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# Lithospheric transdimensional ambient-noise tomography of W-Europe: implications for crustal-scale geometry of the W-Alps

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## SUMMARY

A full understanding of the dynamics of mountain ranges such as the Alps requires the integration of available geological and geophysical knowledge into a lithospheric-scale 3-D geological model. As a first stage in the construction of this geo-model, we derive a new 3-D shear wave velocity model of the Alpine region, with a spatial resolution of a few tens of kilometres, making it possible to compare with geological maps. We use four years of continuous vertical-component seismic noise records to compute noise correlations between more than 950 permanent broad-band stations complemented by ~600 temporary stations from the AlpArray sea-land seismic network and the Cifalps and EASI linear arrays. A specific pre-processing is applied to records of ocean–bottom seismometers in the Liguro-Provençal basin to clean them from instrumental and oceanic noises. We first perform a 2-D transdimensional inversion of the traveltimes of Rayleigh waves to compute group-velocity maps from 4 to 150 s. The data noise level treated as an unknown parameter is determined with a Hierarchical Bayes method. A Fast Marching Eikonal solver is used to update ray path geometries during the inversion. We use next the group-velocity maps and their uncertainties to derive a 3-D probabilistic  $V_s$  model. The probability distributions of  $V_s$  at depth and the probability of presence of an interface are estimated at each location by exploring a set of 130 million synthetic four-layer 1-D  $V_s$  models. The obtained probabilistic model is refined using a linearized inversion. Throughout the inversion for  $V_s$ , we include the water column where necessary. Our  $V_s$  model highlights strong along-strike changes of the lithospheric structure, particularly in the subduction complex between the European and Adriatic plates. In the South-Western Alps, our model confirms the existence of a low-velocity structure at 50 – 80 km depth in the continuation of the European continental crust beneath the subduction wedge. This deep low-velocity anomaly progressively disappears towards the North-Western and Central Alps. The European crust includes lower crustal low-velocity zones and a Moho jump of ~ 8 – 12 km beneath the western boundary of the External Crystalline Massifs of the North-Western Alps. The striking fit between our  $V_s$  model and the receiver function migrated depth section along the Cifalps profile documents the reliability of the  $V_s$  model. In light of this reliability and with the aim to building a 3-D geological model, we re-examine the geological structures highlighted along the Cifalps profile.

**Key words:** Europe; Seismic noise; Seismic tomography; Surface waves and free oscillations; Crustal structure.

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## 1 INTRODUCTION

Since the Late Cretaceous, the geodynamic evolution of Western Europe has been dominated by convergence between the European and African plates. This long-term motion resulted in the formation of the Alps and other peri-Mediterranean mountain ranges (Pyrenees, Dinarides, Apennines, Carpathians, etc.) and also caused the opening of the Western European Cenozoic Rift System and of Mediterranean back-arc oceanic domains (Fig. 1a, Faccenna *et al.* 2014). In the Alps, convergence resulted in oceanic subduction from Late Cretaceous to Early Eocene (e.g. Handy *et al.* 2010), later on followed by continental subduction of the European margin (e.g. Chopin 1984; Duchêne *et al.* 1997; Guillot *et al.* 2009; Zhao *et al.* 2015). From Oligocene onwards, continental collision resulted in thrusting of the subduction wedge towards the European foreland, marked by the development of the Penninic Frontal Thrust (PFT, Simon-Labric *et al.* 2009).

The objective of this study is to improve the resolution of seismic models of the lithospheric structure of Western Europe in order to better constrain 3-D geological structures, particularly in the Alpine mountain belt and its forelands. For this, we build a new 3-D high-resolution shear wave velocity model using ambient-noise records of an exceptional array of seismic stations (Fig. 1b). To assess the robustness and the geological significance of our 3-D  $V_s$  model, we focus its geological interpretation on the Western Alps where the near-surface geology is well documented and a number of high-resolution geophysical investigations have been carried out recently (see review in Malusà *et al.* 2021).

In the internal zones of the Western Alps, the continental subduction of the European crust beneath the Adriatic lithosphere was first highlighted by the observation of preserved UHP minerals (e.g. coesite) in the outcropping subduction wedge (Chopin 1984). Recently, receiver functions (RFs) of the Cifalps experiment have imaged the continuity of the subducted European continental crust down to 75 – 80 km depth (Zhao *et al.* 2015). In this context, the combination of tectonics, petrophysics and numerical thermodynamics has provided conceptual models that are consistent with the formation of a subduction wedge during plate convergence (Burov *et al.* 2014; Liao *et al.* 2018), and with the involvement of serpentinites in the burial/exhumation processes (Schwartz *et al.* 2001; Guillot *et al.* 2009). However, despite the geological structure of the Western Alps having been studied for more than a century, correlating geological units mapped at the surface with geophysical images of the crust remains challenging. This is due to the limited resolution of these images, the lack of accurate uncertainty estimates, and the non-unicity of their geological interpretations regarding lithology, thermicity, fluids and deformation. Moreover, some first order questions remain controversial, such as the nature of the geophysical Ivrea body (IB, e.g. Schmid *et al.* 2017; Solarino *et al.* 2018), and the precise geometry of the Adria mantle wedge (MW) and its role in the partitioning of the present-day strain field (Lardeaux *et al.* 2006; Eva *et al.* 2020).

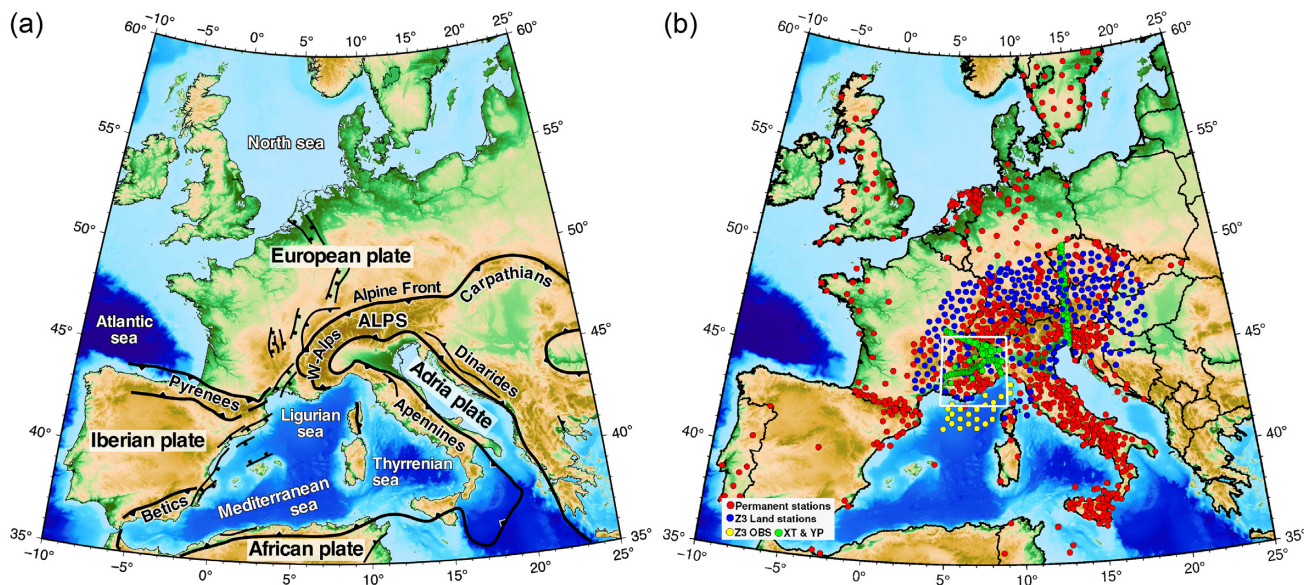
Since the 1990s, the structure of the lithosphere beneath the Western Alps has been probed by a broad spectrum of geophysical techniques. The ECORS-CROP normal-incidence and wide-angle reflection controlled-source seismic survey (CSS) has provided a very high-resolution image of the reflectivity of the crust in the North-Western Alps (Nicolas *et al.* 1990). Local earthquake tomography (LET) studies have yielded local or regional 3-D  $P$ -wave velocity models of the crust (e.g. Paul *et al.* 2001; Diehl *et al.* 2009; Solarino *et al.* 2018). RFs have been used to map Moho depth as

a complement to the 2-D CSS profiles and 3-D LET models (e.g. Spada *et al.* 2013; Zhao *et al.* 2015). These geophysical data have lower resolution than surface geological data. They were however of great help to constrain to the first order the 3-D geometry of the crustal structure under the Western Alps. In particular, they highlighted the underthrusting of the European plate beneath the Adria microplate, with a maximum Moho depth locally reaching 55 km (ECORS-CROP Deep Seismic Sounding Group, Hirn *et al.* 1989; Thouvenot *et al.* 2007).

In the last decade, seismic imaging of the Alpine lithosphere has improved thanks to the densification of permanent broad-band seismological networks and the multiplication of temporary experiments (e.g. Cifalps-2, Liu *et al.* 2022; Hetényi *et al.* 2018a). New passive imaging methods have been introduced that make optimal use of these dense station arrays and are independent of earthquake or active source illumination. Campillo & Paul (2003) and Shapiro & Campillo (2004) have shown that the cross-correlation of long time series of diffuse wavefields including ambient noise recorded at two seismic stations leads to the reconstruction of the surface waves propagating between the two stations as if a virtual seismic source is placed at one station with the emitted signal being recorded at the second station. Shapiro *et al.* (2005) have further shown that ambient-noise tomography (ANT) from continuous noise recordings over dense networks provides high-resolution images of the crust and upper mantle. Since then, several  $S$ -wave velocity models have been derived from ANT to probe the crustal structure under the Alpine arc (e.g. Stehly *et al.* 2009; Verbeke *et al.* 2012; Molinari *et al.* 2015).

Taking advantage of the ever-increasing amount of available data, Kästle *et al.* (2018) used a stochastic inversion of traveltimes of surface waves from noise correlations, regional and teleseismic earthquakes to construct a 3-D  $V_s$  model of the entire Alpine region. Lu *et al.* (2018) used ambient-noise records of 1293 broad-band stations, including the first six months of the AlpArray network, to derive a high-resolution 3-D  $V_s$  model of the European crust from ANT using a Bayesian inversion. This model was further refined in the Alpine region by Lu *et al.* (2020) using wave equation tomography (WET) of Rayleigh-wave phase dispersion data derived from noise correlations. Zhao *et al.* (2020) used Bayesian transdimensional inversion of the group-velocity data set of Lu *et al.* (2020) to derive a shear wave velocity model of the Western Alps that provides a particularly well-resolved image of the subduction. These new approaches highlighted major intracrustal units confirmed by field observations, such as the subduction complex in the internal zone of the Alps, which is either located beneath an MW in the Western Alps or covered by Adriatic crustal units in the Central and Eastern Alps. Since the works by Lu *et al.* (2018, 2020) and Zhao *et al.* (2020), more data have been recorded, in particular in the Western Alps, that should contribute to improving the resolution of ANT.

