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Influence of aperiodic non-tidal atmospheric and oceanic loading deformations on the stochastic properties of global GNSS vertical land motion time series

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Key Points:

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14	• NTAOL deformations cause a latitudinal bias in parameter estimates of the white
15	+ power-law noise model usually describing GNSS time series.
16	• Removing NTAOL deformations corrects the latitudinal bias and dramatically re-
17	duces station velocity uncertainties at high latitudes.
18	• NTAOL deformations also introduce a seasonal spatial correlation in GNSS ver-
19	tical land motion time series.

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20 Abstract

Monitoring vertical land motions (VLMs) at the level of 0.1 mm/yr remains one 21 of the most challenging scientific applications of global navigation satellite systems (GNSS). 22 Such small rates of change can result from climatic and tectonic phenomena, and their 23 detection is important to many solid Earth-related studies, including the prediction of 24 coastal sea-level change and the understanding of intraplate deformation. Reaching a level 25 of precision allowing to detect such small signals requires a thorough understanding of 26 the stochastic variability in GNSS VLM time series. This paper investigates how the ape-27 28 riodic part of non-tidal atmospheric and oceanic loading (NTAOL) deformations influences the stochastic properties of VLM time series. Using the time series of over 10,000 29 stations, we describe the impact of correcting for NTAOL deformation on 5 complemen-30 tary metrics, namely: the repeatability of position residuals, the power-spectrum of po-31 sition residuals, the estimated time-correlation properties, the corresponding velocity un-32 certainties, and the spatial correlation of the residuals. We show that NTAOL deforma-33 tions cause a latitude-dependent bias in white noise plus power-law model parameter es-34 timates. This bias is significantly mitigated when correcting for NTAOL deformation, 35 which reduces velocity uncertainties at high latitudes by 70%. Therefore, removing NTAOL 36 deformation before the statistical analysis of VLM time series might help to detect sub-37 tle VLM signals in these areas. Our spatial correlation analysis also reveals a seasonal-38 ity in the spatial correlation of the residuals, which is reduced after removing NTAOL 39 deformation, confirming that NTAOL is a clear source of common-mode errors in GNSS 40 VLM time series. 41

42 Plain Language Summary

Monitoring vertical land motions (VLMs) at the level of 0.1 mm/yr remains one 43 of the most challenging scientific applications of global navigation satellite systems (GNSS). 44 Such small rates of change can result from climatic and tectonic phenomena, and their 45 detection is important to the prediction of coastal sea-level change and the understand-46 ing of deformation processes acting within tectonic plates. In this paper, we investigate 47 to what extent accounting for the deformation of the Earth's crust due to changes in the 48 distribution of the atmospheric and the oceanic pressure helps to detect small vertical 49 ground deformation. By analyzing the data from over 10,000 globally distributed GNSS 50 stations, we show that accounting for such atmospheric and oceanic deformation can re-51 duce by 70% the uncertainty on the vertical velocity of GNSS stations at high latitudes, 52 which could, for instance, improve sea-level change monitoring in these areas. 53

54 1 Introduction

Analyzing vertical land motion (VLM) time series estimated from global naviga-55 tion satellite systems (GNSS) provides key information about the deformation of the Earth. 56 In particular, estimating GNSS station long-term vertical velocities and their uncertain-57 ties is a crucial step, aiming at giving reliable inputs for the understanding of inter- and 58 intraplate tectonic deformation (Elliott et al., 2016; Kreemer et al., 2020), glacial iso-59 static adjustment (Nocquet et al., 2005; Schumacher et al., 2018; Husson et al., 2018), 60 and coastal sea-level change (Wöppelmann et al., 2007; Pfeffer & Allemand, 2016; Ballu 61 et al., 2019). 62

Measuring vertical deformations at velocities of a few 0.1 mm/yr is one of the most challenging scientific applications of GNSS, required to monitor, for instance, slow intraplate (Masson et al., 2019; Craig & Calais, 2014) and climate change-related Earth deformations (Wöppelmann & Marcos, 2016). This metrological goal could be met by improving the quality of GNSS positioning, the stability of reference frames, and the statistical modeling of geodetic time series.

To estimate long-term velocities, it is customary to decompose the position time 69 series as the sum of two parts: a deterministic component and a stochastic variability 70 (Bos et al., 2020). In practice, the first is described by a trajectory model, whereas the 71 latter is described by a stochastic model. The trajectory model typically consists of known 72 parametric functions – trends, offsets, periodic oscillations, and episodic nonlinear mo-73 tions – depending on unknown deterministic parameters (Bevis & Brown, 2014). The 74 stochastic model consists of a set of stochastic processes with known covariance func-75 tions depending on unknown stochastic parameters – or variance components – (Bos et 76 al., 2020). 77

Since the late 1990s, numerous studies demonstrated that the stochastic variability of GNSS position time series, often referred to as noise, can be approximated by a linear combination of white noise (WN) and power-law (PL) processes (Zhang et al., 1997; Mao et al., 1999; Calais, 1999; Williams et al., 2004; Langbein, 2008; Santamaría-Gómez et al., 2011). Power-law processes are stochastic processes characterized by their power spectra P(f), which follows a power-law function of the time frequency f, such that

$$P(f) \approx P_0 \left(\frac{f}{f_0}\right)^{\kappa},\tag{1}$$

where both P_0 and f_0 are normalization constants, and κ is the so-called spectral index (Agnew, 1992).

The spectral index κ is a scalar identifying the nature of the power-law correlation. Its reported value usually lies within [-3, 1] in Earth sciences, and within [-2, 0] in GNSS time series analysis. Some power-law processes with integer values of κ are known as white - that is, uncorrelated – noise ($\kappa = 0$), flicker noise ($\kappa = -1$), and random walk noise ($\kappa = -2$). The power-law processes with $\kappa < 0$ are time-correlated. For $\kappa \leq -1$, the power-law process is also non-stationary, and the variance of the process diverges with time.

Accounting for the power-law behavior of the stochastic variability in GNSS po-93 sition time series is essential for long-term velocity estimation to obtain realistic estima-94 tion of the uncertainties (Williams, 2003a; Santamaría-Gómez et al., 2011). Although 95 considered in many geodetic studies (Agnew, 1992; Langbein, 2004; Van Camp et al., 96 2005), the presence of power-law processes in GNSS position time series remains largely 97 unexplained (Santamaría-Gómez & Ray, 2021). Given that these power-law processes 98 are the primary source of uncertainty for velocity estimates, it is crucial to identify their 99 origin(s) and to reduce their influence on the position time series. 100

The stochastic variability observed in position time series is also spatially corre-101 lated (Wdowinski et al., 1997). In global solutions, significant spatial correlations are de-102 tected between stations up to a few thousand kilometers apart (Williams et al., 2004; 103 Amiri-Simkooei, 2009, 2013; Amiri-Simkooei et al., 2017; Benoist et al., 2020). The pres-104 ence of large-scale spatial correlations indicates that the stochastic variability cannot be 105 fully attributed to station-specific errors, and that it probably results from a combina-106 tion of large scale positioning errors and non-modelled geophysical or climatic processes 107 (Amiri-Simkooei et al., 2017; Kreemer & Blewitt, 2021). 108

Possible candidates for explaining large-scale structures in the stochastic variability are the non-tidal loading (NTL) deformations, that is, the elastic response of the solid Earth to changes in the distribution of atmospheric, oceanic, and hydrological loads. Following the International Earth Rotation and Reference System Service (IERS) conventions (Petit & Luzum, 2010), NTL corrections are usually not applied during the estimation of position time series, therefore NTL signals remain in VLMs time series products.

