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4-component relativistic Hamiltonian with effective QED potentials for molecular calculations

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We report the implementation of effective QED potentials for all-electron 4-component relativistic molecular calculations using the DIRAC code. The potentials are also available for 2-component calculations, being properly picture-change transformed. The latter point is important; we demonstrate through atomic calculations that picture-change errors are sizable. Specifically, we have implemented the Uehling potential [E. A. Uehling, *Phys. Rev.* 48 , 55 (1935)] for vacuum polarization and two effective potentials [P. Pyykkö and L.-B. Zhao, *J. Phys. B* 36 , 1469 (2003); V. V. Flambaum and J. S. M. Ginges, *Phys. Rev. A* 72 , 052115 (2005)] for electron self-energy. We provide extensive theoretical background for these potentials, hopefully reaching an audience beyond QED-specialists. We report the following sample applications: i) we confirm the conjecture of P. Pyykkö that QED effects are observable for the AuCN molecule by directly calculating ground-state rotational constants B_0 of the three isotopomers studied by MW spectroscopy; QED brings the corresponding substitution Au-C bond length r_s from 0.23 to 0.04 pm agreement with experiment, ii) spectroscopic constants of van der Waals dimers M_2 (M=Hg, Rn, Cn, Og): QED induces bond length expansions on the order of 0.15(0.30) pm for row 6(7) dimers, iii) we confirm that there is a significant change of valence s population of Pb in the reaction $PbH_4 \longrightarrow PbH_2 + H_2$, which is thereby a good candidate for observing QED effects in chemical reactions, as proposed in [K. G. Dyall *et al.*, *Chem. Phys. Lett.* 348 , 497 (2001)]. We also find that whereas in PbH_4 the valence $6s_{1/2}$ population resides in bonding orbitals, it is mainly found in non-bonding orbitals in PbH_2 . QED contributes 0.32 kcal/mol to the reaction energy, thereby reducing its magnitude by -1.27 %. For corresponding hydrides of superheavy flerovium, the electronic structures are quite similar. Interestingly, the QED contribution to the reaction energy is of quite similar magnitude (0.35 kcal/mol), whereas the relative change is significantly smaller (-0.50 %). This curious observation can be explained by the faster increase of negative vacuum polarization over positive electron self-energy contributions as a function of nuclear charge.

I. INTRODUCTION

Relativistic quantum chemistry is the proper framework for the theoretical study of heavy elements.^{1–5} For example, the yellow color of gold,^{6,7} as well as the cell potential of the lead-acid battery⁸ cannot be explained without relativistic effects. Even for light elements, the fine structure of spectra is essentially due to spin-orbit (SO) interaction (e.g. Refs. 9–11).

Improvements in both computational power and methodology nowadays allow highly accurate electronic structure calculations including both relativistic and electron correlation effects. A next challenge for increased accuracy is the inclusion of the effects of quantum electrodynamics (QED), which in principle means going beyond the no-pair approximation.^{5,12,13} We focus on QED effects generating the Lamb shift, roughly described as follows:

- vacuum polarization (VP): a charge in space is surrounded by virtual electron-positron pairs and this

contributes to its observed charge

- the electron self-energy (SE): the electron drags along its electromagnetic field and this contributes to its observed mass.

For hydrogen the splitting between the $^2S_{1/2}$ and $^2P_{1/2}$ states is a mere 4 meV,¹⁴ but for U^{91+} it has grown to a whopping 468 eV.¹⁵ QED effects would possibly constitute the final correction to chemistry concerning the fundamental inter-particle interactions because the next contribution, parity non-conservation (PNC) associated with the weak force, is typically ten orders of magnitude smaller.¹⁶ The magnitude of QED effects has been estimated based on the ionization potential of alkali atoms, and the *rule of thumb* is that QED effects reduce relativistic effects by about one percent.¹⁷

Calculations within the rigorous QED framework have been reported for few-electron systems, and they are in excellent agreement with experiment. Examples are the Lamb shift of Li-like uranium,^{18,19} the hyperfine coupling constant (HFCC) of few-electron atoms,^{20,21} and the anomalous g factor,²² that provides stringent tests of the accuracy of QED.

The rigorous QED approach for few-electron systems cannot be extended to many-electron systems because of

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the high computational cost. A more practical, but approximate approach is the introduction of effective QED potentials (effQED).^{23–27} In the atomic case, some codes for the calculation with effective potentials have been reported (e.g., GRASP,²⁸ QEDMOD,^{29,30} and AMBIT³¹). A nice illustration is the recent work by Pařteka and co-workers,³² which was finally able to bring the calculated ionization potential (IP) and the electron affinity (EA) of the gold atom into meV agreement with experiment, high-order electron correlation being the missing crucial ingredient.

For the case of molecules in chemistry, pioneering works have been done by Kirk Peterson’s group. They added the following parameterized model potentials to the all-electron scalar Douglas–Kroll–Hess (DKH) Hamiltonian:^{33–37} i) an effective SE potential in the form of a single Gaussian function, proposed by Pyykkö and Zhao (PZ),²⁴ and ii) five Gaussian functions fitted by Peterson’s group^{33–35} to a parameterized expression for the Uehling VP potential²³ given in Ref. 17 and corrected in Ref. 24. They then found that the QED effects on the dissociation energy are about 0.6 kcal/mol and 0.4 kcal/mol in closed- and open-shell Hg systems, respectively.³³ A bond length expansion of 0.001 Å was observed for the HgBr molecule.³⁴ Michael Dolg and co-workers have reported pseudopotentials (PPs) fitted to include QED effects.^{38–40} For Cn₂ a bond length expansion due to QED of about 0.003 Å was reported,⁴⁰ in line with the effect observed by Peterson’s group.³⁴ On the other hand, in Ref. 41, the QED effect was found to *shorten* the bond length of TsH⁺, LvH and OgH⁺. The reason for this opposite trend may be that the valence orbitals have p- rather than s-orbital contributions from the heavy atom.

PPs are widely used for the inclusion of relativistic effects, and generally give accurate results for valence properties compared with all-electron calculations.^{42,43} However, the PP approach cannot be applied to molecular core-properties such as NMR and Mössbauer parameters, which bars the possibility to investigate the effect of QED in the nuclear region where such effects are generated.⁴⁴ The effQED approach promoted by the Peterson group can in principle be applied to core-properties, but it should be noted that effQED potentials were added to approximate one-component relativistic Hamiltonians without picture-change.^{45–47} To include QED effects in a more rigorous manner, it seems more appropriate to include effective QED potentials in 4-component relativistic all-electron calculations.

In this work, we report the implementation of effective QED potentials in the DIRAC code for relativistic molecular calculations.⁴⁸ Three potentials have been implemented: the Uehling potential²³ for vacuum polarization, Pyykkö and Zhao’s model SE potential,²⁴ as well as the effective SE potential of Flambaum and Ginges.²⁵ Our implementation is based on numerical routines from the GRASP atomic code²⁸ that have been grafted onto the DFT grid of DIRAC.⁴⁹

As first molecular applications of our implementation we have chosen three case studies:

- the AuCN molecule for which Pekka Pyykkö has suggested QED effects on the bond length.⁵⁰
- the van der Waals dimers M₂ (M = Hg, Rn, Cn, Og) for which one might suspect QED effects to be on par with interaction energies. Interestingly, van der Waals forces have been described in terms of vacuum fluctuations.^{51,52}
- the reaction energy of Pb hydrides, PbH₄ → PbH₂ + H₂, suggested by Dyall *et al.* as a possible candidate for a significant QED effect in chemistry.⁵³ In addition to the Pb system, we have also calculated the heavier analogue, Fl hydrides.

Very recently, Leonid Skripnikov reported the implementation of effective QED potentials for 4-component all-electron molecular calculations, so far with a focus on transition energies.⁵⁴ The initial report has been followed by applications to Ba⁺, BaF, RaF and E120F⁵⁵ as well as the five low-lying excited states of RaF.⁵⁶ The implementation is to some extent complementary to ours in that it uses the effective SE potential proposed by Shabaev and co-workers.^{26,27,29} Interestingly, the implementation is based on the DIRAC code as well.

The paper is organized as follows: in Sec. II we review the effective QED potentials that we have implemented, and in Sec. III we discuss the numerical integration of these potentials. This is followed by Sec. IV which gives the computational details of our calculations. Our results are presented in Sec. V, followed by conclusions in Sec. VI. We also provide an appendix with more extensive theory and reading suggestions. SI-units are used throughout this paper.

II. THEORY

The starting point for our work is an electronic Hamiltonian on the generic form

$$H = V_{\text{NN}} + \sum_i H_D(\mathbf{x}_i) + \frac{1}{2} \sum_{i \neq j} g(\mathbf{x}_i, \mathbf{x}_j) \quad (1)$$

where V_{NN} is the classical repulsion of fixed nuclei. The one-electron part is the Dirac Hamiltonian

$$H_D(\mathbf{x}_i) = (\beta - \mathbf{1}_4)m_e c^2 - i\hbar c \boldsymbol{\alpha} \cdot \nabla_i - e\varphi_N(\mathbf{x}_i), \quad (2)$$

in the electric potential φ_N of the fixed nuclei and shifted by $-m_e c^2$ to align energies with the non-relativistic scale. In the present work, the two-electron interaction g will be the instantaneous Coulomb term supplemented with the Gaunt term.⁵⁷ Further discussion of the resulting Dirac–Coulomb–Gaunt (DCG) Hamiltonian is for instance found in Ref. 3.

Our goal is to introduce QED effects, notably electron self-energy (SE) and vacuum polarization (VP), by extending the one-electron Hamiltonian by the corresponding effective QED potentials

$$H_D \rightarrow H_D - e\varphi_{\text{effQED}}; \quad (3)$$

$$\varphi_{\text{effQED}} = \sum_A (\varphi_A^{\text{SE}} + \varphi_A^{\text{VP}}).$$

Note that the effective QED potentials are formulated as a sum over atomic contributions due to their expected short-range nature (on the order of a reduced Compton wavelength $\lambda = \hbar/m_e c$).⁴⁴

In the following we shall present the effective QED potentials selected for our implementation with some remarks on their construction which may provide indications on their expected performance. We shall proceed within the \hat{S} -matrix (scattering matrix) formalism of QED. Since we hope to address a wider audience than QED specialists, we provide a more extensive theoretical background in Appendix A.

QED is the relativistic quantum field theory that describes the interaction of electromagnetic radiation with relativistic matter (Dirac electrons). The interaction between electrons and photons is given by an interaction Hamiltonian density

$$\hat{\mathcal{H}}_I(x) = -ec\hat{\Psi}(x)\gamma^\mu\hat{\Psi}(x)\hat{A}_\mu(x). \quad (4)$$

Here, $\hat{\Psi}(x)$ and $\hat{\bar{\Psi}}(x)$ are the quantized Dirac field operator and its corresponding adjoint, whereas $\hat{A}_\mu(x)$ is the quantized photon field operator. The job of these operators is to create and annihilate, at the spacetime point $x = (ct, \mathbf{x})$, electrons and photons, respectively. This last expression accounts (explicitly) for the coupling between electron and photon fields, and is obtained through minimal substitution of the four-gradient of the Dirac Lagrangian density, in accordance with the *principle of minimal electromagnetic interaction* (term coined by Gell-Mann⁵⁸). For detailed derivations and discussions, the reader may consult Schweber in Ref. 59 (chapter 10), Peskin and Schroeder in Ref. 60 (chapter 4), as well as Greiner and Reinhardt in Ref. 61 (section 8.6). The scattering matrix is the special case of the time-evolution operator $\hat{U}(t, t_0)$, where the initial t_0 and final times t are at $\mp\infty$, to ensure Lorentz invariance. Upon expansion of the \hat{S} -matrix operator in the fundamental charge e , the n th-order term $\hat{S}^{(n)}$ contains a time-ordered string of n interaction Hamiltonian densities $\hat{\mathcal{H}}_I$, as seen in Eq. (A29). Using Wick's theorem,⁶² a time-ordered string is converted into a linear combination of normal-ordered ones with all possible contractions, which in turn can be translated into the iconic Feynman diagrams.⁶³ We limit attention to systems of n electrons and zero photons (photon vacuum). The latter implies that any string of normal-ordered photon operators $\hat{A}_\mu(x)$ that is not fully contracted will vanish upon taking expectation values, such that the \hat{S} -matrix expansion is ef-

fectively limited to even-ordered contributions, associated with the fine-structure constant $\alpha = e^2/4\pi\epsilon_0\hbar c$ as expansion parameter. To lowest order in α appears five Feynman diagrams, shown in Fig. 1: two of them give state-independent energy-shifts and are usually ignored within a perturbative setting, whereas the remaining three represent electron self-energy, vacuum polarization and single-photon exchange. The latter diagram describes the relativistic electron-electron interaction, mediated by photons, to lowest order and is in line with the statement of Dirac:

Classical electrodynamics, in its accurate (restricted) relativistic form, teaches us that the idea of an interaction energy between particles is only an approximation and should be replaced by the idea of each particle emitting waves, which travel outward with a finite velocity and influence the other particles in passing over them.⁶⁴

In the diagrams of Fig. 1 double electron lines appear to indicate that we are working within the Bound-State QED (BSQED) framework in which the Dirac field operators are expanded in solutions of the Dirac equation in some external (contravariant) four-potential: $A^e = (\varphi^e/c, \mathbf{A}^e)$ (Furry picture⁶⁵), rather than free-particle ones. In the atomic case, this provides us with a second perturbation expansion parameter $Z\alpha$, as will be seen in the next section.

A. Effective QED potentials for vacuum polarization

The four-potential associated with the vacuum polarization effect can be written as

$$\varphi_{\text{VP}}^\mu(\mathbf{x}_1) = \frac{e}{2\pi i} \int_{C_F} dz \int d^3x_2 \frac{\text{Tr}[\gamma^\mu G_{A^e}(\mathbf{x}_2, \mathbf{x}_2; z) \gamma^0]}{4\pi\epsilon_0 |\mathbf{x}_1 - \mathbf{x}_2|}, \quad (5)$$

where the complex z -integral is to be evaluated along the Feynman contour C_F that goes above and below positive- and negative-energy poles, respectively, of the bound electron Green's function G_{A^e} . This function is related to the bound electron propagator $S_{A^e}^F$ by Eq. (A42). This VP potential leads to the following vacuum polarization energy-shift

$$\Delta E_{\text{VP}}^{\alpha,2} = -e \sum_i \int d^3x \bar{\psi}_i(\mathbf{x}) \gamma_\mu \psi_i(\mathbf{x}) \varphi_{\text{VP}}^\mu(\mathbf{x}). \quad (6)$$

From consideration of time-reversal symmetry, one can show that in the case of a purely scalar external potential: $A^e = (\varphi^e/c, \mathbf{0})$, vector components of the vacuum polarization four-potential vanish

$$\varphi_{\text{VP}}^\mu(\mathbf{x}) = 0 \quad \text{for } \mu = 1, 2, 3. \quad (7)$$

The bound Green's function G_{A^e} can be written in terms of the free Green's function G_0 and expanded in powers of the time-independent external potential A^e (hence

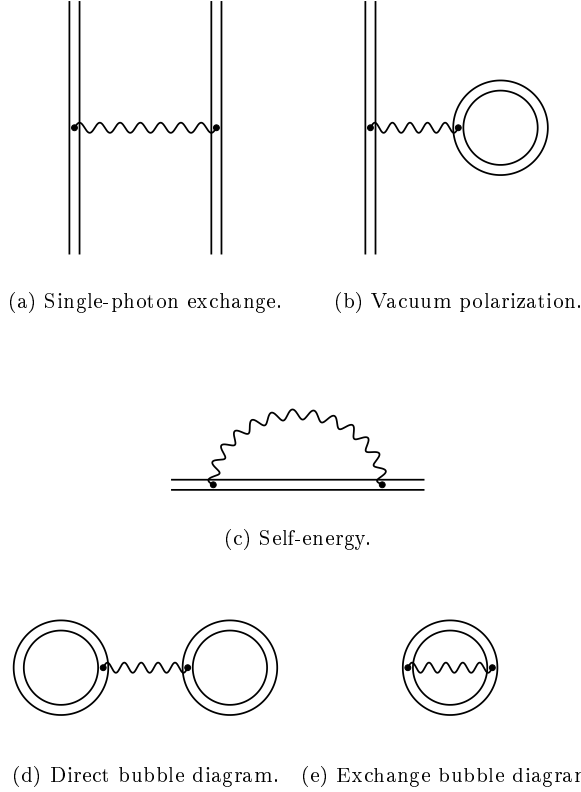


FIG. 1. The lowest-order QED corrections for a many-bound-electron system: of order α .