In the present work, we generate a new 3-D shear wave velocity model of Western Europe by combining ANT with probabilistic inversion. We use four years of continuous seismic noise records at all available broad-band seismic stations including permanent networks, the Cifalps and Cifalps-2 profiles and the AlpArray network (see Fig. 1a, and Section 2). We also include data of the AlpArray ocean–bottom seismometers (OBS) to image the transition between the Ligurian sea and the Alpine region. This data set gives a unique opportunity to image the crust and upper mantle below the Alpine belt and its forelands at an unprecedented resolution. The originality of our approach lies in the use of a transdimensional inversion to compute 2-D group-velocity maps and their uncertainties, which are



**Figure 1.** (a) Topographic map showing the regions, plates and main geological boundaries discussed in the text (modified from Faccenna *et al.* 2014). Black lines with triangles correspond to main thrusts and subduction zones. Black lines with squares represent extensional faults associated with development of the Western European Cenozoic Rift System. (b) Location map of the broad-band seismic networks used in this study. The white frame which includes most of the Western Alps indicates the main focus area of this work.

then used to obtain a 3-D probabilistic shear wave velocity model from a Bayesian inversion (Sections 3 and 4). This model provides a probability distribution of  $V_s$  as well as the probability of presence of an interface at each depth and location. This approach was chosen specifically to image sharp lateral discontinuities, which is crucial in the Western Alps. To the best of our knowledge, this is the first time that a fully probabilistic approach is used at this scale to image the lithosphere. In Section 5, we further assess the model robustness along the Cifalps profile (South-Western Alps, SWA) by comparing our results with an RF depth section. Finally, we propose an updated crustal-scale interpretative geological cross-section based on previous geophysical-geological works (reviewed in Malusà *et al.* 2021), on our  $V_s$  model and the RF section along the Cifalps profile. This comparison illustrates the potential contribution of our  $V_s$  model to the development of a 3-D geological model of the Alpine chain and its forelands.

## 2 AMBIENT-NOISE DATA SET

### 2.1 Station coverage

Our data set consists of four years of vertical-component, daily seismic noise records (2015–2019) of more than 950 permanent broad-band seismic stations located in and around the Greater Alpine region, complemented by 490 temporary stations from the AlpArray (AlpArray Seismic Network 2015), Cifalps (YP network, Zhao *et al.* 2016), and XT network (Zhao *et al.* 2018) and EASI experiments (XT network, AlpArray Seismic Network 2014, Fig. 1b). The average interstation distance is  $\sim 50$  km in the Greater Alpine region.

The on-land component of the AlpArray seismic network (AASN) was operated for 2–4 yr and covered the Greater Alpine area (Hetényi *et al.* 2018a). We also used data from 23 broad-band OBS that were deployed during 8–10 months (2017–2018) in the Ligurian Sea as part of the AASN (Hetényi *et al.* 2018a). The OBS component of the AASN was operated within the framework of a

French–German cooperation that aimed at imaging the lithospheric structure of the Ligurian basin and the transition between the Alps and the Apennines. OBS records are merged into the Z3 data set, which means that data and metadata are archived in the same standard (FDSN) format as data recorded by on-land stations, and they are distributed by the RESIF and GEOFON nodes of the European Integrated Data Archive (EIDA) using the same procedures as on-land data. This integration of data recorded at sea-bottom and on land greatly facilitates their use in massive data processing such as that presented here.

In addition to the AASN, we also used data of the dense (7–10 km spacing) Cifalps (China–Italy–France Alps seismic survey) and Cifalps-2 linear arrays that operated for 14–15 months in the Western Alps in 2012–2013 and 2018–2019, respectively. They provide a very dense coverage of the Western Alps, which are the main focus of this work (Fig. 1b).

### 2.2 Data processing and correlation

Before computing the correlations for each station pair, we pre-processed the noise records in two main steps. First, we applied a generic pre-processing scheme where each daily record was band-pass filtered between 2.5 and 300 s, corrected for the instrumental response, decimated to 1 Hz sampling frequency and split into 4-hr segments. Second, we decreased the contribution of earthquakes and other transient signals by (1) removing signals with amplitude 4 times greater than the standard deviation and (2) removing segments with RMS greater than 1.5 times the daily mean RMS (Boué *et al.* 2014). Then, each daily record was filtered into several period bands (3–5, 5–10, 10–20, 20–40, 40–80 and 80–200 s) and normalized by its envelope. Finally, the six filtered and normalized signals were stacked. We thus obtained a broad-band signal with amplitude normalized in several period bands. This processing is similar to the one used by Soergel *et al.* (2019).

In addition, we used a specific processing for the data of the 23 OBS in order to reduce the tilt and the compliance effects using

frequency-dependent response functions (Crawford & Webb 2000). We also corrected the records of 8 OBS that were affected by periodic glitches of electronic origin. This was achieved using a template matching algorithm (Deen *et al.* 2017).

For each of the 1.1 million station pairs, we computed the cross-correlations of up to 4 yr of vertical component continuous noise records by segments of 4 hr. The resulting cross-correlations were then normalized and stacked. Fig. S1 (Supporting Information) shows distance–time sections of cross-correlations in the 5–10, 10–20 and 20–40 s period bands. We used a specific computation scheme for correlations between ocean–bottom stations in order to enhance the signal-to-noise ratio. It was indeed poorer for OBS pairs than for on-land station pairs or mixed pairs of on-land and ocean–bottom stations. We took advantage of the large number of on-land stations and their high quality by using them as virtual sources for the OBS–OBS correlations. For each OBS pair, we correlated the ballistic Rayleigh waves of the noise correlations between each OBS and a large number of selected on-land stations, and stacked them to obtain the OBS–OBS correlation. This process is known as ‘C2’ (for correlation of correlation) which is a variant that uses the ballistic waves rather than the coda (Stehly *et al.* 2008). Another publication will be dedicated to the processing of OBS records.

### 2.3 Group-velocity measurements

Once cross-correlations were computed for the entire data set, we derived, for each pair of stations, the group-velocity dispersion curves of positive and negative correlation times by using multiple filter analysis (MFA, Dziewonski *et al.* 1969; Herrmann 1973). Similarly to Lu *et al.* (2018), we adapted the width of the filter to the interstation distance to accommodate the trade-off in resolution between the time and frequency domains (Levshin *et al.* 1989). We also corrected our group-velocity measurements for the systematic error that occurs with the MFA technique due to the strong amplitude decrease of the noise spectrum at periods  $> 20$  s (Shapiro & Singh 1999).

Careful selection of group-velocity measurements prior to group-velocity tomography is essential to prevent biases induced by the heterogeneous distribution of noise sources, by interferences of Rayleigh waves in the causal and acausal times and by instrumental problems. To that aim, for each station pair and period, we selected the group-velocity measurements with a signal-to-noise ratio  $> 3$  and the group-velocity difference between causal and acausal Rayleigh waves  $< 0.2 \text{ km s}^{-1}$ . Furthermore, we only kept paths with length of 2–40 wavelengths at each period. By combining these criteria, we keep only the reliable traveltime measurements for the subsequent 2-D group-velocity maps.

## 3 INVERSION FOR 2-D GROUP-VELOCITY MAPS

Our aim is to derive a 3-D probabilistic  $V_s$  model of the Alpine crust and upper mantle. As a first step towards this goal, we compute 2-D group-velocity maps and associated uncertainties using a transdimensional algorithm at discrete periods from 4 to 150 s.

Probabilistic group-velocity maps are derived by exploring millions of 2-D models with different parametrizations using the reversible-jump Markov-chain Monte-Carlo method (*rj-McMC*), first applied in a seismic tomography context by Bodin *et al.* (2012).

The parametrization of the model is treated as part of the inversion without any explicit regularization. This allows the local resolution to self-adapt to the path density and to the variability of the information contained in group-velocity measurements. The model complexity required to fit the data is controlled by the noise level, which is treated as an extra parameter of the inversion and determined within a hierarchical Bayes formalism (Malinverno & Briggs 2004).

Moreover, since the Alpine crust is strongly heterogeneous and displays sharp group-velocity contrasts, the nonlinearity of the forward problem is accounted for by iteratively updating the ray path geometry using the fast marching method (FMM, Rawlinson & Sambridge 2004). Ray bending is more sensitive to phase-velocity changes than to group-velocity changes. However, accounting for ray bending in group-velocity *rj-McMC* tomography has proven to be substantially more accurate in a heterogeneous medium than the straight ray assumption (e.g. Galetti *et al.* 2015).

### 3.1 Method

The 2-D velocity field is parametrized with a set of Voronoi cells with variable number and geometries. The velocity field is described by a vector of model parameters  $\mathbf{m}$  giving the position and group velocity associated with each Voronoi cell.

The inverse problem is treated in a Bayesian framework, where the solution is represented by the posterior probability density function representing the probability of the model  $\mathbf{m}$ , given a set of observed data  $\mathbf{d}$ . The posterior solution is expressed according to Bayes’s theorem (Bodin *et al.* 2012)

$$p(\mathbf{m}|\mathbf{d}) \propto p(\mathbf{d}|\mathbf{m})p(\mathbf{m}) \quad (1)$$

where  $p(\mathbf{m})$  is the *a priori* probability density of the model parameters  $\mathbf{m}$ , that is, what we know about the velocity field independently of the data. The term  $p(\mathbf{d}|\mathbf{m})$  is the likelihood function and represents the probability of observing  $\mathbf{d}$  given a model  $\mathbf{m}$ , and given the statistics of data errors. Assuming normally distributed uncorrelated data errors,  $p(\mathbf{d}|\mathbf{m})$  can be expressed with the general Gaussian form

$$p(\mathbf{d}|\mathbf{m}) = \frac{1}{\prod_{i=1}^N \sqrt{2\pi}\sigma_{d_i}} \times \exp\left(\frac{-\phi(\mathbf{m})}{2}\right) \quad (2)$$

where  $\sigma_{d_i}$  is the standard deviation of data errors on the  $i$ th observation,  $N$  is the number of observations and  $\phi(\mathbf{m})$  is the misfit function for the model  $\mathbf{m}$

$$\phi(\mathbf{m}) = \sum_{i=1}^N \left[ \frac{(g_i(\mathbf{m}) - \mathbf{d}_i)^2}{\sigma_{d_i}^2} \right] \quad (3)$$

The term  $g_i(\mathbf{m})$  represents data computed by the forward problem, that is, the traveltime of the  $i$ th ray predicted by the model  $\mathbf{m}$ , and computed from

$$g_i(\mathbf{m}) = \sum_{j=1}^n \frac{L_{ij}}{v_j} \quad (4)$$

where  $L_{ij}$  is the length of the  $i$ th ray across the Voronoi cell  $j$  of velocity  $v_j$ .

The posterior probability distribution can be estimated with the reversible-jump *McMC* algorithm, which produces a large ensemble of models, whose distribution approximates the posterior solution. The algorithm is based on a Markov-chain Monte-Carlo sampler

where at each iteration, a new velocity model is proposed by perturbing the current model (e.g. perturb the geometry of the Voronoi discretization, change the velocity within one Voronoi cell). The proposed model is then either accepted or rejected in the ensemble depending on an acceptance criteria based on the ratio of posterior values of the current and proposed model. Once the algorithm has been run for enough iterations, statistics can be extracted from the ensemble solution for interpretation. For example, at any geographical location, the mean and standard deviation of velocities can be used to produce a group-velocity map, with associated error estimates. We refer the reader to Bodin *et al.* (2012), for more details.

