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## All-optical matter-wave lens using time-averaged potentials

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The precision of matter-wave sensors benefits from interrogating large-particle-number atomic ensembles at high cycle rates. Quantum-degenerate gases with their low effective temperatures allow for constraining systematic errors towards highest accuracy, but their production by evaporative cooling is costly with regard to both atom number and cycle rate. In this work, we report on the creation of cold matter-waves using a crossed optical dipole trap and shaping them by means of an all-optical matter-wave lens. We demonstrate the trade off between lowering the residual kinetic energy and increasing the atom number by reducing the duration of evaporative cooling and estimate the corresponding performance gain in matter-wave sensors. Our method is implemented using time-averaged optical potentials and hence easily applicable in optical dipole trapping setups.

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Ever since their first realization, atom interferometers<sup>1–4</sup> have become indispensable tools in fundamental physics<sup>5–17</sup> and inertial sensing<sup>18–30</sup>. The sensitivity of such matter-wave sensors scales with the enclosed space-time area which depends on the momentum transferred by the beam splitters as well as the time the atoms spend in the interferometer.

The expansion of the atomic clouds, used in interferometers, needs to be minimized and well controlled to reach long pulse separation times, control systematic shifts, and create ensembles dense enough to detect them after long time-of-flights. Nevertheless, colder ensembles with lower expansion rates typically need longer preparation times. Therefore, matter-wave sensors require sources with a high flux of large cold atomic ensembles to obtain fast repetition rates.

Bose-Einstein condensates (BECs) are well suited to perform interferometric measurements. They are investigated to control systematic effects related to residual motion at a level lower than a few parts in  $10^9$  of Earth's gravitational acceleration<sup>20,31–34</sup>. In addition, due to their narrower velocity distribution<sup>35</sup>, BECs offer higher beam splitting efficiencies and thus enhanced contrast<sup>23,36,37</sup>, especially for large momentum transfer<sup>36,38–43</sup>. Finally, the inherent atomic collisions present in BECs can enhance matter-wave interferometry by enabling (i) ultra-low expansion rates through collective mode dynamics with a recent demonstration of a 3D expansion energy of  $k_B \cdot 38^{+6}_{-7}$  pK<sup>44</sup>, and (ii) ultimately the generation of mode entanglement through spin-squeezing dynamics to significantly surpass the standard-quantum limit<sup>45–48</sup>.

Today's fastest BEC sources rely on atom-chip technology, where near-surface magnetic traps allow for rapid evaporation using radio frequency or microwave transitions. This approach benefits from constant high trapping frequencies during the evaporative cooling process, thus leading to repetition rates on the order of 1 Hz with BECs comprising  $10^5$  atoms<sup>49</sup>.

Anyway, since magnetic traps are not suitable in certain situations optical dipole traps become the tool of choice<sup>50</sup>. Examples are trapping of atomic species with low magnetic susceptibility<sup>51,52</sup>, or molecules<sup>53,54</sup> and composite particles<sup>55,56</sup>. In optical dipole traps external magnetic field allow tuning parameters, e.g., when using Feshbach resonances<sup>57</sup>.

Here, the intrinsic link between trap depth and trap frequencies in dipole traps<sup>58</sup> inhibits runaway evaporation. Cold ensembles can be only produced in shallow traps, leading to drastically increased preparation time  $t_p$ . This long standing problem has been recently overcome through the use of time-averaged potentials, where trap depth and trap frequencies can be controlled independently, thus allowing for more efficient and faster evaporation while maintaining high atom numbers<sup>52,59</sup>.

In this work, we use dynamic time-averaged potentials for efficient BEC generation and demonstrate an all-optical matter-wave lens capable of further reducing the ensemble's residual kinetic energy. Contrary to pulsed schemes of matter-wave lensing<sup>44,60–65</sup>, we keep the atoms trapped over the entire duration of the matter-wave lens<sup>37</sup>, which eases implementation in ground-based sensors. Moreover, we show that with this technique one can short-cut the evaporation sequence prior to the matter-wave lens, which increases the atomic flux by enhancing atom number and reducing cycle time while simultaneously reducing the effective temperature. Our method can largely improve the matter-wave sensor's stability in various application scenarios.

## Results

**Evaporative cooling.** We operate a crossed optical dipole trap at a wavelength of 1960 nm loaded from a <sup>87</sup>Rb magneto-optical

trap (details in the “Methods” section). The time-averaged potentials are generated by simultaneous center-position modulation of the crossed laser beams in the horizontal plane. Controlling the amplitude of this modulation and the intensity of the trapping beams enables the dynamic control and decoupling of the trapping frequencies and depth. We chose the waveform of the center-position modulation to generate a parabolic potential<sup>52</sup>.

