



HAL
open science

Noble gas systematics in new popping rocks from the Mid-Atlantic Ridge (14 °N): evidence for small-scale upper mantle heterogeneities

Sandrine Péron, Manuel Moreira, Mark D. Kurz, Joshua Curtice, Jerzy S. Blusztajn, Benita Putlitz, V. Dorsey Wanless, Meghan R Jones, Samuel Adam Soule, Eric Mittelstaedt

► To cite this version:

Sandrine Péron, Manuel Moreira, Mark D. Kurz, Joshua Curtice, Jerzy S. Blusztajn, et al.. Noble gas systematics in new popping rocks from the Mid-Atlantic Ridge (14 °N): evidence for small-scale upper mantle heterogeneities. *Earth and Planetary Science Letters*, 2019, 519, pp.70-82. 10.1016/j.epsl.2019.04.037 . hal-03468684

HAL Id: hal-03468684

<https://hal.science/hal-03468684v1>

Submitted on 7 Dec 2021

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

1 **Title:** Noble gas systematics in new popping rocks from the Mid-Atlantic Ridge (14 °N):
2 evidence for small-scale upper mantle heterogeneities

3

4 **Authors:** Sandrine Péron^{a*}, Manuel A. Moreira^a, Mark D. Kurz^b, Joshua Curtice^b, Jerzy S.
5 Blusztajn^c, Benita Putlitz^d, V. Dorsey Wanless^e, Meghan R. Jones^f, S. Adam Soule^e, Eric
6 Mittelstaedt^g

7

8 **Affiliations:**

9 ^a Institut de Physique du Globe de Paris - Sorbonne Paris Cité, UMR CNRS 7154, Université
10 Paris Diderot. 1 Rue Jussieu, 75005, Paris, France

11 ^b Marine Chemistry and Geochemistry, MS #25, Clark 421, Woods Hole Oceanographic
12 Institution, Woods Hole, MA 02543, United States

13 ^c Department of Geology and Geophysics, Woods Hole Oceanographic Institution, 02543,
14 Woods Hole, MA, USA

15 ^d Institute of Earth Sciences, University of Lausanne, Géopolis Building, CH-1015 Lausanne,
16 Switzerland

17 ^e Department of Geosciences, Boise State University, 83725, Boise, ID, USA

18 ^f Massachusetts Institute of Technology/Woods Hole Oceanographic Institution Joint Program
19 in Oceanography, 02543, Woods Hole, MA, USA

20 ^g Department of Geological Sciences, University of Idaho, 83844, Moscow, ID

21

22 **Corresponding author:** Sandrine Péron, peron@ipgp.fr

23

24 **Abstract:** The origin and evolution of terrestrial volatiles is a fundamental question in
25 Earth and planetary sciences. To better address this, it is necessary to estimate the present-day
26 isotopic and elemental composition of volatiles in the mantle, which is difficult because
27 degassing during magma transport to the surface changes the initial elemental composition.
28 However, some basalts are considered to be relatively un-degassed and therefore can be used
29 to determine the concentrations of C, N, H₂O and noble gases in the mantle. This is the case
30 of the so-called popping rock sample 2πD43, found near 14 °N on the Mid-Atlantic Ridge in
31 1985, which has been considered as a reference for upper mantle isotopic and elemental
32 volatile compositions due to its high gas content. Such gas-rich samples have been recovered
33 in only a few places on mid-ocean ridges but are crucial to characterizing the heterogeneities
34 in mantle composition and therefore the cycling of volatiles on Earth. In order to better
35 understand the occurrence of popping rocks, mantle heterogeneities and volatile cycling on
36 Earth, cruise AT33-03 on *R/V Atlantis* was conducted in 2016 on the Mid-Atlantic Ridge near
37 14 °N. Popping rocks were found in the same area as sample 2πD43, using the submersible
38 *Alvin*, and their noble gas compositions were determined by step-crushing and laser ablation.
39 A new anaerobic sampling protocol was tested in an attempt to minimize atmospheric
40 contamination, which is ubiquitous in basaltic glass samples used for noble gas analyses, and
41 to evaluate laser extraction techniques. Preliminary step-crushing results suggest that the new
42 anaerobic sampling protocol reduces atmospheric influence on the noble gas measurements.
43 Comparison with laser ablation data shows that there is still some atmospheric contamination,
44 suggesting that contamination begins on the seafloor. Two groups of samples have been
45 identified based on their noble gas compositions: one group with their highest measured ratios
46 similar to those of sample 2πD43 (²⁰Ne/²²Ne of ~ 12.5, ⁴⁰Ar/³⁶Ar of ~ 27,000), and another
47 group with lower maximum noble gas ratios (²⁰Ne/²²Ne of 12.2, ⁴⁰Ar/³⁶Ar of 16,100) and
48 higher argon concentrations. Based on these results, along with Cl/K and H₂O/Ce ratios and

49 preliminary lead, strontium and neodymium isotope data, it is suggested that these two groups
50 of samples reflect separate eruptive events that originated from distinct mantle sources. This
51 provides evidence for upper mantle heterogeneity at small scales and could imply that the
52 mantle source with lower neon and argon isotopic ratios has been more influenced by
53 recycling of atmospheric noble gases.

54

55 **Keywords:** popping rock, mid ocean ridge basalt, noble gases, Mid-Atlantic Ridge,
56 heterogeneous mantle

57 **I) Introduction**

58

59 The isotopic and elemental compositions of mantle volatiles are key to understanding the
60 origin and cycling of volatiles between the mantle and the atmosphere and Earth dynamics
61 since the Hadean (e.g., Holland and Ballentine, 2006; Kurz et al., 2009; Moreira et al., 1998;
62 Mukhopadhyay, 2012; Tieloff et al., 2000). However, the present-day volatile elemental
63 composition of the mantle is difficult to determine because volatiles are sensitive to near-
64 surface degassing (e.g., Burnard et al., 2003; Moreira and Kurz, 2013; Sarda and Moreira,
65 2002). This magmatic process is hard to quantify since it depends on several physicochemical
66 conditions that are not well constrained. Nevertheless, it is considered that some rare samples
67 did not suffer degassing, such as popping rocks (e.g., Sarda and Graham, 1990). Popping
68 rocks are unusually gas-rich oceanic lava samples that pop after recovery from the seafloor
69 due to vesicle gas decompression. Due to their high gas contents, they represent crucial
70 samples for studying the volatile isotopic and elemental composition of the mantle and the
71 associated heterogeneities (Javoy and Pineau, 1991; Moreira et al., 1998; Sarda and Graham,
72 1990).

73 Popping rocks have been found in only a few places, in particular mid-oceanic ridge
74 basalts (MORBs) on the Mid-Atlantic Ridge at 36 °N (sample CH-DR11-315X, Hekinian et
75 al., 1973; Pineau et al., 1976) and the well-studied sample 2πD43 near 14 °N (Bougault et al.,
76 1988; Sarda and Graham, 1990). Due to its high gas content and vesicularity of 16 %, it is
77 often assumed that sample 2πD43 did not suffer vesicle loss and is often considered as
78 representative of the upper mantle noble gas isotopic and elemental compositions (Burnard et
79 al., 1997; Kunz, 1999; Kunz et al., 1998; Moreira et al., 1998; Raquin et al., 2008; Sarda and
80 Graham, 1990; Staudacher et al., 1989) and carbon abundance (Javoy and Pineau, 1991),
81 despite its relative enrichment in incompatible elements (Bougault et al., 1988; Dosso et al.,

82 1991). Chavrit et al. (2014) suggested that such characteristics are more consistent with a
83 process of bubble accumulation, given the higher number of big vesicles compared with other
84 MORBs.

85 However, the location of sample 2πD43 on the Mid-Atlantic Ridge was not precisely
86 known, because it was obtained by dredging, and the tectonic context could not be fully
87 interpreted. A new cruise (AT33-03) was conducted in 2016 to find popping rocks in this area
88 in order to better understand the formation of such gas-rich samples and the cycling of
89 volatiles on Earth.

90 Basaltic glass samples are often contaminated by a seawater- or air-derived component,
91 likely through cracks (Ballentine and Barfod, 2000; Roubinet and Moreira, 2018), which
92 complicates noble gas interpretations. In this study, a new protocol was tested to recover the
93 samples in anaerobic samplers, avoiding contact with air.

94 We report noble gas compositions of new popping rocks recovered in 2016 on the Mid-
95 Atlantic Ridge determined both *via* step-crushing and laser ablation analyses. Laser ablation
96 noble gas measurements have proved to be an effective technique to reduce atmospheric
97 contamination issues (Raquin et al., 2008), so that results of laser ablation and step-crushing
98 can be compared to evaluate the efficiency of the new anaerobic protocol and to assess the
99 mechanisms of atmospheric contamination. The results are also compared with existing and
100 new data for sample 2πD43, which was re-analyzed *via* laser ablation. These new data
101 provide insight into isotopic and elemental heterogeneities of the mantle and the cycling of
102 volatiles on Earth.

103

104

105

106

107

108 **II) Samples and analytical procedure**

109

110 **2.1 AT33-03 cruise and geological settings**

111

112 Cruise AT33-03 of *R/V Atlantis* took place in 2016 on the Mid-Atlantic Ridge and studied
113 the area between 13°45' N and 14°06' N, with the aim of finding popping rocks and relating
114 their occurrence to the tectonic setting (Jones et al., 2019; Parnell-Turner et al., 2018).
115 Popping rock sample 2πD43 was recovered by dredging close to 13°46' N (Bougault et al.,
116 1988) but the location on the ridge was not precisely known due to the uncertainty associated
117 with dredging and navigation.

118 The southern part of the studied area, between 13°45' N and 13°52' N, is a non-transform
119 discontinuity with limited magmatic activity, and topographically irregular flanks, whereas
120 the northern part shows clear volcanic activity (Smith et al., 2008). There is an extinct oceanic
121 core complex in the southern part (at ~ 13°50' N, 44°56' W, figure 1) as identified by Smith
122 et al. (2008) and studied by Parnell-Turner et al. (2018), that is northeast of the approximate
123 location of popping rock 2πD43 (Bougault et al., 1988).

124 High-resolution bathymetric maps of the seafloor were obtained via the autonomous
125 underwater vehicle (AUV) *Sentry* and sampling was performed in-situ with the human
126 occupied vehicle (HOV) *Alvin* (Jones et al., 2019; Parnell-Turner et al., 2018).

127

128 **2.2 Samples and protocol on the ship**

129

130 Five samples from the AT33-03 cruise were analyzed in this study (figure 1), by both step
131 crushing and laser ablation for single vesicle analyses, samples AL4818-003, AL4820-036,

132 AL4820-041, AL4821-055 and AL4821-058 (abbreviated hereafter as 003, 036, 041, 055 and
133 058 respectively). Major and trace element compositions as well as major volatile contents
134 (CO₂, H₂O, Cl, F, S) and averaged helium isotope ratios measured by crushing are presented
135 elsewhere (Jones et al., 2019). These five samples are highly enriched in incompatible
136 elements and have chondrite-normalized La/Sm ratios of about 1.8 (Jones et al., 2019),
137 indicating that they represent E (enriched)-MORB. Samples 003, 036, 055 and 058 were
138 identified as popping rocks because they visibly and audibly popped on the ship. Sample 041
139 was collected from an adjacent pillow mound and has intermediate vesicularity (4.4 % - 6 %,
140 Jones et al. (2019) and table S2), but was not observed to pop on the ship. We also analyzed
141 one piece of the popping rock sample 2πD43 with laser ablation to ensure that we were able
142 to reproduce previous data (Moreira et al., 1998; Raquin et al., 2008).

143 In order to reduce atmospheric contamination of the glass samples for noble gas analyses,
144 we tested a new protocol on the ship. Glass samples can be contaminated by air through
145 cracks and opened vesicles (e.g., Ballentine and Barfod, 2000; Roubinet and Moreira, 2018)
146 and this contaminant mixes with mantle gases when samples are crushed, which complicates
147 noble gas interpretations. During sampling on the seafloor, two of the samples were recovered
148 in sealed containers filled with ambient bottom-water. On the ship, the seawater was flushed
149 out with nitrogen and the container was set inside a glove box filled with nitrogen. The glass
150 from the sample could then be chipped away and stored into a closed stainless-steel tube,
151 which is pumped until the pressure drops to $\sim 10^{-5}$ - 10^{-6} mbar. These samples were never in
152 contact with air. Samples 055 and 058 were recovered in these special sealed containers
153 following this ‘anaerobic’ sampling protocol and were analyzed by step-crushing.

154 We also stored some glass samples (recovered in the sealed sea-floor containers and
155 handled in a nitrogen atmosphere) in plastic bags that were pumped with a commercial
156 culinary vacuum sealing machine to remove as much air as possible. These glass samples

157 were exposed to air for only a few minutes (1-2 min), prior to vacuum packaging. This was
158 the case for pieces of samples 055 and 058 analyzed by laser ablation.