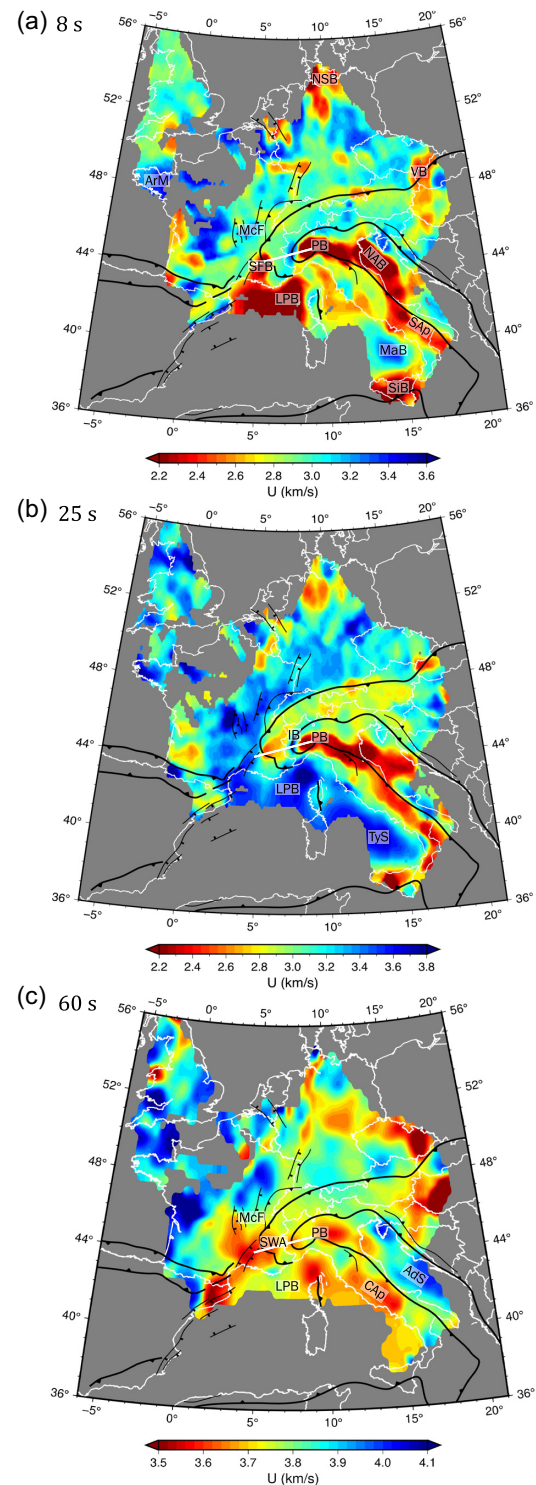
### 3.2 2-D group-velocity maps

We performed the *rj-McMC* tomography to derive group-velocity maps for periods between 4 and 150 s. For each period, 64 Markov chains were run in parallel to explore independently the model space. Each individual process was run for  $180 \times 10^3$  steps in total. Once the convergence was achieved, the ensemble of sampled models was averaged to produce a smooth average solution, that can be interpreted as the mean of the posterior solution. We used this average model to update ray path geometry and recalculate traveltimes using the FMM. The global scheme was run for two iterations, resulting in the group-velocity maps shown in Figs 2 and 3.

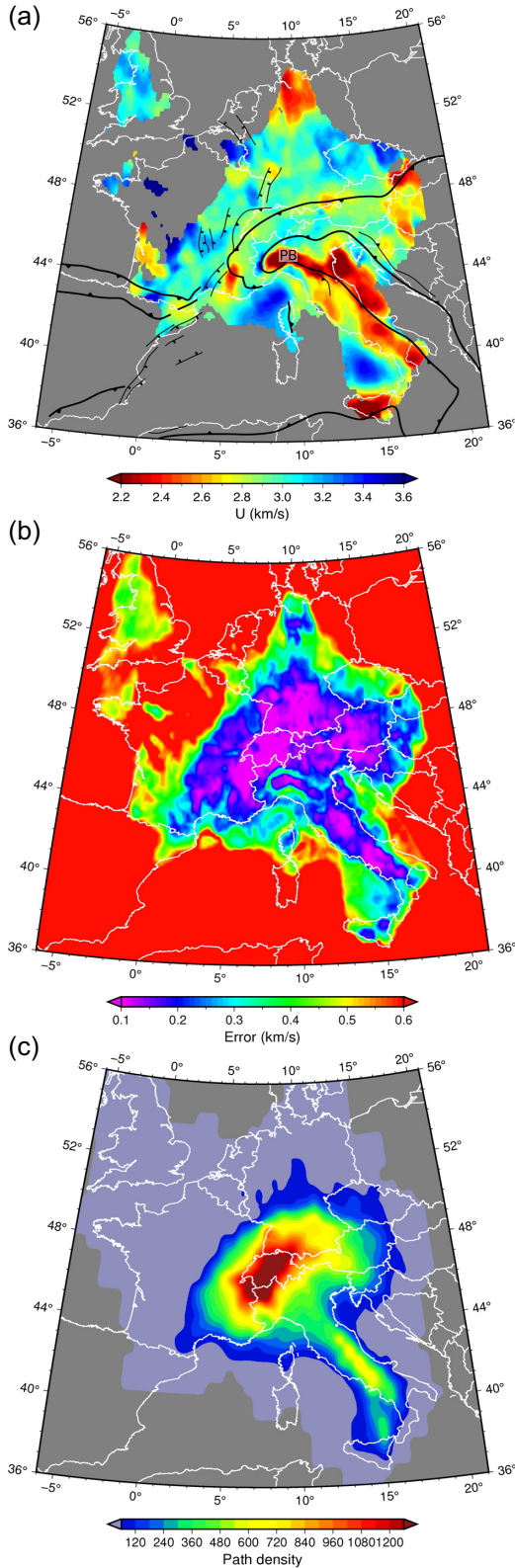
At 8 s period (mainly sensitive to  $\sim 5 - 8$  km depth; Fig. 2a), the clearest features of the group-velocity map are low-velocity anomalies associated with thick sedimentary basins: the Po basin (labelled PB in Fig. 2a) and the North-Adriatic basin (NAB), the Liguro-Provençal basin (LPB), the Southeast-France basin (SFB), the Vienna basin (VB) in the easternmost part of our study area and part of the North-Sea basin in Northern Germany (NSB). The Southern Apennines (SAp) and the Sicily fold-and-thrust belt (SiB) are also characterized by low velocities. The highest velocities ( $> 3.2 \text{ km s}^{-1}$ ) are observed in the Variscan massifs of France such as the Western Massif Central (McF) and the Armorican Massif (ArM), and in the Marsili backarc basin (MaB) of the Tyrrhenian Sea (TyS). Fig. S2(a) (Supporting Information) shows the main sedimentary basins on the 8-s group-velocity map.

At 25 s (sensitive to  $\sim 15 - 30$  km depth; Fig. 2b), the LPB (in Fig. 2b) has velocities  $> 3.5 \text{ km s}^{-1}$  due to the thin crust of the basin. Similar high-velocity anomalies are observed in the TyS and to the west and north of the Alpine front in the Western and Central Alps. The very thick PB is still characterized by very low velocities ( $< 2.3 \text{ km s}^{-1}$ ), while slightly higher velocities ( $\sim 2.5 - 2.6 \text{ km s}^{-1}$ ) prevail in the Apenninic belt and moderately low velocities ( $\sim 2.7 - 3.0 \text{ km s}^{-1}$ ) in the Alpine belt. The strong high-velocity anomaly associated with the shallow mantle flake known as the IB is observed between the PB and the Western Alps (IB in Fig. 2b) in spite of its small size ( $\sim 10 - 15$  km width). This gives a first insight into the good lateral resolution of our 2-D tomography. Fig. S2(b) (Supporting Information) shows the comparison between the 25-s map and the reference Moho depth model of Spada *et al.* (2013).

At 60 s (sensitive to  $\sim 50 - 100$  km depth; Fig. 2c), the upper mantle appears as heterogeneous as for shorter periods with low-velocity anomalies beneath the PB and the central Apennines (CAp), the LPB, the south of the McF and the SWA. High-velocity anomalies are observed beneath the Adriatic Sea (AdS) and in the northwest of the model.



**Figure 2.** Group-velocity maps (average solutions) at (a) 8 s, (b) 25 s and (c) 60 s periods, obtained with the Hierarchical Bayes reversible-jump algorithm. Only areas with uncertainty lower than  $0.5 \text{ km s}^{-1}$  are shown. Black lines show the main geological boundaries as defined in Fig. 1(a). The white line shows the Cifalps profile discussed in Section 5. ArM: Armorican Massif, IB: Ivrea geophysical body, LPB: Liguro-Provençal basin, MaB: Marsili backarc basin, McF: Eastern Massif Central, NAB: North-Adriatic basin, NSB: North-Sea basin, PB: Po basin, CAp: Central Apennines, SAp: Southern Apennines, SWA: South-Western Alps, SFB: Southeast-France basin, SiB: Sicily fold-and-thrust belt, TyS: Tyrrhenian Sea, VB: Vienna basin and AdS: Adriatic Sea.



**Figure 3.** (a) Group-velocity map (average solution) obtained from the *rj-McMC* tomography at 15 s period, showing regions with error  $< 0.5 \text{ km s}^{-1}$ . (b) Estimated error map (standard deviation of the ensemble of sampled velocities). (c) Corresponding path density map (number of paths crossing each  $0.15^\circ \times 0.15^\circ$  cell).

### 3.3 Uncertainty estimates

At each geographical location  $i$ , we extracted posterior uncertainties from the variance of the ensemble of sampled velocities. This uncertainty is given by

$$\sigma_i = \sqrt{\frac{1}{M} \sum_{j=1}^M \left( v(m_j) - \bar{v} \right)^2} \quad (5)$$

where  $M$  is the total number of models in the Markov-chain ensemble and  $\bar{v}$  the average solution model. This uncertainty is primarily related to the complexity of the velocity structure, but also depends on the reliability of the observations and on ray path coverage.

The uncertainty map displayed in Fig. 3(b) for 15 s period shows that our group-velocity model (Fig. 3a) is better constrained in the Greater Alpine region and in the Italian peninsula (uncertainty  $< 0.15 - 0.2 \text{ km s}^{-1}$ ) than in the Ligurian Sea (uncertainty  $\sim 0.25 - 0.4 \text{ km s}^{-1}$ ). This difference is certainly due to a less dense path coverage but also to relatively poorer data quality and the complexity of the medium.

A comparison of Figs 3(b) and (c) shows that although uncertainty is generally low where ray coverage is good, a few areas have high uncertainty in spite of good coverage. This happens in areas of strong velocity contrast, for example along the boundaries of the low-velocity anomaly of the PB (in Fig. 3a). Galetti *et al.* (2015) attributed such loop-like structures in the error maps to the nonlinearity of the forward model. Here instead, we interpret these features as due to the presence of sharp lateral discontinuities in the group-velocity maps. These typical structures in error maps are useful to identify velocity anomalies with uncertain location but well-constrained amplitude.

### 3.4 Benefits of the transdimensional inversion

The inversion scheme that we used to derive the 2-D group-velocity maps is similar to the one used by Bodin *et al.* (2012). This procedure has been successfully adapted to our study region despite its large size and the high volume data set. The dynamic parametrization with Voronoi cells with variable geometries allowed the inversion process to accommodate the irregular distribution of data.

As shown in Fig. S3 (Supporting Information), the 2-D transdimensional inversion outperforms commonly used linearized and regularized inversions. First, this approach allows the computation of group-velocity maps together with their posterior uncertainties that are required to constrain the 3-D  $V_s$  model. Second, linearized inversions require explicit parametrization and regularization that often rely on subjective choices. Such a regularization stabilizes the solution in poorly resolved areas, but tends to smooth lateral discontinuities and to bias the resulting solution. These problems do not arise within a Bayesian framework since the degree of smoothing and damping of the solution is naturally driven by the data noise level as well as the heterogeneity of the medium.

The number of Voronoi cells self-adapts to the level of structure present in the data. We show in Fig. S4 (Supporting Information) the posterior distribution on the number of Voronoi cells. We also show the evolution of the model misfit function along the Markov chain and the posterior distribution of the traveltime misfit (Fig. S5, Supporting Information). We further assess the resolution power of the method by showing results from a synthetic checkerboard test (Fig. S6, Supporting Information).

**Table 1.** Ranges of layer thickness and shear wave velocity used to build the set of four-layer models for the grid search.  $P$ -wave velocity and density are derived from  $V_s$  using Brocher's empirical relationship (Brocher 2005).