$Z\alpha$ in the atomic case) as shown in Eq. (A44). As discussed in Sec. A 6 b, the first non-vanishing term of this expansion is the one that is linear in the external potential $A^e(\mathbf{x})$ (the one-potential term). The potential of Eq. (5) is divergent (as seen in Sec. A 6 b) and calls for regularization and renormalization (see Sec. A 6 d). After employing these techniques, one can extract the physical contribution associated with this vacuum polarization effect, and, in the point nucleus problem, represent it by the following scalar potential²³

$$\varphi_{\text{Ueh.}}^{\text{point}}(\mathbf{x}) = \frac{Ze}{4\pi\epsilon_0 r_x} \frac{2\alpha}{3\pi} K_1\left(\frac{2r_x}{\lambda}\right), \quad (8)$$

that corrects the classical Coulomb potential. Here, $r_x \equiv |\mathbf{x}|$ is the radial distance and expressed in terms of the function⁶⁶

$$K_1(x) = \int_1^\infty d\zeta e^{-x\zeta} \left(\frac{1}{\zeta^2} + \frac{1}{2\zeta^4}\right) \sqrt{\zeta^2 - 1}. \quad (9)$$

(see also Refs. 67 and 68). This potential is named after Uehling who first calculated it in 1935 for a point charge nuclear distribution (as indicated by the superscript “point”). The corresponding potential for an arbitrary nuclear distribution $\rho^{\text{nuc.}}$, normalized to one, is

obtained by the following convolution⁶⁶

$$\varphi_{\text{Ueh.}}^{\text{nuc.}}(\mathbf{x}) = \int d^3\mathbf{y} \rho^{\text{nuc.}}(\mathbf{y}) \varphi_{\text{Ueh.}}^{\text{point}}(\mathbf{x} - \mathbf{y}). \quad (10)$$

In the case of a spherically symmetric nuclear charge distribution, one obtains, after angular integration⁶⁶

$$\begin{aligned} \varphi_{\text{Ueh.}}^{\text{nuc.}}(\mathbf{x}) &= \frac{Ze}{4\pi\epsilon_0 r_x} \lambda \frac{2\alpha}{3} \int_0^\infty r_y dr_y \rho^{\text{nuc.}}(r_y) \\ &\times \left[K_0\left(\frac{2}{\lambda}|r_x - r_y|\right) - K_0\left(\frac{2}{\lambda}|r_x + r_y|\right) \right], \end{aligned} \quad (11)$$

where appears the function

$$K_0(x) = \int_1^\infty d\zeta e^{-x\zeta} \left(\frac{1}{\zeta^3} + \frac{1}{2\zeta^5}\right) \sqrt{\zeta^2 - 1}. \quad (12)$$

The integral functions K_0 and K_1 are related through

$$K_1(x) = -\frac{d}{dx} K_0(x). \quad (13)$$

The Uehling potential generally represents the dominant vacuum polarization effect.^{19,69} The Feynman diagram associated with this process is presented in Fig. 3b, and associated with the $\alpha(Z\alpha)$ perturbation order. The higher-order vacuum polarization potentials, associated with the Wichmann–Kroll:⁷⁰ $\alpha(Z\alpha)^3$ and Källén–Sabry:⁷¹ $\alpha^2(Z\alpha)$ processes, are briefly discussed at the end of Sec. A 6 b.

B. Effective QED potentials for self-energy

The energy-shift associated with the self-energy process, in which an electron emits and absorbs a virtual photon, is given by the following expression

$$\begin{aligned} \Delta E_{\text{SE}}^{\alpha,2} &= -e \sum_i \int d^3x_1 \int d^3x_2 \\ &\times \psi_i^\dagger(\mathbf{x}_2) \varphi_{\text{SE}}(\mathbf{x}_2, \mathbf{x}_1; E_i) \psi_i(\mathbf{x}_1) \end{aligned} \quad (14)$$

This expression probably originated from the work of Baranger *et al.*⁷² (section II). Notice at this point that unlike the vacuum polarization effect that is represented by a local scalar potential, the self-energy effect is represented by a non-local matrix potential

$$\begin{aligned} \varphi_{\text{SE}}(\mathbf{x}_2, \mathbf{x}_1; E_i) &= -\frac{e}{2\pi i} \int_{C_F} dz \alpha^\mu G_{A^e}(\mathbf{x}_2, \mathbf{x}_1; z) \alpha_\mu \\ &\times \frac{\exp(g(\mathbf{x}_2, \mathbf{x}_1; z - E_i))}{4\pi\epsilon_0 |\mathbf{x}_1 - \mathbf{x}_2|}, \\ g(\mathbf{x}_2, \mathbf{x}_1; z) &= +\frac{i}{\hbar} |\mathbf{x}_1 - \mathbf{x}_2| \sqrt{z^2/c^2 + i\epsilon}. \end{aligned} \quad (15)$$

Here, ϵ is a small positive number, and the z -integral is again to be evaluated along the Feynman contour C_F . This expression is obtained using the covariant Feynman

gauge photon propagator. The corresponding expression obtained using Coulomb gauge photon propagator is given by Lindgren in Ref. 73 (section 4.6.1.2) (See also Malenfant in Ref. 74). As in the vacuum polarization case, the self-energy potential of Eq. (15) is divergent (as seen in Sec. A 6 c), and calls for a regularization and renormalization treatment in order to extract the physical (finite) correction; see Sec. A 6 d.

In the next two sections, we shall assume that the non-local potential of Eq. (15) can be written in terms of a local effective potential $\varphi_{\text{SE}}(\mathbf{x}_1)$ as

$$\varphi_{\text{SE}}(\mathbf{x}_2, \mathbf{x}_1; E_i) \approx \varphi_{\text{SE}}(\mathbf{x}_1)\delta(\mathbf{x}_2 - \mathbf{x}_1), \quad (16)$$

and discuss some choices of $\varphi_{\text{SE}}(\mathbf{x}_1)$ that are designed to reproduce some precise self-energy correction calculations, and are employed in our numerical calculations.

1. Pyykkö and Zhao SE potential

In Ref. 24, Pyykkö and Zhao (PZ) proposed a simple local self-energy potential, of the following form

$$\varphi_{\text{SE}}(\mathbf{x}) = B e^{-\beta r_x^2}. \quad (17)$$

The parameters B and β are quadratic nuclear charge (Z) dependent functions

$$B(Z) = -48.6116 + 1.53666 Z + 0.0301129 Z^2 \quad (18)$$

$$\beta(Z) = -12751.3 + 916.038 Z + 5.7797 Z^2 \quad (19)$$

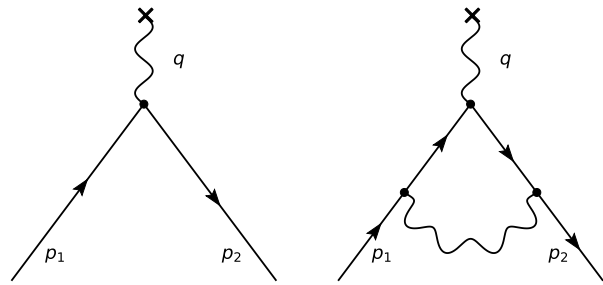
where the six decimal numbers were chosen to fit precise $29 \leq Z \leq 83$ atomic calculations of the renormalized self-energy correction in all orders of $(Z\alpha)^{n \geq 0}$ to the:

1. 2s energy-levels of the hydrogen-like systems, i.e., the renormalized version of Eq. (14), taken from calculations of 1) Beier *et al.*⁶⁹ with nuclear charges $26 \leq Z \leq 110$, using a homogeneously charged sphere nuclear model, and 2) Indelicato and Mohr⁷⁵ with Coulombic nuclear charges of $5 \leq Z \leq 90$.
2. M1 hyperfine splitting for lithium-like atoms taken from calculations⁷⁶ of Boucard and Indelicato⁷⁷ done on stable isotopes with $3 \leq Z \leq 92$.

2. Flambaum and Ginges SE potential

The starting point for the potential proposed by Flambaum and Ginges (FG)²⁵ is associated with the one-potential bound-state self-energy process, of order $\alpha(Z\alpha)$, given in Eqs. (A64) and (A68) and represented by Fig. 4b. However, further modeling, including parametrization, is introduced such that the potential can account for the full self-energy process to all orders in $(Z\alpha)$ and be used in atomic calculations.

In the evaluation of matrix elements over the operator of Eq. (A68), Flambaum and Ginges employ free-particle



(a) Classical scattering process. (b) First radiative correction.

FIG. 2. Momentum-space Feynman diagrams for the lowest-order scattering processes through (the exchange of) momentum-transfer $q = p_2 - p_1$.

solutions rather than atomic bound orbitals. This replacement yields the free-electron vertex-correction (VC) problem. This terminology can be understood from consideration of the scattering of a free electron due to the interaction with a classical external potential (the vertex process). In terms of momentum-space quantities (cf. Eq.(A65)), including *free* electron field operators, the corresponding (non-radiative) \mathcal{S} -matrix is given by

$$\hat{\mathcal{S}}_{\text{scattering}}^{(1)} = -\frac{e}{i\hbar} \int \frac{d^4 p_2}{(2\pi\hbar)^4} \int \frac{d^4 p_1}{(2\pi\hbar)^4} \times : \bar{\Psi}(p_2) \gamma^\mu A_\mu^e(p_2 - p_1) \hat{\Psi}(p_1) : \quad (20)$$

(see for instance section 8.7 in Ref. 78). This process is represented in the left panel of Fig. 2, where the wiggly line ending with a cross \times describes an interaction of a free electron with the classical external potential source through the exchange of a four-momentum $q = p_2 - p_1$. We note that in general a factor $(-e\gamma^\mu)$ is associated with each spacetime point (vertex). The right panel of Fig. 2 represents one of the four lowest-order radiative corrections to the left panel process. The corresponding \mathcal{S} -matrix can be combined with the one of Eq. (20) through the substitution

$$\gamma^\mu \rightarrow \Gamma^\mu = \gamma^\mu + \Lambda^\mu(p_2, p_1), \quad (21)$$

where the vertex-correction function $\Lambda^\mu(p_2, p_1)$ is given by Eq.(A69). After a careful treatment of the divergence when $q = 0$, as done in Refs. 60 (section 6.3) and 79 (section 117), one obtains the regularized (physical) vertex-correction function $\Lambda_R^\mu(p_2, p_1)$. Furthermore, using the fact that the vertex function is sandwiched between free-electron (on-mass-shell) field operators, one can show that this function can be written as

$$\Lambda_R^\mu(q) = \gamma^\mu F_1(q^2) + \frac{i}{2m_e c} \sigma^{\mu\nu} q_\nu F_2(q^2), \quad (22)$$

where $\sigma^{\mu\nu} = \frac{i}{2} [\gamma^\mu, \gamma^\nu]$, and F_1 and F_2 are known as the electric and magnetic form-factors, respectively (corresponding to $f - 1$ and g in Eq. (116.6) of Ref. 79). The

term “form factor” comes from diffraction physics; see for instance Ref. 80.

Since the free-electron vertex function of Eq. (22) only depends on the momentum-transfer $q = p_2 - p_1$, it conveniently yields a local potential in real space. This can be clearly seen from the following relation

$$\begin{aligned} & \int \frac{d^4 p_2}{(2\pi\hbar)^4} \int \frac{d^4 p_1}{(2\pi\hbar)^4} \bar{\Psi}(p_2) \Lambda_R^\mu(q) A_\mu^e(q) \hat{\Psi}(p_1) \\ &= \frac{1}{c} \int d^4 x \hat{\Psi}^\dagger(x) \varphi_{\text{VC}}(x) \hat{\Psi}(x). \end{aligned} \quad (23)$$

When the nucleus is described as a point charge the corresponding Coulomb potential,

$$A_0^e(q) = \delta(q_0) \frac{2\pi\hbar^3 Ze}{c\epsilon_0 \mathbf{q}^2}, \quad (24)$$

generates a vertex-correction potential of the form

$$\begin{aligned} \varphi_{\text{VC}}^{\text{point}}(\mathbf{x}) &= \frac{\hbar^2}{\epsilon_0} \int \frac{d^3 q}{(2\pi\hbar)^3} e^{+\frac{i}{\hbar}\mathbf{q}\cdot\mathbf{x}} \frac{Ze}{\mathbf{q}^2} \\ &\times \left[F_1(-\mathbf{q}^2) + \frac{1}{2m_e c} \boldsymbol{\gamma} \cdot \mathbf{q} F_2(-\mathbf{q}^2) \right] \\ &= \varphi_{\text{elec}}^{\text{point}}(\mathbf{x}) + \varphi_{\text{mag}}^{\text{point}}(\mathbf{x}), \end{aligned} \quad (25)$$

which splits into electric and magnetic scalar potentials. We note that due to the time-independence of the Coulomb potential, the time-like part of the 4-momentum transfer $q = p_2 - p_1$ vanishes. In terms of the variable $t = q^2 = -\mathbf{q}^2$, the form factors are Hermitian analytic functions,⁸¹ that is

$$F(t) = F^*(t^*). \quad (26)$$

This feature, combined with these functions being radial in terms of \mathbf{q} and the clever use of complex analysis techniques, allowed Berestetskii *et al.* to express such functions in coordinate-space using only their imaginary parts in momentum-space

$$F(\mathbf{x}) = \frac{1}{(2\pi\hbar)^2 r_x} \int_{4m_e^2 c^2}^{\infty} dt \Im [F(t)] \exp \left[-\frac{1}{\hbar} r_x \sqrt{t} \right]. \quad (27)$$

(see eq.(114.4) of Ref. 79). It may be noted that the lower limit of integration over t is $4m_e^2 c^2$, corresponding to the threshold of pair creation.⁸² Expressions for the imaginary part of the form factors can be found in Refs. 79 (eq.(117.14-15)) and 83 (eq.(2.12))

$$\begin{aligned} \Im [F_1(t)] &= \frac{\alpha}{\sqrt{t(t-4m_e^2 c^2)}} \left[2m_e^2 c^2 - 3t/4 \right. \\ &\quad \left. + (t/2 - m_e^2 c^2) \log \left(\frac{t-4m_e^2 c^2}{\lambda^2} \right) \right], \end{aligned} \quad (28)$$

$$\Im [F_2(t)] = \frac{\alpha m_e^2 c^2}{\sqrt{t(t-4m_e^2 c^2)}}. \quad (29)$$

Building on the work of Berestetskii *et al.*,⁷⁹ Flambaum and Ginges have evaluated the integral of Eq. (25), and obtained the associated real-space potentials. After the variable substitution $t = 4m_e^2 c^2 \zeta^2$, the magnetic potential was found to be

$$\varphi_{\text{mag}}^{\text{point}}(\mathbf{x}) = \frac{\alpha\hbar}{4\pi m_e c} i\boldsymbol{\gamma} \cdot \nabla_x \left[\frac{Ze}{4\pi\epsilon_0 r_x} \left(K_m \left(\frac{2r_x}{\lambda} \right) - 1 \right) \right], \quad (30)$$

where we have introduced the function

$$K_m(x) = \int_1^\infty d\zeta \frac{e^{-x\zeta}}{\zeta^2 \sqrt{\zeta^2 - 1}}, \quad (31)$$

which can be recognized as the 2nd Bickley–Naylor function Ki_2 (cf. Ref. 68). Note that the same variable ζ is employed in the Uehling potential (cf. Eqs. (9) and (12)). The magnetic contribution gives the first-order correction to the magnetic moment: the anomalous magnetic moment of the electron, first calculated by Schwinger, see for instance Mandl in Ref. 78 (section 10.5).

On the other hand, the electric form factor yields the electric effective potential

$$\varphi_{\text{elec}}^{\text{point}}(\mathbf{x}, \lambda) = -\frac{\alpha}{\pi} \frac{Ze}{4\pi\epsilon_0 r_x} K_e \left(\frac{2r_x}{\lambda} \right), \quad (32)$$

where we have introduced the function

$$\begin{aligned} K_e(x) &= \int_1^\infty d\zeta \frac{e^{-x\zeta}}{\sqrt{\zeta^2 - 1}} \left\{ -\frac{3}{2} + \frac{1}{\zeta^2} \right. \\ &\quad \left. + \left(1 - \frac{1}{2\zeta^2} \right) \left[\ln(\zeta^2 - 1) + 2 \ln \left(\frac{2m_e c^2}{\lambda} \right) \right] \right\}. \end{aligned} \quad (33)$$

These self-energy effective potentials were first derived with respect to a point nucleus (Coulomb potential), and the corresponding generalized expressions for an arbitrary normalized nuclear distribution ρ^{nuc} are obtained by convolution,⁸⁴ as in Eq. (10).

