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**Broadening, shift and narrowing coefficients  
in the  $2\nu_4$  band of  $\text{NH}_3$  perturbed by  $\text{O}_2$ ,  $\text{N}_2$  and Air**

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## Abstract

Collisional broadening and shift coefficients as well as Dicke narrowing parameters have been measured at room temperature in the  $2\nu_4$  band of  $\text{NH}_3$  perturbed by  $\text{O}_2$  and  $\text{N}_2$  for 310 rovibrational lines with rotational quantum numbers  $1 \leq J \leq 11$  and  $0 \leq K \leq 11$ .

The experiments were performed at room temperature with ammonia mixtures in nitrogen and oxygen at different pressures using a high-resolution Fourier transform spectrometer and a mono-spectrum non-linear least squares fitting of Voigt and soft-collision Galatry profiles. Because of Dicke narrowing effect, departure from the Voigt profile is observed that allowed measurement of narrowing parameters for about 100 transitions of the  $2\nu_4$  band. The collisional broadening coefficients obtained with Galatry profile are larger by about 5% than those derived from Voigt profile. Average absolute accuracies of the measurements are estimated to be 5%, 10% and 14% for broadening, shift and narrowing coefficients, respectively.

Air-broadening coefficients were retrieved from the linear combination of  $\text{O}_2$ - and  $\text{N}_2$ -broadening coefficients. The results are in good agreement with previous experimental data available in the literature. The whole set of results is given as [supplementary material](#) for a possible use in spectroscopic databases.

## 1. Introduction

Accurate knowledge of ammonia line parameters is necessary for precise interpretation of high-resolution spectra of the Earth's atmosphere and other planetary spaces. Spectroscopic sensors for ammonia are needed for numerous applications including the environment, detection and control of industrial spaces [1], analysis of human breath [2], and recycling of waste-water [3]. Trace detection at ppm levels is often required for these goals [3]. In addition, several laser devices such as rapid-scan absorption [4], wavelength-modulation spectroscopy [5], cavity ring-down [2], and photoacoustic spectroscopy [6] are employed to detect  $\text{NH}_3$  traces. Also,  $\text{NH}_3$  is an important molecule for astrophysical applications; this interest arises because large quantities of gaseous ammonia have been detected in Jupiter [7] and Saturn [8] conferring to this gas to be a powerful spectroscopic probe of the physico-chemical conditions in the space of these planets. These applications require precise measurements of  $\text{NH}_3$  line spectroscopic parameters performed with appropriate experimental setup and the best line models. Furthermore, the  $\text{NH}_3$  absorption in Jupiter's atmosphere, especially, has been shown to be highly sensitive to the choice of spectral absorption line shape model [9].

Several experimental and theoretical studies of the line-widths and line shifts in infrared fundamental bands of  $\text{NH}_3$  perturbed by several foreign gases were published. However spectroscopy of this molecule is still requested, especially for the overtones and combination bands with the aims to improve spectroscopic databases [10–12] which are largely used for fundamental, planetary, and industrial applications. Among the previous studies, Fabian *et al.* measured  $\text{O}_2$ ,  $\text{N}_2$ , and Air broadenings for 176 rovibrational lines of  $\text{NH}_3$  in the  $\nu_2$  band using a high-resolution Fourier Transform Spectrometer (FTS) [13]. The broadenings due to the same perturbors have also been measured and analyzed in the  $\nu_4$  band, also, with FTS [14,15]. A prediction of these measurements was achieved with a semi classical model in which the inversion motion of  $\text{NH}_3$  is taken into account [16]. Using the same spectrometer, shift coefficients of rovibrational transitions of ammonia perturbed by  $\text{N}_2$  and  $\text{O}_2$  have been reported in Refs. [17] and [18] respectively.

The temperature dependence of pressure broadening in several branches of the  $\nu_4$  and  $2\nu_2$  bands of ammonia perturbed by  $\text{H}_2$  and  $\text{N}_2$  has been measured using a high-resolution FTS [19]. The authors deduced the temperature exponents of the two perturbors. Also, using an FTS, measurements of collision half-widths of  $\text{NH}_3\text{-X}$  ( $\text{X} = \text{H}_2, \text{N}_2, \text{O}_2$ ) systems are reported at three temperatures, lower than 296 K [20]. Air-broadening are deduced at 296 K and the variation of line-widths with rotational quantum numbers  $J$  and  $K$  was studied. In Ref. [21],

O<sub>2</sub>, N<sub>2</sub>, Air-broadening and shift parameters of NH<sub>3</sub> in the  $\nu_2$  band were measured and analyzed for five lines from spectra recorded with diode-laser spectrometer. Temperature effect of N<sub>2</sub>, O<sub>2</sub>, CO<sub>2</sub> and H<sub>2</sub>O broadening of Q-branch transitions near 10.4  $\mu\text{m}$  were studied in Ref. [22] using two quantum cascade lasers and a multi-line fitting Voigt profile procedure. N<sub>2</sub> and O<sub>2</sub> pressure broadening parameters for some transitions in the  $^R P(J,0)$  manifold of the  $\nu_1+\nu_3$  combination band were measured at room temperature using an external cavity tunable diode laser spectrometer [23].

All previous works neglect the Dicke and speed dependent effects [24–26]; and Voigt line-shapes were used to retrieve the line parameters.

In the 3  $\mu\text{m}$  region, pressure broadening coefficients in the  $\nu_1$  and  $\nu_3$  bands of NH<sub>3</sub> self-perturbed and perturbed by He, Ar, O<sub>2</sub>, N<sub>2</sub>, and H<sub>2</sub> were obtained by Markov and Pine [27] using spectra recorded with a difference frequency laser spectrometer. In the low and intermediate pressure range, the authors attributed their fit residual near the line center to the Dicke narrowing.

In Ref. [28] the same authors re-analyze the spectra of Ref. [27] and confront their new results to finer models, which showed clear evidence for Dicke narrowing and speed-dependent effects. A tunable quantum cascade laser was used by Owen *et al.* [29] to measure collisional broadening and Dicke narrowing coefficients of six transitions of NH<sub>3</sub> perturbed by N<sub>2</sub> and O<sub>2</sub> in the spectral region near 1103.46  $\text{cm}^{-1}$ . Their analyses were based on Voigt and soft-collision Galatry profiles. In our previous contributions in the  $2\nu_4$  band in the 3  $\mu\text{m}$  spectral region [30], we analyzed high-resolution FT spectra with a fitting procedure neglecting Dicke and speed-dependent effects since these effects are greatly masked by the strong long-range interactions of NH<sub>3</sub> with a large dipole moment. These effects, which appear in short range interactions near the line centers, arise mostly at low-pressure domains when collisional and Doppler broadening contributions are comparable.

In the present paper, the broadening and shifting coefficients of the same band perturbed by N<sub>2</sub> and O<sub>2</sub> were measured at room temperature using spectra recorded with a high-resolution FTS. We used a retrieval program based on Voigt and soft-collision Galatry profiles. The pressure broadening and shift coefficients as well as Dicke narrowing coefficients were retrieved. Air-broadening parameters were also determined assuming 79% N<sub>2</sub> and 21% O<sub>2</sub> for the Air composition.

Finally, we compare our results with previous measurements performed in other bands of NH<sub>3</sub> with or without accounting for speed dependence and/or Dicke narrowing [10,13-22,28] for some indications of these effects as well as the vibrational and tunneling dependence of the

broadening parameters. [Section 2](#) briefly describes the spectrometer and the experimental conditions, presents also the fitting procedures and the uncertainty analysis. The measurements of N<sub>2</sub>- and O<sub>2</sub>-broadening, shift and narrowing coefficients are described in [Section 3](#), which also describes the generation of NH<sub>3</sub>-Air broadening coefficients. Comparisons with previous measurements are reported in the same section. Conclusions and remarks are provided in [Section 4](#).

