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Pressure dependence of electronic, vibrational and optical properties of wurtzite-boron nitride

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Wurtzite Boron Nitride (wBN) is a wide band gap BN polymorph with unique mechanical properties such as hardness and stiffness. Initially synthesized in 1963 by transforming hexagonal BN (hBN) under high temperature and pressure conditions, wBN can now be stabilized at atmospheric pressure to obtain high-quality samples. Our first-principles study investigates the electronic, vibrational, and optical properties of wBN across a broad range of pressures. We account for the electron-hole interaction in the optical response, revealing that this effect is crucial to interpret the available experimental spectra. We also calculate the coupling between excitons and phonons and provide a phonon-assisted emission spectrum, centered around 6.02 eV. Our results hold significant importance for the potential application of wBN as a dielectric material in BN-based technologies, especially in optoelectronics and harsh environments. We also expect that our prediction could be verified in the future, and it could aid in the identification of wBN through cathodoluminescence experiments.

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I. INTRODUCTION

Recent advances in high pressure physics have enabled the synthesis of a new range of materials that were previously impossible to create at ambient pressure [1]. Pressure has the ability to drastically alter elastic, electronic, magnetic, structural, and chemical properties, transforming materials from simple insulators to superconductors, from amorphous to crystalline solids, and from ionic to covalent compounds. Some of these transformations can be irreversible, such as the transition from graphite to diamond, resulting in the creation of new materials that are stable at ambient pressure.

Among the various materials created by the use of pressure are wide band gap insulators [2]. The term "wide band gap material" generally refers to any semiconductor with an energy band gap much larger than that of conventional semiconductors such as silicon (Si) and gallium arsenide (GaAs). Interest in this class of materials stems from the ongoing quest for smaller, faster, more reliable, and more efficient electronic devices than their Si-based counterparts. Wide band gap materials have promising applications for future generations of high power electronics, deep UV optoelectronics, quantum electronics and harsh environment (high temperature and high pressure) applications.

Wide band gap materials are the optimal choice for efficient green and blue light emitting diodes (LEDs) in opto-electronic applications. For example, gallium nitride (GaN) is used to produce bright blue LEDs [3]. At shorter wavelengths, these materials find applications in UV light germicidal sources for water sanitation and advanced solar blind optical

communication systems. Recently, hexagonal boron nitride (hBN), a wide band gap material, has attracted much attention for its very efficient UV light emission despite being an indirect band gap material [4,5]. This discovery has led to a number of theoretical and experimental studies focusing on its unexpected properties [6–13].

However, compared to currently commercialized semiconductor technologies based on Si, GaAs, and even indium phosphide (InP), applications of wide band gap materials are relatively immature. Nevertheless, significant research and development efforts are underway due to the enormous commercialization opportunities for wide-bandgap semiconductors in next-generation microelectronics and optoelectronics. In some cases, much remains to be done to accurately characterize their electronic and optical properties.

In this context, wurtzite boron nitride (wBN), which is isostructural to other group III nitrides with optoelectronics applications such as GaN, AlN, and InN, stands out due to its exceptional mechanical properties such as hardness and elastic stiffness [14,15], and its electronic properties such as high thermal conductivity and large spontaneous polarization [16,17].

Since its first synthesis in 1963, much effort has been devoted to the production of high quality samples of wurtzite boron nitride (wBN), which is obtained by subjecting hexagonal boron nitride (hBN) to high temperature and pressure conditions. For some time, the dominant synthesis method has been the shock compression technique of hBN, but unfortunately this only produces wBN powders with grain sizes in the μ m range. However, in 2019, researchers successfully synthesized pure single-phase polycrystalline wBN bulk material using wBN powders as a starting point and a well-controlled process under ultrahigh pressure (200 kBar) and high temper-

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ature (1150 °C) conditions [18]. The main technical obstacle in the $h B N \rightarrow w B N$ transformation is the stabilization of the latter at atmospheric pressure, since it tends to revert to the ambient phase due to its low kinetic barrier. Several efforts have been made to overcome this barrier and the wurtzite phase has been stabilized at atmospheric pressure. Finally, w B N has recently been obtained from h B N in an upstroke cycle from ambient pressure conditions up to 200 kBar.

Optical measurements of wurtzite boron nitride (wBN) in the IR-visible spectral energy range have shown that the hexagonal to wurtzite phase transition is complete at 130 kBar [19] and the optical features of the wBN structure are retained in a pressure down-stroke cycle back to ambient conditions. In addition, Chen et al. [20] successfully fabricated millimeter-sized wBN crystals via high-pressure, high-temperature transformation and demonstrated that the stability of wBN at atmospheric pressure is guaranteed by the presence of a 3D high-density network of planar defects.