The potential of Eq. (32) is called the high-frequency term, because it contains an energy parameter λ , already present in Eq. (28), that prevents the obtention of low-frequency contributions. This parameter is associated with the introduction of a small fictitious photon mass, which needs to be plugged in the photon propagator denominator in order to make the divergent (at small momenta) momentum-space integral, associated with the vertex-correction, convergent. Details concerning this problem are discussed by Greiner and Reinhardt in Ref. 85 (eq.(5.91)), Itzykson and Zuber in Ref. 86 (eq.(7.45)), in addition to Peskin and Schroeder in Ref. 60 (pages 195,196). We note that the remaining divergence, occurring in the limit of zero photon mass, or $\lim_{\lambda \rightarrow 0}$ is overcome by taking into account the differential cross section associated with the Bremsstrahlung effect; detailed discussions are found in Refs. 85 (pages 311-313) and 60 (section 6.4). Flambaum and Ginges choose a somewhat different strategy, which furthermore allows them to amend the fact that the used form factors are derived for

the free-electron vertex-correction of order $(Z\alpha)^{n=1}$ only and now take into account complementary self-energy corrections $(Z\alpha)^n$, $n \neq 1$ (diagrams of Figs. 4a, 4c, 4d, and higher orders). They write the high-frequency (HF) contribution as

$$\varphi_{\text{HF}}^{\text{point}}(\mathbf{x}) = A(Z, \mathbf{x}) \varphi_{\text{elec}}^{\text{point}}(\mathbf{x}, \lambda), \quad (34)$$

where $A(Z, \mathbf{x})$ is a fitting function and choose a λ -value that will minimize the low-frequency contribution. They argue that λ should be on the order of electron binding energies, that is $(Z\alpha)^2 m_e c^2$. They finally define it through

$$\ln\left(\frac{2m_e c^2}{\lambda}\right) = 2 \ln\left(\frac{1}{Z\alpha} + \frac{1}{2}\right), \quad (35)$$

though, for better performance. Flambaum and Ginges next argue that the low-frequency (LF) potential should have the range of a $1s$ orbital of hydrogen-like atoms and therefore choose the functional form

$$\varphi_{\text{LF}}^{\text{point}}(\mathbf{x}) = -\frac{B(Z)}{e} Z^4 \alpha^5 m_e c^2 e^{-Zr_x/a_0}, \quad (36)$$

where $a_0 = \lambda/\alpha$ is the Bohr radius and

$$B(Z) = 0.074 + 0.35 \times Z\alpha, \quad (37)$$

is a second fitting function.

The fitting function of the high-frequency contribution is written as

$$A(Z, \mathbf{x}) = \Theta(Z, \mathbf{x}) (1.071 - 1.976y^2 - 2.128y^3 + 0.169y^4), \quad (38)$$

in terms of the variable $x = (Z - 80)\alpha$ and a cutoff-function of the form

$$\Theta(Z, \mathbf{x}) = \frac{r_x}{r_x + 0.07(Z\alpha)^2 \lambda}, \quad (39)$$

which will dampen the contribution of $\varphi_{\text{elec}}^{\text{point}}$ at short distances where the the locality of the effective SE potential breaks down. The coefficients of the A and B fitting functions above were adjusted to reproduce the self-energy corrections to high s - and p -states, respectively, calculated accurately in Refs. 87,88 for Coulombic hydrogen-like atoms of $5 \leq Z \leq 110$. It should be added that Thierfelder and Schwerdtfeger⁸⁹ later modified the fitting function to $A_n(Z, \mathbf{x})$, that is, making it dependent of the principal quantum number n . These potentials with $A_n(Z, \mathbf{x})$ instead of $A(Z, \mathbf{x})$ were used by Pašteka *et al.* to calculate the electron affinity and ionization potential of gold.³² Ginges and Berengut,⁸⁴ on the other hand, made both fitting functions A and B dependent on orbital angular momentum ℓ and further suggest to introduce a κ -dependence as well. The downside of making the effective QED potentials dependent on atomic orbital quantum numbers is that it complicates the extension of these potentials to the molecular regime. This is in fact what motivated us to implement the effective SE potentials of Pyykkö/Zhao²⁴ and Flambaum/Ginges,²⁵ rather the one proposed by Shabaev and co-workers.^{26,27,29}

C. Atomic shift operator

With the above effective QED potentials available in an atomic code (see Section III), we have investigated their extension to molecular calculations by adding to the electronic Hamiltonian, Eq. (1), an operator on the form

$$V_{\text{ASHIFT}} = \sum_i |\psi_i\rangle \omega_i \langle \psi_i|, \quad (40)$$

$$\omega_i = \langle \psi_i | -e\varphi_{\text{effQED}} | \psi_i \rangle,$$

where $\{\omega_i\}$ are expectation values of the effective QED potentials taken from atomic calculations and $\{\psi_i\}$ are pre-calculated atomic orbitals, in practice limited to those that are occupied in the electronic ground state of the atoms constituting the molecule under study, calculated in their proper basis. The import of atomic orbitals into molecular calculations is straightforward in the case of the DIRAC code, since such functionality is already available through projection analysis.^{90,91} There is some overlap between the spectral representation of the self-energy proposed by Dyall⁹² as well as the effective SE operator proposed by Shabaev and co-workers,²⁶ but those approaches are based on hydrogenic orbitals.

III. IMPLEMENTATION

Routines for the radiative potentials used in this work are available in the GRASP atomic code.²⁸ Routines for calculating the Uehling potential were reported as early as 1980.⁹³ McKenzie *et al.* follow the approach suggested by Wayne Fullerton and Rinker.⁶⁶ More precisely, they employ Eq. (11) for the inner grid points until a more approximate form, Eq. (6) of Ref. 66, becomes numerically valid. The latter form is then used until the magnitude of the potential falls below a threshold value. The effective SE potential of Flambaum and Ginges²⁵ was implemented more recently,⁸⁹ as is also the case⁹⁴ of the effective SE potential of Pyykkö and Zhao.²⁴ As already mentioned, the FG potential is in principle that associated with a point nucleus, although fitting parameters have been optimized also to calculations with finite nuclear charge distributions. Thierfelder and Schwerdtfeger⁸⁹ adapted these potentials to finite nuclei by replacing the Coulomb potentials of Eqs. (30) and (32) by the potentials of finite nuclear charge distributions, and we have so far followed this approach which appears to be a reasonable approximation, as can be inferred from Table IV of Ref. 84.

We have adapted the GRASP effective QED potential routines to molecular calculations by using the numerical integration scheme implemented for relativistic Kohn–Sham calculations in the DIRAC molecular code.⁴⁹ The scheme is based on the Becke partitioning⁹⁵ of the molecular volume into atomic ones for which numerical integration is carried out in spherical coordinates. Specifically, we use Lebedev angular quadrature,⁹⁶ by default

setting $\ell = 15$, combined with the basis-set adaptive radial grid proposed by Lindh and co-workers.⁹⁷ It may be noted that the effective QED potentials presented in the previous section are all radial, with the exception of the magnetic contribution to the Flambaum–Ginges SE potential, Eq. (30).

Due to the very local nature of the effective QED potentials⁴⁴ one-electron integrals over a potential associated with atomic center A can be well approximated by

$$V_{\mu\nu}^A \approx \int_0^{R_A} dr_A \int_{\Omega} d\Omega_A [\chi_{\mu} v^A \chi_{\nu}] (\mathbf{r}_A) r_A^2, \quad (41)$$

where $\{\chi_{\mu}\}$ are Gaussian-type basis functions. The most delocal potential is the low-frequency contribution to the electric form factor of the Flambaum–Ginges SE potential, Eq. (36), since it has been designed to have the range of the $1s$ orbital of a hydrogen-like atom. For low Z the potential may thereby overlap significantly with neighbor centers. By default, we therefore deactivate the effective QED potentials for $Z < 19$. We also determine the value of the upper limit of radial integration R_A based on the convergence of the low-frequency term to a very conservative 10^{-50} .

IV. COMPUTATIONAL DETAILS

For all calculations we used a development version of DIRAC code;^{48,98} precise version and build information is found in output files, see Ref. 99. A Gaussian model¹⁰⁰ for the nuclear charge distribution was employed throughout our calculations. Unless otherwise stated, we applied the Uehling VP potential²³ and the SE potential of Flambaum and Ginges,²⁵ added to the Dirac–Coulomb–Gaunt (DCG) Hamiltonian. For correlated calculations we employed the molecular mean-field approximation Hamiltonian (X2Cmmf)¹⁰¹ based on the DCG Hamiltonian, which we denote as ${}^2\text{DCG}^M$. In this approach, the converged Fock matrix obtained with the DCG Hamiltonian, with the effective QED potentials included, is exactly transformed to two-component form, that is, without any picture-change errors.^{45–47} All basis sets were employed in uncontracted form with the small component generated by restricted kinetic balance (see Ref. 48 for details). Electronic structure analysis was carried out using projection analysis⁹¹ where Pipek–Mezey localized MOs^{91,102} are expanded in intrinsic, hence polarized, atomic orbitals.¹⁰³ The analysis was done at the molecular geometries optimized with respect to the employed Hamiltonian, except for DCG with effQED, where the DCG structures were employed.

For the atomic calculations reported in Table I we employed Dyall v3z basis sets;^{104–109} the basis set for Uue was specially optimized by Dyall for this work.¹¹⁰

For van der Waals dimers, the following orbitals were correlated: $5d6s$ for Hg, $5d6s6p$ for Rn, $6d7s$ for Cn, and

$6d7s7p$ for Og. We used an virtual energy cutoff of $40 E_h$. Dyall cv3z basis sets,^{106–108} designed for core-valence correlation, were employed for the Hg and Cn species, whereas Dyall acv3z basis sets,^{111–113} where the Dyall cv3z basis sets have been augmented by diffuse functions, were employed for Rn and Og species. Electronic structure calculations were done at the level of coupled-cluster singles-and-doubles with approximate triples correction (CCSD(T)) using the RELCCSD module.¹¹⁴ We used the counterpoise correction¹¹⁵ to minimize basis set superposition errors (BSSE).

For the calculations of gold cyanide, we used the CCSD(T) method for comparison with experiment. In the CCSD(T) calculation, $4f5s5p5d6s$ for Au, and all electrons of C and N were correlated, which is the same level as the previous work.¹¹⁶ Dyall ae3z and ae4z basis sets,^{106,107,109} designed for correlation of all electrons, were employed in the calculations. We employed a virtual energy cutoff of about $50 E_h$ and $80 E_h$ for dyall.ae3z and dyall.ae4z, respectively, which assures that correlating h and i orbitals, respectively, are included. Effective QED potentials for C and N atoms were not used, as explained in Section III. The potential energy surface (PES) was calculated in the vicinity of the equilibrium structure with a total of 49 points for each basis set, using internal coordinates r_1 (Au–C distance), r_2 (C–N distance). The bond angle was fixed at $\theta = 180^\circ$. The step size for bond distances was $0.1 a_0$. The surface fitting and determination of the equilibrium structure was carried out using the SURFIT program,¹¹⁷ with convergence 3.2×10^{-10} or better on the gradient. In addition, to estimate the relativistic effects we employed the two-component non-relativistic (by using .NONREL keyword), 4c-scalar-relativistic,^{118,119} and the Dirac–Coulomb Hamiltonians at the density functional theory (DFT) level. In these calculations we employed the B3LYP functional^{120,121} and the dyall.3zp basis sets.^{106,107,109}

For the calculation of Pb and Fl hydrides, the DCG Hamiltonian with and without effective QED potentials, as well as the Lévy-Leblond (LL)^{119,122} Hamiltonian were employed. The dyall.3zp basis sets were used for all of the elements. The B3LYP functional was used for both the projection analysis and the geometry optimization.

V. RESULTS

A. Atomic calculations

In Table I we show results of atomic calculations, using average-of-configuration (AOC) HF,^{124,125} that can be directly compared with Table I of Ref. 17 and which provide estimates for the valence-level Lamb shift for group 1 and 11 metal atoms. Pyykkö *et al.* focused on $ns_{1/2}$ orbital energies for estimating ionization energies, albeit, as pointed out by Thierfelder and Schwerdtfeger,⁸⁹ for Roentgenium ($Z = 111$), the first ionization is out of the $6d_{5/2}$ orbital. The VP and SE contributions

TABLE I. Calculated ns orbital energies in eV of group 1 and 11 elements from AOC-HF/dyall.v3z calculations based on the NR and DC Hamiltonians. The VP (Uehling) and SE (Flambaum–Ginges) corrections have been calculated as expectation values.

	NR	DC	VP	SE	Δ QED	SE/VP	SE/VP ¹²³	Δ QED/ Δ R[%]
Li	-5.342	-5.343	-1.373E-06	4.092E-05	3.955E-05	-29.7949	-29.7058	-9.01
Na	-4.955	-4.962	-1.536E-05	2.950E-04	2.796E-04	-19.2057	-18.7963	-4.37
K	-4.013	-4.028	-3.423E-05	5.155E-04	4.813E-04	-15.0615	-14.7030	-3.19
Rb	-3.752	-3.811	-1.309E-04	1.361E-03	1.231E-03	-10.3981	-10.0783	-2.08
Cs	-3.365	-3.490	-2.989E-04	2.304E-03	2.005E-03	-7.7089	-7.4266	-1.61
Fr	-1.740	-3.611	-1.438E-03	6.333E-03	4.895E-03	-4.4038	-4.3351	-0.26
Uue	-2.993	-4.327	-1.034E-02	2.157E-02	1.123E-02	-2.0859	-4.3351	-0.84
Cu	-6.480	-6.649	-2.355E-04	2.840E-03	2.604E-03	-12.0606	-11.7316	-1.54
Ag	-5.985	-6.452	-7.342E-04	6.448E-03	5.714E-03	-8.7825	-8.4755	-1.22
Au	-6.003	-7.923	-4.635E-03	2.374E-02	1.910E-02	-5.1219	-4.9912	-1.00
Rg	-5.441	-11.425	-3.251E-02	8.408E-02	5.157E-02	-2.5863	-2.7223	-0.86

TABLE II. Relativistic and QED effects on the orbital energies $\varepsilon(E_h)$ of the Au atom at the B3LYP/dyall.3zp level. The Uehling VP potential has been combined with two different SE potentials: FG (Flambaum–Ginges) and PZ (Pyykkö–Zhao) in variational calculations. Numbers in parentheses shows the percentage-wise ratio Δ QED/ Δ R for each combination of effective QED potentials.

	NR	DCG	Δ (U+FG)		Δ (U+PZ)	
1s _{1/2}	-2689.451	-2955.841	6.377E00	(-2.39)	6.243E00	(-2.34)
2s _{1/2}	-449.932	-523.020	8.448E-01	(-1.16)	8.586E-01	(-1.17)
2p _{1/2}	-432.492	-500.523	5.895E-02	(-0.09)	5.955E-02	(-0.09)
2p _{3/2}	-432.492	-433.755	1.194E-01	(-9.45)	-3.055E-02	(2.42)
3s _{1/2}	-105.753	-123.999	1.862E-01	(-1.02)	1.924E-01	(-1.05)
3p _{1/2}	-97.515	-114.096	7.320E-03	(-0.04)	1.462E-02	(-0.09)
3p _{3/2}	-97.515	-99.311	2.213E-02	(-1.23)	-8.516E-03	(0.47)
3d _{3/2}	-82.165	-83.236	-1.430E-02	(1.34)	-9.090E-03	(0.85)
3d _{5/2}	-82.165	-80.090	-5.817E-03	(-0.28)	-8.587E-03	(-0.41)
4s _{1/2}	-22.559	-27.178	4.650E-02	(-1.01)	4.833E-02	(-1.05)
4p _{1/2}	-18.993	-22.979	1.124E-03	(-0.03)	3.336E-03	(-0.08)
4p _{3/2}	-18.993	-19.424	4.651E-03	(-1.08)	-2.427E-03	(0.56)
4d _{3/2}	-12.440	-12.630	-3.439E-03	(1.82)	-2.307E-03	(1.22)
4d _{5/2}	-12.440	-11.971	-1.648E-03	(-0.35)	-2.185E-03	(-0.47)
4f _{5/2}	-3.648	-3.228	-2.383E-03	(-0.57)	-1.461E-03	(-0.35)
4f _{7/2}	-3.648	-3.091	-1.824E-03	(-0.33)	-1.425E-03	(-0.26)
5s _{1/2}	-3.253	-4.116	9.003E-03	(-1.04)	9.430E-03	(-1.09)
5p _{1/2}	-2.108	-2.745	-4.569E-05	(0.01)	4.105E-04	(-0.06)
5p _{3/2}	-2.108	-2.139	5.619E-04	(-1.82)	-5.703E-04	(1.85)
5d _{3/2}	-0.346	-0.333	-4.662E-04	(-3.61)	-3.486E-04	(-2.70)
5d _{5/2}	-0.346	-0.276	-2.881E-04	(-0.41)	-3.206E-04	(-0.46)
6s _{1/2}	-0.148	-0.205	6.519E-04	(-1.14)	6.726E-04	(-1.18)

come with opposite sign and are dominated by the latter.¹²³ However, the ratio SE/VP decreases significantly with increasing nuclear charge and, indeed, VP is predicted to eventually overtake SE at very high nuclear charges.⁸⁹ Pyykkö *et al.* calculated the SE contribution as $\langle V^U \rangle * (SE/VP)$ where (SE/VP) is the ratio for 2s_{1/2} of the corresponding hydrogen-like systems, including the nuclear-size effect, tabulated for $1 \leq Z \leq 100$ by Johnson and Soff¹²³ (a more recent compilation is provided by Yerokhin and Shabaev¹²⁶). As confirmed by later calculations¹²⁷ and the numbers in Table I, this is a quite reasonable approximation.

