



**HAL**  
open science

# Controls on generation and accumulation of blended gases (CH<sub>4</sub>/H<sub>2</sub>/He) in the Neoproterozoic Amadeus Basin, Australia

Mahmoud Leila, Keanu Loiseau, Isabelle Moretti

► **To cite this version:**

Mahmoud Leila, Keanu Loiseau, Isabelle Moretti. Controls on generation and accumulation of blended gases (CH<sub>4</sub>/H<sub>2</sub>/He) in the Neoproterozoic Amadeus Basin, Australia. *Marine and Petroleum Geology*, 2022, 140, pp.105643. 10.1016/j.marpetgeo.2022.105643 . hal-03623683

**HAL Id: hal-03623683**

<https://hal.science/hal-03623683v1>

Submitted on 22 Jul 2024

**HAL** is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.



Distributed under a Creative Commons Attribution - NonCommercial 4.0 International License

1 **Controls on generation and accumulation of blended gases (CH<sub>4</sub>/H<sub>2</sub>/He) in the**  
2 **Neoproterozoic Amadeus Basin, Australia.**

3 **Mahmoud Leila<sup>a,b\*</sup>, Keanu Loiseau<sup>a</sup>, Isabelle Moretti<sup>a</sup>**

4 <sup>a</sup>Laboratoire des Fluides Complexes et leurs Réservoirs, Université de Pau et des pays de l'Adour,  
5 France

6 <sup>b</sup>Geology Department, Faculty of Science, Mansoura University, Egypt

7 \*Corresponding author: [Mahmoud\\_lotfy@mans.edu.eg](mailto:Mahmoud_lotfy@mans.edu.eg)

8 **Abstract**

9 Various gases are generated in the subsurface, by diagenetic processes, fluid/rock  
10 interactions, maturation of petroleum source rocks and microbial activity. Often these  
11 gases just leak at the surface but sometimes they may accumulate in a subsurface trap.  
12 The present study investigates the case of the southern part of the Amadeus Basin,  
13 Australia, where exploration wells tested a blend of hydrocarbon (HC), helium (He) and  
14 hydrogen (H<sub>2</sub>). In Mt Kitty-1 well, He- and H<sub>2</sub>-rich gases were reported (He ~9%, H<sub>2</sub>  
15 ~11.4%) with traces of HC, whereas C<sub>1</sub> and C<sub>2+</sub> HC gases associated with He were  
16 discovered in Magee-1 well. Seismic and well data elucidate accumulation of the gas  
17 blend within a classical subsurface system with a structural trap, a permeable reservoir  
18 and a thick salt layer as seal. This salt, the Proterozoic Upper Gillen, is overlying the  
19 Heavitree sandstones and fractured basement reservoirs. The former is dominated by  
20 intergranular porosity (~13%), whereas the latter is commonly associated with complex  
21 patterns of fractures associated with deep-seated faults. The sources of blended gases  
22 are different: thermal cracking of organic matter for hydrocarbons, radiolytic decay of  
23 radioelements for helium. Radiolysis, and water reduction as well as oxidation of  
24 hydrocarbons are open hypotheses for H<sub>2</sub> generation in the study region.  
25 Neoproterozoic and Early Paleozoic tectonic episodes (e.g. Petermann; Alice Springs)  
26 allowed the formation of various structural traps where the gas blend was accumulated.

27 The large quantity of He, light gas slowly generated by the crustal radioactivity  
28 indicates that the seal is very efficient and that accumulations could have been in place  
29 for millions of years very likely with an evolving CH<sub>4</sub>/H<sub>2</sub>/He ratios through time. The  
30 presence of gas chimneys and surface depressions in the basin indicate that the  
31 generation, and/or migration of the gases are still active. All the data are consistent with

32 a long-term accumulation of this gas blend, different from the short term H<sub>2</sub> retention in  
33 the reservoir that was proposed in other contexts such as in Mali.

34 **Keywords:** Natural hydrogen; helium; hydrocarbon gases; Amadeus Basin; Australia.

### 35 **1. Introduction**

36 Generation of hydrocarbon and non-hydrocarbon gases is mostly dependent on the  
37 complex interactions between three main genetic processes (biogenesis,  
38 thermogenesis and abiogenesis) (Bernard, 1978; Schoell, 1983; Whiticar, 1999; Lin et  
39 al., 2005; Vacquand et al., 2018 and references therein). Deciphering the origin of these  
40 gases and various mechanisms of their generation is crucial for a proper exploitation of  
41 the geo-resources. Hydrocarbons (HC) are often expelled from thermally-mature  
42 organic-rich rocks after reaching the oil and gas window maturities, whereas biogenic  
43 processes commonly generate dry HC by consuming the light liquid HC phases at low  
44 temperature levels (Schoell, 1983; Peters and Casa, 1994; Whiticar, 1999; Leila and  
45 Moscariello, 2017). Radiolysis also generates helium (He) and hydrogen (H<sub>2</sub>) gases via  
46 the radiolytic decay of radioelements coupled with a possible dissociation of water  
47 (Ballentine et al., 1991; Warr et al., 2019, Klein et al., 2019 and references therein). The  
48 abiogenic processes can produce light hydrocarbon (mainly C<sub>1</sub> and C<sub>2+</sub>) and non-  
49 hydrocarbon (H<sub>2</sub>, CO<sub>2</sub>) gases at a wide range of temperatures in various geological  
50 environments (Coveney et al., 1987; Holloway and O'Day, 2000; Taran et al., 2010;  
51 Etiope, 2017; Prinzhofer et al., 2019; Moretti et al., 2021a, b; Leila et al., 2021; Radwan  
52 et al., 2022). These processes include redox reactions between water and Fe<sup>+2</sup>-bearing  
53 rocks (mafic, ultramafic rocks and ironstones) (Charlou et al., 2002; McCollom and  
54 Bach, 2009; Vacquand et al., 2018), radiolytic dissociation of water by radioelements  
55 (uranium, thorium) in crustal rocks (Lin et al., 2005; Sherwood Lollar et al., 2014),  
56 volcanism and volcanic degassing (Stefánsson and Arnórsson, 2002) as well as thermal  
57 metamorphism (Orlando et al., 2010).

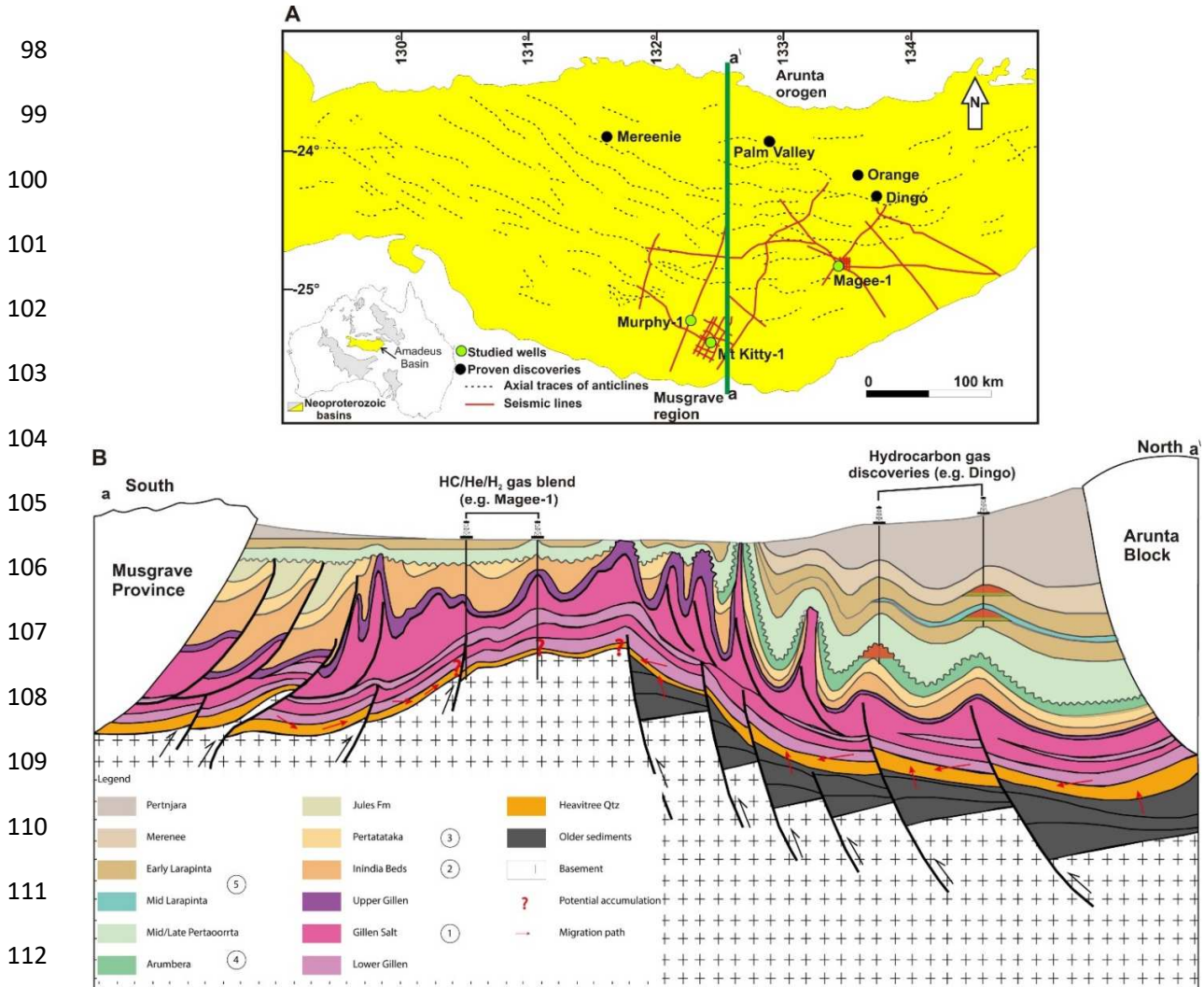
58 Natural H<sub>2</sub> generation via abiogenic processes was recently estimated at 2.8 M-  
59 tonnes/year (Truche et al., 2020), nonetheless, this estimation could be much more  
60 under-estimated (Zgonnik, 2020). Moreover, recent discovery of shallow H<sub>2</sub>-rich (~98  
61 mol%) gas in Bougou-1 well, Taoudeni Basin, in Mali (Prinzhofer et al., 2018), highlights  
62 the potentiality of subsurface natural H<sub>2</sub> exploitation. Just a few years ago, it was

63 claimed that H<sub>2</sub> does not exist as a free gas in the Earth's crust and the fact that a  
64 natural seal did not exist for such a low molecular weight gas was often affirmed and  
65 published (e.g. [Zgonnik et al., 2015](#); [Moretti et al., 2021a, b](#)). The implications of the  
66 subsurface H<sub>2</sub> storage and of the discovery in Mali was not really considered.  
67 Additionally, some geoscientists believe that the subsurface H<sub>2</sub> will be mainly produced  
68 within a dynamic system and a flow of H<sub>2</sub> may exist in a porous medium such as in  
69 Mali ([Prinzhofer et al., 2018](#)) or within the steam of the geothermal plants ([Stefánsson  
70 and Arnórsson, 2002](#); [Combaudon et al., 2022](#)). Hydrogen currently presents a zero  
71 carbon alternative for the traditional fossil fuels and the advances in H<sub>2</sub> exploration and  
72 production may aid in the decarbonization of the global energy supply (e.g. [Gielen et al.,  
73 2019](#)). Similarly, other non-HC gases such as helium is widely used in high-tech  
74 industries such as cryogenics, and the global demand for this important gas is currently  
75 increasing ([Skirrow et al., 2013](#); [Waltenberg, 2015](#); [Haines and Allen, 2019](#)).

76 Blended hydrocarbon and non-hydrocarbon gases were reported from several  
77 places in Australia such as the Cobar copper mine, New South Wales (~75% H<sub>2</sub> and ~  
78 25% CH<sub>4</sub>; [Stack and Rabone, 2018](#)), Yorke Peninsula and Kangaroo Island  
79 ([Woolnough, 1934](#)) as well as in some exploration wells offshore Victoria (Golden  
80 Beach-1 ~38 mol.% H<sub>2</sub>) and in Amadeus Basin (Mt. Kitty-1 ~11 mol.% H<sub>2</sub>) ([McInnes et  
81 al., 2017](#)). In addition, surficial evidence for subsurface gas emanations is recognized  
82 by sub-circular depressions (SCD) that have been reported from all over the globe  
83 (Russia, USA, Brazil, and Ukraine) ([Larin et al., 2015](#); [Zgonnik et al., 2015](#), [Prinzhofer et  
84 al., 2019](#); [Zgonnik, 2020](#); [Moretti et al., 2021a, b](#), [Geymond et al., 2022](#)). They are  
85 visible on satellite images and have been observed in several places in Australia where  
86 seeps of H<sub>2</sub>-rich gases have been reported ([Moretti et al., 2021b](#), [Frery et al., 2021](#)).

87 The present study focuses on the HC and non-HC gas blend discovered in the  
88 Amadeus Basin, central Australia ([Fig. 1](#)). The Amadeus Basin is an extensive  
89 intracratonic structure covering an area of approximately 170,000 km<sup>2</sup>. It is one of the  
90 Australia's most petroliferous provinces with hydrocarbon exploration history that has  
91 been started since the early 1960s ([Young and Ambrose, 2007](#)). The basin contains  
92 proven HC reserves estimated at approximately 37 million barrels of oil and more than  
93 400 BCF (billion cubic feet) of gases ([Jackson et al., 1984](#)). It constitutes a thick (~ 17

94 km) sedimentary cover of Proterozoic to late Devonian rocks (Fig. 2). Five distinct  
 95 petroleum systems have been identified in the Amadeus Basin; the first four are mainly  
 96 hosted in the Neoproterozoic rocks with sub-economic gas discoveries (e.g. Magee-1,  
 97 Mereenie, Orange-1, -2 and Dingo-1 to -5).



113 **Fig. 1.** A) Location map illustrating the location of the studied wells and proven gas  
 114 gas discoveries in Amadeus Basin, central Australia, B) Illustrative N-S cross-section (not to  
 115 scale) in central Amadeus Basin showing the main structural elements, the distribution  
 116 of different stratigraphic units (modified from Johns et al., 2017) and the kind of plays  
 117 that have been targeted.

118 The youngest petroleum system comprises the Ordovician sandstones hosting the  
 119 economic accumulations of large oil and gas such as the Mereenie and Palm Valley

120 fields (Fig. 1; Marshall, 2003). The source rock potential of the Neoproterozoic systems  
121 was considered to be poor or rather negligible (Jackson et al., 1984; Summons and  
122 Powell, 1991). However, the recent organic-geochemical and geologic data have  
123 demonstrated several potential source rocks in the Neoproterozoic successions in some  
124 parts of the basin (Marshall, 2003; 2004). Today, the non-hydrocarbon gas plays (H<sub>2</sub>,  
125 He) in the Neoproterozoic systems are actively explored (Hissey et al., 2016;  
126 Scrimgeour, 2019; Haines and Allen, 2019). The case of HC/He/H<sub>2</sub> gas blend in the  
127 Amadeus Basin is especially interesting since their kinetics are rather different and the  
128 gases have variable residence times in the reservoir (Lin et al., 2005). The factors  
129 controlling the generation, mixing and accumulation of this gas blend remain poorly  
130 constrained but in Australia the subsurface data are publicly available and we took this  
131 opportunity to investigate the surface and subsurface data, seismic and wells for  
132 potential geo-energy exploration plans. Additionally, with proven resources of both HC  
133 and non-HC gases, Amadeus Basin presents a natural laboratory where we can  
134 examine the controls on generation and accumulation of various gas plays.

## 135 **2. Geologic setting**

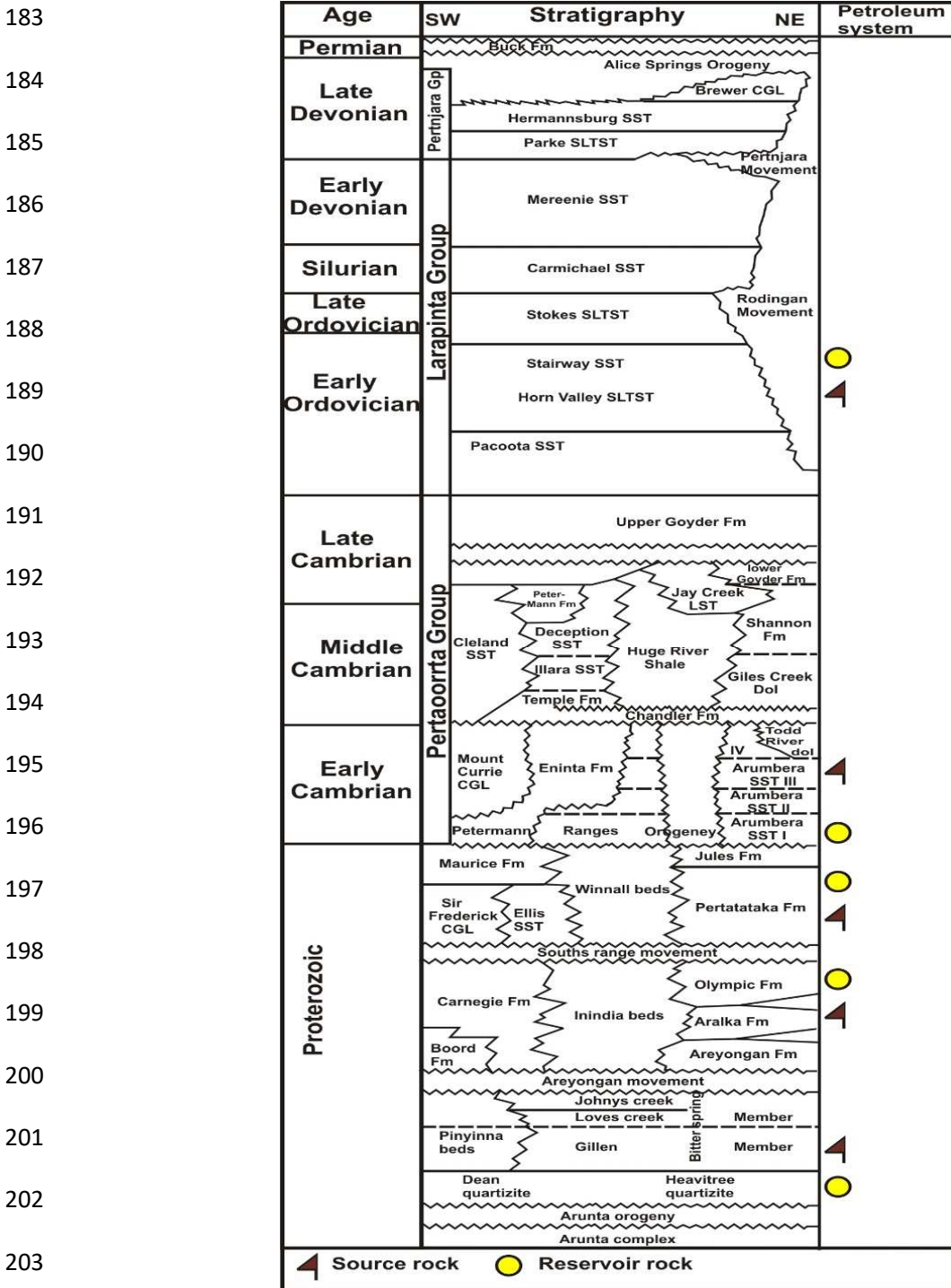
136 The Amadeus Basin presents an E-W trending belt of folded Neoproterozoic-Early  
137 Paleozoic (Cambrian-Devonian) sedimentary successions in central Australia overlying  
138 unconformably the Paleo- to Mesoproterozoic crystalline basement complex (Wells et  
139 al., 1970; Jackson et al., 1984; Marshall et al., 2007; Edgoose, 2013; Carr et al., 2016).  
140 The basement rocks are exposed in the northern (Arunta Orogen) and southern  
141 (Musgrave Complex) parts of the basin due to their uplift during Petermann (Ediacaran-  
142 Early Cambrian) and Alice Springs (Paleozoic) orogenies. Northward, the exposed  
143 basement complex consists of deformed meta-sedimentary and meta-igneous felsic and  
144 intermediate rocks, whereas the southern basement is constituted by high-grade  
145 metamorphic feldspathic gneiss and intrusive igneous bodies (Jackson et al., 1984;  
146 Marshall et al., 2007; Carr et al., 2016). The Amadeus Basin extended westwards  
147 where an NNE-striking shallow ridge of crystalline basement was emplaced during the  
148 Late Paleozoic orogeny. Prior to latest Paleozoic, sediment accumulation in Amadeus  
149 Basin was linked to a continuous deposition southward (Powell et al., 1994). The  
150 sedimentary basin-fill was initiated during the Early Neoproterozoic and the pre-salt

151 sediments comprise a quartzite-rich siliciclastic package of the Dean/Heavitree  
152 Formation which is exposed along the northern margins of the basin (Fig.2; Walter et  
153 al., 1995).