## **2. Experimental analysis**

### **2.1. Experimental details**

Nine spectra of ammonia diluted in oxygen and nitrogen were recorded at a stabilized room temperature of  $295 \pm 1\text{K}$  in the useful spectral domain from 3100 to 3600  $\text{cm}^{-1}$  using the Bruker IFS125HR FTS located at the LISA facility in Créteil (France), the same experimental setup that was used in Ref. [\[30\]](#). The instrument was equipped with a silicon carbide Globar source, a KBr/Ge beamsplitter, and a liquid nitrogen cooled InSb detector. No optical filter was used. A bandpass of 1900–3800  $\text{cm}^{-1}$  is limited by the cut off of the InSb detector (1900  $\text{cm}^{-1}$ ) and by the electronic filter used (3800  $\text{cm}^{-1}$ ). The FTS was continuously evacuated below  $3 \times 10^{-4}$  hPa by a turbomolecular pump to minimize absorption by atmospheric gases. The diameter of the entrance aperture of the spectrometer was set to 1.7 mm to maximize the flux of infrared radiation falling onto the InSb detector without saturation or loss of spectral resolution. Interferograms were recorded with a 40 kHz scanner frequency and a maximum optical path difference (MOPD) of 90 cm. According to the Bruker definition (resolution =  $0.9/\text{MOPD}$ ), this corresponds to a resolution of 0.01  $\text{cm}^{-1}$ . Note that the Doppler width is about 0.0096  $\text{cm}^{-1}$  at the frequency center of the studied region.

The <sup>14</sup>NH<sub>3</sub>, O<sub>2</sub> and N<sub>2</sub> samples were purchased from Sigma Aldrich with stated purities of 99%, 99.99% and 99.98%, respectively. No further sample purification was done. The absorption cell has a base path length of 80 cm and was adjusted for 28 transits in the present experiment, yielding an absorption path length of  $2244.9 \pm 1.1$  cm. The path length includes the distance between the surface of the field mirror and the windows of the cell ( $2 \times 2.45$  cm). The sample pressure in the cell was measured using calibrated MKS Baratron capacitance manometers models 628D (2 Torr full scale) and 627D (100 and 1000 Torr full scale), characterized by a stated reading accuracy of 0.12%. Considering the uncertainty arising from small variations of the pressure during the recording of the interferograms ( $\sim 0.35\%$ ), we estimated the measurement uncertainty on the pressure to be equal to 0.5%.

The following procedure was used for the measurements. A background spectrum was first recorded at a resolution of  $0.01\text{ cm}^{-1}$  while the cell was being continuously evacuated. The cell was then filled with  $\text{NH}_3$  at a given pressure, followed by the perturbing gas ( $\text{O}_2$  or  $\text{N}_2$ ) leading to a series of 9 total pressures. One  $\text{NH}_3/\text{O}_2$  or  $\text{NH}_3/\text{N}_2$  spectrum was recorded for each  $\text{NH}_3$  filling. Transmittance spectra were generated from the ratio of the sample spectra with the background spectrum. The detailed experimental condition is given in [Table 1](#). For the Fourier transform, a Mertz-phase correction with a  $1\text{ cm}^{-1}$  phase resolution, a zero-filling factor of 2 and no apodization (boxcar option) were applied to the averaged interferograms, using the procedure included in the Bruker software OPUS package [\[31,32\]](#). The peak-to-peak signal-to-noise ratio (S/N) in the ratioed spectra is around 800. The spectra were calibrated by matching the measured positions of 37 lines of residual  $\text{H}_2\text{O}$  in the FTS observed therein to reference wavenumbers available in HITRAN [\[10\]](#) with an RMS deviation of  $0.00018\text{ cm}^{-1}$ . The calibration factor was determined with the  $\text{H}_2\text{O}$  positions measured in the  $\text{NH}_3$  spectrum at lower pressure and then applied to all other spectra.

**Table 1:** Experimental conditions and pressures used to record the  $2\nu_4$  band of  $\text{NH}_3$  perturbed by  $\text{O}_2$  and  $\text{N}_2$ .

Experimental conditions					
Resolution ( $\text{cm}^{-1}$ )			0.01		
Maximum optical path difference (cm)			90		
Diaphragm diameter (mm)			1.7		
Focal distance (mm)			418		
Optical cell length (cm)			2244.9		
Useful spectral domain ( $\text{cm}^{-1}$ )			3100 - 3600		
Spectrum number and pressures					
#	$\text{NH}_3$ pressure (Torr)	$\text{N}_2$ pressure (Torr)	#	$\text{NH}_3$ pressure (Torr)	$\text{O}_2$ pressure (Torr)
S1	0.0809(4)	19.94(10)	S1	0.0803(4)	20.46(10)
S2	0.1624(8)	40.02(20)	S2	0.1604(8)	39.94(20)
S3	0.2420(12)	59.97(30)	S3	0.2431(12)	60.52(30)
S4	0.3243(16)	80.00(40)	S4	0.3216(16)	80.01(40)
S5	0.4005(20)	100.0(5)	S5	0.3975(20)	99.70(50)
S6	0.4802(24)	119.9(6)	S6	0.4828(24)	121.1(6)
S7	0.7990(40)	160.0(8)	S7	0.8003(40)	160.3(8)
S8	1.1981(60)	239.8(12)	S8	1.1969(60)	240.1(12)
S9	1.5032(75)	300.0(15)	S9	1.4950(75)	299.9(15)

## 2.2. Data analysis and fitting procedure

Generally, to measure the self-broadening for molecules with large dipole moment, the Dicke and speed-dependent effects can be neglected since they are masked by the strong long-range interactions between molecules. Here, since the ammonia perturbed by oxygen and nitrogen, are rather short range collisions, these effects should be taken into account since at low pressure range, the mean-free path of molecules decreases as the pressure increases to become

smaller than the light wavelength, implying a reduction of the Doppler width [24] and narrower profile (Dicke effect). Therefore, we used the following soft-collision Galatry profile [26] which is based on diffusion model.

$$G(\sigma - \sigma_0) = \frac{A}{\sqrt{\pi}} \operatorname{Re} \left[ \frac{1}{(1/2z) + y - ix} \times M \left( 1; 1 + \frac{1}{2z^2} + \frac{y - ix}{z}; \frac{1}{2z^2} \right) \right] \quad (1)$$

$M(\dots; \dots; \dots)$  is the hypergeometric function;  $x$  and  $y$  are the Doppler normalized shift and pressure broadening given by:

$$A = \frac{S\sqrt{\ln 2}}{\gamma_D \sqrt{\pi}}, \quad x = \sqrt{\ln 2} \frac{\sigma - \sigma_0}{\gamma_D}, \quad y = \sqrt{\ln 2} \frac{\gamma}{\gamma_D} \quad (2)$$

In these equations,  $S$  is the line intensity in  $\text{cm}^{-2}$ ,  $\sigma_0$  its wavenumber in  $\text{cm}^{-1}$  including the collisional shift,  $\gamma$  and  $\gamma_D$  in  $\text{cm}^{-1}$  are the collisional and Doppler line half-widths respectively. The normalized parameter  $z$  is related to the collisional narrowing  $\beta$  in  $\text{cm}^{-1}$  (also called the dynamical friction parameter) by:

$$z = \sqrt{\ln 2} \frac{\beta}{\gamma_D} \quad (3)$$

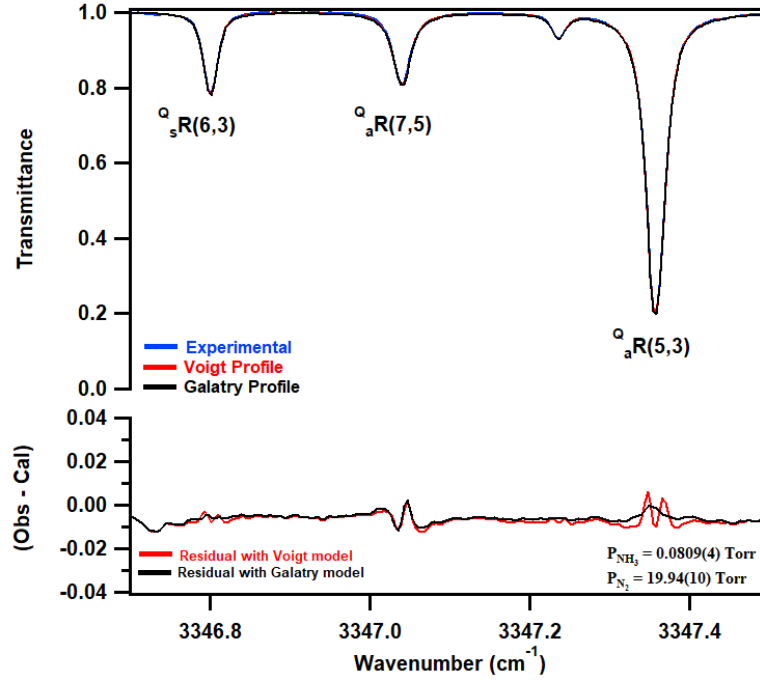
Note that when  $z = 0$ , the Galatry profile is reduced to the Voigt profile.