Motivated by recent experimental progress in obtaining high-quality crystalline samples of wBN and the pressureinduced phase transition from hBN to wBN, we present here a theoretical study of the optical properties (light absorption and emission) of wBN in the UV energy range over a wide pressure range from ambient to 200 kBar. We also investigate the electronic and vibrational properties of wBN as a function of pressure to support the discussion. Our aim is to extend the understanding of the intrinsic light absorption and emission properties of wBN and to complement ongoing research on the cubic phase of boron nitride (cBN), a wide band gap material [21], which is commonly synthesized at high pressure and high temperature using a temperature gradient [22]. It is worth noting that wBN is actually an intermediate product of the phase transformation of hBN to cBN and that the synthesis of high purity cBN is hindered by the presence of various types of defects, including point defects and stacking defects, in the form of very local wBN phase domains [23]. This hypothesis has been proposed due to the existence in cathodoluminescence spectra of phonon replicas incompatible with cBN phonon modes, which on the other hand could be the signal for the presence of wBN domains [21]. For these reasons, theoretical reference spectra of pure wBN are an important contribution and will allow to elucidate the above mentioned aspects of the phase transformation.

The paper is organized as follows. In Sec. II we give an overview of the theoretical methods used. In Sec. III we present all the computational details necessary to reproduce the results discussed in Sec. IV, where we investigate the vibrational, electronic, and optical properties of wBN under hydrostatic pressure. We address the issue of the electronic and optical band gap and compare it with that of other boron nitride phases. Finally, we discuss the phonon-assisted light emission properties of wBN and speculate on possible changes in the light emission behavior with increasing pressure.

II. THEORETICAL METHODS

Using density functional theory (DFT) as implemented in the quantum espresso package [24], we investigated the ground state of the wurtzite phase of boron nitride. To optimize the unit cell and atomic positions at various pressures, we minimized the total enthalpy. Subsequently, we utilized density functional perturbation theory (DFPT) [25] to calculate the vibrational properties based on the optimized atomic structures. These methods allowed us to gain a comprehensive understanding of the structural and vibrational characteristics of the boron nitride wurtzite phase. We used PBEsol exchange-correlation functional [26] and DOJO pseudopotentials [27]. Technical details of these calculations are given in the following section.

Following our investigation of the ground-state properties, we shifted our focus to exploring the electronic structure and optical properties of wBN. To study the electronic bands of wBN, we diagonalized the Kohn-Sham (KS) Hamiltonian. However, as DFT is known to have limitations in accurately describing the electronic structure, including underestimation of the electronic gap [28], we used many-body perturbation theory implemented in the Yambo code [29] to correct the KS eigenvalues. By doing so, we calculated the quasiparticle band structure as:

$$E_{n\mathbf{k}} = \epsilon_{n\mathbf{k}}^{KS} + Z_{n\mathbf{k}} \left[Re\Sigma \left(\epsilon_{n\mathbf{k}}^{KS} \right) - V_{n\mathbf{k}}^{xc} \right], \tag{1}$$

where $E_{n\mathbf{k}}$ are the corrected quasiparticle energies, $Z_{n\mathbf{k}}$ is the renormalization factor, V^{xc} the exchange correlation KS potential, and the self-energy Σ is calculated in the GW approximation $\Sigma = G_0W_0$ [28], at frequency $\omega = \epsilon_{n\mathbf{k}}$.

Starting from the quasiparticle band structure we proceeded to calculate the optical response function, including the electron-hole interaction. To achieve this, we solved the Bethe-Salpeter equation (BSE) [30], which provides a means of describing the behavior of interacting electrons and holes. The BSE can be reformulated as an eigenvalue problem, which is determined by the two-particle Hamiltonian [28] in the following manner:

$$H_{\substack{vck\\v'c'k'}} = (E_{ck} - E_{vk})\delta_{vv'}\delta_{cc'}\delta_{kk'} + (f_{ck} - f_{vk})\left(2\overline{V}_{\substack{vck\\v'c'k'}} - W_{\substack{vck\\v'c'k'}}\right),$$
(2)

where E_{nk} and f_{nk} [where n runs over the valence (v) and conduction bands (c)] are the quasiparticle energies and occupations respectively, \overline{V} is the Coulomb potential derived from the variation of the Hartree term, and W the screened electronhole interaction derived from the screened exchange [30]. After diagonalization of Eq. (2) we obtain the eigenvectors A_{cvk}^S and eigenvalues E_S that are the eigenstates and the energies of the excitons (labeled by the index S), the stationary state of the BSE. These quantities are used to build up the macroscopic dielectric function ϵ_M as