154 The Petermann and Alice Springs orogenies controlled sedimentation and  
155 distribution of the basin-fill sedimentary successions in Amadeus Basin. During  
156 Petermann Orogeny (~550-535 Ma), parts of the Amadeus Basin was either uplifted,  
157 eroded or locally metamorphosed along the southern margin with synchronous  
158 deposition of thick siliciclastic materials that were sourced from the uplifted parts (Bagas  
159 et al., 1996; Haines et al., 2015). The Petermann uplift phase was likely accompanied  
160 by salt movements; however, the irregular patterns of the folds reveal that the salt-  
161 related structures were evolved as earlier growth features prior to their reactivation  
162 during the orogenic uplift. Alice Springs orogeny (~400 Ma) largely affected the  
163 northeastern part of the basin and was accompanied by a contemporaneous deposition  
164 of synorogenic sedimentary facies that are preferentially localized in the northern and  
165 northeastern parts of the basin.

166 The Heavitree Formation comprises quartz-rich sandstones overlying unconformably  
167 the basement complex. The sandstones are intercalated with siltstones and  
168 conglomerate, and are mostly of fluvial to shallow marine origin. The Heavitree  
169 sandstones are followed by carbonates and fine-grained siliciclastic and evaporites of  
170 Bitter Springs Group. The Heavitree-Bitter Springs contact is only conformable in the  
171 eastern parts of the basin. In the northern parts, the Bitter Springs Group comprises,  
172 from bottom to top, the Gillen, Loves Creek and Johnny Creek formations (Fig. 2). The  
173 Gillen Formation consists of thick evaporite deposits (halite, gypsum) interbedded with  
174 dolomite, shale and siltstone. The Loves Creek Formation is made of stromatolitic  
175 dolomite facies unconformably overlying the Gillen facies and conformably followed by  
176 the lacustrine red dolomitic and evaporitic mudstone of the Johnys Creek Formation.  
177 The post-Bitter Springs stratigraphy comprises thick mixed carbonate and siliciclastic  
178 sediments resting unconformably above the evaporitic mudstone and dolomite of the  
179 Johnys Creek Formation. The Amadeus Basin contains thick, well-developed Paleozoic  
180 sedimentary successions (mostly shales, siltstone and minor sand interbeds) except in

181 the western regions where the Paleozoic sediments are comparatively thinner  
 182 (Marshall, 2003; Haines et al., 2010; Normington et al., 2015).





204 **Fig. 2.** Generalized composite stratigraphic column for the Amadeus Basin illustrating  
205 the tectono-stratigraphic framework of the basin and the possible source rocks and  
206 reservoir intervals (Marshall, 2004).

207 Hydrocarbon exploration to date in the Amadeus Basin focused on the Paleozoic  
208 successions principally targeting the structural and combined structural-stratigraphic  
209 plays in the shallow salt-related structures in the north. The discoveries so far are post-  
210 salt and located in the northern part of the basin, however, the subsalt strata are  
211 shallower southward and exploration has been started there since the last decade. For  
212 the pre-salt petroleum systems, Heavitree Formation presents a potential reservoir that  
213 could be sourced from the shale intervals in Bitter Springs Group (Marshall, 2003). The  
214 evaporites of Gillen and Loves Creek Formations are likely to be effective cap rock for  
215 the gases that may accumulate in Heavitree reservoir.

### 216 **3. Data and methods**

217 The Australian policy towards the dissemination and sharing of scientific advances  
218 allows access to a lot of information directly on the government site of the Department  
219 of Geosciences (<https://www.ga.gov.au/>). The available subsurface dataset utilized in  
220 this study includes 2D seismic, wireline logs as well as analytical internal reports.

#### 221 **3.1. Seismic dataset**

222 The subsurface seismic dataset constitutes regional post-stack 2D seismic grid  
223 covering an area of approximately 30,000 km<sup>2</sup> in the eastern and central parts of the  
224 Amadeus Basin. In addition, local 3D seismic blocks around the studied Magee-1 and  
225 Mt Kitty-1 wells have also been utilized in this study (Fig. 1). The seismic data was  
226 interpreted using Petrel Schlumberger 2017 software following the standard seismic  
227 interpretation scheme of Herron (2011). The seismic data (time) was tied with well logs  
228 (depth) using check shots that were available for all the studied wells.

#### 229 **3.2. Wireline logging analysis**

230 A complete set of wireline logs for Magee-1, Mt Kitty-1, and Murphy-1 wells have  
231 been interpreted using the Schlumberger Techlog 2017 software. The available well  
232 logs include gamma ray (GR), bit size (BS), caliper (Cal), spontaneous potential (SP),  
233 shallow and deep resistivity (LLS, LLD), photoelectric factor (PEF), neutron (NPHI) and

234 density (RHOB). The GR and NPHI logs were corrected for environmental effects such  
235 as borehole size, mud weight and salinity. DST data from Magee-1 well was used to  
236 drive  $R_w$  values which is approximately  $0.04 \Omega m$  at  $77^\circ C$ . The cementation exponent  
237 “m” was derived using the triple porosity model (Aguilera and Aguilera, 2004). The  
238 wireline logging interpretation includes a qualitative and semi-quantitative estimation for  
239 the different petrophysical (porosity  $\Phi$ , permeability  $K$ , water saturation  $S_w$ ) and  
240 lithological parameters using the standard logging interpretation technique.

### 241 **3.3. Source rock evaluation and gas analysis**

242 Available analyses for Total Organic Carbon (TOC) and Rock-Eval pyrolysis carried  
243 out by the operators on 12 cutting samples from Upper (2 samples) and Lower Gillen (8  
244 samples) member as well as Inindia Beds (2 samples) in Magee-1 and Murphy-1 wells  
245 were utilized in this study (Table 1, Geotech, 1992). At that time, the cuttings were  
246 analyzed using Rock-Eval 2 apparatus following the procedure described by Espitalié et  
247 al., (1986). The Rock-Eval pyrolysis peaks include  $S_1$ ,  $S_2$  and  $S_3$  values which  
248 respectively reflect the amount of free, pyrolyzed and organically-bound  $CO_2$   
249 hydrocarbons. Additionally, the temperature at  $S_2$  peak ( $T_{max}$ ) reflects the source rock  
250 maturity (Peters and Cassa, 1994). Geochemical gas analyses comprise the  
251 measurement of the molecular gas composition and relative abundance of different  
252 hydrocarbon and non-hydrocarbon gases using chromatographic separation. The  
253 identified hydrocarbon gas phases comprise  $C_1$ ,  $C_{2+}$  components whereas the non-  
254 hydrocarbon gases constitute  $N_2$ , Ar,  $CO_2$ ,  $H_2$ , and He (Table 2).

### 255 **3.4. Basin modelling procedure**

256 Burial and thermal history models were constructed for Magee-1 well using  
257 Schlumberger PetroMod 1D software to simulate the conditions of hydrocarbon  
258 generation and expulsion. The input data for the models comprises a detailed  
259 stratigraphic information, thermal boundary conditions, as well as the present-day  
260 thermal maturity values. The thermal boundary conditions include the sediment-water  
261 interface temperature (SWIT) and the basal heat flow (HF). The paleowater depth  
262 (PWD) was determined based on the depositional environment of the sediments. This is  
263 an important parameter to define the thermal boundary condition (SWIT). The surface

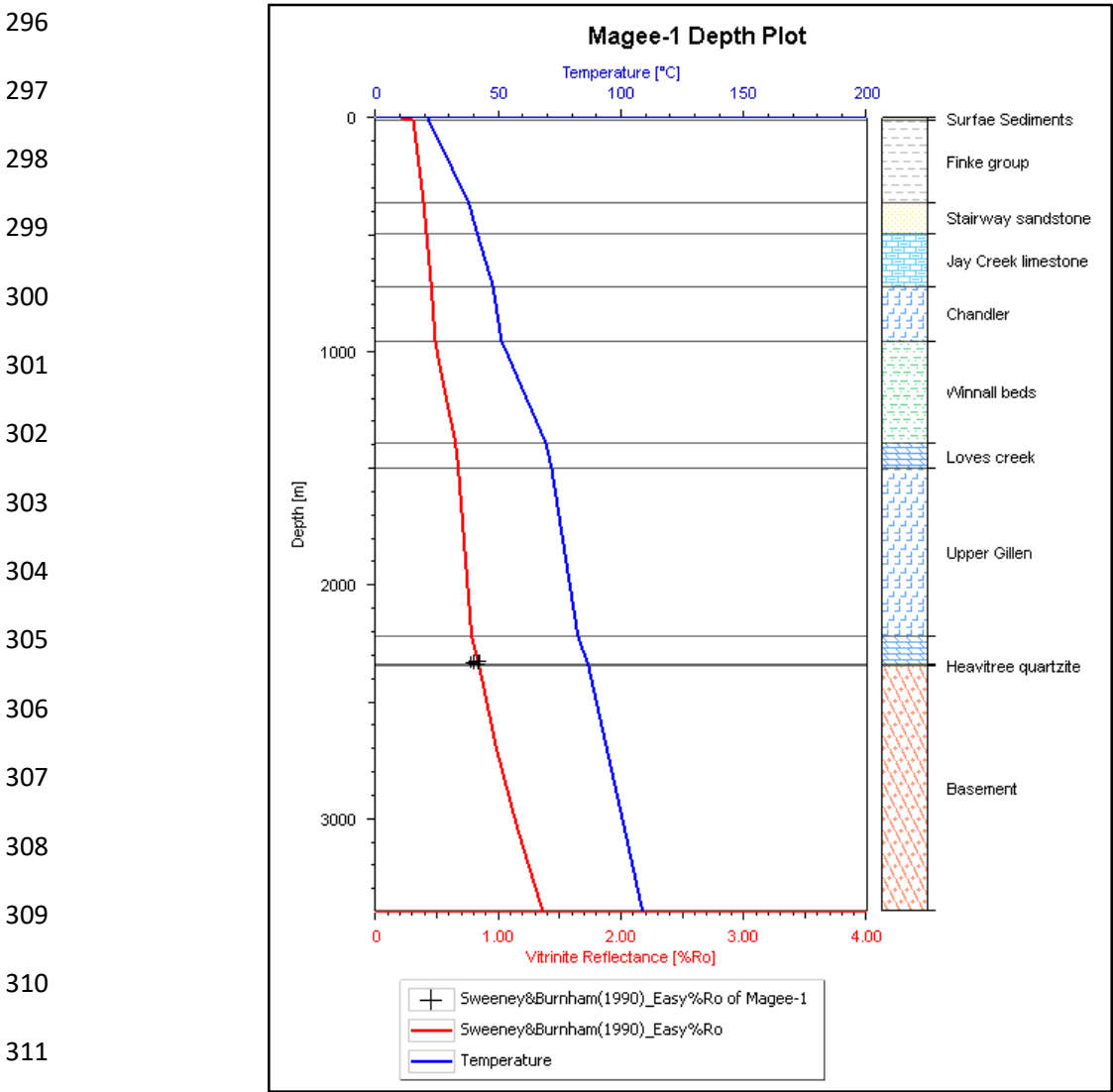
264 temperature trends have been synthesized for Magee-1 well in southern Australia  
 265 (24°S) using the time-latitude diagrams (Wygrala, 1989). The lower thermal boundary  
 266 condition, paleo heat flow (HF, values in mW/m<sup>2</sup>) values were calculated using the  
 267 known plate tectonic framework and crustal evolution models (Fig. 3; Allen and Allen,  
 268 1990). The basal heat flow values follow a general trend reporting a maximum value of  
 269 113 mW/m<sup>2</sup> during the main rifting phase of Areyonga, reducing progressively during the  
 270 compressional tectonic phases of Petemann, Bloodwood, and Rodingan movement and  
 271 finally reached to 50 mW/m<sup>2</sup> in the present-day (Fig. 3).



281 **Fig. 3.** The thermal boundary condition assessment for Magee-1 well including: (top)  
 282 the paleowater depth (PWD, in meter) and (middle) the sediment-water interface  
 283 temperature (SWIT, in °C), and (bottom) the paleo-heat flow (HF, in mW/m<sup>2</sup>). The basal  
 284 heat flow values were specified for each geologic event using the present-day analogs  
 285 and extrapolated through time (Allen and Allen, 1990).

286 The simulation results were calibrated using equivalent vitrinite reflectance data  
 287 (VRE) deduced from the measured Rock-Eval T<sub>max</sub> values (Fig. 4). The calculated  
 288 vitrinite reflectance from temperature histories was carried out using the EASY-%R<sub>o</sub>  
 289 algorithm of Sweeney and Burnham (1990). The predefined kinetic model\_TII(B) of  
 290 Pepper Corvi (1995) was designated based on the source rock type and depositional  
 291 environment. The timing of hydrocarbon expulsion was deduced considering that 20%

292 of the porous volume of source rocks should be filled by hydrocarbons prior to the  
 293 expulsion onset (Ungerer, 1990). The assigned initial TOC and HI ( $TOC_0$ ,  $HI_0$ ) were  
 294 calculated from the measured TOC and HI following the procedure described by (Peters  
 295 et al., 2005).



312 **Fig. 4.** The depth plot of the present-day thermal maturity of the Magee-1 well calibrated  
 313 with the theoretically calculated equivalent vitrinite reflectance (VRE) values.

314 **3.5 Petrography**

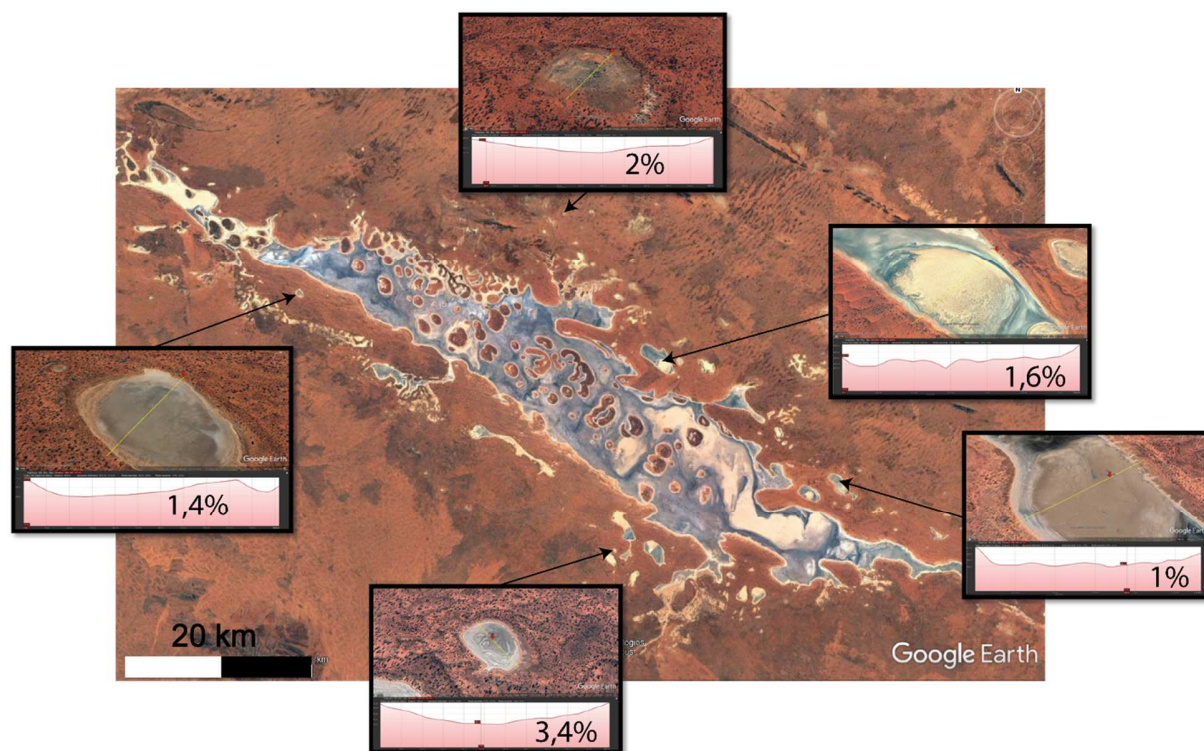
315 Three thin sections were prepared from the cutting and SWC samples retrieved from  
 316 the 7 m thick Heavitree sandstones in Magee-1 well (Geotech, 1992). The cuttings were  
 317 impregnated with araldite and blue dye to facilitate the visualization of pore spaces and

318 quantification of porosity. Additionally, the thin sections were stained with sodium  
319 cobaltinitrite to identify feldspars.

