

# Coupled hydrogen and fluorine incorporation in garnet: New constraints from FTIR, ERDA, SIMS, and EPMA

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1	Revision 1
2 3 4 5 6	Coupled hydrogen and fluorine incorporation in garnet: new constraints from FTIR, ERDA, SIMS, and EPMA
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### 28

### ABSTRACT

29	It is well known that some garnet compositions can incorporate hydrogen and/or fluorine
30	at levels up to several weight percent. However, accurate measurement of these elements can be
31	difficult at trace to minor levels of abundance so they are frequently ignored in routine chemical
32	analysis. Furthermore, the mechanisms of H incorporation are still under debate and only one
33	mechanism for F substitution is commonly considered. We employed infrared spectroscopy
34	(FTIR), elastic recoil detection analysis (ERDA), secondary ion mass spectrometry (SIMS), and
35	electron probe microanalysis (EPMA) to measure H and F concentrations and constrain
36	incorporation mechanisms in ten grossular garnets. We also present SIMS data for 11 spessartine
37	and two andradite garnets. Three grossular garnets were measured with ERDA to obtain an
38	infrared integral molar absorption coefficient ( $\epsilon_i$ ) for H <sub>2</sub> O of 13,470 L $\cdot$ mol <sup>-1</sup> $\cdot$ cm <sup>-2</sup> . Grossular
39	$\rm H_2O$ and F concentrations range from 0.017 to 0.133 wt% and 0.012 to 0.248 wt%, respectively.
40	Correlations between <sup>16</sup> OH and <sup>19</sup> F and interpretation of FTIR spectra prompt us to consider
41	various coupled substitutions of H and F for Si, which can explain some high frequency IR
42	absorption bands that have been attributed previously to "hydrogrossular clusters" (variably-
43	sized clusters in which 4H substitute for Si) or to inclusions of hydrous minerals. A strong
44	correlation between <sup>16</sup> OH and <sup>19</sup> F in spessartine and similar high-frequency IR bands implies a
45	similar role for H-F substitution. Coupled H-F substitution is also probably relevant to some
46	andradite-rich garnets, rare pyrope from the Dora Maira massif, and some synthetic garnets.
47	Improvements in analytical methods for trace to minor H and F open up more possibilities for
48	using these elements to calculate the activities of H <sub>2</sub> O and F-species in fluids that were in

equilibrium with garnet-bearing phase assemblages, as well as constraining the recycling of theseelements into the mantle via study of xenoliths.

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- 52

### INTRODUCTION

53 Considerable progress has been made in the last three decades in understanding how trace 54 amounts of hydrogen can be incorporated in nominally anhydrous minerals, and how this H 55 (most commonly incorporated as structurally bound OH<sup>-</sup> groups) affects geophysical and 56 petrologic processes in the Earth and other planetary bodies. Recent work has also highlighted the potential importance of trace fluorine (substituting for O<sup>2-</sup>) in nominally anhydrous minerals 57 58 (Hervig and Bell 2005; Mosenfelder et al. 2011, 2015; Beyer et al. 2012; Bernini et al. 2012; 59 Dalou et al. 2012; Mosenfelder and Rossman 2013a,b; Crépisson et al. 2014; Roberge et al. 60 2015; Grützner et al. 2017; Klemme and Stalder 2018; Yoshino and Vazhakuttiyakam 2018). In 61 the case of garnets, a diverse supergroup of minerals widely distributed in the crust and mantle of 62 the Earth, it has long been recognized that both H and F can be structurally incorporated at levels 63 up to several weight percent (see Grew et al. 2013 for review). Researchers have taken advantage of this phenomenon to constrain the activities of H<sub>2</sub>O and F-species in fluids in equilibrium with 64 garnet-bearing phase assemblages (Manning and Bird 1990; Visser 1993; Arredondo et al. 2001; 65 66 Chakhmouradian et al. 2008).

Hydrogen and/or F may in fact be almost ubiquitously present in trace to minor amounts
in garnets, but are commonly ignored during chemical analysis because of difficulties inherent in
measuring trace concentrations of light elements with commonly available techniques such as
electron probe microanalysis (EPMA). Note that when we specify "H<sub>2</sub>O" or "H" concentrations

71	in this paper we are primarily referring to the equivalent in structurally bound OH <sup>-</sup> groups,
72	because H is not incorporated in the form of structurally-bound H <sub>2</sub> O groups in garnet.
73	Hydrogen can be detected easily using Fourier transform infrared spectroscopy (FTIR) in
74	transmission mode, which also provides information on incorporation mechanisms, but
75	uncertainties remain in absolute quantification using this method (Maldener et al. 2003; Rossman
76	2006). One of the best microanalytical techniques for trace F is secondary ion mass spectrometry
77	(SIMS), which is highly sensitive owing to the ionization efficiency of both ${}^{19}\text{F}^-$ and ${}^{19}\text{F}^+$ and has
78	benefited from technical advances that lower detection limits to less than 1 $\mu$ g/g (Hauri et al.
79	2002; Koga et al. 2003; Mosenfelder et al. 2011). However, these measurements are most
80	commonly made using F-bearing glasses for calibration, which can result in working curves with
81	large uncertainties (Mosenfelder and Rossman 2013a).
82	SIMS measurements of H in garnets have primarily concentrated on pyrope-rich
83	compositions representative of the mantle (Koga et al. 2003; Aubaud et al. 2007; Tenner et al.
84	2009), with one more recent study on grossular and spessartine (Reynes et al. 2018). Quantitative
85	SIMS analyses of F in garnets are even more sparse (Jamtveit and Hervig 1994; Schingaro et al.
86	2016). In this study, we present new SIMS, FTIR, and EPMA data on the H and F concentrations
87	of ten natural grossular garnets, and SIMS data on 11 spessartine and two andradite crystals.
88	Furthermore, based on elastic recoil detection analysis (ERDA) of three grossular samples, we
89	reassess calibrations for the IR molar absorption coefficient for $H_2O$ derived by Rossman and
90	Aines (1991) and Maldener et al. (2003). We demonstrate that H and F concentrations are
91	strongly correlated in some garnets and suggest coupled H-F incorporation mechanisms
92	analogous to those proposed for olivine by Crépisson et al. (2014). Coupled H-F substitution can
93	rectify unexplained aspects of previous incorporation models for H in grossular, spessartine, and

94	pyrope (Aines and Rossman 1984; Rossman and Aines 1991; Cho and Rossman 1993; Reynes et
95	al. 2018; Geiger and Rossman 2018, 2020a, 2020b).
96	
97	ANALYTICAL METHODS
98	Sample selection and preparation
99	Table 1 lists the localities of the grossular garnets chosen for this study. Six of these
100	garnets, derived from granulites of the Neoproterozoic Mozambqiue belt in Tanzania and Kenya
101	(Giuliani et al. 2008, 2011), are referred to as "East African" in this paper. Additional details on
102	the grossular, spessartine, and andradite samples used in the study are provided in the
103	supplementary material.
104	Samples were prepared for analysis using previously described, epoxy-free polishing
105	methods and cleaning procedures (Mosenfelder et al. 2011). The garnets were fabricated into
106	doubly polished slabs that were first measured by FTIR and then sliced using a wire saw into
107	multiple chips for analysis by different techniques; in all cases, individual chips originated from
108	the same single crystal.
109	
110	FTIR
111	Unpolarized infrared absorption spectra were collected for most of the garnets using the
112	main compartment of either a Nicolet Magna 860 FTIR or Thermo-Nicolet iS50 FTIR at Caltech.
113	All measurements employed a GLOBAR IR source, a KBr beamsplitter, and a MCT-A detector,
114	with at least 512 scans taken at 2 cm <sup>-1</sup> resolution. Analysis areas were selected using 200 to 400
115	$\mu$ m diameter circular apertures and multiple spectra for each crystal were taken to assess the
116	possibility of zoning. Sample thicknesses were measured using a Mitutoyo digital micrometer,

117	with a precision of $\pm 2 \ \mu m$ . JLM83a was measured at UMN using a Hyperion 2000 microscope
118	attached to a Bruker Tensor 37 spectrometer, with a GLOBAR IR source, a KBr beamsplitter,
119	and a MCT-A detector. These spectra were collected on a slab ${\sim}390~\mu m$ thick, with 64 scans at 2
120	$\text{cm}^{\text{-1}}$ resolution, using a 100 $\mu\text{m}^2$ square knife-edge aperture. The video-assisted mapping feature
121	of the OPUS 7.2 software was employed to take regularly spaced data (in a 6 x 8 grid) across the
122	central portion of the sample slab.
123	The low temperature (77 K) spectrum of GRR1386 was obtained in 1994 using a
124	homebuilt vacuum chamber in the main compartment of the Nicolet 60SX FTIR that was
125	operational at that time at Caltech. The room temperature spectrum of the sample taken in 1994
126	closely matches more recently acquired spectra.
127	Details of our baseline correction methods are provided in the supplementary material.
128	After baseline correction, we calculated total integrated band areas in the O-H stretching region,
129	referred to hereafter as $Abs_{tot}$ (considered here as absorbance/cm <sup>2</sup> in one direction, rather than
130	multiplying one measurement by three as in some studies on garnets). Spectra were then curve
131	fitted using the Multipeak Fitting 2 package in Igor Pro software. Table 1 also gives values for
132	the weighted mean wavenumber (as defined by Libowitzky and Rossman 1997) for each
133	spectrum.
134	
135	ERDA
136	In Table 2 we report hydrogen concentrations measured by ERDA for three grossular
137	samples (GRR732, GRR771, and GRR1756), one synthetic forsterite (GRR1017), and three
138	synthetic, rhyolitic glasses (NSL1, N3, and N5). The analyses were conducted at the Laboratoire
139	d'Etudes des Eléments Légers, CEA, Saclay, France using improved methods for micro-ERDA

140 (Raepsaet et al. 2008, Bureau et al. 2009). Analytical methods followed those outlined by 141 Withers et al. (2012) and are detailed in the supplementary material. GRR1017, well-142 characterized as an ultra-dry reference material with  $\leq 0.007 \ \mu g/g H_2O$  (Mosenfelder et al. 2011), 143 was used to assess the H background. Water contents of the rhyolitic glasses were previously 144 reported in Tenner et al. (2009), measured using Karl-Fischer titration as described in Withers 145 and Behrens (1999). 146 147 SIMS SIMS data (Table 3) were obtained on the CAMECA 7f-GEO at Caltech using a  $Cs^+$ 148 149 primary ion beam and previously described methods optimized for collecting low-blank H and F 150 analyses (Mosenfelder and Rossman 2013a; Mosenfelder et al. 2011, 2015). Data are presented 151 from two sessions, conducted in 2012 and 2016. For each analysis, we acquired either 30 cycles (in 2012) or 20 cycles (in 2016) through the mass sequence  ${}^{12}C$ ,  ${}^{16}O^{1}H$ ,  ${}^{18}O$ ,  ${}^{19}F$ , and  ${}^{30}Si$ , 152 collecting negative ions with a mass resolution of ~5500 ( $\Delta M/M$ ), sufficient to separate <sup>16</sup>O<sup>1</sup>H 153 from <sup>17</sup>O and <sup>18</sup>O<sup>1</sup>H from <sup>19</sup>F (Burdo and Morrison 1971). Analyses from 2012 were obtained 154 155 from garnets pressed into the same indium mount as the orthopyroxenes, clinopyroxenes, 156 olivines, and F-bearing glasses (MPI-DING glasses KL2-G and ML3B-G and USGS glasses 157 BCR-2G and BHVO-2G) reported on previously (and analyzed during the same session) by 158 Mosenfelder and Rossman (2013a,b). Some of the garnets were then extracted, re-polished, 159 cleaned, and pressed together into a new indium mount with one additional garnet (JLM83a), the 160 MPI-DING and USGS glasses, and five additional F-bearing glasses (Fba-1, 2, 3, 4 and 5) 161 prepared by Guggino et al. (2011). We used the Fba-series glasses to reevaluate the F 162 concentrations of our MPI-DING and USGS glass splits, as described in Mosenfelder et al.

163 (2020). The revised reference values used to establish working curves for F are given in the164 supplementary material.