Over the last two decades, numerous studies demonstrated that NTL corrections 116 derived from numerical climatic load and elastic Earth models help reduce the stochas-117 tic variability in GNSS VLM time series, whether applied at the GNSS observation level 118 that is, prior to the computation of daily positions – (Gegout et al., 2010; Dach et al., 119 2011; Tregoning & van Dam, 2005; Männel et al., 2019) or at the solution level – that 120 is, after computing daily positions – (Van Dam et al., 1994, 2001; Williams & Penna, 2011; 121 Van Dam et al., 2012; Martens et al., 2020; Klos et al., 2021). Regarding VLMs, the dif-122 ferences between the two correction strategies seem negligible (Tregoning & Watson, 2009). 123

124 As numerical climatic load models perform better at retrieving atmospheric and oceanic than hydrological loads (Mémin et al., 2020), in this work, we focus on assess-125 ing the effect of non-tidal atmospheric and oceanic loading (NTAOL) deformations on 126 the stochastic properties of GNSS VLM time series. NTAOL deformations introduce both 127 periodic and aperiodic deformation in VLM time series. Because annual and semi-annual 128 signals are always accounted for in the trajectory model, for time series longer than a 129 few years, the periodic (i.e., seasonal) part of NTAOL deformations has a negligible in-130 fluence on velocity uncertainties (Blewitt & Lavallée, 2002). This study, therefore, fo-131 cuses on the impact of aperiodic NTAOL deformations. 132

Usually, the influence of NTAOL on the stochastic properties of GNSS time series 133 is investigated through the study of change in root mean squared errors (RMSEs) (Tregoning 134 & van Dam, 2005; Martens et al., 2020), sometimes at different frequency bands (Mémin 135 et al., 2020; Klos et al., 2021). In this work, to further study possible changes in time-136 correlation and try to identify the origin of the stochastic variability in GNSS time se-137 ries, we investigate the influence of correcting for NTAOL deformation (at the solution 138 level) on 5 complementary metrics, namely: the repeatability of position residuals, the 139 power spectrum of position residuals, their time-correlation properties, the correspond-140 ing velocity uncertainties, and the spatial correlation of position residuals. 141

To describe the stochastic variability and the influence of aperiodic NTAOL defor-142 mations, we investigate each of these metrics using the position time series of over 10,000143 globally distributed GNSS stations. Considering as many stations allows describing with 144 an unprecedented resolution the impact of NTAOL deformations on the spatial corre-145 lation of GNSS VLM time series. Besides, systematically analyzing the impact of NTAOL 146 deformations on the stochastic properties of each station allows us to evidence previously 147 unreported stochastic parameter estimation issues, resulting in recommendations for the 148 estimation of velocity uncertainties from global GNSS VLM time series. 149

150 **2 Data**

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2.1 GNSS position time series

We used the North, East, and Up position time series of a total of 10.151 GNSS 152 stations processed and distributed by the Nevada Geodetic Laboratory at the Univer-153 sity of Nevada, Reno (Blewitt et al., 2018). This study focuses on the Up coordinates, 154 however, the North and East coordinates have been used in the preprocessing for the de-155 tection of discontinuities. NGL's position time series have been computed using the Jet 156 Propulsion Laboratory's (JPL) GipsyX 1.0 software (Bertiger et al., 2020) based on single-157 station precise point positioning (PPP) with carrier phase ambiguity resolution (Zumberge 158 et al., 1997). The satellite orbit and clock used for the PPP processing are JPL's Re-159 pro 3.0 products. Though we employ the term GNSS for the sake of generality, NGL's 160 processing only exploited observations from the Global Positioning System (GPS). The 161 final position time series are expressed in the IGS14 reference frame. Incremental zenith 162 wet delays are estimated on top of a priori hydrostatic and wet zenith tropospheric de-163 lays interpolated from the VMF1 grids (Boehm et al., 2006). Both hydrostatic and wet 164 zenith delays are mapped to the observation elevations using the Vienna Mapping Func-165

tion (VMF1). For the ionospheric delays, the first-order effects are removed thanks to
the use of the ionosphere-free L1/L2 combination and second-order effects are corrected
using ionospheric TEC maps together with a model of Earth's magnetic field. For more
details about NGL's processing strategy, we refer to http://geodesy.unr.edu and Kreemer
et al. (2020) or Martens et al. (2020).

2.2 Loading deformation time series

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To remove NTAOL deformations from GNSS position time series, we used the non-172 tidal atmospheric loading (NTAL) and non-tidal oceanic loading (NTOL) deformation 173 time series computed by the Earth System Modelling group of the German Research Cen-174 ter for Geosciences at Potsdam (ESMGFZ) (Dill & Dobslaw, 2013). The vertical defor-175 mations caused by NTAL are calculated by the ESMGFZ using the 3-hourly atmospheric 176 surface pressure from the European Center for Medium-Range Weather Forecast (ECMWF). 177 The vertical deformations caused by NTOL are calculated using the 3-hourly ocean bot-178 tom pressure from the Max-Planck-Institute Meteorology Ocean Model (MPIOM). For 179 both load types, vertical displacements are computed using patched Green's functions 180 computed from the ak135 elastic Earth model (Kennett et al., 1995). For more details 181 about the ESMGFZ processing strategy, we refer to Dill and Dobslaw (2013). 182

NGL's position time series being expressed in the IGS14 reference frame, we use
the deformation time series expressed in the center of Earth's figure frame (CF) (Blewitt,
2003). Because of the higher sampling rate of ESMGFZ's NTAOL deformation time series, we down-sample ESMGFZ's time series at NGL's time series epochs.

¹⁸⁷ **3** Time correlation analysis

3.1 Position selection

For each station, to remove possibly unreliable position estimates, epochs showing a formal error above 5 mm on any of the three coordinates were ignored. Moreover, to eliminate large outliers, the raw position estimates were compared to a running monthly median. Any epoch with a position showing a deviation from the median exceeding 5 times the median absolute deviation, on any of the three coordinates, was discarded (Klos et al., 2015).

Only position estimates prior to January 2018 were considered in this study. Also, as we focus on time correlation, we only used stations presenting time series with over 1090 daily estimates, that is, more than about 3 years of continuous observations. The spans of the considered time series range from 3 to over 20 years, with a median of 8 years. The distribution of the GNSS sites with their number of data points are presented in Figure S1, provided in the Supporting Information.

3.2 Trajectory model

Noting \mathbf{y} the $m \times 1$ observation vector (i.e., a GNSS VLM time series), the trajectory model describes the expectation $\mathsf{E}\{.\}$ of \mathbf{y} . To simplify the processing of thousands of stations, this study only uses linear trajectory models accounting for a linear trend, periodic signals, remaining outliers, offsets, and trend discontinuities. The trajectory model can be expressed as

$$\mathsf{E}\{\mathbf{y}\} = \mathbf{A}\mathbf{x},\tag{2}$$

where **A** is the $m \times n$ design matrix, that is, the Jacobian of the trajectory model, and **x** is the unknown $n \times 1$ deterministic parameter vector. The periodic signals accounted for in this trajectory model include the most commonly observed ones in GNSS time series (Ray et al., 2008; Amiri-Simkooei, 2013), namely, the annual signal with a period of 365.25 days and its first 2 harmonics, the draconitic signal with a period of 351.40 days and its first 7 harmonics, and the 3 fortnightly signals with periods of 13.62, 14.17, and 14.76 days.

To model the offsets, all the offset dates indicated in the NGL database have been considered by default. This database includes dates of known equipment changes and possible earthquake-related discontinuities. This database is not necessarily complete for all stations, and additional offsets must be detected to improve the trajectory model. This issue is addressed using an automatic detection of discontinuities presented in section 3.6.

3.3 Stochastic model

The stochastic model describes the co-variance $var\{.\}$ of the $m \times 1$ observation vector **y**. Under the assumption that the stochastic variability can be approximated by a linear combination of homogeneous white noise – that is, uncorrelated with constant variance – and a power-law process (WN+PL), the stochastic model reads

$$\operatorname{var}\{\mathbf{y}\} = \mathbf{Q}_{\mathbf{y}} = \sigma_{hw}^2 \mathbf{I} + \sigma_{pl}^2 \mathbf{Q}_{\kappa}.$$
(3)

where σ_{hw}^2 is the unknown white noise amplitude, **I** is the $m \times m$ identity matrix, σ_{pl}^2 is the unknown power-law process amplitude, κ is the unknown spectral index, and \mathbf{Q}_{κ} is the $m \times m$ covariance matrix associated to a power-law process of spectral index κ (Williams, 2003a).

3.4 Variance component estimation

Estimating unknown parameters of the stochastic model can be done using Vari-230 ance Component Estimation (VCE) methods. In this study, we use the Least-Squares 231 Variance Component Estimation (LS-VCE) method to estimate the amplitude of each 232 process, namely σ_{hw}^2 and σ_{pl}^2 . Introduced by Teunissen (1988) and further developed by 233 Teunissen and Amiri-Simkooei (2008) and Amiri-Simkooei (2007), the LS-VCE method 234 is an unbiased and minimum-variance alternative to the more commonly used Maximum 235 Likelihood Estimation (MLE) method (Zhang et al., 1997; Williams, 2008; Bos et al., 236 2008; Langbein, 2017) and has already been used in its univariate and multivariate forms 237 for GNSS noise analyses (Amiri-Simkooei et al., 2007; Amiri-Simkooei, 2009; Amiri-Simkooei 238 et al., 2017). 239

Because the LS-VCE method is unconstrained, it can occasionally return negative estimates, which makes no physical sense for stochastic process amplitudes. When encountering negative variance estimates, we systematically applied the Non-Negative LS-VCE method (NNLS-VCE) introduced in Amiri-Simkooei (2016) to obtain non-negative process amplitudes. Though mentioned explicitly in this paper, the non-negativity constraint is the norm, since the nonlinear optimization methods used in most software simply do not explore negative solutions (Williams, 2008; Bos et al., 2008; Langbein, 2017).