The recorded nine spectra were simulated by Voigt and soft-collision Galatry profiles using a mono-spectrum fitting procedure which allows to retrieve line parameters of  $2\nu_4$  band of  $\text{NH}_3$ . This procedure uses the WSpectra code [33] based on a non-linear least-square method in which line position, intensity, width, background, and the narrowing parameter were determined by adjusting a synthetic transmittance to minimize the residual between the observed and simulated profiles. Only the three first parameters are derived from the fit with Voigt profile. Note that in the two fitting procedures, the Doppler width was fixed to its calculated value.

Figure 1 illustrates an example of the measured line shape of the  $Q_sR(6,3)$ ,  $Q_aR(7,5)$  and  $Q_aR(5,3)$  transitions and the calculated profiles described either by the basic Voigt profile or by the Galatry model. The two profiles are quite similar and thus difficult to distinguish on the same plot. The residuals given at the bottom, resulting from the difference between the measured and the calculated profiles show an evident improvement in the fitting obtained



when the Dicke narrowing is taken into account through the Galatry profile. When the Voigt profile is used, one can observe a W-shape residuum characteristic of line narrowing. Examination of differences between the observed and calculated spectra shows that these residuals are generally less than 1%.



**Figure 1:** Experimental line shape of the  $Q_sR(6,3)$ ,  $Q_aR(7,5)$  and  $Q_aR(5,3)$  transitions in the  $2v_4$  band of  $NH_3$  perturbed by  $N_2$  recorded at  $P_{NH_3} = 0.0809$  and  $P_{N_2} = 19.94$  Torr, fitted with Galatry and Voigt profiles; (Obs-cal) shows the residuals using either Voigt or Galatry models.

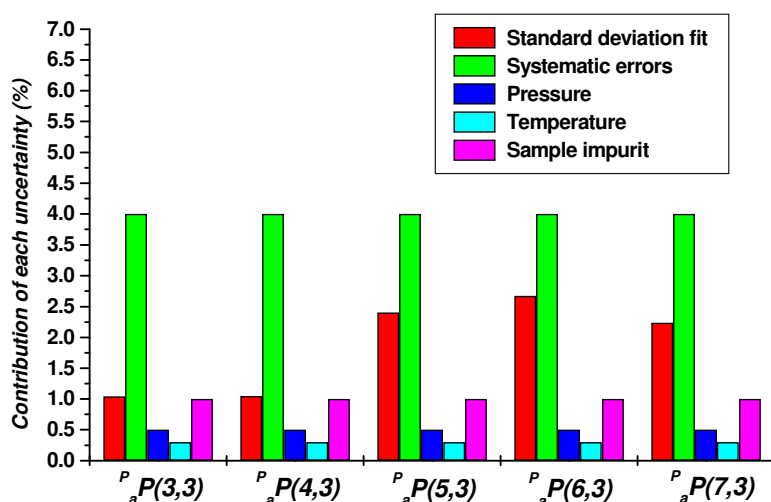
### 2.3. Uncertainty analysis

In the following, the measured values of line broadening and shift coefficient as well as narrowing coefficients are presented with their absolute accuracies estimated from uncertainties on all physical parameters, standard deviation of the fit and contributions from possible systematic errors. Assuming these parameters are uncorrelated quantities, the estimation of each transition accuracy can be calculated as [34]:

$$\varepsilon = \sqrt{\varepsilon_{si}^2 + \varepsilon_t^2 + \varepsilon_p^2 + \varepsilon_{fit}^2 + \varepsilon_{sys}^2} \quad (4)$$

In this equation  $\varepsilon_{si} = 1\%$ ,  $\varepsilon_t = 0.34\%$  and  $\varepsilon_p = 0.5\%$  are the uncertainties associated with the sample purity, the temperature, and the pressure, respectively.  $\varepsilon_{fit}$  and  $\varepsilon_{sys}$  are the uncertainties of the standard deviation of the fit and the systematic errors. The dominant contributions to

systematic errors arise from the location of the full-scale photometric level, channeling, as well as electronic and detector nonlinearities [34,35]. A realistic assessment of systematic errors is not an easy task and probably strongly depends on the setup used. Our present estimate is based upon the arguments developed in Ref. [34] using instruments comparable to ours. We retained 4% as the contribution of systematic errors on retrieved parameters. It is an arbitrary, but conservative value. These various uncertainties evaluated for the broadening coefficients are given in Figure 2 for 5 selected lines, representative of the 310 measured lines, i.e.  $^P_aP(3,3)$ ,  $^P_aP(4,3)$ ,  $^P_aP(5,3)$ ,  $^P_aP(6,3)$  and  $^P_aP(7,3)$ . We can observe that the systematic errors and the standard deviation from the linear fit are the main sources of errors. Considering these various sources of errors, we estimate that, on average, accuracies of the measurements are 5% for broadening coefficients, 10% for shift coefficients and 14% for narrowing coefficients.



**Figure 2:** Comparison between various sources of relative uncertainty on the measured  $O_2$  -broadening coefficients for 5 selected lines:  $^P_aP(3,3)$ ,  $^P_aP(4,3)$ ,  $^P_aP(5,3)$ ,  $^P_aP(6,3)$  and  $^P_aP(7,3)$ .

### 3. Results and discussion

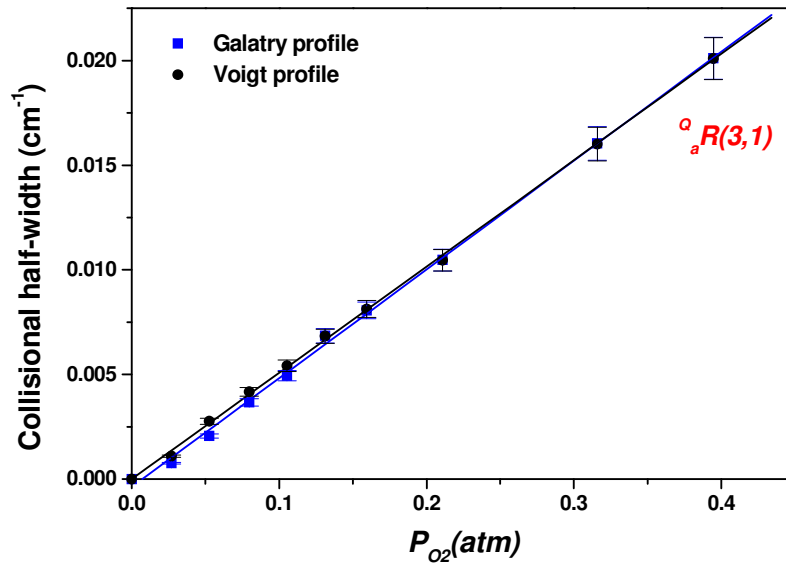
#### 3.1. Broadening coefficients

In collisions between ammonia and the homopolar molecules, nitrogen and oxygen, dipole–quadrupole interaction becomes somewhat dominant, but significantly smaller than the strong long-range interactions of  $NH_3$  self-collision. Then the self-broadening must be considered while determining the foreign-gas-broadened line-widths.

Although the spectra of the NH<sub>3</sub>-O<sub>2</sub> and NH<sub>3</sub>-N<sub>2</sub> mixtures were recorded with at most 0.5% of NH<sub>3</sub>, we subtracted the self-broadening contributions to deduce, at the best, the N<sub>2</sub>- and O<sub>2</sub>-coefficients, according to the equation:

$$\gamma = \gamma_{0O_2/N_2} P_{O_2/N_2} = \gamma_{total} - \gamma_{0Self} P_{NH_3} \quad (5)$$

where  $\gamma$  is the N<sub>2</sub>- or O<sub>2</sub>-collisional half-width in cm<sup>-1</sup>,  $\gamma_{total}$  is the total collisional half-width in cm<sup>-1</sup>, and  $P_{NH_3}$  and  $P_{O_2/N_2}$  are the NH<sub>3</sub> and O<sub>2</sub> or N<sub>2</sub> partial pressures in atm respectively.  $\gamma_{0O_2/N_2}$  is the N<sub>2</sub>- or O<sub>2</sub>-broadening coefficients and  $\gamma_{0Self}$  is the self-broadening coefficient taken from Ref. [30]. The evolution with the O<sub>2</sub> pressure of  $\gamma$  measured for the  $Q_aR(3,1)$  line obtained from the Voigt and Galatry fits is shown in Figure 3.



**Figure 3:** Collisional half-widths of the  $Q_aR(3,1)$  transition versus the O<sub>2</sub> pressure. Data are derived from Voigt (black circle symbols) and Galatry (blue square symbols) profiles. The linear fits are used to retrieve the O<sub>2</sub>-broadening coefficient in units of cm<sup>-1</sup>.atm<sup>-1</sup>.

