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1 **Reservoir depressurization driven by passive gas**
2 **emissions at Ambrym volcano**

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7 **Highlights:**

- 8 • Persistent volcanic degassing explains steady subsidence at Ambrym from 2015–
9 2017
- 10 • Subsidence ends in 2017, as magma influx balances depressurization from degassing
- 11 • We estimate a magma influx of $\sim 0.16 \text{ km}^3$ over 14 months before the 2018 erup-
12 tion

13 **Keywords:** Subsidence, Degassing processes, Magma replenishment

Abstract

14 **Abstract**
15 Despite being a widespread and common process, the impact of passive volcanic degassing
16 on the pressurization state of a magma reservoir is not well understood. If mass loss due
17 to gas emissions results in reservoir depressurization and surface subsidence, the pres-
18 sure difference between a shallow reservoir and deep magma source may result in magma
19 recharge and eventually trigger an eruption. It is therefore important to determine how
20 a simplified reservoir-conduit system responds to such degassing processes. Here we use
21 an extreme example of persistent volcanic degassing—Ambrym—as a case study to re-
22 late sulphur dioxide mass flux with reservoir depressurization and edifice-scale subsidence,
23 both measured from satellite-based remote sensing observations. A geodetic inversion
24 of surface displacements measured with Interferometric Synthetic Aperture Radar mod-
25 eled using the Boundary Element Method provides bounds on the reservoir pressure change
26 during an episode of subsidence at Ambrym from 2015 to 2017. These results are input
27 into a lumped parameter theoretical model developed by Girona et al. (2014), and the
28 free parameters (e.g., reservoir size and conduit radius) are systematically explored. We
29 find that the 2015–2017 subsidence episode is consistent with pressure decreasing at a
30 rate of -5.2 to -2.0 MPa year⁻¹ in a reservoir at ~ 2 km b.s.l., as a result of passive de-
31 gassing. The subsidence episode is observed to end abruptly in October 2017, and no sig-
32 nificant deformation is detected in the 14 months leading up to a rift zone intrusion and
33 submarine eruption in December 2018, despite substantial degassing. We explain this
34 lack of pre-eruptive deformation by an influx of ~ 0.16 km³ of magma into a shallow ($<$
35 2 km b.s.l.) reservoir that counterbalances the depressurization caused by degassing. This

36 recharge volume is comparable with the volume of magma subsequently extracted from
37 Ambrym’s reservoir in December 2018. We conclude that at some open-vent passively
38 degassing volcanoes, deflation caused by degassing may reduce or even cancel any infla-
39 tion signal caused by magma influx. Nonetheless, detection of pre-eruptive recharge can
40 be achieved by monitoring changes in the long-term deformation rate.

41 1 Introduction

42 Persistent passive degassing is observed at volcanoes worldwide (Carn et al., 2017).
43 From 2005–2015, of the 1500 terrestrial volcanoes with possible activity in the Holocene
44 (Siebert et al., 2015), at least 6% (91) were passively degassing above the detection thresh-
45 old of satellite measurements (~ 30 tons day $^{-1}$) (Carn et al., 2017). Degassing-related
46 hazards depend on gas composition, dispersion, and emission duration. In cases of per-
47 sistent passive (non-eruptive) degassing, areas of >1000 km 2 (e.g., Masaya volcano, Nicaragua)
48 can be affected by acid rain or noxious volcanic smog, as well as fluoride contamination
49 of water (Williams-Jones & Rymer, 2015; Cronin & Sharp, 2002; Allibone et al., 2012).

50 Cases of sustained high gas flux are mainly observed at volcanoes characterized by
51 open-system degassing. In such systems, exsolved gas within a magma conduit or reser-
52 voir segregates from the melt and is released into the atmosphere during non-eruptive
53 periods (Cashman, 2004). Theoretical models show that gas emission can result in a pres-
54 sure change of a magmatic reservoir, and that pressure changes can also influence gas
55 exsolution (Tait et al., 1989; Boichu et al., 2008, 2011; Anderson & Segall, 2011; Kaza-
56 haya et al., 2015). As we show in this study, contrary to the inter-eruptive uplift typ-

57 ical of the ‘classic volcano deformation cycle’ (Biggs & Pritchard, 2017), eruptions may
58 be counterintuitively preceded by periods of reservoir depressurization (Girona et al., 2014,
59 2015).

60 From 2005–2015, Ambrym (Vanuatu), Kīlauea (Hawai’i), and Nyiragongo (Demo-
61 cratic Republic of Congo) volcanoes were among the ten strongest passively degassing
62 volcanoes on Earth, influenced by the presence of lava lake activity (Carn et al., 2017).
63 At these open-vent volcanoes, surface deformation can be measured by satellite using
64 interferometric synthetic aperture radar (InSAR). While deformation can be used to con-
65 strain reservoir overpressure through geodetic modeling, estimates must account for un-
66 certainties on the host rock elastic moduli and reservoir geometry. Combining InSAR
67 deformation measurements with remote sensing gas measurements may provide an open-
68 ing for long-term studies linking changes in reservoir pressure (and potentially uplift or
69 subsidence) and volcanic degassing. Our study demonstrates this possibility using a mul-
70 tiparametric approach to investigate the effect of volcanic degassing on reservoir pres-
71 surization at Ambrym volcano in Vanuatu from 2015–2018. Ambrym is a basaltic vol-
72 canic island which hosts a 12-km wide caldera and two main volcanic cones (Marum and
73 Benbow), within which are semi-permanent lava lakes that have been active since at least
74 the 1980’s (Robin et al., 1993; Allard et al., 2015). In particular, we investigate a period
75 of inter-eruptive ground subsidence, coinciding with high gas flux, at Ambrym. The sub-
76 sidence episode preceded a major intrusion crisis that led to the extinction of the lava
77 lakes for more than 2.5 years (at the time of writing, February 2022) (Shreve et al., 2019),
78 after they had been active for at least 3 decades.

79 To explore physical mechanisms explaining ground subsidence at Ambrym, we first
80 estimate reservoir depressurization rates using geodetic modeling. Then, we use a lumped
81 parameter theoretical model that couples gas emissions with reservoir depressurization,
82 developed in Girona et al. (2014), to explain the estimated depressurization rates and
83 their variation over time. This is the first time the theoretical model of Girona et al. (2014)
84 is applied using observations at an active volcano. The effect of magma replenishment
85 will also be considered to explain an abrupt change in depressurization rate that occurred
86 around October 2017, 14 months before a large dike intrusion and submarine eruption
87 that occurred in December 2018 (Shreve et al., 2019; Hamling et al., 2019; Moussallam
88 et al., 2021). We posit that pre-eruptive reservoir pressurization, associated with magma
89 recharge, may have been balanced by reservoir depressurization due to passive degassing,
90 resulting in a net pressure change of zero prior to the 2018 rift zone eruption, explain-
91 ing the lack of pre-eruptive uplift.

92 **2 Data and Methods**

93 **2.1 2015–2017 Subsidence Episode and 2018 Eruption**

94 InSAR measures ground displacements after a VEI 2 eruption in February 2015,
95 when an intra-caldera vent opened and fed a lava flow (Figure 1a) (Hamling & Kilgour,
96 2020; Shreve et al., 2021). A time series is processed with the software MintPy (Yunjun
97 et al., 2019) using images acquired by the European Space Agency’s satellite constella-
98 tion Sentinel-1 (S1) (Figure 1a,b and Section 2.3). The first available Sentinel-1 image
99 was acquired on 30 October 2015. The S1 time series and an interferogram from the Japanese

100 Space Agency’s satellite ALOS-2 spanning 21 March 2015 to 28 October 2017 both es-
101 timate a maximum subsidence velocity of ~ 10 cm year⁻¹. An additional ALOS-2 inter-
102 ferogram indicates the subsidence began between March and June 2015 (Figure S1), af-
103 ter the February 2015 eruption (Section 2.3). The subsidence continued until October
104 2017, after which it ended abruptly (Figure 1b).

105 Starting 14 December 2018, 14 months after the end of the subsidence episode, an
106 intra-caldera eruption began at Ambrym, followed by a rift zone intrusion and subma-
107 rine eruption (Shreve et al., 2019; Hamling et al., 2019; Moussallam et al., 2021). Am-
108 brym’s caldera subsided by up to 3 meters, and the lava lakes were completely drained
109 (Shreve et al., 2019; Hamling et al., 2019). Although >0.4 km³ of magma was intruded
110 into the rift zone (Shreve et al., 2019; Hamling et al., 2019), causing drastic changes to
111 the activity at Ambrym (e.g., lava lake drainage), less than 5 cm of precursory deforma-
112 tion occurred in the year before the event (Figure 1b). Despite a lack of pre-eruptive up-
113 lift, two primary observations indicate an increase in magma and gas supply to the sys-
114 tem, which may have been precursors to the 2018 eruption. First, a new thermal anomaly
115 ~ 1 km to the south of the main lava lake hosted in Marum crater is detected using Sentinel-
116 2 images sometime between 23 June 2016 and 10 March 2018 (Video S1). This indicates
117 either a new vent for gas emissions or the presence of magma close to the surface, although
118 discriminating between the two is ambiguous without additional observations. While the
119 new vent may be a redirection of gas or magma that was already present at depth, its
120 surface manifestation indicates a deviation from background activity. Second, the lava
121 lake level rose rapidly in the weeks before the 2018 eruption, as documented by Moussallam

122 et al. (2021), which can indicate an increase in magma influx to the system, according
123 to Patrick et al. (2019); Burgi et al. (2020); Global Volcanism Program (2017c).