## 320 4. Results

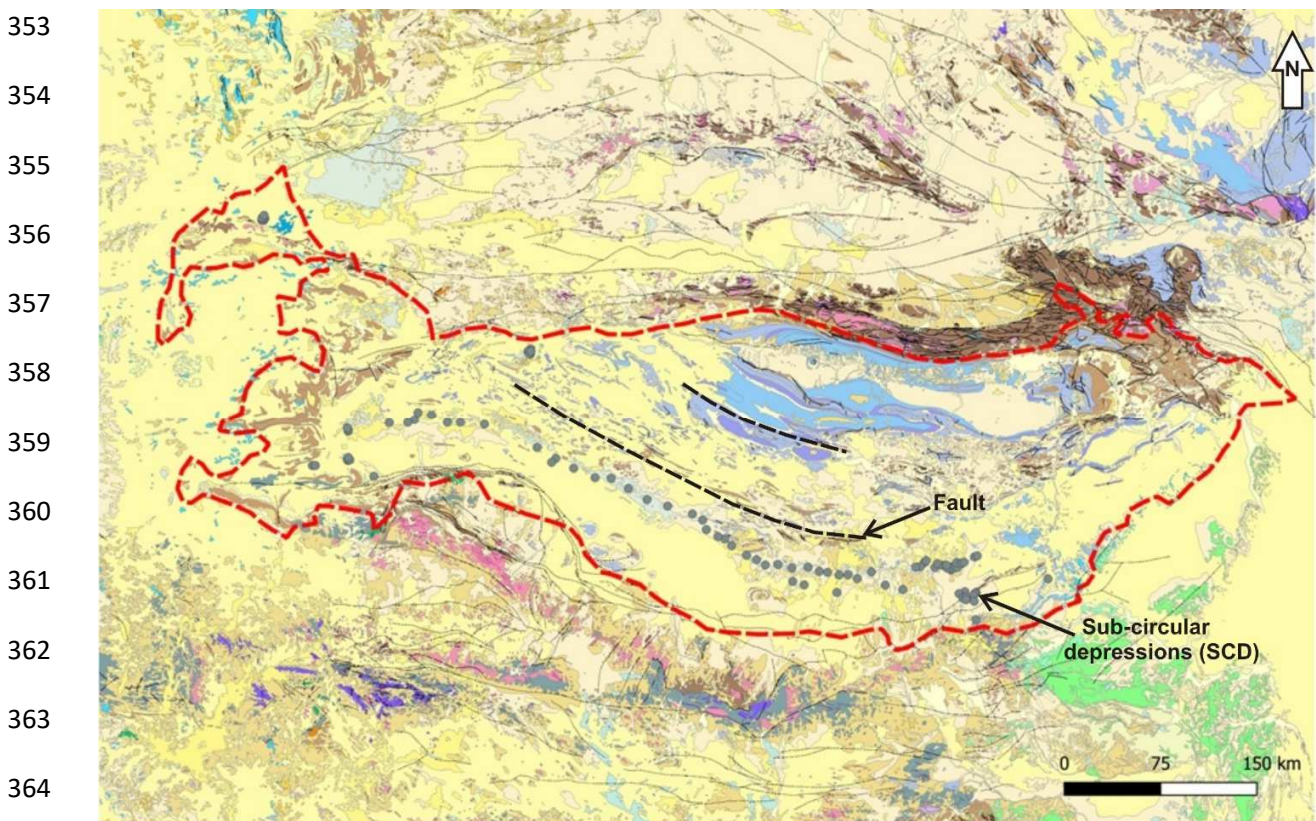
### 321 4.1. Surface emanations

322 Several sub-circular depressions (SCD) have been recognized from the satellite  
323 images in the study area. The observed SCD are similar to those reported for surficial  
324 H<sub>2</sub> emitting structures, the so-called fairy circles, as described in south and west  
325 Australia (Moretti et al., 2021a; Frery et al., 2021, Geymond et al., 2022). High densities  
326 of scattered, slightly elongated SCD are observed near the Amadeus Lake (Fig. 5). The  
327 diameter of the SCD structures varies from several tens to hundreds of meters and  
328 most of them have diameters greater than 400 m. The large SCD often display  
329 concave-up elevation profiles with smooth topography, whereas the small depressions  
330 (< 200 m diameter) have sharp, concave-up and irregular elevation profiles.



342 **Fig. 5.** Google Earth photographs illustrating occurrence of several aligned Sub Circular  
343 Depressions in Amadeus Basin with variable depth/equivalent diameter ratios  
344 highlighted in red.

345 The depressions also contain some local highs usually localized in the center. The  
346 depth/equivalent diameter ratio of the SCD ranges from 1% to more than 3% (Fig. 5).  
347 These values are very similar to the ones reported in other parts of Australia and in the  
348 areas where H<sub>2</sub> emitting structures have been mapped (Moretti et al., 2021b, Geymond  
349 et al., 2022). Moreover, a trend of SCD is observed to the south of the Amadeus Lake.  
350 They are aligned in NW-SE direction parallel to the faults associated with Petermann  
351 Orogeny (Fig. 6) suggesting a possible gas leakage along these faults similar to that  
352 reported elsewhere such as in Russia (Larin et al., 2015).



365 **Fig. 6.** Map illustrating the distribution of the sub-circular depressions (SCD) in the  
366 southern part of Amadeus Basin, grey dots. The depressions are aligned parallel to the  
367 master NW-SE trending fault trends affecting the basin. One may note the highest  
368 density of the SCD eastward near the wells that proven the presence of H<sub>2</sub>.

369

370

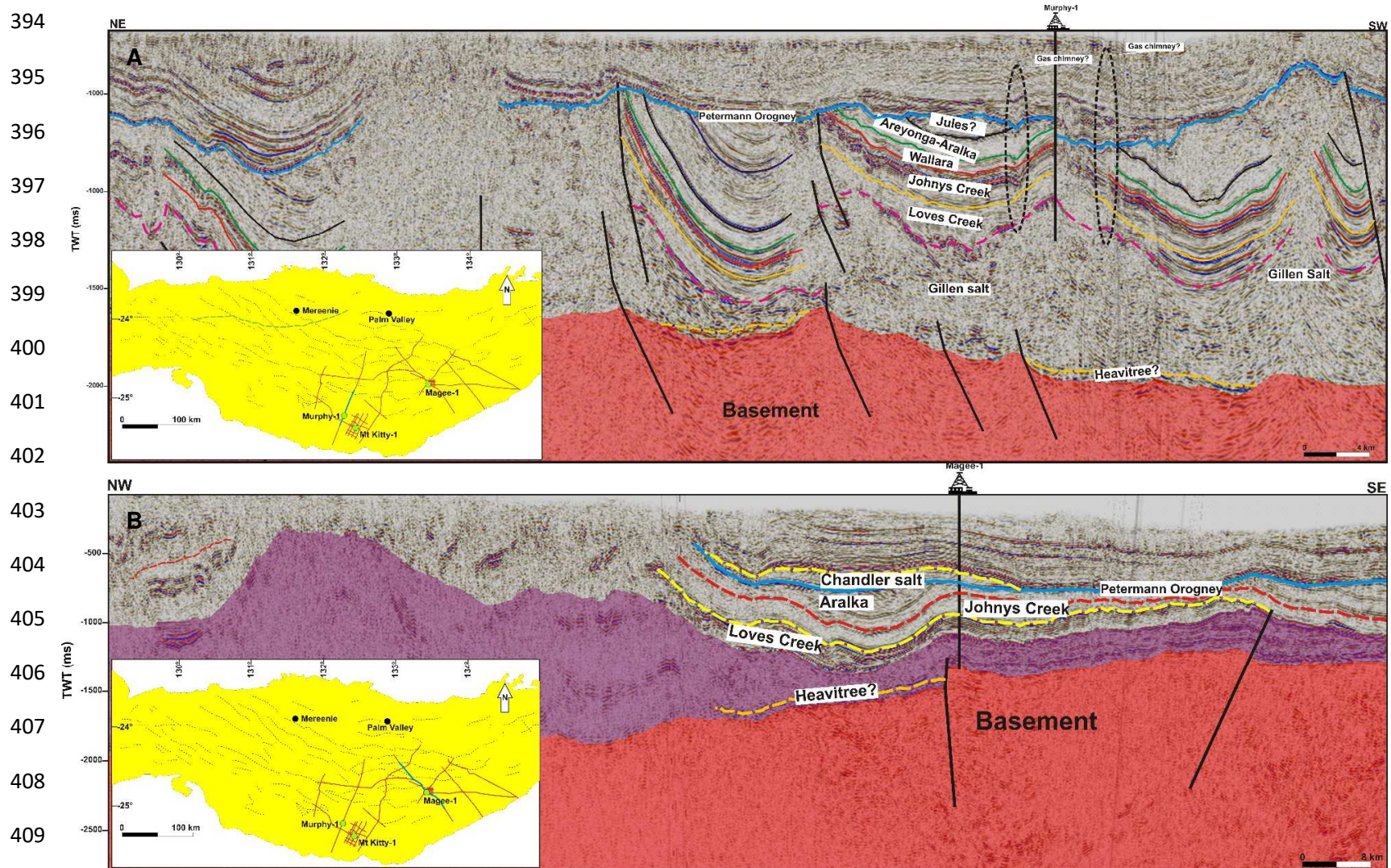
371

372

## 373 **4.2. Seismic interpretation**

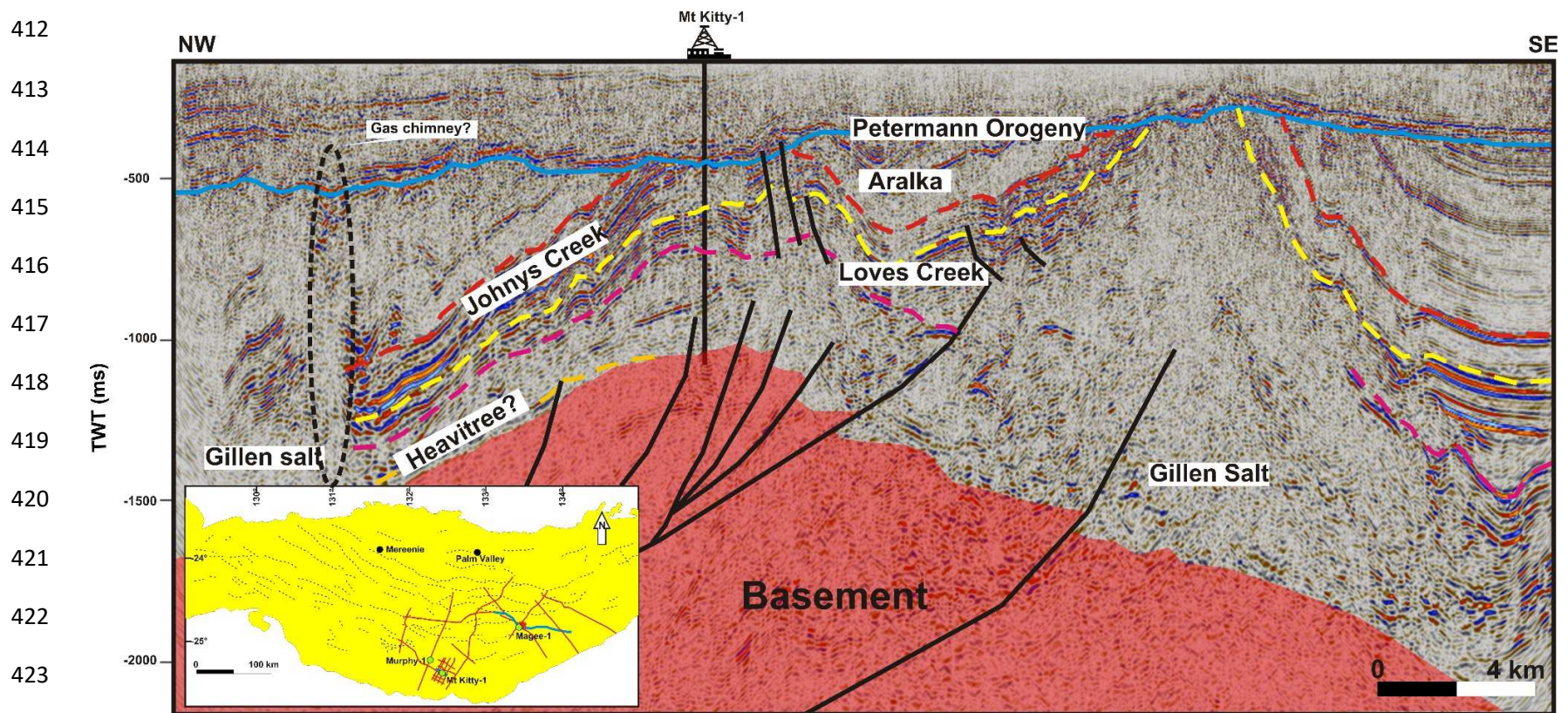
374 The interpreted 2D seismic profiles delineate the complex structural elements that  
375 influenced the stratigraphic architecture of the study region. The seismic profiles  
376 illustrate chaotic and reflection free seismic reflections interpreted as crystalline  
377 basement that are often dissected by large deep-sated faults (Figs. 7A, B and 8). The  
378 basement is picked at relatively shallow depths (< 2000 millisecond two-way time ms  
379 TWT), however several basement highs (< - 1700 ms TWT) are locally observed. The  
380 basement top is defined by irregular, undulated non-conformity surface. Thin-skinned  
381 deformation features probably associated with the Petermann Orogeny with major faults  
382 detached within the basement and the overlying pre-salt and salt layers are identified as  
383 the most prominent structural elements. The faults are mainly dipping westward with  
384 displacements up to 300 ms TWT at the top of the Gillen Formation. The basement  
385 faults were likely reactivated during Petermann Orogeny resulting in a formation of  
386 multiple local antiforms associated with inverted basement blocks (Fig. 8).

387 The pre-salt Heavitree Formation is represented by discontinuous high amplitude  
388 and reflectivity seismic reflections locally onlap the non-conformity surface topping the  
389 basement (Fig. 7A, B). The Heavitree reflections are localized in the basement lows and  
390 depressions adjacent to the deep faults (Fig. 7A), where they attained a maximum  
391 thickness of less than 100 ms TWT. The basement is relatively deeper in the  
392 southwestern part of the study area where its top is picked at more than - 2500 ms TWT  
393 (Fig. 9A).



410 **Fig. 7.** Interpreted 2D seismic profiles showing the distribution of the different stratigraphic horizons and the major  
 411 structural elements. Seismic location: green line on the map.





424 **Fig. 8.** Interpreted 2D seismic profile illustrating the main fault trends in Mt Kitty structure.

425

426

427

428 The Gillen salt is characterized by very low amplitude, transparent to semi-  
429 transparent chaotic reflections (Figs. 7A, B and 8). Lower Gillen reflections are less  
430 chaotic with moderate amplitude and reflectivity, locally displaying discrete patches that  
431 are localized in the basement highs (carbonate build-ups?). In some parts such as in  
432 Magee-1 well area, these patches are followed by thin (~ 15 ms TWT), high amplitude  
433 horizontal, continuous reflections interpreted as halite layers. The top of the Gillen salt is  
434 picked at variable depths in the study area, however, it is relatively shallower and  
435 thicker in the west than in the east (Fig. 9B). Loves Creek reflections are semi-  
436 transparent and display low to moderate amplitude and reflectivity. They grade upward  
437 into the parallel, high amplitude and reflectivity reflections of the Johnys Creek  
438 Formation. The seismic reflections of the Bitter Springs Group are terminated against  
439 Petermann unconformity characterized by high amplitude, continuous, and irregular  
440 reflector erosionally-truncating the pre-Petermann reflections. The Petermann  
441 unconformity is picked at greater depths in the eastern part of the study area (~ -1000  
442 ms TWT, Fig. 10A). The Gillen salt has been mobilized possibly during the Petermann  
443 Orogeny, and therefore its thickness varies greatly in the study region (Fig. 10B).  
444 Reactivation of the basement faults during the Petermann Orogeny is more prominent in  
445 the western part of the study area resulting in the inversion of the basement blocks  
446 associated with formation of flower-like structures (Fig. 8).

447 Several forms of chaotic and disordered vertical disturbances have been observed  
448 on seismic and were interpreted as gas escape structures (gas chimneys). They are  
449 represented by vertically-aligned columnar features associated with an abrupt decrease  
450 in the reflection amplitude and continuity (Figs. 7A, B and 8). The observed gas  
451 chimneys often traced vertically from the basement to the Petermann unconformity  
452 where they often disturb the overlying layers. They range in size on the horizontal scale  
453 from several few tens of meters up to 500 m (Fig. 11). None of these features are  
454 associated with pockmarks that often reflect entrapment of the migrated gases. In this  
455 kind of salt-rich basin, some of these disturbances in seismic reflections may also be  
456 induced from the mobilization of salt forming salt scars.