159 The other samples were recovered in open baskets on the HOV *Alvin* and were hence
160 directly in contact with air after *Alvin* recovery. Some of them were also stored in the plastic
161 bags and pumped 15-20 hours after *Alvin* recovery. This was the case for pieces of samples
162 003 and 036 analyzed by laser ablation.

163

164 **2.3 Laser ablation analyses**

165

166 Laser ablation noble gas analyses were conducted at Institut de Physique du Globe de
167 Paris (IPGP). Pieces of samples 003, 036, 055 and 058 used for laser ablation were stored in
168 plastic bags (Section 2.2) on the ship and were then put into the glove box filled with nitrogen
169 in the IPGP laboratory. The piece of sample 041 used for laser ablation was not stored in
170 vacuum and was therefore fully exposed to air.

171 Before laser ablation analyses, samples were removed from the glove box to be polished
172 and then scanned via X-ray microtomography. This 3D imaging technique allows precise
173 location of the intact bubbles before targeting with the laser for noble gas analyses. The
174 tomography images (Figures S1, S2, S3 and S4) were acquired with a Bruker Skyscan 1173 at
175 the Institute of Earth Sciences, University of Lausanne and the parameters used for the
176 acquisitions are provided in Supplementary Information (Table S1). The protocols for sample
177 preparation, image reconstruction and calculation of vesicularities (table S2) are described in
178 Péron et al. (2016) and Péron et al. (2017). The four popping rock samples have estimated
179 vesicularities between 9 % (sample 003) and 21 % (samples 036 and 055), whereas sample
180 041 has a vesicularity of 6 % (table S2). These estimates are a few percent lower or higher
181 than those of Jones et al. (2019). This could be due to heterogeneity in vesicularity between

182 sub-samples, or it could also be an effect of image processing. In this study, we did not
183 smooth the images, in order to preserve sample characteristics and to distinguish small cracks,
184 even though smoothing refines vesicle contours and is useful for calculation of vesicularities
185 (Péron et al., 2016). The vesicularity of sample 2πD43 was calculated to be 16 %, which is
186 consistent with the values of 16% - 17% found by Sarda and Graham (1990) and Javoy and
187 Pineau (1991).

188 Laser ablation noble gas analyses were conducted following a similar protocol as
189 described by Péron et al. (2016) and Péron et al. (2017) and are summarized briefly here.
190 After X-ray microtomography acquisition, samples were cleaned in ethanol and acetone baths
191 and then with oxalic acid (1 %, at 60-80 °C) if iron-manganese crusts were observed. Samples
192 were then loaded into the laser cell, which was baked at 100-150 °C for at least one month, to
193 minimize line blanks, before starting laser ablation. Sample pieces used for laser ablation
194 were thus exposed to air for several weeks. Laser ablation measurements were started as soon
195 as the line blanks decreased to about 4.9×10^{-10} ccSTP of ^4He , 4.6×10^{-13} ccSTP of ^{22}Ne and
196 5.8×10^{-12} ccSTP of ^{36}Ar , and they kept decreasing thereafter.

197 The laser used is an argon fluoride (excimer) laser, ATLEX-300i instrument, of
198 wavelength 193 nm and the pressure in the laser cell was monitored with a MKS Baratron®
199 manometer. The pressure increase generated when reaching a bubble was calibrated with an
200 air standard, which allows calculation of CO_2 abundances in the vesicles assuming that CO_2 is
201 the major gas (Moore et al., 1977).

202 After a vesicle was breached by laser ablation, gases were purified with two Ti getters,
203 one was heated to 800 °C for five minutes and then set to ambient temperature as the other
204 step, before being trapped (except He) on a cold head at about 20 K. Noble gases were then
205 analyzed on a Noblesse mass spectrometer (Nu instruments ©) (Supplementary Information).
206 To correct for interferences on neon isotopes, as explained in Péron et al. (2017), we carried

207 out a scan of HF before introducing the neon into the mass spectrometer and then measured
208 the CO₂ and ⁴⁰Ar signals during each cycle. The ionization ratios ⁴⁰Ar⁺⁺/⁴⁰Ar⁺ and
209 CO₂⁺⁺/CO₂⁺ were determined on the Noblesse mass spectrometer to be 0.116 and 0.013
210 respectively during the analysis period (Figures S6, S7, S8 and S9).

211 As explained in Péron et al. (2016) and in Péron et al. (2017), laser ablation analyses are
212 corrected with the matrix blanks (Tables S3 and S4). We analyzed the matrix gas composition
213 after laser ablation of vesicle-free glass for a representative time. As described in
214 Supplementary Information, matrix isotopic compositions are atmospheric and do not depend
215 on the ablation time. We used a matrix blank that was measured just before or just after a
216 vesicle for corrections. For the majority of the vesicles, the matrix blank contributions were
217 less than 0.3 % for He and less than 30 % for Ne and Ar.

218 Two standards were used for laser ablation analyses (Supplementary Information), an air
219 standard (²⁰Ne/²²Ne = 9.80, ²¹Ne/²²Ne = 0.029, ³⁸Ar/³⁶Ar = 0.188, ⁴⁰Ar/³⁶Ar = 295.5) and a
220 mixed gas standard for He and Ne (Péron et al., 2016).

221

222 **2.4 Step-crushing analyses**

223

224 Step-crushing analyses were conducted at Woods Hole Oceanographic Institution
225 (WHOI) on clean, centimeter-sized, glass fragments, typically 1-2 grams in size. Helium and
226 neon were analyzed using a MAP 215-50 mass spectrometer in the WHOI Isotope
227 Geochemistry Facility, as described by Kurz et al. (2005) and Kurz et al. (2009). The helium
228 standard has a ³He/⁴He ratio of 8.35 Ra with $\sim 5 \times 10^{-9}$ cm³ STP ⁴He and was calibrated
229 against air aliquots; air was used as the neon and argon standard ($1-2 \times 10^{-10}$ cm³ STP neon).
230 Crushing blanks are typically 1×10^{-11} cm³ STP ⁴He and 2×10^{-12} cm³ STP ²⁰Ne. Interferences
231 at masses 20 and 22 were minimized with a charcoal trap held at liquid nitrogen temperatures

232 within the flight tube volume. Corrections for interferences from $^{40}\text{Ar}^{++}$ and CO_2^{++} on ^{20}Ne
233 and ^{22}Ne peaks were made using measurements at masses 40 and 44; no correction was
234 necessary for interferences from H_2^{18}O and HF because peaks at masses 18 and 19 were
235 negligible. The ratios of doubly charged/singly charged contributions, on masses 20 and 22,
236 for masses 40 and 44 were typically 0.23 and 0.012; the 40 and 44 peaks were generally small
237 relative to the sample gas, making these corrections unimportant (mass 40 and 44 peaks were
238 typically less than 100 counts per second during these measurements). Stated uncertainties are
239 1 sigma, calculated as quadrature sum of estimated analytical uncertainties. Argon isotopes
240 were measured using a dedicated Quadrupole Mass Spectrometer (Hiden), attached to the
241 same extraction line. Helium, neon and argon were separated using two automated cryogenic
242 traps, one nude stainless steel and one charcoal. All measurements were made by crushing *in*
243 *vacuo*, using well established procedures (e.g., Kurz et al., 2009). The quantities of gas
244 introduced into the extraction line and mass spectrometers were controlled by capacitance
245 manometry, followed by a pre-measurement with a quadrupole mass spectrometer, and
246 volumetric splitting as necessary to match sample size with standard size. Pre-gettering
247 capacitance manometry measurements were used to calculate $\text{CO}_2/{}^3\text{He}$ ratios (table 3),
248 assuming that CO_2 constitutes most of the volatiles released by crushing (Moore et al., 1977).
249 Noble gas concentrations are so high in these samples that they could be measured many
250 times by crushing.

251

252

253

254

255

256

257 **III) Results**

258

259 **3.1 Comparison of laser ablation and step-crushing data**

260

261 Of the 27 individual vesicles analyzed by laser ablation, most of the data are similar to the
262 highest neon and argon isotopic values obtained by step-crushing (Figures 2, 3, 4, 5 and tables
263 1, 2 and 3). Overall, there is excellent agreement between the highest step-crushing values
264 and most laser ablation measurements.

265 Six vesicles show anomalous variations in elemental and isotopic compositions, compared
266 to the bulk crushing measurements, which are not easily explained by a single process.
267 Vesicles V5 of sample 055 and V2 of sample 2πD43 have strongly fractionated He, Ne and
268 Ar composition. The neon and argon isotopic ratios are close to air, $^3\text{He}/^{22}\text{Ne}$ and $^4\text{He}/^{40}\text{Ar}^*$
269 ratios are close to 0 and they have lower $^{22}\text{Ne}/^{36}\text{Ar}$ ratios than for other vesicles from the same
270 samples (table 2, figures 2, 3, 4 and 5). Pressure increases were detected by manometry (on
271 laser ablation) in both cases, but are much lower than expected given the vesicle volumes
272 (table 1). Vesicle V4 of sample 041 also yielded a lower-than-expected pressure jump and
273 clearly shows He, Ne and Ar loss, based on very low He, Ne and Ar contents in comparison
274 with other vesicles of similar sizes (table 1). Vesicle V2 of sample 058 has clearly lost
275 helium, based on low ^4He content and extremely low $^3\text{He}/^{22}\text{Ne}$ and $^4\text{He}/^{40}\text{Ar}^*$, but has a neon
276 and argon isotopic compositions similar to the other vesicles of sample 058 (tables 1 and 2).
277 Vesicle V8 of sample 036 seems to have lost helium (based on lower $^4\text{He}/^{40}\text{Ar}^*$, table 2) and
278 was too small to properly measure the neon (table 2).

279 Vesicle V2 of sample 055 also shows some He loss (lower $^4\text{He}/^{40}\text{Ar}^*$ of 0.93, relative to
280 all other measurements, a higher $\text{CO}_2/^3\text{He}$ ratio and a slightly higher $^3\text{He}/^4\text{He}$ ratio) but has

281 neon and argon isotopic ratios similar to the other vesicles of sample 055 and so the neon and
282 argon compositions are used along with data from other vesicles in section 3.3.

283 No clear cracks were identified close to the six anomalous bubbles on the X-ray
284 microtomography images but the image resolutions are insufficient to detect cracks smaller
285 than around 10 μm (table S1). The processes that fractionated the noble gas abundances and
286 isotopic compositions in these six vesicles are not known, but these data are not interpreted as
287 magmatic values.

288 It is also interesting to note that vesicle V9 of sample 036 is a very big vesicle that may
289 have formed by coalescence of smaller vesicles, due to its irregular (non-spherical) shape
290 (Figure S3). Vesicle V9 gas contents (CO_2 , He, Ne, Ar) are lower than expected given its
291 volume but its elemental and isotopic ratios are similar to that of the other vesicles, which
292 could be due to the coalescence process.

293

294 **3.2 CO_2 , helium and radiogenic argon $^{40}\text{Ar}^*$**

295

296 Excluding the six anomalous vesicles mentioned in section 3.1, vesicles from samples
297 003, 036, 055 and 058 have a mean $\text{CO}_2/{}^3\text{He}$ ratio of $3.35 (\pm 0.17) \times 10^9$ (1σ) (table 2), which
298 agrees well with the value of $3.08 (\pm 0.23) \times 10^9$ from bulk step-crushing (table 3). Vesicles
299 from sample 041 have a mean $\text{CO}_2/{}^3\text{He}$ ratio of $3.01 (\pm 0.22) \times 10^9$ (1σ) ($2.53 (\pm 0.21) \times 10^9$
300 with step-crushing) and vesicles of sample 2 π D43 a mean value of $3.17 (\pm 0.23) \times 10^9$ (1σ).
301 These values are higher than the estimated median value of 2.2×10^9 (Marty and Tolstikhin,
302 1998) or the recent average MORB estimate of $1.67 (\pm 0.21) \times 10^9$ (Tucker et al., 2018). They
303 are similar to the value estimated for E-MORB of $3.3 (\pm 1.1) \times 10^9$ (or $4.5 (\pm 1.3) \times 10^9$
304 corrected for fractional degassing) (Marty and Tolstikhin, 1998). This is consistent with these
305 samples being E-MORB, as also suggested by their major and trace element compositions

306 (Jones et al., 2019) and by the preliminary Pb, Sr, Nd data (section 4.3 and table S5). The new
307 data therefore support a higher $\text{CO}_2/{}^3\text{He}$ for E-MORB than for N (normal)-MORB.

