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## Surface Phonon-Polariton Heat Capacity of Polar Nanofilms

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The surface phonon-polariton heat capacity of polar nanofilms is analytically determined and analyzed as a function of their thickness and temperature. It is shown that the polariton heat capacity increases with the square of the nanofilm thickness, such that for thin enough nanofilms at sufficient low temperature  $T$ , its value becomes independent of the material properties and is given by  $C = 1.15\epsilon_0 k_B (k_B T / \hbar c)^2$ , where  $\epsilon_0$  and  $c$  are the respective relative permittivity of the surrounding medium and light speed in vacuum, while  $k_B$  and  $2\pi\hbar$  are the Boltzmann and Planck constants, respectively. The photonlike nature of surface phonon polaritons is found to be responsible for their main contribution to the material heat capacity. As a result of the polariton speed close to  $c$  and hence much higher than that of phonons, the polariton heat capacity of SiO<sub>2</sub>, SiC, and SiN is found to be several orders of magnitude smaller than its corresponding phonon counterpart. Surface phonon polaritons are thus not expected to increase the expansion coefficient of polar nanofilms, which favors their utilization as effective heat dissipators.

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### I. INTRODUCTION

Surface phonon polaritons (SPHPs) are surface electromagnetic waves generated by phonon-photon coupling at the interface of polar materials, such as SiO<sub>2</sub>, SiC, and SiN [1–13]. This coupling blurs the boundary between light and heat, occurs at midinfrared frequencies mainly, and triggers the propagation of SPHPs over distances (approximately 1 cm) orders of magnitude longer than the typical mean free paths of phonons and electrons [14–16]. These long-range surface waves have thus emerged as powerful heat carriers able to enhance the in-plane thermal conductivity of polar materials with high surface-to-volume ratios, as is the case of nanofilms and nanostructures with usually low thermal performances [17,18].

The SPHP contribution to the thermal conductivity of polar nanofilms was modeled and quantified by Chen *et al.* [1], who reported a thermal conductivity of  $4 \text{ W m}^{-1} \text{ K}^{-1}$  for a 40-nm-thick SiO<sub>2</sub> film suspended in vacuum at 500 K. This prediction of the kinetic theory is more than twice the intrinsic phonon thermal conductivity of SiO<sub>2</sub>, was confirmed by means of the fluctuation-dissipation approach [4], and extended for films deposited over a substrate [2], other polar materials [19], layered systems [20–22], and arrangement of particles [23–25]. On the experimental side, on the other hand, Tranchant *et al.* [26]

measured the in-plane thermal conductivity of suspended SiO<sub>2</sub> nanofilms and showed up its clear increase as the film thickness reduces through values smaller than 50 nm. More recently, by measuring the temperature evolution of the in-plane thermal conductivity of SiN nanofilms, Wu *et al.* [5] found that its values doubles up as the temperature raises from 300 to 800 K. This thermal conductivity enhancement for thinner and hotter polar nanofilms represents the fingerprints of the SPHP heat transport [1,2] and hence its experimental observation [5,26] provides decisive evidence for the potential application of SPHPs as long-distance heat dissipators.

Heat conduction under dynamical conditions is not only driven by the thermal conductivity but also by the thermal diffusivity and heat capacity, which is proportional to the expansion coefficient that affects the structural stability of heat dissipation [27]. Experimental measurements showed that the thermal diffusivity of polar nanofilms increases as their thicknesses decrease [6,26]. This thickness dependence is similar to the one of the thermal conductivity of polar nanofilms [5,26], which suggests that their heat capacity (which equals thermal conductivity divided by thermal diffusivity) could be independent or weakly dependent on the film thickness. While the phonon and electron contributions to the heat capacity are well detailed in textbooks [27,28], the one of SPHPs has not been studied yet.

In this work, the SPHP heat capacity of polar nanofilms is analytically modeled, quantified, and analyzed as a function of their thickness and temperature. It is shown that the SPHP heat capacity is independent of the material

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properties and proportional to the square of temperature, for thin enough nanofilms at sufficiently low temperatures.

## II. THEORETICAL MODEL

Let us consider a polar film of thickness  $d$ , length  $a$ , and width  $b$  supporting the propagation of SPHPs, as shown in Fig. 1. These SPHPs can be generated by thermal or optical means [29–33] and their propagation is strongly determined by the permittivity  $\varepsilon$  and thickness  $d$  of the film [1,2]. Given that the generation of SPHPs is driven by the phonon-photon coupling that depends on the external excitation, their propagation and energy transport are determined by the type of source that induces them. While the optical excitation consists in illuminating the material surface with a monochromatic light beam of a given frequency and hence allows exciting a single SPHP mode at a time [9,29], the thermal one is based on heating the material and simultaneously generating SPHPs in a wide frequency spectrum [4,34], which is required to enhance their heat transport [1,2]. In the present work, we assume that the SPHPs are thermally excited to determine the SPHP heat capacity from a broad frequency range supporting their propagation, as shown below.