165	$^{16}O^{1}H/^{18}O$ and $^{19}F/^{18}O$ ratios in Table 3 were blank corrected using $^{16}O^{1}H$ and $^{19}F$
166	backgrounds monitored by analyzing "blank" reference materials: synthetic forsterite GRR1017
167	(Mosenfelder et al. 2011), a natural Cr-rich pyrope from Ugelvik, Norway (GRR332), and
168	laboratory-dehydrated grossular (GRR1122-HT) and spessartine (GRR2215-HT). Details on
169	characterization of these samples are provided in the supplementary material.
170	
171	EPMA
172	Electron microprobe analyses (Table 4) for most samples were acquired using a JEOL
173	JXA-8900R at UMN. We first measured F, using $MgF_2$ as a standard; other elements were
174	subsequently analyzed on the same spots using different analytical conditions. F k $\alpha$ was acquired
175	on three TAP crystals simultaneously and intensities were then aggregated and combined with

177 Despite its lower sensitivity for F Ka, we prefer TAP over LDE1 for this application, because

178 LDE1 has the inherent problem of significant background interferences between the F Kα and Fe

179 K $\alpha$ , Mg K $\beta$ , and Mn K $\alpha$  X-ray lines (e.g., Witter and Kuehner 2004; Zhang et al. 2016). We

180 found that an accelerating voltage of 20 kV optimized fluorescence yields and peak-to-

181 background ratios for F on TAP. For most of the garnets, a beam current of 250 nA and beam

182 diameter of 10 µm was used for this stage of the analysis. Analyses of JLM83a were conducted

183 at either 200, 250, or 400 nA. Calculated detection limits range from ~40 at the highest current to

184  $60 \mu g/g$  at the lowest current.

185	Using the procedure outlined by Donovan et al. (2011), an exponential background was
186	fit to the F k $\alpha$ peak with parameters based on a detailed wavescan of Asbestos grossular.
187	Counting times for each spectrometer were 400 seconds on peak and 350 seconds on each
188	background position, resulting in aggregated times per analysis of 1200 seconds on peak and
189	2100 seconds on the background measurement. Time-dependent intensity (TDI) loss (or gain)
190	was monitored for all elements, but no significant or systematic dependence was found so the
191	TDI correction was not used. Unknown and standard intensities were corrected for deadtime and
192	standard intensities were corrected for standard drift over time.
193	Major and minor elements were measured using an accelerating voltage of 15 kV, a beam
194	current of 20 nA, and a beam diameter of 1 $\mu$ m. Counting times for each element were 20
195	seconds on peak and 10 seconds at each background position. Reference materials used for
196	calibration included synthetic forsterite (GRR1017) and tephroite (GRR392) for Mg and Mn,
197	respectively; a natural almandine (Harvard 112140) for Fe; Asbestos grossular for Ca and Si;
198	chromite (NMNH 117075) for Cr; ilmenite (NMNH 96189) for Ti; Brazilian spessartine (Verma,
199	1960) for Al; and vanadium metal for V. Na and K were also analyzed but were below the
200	detection limit in all samples.
201	Final data processing used the CITZAF matrix correction method (Armstrong 1988), with
202	FFAST mass absorption coefficients (Chantler et al. 2005). The matrix correction takes into
203	account oxygen calculated by cation stoichiometry and the oxygen equivalent from F. Cation
204	proportions (including reassignment of total Fe to $Fe^{2+}$ and $Fe^{3+}$ ) and garnet end members were

205 calculated using the Excel spreadsheet written by Locock (2008).

We also report EPMA data for two samples (GRR732 and GRR1429) that were acquired
at Caltech using a JEOL JXA-8200. We used protocols for these analyses similar to those used at

208	UMN but did not measure F. Reference materials for calibration included synthetic pyrope for
209	Mg, Al, and Si; synthetic tephroite (GRR392) for Mn; synthetic fayalite for Fe; grossular
210	(GRR1386) for Ca; and synthetic rutile, Cr <sub>2</sub> O <sub>3</sub> , and V <sub>2</sub> O <sub>5</sub> for Ti, Cr, and V, respectively.
211	
212	RESULTS
213	FTIR spectroscopy
214	FTIR spectra representing the full variation in O-H stretching vibrations in the grossular
215	samples we studied are displayed in Figure 1. Following Rossman and Aines (1991), these

spectra can be separated into different classes according to the most prominent absorption bands

217 present. Class 2, represented by GRR42 (Fig. 1a), is dominated by strong peaks at 3647, 3656,

218 3676, and 3687 cm<sup>-1</sup>, with the highest peak at 3647 cm<sup>-1</sup>. A subset of spectral class 2 is class 2b,

which exhibits the same bands at high frequency (also with maximum absorbance at  $3647 \text{ cm}^{-1}$ )

but has multiple bands between 3600 and 3500 cm<sup>-1</sup> (Fig. 1a). Class 2b is represented by five of

the East African garnets (GRR229, GRR732, GRR771, GRR1386a, and GRR1386), which have

spectra with similar shapes but varying Abs<sub>tot</sub> (Table 1 and supplemental Fig. S1). Three

grossular garnets (GRR1122, GRR1429, and GRR1756) fall into class 7 in the Rossman and

Aines scheme because their strongest absorption is at  $3599 \text{ cm}^{-1}$  rather than at  $3647 \text{ cm}^{-1}$ , but

otherwise their spectra are similar to class 2b; class 2b and class 7 are compared on the samescale in Figure 1b.

The Asbestos garnet JLM83a (Fig. 1c) is typical of class 3, with a strong peak at 3631 cm<sup>-1</sup> that is not as prominent in the other garnets. JLM83a also has bands centered at  $\sim$ 3621, 3644, 3659, 3665, 3677, and 3689 cm<sup>-1</sup>. Garnets from Asbestos fall into both classes 3 and 4, where class 4 is defined by having its strongest peak at 3621 cm<sup>-1</sup>. Figure 1c compares spectra of

JLM83a to published spectra of some other Asbestos grossular garnets (GRR53 and GRR1537)
and end-member grossular crystals synthesized by Geiger and Armbruster (1997) and Withers et
al. (1998).

234 Whereas the absorbance among different East African grossular varies widely (Table 1 235 and supplemental Fig. S1), we failed to detect H zoning within any individual single crystal from 236 this area. The only crystals in our study that are significantly zoned, based on FTIR, are JLM83a 237 and GRR1429. Abstot in JLM83a varied by a factor of ~2 in the range between 3750 and 3300 238 cm<sup>-1</sup>. All spectra have similar shapes, but relative intensities of the peaks at 3631 and 3621 cm<sup>-1</sup> 239 vary (Fig. 1c). Rossman and Aines (1991) noted a comparable variation in Abs<sub>tot</sub> for GRR53, 240 from the same locality. Systematic zoning patterns (such as core-rim relationships or sector 241 zoning) could not be deciphered in the randomly cut slab of JLM83a that we used for both FTIR 242 and SIMS measurements. Although variations in absorbance in Asbestos garnets have been 243 attributed previously to birefringence (Rossman and Aines 1986; Allen and Buseck 1988), our 244 SIMS measurements (Table 3 and discussion below) demonstrate that actual differences in H 245 concentration are a more important factor in this crystal. 246 GRR1429 is also birefringent, which might be the cause of the variations in Abstot (about 247 5% relative) in this sample. We used the average of seven baseline-corrected, unpolarized 248 spectra taken on separate spots to estimate its H<sub>2</sub>O content. Close examination of the spectrum of GRR1386 (Fig. 1a) reveals 12 peaks between 3700 249 and 3500 cm<sup>-1</sup> and the shoulders on some of the bands indicate that curve fitting should take into 250 251 account more than 12 bands. Figure S5 in the supplementary material displays an example of 252 curve fitting for this garnet. We fit 15 bands with a pseudo-Voigt profile (mixed Gaussian and 253 Lorentzian), at 3531, 3539, 3546, 3567, 3583, 3599, 3608, 3623, 3631, 3643, 3647, 3657, 3664,

3674, and 3688 cm<sup>-1</sup>. Although the fit is obviously non-unique, our peak positions (with initial guesses based on visual examination) are within 1 cm<sup>-1</sup> of those fit for the same sample by Geiger and Rossman (2020a), with the slight difference that we fit an additional band at 3539 cm<sup>-1</sup> not present in their fit. We used these peak positions as a basis to fit spectra for the other grossular garnets. Fitted peak centers in other samples are mostly within 1 cm<sup>-1</sup> of those for GRR1386, with a few deviations up to 3 cm<sup>-1</sup>. A table with the fit results for all samples is provided in the supplementary material.

261 The cooling experiment conducted on GRR1386 demonstrates that nearly all the mid-IR 262 absorption in this garnet can be attributed to structurally incorporated O-H, with most bands 263 shifting to higher energies upon cooling to 77 K (Fig. 2). There is no evidence for band splitting 264 at low T, as observed in many other garnets (Geiger et al. 1991; Geiger and Rossman 2018). However, enlargement of the baseline region reveals a broad absorption band at ~3400 cm<sup>-1</sup> that 265 shifts to  $\sim 3240 \text{ cm}^{-1}$ , typical of the transition from liquid water to ice (also plotted for 266 267 comparison in Fig. 2 are spectra for water and ice from Bertie and Lan 1996 and Johnson and Rossman 2003). The 3400 cm<sup>-1</sup> band is apparently also present in some of the other garnets, but 268 269 quantification is hampered by high-frequency interference fringes in most spectra (not visible at 270 the scales used in Fig. 1). The liquid water in GRR1386 is probably contained in sub-micrometer 271 sized fluid inclusions, as found in other NAMs (e.g., Johnson and Rossman 2003; Mosenfelder et 272 al. 2011). Its contribution to Abs<sub>tot</sub> for this sample amounts to 1.2%. We have elected not to 273 correct any of the estimates for Abs<sub>tot</sub> in Table 1 for the possible presence of H<sub>2</sub>O. Furthermore, 274 we have ignored another possible contribution to Abstot in the O-H stretching region from the lowest energy spin-allowed Fe<sup>2+</sup> transition in grossular with FeO content greater than 1 wt% 275 276 (GRR42, GRR1122, GRR1386a, GRR1429, and JLM83a). This band should occur at lower

277	energy compared to other garnets (Geiger and Rossman 1994), because the larger size of the X
278	cation site in grossular leads to greater splitting of the $Fe^{2+}$ bands (White and Moore 1972).
279	However, its exact position is difficult to determine owing to superposition of O-H stretching
280	vibrations (Slack and Chrenko 1971; White and Moore 1972).
281	
282	ERDA
283	H <sub>2</sub> O concentrations of the three grossular garnets, three rhyolitic glasses, and dry
284	forsterite as determined by ERDA are given in Table 2. ERDA and PIXE detector maps of the
285	analyzed areas for two of the garnets (GRR1756 and GRR732) revealed no significant
286	heterogeneities, but a small, anomalous region of interest with high H and Ca content in GRR771
287	was excluded in the calculation of its bulk H content. This "hot spot" probably represents surface
288	contamination (Withers et al. 2012). Bulk H (ERDA) and Ca (PIXE) maps for this sample are
289	provided in the supplementary material.
290	The nominal H <sub>2</sub> O concentration measured by ERDA in GRR1017 forsterite, established
291	as an ultra-dry "blank" material with $\leq$ 0.007 µg/g H <sub>2</sub> O (Mosenfelder et al. 2011), is 101±14
292	$\mu$ g/g. The bulk H concentrations for grossular are strongly correlated with <i>Abs</i> <sub>tot</sub> (Fig. 3). We
293	applied a York regression to the data (York et al. 2004; Vermeesch 2018), taking into account
294	uncertainties in both Abstot and the ERDA measurements. The regression yields an intercept of
295	100±13 $\mu$ g/g H <sub>2</sub> O, which is the value we subtract to arrive at the blank-corrected H <sub>2</sub> O
296	concentrations. The slope of the fit line provides the following calibration for infrared
297	measurements:
298	

299  $C_{H_2O} (\mu g/g) = 0.370 (\pm 0.035, 2\sigma) \times Abs_{tot}$ 

(1)

300

The integral molar absorption coefficient ( $\varepsilon_i$ , in L · mol<sup>-1</sup> · cm<sup>-2</sup>) can be derived by substituting the above relation into a modified form of the Beer-Lambert law (e.g., Johnson and Rossman, 2004),

304

305 
$$C_{H_20}(\mu g/g) = Abs_{tot} \times 18.015 \times 10^6 / [\rho \cdot \varepsilon_i],$$
 (2)

306

307 where  $\rho$  is the density of the mineral (in g/L, or kg/m<sup>3</sup>). Assuming for simplicity a typical density 308 for V-bearing grossular of 3615 kg/m<sup>3</sup> (Maldener et al. 2003), we derive a value for  $\varepsilon_i$  of 13,470 309 L · mol<sup>-1</sup> · cm<sup>-2</sup>.