This figure shows the linear dependence of the collisional half-widths with pressure for the two models. The O<sub>2</sub>-broadening coefficients (in cm<sup>-1</sup>.atm<sup>-1</sup>) were derived from the slopes of these lines. The results of line broadening and shift coefficients as well as narrowing coefficients, all in the unit of cm<sup>-1</sup>.atm<sup>-1</sup>, along with their absolute uncertainties are given in the [supplementary material](#) listing 310 lines of the 2ν<sub>4</sub> band. [Table 2](#) illustrates a sample of these results for  $P_sP(5,K)$  manifold. Three remarks can be addressed about the trends of broadening coefficients:

- they indicate no systematic inversion symmetry dependence;
- they increase with  $K$  for a given  $J$  and decrease with  $J$  for a fixed  $K$ ;
- in most cases, values retrieved from Galatry profile are relatively higher than the ones derived from Voigt profile by about 5%.

The first two remarks are the same as those observed for self-broadening in Ref. [16,19,20,30].

The measured values range from 0.0458 to 0.1812  $\text{cm}^{-1}.\text{atm}^{-1}$  for  $\text{N}_2$  perturber and from 0.0318 to 0.0753  $\text{cm}^{-1}.\text{atm}^{-1}$  for  $\text{O}_2$  perturber.

**Table 2:** Measured  $\text{N}_2$ -,  $\text{O}_2$ - and Air-broadening, narrowing and shift coefficients in  $\text{cm}^{-1}.\text{atm}^{-1}$  at 295 K for five transitions of the  $^P_sP(5,K)$  manifold retrieved with Voigt profile (VP) and Galatry profile (GP). All the uncertainties are calculated using Eq. (4).

Transition	Position <sup>(a)</sup>	Profile	$\gamma_0$ ( $\text{cm}^{-1}.\text{atm}^{-1}$ )		$\gamma_{0\text{Air}}$ ( $\text{cm}^{-1}.\text{atm}^{-1}$ )	$\beta_0$ ( $\text{cm}^{-1}.\text{atm}^{-1}$ ) <sup>(b)</sup>		$\delta_0$ ( $\text{cm}^{-1}.\text{atm}^{-1}$ ) <sup>(c)</sup>	
			$\text{NH}_3\text{-O}_2$	$\text{NH}_3\text{-N}_2$	$\text{NH}_3\text{-Air}$	$\text{NH}_3\text{-O}_2$	$\text{NH}_3\text{-N}_2$	$\text{NH}_3\text{-O}_2$	$\text{NH}_3\text{-N}_2$
$^P_sP(5,1)$	3154.334810	VP	$0.0491 \pm 0.0032$	$0.0799 \pm 0.0036$	$0.0815 \pm 0.0043$	$0.0196 \pm 0.0047$	$0.0641 \pm 0.0166$	$-0.0046 \pm 0.0004$	$-0.0081 \pm 0.0008$
		GP	$0.0525 \pm 0.0024$	$0.0892 \pm 0.0048$					
$^P_sP(5,2)$	3166.726590	VP	$0.0511 \pm 0.0039$	$0.0863 \pm 0.0038$	$0.0927 \pm 0.0041$	$0.0109 \pm 0.0018$	$0.0324 \pm 0.0044$	$-0.0052 \pm 0.0004$	$-0.0051 \pm 0.0006$
		GP	$0.0519 \pm 0.0024$	$0.1035 \pm 0.0045$					
$^P_sP(5,3)$	3177.714610	VP	$0.0537 \pm 0.0032$	$0.0914 \pm 0.0039$	$0.0878 \pm 0.0038$	$0.0147 \pm 0.0017$	$0.0351 \pm 0.0022$	$-0.0044 \pm 0.0004$	
		GP	$0.0541 \pm 0.0029$	$0.0967 \pm 0.0040$					
$^P_sP(5,4)$	3188.554070	VP	$0.0551 \pm 0.0030$	$0.0954 \pm 0.0041$	$0.0941 \pm 0.0042$	$0.0132 \pm 0.0012$	$0.0342 \pm 0.0049$	$-0.0050 \pm 0.0003$	$-0.0058 \pm 0.0005$
		GP	$0.0558 \pm 0.0026$	$0.1043 \pm 0.0046$					
$^P_sP(5,5)$	3195.033800	VP	$0.0543 \pm 0.0029$	$0.0912 \pm 0.0039$	$0.0938 \pm 0.0043$	$0.0110 \pm 0.0022$	$0.0236 \pm 0.0024$	$-0.0058 \pm 0.0005$	$-0.0043 \pm 0.0003$
		GP	$0.0552 \pm 0.0027$	$0.1040 \pm 0.0047$					

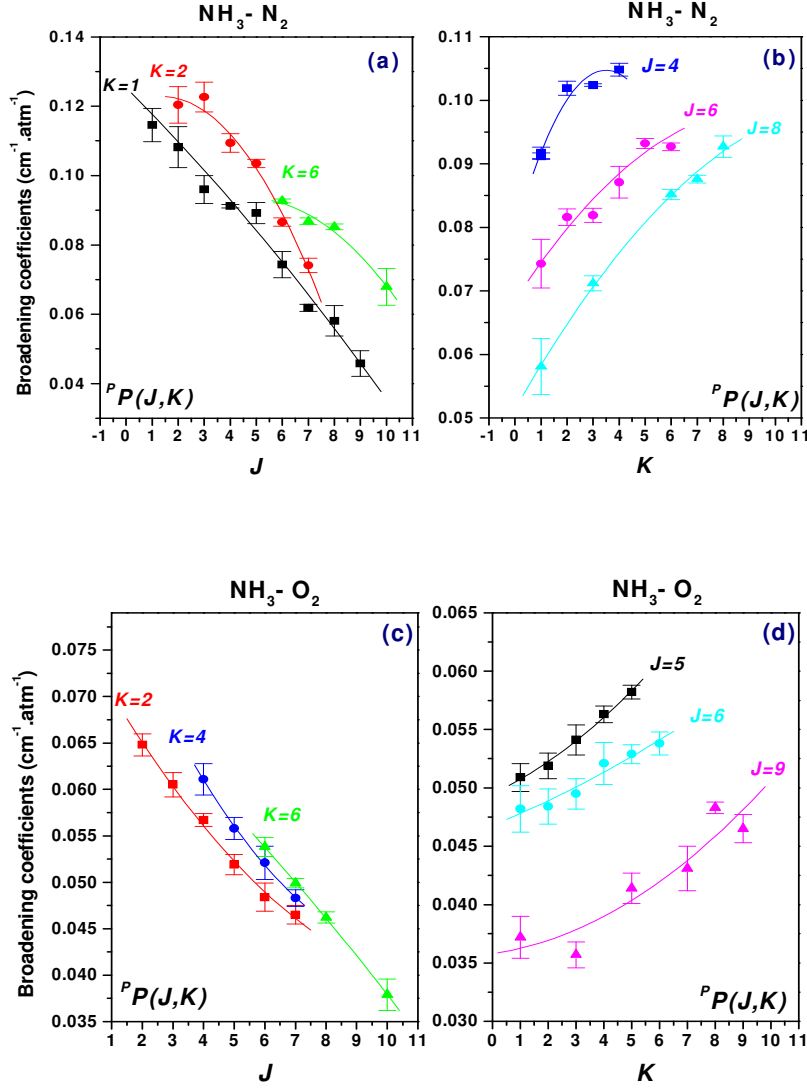
(a) Line positions are from HITRAN.<sup>10</sup>

(b)  $\text{O}_2$ - and  $\text{N}_2$ -narrowing coefficients obtained with Galatry model.

(c)  $\text{O}_2$ - and  $\text{N}_2$ -shift coefficients obtained with Galatry model.

The rotational trends of the broadening coefficients of the two foreign gases are illustrated in Figure 4 for the  $^P P$  sub-branch. The variation of these coefficients with  $J$  for each  $K$  is shown in Figures 4a and 4c, whereas, we show the variation with  $K$  for each  $J$  in Figures 4b and 4d. The plots for  $\text{N}_2$ -broadening (Figures 4a, 4b) and Air-broadening (not shown here) appear to be convex, while they are roughly concave for  $\text{O}_2$ -broadening (Figures 4c, 4d).

Since this work reports the data of  $\text{N}_2$ - and  $\text{O}_2$ -broadening coefficients in the  $2\nu_4$  band for the first time, comparison of values listed in Table 2 and supplementary material will be done with those performed in other bands in Refs. [14,15,20-22,28].