Layer	Thickness (km)	$V_s$ (km s <sup>-1</sup> )
Sediments	0–16	1.6–2.9
Upper crust	0–24	2.6–3.8
Lower crust	2–42	3.3–4.3
Mantle	inf	3.7–4.7

## 4 INVERSION FOR SHEAR WAVE VELOCITY

We used 2-D group-velocity maps and their uncertainties to build a 3-D  $V_s$  model of the crust and upper mantle of the Greater Alpine region in two steps. We first computed a 3-D probabilistic  $V_s$  solution that gives at each location the probability distribution of  $V_s$  and the probability of having an interface as a function of depth. Second, we computed the final  $V_s$  model (with a single  $V_s$  value at each depth) by refining the probabilistic model using a linear inversion at each pixel, as initially introduced by Macquet *et al.* (2014). This is particularly useful in areas with complex structures such as the IB region, where the mean probabilistic model (obtained with an overly simple parametrization) does not fully explain the local dispersion curve.

### 4.1 Construction of the 3-D probabilistic $V_s$ model

To build the 3-D probabilistic model  $V_s$ , we first extracted at each location the local Rayleigh-wave group-velocity dispersion curve from the group-velocity maps presented in the previous section. As discussed further, each of these local dispersion curves is inverted from 4 to 65 s for a local 1-D probabilistic  $V_s$  depth profile that provides at each depth the probability distribution of the shear wave velocity and the probability of presence of an interface. This is done using an exhaustive grid search over a set  $\sim 130$  million of four-layer synthetic models. Each individual model is described with a simple parametrization that includes a sedimentary layer, an upper crust, a lower crust and a half-space representing the uppermost mantle. We assume *PREM* Earth model velocities from the Moho depth to 400 km (Dziewonski & Anderson 1981). Each layer is parametrized by its thickness and  $S$ -wave velocity.  $P$ -wave velocities and densities are converted from  $V_s$  using Brocher's empirical formula (Brocher 2005). As shown in Table 1, we choose to explore a wide range of four-layer models that allows for slow velocity layers. As discussed in Section 4.2, a Bayesian solution (posterior mean) is then extracted by averaging the ensemble of models weighted by their posterior probability.

The 1-D inversion problem is also cast in a Bayesian framework, and we evaluate the probability that each synthetic model  $\mathbf{m}$  explains the measured local dispersion curve  $\mathbf{d}$  using the Gaussian likelihood function

$$p(\mathbf{d}|\mathbf{m}) = \frac{1}{\sqrt{(2\pi)^N |\mathbf{C}|}} \times \exp\left(\frac{-\phi(\mathbf{m})}{2}\right) \quad (6)$$

where  $N$  is the total number of measured periods,  $\mathbf{C}$  the covariance matrix of data errors and  $\phi(\mathbf{m})$  the misfit function. We assume that uncertainties at different periods of the local dispersion curve are uncorrelated. In that case  $\mathbf{C}$  has a diagonal form

$$\mathbf{C} = \begin{bmatrix} \sigma_1^2 & 0 & 0 \\ 0 & \dots & 0 \\ 0 & 0 & \sigma_N^2 \end{bmatrix} \quad (7)$$

where  $\sigma_i$  is the posterior uncertainty on the group velocity at period  $i$ , estimated from the transdimensional group-velocity inversion (see Section 3.3).

The misfit  $\phi(\mathbf{m})$  between the candidate synthetic dispersion curve  $g(\mathbf{m})$  and the local observed dispersion curve is estimated according to the Euclidean distance

$$\phi(\mathbf{m}) = (g(\mathbf{m}) - \mathbf{d})^T \mathbf{C}^{-1} (g(\mathbf{m}) - \mathbf{d}) \quad (8)$$

By substituting eqs (7) and (8) into eq. (6), we obtain the discrete form

$$p(\mathbf{d}|\mathbf{m}) = \frac{1}{\prod_{i=1}^N \sqrt{2\pi} \sigma_i} \times \exp\left(\sum_{i=1}^N \frac{-(g_i(\mathbf{m}) - d_i)^2}{2\sigma_i^2}\right) \quad (9)$$

This likelihood function gives the probability of observing the data given a 1-D profile. Here, we assume all tested models are equally probable *a priori* (uniform prior distribution), which makes  $p(\mathbf{m})$  a constant and allows us to write

$$p(\mathbf{m}|\mathbf{d}) \propto p(\mathbf{d}|\mathbf{m}) \quad (10)$$

At each depth, we therefore estimate the posterior probability of shear wave velocities by simply weighting each of the tested models by its likelihood value  $p(\mathbf{d}|\mathbf{m})$ . Since we perform a uniform grid search in a multidimensional space, a large majority of models are not fitting the data, and have a very low posterior value. For practical reasons, we only used the  $100 \times 10^3$  best-fitting models in the ensemble, and assigned a 0 posterior value to the rest. We verified that this way of approximating the Bayesian solution does not modify the mean and variance of the posterior solution.

This allows us to get for each depth the probability distribution of  $V_s$ . In the same way, the probability of presence of an interface can be derived from the ensemble of sampled models and their associated likelihood value.

### 4.2 Construction of the 3-D final $V_s$ model

The ambient-noise Bayesian scheme described above was applied up to 65 s period to better constrain the crustal part of the model while remaining consistent with the four-layer assumption. To improve the fit to the observed dispersion curve and better constrain the upper-mantle part of the model, at each location we averaged the ensemble of selected models weighted by their likelihood values to build a deterministic 3-D  $V_s$  model that gives a single value of  $V_s$  at each location and depth. This is preferable to only getting the most probable  $V_s$  value, since this may introduce sharp discontinuities related to the four-layer assumption. This deterministic  $V_s$  model was then used as initial model to perform a linearized inversion (Herrmann 2013). Three iterations were computed in the 4 – 150 s frequency band, taking into account the errors on the observed dispersion curves.

This approach has been modified for the marine parts of the model. We used the same probabilistic inversion as for the on-land pixels and we added a surface layer representing the water column to the initial model of the linear inversion. The thickness of the water layer equals the water depth at the given pixel. We kept the parameters of the water layer fixed during the inversion ( $\rho$ ,  $V_s$  and  $V_p$ ).



### 4.3 Strengths of the inversion method for $V_s$

The inversion scheme that we used to derive the 3-D probabilistic  $V_s$  model is similar to that used by Lu *et al.* (2018), but with two significant improvements. First, rather than using a set of 8 million 1-D  $V_s$  models at each location, we explored a broader set of 130 million models that allow for velocity inversions at depth. Second, we did not consider the uncertainties of the local dispersion curves  $\sigma$  to be the same at all periods, nor did we consider them as a parameter of the inversion. Instead,  $\sigma$  was evaluated during the 2-D transdimensional group-velocity inversion. Hence, we formulated the likelihood function so that the uncertainties on group velocities  $\sigma$  were taken into account when computing the misfit  $\phi(\mathbf{m})$  between the local dispersion curve and the dispersion curves associated to each of the 130 million of synthetic  $V_s$  models. Indeed, an underestimation of the noise level contained in the data would limit the range of 1-D  $V_s$  model considered as probable since the algorithm would try to fit the observed local dispersion curve too closely. On the other hand, overestimating the noise level would lead to considering irrelevant 1-D  $V_s$  models as probable, hence widening the probability distribution of  $V_s$ .

Fig. 4 presents an example of a 1-D inversion at a pixel located in the Western Alps (Fig. 4a). In Fig. 4(b), the local dispersion curve (black points) and its uncertainties are compared with the dispersion curves associated with the Bayesian average model obtained from the probabilistic inversion (red line) and the final  $V_s$  model obtained with an additional linear inversion (green line). The probability distribution of  $V_s$  and the probability of presence of an interface are shown on the left- and right-hand panels of Fig. 4(c), respectively. The Bayesian solution (weighted average of the best-fitting models, see Section 4.2) and the final model obtained after a linear inversion are shown with red and black lines, respectively.

We observe a rather sharp  $V_s$  probability distribution over the whole depth range that we investigated (width  $\leq 0.2 \text{ km s}^{-1}$  for probability  $> 50$  per cent). The probability that an interface is present has two main peaks that correspond to the boundary between the sediment layer and the upper crust at 7–8 km and to the Moho at 31–32 km. This results in a multimodal distribution on  $V_s$ , with a very high standard deviation (same effect as the ‘loops’ in 2-D, see Section 3.3).

The dispersion curve associated with the probabilistic average model (red curve) fits well the observed dispersion curve (black points in Fig. 4b) at periods less than 30 s and greater than 50 s. However, between 30 and 50 s where the Rayleigh waves are mostly sensitive to the Moho depth, the posterior mean solution does not completely explain the data, even if it is still compatible with the observed dispersion curve when taking into account the uncertainties. This highlights that in the most complex areas, using a grid search over a library of 130 million four-layer models is insufficient to completely describe the complexity of the medium.

However, in this case, the reference probabilistic model is a relevant initial model for the linear inversion: as shown on Fig. 4(b), the dispersion curve associated with the final model obtained after a linear inversion fits well with the observed dispersion curve in the whole 4–90 s period range.

### 4.4 Results: 3-D $V_s$ model

We have derived a quasi-3-D shear wave velocity model by assembling all 1-D individual models. This new large-scale model has a particularly good resolution over the Greater Alpine region where

the density of stations is highest. We have evaluated the reliability of the  $V_s$  model by computing the data misfit reduction for each pixel.

Fig. 5 presents depth slices at 6, 15, 30 and 60 km in the final 3-D  $V_s$  model. The 6-km depth slice in Fig. 5(a) is consistent with the geological map. We retrieve the main anomalies that we previously discussed on the 8 s map of Fig. 2(a), in particular the thick sedimentary basins with velocities lower than  $2.7 \text{ km s}^{-1}$  (PB and NAB, LPB, SFB and VB).

At 15-km depth (Fig. 5b), the IB is visible as a high-velocity anomaly with  $V_s$  in the range  $4 - 4.1 \text{ km s}^{-1}$ . The LPB is associated with high velocities greater than  $4 \text{ km s}^{-1}$ , indicating a shallow Moho. Another striking feature is that at this depth the Adriatic crust exhibits lower velocities ( $3.1 - 3.2 \text{ km s}^{-1}$ ) than European upper crust ( $3.3 - 3.5 \text{ km s}^{-1}$ ).

