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Cloud-rain predator-prey interactions: analyzing some properties of the Koren-Feingold model and introduction of a new species-competition bulk system with a Hopf bifurcation

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Abstract

This paper deals with some properties of predator-prey cloud-rain models. Focus is put on scaling and on some mathematical features such as stability and limit cycles. Precisely, the Koren-Feingold delay differential equation model is first investigated and it is shown that it has no limit cycles. Then, by considering another point of view (*i.e.* species competition dynamics) for parametrizing cloud-rain processes, a system of ordinary differential equations to model these processes is formulated. Some examples are given to illustrate that this model reproduces in a realistic way the essential macroscopic behavior of a cloud-rain system. The model has a Hopf bifurcation at which certain properties of cloud-rain interactions in the model are represented. This is an important point to prepare for further examination of cloud synchronization in a cloud field by Kuramoto model, for instance.

Keywords: Cloud-rain interactions, Species competition dynamics, Predator-prey model, Cloud physics

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1 1. Introduction

Clouds are complex nonlinear dynamic systems with many degrees of free-2 dom and interactions across a vast range of spatio-temporal scales. Despite 3 this complexity, certain non-trivial aspects of their macrobehavior are pre-4 dictable without a consideration of the full complexity of the dynamic system 5 (Feingold et al., 2010; Koren and Feingold, 2011; Feingold and Koren, 2013; 6 Koren et al., 2017; Mülmenstädt and Feingold, 2018). This implies that mo-7 dels that capture the essential physics have the potential to contribute to our understanding of clouds and their interactions. Such models are thus always 9 welcome, especially in conjunction with more detailed models. This is of im-10 portance in the context of the climate system (e.q. Held, 2005) since, as is 11 well known, clouds play a major role in the climate system (IPCC, 2013). 12

An example of macrobehavior that can be tackled by nonlinear dynamics is the interaction between cloud droplets and rain. Recently, Koren and Feingold (2011) (hereafter KF11) analyzed cloud-rain coupling by means of the following non-linear system of delay differential equations (DDEs):

4

$$\begin{cases} \frac{\mathrm{d}H}{\mathrm{d}t} &= \frac{H_0 - H}{\tau_1} - \frac{\alpha H^2(t - T)}{c_1 N_d(t - T)} \\ \frac{\mathrm{d}N_d}{\mathrm{d}t} &= \frac{N_0 - N_d}{\tau_2} - \alpha c_2 H^3(t - T) \end{cases}$$
(1)

with $c_1 = 2 \times 10^{-6} \,\mathrm{mm}\,\mathrm{m}^{-2}$, $c_2 = 0.3 \,\mathrm{m}^{-1}$, and $\alpha = 2 \,\mathrm{mm}\,\mathrm{m}^{-6}\,\mathrm{d}^{-1}$ (see also 13 Table 1 in appendix). Here, H is cloud depth (in m) and N_d is cloud droplet 14 concentration (in cm^{-3}). These are the two macroscopic degrees of freedom. 15 In addition, t is time and T is an arbitrary constant delay (both in minutes). 16 The constants τ_1 and τ_2 are timescales of the order of ten minutes. H_0 is a 17 height that represents the full environmental potential for cloud development 18 (*i.e.* the maximum cloud depth possible) and N_0 is the background concen-19 tration of aerosol (*i.e.* the maximum concentration of cloud droplets that 20 can be reached). According to KF11, this system exhibits predator-prey like 21 behavior. In particular, two examples of cloud behavior are provided which 22 represent (i) an oscillator limit cycle and (ii) a damped oscillator. 23

In the first part of this paper, we explore some physical and mathematical properties of Eq.1 that we consider important from the perspective of nonlinear dynamics that have thus far not been addressed. More precisely, our analysis, which is presented in Section 2, concentrates on scale analysis, form of the equations, stability analysis, and the question of the existence of
limit cycles by means of the Busenberg theorem applied to DDEs.

In the second part of this paper, we introduce a bulk model which follows 30 the initial idea of modeling the macrobehavior of a cloud-rain system and has 31 properties in common with species competition dynamics. In contrast to the 32 KF11 model, it is a system of ODEs rather than DDEs. However, the model 33 presented here reproduces the behavior of the KF11 model, and is derived 34 from a physically-based parametrization. This model differs from the KF11 35 model not only in its structure, but in the variety of its dynamics which we 36 examine through the lense of nonlinear dynamics, e.g. linear stability analysis 37 and bifurcation theory. This exploration yields regions of parameter space in 38 which clouds grow at the expense of rain and *vice versa*. 39

It must be emphasized that this model does not consider the full com-40 plexity of cloud-precipitation interactions; rather, we examine certain predic-41 table elements of the macrobehavior of the full complex dynamical system. 42 In particular, the model we propose incorporates population dynamics in a 43 natural way, which includes exchanges with the surrounding environment. 44 Our model is presented in Section 3 and examples are given in order to show 45 that it is physically realistic and able to reproduce the macroscopic behavior 46 of a cloud-rain system. Section 4 provides a discussion and conclusions. 47

48 2. Analysis of some physical and mathematical properties of the 49 KF11 model

⁵⁰ 2.1. Scale or order-of-magnitude analysis

The system (1) can be solved numerically by means of the pydelay package of Python (http://pydelay.sourceforge.net/) or MATLAB's dde23 algorithm. The PyDDE solver can also be useful. These methods are based on the Bogacki-Shampine method which is a 3(2) Runge-Kutta scheme adapted to DDEs. However, before diving into numerical computations, we analyze the magnitudes of the different terms in the system of equations.

In order to have a consistent set of equations in terms of units, we replace α and c_2 in Eq. (1) by $\alpha' = 10^{-3} \alpha/(60 \times 24)$ and $c'_2 = 10^{-6} c_2$ (see appendix). Typically, for cloud height, $H_0 \sim 500$ m and so, by definition of H_0 , $0 < H < H_0$. For cloud droplet concentration, $N_0 \sim 400 \text{ cm}^{-3}$ and so, by definition of N_0 , $N_d < N_0$. A reasonable minimum for N_d ($N_{d,min}$) is about 50 cm⁻³. Hence, considering the values of the constants involved in Eq. (1), *i.e.* $\tau_1 =$

 $\tau_2 \sim 60 \text{ min}$, we find:

$$\frac{H_0 - H}{\tau_1} < \frac{H_0}{\tau_1} \approx 8 \,\mathrm{m\,min^{-1}} \qquad \frac{\alpha' H^2}{c_1 N_d} < \frac{\alpha' H_0^2}{c_1 N_{d,min}} \approx 3\,500 \,\mathrm{m\,min^{-1}}$$

and

$$\frac{N_0 - N_d}{\tau_2} < \frac{N_0}{\tau_2} \approx 7 \,\mathrm{cm}^{-3} \,\mathrm{min}^{-1} \qquad \alpha' c_2' H^3 < \alpha' c_2' H_0^3 \approx 0.5 \,\mathrm{cm}^{-3} \,\mathrm{min}^{-1}$$