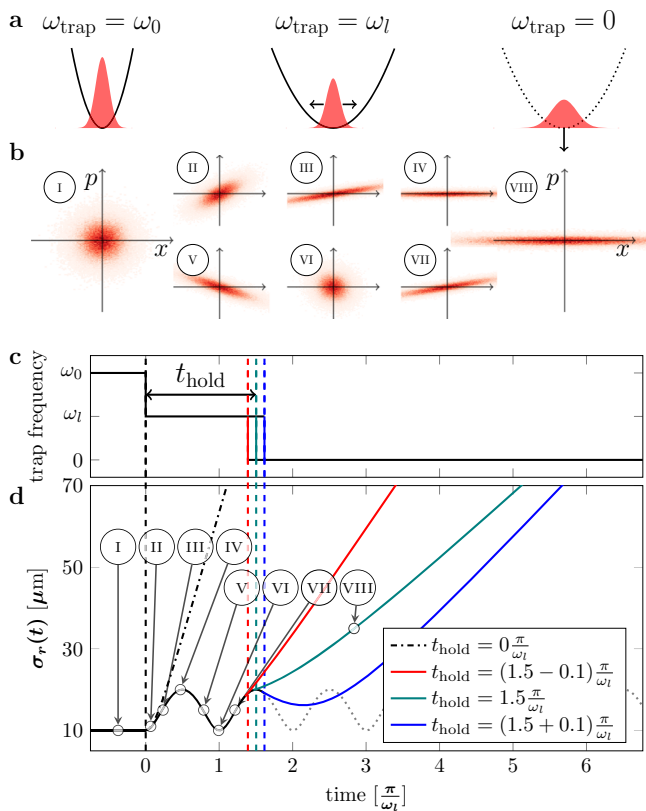
Up to  $2 \times 10^7$  rubidium atoms are loaded into the trap with trapping frequencies  $\omega/2\pi \approx \{140; 200; 780\}$  Hz in  $\{x'; y'; z\}$  direction (definition of coordinate systems in the “Methods” section) with a trap depth of 170  $\mu$ K. For this we operate the trap at the maximum achievable laser intensity of 12 W and the center-position modulation at an amplitude of  $h_0 = 140 \mu\text{m}$ .

We perform evaporative cooling by reducing the trap depth exponentially in time while keeping the trapping frequencies at a high level by reducing the amplitude of the center-position modulation. This method allows us to generate BECs with up to  $4 \times 10^5$  atoms within 5 s of evaporative cooling. By shortening the time constant of the exponential reduction we generate BECs with  $5 \times 10^4$  ( $2 \times 10^5$ ) particles within 2 s (3 s) of evaporative cooling. At the end of the evaporation sequence the trap has frequencies of  $\omega/2\pi \approx \{105; 140; 160\}$  Hz and a depth of about 200 nK. The expansion velocity of the condensate released from the final evaporation trap is  $2 \text{ mm s}^{-1}$ , which corresponds to an effective temperature of 40 nK.

**All-optical matter-wave lens.** Our matter-wave lens can be applied in any temperature regime explorable in our optical trap. We investigate the creation of collimated atomic ensembles for different initial temperatures of the matter-waves. To this aim, the evaporation sequence is stopped prematurely at different times to generate input atomic ensembles at rest with initial trap frequency  $\omega_0$  and initial temperature  $T_0$ . We then initiate the matter-wave lens by a rapid decompression<sup>66</sup> of the trap frequency in the horizontal directions from  $\omega_0$  to  $\omega_l$ . Here we denote by  $\omega_l$  the lensing potential in analogy with the Delta-kick collimation technique. The reduction of the trapping frequencies from the initial  $\omega_0$  to  $\omega_l$  depends on experimental feasibility, such as the maximum achievable amplitude of the center-position modulation and the modulation amplitude right before the rapid decompression. With ongoing evaporative cooling this amplitude is reduced and thus the trap can be relaxed much further for more continued sequences. However, we need to maintain the confinement in the vertical direction by adjusting the dipole trap's intensity to suppress heating or loss of atoms.