Comparing relativistic and QED effects, one sees that

the latter corrects the former by about -1% for the heavier atoms. For the gold atom it is exactly so. In Table II we show the effect of relativity and QED on all orbital energies of the gold atom. Two combinations of effective QED potentials have been used in variational calculations: The Uehling (U) VP potential has been combined either with the Flambaum–Ginges (FG) or Pyykkö–Zhao SE potentials. One sees that for both combinations of effective QED potentials the relativistic effects is, with very few exceptions, reduced with a few percent. For s_{1/2} orbitals the difference in QED shift between the U+FG and U+PZ combinations is below 5%; for other orbitals the difference is generally larger. We note in particu-

TABLE III. First-order QED effects on the orbital energies $\varepsilon(E_h)$ of the Au atom at the B3LYP/dyall.3zp level using the DCG Hamiltonian or the X2C Hamiltonian, the latter with or without picture-change (PC) transformation. The Uehling VP potential has been combined with two different SE potentials: FG (Flambaum–Ginges) and PZ (Pyykkö–Zhao). Numbers in parentheses shows the percentage-wise ratio $\Delta\text{QED}/\Delta$ for each combination of effective QED potentials.

	DCG		X2C - PC		X2C - noPC	
	$\Delta(\text{U+FG})$	$\Delta(\text{U+PZ})$	$\Delta(\text{U+FG})$	$\Delta(\text{U+PZ})$	$\Delta(\text{U+FG})$	$\Delta(\text{U+PZ})$
1s _{1/2}	6.620E00	6.465E00	6.632E00	6.477E00	7.522E00	8.253E00
2s _{1/2}	9.003E-01	9.022E-01	9.012E-01	9.030E-01	1.079E00	1.175E00
2p _{1/2}	1.184E-01	1.051E-01	1.187E-01	1.047E-01	1.971E-01	3.608E-02
2p _{3/2}	1.671E-01	4.922E-03	1.682E-01	5.005E-03	1.279E-01	6.342E-03
3s _{1/2}	2.016E-01	2.040E-01	2.020E-01	2.043E-01	2.441E-01	2.665E-01
3p _{1/2}	2.357E-02	2.674E-02	3.574E-02	1.357E-03	2.671E-02	1.727E-03
3p _{3/2}	3.554E-02	1.335E-03	3.574E-02	1.357E-03	2.671E-02	1.727E-03
3d _{3/2}	-1.234E-03	1.064E-05	-1.220E-03	1.066E-05	3.906E-03	3.605E-07
3d _{5/2}	6.494E-03	-1.680E-07	6.528E-03	-1.675E-07	3.296E-03	-1.132E-07
4s _{1/2}	5.082E-02	5.151E-02	5.097E-02	5.166E-02	6.173E-02	6.741E-02
4p _{1/2}	5.631E-03	6.658E-03	5.638E-03	6.637E-03	9.363E-03	2.339E-03
4p _{3/2}	8.392E-03	3.315E-04	8.444E-03	3.370E-04	6.292E-03	4.295E-04
4d _{3/2}	-1.331E-04	2.846E-06	-1.299E-04	2.851E-06	9.447E-04	9.962E-08
4d _{5/2}	1.461E-03	-4.294E-08	1.468E-03	-4.277E-08	7.944E-04	-2.822E-08
4f _{5/2}	-2.785E-04	-1.968E-10	-2.782E-04	-1.974E-10	1.089E-05	-1.184E-10
4f _{7/2}	2.190E-04	-9.047E-11	2.194E-04	-9.111E-11	9.735E-06	-9.557E-11
5s _{1/2}	1.001E-02	1.015E-02	1.004E-02	1.018E-02	1.217E-02	1.329E-02
5p _{1/2}	9.658E-04	1.152E-03	9.668E-04	1.149E-03	1.605E-03	4.052E-04
5p _{3/2}	1.373E-03	5.478E-05	1.382E-03	5.572E-05	1.029E-03	7.103E-05
5d _{3/2}	-1.078E-05	2.580E-07	-1.048E-05	2.582E-07	8.410E-05	9.077E-09
5d _{5/2}	1.216E-04	-3.635E-09	1.222E-04	-3.618E-09	6.669E-05	-2.375E-09
6s _{1/2}	7.888E-04	8.001E-04	7.908E-04	8.020E-04	9.586E-04	1.047E-03

lar that the shifts have systematically opposite sign for $p_{3/2}$ orbitals. Not surprisingly the largest absolute shifts are observed for inner core orbitals, whereas the largest relative shift -0.33% is seen for the $6s_{1/2}$ orbital.

In Table III we show QED shifts of orbital energies, this time obtained perturbatively as expectation values. Compared to the shifts obtained from variational inclusion of the effective QED potentials, the largest absolute deviations concern the inner core orbitals. The smallest relative deviations are observed for $s_{1/2}$ orbitals and decreasing towards core. The largest relative deviations, on the other hand, are seen for p orbitals; the very largest relative deviation concerns $5p_{1/2}$, but this can probably be attributed to noise, since the QED shift on the energy of this orbital is particularly small with both combinations of effective QED potentials.

Table III also shows perturbative QED shifts of orbital energies obtained with the X2C Hamiltonian. When the effective QED potentials have been correctly picture-changed transformed, deviations from the parent 4c calculation are below 3 %, which clearly validates the use of these potentials in 2-component relativistic calculations. On the other hand, without picture-change, significant errors are observed; the average unsigned error for U+FG and U+PZ is 130 % and 47 %, respectively. This is possibly worrisome since the U+PZ combination, expressed in terms of Gaussians, have been used without picture-change in scalar DKH calculations by Peterson and co-workers.^{33–36}

TABLE IV. M-C bond lengths (in pm) in MCN (M=Cu, Ag, Au) from microwave (MW) spectroscopy and calculations.

	CuCN	AgCN	AuCN	
r_o	183.231(7)	203.324(45)	191.251(16)	MW ^a
r_s	183.284(4)	203.4182(27)	191.22519(84)	MW ^a
r_e	182.36	202.42	191.05	Calc. ¹²⁸
r_e	182.65	202.99	190.71	Calc. ¹¹⁶

^aCuCN: Ref. 129. AgCN, AuCN: Ref. 130

B. Gold cyanide

In 2008 Pyykkö and co-workers reported CCSD(T)/cc-pVQZ calculations on the noble metal cyanides (MCN, M=Cu, Ag, Au).¹²⁸ Small-core scalar-relativistic effective core potentials (SRECP)¹³¹ were used for the metal atoms and spin-orbit corrections added at the PBE-ZORA/QZ4P level. In 2013 Peterson and co-workers reported CCSD(T)-F12/cc-pV5Z calculations on the same compounds, using the same SRECPs as the previous authors and adding a number of corrections.¹¹⁶ As seen from Table IV the newer calculations brought the M-C bond lengths of CuCN and AgCN in better agreement with experiment, but increased the gap between theory and experiment for AuCN. This led Pyykkö to conjecture that this could be the first evidence of the effect of QED on molecular structure.⁵⁰

To possibly verify this conjecture we first carried out

TABLE V. Relativistic and QED effects on the rms radius $\langle r^2 \rangle^{1/2}$ (pm) of the Au atom at the B3LYP/dyall.3zp level. Effective QED potentials: VP(Uehling)+SE(Flambaum-Ginges).

	NR	DCG	DCG+QED	ΔR	ΔQED	$\Delta QED/\Delta R$ (%)
5s _{1/2}	57.71	52.23	52.28	-5.48	0.05	-0.86
5p _{1/2}	63.16	56.78	56.78	-6.38	0.00	-0.02
5p _{3/2}	63.16	62.37	62.38	-0.80	0.01	-1.04
5d _{3/2}	91.07	90.81	90.78	-0.26	-0.03	10.46
5d _{5/2}	91.07	95.75	95.73	4.68	-0.02	-0.36
6s _{1/2}	196.07	167.54	167.79	-28.53	0.25	-0.88

exploratory calculations at the B3LYP/dyall.3zp level. Table V shows the effects of relativity and QED on orbital sizes of the gold atom. For the valence 6s_{1/2} we observe an impressive relativistic contraction of 28.53 pm, whereas QED leads to an orbital expansion of 0.25 pm, roughly -1 % of the relativistic effect.

We next turn to the AuCN molecule. We first, in Table VI, report bonding analysis in localized orbitals.⁹¹ One finds a single σ -type Au-C bond, dominated by carbon 2s_{1/2} and gold 6s_{1/2}, as well as a triple C-N bond. Equilibrium bond lengths r_e with respect to different Hamiltonians are reported in Table VII. We see a very significant scalar relativistic bond-length contraction of 25.31 pm, on par with the 6s_{1/2} orbital contraction observed in Table V. When going from a spin-free (SF) Hamiltonian to the Dirac-Coulomb one, one finds a further contraction of 0.29 pm, which agrees very well with the spin-orbit correction of -0.28 pm obtained by Hill *et al.* taking the same difference, albeit at the CCSD(T) level.¹¹⁶ However, this contraction is almost canceled when adding the Gaunt term, which brings spin-other-orbit interaction³ and which was not considered by Hill and co-workers.¹¹⁶ At this level of theory, the total relativistic effect on the bond length is thereby -25.38 pm; in future work we hope to assert the effect of replacing the Gaunt term by the full Breit one. Finally, adding QED effects, we observe a bond-length extension of 0.19 pm, -0.75 % of the relativistic effect. One may note that the QED effect is of the same order as the effect of adding the Gaunt term.⁸⁹ In passing we note from Table VII that incorporation of QED effects through the atomic shift operator (ASHIFT) described in Section IIC also leads to a bond extension, albeit only capturing half of the full QED effect.

To obtain more accurate bond lengths, we proceeded as indicated in Table VIII: ²DCG^M-CCSD(T) calculations were carried out in the Dyall ae3z and ae4z basis sets and then extrapolated to the basis-set limit,¹³² indicated by “ae ∞ z”. We then added the triples ΔT and quadruples ΔQ corrections reported by Hill *et al.*¹¹⁶ to obtain a Au-C bond length of 190.75 pm, very close to the value 190.71 pm reported by Peterson and co-workers. Finally, we add a QED correction of 0.19 pm, identical to what we obtained at the B3LYP/dyall.3zp level, to obtain our final value of 190.99 pm.

TABLE VI. Gross populations obtained by projection analysis using Pipek-Mezey localized orbitals at the DCG/B3LYP/dyall.3zp level. $\langle \epsilon \rangle$ is the expectation value in E_h with respect to the Kohn-Sham matrix.

		Au		C			N		
ω	$\langle \epsilon \rangle$	5d _{5/2}	6s _{1/2}	2s _{1/2}	2p _{1/2}	2p _{3/2}	2s _{1/2}	2p _{1/2}	2p _{3/2}
3/2	-0.342	-0.01	0.00	0.00	0.00	0.92	0.00	0.00	1.09
1/2	-0.345	0.00	0.00	0.00	0.65	0.28	0.00	0.64	0.43
1/2	-0.586	0.15	0.39	0.92	0.17	0.34	0.01	-0.01	-0.02
1/2	-0.781	0.00	0.01	0.24	0.12	0.38	0.15	0.44	0.65

TABLE VII. Equilibrium bond lengths r_e (in pm) of AuCN calculated at the B3LYP/dyall.3zp level using various Hamiltonians. Numbers in parenthesis indicate the change with respect to the previous line, except ASHIFT, which refers to DCG. SF refers to a spin-free 4-component relativistic Hamiltonian.

Hamiltonian	Au-C	C-N
NR	218.54	115.71
SF	193.23 (-25.31)	115.54 (-0.17)
DC	192.94 (-0.29)	115.56 (+0.02)
DCG	193.16 (+0.22)	115.58 (+0.01)
QED	193.35 (+0.19)	115.57 (+0.00)
ASHIFT	193.25 (+0.09)	115.58 (+0.00)

The devil is, however, in the details: Our Born-Oppenheimer equilibrium bond lengths r_e are not directly comparable to the structural parameters extracted from the rotational spectra recorded by Okabayashi and co-workers.¹³⁰ Experiment gives access to rotational constants B_ν for individual vibrational states. For a linear molecule like AuCN the rotational constant, in units of frequency, is expressed as

$$B = \frac{\hbar}{4\pi I_\perp}; \quad I_\perp = I_{xx} = I_{yy} = \sum_A m_A z_A^2, \quad (42)$$

when the molecular axis is aligned with the z -axis. z_A is the distance of atom A from the center of mass. Effective r_0 and substitution r_s structures are both obtained by assuming identical structures for all isotopomers of

TABLE VIII. Final, recommended equilibrium Au-C bondlength r_e (pm) at the ²DCG^M-CCSD(T) level for the AuCN molecule. ΔQED is the difference between the extrapolated basis set limit ae ∞ z with QED and without QED.

	Au-C	C-N
ae3z	190.89	116.66
ae4z	190.70	116.28
ae ∞ z	190.58	116.07
ΔT ¹¹⁶	0.26	-0.10
ΔQ ¹¹⁶	-0.09	0.19
Final w/o QED	190.75	116.16
ΔQED	0.19	0.00
Final	190.94	116.16

the target molecule observed in experiment.^{133,134} Effective structures r_0 are obtained by least-square fitting of experimental ground-state inertial moments, whereas substitution structures r_s are obtained from observation of how rotational constants (and center of mass) change upon single isotope substitution $A \rightarrow A'$. For a linear molecule one has

$$|z_A| = \sqrt{\frac{\hbar}{4\pi\mu} \left(\frac{1}{B^{A'}} - \frac{1}{B^A} \right)}; \quad \frac{1}{\mu} = \frac{1}{M} + \frac{1}{\Delta m_A}. \quad (43)$$

where M is the total mass of the parent isotopomer. In the case of AuCN $|z_C|$ and $|z_N|$ could be estimated from corresponding single isotope substitutions. However, since gold has a single naturally occurring isotope, ^{197}Au , $|z_{Au}|$ was obtained from the definition of center of mass.¹³⁵

Empirically one typically finds $r_0 \geq r_s \geq r_e$,¹³⁴ which suggests that we should rather compare our recommended $r_e=190.99$ pm for Au-C with the corresponding substitution bond length $r_s=191.22519(84)$ pm reported by Okabayashi and co-workers.¹³⁰ However, a better comparison is provided by calculating the ground-state rotational constant B_0 from B_e . From perturbation theory, excluding Fermi resonances, the rotational constant for a given vibrational state ν of a general molecule is related to B_e as follows:¹³⁴

$$B_\nu^\xi = B_e^\xi - \sum_i \alpha_i^\xi \left(\nu_i + \frac{d_i}{2} \right) + \frac{1}{2} \sum_{i,j} \gamma_{i,j}^\xi \left(\nu_i + \frac{d_i}{2} \right) \left(\nu_j + \frac{d_j}{2} \right) + \dots \quad (44)$$

Here, ξ is the axis of rotation, α^ξ and γ^ξ are vibration-rotation interaction constants of different orders and d_i is the degeneracy of vibration mode i . The series generally converges rapidly, and for AuCN a suitable expression is therefore

$$B_0 \approx B_e - \frac{1}{2} [\alpha_{100} + \alpha_{001} + 2\alpha_{0110}], \quad (45)$$

using the notation $\alpha_{\nu_1\nu_2\nu_3}$, where ν_1 corresponds to the C-N stretch, ν_2 to the doubly degenerate bending mode and ν_3 to the Au-C stretch.

Hill *et al.*¹¹⁶ carried out both perturbative and variational rovibrational calculations. Using their calculated potential surfaces¹³⁶ we have extracted vibration-rotation interaction constants $\alpha_{\nu_1\nu_2\nu_3}$. Combined with our best equilibrium structures from Table VIII, we have calculated the rotational constants of the ground vibrational state of the three isotopomers of AuCN studied by Okabayashi and co-workers.¹³⁰ As can be seen from Table IX, the inclusion of QED corrections brings about a dramatic improvement with respect to experiment. Not surprisingly then, when we extract substitution structures by the same procedure as Okabayashi and co-workers,¹³⁰ correcting for Lamb shift effects bring our calculated substitution bond lengths within 0.05 pm of the experimental ones (cf. Table X). We expect further refinement of

TABLE IX. Calculated and experimental rotational constants (in MHz) for AuCN.

	$^{197}\text{Au}^{12}\text{C}^{14}\text{N}$	$^{197}\text{Au}^{13}\text{C}^{14}\text{N}$	$^{197}\text{Au}^{12}\text{C}^{15}\text{N}$
α_1	14.55	13.53	14.06
α_2	-10.98	-10.25	-10.59
α_3	12.17	11.91	11.30
	w/o QED		
B_e	3237.5	3184.5	3086.4
B_0	3235.1	3182.1	3084.3
	with QED		
B_e	3232.8	3179.9	3082.0
B_0	3230.4	3177.5	3079.9
$B_0(\text{exp.})$ ¹³⁰	3230.21115(18)	3177.20793(13)	3079.73540(12)

TABLE X. Calculated and experimental substitution structure (in pm) for AuCN

	$r_s(\text{Au-C})$	$r_s(\text{C-N})$
w/o QED	190.991	115.910
with QED	191.184	115.909
Exp.	191.22519(84)	115.86545(97)

the potential surface to improve agreement with experiment; we note for instance that our calculated vibration-rotation interaction constant associated with bending for the most abundant isotopomer $^{197}\text{Au}^{12}\text{C}^{14}\text{N}$ is -10.98 MHz, compared to -11.9781 MHz when extracted from experiment.¹¹⁶

C. van der Waals dimers

As a second molecular application of our implementation we consider spectroscopic constants of dimers with van der Waals bonding (M_2 , $M = \text{Hg, Rn, Cn, Og}$). In Table XI we report our calculated the equilibrium bond lengths r_e , harmonic frequencies ω_e , anharmonic constants $\omega_e x_e$ and dissociation energies D_e for these species. We see that the QED effect on bond length is on the order of 0.15 pm for row-6 dimers and approximately doubles when going to the superheavy elements; for Og_2 the QED bond length extension is in line with what was reported by Hangele and Dolg using relativistic effective core potentials.⁴⁰ The QED effect on dissociation energies is rather small: on the order of 0.4 % for the superheavy dimers.

