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Characterization of a New Generation of Silicon Detector: the SIRIUS Side "Strippy-Pad" Detector

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Abstract: SIRIUS (Spectroscopy & Identification of Rare Isotopes Using S3) is a detection system designed for the focal plane of S3 (the Super Separator Spectrometer), which is part of the SPIRAL2 (Système de Production d'Ions RAdioactifs en Ligne de 2e génération) facility at GANIL (France). This study presents the characterization of the Side silicon detector of SIRIUS. This new "Strippy-pad" detector design benefits from an ultra-high resistivity and the windowless technique from Micron Semiconductor Ltd. These detectors were tested at the IPHC Strasbourg on a dedicated test bench with custom fast preamplifiers coupled to digital electronics cards. Combining all these elements, we obtained resolutions as good as 13.6 keV for certain pixels of this detector for 8.8 MeV alpha particles, with an average resolution measured at 14.7 keV over the whole detector.

Keywords: Tracking and position-sensitive detectors, Solid-state detectors, Double Sided silicon Detectors, Charged-particle spectroscopy, alpha particle detection, Electron spectroscopy

1. INTRODUCTION

One of the main challenges in modern nuclear physics is the study of nuclei far from stability, which provides strong constraints for nuclear models. Closed nucleonic shells are clear fingerprints of quantal behaviour of nuclei leading to the socalled magic numbers of neutrons or protons. The region around the doubly magic ¹⁰⁰Sn and the region of very heavy and super heavy nuclei are excellent nuclear structure physics cases to study these extreme magic systems and for the quest of an ultimate island of stability.

The very low production cross-section of these nuclei needs to be compensated by improving the sensitivity and efficiency of the experimental setups. The present development is performed within the S³ [1] (Super Separator Spectrometer) project, which is a part of the new SPIRAL2 (Système de Production d'Ions RAdioactifs en Ligne de 2e génération) facility at GANIL (Caen, France). SIRIUS [1] (Spectroscopy & Identification of Rare Isotopes Using S³) is the detection system designed and optimized for spectroscopic studies of rare exotic isotopes and heavy/super-heavy elements at the final focal plane of S³. Its compact geometry combined with new advanced technologies, such as new detector designs, customized preamplifiers and adapted digital electronics, allows for a large range of energies to be covered (from low-energy electrons, ~50 keV, to heavy fission fragments, ~100-200 MeV) with resolution (from 0.2% for the typical alpha energies and up to 1% for the heavy ions [1]).

Ions enter the SIRIUS focal plane detector array through the large-area tracker detector for time of flight, ion tracking as well as mass measurement. These ions are then implanted into a 10×10 cm² Double-sided Silicon Strip Detector (DSSD, also called implantation detector) segmented in 128 strips on each side. Implanted nuclei, as well as alpha particles and conversion electrons emitted by the subsequent decays, are also partially detected in the DSSD. Four large silicon detectors are placed in a "tunnel" configuration, upstream from the DSSD to detect alpha particles and electrons escaping from the DSSD. An array of five Germanium clover detectors for gamma ray spectroscopy is placed as close as possible around this "silicon box" (one behind the DSSD and 4 around the Side detectors in a cross configuration, Fig. 1 [2]). All detected decay events are individually time-stamped in dedicated digital electronics modules.

picture1.jpg

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1.1 Detector Characteristics and Design

46 The Side detectors of SIRIUS are part of the new generation of silicon detectors produced by Micron Semiconductor Ltd. 47 These detectors are produced using a high resistivity silicon wafers (20-50 K Ω .cm) made from N-Type silicon with the 48 orientation <100>. With a thickness of ~700 μ m and large active area (10 × 10 cm²), they have a surface capacitance of 49 19.71 pF.cm⁻² at full depletion voltage. Finally, due to their dead layer thinner than 50 nm (typically around 30 nm, as provided 50 by Micron Semiconductor Ltd under their 9.5P denomination [3]) are called "windowless". This new design of "Strippy-Pad" 51 (Fig. 2) detectors is composed of 64 pixels (pads) on the ohmic side and one large single windowless pixel side with 3% grid 52 coverage (for good charge collection) on the junction side. 53