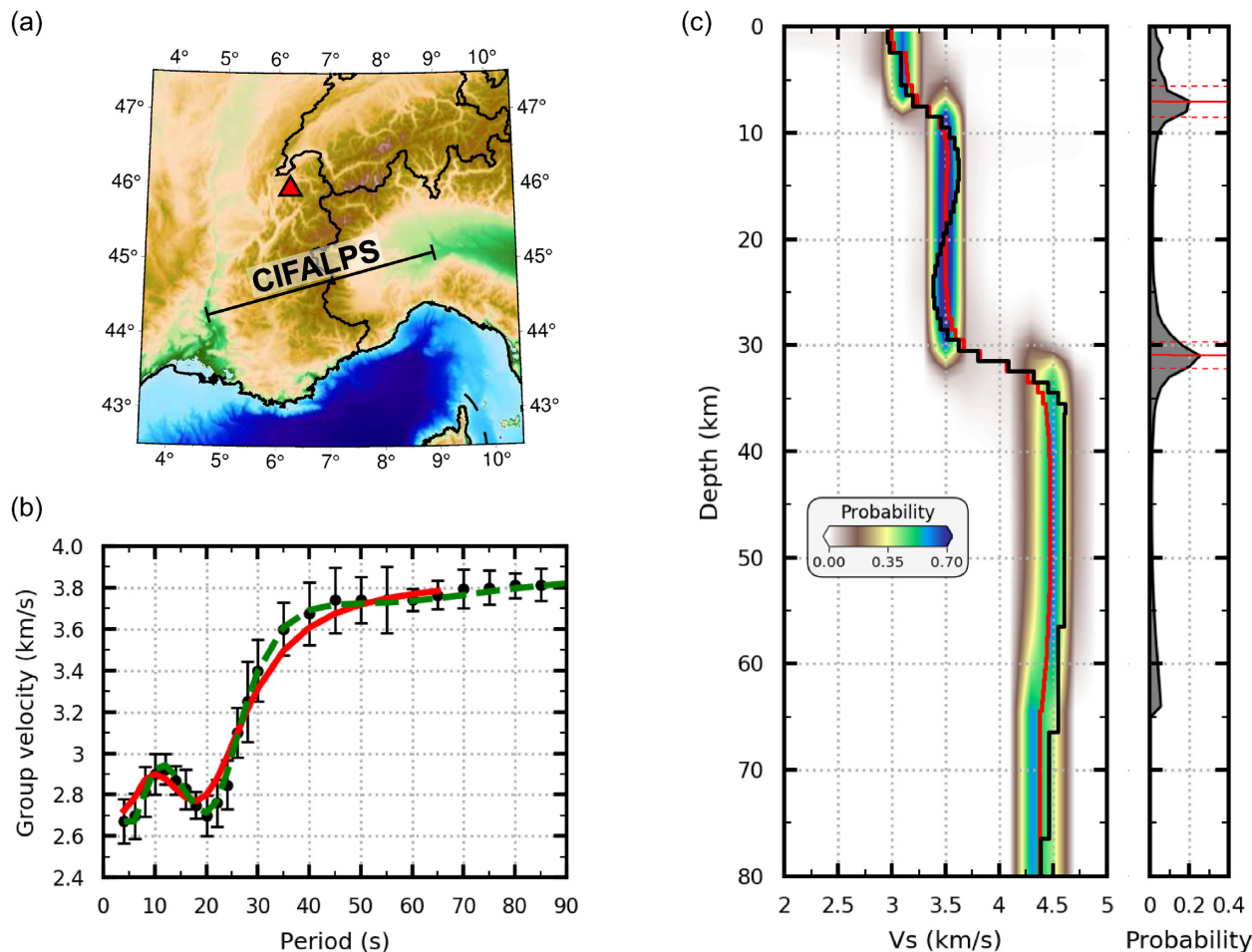
The 30-km depth slice (Fig. 5c) underlines the variations in crustal thickness, with low velocities ( $3.5 \text{ km s}^{-1}$ ) in the mountain belts such as the Alps and Apennines. The 60-km depth slice shows a strong low-velocity anomaly ( $V_s < 4 \text{ km s}^{-1}$ ) at the location of the IB, in sharp contrast to the shallower high-velocity anomaly. The Apennines, the Northern Adriatic and Sicily have higher velocities ( $V_s > 4.6 \text{ km s}^{-1}$ ) than the Alpine belt and its surrounding regions to the north and west. The LPB and the TS have rather low velocities of  $4.1 - 4.3 \text{ km s}^{-1}$ . The low velocities ( $4 - 4.2 \text{ km s}^{-1}$ ) beneath the volcanic regions and the grabens of the French Massif Central (McF) are probably related to high temperature anomalies in the uppermost mantle.

Fig. S7 (Supporting Information) shows three depth sections in our shear wave velocity model, along the Cifalps and ECORS-CROP reference profiles in the Western Alps, and the alpine segment of the EASI profile in the Eastern Alps (Hetényi *et al.* 2018b). These sections document strong along-strike variations in the internal structure of the crust and the Moho geometry.

## 5 DISCUSSION: FOCUS ON THE CIFALPS TRANSECT

This discussion is focused on the Cifalps cross-section in the SWA (location in Fig. 4a), which has been the target of a number of joint geophysical–geological studies in the last few years. The Cifalps transect is therefore a perfect case study to assess the reliability of our new crustal-scale model and to document its potential in terms of geological interpretation.

The Cifalps temporary seismic experiment (2012–2013) aimed at imaging the lithospheric structure of a part of the Alpine arc that had not been previously investigated by major geophysical studies (Zhao *et al.* 2015). It involved a dense quasi-linear profile of 46 broadband stations (5–10 km spacing) and 10 off-line stations (Zhao *et al.* 2016). The data set has been used in a number of seismic tomography and seismotectonic studies, which results have been interpreted in combination with gravity, tectonic, petrological and petrophysical data (see review in Malusà *et al.* 2021). The resulting lithospheric-scale geological–geophysical cross-section has therefore become one of the best-documented sections across the Alps. Based on  $P$  RE, Zhao *et al.* (2015) provided the first seismological evidence of continental subduction in the Alps, with the detection of  $P$ -to- $S$  converted waves at the European Moho at 75–80 km depth beneath the westernmost Po plain. Zhao *et al.* (2015) proposed an interpretative geological cross-section that was later refined based on the  $P$ -wave velocity model of Solarino *et al.* (2018) and on the  $S$ -wave velocity model of Zhao *et al.* (2020). The latter imaged a



**Figure 4.** Example of the inversion process for  $V_s$ . (a) Location map of the selected gridpoint (red triangle); For reference, the black line shows the Cifalps profile discussed in Section 5. (b) Group-velocity dispersion curve of Rayleigh waves (fundamental mode) for the selected grid point; black: observed dispersion curve with its uncertainties; red: predicted dispersion curve for the average Bayesian model and green: predicted dispersion curve after the subsequent linear inversion. (c) Resulting velocity models; left: posterior probability distribution on shear wave velocity from the Bayesian inversion (background colours), weighted average solution model (posterior mean) predicted from the probabilistic scheme (red line) and final solution model predicted from the linear inversion (black line). Right: posterior probability distribution on layer boundaries resulting from the Bayesian inversion. The two peaks outlined by red horizontal bars at 7–8 and 31–32 km depths correspond to the sediment basement and to Moho; the dashed red bars indicate the uncertainty on the boundary depth estimate approximated by the  $1\sigma$  width.

deep low-velocity anomaly at 50–70 km depth in agreement with the deep  $P$ -to- $S$  conversion in RF data.

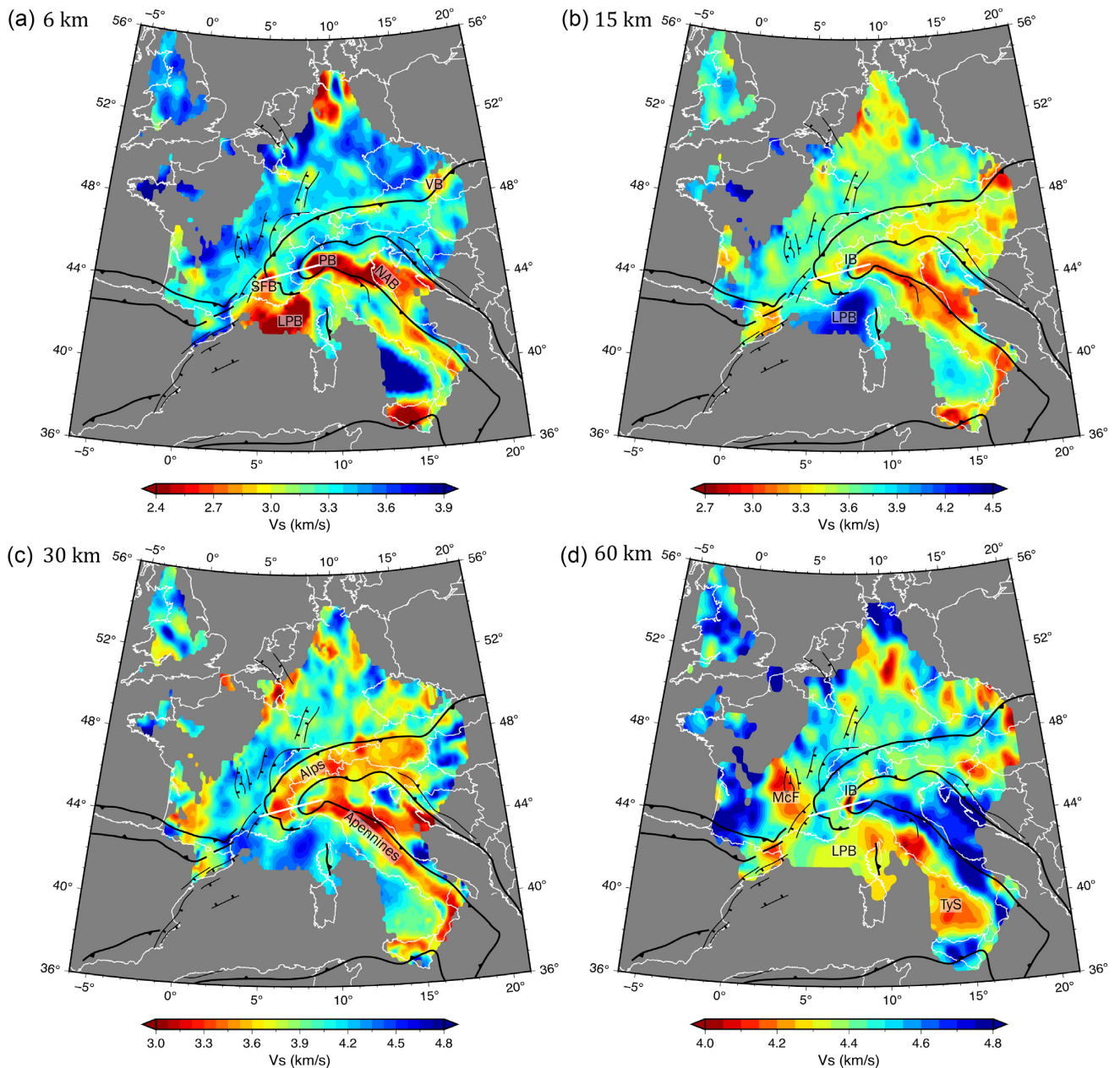
## 5.1 Model robustness

### 5.1.1 Data fit and uncertainty assessment

Figs 6(a) and (b) show sections of group velocity (as a function of period) and shear wave velocity (as a function of depth) along the Cifalps profile. In the western part of the cross-section (40–140 km distance range), the European continental crust is marked by low group velocities at 5–25 s period that result in low  $S$ -wave velocities over the entire crust, including a low-velocity zone (LVZ) in the lower part of the crust (white star in Fig. 6b). By contrast, the 180–270 km distance range is featured by anomalously high values both in group velocity ( $U > 3 \text{ km s}^{-1}$  in a broad period range) and in  $S$ -wave velocity ( $4.0 \text{ km s}^{-1}$  at 10 km depth). These anomalies correspond to the Ivrea high-velocity high-density body (IB) that was discovered in the early 1960s (Closs & Labrouste 1963). Our

$V_s$  model of the IB region is similar to the one computed by Zhao *et al.* (2020) using a different transdimensional inversion scheme. It also exhibits a deep LVZ with  $V_s \leq 3.9 \text{ km s}^{-1}$  at 50–80 km depth, as an extension of the lower part of the European continental crust (black star in Fig. 6b). The step geometry of the velocity contours on the eastern flank of the IB that may correspond to backthrusts of the collision belt is however a new feature that did not appear in the model of Zhao *et al.* (2020). This may indicate that our model is better resolved. Fig. 6(c) shows the probability densities for an interface to be present. The basements of the Southeast basin of France in the WSW and of the PB in the ESE are marked by strong velocity gradients, hence strong probability that an interface is present. This is also the case for most of the European Moho at 25–40 km depth, and for the very shallow Moho on top of the Ivrea mantle body that extends towards the ENE to the normal Adriatic Moho at a depth of  $\sim 35$  km beneath the PB.