D. Reaction energies: Pb and Fl hydrides

As a final case study, we consider the reaction energy of $\text{XH}_4 \rightarrow \text{XH}_2 + \text{H}_2$ to which Dyall *et al.* proposed that the Lamb shift could make a chemically significant contribution.⁵³ Their argument was based on the observation that QED effects are most important for s orbitals, as seen in Tables II and V, and that this is

TABLE XI. Spectroscopic constants of heavy group 12 and 18 dimers obtained at the ${}^2\text{DCG}^M\text{-CCSD(T)}$ level, using the U+FG combination of effective QED potentials and either dyall.cv3z (Hg_2, Cn_2) or dyall.acv3z (Rn_2, Og_2) basis sets. Numbers in parentheses indicate the QED effect.

	r_e/pm	ω_e/cm^{-1}	$\omega_e x_e/\text{cm}^{-1}$	D_e/cm^{-1}
Hg_2	385.71 (0.15)	16.65 (-0.03)	0.232 (-0.002)	277.7 (-0.02)
Cn_2	354.75 (0.36)	22.95 (-0.11)	0.255 (0.001)	532.7 (-2.78)
Rn_2	463.60 (0.14)	13.79 (-0.02)	0.286 (-1.E-04)	174.9 (-0.41)
Og_2	449.97 (0.28)	17.10 (-0.04)	0.210 (0.001)	391.1 (-1.32)

a reaction with a significant change of the valence s population of a heavy element. We have investigated this at the B3LYP/dyall.3zp level and also included the corresponding reaction involving the heavier homologue flerovium. Optimized equilibrium structures are given in Table XII. For the tetrahydrides we assumed T_d symmetry, in line with experiment¹³⁷(PbH_4) and previous calculation¹³⁸(FlH_4).

To monitor valence s populations we carried out bonding analysis in Pipek–Mezey localized MOs.^{91,102} From Table XIII, the change of the valence s population from XH_4 to XH_2 is 0.45 and 0.30 for Pb and Fl systems, respectively. From Table XIV, one sees that in the tetrahydrides the valence s population is contained in the four σ_{XH} bonds. In contrast, in the dihydrides the two σ_{XH} bonds are mediated by the valence p orbitals of the metals, and most of the valence s population is found in a non-bonding (nb) orbital.

Turning next to Tables XV and XVI we see that both reactions are endothermic at the non-relativistic level, but becomes clearly exothermic when adding relativity. For Pb QED reduces the relativistic effect by 1.25%. Its value is 0.32 kcal/mol, which is at the lower end of the perturbation estimate of Dyall *et al.*⁵³ For Fl QED reduces the relativistic effect by 0.50%. Interestingly, its value is very close to that for the Pb reaction, despite Fl being a much heavier atom. The reason for this unexpected result is the cancellation between the SE and VP effects. From Tables XV and XVI, the ratio of VP and SE is $\sim 1:-4.2$ for the Pb system, while it is $\sim 1:-1.8$ for the Fl system. Discussion along these lines is also found in Refs. 55,89. Finally, we note from Tables XV and XVI that the the atomic shift operator (ASHIFT), either using atomic ground state occupations or the effective atomic configuration in the molecules given in Table XIII, is not reliable for describing QED effects in the molecules.

VI. CONCLUSIONS

We have implemented effective QED potentials for relativistic molecular calculations by grafting code from the

numerical atomic code GRASP onto the DFT grid of DIRAC. A general disadvantage of numerical integration is higher computational cost than analytical evaluation, to the extent that such expressions are available, although the implementation itself is easier and considerable savings are achieved by the locality of the effective QED potentials.

We report several applications of the new code, mostly using the the molecular mean-field approximation Hamiltonian (X2Cmmf). We demonstrate (Table II) that with proper picture-change transformation of the effective QED potentials, our 2-component relativistic results reproduce very well 4-component reference data. On the other hand, this transformation is mandatory since picture-change errors are sizable.

We confirm that the discrepancy between the accurate calculations of Kirk Peterson and co-workers¹¹⁶ and experiment¹³⁰ is due to QED by directly calculating the ground-state rotational constants B_0 for the isotopomers investigated in the MW experiment. We then find that QED reduces the discrepancy of the corresponding substitution Au-C bond length r_s from 0.23 to 0.04 pm with respect to experiment.

For the rare-gas dimers Hg_2 and Rn_2 we find that QED increases bond lengths by about 0.15 pm. For the superheavy homologues the bond length increase is on the order of 0.30 pm; the effect on dissociation energies is quite small ($\sim 0.4\%$).

We have also investigated the effect of QED on the reaction energy of $\text{XH}_4 \longrightarrow \text{XH}_2 + \text{H}_2$, ($\text{X}=\text{Pb}, \text{Fl}$). From projection analysis we do find that there is a significant change of valence s population of the metals during the reaction, in line with the proposition of Dyall and co-workers.⁵³ Interestingly, though, we also find that in the tetrahydrides the valence s population essentially resides in bonding orbitals, but in non-bonding ones in the dihydrides. We find for the dissociation of lead tetrahydride that QED reduces the magnitude of the reaction energy by 0.32 kcal/mol (-1.27 %); for the superheavy homologue the magnitude of the QED effect is basically the same (0.35 kcal/mol). This possibly surprising observation is explained by the reduction of the (negative) SE/VP ratio with increasing nuclear charge.

For these metal hydrides, and also AuCN, we have also tried a simpler approach for the incorporation of QED effects in molecular calculations in the form of an atomic shift operator, but we find that this is not a reliable approach.

We would like to stress that our implementation of effective QED potential is general in the sense that they are available in all parts of the code. A natural continuation of our project will therefore be to explore the impact of these potentials on molecular properties probing electron density in the vicinity of nuclei, where the QED effects are generated. Our results so far indicate that QED effects may be more important than for the valence properties reported in the present work. For instance, the QED effect on the parity violation energy of

TABLE XII. Optimized equilibrium structures of Pb and Fl hydrides at the B3LYP/dyall.3zp level based on the DCG Hamiltonian. r_e and α_e refer to the X-H bond length (Å) and H-X-H angle (degree), respectively.

	PbH ₂		FlH ₂		PbH ₄	FlH ₄
	r_e	α_e	r_e	α_e	r_e	r_e
NR	1.879	90.83	2.017	90.85	1.816	1.959
DCG	1.845	91.18	1.920	93.35	1.756	1.825

TABLE XIII. Charge Q and electronic configurations of Pb and Fl atoms in the title compounds obtained by projection analysis at the B3LYP/dyall.3zp level.

	Q	atomic configuration					
PbH ₂	0.39	$5d_{3/2}^{4.00}$	$5d_{5/2}^{5.99}$	$6s_{1/2}^{1.86}$	$6p_{1/2}^{0.90}$	$6p_{3/2}^{0.86}$	
PbH ₄	0.66	$5d_{3/2}^{3.99}$	$5d_{5/2}^{5.98}$	$6s_{1/2}^{1.41}$	$6p_{1/2}^{0.85}$	$6p_{3/2}^{1.10}$	
FlH ₂	0.32	$6d_{3/2}^{3.99}$	$6d_{5/2}^{5.97}$	$7s_{1/2}^{1.91}$	$7p_{1/2}^{1.26}$	$7p_{3/2}^{0.55}$	
FlH ₄	0.46	$6d_{3/2}^{3.98}$	$6d_{5/2}^{5.94}$	$7s_{1/2}^{1.61}$	$7p_{1/2}^{1.17}$	$7p_{3/2}^{0.84}$	

H₂Po₂ is 2.38 %, although it depends on the choice of effective QED potentials.¹³⁹

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AUTHOR’S CONTRIBUTIONS

The overall project was conceived and supervised by TS. All programming and calculations were carried out by AS, whereas MS is the main contributor to the theory sections, notably the appendix.

DATA AVAILABILITY

The data that support the findings of this study are openly available in ZENODO at <https://doi.org/10.5281/zenodo.6874728>, see also Ref. 99. The DIRAC outputs contain information about the precise build of the corresponding executable and, most importantly, the git commit hash, which provides a unique identifier of the precise version of DIRAC generating the output, hence allowing reproduction of the results. We have also included the potential surface generated by Hill *et al.*¹¹⁶ for AuCN, kindly made available by Kirk Peterson.

CONFLICTS OF INTEREST

The authors have no conflicts to disclose.

Appendix A: Theory background

Since we hope to reach a wider audience than QED specialists, we provide in this Appendix a compact, yet accessible introduction (crash course) to QED that would otherwise have necessitated consulting disparate sources. More precisely, in this Appendix, we shall discuss the lowest-order BSQED corrections, and show how the effective potentials associated with these QED processes can be derived within the scattering matrix (\mathcal{S} -matrix) formalism. These effective potentials are to be used in practical relativistic calculations in order to account for the physics that is missing from the Dirac theory. In “conventional” QED one studies how the free-electron field interacts with the free quantized electromagnetic field and/or with an potential source (the scattering problem). On the other hand, BSQED theory studies the same problem but with electrons that are already interacting with some time-independent external field, i.e. their wavefunctions are solutions to the bound-state Dirac equation instead of the free one. This is known as the Furry picture of quantum mechanics; see for instance Refs. 65 and 59 (section 15g).

We shall use unbold symbols for four-quantities such as spacetime points (events): $x = (ct, \mathbf{x})$, here in contravariant coordinates, where \mathbf{x} is the spatial position vector, and the contravariant metric tensor $g^{\mu\nu} = \text{diag}(+1, -1, -1, -1)$. The gamma matrices are defined

TABLE XIV. Gross population obtained by projection analysis of the localized bonding orbitals in the title compounds at the B3LYP level based on the DCG Hamiltonian. $\langle \varepsilon \rangle$ refers to the expectation value with respect to the converged Kohn–Sham matrix (in E_h).

		$\langle \varepsilon \rangle$	X						H_i
			$5s_{1/2}$	$5d_{3/2}$	$5d_{5/2}$	$6s_{1/2}$	$6p_{1/2}$	$6p_{3/2}$	$1s_{1/2}$
PbH ₄	σ_{XH_i}	-0.4192		0.00	0.00	0.35	0.21	0.27	1.19
PbH ₂	σ_{XH_i}	-0.3384	0.00	0.00	0.00	0.03	0.37	0.37	1.24
	nb	-1.0182	0.13	0.00	0.00	1.62	0.14	0.11	-0.03
	$\langle \varepsilon \rangle$		$6s_{1/2}$	$6d_{3/2}$	$6d_{5/2}$	$7s_{1/2}$	$7p_{1/2}$	$7p_{3/2}$	$1s_{1/2}$
FIH ₄	σ_{XH_i}	-0.4524		0.00	0.01	0.40	0.27	0.20	1.11
FIH ₂	σ_{XH_i}	-0.3469	0.00	0.00	0.01	0.03	0.52	0.25	1.18
	nb	-1.2235	0.11	0.04	0.06	1.57	0.15	0.04	-0.03

TABLE XV. Reaction energy of $\text{PbH}_4 \longrightarrow \text{PbH}_2 + \text{H}_2$ (in kcal/mol). Δ DCG refers to the difference between DCG and NR. Other Δ refers to the difference from the DCG value.

	QED effect	reac. energy	Δ (kcal/mol)	Δ (%)
NR	none	16.47		
DCG	none	-8.99	-25.46	
	VP	-9.09	-0.10	0.41
	SE	-8.56	0.42	-1.67
	VP+SE	-8.66	0.32	-1.27
	VP+SE(ASHIFT ^a)	-8.97	0.02	-0.07
	VP+SE(ASHIFT ^b)	-9.06	-0.08	0.31

^a Occupation of atomic fragment was $6s^26p^2$

^b Using the occupations of Table XIII

through their anti-commutation relation

$$\gamma^\mu \gamma^\nu + \gamma^\nu \gamma^\mu = 2g^{\mu\nu} \mathbf{1}_4. \quad (\text{A1})$$

In Dirac basis they are represented by $\gamma^0 = \beta$ and $\boldsymbol{\gamma} = \beta\boldsymbol{\alpha}$. Following Lindgren⁷³ we shall complement the Dirac $\boldsymbol{\alpha}$ matrices with $\alpha^0 = \mathbf{1}_4$ to form a pseudo-4-vector. We finally note that we put hats (^) on quantities that contain creation/annihilation operators acting on occupation number states. Contrary to conventional QED sources, we have decided to express the formalism in full SI units.

1. Electron field operator

The electron field operator is given by the following annihilation expansion over all solutions of the Dirac equation

$$\hat{\Psi}(x) = \sum_i \psi_i(x) c_i, \text{ with } \psi_i(x) = \psi_i(\mathbf{x}) e^{-\frac{i}{\hbar} E_i t}. \quad (\text{A2})$$

In this expression, c_i is the electron annihilation operator obeying the fermionic algebra relations

$$\{c_i, c_j^\dagger\} = \delta_{ij}, \text{ and } \{c_i, c_j\} = \{c_i^\dagger, c_j^\dagger\} = 0, \quad (\text{A3})$$

and associated with the i -th spatial wavefunction $\psi_i(\mathbf{x})$ and energy-level E_i that solve the time-independent

Dirac equation in the presence of a time-independent external four-potential $A^e = (\varphi^e/c, \mathbf{A}^e)$

$$H_D \psi_i(\mathbf{x}) = E_i \psi_i(\mathbf{x}); \quad \text{with}$$

$$H_D = c\boldsymbol{\alpha} \cdot (-i\hbar\nabla + e\mathbf{A}^e(\mathbf{x})) - e\varphi^e(\mathbf{x}) + \beta m_e c^2. \quad (\text{A4})$$

The electron vacuum state is defined to be the one that vanishes after any annihilation:

$$c_i |0_e\rangle = 0, \quad \forall i. \quad (\text{A5})$$

In order to forbid the transition of positive-energy electrons to the negative-energy continuum by the Pauli exclusion principle, and obtain a stable atomic theory, Dirac¹⁴⁰ postulated that this continuum should be totally filled with electrons that are not observed (Dirac sea). This means that the vacuum state is redefined to be the state containing no positive-energy electrons and a fully occupied negative-energy electron sea. Dirac then argued that when a negative-energy electron absorbs enough energy ($E \geq 2m_e c^2$), it becomes real (observable), and leaves, for mass- and charge-conservation reasons, a positron behind (Dirac hole theory).¹⁴¹ This last reasoning allows one to define¹⁴²

$$c_i = a_i, \quad \text{for } E_i > 0,$$

$$c_i = b_i^\dagger, \quad \text{for } E_i < 0. \quad (\text{A6})$$

Here, operators a_i and b_i are introduced to distinguish between the particle (electron) and its hole (positron),

TABLE XVI. Reaction energy of $\text{FIH}_4 \longrightarrow \text{FIH}_2 + \text{H}_2$ (in kcal/mol). Δ DCG refers to the difference between DCG and NR. Other Δ refers to the difference from the DCG value.

	QED effect	reac. energy	Δ (kcal/mol)	Δ (%)
NR	none	9.52		
DCG	none	-60.02	-69.54	
	VP	-60.43	-0.41	0.59
	SE	-59.27	0.75	-1.08
	VP+SE	-59.67	0.35	-0.50
	VP+SE(ASHIFT ^a)	-60.10	-0.07	0.10
	VP+SE(ASHIFT ^b)	-60.23	-0.21	0.30

^a Occupation of the atomic fragment was $7s^27p^2$

^b Using the occupations of Table XIII

and the second line indicates that the annihilation of a negative-energy electron with c_i is equivalent to the creation of a (positive-energy) hole (positron) with b_i^\dagger . The electron field operator of Eq. (A2) can be written, with respect to these definitions, as

$$\hat{\Psi}(x) = \sum_{E_i > 0} \psi_i(x) a_i + \sum_{E_i < 0} \psi_i(x) b_i^\dagger. \quad (\text{A7})$$

Despite its experimental success in predicting the existence of the positron,¹⁴³ the hole theory (its physical implications) was, shortly after its introduction, abandoned. Many physicists including Pauli, Bohr, Weisskopf, Heisenberg and Majorana, opposed this theory, as clearly indicated in Refs. 144 (section 1.6), 145, 146 (section 4.4) and 147. This opposition came mainly from the following flaws of the Dirac hole theory: 1) the existence of a non-observable infinite negative energy and charge and 2) for massive boson systems, whose wavefunctions satisfy the Klein-Gordon equation, the Dirac argument would not hold, and the existence of these bosons is not justified. Modern quantum field theory reached the same mathematical expressions derived with respect to Dirac's hole theory, but provided a more symmetric picture between electrons and positrons, in which 1) one only sees electrons and positrons with positive energies, 2) the infinite negative-energy electron sea assumption is no longer necessary, and 3) operators such as the Hamiltonian, and charge are replaced by their normal-ordered forms. This physical interpretation leads to the modern definition of the vacuum state, that obeys

$$a_i |0_e\rangle = b_i |0_e\rangle = 0, \quad \forall i, \quad (\text{A8})$$

and contains zero positive-energy electrons and positrons. To get a wider and more detailed vision of the historical development of the quantum field theory, the reader may consult Weinberg in Ref. 148 (section 1.2 and chapter 5) and Ref. 149, Mehra in Ref. 150 (chapter 29), Schweber in Ref. 144, Kragh in Ref. 151 and Weisskopf in Ref. 152.