457

458

459

460

461

462

463

464

465

466

467

468

469

470

471

472

473

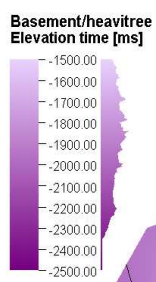
474

475

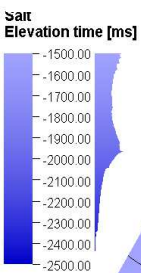
476

477

478



50000m



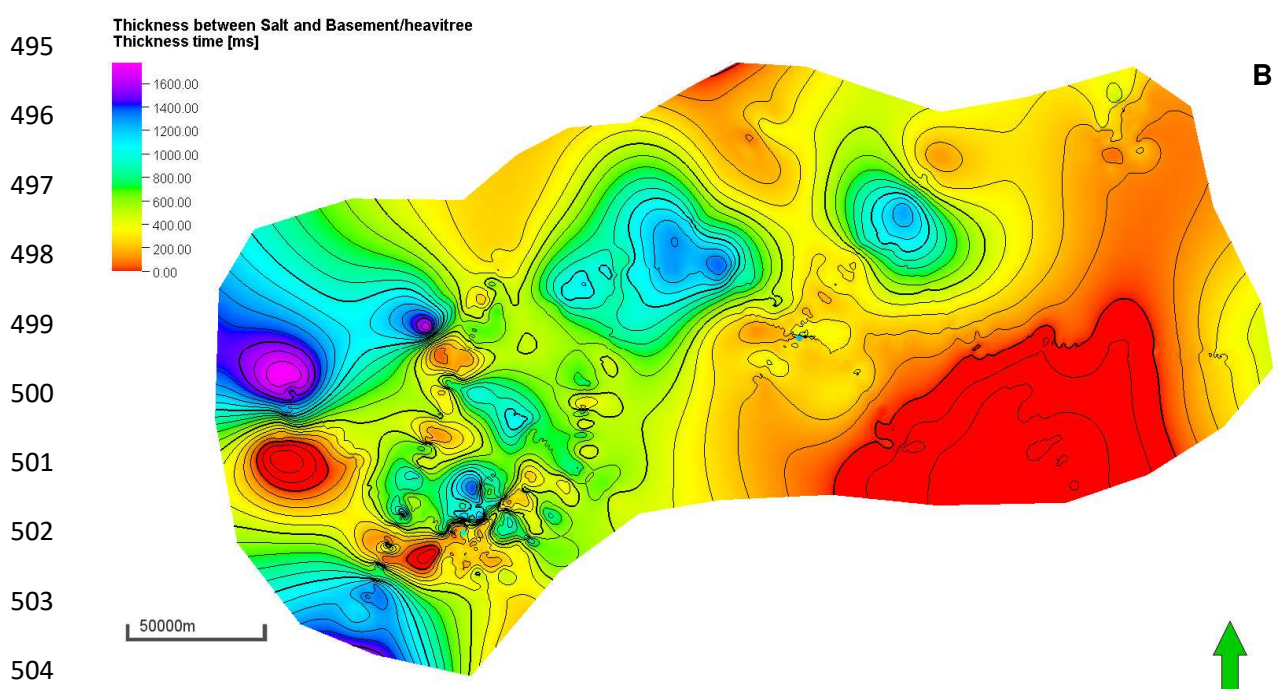
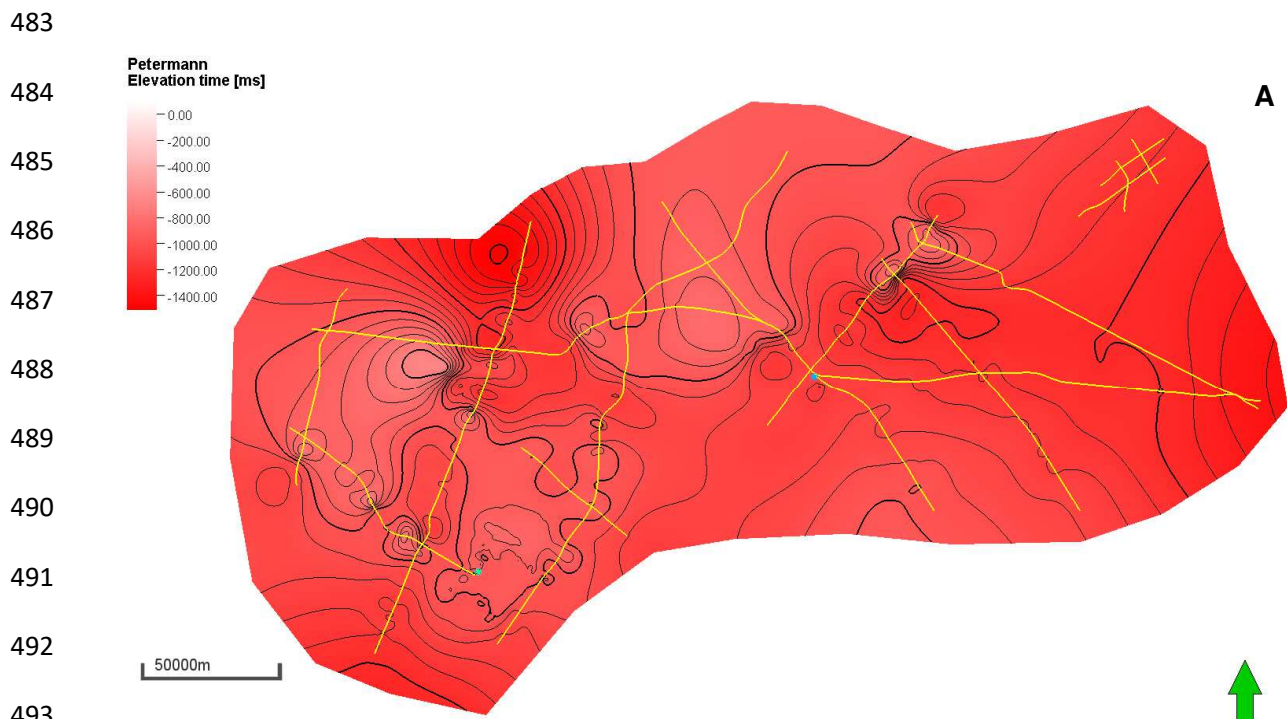
50000m

A

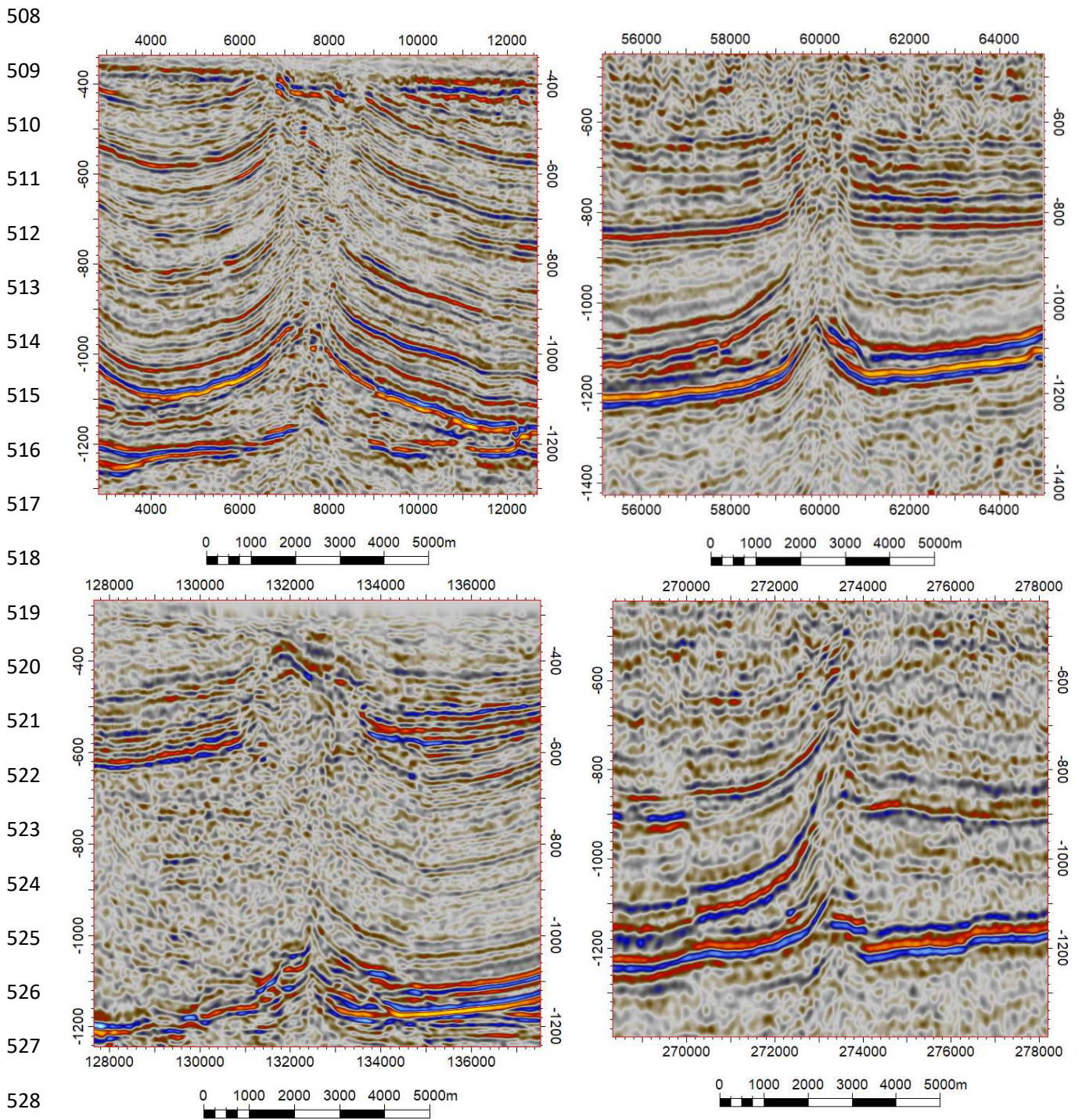
Gas  
chimneys ?

B

479 **Fig. 9.** A) Depth contour maps at the top of Heavtree Formation/basement and Gillen  
480 Salt (B) illustrating relative deepening in the southwestern part of the study area.  
481 Projected gas chimneys often occur in the eastern part where the basement and salt  
482 are relatively shallow.



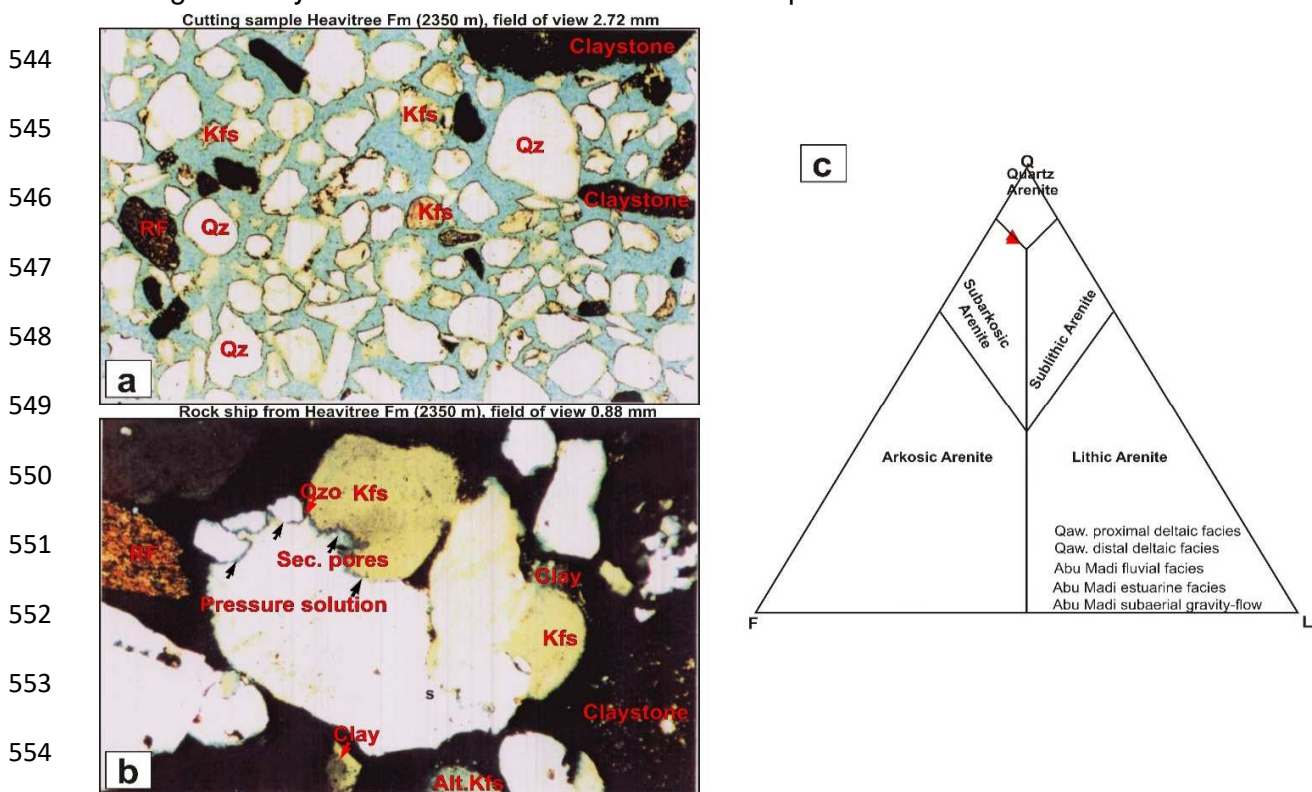
505 **Fig. 10.** A) Depth contour maps at the top of Petermann unconformity; B) Thickness  
 506 map showing variations in salt distribution and its thinning in the eastern part of the  
 507 study area.



529 **Fig. 11.** Local vertical disturbances in the seismic reflections possibly reflect occurrence  
 530 of gas chimneys and/or salt scars.

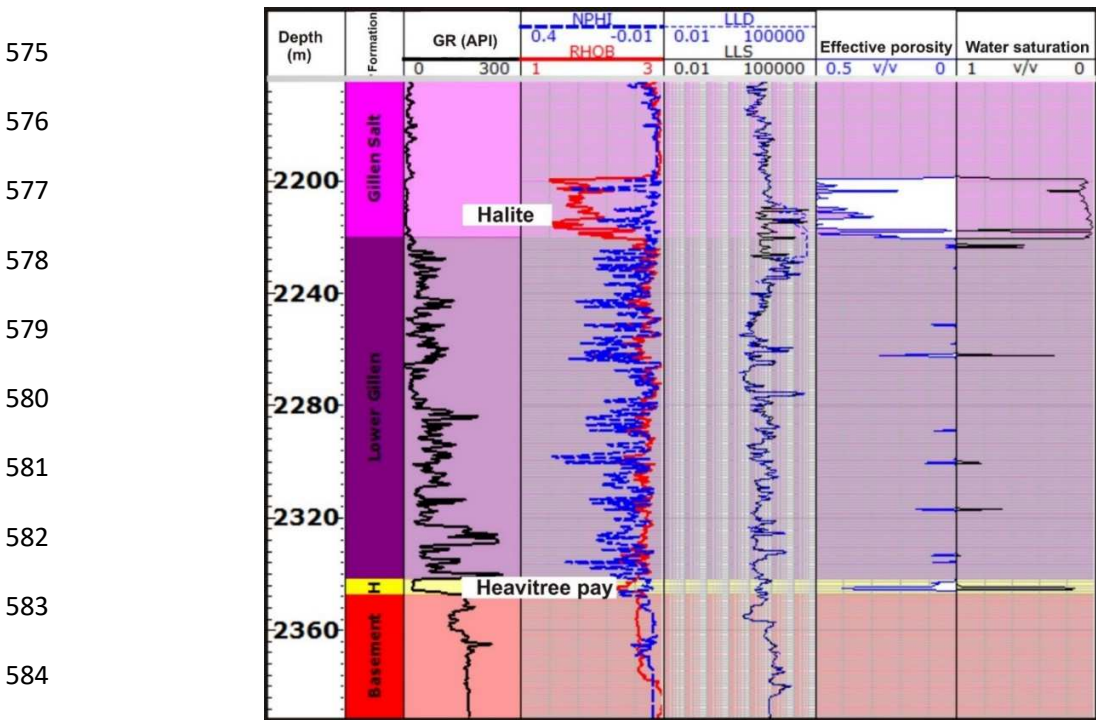
531 **4.3. Lithologic composition of Heavitree reservoir**

532 The Heavitree Formation consists mainly of quartzose sandstone with a framework  
 533 composition dominated by quartz, feldspars and subordinate lithic fragments (Fig. 12A,  
 534 B). The framework components are medium to coarse grained, poorly to moderately  
 535 sorted and sub-rounded to sub-angular. The feldspars are fresh to partly altered and are  
 536 mainly represented by K-feldspars (Av. 8%) with minor plagioclase (Av. 1%). The lithic  
 537 fragments (Av. 5.5%) are mainly sedimentary (claystone) with minor igneous and  
 538 metamorphic fragments. The sandstones have an average composition of Q<sub>77</sub>F<sub>15</sub>L<sub>8</sub>, and  
 539 are therefore classified as quartz to subfeldspathic arenite (Fig. 12C). The framework  
 540 components are tightly-compacted and display sutured and concavo-convex grain  
 541 contacts (Fig. 12B). Moreover, the feldspars are partly dissolved forming secondary  
 542 pore spaces. Quartz overgrowths are often associated with the sutured grain contacts.  
 543 Authigenic clays are observed on the altered feldspars.



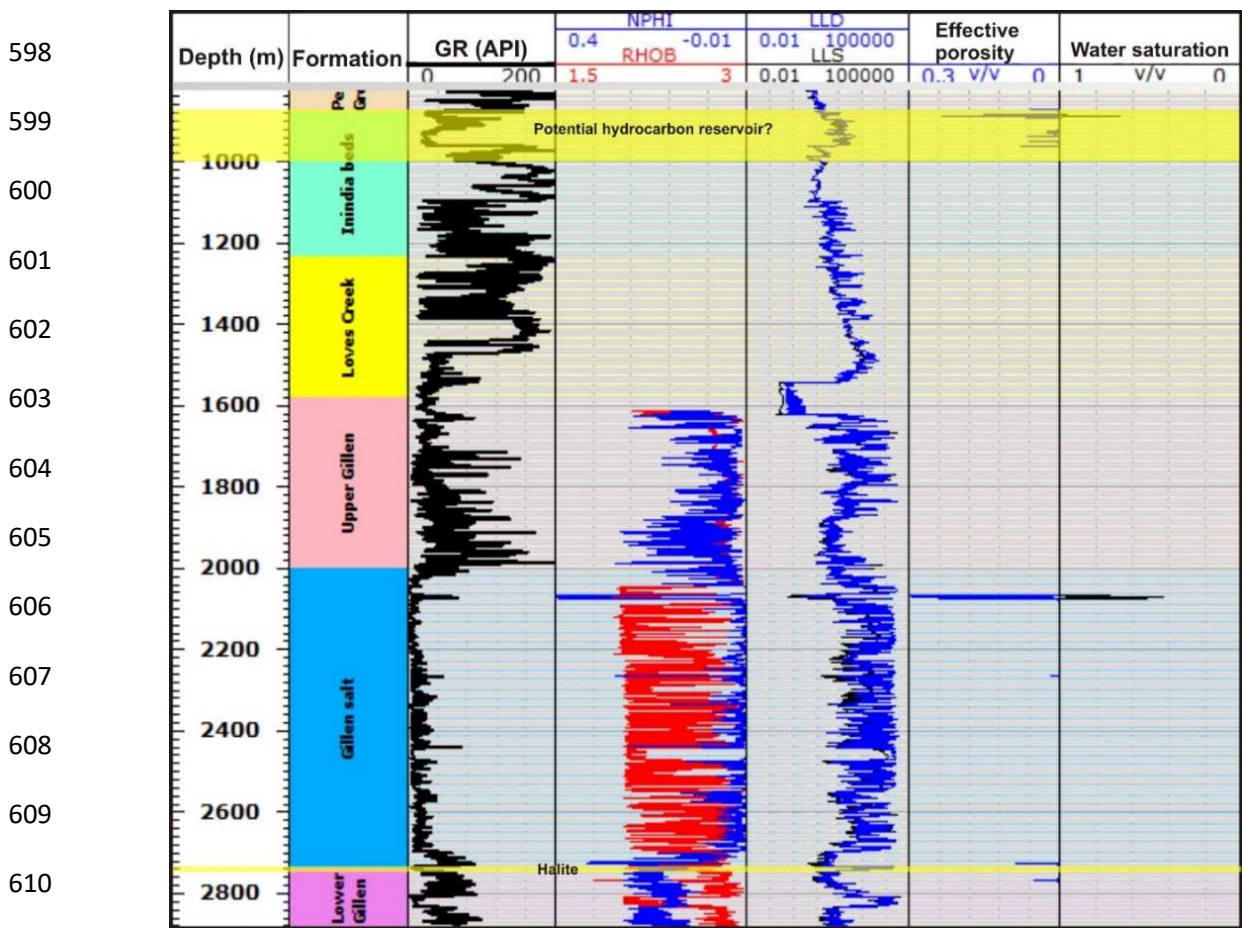
560 **4.4. Wireline log analysis**

561 Interpretation of wireline logs aids in the identification of different petrophysical  
 562 parameters that determine the reservoir potential of the Neoproterozoic succession in  
 563 the study area. In Magee-1 well, porous gas-bearing intervals are scarce and often  
 564 localize above the basement at the base of the Gillen salt (Fig. 13). Heavitree pay (6 m  
 565 thick) is characterized by a blocky gamma ray pattern with very low values (20-30 API).  
 566 The Heavitree pay zone has an average clay volume of 17%, however, most of the clay  
 567 materials are present in structural forms (claystone fragments, Fig. 12A) and therefore  
 568 they do not adversely impact the porosity values (Av. 13%). The pay zone sandstones  
 569 also display low water saturation values (Av. 34%) and a notable neutron-density  
 570 crossover typifying the presence of trapped gas. Another porous, gas-bearing interval is  
 571 observed above the Lower Gillen succession. This interval consists of approximately 20  
 572 m thick halite layers characterized by very low gamma ray values (< 10 API). This halite  
 573 interval is also traced in Murphy-1 well at the base of Gillen Salt (Fig. 14). The halite  
 574 beds have very high porosity (Av. 23%) and very low clay volume (< 5%) values.