$$\epsilon_M(\omega) = 1 - 4\pi \sum_{S} \frac{|T_S|^2}{\omega - E_S + i\eta},\tag{3}$$

where $T_S = \sum_{cv\mathbf{k}} d_{cv\mathbf{k}} A^S_{cv\mathbf{k}}$ are the excitonic dipoles, $d_{cv\mathbf{k}}$ are the dipole matrix elements between the Kohn-Sham states $d_{cv\mathbf{k}} = \langle v\mathbf{k}|\hat{r}|c\mathbf{k}\rangle$, and $i\eta$ is a small broadening term added to simulate experimental spectra.

In addition to studying the optical absorption properties of wBN, we were also interested in investigating its luminescence behavior. However, since wBN is an indirect band

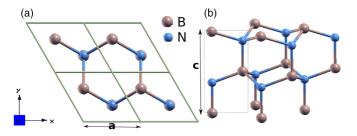


FIG. 1. Top view (a) and three dimensional structure (b) of wurtzite BN. The atoms in the unit cell are framed by a dotted line in panel (b), lattice parameter c is parallel to the z axis.

gap material, the emission of light is only possible if it is assisted by phonon modes, as has been observed in hBN [5]. Thus, it is essential to consider the exciton-phonon coupling when analyzing luminescence in wBN. There are two approaches to account for this coupling: the finite differences approach [11,12], and a direct calculation of the exciton-phonon dipole matrix elements [31]. In this manuscript, we adopted the former method, and the calculation details are provided in the following section. After obtaining the second derivatives of the exciton dipoles, we proceeded with the calculation of the luminescence spectra using a generalization of the Roosbroeck-Shockley relation [32] for the excitonic case, as derived in Refs. [12,33]:

$$R^{\text{sp}}(\omega) = \sum_{\lambda \mathbf{q}} \frac{\omega(\omega + 2\Omega_{\lambda \mathbf{q}})^2}{\pi^2 \hbar c^3} n_r(\omega) \sum_{S} \frac{\partial^2 |T_S|^2}{\partial R_{\lambda \mathbf{q}}^2} \bigg|_{\text{eq}}$$

$$\Im \bigg\{ \frac{1}{\hbar \omega - (E_S - \Omega_{\lambda \mathbf{q}}) + i\eta} \bigg\} B(E_S, T_{\text{exc}}), \tag{4}$$

where λ , \mathbf{q} are the phonon mode indexes, $n_r(\omega)$ is the refractive index, $B(E_S, T_{\rm exc})$ is the Boltzmann occupation of the lowest excitons and $\Omega_{\lambda\mathbf{q}}$ are the phonon frequencies. The formula we used provides a luminescence spectrum that accounts for the coupling with phonon modes under equilibrium conditions, where the excitons have relaxed at the minimum of their band structure.

III. COMPUTATIONAL DETAILS

In this section, we present all the necessary computational details to reproduce the results of our work using the theoretical methods outlined in the previous section. *Atomic structure*

The wurtzite phase has hexagonal symmetry. The crystal structure can be seen as two interpenetrating hexagonal close-packed (hcp) sublattices, each of them formed by one atom type. The two individual atom types are displaced along the c axis of amount z in an ABAB... stacking sequence. Boron atoms sit in the $(^{1}/_{3}, ^{2}/_{3}, 0)$, $(^{2}/_{3}, ^{1}/_{3}, ^{1}/_{2})$ crystallographic positions, while nitrogen atoms sit in $(^{1}/_{3}, ^{2}/_{3}, z)$, $(^{2}/_{3}, ^{1}/_{3}, ^{1}/_{2} + z)$ ones. The unit cell is shown in Fig. 1. We used the lattice parameters a = 2.55 Å, c = 4.22 Å taken from Ref. [34] as starting point parameters to relax the structure at the different pressures. The internal parameter z is defined as the length of the BN bond along the z axis (in the present case 1.56 Å) in units of the c lattice parameter.

We sampled the pressure range 0–200 kBar at a pace of 50 kBar. Then we relaxed the lattice parameters and the atomic positions at each pressure by sampling the Brillouin zone with a $12 \times 12 \times 8$ k point grid and 70 Ry for the kinetic energy cutoff of the plane wave expansion. The forces acting on the cell and on the atoms were converged to be lower than 10^{-5} a.u. Lattice parameters as a function of pressure are summarized in Fig. S1 and Table S1 in the SM [35]. We used the final coordinates and lattice parameters as starting points for investigating the vibrational, electronic and optical properties of wBN, that we are going to discuss in the following.