**Figure 4:** Variation of N<sub>2</sub>- and O<sub>2</sub>-broadening coefficients (cm<sup>-1</sup>.atm<sup>-1</sup>) in 2ν<sub>4</sub> band of NH<sub>3</sub> with J and K in the P sub-branch at 295 K.

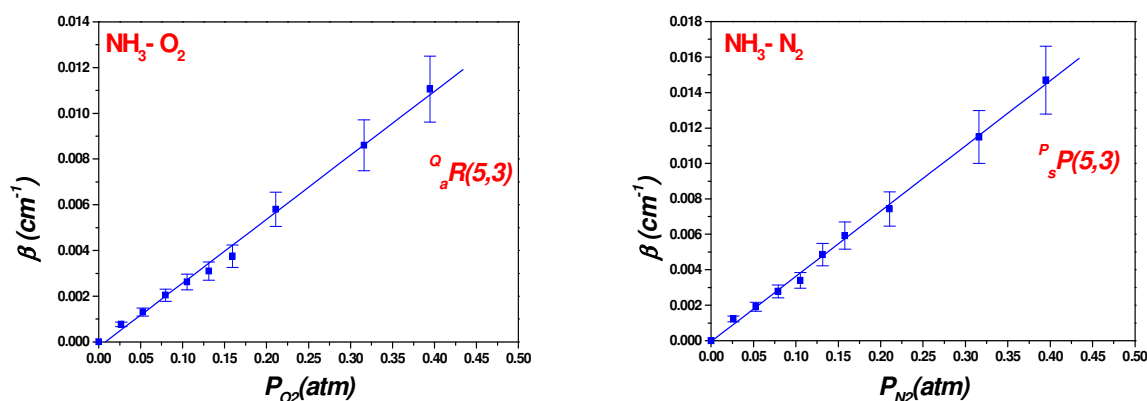
For broadening coefficients, the values retrieved from the Galatry profile, differ by about 2% and 4% for O<sub>2</sub> and N<sub>2</sub> perturbers respectively from the values taken from Pine *et al.* [28] for 63 lines studied in common. These data were obtained mostly in the Q- and R-branch of the ν<sub>1</sub> band. Also, we compared our data with those in the ν<sub>2</sub> band reported in Ref. [20] also obtained with a FTS using the same spectral resolution of 0.01 cm<sup>-1</sup> as the present experiment. Our data agreed within 5% for most of the lines reported in this reference. In the same band, O<sub>2</sub>- and N<sub>2</sub>-broadening coefficients in Ref. [21] agree with our data within 2% and 5% for O<sub>2</sub> and N<sub>2</sub> perturbers respectively. When the present values were compared with those reported in Ref. [22] for N<sub>2</sub>-broadening in the Q-branch of the ν<sub>2</sub> band, the agreement is less satisfactory with a difference of about 10%. This difference is reduced to 4% if we compare

our O<sub>2</sub>-broadening data with those reported in the same reference. Moreover, our results are in excellent agreement with the values obtained in Refs. [14,15]; the average difference is about 2% for O<sub>2</sub> and N<sub>2</sub> perturbers.

### 3.2. Collisional narrowing parameters

The Dicke parameters,  $\beta$ , are measured within the spectra adjusted by the Galatry profile that includes the narrowing confinement parameter arising from velocity changing collisions. In our fitting procedure, these effects occur at low pressure and we observed a convergence of the Galatry fit with positive values of this parameter for all transitions for which this narrowing is observed. Note that in the observed line profiles these effects can occur themselves as a shift of amplitude transmittance at the line center when a Voigt profile is used.

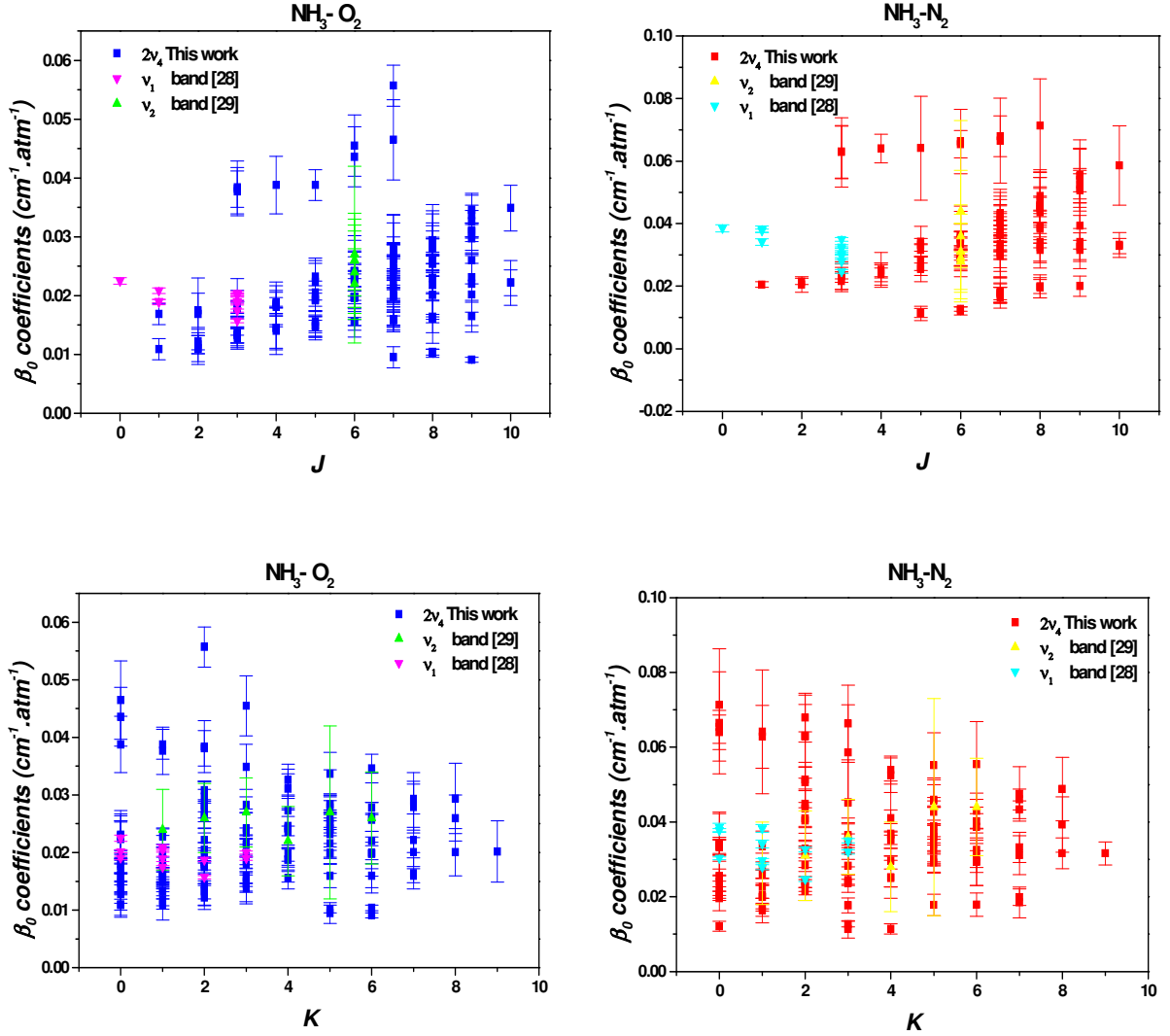
Collisional narrowing coefficients  $\beta_0$  for NH<sub>3</sub> perturbed by O<sub>2</sub> and N<sub>2</sub> were derived from the plots of the parameter  $\beta$  as a function of the gas pressure. The slope of the straight-line gives the coefficient  $\beta_0$ . A typical plots of  $\beta$  versus  $P$  are presented in Figure 5 for the  $^Q_aR(5,3)$  and  $^P_sP(5,3)$  transitions which show a good linearity of  $\beta$  with the pressure.



**Figure 5:** Narrowing parameter  $\beta$  of the  $^Q_aR(5,3)$  and  $^P_sP(5,3)$  transitions in the  $2\nu_4$  band of NH<sub>3</sub> versus the pressure retrieved with the Galatry profile. The corresponding  $\beta_0$  coefficient is deduced from the slope of the linear fit.