It follows that, since H and N_d are physically expected to maintain or-57 ders or magnitude as mentioned above during the evolution of the cloud 58 rain system, that $\alpha' H^2/(c_1 N_d)$ and $(N_0 - N_d)/\tau_2$ are the dominant terms 59 in Eq. (1). This is particularly clear for the equation governing the evolu-60 tion of H. Concerning the equation for dN_d/dt , having a ratio close to 1 61 implies that c_2 be 10⁴ or 10⁵ higher than the value given in Table 1. Conse-62 quently, the evolution of N_d is either an increasing or a decreasing expo-63 nential function according to the sign (positive or negative, respectively) 64 of $N_0 - N_d$. The examples chosen by KF11 have positive signs. Precisely, 65 $N_d(t) - N_0 = [N_d(0) - N_0] \exp(-t/\tau_2)$, where $N_d(0) = N_d(t=0)$. It ensues 66 that N_d evolves exponentially towards N_0 . With respect to dH/dt, the do-67 minant term is always negative and, unless N_d becomes very small, it will 68 remain the dominant term in the evolution of H. Hence, H is a decreasing 69 function which must stop at H = 0 by definition. This scale analysis, based 70 only on the physical realistic orders of magnitude of H and N_d is independent 71 of any constant delay in the equation. Some numerical computations with the 72 numerical scheme mentioned above have confirmed the evolution described 73 above (Fig. 1). Nonetheless, we have obtained some oscillatory behavior in 74 rare cases by varying arbitrarily some constants in Eq. (1) as the constant 75 c_1 around 10^{-3} mm m⁻¹ and the constant c_2 around 10^4 m⁻¹ or 10^5 m⁻¹. We 76 have observed that these oscillations are, in addition, extremely sensitive to 77 the value of c_2 chosen since they disappear if c_2 is changed even slightly (we 78 recover the evolution described above). 79

⁸⁰ 2.2. Form of the system of equations

It is worth noting that the system (1) of two equations is not formally similar to a predator-prey system, *stricto sensu*. Indeed, in its basic form, the Lokta-Volterra system is the following (x and y are the two degrees of) freedom):

$$\begin{cases} \frac{\mathrm{d}x}{\mathrm{d}t} = ax - bxy\\ \frac{\mathrm{d}y}{\mathrm{d}t} = -cy + dxy \end{cases}$$
(2)

with $\{a, b, c, d\}$ a set of four positive constants. The signature (+, - | -, +), or sometimes (+, - | +, -), is also present in more elaborate versions of Lokta-Volterra models (*e.g.* Lipowski and Lipowska, 2000) since it is a characteristic of species-competition models, with or without predation. However, the system 1 has the form: $dH/dt = -AH - BH^2/N_d$ and $dN_d/dt =$ $-CN_d - DH^3$ with $\{A, B, C, D\}$ a set of four positive constants. So, system (1) does not have the general form of a Lokta-Volterra system since the signature is fully negative, (-, - | -, -). It follows that \dot{H} and \dot{N}_d are negative if (H, N_d) always satisfy the two inequalities:

$$\frac{H}{\tau_1} + \frac{\alpha' H^2}{c_1 N_d} > \frac{H_0}{\tau_1} \sim 10 \,\mathrm{m}\,\mathrm{min}^{-1}$$

and

$$\frac{N_d}{\tau_2} + \alpha' c_2' H^3 > \frac{N_0}{\tau_2} \sim 1 - 10 \,\mathrm{cm}^{-3} \,\mathrm{min}^{-1}$$

⁸¹ So, with regard to the typical orders of magnitude of the two degrees of ⁸² freedom ($H \sim 400 \,\mathrm{m}, N_d \sim 100 \,\mathrm{cm}^{-3}$), care is needed in the choice of the ⁸³ constants (*i.e.* in the model parametrization) in order to avoid always de-⁸⁴ creasing functions H(t) and $N_d(t)$.

85 2.3. Stability analysis

To determine further some of the properties of the system (1), we briefly discuss the stability of the model (with no delay) near its equilibrium points. These are independent of any delays, so that a system of DDEs has the same equilibrium points as the corresponding system (zero delay) of ODEs. However, investigating the stability for DDEs is a more complex task (*e.g.* Engelborghs et al., 2000, and references therein) which is beyond the scope of this paper (some elements can be found in Koren et al. (2017, Section IV)). Nonetheless, it is still instructive to analyze stability in the simplest case of no delay. A quick calculation yields that the KF11 system has an equilibrium point given by

$$N_{d}^{e} = N_{0} - \alpha' c_{2}' \tau_{2} (H^{e})^{3}$$

and

$$H^{e} = \frac{c_{1}N_{d}^{e}}{2\alpha'\tau_{1}} \left[-1 + \left(1 + \frac{4\alpha'\tau_{1}H_{0}}{c_{1}N_{d}^{e}}\right)^{1/2} \right]$$
(3)

Physically, because N_d^e (and H^e) must be positive, the value of $\alpha' c'_2$ is constrained: $\alpha' c'_2 < N_0/[\tau_2(H^e)^3]$. The behavior of Eq. (1) near this equilibrium point can be obtained by evaluating the Jacobian matrix J at (H^e, N_d^e) and finding its eigenvalues. These are given by the well-known formula:

$$\lambda_{\pm} = \frac{1}{2} \left[T_J \pm \left(T_J^2 - 4\Delta_J \right)^{1/2} \right]$$

where T_J and Δ_J are respectively the trace and the determinant of [J]. A straightforward calculation yields:

$$T_J = -\left(\frac{1}{\tau_1} + \frac{1}{\tau_2} + \frac{2\alpha' H^e}{c_1 N_d^e}\right)$$

and

$$\Delta_J = \frac{1}{\tau_2} \left(\frac{1}{\tau_1} + \frac{2\alpha' H^e}{c_1 N_d^e} \right) + 3\alpha'^2 c_2' \frac{(H^e)^4}{c_1 (N_d^e)^2}$$

Since $T_J < 0$ and $\Delta_J > 0$, the equilibrium point (H^e, N_d^e) is a stable node or a stable focus (spiral) according to the sign of $(T_J^2 - 4\Delta_J)$. We note that if the value of $\alpha' c'_2$ does not satisfy the above constraint, then T_J can be negative, which means that Eq. (1) describes an unstable dynamical system. In addition, since the sign of T_J never changes, there is no local bifurcation. A similar analysis that corroborates the above result has been performed recently by Jiang and Wang (2014).