A check on the accuracy of our ERDA data on grossular is provided by the results on the three rhyolitic glasses, which are in excellent agreement with previously published total H<sub>2</sub>O concentrations determined for these samples using Karl-Fischer titration (Tenner et al. 2009). The H<sub>2</sub>O contents of the glasses measured by ERDA are within 5% of the reference values.

314

### 315 SIMS and EPMA measurements of H and F

SIMS data for the sessions conducted in 2012 and 2016 are summarized in Table 3, with complete data for all individual analyses provided in the supplementary material. Statistics are reported in this paper using the symbology of Fitzsimons et al. (2000). Internal precision in  $^{16}$ OH/ $^{18}$ O (calculated from  $2s_{\bar{x}}$ , the standard error of the mean for 20 or 30 cycles through the mass sequence) ranged from 0.24 to 1.35% relative for all garnets except GRR1122-HT (the blank) and JLM83a. Measured  $^{19}$ F/ $^{18}$ O ratios were similarly precise (0.15 to 0.52%). The high variability in internal precision as well as external precision (reproducibility) for JLM83a isdiscussed below.

 $^{16}$ OH/ $^{18}$ O for each garnet with the exception of JLM83a is plotted in Figure 4 as a 324 325 function of H<sub>2</sub>O concentration determined from FTIR data using the ERDA calibration. Similar plots are derived if <sup>30</sup>Si is used instead of <sup>18</sup>O as the reference mass. Uncertainties in H<sub>2</sub>O 326 327 concentration were estimated by propagating the error in  $Abs_{tot}$  (Table 1) and the uncertainty in 328 the fit to the ERDA data (Fig. 3 and Equation 1). York regressions applied to the data for East 329 African garnets (GRR229, GRR732, GRR771, GRR1386, GRR1386a, and/or GRR1756) have 330 indistinguishable slopes for the two sessions. These regressions also fit well to the data for 331 GRR1429 and GRR1122 measured in 2012. However, the 2016 data for GRR1122 are 332 significantly offset from the regression line, as are both the 2012 and 2016 data for GRR42. 333 In the case of GRR1122 we provisionally attribute this discrepancy to heterogeneity in 334 the sample not recognized from FTIR data and insufficiently probed by the six SIMS analyses 335 that were acquired. The heterogeneity could be related to the presence of fluid inclusions, indicated by a weak broad band at 3420 cm<sup>-1</sup> similar to that seen in GRR1386 (Fig. 2). As for 336 337 GRR42, the consistency in values for seven analyses from the two sessions suggests a different 338 explanation for the offset from the regression. In the discussion section, we explore the 339 possibility that the H concentration of GRR42 is significantly underestimated owing to the 340 frequency dependence of the IR molar absorption coefficient, which cannot be accurately 341 constrained from our ERDA data. 342 Fluorine concentrations measured by SIMS and EPMA are compared in supplemental

Figure S8. The SIMS concentration values are tied to our redetermination of reference values forthe F-bearing basaltic glasses we regularly use for calibration, not taking into account any

345 possible SIMS matrix effect. The close correspondence between SIMS and EPMA values (with 346 all values reasonably close to the 1:1 line) suggests that any such matrix effect is minimal in this 347 case.

348 The heterogeneity in apparent H<sub>2</sub>O concentrations measured by FTIR in JLM83a was 349 confirmed by the grid map of SIMS analyses taken over approximately the same area (Fig. 5). 350 The SIMS analyses also reveal heterogeneity in F, which was additionally confirmed by EPMA 351 (see supplementary material). The measurements using SIMS, FTIR, and EPMA cannot be 352 compared directly to determine a zoning pattern because the FTIR analyses are taken in 353 transmission, whereas the SIMS and EPMA measurements sample a volume that is close to the 354 surface (~1-2  $\mu$ m deep for SIMS, and ~3  $\mu$ m deep for EPMA at the analytical conditions used). 355 However, the range in apparent H<sub>2</sub>O concentrations measured by FTIR (256-571  $\mu$ g/g) is similar 356 to the range measured by SIMS (170-576  $\mu$ g/g), calculated by applying the York regression in 357 Figure 4b. For the sake of this comparison we discard four analyses (marked by asterisks in Fig. 358 5) that were affected by the spike in background contaminant levels in the vacuum caused by the cold trap running out of liquid N<sub>2</sub> (see supplementary material). These analyses had  ${}^{12}C/{}^{18}O$ 359 360 ratios one to two orders of magnitude higher than the other analyses.

Following the 2012 session, it became apparent that <sup>19</sup>F and <sup>16</sup>OH are correlated for some grossular garnets, particularly the East African samples with similar class 2b and class 7 FTIR spectra (Fig. 6a). A similar correlation was noted for spessartine (Fig. 6b). An even more robust correlation exists for the data taken on JLM83a alone (Fig. 6a). The values in Figure 6a for JLM83a represent average <sup>16</sup>OH/<sup>18</sup>O and <sup>19</sup>F/<sup>18</sup>O for each analysis. However, the internal precision of the analyses varies widely, with  $s_{\bar{x}}/\hat{s}_{\bar{x}}$  (where  $\hat{s}_{\bar{x}}$  is the precision predicted by Poisson statistics; Fitzsimons et al., 2000), ranging from 1 (ideal) to 30 (poor). We would

368 normally discriminate against SIMS measurements with such high  $s_{\bar{x}}/\hat{s}_{\bar{x}}$  values; for instance, in 369 Mosenfelder et al. (2011) all analyses of olivine with  $s_{\bar{x}}/\hat{s}_{\bar{x}} > 5$  were filtered out. High  $s_{\bar{x}}/\hat{s}_{\bar{x}}$  in 370 olivine and orthopyroxene has been attributed to the presence of sub-micrometer to micrometer-371 sized inclusions of hydrous minerals (such as serpentine and amphibole) and/or fluid inclusions 372 (Mosenfelder et al. 2011; Mosenfelder and Rossman 2013a). In the present case, however, we 373 attribute the high variability to zoning in H and F within the garnet itself. Analyses with high  $s_{\bar{x}}/\hat{s}_{\bar{x}}$  show strong cycle-to-cycle covariation between <sup>19</sup>F and <sup>16</sup>OH (Supplementary Fig. S8), 374 375 indicating that F and H are coupled on a sub-micrometer scale. 376 377 DISCUSSION 378 Molar IR absorption coefficient for H<sub>2</sub>O in grossular 379 In Figure 7, we compare our ERDA data to  $H_2O$  concentrations measured (as a function 380 of Abstot) in grossular-rich garnets using other "absolute" techniques: NRA (Rossman and Aines 381 1991; Maldener et al. 2003), P<sub>2</sub>O<sub>5</sub> cell coulometry using a moisture evolution analyzer (MEA, 382 Aines and Rossman 1984), H manometry (Aines and Rossman 1984), and continuous-flow mass 383 spectrometry (O'Leary et al., 2007). Also displayed in the graph is the regression line for all of 384 the data considered by Rossman and Aines (1991) for grossular and hydrogrossular with up to 385 12.75 wt% H<sub>2</sub>O (with H<sub>2</sub>O contents on the high end well-constrained using X-ray diffraction). 386 We have no definitive explanation for the discrepancy between H concentrations 387 measured in GRR732 by ERDA (599  $\mu$ g/g H<sub>2</sub>O) and continuous-flow mass spectrometry 388 (CFMS; 1062  $\mu$ g/g H<sub>2</sub>O); we can only speculate that the blank for that particular CFMS 389 measurement was higher than recognized. Our data show excellent agreement, however, with 390 some of the NRA data. In particular, the analyses by Maldener et al. (2003) on their TSAV (480

391	$\pm$ 40 $\mu g/g$ H_2O) and thin (0.0165 cm) slab of HESS1 (870 $\pm$ 90 $\mu g/g$ H_2O) samples lie close to
392	our regression line. These East African grossular garnets have IR spectra similar to those from
393	the same area that we measured with ERDA. Our data are in poorer agreement with Maldener et
394	al.'s analyses of a thicker (0.0312 cm) slab of HESS1 (950 $\pm$ 80 $\mu g/g$ H_2O) and with MALI (170
395	$\pm$ 20 $\mu g/g$ H_2O) and GRMALI (190 $\pm$ 20 $\mu g/g$ H_2O). The two NRA measurements of HESS1 are
396	in agreement with each other within mutual uncertainties, which suggests that there is a
397	discrepancy in the FTIR measurements (e.g., non-linearity of the detector for linear absorbance
398	higher than 1.5 in the thick HESS1 sample) or perhaps unrecognized H zoning in the crystal. The
399	discrepancies between our measurements and those on MALI and GRMALI might derive from
400	errors made in correcting the sample thicknesses for beam convergence in the microscope.
401	Furthermore, any inaccuracies in blank correction are also more likely to affect these low H
402	content samples.
403	The H concentrations of Asbestos grossular garnets GRR53 and GRR53F were measured
404	using three different techniques: H manometry, MEA, and NRA (Aines and Rossman 1984;
405	Rossman and Aines 1991). The poor agreement of our data fit with the values determined with
406	those techniques might be related to H zoning comparable to that in JLM83a. Hydrogen

407 manometry and MEA are bulk methods that sample a much larger volume than the FTIR

408 measurements and thus cannot account for fine-scale zoning. Even though NRA is a near-surface

409 technique, the beam size employed at that time was much larger (several mm) than the apparent

410 scale of H zoning in Asbestos grossular. Furthermore, the NRA measurement of GRR53F

411 employed a reaction with <sup>19</sup>F ions that was subsequently abandoned in favor of a more

412 reproducible reaction employing a <sup>15</sup>N beam (Rossman 2006).

413	There is a large discrepancy (Fig. 7) between our data fit on relatively low H content
414	grossular and the regression applied by Rossman and Aines (1991), which was largely
415	constrained by hydrogrossular garnets with high H contents. The discrepancy cannot be
416	attributed to the lower sample density of hydrogrossular; for instance, using Equation 2,
417	Crestmore hydrogrossular GRR1358, which has 12.75 wt% H <sub>2</sub> O, Abs <sub>tot</sub> of 158,750 cm <sup>-2</sup> , and a
418	calculated density of 3050 kg/m <sup>3</sup> (Basso et al. 1983), would have an $\varepsilon_i$ of 7354 L $\cdot$ mol <sup>-1</sup> $\cdot$ cm <sup>-2</sup> ,
419	lower than our new value by a factor of almost two. The data on hydrogrossular samples could
420	have been compromised by uncertainties in sample thickness or unrecognized H zoning, but
421	these factors are speculative and more work is needed to ascertain the reason for the large
422	apparent difference between low-H grossular and hydrogrossular.
423	The possibility of a difference in calibration factor for grossular and hydrogrossular was
424	already noted by Rossman and Aines (1991), and Rossman (2006) recommended a calibration
425	factor for low H content grossular of 0.14 (for absorbance in three directions). That value,
426	equivalent to 0.42 for the convention for $Abs_{tot}$ used in this study, compares reasonably well to
427	our new value of $0.370 \pm 0.035$ (Equation 1). We conclude that our new calibration is more
428	appropriate than the Rossman and Aines calibration for grossular with less than ~1 wt% H <sub>2</sub> O.
429	Empirical and theoretical studies consistently show an increase in $\varepsilon_i$ with decreasing
430	frequency for O-H absorption bands (Paterson 1982; Libowitzky and Rossman 1997; Balan et al.
431	2008; Mosenfelder et al., 2015). This means that an IR calibration based on grossular with
432	relatively low mean wavenumber might underestimate the H <sub>2</sub> O contents of grossular with higher
433	mean wavenumber. Whereas our ERDA calibration was performed on grossular with mean
434	wavenumbers between 3595 and 3622 cm <sup>-1</sup> (Table 1), GRR42 is dominated by bands with
435	significantly higher frequency (with a mean wavenumber of 3647 cm <sup>-1</sup> ; Fig. 1a, Table 1). A