Subsequent oscillations in the trap result in a manipulation in phase space (Fig. 1a, b) for focusing, diffusion, and, importantly collimation of the matter-wave (Fig. 1c). Figure 1c depicts the expansion of a thermal ensemble in 1D for three different holding times ( $t_{\text{hold}}$ ) to highlight the importance of a well chosen timing for the lens. Figure 2 shows exemplary expansion velocities (colored circles) depending on the holding time  $t_{\text{hold}}$ . The colored curves in this graph display the simulated behavior following the scaling ansatz (details in the “Methods” section) with an error estimation displayed by shaded areas. Only for the final measurement (also shown in the inset in Fig. 2) we create a BEC with a condensed fraction of 92.5% of the total atom number and apply the matter-wave lens to it.

With the presented method we observe oscillations of the expansion rate, which are in good agreement with the simulations for different ensemble temperatures. For all investigated temperatures an optimal holding time exists for which the final expansion rate is minimized (Fig. 3a). The ratios of  $\sigma_{v,l}/\sigma_{v,0}$  and  $\omega_l/\omega_0$  for each measurement is shown in Fig. 3b.

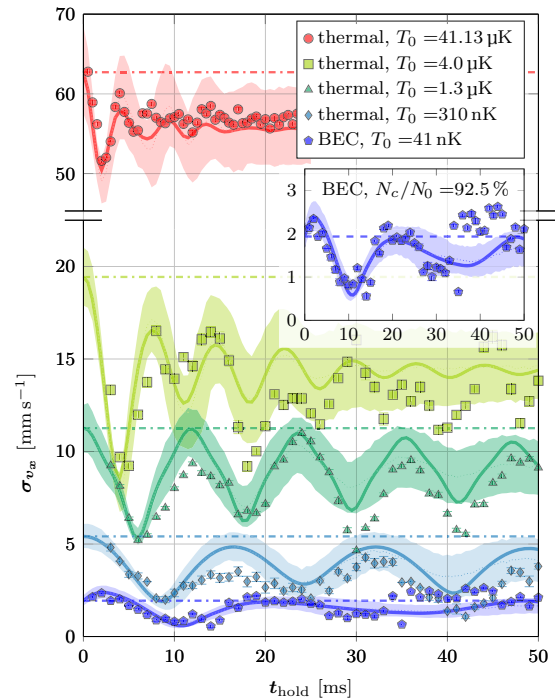


**Fig. 1 Scheme of the matter-wave lens.** The drawing in **a** shows the three trap configurations and the distribution of the atomic ensemble during the matter-wave lens. At the beginning the trap frequency ( $\omega_{\text{trap}}$ ) corresponds to the initial trap frequency ( $\omega_0$ ), which is then decompressed to the lens trap frequency ( $\omega_l$ ). After the holding time ( $t_{\text{hold}}$ ) the trap is switched off ( $\omega_{\text{trap}} = 0$ ). The phase-space ( $p-x$ ) diagrams in **b** show the atomic distributions at different timings during the matter-wave lens. Panel **c** shows the behavior of the time dependency of the trapping frequency. After the holding time in the initial trap the trapping frequency is rapidly decreased at time  $t = 0$ . Panel **d** shows the time evolution of the atomic ensemble's size ( $\sigma_r$ ), which starts to oscillate (solid and dotted black line) after the trap is relaxed. At time  $t_{\text{hold}} = (n + 0.5) \times \pi / \omega_l$ , with  $n \in \mathbb{N}$ , this oscillation reaches an upper turning point (teal curve). The atomic ensemble is released at its maximum size to minimize its later expansion rate. The labeled points (I-VIII) correspond to the phase-space diagrams in **b**. If the release time ( $t_{\text{hold}}$ ) does not match this condition the expansion rate is not minimized (red and blue curve). The dashed-dotted black curve displays the size of a free falling ensemble without lensing, released at time  $t = 0$ .

The change in atom number from the initial to the lensing trap (Fig. 3a) lies within the error bars and arises mainly due to pointing instabilities of the crossed optical dipole trap beams. The lowest expansion rate is achieved with  $553(49) \mu\text{m s}^{-1}$  with a related effective temperature of  $3.2(0.6) \text{ nK}$  and an atom number of  $4.24(0.02) \times 10^5$ . With this method we achieve a more than one order of magnitude lower effective temperature while maintaining a comparable atom number compared to evaporative cooling.