Taking into account that the SPHPs can be treated like a gas of bosonic particles with energy  $\hbar\omega$ , as is the case of photons and phonons, their contribution to the internal thermal energy per unit area of the film at temperature  $T$  is given by [27]

$$U = \int \hbar\omega f(T, \omega) D(\omega) d\omega, \quad (1)$$

where  $\hbar$  is Planck's constant divided by  $2\pi$ ,  $f = [\exp(\hbar\omega/k_B T) - 1]^{-1}$  is the Bose-Einstein distribution function, and  $D(\omega)$  is the SPHP density of states per unit frequency  $\omega$ . Since SPHPs mainly propagate along the interface of the film and they span over its surface, their density of states is two dimensional (2D) and is given by [1,34]  $D(\omega) = \beta_R/(2\pi V)$ , with  $\beta_R = \text{Re}(\beta)$  and  $V = \partial\omega/\partial\beta_R$  being the real part of the in-plane wave vector  $\beta$

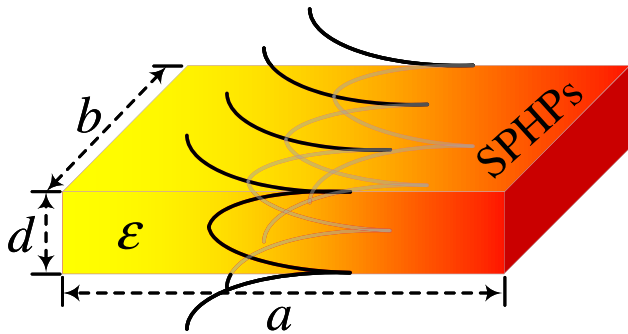


FIG. 1. Scheme of a polar nanofilm supporting the propagation of surface phonon polaritons.

and group velocity of SPHPs, respectively. The SPHP contribution to the heat capacity per unit surface of the film can then be determined by  $C = \partial U/\partial T$  [27,28], which in combination with Eq. (1) yields

$$C = \frac{1}{2\pi} \int \frac{\hbar\omega\beta_R}{V} \frac{\partial f}{\partial T} d\omega. \quad (2)$$

In writing Eq. (2), we assume that the wave vector  $\beta_R = \beta_R(\omega, \varepsilon, d)$  is independent of temperature. This assumption is valid within a wide interval of temperatures below and above room temperature, as is the case of the permittivity  $\varepsilon$  of SiO<sub>2</sub>, SiC, and SiN [5,35]. Equation (2) thus establishes that  $C$  depends on the material properties through the ratio  $\beta_R/V$ , which can be calculated from the dispersion relation of SPHPs propagating along the polar film, as shown in Fig. 1. When the film thickness  $d$  is small enough ( $\lambda = k_0 d/2 \ll 1$ ), which is the case of interest for enhancing their thermal energy transport [1,2,5,26], one can show that the real and imaginary parts of the SPHP wave vector  $\beta = \beta_R + i\beta_I$  are well approximated by [2,5]

$$\beta_R = k_0 \sqrt{\varepsilon_0} (1 + A\lambda^2), \quad (3a)$$

$$\beta_I = k_0 \sqrt{\varepsilon_0} \frac{\varepsilon_0^2 \text{Im}[\varepsilon(\omega)]}{|\varepsilon(\omega)|^2} J\lambda^2, \quad (3b)$$

where  $\varepsilon_0$  is the relative permittivity of the medium surrounding the film,  $k_0 = \omega/c$ ,  $c$  is the speed of light in vacuum, and

$$A = \frac{\varepsilon_0}{2} \left( 1 - \frac{2\varepsilon_0 \text{Re}[\varepsilon(\omega)]}{|\varepsilon(\omega)|^2} + \frac{\varepsilon_0^2 \{\text{Re}[\varepsilon(\omega)]^2 - \text{Im}[\varepsilon(\omega)]^2\}}{|\varepsilon(\omega)|^4} \right), \quad (4a)$$

$$J = 1 - \frac{\varepsilon_0 \text{Re}[\varepsilon(\omega)]}{|\varepsilon(\omega)|^2}. \quad (4b)$$

Taking into account that the SPHP propagation length is given by [2,16]  $\Lambda = (2\beta_I)^{-1}$ , the key function  $J(\omega)$  establishes that the SPHPs propagate only for frequencies that satisfy the condition  $J > 0$ . The positive values of  $J(\omega)$  also determine the confinement of SPHPs to the film interface [5], and therefore they can be used to quantify the frequency spectrum ( $J > 0$ ) supporting the existence and propagation of SPHPs in different polar nanofilms. Based on Eq. (3a), one can show that  $\beta_R/V = (k_0 \varepsilon_0/c) [1 + (4A + \omega \partial A/\partial \omega) \lambda^2]$ , which allows rewriting Eq. (2) as follows:

$$C = C_0 + C_2 d^2, \quad (5)$$

where the coefficients  $C_0$  and  $C_2$  are defined by

$$C_0 = \frac{\varepsilon_0}{2\pi c^2} \int \hbar\omega^2 \frac{\partial f}{\partial T} d\omega, \quad (6a)$$

$$C_2 = \frac{\varepsilon_0}{8\pi c^4} \int \hbar\omega^4 \left( 4A + \omega \frac{\partial A}{\partial \omega} \right) \frac{\partial f}{\partial T} d\omega. \quad (6b)$$