In their linear forms, both the LS-VCE and the NNLS-VCE methods do not allow estimating the unknown spectral index κ . Hence, to assess both the process amplitudes $\hat{\sigma}_{hw}^2$ and $\hat{\sigma}_{pl}^2$, and the spectral index $\hat{\kappa}$, we systematically tested for each time series 30 stochastic models with discrete values of κ ranging from -1.8 to -0.3 with a step of +0.05. The most likely $\hat{\kappa}$ and the corresponding process amplitudes, were chosen as those maximizing the restricted likelihood defined in Koch (1986). ²⁵³ When VLM is expressed in millimeters (mm), the power-law amplitude estimate $\hat{\sigma}_{pl}$ is expressed in (mm yr^{κ /4}) if the scaling of Williams (2003a) is used. Therefore, two σ_{pl} obtained with different spectral indices are not comparable. To overcome this issue, we use a modified standard deviation, noted $\hat{\sigma}'_{pl}$, expressed in mm and representing the square root of the expected empirical variance of a power-law process over 8 years, which is the median cumulative span of the considered VLM time series.

Noting **r** a random vector such that $\mathbf{r} \sim \mathcal{N}(\mathbf{0}, \hat{\sigma}_{pl}^2 \mathbf{Q}_{\kappa})$, the amplitude $\hat{\sigma}'_{pl}$ is defined as

$$\hat{\sigma}_{pl}' = \sqrt{\mathsf{E}\left\{\frac{1}{m'}\sum_{i=1}^{m'} (r_i - \bar{r})^2\right\}}$$
(4)

where r_i is an element of the vector \mathbf{r} , m' is the number of observations for an 8-year long time series, and \bar{r} is the arithmetic mean of the r_i .

As power-law processes with $\kappa \leq -1$ are non-stationary, their modified standard deviations $\hat{\sigma}'_{pl}$ increase with m'. Therefore, to allow comparisons between $\hat{\sigma}'_{pl}$ values, m'is kept fixed for all time series.

One can show that
$$\hat{\sigma}'_{pl}$$
 can be expressed as

$$\hat{\sigma}_{pl}' = \hat{\sigma}_{pl} \sqrt{\frac{1}{m'}} \operatorname{tr} \{ \mathbf{Q}_{\kappa}' \} - \frac{1}{{m'}^2} \mathbf{u}^T \mathbf{Q}_{\kappa}' \mathbf{u}$$
(5)

with **u** a $m' \times 1$ vector defined as $\mathbf{u}^T = \begin{bmatrix} 1 & 1 & \cdots & 1 \end{bmatrix}$ and \mathbf{Q}'_{κ} the covariance matrix of a power-law process with spectral index κ and length m'.

3.5 Residual estimation

Part of the investigations presented in this work are based on the analysis of the
residual VLM time series. Residual time series are computed as the difference between
the observations and the trajectory model estimated by weighted least-squares (Teunissen,
2000a). In particular, the residual vector ê is computed from the observation vector y
by

$$\hat{\mathbf{e}} = \mathbf{P}_{\mathbf{A}}^{\perp} \mathbf{y},\tag{6}$$

in which the $m \times m$ matrix $\mathbf{P}_{\mathbf{A}}^{\perp}$ is the weighted least-squares orthogonal projector defined by

$$\mathbf{P}_{\mathbf{A}}^{\perp} = \mathbf{I} - \mathbf{A} (\mathbf{A}^T \mathbf{W} \mathbf{A})^{-1} \mathbf{A}^T \mathbf{W}.$$
 (7)

where **W** denotes the $m \times m$ weight matrix. Hereafter, unless specified otherwise, the weight matrix **W** is taken as the inverse of the covariance matrix of the observations: $\mathbf{W} = \mathbf{Q}_{\mathbf{y}}^{-1}$.

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3.6 Trajectory model improvement

As undetected offsets can bias stochastic parameter estimates (Williams, 2003b), it is necessary to identify possible additional offsets to the NGL database and to include them in the trajectory models. In practice, the offset detection is often done manually, since experimented operators tend to perform better than most algorithms (Gazeaux et al., 2013). However, when considering thousands of time series, an automatic detection
 is necessary.

Amiri-Simkooei et al. (2019) recently showed that one can improve the automatic 287 detection of offsets by accounting for the time-correlated noise, and by analyzing the North 288 (n), East (e), and Up (h) coordinates simultaneously. The multivariate LS-VCE used 289 in Amiri-Simkooei et al. (2019) assumes constant $\sigma_{hw}^2/\sigma_{pl}^2$ ratios for all coordinates, which 290 might be a risky hypothesis. Here, we relax this hypothesis and adapt the method of Amiri-291 Simkooei et al. (2019) to allow process amplitudes to have different amplitude ratios for 292 each coordinate. We then check for discontinuities, including offsets, outliers, and veloc-293 ity changes. 294

The detection method is derived from the Detection Identification Adaptation (DIA) procedure introduced by Baarda (1968); Teunissen (2000b). Applied to the detection of offsets, outliers, or velocity changes, the DIA procedure consists in testing a null hypothesis $H_0: \mathsf{E}\{\mathbf{y}\} = \mathbf{A}\mathbf{x}$ against an alternative hypothesis $H_a: \mathsf{E}\{\mathbf{y}\} = \mathbf{A}\mathbf{x} + \mathbf{c}_a(t_k)x_a$, in which $\mathbf{c}_a(t_k)$ is the known $m \times 1$ Jacobian vector of a discontinuity of type a (either an offset, an outlier or a velocity change) at time t_k with an unknown amplitude x_a . According to Teunissen (2000b), accepting or rejecting H_0 depends on the value of a T-statistic defined by

$$T_a(t_k) = \frac{(\mathbf{c}_a^T(t_k)\mathbf{Q}_{\mathbf{y}}^{-1}\hat{\mathbf{e}})^2}{\mathbf{c}_a^T(t_k)\mathbf{Q}_{\mathbf{y}}^{-1}\mathbf{P}_{\mathbf{A}}^{\perp}\mathbf{c}_a(t_k)}.$$
(8)

Under H_0 , the random variable $T_a(t_k)$ follows a central chi-squared distribution with 1 degree of freedom: $T_a(t_k) \sim \chi^2(1,0)$. Hence, for a given confidence level α , if $T_a(t_k) > \chi^2_{\alpha}(1,0)$, the null hypothesis can be rejected and \mathbf{c}_a can be added to the columns of the design matrix \mathbf{A} .

To combine the information of all coordinates, we define the power $P_a(t_k)$ of a discontinuity $\mathbf{c}_a(t_k)$ at time t_k as the sum of the *T*-statistics estimated on each coordinate, that is

$$P_a(t_k) = \sum_{i \in \{n,e,h\}} T_a^{(i)}(t_k),$$
(9)

where the residual vector $\hat{\mathbf{e}}_{(i)}$, the covariance matrix $\mathbf{Q}_{\mathbf{y}_{(i)}}$, and the orthogonal projector $\mathbf{P}_{\mathbf{A}_{(i)}}^{\perp}$ are specific to each coordinate (*i*).

Assuming that the coordinates are uncorrelated (Amiri-Simkooei, 2009; Benoist et al., 2020), under H_0 , the power $P_a(t_k)$ follows a central chi-squared distribution with 3 degrees of freedom, that is $P_a(t_k) \sim \chi^2(3,0)$.

The chi-squared distribution of $T_a(t_k)$ and $P_a(t_k)$ is based on the assumption that the covariance matrices $\mathbf{Q}_{\mathbf{y}(i)}$ are known. However, the two variance components σ_w^2 and σ_{pl}^2 are unknown and must be estimated. In such a case, the exact distribution of $T_a(t_k)$ becomes intractable. Nonetheless, the chi-squared distribution remains a good approximation, as the redundancy of the trajectory model m-n is large (Amiri-Simkooei et al., 2019).