124 2.2 2015–2018 SO₂ Gas Emissions

125 To estimate temporal variation of the sulphur dioxide (SO₂) budget emitted by Am-
126 brym, we calculate the daily SO₂ mass by integrating OMI/Aura SO₂ Total Column over
127 a 150 km-radius circular area centered on Ambrym. We use the Level 2 Version 003 lower
128 tropospheric (TRL) products (corresponding to a center of mass altitude at 3 km (Li et
129 al., 2020)), keeping pixels with a cloud fraction lower than 30% and a column amount
130 greater than 0.3 DU, and removing acquisitions affected by OMI row anomalies. We also
131 remove acquisitions affected by emissions from nearby Ambae volcano in 2017–2018, by
132 discarding dates coinciding with a thermal anomaly at Ambae by MODIS (from MOD-
133 VOLC (Wright et al., 2004)), and by applying an additional automated procedure re-
134 lying on the ratio between the gas mass around Ambrym and Ambae (see details in Text
135 S1).

136 We observe that the daily SO₂ mass burden released passively by Ambrym remains
137 nearly constant between the February 2015 and December 2018 eruptions (Figure 1b).
138 As shown in Figure S2, this observation is consistent with the steadiness of the annual
139 mean SO₂ flux over the 2015–2018 time interval according to the NASA database ([https://](https://so2.gsfc.nasa.gov/kml/OMI_Catalogue_Emissions_2005-2019.xls)
140 so2.gsfc.nasa.gov/kml/OMI_Catalogue_Emissions_2005-2019.xls, updated follow-
141 ing Carn et al. (2017)). The consistency between SO₂ mass and flux also holds from 2005
142 through to 2019, even if SO₂ flux varied substantially before 2015 (Figure S2). This ex-

143 cellent consistency indicates that our estimates of SO_2 mass burden at Ambrym, once
 144 averaged over several months, are a good proxy of, and nearly proportional to, the SO_2
 145 flux. Since January 2019, SO_2 emissions have reduced to below the detection threshold
 146 of satellites, in stark contrast with Ambrym’s persistent degassing over the past decade.

147 We conclude that the SO_2 flux at Ambrym remained essentially constant in the 2015–
 148 2018 time interval. Because the major volcanic gas species is H_2O , we assume that the
 149 mean SO_2 gas flux in the 2015–2018 time interval (7 kt day^{-1}) can reasonably approx-
 150 imate the mean water vapor gas flux using the mass ratio of $\frac{\text{H}_2\text{O}}{\text{SO}_2} \approx 15$ from the aver-
 151 age molar ratios obtained during two field campaign measurements of Ambrym’s bulk
 152 gas emissions in October 2007 (Table 1 in Allard et al. (2015)). Since there are no es-
 153 timates of this ratio during the time period of interest, we assume it is constant and dis-
 154 cuss the possible effects of a 50% increase or decrease in H_2O flux on our results in Sec-
 155 tion 4.1.1. We proceed under the assumption that $\hat{Q} = \hat{Q}_{\text{H}_2\text{O}} \approx 110 \text{ kt day}^{-1}$.

156 2.3 Geodetic Inversion

157 To estimate magma reservoir depressurization, we invert surface displacements mea-
 158 surements spanning March 2015 to October 2017 measured by ALOS-2 and S1. Inter-
 159 ferograms were processed using the Interferometric SAR scientific computing environ-
 160 ment (ISCE) (Rosen et al., 2012). We multilook ascending S1 interferograms (Track 81)
 161 36 times in range and 12 times in azimuth, filter with a power spectrum filtering strength
 162 of 0.3 and unwrap with SNAPHU (Chen & Zebker, 2001). The ascending ALOS-2 in-
 163 terferograms (Track 101) are multilooked 8 times in range and 16 times in azimuth, fil-

164 tered with a power spectrum filtering strength of 0.2, and unwrapped with a coherence-
165 based NSBAS module (“MPD”) described in Shreve et al. (2021). A 12 m TanDEM-X
166 Global DEM is used to remove the topographic phase contribution (Wessel, 2016).

167 We calculate an S1 InSAR small baseline time series using MintPy with 63 acqui-
168 sitions to generate and invert 169 interferograms in radar geometry, with temporal base-
169 lines ranging from 12 days to 1 year (interferogram network shown in Figure S3) (Yunjun
170 et al., 2019). It is common practice to fit and remove a phase ramp, which may be due
171 to orbital or timing inaccuracies. However, we skip this step because the lack of a non-
172 deforming reference area results in an unstable ramp estimate, which biases the displace-
173 ment measurements. A comparison of different inversion strategies indicates no mean-
174 ingful difference between an inversion with uniform weighting and an inversion weighted
175 by the inverse of the phase variance, most likely due to the redundancy of the dense in-
176 terferometric network (Yunjun et al., 2019) (Figures S3, S4, and Text S3). The two in-
177 version strategies result in different temporal coherence estimates, which determine pixel
178 masking (Figure S4), and we proceed with the time series weighted by the inverse of the
179 phase variance. Finally, we correct the time series for atmospheric effects using the ERA-
180 5 atmospheric model implemented in MintPy with the PyAPS software (Yunjun et al.,
181 2019). The ERA-5 correction is an improvement to the empirical correction based on
182 the relationship between phase delay and elevation (also available in MintPy), which causes
183 spurious jumps in the time series (Figure S5).

184 Figure 1b shows the displacement time series (acquisitions every 12 days) for var-
185 ious locations at Ambrym from 30 October 2015 to 13 December 2018. Subsidence is mea-

186 sured, elongated SE-NW within the caldera. The average velocity of the point of max-
187 imum subsidence is ~ 10 cm year⁻¹, calculated from 30 October 2015 to 31 October 2017,
188 after which subsidence ends. We note that the subsidence stops two months after a seis-
189 micity increase on 25 August 2017 (accompanied by a change in eruptive activity), ac-
190 cording to a Vanuatu Meteorology and Geo-Hazards Department report (Global Volcan-
191 ism Program, 2017a) (green vertical line, Figure 1b). Seismicity declined and eruptive
192 activity stabilized by the end of November 2017 (Global Volcanism Program, 2017b). Us-
193 ing the *Defvolc* 3D Mixed Boundary Element Method (BEM) code (Cayol & Cornet, 1997;
194 Fukushima et al., 2005) and Neighborhood Algorithm (Sambridge, 1999b, 1999a) (See
195 Shreve et al. (2021) for a summary of the BEM and Text S2 for a description of the in-
196 version scheme and uncertainty estimation), we perform a joint inversion of the Sentinel-
197 1 velocity map and ALOS-2 ascending interferogram (scaled to obtain displacement in
198 cm year⁻¹). See Text S3 for a detailed description of data preparation for the geodetic
199 modeling, which includes downsampling, estimating the data covariance matrix, and cre-
200 ating the topography mesh.

201 **2.4 Model Coupling Gas Emissions and Reservoir Depressurization**

202 Numerous theoretical models exist to explain pressure change due to degassing pro-
203 cesses (i.e., volatile exsolution and escape from a magmatic system) (Tait et al., 1989;
204 Boichu et al., 2008, 2011; Anderson & Segall, 2011). In particular, a theoretical model
205 developed by Girona et al. (2014) addresses the relationship between passive degassing
206 and inter-eruptive reservoir depressurization on timescales of years. Girona et al. (2014)
207 assume the reservoir is connected to the surface by an open conduit, and that the pres-