Equation (5) thus establishes that the SPHP heat capacity generally increases with the square of the film thickness, within the thin-film limit ( $k_0 d/2 \ll 1$ ). In practice, this condition is typically fulfilled for  $d < 300$  nm [2,3], which indicates that  $C$  can become independent of the properties of films with thicknesses of tens of nanometers, as is established by Eqs. (5) and (6a) ( $C \rightarrow C_0$ ). Note that both parameters  $C_0$  and  $C_2$ , and hence  $C$ , increase with the permittivity  $\varepsilon_0$  of the surrounding medium. This fact indicates that  $C$  exhibits an opposite dependence on  $d$  and  $\varepsilon_0$  than its SPHP thermal conductivity, which is lower for greater values of  $d$  and  $\varepsilon_0$  [2]. According to its definition ( $C = \partial U/\partial T$ ), the SPHP heat capacity per unit area  $C$  ( $\text{J m}^{-2} \text{K}^{-1}$ ) is the amount of heat required to raise the temperature of the film material with an arbitrary thickness by 1 K, and therefore its increase with  $d$  is reasonable, as established by Eq. (5). By contrast, the corresponding SPHP specific heat capacity  $c = abC/\text{mass} = abC/(\rho abd) = C/(\rho d) = (C_0 d + C_2 d)\rho$ , with  $\rho$  being the film density, takes higher values for thinner films, as is the case of the SPHP thermal conductivity. The surface nature of SPHPs can thus be seen in both its thermal conductivity and specific heat capacity (intensity property), whose values are higher for polar materials with higher surface-to-volume ratios.

In order to show up the temperature behavior of the fundamental component  $C_0$ , Eq. (6a) is rewritten as follows:

$$C_0 = \frac{\varepsilon_0 k_B}{2\pi} \left( \frac{k_B T}{\hbar c} \right)^2 \int_0^{\theta/T} \frac{x^3 e^x}{(e^x - 1)^2} dx, \quad (7)$$

where  $x = \hbar\omega/k_B T$  and  $\theta = \hbar\omega_{\text{max}}/k_B$  is a characteristic temperature determined by the maximum frequency  $\omega_{\text{max}}$  supporting the propagation of SPHPs. The integrand in Eq. (7) establishes that the heat capacity spectrum of a polar nanofilm at temperature  $T$  peaks at the frequency  $\omega_p$  given by  $\hbar\omega_p/k_B T = 2.5757$ , that is to say,  $\omega_p/2\pi = 0.05367T$  ( $\text{THz K}^{-1}$ ), which takes the value of 16.1 THz at room temperature. Note that Eq. (7) is analogous to the Debye model for the phonon heat capacity, with  $\theta$  being the Debye temperature for SPHPs. At high temperature ( $T \gg \theta$ ), the integral in Eq. (7) reduces to  $0.5(\theta/T)^2$  and  $C_0$  takes its maximum value

$$C_{0,\text{max}} = \frac{\varepsilon_0 k_B}{4\pi} \left( \frac{k_B \theta}{\hbar c} \right)^2 = \frac{\varepsilon_0 k_B}{4\pi} \left( \frac{\omega_{\text{max}}}{c} \right)^2. \quad (8)$$

The dependence of  $C_{0,\text{max}}$  on  $\omega_{\text{max}}^2$  is the result of the 2D density of states of SPHPs, which is analog to the dependence on  $\omega_{\text{max}}^3$  of the high-temperature heat capacity of phonons, due to their three-dimensional (3D) density of

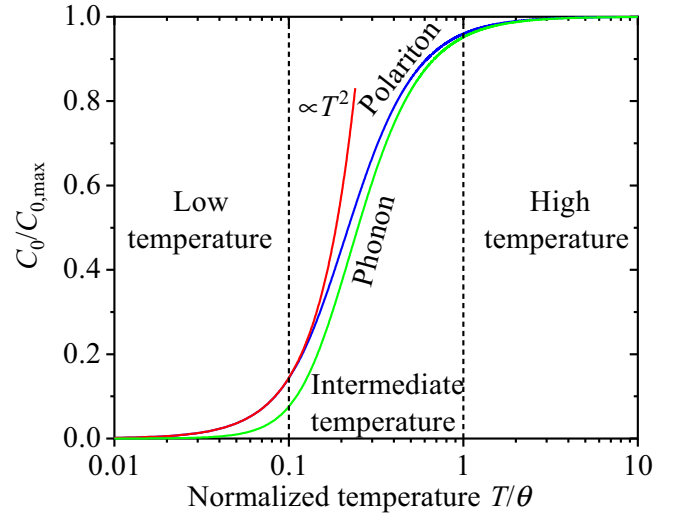


FIG. 2. Fundamental component of the SPHP heat capacity of a polar nanofilm as a function of its temperature. The red line represents the low-temperature limit predicted by Eq. (9) and the green one stands for the prediction of the Debye model for phonons.