As shown in Fig. 6(d), the misfit reduction between the observed dispersion curves and the dispersion curves computed from the final  $V_s$  model (dashed red line) is everywhere higher than the misfit reduction computed for the Bayesian model, used as initial model

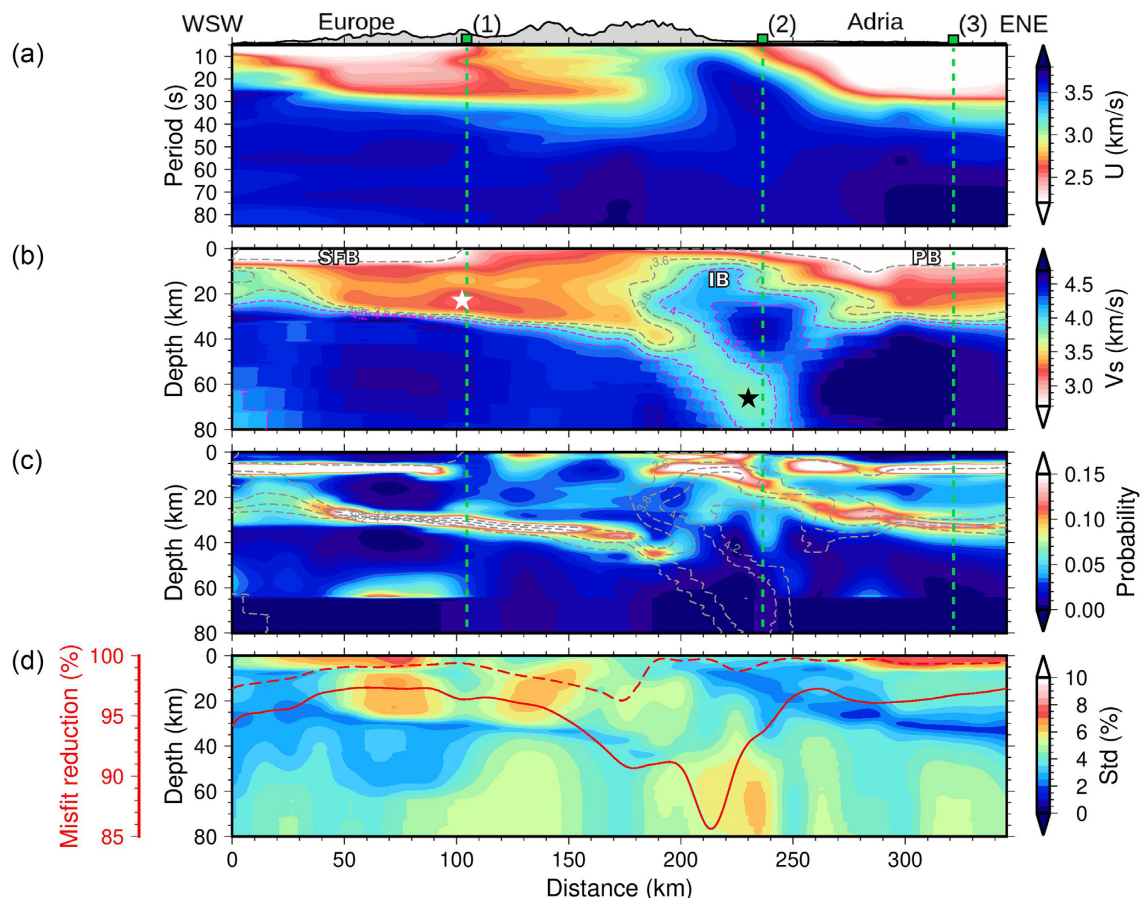


**Figure 5.** Depth slices in the final  $V_s$  model at 6, 15, 30 and 60 km. Only regions with  $1\sigma$  error < 8 per cent are shown. Black lines show the main geological boundaries as defined in Fig. 1(a). The white line shows the Cifalps profile discussed in Section 5. IB: Ivrea geophysical body, LPB: Liguro-Provençal basin, McF: Massif Central, NAB: North-Adriatic basin, PB: Po basin, SFB: Southeast-France basin, TyS: Tyrrhenian Sea and VB: Vienna basin.

for the linearized inversion (solid red line). The misfit reduction is strongest in the IB region (distance 170–240 km), where the uncertainty on the Bayesian  $V_s$  model is strongest at large depth (50–80 km). Fig. 6(d) also shows that sedimentary basins exhibit the largest uncertainties of  $\sim 7$ –8 per cent, while the smallest uncertainties are observed along the Adriatic Moho. A similar small uncertainty is observed along the European Moho up to 180 km distance where the Moho dip increases abruptly. These lateral changes in model uncertainty and misfit differences between the Bayesian and the final  $V_s$  model indicate that in the most complex areas none of the 130 million four-layer models that have been explored during the probabilistic inversion can explain completely the observed dispersion curve. In this case the linearized inversion is required to refine the  $V_s$  model.

To further investigate the robustness of our final model and its lateral variations, we analyse in Fig. 7 the outputs of 1-D inversions for  $V_s$  at three representative locations labelled (1), (2) and (3) in Figs 6(a) and (b). Fig. 7 shows first of all that dispersion curves computed from the final models (green curves) match observations well. The gain with respect to the results of the Bayesian inversion (red curves) is higher at pixels (1) and (2) (Figs 7a and b), which have a complex crustal structure that deviates more from the three-layer crust assumption than pixel (3) (Fig. 7c).

At pixel (1), the linearized inversion results in an LVZ at 15–30 km depth that better explains low group velocities at 15–25 s period. Pixel (2) is located in the westernmost Po plain on the eastern flank of the Ivrea high-velocity body. Compared to the average probabilistic model (red curve), the linearized inversion leads to



**Figure 6.** Depth sections of observations and inversion results along the Cifalps WSW-ENE profile (location in Fig. 4a). (a) Rayleigh-wave group velocities obtained from the Hierarchical Bayes reversible-jump tomography at gridpoints located along the profile. Green vertical dashed lines indicate locations of pixels (1), (2) and (3) discussed in the text. (b) Final shear wave velocity section. The 2.7, 3.6 and 3.8  $\text{km s}^{-1}$  velocity contours are shown as grey dashed lines while the pink dashed lines indicate the 4.0 and 4.2  $\text{km s}^{-1}$  velocity contours. The white and black stars mark low-velocity anomalies discussed in the text. (c) Posterior probability density of presence of a layer boundary obtained from the Bayesian inversion. The dashed lines show the same velocity contours as in (b). (d) Colour map: model uncertainties approximated for each location and at each depth by the  $1\sigma$  value of the ensemble of most probable models normalized by the  $V_s$  value at this gridpoint/depth. The thick red line indicates the total misfit reduction after the Bayesian inversion, while the red dashed line indicates the total misfit reduction after the linearized inversion. IB: Ivrea geophysical Body, PB: Po basin and SFB: Southeast-France basin.

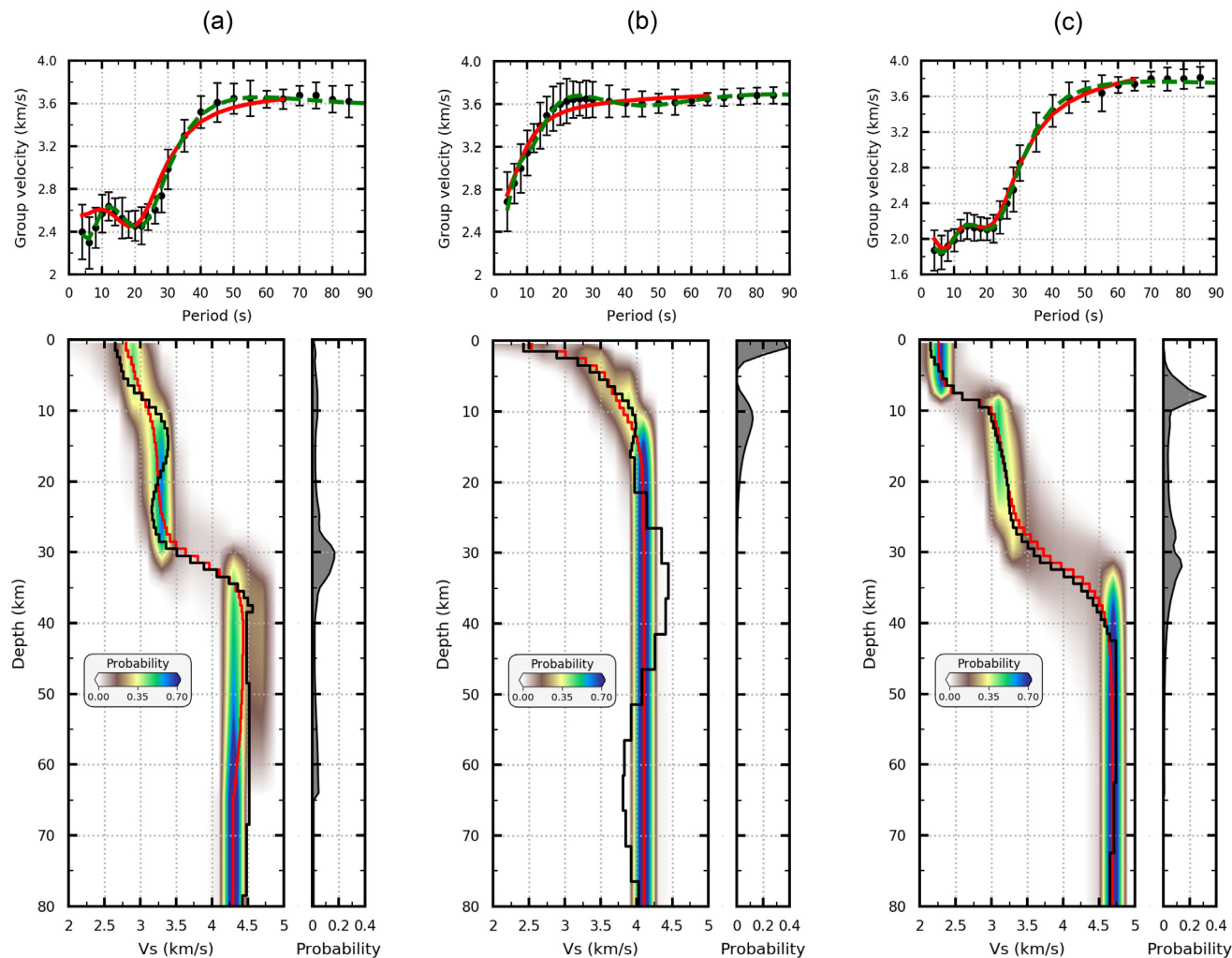
slightly higher velocities of  $4.0 \text{ km s}^{-1}$  at  $\sim 10 \text{ km}$ , even higher velocities at  $25 - 45 \text{ km}$  depth ( $4.4 \text{ km s}^{-1}$ ) and lower velocities ( $3.8 \text{ km s}^{-1}$ ) at  $52 - 77 \text{ km}$  depth. The final modelled dispersion curve (green curve in Fig. 7b, top) fits the observations better than the dispersion curve computed from the average probabilistic model (red curve). The example pixels (1) and (2) further confirm that the Bayesian inversion and its four-layer model assumption fail to explain such complex group-velocity data, while the final step of linearized inversion does a better job thanks to its fair initial model.

Such strong differences between the average probabilistic model and the final model are not observed at pixel (3) because its  $V_s$  model has a ‘normal’ layered structure with a positive velocity gradient (Fig. 7c). The very low velocities ( $\leq 2.2 \text{ km s}^{-1}$ ) from 0 to 7–8 km depth are the signature of the thick PB.

### 5.1.2 Benefits of our hybrid inversion scheme

Unlike Lu *et al.* (2018) and most ANTs, we used a 2-D transdimensional inversion for group-velocity maps to avoid biases arising from

explicit regularization. The Bayesian nature of this inversion also allows for a robust quantification of the uncertainties on group velocities and hence local dispersion curves. Unlike Zhao *et al.* (2020) who used a full transdimensional approach in the inversion for  $V_s$ , we performed an exhaustive grid search over all possible four-layer models, followed by a linear inversion. This hybrid methodology limits the complexity of resulting  $V_s$  model while giving enough freedom to the inversion to actually fit the data. It also provides uncertainties on  $V_s$  and the probability of having an interface at given depths. To document the benefits of our approach, Fig. S8 (Supporting Information) and its comments compare our model to the one by Zhao *et al.* (2020) on depth sections along the Cifalps and ECORS-CROP reference profiles. The two models exhibit differences that can be attributed to the improved data coverage in the western part of the study region, or to the inversion scheme. The results of our inversion is more in line with previous geophysical data. Moreover, the hybrid character of our inversion makes it much less computationally expensive than pure transdimensional inversion. Therefore, we could compute a robust 3-D  $V_s$  model at the scale of Western Europe from such a large volume of data using reasonable computational resources.