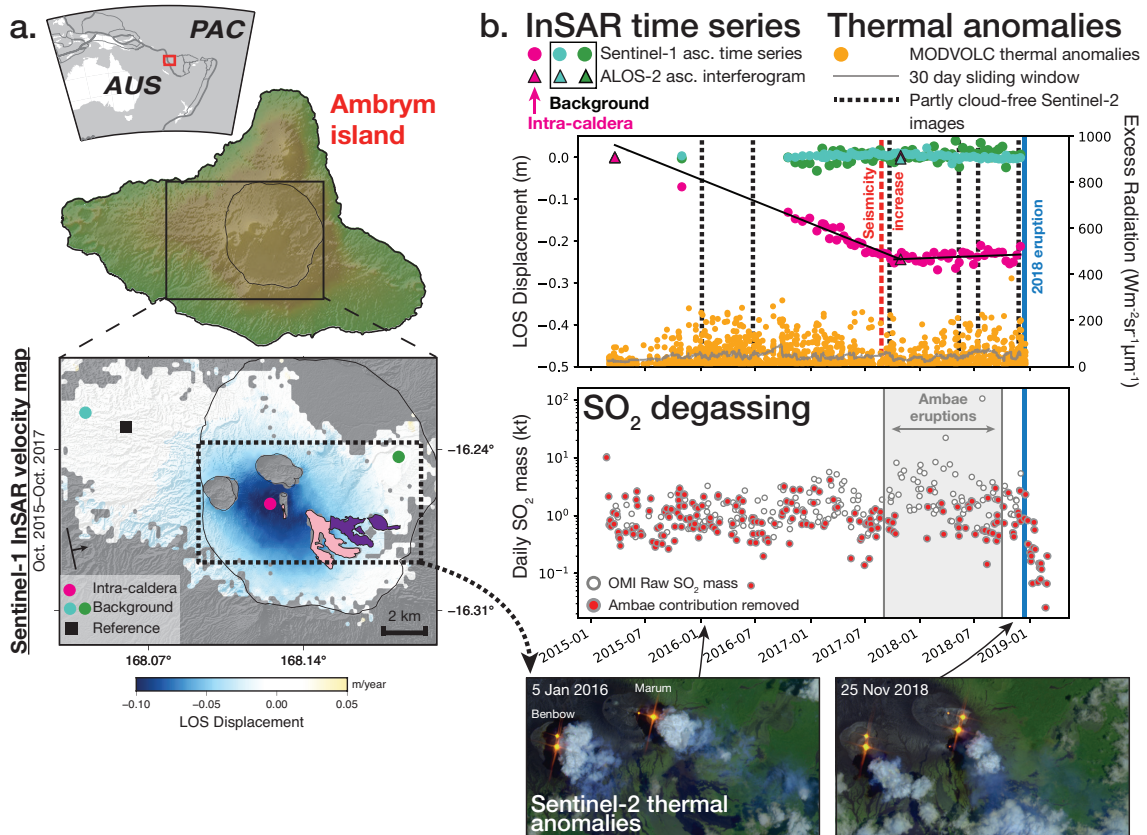


Figure 1: Time series of InSAR LOS displacement, thermal anomalies, and SO₂ emissions. **a.** A tectonic map showing the location of Ambrym island (from Shreve et al. (2019)) and a zoom on Ambrym's caldera, showing the velocity map of the Sentinel-1 ascending time series, assuming deformation spans 30 October 2015–31 October 2017. The caldera is outlined in black, and the extent of the lava flows emplaced in 2015 and 2018 are shown in pink and purple, respectively. The points plotted in the LOS displacement time series are identified with the colored circles (magenta measures displacement inside the deforming caldera, while green and blue measure background displacement near the borders or outside the caldera). The reference point is the black square. **b.** Top figure: The colored dots correspond to the cumulative LOS displacement calculated from a Sentinel-1 ascending time series in three different locations. The triangles outlined in black indicate the cumulative LOS displacement of an ALOS-2 interferogram.

Figure 1: (cont.) The three triangles plotted on 21 March 2015 overlap at 0 m. At each location, the mean of a 3×3 pixel box ($\sim 0.6 \times 0.6$ km) is calculated. The black dotted lines indicate the (at least partly) cloud-free Sentinel-2 images available during this time period (see Video S1). The orange dots are the excess radiation of thermal anomalies detected within Benbow and Marum craters by the Moderate Resolution Imaging Spectroradiometer (MODIS), which has a pixel size of $\sim 1 \times 1$ km at nadir (Wright et al., 2004; Wright, 2016), and processed using the MODVOLC algorithm. The grey line shows temporal averages calculated using a sliding window of 30 days. Bottom figure: Empty grey dots represent the time-series of raw SO_2 mass burden integrated in an 150 km radius disk around Ambrym from OMI data (TRL product). Red-filled dots represent selected SO_2 mass burden estimates, after removing measurements affected by emissions from the nearby Ambae volcano. The period of elevated volcanic activity at Ambae is indicated by the grey area. Figure adapted from Shreve (2020).

Model	Center (UTM)	Depth (m b.s.l.)	Axis 1 (m)	Ratio Axis 1:Axis 2	Ratio Axis 2:Axis 3	Rotation ($^{\circ}$)	V_r (km 3)	ΔV ($\times 10^6$ m 3 year $^{-1}$)	ΔP (MPa year $^{-1}$)
Best-fit	192210 E, 8199170 N	-1880	4700	10	3.6	147	1.2	-3	-2.15
Mean	192367 E \pm 1571, 8199425 N \pm 1228	-2075 \pm 1018	4597 \pm 1159	7.9 \pm 1.5	5.1 \pm 1.9	142 \pm 25	N/A	N/A	-3.56 \pm 1.63

Table 1: The estimated best-fit ellipsoid model parameters, as well as the mean model and uncertainties (one standard deviation from the mean). Obtained from jointly inverting the Sentinel-1 time series and ALOS-2 interferogram with *Defvolc*. The best-fit ellipsoid is visualized in Figure 2.

Symbol	Description	Range used in this study	Reference
$\frac{d\Delta P}{dt}$	pressure change rate during quiescence	-5.2 to -2.0 MPa year ⁻¹	This study
\tilde{Q}	total mean gas flux	110 kt d ⁻¹	Carn et al. (2017), Allard et al. (2015)
t	time of passive degassing	~4 years	This study
R_c	volcanic conduit radius	≤ 300 m	Allard et al. (2016)
V_r	volume of the reservoir	$5 \times 10^8 - 3 \times 10^{10}$ m ³	Shreve et al. (2021)
L	length of the magma column (up to the average reservoir depth)	~2 km	This study
α	mass fraction of dissolved volatiles in parent melt	1.3 wt%	Allard et al. (2015)
k_r	bulk modulus accounting for chamber compressibility (Text S4)	$5 \times 10^8 - 1 \times 10^{10}$ Pa	Heap et al. (2019)
g	gravity	9.8 m s ⁻²	
$\hat{\rho}_{m,c}$	mean density of melt inside the magma column	2550 kg m ⁻³	Girona et al. (2014)
$\hat{\rho}_{g,c}$	mean density of gas inside the magma column	≤ 200 kg m ⁻³	Girona et al. (2014)
ρ_c	density of fully degassed melt in the column	2670 kg m ⁻³	Girona et al. (2014)
ρ_{nd}	density of the parent undegassed magma	2430 kg m ⁻³	Girona et al. (2014)
γ_c	volume fraction of degassed melt in the magma column	0.5	Girona et al. (2014)

Table 2: A description of the symbols used in Cases 1–3, as well as the ranges chosen for the parameters in this study. Symbols with no range presented are calculated from the defined parameters.

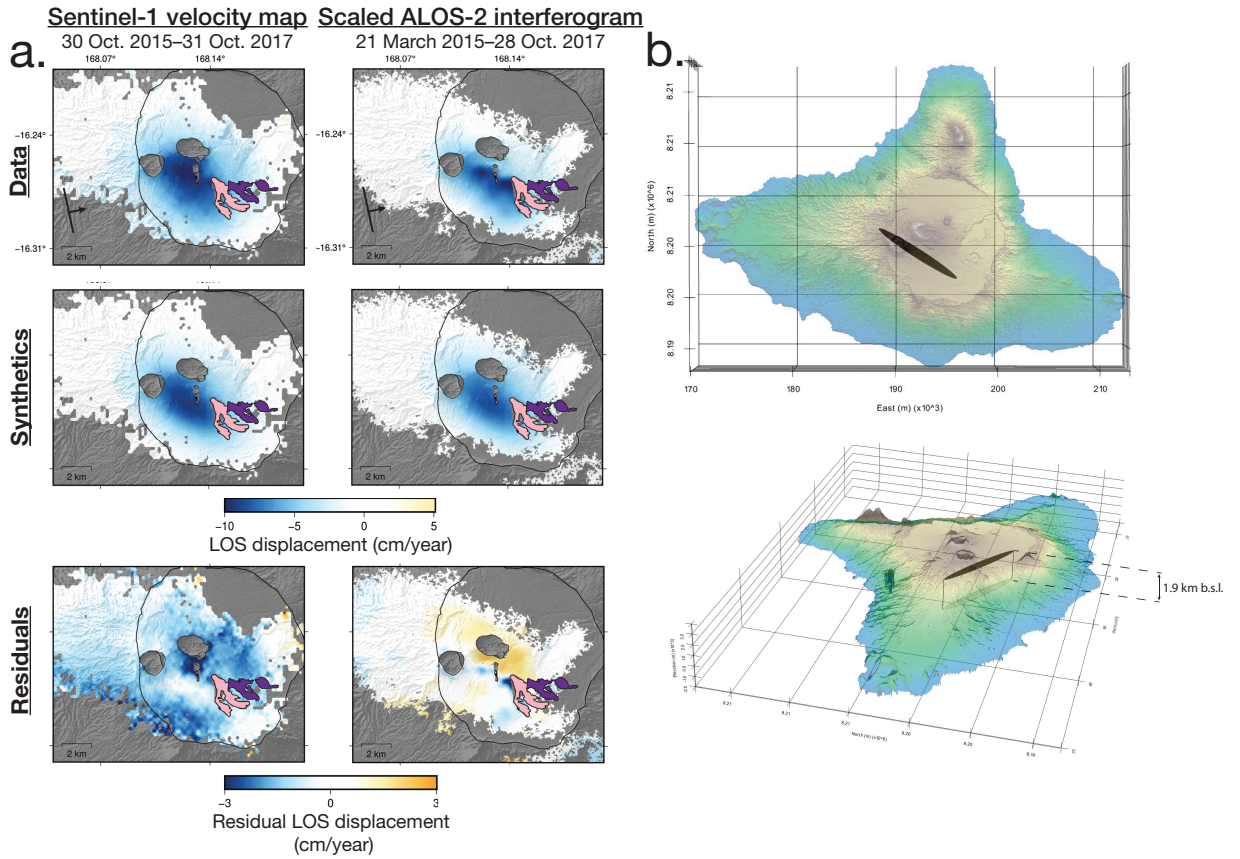


Figure 2: **Inversion results.** **a.** Geodetic observations (top row), synthetic displacements (middle row) and residuals (bottom row) of an ellipsoid at ~ 1.9 km b.s.l (2.8 km beneath the vents). Left column is the Sentinel-1 velocity map, and right column is the ALOS-2 interferogram scaled to obtain displacement in centimeters year $^{-1}$. The caldera and craters are outlined in black, and the 2015 and 2018 lava flows are shown in pink and purple, respectively. **b.** A map and aerial view of the geometry of the depressurized ellipsoid estimated using *Defvolc*. Plot created with PyVista (Sullivan & Kaszynski, 2019).