Numerous weak and/or blended lines illustrate significant dispersion of points in the graph of  $\beta$  versus the pressure, so it does not allow to determine any slope. Consequently, the measurement of the narrowing coefficients  $\beta_0$  was only possible for 120 transitions for O<sub>2</sub> perturber and 100 transitions for N<sub>2</sub> one. The values of these coefficients with their uncertainties determined by Eq. (4) are listed in Table 2 and in the supplementary material,

and are shown in Figure 6 versus  $K$  for all values of  $J$ , illustrating an agreement with Refs. [28] and [29], which report data for the  $\nu_1$  and  $\nu_2$  bands respectively. The average values are  $0.0226$  and  $0.0357 \text{ cm}^{-1} \cdot \text{atm}^{-1}$  for  $\text{O}_2$  and  $\text{N}_2$  respectively. The corresponding accuracies are about 14% for the two perturbers.



**Figure 6:** Narrowing coefficients  $\beta_0$  in  $\text{cm}^{-1} \cdot \text{atm}^{-1}$  in the  $2\nu_4$  band of  $\text{NH}_3$  at 295 K obtained with Galatry model as function of quantum numbers  $J$  and  $K$ . Measured values of Pine et al. [28] and Owen et al. [29] are also plotted for comparison.

It is difficult to measure the narrowing parameters with high precision since they are generally small. Two reasons can be addressed; firstly, it may be related to the strong correlation between the broadening and narrowing parameters [36]; secondly, it concerns the difficulty of observing fine effects such as Dicke collisional narrowing and speed-dependent using Fourier transform spectroscopy.

The measured collisional narrowing coefficients can be compared to those measured in the  $\nu_1$  band in Ref. [28] using the Galatry model. This reference reported an average values of  $0.0226 \pm 0.0029 \text{ cm}^{-1}.\text{atm}^{-1}$  for  $\text{NH}_3\text{-O}_2$  mixture and  $0.0357 \pm 0.0051 \text{ cm}^{-1}.\text{atm}^{-1}$  for  $\text{NH}_3\text{-N}_2$  one. Owen *et al.* [29] measured line narrowing coefficients  $\beta_0$  when fitting ro-vibrational lines of the  $sR(6,K)$  manifold in the  $\nu_2$  band of  $\text{NH}_3$  mixed with  $\text{O}_2$  and  $\text{N}_2$ . The average values on studied transitions are listed in Table 3 which illustrates a reasonable agreement between narrowing parameters measured in different wavenumber regions in the  $\nu_1$  [28],  $\nu_2$  [29] and  $2\nu_4$  bands. The largest difference is observed for  $\text{N}_2$  with a value of 12% between the present work and Ref. [28] for  $\text{O}_2$  the agreement is better with a difference of about 7%. When we compare with those of the  $\nu_2$  band [29], our value for  $\text{N}_2$  is in perfect agreement, but we found a difference of about 11% for  $\text{O}_2$ . These differences appear to be large but are in the margin of errors of Ref. [29], which are about 41% and 32% for  $\text{N}_2$  and  $\text{O}_2$  respectively; those of the present work are 14% and 13% for  $\text{N}_2$  and  $\text{O}_2$  respectively. The two last uncertainties account for all sources of errors, while that of Ref. [28], 1% for the two perturbers, is equal to the standard deviation. Considering these uncertainties, we can state a reasonable agreement between these works.

**Table 3:** Average narrowing coefficients  $\beta_0$  in  $\text{cm}^{-1}.\text{atm}^{-1}$  measured for  $\text{NH}_3$  perturbed by  $\text{N}_2$  and  $\text{O}_2$  with the Galatry model. Data are compared with measurements of Refs. [28,29].

	$\beta_0 \text{ NH}_3\text{-N}_2$	$\beta_0 \text{ NH}_3\text{-O}_2$
This work	$0.0357 \pm 0.0051$	$0.0226 \pm 0.0029$
Pine <i>et al.</i> <sup>28</sup>	$0.0412 \pm 0.0003$	$0.0243 \pm 0.0002$
Owen <i>et al.</i> <sup>29</sup>	$0.0353 \pm 0.0145$	$0.0253 \pm 0.0080$

Our values of  $\beta_0$  allow to calculate the optical diffusion constants  $D_{diff}$  of the  $\text{NH}_3\text{-O}_2$  and  $\text{NH}_3\text{-N}_2$  mixtures using the formula:

$$D_{diff} = \frac{kT}{2\pi cM \beta_0} \quad (6)$$

where  $k = 1.38064852 \times 10^{-23} \text{ J.K}^{-1}$  is the Boltzmann constant,  $T$  is the temperature of the mixture,  $c$  is the speed of light and  $M$  is the mass of the absorbing molecule. We found  $0.218 \text{ cm}^2.\text{s}^{-1}$  and  $0.347 \text{ cm}^2.\text{s}^{-1}$  for  $\text{NH}_3\text{-O}_2$  and  $\text{NH}_3\text{-N}_2$  binary mixture, respectively. When comparing these values with Ref. [37], we found an excellent agreement for  $\text{N}_2$  with a difference of 1%, but a large discrepancy ( $\approx 34\%$ ) is observed for  $\text{O}_2$ . This reference reported



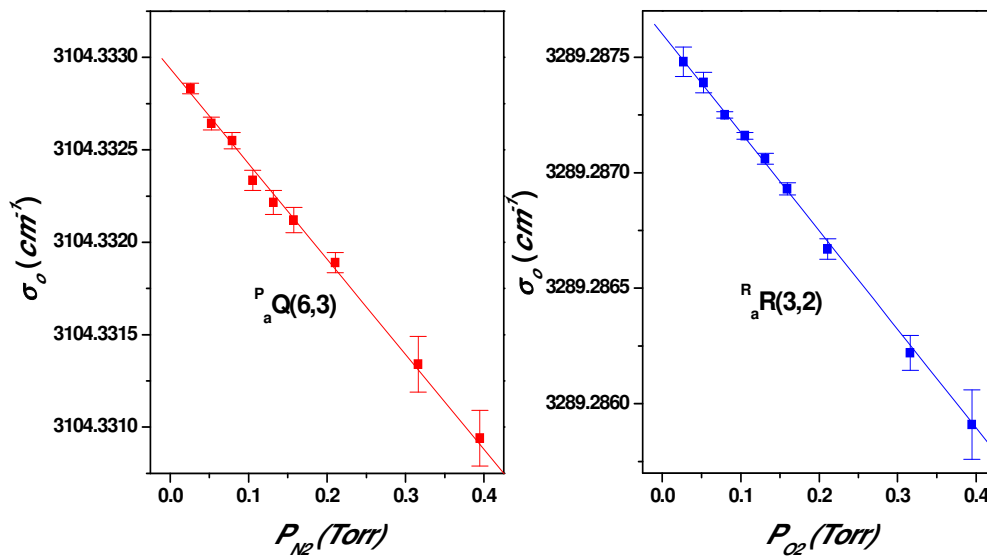
the values  $0.220 \pm 0.005 \text{ cm}^2.\text{s}^{-1}$  and  $0.227 \pm 0.004 \text{ cm}^2.\text{s}^{-1}$  for  $\text{NH}_3\text{-O}_2$  and  $\text{NH}_3\text{-N}_2$  binary mixture, respectively.

### 3.3. $\text{O}_2$ and $\text{N}_2$ pressure-induced shift coefficients

The pressure shifts are generally much smaller than the widths and are difficult to measure with high accuracy. The pressure-induced shift coefficients  $\delta_0$  in  $\text{cm}^{-1}.\text{atm}^{-1}$  of  $\text{O}_2$ - and  $\text{N}_2$ -perturber were determined using the following expression:

$$\sigma_0 = \sigma_{00} + \delta_0 P_{N_2/O_2} \quad (7)$$

where  $\sigma_{00}$ , in  $\text{cm}^{-1}$ , is the unperturbed wavenumber and  $\sigma_0$  ( $\text{cm}^{-1}$ ) the line position corresponding to a given total pressure. A typical plot of the collisional wavenumber  $\sigma_0$  as a function of the pressure  $P_{N_2/O_2}$  is shown in Figure 7.