585 **Fig. 13.** Vertical distribution of the different wireline logs data including the calculated  
 586 values of effective porosity and water saturation in Magee-1 well.

587 The abnormally high porosity of the halite possibly reflects a gas entrapment in the  
 588 pore spaces during halite crystallization (Blamey et al., 2016). The elevated resistivity  
 589 values (> 200 ohm.m) and the prominent neutron-density crossover coupled with the  
 590 very low water saturation values (Av. 9%) are evidence for gas entrapment in the halite  
 591 crystal lattice. In Mt Kitty-1 well, porous intervals are locally observed in the basement  
 592 rocks where the effective porosity values are up to 13% with an average value of 4%  
 593 (Fig. 15). The basement contains several fracture zones characterized by anomalously  
 594 high resistivity values with a notable separation between shallow and deep resistivity  
 595 curves characteristics for gas-filled fracture network (Sw < 50%). The fracture zones  
 596 also display very low formation matrix density and high neutron values typifying a  
 597 possible invasion of mud filtrate.

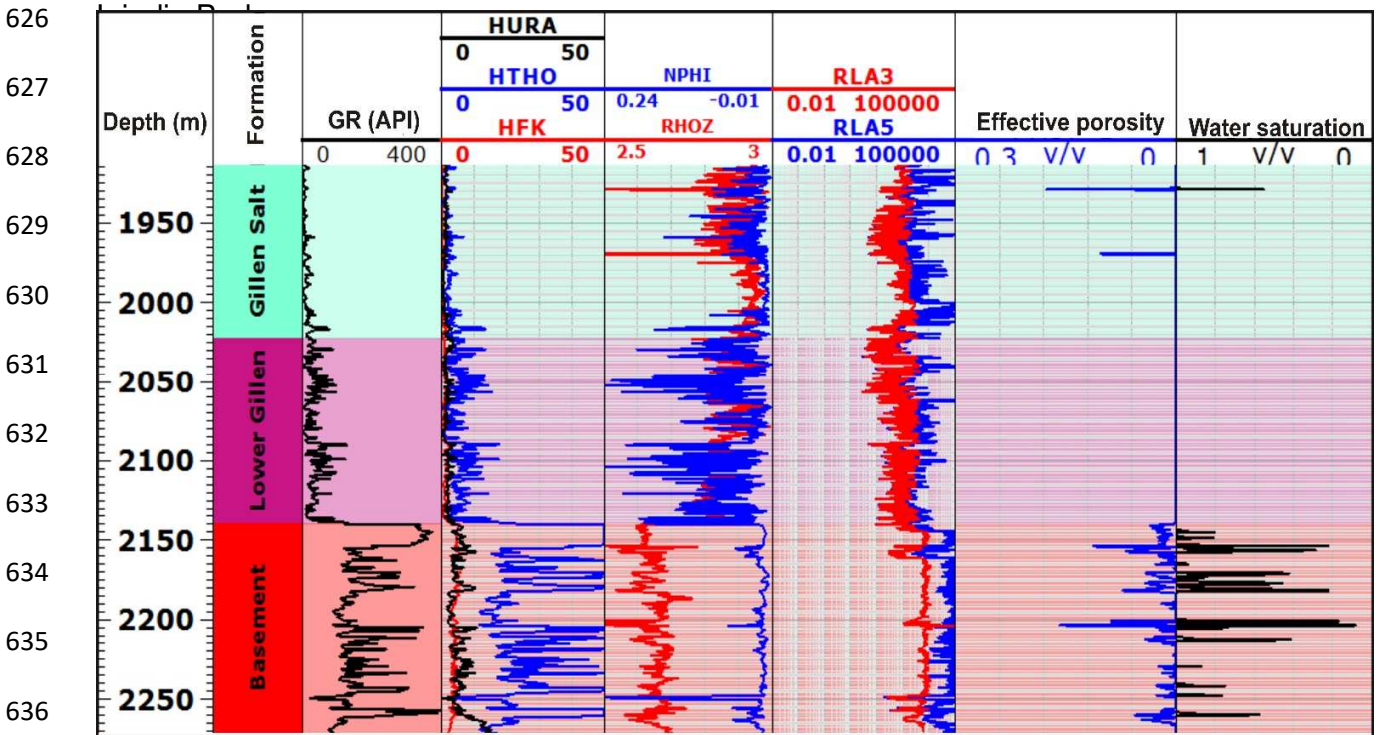


611 **Fig. 14.** Vertical distribution of the different wireline logs and calculated values of  
 612 effective porosity and water saturation showing some potential above salt reservoir  
 613 intervals in Murphy-1 well.



614 **4.5. Source rock characteristics**

615 The investigated shales from Lower and Upper Gillen have low TOC values < 1 wt%.  
 616 Those of Lower Gillen have slightly higher TOC values (Av. 0.57 wt%) than the Upper  
 617 Gillen samples (Av. 0.50 wt%). Elevated TOC values (however lower than 2 wt%) are  
 618 observed in the shales of Inindia Beds. Most of the TOC content is attributed to the  
 619 residual carbon (RC), whereas the pyrolyzed carbon (PC) values are very low (Table 1).  
 620 The Rock-Eval S<sub>1</sub> values are low in the Lower Gillen and Inindia Beds shales (Av. 0.1  
 621 mg HC/rock) and are comparatively higher in the Upper Gillen shale samples (Av. 1.5  
 622 mg HC/rock). Rock-Eval S<sub>2</sub> values are lower than 1 in all the studied rocks. Similar to  
 623 S<sub>2</sub>, the Rock-Eval S<sub>3</sub> values are less than 1 with values ranging from 0.04 to 0.60.  
 624 Additionally, the T<sub>max</sub> values vary in a wide range in lower and Upper Gillen samples  
 625 (393-445°C and 350-393°C, respectively), and they range between 433°C and 435°C in



637

638 **Fig. 15.** Vertical distribution of the different wireline logs showing potential reservoir  
 639 intervals only in the sub-salt basement of Mt Kitty-1 well.

640 The production index (PI) is high in the Upper Gillen samples with a maximum value  
641 of 0.77, while it is relatively lower ( $< 0.4$ ) in the Lower Gillen and Inindia beds samples  
642 (Table 1). The generative potential (GP) is less than 1 except in one sample of the  
643 Upper Gillen. The hydrogen index (HI) values are very low in Lower Gillen and Inindia  
644 Beds shales (31-60 and 17-21 mg HC/g TOC, respectively) and are higher in the Upper  
645 Gillen with values ranging between 90 and 159 mg HC/g TOC. Similarly, the oxygen  
646 index (OI) values are very low in all the studied samples with values less than 50 and 5  
647 mg CO<sub>2</sub>/g TOC in the Lower Gillen and Inindia Beds, respectively, whilst elevated OI  
648 values are only observed in the Upper Gillen samples (Table 1).

#### 649 **4.6. Hydrocarbon and non-hydrocarbon gas analysis**

650 The Magee-1 well gases consist mainly of N<sub>2</sub> and CH<sub>4</sub> comprising approximately  
651 75% of the whole gas composition. Other gas phases comprise He (6.20%) with  
652 subordinate contents of C<sub>2</sub> and C<sub>2+</sub> components (Table 1). The Magee-1 gas is  
653 characterized by higher IC<sub>5</sub>/nC<sub>5</sub> than IC<sub>4</sub>/nC<sub>4</sub>, and a dryness coefficient (DC) value of  
654 0.80. Mt Kitty-1 gases are characterized by elevated N<sub>2</sub> contents that range from 61% to  
655 approximately 100%. They contain variable contents of CH<sub>4</sub> (0.04-13.14%) with  
656 subordinate quantities of He (0-9%) and H<sub>2</sub> (0-11.4%). C<sub>2+</sub> phases are also significant in  
657 Mt Kitty-1 gas samples where the DC values are often less than 0.75. Notably, Mt Kitty-  
658 1 gases display much greater IC<sub>5</sub>/nC<sub>5</sub> relative to IC<sub>4</sub>/nC<sub>4</sub> (Table 1). In Murphy-1 gas  
659 samples, the hydrocarbon gases mostly constitute CH<sub>4</sub> with variable contents of C<sub>2+</sub>  
660 components. Murphy-1 gases also display a wide range of DC values between 0.79-  
661 0.90 and show relatively elevated IC<sub>4</sub>/nC<sub>4</sub> than IC<sub>5</sub>/nC<sub>5</sub> values.

662

663

664

665

666

667

668

669

670

#### 4.7. Burial model

671 The burial model of Magee-1 well highlights five major tectonic phases that have  
672 controlled the sediment distribution and architecture in the study region. These tectonic  
673 events comprise: the earliest Areyonga rifting phase that is followed by four successive  
674 younger compressional events. These are the Petemann, Bloodwood, Rodingan and  
675 Alice Springs oroginies. The heat flow corresponding to these tectonic phases are  
676 assigned as follows: 113, 75, 72, 70, and 65 mW/m<sup>2</sup>, respectively. These tectonic  
677 phases likely controlled the thermal maturity of the Lower Gillen organic-rich shales  
678 (Fig. 16). The pre-rift sediments were subjected to a moderate subsidence which  
679 created an accommodation space for the accumulation of the thick salt layers (Gillen  
680 salt) and the massive white dolomite (Loves Creek) above the Lower Gillen sediments.  
681 The syn-rift phase (Areyonga rifting) was characterized by a lithospheric stretching  
682 coupled with elevated levels of heat flux. Contemporaneous to the post-rift thermal  
683 subsidence (Cambrian-Ordovician), very thick sedimentary successions of post-rift  
684 clastic facies were deposited. The post-rift thermal subsidence has been ended with the  
685 onset of the successive compressive phases (Petemann, Bloodwood, Rodingan and  
686 Alice Springs oroginies) during which the subsidence rate has significantly decreased  
687 (Fig. 16). The top of oil window was reached at relatively shallow depths (~ 1350 m).  
688 Additionally, the expulsion onset and the gas window maturity were observed at greater  
689 depths (~ 2338 m) and higher temperatures (~ 117 °C, Fig. 17) compared to the oil  
690 window.

691

692

693

694

695

696

697

698

699

700

701

702

703

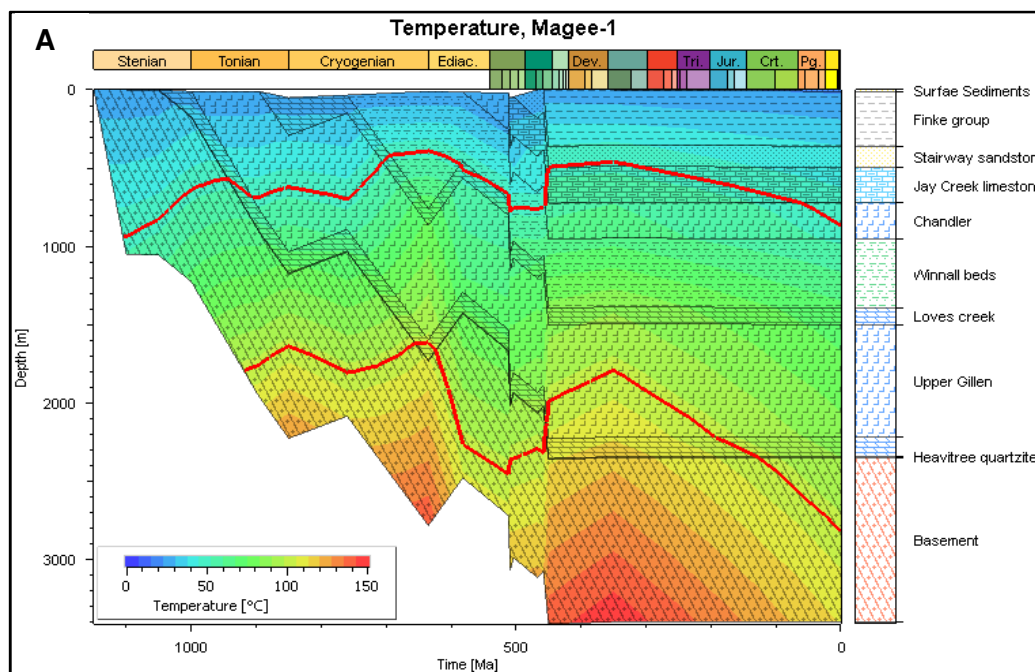
704

705

706

707

708



709

710

711

712

713

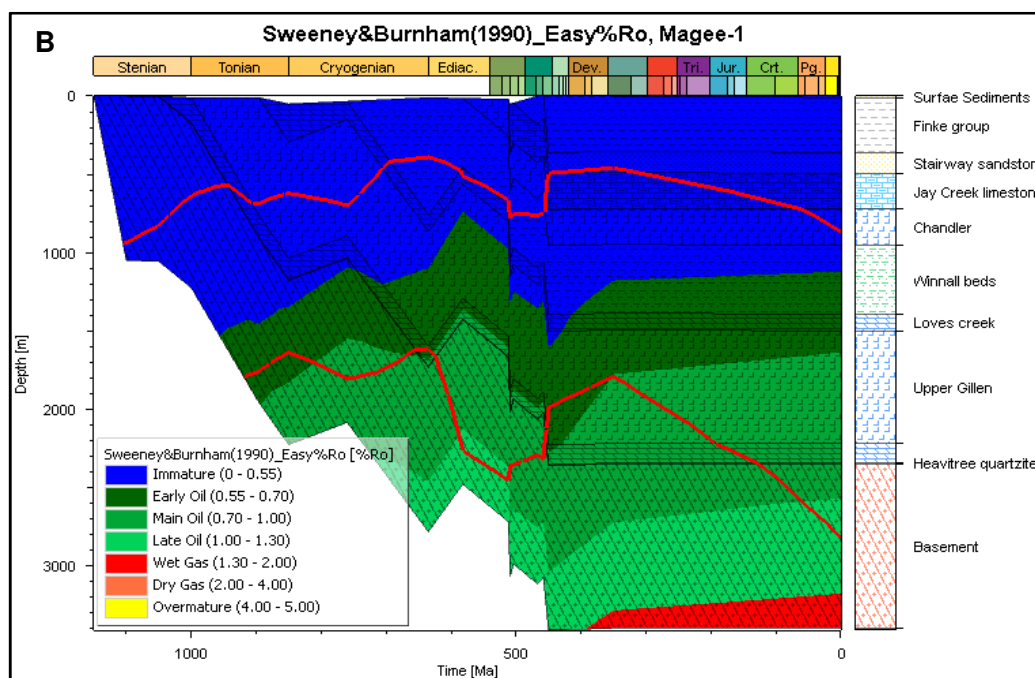
714

715

716

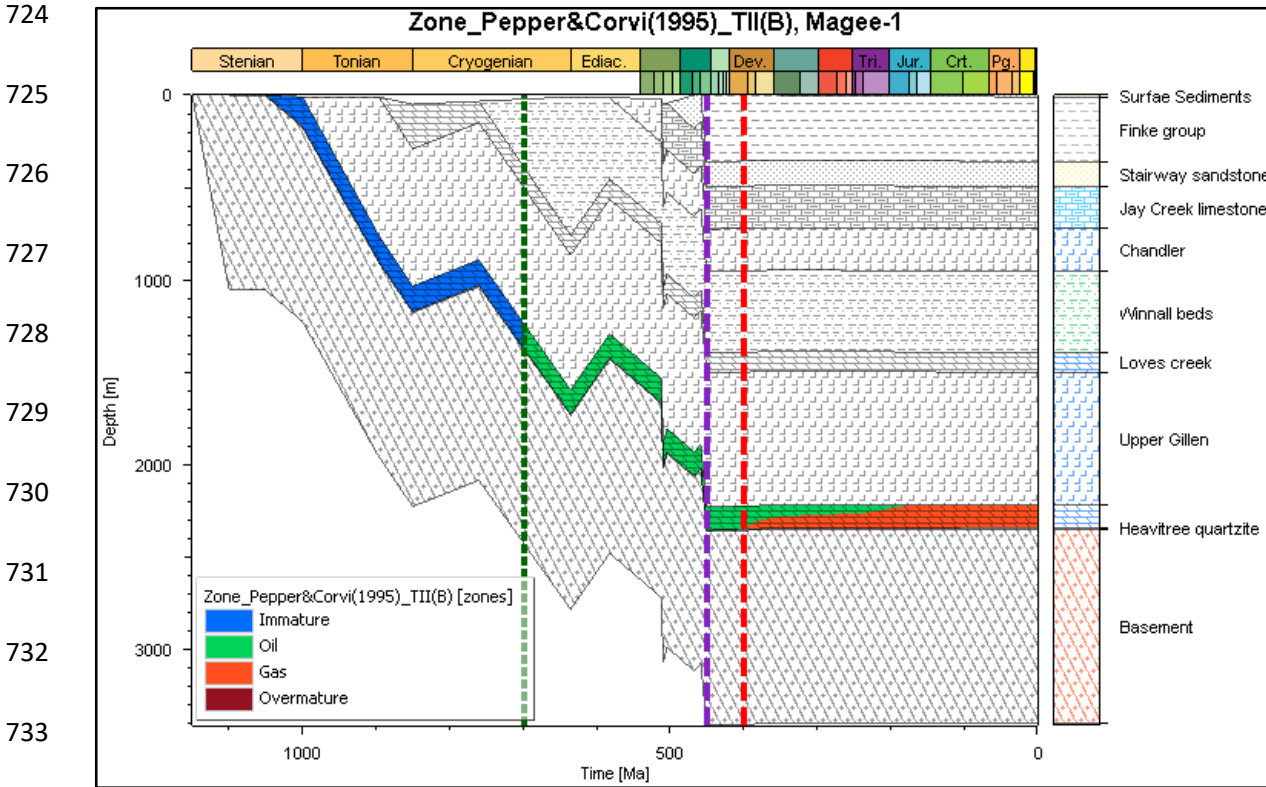
717

718



719 **Fig. 16.** Illustrative burial models for the sedimentary succession in Magee-1 well. The  
720 temperature overlay (A) is added to demonstrate the thermal history. The simulated  
721 vitrinite reflectance maturity overlay (B) was deduced using the EASY%\_Ro algorithm

722 (Sweeney and Burnham, 1990) as a function of time and space. The red lines  
 723 represents the temperature intervals for 50, and 100°C.