208 sure in the reservoir equilibrates the weight of the magma in the conduit. Consequently,
 209 mass changes within the system (either in the conduit or reservoir) due to degassing re-
 210 sult in reservoir depressurization, causing ground subsidence. This model takes into ac-
 211 count whether volatiles are degassed from the conduit or reservoir, as well as the viscoelas-
 212 tic behavior of the host rock, magma replenishment, and magma density changes. We
 213 proceed with this theoretical model to couple passive degassing and reservoir depressur-
 214 ization at Ambrym volcano.

215 This model simplifies the relationship between mass loss by degassing and reser-
 216 voir depressurization by assuming the following:

- 217 1. A magma reservoir is connected to an open, magma-filled cylindrical conduit with
 218 a fixed length L . The reservoir's magma pressure is magmastatic (Figure 3a).
- 219 2. Gas mass in the conduit is much smaller than the (incompressible) melt (liquid
 220 and solid phase) mass in the conduit (i.e., $m_{m,c}(t) \gg m_{g,c}(t)$).
- 221 3. The melt in the conduit is a mixture of an undegassed parent melt and a denser,
 222 degassed melt.
- 223 4. Gas separation may occur in the conduit or reservoir. Without magma recharge,
 224 degassing can continue until the total degassed volatile mass equals that of the
 225 initial mass of bulk dissolved volatiles in the undegassed magma.
- 226 5. We assume a mean degassing rate \hat{Q} that is constant over months to years. In Sec-
 227 tion 2.2, we demonstrate that the mean SO_2 flux is constant, while temporal vari-
 228 ations in the $\frac{\text{H}_2\text{O}}{\text{SO}_2}$ ratio are unknown and henceforth assumed constant.

- 229 6. Crystal content is neglected, temperature (T) of the system is constant, gas sol-
 230 ubility depends only on pressure, and the gas mass within the reservoir is calcu-
 231 lated at a mean depth.
- 232 7. Melt volume decreases during degassing of H_2O . No other volatiles are considered.
- 233 8. The host rock of the system is a half-space with a Maxwell viscoelastic rheology
 234 (with a bulk modulus k and an effective viscosity μ).

235 Using these assumptions, this model can be used to gain a first-order understand-
 236 ing of the relationship between a system's degassing flux and depressurization. We in-
 237 vestigate whether this model can explain the reservoir depressurization estimated from
 238 geodetic models at Ambrym. We then discuss the role of magma replenishment and its
 239 effect on the reservoir pressurization rate.

240 Based on the first assumption above, the pressure change at the base of the con-
 241 duit at time t can be written as

$$\Delta P(t) = P(t) - P(t_0) = \frac{g(m_{m,c}(t) + m_{g,c}(t))}{\pi R_c(t)^2} - P(t_0), \quad (1)$$

242 where $P(t)$ and $P(t_0)$ are the pressure within the reservoir at time t and t_0 , $m_{m,c}(t)$
 243 is the melt mass within the conduit of constant length L (Figure 3a), $m_{g,c}(t)$ is the gas
 244 mass within the conduit, $R_c(t)$ is the conduit radius, and g is gravity. Equation 1 de-
 245 scribes a pressurized reservoir which is sustaining the weight of the magma-filled cylin-
 246 drical conduit, which acts on a surface area of $\pi R_c(t)^2$ (Girona et al., 2014).

247 Taking the time derivative of this equation, we can obtain the rate of pressure change:

$$\frac{d\Delta P(t)}{dt} = \frac{g\hat{\rho}_{m,c}(t)V_{m,c}(t)}{\pi R_c(t)^2} \left\{ \frac{1}{\hat{\rho}_{m,c}(t)V_{m,c}(t)} \left[\hat{\rho}_{g,c}(t) \frac{dV_{g,c}(t)}{dt} + V_{g,c}(t) \frac{d\hat{\rho}_{g,c}(t)}{dt} \right. \right. \\ \left. \left. + \hat{\rho}_{m,c}(t) \frac{dV_{m,c}(t)}{dt} + V_{m,c}(t) \frac{d\hat{\rho}_{m,c}(t)}{dt} \right] - \frac{2}{R_c(t)} \frac{dR_c(t)}{dt} \right\}. \quad (2)$$

248 We have substituted $m_{m,c}(t) = \hat{\rho}_{m,c}(t)V_{m,c}(t)$ and $m_{g,c}(t) = \hat{\rho}_{g,c}(t)V_{g,c}(t)$, where
 249 $\hat{\rho}_{m,c}(t)$, $V_{m,c}(t)$, $\hat{\rho}_{g,c}(t)$, $V_{g,c}(t)$ are the mean melt and gas density and volume in the con-
 250 duit, respectively.

251 We further simplify Equation 2 by assuming an elastic host rock, a constant gas
 252 and melt density in the conduit ($\frac{d\hat{\rho}_{g,c}(t)}{dt} = 0$ and $\frac{d\hat{\rho}_{m,c}(t)}{dt} = 0$), a conduit radius that
 253 does not change significantly over time ($\frac{dR_c(t)}{dt} = 0$), and that the mean gas density
 254 in the conduit is significantly smaller than the mean melt density ($\hat{\rho}_{g,c}(t) \ll \hat{\rho}_{m,c}(t)$).

255 We then obtain:

$$\frac{d\Delta P(t)}{dt} = \frac{g\hat{\rho}_{m,c}(t)}{\pi R_c(t)^2} \frac{dV_{m,c}(t)}{dt}. \quad (3)$$

256 After considering the relationship between the volume of magma in the reservoir
 257 and the partial density of water in a silicate melt, ρ_w , we can rewrite the above equa-
 258 tion as:

$$\frac{d\Delta P(t)}{dt} \left(\frac{\pi R_c(t)^2}{g\hat{\rho}_{m,c}(t)} + \frac{V_r(t)}{k_r} \right) = -\frac{\hat{Q}}{\rho_w}, \quad (4)$$

259 which yields, by integration, assuming \hat{Q} is constant, Equation 45 from Girona et
 260 al. (2014):

$$\Delta P(t) = -\frac{g\hat{\rho}_{m,c}(t_0)k_r\hat{Q}t}{\pi R_c(t_0)^2\rho_w k + g\rho_w\hat{\rho}_{m,c}(t_0)V_r(t_0)}. \quad (5)$$

261 The full derivation can be found in Text S4.

262 Equation 5 can be used to compare the depressurization measured with geodetic
 263 modeling (average $\Delta P \propto$ surface deformation) at Ambrym to the theoretical depres-
 264 surization of the magmatic system by passive degassing (using \hat{Q} as input, using OMI
 265 to measure SO_2 mass, then converting to H_2O mass). The remaining parameters nec-
 266 essary to calculate $\Delta P(t)$ are fixed based on previous studies (Table 2—without magma
 267 replenishment, Cases 1–3— and Table S1—with magma replenishment, Case 4).

268 3 Results

269 3.1 Geodetic Source Estimation

270 We invert for the location, depth, axes lengths, and pressure change per year of a
 271 sub-horizontal ellipsoid beneath Ambrym’s caldera (dip fixed to 0°). See Table 1 for the
 272 inverted parameters and final geometries. The final estimated source geometry is an el-
 273 lipsoid elongated $\text{N}^\circ 147$ at a depth of ~ 2.8 km beneath the craters (1.9 km b.s.l., Fig-
 274 ure 2). The northwest tip of the source is located beneath Benbow, and the source ex-
 275 tends beneath the caldera for more than 9 km to the SE, oblique to the $\text{N}^\circ 110$ rift zone.

276 The mean pressure change per year of the ellipsoid is $-3.56 \text{ MPa year}^{-1}$, ranging be-
 277 tween -5.2 to $-2.0 \text{ MPa year}^{-1}$ when including an error of one standard deviation from
 278 the mean (Table 1), given a Young’s modulus of 5 GPa and Poisson’s ratio of 0.25. The

279 best-fit model has a pressure change of $-2.15 \text{ MPa year}^{-1}$ (Table 1). The Young's mod-
280 ulus is estimated from laboratory measurements, according to Heap et al. (2019), yet this
281 parameter is still unconstrained at the scale of the volcanic edifice. These errors are prop-
282 agated into both the geodetic and theoretical modeling. The Young's modulus may re-
283 alistically range from 0.5 to 10 GPa. This will affect the estimated pressure change, due
284 to the relationship $\Delta V_r \propto \frac{V_r \Delta P}{G}$ (Anderson & Segall, 2011), where V_r is the volume of
285 the reservoir, ΔV_r and ΔP are its volume and pressure change, respectively, and G is
286 the host rock shear modulus. It follows that the pressure change of the best-fit model
287 may range from -4.2 to $-0.2 \text{ MPa year}^{-1}$. A pressure change of -4.2 MPa is within the range
288 of uncertainties from the theoretical model discussed in Section 3.2, while -0.2 MPa could
289 only be obtained with unreasonably large ($V_r > 300 \text{ km}^3$) reservoir volumes (Section
290 3.2 and Figure S8). Such uncertainties emphasize the sensitivity of this model to the vol-
291 canic conduit and reservoir geometries, which are often unknown. Future studies using
292 seismic wave velocities to calculate the dynamic shear modulus, then converted to the
293 static shear modulus (e.g., Grandin et al., 2010), or using experimental studies of rock
294 samples from the edifice (e.g., Heap et al., 2019), could further constrain the host rock
295 properties.