93 2.4. Existence of limit cycles

One of the important points of this paper is the possibility of limit cycle 94 solutions of the KF11 delay differential equations (DDEs). This needs ri-95 gorous examination. First, we note that the KF11 equations with constant 96 delay T do not have any limit cycle solutions with period T. To see this, we 97 consider either the Bendixson-Dulac theorem or the special case known as 98 Bendixson's criterion (e.g. Minorsky, 1962; Glansdorff and Prigogine, 1971). 99 To illustrate the use and consequences of this theorem to the problem at 100 hand, we first rewrite the KF11 system: $dH/dt = f[H(t-T), N_d(t-T)]$ 101 and $dN_d/dt = g[H(t-T), N_d(t-T)]$, where the functions f and g are 102

given by the rhs (right hand sides) in Eq. (1). The Bendixson-Dulac theorem may be applied to this system by considering the sign of the expression $\partial_{H}(\phi f) + \partial_{N_{d}}(\phi g)$, where $\phi(H, N_{d}) = 1$ (this case is Bendixson's criterion).

Now, suppose that this set of equations has a non-constant *T*-periodic solution $[H(t), N_d(t)]$. Then, it follows that $dH/dt = f[H(t), N_d(t)]$ and $dN_d/dt = g[H(t), N_d(t)]$. A simple calculation shows that $\partial_H(f) + \partial_{N_d}(g)$ is always negative. We can therefore conclude by the Bendixson-Dulac theorem that no solutions of period *T* exist in the simply connected region defined by positive values of *H* and N_d .

Since the KF11 equations are DDEs, a customary application of the 112 Bendixson-Dulac theorem to prove the non-existence of limit cycles is inade-113 quate. However, this theorem has been generalized (Busenberg and van den 114 Driessche, 1993, Section 4) to higher dimensional ODEs, including DDEs like 115 those proposed by KF11, for instance. This generalized "Bendixson-Dulac" 116 theorem extends the concept of limit cycle to that of simple loop solutions, 117 *i.e.* any continuous solution of the system of equations whose orbit contains 118 a closed curve. Here, we apply this generalized theorem to the KF11 DDEs 119 in order to demonstrate the absence of loop solutions. 120

To illustrate the use and consequences of this generalized theorem to the problem at hand, especially theorem 4.1 in Busenberg and van den Driessche (1993), we first rewrite the KF11 system:

$$\begin{cases} \frac{\mathrm{d}H}{\mathrm{d}t} &= F_1(H, N_d) \, h_t(H_t, N_{d,t}) + l_1(H, N_d) \\ \frac{\mathrm{d}N_d}{\mathrm{d}t} &= F_2(H, N_d) \, k_t(H_t, N_{d,t}) + l_2(H, N_d) \end{cases}$$

where $F_1(H, N_d) = -\alpha/c_1$, $F_2(H, N_d) = -\alpha c_2$, $h_t(H_t, N_{d,t}) = H_t^2/N_{d,t}$ and $k_t(H_t, N_{d,t}) = H_t^3$ are time-delayed functions (the subscript "t" stands for "time-delayed"), $l_1(H, N_d) = (H_0 - H)/\tau_1$, and $l_2(H_d, N_d) = (N_0 - N_d)/\tau_2$. Theorem 4.1 in Busenberg and van den Driessche (1993) gives conditions under which the KF11 DDEs have no simple loop solution. To verify the first condition, we define the vector function $g(H, N_d) = (g_1, g_2, 0)$. Here, $g_1 = \alpha c_2 A^3 + (N_0 - N_d)/\tau_2$, $g_2 = -\alpha A^2/(c_1 B) - (H_0 - H)/\tau_1$, A and B being constants. For notational simplicity, let (u, v) represent the coordinates of the

curve in Theorem 4.1 of Busenberg and van den Driessche (1993). Then:

$$g(H, N_d) \cdot \left(\frac{-\alpha u^2}{c_1 v} + l_1, -\alpha c_2 u^3 + l_2, 0\right)$$

= $\frac{\alpha^2 c_2 A^2 u^2}{c_1} \left(\frac{u}{B} - \frac{A}{v}\right) + \alpha c_2 l_1 \left(u^3 + A^3\right)$
 $- \frac{\alpha l_2}{c_1} \left(\frac{A^2}{B} + \frac{u^2}{v}\right)$
< 0

This follows from the following observations. The third term always contri-121 butes negatively, the second positively, to the sum. The first contributes 122 negatively whenever uv < AB. Because c_1 is small, the first and third terms 123 dominate the second so that, whenever uv < AB, the entire sum is negative. 124 Since A and B can be taken as large as wished, the region is finally \mathbb{R}^2_+ . To 125 verify the second condition, we calculate $\partial_H g_2 - \partial_{N_d} g_1 = 1/\tau_1 + 1/\tau_2 > 0$ for 126 any solution (H, N_d) in the region specified above. So, an application of the 127 Busenberg-van den Driessche theorem implies that there is no simple loop 128 solution traversed in the clockwise sense (the analogue for DDEs of the limit 129 cycle of ODEs) for the positive values of H and N_d . A similar argument with 130 vector function -q works in the opposite sense as well. In other words, there 131 is no solution of the KF11 system which enclose a smooth oriented simple 132 closed curve. 133

¹³⁴ 3. A new species-competition bulk model

135 3.1. Description

In order to retain the simplicity of the KF11 model with ODEs instead of 136 DDEs, we propose an approach which incorporates population competition, 137 interaction, and carrying capacities in a natural way. Here, the populations 138 are given by cloud water content L_c (g cm⁻³), rain water content L_r (g cm⁻³), 139 and cloud droplet number concentration N_d (cm⁻³). We emphasize that these 140 variables are considered here as macroscopic averages and that our model 141 captures some of the macroscale properties of cloud-precipitation interactions 142 by means of three relatively simple equations in three variables, rather than a 143 fine-scale description with complete microphysical details. The model consists 144 of three coupled differential equations, one each for L_c , L_r , and N_d . The 145 system includes three main processes that contribute to the cloud budgets of 146

¹⁴⁷ L_c , L_r , and N_d , namely: (1) sources/sinks which represent exchanges with the ¹⁴⁸ surroundings, and two internal processes, (2) autoconversion of cloud water ¹⁴⁹ to rain water, and (3) accretion of cloud water by rain water.

