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## Locating rockfalls using inter-station ratios of seismic energy at Dolomieu crater, Piton de la Fournaise volcano

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Key Points:

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14	•	Rockfalls are located using generated seismic signals at high frequencies for highly
15		resolved spatial and temporal tracking
16	•	Rockfall location is improved using the signature of surface topography on seis-
17		mic signals simulated with the 3D Spectral Element Method
18	•	By accounting for topography, all signal components can be used, critical in the
19		case of sparse station networks or noise

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### 20 Abstract

Rockfalls generate seismic signals that can be used to detect and monitor rockfall ac-21 tivity. Event locations can be estimated on the basis of arrival times, amplitudes or po-22 larization of these seismic signals. However, surface topography variations can signifi-23 cantly influence seismic wave propagation and hence compromise results. Here, we specif-24 ically use the signature of topography on the seismic signal to better constrain the source 25 location. Seismic impulse responses are predicted using Spectral Element based simu-26 lation of 3D wave propagation in realistic geological media. Subsequently, rockfalls are 27 located by minimizing the misfit between simulated and observed inter-station energy 28 ratios. The method is tested on rockfalls at Dolomieu crater, Piton de la Fournaise vol-29 cano, Reunion Island. Both single boulder impacts and distributed granular flows are 30 successfully located, tracking the complete rockfall trajectories by analyzing the signals 31 in sliding time windows. Results from the highest frequency band (here 13-17 Hz) yield 32 the best spatial resolution, making it possible to distinguish detachment positions less 33 than 100 m apart. By taking into account surface topography, both vertical and hori-34 zontal signal components can be used. Limitations and the noise robustness of the lo-35 cation method are assessed using synthetic signals. Precise representation of the topog-36 raphy controls the location resolution, which is not significantly affected by the assumed 37 impact direction. Tests on the network geometry reveal best resolution when the seis-38 39 mometers triangulate the source. We conclude that this method can improve the monitoring of rockfall activity in real time once a simulated database for the region of inter-40 est is created. 41

#### 42 **1** Introduction

Seismology is increasingly used to study and monitor dynamic processes at the in-43 terface between the Earth and its fluid envelopes, a field often more specifically referred 44 as environmental seismology (Larose et al., 2015; K. E. Allstadt et al., 2018). Surface 45 processes can include natural phenomena such as storms (e.g. Ebeling & Stein, 2011; Stutz-46 mann et al., 2012), glaciers (e.g. Tsai et al., 2008; Podolskiy & Walter, 2016; Sergeant 47 et al., 2016, 2019), rivers (e.g. Gimbert et al., 2014), debris flow (e.g. Burtin et al., 2009), 48 snow avalanches (e.g. Norris, 1994; Leprettre & Navarre, 1998; Suriñach et al., 2000, 2001) 49 as well as landslides and rockfalls (e.g. Suriñach et al., 2005; Favreau et al., 2010; Hi-50 bert et al., 2011; K. Allstadt, 2013; Bottelin et al., 2014; Vouillamoz et al., 2018). 51

In the context of landslides (used here as the most general term for gravitational 52 mass movements), seismic signals can be used to identify hazards. Growing networks of 53 seismic stations offer the opportunity to continuously monitor large regions of interest. 54 Landslide events can be detected, characterized, and located using the seismic signals 55 they generate (e.g. Suriñach et al., 2005; Hibert et al., 2014; Provost et al., 2017; E. J. Lee 56 et al., 2019). This helps in creating catalogs of landslides that allow statistical analysis 57 of their spatial and temporal activity and estimation of their probability of occurrence. 58 In this way, triggering mechanisms can be studied by correlating landslide catalogs with 59 meteorological data (Burtin et al., 2009; Helmstetter & Garambois, 2010; Durand et al., 60 2018) or with volcanic seismicity data (Hibert, Mangeney, et al., 2017; Durand et al., 2018). 61 On volcanoes, rockfall locations can provide insight into volcano summit deformation (Durand 62 et al., 2018), and seismic signals are also used to monitor other processes such as lahars 63 (e.g. Zobin et al., 2009; Zobin, 2012; Vázquez et al., 2016; Coviello et al., 2018) as well 64 as magma migration (e.g. Taisne et al., 2011; Lengliné et al., 2016; Duputel et al., 2019). 65