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57 Each pad has a peripheral metallization (Fig. 3) for charge collection associated with a 30 µm wide strip connected to a 58 bonding pad. The bonding pads are gathered into four groups (two columns of 8 pixels in each group), and are located on the 59 same edge of the detector (Fig. 2). To avoid charge trapping under the usual insulation layer between the wafer and the metallic 60 collection strip, the signal-tracking were designed as active strips (i.e., metallic deposit on an active volume of silicon Fig. 3, 61 rather than the metallic deposit being insulated from the wafer by an oxide layer). These metallic strips are located between 62 every two columns of pixels. Due to this design, the surface of each pixel, as well as the tracking length, differs from one pixel 63 to the next. Thus, the measured pixel capacitances range from 51 pF (large pixels area and long tracking: 1.4484 cm² and 64 8.51 cm of tracking) down to 32 pF (smaller active area: 1.356 cm² and no tracking next to the bounding pads). 65

Picture3.jpg

68 This "Strippy-Pad" concept [4,5] enables windowless operating for a pad detector by avoiding the dead layer associated with 69 the double metal technique usually used for pads, as well as the insulation oxide layer for the metallic strip. This design also 70 allows the connectors to be located only on one side, on the "bottom"-side of the detector frame, to ensure a very compact 71 geometry of SIRIUS. 72

An innovative hybrid support frame, combining an anodized aluminium frame for an efficient cooling and a PCB for the 73 connectors and signal transmission, were developed at the IPHC Strasbourg (Fig. 2). Good cooling and grounding properties of 74 this detector are provided by the aluminium frame. In addition to the opening dedicated to the silicon detector wafer, and to 75 minimize the absorption of low-energy gamma- and X-rays, a shallow volume is machined at the bottom of the aluminium 76 frames for the positioning of a double-sided PCB board. This board bridges the signals from the detector to the output cables. Bonding pads are placed facing the detector field plate, guard rings and pixel bonding pads. Connection to the preamplifier is 78 made via two 40-channel compact cables onto SAMTEC connectors. The PCB board is glued onto a non-anodized recess into the aluminium frame using a silver-loaded glue to ensure good grounding connection. This hybrid frame allows for a compact 80 mounting of the five detectors (Four Side detectors and one implantation detector) into a silicon-box geometry.

1.2 SIRIUS Detectors Test Bench

84 The characterization of the Side detectors of SIRIUS was performed at the IPHC Strasbourg using a dedicated test bench [[6], 85 page 151-155]. It is composed of a chamber for operation under controlled vacuum/temperature and a digital acquisition chain 86 based on TNT2 cards [7]. A copper cooling plate was designed to be able to either place the ohmic (or the junction) side in front 87 of a radioactive source inside the vacuum chamber. Four radioactive sources were available (²⁰⁷Bi, ¹³³Ba, ²¹²Pb and a mixed 88 ²³⁹Pu/²⁴¹Am/²⁴⁴Cm) and could precisely be positioned using a SMARACT piezoelectric XY-O table combined with a set of 89 collimators.

90 The acquisition chain is composed of custom CR110 CREMAT preamplifiers [8] combined with digital electronics composed 91 92 TNT2 cards [7]. The custom gain of these custom preamplifiers is 173 mV/MeV (rather than 62mV/MeV, both of corresponding to unloaded gain) was chosen to get a 14.5 MeV full-scale energy range (i.e. 1.25 Vpp signal) while retaining 93 94 their very fast response time. For pixels with capacitances in the 30-50 pF range, the response-time of the custom CR110 is 13.2-16.6 ns (0.17 x C_d (pF) + 3 ns as given by the CREMAT specification sheet [8], where C_d is accounting for detector and 95 cables capacitance (30 pF)) which is much faster than the ~100-150 ns needed for charge collection our silicon and could enable 96 to probe the discrimination between alpha particle and electron through their different collection time. A dedicated 26 channel 97 PCB motherboard was developed at the IPHC to host these preamplifiers. After the pre-amplification stage, the signals are

digitized at a sampling rate of 100 MHz rate using 7 TNT2 cards [7] developed at IPHC Strasbourg. These 4 channels TNT2 boards are based on 14-bit flash ADC's operating at 100 MHz. Each channel is treated independently and in parallel. Moreover, they can generate online spectra as well as to store individual oscillograms (traces) to disc for offline treatment. The energy measurement is performed using trapezoidal shaping provided by a Jordanov algorithm [9] implemented on the FPGA of the TNT2 board. After setting the proper pole zero correction for each individual channel, this pipe-line data processing of the energy is defined by 2 parameters: integration time, and flat top duration of the trapezoid (defined for each channel individually).