states, as predicted by the Debye model [27,28]. For low temperatures ( $T \ll \theta$ ), on the other hand, the integral in Eq. (7) has the value  $6z(3)$  of 6 times the Riemann  $\zeta$  function  $z(3)$  and  $C_0$  increases with the square of temperature, as follows:

$$C_0 = \frac{3z(3)}{\pi} \varepsilon_0 k_B \left( \frac{k_B T}{\hbar c} \right)^2. \quad (9)$$

Equations (7)–(9) thus indicate that the SPHP heat capacity of a polar nanofilm is strongly determined by the relative values of its temperature with respect to the characteristic one  $\theta$ , as shown in Fig. 2. Note that the normalized SPHP heat capacity follows pretty much the same behavior as its phonon counterpart [27,28], through the three regimes determined by the ratio  $T/\theta$ . The  $T^2$  regime in which  $C_0$  is independent of the material properties, shows up for  $T/\theta \leq 0.1$ , while  $C_0$  reaches its maximum value for  $T/\theta > 1$ .

### III. RESULTS AND DISCUSSION

According to Eqs. (3b) and (4b), the range of frequencies supporting the propagation of SPHPs is determined by the condition ( $J > 0$ )

$$\text{Re}[\varepsilon(\omega)] < \frac{|\varepsilon(\omega)|^2}{\varepsilon_0}, \quad (10)$$

which can be satisfied not only by negative values of  $\text{Re}(\varepsilon)$ , but also by positive ones. The key condition in Eq. (10) thus indicates that the SPHP propagation along the

interface of polar nanofilms is totally determined by its permittivity ( $\varepsilon$ ) and that ( $\varepsilon_0$ ) of its surrounding medium, while being independent of its thickness. Note that the range of allowed frequencies reduces to those within the negative peak of  $\text{Re}[\varepsilon(\omega)] < 0$  (Reststrahlen band) for  $\varepsilon_0 \rightarrow \infty$  only. For any other case of practical interest, the range of frequencies supporting the propagation of SPHPs is usually much broader than the Reststrahlen band, as was reported experimentally [14,15] and is predicted by the function  $J > 0$ . This frequency range is independent of the film thickness [Eq. (10)], provided that the thin-film limit ( $\lambda \ll 1$ ) is valid, which ensures a strong coupling of the SPHPs propagating along the upper and lower interfaces of the film. In fact, the SPHP modes propagating along lossy nanofilms, as is the case of our present work, are called Zenneck modes, which exist and propagate with frequencies inside and outside the Reststrahlen band, as detailed in the literature [3,29] and are shown in Fig. 4. Furthermore, in absence of energy absorption [ $\text{Im}(\varepsilon) = 0$ ], Eq. (10) reduces to  $\text{Re}(\varepsilon) = \varepsilon > \varepsilon_0$ , which is totally different than the well-known condition  $\varepsilon < -\varepsilon_0$  determining the propagation of SPHPs along a single interface of a bulk lossless material.

As the function  $J$  changes of sign at  $\text{Re}[\varepsilon(\omega)] = |\varepsilon(\omega)|^2 / \varepsilon_0$ , the cutoff frequency  $\omega = \omega_{\text{max}}$  is given by the solution of this latter relation, which can be rewritten as the circle equation  $\{\text{Re}[\varepsilon(\omega)] - \varepsilon_0/2\}^2 + \{\text{Im}[\varepsilon(\omega)]\}^2 = (\varepsilon_0/2)^2$ . This circle of radius  $\varepsilon_0/2$  represents the transition between the range of frequencies that allow (outside the circle) or forbid (inside the circle) the propagation of SPHPs. Therefore, if there is more than one value for  $\omega_{\text{max}}$ , the integrals in Eqs. (6a) and (6b) have to be integrated skipping the frequency band gaps in which SPHPs do not propagate ( $J < 0$ ).

The frequency dependence of the real  $\varepsilon_R(\omega)$  and imaginary  $\varepsilon_I(\omega)$  parts of the relative permittivity of  $\text{SiO}_2$ ,  $\text{SiC}$ , and  $\text{SiN}$  are shown in Figs. 3(a)–3(c), respectively [2,3,5]. Note that  $\text{SiO}_2$  has two Reststrahlen bands [ $\varepsilon_R(\omega) < 0$ ] determined by the frequency intervals (13.8; 15.1) THz and (32.1; 37.3) THz. On the other hand,  $\text{SiC}$  and  $\text{SiN}$  have a single Reststrahlen band each, within the frequency ranges (23.7; 28.9) THz and (26.6; 31.8) THz, respectively. For the three considered materials, these frequency bands contain the resonance peaks of  $\varepsilon_R(\omega)$  and occur in the vicinity of the maxima of  $\varepsilon_I(\omega)$ . This fact indicates that the materials absorb the highest amount of energy from the electromagnetic field at frequencies near the ones defining their corresponding Reststrahlen bands. In these bands, the function  $J$  (Fig. 4) is positive and therefore they support the existence and propagation of SPHPs. Note that both  $\text{SiO}_2$  and  $\text{SiC}$  exhibit frequency band gaps (colored areas) just above their corresponding Reststrahlen bands. By contrast, these band gaps are absent for the  $\text{SiN}$  film, which hence is capable of supporting the SPHP propagation for all considered frequencies.