## Discussion

In this paper, we demonstrate a technique to reduce the expansion velocity of an atomic ensemble by rapid decompression and subsequent release from an dipole trap at a well-controlled time. The efficiency of the matter-wave lens for higher temperatures is mainly limited experimentally by the limited ratio between the

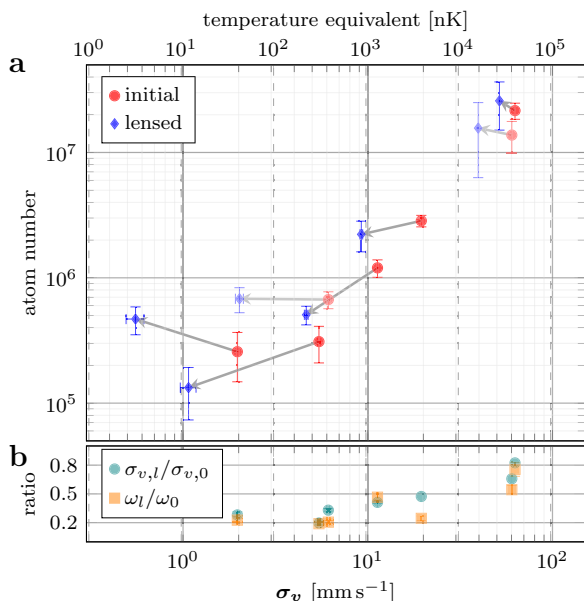


**Fig. 2 Oscillations of expansion velocity.** Expansion velocity ( $\sigma_v$ ) after 30 ms of time of flight in dependence of the holding time. The graph shows the expansion velocity for thermal ensembles with different initial temperatures  $T_0$  and the Bose-Einstein condensate (BEC) with an initial expansion temperature of  $T_0 = 41 \text{ nK}$  and a condensed fraction of  $N_c/N_0 = 92.5\%$ . The markers show the measurements, the dashed-dotted lines the expansion rate from the initial traps (cf. black dashed-dotted line in Fig. 1). The simulations use the scaling ansatz and are depicted as lines with a shaded 1-sigma error estimation for the used trap parameters. The inset displays the size oscillation of the Bose-Einstein condensate on an enlarged scale for better readability.

initial and the lensing trap frequency  $\omega_l/\omega_0$  (Fig. 3b) which is constrained by the maximum possible spatial modulation amplitude of the trapping beams. In general, according to the Liouville theorem, the expansion speed reduction of the matter-wave is proportional to  $(\omega_0/\omega_l)^2$  where a large aspect ratio enables a better collimated ensemble. The atoms are loaded into the time-average potential with an optimized center-position modulation amplitude of  $140 \mu\text{m}$ , while the maximum is  $200 \mu\text{m}$ . During the evaporative sequence this amplitude is decreased. Consequently, the relaxation of the trap is less efficient at the beginning of the evaporative sequence or directly after the loading of the trap.

Another constraint is that the trap's confinement in the unpainted vertical direction is required to remain constant. If the vertical trap frequency is increased we observe heating effects and suffer from atom loss when it is decreased. To compensate for the trap depth reduction during the switch from the initial to the lensing trap we increase the dipole trap laser's intensity accordingly.

An additional modulation in the vertical direction, e.g., by means of a two-dimensional acousto-optical deflector, as well as an intersection angle of  $90^\circ$  would enable the generation of isotropic traps. In such a configuration, the determination of the optimal holding time will benefit from the in-phase oscillations of the atomic ensemble's size<sup>67</sup>. When applying our matter-wave lens in a dual-species experiment, isotropy of the trap will also improve the miscibility of the two ensembles<sup>68</sup>.



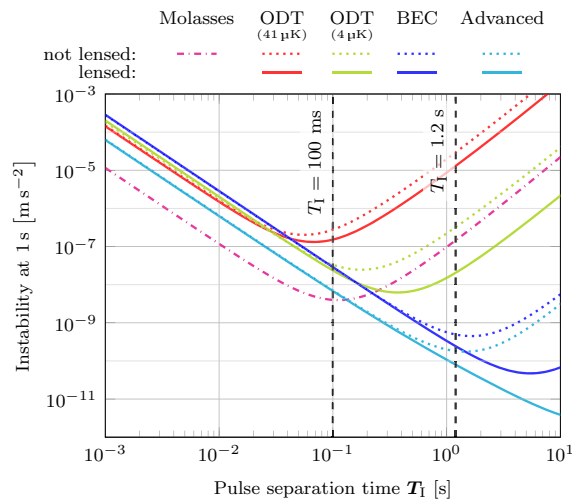
**Fig. 3 Expansion velocity dependent performance.** Panel **a** shows the atom number depending on the expansion velocity ( $\sigma_v$ ) for the resulting expansion velocity (blue circles) after the matter-wave lens and starting points (red circles). The lines connect the corresponding data points. The error bars arise from the time of flight measurements used to determine the expansion velocity. We use the 1-sigma fitting error of the ballistic expansion for  $\sigma_v$  and the statistical error of measured atom numbers. The two grayed out data points associated with the starting expansion velocities 60 mm s<sup>-1</sup> and 6 mm s<sup>-1</sup> are not displayed in Fig. 2 while all the others are included. Panel **b** shows the ratio of the lensed and the initial ensembles expansion velocity ( $\sigma_{v,l}/\sigma_{v,0}$ ) as well as the trapping frequencies ( $\omega_l/\omega_0$ ) depending on the starting expansion velocity.