The various terms in the equations were suggested by the Seifert and 150 Beheng (2001) parametrization for simulating autoconversion and accretion. 151 Other terms can be added to the equations, such as self-collection for ins-152 tance, but to retain some measure of simplicity we use only the terms al-153 ready mentioned and which we describe further. This parametrization is a 154 straightforward consequence of the stochastic collection equation (STE) with 155 polynomial kernel. For this reason, it seems to us that this approach is more 156 rigorous than many other methods which consist in using heuristic (or em-157 pirical) parametrizations. Details on the STE can be found in Seifert and 158 Beheng (2001) and Pruppacher and Klett (1997) and references therein. It is 159 worth mentioning that the goal of Seifert and Beheng (2001) has been to fill 160 a gap between heuristic parametrizations and detailed microphysical (com-161 putationally expensive) schemes. The terms representing autoconversion and 162 accretion are given by the following expressions (for a comprehensive over-163 view see the references): 164

- Autoconversion: the contribution to dL_c/dt and dL_r/dt is given by

$$\left. \frac{dL_c}{dt} \right|_{auto} = -KL_c^4 N_d^{-2} = -\left. \frac{dL_r}{dt} \right|_{auto} \tag{4}$$

- Accretion: the contribution to dL_c/dt and dL_r/dt is given by

$$\left. \frac{dL_c}{dt} \right|_{accret} = -k_r L_c L_r = -\left. \frac{dL_r}{dt} \right|_{accret}.$$
(5)

We note that these expressions are adapted from the Seifert and Beheng 165 (2001) parametrization by considering the variables as bulk, cloud-scale va-166 riables. The various parameters in the equations above are defined as follows: 167 $K = [k_c/(20x^*)](\nu + 2)(\nu + 4)(\nu + 1)^{-2}$, where x^* is a cloud drop mass se-168 parating the cloud droplets from raindrops, k_c (k_r) is a constant from the 169 cloud (rain) water kernel, and ν is the shape parameter of the gamma dis-170 tribution. See Seifert and Beheng (2001) for complete details. By combining 171 these equations, neglecting spatial dependence and adding source/sink terms 172 for cloud and rain water contents, we arrive at the following system of bulk 173

¹⁷⁴ coupled first-order differential equations for L_c , L_r and N_d :

$$\begin{pmatrix}
\frac{dL_c}{dt} = A_c L_c - B_c (N_d) L_c^4 - k_r L_c L_r \\
\frac{dL_r}{dt} = -A_r L_r + B_c (N_d) L_c^4 + k_r L_c L_r \\
\frac{dN_d}{dt} = f(N_d, L_c, L_r)$$
(6)

Here, $B_c(N_d) = K N_d^{-2}$, and A_c and A_r are two positive constants that can be chosen so that the source/sink terms represent the background meteorological conditions. In particular, A_r represents the rain out process. These two constants give the timescales, $\tau_c = 1/A_c$ and $\tau_r = 1/|A_r|$, of cloud water content and rain water content evolutions respectively. In addition, $f(N_d, L_c, L_r)$ is a function which can be defined according to the question being investigated. Here, we choose to define f as follows:

$$\frac{dN_d}{dt} = A_c(N_0 - N_d) - \frac{4}{3}k_cL_c^2 - k_rL_rN_d.$$
(7)

On the rhs, the first term represents a supply of cloud droplets from the 175 surroundings, N_0 being the background aerosol concentration that feeds the 176 system (nucleation). Considered alone, this term causes an exponential in-177 crease (decrease) of N_d with N_0 as horizontal asymptote if initially N_d is lo-178 wer (higher) than N_0 . This represents cloud droplet concentration and when 179 considered alone it tends toward a constant background aerosol population 180 concentration. Moreover, this term (or similar ones) should be dominant in 181 the very early stages of the cloud system evolution. Indeed, at the initial time 182 (and close to t = 0), N_d is zero or very small (clear air condition) and must 183 increase enough in order that the cloud can further evolve and produce rain. 184 The two other terms in Eq. (7) come from Seifert and Beheng (2001, Eqs. 185 A-5, A-6, and A-9) and represents respectively cloud droplet self-collection 186 and accretion. 187

It has to be recalled that the state variables considered $(L_c, L_r, \text{ and } N_d)$ in this bulk model are macroscopic averages over the whole cloud. They can also be seen as local variables over a volume element of the cloud for which they are more or less uniform. A further step would be to introduce in the equation an explicit inside-cloud location (*e.g.* altitude *z*) dependence. Nonetheless, as showed below, the present model describes some macroscale features of could-precipitation coupling.

¹⁹⁵ 3.2. Comment on the form of the system of equations

We note that the first two equations in (6) constitute a predator-prey 196 system modified to include autoconversion. To see this, observe that the 197 constants $\{a, b, c, d\}$ in the Lotka-Volterra equations (Eq. 2) have the follo-198 wing formal correspondence with the model presented here: $\{a = A_c, b =$ 199 $-d = k_r, c = A_r$. Hence, the source/sink and accretion terms form a stan-200 dard predator-prey system when considered alone. The quartic terms not 201 only represent autoconversion of cloud droplets to rain given in the Seifert 202 and Beheng (2001) parametrization, but also ensure that the cloud droplet 203 population does not exhibit unbounded growth in the absence of competition. 204 Thus, it plays a similar role to the carrying capacity term used in population 205 dynamics to modify some of the unrealistic features of the original Lotka-206 Volterra system. 207

The entire system of equations is not rigorously equivalent to a threespecies competition model. As already discussed, the first two equations constitute a modified Lotka-Volterra system. The autoconversion terms which modify the traditional predator-prey system have coefficients driven by the third equation, *viz.* that for N_d . The interpretation of the varying coefficients is that the carrying capacities of the water contents vary according to the state of N_d , which is more realistic for clouds.

It is worth mentioning that L_c and L_r are, for physical reasons, necessa-215 rily bounded. So, in virtue of the Poincaré-Bendixson theorem, the first two 216 equations of system (6) have a solution that either converges towards a limit 217 or presents an asymptotic behaviour that can take the form of a limit cycle. 218 Boundedness has not been proved in a mathematical sense but, rather, we 219 rely on numerical evidence (see some examples below, in Subsection 3.4) and 220 on the fact that the present system of equations has much in common with 221 other predator-prey models which are known to have bounded solutions. Ob-222 viously, if k_r and B_c are zero, the evolution of L_c is exponentially divergent. 223 The terms involving non-zero values of k_r and B_c prevent unbounded evolu-224 tion within certain parameter ranges, as is the case in predator-prey models. 225 It is far beyond the scope of this paper to address rigorously (*i.e.* mathema-226 tically speaking) boundedness of solutions, but that could be formulated and 227 explored deeply in a future specific work. 228

229 3.3. Stability analysis

230 3.3.1. Equilibrium points

The system

$$\begin{cases}
\frac{dL_c}{dt} = A_c L_c - B_c (N_d) L_c^4 - k_r L_c L_r \\
\frac{dL_r}{dt} = -A_r L_r + B_c (N_d) L_c^4 + k_r L_c L_r
\end{cases}$$
(8)

has two equilibrium points $X_e = \{L_c^{(e)}, L_r^{(e)}\}$. In this section we treat N_d as a parameter. The first one is trivial: $X_e^{(1)} = 0$. There is no cloud nor *a fortiori* rain. Using Eq. (7), $N_d^{(e)} = N_0$, which means that the atmospheric volume considered has a constant CCN (cloud condensation nuclei) loading (that does not allow cloud formation). The second equilibrium point $X_e^{(2)}$ is such that:

$$A_c L_c^{(e)} = |A_r| L_r^{(e)} \quad \text{and} \quad B_c(N_d) \frac{|A_r|^3}{A_c^4} \left(L_r^{(e)}\right)^3 + \frac{k_r}{A_c} L_r^{(e)} - 1 = 0 \qquad (9)$$