296 Because the subsidence rate is approximately constant, we assume that the depres-
297 surization rate bounds of -5.2 to $-2.0 \text{ MPa year}^{-1}$ and the source geometry also remain
298 constant. By combining the depressurization rate with the mean degassing SO_2 flux es-
299 timated in Section 2.2, we investigate whether a model coupling mass loss by passive de-
300 gassing and reservoir depressurization reasonably explains these estimates. We acknowl-

301 edge that other physical mechanisms may drive subsidence on time scales of months to
 302 years, including viscoelastic response of the reservoir tapped during the 2015 eruption
 303 (depending on the viscoelastic shell radius and whether the reservoir is refilled) (Segall,
 304 2019), as well as cooling and crystallization of magma (Caricchi et al., 2014; Townsend,
 305 2022). In addition, deformation rates can change abruptly due to energy released from
 306 seismic swarms, although typically this occurs during uplift episodes. We conclude that
 307 these mechanisms are unlikely because of the relatively short time period of subsidence
 308 (2.5 years), the high deformation rate (~ 1 cm month⁻¹), and the abrupt end to the sub-
 309 sidence.

310 **3.2 Reservoir Depressurization Rate With No Magma Replenishment**

311 We explore various cases to investigate the effect of the fixed model parameters on
 312 reservoir depressurization rate. This includes fixing whether the host-rock is either elas-
 313 tic or viscoelastic, or if gas exsolution occurs in the conduit or reservoir. Finally, we dis-
 314 cuss the effect of magma replenishment on the reservoir depressurization rate.

315 **3.2.1 Elastic Rheology (Case 1)**

316 We calculate the theoretical reservoir depressurization rate at Ambrym in an elas-
 317 tic halfspace, for a range of reservoir volumes $V_r(t_0)$ and conduit radii $R_c(t_0)$, given a
 318 water vapor flux of $\hat{Q} \approx 110$ kt day⁻¹. The higher the gas flux \hat{Q} , the higher the depres-
 319 surization rate. We also explore a range of reasonable values for the free parameters. For
 320 example, the reservoir bulk modulus k_r may realistically vary from 5×10^8 to 1×10^{10}
 321 MPa. As k_r increases, the host rock becomes more rigid. On the contrary, if the reser-

322 voir is compliant, reservoir deformation helps sustain the reservoir’s internal pressure.
 323 Finally, we fix the initial bulk volatile content, α (used to calculate ρ_w ; see Text S4), to
 324 1.3 wt% (Allard et al., 2015). For a given \hat{Q} , the lower the value of α , the higher the de-
 325 pressurization rate. This is because for a lower α , a greater volume of magma will need
 326 to lose gas to sustain the same \hat{Q} . However, for a lower α , the melt density change af-
 327 ter degassing would be smaller, which could partly compensate, or even counterbalance,
 328 the higher depressurization rate. The density of the parent melt is assumed to be $\rho_{nd} =$
 329 2430 kg m^{-3} and the density of the degassed melt is $\rho_c = 2670 \text{ kg m}^{-3}$ (values imple-
 330 mented in Girona et al. (2014)). The influence of ρ_{nd} and ρ_c on depressurization rate
 331 is shown in Figure S6. Following Girona et al. (2014), we assume steady-state convec-
 332 tion in a conduit that consists of 50% parent melt (undegassed) and 50% degassed melt,
 333 and the constant mean melt density is $\hat{\rho}_{m,c}(t) = \hat{\rho}_{m,c}(t_0) = 2550 \text{ kg m}^{-3}$.

334 Figure 3b shows the theoretical depressurization rates for Case 1 when $k_r = 5 \times$
 335 10^8 Pa , $\hat{Q} = 110 \text{ kt day}^{-1}$, $\alpha = 1.3 \text{ wt}\%$. According to Shreve et al. (2021), Ambrym’s
 336 reservoir has a minimum size of 1 km^3 . For completeness, we assume a range of reser-
 337 voir volumes from $1\text{--}35 \text{ km}^3$ and conduit radii from $10\text{--}700 \text{ m}$. While Allard et al. (2016)
 338 inferred a 3 m radius for the upper section of Benbow’s conduit to sustain the magma
 339 convection needed to explain the SO_2 flux, there have been no direct measurements of
 340 the conduit diameter. We therefore investigate a wide range of conduit radii to explore
 341 the effect of this parameter on the model outputs. Our models indicate that the depres-
 342 surization rates at Ambrym inferred from geodesy (between -5.2 and -2 MPa) can be reached
 343 for large conduit radii ($R_c > 400 \text{ m}$) when reservoir volumes are small ($V_r \approx 1 \text{ km}^3$).

344 For larger reservoirs ($V_r > 12 \text{ km}^3$), the conduit radius connecting the lava lakes to the
 345 reservoir may be as low as 10 m.

346 *3.2.2 Viscoelastic Rheology, Gas Exsolves in Conduit (Case 2)*

347 Case 1 assumes an elastic half-space, and we wish to investigate how viscoelastic-
 348 ity affects the response of the host rock to mass loss by degassing. We still assume that
 349 there is no exsolved gas in the reservoir and introduce a host rock with a Maxwell vis-
 350 coelastic rheology. Gas is exsolved in and degassed from the conduit of length L , which
 351 has an initial gas volume fraction $\beta_c = 0.1$. As in Case 1, we impose steady-state magma
 352 convection in the conduit.

353 Despite the addition of viscoelasticity, for the range of parameters of interest ($V_r =$
 354 $1\text{--}35 \text{ km}^3$ and $R_c = 10\text{--}700 \text{ m}$), no meaningful difference is noted between Case 1 and
 355 Case 2 (Figure S7). For larger values of V_r , the solutions begin to diverge when the host
 356 rock becomes less viscous ($\mu \approx 10^{17} \text{ Pa s}$, Figure S7c). Aside from the minimum vis-
 357 cosity necessary to maintain topography, we have no constraints on this parameter for
 358 the host rock at Ambrym. Therefore, we proceed with Case 3 by setting the effective vis-
 359 cosity of the crust to $\mu = 10^{18} \text{ Pa s}$ (Table S1) and allowing for gas exsolution in the
 360 reservoir.

361 *3.2.3 Viscoelastic Rheology, Gas Exsolves in Reservoir (Case 3)*

362 We assume that gases may exsolve within the reservoir according to Henry's Law
 363 ($n_r(t) = \alpha - S[P(t_0) - \Delta P(t)]^{\frac{1}{2}}$ when $\alpha > S[P(t_0) + \Delta P(t)]^{\frac{1}{2}}$) (e.g., Huppert & Woods,

2002). Assuming the conduit is completely degassed ($\hat{\rho}_{m,c} = 2670 \text{ kg m}^{-3}$), and given that $\alpha = 1.3 \text{ wt}\%$, $S = 4 \times 10^{-6} \text{ Pa}^{-1/2}$ for water, and that $\Delta P(t) \ll P(t_0) = \rho_{m,c} g L \approx 52 \text{ MPa}$, the depth of the reservoir when gas exsolution occurs is $L < \frac{1}{g \rho_{m,c}} \left(\frac{\alpha}{S}\right)^2 \approx 400 \text{ m}$.

If we assume a slightly larger α (e.g., 2.1 wt% as estimated by Moussallam et al. (2021)), gas exsolution may occur at deeper levels, $\sim 1 \text{ km}$. This is still too shallow to exsolve bubbles in a reservoir residing at 2–4 km depth, as estimated for the shallowest levels of Ambrym’s magmatic system (Shreve et al., 2019, 2021). Geodetic inversions estimate a reservoir depth of $\sim 2 \text{ km}$ beneath Ambrym’s lava lakes. We can therefore conclude, according to Henry’s Law, that there is no exsolved H_2O in the reservoir. In this case, Case 3 will have the same result as Cases 1 and 2, where gas exsolution occurs within the conduit.

3.3 Reservoir Depressurization Rate With Magma Replenishment (Case 4)

In Section 3.2, depressurization rates estimated at Ambrym from geodetic observations can be obtained using the theoretical model for reservoir volumes $> 12 \text{ km}^3$. Using an end member scenario that is consistent with the best-fit mean depressurization rate of $-2.15 \text{ MPa year}^{-1}$, we investigate the change in depressurization rate that occurred in October 2017. Between October 2017 (t_0) and December 2018 (t_1), we assume no net depressurization of the magmatic system ($\frac{d\Delta P(t_1 - t_0)}{dt} = 0$) because the time series displacements are within the data uncertainty ($\sim 1 \text{ cm}$).