Several methods for locating landslides from seismic signals have been proposed and can be divided into two main groups. In the first group, the source location is inferred geometrically by pointing to it from several stations and determining the intersection. This can be done by polarization analysis with three-component seismometers (Vilajosana et al., 2008) or by array analysis methods that estimate the apparent slowness vector (Almendros et al., 2002). In the second group, seismic signal properties are back-projected, optimizing correlation between multiple stations. The back-projection ESSOAr | https://doi.org/10.1002/essoar.10503303.3 | CC\_BY\_

relies either on the decay of amplitudes with distance, using methods such as amplitude 73 source location (ASL, e.g. Battaglia & Aki, 2003; Battaglia et al., 2005; Walter et al., 74 2017; Morioka et al., 2017; Walsh et al., 2017; Pérez-Guillén et al., 2019; Walsh et al., 75 2020) and seismic intensity ratios (e.g. Taisne et al., 2011), or on travel time differences 76 between stations pairs, using cross-correlation of signal envelopes (Burtin et al., 2009; 77 Lacroix & Helmstetter, 2011; Yamada et al., 2012; Bottelin et al., 2014; Dietze et al., 2017) 78 or inversion of first arrival times (Hibert et al., 2014; Gracchi et al., 2017; Fuchs et al., 79 2018). Li et al. (2020) reviews recent advances of back-projection methods to locate seis-80 mic sources, including wavefield migration, waveform inversion, semblance and template 81 matching. 82

As landslides predominantly occur in mountainous regions, generated seismic waves
are prone to interact with rough surface topography variations. The influence of topography on seismic wave propagation has long been a subject of study (Geli et al., 1988).
Topography can affect the wave path, wave polarization (e.g. Ripperger et al., 2003; Métaxian
et al., 2009) and seismic amplitudes (e.g. S. J. Lee et al., 2009; Maufroy et al., 2015).
If not taken into account correctly, topographic effects compromise location methods and
decrease their accuracy.

Assuming elongated wave paths along the topography, back-projection methods can take topography into account by adjusting source-receiver distances and thus travel times. This was done for example by Hibert et al. (2014) to locate rockfalls at Dolomieu crater, Reunion Island, and by Levy et al. (2015) to locate granular flows at Soufrière Hills volcano, Montserrat. However, adjusting the source-receiver distance does not account for diffraction or scattering during the propagation of the seismic wave along the topography.

In the following we propose a new location method that accounts for the cumula-97 tive effect of the topography on the recorded signal. The method is based on the work 98 of Kuehnert et al. (2020), in which topography effects on the seismic wave field were investigated using the 3D Spectral Element Method (SEM, e.g. Komatitsch & Vilotte, 1998; 100 Chaljub et al., 2007) in combination with a realistic geological domain. By calculating 101 seismic energy ratios between stations pairs and hence removing the signature of the seis-102 mic source, they concluded that observed energy ratios from recorded rockfall signals can 103 be reproduced when topography is considered in the simulations and site effects are re-104 moved from the observations. This is used here for locating seismic sources by construct-105 ing a database of simulated energy ratios from a grid of potential source positions with 106 10 m spacing which are then compared to the observed energy ratios after site effect cor-107 rection using spectral amplification functions estimated from volcano-tectonic (VT) events. 108

The method is tested on seismic signals generated by rockfalls at Piton de la Four-109 naise volcano, Reunion Island. After analyzing one rockfall in detail to tune the method 110 for best resolution, a variety of diverse rockfall events are located. As the method as-111 sumes single sources, its performance for largely distributed sources such as granular flows 112 is evaluated. Finally, to investigate the limitations of the method, synthetic rockfall sig-113 nals are constructed from single as well as multiple source positions. A resolution proxy 114 is defined to test the station coverage and identify network geometries with enhanced 115 resolution. Furthermore, the sensitivity of the locating method to the ambient noise level 116 as well as to the underlying model assumptions such as the topography resolution and 117 the source impact direction is assessed. 118