2. Photon field operator

The photon field operator is written as a sum over positive and negative plane-wave Fourier modes

$$\hat{A}_\mu(x) = \hat{A}_\mu^+(x) + \hat{A}_\mu^-(x), \quad (\text{A9})$$

$$\hat{A}_\mu^+(x) = \sum_{r=0}^3 \sum_{\mathbf{k}} N_{\mathbf{k}} a(\mathbf{k}, r) \varepsilon_\mu(\mathbf{k}, r) e^{-ik \cdot x}, \quad (\text{A10})$$

$$\hat{A}_\mu^-(x) = \sum_{r=0}^3 \sum_{\mathbf{k}} N_{\mathbf{k}} a^\dagger(\mathbf{k}, r) \varepsilon_\mu(\mathbf{k}, r) e^{+ik \cdot x}, \quad (\text{A11})$$

where $N_{\mathbf{k}} = \sqrt{\hbar/(2\epsilon_0\omega_{\mathbf{k}}V)}$ is the normalization constant, the zeroth component four-wave vector is $k_0 = |\mathbf{k}| = \omega_{\mathbf{k}}/c$, $\varepsilon_\mu(\mathbf{k}, r)$ are the four polarization vectors, and $a(\mathbf{k}, r)$ (and $a^\dagger(\mathbf{k}, r)$) is the annihilation (creation) operator that annihilates (creates) a photon with wave vector \mathbf{k} and polarization r , respectively (see Refs. 78 (eqs.(5.16a-c)) and 153 (section 8.4)). The choice of k_0 is imposed by the fact that the photon field operator must satisfy the Maxwell equation

$$\square \hat{A}_\mu(x) = 0, \quad \text{with} \quad \square = \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \nabla^2, \quad (\text{A12})$$

obtained after setting the Lorenz gauge condition ($\partial^\mu \hat{A}_\mu = 0$). This equation leads to the (massless) photon energy-momentum relation

$$k^2 = 0. \quad (\text{A13})$$

The boson creation and annihilation operators do satisfy the following bosonic commutation relations :

$$[a(\mathbf{k}, r), a^\dagger(\mathbf{k}', s)] = \delta_{rs} \zeta_r \delta_{\mathbf{k}, \mathbf{k}'} \quad (\text{A14})$$

$$[a^\dagger(\mathbf{k}, r), a^\dagger(\mathbf{k}', s)] = [a(\mathbf{k}, r), a(\mathbf{k}', s)] = 0. \quad (\text{A15})$$

Here, ζ_r is a function defined by the following relation

$$\zeta_r = \begin{cases} +1 & r = 0 \\ -1 & r = 1, 2, 3 \end{cases}, \quad (\text{A16})$$

and the polarization vectors satisfy the following completeness relation

$$\sum_{r=0}^3 \zeta_r \epsilon^\mu(\mathbf{k}, r) \epsilon^\nu(\mathbf{k}, r) = -g^{\mu\nu}. \quad (\text{A17})$$

Finally, we note that the photon vacuum state is defined to be the state that satisfies the following relation

$$\begin{aligned} a(\mathbf{k}, r) |0_p\rangle &= 0, \quad \forall \mathbf{k}, r \\ \rightarrow \hat{A}_\mu^+(x) |0_p\rangle &= 0, \quad \forall \mu, x. \end{aligned} \quad (\text{A18})$$

We shall now consider the interaction between the non-interacting electron and photon fields and show how one can derive QED corrections using perturbation theory.

3. Perturbation theory

As in conventional perturbation theory, one wants to get the eigensolutions of the following total Hamiltonian

$$\hat{H}_S = \hat{H}_S^0 + \lambda \hat{H}_S^1. \quad (\text{A19})$$

The zeroth-order Hamiltonian

$$\hat{H}_S^0 = \hat{H}_{\text{electron}}^0 + \hat{H}_{\text{photon}}^0, \quad (\text{A20})$$

represents the free electron and photon fields. The electronic part is given by a spatial integral over the normal-ordered Dirac Hamiltonian density

$$\begin{aligned} \hat{H}_{\text{electron}}^0 &= \int d^3x : \hat{\Psi}^\dagger(x) H_D(x) \hat{\Psi}(x) : \\ &= \sum_{E_i > 0} E_i a_i^\dagger a_i - \sum_{E_i < 0} E_i b_i^\dagger b_i, \end{aligned} \quad (\text{A21})$$

where in BSQED, H_D is the Dirac Hamiltonian in the presence of the external four-potential $A^e = (\varphi^e/c, \mathbf{A}^e)$, given in Eq. (A4), and where normal-ordering is indicated by double dots. The free photon Hamiltonian is written as an integral of the electromagnetic Hamiltonian density

$$\begin{aligned} \hat{H}_{\text{photon}}^0 &= \frac{1}{\mu_0} \int d^3x : \left[-(\partial^0 \hat{A}^\mu(x)) (\partial_0 \hat{A}_\mu(x)) \right. \\ &\quad \left. + \frac{1}{2} (\partial^\nu \hat{A}^\mu(x)) (\partial_\nu \hat{A}_\mu(x)) \right] : \\ &= \sum_{\mathbf{k}} \sum_{r=0}^3 \hbar \omega_{\mathbf{k}} \zeta_r a^\dagger(\mathbf{k}, r) a(\mathbf{k}, r). \end{aligned} \quad (\text{A22})$$

For further details and discussions on the photon Hamiltonian, the reader may consult Greiner and Reinhardt in Ref. 61 (section 7.3) and Mandl and Shaw in Ref. 78 (chapter 5).

The perturbation Hamiltonian \hat{H}_S^1 complicates the problem, and prevents us from obtaining eigensolutions of the full Hamiltonian \hat{H}_S . λ is a dimensionless parameter that can be varied between 0 and 1, and which keeps

track of the perturbation-order. This parameter is to be taken to 1 in order to account for the full perturbation by the end of the calculation. Notice that so far our Hamiltonians have an S subscript; this is made to indicate that they are in the Schrödinger picture of quantum mechanics. Assuming that we know the eigensolutions of the unperturbed time-independent problem equation

$$\begin{aligned} \hat{H}_S^0 |\Phi_0^\alpha\rangle_S &= E_0^\alpha |\Phi_0^\alpha\rangle_S; \\ |\Phi_0^\alpha(t)\rangle_S &= e^{-iE_0^\alpha t/\hbar} |\Phi_0^\alpha\rangle_S, \end{aligned} \quad (\text{A23})$$

where the α superscript labels solutions (states and associated energy-levels), the ultimate goal is to find eigensolutions of the perturbed problem

$$\begin{aligned} \hat{H}_S |\Phi^\alpha\rangle_S &= E^\alpha |\Phi^\alpha\rangle_S; \\ |\Phi^\alpha(t)\rangle_S &= e^{-iE^\alpha t/\hbar} |\Phi^\alpha\rangle_S. \end{aligned} \quad (\text{A24})$$

Gell-Mann and Low provided a closed form of the perturbed eigensolutions ($E^\alpha, |\Phi^\alpha\rangle_S$) in terms of the unperturbed ones ($E_0^\alpha, |\Phi_0^\alpha\rangle_S$) and the time-evolution operator,¹⁵⁴ see also Refs. 155 (pages 61-64) and 59 (section 11f.). A few years later, Sucher¹⁵⁶ provided an expression of the perturbation energy-shift that is more symmetric in time

$$\begin{aligned} \Delta E^\alpha &= E^\alpha - E_0^\alpha \\ &= \lim_{\lambda \rightarrow 1} \frac{i\epsilon\lambda}{2} \frac{\partial}{\partial \lambda} \log \left\langle \Phi_0^\alpha \left| \hat{\mathcal{S}}(\epsilon, \lambda) \right| \Phi_0^\alpha \right\rangle, \end{aligned} \quad (\text{A25})$$

where ϵ is an energy-parameter, to be shortly discussed. This energy-shift expression contains the \mathcal{S} -matrix operator that is defined to be the time-evolution operator that takes the interaction state from the very past $t = -\infty$ to the very future $t = +\infty$, and can be written as (see Dyson in Ref. 157 eq.(4))

$$\hat{\mathcal{S}}(\epsilon, \lambda) = \text{T} \left[\exp \left(\frac{\lambda}{i\hbar c} \int d^4x e^{-\frac{\epsilon}{\hbar} |t|} \hat{\mathcal{H}}_I(x) \right) \right]. \quad (\text{A26})$$

In this expression, T stands for time-ordering, i.e., it re-orders the inside operators such that those associated with earlier times act first. In the simplest case of two operators, the time-ordering operation is defined to be

$$\begin{aligned} \text{T} \left[\hat{A}(x_1) \hat{B}(x_2) \right] &\equiv \Theta(t_1 - t_2) \hat{A}(x_1) \hat{B}(x_2) \\ &\quad \pm \Theta(t_2 - t_1) \hat{B}(x_2) \hat{A}(x_1), \end{aligned} \quad (\text{A27})$$

where the minus sign applies when both operators \hat{A} and \hat{B} are of fermionic nature. Furthermore, the \mathcal{S} -matrix in Eq. (A26) is a functional of the interaction-Hamiltonian density $\hat{\mathcal{H}}_I(x)$, that is related to the interaction Hamiltonian $\hat{H}_I^1(t)$ by the following integral

$$\hat{H}_I^1(t) = \int d^3x \hat{\mathcal{H}}_I(x). \quad (\text{A28})$$

Recall that $\hat{H}_I^1(t)$ is the interaction-picture version of the Schrödinger-picture interaction-Hamiltonian \hat{H}_S^1 of Eq. (A19).

We shall note that the interaction density is multiplied by a damping factor $e^{-\frac{\epsilon}{\hbar}|t|}$, cf. Eq. (A26), where ϵ is a small positive quantity that has energy dimensions. This term is known as the “adiabatic switch” that al-

lows the interpolation between the perturbed and unperturbed problems ($t = 0, \pm\infty$), and was first introduced by Gell-Mann and Low in Ref. 154 (Appendix A) while extending the \mathcal{S} -matrix formalism to cover the bound-electron problem (see also Ref. 158 (section 1.3)). The scattering matrix of Eq. (A26) may be expanded in powers of the perturbation parameter λ as

$$\begin{aligned}\hat{\mathcal{S}}(\epsilon, \lambda) &= \sum_{n=0}^{\infty} \hat{\mathcal{S}}^{(n)}(\epsilon, \lambda) \\ \hat{\mathcal{S}}^{(n)}(\epsilon, \lambda) &= \frac{1}{n!} \left(\frac{\lambda}{i\hbar c} \right)^n \int d^4x_1 \dots \int d^4x_n e^{-\frac{\epsilon}{\hbar}(|t_1| + \dots + |t_n|)} \text{T} \left[\hat{\mathcal{H}}_I(x_1) \dots \hat{\mathcal{H}}_I(x_n) \right].\end{aligned}\tag{A29}$$

This form of the $\hat{\mathcal{S}}$ -matrix expansion is known as the Dyson series, and originated from the works of Dyson^{157,159} and Schwinger.¹⁶⁰ Detailed derivations of the time-evolution and $\hat{\mathcal{S}}$ -matrix operators can be found in Fetter and Walecka Ref. 155 (pages 54-58), Mandl and Shaw Ref. 78 (section 6.2), as well as Bjorken and Drell Ref. 161 (section 17.2). In QED, the (perturbation) interaction-Hamiltonian density is given by

$$\begin{aligned}\hat{\mathcal{H}}_I(x) &= \hat{J}_\mu(x) \hat{A}^\mu(x), \\ \text{with } \hat{J}_\mu(x) &= -ec\bar{\hat{\Psi}}(x)\gamma_\mu\hat{\Psi}(x),\end{aligned}\tag{A30}$$

that explicitly couples the quantized electron-current field operator \hat{J}_μ to the photon field operator \hat{A}^μ . Some authors, starting with Schwinger in Ref. 160 (Eq.(1.14)), use the symmetrized form

$$\hat{J}_\mu(x) = -\frac{ec}{2}[\bar{\hat{\Psi}}_\alpha(x), \hat{\Psi}_\beta(x)][\gamma_\mu]_{\alpha\beta},\tag{A31}$$

for the electron-current field operator, but the two forms are equivalent under time-ordering (see Eq. (29) of Ref. 162). We recall that the electron and photon field operators are given in Eqs. (A7) and (A9), respectively. We note that the Dirac field operator with a bar on the top represents the Dirac adjoint field: $\bar{\hat{\Psi}}(x) = \hat{\Psi}^\dagger(x)\gamma^0$. At this point, the reader can see, from the last two equations, that the QED theory treats the electron-photon field (interaction) coupling perturbatively, in powers of the elementary charge e .

We next consider how to expand the time-ordered product of the $\hat{\mathcal{S}}$ -matrix, and assign each of the obtained normal-ordered terms to a specific Feynman diagram.

4. Wick's theorem, field contractions and propagators

Wick's theorem⁶² allows writing the time-ordered products of Eq. (A29) in terms of normal-ordered products of all possible contractions, as given in the following

equation

$$\begin{aligned}&\text{T} \left[\hat{\mathcal{O}}(x_1) \hat{\mathcal{O}}(x_2) \hat{\mathcal{O}}(x_3) \hat{\mathcal{O}}(x_4) \dots \right] \\ &= : \hat{\mathcal{O}}(x_1) \hat{\mathcal{O}}(x_2) \hat{\mathcal{O}}(x_3) \hat{\mathcal{O}}(x_4) \dots : \\ &+ : \overbrace{\hat{\mathcal{O}}(x_1) \hat{\mathcal{O}}(x_2)} \hat{\mathcal{O}}(x_3) \hat{\mathcal{O}}(x_4) \dots : + \dots \\ &+ : \overbrace{\hat{\mathcal{O}}(x_1) \hat{\mathcal{O}}(x_2) \hat{\mathcal{O}}(x_3)} \hat{\mathcal{O}}(x_4) \dots : + \dots\end{aligned}\tag{A32}$$

Contracted operators are moved next to each other, noting that under normal-ordering (fermionic) bosonic operators can be permuted as if they (anti)commuted. A contraction is represented by a line that links two operators and is defined to be the vacuum expectation value of the time-ordered product

$$\overbrace{\hat{\mathcal{O}}(x_1) \hat{\mathcal{O}}(x_2)} \equiv \langle 0 | \text{T} [\hat{\mathcal{O}}(x_1) \hat{\mathcal{O}}(x_2)] | 0 \rangle.\tag{A33}$$

Since our QED interaction-Hamiltonian density contains electron and photon operators, the time-ordered product in our \mathcal{S} -matrix of Eq. (A29) will be expanded with two types of contractions: electronic and photonic. The contraction of two electron field operators (of Eq. (A7)) components α and β is defined with respect to the last formula by

$$\begin{aligned}\overbrace{\hat{\Psi}_\alpha(x_1) \hat{\Psi}_\beta(x_2)} &\equiv \langle 0_e | \text{T} [\hat{\Psi}_\alpha(x_1) \hat{\Psi}_\beta(x_2)] | 0_e \rangle \\ &= i\hbar [S_{A^e}^F(x_1, x_2)]_{\alpha\beta},\end{aligned}\tag{A34}$$

where $[S_{A^e}^F(x, y)]_{\alpha\beta}$ is the α, β matrix component of the Feynman electron propagator, which in turn satisfies the Dirac propagator equation

$$[\gamma^\mu (i\hbar\partial_\mu + eA_\mu^e(\mathbf{x})) - m_e c] S_{A^e}^F(x, y) = \mathbb{1}_4 \delta(x - y);\tag{A35}$$

cf. Eq. (A4). Furthermore, one can show that the following identities

$$\begin{aligned}\overbrace{\hat{\Psi}_\beta(x_2) \hat{\Psi}_\alpha(x_1)} &= -\overbrace{\hat{\Psi}_\alpha(x_1) \hat{\Psi}_\beta(x_2)}, \\ \overbrace{\hat{\Psi}_\alpha(x_1) \hat{\Psi}_\beta(x_2)} &= \overbrace{\hat{\Psi}_\alpha(x_1) \hat{\Psi}_\beta(x_2)} = 0,\end{aligned}\tag{A36}$$

hold. These relations show that the only non-zero contractions are between electron field operators and their adjoints. The free Feynman electron propagator $S_0^F(x, y)$, corresponding to the case $A_\mu(\mathbf{x}) = 0_\mu$, can be written as

$$S_0^F(y, x) = \lim_{\epsilon \rightarrow 0} \int \frac{d^4 p}{(2\pi\hbar)^4} e^{-\frac{i}{\hbar} p \cdot (y-x)} S_0^F(p), \quad (\text{A37})$$

$$\text{with } S_0^F(p) = \frac{\gamma^\mu p_\mu + m_e c}{p^2 - m_e^2 c^2 + i\epsilon},$$

where $S_0^F(p)$ is the Fourier transformed free-electron propagator. The role of the small positive number ϵ is to shift energy-poles (at the energy-momentum relation) with respect to the Feynman prescription.