To find a candidate for the alternative hypothesis $\mathbf{c}_a(t_k)$, offsets $\mathbf{c}_o(t_k)$, outliers $\mathbf{c}_t(t_k)$, and velocity changes $\mathbf{c}_v(t_k)$ were systematically tested for each epoch t_k . The most-likely discontinuity date \hat{t}_k and type \hat{a} , were chosen as

$$\hat{a}, \hat{t}_k = \underset{a,t_k}{\operatorname{arg\,max}} P_a(t_k). \tag{10}$$

When the power of the most likely alternative $P_{\hat{a}}(\hat{t}_k)$ exceeded a given threshold 324 $P_{\hat{a}}(\hat{t}_k) > 25$, the discontinuity $\mathbf{c}_{\hat{a}}(\hat{t}_k)$ was added to the trajectory model. The thresh-325 old value 25 is arbitrarily chosen and corresponds to a confidence level $\alpha > 0.9999$ un-326 der the $\chi^2(3,0)$ hypothesis. This detection procedure was repeated until no significant 327 discontinuities were found anymore. At each iteration, the variance components of the 328 stochastic model were (re)estimated for all coordinates. To reduce the computational bur-329 den, and for this detection step only, we used a fixed spectral index $\kappa = -0.80$ and the 330 NNLS-VCE method by default. 331

332 In the end, an average of 2.49 offsets, 0.67 outliers, and 0.97 velocity changes were added to the trajectory model of each station. To compensate for the linearity of the tra-333 jectory model, stations showing nonlinear behaviors such as post-seismic deformation present 334 more offsets and velocity changes than the others. But the median number discontinu-335 ity per station is actually 1.0 for the offsets and 0.0 for both the outliers and the veloc-336 ity changes. The imperfect nature of this detection procedure certainly results in an over-337 parameterization of the trajectory model of some stations. However, considering that 338 we focus on large scale patterns obtained from a large number of independently processed 339 stations, a possible over-parameterization of a few trajectory models is unlikely to af-340 fect the results discussed in section 5. 341

³⁴² 4 Spatial correlation analysis

Williams et al. (2004) computed the variation of the cross-correlation between pairs 343 of residual time series with the station's separation distance. Amiri-Simkooei (2009); Amiri-344 Simkooei et al. (2017) did a similar analysis but used a particular metric to take time 345 correlation into account. However, this metric, based on the multivariate LS-VCE (Amiri-346 Simkooei, 2009), assumes an identical white noise to power-law amplitude ratio and an 347 identical spectral index for all the time series. Benoist et al. (2020) assumed a second-348 order stationary spatial distribution of residuals and investigated the spatial correlation 349 using variograms (Wackernagel, 2013). 350

In these studies, the spatial correlation coefficients were systematically estimated using all observations over a span of a few years. In results, the estimated coefficient was an average over the considered span. In practice, for short time scales, let us say a few weeks, one would expect the spatial correlation structure to show a time-variable behavior, especially if position residuals are linked to climate-related mass redistribution.

Hence, to investigate the time dependence of the spatial correlation, we computed 356 the Pearson correlation coefficients between pairs of residual time series over running win-357 dows of 30 days. We then averaged these pairwise correlation coefficients for 200 inde-358 pendent classes of separation distances, ranging from [0, 100] to [19900, 20000] km. Cor-359 relation coefficients obtained from 30-day windows are less precise than those based on 360 several years. However, given the number of stations analyzed in this study, the final spa-361 tial correlation estimates result from an average of over a few hundred thousands pairs 362 in each distance class, making them robust and interpretable nevertheless. 363

³⁶⁴ 5 Influence of aperiodic NTAOL deformations

5.1 Repeatability

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To measure the change in repeatability resulting from the correction of either NTAL or NTOL deformations, we analyze RMSE reductions. In this part, residuals are computed using the unweighted least-squares adjustment, that is, using W = I in Equation 7. As the trajectory model accounts for annual and semi-annual signals, the changes in RMSE do not reflect changes of the seasonal variations in the series, as those are not present in the residuals.

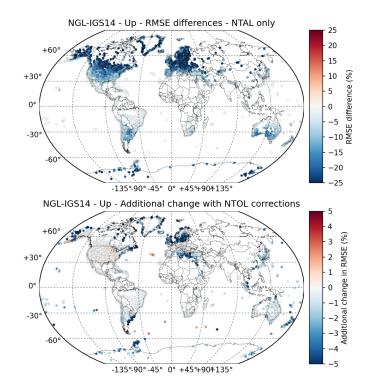


Figure 1. Global distribution of change in RMSE after removing ESMGFZ's NTAL deformation predictions only (top), and additional change in RMSE after also removing NTOL deformation predictions (bottom).

The global distribution of the change in RMSE obtained after removing NTAL de-372 formations, and the additional change resulting from also removing NTOL deformations, 373 are presented in Figure 1. NTAL corrections result in an RMSE reduction for 97.5% of 374 the stations. The global-average change in RMSE is -11.4%, but the differences are more 375 pronounced at latitude over $\pm 45.0^{\circ}$ than near the equator, where the atmospheric pres-376 sure is less variable (Van Dam & Wahr, 1987). NTOL corrections result in an additional 377 RMSE reduction for 72.8% of the stations. Though lower by a factor 5, the impact of 378 NTOL, like that of NTAL, is higher near the poles, but logically confined to coastal ar-379 eas. With an overall reduction of about 30% at high latitudes and about 0% at the equa-380 tor, the high latitude reduction observed here is about 10% better than that reported 381 in Fig. 7 in Mémin et al. (2020), probably because their dataset mostly contains coastal 382 GNSS stations. 383

Our results show that the correction of either NTAL or NTOL deformations results in a quasi-systematic scatter reduction of the aperiodic VLM residual time series. Hence, though this dataset aggregates observation from GNSS stations with variable geodetic quality, the recently reprocessed NGL's position time series are, nonetheless, of particular interest for NTL deformation studies.

Because removing NTAL and NTOL deformations results in RMSE reduction for the wide majority of stations, in the following, we focus on their combined effect, that is, the impact of the non-tidal atmospheric and oceanic loading (NTAOL=NTAL+NTOL) on the spectral, temporal, and spatial properties of the stochastic variability in VLM time series.

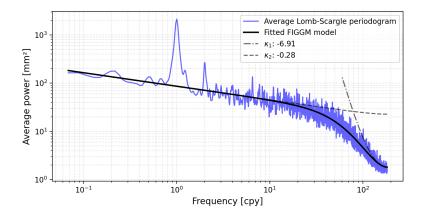


Figure 2. Average Lomb-Scargle periodogram of the ESMGFZ's vertical NTAOL deformation time series (blue solid line). The NTAOL deformation time series have the same sampling and gaps as the GNSS time series. The average power is computed using only time series with a span longer than the considered period. The average power spectrum is approximated by a Fractionally Integrated Generalized Gauss-Markov model (FIGGM, black solid line). The dash dotted black line represents the spectral index $\kappa_1 = -6.91$ at high frequencies. The dashed black line represents the spectral index $\kappa_2 = -0.28$ at low frequencies.

5.2 Average power spectrum

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The improved repeatability observed in the previous section does not indicate whether NTAOL deformations are time-correlated. One way to investigate this effect is to analyze the average power spectrum of the NTAOL deformation time series. In Figure 2, we present the average Lomb-Scargle periodogram of the ESMGFZ's vertical NTAOL deformation time series at the GNSS sites.

The energy of the spectrum is mainly concentrated at low frequencies and rapidly 400 decreases at high frequencies. This indicates that NTAOL deformations are time-correlated. 401 However, the shape of the power-spectrum does not resemble that expected from the com-402 bination of white noise and power-law processes, as it collapses instead of flattening at 403 high frequencies. More precisely, the average power-spectrum shows two distinct slopes 404 above, and below, 20.0 cycles per year [cpy]. Time series presenting such a behavior can 405 be well described by a fractionally integrated generalized Gauss-Markov (FIGGM) pro-406 cess (Bos et al., 2020). 407

The power spectrum P(f) of a discrete FIGGM process reads

$$P(f) = 2\frac{\sigma^2}{f_s} [1 + \phi^2 - 2\phi \cos(2\pi f/f_s)]^{\kappa_1/2} \times [2\sin(\pi f/f_s)]^{\kappa_2},$$
(11)

⁴⁰⁹ in which σ^2 denotes the amplitude of the process, f_s denotes the sampling frequency in ⁴¹⁰ cpy, κ_1 denotes the spectral index at high frequencies, κ_2 denotes that at low frequen-⁴¹¹ cies, and ϕ denotes the Gauss-Markov parameter controlling the transition between both.