436	frequency dependence to $\varepsilon_i$ can therefore explain the offset of SIMS data for this sample in
437	Figure 4, but the magnitude of the dependence is poorly constrained. For the sake of calculating
438	formula units and garnet end-members (Table 4), we assume a H <sub>2</sub> O concentration for this sample
439	based on the York regression shown in Figure 4a. Further work is needed to constrain the
440	frequency dependence of $\varepsilon_i$ to determine whether this assumption is justified. Our calibration
441	also might underestimate H <sub>2</sub> O concentrations for JLM83a and other Asbestos grossular garnets,
442	which also have a higher mean wavenumber (Table 1). Unfortunately, the zoning in some (or
443	all?) of these class 3 and 4 garnets will make this goal challenging, and class 2 garnets with
444	spectra like that of GRR42 are rare in our experience. A better way forward might entail
445	synthesizing grossular crystals at high P-T (Geiger and Armbruster 1997; Withers et al. 1998)
446	and establishing their homogeneity using a technique such as ERDA.
447	
448	Previous models of H and F incorporation mechanisms
449	Hydrogen incorporation in garnets is classically assumed to take place via the
450	"hydrogarnet" substitution, whereby four H atoms bond to the O atoms surrounding a vacant Si
451	site. Hereafter, we refer to this substitution as $(4H)_{Si}^{\times}$ , which is a commonly used, abbreviated
452	form of the following defect associate written in Kröger-Vink (1956) notation:
453	
454	$\left\{V_{\rm Si}^{\mu\nu} - 4(\rm OH)_{0}^{\cdot}\right\}^{\times}$ . (3)
455	
456	Here V represents a vacancy: the superscripts $\times$ and 'represent null positive and pagative

Here, V represents a vacancy; the superscripts  $\hat{}$ , and represent null, positive, and negative 456 charges relative to the perfect structure, respectively; and the subscripts refer to Si and O 457 458 crystallographic sites. We recognize that some mineralogists might object to using point defect

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- 459 symbolism to refer to both trace and high concentrations of elements, but we use it in the
- 460 following discussion for the sake of convenience.

The  $(4H)_{Si}^{\times}$  mechanism is well established for H-rich garnets along the grossular-katoite 461 462 join (for review, see Beran and Libowitzky 2006; Grew et al. 2013; and Geiger and Rossman 463 2020a), not only from chemistry but from NMR, X-ray diffraction (XRD), and neutron 464 diffraction (Cohen-Addad et al. 1987; Lager et al. 1987, 1989; Cho et al. 1993). An analogous 465 substitution of four F atoms substituting for O surrounding a vacant Si site is well documented 466 for F-rich garnets, based on XRD (Smyth et al. 1990; Chakhmouradian et al. 2008) and negative 467 correlations between Si and F concentrations (Flohr and Ross 1989; Visser 1993; 468 Chakmouradian et al. 2008). This substitution can be written as: 469  $\{V_{Si}^{''''} - 4(F)_0^{\cdot}\}^{\times}$ . 470 (4)471 This substitution differs structurally from  $(4H)_{Si}^{\times}$  in the sense that F atoms can occupy the 472

regular site of O atoms, whereas the H atoms in  $(4H)_{Si}^{\times}$  are bonded to O that is displaced from its 474 normal position owing to expansion of the tetrahedron (e.g. Lager et al. 1989).

475 Other substitution mechanisms for H and particularly for F in garnet have received 476 somewhat less attention. Valley et al. (1983) considered the possibility, in addition to 477 mechanisms (3) and (4), of F and H incorporation in grossular via the coupled substitution:

478

473

479 
$$(Al, Fe)^{3+} + 0^{2-} \leftrightarrow (Mg, Fe)^{2+} + (F, OH)^{-},$$
 (5)

480

481 where the cations are located in the six-coordinated Y site. They recognized that verification of this substitution relies on accurate determination of  $Fe^{2+}/Fe^{3+}$  as well as H and F content, beyond 482 the scope of their study. Various mechanisms involving coupled substitution of Ti and H have 483 484 also been considered, primarily for pyrope (Bell and Rossman 1992; Khomenko et al. 1994; 485 Geiger et al. 2000) but also for grossular (Reynes et al. 2020). Partial substitution of H for the Y-486 cation in a hydrogrossular was inferred by Kalinichenko et al. (1987), based on NMR, and at least two studies have inferred substitutions other than  $(4H)_{Si}^{\times}$  in H-rich garnets based on crystal 487 488 chemical arguments (Birkett and Trcienski 1984; Basso and Cabella 1990). In principle OH<sup>-</sup> or 489  $F^{-}$  can replace O in the dodecahedral X-site in garnet, if it is vacant or contains a monovalent 490 cation (e.g., Andrut et al. 2002). For instance, a Ca vacancy in grossular can be compensated by two H, which we notate as  $(2H)_{Ca}^{\times}$ . A similar substitution for the Y cation can be notated as 491  $(3H)_{Al}^{\times}$  or  $(3H)_{Fe}^{\times}$ , depending on the garnet composition. 492 493 Early on in the study of nominally anhydrous minerals it was recognized that the mid-IR

494 spectra of some garnets are incompatible with the  $(4H)_{Si}^{\times}$  mechanism, because too many peaks are present (Aines and Rossman 1984). The (4H)<sup>×</sup><sub>Si</sub> substitution sensu stricto (cf. Geiger and 495 496 Rossman 2020a; see below) can result in a maximum of four O-H stretching vibrations, or two if 497 the symmetry of the garnet is maintained (Harmon et al. 1982; Aines and Rossman 1984). The 498 IR spectra of many pyrope, almandine, and andradite garnets are dominated by a broad, 499 asymmetric band that splits into two bands at low temperature (cooled by liquid  $N_2$ ), thus 500 consistent with (4H)<sup>×</sup><sub>Si</sub> (Geiger et al. 1991; Geiger and Rossman 2018). Differences in band 501 positions as a function of composition can be explained by the presence of different neighboring 502 X and Y cations, and the broad width of the bands can be explained by a combination of proton 503 disorder and variations in local cation configuration. On this basis, Geiger and Rossman (2018)

504	assigned bands at 3629, 3613, 3598, and 3563 cm <sup>-1</sup> in pyrope, almandine, grossular, and
505	and radite respectively to $(4H)_{Si}^{\times}$ . However, in grossular, as well as many spessartine garnets
506	(Aines and Rossman 1984; Arredondo et al. 2001) and some rare pyrope garnets from the Dora
507	Maira massif (Rossman et al. 1989), the multitude of sharp peaks cannot be explained by this
508	mechanism alone.
509	Additional constraints on H incorporation in grossular and garnets on the grossular-
510	andradite join come from experimental studies of H diffusion (Kurka et al. 2005;
511	Phichaikamjornwut et al. 2012; Reynes et al. 2018). All of these studies support the idea that
512	there are multiple sites for O-H incorporation, because various O-H bands decrease at different
513	rates during dehydration. In general, dehydration kinetics for bands at higher frequency have a
514	lower activation energy ( $E_a$ ) than those at low frequency. For instance, Reynes et al. (2018)
515	measured $E_a$ for a grossular-rich garnet in the range between 157 and 185 kJmol <sup>-1</sup> for bands at
516	3628, 3645, 3657, and 3686 $\text{cm}^{-1}$ , and in the range between 216 and 275 kJmol <sup>-1</sup> for bands at
517	3533, 3576, and 3604 cm <sup>-1</sup> . A positive dependence of the diffusion coefficient on $f_{O_2}$ as well as
518	observations of color changes in sample rims in the experiments by Reynes et al. also support a
519	key role for Fe <sup>2+</sup> in dehydration. Reynes et al. identify a fast diffusion mechanism, analogous to
520	that identified in many other nominally anhydrous minerals (Ingrin and Blanchard 2006), that is
521	limited by the availability of $Fe^{2+}$ :

522

523 
$$M^{2+} + OH^{-} + \frac{1}{4}O_2 \to M^{3+} + O^{2-} + \frac{1}{2}H_2O,$$
 (6)

524

where M represents either Fe or Mn, which may or may not be directly tied to a defect
responsible for incorporating H. Our failed attempt to fully dehydrate GRR1122 (supplementary

Fig. S4) at low  $f_{O_2}$  (with virtually no change in spectra between dehydration steps 1 and 2) is consistent with this interpretation. We will return to this point when we discuss our model for coupled H-F substitution in the next section.

530 Another critical constraint on H incorporation in grossular comes from the study of Cho and Rossman (1993), who performed solid-state <sup>1</sup>H NMR measurements on a synthetic end-531 532 member katoite (Ca<sub>3</sub>Al<sub>2</sub>• <sub>3</sub>(OH)<sub>12</sub>) powder and single crystals of GRR1386 (Fig. 1a), GRR53 533 (Fig. 1c), and GRR1537 (Fig. 1c). From their analysis of multiple-quantum spectra, Cho and 534 Rossman concluded that the protons in katoite occur in closely spaced groups of four, as 535 expected for  $(4H)_{si}^{\times}$ . However, the best fit to the multi-quantum spectrum of GRR1386 was for a 536 cluster size of two protons, and the best fit to the data for GRR1537 was for two separate clusters 537 consisting of two and four protons each. The class 4 IR spectrum of GRR1537 is dominated by a band at 3621 cm<sup>-1</sup> (Fig. 1c), which is the position of the main band in a synthetic hydrogrossular 538 539 with composition Ca<sub>3</sub>Al<sub>2</sub>(SiO<sub>4</sub>)<sub>2.28</sub>(O<sub>4</sub>H<sub>4</sub>)<sub>0.72</sub> (Rossman and Aines 1991). Therefore, Cho and Rossman assigned this band to  $(4H)_{Si}^{\times}$  and other, less intense bands to a different, unspecified 540 541 defect containing two closely spaced protons.

Geiger and Rossman (2020a, 2020b) offered a novel explanation for the sharp, highfrequency bands in natural grossular. They hypothesized that bands at 3599, 3612, 3622, 3634, and 3641 cm<sup>-1</sup> correspond to clustered groups of  $(4H)_{Si}^{\times}$  with cluster sizes of one, two, three, four and five, respectively. A band at 3657 cm<sup>-1</sup> was tentatively assigned to a cluster size of six, and a band at 3660 cm<sup>-1</sup> was assigned to a "finite size katoite cluster" based on the fact that the main absorption in pure katoite (GRR1059; mislabeled as "Si 2.28" in Fig. 7 of Geiger and Rossman 2020a) occurs at this frequency (Rossman and Aines 1991). Bands at 3674 and 3688 cm<sup>-1</sup> were

attributed to inclusions of hydrous minerals. This model explains all the features in IR spectra of class 2, 3, and 4 grossular and all absorptions above  $\sim$ 3600 cm<sup>-1</sup> in class 2b.

551 It is well known from studies of hydrous minerals that the strengths of O-H bonds (and 552 thus their vibrational frequencies) are affected by different cations sharing bonds with the same 553 O atom (i.e., second nearest neighbors to the H atom). For example, the frequencies of O-H bonds in amphibole, which vary by up to  $\sim 50$  cm<sup>-1</sup>, can be reliably assigned to different cation 554 555 environments (e.g., Burns and Strens 1966). Recent work has also explored the effects of 556 neighboring cations on O-H stretching frequencies in olivine (Blanchard et al. 2017). The cluster 557 model of Geiger and Rossman (2020a,b), however, relies on an assumption that comparable shifts in O-H bond frequencies are influenced by bonds in other  $(4H)_{Si}^{\times}$  that are not directly 558 559 adjacent, because the tetrahedra in garnet are isolated from each other. Geiger and Rossman 560 justify this assumption as a consequence of the increase in Ca-O bond lengths between the end 561 members grossular (Geiger and Armbruster 1997) and katoite (Lager et al. 1987). It remains to 562 be shown by first-principles modeling or other methods whether or not this relaxation in Ca-O 563 distances for the end members also operates on a localized scale within small, isolated clusters of  $(4H)_{Si}^{\times}$  in garnets with trace amounts of H. It should also be noted that the model of Geiger and 564 565 Rossman ignores the results of Cho and Rossman (1993) that indicate some grossular garnets 566 contain a significant fraction of defects with only two protons. Furthermore, the assignment of bands at 3674 and 3688 cm<sup>-1</sup> to inclusions of hydrous phases is problematic, owing to geological 567 568 and spectroscopic reasons that we elaborate in the supplementary material. In the following 569 sections, we offer alternative explanations for some of the absorption bands that call on coupled 570 H and F substitution for Si as well as substitution of H (and possibly F) for cations in octahedral 571 and/or dodecahedral sites.