To illustrate the relevance for atom interferometers, we discuss the impact of our source in different regimes (details in the “Methods” section) operated at the standard quantum limit for an acceleration measurement. In a Mach-Zehnder-like atom interferometer<sup>1,18</sup>, the instability reads

$$\sigma_a(\tau) = \frac{1}{C\sqrt{N}n k_{\text{eff}} T_I^2} \cdot \sqrt{\frac{t_{\text{cycle}}}{\tau}} \quad (1)$$

after an averaging time  $\tau$ , neglecting the impact of finite pulse durations on the scale factor<sup>69–71</sup>. Eq. (1) scales with the interferometer contrast  $C$ , the atom number per cycle  $N$ , the effective wave number  $n\hbar k_{\text{eff}}$  indicating a momentum transfer during the atom-light interaction corresponding to  $2n$  photons, and the separation time between the interferometer light pulses  $T_I$ . The cycle time of the experiment  $t_{\text{cycle}} = t_p + 2T_I + t_D$  includes the time for preparing the ensemble  $t_p$ , the interferometer  $2T_I$ , and the detection  $t_D$ . In Eq. (1), the contrast depends on the beam splitting efficiency. This, in turn, is affected by the velocity acceptance and intensity profile of the beam splitting light, both implying inhomogeneous Rabi frequencies, and consequently a reduced mean excitation efficiency<sup>35,72,73</sup>. Due to expansion of the atomic ensemble and inhomogeneous excitation, a constrained beam diameter implicitly leads to a dependency of the contrast  $C$  on the pulse separation time  $T_I$ , which we chose as a boundary for our discussion. We keep the effective wave-number fixed and evaluate  $\sigma_a(1\text{ s})$  for different source parameters when varying  $T_I$ .

Figure 4 shows the result for collimated (solid lines) and uncollimated (dotted lines) ensembles in our model (details in the “Methods” section) and compares them to the instability under



**Fig. 4 Instability comparison.** Behavior of the instability for shot-noise limited atom interferometers after an integration time of  $\tau = 1\text{ s}$  for different sources (Tab. 1) over the pulse separation time  $T_I$ . The instability is estimated for molasses cooled ensembles, thermal ensembles released from the optical dipole trap (ODT), and Bose-Einstein condensates (BEC and Advanced) also released from the optical dipole trap. The colors of the curves, except for the molasses and advanced case, agree with the measurements displayed in Fig. 2. The different source parameters are described in more detail in the “Methods” section.

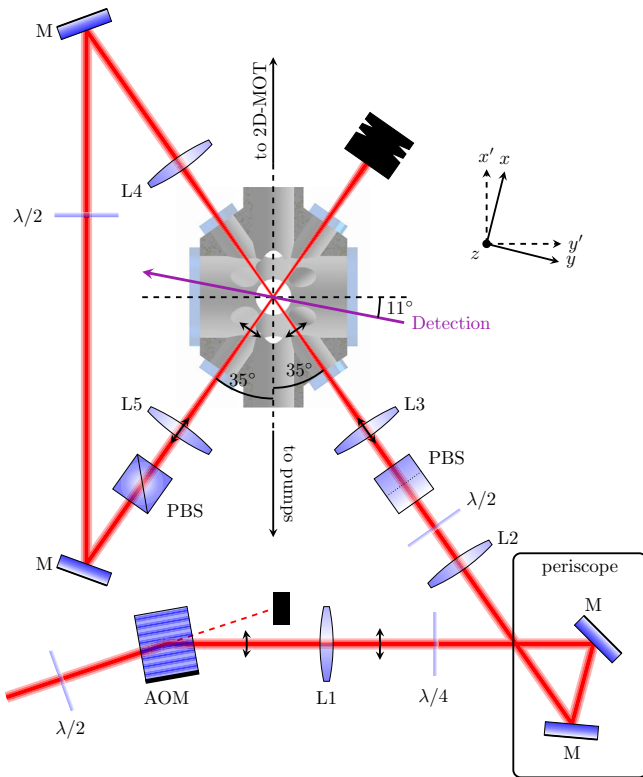
use of a molasses-cooled ensemble (dash-dotted line). Up to  $T_I = 100\text{ ms}$  and  $\sigma_a(1\text{ s}) = 10^{-8}\text{ m s}^{-2}$ , the molasses outperforms evaporatively cooled atoms or BECs due the duration of the evaporation adding to the cycle time and associated losses. In this time regime, the latter can still be beneficial for implementing large momentum transfer beam splitters<sup>36,38–40,42,43</sup> reducing  $\sigma_a(\tau)$  or suppressing systematic errors<sup>20,31–34,74</sup> which is not represented in our model and beyond the scope of this paper. According to the curves, exploiting higher  $T_I$  for increased performance requires evaporatively cooled atoms or BECs. This shows the relevance for experiments on large baselines<sup>23,37,74–77</sup> or in microgravity<sup>78,79</sup>. We highlight the extrapolation for the Very Long Baseline Atom Interferometer (VLBAI)<sup>76,80</sup>, targeting a pulse separation time of  $T_I = 1.2\text{ s}$ <sup>81</sup>. Here, the model describing our source gives the perspective of reaching picokelvin expansion temperatures of matter-wave lensed large atomic ensembles.