**Figure 7.** Results of the inversion for  $V_s$  at pixels 1 (panels a), 2 (panels b) and 3 (panels c). Pixel locations are indicated in Fig. 6a). The top panels show observed (black) and predicted dispersion curves (red: from the Bayesian inversion; green: from the linear inversion). The bottom panels show: (left) probability density of  $V_s$  (colour map) and predicted  $V_s$  models (red: probabilistic average model; black: final model) and (right) probability density of interface presence.

## 5.2 Comparison with other geophysical data

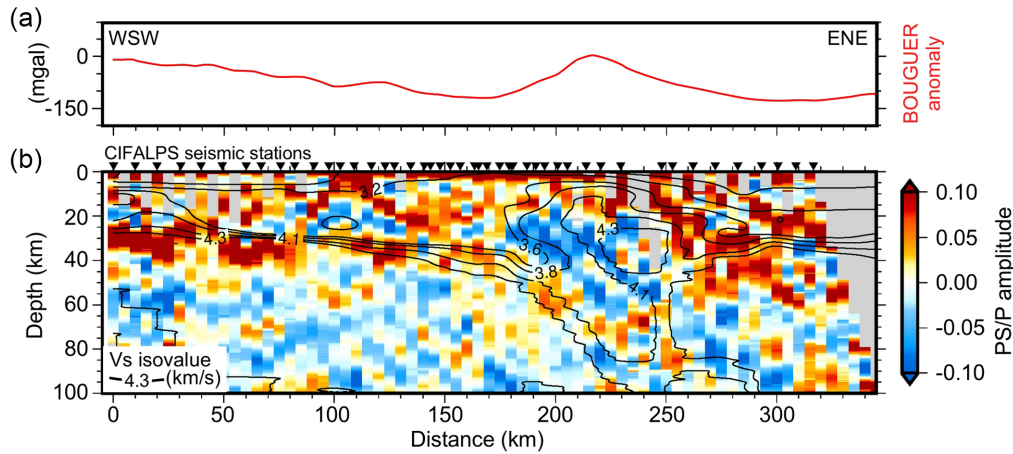
In Fig. 8, we compare the depth section through our  $V_s$  model along the Cifalps profile with the Bouguer anomaly (Fig. 8a) and the RF CCP migrated depth section (common conversion point; Zhu 2000) computed from the Cifalps data set (Zhao *et al.* 2015) and migrated using our  $V_s$  model (Fig. 8b).

An RF section displays teleseismic  $P$ -to- $S$  converted waves ( $P_s$ ) at velocity discontinuities beneath the stations, with positive amplitudes indicative of velocity increase with depth, and negative polarities indicative of velocity decrease with depth (Vinnik 1977). Unlike (Zhao *et al.* 2015) who used a four-segment 1-D migration model, we used the 2-D  $V_s$  model derived from our ANT (shown as velocity contours in Fig. 8b) to perform the time-to-depth CCP migration of RFs.

The maximum of the Ivrea gravity anomaly (positive Bouguer anomaly in Fig. 8a) coincides well with the top of the high- $V_s$  anomaly (3.8 and 4.1  $\text{km s}^{-1}$  contours in Fig. 8b). The consistency between the  $V_s$  model and the CCP section is outstanding, considering that the sensitivity of surface waves to sharp velocity boundaries is rather weak. The 4.2 and 4.3  $\text{km s}^{-1}$  velocity contours are fairly consistent with the European Moho, which is characterized

by a clear  $P_s$  phase of positive polarity at distances 0–200 km. The agreement is not that good at 200–250 km distance where the 4.2–4.3  $\text{km s}^{-1}$  contours are deeper than the deepest  $P_s$  phases at 70–80 km depth that were interpreted by Zhao *et al.* (2015) as evidence of continental subduction of Europe beneath Adria. The RF data quality is poorer under the PB, making comparison with  $V_s$  contours difficult. In addition, the 3.6 and 3.8  $\text{km s}^{-1}$  contours match the negative polarity  $P_s$  phases observed above the European Moho at distances 180–250 km, which were interpreted by Zhao *et al.* (2015) as traces of an ‘inverted’ Moho between the Ivrea mantle body on top and the European crust below. We conclude from the match of the 4.2–4.3  $\text{km s}^{-1}$  velocity contours with the European Moho that these contours are an accurate proxy of the autochthonous Moho, while the 3.8 or 4.0  $\text{km s}^{-1}$  contours are more adequate to characterize the Moho at the top of the serpentinized IB.

We have seen in Fig. 6(c) that the probability of presence of an interface is a reliable marker of a strong velocity contrast. For instance, the probability is lower for the Adriatic Moho than for the European Moho, due to a lower velocity gradient on the Adriatic side (Figs 7a and c). This correlation with the velocity gradient is however not valid everywhere, as for instance at the top of the IB



**Figure 8.** Comparison of our  $V_s$  model along the Cifalps line with independent geophysical data. (a) Bouguer gravity anomaly extracted from the AAGRG model (AlpArray Gravity Research Group; Zahorec *et al.* 2021). (b) RF CCP depth section migrated using our  $V_s$  model; the black plain lines are the 2.7, 3.2, 3.6, 3.8, 4.1 and 4.3  $\text{km s}^{-1}$  contours of our  $V_s$  model; Locations of stations projected onto the profile are shown as black inverted triangles (station map shown in Fig. 9a).

anomaly where the probability is high while the gradient is rather smooth (Fig. 7b). Hence, the probability of presence of an interface is an accurate proxy not only for the base of sedimentary basins, but also for delineating the crust–mantle boundary of Adria and Europe, and highlighting interfaces with weak velocity gradient such as the top of the IB. However, Fig. 6(c) shows that the probability section is almost blind inside the subduction complex due to its very complex structure, which means that interface probability cannot be considered as a universal proxy for Moho. By contrast, the comparison between the  $V_s$  section and the RF section has confirmed that the 4.2 or 4.3  $\text{km s}^{-1}$  iso-velocity contour is better suited to delineate deep discontinuities associated with the subduction of the European lithosphere. In future studies, we will therefore rather exploit the 4.3  $\text{km s}^{-1}$  contour as a proxy of Moho to delineate its 3-D depth variations.

### 5.3 Crustal-scale geological interpretations

In Fig. 9, we propose a geological interpretation of the shear wave velocity section along the Cifalps profile that also builds on surface geological observation, on previous joint geophysical–geological works summarized in Malusà *et al.* (2021), on the RF section of Fig. 8(b) and on hypocentre locations by Eva *et al.* (2015) and Malusà *et al.* (2017).

In the following subsections, we emphasize a number of first-order features related to the geometry of structures at crustal scale, to the deformation and thermicity of the lithosphere, and to the link between the lithospheric structure and the strain field.

#### 5.3.1 Geometry of structures at lithospheric scale

The geology of the Western Alps (Fig. 9a) results from the convergence of two lithospheric plates (Polino *et al.* 1990; Dumont *et al.* 2012). To the west, the European margin corresponds to the lower plate, which subducted beneath the Adriatic upper plate to the east. Between these two continental domains, the subduction wedge developed by the juxtaposition of oceanic sediments and crustal fragments (Lardeaux *et al.* 2006; Schwartz *et al.* 2009; Agard 2021) with units originating from the two continental margins (Fig. 9a). This prism is bounded by two major crustal-scale faults: to the west

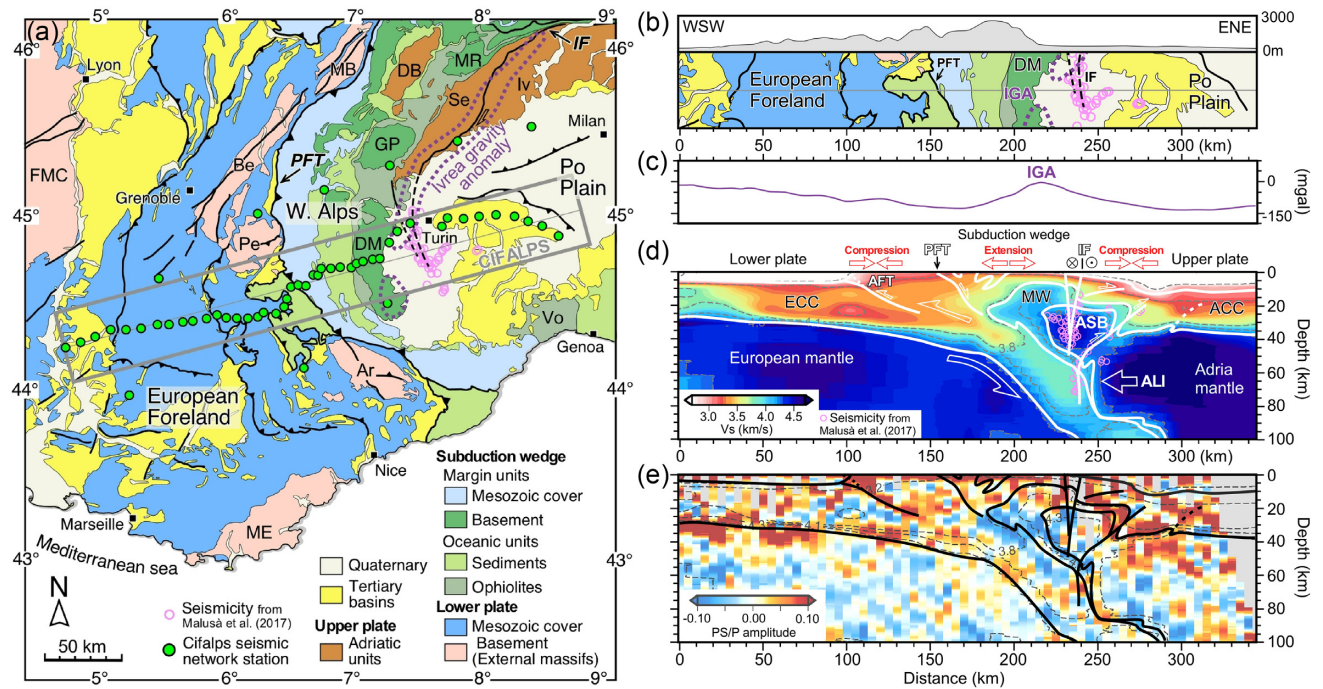
the PFT and to the east the dextral strike-slip Insubric Fault (IF). The European foreland accommodates the convergence through the propagation of a fold-and-thrust-belt system associated with the development of Cenozoic sedimentary basins (Ford *et al.* 1999). These basins are further deformed, uplifted and partially eroded during the westward propagation of compressional structures. This compressional deformation is localized at depth along crustal-scale thrusts leading to the exhumation of the External Crystalline Massifs in thick skin tectonic mode (e.g. Jourdon *et al.* 2014; Schwartz *et al.* 2017). To the East, the Po plain is affected by top-to-northeast back-thrusts allowing the uplift and erosion of Cenozoic basins (Mosca *et al.* 2010).

The  $V_s$  model documents the subduction of the European lithosphere towards the east below the Adria lithospheric mantle, marked by a continuous European continental crust down to a minimum depth of 100 km. This continental crust is featured by a progressive increase of  $V_s$  with depth from 3.4–3.5  $\text{km s}^{-1}$  at 20–30 km to 4.2  $\text{km s}^{-1}$  at 100 km interpreted as resulting from the densification of continental rocks during their progressive eclogitization.