A. Vibrational Properties

Starting from the atomic geometries obtained for each pressure chosen in the range 0–200 kBar, the vibrational properties (phonon frequencies, eigenvectors, dielectric constants, and Born effective charges) of wBN as a function of pressure have been calculated. The dynamical matrices have been sampled on a regular $9 \times 9 \times 6$ **q** point grid. Then a Fourier interpolation is used to obtain the entire phonon dispersion along a **q** path connecting high symmetry points of the Brillouin zone. Being four atoms in the unit cell, 12 phonon modes appear for each **q** point.

B. Electronic band structure and quasiparticle corrections

The quasiparticle band structure was obtained starting from the KS Hamiltonian and correcting the eigenvalues within the G_0W_0 approximation [36]. We used a $12 \times 12 \times 8$ k points grid, and 75 bands for the expansion of G and G, a cutoff of 3 Ha for the dielectric constant, and the Godby-Needs plasmon-pole model [37] for the dynamical part of G.

C. Light Absorption

Optical properties of wBN have been investigated by solving the Bethe-Salpeter equation as explained in the previous section [30]. We used a $12 \times 12 \times 8$ k point grid and included three valence and three conduction bands, and the same screened Coulomb interaction of the GW calculations.

D. Phonon assisted luminescence

As it will be discussed in Sec. IV, wBN is an indirect gap material with the lowest indirect gap occurring between points Γ and K. Therefore the momentum responsible for the indirect emission would then be $\mathbf{q} = K - \Gamma = K = (1/3, 1/3, 0)$. In order to get the phonon-assisted luminescence spectra we apply the following strategy:

- (1) We diagonalized the Bethe Salpeter equation at the transferred momentum $\mathbf{q} = K$, first including and then discarding the quasiparticle corrections. In this way we estimated a scissor correction that exactly reproduces the position of the exciton at K. In Table I we report the position of the exciton at $\mathbf{q} = K$, with and without GW corrections, and the obtained energy shift.
- (2) We built a supercell which maps the point K at Γ (see Fig. 2) [38] and we calculated the dielectric constant that enters in BSE with the parameters of the primitive cell (polarization function number of bands) rescaled to the

TABLE I. Lowest indirect excitons energies at KS+BSE (1st column) and *GW*+BSE level (2nd column) and the estimated scissor correction (3rd column).

	Lowest exciton at $\mathbf{q} = K$	
Exc. En. (IP)	Exc. En. (GW)	Est. scissor corr.
(eV)	(eV)	(eV)
4.73	6.14	1.41

supercell size. Then we solved the BSE on the supercell at transferred momentum $\mathbf{q}=0$, applying the scissor operator estimated in Table I. In this way we found the excitons already present in the unit cell and new lower four times degenerate dark excitons, corresponding to the indirect excitons mapped at Γ . Being dark, the indirect excitons do not contribute to the optical response if the coupling with phonons was not explicitly taken into account.

- (3) We map the phonon-modes at $\mathbf{q} = K$ in the supercell and generate 12 supercells, one for each phonon mode, with atoms displaced along the phonon eigenmodes times 0.1 Bohr.
- (4) For each displaced supercell we diagonalized the BSE where single particle energies have been corrected using the previously estimated scissor operator to reproduce the exact exciton energy at momentum K in the primitive cell. Notice that we repeated the calculation by orienting the electric field, which enter into the exciton dipole matrix elements T_S , in each of the three directions of space ($\hat{\mathbf{r}} = \{\hat{x}, \hat{y}, \hat{z}\}$) in order to subsequently calculate a spatial average of the luminescence signal. Moreover, in the displaced supercells we do not recalculate the dielectric constant but we use the equilibrium one. This approximation allows us to speed up calculation and it has been shown not to produce any visible error in the final luminescence spectrum [39].
- (5) Using the BSE results of the displaced supercells we calculated the second order derivatives of the excitonic dipole matrix elements $\frac{\partial^2 |T_S|^2}{\partial R_{\lambda q}^2}$ by finite differences, where *S* is the exciton index and λ , **q** the phonon mode, as pointed out in Eq. (4) [40].
- (6) Finally, we applied Eq. (4) to get the luminescence spectra at 0 kBar, and spatially averaged on the three Cartesian directions. In Eq. (4) we used an excitonic temperature of 75 K and a lattice one of 55 K in analogy with the measurements

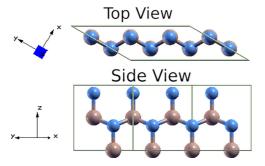


FIG. 2. Top view and side view of the supercell mapping the K point at Γ .