## Methods

**Experimental realization.** The experimental apparatus is designed to operate simultaneous atom interferometers using rubidium and potassium and is described in detail in references<sup>9,10,82</sup>.

For the experiments presented in this article only rubidium atoms were loaded from a two dimensional to a three dimensional magneto-optical trap (2D/3D-MOT) situated in our main chamber. After 2 s we turn off the 2D-MOT and compress the atomic ensemble by ramping up the magnetic field gradient as well as the detuning of the cooling laser in the 3D-MOT. Subsequent to compression, the atoms are loaded into the crossed dipole trap by switching off the magnetic fields and increasing the detuning of the cooling laser to about  $-30\Gamma$ , with  $\Gamma$  being the natural linewidth of the  $D_2$  transition.

Figure 5 depicts the setup of our crossed optical dipole trap. The center-position modulation of the trapping beams is achieved by modulating the frequency driving the acousto-optical modulator (AOM) (Polytec, ATM-1002FA53.24). A voltage-controlled oscillator (Mini-Circuits, ZOS-150+) generates the signal for this, which is driven by a programmable arbitrary-waveform generator (Rigol, DG1022Z). We chose the waveform to generate a large-volume parabolic potential based on the derivation shown by Roy et al.<sup>52</sup>. The amplitude of the displacement of the center-position of the dipole trap beam,  $h_0$ , is controlled by regulating the amplitude of the AOM’s frequency modulation. This yields a maximum beam displacement of  $h_0 = 200\text{ }\mu\text{m}$  ( $300\text{ }\mu\text{m}$ ) at the position of the atoms for the initial (recycled) beam.



**Fig. 5 Experimental setup.** Optical setup of the dipole trap and alignment through the vacuum chamber. Vertical arrows point towards the 2D magneto-optical trap (2D-MOT) and the vacuum pumps. The acousto-optical modulator (AOM) is used for modulating the center-position of the laser beams and intensity control. A  $\lambda/2$  wave plate rotates the polarization of the beam for best diffraction efficiency of the AOM. Lenses L1 ( $f_1 = 100$  mm) and L2 ( $f_2 = 300$  mm) magnify the beam radius to about 3 mm, 4 mm (vertical, horizontal). Downstream lenses L3, L4, and L5 ( $f_{3,4,5} = 150$  mm) focus, re-collimate and re-focus the beam into the center of the chamber. The  $\lambda/2$  and the  $\lambda/4$  wave plates set the polarization for maximum transmission at the orthogonal oriented polarization beam splitters (PBS). The mirrors (M) are used to guide the beam through the chamber. The purple arrow indicates the direction of absorption detection along the  $y$ -direction.

**Data acquisition and analysis.** We apply our matter-wave lens subsequent to loading the dipole trap and evaporative cooling. The duration of the complete evaporative sequence is 5 s for the measurements presented here we interrupt this sequence after 0 s, 0.2 s, 1 s, 2 s, 3.5 s, 4.3 s and 5 s. Before the step-wise change of the trap frequency ( $\omega_0 \rightarrow \omega_i$ ) we hold the ensemble in the trap given by the respective evaporation step configuration for 50 ms.