Since the two first terms of this third order polynomial are positive, there is only one real positive solution $L_r^{(e)}$ – the other two are complex conjugates. The expression of $N_d^{(e)}$ follows from those of $L_c^{(e)}$ and $L_r^{(e)}$ using Eq. (7) Plugging $L_c^{(e)}$ and $L_r^{(e)}$ into Eq. (7) yields:

$$N_d^{(e)} = \left(A_c + \frac{k_r A_c}{|A_r|} L_c^{(e)}\right)^{-1} \left[A_c N_0 - \frac{4}{3} k_c \left(L_c^{(e)}\right)^2\right]$$
(10)

The eigenvalues of the jacobian matrix J of the system (6) are:

$$\lambda_{\pm} = \frac{1}{2} \left[T_J \pm \left(T_J^2 - 4\Delta_J \right)^{1/2} \right]$$

where the trace T_J and the determinant Δ_J of [J] can be written as follows:

$$T_J = \text{Tr}(J) = A_c - |A_r| + k_r \left(L_c^{(e)} - L_r^{(e)} \right) - 4B_c(N_d) \left(L_c^{(e)} \right)^3$$
(11)

and

$$\Delta_J = \det(J) = -A_c |A_r| + k_r \left(A_c L_c^{(e)} + |A_r| L_r^{(e)} \right) + 4|A_r| B_c(N_d) \left(L_c^{(e)} \right)^3$$
(12)

²³¹ They must be evaluated for each of the two equilibrium points.

²³² 3.3.2. Stability analysis around $X_e^{(1)}$

In this case, $T_J = A_c - |A_r|$ and $\Delta_J = -A_c |A_r| < 0$. The two eigenvalues have opposite sign, $\lambda_+\lambda_- < 0$, with $\lambda_+ = A_c > 0$ and $\lambda_- = -|A_r| < 0$. The state $X_e^{(1)}$ is thus a saddle point.

²³⁶ 3.3.3. Stability analysis around $X_e^{(2)}$ In this case

$$T_J = (3A_c + |A_r|) \left(\frac{k_r L_r^{(e)}}{A_c} - 1\right) \quad \text{and} \quad \Delta_J = -2A_c |A_r| \left(\frac{k_r L_r^{(e)}}{A_c} - \frac{3}{2}\right)$$

It follows that the equilibrium point $X_e^{(2)} = (L_c^{(e)}, L_r^{(e)})$ is stable if:

$$\frac{k_r L_r^{(e)}}{A_c} < 1 \quad \text{i.e.} \quad L_r^{(e)} < \frac{A_c}{k_r} \quad \text{and} \quad L_c^{(e)} < \frac{|A_r|}{k_r}$$
(13)

²³⁷ Otherwise, for $k_r L_r^{(e)}/A_c > 1$, $X_e^{(2)}$ is an unstable equilibrium point.

The identity $k_r L_r^{(e)} / A_c = 1$, which separates a stable regime from an 238 unstable one, has the straightforward consequence that $B_c \left[N_d^{(e)} \right] \left(L_c^{(e)} \right)^4 =$ 239 0, *i.e.* $B_c = 0$ since $L_c^{(e)} \neq 0$. Physically, this critical value means that 240 the net rate of the autoconversion process is zero: formation of a raindrop 241 from cloud droplets only is compensated over a given time interval by the 242 formation of cloud droplets from a raindrop only. Thus, the system maintains 243 constant rain and cloud water contents whose values — see Eq. (13) — 244 depend only on the exchange rates with the surroundings $(A_c \text{ and } A_r)$ and 245 on the cloud-to-rain reaction constant (k_r) . Below the critical value $(B_c < 0)$, 246 the autoconversion process results in net production of cloud droplets from 247 raindrops. It constitutes a nonlinear amplification term for cloud droplets and 248 a nonlinear damping term for raindrops, *i.e.* disturbances from equilibrium 249 grow. In other words, the cloud is growing such that cloud droplets become 250 more numerous while the number of raindrops vanishes. This behavior is 251 reversed above the critical value $(B_c > 0)$: the non linear term compensates 252 any deviation from the equilibrium point (stability). In the case $B_c < 0$, 253 without any cloud-rain interaction $(k_r = 0)$, cloud (rain) water content is an 254 increasing (decreasing) function. 255

The other characteristic value is $k_r L_r^{(e)}/A_c = 1.5$. Below (above) this value, Δ_J is positive (negative). So $X_e^{(2)}$ is a saddle point if $L_r^{(e)} = 1.5A_c/k_r$. The nature (node, focus or spiral) of $X_e^{(2)}$ is provided by the sign of $T_J^2 - 4\Delta_J$:

$$T_J^2 - 4\Delta_J = \left[(3A_c + |A_r|)^2 \left(\frac{k_r L_r^{(e)}}{A_c} - 1 \right)^2 \right] + 4|A_r|A_c \left(2\frac{k_r L_r^{(e)}}{A_c} - 3 \right)$$

If $k_r L_r^{(e)}/A_c = 1$, $T_J^2 - 4\Delta_J = -4|A_r|A_c < 0$ which means that $X_e^{(2)}$ is a focus. For $k_r L_r^{(e)}/A_c \neq 1$, the sign of $T_J^2 - 4\Delta_J$ depends on the number $\mathcal{N}_{\tau} = |A_r|/A_c$ through the second order polynomial equation:

$$(3A_c + |A_r|)^2 \left(\frac{k_r L_r^{(e)}}{A_c}\right)^2 - 2\left(\frac{k_r L_r^{(e)}}{A_c}\right) \left[(3A_c + |A_r|)^2 - 4A_c |A_r|\right] + (|A_r| - 3A_c)^2$$

The two solutions are easy to find:

$$Z_{\pm} = \frac{\left(3A_c + |A_r|\right)^2 - 4|A_r|A_c}{\left(3A_c + |A_r|\right)^2} \pm \frac{2\left[\left(9A_c^2 + |A_r|A_c\right)\left(|A_r|^2 + A_c|A_r|\right)\right]^{1/2}}{\left(3A_c + |A_r|\right)^2}$$

²⁵⁹ Three cases can be distinguished at first sight:

²⁶⁰ $-|A_r| = A_c$: so $Z_{\pm} = (3 \pm \sqrt{5})/4 > 0$. Between these two values, $T_J^2 - 4\Delta_J < 0$, so that $X_e^{(2)}$ is a focus. Outside of this interval, we have a node

$$_{263} - |A_r| \gg A_c$$
: so $Z_{\pm} = 1 \pm 2 (A_c/|A_r|)^{1/2} > 0$. Same as previously.

$$_{264}$$
 - $|A_r| \ll A_c$: so $Z_{\pm} = 1 \pm (2/3) (|A_r|/A_c)^{1/2} > 0$. Same as previously.