572

## 573 Coupled substitution of H and F

574	Our first clue that incorporation mechanisms of H and F in grossular might be entangled
575	came from correlations between <sup>16</sup> OH and <sup>19</sup> F in grossular (Fig. 6a) and spessartine (Fig. 6b),
576	prompting us to consider coupled substitution of H and F for Si (or other cations). While it is
577	possible that these correlations are coincidental and reflect nothing about coupled H-F
578	substitution, interpretation of the IR spectra combined with quantitative data on H and F
579	concentrations (Table 4) suggests otherwise.
580	Consider first the comparison between GRR42 and GRR1386: these crystals have the
581	same F content (0.25 wt%), but GRR1386 has nearly twice as much $H_2O$ (0.13 wt%) as GRR42
582	(0.07 wt%; based on SIMS). The other obvious difference (Fig. 1a) is that both garnets have
583	significant absorbance above $\sim$ 3620 cm <sup>-1</sup> , but GRR42 has much lower absorbance at lower
584	frequencies. This suggests that some of the high-frequency bands are related to defects
585	containing both H and F, unless all F is bound in anhydrous defects (e.g., via Equation 4). We
586	draw a similar conclusion from the comparison between GRR732 and GRR1122 in Figure 1b:
587	these garnets have nearly the same $H_2O$ content (0.052-0.054 wt%), but GRR732 has almost
588	twice as much F (0.11 wt%) as GRR1122 (0.06 wt%) and has stronger absorbance above $\sim$ 3620
589	cm <sup>-1</sup> , whereas GRR1122 has stronger absorbance at lower frequencies. Furthermore, the
590	strongest correlation between <sup>16</sup> OH and <sup>19</sup> F is evident in JLM83a, which has virtually no
591	absorbance at frequencies lower than $\sim 3600 \text{ cm}^{-1}$ (Fig. 1c) – again suggesting that the higher-
592	frequency O-H bands are related to F incorporation. An attempt to quantify these general
593	observations is summarized in Figure 8, where we plot F (per formula unit, Table 4) against the
594	band areas derived by spectral curve fitting for nine out of ten of the grossular garnets in the

- 595 study (JLM83a is excluded from this analysis because it is difficult to compare any given
- 596 spectrum to a precise F content in this zoned garnet).

597 Motivated partly by these observations and partly by the experiments and first-principles 598 density functional theory (DFT) calculations of Crépisson et al. (2014) on forsterite, we propose that some of the bands at frequencies  $> 3600 \text{ cm}^{-1}$  represent coupled H-F substitution for Si. 599 600 Following Cho and Rossman (1993) and Withers et al. (1998), we start with the assumption that the band at 3621 cm<sup>-1</sup>, with a shoulder at ~3610 cm<sup>-1</sup>, represents  $(4H)_{Si}^{\times}$ . The asymmetry of the 601 602 stretching vibration in this region is consistent with the band structure in other garnet 603 compositions previously assigned to  $(4H)_{Si}^{\times}$  (Geiger and Rossman 2018), although the bands in 604 grossular are narrower. Some or all of the bands at higher frequencies are then assigned to defect 605 associates involving both OH<sup>-</sup> and F<sup>-</sup>:

606

607 
$$\left\{ V_{\text{Si}}^{'''} - (\text{OH})_0^{\cdot} - 3(\text{F})_0^{\cdot} \right\}^{\times}$$
 (7)

$$608 \quad \left\{ V_{\rm Si}^{'''} - 2(\rm OH)_0^{\cdot} - 2(\rm F)_0^{\cdot} \right\}^{\times} \tag{8}$$

$$609 \quad \left\{ V_{\rm Si}^{'''} - 3(0{\rm H})_0^{\cdot} - ({\rm F})_0^{\cdot} \right\}^{\times} \tag{9}$$

610

611 Calculations made by Crépisson et al. (2014) indicate that some of these clumped OH-F 612 defects in forsterite lead to O-H vibrations at higher frequencies than those assigned by previous 613 studies to the  $(4H)_{s_i}^{\times}$  defect. For instance, the single O-H stretching vibration corresponding to 614 defect (7) is at 3674 cm<sup>-1</sup>, as opposed to the four O-H vibrations predicted for the most stable  $(4H)_{Si}^{\times}$  configurations, which are between ~3400 and 3625 cm<sup>-1</sup> (Balan et al. 2011; Umemoto et 615 616 al. 2011). In general, the frequencies of bands assigned to the various clumped OH-F defects 617 decrease as the number of F atoms decreases, but complexities arise owing to the three non-

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equivalent crystallographic sites for O in the olivine structure. A positive frequency shift is also 618 619 predicted for a clumped OH-F defect in an octahedral site, compared to a fully protonated vacancy. The positive O-H frequency shifts for sites where  $F^{-}$  replaces  $O^{2-}$  are presumably a 620 consequence of the influence of  $F^-$  (which is more electronegative than  $O^{2-}$ ) on the geometry of 621 622 the polyhedron (Crépisson et al. 2014). The theoretically derived frequencies of the clumped 623 OH-F defects represented by Equations 7-9 are consistent with IR spectra of natural olivines 624 known to contain relatively high F concentrations (Sykes et al. 1994; Libowitzky and Beran 1995; Mosenfelder et al. 2011), which have bands at 3670, 3637, and 3620 cm<sup>-1</sup> that are not 625 626 present in olivines with lower F contents. 627 Precise assignment of individual high-frequency bands in grossular to defects represented by Equations 7-9 is difficult. We tentatively assign the band at 3687-3688 cm<sup>-1</sup> to Equation (7), 628 629 by analogy to the corresponding defect in olivine. This attribution helps to reconcile the H 630 diffusion behavior observed by Kurka et al. (2005), who found that this band and the band at 3600 cm<sup>-1</sup> decrease more rapidly than bands at 3645 and 3568 cm<sup>-1</sup>. Specifically, H diffusion of 631 632 the defect associated with Equation 7 can take place rapidly via the reaction: 633

634 
$$\left\{V_{\text{Si}}^{\mu\nu} - (\text{OH})_{0}^{\cdot} - 3(\text{F})_{0}^{\cdot}\right\}^{\times} + \text{Fe}_{\text{Ca}}^{\times} + \frac{1}{4}O_{2} = \left\{V_{\text{Si}}^{\mu\nu} - (0)_{0}^{\times} - 3(\text{F})_{0}^{\cdot}\right\}^{\prime} + \text{Fe}_{\text{Ca}}^{\cdot} + \frac{1}{2}H_{2}O_{2}$$
 (10)

635

This reaction is similar to the generalized redox reaction represented by Equation (6) but is defect-specific. In this case, F diffusion – which is inherently slower than H diffusion – is not needed to maintain charge balance because dehydration involves oxidation of  $Fe^{2+}$  in the dodecahedral site. Many studies have found that oxidation reactions of this type are faster than reactions that require counter-diffusion of cations to maintain charge balance (Ingrin and 641 Blanchard 2006), and Reynes et al. (2018) provide direct evidence from color changes in 642 dehydrated crystals and experiments at variable  $f_{O_2}$  that oxidation of Fe<sup>2+</sup> plays a key role in 643 dehydration of grossular.

644 The effect of dilute F substitution on local symmetry of the grossular structure is 645 unknown, so it is not straightforward to use group theory to predict how many bands should be 646 present for defects containing more than one O-H bond (Equations 8 and 9). Significant 647 proportions of cations other than Ca and Al in neighboring X and Y sites, such as V or Fe 648 (present in some of the grossular garnets in this study; Table 4), could also shift the frequencies 649 of the O-H bands, because these frequencies largely depend on the masses, ionic radii, and 650 electronic states of the neighboring cations (Burns and Strens 1966; Berry et al. 2007; Geiger and 651 Rossman 2018, 2020a). This effect is likely to be minor but may explain the small shifts up to 3 cm<sup>-1</sup> for some peaks in our curve fits for different grossular spectra. The imperfect correlation of 652 653 F to the integrated absorbance of the high-frequency bands (Fig. 8a) could be related to the 654 additional presence of F in sites with no OH (e.g., Equation 4), vagaries of the fitting exercise, or 655 perhaps other mechanisms completely unrelated to F – such as the cluster model of Geiger and 656 Rossman (2020a,b) or coupled substitution of H and Ti (Reynes et al. 2020), which we discuss 657 below. Therefore, more work including ab initio calculations is needed to verify our hypothesis 658 that clumped OH-F defects are present in grossular. One attractive feature of our model is that 659 the presence of a significant proportion of defect (8) could partially explain the finding of Cho 660 and Rossman (1993) that H is predominantly present in clusters of two closely spaced protons in 661 GRR1386, and by extension to other class 2b grossular garnets.

662

### 663 Low frequency O-H absorption bands

No study has yet offered a satisfactory explanation for all the bands below  $\sim 3600$  cm<sup>-1</sup> in 664 665 class 2a and class 7 grossular (some of the bands are also present in class 5 and 6 grossular; Rossman and Aines 1991). Geiger and Rossman (2020a) assigned the band at 3563 cm<sup>-1</sup> to a 666 localized "hydroandradite group" consisting of a single  $(4H)_{s_i}^{\times}$  adjacent to Fe<sup>3+</sup> in the octahedral 667 668 site, based on comparison to the most intense O-H band frequency in andradite IR spectra 669 (Amthauer and Rossman 1998; Geiger and Rossman 2018). Furthermore, they speculated that 670 their model for hydrogrossular clusters could be extended to assign other bands at 3581 and 3594 671 cm<sup>-1</sup> to different sized hydroandradite clusters. This hypothesis is difficult to assess because the total Fe content in many grossular garnets is too low to accurately determine  $Fe^{2+}/Fe^{3+}$  using 672 673 established methods. Our cation assignments (Table 4) based on stoichiometry using the calculation scheme of Locock (2008) indicate that all of the Fe is  $Fe^{2+}$  for our garnets with 0.06 674 to 1.8 wt% FeO, but this method for calculating  $Fe^{2+}/Fe^{3+}$  suffers from considerable uncertainty 675 and Fe<sup>3+</sup> may also be present. 676 677 The role of Ti in H incorporation in grossular and andradite was recently studied by

679 follows:

...