2. VALIDATION OF THE DETECTORS

107 This paper is focused on the validation of the concept of the "Strippy-Pad" detector. Two detectors were characterized during this work. The results presented here report on the first one, the characteristic of the second one being very similar.

109 2.1 Reduced Number of Channels by Pixel Grouping

Even though the "Strippy-Pad" detector was designed and produced with 64 pixels, it was foreseen from the beginning to reduce this number by making macro-pixels. Indeed, in order to meet the specifications of the SIRIUS collaboration, the number of channels per detector needed to be reduced, with the condition that the maximum energy resolution should remain below 20 keV for 8 MeV alpha particles. A good pixel granularity near the DSSD is required to limit the dead layer effect on the escaped alpha particles [Courtesy of A. Lopez-Martens, [5]], and lower granularity can be accepted further away from the DSSD. It is worth pointing out that these constraints lead to the grouping of lower capacitance pixels rather than higher capacitance ones.

We established that the capacitance of the macropixels should remain under 145 pF in order to minimize the influence of the capacitance noise from the CREMAT preamplifier on the measured resolution (less than 10% of the 20 keV of the specification); considering the equivalent noise charge slope (3.8 e⁻RMS/pF) and the average energy for a pair creation in the silicon (3.62 eV). Moreover, the intrinsic noise of the system at 0 pF was less than 5 keV (1.7 keV coming from the preamplifier [8])

Several possible schemes of macropixels structures corresponding to 16, 20 or 24 groups were studied. Since the structure of the "Strippy-Pad" presents four identical blocks composed of 2 columns of 8 pixels with their respective strips, the reduced schemes were designed around this elementary block. Several reduced configurations were selected by the collaboration and fully characterized. However, two were selected, based on their performances and have been used for the present work. Fig. 4 illustrates both the mapping of the macropixels used during this work (Fig. 4 (a)) and the two macropixels configurations selected (Fig. 4 (b)). The final choice between these two configurations was left to the collaboration.

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Practically, the number of pixels is reduced physically by using a PCB board plugged between the connector and the Kaptoncable (Fig. 5) grouping the pixels into the macropixels.

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136 2.2 Testing Protocol

A two-step protocol was established to test the detectors upon delivery and to precisely determine the optimal operating conditions for each detector. This protocol will be applied for the validation tests of every Side detector that the collaboration order.

2.2.1 Depletion Voltage and Operating Voltage Determination

After a visual inspection under binocular lenses, the detector is characterized in a clean room. During this step, done in a dark room, one measures pixel by pixel the current (Keithley 6487 (set V and read I) Fig. 6.b) and capacitance (Keithley 6487 (set V) and Boonton 7200 (read C) Fig. 6.a) evolution with applied voltage. This characterization enables us to determine the full depletion voltage of the detector (VD), the possible operating voltage range as well as the good behaviour of every single pixel (not macropixel).

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For this particular detector, the full depletion voltage (VD) is at 50 V from the capacitance plots (Fig. 6 (a)); Above VD, there are no more changes in the measured capacitances, capacitance stabilizes around a constant value. This value is consistent with the value given by Micron Semiconductor Ltd. as well as the value obtained from the current versus voltage plot (start of the plateau observed in Fig. 6 (b)). Thus, the range of operation of the detector was chosen to be above 1.5xVD and below 150V (according to the recommendation of MICRON Semiconductor Ltd.).

2.2.2 Optimal Working Point

160 In a second step, we tune the applied voltage and the detector operating temperature by monitoring the energy resolution of 161 detected alpha particles. A resolution versus temperature and voltage matrix is first constructed with data from the TNT2 energy 162 spectra using the trapezoidal rising time and flat top duration known to give good resolution from previous detectors; The pole 163 zero compensation was optimized channel by channel once and for all at the beginning of the characterization.