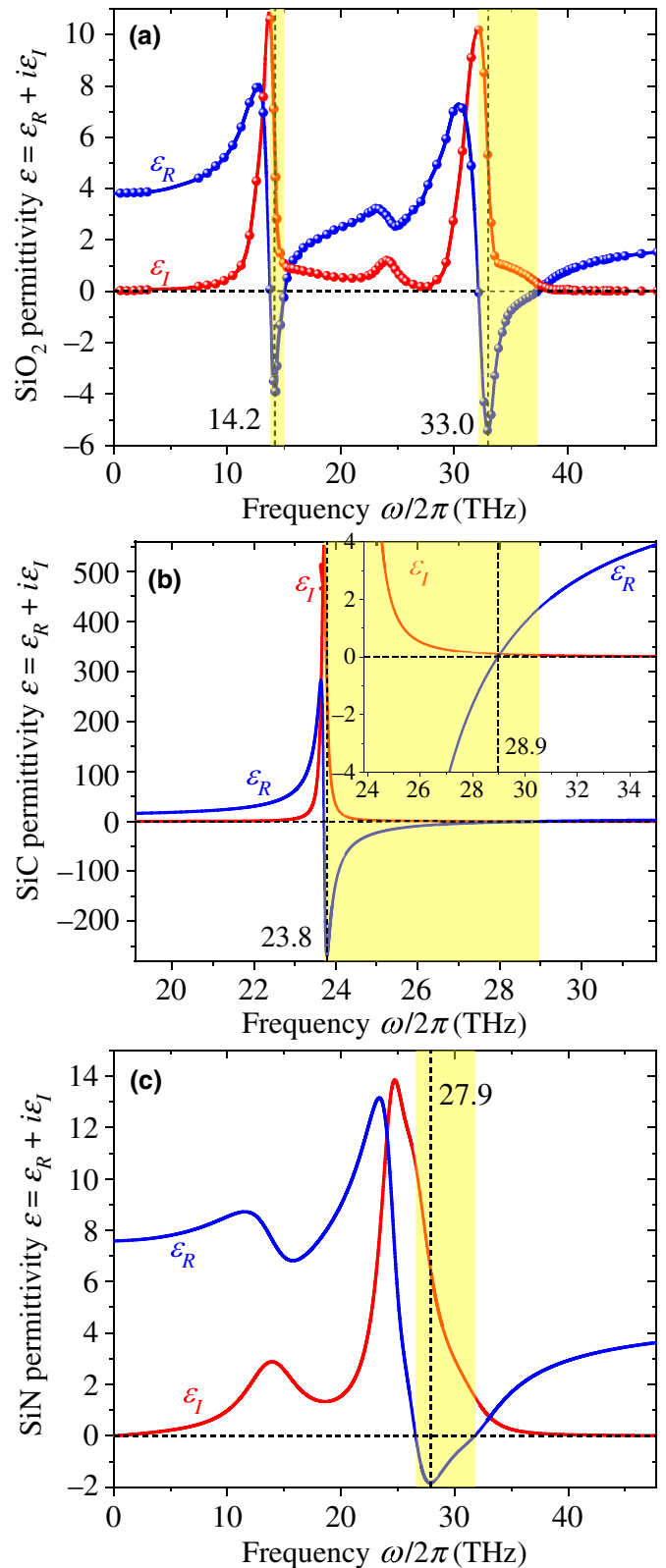


FIG. 3. Real and imaginary parts of the relative permittivity  $\varepsilon(\omega) = \varepsilon_R(\omega) + i\varepsilon_I(\omega)$  of (a)  $\text{SiO}_2$ , (b)  $\text{SiC}$ , and (c)  $\text{SiN}$  as a function of frequency [2,3,5]. The yellow zones stand for the Reststrahlen bands [ $\varepsilon_R(\omega) < 0$ ] of each material.