Under the westernmost part of the European foreland, high crustal  $V_s$  values of 3.6–3.8  $\text{km s}^{-1}$  and a slight Moho uprise (iso-velocity contour  $V_s = 4.3 \text{ km s}^{-1}$ ) indicate crustal thinning beneath the Rhône valley consistently with the formation of the Western European Cenozoic Rift System (Ziegler 1992).

One of the most striking features of the section is a standard crustal thickness of  $\sim 30$  km below the frontal part of the External Massifs, increasing to  $\sim 50$  km below the PFT. This observation is in line with other recent geophysical models along the Cifalps profile that exhibit very little crustal thickening of the External Alps (e.g. Malusà *et al.* 2021).

Beneath the Ivrea Gravity Anomaly (IGA in Fig. 9c), located along the IF and extending southwards beneath the Internal Crystalline Massif of Dora Maira (Figs 9 a and b), the upper part of the subduction wedge is likely formed by stacking of crustal slices extracted from both the continental and oceanic crusts. This subduction wedge overlies the MW (Fig. 9d) at a depth of about 10 km. This MW, documented by shear wave velocities ranging between  $\sim 3.8$  and 4.0  $\text{km s}^{-1}$ , may be composed of partially serpentinized mantle of Adriatic origin produced by hydration above the subducted oceanic units. Below the subduction wedge, a superposition of several mantle bodies with different geophysical signatures is observed,



**Figure 9.** Geological and tectonic settings of the study area along the Cifalps transect. (a) Geological map of the W-Alps with location of the Cifalps seismic network (white rectangle in Fig. 1b). Earthquake epicentres with depth  $> 20$  km beneath the westernmost Po plain are shown as pink open circles (after Eva *et al.* 2015; Malusà *et al.* 2017). The Ivrea gravity anomaly is represented by the 0 mGal contour (dashed purple line) from Bigi *et al.* (1990). The continental crust units of the European plate (lower plate) are indicated by: Ar, Argentera; Be, Belledonne; MB, Mont-Blanc; Pe, Pelvoux; ME, Maures-Estérel; FMC, French Massif Central. The Adriatic units (upper plate) are indicated by: DB, Dent Blanche; Iv, Ivrea-Verbanò; Se, Sesia. The internal crystalline massifs of the subduction wedge are indicated by: DM, Dora Maira; GP, Gran Paradiso; MR, Monte Rosa; two major tectonic limits are also indicated: IF, Insubric Fault; PFT: Penninic Frontal Thrust. (b) Topographic profile and geological map [extracted from (a)] along the Cifalps transect. (c) Bouguer gravity anomaly extracted along the profile (same as Fig. 8a). IGA: Ivrea Gravity Anomaly. (d) Crustal-scale interpretation of the  $V_s$  depth section along the Cifalps transect. The main geological interfaces corresponding to lithological units and tectonic structures are underlined in white (see interpretation in text). This interpretation shows the subduction of the European lithosphere toward the East below the Adria lithospheric mantle. The European continental crust (ECC) is continuous down to 100 km depth. We observe separation of the Adria lithospheric mantle into two main bodies: (1) the deeper part ALI acts as a horizontal indenter that deformed the European slab and enabled the activation of the AFT. (2) The upper part ASB acts as a vertical indenter controlling the exhumation of the MW and locating the deep seismicity along the IF system. In this interpretation, the ACC is deformed by top-to-northeast thrusts. In the global tectonic convergence, the crustal-scale geometry deduced from the  $V_s$  model controls the partitioning of the strain field (red arrows). The crustal compartment located between the IF and PFT tectonic structures records extension while the two other compartments record compression. (e) Crustal-scale interpretation plotted as black lines on the RF CCP depth section of Fig. 8(b).

which suggests the superposition of two petrological Mohos: a first one characterized by a reverse  $V_s$  gradient with a negative polarity in the RF section (Fig. 9e). This ‘inverted Moho’ corresponds to the subduction interface between the European slab and the MW. The second, deeper Moho corresponds to the European Moho itself, steeply dipping due to subduction.

### 5.3.2 Deformation and thermicity

The deformation is dominated by compressive structures occurring in both forelands of the chain, with top-to-the-west thick-skinned propagation on the western side of the belt, rooted in the lower crust. A major thrust is well marked in the RF section as a set of positive-polarity signals down to 25 km depth that connects at the surface with the thrust (Fig. 9e). We name it the Alpine Frontal Thrust (AFT). Similar structures occur on the eastern side of the Alps, below the Po plain, which either do not outcrop or may merge with the frontal Po thrust system. The rooting of these peripheral thrusts appears to occur at different depth levels. We propose that the AFT becomes horizontal in the lower crust, where  $V_s$  has very low values  $< 3.4$  km s $^{-1}$  (Fig. 9d), which would be consistent with

the more ductile rheology of the European lower crust. In contrast, the step-like shape of the  $V_s$  contours in the eastern side suggests that the Adriatic Moho and continental crust (ACC, Fig. 9d) are significantly offset by the main top-to-northeast thrusts. This difference in deformation style between the European crust and the Adria crust is in agreement with a different thermal state, revealed by different  $V_s$  layouts in the lithospheric mantle domains. The higher and more homogeneous shear wave velocities observed in the Adria mantle may indicate a colder state leading to a more rigid behaviour. In this context, the Adria lithospheric mantle behaves as a deep horizontal indenter (the Adria Lithospheric Indentor or ALI, Fig. 9d) buttressing the European slab along a 25 km-high vertical boundary.

As a result of this deep indentation, the MW is vertically indented by an intermediate, non-serpentinized rigid mantle body with  $V_s \sim 4.2 - 4.4$  km s $^{-1}$  that localizes most of the deep seismicity ( $> 20$  km) below the westernmost Po plain (pink open circles in Figs 9a, b and d; Eva *et al.* 2015; Malusà *et al.* 2017). This rigid body is named Adria Seismic Body (ASB). We relate the vertical seismic swarm in the ASB to the rooting zone of the IF.

### 5.3.3 Link between the geometry of lithospheric structures and the strain field

In the framework of plate tectonics and convergence between Europe and Adria, the observed split of the Adriatic lithospheric mantle into two units (ASB and ALI) would be responsible for the partitioning of the deformation at lithospheric scale. The deeper part of the Adria mantle (ALI) horizontally indents the European slab. This results both in lithospheric loading and in lithospheric shear deformation propagating westwards to reach the upper crust beneath the European foreland, corresponding to the activation of the AFT. The upper part of the Adria mantle (ASB) was pushed upwards and maintained in a high structural position by the ALI horizontal indentation. In turn, the ASB acted as a vertical indenter responsible for the deformation and exhumation of the subduction wedge units.

This geometry leads to a partitioning of the active deformation on both sides of the IF (Fig. 9d). To the east of the IF, the subduction wedge is backthrust onto the Adria crust and the Adria lithosphere is significantly shortened (Mathey *et al.* 2020). In contrast, to the west of the IF, extensional deformation dominates in a crustal compartment bounded by the IF and the PFT (Mathey *et al.* 2020), as a result of vertical indentation. This latter thrust was reactivated since at least 3–4 Ma, and it is still currently undergoing extensional deformation (Bilau *et al.* 2021).

## 6 CONCLUSION

The high density and rather homogeneous coverage of permanent European broadband networks complemented with the AlpArray temporary network and quasi-linear dense arrays such as Cifalps, Cifalps-2 and EASI were key in our derivation of a high-resolution shear wave velocity model of the Alpine mountain belt and its forelands. We used four years of continuous seismic noise recorded by ~1550 broad-band seismic stations, which led to 1.1 million cross-correlations of vertical noise records. Rayleigh-wave group-velocity dispersion curves were measured in the 4–150 s period band and 2-D group-velocity maps and their uncertainties were computed using a data-driven transdimensional inversion that allows the local resolution to self-adapt to the path density and to the variability of information contained in group-velocity measurements. At each location of our  $0.15^\circ \times 0.15^\circ$  grid, the local group-velocity dispersion curve and its associated uncertainties were inverted jointly in a Bayesian approach to derive a quasi-3-D probabilistic shear wave velocity model. This model was further refined using a linear inversion.

The best-resolved part of our 3-D  $V_s$  model covers the Alpine range, its forelands and the Ligurian Sea. We showed on the Cifalps traverse of the SWA that the resolution of our model makes it possible to correlate features of the geophysical model with geological structures mapped at the surface. This enabled an integrated geological–geophysical interpretation that highlights first-order features related to the subduction of the European lithosphere under the Adriatic lithosphere in the SWA. One of these key features is the IB, which is likely made of a subduction wedge formed by stacking of crustal slices of continental and oceanic origin overlying an MW likely composed of partially serpentinized Adriatic mantle. Another key feature is the split of the Adriatic mantle lithosphere in two different units. This split could be responsible for the partitioning of the deformation at lithospheric scale.

Following the model verification at the location of the 2-D Cifalps cross-section, our velocity model makes it possible to construct a

3-D, lithospheric-scale geological models of the Alpine belt, extending the geological maps in depth. These models are expected to provide key constraints to conceptual geological–petrophysical models of subduction wedge systems, for example. Such interpretation work has already started for the Western Alps.

Our  $V_s$  model can also be used in further geophysical studies, including waveform modelling and full waveform inversion. Our next project is to use this model as an initial model to simultaneously build a 3-D  $V_p$  and  $V_s$  model of the Alpine lithosphere by joint inversion of surface-wave dispersion data from noise correlations and traveltimes of body waves emitted by local earthquakes using a 3-D wave-equation tomography. This combination of  $P$ - and  $S$ -wave velocities is required to establish the lithology of rocks buried at large depths.

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#### DATA AVAILABILITY STATEMENT

Our tomographic 3-D model will be available in the Résif products repository (Réseau sismologique et géodésique Français; <https://www.resif.fr/en/data-and-products/products-repository/>). The inversion code is available on request to the first author.

Waveform data used in this paper belong to the permanent networks with codes AC, BE (Royal Observatory Of Belgium 1985), CA (Institut Cartogràfic I Geològic De Catalunya—Institut D’Estudis Catalans 1984), CH [Swiss Seismological Service (SED) at Swiss Seismological Service (SED) At ETH Zurich (1983)], CR (University Of Zagreb 2001)), CZ (Institute Of Geophysics Of The Academy Of Sciences Of The Czech Republic 1973), ES (Instituto Geográfico Nacional, Spain 1999), FR ((RESIF, RESIF 1995), G [(Institut de Physique du Globe de Paris (IPGP) & Ecole et Observatoire des Sciences de la Terre de Strasbourg (EOST) 1982)], GB [(GEOFON Data Centre 1993)], GR [(Federal Institute for Geosciences & Natural Resources (BGR) 1976)], IV ((INGV Seismological Data Centre 2006)), NL (KNMI 1993), OE (ZAMG-Zentralanstalt Für Meteorologie Und Geodynamik 1987), SL (Slovenian Environment Agency 2001) and UP (SNSN 1904). We also used data of the temporary AlpArray network (AlpArray Seismic Network 2015), Cifalps (YP network, Zhao *et al.* 2016) and XT network Zhao *et al.* 2018) and EASI experiments (XT network, AlpArray Seismic Network 2014).

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## SUPPORTING INFORMATION

Supplementary data are available at [GJI](https://doi.org/10.1016/j.gji.2022.06.022) online.

**Figure S1.** Distance–time plot of noise correlations.

**Figure S2.** Comparison of the 8-s and 25-s group-velocity maps with sedimentary basins and Moho depth maps.

**Figure S3.** Comparison with a matrix-based inversion.

**Figure S4.** Posterior distribution on the number of Voronoi cells.

**Figure S5.** Inversion misfit.

**Figure S6.** Checkerboard tests.

**Figure S7.** Depth sections through the  $V_s$  model along reference profiles.

**Figure S8.** Comparison with the  $V_s$  model by Zhao *et al.* (2020).

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