⁴¹² The estimated low-frequency spectral index of NTAOL deformation is $\hat{\kappa}_2 = -0.28$ ⁴¹³ whereas, at high frequencies, $\hat{\kappa}_1 = -6.91$. The power-spectrum at high frequencies does ⁴¹⁴ not closely follow the steep $\hat{\kappa}_1$ slope because of the early transition to $\hat{\kappa}_2$, due to a small ⁴¹⁵ Gauss-Markov parameter ($\hat{\phi} = 0.18$) corresponding to a crossover frequency of 78.8 cpy.

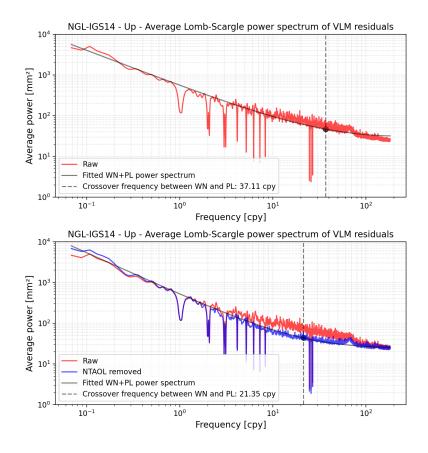


Figure 3. Average Lomb-Scargle periodogram of residual GNSS position time series before (top) and after removing the NTAOL deformations (bottom). The average power is computed using only time series with a span longer than the considered period.

To investigate how the FIGGM nature of NTAOL deformations affects the spectral properties of the stochastic variability in GNSS time series, we computed the average Lomb-Scargle periodogram of VLM residuals before and after removing NTAOL deformations. Unlike in the previous section, hereafter, the VLM residuals are estimated using the inverse of the estimated covariance matrix as the weight matrix, that is, using $\mathbf{W} = \mathbf{Q}_{\mathbf{v}}^{-1}$ in Equation 7. Both periodograms are presented in Figure 3.

In each plot, we observe power drops at the seasonal, draconitic, and fortnightly 422 frequencies accounted for in the trajectory model because the power at these frequen-423 cies is absorbed by the deterministic parameters (see Section 3.2). Besides these frequency-424 specific drops, both power spectra tend to follow that expected from a WN+PL model, 425 that is, a slope close to -1.00 (flicker noise) at low frequencies and a flattening at high 426 frequencies. Thus, the WN+PL stochastic model used for variance component estima-427 tion (section 5.3) seems well-suited to both cases. To illustrate this, in Figure 3, WN+PL 428 power spectra have been adjusted to each average Lomb-Scargle periodogram. The crossover 429 frequency – that is, the frequency where the power resulting from white noise equals that 430 of the power-law process – is also presented in both cases. 431

Before NTAOL correction, although the WN+PL model approximates the overall shape of the periodogram, it fails to describe the increase in power due to NTAOL
deformations between 10 cpy and 70 cpy. On the other hand, after NTAOL correction,
there is a much better agreement between the average periodogram and the WN+PL

model. We also observe a more pronounced flattening at high frequencies, which moves
the estimated crossover frequency between WN and PL from 37.11 cpy to 21.35 cpy. A
wide spectral peak at 66.4 cpy, already reported in Amiri-Simkooei (2013) and Ray et
al. (2013), possibly related to the PPP processing (Selle et al., 2014; Amiri-Simkooei et al., 2017), is also more visible after NTAOL corrections.

The difference in power spectra in Figure 3 differs from that in Fig. 8 in Männel 441 et al. (2019), which shows meager differences at high-frequencies and more pronounced 442 ones at low-frequencies. There are two possible reasons for the differences at high-frequencies. 443 The first one is that we use about 65 times more stations, most of them located in Europe and North America. Thus, by averaging, Figure 3 might better reveal NTAOL sig-445 nals in these areas. Another possibility is the use of the Global Mapping Function (GMF) 446 (Böhm et al., 2006) by Männel et al. (2019), which might result in tropospheric mod-447 eling errors that partially compensate for the atmospheric loading signal (Steigenberger 448 et al., 2009) and reduce the influence of NTAOL corrections on their VLMs time series 449 (Martens et al., 2020). The high-frequency flattening of the power-spectrum after NTAOL 450 correction presented in this work (Figure 3) seems in better agreement with that pre-451 sented in Fig. 8 of Tregoning and Watson (2009). 452

Regarding the low-frequencies, the more significant differences in Männel et al. (2019) are likely due to the correction of hydrology loading, not applied in this work. Also, as the spectral properties of the estimated residuals depend on the estimated stochastic model (through the weight matrix in Equation 7), the change of slope at low-frequencies in Figure 3 is undoubtedly due to the significant changes in spectral indices at mid-latitudes, which we present in the following section.

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5.3 Spatial variability of variance components

⁴⁶⁰ In this section, we present the influence of removing NTAOL deformation on the ⁴⁶¹ spatial variability of the estimated white noise amplitudes $\hat{\sigma}_{hw}$, modified power-law am-⁴⁶² plitudes $\hat{\sigma}'_{pl}$, spectral indices $\hat{\kappa}$, and RMSEs.

The global distribution of the estimated variance components using raw VLM time 463 series are presented in Figure 4. That obtained using VLM time series with NTAOL de-464 formations removed are presented in Figure 5. All stochastic parameters show a strong 465 global scale variability, mostly characterized by a latitude dependence, which Figure 6 evidences and compares before and after removal of NTAOL deformation, showing also 467 the RMSE changes in parallel. To improve the legibility of Figure 6, we represent run-468 ning medians and interquartile ranges over $+10.0^{\circ}$ latitude windows. In the following, 469 the quantitative description of the latitude dependence refers to these running medians 470 and not the individual station estimates. 471

Before removing NTAOL deformations, the estimated white noise amplitudes reach 472 a maximum of about 4.0 mm near the equator and are almost always equal to 0.0 mm 473 for latitudes over $\pm 40.0^{\circ}$, forming an equatorial bulge. After the removal of NTAOL de-474 formations, the spatial variability of the stochastic parameters is strongly modified (Fig-475 ure 5). Instead of being null everywhere but at low latitudes, the white noise amplitude 476 is now non-zero everywhere. This does not indicate that NTAOL corrections introduce 477 additional white noise in the series, but rather that white noise amplitude estimates based 478 on the uncorrected series were biased low, due to NTAOL deformations partially hid-479 ing white noise (see Figure 3). The equatorial bulge in white noise amplitudes remains 480 visible, with a minimum of 2.0 mm near the poles and a maximum of about 4.0 mm at 481 482 the equator.

This bulge was originally pointed out by Williams et al. (2004) but was also more recently observed in IGS station position time series by Klos et al. (2019). Williams et al. (2004) suggested that the origin of this bulge could reside in mis-modeling of the at-

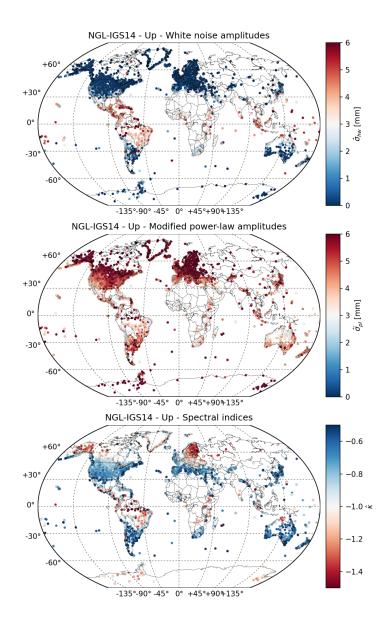


Figure 4. Global distribution of stochastic parameters estimated from raw VLM time series. White noise amplitudes (top). Modified power-law amplitudes (middle). Spectral indices (bottom).