During the matter-wave lens, the rapid decompression of the trap causes oscillations of the ensemble's radius in the lensing trap. Depending on the release time we observe oscillations by performing absorption imaging with iterating  $t_{\text{hold}}$  for different times after the release from the trap. For each holding time the expansion velocity is extracted by fitting a ballistic expansion. This expansion can be transformed into an effective temperature using:

$$\sigma_{v_i}^2 = \frac{k_B T_i}{m}, \quad (2)$$

along each direction. The measurement is performed for different starting temperatures in the thermal regime as well as the BEC.

The simulations shown in Fig. 2 use the scaling ansatz as described in the ‘‘Scaling Ansatz’’ section. Here, the trapping frequencies of the lens potential in  $x$ - and  $y$ -direction have been extracted by fitting two damped oscillations to the measured data. The starting expansion velocity was set by choosing a reasonable initial radius of the ensemble (Table 1). The other parameters arise from the measurements or simulations of the trapping potentials. The shaded areas in fig. 2 depict an error estimation of the expansion velocity oscillations obtained from performing the simulation by randomly choosing input parameters from within the error bars for 1000 simulation runs and calculating the mean value as well as the standard deviation for each  $t_{\text{hold}}$ .

**Table 1 Source parameters for the instability estimation.**

	$\sigma_r$ (h/v) [ $\mu\text{m}$ ]	$\sigma_v$ (h/v) [ $\text{mm s}^{-1}$ ]	$N$	$t_p$ [s]
Molasses	750/750	30.9/30.9	$4 \times 10^8$	2
ODT (41 $\mu\text{K}$ )	65/6.2	62.7/44.5	$2.3 \times 10^7$	2.7
ODT (41 $\mu\text{K}$ ) lensed	86/7.6	51.5/40.8	$2.2 \times 10^7$	2.7
ODT (4 $\mu\text{K}$ )	15.2/12.2	19.6/12	$2.7 \times 10^6$	4.7
ODT (4 $\mu\text{K}$ ) lensed	49/12.2	9.2/12.9	$2.6 \times 10^6$	4.7
BEC	3.8/3.3	2/2	$4.3 \times 10^5$	8.2
BEC lensed	16.9/3.3	0.55/2.2	$4.2 \times 10^5$	8.2
Advanced	5/5	2/2	$1 \times 10^6$	1
Advanced lensed	46.1/46.1	0.14/0.14	$1 \times 10^6$	1

Source parameters for the instability estimation for molasses cooled, released from the optical dipole trap (ODT) with and without evaporation, and Bose-Einstein condensates (BEC and Advanced). The values for the ensemble radius ( $\sigma_r$ ) and expansion velocity ( $\sigma_v$ ) are given for the horizontal (h) and vertical (v) direction, which corresponds to the transverse and longitudinal direction of the beam splitter respectively.  $N$  is the number of atoms and  $t_p$  the preparation time to calculate the cycle time  $t_{\text{cycle}}$ .

**Scaling Ansatz.** In the case of a thermal ensemble in the collision-less regime, the dynamics of a classical gas can be described using the scaling ansatz<sup>83,84</sup>, which we briefly recall here for sake of simplicity. Here, the size of the ensemble scales with the time dependent dimensionless factor  $b_i(t)$ .

$$\ddot{b}_i(t) + \omega_i^2(t)b_i(t) - \omega_{0,i}^2 \frac{\theta_i(t)}{b_i(t)} + \omega_{0,i}^2 \xi \left( \frac{\theta_i(t)}{b_i(t)} - \frac{1}{b_i(t) \prod_j b_j(t)} \right) = 0 \quad (3)$$

$$\dot{\theta}_i(t) + 2 \frac{\dot{b}_i(t)}{b_i(t)} \theta_i(t) + \frac{1}{\tau} \left( \theta_i(t) - \frac{1}{3} \sum_j \theta_j(t) \right) = 0, \quad (4)$$

where  $\theta_i$  acts as an effective temperature in the directions  $i \in x, y, z$ . Here  $\omega_{0,i}$  denotes the initial angular trap frequency and  $\omega_i(t)$  denotes the time-dependent angular trap frequency defined such as:  $\omega_i(t) = \omega_{i,j}$  for  $0 < t < t_{\text{hold}}$ , with  $\omega_{i,j}$  being the lensing potential, and  $\omega_i(t) = 0$  after the release (see Fig. 1). This system of coupled differential equations contains the mean field interaction, given by the factor:

$$\xi = \frac{E_{\text{mf}}}{E_{\text{mf}} + k_B T}, \quad (5)$$

with

$$E_{\text{mf}} = \frac{4\pi\hbar^2 a_s n_0}{m}, \quad (6)$$

where  $a_s$  is the  $s$ -wave scattering length,  $n_0$  the peak density and  $m$  the mass of a single particle. Collision effects are also taken into account through

$$\tau = \tau_0 \times \left( \prod_j b_j \right) \times \left( \frac{1}{3} \sum_k \theta_k \right) \quad (7)$$

with the relaxation time

$$\tau_0 = \frac{5}{4\gamma} \quad (8)$$

and<sup>84</sup>

$$\gamma = \frac{2}{\sqrt{2\pi}} n_0 \sigma_{\text{coll}} \sqrt{\frac{k_B T}{m}}. \quad (9)$$

In the special case of a BEC, the mean field energy is large compared to the thermal ensemble's energy ( $\xi \approx 1$ ) and the time scale on which collisions appear goes to zero ( $\tau = 0$ ). In this case the time dependent evolution of the matter-wave can be described following Castin & Dum<sup>85</sup>. Here, the evolution of the BEC's Thomas-Fermi radius,  $R_i(t) = b_i(t)R_i(0)$ , is described by the time-dependent evolution of the scaling parameter:

$$\ddot{b}_i(t) + \omega_i^2(t)b_i(t) = \frac{\omega_i(0)}{b_i(t)b_x(t)b_y(t)b_z(t)} \quad (10)$$

and  $R_i(0)$  is the initial Thomas-Fermi radius of the BEC along the  $i$ -th direction. It is worth to notice that recent studies<sup>86,87</sup> extend the analysis of Guéry-Odelin<sup>83</sup> and Pedri et al.<sup>84</sup> to the BEC regime described by Castin & Dum<sup>85</sup>.

With this set of equations the time evolution of the ensemble's size ( $\sigma_r$ ) and velocity distribution ( $\sigma_v$ ) is determined during the entire sequence of our matter-wave lensing sequence by

$$\sigma_r(t) = \sigma_r(0) \times b_i(t) \quad (11)$$

and

$$\sigma_{v_i}(t) = \frac{d\sigma_{v_i}(t)}{dt}. \quad (12)$$

The scaling parameter  $b_i$  can be applied either on the radius of a gaussian distributed thermal ensemble or the Thomas-Fermi radius of a BEC.

**Estimation of instability in matter-wave sensors.** The instability of a matter-wave sensor operating at the standard quantum limit can be estimated using Eq. (1). Here we assume Raman beam splitters ( $n = 1$ ) with a  $1/e^2$ -radius of 1.2 cm and a pulse duration of  $t_\pi = 15 \mu\text{s}$ . The contrast ( $C$ ) is taken into account as the product of the excitation probabilities of the atom-light interactions during the Mach-Zehnder type interferometer following Loriani et al.<sup>72</sup> Table 1 shows the source parameters used for the estimation of the instability. We chose three parameter sets from the here presented measurements of two thermal ensembles released from the optical dipole trap with starting temperatures of  $T_0 = 41 \mu\text{K}$  and  $4 \mu\text{K}$  and the BEC. Besides that we simulated the performance of the interferometer operated with a molasses cooled ensemble combined with a velocity selective Raman pulse of  $30 \mu\text{s}$ <sup>73</sup>, based on typical parameters in our experiment, and an advanced scenario. For this we assume a BEC with  $1 \times 10^6$  atoms after a preparation time  $t_p = 1 \text{ s}$  with a starting expansion velocity of  $2 \text{ mm s}^{-1}$ , as anticipated for the VLBAI setup<sup>76,80</sup>. We extrapolate the performance of our matter-wave lens for this experiment, resulting in expansion velocities of  $0.135 \text{ mm s}^{-1}$  corresponding to an equivalent 3D temperature of  $200 \text{ pK}$ .

### Data availability

The data used in this manuscript are available from the corresponding author upon reasonable request.

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## Author contributions

W.E., E.M.R., and D.S. designed the experimental setup and the dipole trapping laser system. H.A., A.H., A.R., and D.S. contributed to the design, operation, and maintenance of the laser system and the overall setup. R.C., E.C. and N.G. set the theoretical framework of this work. H.A., R.C., C.S., and D.S. drafted the initial manuscript. H.A., and R.C. performed the analysis of the data presented in this manuscript. H.A., and R.C. under lead of N.G. and C.S. performed the instability study. C.V., M.W., C.L., S.H. together with the other authors discussed and evaluated the results and contributed to, reviewed, and approved of the manuscript.

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The authors declare no competing interests.

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