308 The helium isotopic compositions are quite homogenous among all the vesicles, excluding
309 the six anomalous vesicles, which have ${}^3\text{He}/{}^4\text{He}$ both lower and higher than other
310 measurements (Table 4). The mean ${}^3\text{He}/{}^4\text{He}$ ratio is 8.33 ± 0.17 Ra (1σ) for samples 003, 036,
311 055 and 058, with Ra indicating the atmospheric ${}^3\text{He}/{}^4\text{He}$ ratio (1.384×10^{-6} , (Clarke et al.,
312 1976)). The mean ${}^3\text{He}/{}^4\text{He}$ ratio is 8.48 ± 0.11 Ra (1σ) for sample 041 and 8.48 ± 0.15 Ra
313 (1σ) for sample 2 π D43. The averages are lower, but within uncertainties for step-crushing
314 data, with ${}^3\text{He}/{}^4\text{He}$ ratio of 8.16 ± 0.18 Ra (1σ) for samples 003, 036, 055 and 058 and $8.32 \pm$
315 0.13 Ra (1σ) for sample 041. The ${}^3\text{He}/{}^4\text{He}$ from the six anomalous vesicles range from 0.98 to
316 10 Ra, suggesting extensive isotope fractionation by helium migration into and out of the
317 vesicles.

318 Except for the six anomalous vesicles, the mean vesicle ${}^4\text{He}/{}^{40}\text{Ar}^*$ ratio obtained by laser
319 ablation is 1.14 ± 0.04 (1σ) for samples 003, 036, 055 and 058, whereas it is 1.39 ± 0.06 for
320 sample 041 and 1.34 ± 0.05 for sample 2 π D43. The average values with step-crushing
321 measurements are 1.06 ± 0.07 (1σ) for samples 003, 036, 055 and 058 and 1.32 ± 0.07 (1σ)
322 for sample 041. The laser ablation and step-crushing values are therefore in good agreement.

323

324 **3.3 Neon and argon isotopic ratios**

325

326 Figures 2 and 3 show the neon and argon isotopic data for step-crushing and laser ablation
327 measurements. Sample 041 has neon and argon isotopic compositions similar to sample
328 2 π D43 (green dots and results from Moreira et al. (1998)) with maximum ${}^{20}\text{Ne}/{}^{22}\text{Ne}$ values of
329 around 12.5, ${}^{21}\text{Ne}/{}^{22}\text{Ne}$ of approximately 0.060 and ${}^{40}\text{Ar}/{}^{36}\text{Ar}$ of 26,000-27,000.

330 The other samples (003, 036, 055 and 058) show distinctly lower (maximum) neon and
331 argon isotopic ratios. The mean $^{20}\text{Ne}/^{22}\text{Ne}$ ratios are 12.17 ± 0.06 (1σ) for step-crushing (steps
332 with $^{20}\text{Ne}/^{22}\text{Ne}$ ratio > 12) and 12.24 ± 0.22 (1σ) for single vesicle analyses (vesicles with
333 $^{20}\text{Ne}/^{22}\text{Ne}$ ratio > 12). Similarly, $^{21}\text{Ne}/^{22}\text{Ne}$ for crushing steps and laser analyses (each
334 restricted to $^{20}\text{Ne}/^{22}\text{Ne}$ ratio > 12) display ratios of around 0.0567 and 0.0578 respectively,
335 and the mean $^{40}\text{Ar}/^{36}\text{Ar}$ ratio is $16,088 \pm 262$ (1σ) for the single vesicle analyses. The latter
336 value is consistent with the highest $^{40}\text{Ar}/^{36}\text{Ar}$ ratios obtained by step-crushing. It is important
337 to note that strong CO_2 memory effects in the mass spectrometer were observed on the neon
338 analysis of vesicle V4 from sample 2 π D43 because it was analyzed just after vesicle V3 and
339 so ^{22}Ne was poorly corrected for CO_2 , which makes this data point unreliable (table 2 and
340 figure S7). Vesicle V2 of sample 041 was too small to measure reliable neon isotopic ratios.

341 In figure 3, it is interesting to note that samples 055 and 058, recovered in the anaerobic
342 samplers, plot on a hyperbola with a higher curvature r (r is $(^{36}\text{Ar}/^{22}\text{Ne})_{\text{air}}$ or
343 $_{\text{seawater}}/(^{36}\text{Ar}/^{22}\text{Ne})_{\text{mantle}}$), compared with sample 003 (which was fully air exposed). Vesicles
344 V5 of sample 055 and V2 of sample 2 π D43 plot on a hyperbola with a lower curvature (figure
345 3), similar to that suggested by Raquin et al. (2008).

346

347 3.4 Elemental ratios

348

349 Figures 4 and 5 present the two elemental ratios $^3\text{He}/^{22}\text{Ne}$ and $^{36}\text{Ar}/^{22}\text{Ne}$ plotted against
350 $^{20}\text{Ne}/^{22}\text{Ne}$. The $^3\text{He}/^{22}\text{Ne}$ and $^{36}\text{Ar}/^{22}\text{Ne}$ ratios of sample 041 are similar to sample 2 π D43 and
351 are about 4.52 ± 0.09 (1σ , steps with $^{20}\text{Ne}/^{22}\text{Ne}$ ratio > 12) and 12.10 ± 0.34 (1σ , vesicles)
352 respectively. The other samples (003, 036, 055 and 058) show different $^3\text{He}/^{22}\text{Ne}$ and
353 $^{36}\text{Ar}/^{22}\text{Ne}$ ratios of 4.04 ± 0.19 (1σ , steps with $^{20}\text{Ne}/^{22}\text{Ne}$ ratio > 12) and 23.7 ± 2.3 (1σ ,
354 vesicles) respectively.

355 Most of the vesicle data have higher $^3\text{He}/^{22}\text{Ne}$ ratios than the highest step-crushing data
356 but the uncertainties are relatively large and so the step-crushing data seem to best determine
357 the $^3\text{He}/^{22}\text{Ne}$ ratio.

358 In figure 5 and S5 in Supplementary Information, step-crushing data for sample 003
359 suggest mixing between air and a mantle component, whereas step-crushing data for samples
360 041, 055 and 058 show mixing with a seawater-derived component, based on the different
361 abundance ratios. Vesicles V5 of sample 055 and V2 of sample 2πD43 show mixing between
362 air and seawater.

363

364 **IV) Discussion**

365

366 **4.1 Sample contamination processes and the use of laser ablation measurements**

367

368 Our results suggest that the protocol used to collect and analyze anaerobic samples (those
369 without contact with air) is effective in reducing shallow air contamination. It is particularly
370 evident in the case of neon (figures 2, 3, 4, 5 and S5), as observed in step-crushing data for
371 anaerobic samples 055 and 058 compared with sample 003, which was fully exposed to air.
372 The majority of the crushing steps for samples 055 and 058 are close to the highest values, in
373 contrast to the crush steps for sample 003. Sample 041 was also fully exposed to air. The
374 crushing measurements of sample 041 show limited neon and argon contamination, which
375 could be related to its lower vesicularity (table S2). This limited sample suite shows promise
376 for the method, but additional testing on more samples is needed.

377 As shown in figures 5 and S5, step-crushing data for samples 041, 055 and 058 show
378 mixing with a seawater-derived component rather than with air, as is observed for sample
379 003, even if most of the steps for samples 041, 055 and 058 show limited contamination.

380 Interaction with a seawater-derived component is also suggested by the mixing hyperbola of
381 higher curvature ($r \sim 5$) in figure 3 because the seawater $^{36}\text{Ar}/^{22}\text{Ne}$ ratio is higher than that of
382 air (figure 5). This suggests that seawater contamination begins on the seafloor during or after
383 emplacement, likely through cracks (Ballentine and Barfod, 2000), and that air quickly
384 replaces this seawater-derived component after collection. Indeed, sample 003, which was
385 fully air exposed after collection and has a high vesicularity, is more influenced by air than
386 seawater (figures 3 and 5), in contrast to the anaerobically collected samples. Additional
387 results from this new protocol will be provided in a subsequent paper.

388 The excellent agreement between the laser ablation and the highest step-crushing data
389 (section 3.1) suggests that both techniques are effective in determining the mantle
390 composition. A few laser ablation data show variability from the step-crushing results (figures
391 2, 3, 4, 5), highlighting that processes of contamination are complex, as also discussed by
392 Roubinet and Moreira (2018). Hence, laser ablation analyses represent a powerful method for
393 better understanding this contamination process compared to step-crushing. It is important to
394 note that some bubbles could not be analyzed, either because the laser spot was positioned
395 incorrectly or more likely because no pressure increase was detected. The latter case would
396 indicate that these bubbles were completely contaminated with atmosphere/seawater gases,
397 maybe through cracks, prior to analysis and that no magmatic CO_2 was left. No pressure jump
398 can be detected with our system if a bubble is at atmospheric pressure.

399 As mentioned in section 3.3, vesicles V5 of sample 055 and V2 of sample 2 π D43 show
400 mixing between air and seawater given their $^{22}\text{Ne}/^{36}\text{Ar}$ ratios, and plot on the hyperbola with a
401 low curvature in figure 3. Vesicle V2 of sample 058 and vesicle V4 of sample 041 also seem
402 to plot on hyperbolas with low curvatures ($r < 1$). Other vesicles (V2 and V3 of sample 003,
403 V2 of sample 036, V1 and V4 of sample 058) like V2 of sample 058 also seem to have a
404 $^{20}\text{Ne}/^{22}\text{Ne}$ ratio slightly lower than 12 and fall on the hyperbola with a low curvature in figure

405 3. We can exclude analytical issues for these bubbles (Supplementary Information and figures
406 S6, S7, S8, S9). All these vesicles are slightly to heavily contaminated by air/seawater or have
407 lost gases (e.g., V4 of sample 041). The fact that they all fall on hyperbolas with low
408 curvatures suggests that they were contaminated in different proportions. This contamination
409 process may occur through small cracks that cannot be observed on X-ray microtomography
410 images given the image resolutions (Table S1).

411 For vesicles that have lower $^{20}\text{Ne}/^{22}\text{Ne}$ ratios (V2 and V3 of sample 003, V2 of sample
412 036, V1 and V4 of sample 058) and also for vesicle V2 of sample 058 with a fractionated He
413 isotopic ratio ($^3\text{He}/^4\text{He}$ of 5.60 ± 0.37 Ra), it is important to note that they contain CO_2 in the
414 expected amount given their volumes. No influence of the ablation depth below the sample
415 polished surface can be observed. All these observations are hard to reconcile. When bubbles
416 are only slightly contaminated, helium should not be easily affected because there is so little
417 He in air and seawater. Only vesicle V2 of sample 058 seems to have a higher $^{36}\text{Ar}/^{22}\text{Ne}$ ratio
418 (figure 5). Previous studies have noted that the complexities of the isotopic and abundance
419 ratio mixing trends require several different processes (e.g., Burnard et al., 2003).

420 The two strongly fractionated vesicles, V5 of sample 055 and V2 of sample 2 π D43, with
421 low CO_2 abundances (table 1), may be connected to bigger cracks than the other vesicles and
422 so CO_2 could have escaped with noble gases. Seawater-derived noble gases could enter the
423 bubbles, followed by atmospheric noble gases when samples are exposed to air after recovery.
424 This may be a relatively slow process because the two vesicles still show influence of
425 seawater-derived noble gases, consistent with contamination via very small cracks.

426 Vesicle V4 of sample 041 seems to be a special case because it has lost CO_2 , He, Ne and
427 Ar but does not have atmospheric isotopic ratios like V5 of sample 055 and V2 of sample
428 2 π D43. This suggests that this piece of sample 041 was degassed.

429 To summarize, the new protocol to collect samples anaerobically appears to reduce
430 air/seawater contamination and is promising for future missions to better characterize mantle
431 composition. Results from this new protocol highlight that contamination starts on the
432 seafloor and that air replaces this seawater component relatively quickly after recovery.
433 However, as shown by the laser ablation data, the contamination processes are complex and
434 not yet fully understood.

435

436 **4.2 The $^4\text{He}/^{40}\text{Ar}^*$ ratio**

437

438 The $^4\text{He}/^{40}\text{Ar}^*$ ratio is a useful tracer of degassing. Indeed, both ^4He and $^{40}\text{Ar}^*$ ($^{40}\text{Ar}^*$ is
439 radiogenic ^{40}Ar , calculated by subtracting the atmospheric contribution assuming ^{36}Ar is fully
440 atmospheric) are radiogenic isotopes produced in the mantle, with an estimated production
441 ratio of 2 to 5 (Moreira and Kurz, 2013). This production ratio is a function of assumed
442 mantle K, Th, and U abundances and the assumption of a closed system evolution, so there is
443 inherent uncertainty, but serves as a useful reference for mantle values and the influence of
444 degassing. Since argon is less soluble in melts than helium, the $^4\text{He}/^{40}\text{Ar}^*$ ratio would increase
445 if a magma was degassed (Moreira and Kurz, 2013; Sarda and Moreira, 2002).