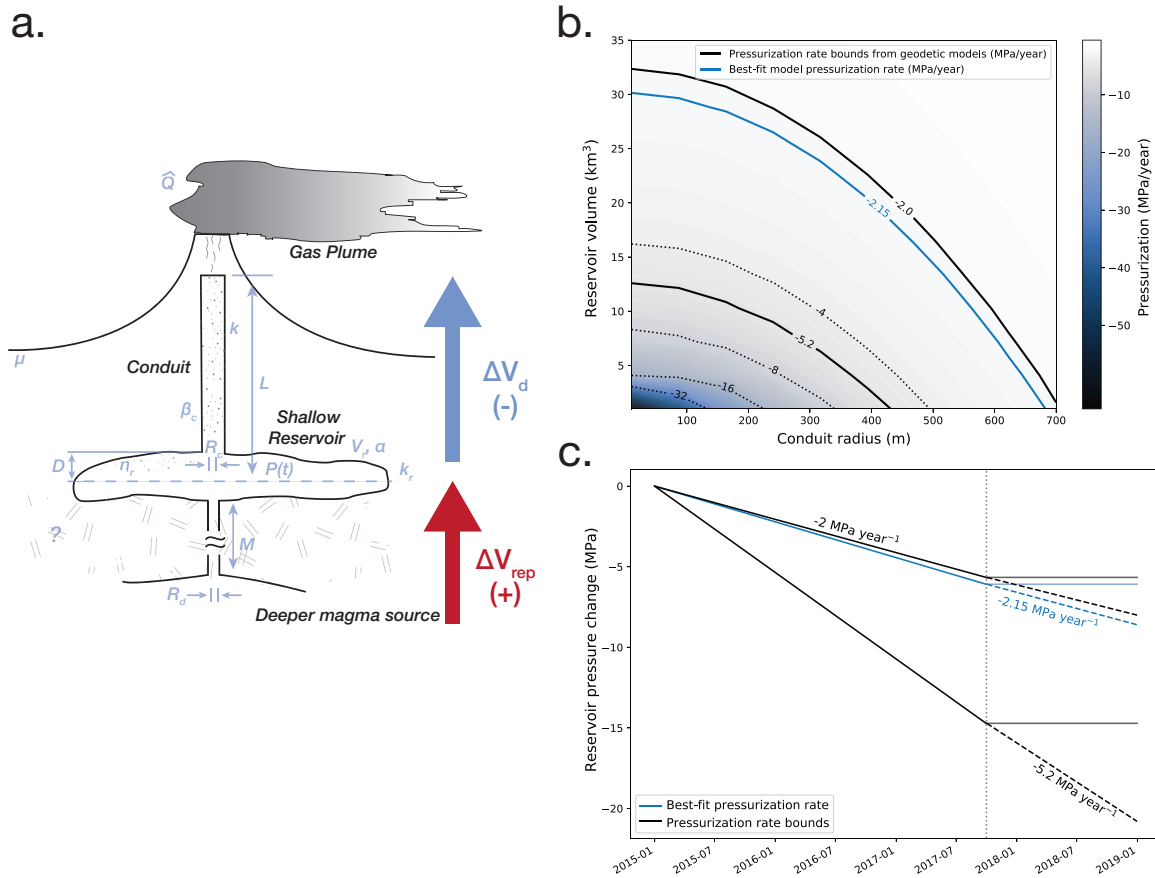


Figure 3: **Degassing-induced depressurization theoretical model.** **a.** A schematic showing the parameters included in the theoretical model. Modified after Girona et al. (2014). **b.** The pressurization rate change in MPa year^{-1} for a variety of conduit radii and reservoir volumes. The heavy black lines show the upper and lower bounds of the pressure change estimated in the geodetic inversion, and the heavy blue line shows the pressure change of the best-fit model. The dotted contour lines show pressure changes of -32, -16, -8, and -4 MPa year^{-1} . **c.** The reservoir pressure change evolution over time (black and blue corresponding to the upper and lower bounds and best-fit pressure change, respectively), as estimated from geodetic inversions. The vertical dotted line indicates the end of the subsidence episode. The dotted lines show the total reservoir pressure change if no change in the depressurization rate occurred in October 2017.

385 One possible mechanism that could change the depressurization rate is a decrease
 386 in passive degassing. There is, however, no meaningful change in the gas flux measured
 387 by satellite during this time period (Figure 1b and Figure S2). Therefore, we hypoth-
 388 esize that the depressurization rate change is due to magma replenishment from a deeper
 389 reservoir into the shallow reservoir, counterbalancing the degassing-induced depressur-
 390 ization and resulting in no net pressurization of the shallow reservoir. Building on the
 391 previous findings of Cases 1–3, we attempt to constrain the magma replenishment rate.
 392 We address two possible scenarios:

- 393 1. The deeper source is not overpressurized ($\Delta P_s(t) = 0$). Shallow reservoir depres-
 394 surization results in a pressure gradient between the shallow and deeper sources.
 395 Passive magma replenishment begins as soon as the depressurization began, at the
 396 latest in June 2015.
- 397 2. Alternatively, the deeper source is overpressurized (i.e., active replenishment, $\Delta P_s(t) >$
 398 0). The deep and shallow sources become (instantaneously) connected in Octo-
 399 ber 2017, and the deep reservoir overpressure remains constant.

Both scenarios will consider a viscoelastic behavior of the host rock. Therefore, we
 can no longer use the simplified Equation 5 to calculate the reservoir pressure change.

Without the simplifications of Equation 2 outlined in Section 2.4, $\frac{d\Delta P(t)}{dt}$ is instead ex-
 pressed as

$$\frac{d\Delta P(t)}{dt} = \frac{C_1(t) + C_2(t)\Delta P(t)}{C_3(t)}, \quad (6)$$

400 where the constants $C_1(t)$, $C_2(t)$, and $C_3(t)$ are functions of the parameters in Equation
 401 2 (See also Equation 34 in Girona et al. (2014), which includes the rate of magma re-

402 plenishment). The full expansion of these constants and a table of parameters used in
 403 this study can be found in Table 2, Text S5, and Table S1.

Following Girona et al. (2014), the pressure change in the shallow reservoir can be expressed with an equation of the form

$$\Delta P(t) = -\Delta P_\infty(1 - e^{-\Gamma t}), \quad (7)$$

404 where $\Delta P_\infty = \frac{C_1(t_0)}{C_2(t_0)}$ and $\Gamma = -\frac{C_2(t_0)}{C_3(t_0)}$, assuming $C_1(t)$, $C_2(t)$ and $C_3(t)$ do not vary
 405 significantly with time from their initial values.

The volumetric rate of magma replenishment is

$$\frac{dV_{\text{rep}}}{dt} = \lambda(t)(\Delta P_s(t) - \Delta P(t)), \quad (8)$$

406 where $\Delta P(t)$ is the pressure change in the shallow reservoir, $\Delta P_s(t)$ is the deep source
 407 overpressure, and $\lambda(t)$ is, as defined by Girona et al. (2014), the hydraulic strength (or
 408 strength of connectivity between the shallow and deep reservoirs), as derived from the
 409 Hagen-Poiseuille law for the volumetric flow rate, given laminar fluid flow in a conduit
 410 (Anderson & Segall, 2011; Le Mével et al., 2016).

$$\lambda(t) = \frac{\pi R_d(t)^4}{8M\mu_{nd}}, \quad (9)$$

411 where $R_d(t)$ is the radius of the dike connecting the deeper magma source to the shal-
 412 low reservoir, M is the length of the dike, and μ_{nd} is the viscosity of the replenished, un-
 413 degassed magma. The maximum Reynold's number $Re = \frac{\rho_m M V_{\text{rep}}}{\pi R_d^2 \mu_m}$, given $\rho_m < 2700$
 414 kg m^{-3} , $V_{\text{rep}} < 5 \text{ m}^3\text{s}^{-1}$, $M < 15 \text{ km}$, $\mu_m > 10^4 \text{ Pa s}$ and $R_d > 2 \text{ m}$, is ~ 1600 . This is

415 smaller than the $Re_{\text{crit}} \sim 2000$ that defines the bounds between laminar and turbulent flow
 416 (Turcotte & Schubert, 2014).

417 The first scenario addresses passive replenishment, or when the deeper source is
 418 not overpressured ($\Delta P_s(t) = 0$). Regardless of the input parameters, the depressuriza-
 419 tion rate decays exponentially (Equation 7). This decrease is not consistent with the abrupt
 420 pressure rate change observed in October 2017 (Figure S9). Therefore, we proceed with
 421 the model of active replenishment, considering instead an overpressured deep source ($\Delta P_s(t) >$
 422 0).

423 As mentioned previously, we assume that a hydraulic connection is created between
 424 the deep and shallow reservoirs in October 2017 (t_0), and the connection remains steady
 425 until the eruption in December 2018 (t_1). This hydraulic connection may occur when
 426 the pressure difference between the shallow and deep reservoirs reaches a critical value,
 427 after which any crystallized magma (i.e., a stiff plug) in the connecting dike begins to
 428 flow upwards (Girona et al., 2015). As hotter material from the deep reservoir flows into
 429 the dike, the dike unplugs and the two reservoirs become hydraulically connected. This
 430 unplugging also occurs within two months of the seismicity increase noted by the VMGD
 431 (Figure 1b) (Global Volcanism Program, 2017a). As volume is lost during gas emission,
 432 the resulting pressure decrease in the shallow reservoir is compensated by a renewed magma
 433 inflow from a deeper reservoir, and a pressure balance is obtained in the shallow reser-
 434 voir ($\Delta P(t_1) - \Delta P(t_0) = 0$). The replenished volume is calculated for each consecu-
 435 tive time step t_{n+1} using Equation 8, given the pressure state of the shallow reservoir
 436 at time t_n , a fixed $\lambda(t_0)$, and a $\Delta P_s(t > t_0) > 0$.