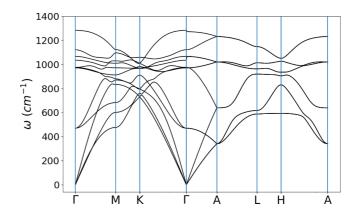


FIG. 3. Phonon dispersion of wBN at 0 kBar calculated in the framework of DFPT.

on hBN [12]. Notice that in our case the sum over \mathbf{q} points reduces to the single point $\mathbf{q} = K$.

This procedure is explained with a working example on the wiki webpage of the Yambo code [41].

IV. RESULTS

In this section we present results for the different properties of wBN as a function of the pressure between 0 and 200 kBar.

A. Vibrational Properties

The phonon dispersion at 0 kBar is represented in Fig. 3. The four atoms in the unit cell give rise to twelve phonon branches for a given \mathbf{q} point. According to the analysis of the wBN symmetry group C_{6v} , phonon modes can be decomposed into $N(A_1 \oplus B_1 \oplus E_1 \oplus E_2)$ modes, where N=2, whose modes $A_1 \oplus E_1$ are acoustic and $A_1 \oplus 2B_1 \oplus E_1 \oplus 2E_2$ are optic modes (nine in total). Except for B_1 modes which are infrared and Raman forbidden, the others are Raman active modes. Out of those, A_1 and E_1 modes are also infrared active. The A_1 and E_1 modes are each split into LO and TO components because of the macroscopic electric field associated with

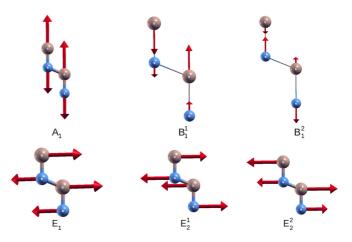
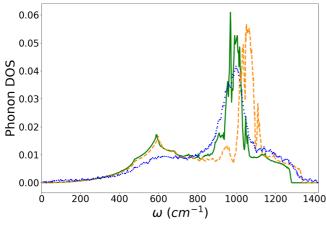


FIG. 4. Phonon eigenmodes at Γ point and their irreducible representations. The A_1 eigenmode is intended to be the same for TO and LO modes, while the represented E_1 mode corresponds to the TO one, the LO being out of plane with respect to the TO mode.



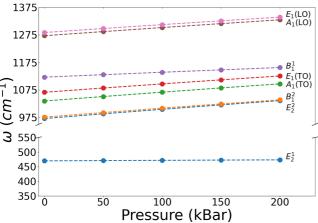


FIG. 5. Top panel: Calculated phonon density of states for 0 (green, solid line) and 200 kBar (orange, dashed line) and experimental phonon density of states (blue, dotted line) extracted from inelastic x-ray scattering in Refs. [43], multiplied by two to make the comparison easier. Bottom panel: Optical phonon frequencies at Γ point in wBN as a function of hydrostatic pressure. The first three acoustic modes are omitted.

the motion of the longitudinal-optical phonon vibrations and the consequent dependence on the \mathbf{q} point of the Born effective charges. Phonon modes are identified in Fig. 4 according to the notation used in Ref. [42].

We are going to examine now the evolution of phonon modes as a function of pressure. As hydrostatic pressure increases from 0 to 200 kBar, with a consequent compression of the entire crystal structure, phonon frequencies in general increase as well as expected. The general tendency is captured by the phonon density of state represented in Fig. 5 (top panel), obtained by sampling the Brillouin zone on a uniform $30 \times 30 \times 20$ **q** point grid. We observe that the phonon density of states, calculated at 0 kBar, agrees with the experimental DOS and that low energy optical phonons (in the frequency range 400 - 800 cm⁻¹) are slightly affected when the pressure is increased from 0 to 200 kBar, while the high energy optical phonons are significantly shifted.

This point is highlighted in Fig. 5 (bottom panel) where we report the phonon energies at the Γ point as a function of pressure (with the exception of the first three acoustic modes). Note that, in contrast to all the other modes which have a similar slope, the double degenerate E_2^1 mode is less sensitive

TABLE II. Longitudinal and transverse zone-center phonon frequencies of wBN at 0 kBar in cm⁻¹, compared with Refs. [19,44]. The A_1 mode LO-TO splitting is obtained by sampling the Brillouin zone along the Γ -A and Γ -K directions, respectively.

