisms endowed of a higher
1135	organization degree may also have harnessed such networks by improving them
1136	through the activity of genetically encoded biocatalysts.
1137	

Figure X. 2 Features of self-organization including loops of positive feedback can be selected as a result of their kinetic advantage in way reminiscent of how natural selection proceeds. Catalysis is likely to constitute a strong player acting on kinetic selection. Simple catalysis by metals or inorganic surface may play an early role. An autocatalytic behaviour can emerge from organocatalysis and natural selection as a result, for instance, of polymer catalysts encoded by information carriers reproduced by template replication.

1145

Figure X. 3 Biological reproduction essentially proceeds irreversibly (left), whereas
an autocatalytic loop must be held far from equilibrium by feeding with low-entropy
energy to become irreversible. Reproduced from Ref. 88 with permission from
Springer Nature, Copyright 2020.



Figure X. 4 Examples of autocatalytic cycles. Autocatalytic properties can result from the catalytic ability of one component on a chemical step involved in the cycle (left).

1154

1155	Figure X. 5 An example of autocatalytic loop in the architecture of the reaction
1156	network responsible for the formose reaction ¹¹¹ . The system meets a first condition
1157	for complexity by the occurrence of kinetic barriers that hinder the direct dimerization
1158	of formaldehyde into glycoaldehyde and hold the system in a far from equilibrium
1159	state. The cleavage of tetramers or pentamers into two shorter reactive
1160	carbohydrates (retroaldolization) able to undergo the reaction with formaldehyde
1161	(aldolization) allow the system to overcome this blockade through an autocatalytic
1162	behaviour.
1163	Table X. 1. List of observed interstellar molecules. Although identified gas phase
1164	molecules dominate this list, some molecules are also observed in the solid state
1165	under the form of ices (bold). Adapted from Ref. 53 with permission from the
1166	author, Copyright [2013].
1167	

1168





1171 1172









1177





1179

1180



1182





1184 Table X. 1.

2	3	4 atoms	5 atoms	6 atoms	7 atoms	8 atoms	> 8 atoms
atoms	atoms						
AICI	AINC	C₃H	C ₅	H_2C_3O	C ₂ H ₄ O	H ₃ C ₃ N	CH ₃ C ₄ H
AIF	AIOH	C₃N	CH₄	C_2H_4	CH ₃ C ₂ H	OHCH ₂ CHO	H ₃ COCH ₃
AIO	C ₃	C₃N⁻	C_3H_2	CH₃CN	CH ₃ NH ₂	HCOOCH ₃	CH ₃ CH ₂ CN
C ₂	C ₂ H	C ₃ O	H ₂ CCN	CH₃NC	CH₂CHCN	CH₃COOH	CH ₃ CONH ₂
CF⁺	C2O	C₃S	H_2C_2O	CH₃OH	H ₂ CHCOH	H_2C_6	CH ₃ CH ₂ OH
CH	C_2S	H₃O⁺	H₂CNH	CH₃SH	C ₆ H	CH₂CHCHO	CଃH
CN	C_2P	C ₂ H ₂	H ₂ COH+	C_4H_2	C ₆ H [−]	CH₂CHCN	HC ₇ N
CN ⁻	CO2	H ₂ CN	C₄H	HC₃NH⁺	HC₅N	C7H	CH ₃ CHCH ₂
CO	H_3^+	H₂CN⁺	C₄H–	HCONH ₂	CH₃CHO	NH ₂ CH ₂ CN	(CH ₃) ₂ CO
CO⁺	H ₂ C	H₂CO	HC₃N	C₅H		C ₂ H ₆	CH ₃ CH ₂ CHO
CP	H₂O	H₂CS	HC₂NC	HC ₂ CHO		CH₃CHNH	CH₃C₅N
CS	H_2O^+	HCCN	нсоон	HC₄N			CH ₃ C ₆ H
H_2	HO ₂	HCNH⁺	NH₂CN	CH₂CNH			C₂H₅OCHO
HCI	H_2S	HOCO⁺	HC(O)CN	HNCHCN			$H_3COC_2H_5$
HF	H_2CI^+	HOCN	SiC₄				C ₆ H ₆
HN	HCN	HOOH	SiH₄				C ₃ H ₇ CN
HO	HCO	HCNO	CH₃O				HC ₁₁ N
KCI	HCO⁺	HNCO	HNCNH				$C_{14}H_{10}^{+}$
N_2	HCP	HNCS					C ₆₀
NO	HCS⁺	NH₃					C ₇₀
NS	HNC	SiC₃					CH ₃ COOCH ₃
NaCl	HN_2^+	HCOO ⁻					HC₀N
O ₂	HNO						
PN	HCO⁺						
PO	FeCN						
SH	KCN						
SH⁺	MgCN						
SO	MgNC						
SO	NH_{2}						
SiC	NH ₂						
SiN	N2O						
SiO	NaCN						
SiS	OCS						
OH+	SO ₂						
FeO	SiC ₂						
HCI+	SiCN						
TiO	SiNC						
	OCN ⁻						
	TiO ₂						