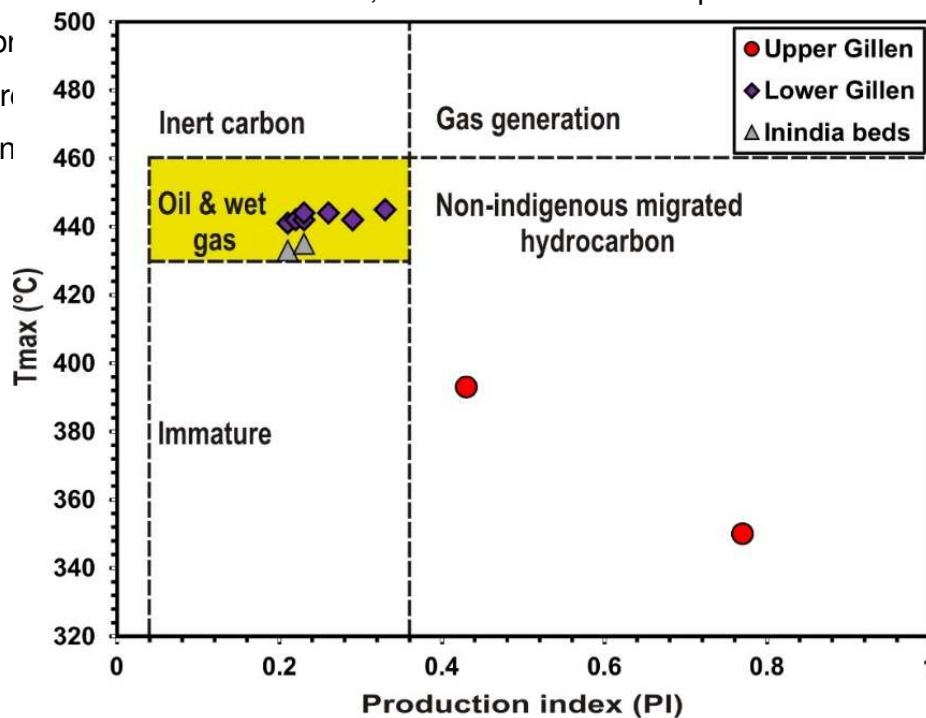
734 **Fig. 17.** The burial history chart of Magee-1 well with the hydrocarbon generation zone  
 735 overlay according to TII-B kinetic model of [Pepper and Corvi \(1995\)](#) through geologic  
 736 time. The vertical colored dashed lines represent the onset of oil generation at ~ 700 Ma  
 737 (green dashed line), the vertical purple dashed line represents the expulsion onset at  
 738 448 Ma and the red dashed line points to the gas generation window at 400 Ma.

739 **5. Discussion**

740 **5.1. Hydrocarbon gas origin, generation and expulsion**

741 The TOC and Rock-Eval pyrolysis results are frequently used to investigate and  
 742 assess the capability of the organic-rich sediments to generate thermogenic  
 743 hydrocarbons ([Peters, 1986](#); [Peters and Cassa, 1994](#), [Leila and Moscariello, 2017](#)).  
 744 Based on the measured TOC values, Lower and Upper Gillen shales contain fair  
 745 contents of organic matter, whereas the samples of Inindia Beds are richer in the  
 746 organic content. However, the Rock-Eval S<sub>2</sub> values are below 0.8 that is often

747 considered as a lower limit to interpret the values in terms of generation potentiality,  
 748 especially the  $T_{max}$  as valid for maturity assessment. The measured  $S_2$  values in studied  
 749 shales reflect their current poor potentiality to generate hydrocarbons upon thermal  
 750 cracking of the kerogen. The poor  $S_2$  values in all the shale samples are consistent with  
 751 their depleted contents of pyrolyzed carbon (Pc), however, most of these samples may  
 752 have had much higher initial TOC contents than that currently preserved. Based on  
 753 measured present-day TOC, the initial TOC values were deduced following theoretical  
 754 calculations proposed by Peters et al. (2005) (Table 1). The maximum initial TOC and  
 755 HI values were theoretically deduced as 1.3 % and 500 mg HC/g TOC, respectively.  
 756 Accordingly, potential source rock intervals that could have generated hydrocarbon  
 757 gases are hosted in the Lower Gillen and Inindia Beds (Fig. 18). However, more data  
 758 are required to better define the source rock in the deepest part in the drainage area of  
 759 the drilled structures. In addition, no work has been published for such an old,



and in term  
 depend on  
 the maturity

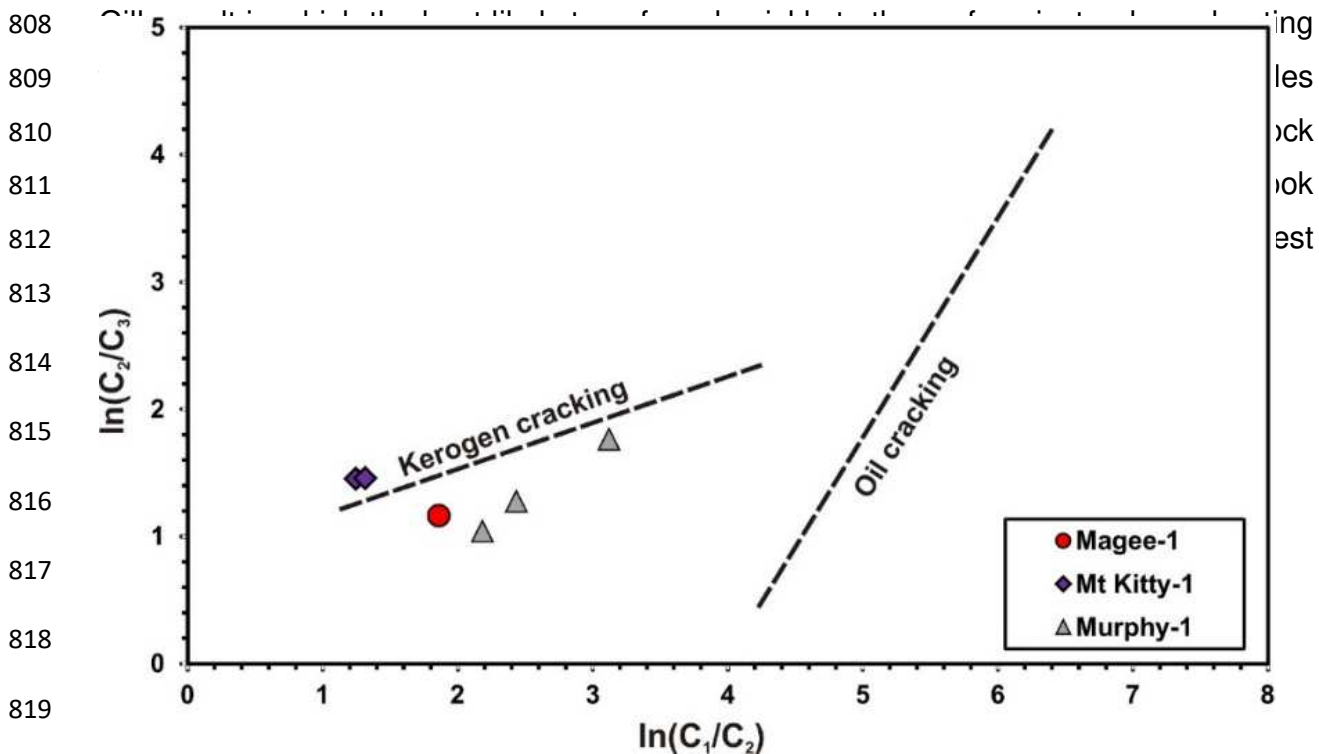
760 Precambrian  
 761 of interpretation  
 762 equivalent  
 763 level.

764  
 765  
 766  
 767  
 768  
 769  
 770  
 771  
 772  
 773

774 **Fig. 18.** Production index (PI) versus Rock-Eval  $T_{\max}$  plot illustrating the potentiality of  
775 some intervals in Lower Gillen and Inindia Beds to generate hydrocarbon gases. The  
776 investigated shales of Lower Gillen Formation likely generated most of the hydrocarbon  
777 gases in the basin.

778 Dominance of  $C_2$  and  $C_{2+}$  gas phases in all the investigated HC gases points to a  
779 prominent contribution from thermogenic sources (e.g. [Tissot and Welte, 1978](#); [Hunt,](#)  
780 [1996](#)). Furthermore, the calculated dryness coefficient (DC) values for the HC gases are  
781 less than 0.95 which is much lower than those reported for pure biogenic gases ( $\sim 0.99$ ;  
782 [Radwan et al., 2022](#)), thus confirming a dominant thermogenic origin for the studied HC  
783 gases ([Table 2](#)). Genesis of hydrocarbon gases via thermogenic cracking of organic-  
784 rich kerogen is often associated with generation of wet gas components ( $DC < 0.9$ ),  
785 however, mixing with biogenic gas phases would increase the DC values to more than  
786 0.9 ([Schoell, 1983](#); [Whiticar, 1999](#)). Accordingly, mixed biogenic and thermogenic  
787 hydrocarbon gases are common in Murphy-1 well at different depths, whereas gases  
788 collected from the other wells only contain thermogenic hydrocarbon gases. Occurrence  
789 of some alkanes and their relative proportion such as  $iC_4/nC_4$  and  $iC_5/nC_5$  often infers  
790 on the conditions of HC gas generation from mature kerogens ([Wang et al., 2017](#), [Qin](#)  
791 [et al., 2018](#)). With increasing thermal maturity, the concentration of the less stable  
792 isobutane decreases relative to the more stable n-butane ([Qin et al., 2018](#)). The good  
793 correlation between the dryness coefficient and  $iC_4/nC_4$  ( $R^2= 0.75$ ) suggests generation  
794 of studied HC gases via primary cracking of an organic-rich kerogen at the gas window.  
795 The relationship between natural logarithms of  $C_1/C_2$  and  $C_2/C_3$  ratios commonly reflects  
796 the mechanism of HC gas generation (primary versus secondary cracking).  
797 Hydrocarbon gases generated via secondary cracking of liquid HC phases often display  
798 a progressive increase in both  $C_1/C_2$  and  $C_2/C_3$  ratios ([Tao et al., 2014](#)). The studied  
799 hydrocarbon gases show a slight increase in the natural logarithms of  $C_1/C_2$  and  $C_2/C_3$   
800 ratios typifying their generation from primary cracking of a mature kerogen ([Fig. 19](#)).  
801 This is consistent with our modelling results and the hypothesis of thermogenic cracking  
802 of the Lower Gillen organic-rich intervals that have entered the gas window during the  
803 earliest Mesozoic.

804 Burial modeling results point out that the Lower Gillen had reached the gas window  
 805 about 200 Ma with a simulated present-day temperature of 86°C (Fig. 16). The initial  
 806 rifting phase (850 Ma to 760 Ma) was characterized by high heat flow value; however, it  
 807 did not contribute to the HC generation due to the high thermal conductivity of the Upper



825 **Fig. 19.** Plot of  $\ln C_1$ ,  $C_2$  and  $C_3$  values typifying generation of the studied HC gas  
 826 phases via primary cracking of the kerogen (Tao et al., 2014).

827 The transition from liquid to gaseous hydrocarbon generation was gradual with a  
 828 generation phase of wet gases within the oil window. The 1D modeling results  
 829 elucidated that the gas window occurred at a burial depth of 2340m in Magee-1 well.  
 830 The first two tectonic events have induced a minimal impact on the hydrocarbon



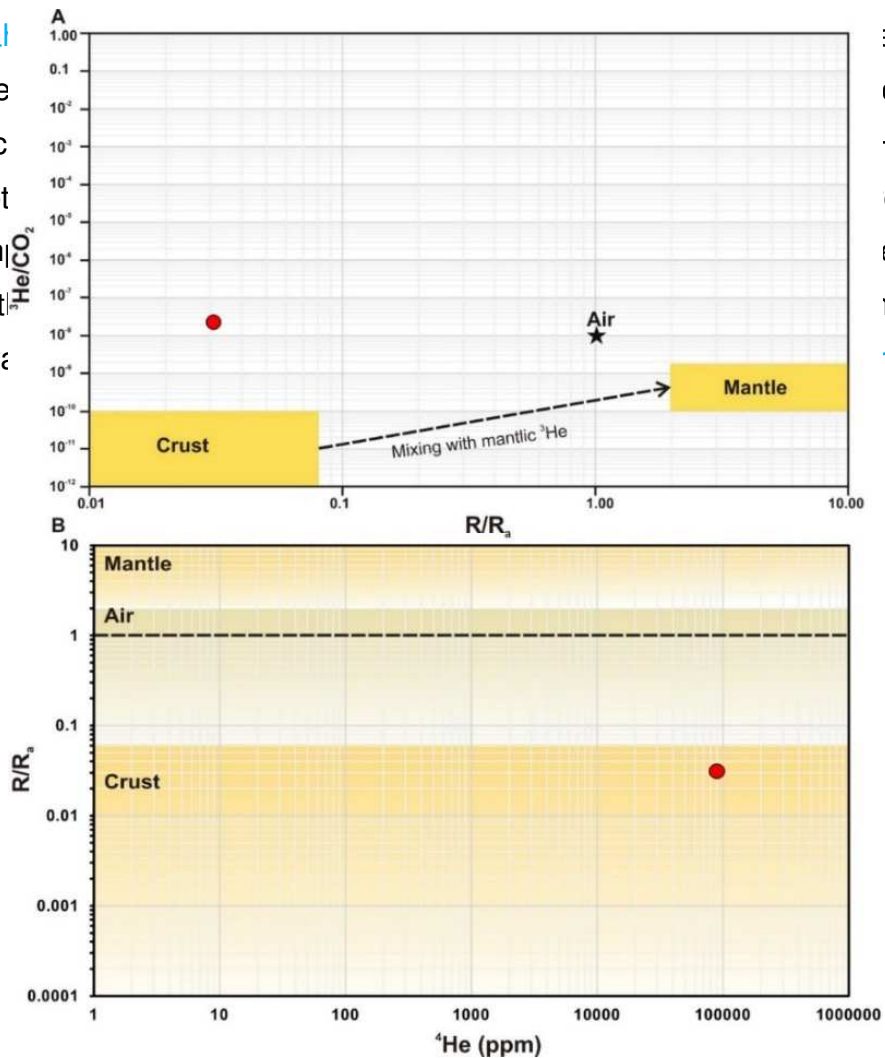
831 generation due to the cooling effect of the overlying Gillen salt layers. Despite the fact  
832 that the burial temperature did not exceed 110°C and was even constant since the Late  
833 Paleozoic, the prolonged exposure to this temperature resulted in a slow maturation of  
834 the organic matter and kerogen transformation into hydrocarbons (e.g. [Loptain, 1971](#);  
835 [Waples, 1980](#); [Wood, 2018](#)). We hypothesize that the kerogen type of the initial organic  
836 matter, the interplay between burial temperature, exposure time as well as the cooling  
837 effect induced by the Upper Gillen salt control the type and composition of the  
838 generated hydrocarbons.

## 839 **5.2. Origin of H<sub>2</sub> and He**

840 Helium (He) occurrence in the study area with a maximum content of 9% in Mt Kitty-  
841 1 well presents one of the highest recorded subsurface concentrations ever recorded on  
842 Earth ([McInnes et al., 2017](#), [Boreham et al., 2018](#)). Furthermore, the He-rich gases are  
843 mixed with a significant concentration of H<sub>2</sub> (~ 11%) as well as wet hydrocarbon gases  
844 (C<sub>1</sub>/C<sub>2</sub> ~ 3). This unusual gas blend reveals a complex interplay between various gas  
845 generation processes in the subsurface. The occurrence of He in the subsurface is  
846 often associated with either mantellic or crustal gases as well as the dissolution of  
847 atmospheric gases in the groundwater ([Poreda et al., 1986](#); [Ni et al., 2014](#); [Moore et al.,](#)  
848 [2017](#)). Except mantellic gases, all the other sources are associated with an enrichment  
849 of <sup>4</sup>He relative to <sup>3</sup>He. Therefore, a useful tool to discriminate the origin of He-rich gases  
850 is the relative proportion of He isotopic composition ( $R = \frac{^3\text{He}}{^4\text{He}}$ ) relative to that in the  
851 air ( $R_a = 1.4 \times 10^{-6}$ ). For example, the He derived from the upper mantle displays  
852 average  $R/R_a$  values between 2 and 8, whereas those from the lower mantle show  
853 elevated  $R/R_a$  values up to 40 ([Ni et al., 2014](#); [Moore et al., 2017](#)). Helium derived from  
854 the Mid-Ocean Ridge Basalt (MORB) displays high isotope ratios with an average  
855 ( $R/R_a$ ) value of about 8. On the other hand, atmospheric He dissolved in the  
856 groundwater has an average ( $R/R_a$ ) value of approximately 1, whereas the crustal-  
857 derived He is largely depleted in <sup>3</sup>He and therefore has an average value of 0.02 or less  
858 ([Ballentine et al., 1991](#)).

859 The relative abundance of some elements such as uranium and thorium in the rock  
860 forming minerals determines the extent of He generation from the crustal rocks and

861 hence controls its average  $R/R_a$  values (e.g. Poreda et al. 1986; Darrah et al. 2015;  
 862 Moore et al. 2017). Pure crustal-derived He commonly shows  $R/R_a$  values greater than  
 863 0.05 and mixing with mantle-derived He would increase the values up to 0.5 (Ni et al.,  
 864 2014; Darrah et al. 2018).  
 865 1 well revealed by the  $^3\text{He}/^4\text{He}$  ratio of Mt-Kitty, possibly with  
 866 atmospheric He, reflecting the depletion of  $^3\text{He}$  in the atmosphere  
 867 relative to other sources. The  $^3\text{He}/^4\text{He}$  ratio of Mt-Kitty is within the  
 868 dominant in the range of 0.01 to 0.1 (Fig. 20A) reflecting the  
 869 coupled with the formation of the mantle-derived He (Fig. 20A) within the  
 870 formation with the mantle-derived He (Fig. 20A) within the  
 871 2018).



872  
 873  
 874  
 875  
 876  
 877  
 878  
 879  
 880  
 881  
 882  
 883  
 884  
 885  
 886

887 **Fig. 20.** (A) Plot of R/R<sub>a</sub> ratio versus <sup>3</sup>He/CO<sub>2</sub>, (B) <sup>4</sup>He versus R/R<sub>a</sub> plot illustrating the  
888 crustal origin of helium in the studied gases. Data are from [McInnes et al. \(2017\)](#).

889 The log-log plot of R/R<sub>a</sub> versus <sup>4</sup>He concentration ([Fig. 20B](#)) confirms the dominant  
890 crustal origin of He in Mt Kitty-1 well. Thus, we hypothesize that He gas accumulation in  
891 Mt Kitty-1 well is mostly derived from the sub-salt granitic basement rocks. The spectral  
892 gamma ray logs typified that these rocks are rich in uranium and thorium ([Fig. 15](#)), thus  
893 they are capable of generating helium via radiogenic reactions and radioactive decay  
894 these elements. This crustal derivation of He is often coupled with a concomitant  
895 generation of H<sub>2</sub> via a radiolytic dissociation of water (e.g. [Lin et al., 2005](#); [Warr et al.,](#)  
896 [2019](#)). Alternatively, H<sub>2</sub> may also be generated by oxidation of iron rich materials but we  
897 do not have data on the deepest layer mineralogy to discuss this hypothesis. Another  
898 plausible mechanism for H<sub>2</sub> generation is the oxidation of hydrocarbons, however the  
899 negative correlation between H<sub>2</sub> and hydrocarbon phases such as CH<sub>4</sub> (R<sup>2</sup>= -0.75) as  
900 well as the complete absence of H<sub>2</sub> in the above-salt layers of Murphy-1 well may rule  
901 out a significant H<sub>2</sub> input via oxidation of HC gases.