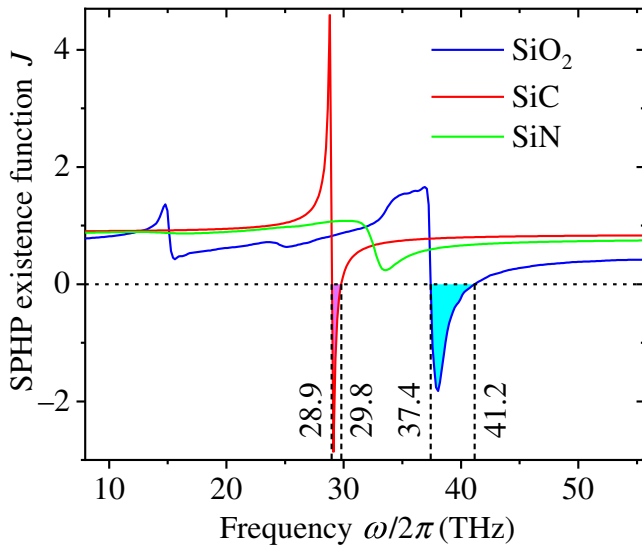


FIG. 4. Frequency spectrum of the SPHP existence function of three common polar materials surrounded by vacuum ( $\epsilon_0 = 1$ ). The colored areas represent the band gaps for the existence of SPHPs and the calculations are done with the complex permittivities of SiO<sub>2</sub>, SiC, and SiN reported in the literature [2,3,5].

The frequency dependence of the dispersion relation (real part  $\beta_R$  of the in-plane wave vector) and propagation length  $\Lambda$  of SPHPs propagating along the interfaces of 100-nm-thick films of SiO<sub>2</sub>, SiC, and SiN suspended in air, are shown in Figs. 5(a) and 5(b), respectively. For the three materials,  $\beta_R$  linearly increases with frequency through values very similar to those of the light line, in particular, for  $\omega/2\pi < 250$  THz. In this frequency regime, the SPHPs thus show a photonlike behavior characterized by a group velocity close to that of light in vacuum ( $V \rightarrow c$ ). For higher frequencies, on the other hand, the SPHP velocity decreases as a result of the values of  $\beta_R$  higher than those of the light line. The fact that  $\beta_R(\text{SiC}) > \beta_R(\text{SiN}) > \beta_R(\text{SiO}_2)$  establishes that SPHPs propagate in SiC slower than in SiN and SiO<sub>2</sub>. However, SiC supports SPHPs with longer propagation lengths ( $\Lambda > 1$  mm) than these latter two materials, within a representative range of frequencies, as shown in Fig. 5(b). Note that both SiC and SiO<sub>2</sub> exhibit frequency band gaps (colored regions) for which there is no propagation of SPHPs, in agreement with the predictions of the SPHP existence function shown in Fig. 4. For the three materials, the maxima and minima of  $\Lambda$  appear at frequencies close to the corresponding ones determining the minima and maxima of the imaginary part of their permittivity [2,3,5], which drives the absorption of energy by the film material. This slight displacement of the peaks and dips of  $\Lambda$  and  $\text{Im}(\epsilon)$  is reasonable, given that the SPHP propagation length is not determined only by  $\text{Im}(\epsilon)$ , but rather by  $\text{Re}(\epsilon)$ , as established by Eq. (3b). The absence of SPHP propagation for the band gaps of SiN and SiO<sub>2</sub>

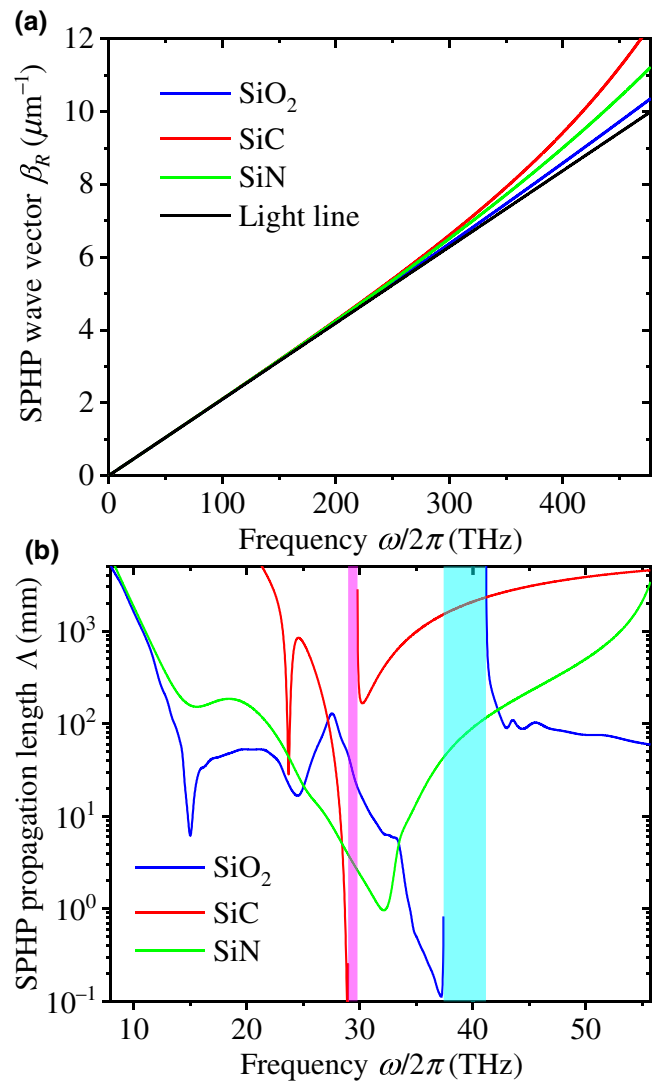


FIG. 5. (a) SPHP dispersion relation and (b) propagation length as functions of frequency, for 100-nm-thick films of SiO<sub>2</sub>, SiC, and SiN suspended in air ( $\epsilon_0 = 1$ ). The colored regions represent the range of frequencies for which there is no propagation of SPHPs.

indicates that the SPHP heat capacity of these materials must be determined by excluding their frequency gaps in the integral of Eq. (2).