As stated in Equation 8, the volume of replenished magma (Figure 4a), is a function of the deep reservoir overpressure $\Delta P_s(t)$, the dike radius and length, and the magma viscosity. Because we have no constraints on these parameters, we instead explore the tradeoff between the lumped parameter $\lambda(t_0)$ (hydraulic strength) and the deep source overpressure $\Delta P_s(t_0)$ (Figure 4c, assuming a $V_r = 29 \text{ km}^3$ and $R_c = 10 \text{ m}$, resulting in a shallow reservoir pressure change of $\sim 2.15 \text{ MPa year}^{-1}$). A pressure balance occurs for a wide range of $\Delta P_s(t_0)$ and $\lambda(t_0)$. However, the volume of replenished magma does not vary based on the depressurization rate measured, because the melt volume decrease in the conduit-reservoir system when $\Delta P(t_n) = \Delta P(t_{n+1})$ (i.e., no net pressure change) is also constant and defined by $\frac{dV_m(t)}{dt} = \frac{dV_d(t)}{dt} + \frac{dV_{\text{rep}}(t)}{dt} = \frac{-\hat{Q}}{\rho_w} + \lambda(t_0)(\Delta P_s(t_0) - \Delta P(t))$ (Figure 4a). In the case of no net pressure change, Equation 6 can be simplified to

$$\frac{d\Delta P(t)}{dt} = 0 \implies \frac{\hat{Q}}{\rho_w} - \lambda(t_0)(\Delta P_s(t_0) - \Delta P(t)) - \frac{\Delta P(t)}{\mu}(V_r(t_0) + V_{m,c}(t_0)) = 0, \quad (10)$$

437 where volume change is determined by the balance between volume loss due to de-
 438 gassing, volume increase due to magma influx, and the viscous response of the conduit
 439 and reservoir. As shown in Figure 4a, the total replenished volume necessary to create
 440 a pressure balance $\Delta P(t_{n+1}) - \Delta P(t_n) = 0$ within the shallow reservoir is $\sim 0.16 \text{ km}^3$
 441 over 14 months, or a magma supply rate into the shallow reservoir of $\sim 4.3 \text{ m}^3 \text{ s}^{-1}$.

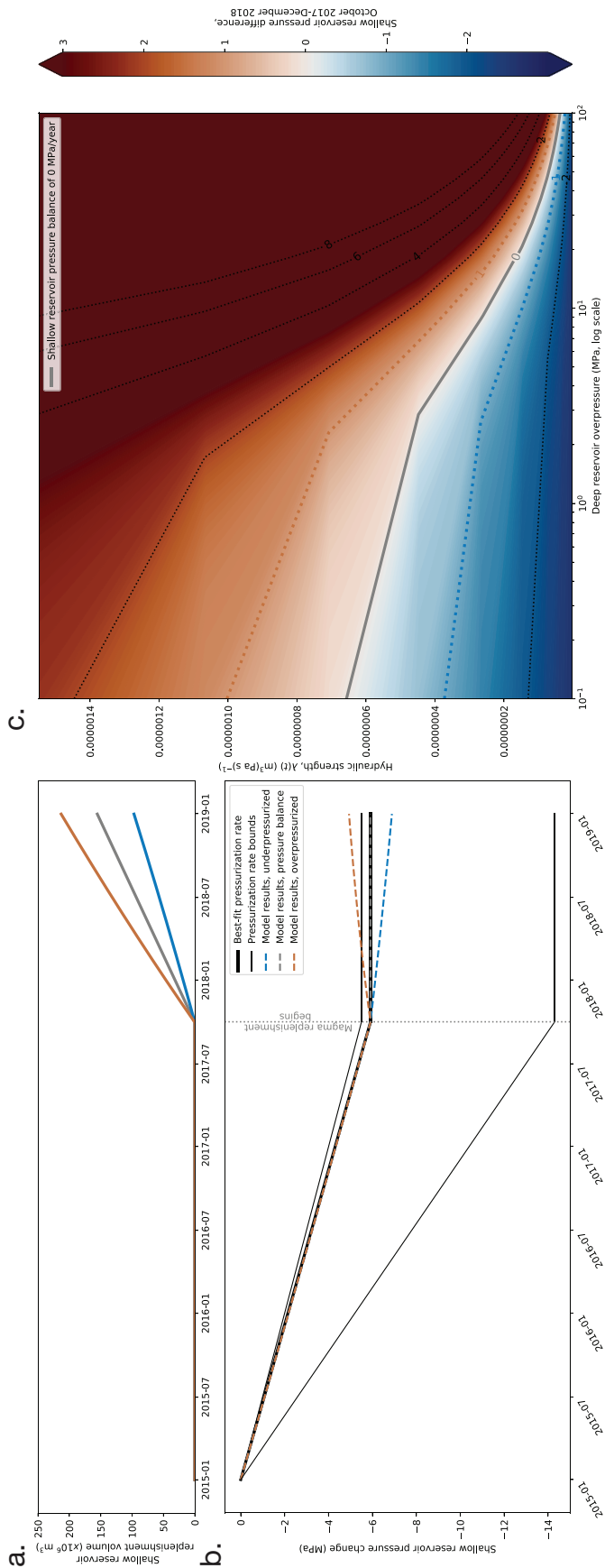


Figure 4: Effect of magma replenishment. a. The replenishment volume over time for three cases: an overpressurized and underpressurized reservoir (pink and blue lines, respectively) at t_1 (December 2018) and a reservoir with a pressure balance (grey line). **b.** The lower and upper bounds and best-fit shallow reservoir pressurization bounds (thin and thick black lines, respectively) over time as estimated from the theoretical model (Figure 3b) for reasonable V_r and R_c that correspond to these values). The vertical dotted line represents when the deeper reservoir becomes instantaneously overpressurized. **c.** The relationship between shallow reservoir pressure difference between t_0 (October 2017) and t_1 (December 2018), hydraulic strength, and deep reservoir overpressure $\Delta P_s(t_0)$. The parameters that lead to a pressure balance in the shallow reservoir ($\Delta P = 0$) are indicated by the black line, and an overpressurized ($\Delta P > 0$) and underpressurized ($\Delta P < 0$) shallow reservoir by the pink and blue dotted lines, respectively.

442 **4 Discussion**

443 **4.1 Volume of Ambrym’s Reservoir**

444 Using a simplified conceptual model of a volcanic plumbing system (a single con-
 445 duit hydraulically connected to a shallow reservoir fed by a deep source), the subsidence
 446 rates at Ambrym estimated with geodetic models can be obtained with a smaller reser-
 447 voir volume ($V_r \approx 1 \text{ km}^3$) and a conduit radius of $>400 \text{ m}$. We consider this unreal-
 448 istic based on studies showing that conduit radii are $<200 \text{ m}$ (e.g., Stephens et al., 2017).
 449 As mentioned previously, the upper section of Benbow’s conduit is estimated to have a
 450 radius of 3 m (Allard et al., 2016). This narrow conduit would be consistent, given our
 451 calculations, with a magma reservoir volume $>12 \text{ km}^3$.

452 Previous estimates of Ambrym’s shallow magmatic reservoir have ranged from 0.5
 453 (Allard et al., 2015) to 9.5 km^3 (Shreve et al., 2021), with no estimates in the range of
 454 $>12 \text{ km}^3$. This may imply one of the following:

- 455 1. The model is oversimplified, and does not represent Ambrym’s plumbing system,
 456 or
 457 2. previous reservoir volume estimates are underestimated.

458 We address each of these possibilities in turn.

459 ***4.1.1 Model Parameters Affecting Depressurization Rates***

460 Given the high gas flux at Ambrym, the geodesy-derived depressurization rates re-
 461 quire a large reservoir because it will need to be subjected to a smaller pressure change

462 to result in the same volume change due to gas emissions (recall $\Delta V_r \propto \frac{V_r \Delta P}{G}$). In this
463 study, the geometry of the reservoir-conduit system is simplified. We assume a single con-
464 duit extending from the surface to 2 km b.s.l., from which gas can exsolve and escape,
465 causing magma convection in the conduit. Volcanic conduits have been imaged to depths
466 of up to a few hundred meters within the edifice using muon tomography (Tanaka et al.,
467 2009), but at deeper depths (>1 km) magma ascent is thought to occur by porous flow
468 through the thermally weakened, yet crystal-rich host rock (McKenzie, 1984). In addi-
469 tion, Ambrym has two active craters with multiple vents that were passively degassing
470 during the time period of interest, and their connection at depth is not understood. Through-
471 out this study, we assume a single conduit connects the reservoir to the surface, which
472 may oversimplify reality.

473 Another factor that affects the pressure change within the reservoir, but is not con-
474 sidered in this model, is magma compressibility. Withdrawal of a compressible magma
475 from the system mutes reservoir pressure changes, because magma volume increases as
476 the magma ascends and decompresses. Therefore, the magma density in the reservoir
477 should be considered to account for this volume change, similar to Case 3 (when gas ex-
478 solves in the reservoir). However, using the pressure-dependent model based on Henry's
479 law, no water vapor exsolves at depths of ~ 2 km. Past studies of Ambrym indicate that
480 magma compressibility in the reservoir may be higher than for gas-poor basalts (Shreve
481 et al., 2021; Hamling & Kilgour, 2020). These results imply that magma compressibil-
482 ity at Ambrym may be high enough to cause larger reservoir volume estimates, but the
483 full implications of magma compressibility are outside this study's scope.