### 902 **5.3. Gas accumulation versus gas seepage**

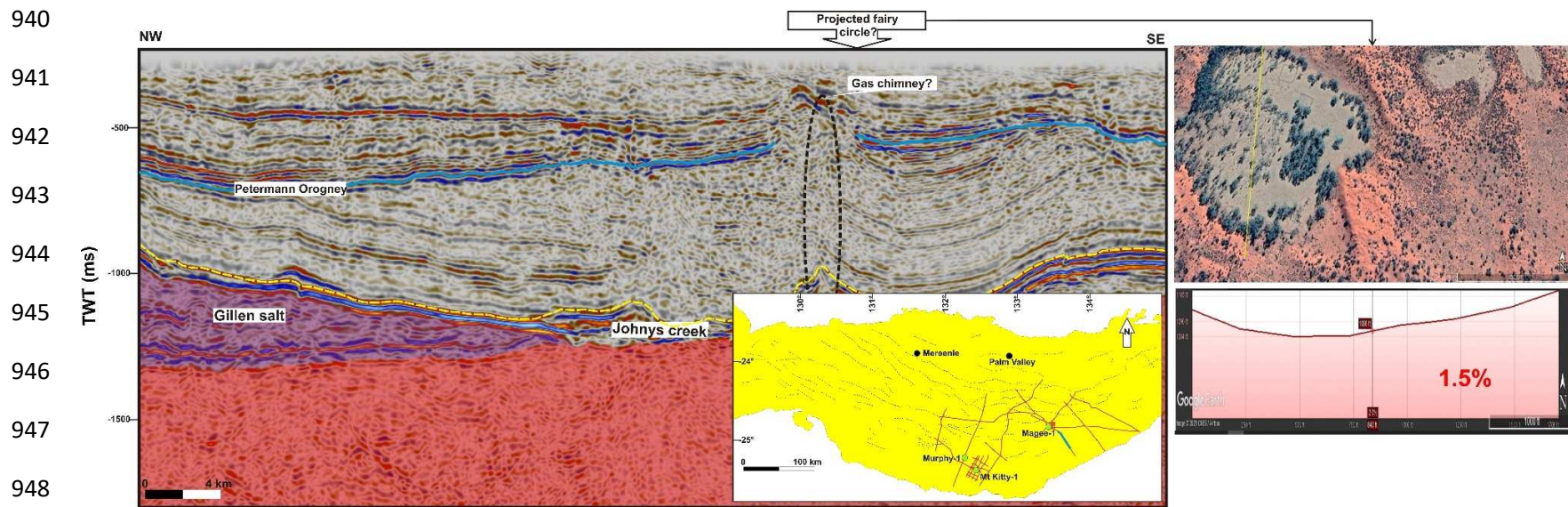
903 In the study region, the studied gases are reported from the sub-salt basement and  
904 Heavitree Formation. Fractures in the granitic basement present the main gas-bearing  
905 zones in Mt Kitty-1 well. Fractures are ubiquitous in the Proterozoic basement rocks  
906 paving the base of the Amadeus Basin. The fracture network is commonly controlled by  
907 the faults and the high porosities and permeabilities are often associated with the  
908 tectonically induced fractures ([Heap and Kennedy, 2016](#)). Mt-Kitty-1 well is located in a  
909 tectonically-high region due to the tectonic uplift of the faulted basement associated with  
910 the fracture patterns and deep-seated faults which are commonly infilled with gases  
911 (high resistivity zones) that are sealed beneath the thick Gillen salt layers ([Fig. 6](#)).  
912 Therefore, gas accumulation in Mt Kitty-1 area are a classical structural trap sealed by  
913 the thick salt layers.

914 The porous sandstones of Heavitree Formation present a potential gas pay zone in  
915 Magee-1 well. Heavitree Formation has minimal contents of matrix and pore-filling  
916 cements and therefore preserves a very good intergranular porosity. The intergranular  
917 pores are well-connected and hence display a good effective porosity likely infilled with

918 gases (Fig. 11). But the Heavitree Formation is very thin in Magee-1 well (< 6 m), and  
919 thus hindering any economic accumulation of gases. Moreover, economic gas  
920 accumulations should occur in the western and southwestern parts of the basin where  
921 the Heavitree Formation is believed to be thicker (Fig. 5B). Additionally, the thickness of  
922 the Gillen salt layers also increases westward and southwestward and therefore will  
923 provide effective seal for the trapped gasses in the Heavitree pay zone (Fig. 8B).

924 The above-salt reservoir intervals are scarce and the wireline log results reveal  
925 occurrence of thick non-porous and impermeable intervals. However, thin porous and  
926 permeable sandstone zones were observed in the Inindia Beds where the effective  
927 porosity reaches up to 20%. These zones likely present potential reservoirs for the  
928 gaseous hydrocarbons that were generated from the organic-rich zones in Inindia Beds,  
929 however its low thickness may hinder any economic gas accumulations in this region.

930 Gas seepage zones such as SCD are correlated with distribution of the thin salt  
931 layers. Observed discrepancies in the seismic reflections (gas chimneys) often coincide  
932 with the regions where the thickness of the salt layers is minimal (Fig. 7). Thick  
933 accumulation of salt layers commonly provides effective seal for fluids preventing their  
934 leakage and hence allows the economic gas accumulation and entrapment in the  
935 reservoir rocks (Grunau, 1987; Schlomer and Krooss, 1997). Furthermore, for gases  
936 with a relatively small atomic radius such as He and H<sub>2</sub>, thick impervious salt layers are  
937 essential for to their entrapment in the reservoir. On the other hand, a clear link is  
938 observed between the surficial SCD distribution and the subsurface gas chimneys (Fig.  
939 22), and both are correlated with regions where thin layers of Gillen salt exist.



949 **Fig. 21.** Interpreted 2D seismic profile illustrating possible occurrence of gas chimneys where the Gillen Salt layers are  
 950 thinning. The subsurface gas chimneys are well-correlated with SCD typifying continuous gas emanations.

951

952 **6. Conclusions**

- 953 • In the Amadeus Basin, hydrocarbon (HC) and non-hydrocarbon gas blend was  
954 discovered in some exploratory wells. The gases range from HC-He-rich in  
955 Magee-1 to HC-He-H<sub>2</sub> blend in Mt Kitty-1 well.
- 956 • The HC gases are wet gas with a significant share (> 10%) of C<sub>2+</sub>  
957 components.
- 958 • Potential HC-source rock intervals likely occur in the Neoproterozoic Lower  
959 Gillen shales, however, the Rock-Eval data revealed a remnant inert carbon  
960 with minimal capability to generate HC gases.
- 961 • Modeling results shows that wet HC-gases may have been generated from the  
962 Lower Gillen shales that had an initial organic carbon of at least 1.3 wt% and  
963 entered the gas generation window during Late Devonian at the end of Alice  
964 Springs orogeny. The temperatures remain rather low but the 400 years  
965 allowed this gas generation.
- 966 • He-rich gases display a crustal signature with elevated <sup>3</sup>He contents and R/R<sub>a</sub>  
967 value of 0.031 typifying their generation via radiolytic decay of the thorium and  
968 uranium-rich basement potentially synchronous with H<sub>2</sub> generation from  
969 radiolysis. Other source of H<sub>2</sub> by water reduction cannot be precluded.

970 The Amadeus Basin case study allows to open new horizons since the gas is very  
971 likely a classical accumulation, in a structural trap sealed by a salt layer. The reservoir  
972 has been charged by these three gases at different rates but probably all over hundreds  
973 of millions of years since the He generation is a low process. The presence of this  
974 mixture shows that the salt may retain these light gases for millions of years, which  
975 opens new exploration perspectives.

976 **Acknowledgment**

977 The authors gratefully acknowledge Emanuelle Frery for the collaboration and the  
978 help to collect the Amadeus basin data set. Moreover, some datasets (e.g. 2D seismic)  
979 were kindly provided by Geoscience Exploration and Mining Information System  
980 (GEMIS) of the Northern Territory, Australia. We thank Dan Levy and Gabriel Pasquet,  
981 from E2S UPPA, for many interesting discussions. We appreciate the comments of

982 Branimir Segvic and another anonymous reviewer that significantly improved the  
983 manuscript. This work is a part of Keanu Loiseau master project that has been funded  
984 by Engie. A part of the work has also been funded by Campus France in the frame of  
985 Make our Planet Great Again “MOPGA” program (project number mopga-pdf-  
986 0000000089).

## 987 **References**

- 988 Aguilera, R.F., Aguilera, R., 2004. A triple-porosity model for petrophysical analysis of naturally  
989 fractured reservoirs”, *Petrophysics* 45, 157-166.
- 990 Allen, A.P., Allen, R.J., 1990. *Basin Analysis: Principles and Applications* (First ed.). 451 pp.  
991 Oxford, Blackwell Scientific Publications.
- 992 AMDEL CORE SERVICES, 1993. Petrology and core analysis report of Heavitree sandstone.  
993 Prepared for oil Pacific Oil & Gas Ltd., 11 pp, internal report.
- 994 Bagas, L., Grey, K., Williams, I.R., 1996. Reappraisal of the Paterson Orogen and Savory  
995 Basin, in *Geological Survey of Western Australia Annual Review*, 55-63.
- 996 Ballentine, C. J., O’Nions, R. K., Oxburgh, E. R., Horvath, F., Deak, J., 1991. Rare gas  
997 constraints on hydrocarbon accumulation, crustal degassing and groundwater flow in the  
998 Pannonian Basin. *Earth and Planetary Science Letters* 105, 229-246.
- 999 Bernard, B., Brooks, J., Sackett, W., 1978. Light hydrocarbons in recent Texas continental shelf  
1000 and slope sediments. *J. Geophys. Res.* 83, 4053-4061.
- 1001 Blamey, N., Brand, U., Parnell, J., Spear, N., Lécuyer, C., Benison, K., Meng, F., Ni, P., 2016.  
1002 Paradigm shift in determining Neoproterozoic atmospheric oxygen. *Geology* 44 (8), 651-  
1003 654.
- 1004 Blatt, H., Tracy, R. 1996. *Petrology: Igneous, Sedimentary, and Metamorphic*, Freeman 2<sup>nd</sup> ed.,  
1005 pp. 518.
- 1006 Boreham, C., Edwards, D., Poreda, R., Darrah, T., Zhu, R., Grosjean, E., Main, P., Waltenberg,  
1007 K., Henson, P., 2018. Helium in the Australian liquefied natural gas economy. *The*  
1008 *APPEA Journal* 58, 209-237.
- 1009 Carr, L.K., Korsch, R.J., Palu, T.J., Reese, B., 2016. Onshore basin inventory: the McArthur,  
1010 South Nicholson, Georgina, Wiso, Amadeus, Warburton, Cooper and Galilee Basins,  
1011 central Australia, *Geoscience Australia Record* 2016/04.
- 1012 Charlou, J.L., Donval, J.P., Fouquet, Y., Jean-Baptiste, P., Holm, N., 2002. Geochemistry of  
1013 high H<sub>2</sub> and CH<sub>4</sub> vent fluids issuing from ultramafic rocks at the Rainbow hydrothermal  
1014 field (36°14' N, MAR). *Chem. Geol.* 191, 345-359.

1015 Combaudon, V., I Moretti, B. Kleine, A Stefánsson. 2022. Natural hydrogen emissions in Iceland  
1016 and comparison with the Mid-Atlantic Ridge. *Int. J. Hydrog. Energy* 47, 10217-10227.

1017 Coveney, R., Goebel, E., Zeller, E., Dreschhoff, G.A.M., Angino, E.E., 1987. Serpentinization  
1018 and origin of hydrogen gas in Kansas. *AAPG Bull.* 71, 39-48.

1019 Darrah, T. H., Jackson, R. B., Vengosh, A., Warner, N. R., Whyte, C. J., Walsh, T. B., Kondash,  
1020 A. J., Poreda, R. J., 2015. The evolution of Devonian hydrocarbon gases in shallow  
1021 aquifers of the northern Appalachian Basin: insights from integrating noble gas and  
1022 hydrocarbon geochemistry. *Geochimica et Cosmochimica Acta* 170, 321-355.

1023 Edgoose, C.J., 2013. Amadeus Basin, In Ahmad, M., Munson, T.J. (Eds.), *Geology and mineral  
1024 resources of the Northern Territory*, Northern Territory Geological Survey, Special  
1025 Publication 5, 23-70.

1026 Espitalié, J., Deroo, G., Marquis, F., 1986. Rock–Eval pyrolysis and its applications. Part III.  
1027 *Revue de l'Institut Français du Pétrole* 41, 73-89.

1028 Etiope, G., 2017. Abiotic methane in continental serpentinization sites: an overview. *Proc. Earth  
1029 Planet. Sci.* 17, 9-12.

1030 Frery, E., Langhi, L., Maison, M., Moretti, I., 2021. Natural hydrogen seeps identified in the  
1031 North Perth Basin, Western Australia. *Int. J. Hydrog. Energy* 46, 31158-31173.

1032 Geotech, 1992. Source rock characterization study Magee-1. Prepared for oil Pacific Oil & Gas  
1033 Ltd., 11 pp, internal report.

1034 Geymond, U., Ramanaidou, E., Lévy, D., Ouaya, A., Moretti, I., 2022. Can weathering of  
1035 Banded Iron Formations generate natural hydrogen? Evidences from Australia, Brazil  
1036 and South Africa. *Minerals*, 12, 163. <https://doi.org/10.3390/xxxxx>

1037 Gielen, D., Boshell, F., Saygin, D., Bazilian, M., Wagner, N., Gorini, R., 2019. The role of  
1038 renewable energy in the global energy transformation. *Ener. Strat. Rev.* 24, 38-50.

1039 Grunau, H.R., 1987. A worldwide look at the caprock problem. *J. Petrol. Geol.* 10, 245-266.

1040 Haines, P.W., Allen, H-J. A., 2019. Hydrocarbon and helium prospectivity of the Amadeus and  
1041 Murraba basins in Western Australia, in Keep, M., Moss, S.J. (Eds.), *The Sedimentary  
1042 Basins of Western Australia V: Proceedings of the Petroleum Exploration Society of  
1043 Australia Symposium*, Perth, WA, 2019, 17 pp.

1044 Haines, P.W., Wingate, M.T.D., Kirkland, C.L., Allen, H.J., 2015, Detrital zircon geochronology  
1045 of upper Ediacaran to lower Cambrian deposits (Supersequence 4), western Amadeus  
1046 Basin: Testing revised stratigraphic correlations, *Geological Survey of Western Australia  
1047 Record* 2015/8.



- 1048 Heap, M., Kennedy, B., 2016. Exploring the scale-dependent permeability of fractured andesite.  
1049 Ear. Plan. Sci. Lett. 447, 139-150.
- 1050 Herron, D.A., 2011. First Steps in Seismic Interpretation. Society of Exploration Geophysicists.
- 1051 Hissey, E., Clifford, J., Debacker, T., 2016, Exploring the sub-salt play in the frontier Amadeus  
1052 Basin –insights from regional 2D seismic and potential field data, in 25<sup>th</sup> International  
1053 Geophysical Conference and Exhibition, Adelaide, South Australia, Australian Society of  
1054 Exploration Geophysicists, Extended Abstracts, 111-118.
- 1055 Holloway, J., O'Day, P., 2000. Production of CO<sub>2</sub> and H<sub>2</sub> by dike-eruptive events at Mid-  
1056 Ocean Ridges: Implications for abiotic organic synthesis and global geochemical cycling.  
1057 Int. Geol. Rev. 42, 673-683.
- 1058 Hunt, J.M., 1996. Petroleum Geochemistry and Geology, second ed. W.H. Freeman and  
1059 Company, New York, p. 743.
- 1060 Jackson, K.S., McKirdy, D.M., Deckelman, J.A., 1984. Hydrocarbon generation in the Amadeus  
1061 Basin, central Australia. The APEA Journal 24(1), 42-65.
- 1062 Johns, D. R., Menpes, S. A., Walshe, P., Bache, F., 2017. Exploration of a sub-salt play in the  
1063 southern Amadeus Basin, Central Australia – searching for big gas in Proterozoic  
1064 reservoirs. Proceedings of the 2017 South East Asia Petroleum Exploration Society  
1065 (SEAPEX) Conference, 2017, Abstract: [http://archives.datapages.com/data/southeast-](http://archives.datapages.com/data/southeast-asia-petroleumexploration-society/data/028/028001/3a_seapex0280003a.htm)  
1066 [asia-petroleumexploration-society/data/028/028001/3a\\_seapex0280003a.htm](http://archives.datapages.com/data/southeast-asia-petroleumexploration-society/data/028/028001/3a_seapex0280003a.htm). Available  
1067 at [https://centralpetroleum.com.au/wp-content/uploads/2016/10/ Santos-Seapex-](https://centralpetroleum.com.au/wp-content/uploads/2016/10/Santos-Seapex-Presentation-AMADEUS-Johns-et-al-2017.pdf)  
1068 [Presentation-AMADEUS-Johns-et-al-2017.pdf](https://centralpetroleum.com.au/wp-content/uploads/2016/10/Santos-Seapex-Presentation-AMADEUS-Johns-et-al-2017.pdf).
- 1069 Klein F, Tarnas J, Bach W. Abiotic sources of molecular hydrogen on earth. Elements  
1070 2020;16:19e24. [https://doi.org/ 10.2138/gselements.16.1.19](https://doi.org/10.2138/gselements.16.1.19).
- 1071 Larin, N., Zgonnik, V., Rodina, S., Deville, E., Prinzhofer, A., Larin, V.N., 2015. Natural  
1072 molecular hydrogen seepage associated with surficial, rounded depressions on the  
1073 European craton in Russia. Nat. Resour. Res. 24(3), 369-383.
- 1074 Leila, M., Levy, D., Battani, A., Piccardi, L., Segvic, B., Badurina, L., Pasquet, G., Combaudon,  
1075 V., Moretti, I., 2021. Origin of continuous hydrogen flux in gas manifestations at the  
1076 Larderello geothermal field, Central Italy. Chem. Geol. 585, 120564.
- 1077 Leila, M., Moscariello, A., 2017. Organic geochemistry of oil and natural gas in the West Dikirnis  
1078 and El-Tamad fields onshore Nile Delta, Egypt: Interpretation of potential source rocks.  
1079 J. Petrol. Geol., 40, 37-58.