Figure 6(a) shows the SPHP heat-capacity spectrum established by Eq. (2) and the results shown in Fig. 5 for the films of SiO<sub>2</sub>, SiC, and SiN. Note that the spectra for these three materials are pretty much the same at room temperature ( $T = 300$  K) and therefore their SPHP heat capacity  $C$  becomes independent of the material properties, as predicted by Eq. (9) and is shown in Figs. 6(b) and 6(c), for thin enough ( $d < 500$  nm) films at low temperature ( $T < 300$  K). By contrast, at high temperature ( $T = 1000$  K), the spectrum for SiC is higher than those of SiN and SiO<sub>2</sub> due to its relatively high wave vector  $\beta_R$

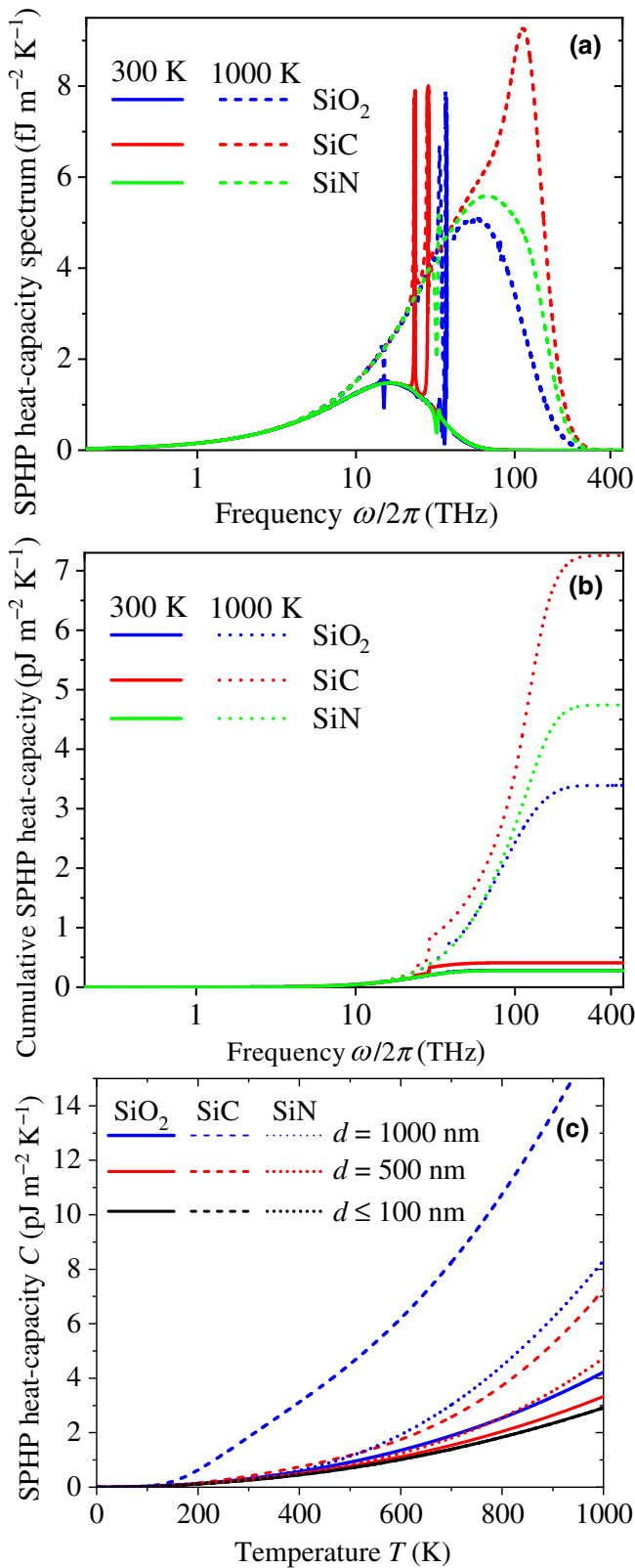


FIG. 6. (a) SPHP heat-capacity spectrum along with its (b) cumulative and (c) integrated values as functions of the temperature of a film suspended in air. Calculations in (a),(b) are done for  $d = 500$  nm, two representative temperatures, and three materials.