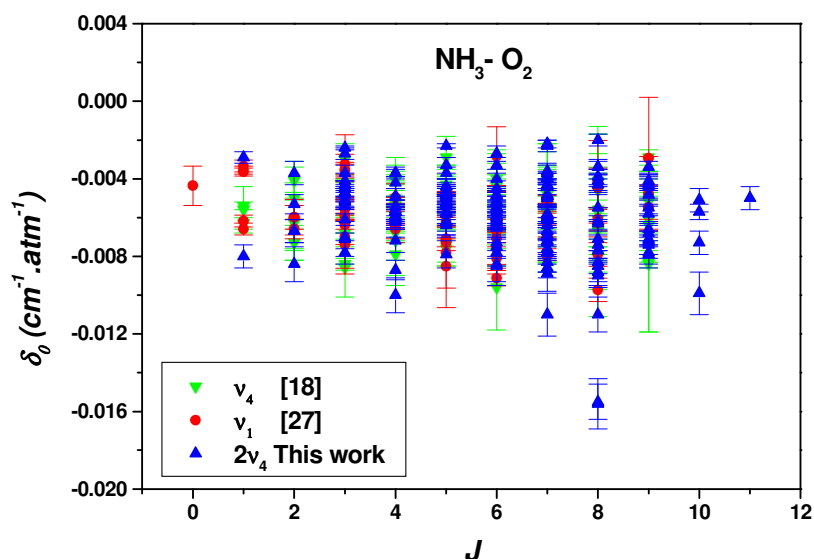


**Figure 7:** Dependence of measured wavenumber  $\sigma_0$  with the  $\text{N}_2$  and  $\text{O}_2$  partial pressures  $P_{N_2}$  and  $P_{O_2}$  for the  $P_a Q(6,3)$  and  $R_a R(3,2)$  lines. The collisional line shift coefficient  $\delta_0$  is derived from the slope of the best fit line. (■) correspond to experimental data and solid line to the linear regression.

The straight line obtained from the fitting illustrates the linearity of the measured wavenumber against pressure for the  $P_a Q(6,3)$  and  $R_a R(3,2)$  lines in the  $2\nu_4$  band of  $\text{NH}_3$ . Note that due to the low pressure values of  $\text{NH}_3$  ( $\leq 0.5\%$  of the total pressure), the self-shifts were neglected.

The N<sub>2</sub>- and O<sub>2</sub>-shift coefficients  $\delta_0$ , in cm<sup>-1</sup>.atm<sup>-1</sup>, are derived from the slope of these straight lines. The whole set of the data shift for the two buffer gases are listed in the [supplementary material](#), and few transitions are collected in [Table 2](#), together with their absolute uncertainty estimated using Eq. (4).

[Figure 8](#) shows a comparison of our experimental shift coefficients  $\delta_0$  for the 2 $\nu_4$  band perturbed by O<sub>2</sub> with the values obtained for the  $\nu_1$  band of NH<sub>3</sub> in Ref. [27] and those of the  $\nu_4$  band in Ref. [18]. This figure and [supplementary material](#) illustrate an excellent agreement, with an average difference of about 3% with Ref. [27] and about 2% with Ref. [18]. For the NH<sub>3</sub>-N<sub>2</sub> system, we obtain an average difference of 6% with Ref. [27], and 8% with Ref. [17]. All the O<sub>2</sub> shift coefficients are negative illustrating a decrease of line position with pressure. They vary between -0.0020 and -0.0267 cm<sup>-1</sup>.atm<sup>-1</sup> with an average value of -0.0061 cm<sup>-1</sup>.atm<sup>-1</sup>. For N<sub>2</sub>, 12 transitions out of 155 lines have positive shifts; they range from -0.0551 and 0.0303 cm<sup>-1</sup>.atm<sup>-1</sup> with a mean value of -0.0059 cm<sup>-1</sup>.atm<sup>-1</sup>. The same trend of signs is observed by Pine *et al.* in Ref. [27] and Dhib *et al.* in Ref. [17]. These shifts do not exhibit any distinguishable  $J$  and  $K$  dependences. On the other hand, we can also observe a satisfactory agreement between our values and those of Refs. [27] and [17,18] for the  $\nu_1$  and  $\nu_4$  band respectively, indicating that they are insensitive to vibrational excitation.



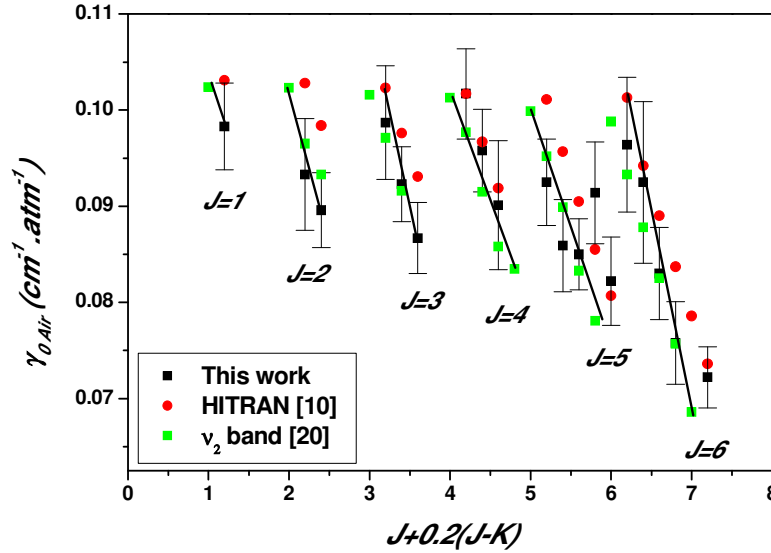
**Figure 8:** Comparison of our values of O<sub>2</sub> -shift coefficients  $\delta_0$  obtained in the 2 $\nu_4$  band of NH<sub>3</sub> with results obtained by Pine *et al.* in the  $\nu_1$  band [27] and Aroui *et al.* in the  $\nu_4$  band [18].

### 3.4. Air-broadening coefficients

Since  $N_2$  and  $O_2$  are the major constituents of the Earth's atmosphere, it is possible to determine Air-broadening coefficients from the values for nitrogen- and oxygen-broadening given in the [supplementary material](#) using the following equation:

$$\gamma_{0,Air} = 0.79 \times \gamma_{0,N_2} + 0.21 \times \gamma_{0,O_2} \quad (8)$$

The derived values of the Air-broadening coefficients range from 0.0434 to 0.1544  $\text{cm}^{-1} \cdot \text{atm}^{-1}$ . These coefficients are listed for a few transitions in [Table 2](#). The whole set of these coefficients is presented in the [supplementary material](#).



**Figure 9:** Air-broadening coefficients in the  $2\nu_4$  band of  $NH_3$  at 295 K as function of  $J+0.2(J-K)$  for the manifolds  $J=1 \rightarrow 6$ . Values from Refs. [10,20] are also plotted.

The present results of Air-broadening coefficients reproduce the systematic  $J$  and  $K$  quantum numbers dependencies of Refs. [10,13,20] namely a decrease as a function of  $J$  for a fixed  $K$  and an increase versus  $K$  for a given  $J$ . It is illustrated in [Figure 9](#) for the  $^Q P$  and  $^R R$  sub-branches of the  $2\nu_4$  and  $\nu_2$  bands, respectively. The Air-broadening coefficients of HITRAN [10] derived from the polynomial model of  $\nu_2$  band measurements of Nemtchinov *et al.* [20], the difference with previous works are about 6% when the present data are compared to those of the  $\nu_2$  band of Ref. [20].

#### 4. Conclusions

This work presents the first measurements of pressure broadening and shift coefficients as well as narrowing coefficients for 310 transitions in the  $2\nu_4$  band of  $\text{NH}_3$  perturbed by  $\text{N}_2$  and  $\text{O}_2$  at 295 K in the  $3100 - 3600 \text{ cm}^{-1}$  spectral region. Voigt and soft-collision Galatry profiles were used to fit the measured line shape of each individual transition at various pressures. For a significant number of transitions, deviations from the Voigt profile were observed, highlighting the Dicke narrowing effect. We have quantified this effect for more than 100 lines, thanks to the high signal-to-noise ratio obtained in our Fourier Transform spectra. In all cases, Galatry model gives slightly higher values compared to Voigt one. Our data are compared with previous work when it's possible.

Using the measured broadening coefficients of the  $\text{NH}_3\text{-O}_2$  and  $\text{NH}_3\text{-N}_2$  systems, we have produced the  $\text{NH}_3$  Air-broadening coefficients that were also compared with data in the literature. The measured narrowing and shift coefficients are consistent with the rare previous works in the literature for other bands. The data obtained in this work are available for use in the HITRAN and GEISA databases.