680

678

$$681 \quad \left\{ V_{\text{Si}}^{\text{I}^{\text{IIII}}} - 2(\text{OH})_{0}^{\cdot} \right\} + 2\text{Ti}_{\text{Al,Fe}}^{\cdot}$$
(11)

Reynes et al. (2020). They proposed coupled substitution of Ti on the Y site and 2H for Si, as

682

FTIR spectra of the grossular studied by Reynes et al. are unusual compared to most grossular garnets surveyed by Rossman and Aines (1991), containing only three bands at 3546, 3595, and 3652 cm<sup>-1</sup>; the closest match is to GRR1359, a grossular from a metarodingite in the Gruppo di Voltri, Italy (probably from a similar geologic setting as the rodingite grossular studied by

687	Reynes et al.). Bands at similar frequencies are also present in spectra of the class 2b grossular
688	samples in this study, centered at 3546, 3599, and 3657 cm <sup>-1</sup> in GRR1386. Therefore, this
689	mechanism may also be important in these garnets, which contain between 0.3 to 0.6 wt% $\rm TiO_2$
690	(Table 4). The presence of a band at 3652-3657 cm <sup>-1</sup> related to defect (11) could also partially
691	explain the imperfect correlation of F content to high frequency absorption in Figure 8a.
692	Another possibility is that some of the low frequency bands represent H substitution in
693	dodecahedral and/or octahedral sites via defects such as $(2H)_{Ca}^{\times}$ or $(3H)_{Al}^{\times}$ , respectively. This
694	attribution could further help to explain the identification of two-proton clusters in GRR1386 by
695	Cho and Rossman (1993), if a large fraction of the H is located in the defects (8), (11), and
696	$(2H)_{Ca}^{\times}$ combined. Substitution of H via defects such as $(2H)_{Ca}^{\times}$ would also be consistent with
697	the higher $E_a$ measured for lower frequency O-H bands by Reynes et al. (2018), if dehydration
698	requires counter-diffusion of Ca to take place (as opposed to dehydration via a fast-redox
699	mechanism). Although this hypothesis is highly speculative, DFT calculations for nominally
700	anhydrous minerals to date indicate that the $(4H)_{Si}^{\times}$ defect gives rise to O-H absorptions at higher
701	frequencies than substitutions involving octahedral cation vacancies. This phenomenon applies
702	to olivine (Balan et al. 2011; Umemoto et al. 2011), enstatite (Balan et al. 2013), wadsleyite
703	(Blanchard et al. 2013), and ringwoodite (Blanchard et al. 2009).
704	In principle, clumped OH-F defects in octahedral or dodecahedral sites are also possible,
705	as are coupled substitution mechanisms like those put forth by Valley et al. (1983) (Equation 5).
706	The intensity of the low-frequency bands in East African grossular also correlates positively with
707	F content (Fig. 8b). By analogy to our argument with the high frequency bands, this might
708	suggest a role for such defects in addition to clumped OH-F in tetrahedral sites. One weakness
709	with this idea is that the presence of isolated, clumped OH-F defects in octahedral or

710	dodecahedral sites would require the presence of Si-F bonds, because all O sites in the garnet
711	structure are shared between the tetrahedral site, two dodecahedral sites, and the octahedral site.
712	As noted by Smyth et al. (1990), there are no known mineral structures in which F bonds directly
713	to tetrahedral Si. However, some NMR studies provide evidence for Si-F bonding in
714	aluminosilicate glasses (e.g., Zeng and Stebbins 2000; Liu and Nekvasil 2002). It is probably
715	more likely that any defect associate involving clumped OH-F in an octahedral or dodecahedral
716	site would be paired with a neighboring defect associate in the tetrahedral site (e.g., three H and
717	one F substituting for Si).
718	
719	Coupled H and F substitution in other garnet compositions
720	Our model of coupled H and F substitution for Si in grossular likely applies to garnets
721	with other compositions. Data from our 2012 session show strong correlations between <sup>16</sup> OH and
722	<sup>19</sup> F in 11 spessartine garnets (Fig. 6b), similar to the correlations seen for East African grossular
723	(Fig. 6a). These spessartine garnets have peaks at $\sim$ 3582, 3595, 3615, 3625, 3640, and 3649 cm <sup>-1</sup>
724	(Arredondo et al. 2001). Bands at similar frequencies are also present in spectra of a spessartine
725	with up to 3.8 wt% F studied by Smyth et al. (1990). Aines and Rossman (1984) recognized that
726	these six peaks (which can be fit with even more bands) cannot be explained by $(4H)_{Si}^{\times}$ alone. It
727	is tempting to assign the band at 3595 cm <sup>-1</sup> (with a shoulder at 3582 cm <sup>-1</sup> ) to $(4H)_{Si}^{\times}$ . The lower
728	frequency compared to katoite or grossular can be explained by the high mass of $Mn^{2+}$
729	(compared to Ca <sup>2+</sup> ) in the X-site (Geiger and Rossman 2018). However, relative peak heights at
730	3582 and 3595 cm <sup>-1</sup> vary considerably in garnets from different parts of the pegmatite and in
731	spessartine from the Himalayan Mine (Arredondo et al. 2001), which is not easily reconcilable
732	with a single O-H site being responsible for both bands. Taken together, these observations

suggest the possibility that some or all of the bands represent clumped OH-F defects analogousto those we propose for grossular.

735	The presence of high frequency bands in some garnets along the grossular-andradite join
736	(Phichaikamjornwut et al. 2012; Geiger and Rossman 2020a) was already noted. In the
737	supplementary material, we provide IR spectra of a Grs <sub>26</sub> Adr <sub>74</sub> garnet (GRR1830) and a nearly
738	end-member andradite (GRR2103) that we analyzed by SIMS in 2012. GRR1830 has at least
739	eight bands, with peaks at 3560, 3580, 3589, 3617, 3640, 3653, 3660, and 3677 cm <sup>-1</sup> . The band
740	locations are similar to those in the garnets from Thailand studied by Phichaikamjornwut et al.
741	(2012). The spectrum of GRR2103 is dominated by a strong peak at 3563 cm <sup>-1</sup> (with a shoulder
742	at about 3547 cm <sup>-1</sup> ) and has only weak absorbance above 3600 cm <sup>-1</sup> . Published spectra of other
743	garnets close to end-member andradite also typically do not have these bands (e.g., Amthauer
744	and Rossman 1998; Geiger and Rossman 2018). However, Amthauer and Rossman (1998) noted
745	spectral complexity in a near-end member andradite from Stanley Butte, AZ, USA (GRR1137).
746	Again, the number of bands in GRR1830 and the grossular-andradite garnets studied by
747	Phichaikamjornwut et al. (2012) raises the possibility that some of them represent clumped OH-F
748	defects. This inference is also consistent with SIMS data. Whereas GRR1830 yielded an average
749	$^{19}\text{F}/^{18}\text{O}$ ratio of 0.191, equivalent to 76 $\mu$ g/g F, measurements during the same session on
750	GRR2103 yielded $^{19}\text{F}/^{18}\text{O}$ ratios just barely above the detection limit (<1 µg/g F; Mosenfelder
751	and Rossman 2013a).
752	Coupled H-F substitution is less likely to be important in pyrope, if only because it
753	typically contains significantly lower concentrations of both elements compared to other garnet
754	compositions. Mosenfelder and Rossman (2012) report 0.2 to 10 $\mu$ g/g F in pyrope garnets from
755	mantle xenoliths that contain up to 135 $\mu$ g/g H <sub>2</sub> O (using the IR calibration of Bell et al. 1995).

756	These xenoliths also contain other phases (olivine, orthopyroxene, and/or clinopyroxene) that
757	can partition much of the available F (Mosenfelder et al. 2011; Mosenfelder and Rossman
758	2013a,b). Nearly end-member pyrope from the Dora Maira massif represents an unusual case; its
759	IR spectra are unlike those of other pyrope-almandine garnets, with sharp peaks at 3602, 3641,
760	3651, and 3662 cm <sup>-1</sup> (Rossman et al. 1989). As noted above similar bands are present in a pyrope
761	synthesized by Geiger et al. (1991).
762	On the one hand, Lu and Keppler (1997) attributed the bands in Dora Maira pyrope to O-
763	H defects associated with Li and B substitution, based on Li <sub>2</sub> O and B <sub>2</sub> O <sub>3</sub> concentrations
764	measured by ICP-AES (inductively coupled plasma atomic emission spectroscopy). On the other
765	hand, our attribution of high-frequency bands to clumped OH-F defects is particularly apropos in
766	this case because some Dora Maira garnets contain inclusions of phlogopite with up to 1.6 wt%
767	F as well as other F- and Cl-bearing phases, unequivocally indicating that they formed in a
768	halogen-rich environment (Philippot et al. 1995). Furthermore, SIMS analyses of Dora Maira
769	pyrope by Tenner et al. (2009) – while uncalibrated for F concentration – yielded a $^{19}$ F/ $^{30}$ Si ratio
770	of 0.215, approximately one order of magnitude higher than ratios measured in pyrope garnets
771	from mantle xenoliths (0.008-0.025). A rough estimate of 30 $\mu$ g/g F can be derived by
772	comparing to our SIMS data on pyrope (Mosenfelder and Rossman 2012), including MON9
773	from Bell et al. (1995), which was also measured by Tenner et al.
774	
775	IMPLICATIONS
776	This work highlights technical advances in analytical methods for F and H that could
777	facilitate future studies using garnets to constrain the activities of H <sub>2</sub> O and F-species in fluids in
778	equilibrium with garnet-bearing phase assemblages. The present study – including our new

calibration for the IR molar absorption coefficient in grossular - most directly applies to geologic 779 780 problems in the Earth's crust. However, measurements of trace F and H in garnets from mantle 781 xenoliths may also provide additional constraints on recycling of these elements into the mantle, 782 already a topic of interest addressed by studies of F and H in olivines and pyroxenes. 783 Fine-scale zoning of H in garnets such as those from Asbestos or the garnets studied by 784 Jamtveit and Hervig (1994) or Reynes et al. (2020) is probably best studied using micro-785 analytical techniques such as SIMS or FTIR conducted with a focal plane array (FPA) detector, 786 which is becoming more widely available (the FPA detector has a higher spatial resolution than 787 normal MCT detectors such as those we used; e.g., Reynes et al. 2020). We have also 788 demonstrated that F can be measured reliably with EPMA at levels down to  $\sim$ 50 µg/g, and it 789 should be possible to achieve higher analytical throughput at lower cost with EPMA compared to 790 SIMS. Furthermore, in principle, minimum estimates for F concentrations in garnet could be 791 constrained from FTIR spectra alone, once band assignments for coupled H-F defects are 792 confirmed and controversy over incorporation mechanisms is settled. We expect that theoretical 793 calculations will shed additional light on coupled H-F substitution in garnets and other nominally 794 anhydrous minerals. 795 796 **ACKNOWLEDGMENTS** 797 Financial support for this research is gratefully acknowledged from: NSF grants EAR-798 0947956 and EAR-1322082 to George Rossman, EAR-1347908 to Jed Mosenfelder, NASA 799 80NSSC19K0959 to Marc Hirschmann, the Gordon and Betty Moore Foundation, and the 800 White Rose Foundation. The manuscript was improved by helpful reviews from Andrew Locock,

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- 1087

1088	TABLES
1089	Table 1. Garnet sample localities and FTIR data
1090	<b>Table 2.</b> Results of elastic recoil detection analysis compared to other H calibration techniques
1091	Table 3. SIMS data
1092	Table 4. EPMA data and calculated garnet end members
1093	
1094	FIGURE CAPTIONS
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1096	FIGURE 1. Representative FTIR spectra for grossular spectral classes 2, 2b, 3, 4, and 7. Class
1097	designations follow Rossman and Aines (1991). Spectra are plotted without baseline correction
1098	and offset from each other for clarity. Spectra in <b>a</b> and <b>b</b> are normalized to 1 cm; spectra in <b>c</b> are
1099	normalized to different thicknesses (as labeled) to facilitate comparison. (a) Class 2 (GRR42),
1100	with the strongest peak at 3647 cm <sup>-1</sup> , and class 2b (GRR1386). Class 2b has peaks below 3620
1101	cm <sup>-1</sup> that are almost absent in GRR42. Room temperature spectrum of GRR1386 (solid black
1102	curve) is directly overlain by spectrum taken at 77 K (grey dashed curve). (b) Class 2b
1103	(GRR732) and class 7 (GRR1122). Although these spectra have bands at similar positions, in
1104	class 7 the dominant peaks are at lower frequencies and the strongest peak is at 3599 cm <sup>-1</sup> rather
1105	than 3647 cm <sup>-1</sup> . (c) Class 3 and 4 spectra of natural and synthetic grossular garnets. Three
1106	representative spectra of JLM83a are displayed, with integrated absorbances (Abstot) equal to 691
1107	(dotted line), 1118 (dashed line), and 1526 cm <sup>-2</sup> (solid line). The spectrum of GRR53 is from the
1108	slice labeled "GRR53B-F" displayed in Figure 6 of Rossman and Aines (1991). Spectra of
1109	synthetic, end-member grossular crystals are from Geiger and Armbruster (1997) and Withers et
1110	al. (1998). Spectral classes 3 and 4 are nominally distinguished by whether the strongest peak is
1111	at 3631 or 3621 cm <sup>-1</sup> , respectively. Three of these garnets (GRR53, JLM83a, and the grossular

1112 n	nade bv (	Geiger and A	(Armbruster)	have strong	peaks at	3689.	3677.	3665.	and/or	3559	$cm^{-1}$	that
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are weak or absent in the other two samples (GRR1537 and the grossular synthesized by Withers

1114 et al.).

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1117 **FIGURE 2.** Detailed comparison near the baselines of 77 K and 298 K spectra of GRR1386.