484 In addition, due to the unconstrained depth of the deeper reservoir, we assume that
 485 InSAR surface displacements are only influenced by pressurization of the shallow reser-
 486 voir (Du et al., 1992; Rivalta & Segall, 2008). Petrological constraints from previous work
 487 (e.g., Moussallam et al. (2021)) based on melt inclusion entrapment pressures from the
 488 2018 eruptive products indicate that magma may be stored at depths of up to 9.5 km
 489 b.s.l. Clinopyroxene barometry estimates a magma plumbing system that reaches depths
 490 of 14 km b.s.l. (Firth et al., 2016; Sheehan & Barclay, 2016). However, given that only
 491 one melt inclusion from the 2018 eruption was used to estimate the depth of 9.5 km b.s.l.
 492 (Moussallam et al., 2021), we cannot robustly constrain the deeper reservoir location.
 493 Assuming a spherical point pressure source (Mogi, 1958), an elastic medium, and the es-
 494 timated magma influx ($\sim 0.16 \text{ km}^3$), the reservoir would need a depth $>25 \text{ km}$ for sub-
 495 sidence to be $<5 \text{ cm}$ in the satellite LOS. This reservoir may be shallower than 25 km
 496 if magma compressibility results in muted ground displacements, however, as mentioned
 497 previously, we do not further explore effects of magma compressibility.

498 Finally, as mentioned in Section 2.2, the H_2O flux used to estimate the depressur-
 499 ization rate is based on the assumption of a constant $\frac{\text{H}_2\text{O}}{\text{SO}_2}$ ratio, as well as a SO_2 CMA
 500 at 3 km. If these assumptions are incorrect, the H_2O flux used as input for the theoret-
 501 ical model would be biased. Therefore, we run the model for H_2O fluxes $\pm 50\%$ (55 or
 502 165 kt; Figure S10). The estimated reservoir volume is $>5 \text{ km}^3$ for a conduit radius $R_c \sim$
 503 10 m, and the total replenished volume is $V_{rep} = 0.077 \text{ km}^3$ when $\hat{Q} = 55 \text{ kt}$ and 0.232
 504 km^3 when $\hat{Q} = 165 \text{ kt}$.

505 *4.1.2 Underestimation of Reservoir Volume Estimates*

506 Previous studies have put bounds on Ambrym’s reservoir volume using either magma
507 chamber residence times (Allard et al., 2015), or with geodesy (Shreve et al., 2021). The
508 former estimated a minimum reservoir volume of 0.5 km^3 at depths of $\sim 3.8 \text{ km}$ below
509 the summit according to the magma influx rate needed to explain the mean SO_2 plume
510 flux ($\sim 7800 \text{ tons day}^{-1}$) (Allard et al., 2015; Bani et al., 2012). Allard et al. (2015) ac-
511 knowledge that this reservoir is connected to the lava lakes by a shallower plumbing sys-
512 tem, which we hypothesize is the depressurized reservoir imaged with geodesy in this study.
513 In addition, they conclude that in order to accommodate the recycling of degassed magma,
514 the magma plumbing system must be larger than 0.5 km^3 .

515 By relating geodetic models to the SO_2 flux measured by OMI satellite-based UV
516 spectrometers, Shreve et al. (2021) estimates that the minimum reservoir size is 1 km^3 ,
517 with possible reservoir volumes as high as 9.5 km^3 . However, these may also be under-
518 estimates, because they depend on the assumed host rock shear modulus, which is not
519 well constrained at Ambrym. This study cannot further constrain the upper bound on
520 the size of the degassed reservoir at Ambrym. Nonetheless, the results presented here
521 are consistent with the conclusions that previous studies have underestimated the reser-
522 voir volume, and that it may be as large as 12 km^3 .

4.2 Magma Replenishment Prior to the 2018 Eruption

4.2.1 Magma Supply Rate

According to our study, approximately 0.16 km^3 of magma would have been needed to replenish the shallow reservoir to counteract the subsidence from degassing, equivalent to a magma supply rate into the shallow reservoir $\frac{dV_{rep}}{dt} \approx 4.3 \text{ m}^3 \text{ s}^{-1}$. In a previous study, Allard et al. (2015) estimates that $25 \text{ m}^3 \text{ s}^{-1}$ of magma is necessary to sustain the SO_2 flux released at the surface. However, the estimate from Allard et al. (2015) corresponds to the volume of the magma batch that loses its gas and is recycled deeper in the system per unit time. On the other hand, our estimate corresponds to the volume lost in this recycling process, due to density changes during degassing, hence representing a fraction of the recycled volume of Allard et al. (2015). The ratio between the two estimates is approximately equal to a “shrinkage ratio” $\beta = \frac{\rho_e - \rho_{nd}}{\rho_{nd}} = \frac{\Delta\rho}{\rho_{nd}} \approx 10\%$, defined as the density increase of the magma due to degassing (see Text S6 for a detailed derivation).

The magma supply rate estimated in this study ($4.3 \text{ m}^3 \text{ s}^{-1}$) is on the same order of magnitude as the mantle-driven surge in magma supply rate at Kīlauea in 2006, during the eruption from the Pu’u ’Ō’ō–Kupaianaha vent system which began in 1983 (Poland et al., 2012). The supply rate was estimated to range between $6\text{--}7 \text{ m}^3 \text{ s}^{-1}$, depending on the assumed magma compressibility (Poland et al., 2012). This estimate was made using a calculation similar to Allard et al. (2015), with SO_2 flux used as a proxy for lava effusion rates and consequently magma supply rate (Poland et al. (2012) also includes volume change due to geodetic modeling of uplift that occurred in 2006). However, at

545 Kīlauea, the degassed magma is assumed to have erupted from the Pu'u Ō'ō vent (as
546 opposed to being recycled) and been replenished at the same rate with magma from depth.
547 Therefore, these calculations are more comparable to our estimate than the supply rate
548 estimated by Allard et al. (2015).

549 *4.2.2 Stealth Magma Replenishment*

550 Recent conceptual models describe volcanic unrest and eruption (in particular ini-
551 tial conduit formation) in three phases. The first is “staging”, which can occur either at
552 the end of a previous eruption, during intereruptive time periods, or immediately pre-
553 ceding an eruption (Roman & Cashman, 2018). In the case of replenishment at Ambrym,
554 the latter seems the most likely. Earthquakes detected in mid-2017 were the first indi-
555 cation of magma replenishment to shallower levels (Global Volcanism Program, 2017a),
556 possibly indicating the shear failure of host rocks due to a magma intrusion. Further anal-
557 ysis of seismic data is necessary to discriminate between the processes causing seismic-
558 ity in August 2017.

559 Around the same time, the number of active vents increased, also indicating a magma
560 influx into the system (Video S1). After an increase in lava lake level (Moussallam et al.,
561 2021), the shallow reservoir destabilized, causing the initial intra-caldera eruption of 14
562 December 2018, and eventually tapped the deeper magmatic system the following day
563 (Shreve et al., 2019). The deeper portion of the magmatic system (>6.5 km b.s.l.) fed
564 the rift zone intrusion and submarine eruption, according to melt inclusion entrapment
565 pressures (Moussallam et al., 2021).

566 The pressure balance due to the combined effects of degassing and magma replen-
567 ishment results in a low stress change on the host rocks, effectively equivalent to a “rate
568 or volume that is too low/small to produce detectable seismicity” (Roman & Cashman,
569 2018). This demonstrates the possibility of a replenishment volume and rate that is large
570 compared to other geodetically-derived magma replenishment rates (e.g., $187 \times 10^6 \text{ m}^3$
571 over 7.3 years at Laguna del Maule (Le Mével et al., 2016)) without notable seismicity
572 in the months prior to the eruption (according to VMGD reports).

573 5 Conclusion

574 This study attempts to quantify the influence of persistent degassing on the pres-
575 surization state of a magma reservoir, ultimately allowing for more accurate forecasts
576 of future activity. The theoretical model is a simplistic view of reality (e.g., only one shal-
577 low reservoir and conduit, assumes steady conduit convection, etc.), but the uncertain-
578 ties and tradeoffs presented in this study may be further reduced at a volcanic system
579 with more instrumentation, or a longer time series of passive degassing and ground dis-
580 placement.

581 Despite the theoretical model’s simplifications, we can conclude that:

- 582 1. continuous passive degassing is a plausible mechanism to drive ~ 2.5 years of reser-
583 voir depressurization and subsidence at Ambrym volcano, and
- 584 2. “stealth” replenishment of magma may have occurred in the 14 months prior to
585 the 2018 eruption, resulting in a net pressure balance in the shallow magmatic sys-

586 tem, explaining the abrupt cessation of subsidence in October 2017 and the ab-
587 sence of pre-eruptive uplift prior to the December 2018 eruption.

588 These findings have implications for the monitoring of Ambrym volcano. InSAR
589 geodesy demonstrates a lack of substational uplift (<5 cm, Figure 1b) in the 12 days prior
590 to the 2018 eruption. However, InSAR detected changes in the long-term deformation
591 trend, in particular the abrupt end of a subsidence episode after 2.5 years. We interpret
592 the subsequent lack of subsidence as indirect evidence for magma influx from depth in
593 the 14 months preceding the 2018 eruption, a conclusion that cannot be drawn from de-
594 gassing and thermal anomalies alone. As quantitative in-situ measurements (such as GNSS
595 or gravity measurements and lava lake level heights) are seldom available, multiparamet-
596 ric remote sensing observations (radar, gas and thermal satellite imagery) provide an al-
597 ternative means of detecting magma replenishment over time scales of months to years
598 before eruptions.

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