The Fig. 7 illustrates such evolution of the measured energy resolution (all energy resolutions in this article will be given in FWHM). For this detector, the best resolution is obtained at 80V and -20 °C. However, this graph does not show the error bars on the measured resolution: typically below 0.5 keV on the FWHM. Thus, once the applied voltage and temperature reach the plateau (above 80 V and below -5 °C), every point is expected to behave the same regarding the measured resolution within the error bars. However, the evolution of the energy resolution is quite sensitive to change of both voltage and temperature at the edges of this operating region. Therefore, a slightly higher applied voltage was preferred and enabled higher operating temperature. The working point was then set at 90 V and -15 °C for this detector.

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Once this working point is determined, we accumulated a significant number of preamplifier signal traces (~10000 per alpha line) in order to get a statistical data set for the full detector. An offline analysis code processed the recorded signal traces and was used to optimize the Jordanov [9] parameters to obtain the best energy resolution using a 100 ns increments. Fig. 8 shows the resolution for the 8.8 MeV ²¹²Po alpha line as a function of the trapezoidal integration time and flat top duration. The best resolution is obtained for a quasi-triangular configuration: 4.8 μs integration time and 350-400 ns of flat top duration for this particular macropixel.

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For flat top durations of 0 to 200 ns (cf. Fig. 8, corresponding to quasi-triangular shape) a drastic increase in the measured FWHM was observed as expected. This width corresponds to the signal rise time of ~100-150 ns. This effect has already been observed and discussed in [10]. Fig. 8 corresponds to a typical macropixel within the average resolution measured. Such an optimization has been performed for each macropixel within the same run, since our data file contains traces from the full detector. This allows us to determine the optimal parameters over the whole detector. According to this measurement, the optimal working point is for a rise time of 5 µs and a flat top time of 0.4 µs for this particular detector.

2.3 Macropixel Resolution

Macropixels were characterized with a mixed ²¹²Po/²¹²Bi/²¹²Po alpha emitting source obtained from the deposition of ²¹⁶Po onto a gold foil during the decay of the ²³²Th through the gaseous alpha decay of the ²²⁰Rn to avoid any source thickness effect on the alpha energy. A typical ²¹²Pb alpha spectrum can be seen in Fig. 9 with very good resolution within the whole alpha energy range. This spectrum was obtained at (90V,-15 °C) following the optimization of the Jordanov parameters, with an energy resolution of 13.66 keV for the 8.783 MeV ²¹²Po line. The fitting function used during this work is the derived from the SAMPO gamma-ray program, which is a modified Gaussian distribution [11].

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205 Resolutions are reported in Fig. 10 as a function of the capacitance of the macropixel for the two selected reduced pixel 206 configurations. The average value of energy resolutions obtained is about 14.7 keV over the full detector for the ²¹²Po 8.783 207 MeV alpha line. This figure shows a slight increase in the measured average FWHM with capacitance of about 0.6-0.7 keV on 208 average from the low capacitance macropixel to the high capacitance ones (from 14.4 keV to 15-15.1 keV). The expected 209 contribution of the intrinsic noise of the preamplifier is 3.8 e-(RMS)/pF, thus leading to a 3.1 keV contribution to the FWHM at 210 95 pF and 4.7 keV at 146 pF. Moreover, the noise contributions add in quadrature. Thus, if we take the 14.4 keV FWHM at 95 211 pF, the "intrinsic" FWHM will be 14.1 keV leading to a 14.9 keV at 146 pF just based on the contribution from the preamplifier 212 which is in good agreement with the measurements.

213 In addition, all the FWHM values measured during these characterizations are within the specifications of the project, with 214 every single macropixel below the 20 keV upper limit.