Irr. repr E_{Γ}	This work	Ref. [19]	Ref. [44]
$\overline{E_2^1}$	469.96	475.1	
$E_2^1 \\ E_2^2 \\ B_1^1$	971.08	978.3	
$B_1^{\tilde{1}}$	976.17	981.3	
$A_1(TO)$	1034.7	1043.4	1006
$E_1(TO)$	1066.4	1075.9	1053
B_1^2	1121.3	1131.3	
$A_1(LO)$	1271.8	1278.9	1258
$E_1(LO)$	1283.6	1291.5	1281

to changes in pressure. This is consistent with the fact that this mode is associated with the phase transition to the zincblende phase (cBN). The E_2^1 mode is in fact the result of the folding of the cubic phase transverse acoustic phonon branch onto Γ in the wurtzite phase. Interestingly, the frequency associated with this mode is compatible with the phonon replica (~60 meV) observed in the cathodoluminescence spectra of cBN, supporting then the hypothesis that wBN domains are present in cBN after phase transformation under high pressure conditions [21]. In Table II we report the calculated phonon frequencies at Γ point at 0 kBar, which are in agreement with previous results [2,44]. Here we observed that the E_1 LO-TO splitting amounts to 217.2 cm⁻¹ at 0 kBar while it decreases up to 213.7 cm⁻¹ at 200 kBar (see Table S2 in the SM [35]), accordingly to the decrease of the Born effective charges and the dielectric tensor matrix elements (see Table S3 in the SM [35]). This finding goes in the same direction as what happens for wurtzite phase GaN and AlN [42,45,46] and other III-V compounds [47]. As pressure increases such splitting is always more pronounced than the E_1 - A_1 splitting (which amounts to 11.7 cm⁻¹ at 0 kBar and 9.7 cm⁻¹ at 200 kBar, see Table S2 in the SM [35]). Therefore wBN belong to crystals of class I, according to Refs. [48,49], in which the electrostatic Coulomb forces dominate over the anisotropy of the short-range interatomic forces.

B. Electronic band structure and quasiparticle properties

Here we focus on the evolution of electronic properties as a function of pressure. At 0 kBar we found that at the DFT level, the lowest energy gap occurs for a transition between Γ and K points of the Brillouin zone, and amounts to 5.01 eV. The direct gap of 8.3 eV, in agreement with previous results [50], occurs at a k point close to Γ that we will call hereafter $X_{\rm DFT} = (0, 1/12, 0)$. The quasiparticle (QP) correction in GW approximation widens the indirect and direct gap of 1.4 and 1.72 eV, respectively, at 0 kBar. The k point where the direct transition occurs (direct gap transition) shifts a little bit further away from the Γ point to the new **k** point of coordinates $X_{GW} = (0, 1/6, 0)$. The DFT and QP band structures together with direct and indirect band gap transitions are highlighted in Fig. 6. While in Fig. 7, the QP corrected band structure at the extreme values of the considered pressure range are shown. While the first valence bands are essentially unaffected, along

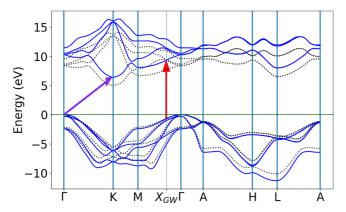


FIG. 6. DFT (black points) and QP-corrected electronic (solid blue line) band structure of wBN at 0 kBar. Arrows highlight the transitions where direct (red) and indirect (violet) gaps occur at QP level. Only four bands above and below the band gap are shown.

the Γ -M line (where the direct transition occurs), the lowest conduction band shifts up, while at K it is slightly shifted down. Such an observation opens the discussion on the evolution of the band gaps with the applied hydrostatic pressure.

As pressure increases, the **k**-space positions (X_{DFT} and X_{GW}) of the direct gap do not change, and in the 0–200 kBar range the QP correction itself is almost constant. On the other hand, direct and indirect band gaps undergo a different behavior with pressure (see Fig. 8). We observe that while the first tends to increase the latter does the opposite. This is in contrast of what is observed for direct and indirect band gaps in hBN under pressure [2]. We ascribed this behavior to the different character of the band edges, that we deduced by projecting the Kohn-Sham states on selected atomic orbitals of B, N atoms (see Fig. S3 in the SM [35]). The states taken into consideration are 2s and 2p for each of the four atoms in the primitive cell. The electrons contributing to conduction bands are mostly from the B atoms while the electrons contributing to the valence ones come from the N atoms; this is in particular true for **k** points very close Γ point. However at the K point the contribution comes from a mixture of the orbitals of two atoms, making the behavior of the indirect gap different from the direct one.