446 As indicated in table 2 and table 3, the $^4\text{He}/^{40}\text{Ar}^*$ ratios are lower than the production ratio.
447 The $^4\text{He}/^{40}\text{Ar}^*$ ratio of 1.37 ± 0.05 for samples 041 and 2 π D43 obtained here is consistent with
448 previous studies of the popping rock 2 π D43, which found values of 1.36 (Javoy and Pineau,
449 1991), 1.5 ± 0.1 (Moreira et al., 1998) and 1.4 ± 0.1 (Raquin et al., 2008). Burnard et al.
450 (1997) determined a lower value of 0.99 but Raquin et al. (2008) suggested that this could be
451 due to different extent of degassing between vesicles, or that Burnard et al. (1997) did not
452 correct argon isotopic ratios for mass discrimination. Staudacher et al. (1989) calculated a
453 higher $^4\text{He}/^{40}\text{Ar}^*$ ratio, between 2.08 and 2.72, for sample 2 π D43.

454 The low $^4\text{He}/^{40}\text{Ar}^*$ ratio of the popping rocks is enigmatic. Lower $^4\text{He}/^{40}\text{Ar}^*$ cannot be
455 explained via degassing because one would expect $^4\text{He}/^{40}\text{Ar}^*$ ratios higher than the
456 production ratio due to lower Ar solubility (Moreira and Kurz, 2013). In case of very low
457 vesicularity, vesicles could potentially have low $^4\text{He}/^{40}\text{Ar}^*$ ratios because most of the gas
458 remains in the melt and argon is less soluble than helium (Sarda and Moreira, 2002), leading
459 to bubbles enriched in argon. But this is not the case for the highly vesicular samples analyzed
460 in this study. It was suggested that popping rocks reflect a process of bubble accumulation
461 (Chavrit et al., 2014; Jones et al., 2019). Such a process should not influence the $^4\text{He}/^{40}\text{Ar}^*$
462 ratio, unless first-generated vesicles, with lower $^4\text{He}/^{40}\text{Ar}^*$ ratios due to the lower solubility of
463 argon in melts, are accumulated. Another explanation would be diffusive loss of helium from
464 the magma before vesiculation since helium diffuses faster than argon.

465 An interesting point is that samples 003, 036, 055 and 058 have systematically (both laser
466 ablation and step-crushing measurements) lower $^4\text{He}/^{40}\text{Ar}^*$ ratios than the samples 041 and
467 $2\pi\text{D}43$. There is no relation between the vesicularities and the $^4\text{He}/^{40}\text{Ar}^*$ ratios (tables [S2](#) and
468 [2](#)). This difference in the $^4\text{He}/^{40}\text{Ar}^*$ ratio could be related to the enrichment in argon observed
469 in samples 003, 036, 055 and 058 compared with samples 041 and $2\pi\text{D}43$ (figure [6](#)): for a
470 same vesicle volume, samples with lower $^4\text{He}/^{40}\text{Ar}^*$ ratio show higher ^{36}Ar content. Such a
471 trend is not observed for helium and neon. This could reflect that the magmatic source of
472 samples 003, 036, 055 and 058 is enriched in argon compared with the source of samples 041
473 and $2\pi\text{D}43$.

474

475

476

477

478

4.3 Upper mantle noble gas isotopic and elemental compositions

One goal of this study is to better characterize upper mantle noble gas compositions. The difference in isotopic and elemental compositions between samples 003, 036, 055 and 058 and samples 041 and 2 π D43 is significant (figures 2, 3, 4 and 5).

One hypothesis is that samples 003, 036, 055 and 058 represent assimilation of altered oceanic crust or of sedimentary material, leading to lower $^{40}\text{Ar}/^{36}\text{Ar}$, $^{20}\text{Ne}/^{22}\text{Ne}$ and $^{22}\text{Ne}/^{36}\text{Ar}$ ratios. The Cl/K and H₂O/Ce ratios are useful tracers of assimilation of hydrothermally altered material and seawater-derived components (Dixon et al., 2002; Michael and Cornell, 1998; Stroncik and Niedermann, 2016). However, all the samples have Cl/K ratios in the range 0.050-0.054 and H₂O/Ce ratios between 160 and 174 (table 3, Jones et al., 2019), which is in the range of 0.05-0.08 for Cl/K suggested for uncontaminated E-MORB (Michael and Cornell, 1998) and is typical of MORB for H₂O/Ce (Dixon et al., 2002). Therefore, assimilation of seawater-altered oceanic crust during the transport to the surface seems not to be reflected in the Cl/K and H₂O/Ce ratios; even though, noble gas ratios (e.g., $^{40}\text{Ar}/^{36}\text{Ar}$, $^{22}\text{Ne}/^{36}\text{Ar}$) can be more sensitive to these processes than Cl/K as shown by Stroncik and Niedermann (2016).

The popping rock samples 003, 036, 055 and 058 exhibit very similar noble gas compositions, whereas one could expect that they would have been contaminated to various degrees, if they were influenced by crustal contamination. As discussed in section 4.2, samples are unlikely to be degassed given their low $^4\text{He}/^{40}\text{Ar}^*$ ratios. However, the magmatic source of samples 003, 036, 055 and 058 seems to be enriched in argon (figure 6 and section 4.2) compared with the magma of samples 041 and 2 π D43.

Preliminary lead, strontium and neodymium isotopic compositions for samples from the popping rock area show measurable differences (table S5), providing independent evidence

504 that these samples originate from two distinct mantle sources. The $^{206}\text{Pb}/^{204}\text{Pb}$ ratio is 18.962
505 for sample 041 and 19.026 ± 0.002 for samples 003, 055 and 058, which is well beyond
506 measurement uncertainty (section S5 in Supplementary Information). Small but measurable
507 differences are also observed in Sr and Nd isotopes. The mantle source with low $^{20}\text{Ne}/^{22}\text{Ne}$,
508 $^{40}\text{Ar}/^{36}\text{Ar}$ and $^{22}\text{Ne}/^{36}\text{Ar}$ ratios (samples 003, 036, 055, 058) has more radiogenic lead than
509 sample 041. Sample 041 was collected from the easternmost location, potentially indicating a
510 geographic boundary. Further measurements are ongoing on a larger set of samples from the
511 area and will be published elsewhere.

512 The popping rock area (figure 1) is part of the strong geochemical (E-MORB) anomaly
513 identified at 14° N on the Mid-Atlantic Ridge, devoid of plume influence (Bougault et al.,
514 1988; Dosso et al., 1991; Hémond et al., 2006). Here, the preliminary Pb, Nd and Sr data
515 show that the upper mantle is heterogeneous at a much smaller scale. All the samples
516 analyzed in this study from the 2016 cruise (figure 1) have indistinguishable major and trace
517 element compositions but distinct Pb, Sr and Nd isotopic compositions, along with different
518 elemental and isotopic noble gas compositions (in particular distinct $^{20}\text{Ne}/^{22}\text{Ne}$, $^{40}\text{Ar}/^{36}\text{Ar}$,
519 $^{22}\text{Ne}/^{36}\text{Ar}$ and $^4\text{He}/^{40}\text{Ar}^*$). Several studies suggested that seawater noble gases (Ne, Ar, Kr,
520 Xe) could be recycled into the mantle through incorporation into hydrated minerals of the
521 subducted oceanic lithosphere despite their inertness (e.g., Holland and Ballentine, 2006;
522 Kendrick et al., 2018; Smye et al., 2017). Therefore, the mantle source of samples 003, 036,
523 055 and 058 could have been more affected by recycling of oceanic crust and so by noble gas
524 recycling, leading to a more radiogenic lead isotopic composition, argon enrichment and
525 lower $^{20}\text{Ne}/^{22}\text{Ne}$, $^{40}\text{Ar}/^{36}\text{Ar}$ and $^{22}\text{Ne}/^{36}\text{Ar}$.

526 Several studies have shown that the upper mantle is chemically heterogeneous on different
527 scales (e.g., Agranier et al., 2005; Allègre and Turcotte, 1986; Bougault et al., 1988; Chavrit
528 et al., 2014; Dosso et al., 1991; Graham et al., 1992; Hémond et al., 2006; Parai et al., 2012;

529 Tucker et al., 2012; Wilson et al., 2013). Recycling of atmospheric noble gases was invoked
530 to account for some of these heterogeneities (e.g., Parai et al., 2012; Sarda et al., 1999; Tucker
531 et al., 2012). Based on lead-argon isotopic correlations observed for MORBs from the
532 Atlantic ocean, Sarda et al. (1999) suggested that atmospheric noble gases are recycled into
533 the mantle and that argon would be a tracer of subducted material which shows $^{40}\text{Ar}/^{36}\text{Ar}$
534 close to air. Tucker et al. (2012) discussed heterogeneities observed in the noble gas
535 compositions of MORBs from the equatorial Mid-Atlantic Ridge, in particular large $^{40}\text{Ar}/^{36}\text{Ar}$
536 variations despite constant helium and neon isotopic compositions, which could also be partly
537 related to atmospheric noble gas recycling.

538 Based on these new data as well as on recent studies and compilations (Moreira, 2013;
539 Parai et al., 2012; Tucker et al., 2012), it seems that the highest (measurable) noble gas
540 isotopic ratios for the MORB mantle devoid of plume influence are 12.5 for $^{20}\text{Ne}/^{22}\text{Ne}$ and
541 30,000 for $^{40}\text{Ar}/^{36}\text{Ar}$. Lower maximum values (e.g., excluding shallow atmospheric
542 contamination) for these ratios could be related to interaction with a HIMU-plume component
543 (e.g., Tucker et al., 2012) and/or to different degrees of incorporation of recycled material in
544 mantle sources in the context of a marble-cake mantle (Agranier et al., 2005; Allègre and
545 Turcotte, 1986). Therefore, the noble gas isotopic composition of popping rock 2πD43 can be
546 considered as a reference for the upper mantle, even if this sample is an E-MORB (Bougault
547 et al., 1988; Dosso et al., 1991). A better understanding of the popping rock $^4\text{He}/^{40}\text{Ar}^*$ ratios
548 (section 4.2) is needed to constrain the upper mantle elemental noble gas ratios.

549 One explanation for the new observations is that recycled oceanic lithosphere into the
550 mantle can introduce small-scale heterogeneities, in particular for the noble gas elemental and
551 isotopic composition. A detailed study of the lithophile element and noble gas isotopic
552 compositions on a larger set of samples from the popping rock area will further constrain the
553 extent and scale of these heterogeneities.

554 **V) Conclusion**

555

556 Determining the volatile isotopic and elemental compositions of the mantle is crucial to
557 better understand volatile cycling on Earth. Undegassed MORB, such as popping rocks,
558 represent important samples to address these issues. We analyzed the noble gas compositions
559 of five samples from the cruise AT33-03 on the Mid-Atlantic Ridge, four of which were
560 popping rocks, with step-crushing and laser ablation, as well as the popping rock 2πD43 that
561 was recovered in the same area in 1985.

562 A new protocol was tested to recover lavas with anaerobic samplers. It seems very
563 efficient to reduce sample contamination, but the presence of a seawater- or air-derived
564 component is still observed, suggesting that samples start being contaminated after eruption
565 on the seafloor through cracks and opened vesicles and that air replaces this seawater
566 component after recovery. This highlights that contamination processes are complex and that
567 this contaminant is hard to remove (Ballentine and Barfod, 2000; Roubinet and Moreira,
568 2018).

569 Four samples (003, 036, 055 and 058) show distinct neon and argon isotopic and
570 elemental compositions ($^{20}\text{Ne}/^{22}\text{Ne}$ of about 12.2 and $^{40}\text{Ar}/^{36}\text{Ar}$ of about 16,100) compared
571 with sample 2πD43, whereas sample 041, the easternmost sample studied, has a composition
572 similar to the popping rock 2πD43 ($^{20}\text{Ne}/^{22}\text{Ne}$ of 12.5 and $^{40}\text{Ar}/^{36}\text{Ar}$ of 27,000). The two
573 groups of samples also have distinct $^4\text{He}/^{40}\text{Ar}^*$ ratios, which is consistent with the source of
574 samples 003, 036, 055 and 058 being enriched in argon compared with the source of samples
575 041 and 2πD43 based on laser ablation data.

576 Preliminary lead, strontium and neodymium isotopic compositions indicate that these two
577 groups of samples were produced by separate eruptive events from two distinct mantle
578 sources. This provides evidence for upper mantle heterogeneity at small scale. The mantle

579 source of the samples with lower neon and argon isotopic ratios could have been affected by
580 recycled oceanic crust and by recycled atmospheric noble gases, in a context of a marble-cake
581 mantle (Allègre and Turcotte, 1986).