and slow speed of SPHPs, as shown in Fig. 5(a). According to the cumulative heat capacity in Fig. 6(b), the major contribution to  $C$  arises from frequencies  $\omega/2\pi < 200$  THz in which the SPHPs propagate with pretty much the same wave vector (and speed) than light, as shown in Fig. 5(a). The photonlike nature of SPHPs is thus responsible of their main contribution to the material heat capacity. In addition, given that  $C$  depends on the material properties through the ratio  $\beta_R/V = \beta_R \partial \beta_R / \partial \omega$ , whose values for SiC, are higher than the corresponding ones obtained for SiN and SiO<sub>2</sub> [Fig. 5(a)] at high frequencies ( $\omega/2\pi > 200$  THz), their impact on  $C$  shows up at high temperature mainly, as established by  $\partial f / \partial T$  involved in Eq. (2) and is shown in Figs. 6(a)–6(c) for  $T = 1000$  K. The sharp changes of the spectra for SiC and SiO<sub>2</sub> appear for the frequencies that define the boundaries of the band gaps shown in Fig. 5(b), as expected. For thicknesses  $d < 500$  nm, the maximum contribution to the SPHP heat capacity arises from the vicinity of the frequency  $\omega_p/2\pi = 0.053 67T$ , as discussed just below Eq. (7). This linear relation of  $\omega_p$  with  $T$  drives the increase of the SPHP heat capacity  $C$  with temperature, as shown in Fig. 6(c). One can see that  $C$  increases with the film thickness, such that its values for SiC are generally higher than those of SiN and SiO<sub>2</sub>. However, as the film thickness decreases through values  $d \leq 100$  nm, the SPHP heat capacity becomes independent of  $d$  and takes its smallest value  $C = C_0$ , as established by Eqs. (5) and (6a). This fundamental contribution is independent of the material properties and proportional to the square of temperature, as predicted by Eq. (9) and is shown in Fig. 6(c) for SiO<sub>2</sub>, SiC, and SiN. In terms of absolute values, the SPHP heat capacity  $C/d \approx 10^{-11}$  Jcm<sup>-3</sup> K<sup>-1</sup> (for SiC at  $T = 800$  and  $d = 1 \mu\text{m}$ ) is, however, 11 orders of magnitude smaller than the corresponding typical values of their phonon counterparts (approximately 1 Jcm<sup>-3</sup> K<sup>-1</sup>) [27,28] and therefore its values are not expected to be observable in an experiment. The relatively small values of the SPHP heat capacity are generated by the speed of SPHPs, which is close to that of light in vacuum and hence much higher than that of phonons, as established by Eqs. (6a) and (6b). In practice, the contribution of SPHPs to the heat capacity of a polar nanofilm can thus be neglected. For solid materials, such as the ones considered in this work, (SiO<sub>2</sub>, SiC, and SiN), the volumetric heat capacity is usually proportional to the thermal expansion coefficient [36], as established by the experimentally validated Gruneisen relation [37,38]. Therefore, the propagation of SPHPs along a thin film is not expected to affect its expansion coefficient, which favors their utilization as heat dissipators.

Finally, note that the considered heat capacity is derived from the variation of the internal energy only (without considering any work [Eq. (2)]) and therefore its values  $C = C_V(\text{SPHP})$  correspond to the SPHP heat capacity at constant volume. The SPHP heat capacity at constant

pressure  $C_P$ (SPHP), on the other hand, can be deduced from Mayer's relation, which establishes that the difference between the total heat capacities at constant pressure ( $C_P$ ) and volume ( $C_V$ ) is proportional to the isothermal compressibility  $\beta_T$  ( $C_P - C_V \propto \beta_T$ ) [27,39]. Given that solids are nearly incompressible ( $\beta_T \rightarrow 0$ ), this relation establishes that  $C_P \approx C_V$ , for materials as the ones considered in this work (SiO<sub>2</sub>, SiC, and SiN). Taking into account that the phonon contributions to the heat capacity at constant volume and pressure of these dielectric materials are nearly identical [ $C_P(\text{phonon}) \approx C_V(\text{phonon})$ ] [27,39], this latter result indicates that their corresponding polariton counterparts are also expected to be nearly equal [ $C_P(\text{SPHP}) \approx C_V(\text{SPHP})$ ].

#### IV. CONCLUSIONS

The surface phonon-polariton heat capacity of polar nanofilms is quantified and analyzed as a function of their thickness and temperature. Based on the two-dimensional density of states of surface phonon polaritons, it is shown that their contribution to the heat capacity of nanofilms ( $d \leq 100$  nm) at sufficiently low temperature ( $T < 300$  K) becomes independent of the material properties and is given by  $C = 1.15\epsilon_0 k_B (k_B T / \hbar c)^2$ , where  $\epsilon_0$  and  $c$  are the respective relative permittivity of the surrounding medium and light speed in vacuum, while  $k_B$  and  $2\pi\hbar$  are the Boltzmann and Planck constants, respectively. Furthermore, it is found that the polariton heat capacity of SiO<sub>2</sub>, SiC, and SiN is several orders of magnitude smaller than its corresponding phonon counterpart, as a result of the relatively high speed of surface phonon polaritons compared to that of phonons.

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