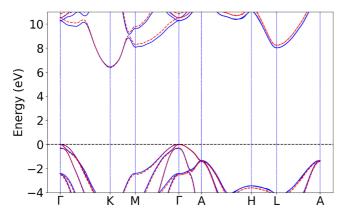


FIG. 7. QP corrected band structure at 0 kBar (blue solid line) and 200 kBar (red dashed line).

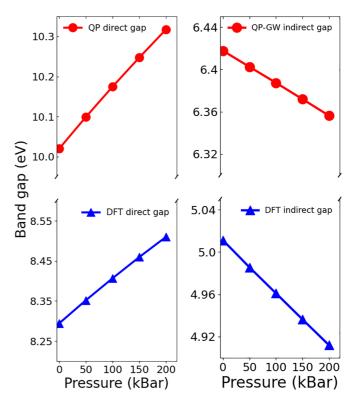


FIG. 8. DFT and QP-corrected direct and indirect gaps as a function of pressure.

C. Light absorption

In Fig. 9 we plot a comparison between the imaginary part of the dielectric function calculated at the independent particle (IP) level, both from DFT and *GW* band structure, and by including the electron-hole interaction. Such interaction modifies the absorption signal, resulting in both a redistribution of the main features intensity and the appearance of an excitonic peak below the *GW*-corrected band gap at 9.93 eV, absent at the IP level.

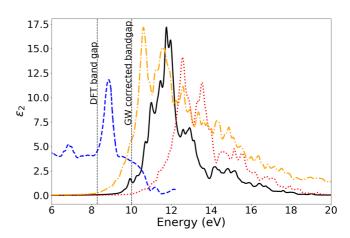


FIG. 9. Quasiparticle (solid black), DFT based independent particle (dashed dotted orange), GW based independent particle (dotted red) spectra of the imaginary part of ϵ_M at 0 kBar. Dashed vertical lines mark the position of the GW-corrected and the DFT band gaps in both main figure and inset. Experimental data (dashed blue) are extracted from Ref. [17].

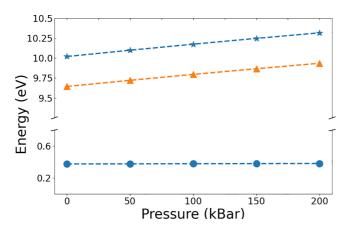


FIG. 10. Quasiparticle-corrected direct band gap (blue stars line), first bright exciton level (orange triangles line), and the binding energy (blue dots line) as a function of pressure.

Our analysis of the excitonic structure of the absorption spectrum revealed the presence of a low energy dark exciton and two degenerate excitons at 9.59 eV with a small dipole. The first bright exciton, which determines the optical gap at 0 kBar, is also doubly degenerate and has an energy of 9.64 eV, a value quite close to that one calculated for cBN [21]. In Fig. 9 we also plot the experimental dielectric constant measured in Ref. [17]. The authors affirm that the optical gap was found to be at 8.7 ± 0.5 eV. Within the error bar such a value could correspond to our first exciton, which has a nonnegligible dipole at 9.64 eV. Overall, the comparison is rather delicate, the experimental spectrum does not extend to higher energies. The discrepancy in energy between our theoretical predictions and experimental results can only partially be attributed to the absence of electron-phonon interaction in our theoretical approach [51]. However, it is also possible that the distinct features observed in the experimental spectrum are caused by the presence of defects, which cannot be ruled out at this time. In this case, further studies are necessary to determine the source of these signals, and luminescence measurements could easily discriminate between defect and bulk states due to the indirect nature of this material, see next section.

We then proceed to analyze the effect of pressure on the lowest excitonic peaks. As shown in Fig. 10, optical excitation shift linearly toward high energies from 0 kBar to 200 kBar, closely following the linear behavior of the quasiparticle corrected direct band gap. Consequently, the exciton binding energy ($E_{\rm bind} = E_{\rm gap} - E_{\rm exc}$) remains almost constant with pressure, with a value of 0.38 eV at 0 kBar. This behavior can be explained by considering how screening affects the electron-hole pair. In fact, excitons in wBN are highly delocalized well beyond the boundaries of the unit cell. Thus, changing the geometry of the system has less effect on the screening which affects their electron-hole interaction.

D. Phonon assisted luminescence

In Fig. 11 we report phonon-assisted luminescence calculated using the theoretical methods presented in Secs. II and III.