582

583 **Acknowledgements**

584

585 We thank Sujoy Mukhopadhyay and an anonymous reviewer for their insightful comments
586 that helped improving this contribution as well as Tamsin Mather for editorial handling and
587 constructive comments. Javier Escartín is greatly thanked for assistance with cruise planning
588 and for his constructive comments on the first versions of this paper. We thank Dan Fornari
589 and the rest of the science team for their excellent work at sea. The authors thank the captain
590 and the crew of the *R/V Atlantis*, the AUV *Sentry* team, and the HOV *Alvin* team, for their
591 invaluable help during the cruise. The authors are also grateful to Déborah Chavrit for fruitful
592 discussions. The cruise and the WHOI noble gas measurements were supported by NSF OCE-
593 1259218. S.P. acknowledges the financial support from the L'Oréal-UNESCO *For Women in*
594 *Science* 2017 fellowship.

595

596 **References**

597

- 598 Agranier, A., Blichert-Toft, J., Graham, D.W., Debaille, V., Schiano, P., Albarède, F., 2005.
599 The spectra of isotopic heterogeneities along the mid-Atlantic Ridge. *Earth and*
600 *Planetary Science Letters* 238, 96-109.
- 601 Allègre, C.J., Turcotte, D., 1986. Implications of a two-component marble-cake mantle.
602 *Nature* 323, 123-127.
- 603 Ballentine, C., Barfod, D., 2000. The origin of air-like noble gases in MORB and OIB. *Earth*
604 *and Planetary Science Letters* 180, 39-48.
- 605 Bougault, H., Dmitriev, L., Schilling, J.-G., Sobolev, A., Joron, J.-L., Needham, H.D., 1988.
606 Mantle heterogeneity from trace elements: MAR triple junction near 14°N. *Earth and*
607 *Planetary Science Letters* 88, 27-36.
- 608 Burnard, P., Graham, D., Turner, G., 1997. Vesicle-specific noble gas analyses of "popping
609 rock": implications for primordial noble gases in the Earth. *Science* 276, 568-571.
- 610 Burnard, P., Harrison, D., Turner, G., Nesbitt, R., 2003. Degassing and contamination of
611 noble gases in Mid-Atlantic Ridge basalts. *Geochemistry Geophysics Geosystems* 4.

612 Chavrit, D., Humler, E., Grasset, O., 2014. Mapping modern CO₂ fluxes and mantle carbon
613 content all along the mid-ocean ridge system. *Earth and Planetary Science Letters* 387,
614 229-239.

615 Clarke, W.B., Jenkins, W.J., Top, Z., 1976. Determination of tritium by mass spectrometric
616 measurement of ³He. *The International Journal of Applied Radiation and Isotopes* 27,
617 512-522.

618 Dixon, J.E., Leist, L., Langmuir, C., Schilling, J.-G., 2002. Recycled dehydrated lithosphere
619 observed in plume-influenced mid-ocean-ridge basalt. *Nature* 420, 385.

620 Dosso, L., Hanan, B.B., Bougault, H., Schilling, J.-G., Joron, J.-L., 1991. Sr-Nd-Pb
621 geochemical morphology between 10° and 17°N on the Mid-Atlantic Ridge: A new MORB
622 isotope signature. *Earth and Planetary Science Letters* 106, 29-43.

623 Graham, D.W., Jenkins, W.J., Schilling, J.G., Thompson, G., Kurz, M.D., Humphris, S., 1992.
624 Helium isotope geochemistry of mid-ocean ridge basalts from the South Atlantic. *Earth
625 and Planetary Science Letters* 110, 133-147.

626 Hekinian, R., Chaigneau, M., Cheminee, J.L., 1973. Popping Rocks and Lava Tubes from
627 the Mid-Atlantic Rift Valley at 36° N. *Nature* 245, 371-373.

628 Hémond, C., Hofmann, A.W., Vlastélic, I., Nauret, F., 2006. Origin of MORB enrichment
629 and relative trace element compatibilities along the Mid-Atlantic Ridge between 10° and
630 24°N. *Geochemistry, Geophysics, Geosystems* 7.

631 Holland, G., Ballentine, C.J., 2006. Seawater subduction controls the heavy noble gas
632 composition of the mantle. *Nature* 441, 186-191.

633 Javoy, M., Pineau, F., 1991. The volatiles record of a « popping » rock from the Mid-
634 Atlantic Ridge at 14°N : chemical and isotopic composition of gas trapped in the vesicles.
635 *Earth and Planetary Science Letters* 107, 598-611.

636 Jones, M.R., Wanless, V.D., Soule, S.A., Kurz, M.D., Mittelstaedt, E., Fornari, D.J., Curtice, J.,
637 Klein, F., Le Roux, V., Brodsky, H., Péron, S., Schwartz, D.M., 2019. New constraints on
638 mantle carbon from Mid-Atlantic Ridge popping rocks. *Earth and Planetary Science
639 Letters* 511, 67-75.

640 Kendrick, M.A., Scambelluri, M., Hermann, J., Padrón-Navarta, J.A., 2018. Halogens and
641 noble gases in serpentinites and secondary peridotites: Implications for seawater
642 subduction and the origin of mantle neon. *Geochimica et Cosmochimica Acta* 235, 285-
643 304.

644 Kunz, J., 1999. Is there solar argon in the Earth's mantle? *Nature* 399, 649-650.

645 Kunz, J., Staudacher, T., Allègre, C.J., 1998. Plutonium-Fission Xenon Found in Earth's
646 Mantle. *Science* 280, 877-880.

647 Kurz, M.D., Curtice, J., Fornari, D., Geist, D., Moreira, M., 2009. Primitive neon from the
648 center of the Galápagos hotspot. *Earth and Planetary Science Letters* 286, 23-34.

649 Kurz, M.D., Moreira, M., Curtice, J., Lott III, D.E., Mahoney, J.J., Sinton, J.M., 2005.
650 Correlated helium, neon, and melt production on the super-fast spreading East Pacific
651 Rise near 17°S. *Earth and Planetary Science Letters* 232, 125-142.

652 Marty, B., Tolstikhin, I.N., 1998. CO₂ fluxes from mid-ocean ridges, arcs and plumes.
653 *Chemical Geology* 145, 233-248.

654 Michael, P.J., Cornell, W.C., 1998. Influence of spreading rate and magma supply on
655 crystallization and assimilation beneath mid-ocean ridges: Evidence from chlorine and
656 major element chemistry of mid-ocean ridge basalts. *Journal of Geophysical Research:
657 Solid Earth* 103, 18325-18356.

658 Moore, J.G., Batchelder, J.N., Cunningham, C.G., 1977. CO₂-filled vesicles in mid-ocean
659 basalt. *Journal of Volcanology and Geothermal Research* 2, 309-327.

660 Moreira, M., 2013. Noble gas constraints on the origin and evolution of Earth's volatiles.
661 *Geochemical Perspectives* 2, 229-403.

662 Moreira, M., Kunz, J., Allègre, C.J., 1998. Rare gas systematics on popping rock : estimates
663 of isotopic and elemental compositions in the upper mantle. *Science* 279, 1178-1181.

664 Moreira, M.A., Kurz, M.D., 2013. Noble Gases as Tracers of Mantle Processes and
665 Magmatic Degassing, in: Burnard, P. (Ed.), *The Noble Gases as Geochemical Tracers*.
666 Springer Berlin Heidelberg, pp. 371-391.

667 Mukhopadhyay, S., 2012. Early differentiation and volatile accretion recorded in deep
668 mantle Neon and Xenon. *Nature* 486, 101-104.

669 Ozima, M., Podosek, F.A., 2002. *Noble Gas Geochemistry*. Cambridge University Press.

670 Parai, R., Mukhopadhyay, S., Standish, J.J., 2012. Heterogeneous upper mantle Ne, Ar and
671 Xe isotopic compositions and a possible Dupal noble gas signature recorded in basalts
672 from the Southwest Indian Ridge. *Earth and Planetary Science Letters* 359-360, 227-
673 239.

674 Parnell-Turner, R.E., Mittelstaedt, E., Kurz, M.D., Jones, M.R., Soule, S.A., Klein, F.,
675 Wanless, V.D., Fornari, D.J., 2018. The Final Stages of Slip and Volcanism on an Oceanic
676 Detachment Fault at 13° 48' N, Mid-Atlantic Ridge. *Geochemistry, Geophysics,*
677 *Geosystems* 0.

678 Péron, S., Moreira, M., Colin, A., Arbaret, L., Putlitz, B., Kurz, M.D., 2016. Neon isotopic
679 composition of the mantle constrained by single vesicle analyses. *Earth and Planetary*
680 *Science Letters* 449, 145-154.

681 Péron, S., Moreira, M., Putlitz, B., Kurz, M.D., 2017. Solar wind implantation supplied light
682 volatiles during the first stage of Earth accretion. *Geochemical Perspectives Letters* 3,
683 151-159.

684 Pineau, F., Javoy, M., Bottinga, Y., 1976. ¹³C/¹²C ratios of rocks and inclusions in popping
685 rocks of the mid Atlantic ridge and their bearing on the problem of isotopic composition
686 of deep-seated carbon. *Earth and Planetary Science Letters* 29, 413-421.

687 Raquin, A., Moreira, M., Guillon, F., 2008. He, Ne and Ar systematics in single vesicles:
688 Mantle isotopic ratios and origin of the air component in basaltic glasses. *Earth and*
689 *Planetary Science Letters* 274, 142-150.

690 Roubinet, C., Moreira, M.A., 2018. Atmospheric noble gases in Mid-Ocean Ridge Basalts:
691 Identification of atmospheric contamination processes. *Geochimica et Cosmochimica*
692 *Acta* 222, 253-268.

693 Sarda, P., Graham, D.W., 1990. Mid-ocean ridge popping rocks: implications for
694 degassing at ridge crests. *Earth and Planetary Science Letters* 97, 268-289.

695 Sarda, P., Moreira, M., 2002. Vesiculation and vesicle loss in Mid Oceanic Ridge basalt
696 glasses: He, Ne, Ar elemental fractionation and pressure influence. *Geochimica et*
697 *Cosmochimica Acta* 66, 1449-1458.

698 Sarda, P., Moreira, M., Staudacher, T., 1999. Argon-lead isotopic correlation in Mid-
699 Atlantic Ridge Basalts. *Science* 283, 666-668.

700 Smith, D.K., Escartín, J., Schouten, H., Cann, J.R., 2008. Fault rotation and core complex
701 formation: Significant processes in seafloor formation at slow-spreading mid-ocean
702 ridges (Mid-Atlantic Ridge, 13°-15°N). *Geochemistry, Geophysics, Geosystems* 9,
703 Q03003.

704 Smye, A.J., Jackson, C.R.M., Konrad-Schmolke, M., Hesse, M.A., Parman, S.W., Shuster, D.L.,
705 Ballentine, C.J., 2017. Noble gases recycled into the mantle through cold subduction
706 zones. *Earth and Planetary Science Letters* 471, 65-73.