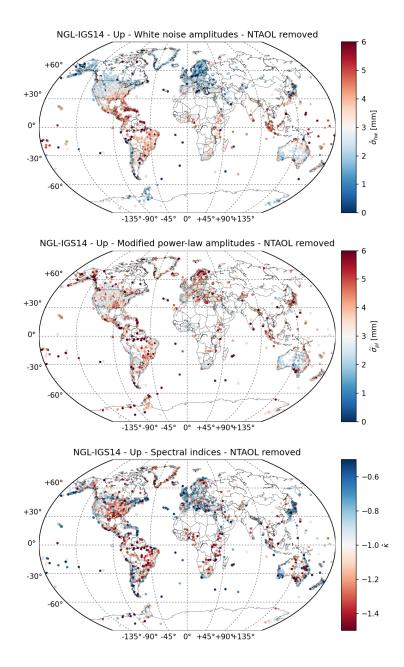


Figure 5. Global distribution of stochastic parameters estimated from VLM time series with NTAOL deformations removed. White noise amplitudes (top). Modified power-law amplitudes (middle). Spectral indices (bottom).

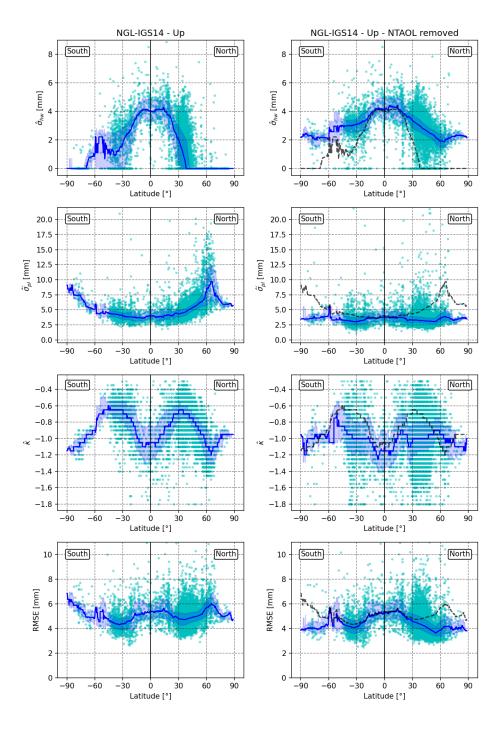


Figure 6. Latitude dependence of white noise amplitudes $\hat{\sigma}_{hw}$ (top), spectral indices $\hat{\kappa}$ (upper middle), power-law amplitudes $\hat{\sigma}'_{pl}$ (lower middle), and RMSEs (bottom), in the presence of NTAOL deformations (left) and with NTAOL deformations removed (right). The dots in cyan represent the station-specific estimates. The continuous blue lines represent running medians over 10° latitude windows. The shaded areas represent the corresponding interquartile ranges. The dashed black lines are the running medians before NTAOL correction, reported in the right plots for comparison with results after NTAOL correction.

mospheric delays affecting GNSS observations. Although the NGL-IGS14 solution used
in this study benefits from recent improvements in the modeling of both tropospheric
and second-order ionospheric effects, the bulge remains visible in both datasets. Hence,
further investigations are needed to identify its causes and to reduce its amplitude.

Unlike white noise amplitudes, the modified power-law amplitudes before remov-490 ing NTAOL deformations reach a minimum amplitude of 3.5 mm near the equator and 491 are maximum at high latitudes, especially near the latitude $+65.0^{\circ}$, where they peak at 492 10.0 mm. The estimated spectral indices show a bi-modal distribution with a maximum 493 of about -0.65 near $\pm 40.0^{\circ}$ and a minimum of about -1.10 near latitudes -90.0° , 0.0° , and $+65.0^{\circ}$. In other words, the estimated stochastic variability shows longer memory 495 effects near the equator and near the poles than at mid-latitudes. After the removal of 496 NTAOL deformations, the modified power-law amplitudes are dramatically reduced at 497 middle and high latitudes, making their latitude dependence almost disappear. The power-498 law amplitudes now uniformly lie in a narrow range centered around 3.5 mm. For spec-499 tral indices, as for the other parameters, the latitude dependence is also affected. After 500 removing NTAOL deformations, the latitudinal medians of the estimated spectral indices 501 no longer show a bi-modal distribution, but a flatter latitude dependence centered on 502 -1.00, that is, flicker noise. In addition, their interquartile range shows an increase at 503 mid-latitudes, indicating that the spatial distribution is more complex than just a lat-504 itude dependence (see Figure 5). 505

The striking latitude dependence of power-law amplitudes and spectral indices $\hat{\sigma}'_{pl}$ 506 visible in Figure 4 was not observed by Williams et al. (2004). This difference is likely 507 due to the processing improvements having occurred over the past 17 years. Such global 508 scale patterns have however been pointed out in Klos et al. (2019, 2020). In Europe, Klos 509 and Bogusz (2017); Gruszczynski et al. (2019) and Klos et al. (2021), suggested that NTL 510 deformations could be the cause of the spatial variability of stochastic parameters ob-511 served in the Up component. The observed impact of NTAOL corrections at the global 512 scale from Figures 5 and 6 is consistent with this hypothesis. 513

Before removing NTAOL deformations, the RMSEs are centered around 5.1 mm and reach maxima of about 6.0 mm near the poles and about 5.6 mm near the equator. After the removal of NTAOL deformations, they now show a slightly subdued latitudinal profile mainly resulting from their significant decrease, from 6.0 mm to 4.0 mm, at high latitudes, while their values at low latitudes stay unchanged.

Note that, at high latitudes, before removing NTAOL deformations, the amplitude 519 of power-law processes alone exceeds the RMSE of the residuals (Figure 6). As the RMSE 520 accounts for the total effect of white noise plus power-law processes, this indicates that 521 there is a bias in the estimation of the power-law parameters. To illustrate how the es-522 timated stochastic parameters change with NTAOL correction, Figure 7 presents the Lomb-523 Scargle periodogram of the VLM residuals along with the WN+PL power spectrum ex-524 pected from the estimated stochastic parameters for station ONSA (Onsala, Sweden), 525 both before and after NTAOL correction. 526

Before NTAOL correction, the estimated power spectrum differs from the Lomb-527 Scargle periodogram of the residuals because of the non-modelled variability at sub-annual 528 periods resulting from the presence of NTAOL deformations (Figure 2). This illustrates 529 that, at high latitudes, where the scatter of the NTAOL deformation is the largest (Tregoning 530 & van Dam, 2005), the usual WN+PL stochastic model is not appropriate and the as-531 sociated parameter estimates are not reliable. However, after NTAOL correction, the non-532 modelled sub-annual variability is removed, the stochastic variability resembles more that 533 of the WN+PL model, which results in more interpretable parameter estimates. 534

In principle, considering a WN+PL+FIGGM stochastic model instead of a WN+PL one could also be a solution to this problem. However, adding a FIGGM process to the

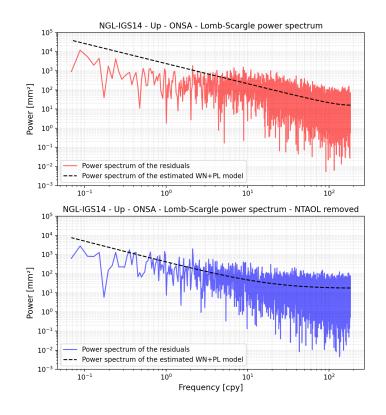


Figure 7. Power spectra of the VLM time series of station ONSA (Onsala, Sweden), before (top) and after (bottom) removing NTAOL deformations. The colored lines are the Lomb-Scargle periodograms of the VLM residuals. The black dashed lines are the WN+PL power spectra expected from the estimated stochastic parameters.

stochastic model would lead to a more complex and nonlinear variance component es timation problem. Correct for the NTAOL deformations before stochastic parameter es timation is a more efficient and pragmatic solution.

Figures 4, 5 and 6 show that all stochastic parameters show a spatial variability, both before and after NTAOL correction. Hence, considering spatially variable stochastic process amplitudes and spectral indices is preferable when analyzing VLM time series. In particular, the assumption of a constant spectral index and of a constant powerlaw to white noise amplitude ratio for all stations, as done in Amiri-Simkooei (2013); Amiri-Simkooei et al. (2017), seems sub-optimal in global solutions.

Although we demonstrated that a large part of the spatial variability of stochastic parameters can be attributed to the influence of NTAOL deformations, Figure 5 clearly shows that non-latitudinal patterns remain. This residual variability could be further reduced by improving the accuracy of NTAOL prediction or by accounting also for hydrological loading deformations. Once the strong impact of NTAOL deformation is removed, detecting the influence of the monumentation quality on the stochastic parameters of individual stations in global solutions might become easier (Langbein & Svarc, 2019).