- 1080 Lin, L., Slater, G.F., Lollar, B.S., Lacrampe-Couloume, G., Onstott, T.C., 2005. The yield and  
1081 isotopic composition of radiolytic H<sub>2</sub>, a potential energy source for the deep subsurface  
1082 biosphere. *Geochim. Cosmochim. Acta* 69 (4), 893-903.
- 1083 Lopatin, N. V., 1971. Temperature and geologic time as factors in coalification (in Russian).  
1084 *Akademiya Nauk SSSR Izvestiya, Seriya Geol.* 3, 95-106.
- 1085 Marshall, T.R., 2003. Petroleum systems and source rocks in the Amadeus Basin, Northern  
1086 Territory, In Cowie, M., (Ed.), Queensland 2003 exploration and development. Petroleum  
1087 Exploration Society of Australia (Qld/NT) 2003 Symposium, 37-42.
- 1088 Marshall, T.R., 2004, A review of source rocks in the Amadeus Basin. Northern Territory  
1089 Geological Survey, Record 2004-008.
- 1090 Marshall, T.R., Dyson, I.A., Liu, K., 2007, Petroleum systems in the Amadeus Basin, central  
1091 Australia: Were they all oil prone?, In Munson, T.J., Ambrose, G.J. (Eds.), Central  
1092 Australian Basins Symposium, Northern Territory Geological Survey Special Publication  
1093 2, 136-146.
- 1094 McCollom, T.M., Bach, W., 2009. Thermodynamic constraints on hydrogen generation during  
1095 serpentinization of ultramafic rocks. *Geochim. Cosmochim. Acta* 73, 856-875.
- 1096 McInnes, B., Darrah, T., Hilton, R., Boreham, C., Edwards, D., Gusterhuber, J., 2017.  
1097 Investigations into the highest reported He concentration in a natural gas sample: Mt  
1098 Kitty, Amadeus Basin, Northern Territory. In: Abstract at Goldschmidt 2017 Conference,  
1099 Paris, France, August 13-18, 2017. <https://goldschmidt.info/2017/abstracts/abstractView?id=2017003781>.
- 1100
- 1101 Moore, M. T., Vinson, D. S., Whyte, C. J., Eymold, W. K., Walsh, T. B., Darrah, T. H., 2017.  
1102 Differentiating between biogenic and thermogenic sources of natural gas in coalbed  
1103 methane reservoirs from the Illinois Basin using noble gas and hydrocarbon  
1104 geochemistry. In Lawson, M., Formolo, M., J.M. Eiler, J.M., (Eds.), *From Source to Seep: Geochemical Applications in Hydrocarbon Systems*. Geological Society, London,  
1105 Special Publications, 468.
- 1106
- 1107 Moretti, I., Brouilly, E., Loiseau, K., Prinzhofer, A., Deville, E., 2021b. Hydrogen emanations in  
1108 intracratonic areas: New guide lines for early exploration basin screening. *Geosci.*, 11  
1109 (3), 145.
- 1110 Moretti, I., Prinzhofer, A., Francolin, J., Pacheco, C., Rosanne, M., Rupin, F., Mertens, J.,  
1111 2021a. Long-term monitoring of natural hydrogen superficial emissions in a brazilian  
1112 cratonic environment. Sporadic large pulses versus daily periodic emissions. *Int. J.*  
1113 *Hydrog. Energy* 46, 3615-3628.

- 1114 Ni, Y., Dai, J., Tao, S., Wu, X., Liao, F., Wu, W., Zhang, D., 2014. Helium signatures of gases  
1115 from the Sichuan Basin, China. *Org. Geochem.* 74, 33-43.
- 1116 Orlando, A., Maria Conte, A., Borrini, D., Perinelli, C., 2010. Experimental investigation of CO<sub>2</sub>-  
1117 rich fluids production in a geothermal area: the Mt Amiata (Tuscany, Italy) case study.  
1118 *Chem. Geol.* 274, 177-186.
- 1119 Pepper, S.A., Corvi, J.P., 1995. Simple kinetic models of petroleum formation. Part I: oil and gas  
1120 generation from kerogen. *Mar. Petrol. Geol.* 12, 291-319.
- 1121 Peters, K.E., Cassa, M.R., 1994. Applied source rock geochemistry. In Magoon, L.B., and Dow,  
1122 W.G. (Eds.), *The petroleum system from source to trap* (93-117), AAPG, Tulsa,  
1123 Oklahoma.
- 1124 Peters, E.K., Walters, C.C., Moldowan, M.J., 2005. *The Biomarker guide 2nd Edition. Volume I:  
1125 Biomarkers and Isotopes in the Environment and Human History*. 471 pp. Cambridge,  
1126 Cambridge University Press.
- 1127 Peters, K.E., 1986. Guidelines for evaluating petroleum source rock using programmed  
1128 pyrolysis. *AAPG Bull.* 70, 318-329.
- 1129 Powell, C., Preiss, W.V., Gatehouse, C.G., Krapez, B., Li, Z.X., 1994, South Australian record of  
1130 a Rodinian epicontinental basin and its mid-Neoproterozoic breakup (~700 Ma) to form  
1131 the Palaeo-Pacific Ocean, *Tectonophysics* 237, 113–140. [https://doi.org/10.1016/0040-  
1132 1951\(94\)90250-X](https://doi.org/10.1016/0040-1951(94)90250-X).
- 1133 Poreda, R. J., Jenden, P. D., Kalpan, I. R., Craig, H., 1986. Mantle helium in Sacramento Basin  
1134 natural gas wells. *Geochimica et Cosmochimica Acta* 50, 2847-2853.
- 1135 Prinzhofer, A., Moretti, I., Françolin, J., Pacheco, C., D'Agostino, A., Werly, J., Rupin, F., 2019.  
1136 Natural hydrogen continuous emission from sedimentary basins: the example of a  
1137 Brazilian H<sub>2</sub>-emitting structure. *Int. J. Hydrog. Energy* 44, 5676-5685.
- 1138 Prinzhofer, A., Tahara Cissé, C.S., Diallo, A.B., 2018. Discovery of a large accumulation of  
1139 natural hydrogen in Bourakebougou (Mali). *Int. J. Hydrog. Energy* 43, 19315-19326.
- 1140 Qin, S., Zhang, Y., Zhao, C., Zheng Zhou, Z., 2018. Geochemical evidence for in situ  
1141 accumulation of tight gas in the Xujiahe Formation coal measures in the central Sichuan  
1142 Basin, China. *Int. J. Coal Geol.* 196, 173-184.
- 1143 Qin, S., Chunhu, H., Benjian, Z., Yu, Y., Miao, Y., 2019. Relationships of the iC<sub>4</sub>/nC<sub>4</sub> and  
1144 iC<sub>5</sub>/nC<sub>5</sub> ratios with maturity of coal-derived gases of Triassic Xujiahe Formation in  
1145 central Sichuan Basin, SW China. *Petrol. Expl. Develop.* 46(3), 496-503.

- 1146 Radwan, A., Nabawy, B., Shihata, M., Leila, M., 2022. Seismic interpretation, reservoir  
1147 characterization, gas origin and entrapment of the Miocene-Pliocene Mangaa C  
1148 sandstone, Karewa Gas Field, North Taranaki Basin, New Zealand
- 1149 Schlomer, S., Krooss, B., 1997. Experimental characterization of the hydrocarbon sealing  
1150 efficiency of cap rocks. *Mar. Petrol. Geol.* 14, 565-580.
- 1151 Schoell, M., 1983. Genetic characterization of natural gases. *AAPG Bull.* 67 (12), 2225-2238.
- 1152 Scrimgeour, I.R., 2019, Overview of mineral and petroleum exploration and production in 2018,  
1153 Annual Geoscience Exploration Seminar Proceedings, Northern Territory Geological  
1154 Survey, 1-14.
- 1155 Sherwood Lollar, B., Onstott, T.C., Lacrampe-Couloume, G., Ballentine, C.J., 2014. The  
1156 contribution of the Precambrian continental lithosphere to global H<sub>2</sub> production. *Nature*  
1157 516, 379-382.
- 1158 Skirrow, R. G., Huston, D. L., Mernagh, T. P., Thorne, J. P., Dulfer, H., Senior, A. B., 2013.  
1159 Critical commodities for a high-tech world: Australia's potential to supply global demand.  
1160 Geoscience Australia, Canberra. Available at [https://ecat.ga.gov.au/geonetwork/srv/eng/  
1161 search#!dfdb8e16-3c25-0990-e044-00144fdd4fa6](https://ecat.ga.gov.au/geonetwork/srv/eng/search#!dfdb8e16-3c25-0990-e044-00144fdd4fa6).
- 1162 Stack, S., Rabone, H., 2018. Cobar Management Pty Ltd: the CSA mine annual environmental  
1163 management report. AEMR 2017. [http://www.csamine.com.au/u  
1164 ploads/51687/ufiles/CSA-AEMR\\_2017\\_-\\_Final\\_opt.pdf](http://www.csamine.com.au/uploads/51687/ufiles/CSA-AEMR_2017_-_Final_opt.pdf).
- 1165 Stefánsson, A., Arnórsson, S., 2002. Gas pressures and redox reactions in geothermal fluids in  
1166 Iceland. *Chem. Geol.* 190, 251-271.
- 1167 Summons, R.E., Powell, T.G., 1991. Petroleum source rocks of the Amadeus Basin. Bureau of  
1168 Mineral Resources, Australia, Bull. 236, 511-523.
- 1169 Sweeney, J.J., Burnham, K.A., 1990. Evaluation of a simple model of vitrinite reflectance based  
1170 on chemical kinetics. *AAPG Bull.* 74(10), 1559-1570.
- 1171 Tao, S., Zou, C., Mi, J., Gao, X., Yang, C., Zhang, X., Fan, J., 2014. Geochemical comparison  
1172 between gas in fluid inclusions and gas produced from the Upper Triassic Xujiahe  
1173 Formation, Sichuan Basin, and SW China. *Org. Geochem.* 74, 59-65.
- 1174 Taran, Y.A., Varley, N.R., Inguaggiato, S., Cienfuegos, E., 2010. Geochemistry of H<sub>2</sub>- and CH<sub>4</sub>-  
1175 enriched hydrothermal fluids of Socorro Island, Revillagigedo Archipelago, Mexico.  
1176 Evidence for serpentinization and abiogenic methane. *Geofluids* 10, 542-555.
- 1177 Tissot, B., Welte, D., 1984. *Petroleum Formation and Occurrence*. Springer-Verlag, New York,  
1178 p. 699.

- 1179 Truche, L., McCollom, T., Martinez, I., 2020. Hydrogen and Abiotic Hydrocarbons: Molecules  
1180 that Change the World. *Elements* 16 (1), 13-18.
- 1181 Ungerer, P., 1990. State of the art research in kinetic modeling of oil formaion and expulsion.  
1182 *Org. Geochem.*, 16(1-3), 1-25.
- 1183 Vacquand, C., Deville, E., Beaumont, V., Guyot, F., Sissmann, O., Pillot, D., Arcilla, C.,  
1184 Prinzhofer, A., 2018. Reduced gas seepages in ophiolitic complexes: evidences for  
1185 multiple origins of the H<sub>2</sub>-CH<sub>4</sub>-N<sub>2</sub> gas mixtures. *Geochim. Cosmochim. Acta* 223, 437-  
1186 461.
- 1187 Waltenberg, K., 2015, Helium as a critical commodity: opportunities in the Northern Territory,  
1188 Annual Geoscience Exploration Seminar Proceedings, Northern Territory Geological  
1189 Survey Record 2015-002, p. 92.
- 1190 Walter, M.R., Veevers, J.J., Calver, C.R., Grey, K., 1995, Neoproterozoic stratigraphy of the  
1191 Centralian Superbasin, Australia, *Precambrian Research* 73, 173-195.  
1192 [https://doi.org/10.1016/0301-9268\(94\)00077-5](https://doi.org/10.1016/0301-9268(94)00077-5).
- 1193 Wang, P., Shen, Z., He, C., 2017. The geochemical characteristics of butane and its application.  
1194 *Natural Gas Geoscience* 28(4), 529-538.
- 1195 Waples, D., 1980. Time and temperature in petroleum generation and application of Lopatin's  
1196 technique to petroleum exploration. *AAPG Bull.* 64, 916-926.
- 1197 Watson, M. N., Boreham, C. J., Tingate, P. R., 2004. Carbon dioxide and carbonate cements in  
1198 the Otway Basin: implications for geological storage of carbon dioxide. *The APPEA*  
1199 *Journal* 44(1), 703-720.
- 1200 Warr, O., Giunta, T., Ballentine, C., Sherwood, B., 2019. Mechanisms and rates of <sup>4</sup>He, <sup>40</sup>Ar,  
1201 and H<sub>2</sub> production and accumulation in fracture fluids in Precambrian Shield  
1202 environments. *Chem. Geol.* 530, 119322. <https://doi.org/10.1016/j>.
- 1203 Wells, A.T., Forman, D.J., Ranford, L.C., Cook, P.J., 1970, Geology of the Amadeus Basin,  
1204 central Australia. Bureau of Mineral Resources, Geology and Geophysics, Bull. 100, pp.  
1205 282.
- 1206 Whiticar, M., 1999. Carbon and hydrogen isotope systematics of bacterial formation and  
1207 oxidation of methane. *Chem. Geol.* 161, 291-314.
- 1208 Wood, D., 2018. Thermal maturity and burial history modelling of shale is enhanced by use of  
1209 Arrhenius time-temperature index and memetic optimizer. *Petroleum* 4, 25-42.
- 1210 Woolnough, W.G., 1934. Natural gas in Australia and New Guinea. *AAPG bull.*, 18, 226-242.

- 1211 Wygrala, B.P., 1989. Integrated study of an oil field in the southern Po basin, northern Italy.  
1212 Unpublished Dissertation thesis, Berichte Kernforschungsanlage Jülich, University of  
1213 Köln, Köln. 217 pp.
- 1214 Young, I.F., Ambrose, G.J., 2007, Petroleum geology of the southeastern Amadeus Basin: The  
1215 search for sub-salt hydrocarbons, In MUNSON, T.J., AMBROSE, G.J. (Eds.), Central  
1216 Australian Basins Symposium, Northern Territory Geological Survey Special Publication  
1217 2, 183-204.
- 1218 Zgonnik, V., 2020. The occurrence and geoscience of natural hydrogen: a comprehensive  
1219 review. *Earth Sci. Rev.* 203, 103140.
- 1220 Zgonnik, V., Beaumont, V., Deville, E., Larin, N., Pillot, D., Farrell, K.M., 2015. Evidence for  
1221 natural molecular hydrogen seepage associated with Carolina bays (surficial, ovoid  
1222 depressions on the Atlantic Coastal Plain, Province of the USA). *JPiE and Science P*,  
1223 2(1), 31.

1224 **Table 1.** Total organic carbon (TOC) and Rock-Eval pyrolysis results for the analyzed  
 1225 shale samples retrieved from the studied wells in Amadeus Basin.

Well name	Formation	Depth (m)	TOC (wt% )	PC (wt%)	S1 mg/g	S2 mg/g	S3 mg/g	T <sub>max</sub> (C°)	HI	OI	PI	GP
Magee-1	U. Gillen	1927	0.52	0.07	0.35	0.47	0.47	393	90	903	0.43	0.82
Magee-1	U. Gillen	2217	0.49	0.28	2.65	0.78	0.60	350	159	122	0.77	3.43
Magee-1	L. Gillen	2323	0.79	0.05	0.18	0.45	0.27	442	57	34	0.29	0.63
Magee-1	L. Gillen	2326	0.81	0.05	0.20	0.41	0.27	445	51	33	0.33	0.61
Magee-1	L. Gillen	2329	0.53	0.03	0.10	0.28	0.24	444	53	45	0.26	0.38
Magee-1	L. Gillen	2332	0.63	0.03	0.09	0.33	0.24	441	52	38	0.21	0.42
Magee-1	L. Gillen	2335	0.49	0.02	0.06	0.21	0.20	442	43	41	0.22	0.27
Magee-1	L. Gillen	2338	0.60	0.04	0.11	0.36	0.21	442	60	35	0.23	0.47
Magee-1	L. Gillen	2341	0.50	0.03	0.09	0.31	0.15	444	62	30	0.23	0.40
Magee-1	L. Gillen	2344	0.45	0.02	0.05	0.14	0.18	393	31	40	0.26	0.19
Murphy-1	Inindia Beds	1005	2.29	0.05	0.13	0.49	0.04	433	21	2	0.21	0.62
Murphy-1	Inindia Beds	1056	3.19	0.06	0.16	0.53	0.06	435	17	2	0.23	0.69

1226 *TOC = total organic carbon. S<sub>1</sub> = volatile hydrocarbon content (mg HC/g rock); S<sub>2</sub> = remaining*  
 1227 *hydrocarbon generative potential (mg HC/g rock); S<sub>3</sub> = carbon dioxide yield (mg CO<sub>2</sub>/g rock); T<sub>max</sub> =*  
 1228 *temperature at maximum of S<sub>2</sub> peak; HI = hydrogen index (S<sub>2</sub> × 100/TOC) (mg HC/g TOC); OI = oxygen*  
 1229 *index (S<sub>3</sub> × 100/TOC) (mg CO<sub>2</sub>/g TOC); PI = production index (S<sub>1</sub>/ (S<sub>1</sub> + S<sub>2</sub>)); GP = Generation potential*  
 1230 *(S<sub>1</sub> + S<sub>2</sub>) (mg/g).*

1231

1232

1233

1234

1235

1236

1237 **Table 2.** Composition of the analyzed gas blend collected from the studied wells in Amadeus Basin. isotope gas analyses  
 1238 are reported from [McInnes et al. \(2017\)](#).

Well name	Depth (m)	N <sub>2</sub>	Ar	He	H <sub>2</sub>	CO <sub>2</sub>	C <sub>1</sub>	C <sub>2</sub>	C <sub>2+</sub>	IC <sub>4</sub> /nC <sub>4</sub>	IC <sub>5</sub> /nC <sub>5</sub>	DC	R/R <sub>a</sub>	CO <sub>2</sub> / <sup>3</sup> He	<sup>84</sup> Kr/ <sup>36</sup> Ar	<sup>132</sup> Xe/ <sup>84</sup> Kr
Magee-1	2349	43.61	0.46	6.20	0.03	0.82	39.26	6.10	3.52	0.63	1.11	0.80	-	-	-	-
Mt Kitty-1	2144	79.62	-	4.23	5.05	0.33	7.29	2.10	2.85	0.41	8.00	0.71	-	-	-	-
Mt Kitty-1	2156	61.04	0.57	8.96	11.4	0.09	13.14	3.53	1.24	0.36	6.50	0.73	0.031*	4.46 x 10 <sup>7</sup> *	0.0062*	0.142*
Mt Kitty-1	2253	99.67	-	0.0	0.0	0.08	0.04	0.02	0.01	0	0	-				
Murphy-1	1647	-	-	-	-	-	1.18	0.01	0.31	0	0	0.79				
Murphy-1	1650	-	-	-	-	-	32.69	2.87	3.90	1.08	0.96	0.83				
Murphy-1	1653	-	-	-	-	-	72.61	3.21	0.99	1.10	0.25	0.94				
Murphy-1	1656	-	-	-	-	-	27.81	3.14	3.61	1.30	0.82	0.81				
Murphy-1	1764	-	-	-	-	-	1.48	0.08	0.16	0	0	0.90				

1239 *Dc is the dryness Coefficient (C<sub>1</sub>/ΣC<sub>1</sub>-C<sub>3</sub>); \* Noble gas data is from (McInnes et al., 2017).*

1240