707 Staudacher, T., Sarda, P., Richardson, S.H., Allègre, C.J., Sagna, I., Dmitriev, L.V., 1989.
708 Noble gases in basalt glasses from a Mid-Atlantic ridge topographic high at 14°N:
709 geodynamic consequences. *Earth and Planetary Science Letters* 96, 119-133.
710 Stroncik, N.A., Niedermann, S., 2016. Atmospheric contamination of the primary Ne and
711 Ar signal in mid-ocean ridge basalts and its implications for ocean crust formation.
712 *Geochimica et Cosmochimica Acta* 172, 306-321.
713 Trieloff, M., Kunz, J., Clague, D.A., Harrison, D., Allègre, C.J., 2000. The Nature of pristine
714 noble gases in mantle plumes. *Science* 288, 1036-1038.
715 Tucker, J.M., Mukhopadhyay, S., Gonnermann, H.M., 2018. Reconstructing mantle carbon
716 and noble gas contents from degassed mid-ocean ridge basalts. *Earth and Planetary
717 Science Letters* 496, 108-119.
718 Tucker, J.M., Mukhopadhyay, S., Schilling, J.-G., 2012. The heavy noble gas composition of
719 the depleted MORB mantle (DMM) and its implications for the preservation of
720 heterogeneities in the mantle. *Earth and Planetary Science Letters* 355-356, 244-254.
721 Wilson, S.C., Murton, B.J., Taylor, R.N., 2013. Mantle composition controls the
722 development of an Oceanic Core Complex. *Geochemistry, Geophysics, Geosystems* 14,
723 979-995.
724
725
726
727

728 **Table 1.** Laser ablation CO₂ and noble gas abundances (in cm³ STP) for individual pierced
729 vesicles. The ablation duration is indicated in brackets and the symbol ≠ denotes anomalous
730 vesicles (section 3.1). The piece number (#) is detailed in Table S1 and the depth indicates the
731 distance between the sample surface and the vesicle. Vesicle volumes were calculated using
732 the 3D Object Counter plugin of ImageJ. Errors are 1σ uncertainties, n.d. stands for not
733 determined. Longitude, latitude and depth of sample recovery are indicated in brackets (figure
734 1).
735

Sample	#	Depth μm	Volume x10 ⁶ μm ³	Pressure Torr	CO ₂ x10 ⁻³ cm ³	s	⁴ He x10 ⁻⁷ cm ³	s	²² Ne x10 ⁻¹³ cm ³	s	³⁶ Ar x10 ⁻¹² cm ³	s
Sample AL4818-003 (-45.011, 13.774, -3672 m)												
V1 (2'50")	1	63	68.6	0.2270	8.21	0.77	2.24	0.10	4.56	0.26	12.5	0.1
V2 (50")	1	102	276.6	0.9990	36.1	3.4	9.36	0.40	23.9	0.57	53.3	0.3
V3 (35")	1	70	96.9	0.3235	11.7	1.1	3.10	0.13	7.36	0.29	17.3	0.1
V4 (1'47")	1	63	105.4	0.3226	11.7	1.1	2.94	0.13	6.33	0.28	16.7	0.1
V5 (40")	1	78	77.4	0.2292	8.29	0.78	2.41	0.10	6.40	0.28	13.7	0.1
V6 (2'10")	1	157	109.9	0.3103	11.2	1.1	2.97	0.02	6.95	0.46	16.4	0.2
Sample AL4820-036 (-45.010, 13.767, -3718 m)												
V1 (23")	1	80	84.0	0.2140	7.74	0.73	1.96	0.02	5.31	0.44	11.0	0.1
V2 (23")	1	46	78.6	0.2809	10.2	1.0	2.56	0.03	6.82	0.52	14.4	0.2
V3 (1'05")	1	86	267.4	0.9489	34.3	3.2	8.40	0.09	22.0	1.5	48.0	0.6
V4 (27")	1	17	47.3	0.1733	6.27	0.59	1.52	0.05	3.48	0.28	8.47	0.21
V5 (2'10")	1	108	75.3	0.2550	9.23	0.87	2.29	0.08	5.24	0.30	13.2	0.3
V6 (2'35")	1	80	48.8	0.1501	5.43	0.51	1.34	0.05	2.87	0.27	7.56	0.19
V7 (35")	1	23	36.5	0.1361	4.92	0.46	1.17	0.04	2.49	0.26	6.82	0.17
V8 (5'30")≠	1	8.6	6.1	0.0187	0.68	0.07	0.11	0.01	n.d.	n.d.	1.13	0.05
V9 (1'35")	1	112	2896.8	2.6227	94.9	8.9	21.8	0.07	54.9	1.8	127	3
Sample AL4820-041 (-45.002, 13.770, -3645 m)												
V1 (1'18")	2	60	422.3	1.0920	39.5	3.7	10.5	0.1	26.1	1.1	31.5	0.2
V2 (4'39")	2	17	12.6	0.0437	1.58	0.15	0.47	0.01	0.98	0.25	1.22	0.04
V3 (5'12")	1	87	133.8	0.4420	16.0	1.5	4.68	0.03	10.6	0.51	12.5	0.1
V4 (4'47")≠	3	127	539.2	0.0284	1.03	0.10	0.005	0.008	0.38	0.24	0.96	0.04
Sample AL4820-055 (-45.010, 13.773, -3665 m)												
V1 (5'45")	1	94	42.0	0.1375	4.97	0.47	1.45	0.01	3.78	0.33	7.85	0.09
V2 (1'25")≠	1	37	316.1	1.0827	39.2	3.7	8.43	0.04	24.6	1.4	56.5	0.6
V3 (2'16")	1	102	361.2	1.2069	43.7	4.1	11.3	0.3	26.6	1.4	62.4	0.6
V4 (4'57")	1	78	259.7	0.9467	34.3	3.2	9.47	0.28	21.6	1.2	47.8	0.5
V5 (2'48")≠	1	31	47.8	0.0253	0.92	0.09	0.04	0.01	10.2	0.5	44.7	0.4
Sample AL4820-058 (-45.009, 13.774, -3684 m)												
V1 (2'15")	1	22	30.5	0.0932	3.37	0.32	0.92	0.03	2.22	0.26	5.10	0.57
V2 (37")≠	1	17	89.8	0.2932	10.6	1.0	0.05	0.01	5.72	0.33	17.0	0.2
V3 (2'08")	1	29	95.5	0.3102	11.2	1.1	3.01	0.09	6.80	0.35	16.2	0.1
V4 (2'43")	1	28	43.1	0.1613	5.84	0.55	1.56	0.05	3.60	0.28	8.45	0.08
V5 (4'20")	1	29	28.9	0.0812	2.94	0.28	0.78	0.03	1.62	0.25	4.28	0.05
Sample 2πD43												
V1 (4'42")	1	24	525.9	1.6844	60.9	5.7	15.2	0.56	40.1	2.0	46.0	1.3
V2 (3")≠	1	61	436.2	0.0130	0.47	0.05	0.01	0.01	5,930	397	13,221	167
V3 (30")	1	66	58.3	0.2355	8.52	0.80	2.24	0.03	5.28	0.44	6.18	0.09
V4 (1'01")	1	72	74.4	0.2327	8.42	0.79	2.50	0.03	5.02	0.42	7.03	0.10

736

737

738

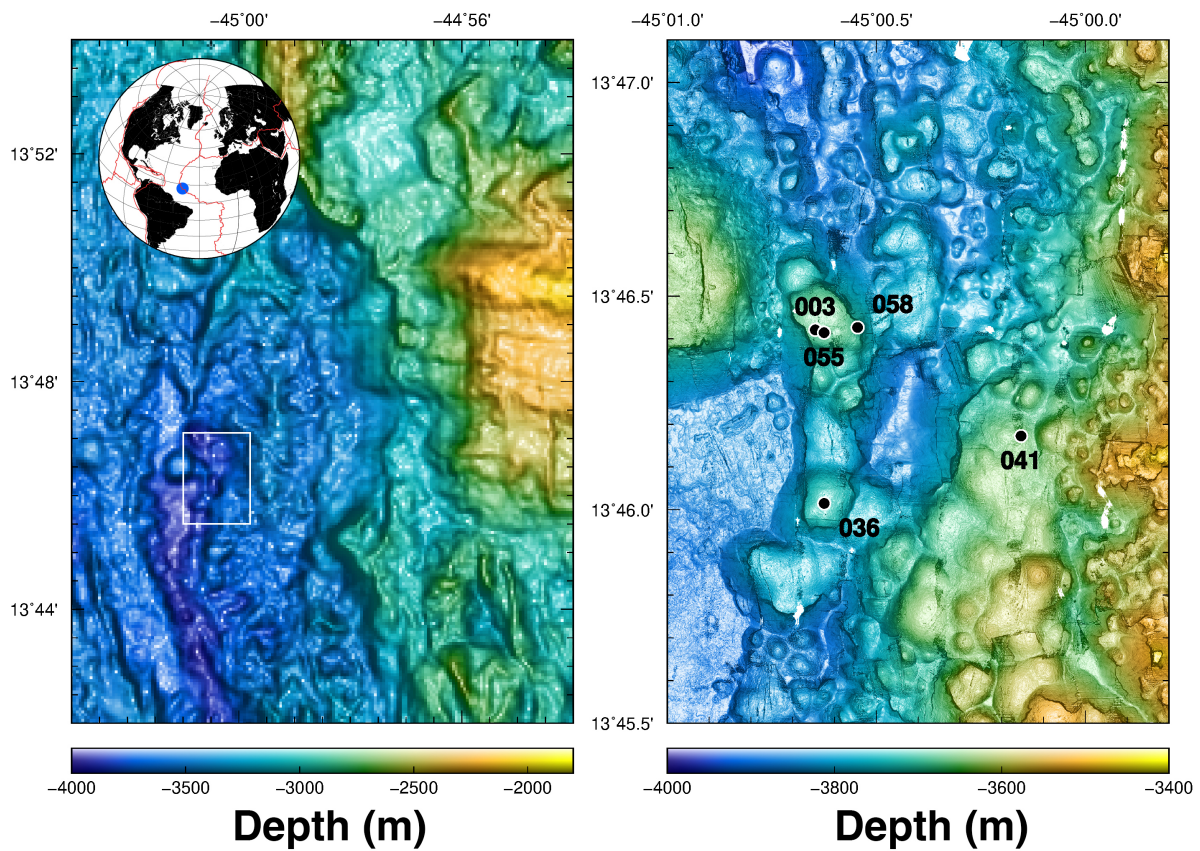
739 **Table 2.** Vesicle elemental and isotopic ratios obtained by laser ablation. The symbol ≠
740 denotes anomalous vesicles (section 3.1). $^3\text{He}/^4\text{He}$ ratio is presented relative to atmosphere
741 (Ra), where 1 Ra is equal to 1.384×10^{-6} (Clarke et al., 1976). $^{40}\text{Ar}^*$ is radiogenic ^{40}Ar . We use
742 the value of 295.5 for the atmospheric $^{40}\text{Ar}/^{36}\text{Ar}$ ratio. Errors are 1σ uncertainties, n.d. stands
743 for not determined.
744

Sample	CO_2/H_2	s	$^3\text{He}/^4\text{He}$	s	$^{20}\text{Ne}/^{22}\text{Ne}$	s	$^{21}\text{Ne}/^{22}\text{Ne}$	s	$^{38}\text{Ar}/^{36}\text{Ar}$	s	$^{40}\text{Ar}/^{36}\text{Ar}$	s	$^4\text{He}/^{40}\text{Ar}^*$	s
	$\times 10^9$		(xRa)											
Sample AL4818-003														
V1	3.19	0.34	8.29	0.21	12.25	0.15	0.0621	0.0020	0.1879	0.0006	16040	247	1.13	0.05
V2	3.36	0.36	8.30	0.21	11.85	0.05	0.0539	0.0011	0.1893	0.0008	15275	80	1.17	0.05
V3	3.26	0.35	8.36	0.21	11.81	0.08	0.0541	0.0016	0.1861	0.0005	16052	171	1.14	0.05
V4	3.38	0.36	8.48	0.22	12.19	0.09	0.0566	0.0018	0.1862	0.0005	16342	179	1.10	0.05
V5	2.98	0.32	8.35	0.21	12.04	0.08	0.0535	0.0024	0.1876	0.0006	16303	210	1.10	0.05
V6	3.36	0.32	8.11	0.11	12.05	0.09	0.0566	0.0013	0.1882	0.0006	16256	199	1.14	0.02
Sample AL4820-036														
V1	3.46	0.34	8.28	0.20	12.08	0.14	0.0565	0.0017	0.1892	0.0010	15707	439	1.15	0.04
V2	3.53	0.35	8.13	0.20	11.96	0.10	0.0555	0.0013	0.1900	0.0009	16251	349	1.11	0.03
V3	3.58	0.35	8.24	0.21	12.10	0.06	0.0572	0.0010	0.1906	0.0009	15851	147	1.12	0.02
V4	3.51	0.38	8.50	0.37	12.50	0.16	0.0572	0.0022	0.1869	0.0015	16267	545	1.12	0.06
V5	3.44	0.38	8.48	0.37	12.44	0.12	0.0602	0.0020	0.1903	0.0014	16229	362	1.08	0.05
V6	3.45	0.38	8.51	0.38	12.76	0.25	0.0586	0.0030	0.1867	0.0014	16109	601	1.12	0.06
V7	3.51	0.38	8.69	0.38	12.28	0.24	0.0606	0.0033	0.1891	0.0015	15940	638	1.09	0.06
V8 [≠]	4.64	0.64	10.00	0.48	n.d.	n.d.	n.d.	n.d.	0.1910	0.0041	13327	3132	0.71	0.18
V9	3.64	0.40	8.63	0.38	12.14	0.04	0.0583	0.0009	0.1884	0.0014	16145	133	1.08	0.05
Sample AL4820-041														
V1	3.26	0.31	8.36	0.10	12.31	0.07	0.0573	0.0010	0.1880	0.0008	25173	336	1.34	0.02
V2	2.89	0.28	8.48	0.16	13.38	0.43	0.0632	0.0050	0.1786	0.0035	26470	4928	1.45	0.28
V3	2.88	0.27	8.58	0.11	12.44	0.07	0.0591	0.0016	0.1915	0.0006	27073	509	1.40	0.03
V4 [≠]	4932	21599	0.33	1.32	11.70	0.51	0.0518	0.0074	0.2006	0.0040	26089	6300	0.02	0.03
Sample AL4821-055														
V1	3.05	0.29	8.17	0.12	12.13	0.17	0.0574	0.0023	0.1886	0.0007	16149	373	1.16	0.03
V2 [≠]	3.77	0.36	8.91	0.12	12.22	0.06	0.0585	0.0008	0.1885	0.0007	16311	88	0.93	0.01
V3	3.33	0.34	8.38	0.21	12.07	0.06	0.0578	0.0007	0.1882	0.0007	16089	84	1.15	0.04
V4	3.16	0.32	8.28	0.20	12.15	0.06	0.0573	0.0010	0.1884	0.0009	15915	90	1.27	0.04
V5 [≠]	18.0	4.4	9.33	0.56	9.71	0.08	0.0296	0.0011	0.1884	0.0004	2418	14	0.04	0.01
Sample AL4821-058														
V1	3.28	0.34	8.11	0.21	11.85	0.16	0.0471	0.0027	0.1872	0.0010	15721	551	1.16	0.06
V2 [≠]	274	56	5.60	0.37	11.83	0.11	0.0547	0.0015	0.1886	0.0006	16146	179	0.02	0.00
V3	3.31	0.34	8.14	0.20	12.02	0.11	0.0576	0.0010	0.1894	0.0006	16261	189	1.17	0.04
V4	3.26	0.33	8.25	0.22	11.86	0.15	0.0568	0.0018	0.1893	0.0007	16220	351	1.16	0.05
V5	3.31	0.36	8.25	0.27	12.65	0.34	0.0565	0.0036	0.1881	0.0011	16446	757	1.13	0.07
Sample 2πD43														
V1	3.36	0.35	8.61	0.17	12.47	0.10	0.0596	0.0019	0.1916	0.0011	25866	330	1.29	0.06
V2 [≠]	287	241	0.98	0.49	9.67	0.04	0.0289	0.0004	0.1882	0.0008	299	2	0.03	0.02
V3	3.23	0.32	8.51	0.21	12.53	0.16	0.0586	0.0016	0.1880	0.0011	26809	1443	1.37	0.08
V4*	2.92	0.29	8.32	0.21	(13.12*)	0.16	(0.0661*)	0.0024	0.1889	0.0012	26191	1243	1.38	0.07