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2.4 Determination of the Diode Orientation

The tunnel detectors collect electrons from internal conversion electrons (ICE) as well as escaping alpha particles from decays occurring within DSSD detector. The energy range of these particles is in principle significantly different, enabling a selection with a single energy deposit threshold. Unfortunately, due to the energy deposition in the DSSD while escaping and the dead layer effect in the Side detector, the residual energy from alpha particles can be pushed down to the energy range of ICE. This overlapping is particularly damaging for energies below 300 keV. Fortunately, the range of both particles in silicon differs, thus inducing noticeable difference in collection timings even with similar number of charges to be collected. The main key points in the discrimination between the electron signals and escaped alpha signals when they have similar amplitude in a Side detector (100-300 keV energy range) is the optimization of the physical characteristic of the pulse shape of the signal. In that sense, a correct choice of the orientation of the silicon detector with respect to the alpha emission (decay of implanted heavy ions in DSSD) is mandatory. In order to investigate this point, our Strippy-pad detectors were designed with zero dead layer on both junction and ohmic

In order to investigate this point, our Strippy-pad detectors were designed with zero dead layer on both junction and ohmic sides. Fig. 11 illustrates in that sense the difference of pulse shapes observed for detection of alpha particles on the ohmic side (blue) and on the junction side (red). A difference of rise-time of about 50 ns is observed, as expected. This is due to the difference in mobility of the charge carrier inside the silicon, as well as the very small range of alpha particles in silicon.

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As alpha particles are being stopped in the first few micrometers of the detector, the charge collection time will mainly be driven by one charge carrier, the opposite one being created very close to its collecting electrode. Therefore, only one type of charge carrier will drive the rise-time of the signal. Indeed, in the case of the junction-side detection the signal rise-time is shorter, as the signal will be mainly driven by the electron, compared to the ohmic implantation where the signal is due to the motion of holes through the entire thickness of the detector (Fig. 11).

We can confirm the difference in rise-time for the alpha particles: as predicted, the one on the ohmic side is slower. Indeed, the electrons have a deeper implantation depth in the silicon (~50 µm to a few hundreds of micrometers depending on their energy) than the alpha particle (few micrometers only). Thus, by using the slower signal for the alpha particle, we should have a better chance to discriminate between the two signals in the low-energy range (100-300 keV). This observation in the difference in mobility for holes and electrons inside silicon firmly points out the orientation of our detector inside the silicon box.

Moreover, the fastest collection time in the detector is not at the surface of the junction side where the holes are collected immediately, but slightly deeper. In our case (90 V and -10 °C), the fastest collection is about 120 µm from the Junction side, where electrons and holes have the same collection time. Thus, if the junction side was selected to face the DSSD, the difference between alpha and electron signal of low energy would have been less than 5 ns. But, in the case of the ohmic side, this difference should be around 10-20 ns due to the slow collection time of the holes.

The optimal orientation was determined by setting the inner face inside our 3D assembly to the ohmic side, which that should give the best potential discrimination properties (slower signal) between alpha particles and electrons. This decision triggered the development of the hybrid frame for the final assembly (Fig. 1.); the design of the frame and double-sided PCB depended on the choice of orientation.

257 **CONCLUSION:**

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Two new generation silicon detectors for the S3-SIRIUS focal plane Side detectors were characterized during this work. These detectors have 64 pixels on the ohmic side and a wide pixel on the junction side. A macropixel grouping of the 64 individual pixels was used to determine two 24-channel configurations for the collaboration, while maintaining the possibility to operate in single pixel mode in the future. Very good alpha energy resolution was achieved with overall resolutions (FWHM) better than 18 keV for 8.8 MeV alpha line and the best macropixels as low as 14.8 keV. These innovative detectors largely meet the SIRIUS collaboration specifications. They will be used for the physics program of S3.

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11 (98 pF)		0 (98 pF)		23 (99 pF)		12 (98 pF)	
0	10	1	2	22 (96 pF)		13 (95 pF)	
9 (93 pF)	(95 pF)	(94 pF)	2 (94 pF)				
				21 (132 pF)	20 (132 pF)	15 (132 pF)	14 (130 pF)
<mark>8</mark> (125 pF)	7 (127 pF)	4 (127 pF)	3 (126 pF)				
6				19 (112 pF)	18 (113 pF)	17 (113 pF)	16 (111 pF)
(145 pF)		ر (145 pF)					



2 pixels					
2 pixels					
3 pixels	3 pixels				
3 pixels	3 pixels				

(a)

(b)

Kapton connectors

Macro Pixel Reduction PCB Detector PCB

Silicon detector









a)

b)