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## References

- [1] Schilt S, Thévenaz L, Niklès M, Emmenegger L, Hügli C. Ammonia monitoring at trace level using photoacoustic spectroscopy in industrial and environmental applications. *Spectrochim Acta Part A Mol Biomol Spectrosc* 2004;**60**:3259–68.
- [2] Manne J, Sukhorukov O, Jäger W, Tulip J. Pulsed quantum cascade laser-based cavity ring-down spectroscopy for ammonia detection in breath. *Appl Opt* 2006;**45**:9230.
- [3] Webber ME, Claps R, English FV, Tittel FK, Jeffries JB, Hanson RK. Measurements of NH<sub>3</sub> and CO<sub>2</sub> with distributed-feedback diode lasers near 2.0 μm in bioreactor vent gases. *Appl Opt* 2001;**40**:4395.
- [4] Claps R, English FV, Leleux DP, Richter D, Tittel FK, Curl RF. Ammonia detection by use of near-infrared diode-laser-based overtone spectroscopy. *Appl Opt* 2001;**40**:4387–94.
- [5] Owen K, Farooq A. A calibration-free ammonia breath sensor using a quantum cascade laser with WMS 2f/1f. *Appl Phys B* 2013;**116**:371–83.
- [6] Filho MB, da Silva MG, Sthel MS, Schramm DU, Vargas H, Miklós A and *al.* Ammonia detection by using quantum-cascade laser photoacoustic spectroscopy. *Appl Opt* 2006;**45**:4966–71.
- [7] Kunde V, Hanel R, Maguire W, Gautier D, Baluteau JP, Marten A, Chedin A, Husson H, and Scott NS, *J Astrophys* 1982;**263**:443–67.
- [8] Ho PTP and Townes CH, *Ann Rev Astron Astrophys* 1983;**21**:239–70.
- [9] Encrenaz T and Moreno R. The microwave spectra of planets, in: *AIP Conference Proceedings – Experimental cosmology at millimetre wavelengths: 2K1BC Workshop*. 2002;**616**:330–37.
- [10] Gordon IE, Rothman LS, Hill C, Kochanov RV, Tan Y, Bernath PF and *al.* The HITRAN 2016 molecular spectroscopic database. *JQSRT* 2017;**203**:3–69.
- [11] Jacquinet-Husson N, Scott NA, Chedin A, Crepeau L, Armante R, Capelle V *et al.* The 2009 edition of the GEISA spectroscopic database. *JQSRT* 2011;**112**:2395–445.
- [12] Pickett HM, Poynter RL, Cohen EA, Delitsky ML, Pearson JC, Muller HSP. Submillimeter, millimeter, and microwave spectral line catalog. *JQSRT* 1998;**60**: 883–90.
- [13] Fabian M. Ito F and Yamada KMT. N<sub>2</sub>, O<sub>2</sub>, and Air broadening of NH<sub>3</sub> in ν<sub>2</sub> band measured by FTIR Spectroscopy. *J Mol Spectrosc* 1995;**173**:591–602, corrected later in 2006 in *J of Mol Spectros* 2006;**236**:150.
- [14] Aroui H, Chevalier M, Broquier M, Picard-Bersellini A, Gherissi S, Legay Sommaire N. Line mixing parameters in the ν<sub>4</sub> rovibrational band of NH<sub>3</sub> perturbed by N<sub>2</sub>. *J Mol Spectrosc* 1995;**169**:502–10.
- [15] Aroui H, Broquier M, Picard-Bersellini A, Bouanich JP, Chevalier M, Guerissi S. Absorption intensities, pressure-broadening and line mixing parameters of some lines of NH<sub>3</sub> in the ν<sub>4</sub> band. *JQSRT* 1998;**60**:1011–23.
- [16] Dhib M, Bouanich JP, Aroui H, Picard-Bersellini A. Analysis of N<sub>2</sub>, O<sub>2</sub>, CO<sub>2</sub>, and Air broadening of infrared spectral lines in the ν<sub>4</sub> band of NH<sub>3</sub>. *JQSRT* 2001;**68**:163–78.
- [17] Dhib M, Aroui H, Orphal J. Experimental and theoretical study of line shift and mixing in the ν<sub>4</sub> band of NH<sub>3</sub> perturbed by N<sub>2</sub>. *JQSRT* 2007;**107**:372–84.
- [18] Aroui H, Chelin P, Echargui MA, Fellows CE, Orphal J. Line shifts in the ν<sub>2</sub>, 2ν<sub>2</sub> and ν<sub>4</sub> bands of NH<sub>3</sub> perturbed by O<sub>2</sub>. *J Mol Spectrosc* 2008;**252**:129–35.
- [19] Nouri S, Orphal O, Aroui H and Harmann JM. Temperature dependence of pressure broadening of NH<sub>3</sub> perturbed by H<sub>2</sub> and N<sub>2</sub>. *J Mol Spectrosc* 2004;**227**:60–66.
- [20] Nemtchinov V, Sung K, Varanasi P. Measurements of line intensities and half-widths in the 10-μm bands of <sup>14</sup>NH<sub>3</sub>. *JQRST* 2004;**83**:243–65.

- [21] Dhib M, Ibrahim N, Chelin P, Echargui M.A, Aroui H, Orphal J. Diode-laser measurements of O<sub>2</sub>, N<sub>2</sub> and Air-pressure broadening and shifting of NH<sub>3</sub> in the 10 μm spectral region. *J Mol Spectrosc* 2007;**242**:83–89.
- [22] Sur R, Spearrin RM, Peng WY, Strand CL, Jeffries JB, Enns GM, Hanson RK. Line intensities and temperature-dependent line broadening coefficients of Q-branch transitions in the ν<sub>2</sub> band of ammonia near 10.4 μm. *JQSRT* 2016;**175**:90-99.
- [23] Koshelev MA, Tretyakov MY, Lees RN, Xu LH. ECTDL study of N<sub>2</sub>- and O<sub>2</sub>-pressure broadening of a series of ammonia lines in the 1.5 μm (ν<sub>1</sub>+ν<sub>3</sub>) combination band. *Appl Phys B* 2006;**85**:273-77.
- [24] Dicke RH. The effect of collisions upon the Doppler width of spectral Lines. *Phys Rev* 1953;**89**:472–3.
- [25] Rautian SG, Sobelman II. The effect of collisions on the Doppler broadening of spectral lines. *Sov phys Uspheki* 1967;**9**:701–16.
- [26] Galatry L. Simultaneous effect of Doppler and foreign gas broadening on spectral lines. *Phys Rev* 1961;**122**:1218–23.
- [27] Pine AS, Markov VN, Buffa G, Tarrini O. N<sub>2</sub>, O<sub>2</sub>, H<sub>2</sub>, Ar and He broadening in the ν<sub>1</sub> band of NH<sub>3</sub>. *JQSRT* 1993;**50**:337–48.
- [28] Pine AS, Markov VN. Self and foreign-gas-broadened line shapes in the ν<sub>1</sub> band of NH<sub>3</sub>. *J Mol Spectrosc* 2004;**228**:121-42.
- [29] Owen K, Es-sebbar E-t, Farooq A. Measurements of NH<sub>3</sub> line strengths and collisional broadening coefficients in N<sub>2</sub>, O<sub>2</sub>, CO<sub>2</sub>, and H<sub>2</sub>O near 1103.46cm<sup>-1</sup>. *JQSRT* 2013;**121**:56–68.
- [30] Maaroufi N, Jalleli. C, Kwabia Tchana F, Landsheere X and Aroui H. Absolute line intensities and first measurements of self-collisional broadening and shift coefficients in the 2ν<sub>4</sub> band of NH<sub>3</sub>. *J Mol Spectrosc* 2018;**354**:24-31.
- [31] Wartewig S. IR and Raman spectroscopy: fundamental processing. Weinheim: Wiley-VCH 2003.
- [32] Mertz L. Transformations in optics. New York: Wiley:1965.
- [33] Carleer MR. A Windows program to accurately measure the line intensities of high-resolution Fourier transform spectra, Remote sensing of clouds and the atmosphere V 2001;**4168**:337-43.
- [34] Sharpe S, Johnson T, Sams R, Chu RP, Rhoderick G, Johnson P. Gas-phase databases for quantitative infrared spectroscopy. *Appl Spectrosc* 2004;**58**:1452–61.
- [35] Ballard J, Knight RJ, Vander Auwera J, Herman M , Di Lonardo G, Masciarelli G, Nicolaisen FM, Beukes JA, Christensen LK, McPheat R, Duxbury G, Freckle- ton R, Shine KP. An intercomparison of laboratory measurements of absorption cross-sections and integrated absorption intensities for HCFC-22. *JQSRT* 2000;**66**:109–28 .
- [36] Pine AS. N<sub>2</sub> and Ar broadening and line mixing in the P and R branches of the ν<sub>3</sub> band of CH<sub>4</sub>. *JQSRT* 1997;**57**:157-76.
- [37] Weissman S. Estimation of diffusion coefficients from viscosity measurements: Polar and polyatomic gases. *J Chem Phys* 1964;**40**:3397-3406.