745 *CO₂ interference is overcorrected. This bubble was targeted just after vesicle V3 of
746 sample 2πD43 and strong CO₂ memory effects were observed when analyzing neon from
747 vesicle V4.
748

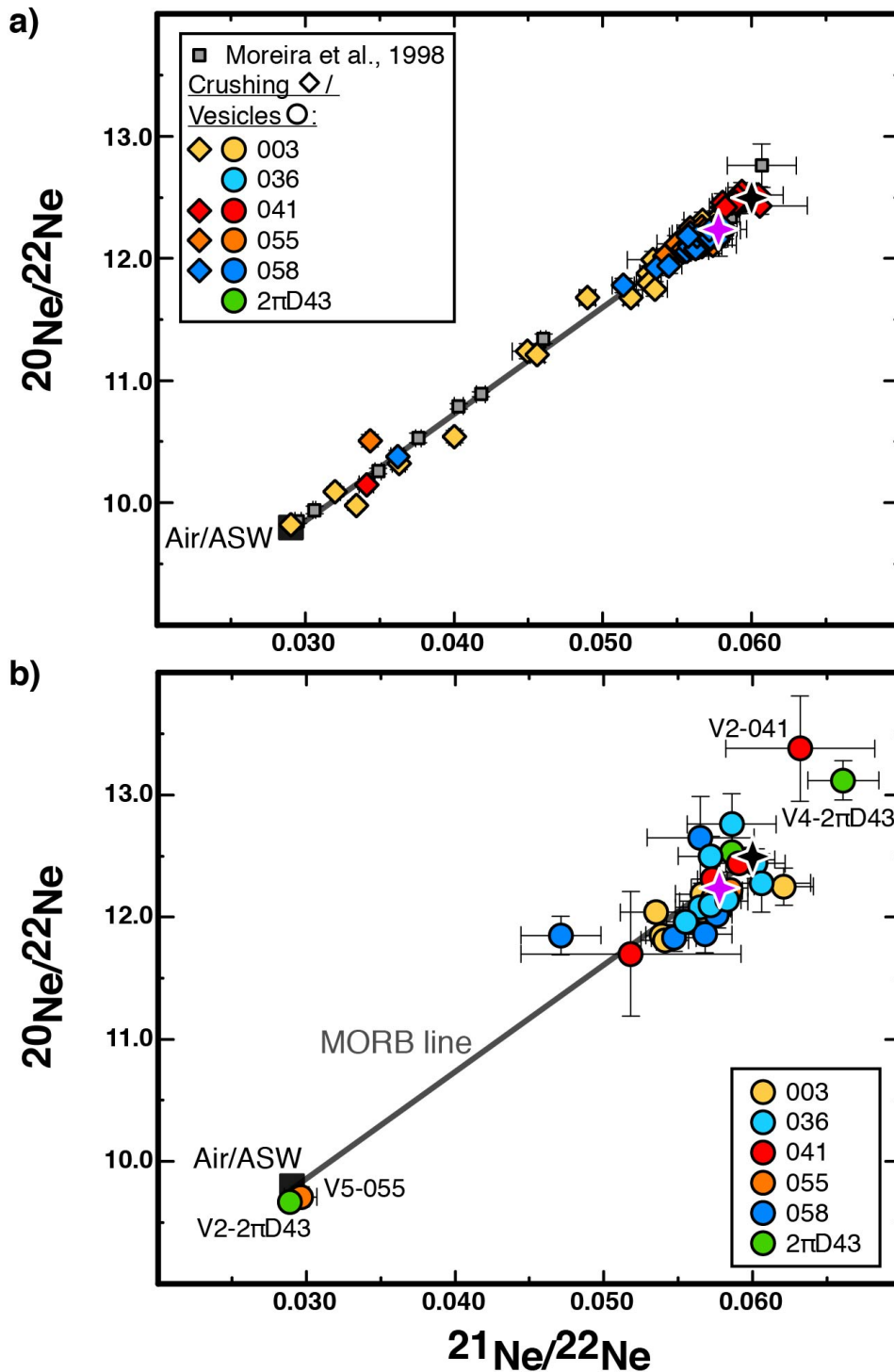
749 **Table 3.** Noble gas abundances and isotopic compositions obtained by crushing in vacuum.
 750 $^3\text{He}/^4\text{He}$ ratio is presented relative to atmosphere (Ra), where 1 Ra is equal to 1.384×10^{-6}
 751 (Clarke et al., 1976). $^{40}\text{Ar}^*$ is radiogenic ^{40}Ar . We use the value of 295.5 for the atmospheric
 752 $^{40}\text{Ar}/^{36}\text{Ar}$ ratio. Errors are 1σ uncertainties, n.d. stands for not determined. The weights of
 753 samples 055 and 058 are less accurate because they were determined inside the glove box.
 754 The Cl/K and $\text{H}_2\text{O}/\text{Ce}$ ratios are from Jones et al. (2019).
 755

756 **Figure 1.** Bathymetric map showing the location of the samples from cruise AT33-03 of *R/V*
 757 *Atlantis* on the Mid-Atlantic Ridge analyzed in this contribution. The extinct oceanic core
 758 complex (Smith et al., 2008) is visible in the east of the left image. The popping rock $2\pi\text{D}43$
 759 was recovered by dredging close to $13^\circ 46' \text{N}$ (Bougault et al., 1988). The image on the right is
 760 a zoom of the white rectangle.
 761
 762



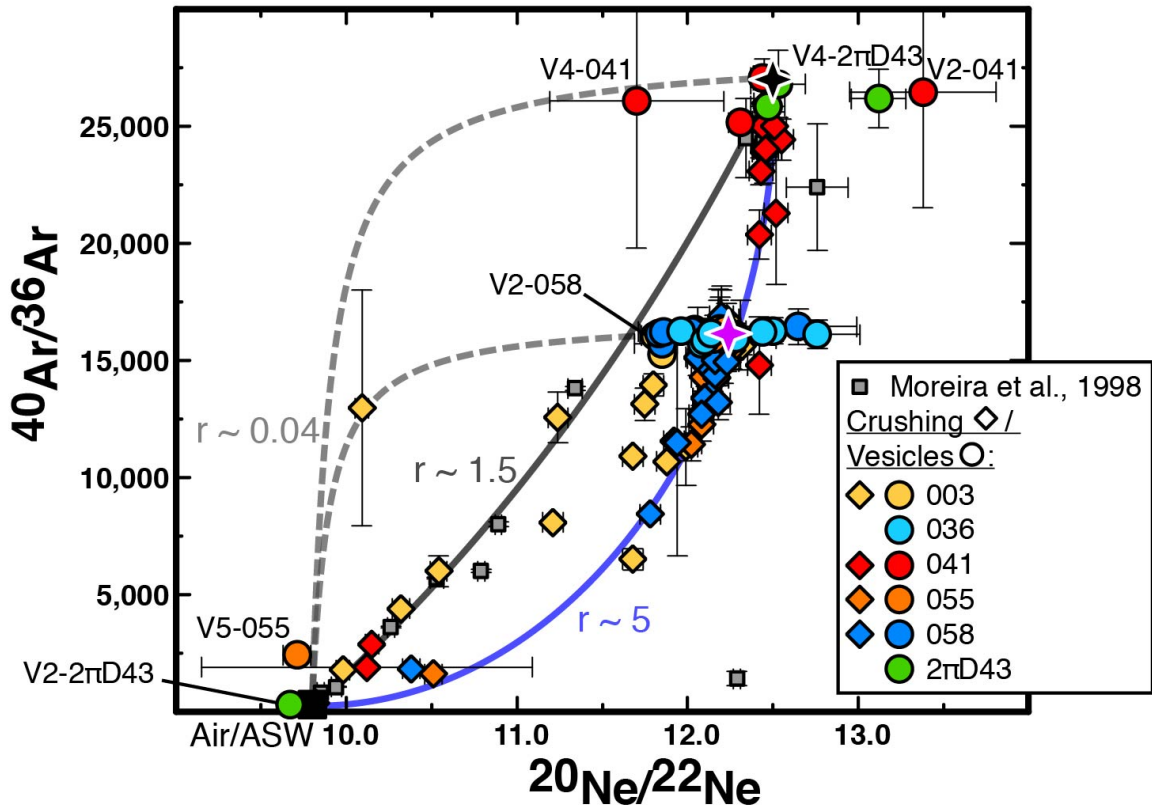
763
 764
 765
 766
 767
 768
 769
 770
 771
 772
 773
 774
 775

776 **Figure 2.** Neon-neon isotope diagram comparing (a) step-crushing data and (b) laser ablation
 777 data. Data from Moreira et al. (1998) for the popping rock 2πD43 are indicated with the grey
 778 squares. The black star indicates the mantle composition from the popping rock 2πD43, the
 779 mauve star corresponds to the mean vesicle composition (vesicles with $^{20}\text{Ne}/^{22}\text{Ne} > 12$).
 780 ASW, *air-saturated seawater*. Strong CO₂ memory effects were observed during analysis of
 781 vesicle V4-2πD43. Vesicle V2-041 is one of the smallest vesicles analyzed. The neon
 782 compositions of these vesicles are thus unreliable. *MORB line* as defined by the 2πD43 data
 783 of Moreira et al. (1998).
 784