5.4 Velocity uncertainties

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⁵⁵⁵ Velocity uncertainties depend on the estimated stochastic properties, but also on ⁵⁵⁶ each station's functional model and span. To isolate the influence of the stochastic prop-⁵⁵⁷ erties, for a given stochastic model, we define the minimum 8-year velocity uncertainty, ⁵⁵⁸ noted σ_v^* , as the expected standard deviation of the velocity estimated assuming a con-⁵⁵⁹ tinuous 8-year long time series with a trajectory model consisting of a linear trend and ⁵⁶⁰ the periodic signals presented in section 3.2. σ_v^* is a lower bound estimate because miss-⁵⁶¹ ing values and additional deterministic parameters would increase the velocity uncertainty.

Figure 8 shows the global distribution of σ_v^* obtained from the stochastic parameters estimated before and after NTAOL correction. Figure 9 depicts their latitudinal distributions along with the running medians and the interquartile ranges over $+10.0^{\circ}$ latitude windows. In the Supporting Information, Figure S2 presents an alternative to Figure 8, in which the velocity uncertainties account for the full span of each VLM time series.

Before removing NTAOL deformations σ_v^* is maximum at high latitudes, namely, in Alaska, Canada, around the Baltic Sea, Russia, Antarctica, but also within the Amazon basin. More precisely, the median σ_v^* reaches a maximum of 1.42 mm/yr around the latitude +65.0°, and a minimum of about 0.24 mm/yr near ±30° latitudes.

After NTAOL correction, the latitude dependence of σ_v^* is dramatically reduced. 572 The median σ_v^* are now centered around 0.37 mm/yr at all latitudes, with a maximum 573 of 0.55 mm/yr near the equator and a minimum around 0.26 mm/yr near $\pm 45^{\circ}$ latitudes. 574 The most dramatic reduction of σ_v^* is observed around the latitude +65.0°, where σ_v^* drops 575 from 1.42 mm/yr to 0.43 mm/yr, that is, a 69.7% decrease. However, the reduction in 576 uncertainty is not systematic and concerns only 49.6% of the stations. The other 50.4%577 show a moderate increase in σ_v^* (Figure 9) because of the variance component estima-578 tion bias caused by uncorrected NTAOL deformations, which results in overestimated 579 spectral indices at mid-latitudes (e.g., Figure 6). 580

The other stations, mostly localized at mid-latitudes, show a moderate increase in uncertainty. An apparent increase in outliers is also observed at mid-latitudes, mainly caused by groups of stations near the northern coasts of the Gulf of Mexico, the south of Brazil, and Central Africa showing both strong RMSEs and low spectral indices. But overall, after NTAOL correction, most stations show more homogeneous σ_v^* estimates

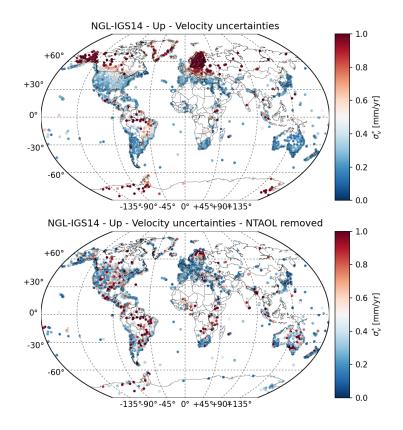


Figure 8. Global variability of the minimum velocity uncertainty for 8-year long time series before removing NTAOL deformations (top) and after removing NTAOL deformations (bottom).

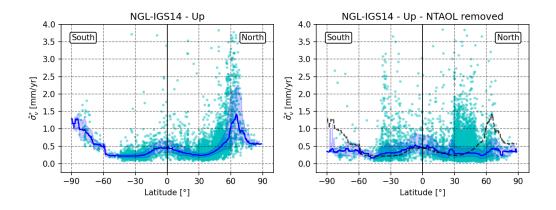


Figure 9. Latitude dependence of the minimum velocity uncertainty for 8-year long time series before (left) and after (right) removing NTAOL deformations. The dots in cyan represent the station-specific estimates. The continuous blue lines represent running medians over 10° latitude windows. The shaded areas represent the corresponding interquartile ranges. The dashed black lines are the running medians before NTAOL correction, reported on the right plot for comparison with results after NTAOL correction.

Dataset	$\hat{\bar{ ho}}_c$	$\hat{\bar{ ho}}_0$	$\hat{\bar{\rho}}_{\max}$	$\hat{\bar{r}}$ [km]
Raw	0.01	0.40	0.41	2470
NTAOL removed	0.02	0.27	0.29	2485

Table 1. Parameter estimates obtained by fitting the Gaussian correlation model in Equation12 to the average spatial correlation estimates.

and the observed change in σ_v^* mostly highlights the incapacity of the WN+PL model to retrieve realistic stochastic parameter estimates, hence realistic velocity uncertainties, in the presence of NTAOL deformations.

We note that the dramatic change in velocity uncertainty resulting from NTAOL corrections at high latitudes does not match the minor shift in uncertainty predicted in Santamaría-Gómez and Mémin (2015) from NTAL deformation alone. This difference is because the inadequacy of the WN+PL model at high latitudes, and the resulting variance component estimation bias, cannot be observed from the analysis of NTAL deformation time series alone.

Given the dramatic reduction in velocity uncertainty observed at high latitudes, correcting for NTAOL deformation before the estimation of stochastic parameters in GNSS VLM time series will allow a better detection of subtle VLM signals related to glacial isostatic adjustment, present-day ice-melting, and improve sea-level change estimates in these areas.

5.5 Spatial correlation

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For both the raw and the NTAOL-corrected VLM time series, we computed aver-601 aged spatial correlation coefficients $\hat{\rho}(t_k, d_l)$ for 200 classes of separation distances d_l rang-602 ing from [0, 100] km to [19900, 20000] km and for 240 epochs t_k between 2008 and 2018. 603 Correlation coefficients between all possible pairs of stations are computed over time win-604 dows of 30 days and averaged within each distance class. The resulting time-variable spa-605 tial correlation estimates $\hat{\rho}(t_k, d_l)$ are presented, for both cases, in Figure 10. To get a 606 clearer picture of the spatial pattern, the spatial correlation averaged over the 2008-2018 607 period and denoted $\bar{\rho}(d_l)$ is also presented in Figure 10. 608

Prior to removing NTAOL deformations, the average spatial correlation $\bar{\rho}(d_l)$ is close to +0.45 for short separation distances (that is, $d_l < 100$ km) and decreases close to zero for larger separation distances (that is, $d_l \geq 5000$ km). To model this behavior, we approximated the average correlogram $\bar{\rho}(d_l)$ by a Gaussian correlation function defined by

$$\bar{\rho}(d_l) = \begin{cases} 1, & d_l = 0\\ \bar{\rho}_c + \bar{\rho}_0 \cdot \exp\left(-\left(d_l/\bar{r}\right)^2\right), & d_l > 0 \end{cases}$$
(12)

where $\bar{\rho}_c$ denotes the distance-independent correlation, $\bar{\rho}_0$ the distance-variable correlation, and \bar{r} the distance range parameter. According to this Gaussian correlation function, the maximum correlation for $d_l > 0$ is $\bar{\rho}_{max} = \bar{\rho}_c + \bar{\rho}_0$.

The adjusted parameters of the Gaussian correlation function, with and without NTAOL deformations, are presented in Table 1. The graphs corresponding to the adjusted correlation functions are depicted in Figure 10. Although, for the raw VLM resid-

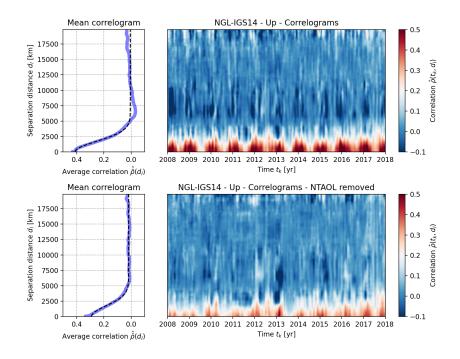


Figure 10. Time-variable spatial correlation estimates $\hat{\rho}(t_k, d_l)$ before (top) and after (bottom) removing NTAOL deformations. The blue curves in the left plots represent the average correlograms over the 2008-2018 period. The dashed black lines represent the adjusted Gaussian correlation models (Equation 12).

⁶²⁰ ual time series, small departures from the observations, inferior to 0.05, are observed for ⁶²¹ $d_l \geq 5000$ km, the Gaussian correlation function is overall a good approximation of the ⁶²² correlation curve for both datasets. After the removal of NTAOL deformations, Table ⁶²³ 1 shows that the maximum of correlation for $d_l > 0.0$ km changed from +0.41 to +0.29, ⁶²⁴ that is, a reduction of 29.27%. The average range parameter \bar{r} is hardly affected by the ⁶²⁵ NTAOL corrections.