265 3.3.4. Andronov-Hopf bifurcation

It follows from what precedes that the eigenvalues of the jacobian matrix 266 cross the imaginary axis for $B_c = 0$. At this point, they are pure imaginary 267 numbers, $\lambda_{\pm} = i(A_c|A_r|)^{1/2}$. There, the system has a limit cycle. Just after 268 $(B_c < 0)$, the spiral is unstable whilst it is stable just before $(B_c > 0)$. 269 This behavior is precisely a *Hopf bifurcation*: an unstable focus gives birth 270 to a stable focus (and vice versa) through a limit cycle. The existence of 271 such a bifurcation is an important point for further investigations of cloud 272 organization in a coupled-cloud field at the mesoscale (Kuramoto, 2003). 273

274 3.4. Examples

We now illustrate the behavior of this set of three equations with some examples. The values chosen for k_r and k_c are those given in Seifert and Beheng (2001) which follow from approximations to the collection kernel in Long (1974), *i.e.* $k_c \approx 5.66 \times 10^{11} \,\mathrm{cm}^3 \,\mathrm{g}^{-2} \,\mathrm{min}^{-1}$ and $k_r \approx 3.47 \times 10^5 \,\mathrm{cm}^3 \,\mathrm{g}^{-1} \,\mathrm{min}^{-1}$. The value of $x^* = 2.6 \times 10^{-7} \,\mathrm{g}$, following Seifert and Beheng (2001). We take $\nu = 2$, which is a typical value for warm clouds (*e.g.*, Pruppacher and Klett, 1997, chap. 2). With these values, the parameter $K \approx 2.905 \times 10^{17} \,\mathrm{cm}^3 \,\mathrm{g}^{-3} \,\mathrm{min}^{-1}$.

In addition, we take $A_c = 0.02 \text{ min}^{-1}$ and $A_r = 0.1 \text{ min}^{-1}$. So, cloud water content and rain water content are expected to evolve on the timescales $\tau_c = 50 \text{ min}$ and $\tau_r = 10 \text{ min}$. The latter is lower than the former since raining out is a rapid process compared to the evolution of cloud water content. These values are chosen for illustrative purposes.

288 3.4.1. First example: periodic behavior

We choose as initial conditions $N_d(0) = 10 \,\mathrm{cm}^{-3}, L_c(0) = 10^{-9} \,\mathrm{g} \,\mathrm{cm}^{-3},$ 280 $L_r(0) = 0$, and $N_0 = 50 \,\mathrm{cm}^{-3}$. The system exhibits oscillatory behavior after 290 approximately 300 min, with a period of about 200 min, which is physically 291 realistic (Fig. 2a). Both N_d and L_c increase initially (the former faster than 292 the latter) until rain appears. Soon thereafter, rain water content peaks and 293 droplet number and mass concentration decrease drastically in a short in-294 terval of time. Minimum values of L_c are reached and rain water content 295 decreases to zero. Then N_d and L_c start to increase again, and so on. This 296 periodicity appears clearly as a limit cycle on the plot (L_c, L_r) displayed in 297 Fig. 2b. It is interesting to mention that peaks in rain water content are de-298 layed from L_c peaks by about $20 - 30 \min$, which is a very realistic value for 299 rain to form from a cloud droplet by collection (Fig. 2c). We also note that 300 similar oscillations and limit cycles are obtained for other initial conditions 301 (not shown). Note that the model presented here represents well a timescale 302 of precipitation production that has to be explicitly and arbitrarily accounted 303 for via delays in the more complicated DDE framework. 304

305 3.4.2. Second example: damped behavior

If N_0 is lower than $N_d(0)$, the feeding of the system is not sufficient to maintain the cloud-rain system in a harmonic (or quasi) oscillation and a damping is observed (Fig. 3): N_d , L_c , and L_r still oscillate but their respective amplitudes decrease progressively until a steady state is reached. Initially N_d decreases exponentially as expected, and then decreases more rapidly when L_c and L_r peak for the first time.

³¹² Damping can also be obtained by introducing in the model a decrea-³¹³ sing evolution for the CCN supply, for instance $N_0 = 50 \exp(-0.002 t)$. This

means that the initial CCN burden is divided by around 2.72 in 500 min, 314 *i.e.* 8 hours. Keeping the same initial conditions as previously, we get the 315 behavior displayed in Figure 4, which is typical of a damped oscillator: N_d 316 increases until a certain time t (approximately $120 \min$) at which the expo-317 nential function becomes significant. Then N_d decreases dramatically. The 318 same behavior is seen for L_c and L_r which increase until t = 500 min and 319 then exhibit decaying amplitude oscillations that eventually (t > 1500 min)320 die out. Again, we note that the peaks of L_r are delayed with respect to those 321 of L_c by about 20 - 30 min. 322

The same damping behavior is observed (not shown) for different initial conditions, in particular if $L_r \neq 0$, provided the values are realistic.

4. Discussion and conclusion

Certain aspects of a system's bulk behavior arise out of many small-326 scale interactions. These aspects may be investigated with simple models 327 that faithfully represent the bulk behavior, but do not consider directly the 328 full complexity of the system. In conjunction with more detailed models and 329 simulations, models which represent bulk behavior can also enhance our un-330 derstanding of complex dynamical systems. The model presented here, which 331 has aspects of species competition in biology, is a tool with which to examine 332 the complexity of cloud-rain interactions in a way that highlights certain as-333 pects of their behavior in a simple but realistic manner. This model is not 334 meant to supplant more detailed simulations of cloud-precipitation interac-335 tions but to represent certain predictable elements of the macrobehavior of 336 the full complex dynamical system. 337

In the first part of the paper, we examined some important mathematical 338 properties of the KF11 DDEs model which is a similar model devoted to 339 analyze the bulk behavior of a cloud-rain system. The focus has been put on 340 scaling and limit cycles or loop solutions. In particular, it has been shown, 341 through the Busenberg theorem that this model has no proper limit cycle so-342 lutions. The model we have introduced in this paper (second part) is another 343 point of view which retains the idea of modeling cloud-precipitation interac-344 tion through population dynamics and exhibits predator-prey behavior with 345 rain as predator and cloud droplets as prey. The stability analysis has shown 346 that our model has limit cycles and a Hopf bifurcation. 347

In particular, the model formulated here is a system of ODEs rather than DDEs. It is derived from a physically-based parametrization. Several examples illustrate that our model reproduces realistically the behaviour of an unique cloud-rain system. In particular, rain water content peaks are delayed by about 20 – 30 min with respect to cloud water content peaks. Some of the novel properties of the model include regions of parameter space in which clouds grow at the expense of rain and *vice versa* as determined from bifurcation analysis.

