

Froude scaling for rovers on small body surfaces

Cecily Sunday, Naomi Murdoch, Simon Tardivel, Patrick Michel

▶ To cite this version:

Cecily Sunday, Naomi Murdoch, Simon Tardivel, Patrick Michel. Froude scaling for rovers on small body surfaces . Europlanet Science Congress 2022, Sep 2022, Grenade, Spain. $10.5194/{\rm epsc2022-525}$. hal-03835761

HAL Id: hal-03835761 https://hal.science/hal-03835761

Submitted on 1 Nov 2022

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers. L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.



EPSC Abstracts Vol. 16, EPSC2022-525, 2022 Europlanet Science Congress 2022 © Author(s) 2022. This work is distributed under the Creative Commons Attribution 4.0 License.



Froude scaling for rovers on small body surfaces

Cecily Sunday^{1,2}, Naomi Murdoch¹, Simon Tardivel³, and Patrick Michel² ¹Institut Supérieur de l'Aéronautique et de l'Espace (ISAE-SUPAERO), Toulouse, France (cecily.sunday@isae-supaero.fr) ²Université Côte d'Azur, Observatoire de la Côte d'Azur, Centre National de la Recherche Scientifique (CNRS), Laboratoire Lagrange, Nice, France

³Centre National d'Études Spatiales (CNES), Toulouse, France

Abstract

We study rover-regolith interactions in low-gravity environments by conducting soft-sphere discrete element method (SSDEM) simulations with a simplified rover wheel and a bed of spherical particles. The simulations reveal that rover performance scales according to the Froude number, or a dimensionless parameter which accounts for the size of the wheel, the rotational velocity of the wheel, and gravity. This relationship provides valuable insight into how to operate rovers and analyze wheel-regolith interactions during future rover missions.

Introduction

Wheeled-rovers are useful tools for identifying the surface material properties of planetary surfaces [1-3]. The sinkage and traction of a rover can be used to assess the bearing strength of a material, and the tracks left behind by a rover's wheels can provide information regarding the density, the friction angle, and the cohesion of the regolith [3]. To maximize the scientific return from a rover mission, however, the vehicle must first be able to drive on a granular terrain in a reduced-gravity environment. Parabolic flight experiments and numerical simulations have shown that, as gravity decreases, wheel slip increases and rover traction decreases [4-6]. At the same time, wheel sinkage remains comparable for different gravity levels, at least for certain types of materials [4]. While a few models have been proposed to help explain these findings [7,8], an explicit scaling relationship between rover performance and gravity has yet to be established.

In this work, we analyze rover-regolith interactions on Earth versus a small moon, Phobos. Based on the findings of [8] and [9], we hypothesize that rolling behavior scales with gravity according to the dimensionless Froude number, $Fr = L\omega^2/g$, where L is a characteristic length (e.g., the radius of a rover wheel), ω is the rotational velocity of the wheel, and g is gravity.

Simulations

We conducted SSDEM simulations using a simplified rover wheel and the MULTICORE module of the open-source code CHRONO [10, 11]. The rover wheel is 214 mm in diameter and 53 mm in width and has 9 thin grousers that are equally spaced around a cylindrical hub. The wheel rotates at a constant speed through a bed of spherical grains and can translate freely in the horizontal and vertical directions. The grains are 6 +/- 0.5 mm in diameter and are contained within a 800x250x150 mm box. Simulations were performed using different material cases (e.g., glass beads, rough glass beads, and rough glass beads with cohesion), two gravity levels (g = 9.81 and

0.006 m/s²), and several wheel speeds ($\omega = 0.065$ -2.65 rad/s). Fig. 1 provides a snapshot of a typical simulation. We evaluate driving performance by comparing the travel distance of the wheel Δx as a function of its angular displacement $\Delta \theta$. We also compare the static and dynamic sinkage of the wheel for the different simulation cases.



Fig. 1: Snapshots from a simulation where gravity is 9.81 m/s^2 and the rotational velocity is 0.65 rad/s. The container is filled with 6 +/- 0.5 mm spherical particles which are colored by their vertical positions at the start of the simulation [12].

Results

We find that static wheel sinkage is independent of gravity for simulations with cohesionless materials. Like previous works, we also find that driving performance decreases as gravity decreases for configurations with similar driving speeds (Fig. 2a). We also observe, however, that rover performance is comparable for configurations with similar Froude numbers (Fig. 2b). This means that driving on a cohesionless granular material on Earth at 0.2 m/s is like driving on the same material at 0.005 m/s on Phobos.



Fig. 2: Travel distance Δx by angular displacement $\Delta \theta$ for a wheel rotating through rough glass beads on Earth and Phobos. Plot (a) shows results for similar wheel speeds and plot (b) shows results for similar Froude numbers [12].

To understand these findings, we consider wheel-regolith interactions in terms of granular regimes. When a wheel initially sinks into a granular material, the system is in a quasi-static state. Sinkage is independent of gravity, because the reduced strength of the material is offset by the reduced weight of the wheel [4,7]. Then, as the wheel turns, the grains become fluidized, and above a certain velocity, the system enters the inertial regime. For similar wheel velocities, the system might be in a quasi-static state on Earth but in an inertial regime on Phobos. This explains why traction decreases as gravity decreases. For similar Froude numbers, however, the system will be in the same regime regardless of gravity level. Fig. 3 illustrates this point further by comparing the velocity fields of the particles in the different simulations. Overall, we demonstrate how the Froude relationship can be a powerful tool for planning rover operations and for interpreting surface interactions during future rover missions.



Fig.3: Velocity field of the particles beneath a rover wheel for simulations in rough glass beads on Earth and Phobos. Plots (a) and (b) show results for similar wheel speeds and plots (c) and (d) show results for similar Froude numbers.

Acknowledgements

This work was funded by ISAE-SUPAERO and CNES through a PhD research grant. The simulations were conducted using HPC resources from CALMIP under grant 2019-P19030. NM acknowledges funding support from CNES.

References

[1] Sullivan, R., et al. Journal of Geophysical Research: Planets, 116(E2), 2011. [2] Gao, Y., et al. Advances in Space Research, 58(9):1893–1899, 2016. [3] Ding, L., et al. Preprint, 2021. [4] Kobayashi, T., et al. Journal of Terramechanics, 47(4):261–274, 2010. [5] Kovács, L., et al. Journal of Field Robotics, 37(5):754–767, 2020. [6] Nakashima, H. and Kobayashi, T. Journal of Terramechanics, 53:37–45, 2014. [7] Wong, J. Y. and Kobayashi, T. Journal of Terramechanics, 49(6):349–362, 2012. [8] Slonaker, J., et al. Physical Review E, 95(5):052901, 2017. [9] Sunday, Cecily, et al. Astronomy & Astrophysics, 658 A118, 2022. [10] Tasora, A., et al. High Performance Computing in Science and Engineering. Springer International Publishing, 2016. [11] Sunday, C., et al. Monthly Notices of the Royal Astronomical Society, 498(1):1062–1079, 2020. [12] Sunday, C. PhD Thesis, ISAE-SUPAERO, 2022.