1118 The band at  $\sim$ 3240 cm<sup>-1</sup> at 77 K is attributed to frozen H<sub>2</sub>O in fluid inclusions. Bands attributed

1119 to structural OH<sup>-</sup> groups are off scale. Spectra of liquid water from Bertie and Lan (1996) and ice

1120 from Johnson and Rossman (2003) also plotted for comparison. GRR1386 spectra are

1121 normalized to 1 cm thickness; spectra of ice and water are arbitrarily scaled.

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1123 FIGURE 3. Raw (uncorrected) H<sub>2</sub>O concentrations determined by ERDA for three East African

grossular garnets (GRR732, GRR771, and GRR1756) and blank forsterite (GRR1017), as a

1125 function of integrated absorbance (*Abs*<sub>tot</sub>). Solid line is a York regression. Dashed lines represent

1126 95% confidence intervals calculated from *a priori* errors. Equation and mean squared weighted

deviation (MSWD) of the fit are also shown.

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**FIGURE 4.** <sup>16</sup>OH/<sup>18</sup>O data from two SIMS sessions. Error bars not shown are within the symbol size. Solid lines are York regressions fit to data (circles) on the blank and East African samples only, with 95% confidence intervals (calculated from *a priori* errors) displayed as dashed curves. The equation and MSWD for each regression are also displayed in the graphs. (a) 2012 data. Each point represents the average of three to five measurements, with  $2s_x$  (sample standard deviation) error bars displayed for <sup>16</sup>OH/<sup>18</sup>O. (b) 2016 data. In this plot, each point represents a

- single analysis. Consequently, uncertainties in  ${}^{16}\text{OH}/{}^{18}\text{O}$  (2 $s_{\bar{x}}$ , standard error of the mean) are all within the symbol size in this plot. Some data points are obscured by overlap with other data; all original data are given in the supplementary material.
- 1138
- 1139 FIGURE 5. Backscattered electron image of JLM83a after SIMS analysis. H<sub>2</sub>O and F
- 1140 concentrations (in  $\mu$ g/g) are displayed above (up and to the left) each SIMS crater (black spot).
- 1141 Analysis numbers in order of collection are labeled to the lower right of each spot.
- 1142 The analyses with asterisks (analyses 15-18) were compromised by elevated background <sup>16</sup>OH<sup>-</sup>
- and <sup>19</sup>F<sup>-</sup>, as discussed in the text. Cycle profiles (uncalibrated depth profiles) for analyses 12 and
- 1144 20 are displayed in supplementary Figure S8.
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FIGURE 6. Correlations between <sup>16</sup>OH and <sup>19</sup>F for grossular and spessartine. (a) SIMS data for 1146 1147 grossular from sessions in 2012 (open symbols) and 2016 (closed symbols). Each symbol 1148 represents a single analysis ( $2s_{\bar{x}}$  uncertainties are within the symbol size for each point). Solid lines are ordinary least-squares (OLS) regressions to data for East African grossular garnets ( $r^2 =$ 1149 0.98) and multiple analyses on Asbestos grossular JLM83a ( $r^2 = 0.99$ ). Data on JLM83a that were 1150 1151 compromised by contamination (Fig. 5) are not included in this plot. (b) Spessartine data from 1152 2012 session. Each data point on this plot represents an average of between four and 11 analyses for one sample ( $2s_x$  uncertainties are within the symbol size). Solid line is an OLS regression ( $r^2$ 1153 1154 = 0.99) to all of the data.

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1157	FIGURE 7.	Comparison	of ERDA res	sults (circles)	to previous	calibration	data on Asbestos
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- grossular (GRR53 and GRR53B, triangles; Aines and Rossman 1984; Rossman and Aines 1991),
- 1159 Maldener et al. (2003, diamonds), and O'Leary et al. (2007, square). Only two data points from
- the compilation of Rossman and Aines (1991) are plotted (for Asbestos garnets GRR53 and
- 1161 GRR53F), but the original regression determined for their entire data set of grossular and
- 1162 hydrogrossular garnets over a much wider range of H contents (up to 12.75 wt% H<sub>2</sub>O) is
- displayed as a dashed line. The ERDA data have been blank corrected in this graph by
- subtracting 100  $\mu$ g/g from the raw values plotted in Figure 3.
- 1165
- 1166 **FIGURE 8.** Results of spectral curve fitting plotted as function of calculated F (per formula unit).
- 1167 Frequencies of fitted bands vary by up to 3 cm<sup>-1</sup>. Symbols (see inset legend in **a**) correspond to
- those in Figure 7. (a) Integrated absorbance vs. F for fitted bands centered at ~3631, 3643, 3647,
- 1169 3657, 3664, 3674, and 3688 cm<sup>-1</sup>. (**b**) Integrated absorbance vs. F for fitted bands centered at
- 1170  $\sim$  3530, 3539, 3545, 3560, 3567, 3582, 3600, 3608, and 3622 cm<sup>-1</sup>.
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### FIGURES



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1182	FIGURE 1. Representative FTIR spectra for grossular spectral classes 2, 2b, 3, 4, and 7. Class
1183	designations follow Rossman and Aines (1991). Spectra are plotted without baseline correction
1184	and offset from each other for clarity. Spectra in <b>a</b> and <b>b</b> are normalized to 1 cm; spectra in <b>c</b> are
1185	normalized to different thicknesses (as labeled) to facilitate comparison. (a) Class 2 (GRR42),
1186	with the strongest peak at 3647 cm <sup>-1</sup> , and class 2b (GRR1386). Class 2b has peaks below 3620
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1188	curve) is directly overlain by spectrum taken at 77 K (grey dashed curve). (b) Class 2b
1189	(GRR732) and class 7 (GRR1122). Although these spectra have bands at similar positions, in
1190	class 7 the dominant peaks are at lower frequencies and the strongest peak is at 3599 cm <sup>-1</sup> rather
1191	than 3647 cm <sup>-1</sup> . (c) Class 3 and 4 spectra of natural and synthetic grossular garnets. Three
1192	representative spectra of JLM83a are displayed, with integrated absorbances ( $Abs_{tot}$ ) equal to 691
1193	(dotted line), 1118 (dashed line), and 1526 cm <sup>-2</sup> (solid line). The spectrum of GRR53 is from the
1194	slice labeled "GRR53B-F" displayed in Figure 6 of Rossman and Aines (1991). Spectra of
1195	synthetic, end-member grossular crystals are from Geiger and Armbruster (1997) and Withers et
1196	al. (1998). Spectral classes 3 and 4 are nominally distinguished by whether the strongest peak is
1197	at 3631 or 3621 cm <sup>-1</sup> , respectively. Three of these garnets (GRR53, JLM83a, and the grossular
1198	made by Geiger and Armbruster) have strong peaks at 3689, 3677, 3665, and/or 3559 cm <sup>-1</sup> that
1199	are weak or absent in the other two samples (GRR1537 and the grossular synthesized by Withers
1200	et al.).
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1209 The band at  $\sim$ 3240 cm<sup>-1</sup> at 77 K is attributed to frozen H<sub>2</sub>O in fluid inclusions. Bands attributed

1210 to structural OH<sup>-</sup> groups are off scale. Spectra of liquid water from Bertie and Lan (1996) and ice

1211 from Johnson and Rossman (2003) also plotted for comparison. GRR1386 spectra are

1212 normalized to 1 cm thickness; spectra of ice and water are arbitrarily scaled.

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FIGURE 3. Raw (uncorrected) H<sub>2</sub>O concentrations determined by ERDA for three East African
grossular garnets (GRR732, GRR771, and GRR1756) and blank forsterite (GRR1017), as a

1217 function of integrated absorbance (*Abs*tot). Solid line is a York regression. Dashed lines represent

1218 95% confidence intervals calculated from *a priori* errors. Equation and mean squared weighted



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**FIGURE 4.** <sup>16</sup>OH/<sup>18</sup>O data from two SIMS sessions. Error bars not shown are within the symbol 1225 1226 size. Solid lines are York regressions fit to data (circles) on the blank and East African samples 1227 only, with 95% confidence intervals (calculated from *a priori* errors) displayed as dashed curves. 1228 The equation and MSWD for each regression are also displayed in the graphs. (a) 2012 data. Each point represents the average of three to five measurements, with  $2s_x$  (sample standard 1229 deviation) error bars displayed for  ${}^{16}$ OH/ ${}^{18}$ O. (b) 2016 data. In this plot, each point represents a 1230 single analysis. Consequently, uncertainties in <sup>16</sup>OH/<sup>18</sup>O ( $2s_{\bar{x}}$ , standard error of the mean) are all 1231 1232 within the symbol size in this plot. Some data points are obscured by overlap with other data; all 1233 original data are given in the supplementary material. 1234

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1242 FIGURE 5. Backscattered electron image of JLM83a after SIMS analysis. H<sub>2</sub>O and F

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1246 and <sup>19</sup>F<sup>-</sup>, as discussed in the text. Cycle profiles (uncalibrated depth profiles) for analyses 12 and

1247 20 are displayed in supplementary Figure S8.

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FIGURE 6. Correlations between <sup>16</sup>OH and <sup>19</sup>F for grossular and spessartine. (a) SIMS data for 1250 1251 grossular from sessions in 2012 (open symbols) and 2016 (closed symbols). Each symbol 1252 represents a single analysis ( $2s_{\bar{x}}$  uncertainties are within the symbol size for each point). Solid lines are ordinary least-squares (OLS) regressions to data for East African grossular garnets ( $r^2 =$ 1253 0.98) and multiple analyses on Asbestos grossular JLM83a ( $r^2 = 0.99$ ). Data on JLM83a that were 1254 1255 compromised by contamination (Fig. 5) are not included in this plot. (b) Spessartine data from 1256 2012 session. Each data point on this plot represents an average of between four and 11 analyses for one sample ( $2s_x$  uncertainties are within the symbol size). Solid line is an OLS regression ( $r^2$ 1257 1258 = 0.99) to all of the data. 1259 1260 1261

- 1262
- 1263
- 1264



1265



1276





0

0.00

0.01

0.02

0.03

0.04

F (pfu)

0.05

1280 **FIGURE 8.** Results of spectral curve fitting plotted as function of calculated F (per formula unit).

0.07

.....

0.01

0.02

0.03

F (pfu)

0.04

0.05

0.06

0.07

0

0.00

1281 Frequencies of fitted bands vary by up to 3 cm<sup>-1</sup>. Symbols (see inset legend in **a**) correspond to

0.06

1282 those in Figure 7. (a) Integrated absorbance vs. F for fitted bands centered at ~3631, 3643, 3647,

1283 3657, 3664, 3674, and 3688 cm<sup>-1</sup>. (**b**) Integrated absorbance vs. F for fitted bands centered at

1284  $\sim$  3530, 3539, 3545, 3560, 3567, 3582, 3600, 3608, and 3622 cm<sup>-1</sup>.