Further properties of our model such as ability to represent various physical situations, and sensitivity tests to parametrization schemes similar to those performed by Wacker (1995) for mixed and ice clouds, will be explored in more detail in future work. Due to its properties, this cloud-rain model is intended to be further developed for examination of cloud organization at larger scales in a coupled-cloud field.

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³⁶⁷ Appendix 1: Unit analysis

The system (1) is derived from four empirical equations which we have summarized in Table 1 together with their numbering in KF11.

Eq. number	Equation	Constants
in KF11		
[2]	$LWP = 0.5 c_1 H^2$	$c_1 = 2 \times 10^{-6} \mathrm{mm}\mathrm{m}^{-2}$
[5]	$\mathrm{d}LWP/\mathrm{d}t = -R$	None
[4]	$R = \alpha H^3 N_d^{-1}$	$\alpha = 2 \operatorname{mm} \operatorname{m}^{-6} \operatorname{d}^{-1}$
[9]	$\mathrm{d}N_d/\mathrm{d}t = -c_2 N_d R$	$c_2 = 0.3 \mathrm{m}^{-1}$

TAB. 1 – Equations and constants in KF11 discussed here. In KF11, LWP and R are liquid water path and rainrate, respectively. The corresponding units are explicitly $g m^{-2}$ and $mm d^{-1}$.

Eq. [2] connects liquid water path $(LWP, \text{ in } \text{g m}^{-2})$ with cloud depth (H, in m). The dimension of LWP is $[M][L]^{-2}$ (where [M] and [L] are mass and length dimensions), which implies that the dimension of c_1 is $[M][L]^{-4}$

and not $[L]^{-1}$ (cf. Table 1). Indeed, the well-established relationship is given by LWP (g m⁻²) = 0.5 $c_w H^2$, where $c_w = 2 \times 10^{-6} \text{ kg m}^{-4}$ (e.g. Geoffroy et al., 2008). Actually, $c_w = \rho_w c_1$, where $\rho_w \approx 10^3 \text{ kg m}^{-3}$ is liquid water density and $c_1 = 2 \times 10^{-9} \text{ m}^{-1}$ — in Tab. 1, $c_1 = 2 \times 10^{-6} \text{ mm m}^{-2}$, *i.e.* $c_1 (\text{m}^{-1}) = 10^{-3} c_1 (\text{mm m}^{-2})$. The relationship between LWP and H we get is thus:

$$LWP(g m^{-2}) = \frac{1}{2} c_w (kg m^{-4}) H^2(m^2)$$

= $\frac{1}{2} \rho_w (kg m^{-3}) c_1(m^{-1}) H^2(m^2)$ (14)

If we look now at Eq [5] and the units chosen, the decreasing rate of LWP due to rain cannot equal exactly the rainrate R (in mm d⁻¹). In fact, dLWP/dtis proportional to the precipitation flux density (or rain current $J_r = -\rho_w R$). The coefficient of proportionality is determined by the choice of units. If, as in KF11, we express time in minutes (it is the most natural timescale for cloud macroscopic physics), and keep LWP in g m⁻² as above, we obtain:

$$\frac{\mathrm{d}LWP}{\mathrm{d}t} (\mathrm{g}\,\mathrm{m}^{-2}\,\mathrm{min}^{-1}) = -\rho_w (\mathrm{g}\,\mathrm{m}^{-3}) \,R (\mathrm{m}\,\mathrm{min}^{-1})
= -10^3 \,\rho_w (\mathrm{kg}\,\mathrm{m}^{-3}) \,R (\mathrm{m}\,\mathrm{min}^{-1})
= -\frac{10^3}{60 \times 24} \,\rho_w (\mathrm{kg}\,\mathrm{m}^{-3}) \,R (\mathrm{m}\,\mathrm{d}^{-1})
= -\frac{1}{60 \times 24} \,\rho_w (\mathrm{kg}\,\mathrm{m}^{-3}) \,R (\mathrm{mm}\,\mathrm{d}^{-1})$$
(15)

since $R(\text{mmin}^{-1}) = 10^{-3} R(\text{mm} \text{d}^{-1})/(60 \times 24)$. Combining Eqs. (14) and (15) yields:

$$\frac{dH}{dt} (m \min^{-1}) = \frac{dLWP/dt (g m^{-2} \min^{-1})}{c_w (kg m^{-4}) H (m)}
= -\frac{\rho_w (kg m^{-3}) R (mm d^{-1})}{(60 \times 24) c_w (kg m^{-4}) H (m)}
= -\frac{1}{60 \times 24} \frac{R (mm d^{-1})}{c_1 (m^{-1}) H (m)}
= -\frac{10^3}{60 \times 24} \frac{R (mm d^{-1})}{c_1 (mm m^{-2}) H (m)}$$
(16)

It is possible, although not common, to divide LWP by ρ_w and then to define a volumetric liquid water path: LWP_v (m) = $10^{-3} LWP$ (g m⁻²)/ ρ_w (kg m⁻³). In this case, Eqs. [2] and [5] of Table 1 are correct provided that, for Eq. [5], Rbe expressed in m min⁻¹, or equivalently, considering $10^{-3} R (\text{mm d}^{-1})/(60 \times 24)$. However, we have not found any explicit mention of LWP in m or mm in KF11 and further papers; there LWP is explicitly expressed in g m⁻² several times.

The KF11 model also employs two parametrizations which relate R, N_d , and dN_d/dt (see Eqs. [4] and [9] in Table 1). The units of the latter two quantities are naturally cm⁻³ and cm⁻³ min⁻¹, respectively. According to the units used in KF11, the rhs of Eq. [4] is in mm m⁻³ cm³ d⁻¹, while the lhs (left hand side) is in mm d⁻¹. Although Eq. [4] is dimensionally correct $([L][T]^{-1}$, where [T] is the time dimension), the units of the rhs and lhs are not the same. Indeed, the relation $R = \alpha H^3/N_d$ should have N_d in m⁻³ when H is in m. If we choose to express N_d in cm⁻³, then we must multiply H^3/N_d by 10⁻⁶. With the unit of α , R is thus in mm d⁻¹:

$$R(\mathrm{mm\,d^{-1}}) = 10^{-6} \,\alpha(\mathrm{mm\,m^{-6}\,d^{-1}}) \,H^3(\mathrm{m^3}) \,N_d^{-1}(\mathrm{cm^3}) \tag{17}$$