Similarly, the contraction of two photon operators (of Eq. (A9)) is defined by the following expression

$$\overline{\hat{A}_\mu(x_1) \hat{A}_\nu(x_2)} \equiv \langle 0_p | T[\hat{A}_\mu(x_1) \hat{A}_\nu(x_2)] | 0_p \rangle \quad (\text{A38})$$

$$= i\hbar D_{\mu\nu}^F(x_1, x_2),$$

where $D_{\mu\nu}^F(x, y)$ is the photon propagator in the Feynman gauge, is given by the following expression

$$D_{\mu\nu}^F(x, y) = \lim_{\epsilon \rightarrow 0} \int \frac{d^4 p}{(2\pi\hbar)^4} e^{-\frac{i}{\hbar} p \cdot (x-y)} D_{\mu\nu}^F(p), \quad (\text{A39})$$

$$\text{with } D_{\mu\nu}^F(p) = g_{\mu\nu} D^F(p) = -\frac{\hbar^2}{c\epsilon_0} \frac{g_{\mu\nu}}{p^2 + i\epsilon}$$

and satisfies the Maxwell Green's-type equation

$$\partial_\sigma \partial^\sigma D_{\nu\theta}^F(x, y) = \frac{g_{\nu\theta}}{\epsilon_0 c} \delta(x - y). \quad (\text{A40})$$

This equation is obtained after imposing the Lorenz gauge condition, otherwise this propagator will not be invertible; see Schwartz in Ref. 163 (section 8.5). We should finally note that the F superscript on both propagators is added to indicate that these are Feynman propagators. This means that when writing the propagators as Fourier transforms, the energy-integrals are to be taken along the Feynman contour. Different choices of paths (contours) lead to different propagators (retarded and advanced), but they all satisfy the corresponding Dirac and Maxwell equations.

5. Bound electron propagator expansion

The bound Feynman propagator $S_{A^e}^F(x_2, x_1)$ of Eq. (A34) can be expanded in powers of the external potential as (Refs. 86 eq.(2-119) and 164 eq.(16)):

$$S_{A^e}^F(x_2, x_1) = S_0^F(x_2, x_1) - \int d^4 x_3 S_0^F(x_2, x_3) e A_\mu^e(\mathbf{x}_3) \gamma^\mu S_0^F(x_3, x_1) + \dots, \quad (\text{A41})$$

and written in terms of the free-electron propagator $S_0^F(x_2, x_1)$ of Eq. (A37). It is worth noting that the bound Feynman propagator can be related to the bound Dirac Green's function G_{A^e} by the relation of Ref. 162 (eq.(32)):

$$S_{A^e}^F(x_2, x_1) = \frac{1}{i\hbar} \frac{1}{2\pi i} \int_{C_F} dz G_{A^e}(\mathbf{x}_2, \mathbf{x}_1; z) \gamma^0 e^{-\frac{i}{\hbar} z(t_2 - t_1)}, \quad (\text{A42})$$

This Green's function satisfies

$$[H_D(\mathbf{x}_2) - z] G_{A^e}(\mathbf{x}_2, \mathbf{x}_1; z) = \mathbb{1}_4 \delta(\mathbf{x}_2 - \mathbf{x}_1); \quad (\text{A43})$$

cf. Eqs. (A4) and (A35). Using Eqs. (A41) and (A42), and integrating over time variables, one obtains the potential expansion associated with the Green's function

$$G_{A^e}(\mathbf{x}_2, \mathbf{x}_1; z) = G_0(\mathbf{x}_2, \mathbf{x}_1; z) + ec \int d^3 x_3 G_0(\mathbf{x}_2, \mathbf{x}_3; z) A_\mu^e(\mathbf{x}_3) \alpha^\mu G_0(\mathbf{x}_3, \mathbf{x}_1; z) + \dots \quad (\text{A44})$$

where the free Dirac Green's function is given by: $G_0 = \lim_{A^e \rightarrow 0} G_{A^e}$. These two expansions are known as the potential expansion, where consecutive terms are known as the zero- one- and many-potential terms. The main utility of this expansion is that it allows the isolation of ultraviolet divergent integrals encountered when evaluating loop integrals, as done by Baranger *et al.*,⁷² and later by many authors working within BSQED theory.

6. No-photon BSQED energy-shifts

Using the \mathcal{S} -matrix expansion of Eq. (A29), one can expand Sucher's energy-shift expression of Eq. (A25) in powers of the interaction-Hamiltonian density and write, following Mohr in Ref. 165 (eqs.(18) and (31)),

$$\Delta E^\alpha = \lim_{\substack{\epsilon \rightarrow 0 \\ \lambda \rightarrow 1}} \frac{i\epsilon\lambda}{2} \left[\left\langle \Phi_0^\alpha \left| \hat{\mathcal{S}}^{(1)}(\epsilon, \lambda) \right| \Phi_0^\alpha \right\rangle + 2 \left\langle \Phi_0^\alpha \left| \hat{\mathcal{S}}^{(2)}(\epsilon, \lambda) \right| \Phi_0^\alpha \right\rangle - \left\langle \Phi_0^\alpha \left| \hat{\mathcal{S}}^{(1)}(\epsilon, \lambda) \right| \Phi_0^\alpha \right\rangle^2 + \mathcal{O}(\lambda^3) \right], \quad (\text{A45})$$

where $\hat{\mathcal{S}}^{(n)}$ is given in Eq. (A29). We shall now consider a system of n electrons and zero photons (photon vacuum), represented by the following electron-photon state, labeled by α :

$$|\Phi_0^\alpha\rangle = |n_e^\alpha, 0_p^\alpha\rangle. \quad (\text{A46})$$

We remind the reader that the electron field operators entering in our expressions describe non-interacting electrons, in the presence of an external potential, as also seen in the zeroth-order electron Hamiltonian of Eq. (A21). As already pointed out in Section II, the electron-electron interaction arises from terms describing exchange of virtual photons between bound electrons.

Since the photon state is chosen to be the vacuum one, this means that any string of photon operators that is not fully contracted, will vanish under the photon vacuum expectation value of Eq. (A25). Following this reasoning one concludes that the first non-vanishing QED correction comes from the second-order $\hat{\mathcal{S}}^{(2)}$ -matrix

$$\begin{aligned} \hat{\mathcal{S}}^{(2)}(\epsilon, \lambda) &= -\frac{\lambda^2}{2\hbar^2 c^2} \int d^4 x_1 \int d^4 x_2 \\ &\times e^{-\frac{\epsilon}{\hbar}(|t_1|+|t_2|)} \text{T} \left[\hat{\mathcal{H}}_I(x_1) \hat{\mathcal{H}}_I(x_2) \right]. \end{aligned} \quad (\text{A47})$$

Using Wick's theorem, we expand the electron and photon time-ordered products, and replace operator contractions by corresponding propagators, following the contraction definitions of Eqs. (A34) and (A38). Fur-

thermore, using the symmetry properties of the photon propagator of Eq. (A39):

$$D_{\mu\nu}^F(x, y) = D_{\mu\nu}^F(y, x) = D_{\nu\mu}^F(x, y), \quad (\text{A48})$$

the second-order \mathcal{S} -matrix of Eq. (A47) can be shown to reduce to the following expression

$$\begin{aligned} \hat{\mathcal{S}}^{(2)}(\epsilon, \lambda) &= -\frac{\lambda^2 e^2}{2\hbar^2} \int d^4 x_1 \int d^4 x_2 \\ &\times e^{-\frac{\epsilon}{\hbar}(|t_1|+|t_2|)} \hat{F}(x_1, x_2), \end{aligned} \quad (\text{A49})$$

where the operator $\hat{F}(x_1, x_2)$ contains the following five QED corrections to the non-interacting problem:

$$\begin{aligned} &\hat{F}(x_1, x_2) \\ &= i\hbar D_{\mu_1\mu_2}^F(x_1, x_2) : \tilde{\Psi}(x_1)\gamma^{\mu_1}\hat{\Psi}(x_1)\tilde{\Psi}(x_2)\gamma^{\mu_2}\hat{\Psi}(x_2) : \quad \text{SP} \\ &+ 2\hbar^2 D_{\mu_1\mu_2}^F(x_1, x_2) \text{Tr}[S_{A^e}^F(x_2, x_2)\gamma^{\mu_2}] : \tilde{\Psi}(x_1)\gamma^{\mu_1}\hat{\Psi}(x_1) : \quad \text{VP} \\ &- 2\hbar^2 D_{\mu_1\mu_2}^F(x_1, x_2) : \tilde{\Psi}(x_1)\gamma^{\mu_1}S_{A^e}^F(x_1, x_2)\gamma^{\mu_2}\hat{\Psi}(x_2) : \quad \text{SE} \\ &- i\hbar^3 D_{\mu_1\mu_2}^F(x_1, x_2) \text{Tr}[S_{A^e}^F(x_1, x_1)\gamma^{\mu_1}] \text{Tr}[S_{A^e}^F(x_2, x_2)\gamma^{\mu_2}] \quad \text{D1} \\ &+ i\hbar^3 D_{\mu_1\mu_2}^F(x_1, x_2) \text{Tr}[S_{A^e}^F(x_2, x_1)\gamma^{\mu_1}S_{A^e}^F(x_1, x_2)\gamma^{\mu_2}] \quad \text{D2} \end{aligned} \quad (\text{A50})$$

Finally, using Sucher's energy expression of Eq. (A45), the second-order energy-shift becomes

$$\begin{aligned} \Delta E^{\alpha,2} &= \lim_{\substack{\epsilon \rightarrow 0 \\ \lambda \rightarrow 1}} i\epsilon\lambda \langle \Phi_0^\alpha | \hat{\mathcal{S}}^{(2)}(\epsilon, \lambda) | \Phi_0^\alpha \rangle \\ &= \Delta E_{\text{SP}}^{\alpha,2} + \Delta E_{\text{VP}}^{\alpha,2} + \Delta E_{\text{SE}}^{\alpha,2} \\ &+ \Delta E_{\text{D1}}^{\alpha,2} + \Delta E_{\text{D2}}^{\alpha,2}. \end{aligned} \quad (\text{A51})$$

Each of these terms will be discussed in the next sections, and is represented by a Feynman diagram in Fig. 1. The elements of these diagrams are the following:

1. Double external lines represent bound-electrons, i.e., with wavefunctions and energies satisfying the interacting Dirac equation of Eq. (A4), in the presence of a classical time-independent external potential $A^e(\mathbf{x})$.
2. Double internal lines represent a virtual bound-electron propagation between the two vertices $S_{A^e}^F(x_2, x_1)$ and arise from a single contraction of two electron field operators.
3. Internal wiggly-lines connecting two vertices represent propagations of virtual-photons $D_{\mu_2\mu_1}^F(x_2, x_1)$ and come from a single contraction of two photon field operators.

The last two contributions from Eq. (A50) correspond to fully contracted products, and they are thus free of creation and annihilation operators. This means that their corresponding energy-shifts $\Delta E_{\text{D1}}^{\alpha,2}$ and $\Delta E_{\text{D2}}^{\alpha,2}$ are state-independent and hence do not contribute to energy-differences. They are therefore discarded from further consideration; see for instance Mohr in Ref. 165. On the other hand, the first three contributions correspond to partially contracted products, associated with the following physical processes:

a. SP: Single-photon exchange

This process, coming from the SP term in Eq. (A50) and represented in Fig. 1a, describes electron-electron interaction in its lowest-order, where an electron feels the existence of the other electron through the exchange of a single virtual-photon. After integrating over times t_1 and t_2 in Eq. (A49), and taking limits in Eq. (A51), this correction yields an instantaneous direct interaction-term, in addition to a retarded exchange interaction-term, analogous to the direct and exchange terms in the Hartree-Fock theory:

$$\begin{aligned}
 \Delta E_{\text{SP}}^{\alpha,2} &= \frac{e^2}{2} \sum_{i,j} \int d^3x_1 \int d^3x_2 \bar{\psi}_i(\mathbf{x}_1) \gamma^\mu \psi_i(\mathbf{x}_1) \frac{1}{4\pi\epsilon_0 |\mathbf{x}_1 - \mathbf{x}_2|} \bar{\psi}_j(\mathbf{x}_2) \gamma_\mu \psi_j(\mathbf{x}_2) \quad \text{Direct} \\
 &- \frac{e^2}{2} \sum_{i,j} \int d^3x_1 \int d^3x_2 \bar{\psi}_i(\mathbf{x}_1) \gamma^\mu \psi_j(\mathbf{x}_1) \frac{e^{+\frac{i}{\hbar}|E_i - E_j||\mathbf{x}_1 - \mathbf{x}_2|}}{4\pi\epsilon_0 |\mathbf{x}_1 - \mathbf{x}_2|} \bar{\psi}_j(\mathbf{x}_2) \gamma_\mu \psi_i(\mathbf{x}_2) \quad \text{Exchange}
 \end{aligned} \tag{A52}$$

as noted by Mohr in Ref. 166 (section IV). Notice that for $\mu = 0$, and $\mu = 1, 2, 3$ these integrals account for the Coulomb and Gaunt interaction, respectively. On the other hand, if we used the Coulomb gauge photon propagator instead of the Feynman one, we would get the retarded Breit interaction, as noted by Lindgren in Ref. 167 (page 262).

b. VP: Vacuum polarization

This process, presented in Fig. 1b, accounts for the instantaneous interaction of a bound-electron with the electron-positron pair cloud, polarized by the presence of a classical potential source. After plugging the VP term of Eq. (A50) in the second-order scattering matrix expression, one can use Sucher's formula of Eq. (A45) to write the energy-shift associated with the vacuum polarization process as Ref. 59 (chapter 15 eq.(205)):

$$\Delta E_{\text{VP}}^\alpha = -e \sum_i \int d^3x_1 \bar{\psi}_i(\mathbf{x}_1) \gamma_\mu \psi_i(\mathbf{x}_1) \varphi_{\text{VP}}^\mu(\mathbf{x}_1). \tag{A53}$$

We note that the vacuum polarization effect is local, i.e., it can be written as an expectation value of a local vacuum polarization four-potential

$$\varphi_{\text{VP}}^\mu(\mathbf{x}_1) = ie\hbar \int d^3x_2 \frac{\text{Tr}[\gamma^\mu S_{A^e}^F(x_2, x_2)]}{4\pi\epsilon_0 |\mathbf{x}_1 - \mathbf{x}_2|}. \tag{A54}$$

The energy expression of Eq. (A53) (as well as the last potential) is divergent due to the fact that

$$\lim_{x_1 \rightarrow x_2} S_{A^e}^F(x_2, x_1) = \infty, \tag{A55}$$

as mentioned in Ref. 168. The isolation of the divergent terms in this expression can be done by expanding the propagator inside the trace using Eq. (A41), and write the energy as

$$\Delta E_{\text{VP}}^{\alpha,2} = \Delta E_{\text{VP},0}^{\alpha,2} + \Delta E_{\text{VP},1}^{\alpha,2} + \Delta E_{\text{VP},2}^{\alpha,2} + \dots \tag{A56}$$

where $\Delta E_{\text{VP},i}^{\alpha,2}$ represents the term that corresponds to an i number of interactions with the external potential $(Z\alpha)^i$. The first four terms are presented in Figs. 3a to 3d. Notice that the double-line loop is replaced by a single-line one. This is made to indicate that the propagators between these vertices are the free ones S_0^F , instead of the bound-ones $S_{A^e}^F$. Using Furry's theorem,¹⁶⁹

that is based on a charge conjugation symmetry argument, one can show that any diagram containing a free-electron loop with an odd number of vertices does not contribute. This means that the above energy expression reduces to

$$\Delta E_{\text{VP}}^{\alpha,2} = \Delta E_{\text{VP},1}^{\alpha,2} + \Delta E_{\text{VP},3}^{\alpha,2} + \dots \tag{A57}$$

A naive estimation of the degree of divergence of a QED integral can be done by calculating the superficial degree of divergence S that simply counts overall momentum powers of the integral in question (in momentum-space):

$$S \equiv 4 - N_e - 2N_p, \tag{A58}$$

where 4 are the spacetime dimensions and N_e and N_p are the number of electron and photon propagators, respectively, in the loop in question; see for instance Refs. 60 (section 10.1) and 86 (sections 7-1-4 and 8-1-3). The integral is said to be superficially divergent if $S \geq 0$. The possible cases are:

$$\begin{aligned}
 S \leq 0 & \quad \text{convergence} \\
 S = 0 & \quad \text{logarithmic divergence} \\
 S = 1 & \quad \text{linear divergence} \\
 S = 2 & \quad \text{quadratic divergence}
 \end{aligned} \tag{A59}$$

This naive estimation usually overestimates the effective (divergence), that we shall call E , and this can be seen after further analysis of the integral in question. As a consequence, some superficially divergent integral can be effectively less divergent, or hopefully convergent. In the next Table XVII, we list the superficial and effective divergences of the first vacuum polarization terms. The reader should notice that with higher-order terms, more propagators are included in the momentum-space integral, meaning that more denominator powers are added, and as a consequence, the integral becomes less divergent.