- 1285
- 1286
- 1287
- 1288
- 1289

Table 1: Garnet	t sample localities and <b>F</b>	TIR data			
Sample No.	Locality	Spectral class <sup>a</sup>	Abs tot (cm <sup>-2</sup> )	Mean wavenumber (cm <sup>-1</sup> )	
GRR42	Ramona Mine, California, U.S.A.	2	1430(14)	3647	
	Stream gravel, Merelani Hills,				
GRR229	Tanzania	2b	1909(38)	3623	
GRR732	Mindi Hills, Kenya	2b	1403(28)	3617	
	Merelani Hills,				
GRR771	Tanzania	2b	2441(73)	3622	
	Crestmore Mine,				
GRR1122	California, U.S.A.	7	1471(44)	3597	
	Stream gravel, Merelani Hills,		2.500 (52)	2 (22)	
GRR1386	Tanzania	26	3589(72)	3633	
0001307	Stream gravel, Merelani Hills,	21	1502(40)	2(10	
GRR1386a	l anzania	26	1592(48)	3619	
GRR1429	New York, U.S.A.	7	493(38)	3570-3579 <sup>b</sup>	
	#2 reef, Scorpion Mine, Voi, Taita,				
GRR1756	Kenya	7	591(12)	3595	
JLM83a	Asbestos, Quebec, Canada	3	691-1544°	3630-3631°	
Notes: uncertaint	ties in parentheses are est	timated as discus	sed in text.		
<sup>a</sup> Classification se	cheme of Rossman and A	Aines (1991)			
<sup>b</sup> Range of values	s for 7 spectra				
<sup>c</sup> Range measure	d in 48 spectra				

Sample No.	Phase	Abs <sub>tot</sub> (cm <sup>-2</sup> )	$H_2O, \mu g/g$		Reference	Method
			Uncorrected	Blank corrected		
GRR1756	Grossular	591(12)	308(40)	208(40)	This study	ERDA
GRR732	Grossular	1403(28)	599(76)	499(76)	This study	ERDA
			—	1062(31)	1	CFMS
GRR771	Grossular	2441(73)	1064(130)	964(130)	This study	ERDA
NSL1	Rhyolitic glass		11,133(1399)	11,033 (1399)	This study	ERDA
			_	11,600	2	KFT-FTIR
N3	Rhyolitic glass		33,658 (4209)	33,558 (4209)	This study	ERDA
			_	33,000	2	KFT-FTIR
N5	Rhyolitic glass		50,549 (6313)	50,449 (6313)	This study	ERDA
			—	50,600	2	KFT-FTIR
GRR1017 <sup>a</sup>	Forsterite	0	101(14)	0	This study	ERDA
GRR53	Grossular	2877	—	1800	3, 4	MEA, MAN
GRR53F	Grossular	4796	_	2800	4	<sup>19</sup> F NRA
HESS1	Grossular	2510	_	870	5	<sup>15</sup> N NRA
		2006	_	950	5	<sup>15</sup> N NRA
TSAV	Grossular	1208	_	480	5	<sup>15</sup> N NRA
MALI	Grossular	168	_	170	5	<sup>15</sup> N NRA
GRMALI	Grossular	77	_	190	5	<sup>15</sup> N NRA

#### Table 2: Results of elastic recoil detection analysis (ERDA) compared to other H calibration techniques

*Notes* : References: 1 = O'Leary et al. (2007); 2= Tenner et al. (2009); 3 = Aines and Rossman (1984); 4 = Rossman and A CFMS = continuous-flow mass spectrometry; KFT-FTIR = Karl-Fischer titration and FTIR; MEA =  $P_2O_5$  coulometry (mo Uncertainties in parentheses are 2s in terms of least units cited

<sup>a</sup>Blank reference material used to correct ERDA data.

.ines (1991); 5 = Maldener et al. (2003) visture evolution analyzer); MAN = evolved  $H_2$  gas volume (manometry); NRA = nuclear reaction analysis

Sample no.	SIMS session	No. of analyses	$H_2O$ , $\mu g/g^a$	<sup>16</sup> O <sup>1</sup> H/ <sup>18</sup> O
GRR1122-HT	2012	4	0	0.000338 - 0.000450 <sup>c</sup>
	2016	3	0	0.001365 - 0.001904 <sup>c</sup>
GRR42	2012	5	529(50)	0.1244(14)
	2016	$2^{d}$		0.1300, 0.1302
GRR229	2012	4	706(68)	0.1262(16)
	2016	3		0.1304(49)
GRR732	2012	4	519(50)	0.0885(16)
GRR771	2012	4	903(90)	0.155(13)
	2016	3		0.150(15)
GRR1122	2012	4	544(54)	0.0956(12)
	2016	$2^d$		0.1183, 0.1247
GRR1386	2012	4	1328(128)	0.2283(31)
	2016	3		0.2302(6)
GRR1386a	2012	4	589(58)	0.1007(23)
	2016	3		0.1016(10)
GRR1429	2012	4	182(22)	0.0303(34)
GRR1756	2012	4	219(21)	0.0349(2)
	2016	$2^d$		0.349, 0.349 <sup>d</sup>
JLM83a <sup>e</sup>	2016	16 <sup>f</sup>	256-571	0.0296-0.1002

### Table 3: SIMS data

Notes: Numbers in parentheses are 2s uncertainties in terms of least units cited.

<sup>a</sup> Values calculated from FTIR using calibration factor of 0.37 (Eq. 1), with uncertainties propag

<sup>b</sup> Values calculated from <sup>16</sup>O<sup>1</sup>H/<sup>18</sup>O using working curve for F-bearing glasses.

<sup>c</sup> Uncorrected ratios, given as a range.

<sup>d</sup> For samples with only two analyses, each analysis is given (separated by comma).

<sup>e</sup>Values for this sample are given as a range (for both FTIR and SIMS analyses).

<sup>f</sup> Four analyses were rejected owing to a spike in the blank, as described in the text.

$^{19}\text{F}/^{18}\text{O}$	F, $\mu g/g^{b}$
1.279(21)	508 (8)
1.325(40)	537(84)
6.563(69)	2606(27)
6.529, 6.584	2559, 2581
3.713(48)	1474(19)
3.751(57)	1470(22)
2.819(49)	1119(19)
4.14(19)	1644(75)
3.99(10)	1565(39)
1.281(16)	508(6)
1.295, 1.447	508, 567
6.824(49)	2709(19)
6.694(64)	2624(25)
2.422(48)	962(19)
2.423(16)	950(6)
0.465(34)	185(13)
0.602(4)	239(2)
$0.611, 0.615^{d}$	239, 241
0.3502-0.8789	120-345

gated as described in text.

Sample No.	GRR42	GRR229	GRR732	GRR771	GRR1122
SiO <sub>2</sub>	38.59(24)	38.77(24)	40.15(39)	39.84(18)	38.69(31)
TiO <sub>2</sub>	_	0.40(3)	0.34(3)	0.54(2)	0.57(6)
$AI_2O_3$	21.43(12)	20.94(13)	21.97(28)	22.34(8)	17.64(7)
$Cr_2O_3$	_	0.28(2)	0.04(3)	—	_
$V_2O_3$	_	1.90(3)	0.88(1)	0.05(2)	_
FeO <sup>a</sup>	3.23(2)	0.08(1)	0.06(3)	0.14(2)	7.53(26)
MnO	0.10(2)	0.90(5)	1.15(1)	0.27(3)	0.38(4)
MgO	—	0.47(2)	0.54(3)	0.34(1)	0.07(2)
	35.33(6)	35.10(14)	35.44(33)	36.25(10)	34.61(20)
H <sub>2</sub> O	0.0723	0.0700	0.0519	0.0903	0.0344
F	0.2477(36)	0.1516(30)	0.1119°	0.1470(62)	0.0586(62)
	-0.1	-0.06	-0.05	-0.06	-0.02
TOTAL	90.90	99.00	100.00	99.95	99.50
Recalculated					
FeO	1.16	0.08	0.06	0.14	2.09
$Fe_2O_3$	2.30	-	-	_	6.05
$Recalculated Total^{f}$	99.13	99.00	100.68	99.95	100.19
Si	2 9457	2 9576	3 0015	2 9906	2 9804
Ti	_	0.0230	0.0191	0.0305	0.0330
$AI^{VI}$	1.8978	1.8583	1.9357	1.9764	1.5924
$Al^{\mathrm{IV}}$	0.0301	0.0243	_	_	0.0091
Cr	_	0.0169	0.0024	_	_
V	_	0.1162	0.0527	0.0030	_
Fe <sup>2+</sup>	0.0738	0.0051	0.0038	0.0088	0.1346
Fe <sup>3+</sup>	0.1324	_	_	_	0.3506
Mn	0.0065	0.0582	0.0728	0.0172	0.0248
Mg	-	0.0534	0.0602	0.0380	0.0080
La H	2.0090	2.0009	2.0307	2.9155	2.0000
	11 0402	11 0600	12 0027	11 0050	11 9857
F	0.0598	0.0366	0.0265	0.0349	0.0143
Fnd Members					
Katoite	0.31%	0.30%	0.22%	0.38%	0.23%
FCa garnet	0.50%	0.30%	0.22%	0.29%	0.12%

## Table 4: EPMA data and calculated garnet end members

Schorlomite-Al	_	1.15%	_	_	0.45%
Morimotoite	_	_	_	_	2.40%
Majorite	_	_	_	_	_
Goldmanite	_	5.81%	2.64%	0.15%	_
Uvarovite	_	0.84%	0.12%	_	_
Spessartine	0.22%	1.94%	2.43%	0.57%	0.83%
Pyrope	_	1.78%	2.01%	1.27%	0.27%
Almandine	2.46%	0.17%	0.13%	0.29%	3.69%
Grossular	91.41%	87.22%	91.43%	96.02%	74.49%
Andradite	4.11%	_	_	_	17.53%
Remainder	1.00%	0.48%	0.82%	1.03%	0.00%
Total	100.01%	99.99%	100.02%	100.00%	100.01%

*Notes:* Cation assignments and garnet end members calculated using Locock (2008), on an 8 See Locock (2008) for definition of end members. "Remainder" represents the proportion of ca Dashed lines denote oxides, elements or end members below detection limit or calculated to t EPMA data represent averages (in wt%, with 1*s* uncertainties in parentheses) of 5 to 6 analys

<sup>a</sup> Total Fe calculated as FeO.

<sup>b</sup>H<sub>2</sub>O calculated from 2012 SIMS data, using fit to data for East African grossulars.

<sup>c</sup>H<sub>2</sub>O calculated from FTIR data (Table 3).

<sup>d</sup>H<sub>2</sub>O and F based on 2016 SIMS data, for two analyses at extreme ends of range.

<sup>e</sup>F based on 2012 SIMS data.

<sup>f</sup>Recalculated wt% total after calculation of  $Fe^{2^+}/Fe^{3^+}$ .

GRR1386	GRR1386a	GRR1429	GRR1756	JLM83a	A
				low	high
39.19(23)	39.96(5)	39.48(16)	39.70(11)	39.95(18)	39.95(18)
0.64(3)	0.36(3)	0.42(3)	0.40(1)	_	_
22.29(7)	21.90(5)	18.62(13)	20.90(12)	22.23(13)	22.23(13)
_	_	_	0.29(2)	_	_
_	_	_	1.77(5)	_	_
0.20(3) 0.14(2) 0.43(1) 36.27(11)	1.36(5) 0.33(2) 0.31(1) 35.78(6)	7.27(9) 0.21(2) 0.14(1) 34.11(9)	0.06(1) 0.97(1) 0.57(1) 35.02(5)	1.82(10) 0.24(3) 	1.82(10) 0.24(3) 
0.1328	0.0589	0.0182	0.0219	0.0171	0.0576
0.2478(24) -0.10 99.44	0.0924(29) -0.04 100.11	0.0185 <sup>°</sup> -0.01 100.28	0.0343(42) -0.01 99.73	0.012ª -0.01 99.77	0.0345 <sup>ª</sup> -0.01 99.83
0.2	1.36 _	3.48 4.22	0.06	1.82 _	1.82
99.44	100.11	100.71	99.73	99.77	99.83
2 9536	3 0037	3 0137	3 0079	3 0162	3 0138
0.0363	0.0204	0.0241	0.0228	_	_
1.9650	1.9401	1.6752	1.8663	1.9781	1.9765
0.0149	_	_	_	_	_
			0.0174 0.1075		-
0.0126	0.0855	0.2219	0.0038	0.1149	0.1148
_	_	0.2422	_	_	_
0.0089 0.0483 2.9289	0.0210 0.0310 2.8817	0.0136 0.0022 2.7899	0.0622 0.0644 2.8429	0.0153 _ 2.8726	0.0153  2.8703
0.0167	0.0074	0.0023	0.0028	0.0022	0.0072
11.9523 0.0591	11.9850 0.0220	11.9955 0.0045	12.0229 0.0082	12.0053 0.0029	12.0031 0.0082
0.56% 0.49%	0.25% 0.18%	0.08% 0.04%	0.09% 0.07%	0.07% 0.02%	0.24% 0.07%

_	_	_	_	_
2.04%	2.41%	_	_	_
0.37%	0.07%	_	_	_
_	_	5.38%	_	_
_	_	0.87%	_	_
0.70%	0.45%	2.07%	0.51%	0.51%
0.66%	_	2.15%	_	_
1.78%	6.36%	0.13%	3.64%	3.50%
93.43%	76.83%	88.36%	94.66%	94.50%
_	12.11%	_	_	_
0.59%	1.65%	0.89%	1.10%	1.18%
100.00%	100.00%	100.01%	100.00%	100.00%
	- 2.04% 0.37% - - 0.70% 0.66% 1.78% 93.43% - 0.59% 100.00%	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$

cation basis (treating H as  $H_4^{4+}$ ).

ations that could not be allocated.

)e 0.

es for all samples except JLM83a (average of 36 analyses).