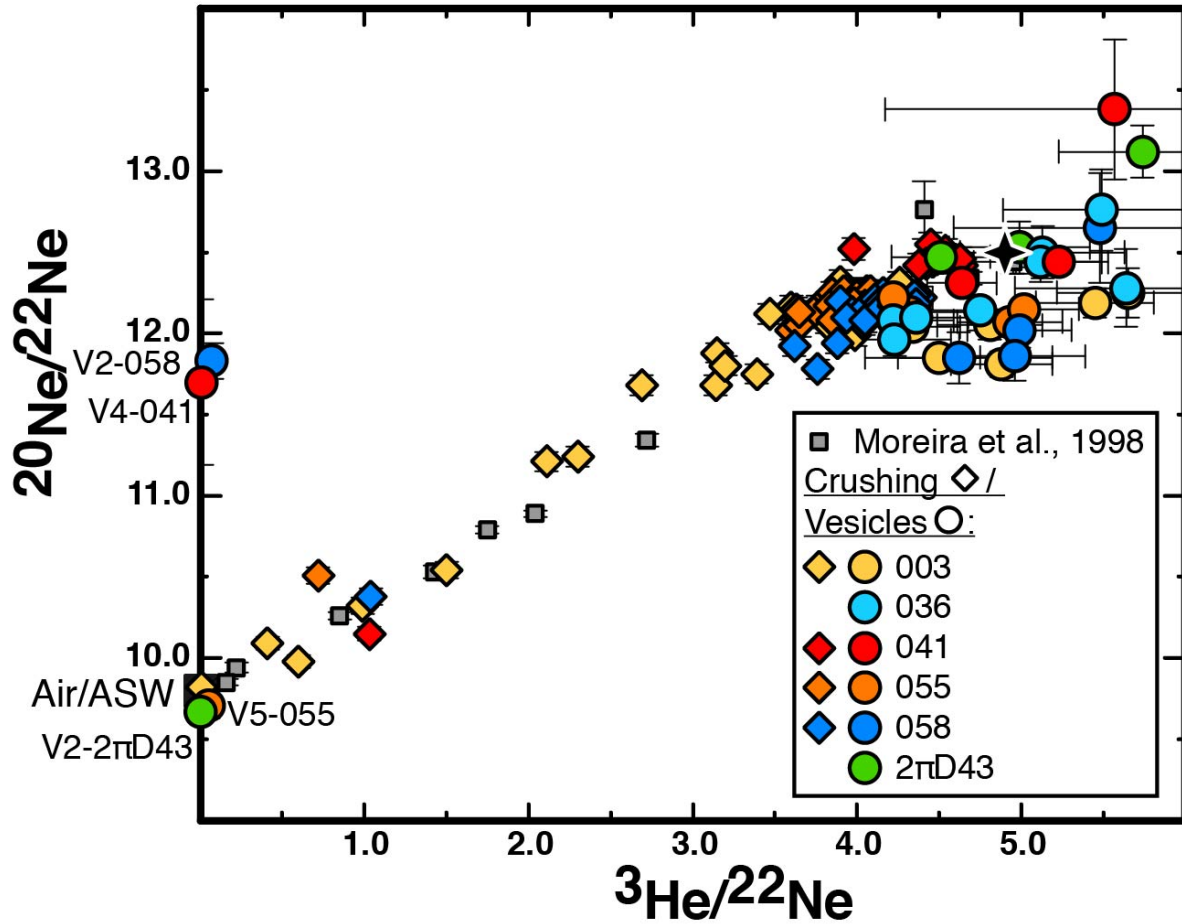
785
 786
 787

788 **Figure 3.** $^{40}\text{Ar}/^{36}\text{Ar}$ ratios vs. $^{20}\text{Ne}/^{22}\text{Ne}$ ratios for the step-crushing and laser ablation data.
 789 Data from Moreira et al. (1998) for the popping rock 2 π D43 are indicated with the grey
 790 squares. The black star indicates the mantle composition from the popping rock 2 π D43, the
 791 mauve star corresponds to the mean vesicle composition (vesicles with $^{20}\text{Ne}/^{22}\text{Ne} > 12$).
 792 ASW, *air-saturated seawater*. The curvatures r of the hyperbolas are defined by the ratio
 793 $(^{36}\text{Ar}/^{22}\text{Ne})_{\text{air/SW}} / (^{36}\text{Ar}/^{22}\text{Ne})_{\text{mantle}}$. The neon compositions of V2-041 and V4-2 π D43
 794 are unreliable (section 3.3).
 795



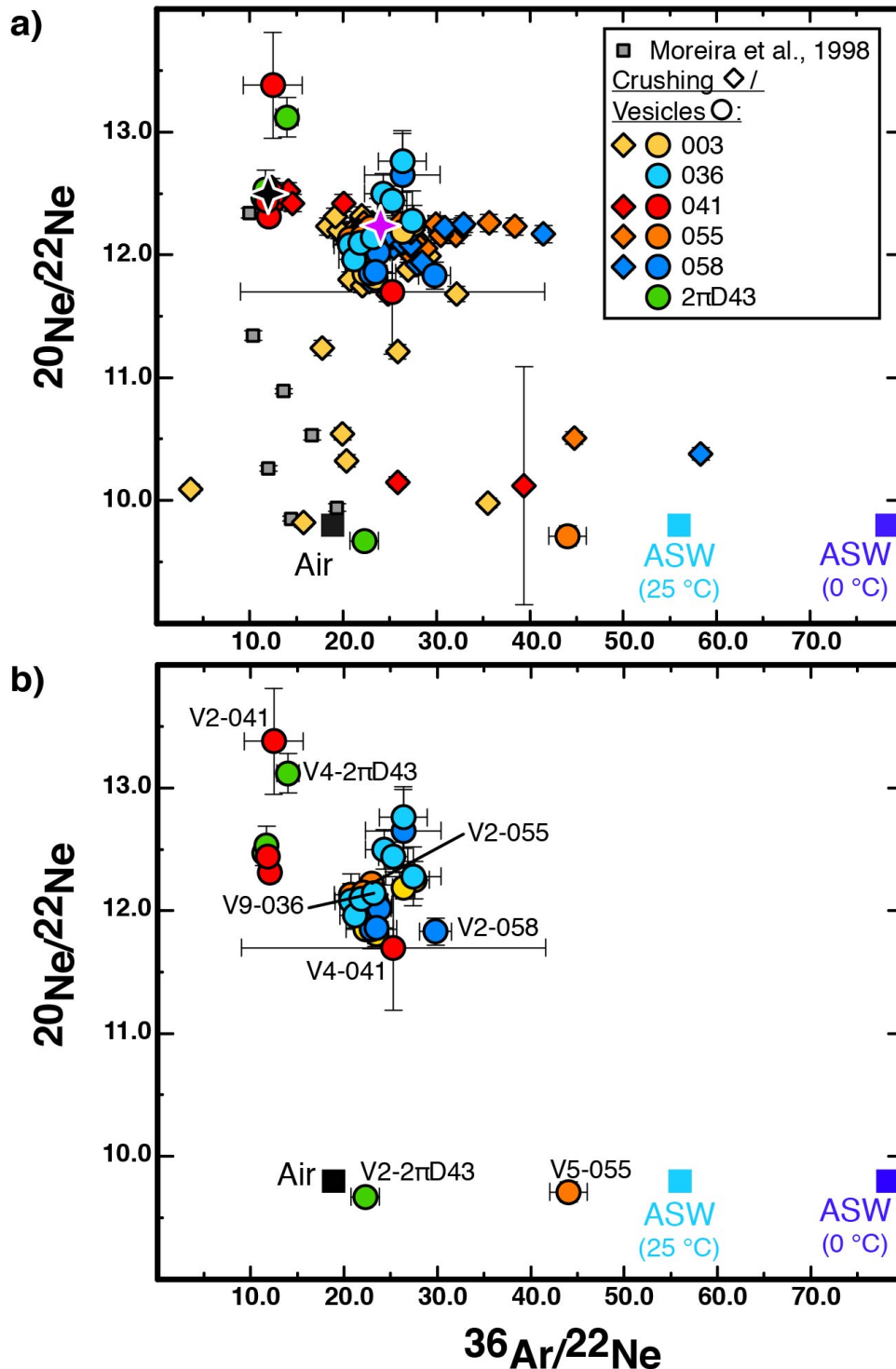
796
 797
 798
 799
 800
 801
 802
 803
 804
 805
 806
 807
 808
 809
 810
 811
 812
 813
 814
 815

816 **Figure 4.** $^{20}\text{Ne}/^{22}\text{Ne}$ ratios vs. $^3\text{He}/^{22}\text{Ne}$ ratios for the step-crushing and laser ablation data.
 817 Data from Moreira et al. (1998) for the popping rock 2πD43 are indicated with the grey
 818 squares. The black star indicates the mantle composition from the popping rock 2πD43
 819 (Moreira et al., 1998). ASW, *air-saturated seawater*.
 820
 821



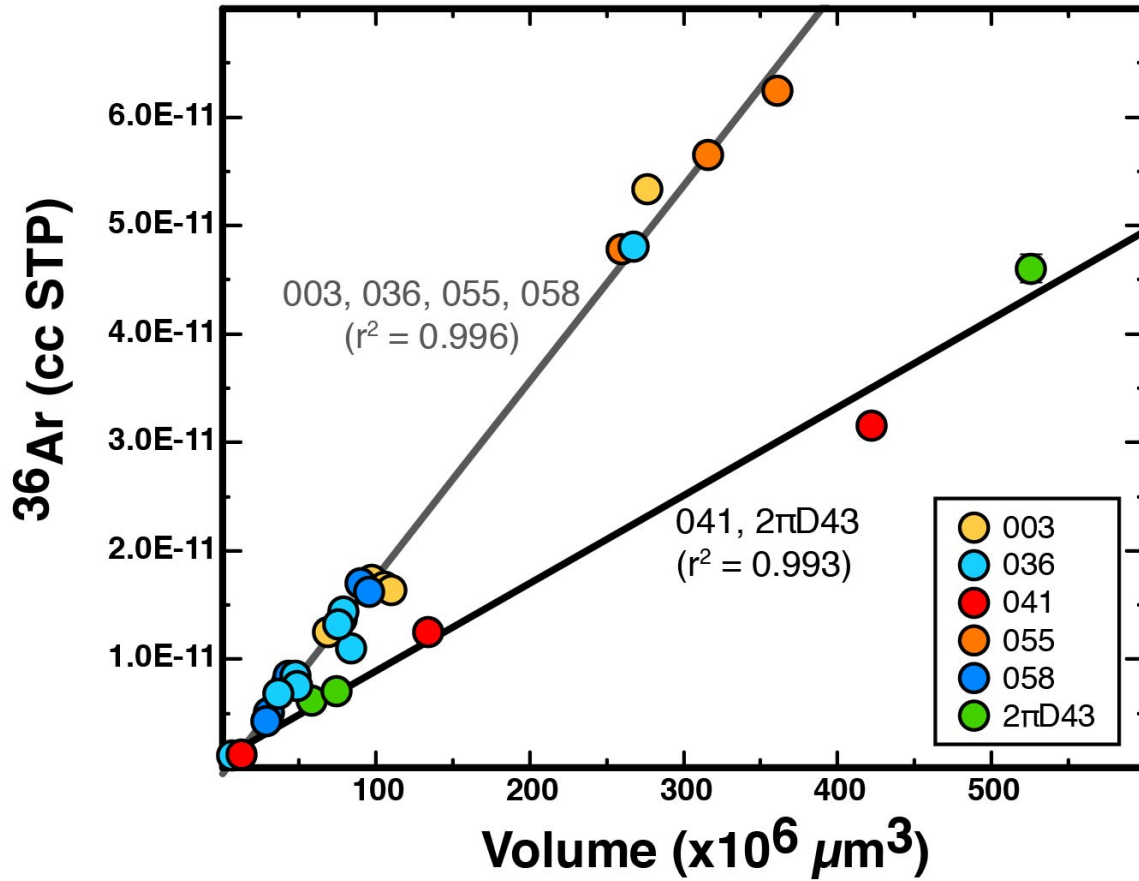
822
 823
 824
 825
 826
 827
 828
 829
 830
 831
 832
 833
 834
 835
 836
 837
 838
 839
 840

841 **Figure 5.** $^{20}\text{Ne}/^{22}\text{Ne}$ ratio vs. $^{36}\text{Ar}/^{22}\text{Ne}$ ratio for the step-crushing and laser ablation data (a).
 842 The bottom panel (b) shows only the vesicles. Data from Moreira et al. (1998) for the popping
 843 rock $2\pi\text{D43}$ are indicated with the grey squares. Air (black square), air-saturated seawater at
 844 25°C (light blue square), air-saturated seawater at 0°C (dark blue square), air-saturated
 845 seawater data are from Ozima and Podosek (2002). The black star indicates the mantle
 846 composition from the popping rock $2\pi\text{D43}$, the mauve star corresponds to the mean vesicle
 847 composition (vesicles with $^{20}\text{Ne}/^{22}\text{Ne} > 12$). The crushing data with $^{36}\text{Ar}/^{22}\text{Ne} > 30$ and
 848 $^{20}\text{Ne}/^{22}\text{Ne} > 12$ correspond to analyses with low ^{36}Ar contents and so these $^{36}\text{Ar}/^{22}\text{Ne}$ ratios
 849 may be poorly constrained.
 850



851

852 **Figure 6.** ^{36}Ar contents (in cm^3 STP) of the vesicles associated with their volumes. Samples
 853 041 and $2\pi\text{D43}$ plot on a different line than samples 003, 036, 055 and 058. The linear
 854 regression coefficients are indicated for the two lines. Vesicles V5 055 and V2 $2\pi\text{D43}$ are
 855 strongly contaminated, vesicle V4 041 has lost gases and vesicle V9 036 is special due to
 856 coalescence (section 3.1) and so they are not represented.
 857



858
 859
 860
 861
 862
 863
 864
 865
 866
 867
 868
 869
 870
 871