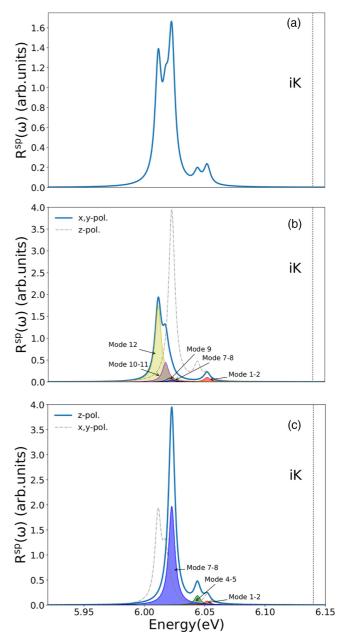


FIG. 11. (a) Phonon-assisted luminescence at 0 kBar, when the luminescence has been averaged across the three Cartesian directions. The dotted vertical line indicates the energy of the indirect exciton (iK). In panels (b) and (c) we separated the contributions of the basal plane Cartesian coordinates (average over the x and y directions) and the z axis. In both cases the contributions of single phonon modes have been specified. Note that a couple of phonon modes (the third and the sixth) have a much lower contribution compared to the others and are not shown in the figure.

The luminescence is calculated at 0 kBar, and spatially averaged on the three Cartesian directions, panel (a) in Fig. 11. One observes four main peaks, at 6.012, 6.022 (plus a shoulder at 6.018 eV), 6.044, and 6.053 eV. We first decomposed the spectrum into the contribution coming from the basal plane components and that from the z axis as shown in panels (b) and (c) of Fig. 11. This allows us to interpret the origin of the four main peaks. Additionally, the contribution of each phonon mode at K point of the Brillouin Zone is highlighted,

allowing to identify each peak as the result of the coupling of the lowest indirect exciton iK and a precise phonon mode at the point K of the BZ. We found that for light polarized in the basal plane the indirect exciton couples mostly with the highest-energy optical modes and to the lowest energy phonon modes to give origin to the peak at 6.012 eV, the shoulder at 6.018 eV, and the peak at 6.053 eV. On the other hand when light is polarized along the z axis, the coupling occurs with the 7th-8th, the 4th-5th, and slightly with the 1st-2nd phonon modes, and originate to the intense signal at 6.022 eV, to that at 6.044 and 6.053 eV, respectively.

On the basis of the evolution as a function of pressure of the QP direct band gap and of the direct exciton energy (see Figs. 6–10), we deduce that the indirect exciton energy will decrease as pressure increases. Beside that, the phonon frequencies at K point (see Fig. S2 and Table S4 in the SM [35]) tend to increase with pressure even if with a relatively different rates when optic and acoustic modes are compared. Putting all this information together we expect a slight red shift with pressure of the luminescence spectrum accompanied by a small deviation of the two main peaks, since the originally located peak at 6.02 eV will red shift more than the other at 6.05 eV, because of the contribution coming from different phonon modes. As mentioned above, the remarkable light-emitting properties of hBN have been the subject of extensive research in recent years. Recent studies have shown that its high light-emitting efficiency is due to its flat band nature at the band edges [52]. Although experimental luminescence data are not currently available, it is possible to speculate that the light-emitting properties of wBN would be more similar to those of diamond, given its more parabolic band nature [10]. As a result, the inclusion of wBN in hBN may reduce its luminescence efficiency. We expect that this prediction could be verified in future work.

V. CONCLUSIONS

This study presents an in-depth analysis of the properties of wBN over a wide range of pressures. Our findings reveal that

wBN is a large indirect gap material, with a predicted direct gap near the Γ point, which increases with pressure, while the indirect gap between $\Gamma \to K$ decreases. We also explored the vibrational properties of wBN and have discovered that the frequency associated with the E_2^1 mode aligns with the phonon replica observed in the cathodoluminescence spectra of cBN. In terms of optical properties, our results show that the first optically active excitations occur at 9.64 eV, below the GW gap. It is important to consider the electron-hole interaction to interpret the experimental measurements, which display peaks below the GW gap. At the IP level the onset of the light absorption is high in energy, and not compatible with the experimental observations. We also studied light emission in wBN, and found that phonon-mediated emission is expected to occur at around 6 eV, but it will be less intense than in hBN due to the nonflat bands of wBN. Overall, this study, along with new techniques for the synthesis of wBN, can aid in the identification of wBN through luminescence measurements and promote its use in optoelectronics in the

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