Our results agree with Williams et al. (2004); Amiri-Simkooei et al. (2017) and Benoist et al. (2020), evidencing a spatial correlation of the VLM residual time series. However, our spatial correlations are slightly smaller, most certainly because we analyze over ten thousand recently reprocessed position time series instead of a few hundred (at most).

In addition to the approximately Gaussian average spatial correlation, the time-630 variable correlograms in Figure 10 reveal an annual oscillation for short separation dis-631 tances, that is $d_l < 4000$ km, most visible when NTAOL deformations are not removed. 632 The spatial correlation time series $\hat{\rho}(t_k, d_l)$ for separation distances of 50 km, 950 km, 633 1950 km, and 3950 km are presented in Figure 11 to highlight this seasonality. The am-634 plitude of the annual oscillation is maximal for short distances ($d_l < 100$ km) and syn-635 chronous for all distance classes. Also, both the mean correlation and the annual am-636 plitude get closer to 0.0 for increasing separation distances. To describe the oscillation 637 of $\hat{\rho}(t_k, d_l)$, we added an annual modulation to the distance-variable part of the mean 638 Gaussian correlation function of Equation 12, that is 639

$$\rho(t_k, d_l) = \begin{cases} 1, & d_l = 0\\ \rho_c + [\rho_0 + \rho_A \cos(\omega_A t_k - \phi_A)] \cdot \exp\left(-\left(d_l/\bar{r}\right)^2\right), & d_l > 0 \end{cases} \tag{13}$$

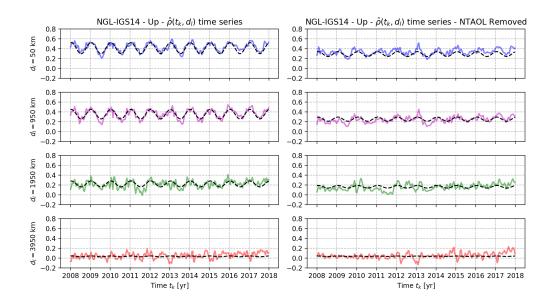


Figure 11. Time-variable spatial correlation estimates $\hat{\rho}(t_k, d_l)$ for 4 classes of distances d_l before (left) and after (right) removing NTAOL deformations. The colored lines represent the estimates for each class of distance d_l . The dashed black lines represent the corresponding adjusted time-variable Gaussian correlation function (Equation 13).

Dataset	$\hat{ ho}_c$	$\hat{ ho}_0$	$\hat{ ho}_A$	$\hat{\phi}_A$ [day]	$\hat{ ho}_{min}$	$\hat{ ho}_{max}$	\hat{r} [km]
Raw	0.01	0.40	0.11	13.8	0.30	0.52	2469
NTAOL removed	0.02	0.27	0.05	22.0	0.24	0.34	2485

 Table 2.
 Parameter estimates obtained by fitting the seasonal Gaussian spatial correlation

 model in Equation 13 to the time-variable spatial correlation estimates.

where ω_A is the annual angular frequency, ρ_A is the annual amplitude, and ϕ_A is the corresponding phase. From these estimates, one can define a minimum and a maximum correlation for $d_l \approx 0$ as $\rho_{\min} = \rho_c + \rho_0 - \rho_A$ and $\rho_{\max} = \rho_c + \rho_0 + \rho_A$.

The adjusted parameters of the time-variable correlation model $\hat{\rho}_c$, $\hat{\rho}_0$, $\hat{\rho}_A$, ϕ_A , ρ_{\min} and ρ_{\max} , estimated with and without NTAOL correction, are presented in Table 2. The corresponding graphs are visible in Figure 11.

Prior to NTAOL correction, the spatial correlation at short distances shows an an-646 nual amplitude of $\hat{\rho}_A = 0.11$ and oscillates between +0.30 and +0.52. According to the 647 estimated phase ϕ_A , the minimum of correlation occurs in July and the maximum in Jan-648 uary. Since most of the stations analyzed are located in the United States and Europe, 649 the estimated amplitude and phase of this periodicity are certainly more representative 650 of the stochastic variability in these areas but less surely in other parts of the world. Af-651 ter removing NTAOL deformations, we observe a change in annual amplitude, from 0.11 652 to +0.05, that is, a 54.55% reduction. The spatial correlation at short distances then os-653 cillates between +0.24 and +0.34. The estimated phase ϕ_A remains similar, and the max-654 imum still occurs in January. 655

According to the observed reductions, NTAOL deformation is a significant cause
 of seasonal spatial correlation in GNSS VLM time series. These results support that of
 Gruszczynski et al. (2019), Männel et al. (2019), and Kreemer and Blewitt (2021), who
 pointed out that common mode errors in Europe could be due to NTL deformations.

Note that after the NTAOL correction, the spatial correlation at separation distances shorter than 100 km still oscillates between +0.24 in July and +0.34 in January (e.g., Figure 11 and Table 2). This is expected, as hydrological loading, not corrected in this work, also result in spatially correlated deformation (Van Dam et al., 2001; Tregoning et al., 2009; Chanard et al., 2018). Besides, it is also possible that improving the quality of NTAOL deformation predictions for instance by computing the dynamic ocean response to changes in atmospheric pressure, as done by Mémin et al. (2020), could further reduce the stochastic variability in VLM time series.

Nonetheless, considering that orbit and atmospheric positioning errors are other
known sources of spatially correlated stochastic variability (Amiri-Simkooei et al., 2017),
it is likely that physical models cannot fully reproduce the spatial correlation in GNSS
position time series. Hence, further investigating how the remaining spatial correlation
and its seasonality affect the covariance between GNSS station velocities, similarly as
Amiri-Simkooei (2009); Razeghi et al. (2016) or Benoist et al. (2020), will undoubtedly
benefit the study of small deformation signals in global solutions.

675 6 Conclusion

Characterizing and modeling the stochastic variability in GNSS position time series is essential to reduce the uncertainty on the estimated station velocities. In this study,
we investigated how aperiodic non-tidal atmospheric and oceanic loading (NTAOL) deformations influence the stochastic properties of vertical land motion (VLM) time series. To do so, we used 10,151 time series processed by the Nevada Geodetic Laboratory
and NTAOL deformation time series provided the Earth System Modelling team at GFZ
Potsdam.

We first showed that NTAOL deformations exhibit time correlation describable by 683 a fractionally integrated Gauss-Markov process. Because this behavior cannot be described 684 by the commonly used stochastic model consisting of linear combination of white noise 685 and power-law process, the presence of NTAOL deformations in uncorrected VLM time 686 series results in biases in white noise amplitude, power-law amplitude, and spectral in-687 dex estimates. In particular, NTAOL deformations cause the power-law parameters to 688 display a strong latitude dependence, resulting in a dramatic increase in velocity uncer-689 tainty, exceeding 1.4 mm/yr for 8-year long time series, at high latitudes. 690

Once NTAOL deformations are removed from the VLM time series, we observe a 691 drastic reduction of the latitude dependence of all stochastic parameters. This change 692 is followed by a significant reduction of velocity uncertainty reaching about 70% in North-693 ern America, Greenland, Fennoscandia, and Antarctica. Therefore, to avoid coping with 694 unrealistically high velocity uncertainties, we advise correcting for NTAOL deformation 695 prior to the estimation of stochastic parameters in GNSS VLM time series. Doing so will 696 certainly allow a better detection of subtle VLM signals related to glacial isostatic ad-697 justment, as well as present-day ice-melting, and improve sea-level change estimates at 698 high latitudes. 699

In addition to time-correlation, we investigated how NTAOL deformations influence the spatial correlation of aperiodic VLM residuals. First, we demonstrated that, before removing NTAOL deformations, VLM residuals are spatially correlated and that this spatial correlation shows a Gaussian decay with the separation distance, along with a strong seasonality. After removal of NTAOL deformations, the average correlation is reduced by 29.3% and its seasonality by 54.4%, showing that NTAOL deformations are
 an important source of spatial correlation.

Although NTAOL deformation plays a major role in the stochastic variability in
 the VLM time series, it does not fully correct for its power-law behavior or its spatial
 correlation. Therefore, further investigations are still needed to identify the physical and
 artificial components of the remaining stochastic variability in VLM time series.

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