We shall now focus on the first non-vanishing vacuum polarization contribution $\Delta E_{\text{VP},1}^\alpha$. As seen in Table XVII, this term is of superficial quadratic divergence, but it is, effectively, only logarithmic. This can be shown using the Ward identity, as mentioned by Peskin and Schroeder in Ref. 60 (section 7.5). After the employment of regularization, followed by renormalization (discussed in Sec. A 6d), one may extract the physical contribution out of the divergent $\Delta E_{\text{VP},1}^{\alpha,2}$. In the case where the Hamiltonian is invariant under time-reversal symmetry,

Terms	S	E
$\alpha (Z\alpha)^1$	2	0
$\alpha (Z\alpha)^3$	0	< 0
$\alpha (Z\alpha)^5$	< 0	< 0
\vdots	\vdots	\vdots

TABLE XVII. Superficial and effective degrees of divergence for the bound-state vacuum polarization contributions.

i.e., if the external vector potential $\mathbf{A}^e(\mathbf{x})$ vanishes, cf. Greiner in Ref. 170 (eqs. (12.52-53)), then only the time-component potential $\varphi_{\text{VP},1}^0$ survives, and one obtains the Uehling potential,²³ given in Eq. (8).

For detailed discussions and derivations of the one-potential bound-state vacuum polarization correction, the reader may consult the calculation of Greiner and Reinhardt in Ref. 85 (section 5.2) where the authors used Pauli–Villars regularization, in addition to Peskin and Schroeder in Ref. 60 (section 7.5) who used dimensional-regularization to treat the occurring divergences; see also Mandl and Shaw in Ref. 78 (section 10.4), in addition to Schwartz in Ref. 163 (section 16.2.2). Contrary to the conventional momentum-space approach to evaluate QED corrections, Indelicato and Mohr in Ref. 168 (section B) considered the vacuum polarization problem in coordinate space, and derived the physical Uehling contribution using coordinate-space Pauli–Villars regularization.

The second non-vanishing vacuum polarization effect, associated with $\Delta E_{\text{VP},3}^{\alpha,2}$ and presented in Fig. 3d, is known as the Wichmann–Kroll effect.⁷⁰ As seen in Table XVII and noted by Gyulassy,¹⁷¹ this contribution is free of divergences. Wichmann and Kroll calculated the effective potential associated with the $\Delta E_{\text{VP},3}^{\alpha,2}$ correction in Laplace space. On the other hand, in Ref. 172 (section 4), Blomqvist has evaluated the inverse Laplace-transform, and obtained the real-space potential expression for a point nuclear charge distribution. The last reference presents a relatively complex analytical expression for this $\alpha (Z\alpha)^3$ potential, and this motivated Fainshtein *et al.*¹⁷³ to provide an approximation that facilitates the numerical computation, yet conserving precision.

We finally note that in the fourth-order BSQED correction, one finds the Källén–Sabry potentials⁷¹ of order $\alpha^2 (Z\alpha)$ that can be obtained by expanding the bound propagators of Ref. 162 (fig. 25 diagrams b,c VPVP). In order to make this momentum-space potential usable in practical calculations, in Ref. 172 (section 3) Blomqvist derived its real-space version for a point nucleus distribution, whereas Wayne Fullerton and Rinker generalized this result to account for an extended nuclear charge distribution; see Ref. 66 (eq. (9)). We finally note that the latter authors provided a good approximation of the corresponding potential, in order to render the numerical evaluation more practical.

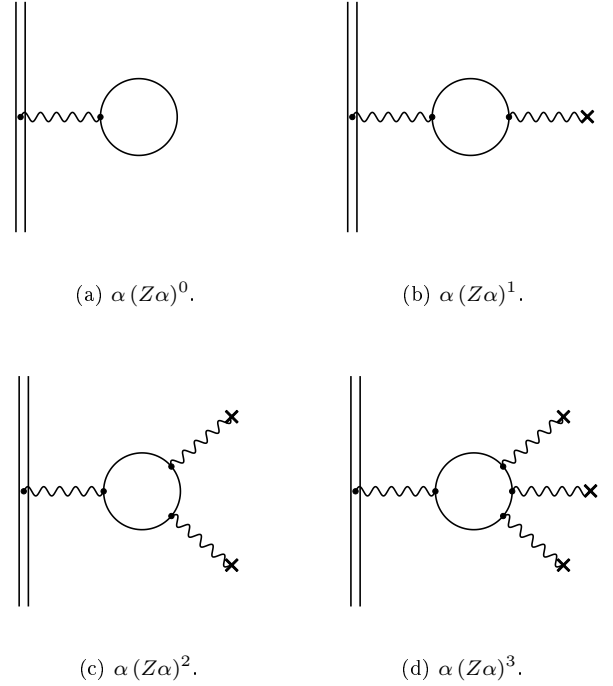


FIG. 3. First four bound-state vacuum polarization processes, obtained after expanding the bound propagator in powers of the external potential. A wiggly line ending with a cross \times indicates an interaction with the external field.

c. SE: Self-energy

The self-energy process, presented in Fig. 1c, is the dominant radiative QED correction in electronic atoms, as seen in the work of Johnson and Soff of Ref. 123 (fig.2). This process describes the interaction of the bound-electron with itself, by emitting and absorbing a virtual-photon. The first calculation for this correction was made in 1947 by Hans Bethe in a purely non-relativistic framework,¹⁷⁴ where he used a renormalization technique (by subtracting the free self-energy) to render the integral less divergent, and introduced a reasonable virtual-photon energy cutoff at $E = m_e c^2$. This simple calculation gave hope in digging for the physical Lamb shift in the frustrating non-physical divergences in the QED theory. Using Sucher’s energy formula of Eq. (A45), the SE term of Eq. (A50) leads to the following energy-shift

$$\Delta E_{\text{SE}}^{\alpha,2} = -e \sum_i \int d^3 x_1 \int d^3 x_2 \times \psi_i^\dagger(\mathbf{x}_2) \varphi_{\text{SE}}(\mathbf{x}_2, \mathbf{x}_1; E_i) \psi_i(\mathbf{x}_1) \quad (\text{A60})$$

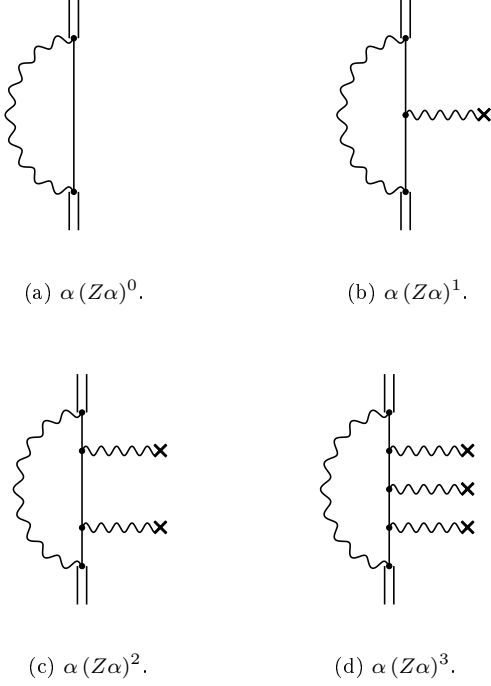


FIG. 4. First four bound-state self-energy processes, obtained after expanding the bound propagator in powers of the external potential.

where the self-energy potential is given by

$$\begin{aligned} \varphi_{\text{SE}}(\mathbf{x}_2, \mathbf{x}_1; E_i) &= -\frac{e}{2\pi i} \int_{C_F} dz \alpha^\mu G_{A^e}(\mathbf{x}_2, \mathbf{x}_1; z) \alpha_\mu \\ &\times \frac{\exp\left(+\frac{i}{\hbar} |\mathbf{x}_1 - \mathbf{x}_2| \sqrt{(z - E_i)^2 / c^2 + i\epsilon}\right)}{4\pi\epsilon_0 |\mathbf{x}_1 - \mathbf{x}_2|}. \end{aligned} \quad (\text{A61})$$

The next step is to transform these two real-space integral \mathcal{S} -matrices into Fourier-space ones. We first use the electron and photon propagators of Eqs. (A37) and (A39) and write electron field operators and the external

Similar forms of this equation are provided by Schwinger in Ref. 59 (eq.(205)) and Mohr in Ref. 175 (eq.(2.6)). Notice at this point that unlike the vacuum polarization case, the self-energy is a non-local effect, as seen from Eq. (A60), and this is the reason behind the complexity of its analytical and numerical evaluation. As in the vacuum polarization case, the self-energy potential is a divergent quantity and needs to be regularized. In order to isolate divergent terms, one can use the Green's function (propagator) expansion of Eq. (A44) and write the total energy-shift as

$$\Delta E_{\text{SE}}^{\alpha,2} = \Delta E_{\text{SE},0}^{\alpha,2} + \Delta E_{\text{SE},1}^{\alpha,2} + \Delta E_{\text{SE},2}^{\alpha,2} + \dots, \quad (\text{A62})$$

where $\Delta E_{\text{SE},i}^{\alpha,2}$ represents the process in which the internal electron interacts i times with the external potential, it is thus associated with the $\alpha(Z\alpha)^i$ order. The first four terms of the last expansion are represented in Figs. 4a to 4d. The zero- and one-potential terms: $\Delta E_{\text{SE},0}^{\alpha,2}$ and $\Delta E_{\text{SE},1}^{\alpha,2}$ are known as the (free) self-energy and the vertex-correction processes. These two contributions are

Terms	S	E
$\alpha(Z\alpha)^0$	1	0
$\alpha(Z\alpha)^1$	0	0
$\alpha(Z\alpha)^2$	< 0	< 0
\vdots	\vdots	\vdots

TABLE XVIII. Superficial and effective degrees of divergence for the bound-state self-energy contributions.

logarithmically divergent (in momentum-space), while all higher-order ones are convergent, as presented in Table XVIII. A coordinate-space treatment of these quantities has been provided by Indelicato and Mohr.^{75,168,175} Using the bound propagator expansion of Eq. (A41), one obtains the scattering matrices associated with these two processes:

$$\hat{\mathcal{S}}_{\text{SE},0}^{(2)}(\epsilon, \lambda) = \lambda^2 e^2 \int d^4 x_1 \int d^4 x_2 e^{-\frac{\epsilon}{\hbar}(|t_1|+|t_2|)} D_{\mu_1 \mu_2}^F(x_1, x_2) : \bar{\Psi}(x_1) \gamma^{\mu_2} S_0^F(x_1, x_2) \gamma^{\mu_1} \hat{\Psi}(x_2) : \quad (\text{A63})$$

$$\begin{aligned} \hat{\mathcal{S}}_{\text{SE},1}^{(2)}(\epsilon, \lambda) &= -\lambda^2 e^2 \int d^4 x_1 \int d^4 x_2 e^{-\frac{\epsilon}{\hbar}(|t_1|+|t_2|)} D_{\mu_1 \mu_2}^F(x_1, x_2) \\ &\times : \bar{\Psi}(x_1) \gamma^{\mu_1} \int d^4 x_3 S_0^F(x_1, x_3) e A_\mu^e(\mathbf{x}_3) \gamma^\mu S_0^F(x_3, x_2) \gamma^{\mu_2} \hat{\Psi}(x_2) : \end{aligned} \quad (\text{A64})$$

(classical) potential in their Fourier-integral forms

$$\begin{aligned} \hat{\Psi}(x) &= \int \frac{d^4 p}{(2\pi\hbar)^4} e^{-\frac{i}{\hbar} p \cdot x} \hat{\Psi}(p), \\ A^e(x) &= \int \frac{d^4 p}{(2\pi\hbar)^4} e^{-\frac{i}{\hbar} p \cdot x} A^e(p). \end{aligned} \quad (\text{A65})$$

We finally note that variable dependence indicates in which space the corresponding physical quantity is: We use x variables for spacetime points (coordinate-space), and p and q variables for four-momentum points (in momentum-space). The first \mathcal{S} -matrix $\hat{\mathcal{S}}_{\text{SE},0}^{(2)}(\epsilon, \lambda)$ (the

zero-potential bound-state self-energy) becomes

$$\hat{\mathcal{S}}_{\text{SE},0}^{(2)}(0, 1) = \int \frac{d^4 q}{(2\pi\hbar)^4} : \bar{\Psi}(q)\gamma^\mu \Sigma(q)\gamma_\mu \hat{\Psi}(q) : \quad (\text{A66})$$

$$\Sigma(q) = e^2 \int \frac{d^4 p}{(2\pi\hbar)^4} S_0^F(q-p) D^F(p). \quad (\text{A67})$$

Here $\Sigma(q)$ is the so-called self-energy matrix function (see, for instance, Mandl and Shaw in Ref. 78 (eq.(9.20)). Notice that in the limit of large momentum p , the integrand behaves as $\propto \frac{1}{\gamma^\mu(q-p)_\mu} \frac{1}{p^3}$, which indicates a superficial linear divergence. However, with further investigation, one can show that this divergence is reduced by one degree, as noted by Schweber in Ref. 59 (section 15a), and presented in Table XVIII. Following the same steps, one can show that the second scattering matrix, associated with the one-potential bound-state self-energy process $\hat{\mathcal{S}}_{\text{SE},0}^{(2)}(0, 1)$ can be written as

$$\hat{\mathcal{S}}_{\text{SE},1}^{(2)}(0, 1) = -\frac{e}{i\hbar} \int \frac{d^4 p_2}{(2\pi\hbar)^4} \int \frac{d^4 p_1}{(2\pi\hbar)^4} A_\mu^e(p_2 - p_1) : \bar{\Psi}(p_2)\Lambda^\mu(p_2, p_1)\hat{\Psi}(p_1) : \quad (\text{A68})$$

$$\Lambda^\mu(p_2, p_1) = i\hbar e^2 \gamma^\nu \int \frac{d^4 q}{(2\pi\hbar)^4} D^F(q) S_0^F(p_2 - q) \gamma^\mu S_0^F(p_1 - q) \gamma_\nu, \quad (\text{A69})$$

where in the last equation, $\Lambda^\mu(p_2, p_1)$ is the so-called vertex-correction function; see, for instance, Mandl and Shaw⁷⁸ (eq.(9.48)). A detailed study of these momentum-space expressions and associated energy-shifts in the bound-electron problem was first considered by Snyderman in Ref. 176 (section 4) (see also Yerokhin and Shabaev¹⁷⁷).

d. Regularization and renormalization

When computing integrals associated with QED corrections, one finds (as already seen) that some of these integrals are divergent. How can one extract the meaningful finite (physical) from the meaningless infinite? This is done through regularization and renormalization.

Regularization is a technique for rendering a divergent integral convergent, albeit still dependent on the regularization-parameter. The main regularization techniques are: sharp momentum-cutoff, Pauli-Villars,¹⁷⁸ dimensional-regularization¹⁷⁹ and analytic continuation¹⁸⁰ regularization. The reader may also consult Zeidler in Ref. 181 (chapter 2) for a general conceptual formulations of regularization schemes. The sharp momentum-cutoff regularization consists of cutting off momentum contributions higher than a some $p_{\text{max}} = \Lambda \gg m_e c$. Unfortunately, this intuitive regularization breaks Lorentz- and gauge-invariance and the solution

is to use the other regularization schemes. The Pauli-Villars regularization consists of modifying the photon and electron propagators by introducing new propagators, associated with auxiliary masses (entering in propagators), for the self-energy and vacuum polarization processes. Finally, one can use dimensional regularization, which is based on the fact that logarithmically divergent integrals (as for the divergences associated basic QED processes) are convergent if one modifies the spacetime dimensions through $d = 4 \rightarrow d = 4 - \epsilon$ where ϵ is a small positive number. In all cases, after regularization, the divergent integrals are parameterized by the regularization parameters and are still divergent in the limit $\Lambda \rightarrow \infty$ or $\epsilon \rightarrow 0$, for instance. This is where renormalization comes into play.

Renormalization is a mathematical technique that consists of redefining the electron mass and charge (in addition to fields) such that the divergences, that come from including the QED corrections, are eliminated: absorbed by the bare physical quantities. It is needed at this point to note that the experimentally observed mass and charge, m_{exp} and e_{exp} are results of experiments that already include QED corrections. On the other hand, one can imagine a world in which the QED interaction is switched off; in this world, one would measure what is known as bare mass and charge: m_0 and e_0 . This distinction clearly shows that the electron mass and charge that we start with (before switching QED on: before

taking it into consideration) should be the bare ones, instead of the measured ones. This awareness played a crucial role in formulating the renormalization theory. Since we do not have access to bare quantities, and since infinity is not natural (not measurable), the renormalization theory says that we are allowed to redefine our physical constants such that the bare ones absorb the emerging divergences and lead to overall values of the physical constants that correspond to the experimentally observed ones. Detailed discussions on renormalization in the quantum field theory are provided by Collins in Ref. 182, Greiner and Reinhardt in Ref. 85 (chapter 5), Peskin and Schroeder in Ref. 60 (chapter 7), Itzykson and Zuber in Ref. 86 (section 7.1), and Huang in Ref. 183 (chapter 13) and 184.

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