Thus, for Eq. (16), we obtain:

$$\frac{\mathrm{d}H}{\mathrm{d}t} \left(\mathrm{m\,min^{-1}}\right) = -\frac{10^{-3}}{60 \times 24} \frac{\alpha \left(\mathrm{mm\,m^{-6}\,d^{-1}}\right) H^2 \left(\mathrm{m}^2\right)}{c_1 \left(\mathrm{mm\,m^{-2}}\right) N_d \left(\mathrm{cm^{-3}}\right)}$$
(18)

In addition, Wood (2006) (ref. 20 in KF11) gives a value of c_2 which does not correspond to that used in KF11. Aside from the conflation of rain current and rainrate in Wood (2006) — this article seems to use P as rainrate and divides it by liquid water density (ρ_w), whereas P should be the rain current and P/ρ_w the rainrate —, the coefficient of proportionality between dN_d/dt and R is $3E_0/4 = 3 \times 10^3 \text{ m}^{-1}$ since $E_0 = 4 \times 10^3 \text{ m}^{-1}$. This value is different from the 0.3 m^{-1} used in KF11 (assuming that P, in Wood (2006), is rainrate, and not rain current, which is very few probable due to a dimensional analysis of the Eq. 14 of Wood (2006), c_2 would be 3 m^{-1} , *i.e.* the previous value of c_2 divided by ρ_w , which is again different from the value of Table 1). Because, in Eq. [9], dN_d/dt has to be in cm⁻³ min⁻¹, we have:

$$\frac{\mathrm{d}N_d \,(\mathrm{cm}^{-3}\,\mathrm{min}^{-1})}{\mathrm{d}t} = c_2 \,N_d \,(\mathrm{cm}^{-3}) \,R \,(\mathrm{m}\,\mathrm{min}^{-1}) \tag{19}$$

with $c_2 = 3 \times 10^3 \,\mathrm{m}^{-1}$.

Then, because $R (\text{m min}^{-1}) = 10^{-3} R (\text{mm d}^{-1})/(60 \times 24)$, we obtain, by substituting the rainrate of Eq. (17) into Eq. (19):

$$\frac{\mathrm{d}N_d \left(\mathrm{cm}^{-3}\,\mathrm{min}^{-1}\right)}{\mathrm{d}t} = \frac{10^{-9}}{60 \times 24} \,c_2 \,(\mathrm{m}^{-1}) \,\alpha \left(\mathrm{mm}\,\mathrm{m}^{-6}\,\mathrm{d}^{-1}\right) H^3 \,(\mathrm{m}^3) \tag{20}$$

Finally, with this set of units, *i.e.* with H and N_d in m and cm⁻³ respectively, and time in minutes, the consistent (in terms of units) system of equations we get is:

$$\begin{cases} \frac{\mathrm{d}H}{\mathrm{d}t} = \frac{H_0 - H}{\tau_1} - \frac{10^{-3}}{60 \times 24} \frac{\alpha H^2(t - T)}{c_1 N_d(t - T)} \\ \frac{\mathrm{d}N_d}{\mathrm{d}t} = \frac{N_0 - N_d}{\tau_2} - \frac{10^{-9}}{60 \times 24} \alpha c_2 H^3(t - T) \end{cases}$$
(21)

with $c_1 = 2 \times 10^{-6} \,\mathrm{mm \, m^{-2}}$, $c_2 = 3 \times 10^3 \,\mathrm{m^{-1}}$, $\alpha = 2 \,\mathrm{mm \, m^{-6} \, d^{-1}}$. If we consider that LWP is a volumetric liquid water path, as it might be assumed as suggested above, and R in $\mathrm{mm \, d^{-1}}$, the factor 10^{-3} is replaced by 10^3 .

³⁸¹ Appendix 2: Other empirical parametrizations

Other empirical parametrizations can also be used to relate H, N_d , and 382 R. Here, we would like to suggest three. In particular, we can introduce 383 the following three alternative parametrizations for the delay term dN_d/dt 384 instead of the one in Wood (2006). (i) The empirical relationship proposed in 385 Mechem et al. (2006, Eq. 9): $dN_d/dt (cm^{-3} d^{-1}) = -69.4 (N_d R)^{0.668}$, with N_d 386 in cm^{-3} and R in $cm d^{-1}$. (ii) The following empirical relationships given 387 in Geoffroy et al. (2008): $J_r (10^{-6} \text{ kg m}^{-2} \text{ s}^{-1}) = a (LWP^{\alpha}/N_d) - b$. Here, 388 LWP is in $g m^{-2}$, a, b, and α are positive constants that take the values 389 summarized in Table 2. 390

a	α	b
0.3×10^{6}	2	10^{-6}
24.37×10^{9}	1	0
21.5×10^{3}	1.5	2.3×10^{-6}

TAB. 2 – Values of the constants in Geoffroy et al. (2008)

³⁹¹ When using this parametrization, it is necessary to first use Eq. (14) to ³⁹² change LWP into H.

(*iii*) Finally, a parametrization of the delay term dN_d/dt by taking

$$\frac{N_d}{LWP}\frac{\mathrm{d}LWP}{\mathrm{d}t} = \frac{2N_d}{H_r}\frac{\mathrm{d}H_r}{\mathrm{d}t}$$
(22)

This follows from Eq. (14). The equation for dH_r/dt can then be substituted into this equation to complete the parametrization. This equation with delay is finally substituted for the delay term in the second equation in the KF11system.

When using these different parametrizations, a set of coupled first-order 397 DDEs similar to Eq. (1) is obtained. The units must be made consistent 398 when the various empirical relationships are employed. Namely, H is in m, 399 N_d is in cm⁻³, and time is expressed in minutes. For the numerical tets we 400 have performed, these parametrizations have given results similar to those 401 in Figure 1. However, we have not made a complete test of sensitivity of 402 the model to changes in parametrization since it is not the purpose of this 403 paper. Our goal here is to suggest other possibilities to parametrize cloud-404 rain interaction processes that could be investigated more deeply in future 405 research. 406

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FIG. 1 – Evolution of H and N_d according to system (1) with $H_0 = 530 \text{ m}$, $N_0 = 180 \text{ cm}^{-3}$, $\tau_1 = \tau_2 = 60 \text{ min}$, and T = 12 min. Initial conditions are H(t = 0) = 200 m and $N_d(t = 0) = 50 \text{ cm}^{-3}$. We have used the pydelay package of Python.



FIG. 2 – Evolution of N_d , L_c and L_r (a) and limit-cycle in the configuration space (L_c, L_r) (b) for the initial conditions and values of parameters given in the text. On the limit cycle, initial and final time of integration are at the origin and (0.25,0) respectively. (c) Zoom of (a).



FIG. 3 – Same as Fig. 2 but for $N_0 = 5 \text{ cm}^{-3}$. On the limit cycle, initial time of integration is at the origin and final time of integration is represented by the black small ball at about (0.28, 0.05).



FIG. 4 – Another kind of damped cloud system. On the limit cycle, initial time of integration is at the origin and final time of integration is represented by the black small ball at about (0.2, 0.05).