## <sup>119</sup> 2 Rockfall seismic signals at Dolomieu crater, Reunion Island

The study site is located on Piton de la Fournaise volcano, Reunion Island, shown in Figure 1a. Rockfalls occur frequently on the unstable flanks of Dolomieu crater, which was formed during the caldera collapse in 2007 (Michon et al., 2007). The volcano is monitored by the *Observatoire Volcanologique du Piton de la Fournaise* (OVPF). The instrumentation, which includes both seismic stations and cameras, allows rockfall analysis by correlating recorded seismic signals with video recordings. For the present study, four

stations around the Dolomieu crater with a sampling frequency of 100 Hz are used, namely 126 the three-component stations BON, BOR, and SNE and the vertical component station 127 DSO. BON and SNE are broadband (i.e. corner frequency  $< 0.1 \, \text{Hz}$ ), BOR and DSO are 128 short-period (i.e. corner frequency  $> 0.1 \,\mathrm{Hz}$ ) stations. The three cameras CBOC, DOEC, 129 and SFRC are located on the summit of Piton de la Fournaise, look into the Dolomieu 130 crater and continuously sample two images per second. The supporting information of 131 this article provides the seismograms and the videos of all analyzed rockfalls. To eval-132 uate the results of the present location method, rockfall trajectories are manually esti-133 mated from the videos by determining landmarks visible on both the videos and the avail-134 able Digital Elevation Model (DEM) of 10 m resolution. This way, the trajectory of a 135 rockfall on December 13, 2016, is reconstructed and marked in red in Figure 1a. An un-136 certainty of  $\pm 50 \,\mathrm{m}$  is assumed. 137

The recorded ground velocity generated by this rockfall on the southwestern crater wall is shown in Figure 1b. The most abrupt signals are observed at BOR and DSO, the two closest stations, whereas the signals at stations BON and SNE at larger distances slowly rise in amplitude. The temporal evolution of the recorded signals can be characterized by a proxy of the seismic energy  $E_{ij}$  measured at each station *i* and component *j* that we define as the square of the recorded ground velocity  $v_{ij}^2(t)$ , integrated over a sliding window of width *d*:

$$E_{ij} = \int_d v_{ij}^2(t) \,\mathrm{d}t. \tag{1}$$

Energy proxies  $E_{iz}$ , calculated from vertical component j = z, are shown in Figure 1b. Their inter-station ratios are shown in Figure 1c, where BON is chosen as the reference station. The beginning of the seismic signal generated by the rockfall is marked by an abrupt increase of the ratios BOR/BON and DSO/BON, whereas the ratio SNE/BON decreases. Subsequently, the ratios evolve differently as the rockfall moves towards the bottom of the crater.

As the seismic source is identical for all stations, the temporal evolution of energy 151 ratios is caused by the wave propagation path. First of all, as the source position moves, 152 the source-receiver distances vary, which modifies the amplitude of the signals because 153 of geometric spreading and attenuation. Moreover, soil heterogeneities and topography 154 can affect the wave propagation between the source and the receiver. By modeling the 155 influence of the topography on the energy ratios through direct numerical wave simu-156 lation and by taking into account local heterogeneities using empirical site amplification 157 factors, the present study aims to locate rockfalls with high spatial and temporal res-158 olutions. 159

## 160 3 Methodology

The proposed methodology for estimating rockfall locations uses energy ratios be-161 tween stations to predict source positions. The ratios characterize the path effects, while 162 the energy of the source itself can be ignored. This was used for example by Taisne et 163 al. (2011) to map magma propagation. Here we use this strategy to compare the observed 164 energy ratios with simulated ones, which was shown by Kuehnert et al. (2020) to be pos-165 sible considering the topography. Instead of using spectral ratios at single frequencies, 166 we average the energy ratios within frequency bands of 4 Hz. This makes the method more 167 robust to fluctuations in spectral values. In order to explore all potential rockfall sources, 168 reciprocal simulations are carried out, where the synthetic source is placed at the po-169 sition of the seismometer and the wave field is recorded over the entire crater area. A 170 grid search is then performed to find the source positions that best fit the observed en-171 ergy ratios. 172

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Figure 1. (a) Map of Reunion Island (top) and hillshaded elevation map of Dolomieu crater on Piton de la Fournaise volcano (bottom). The smaller Bory and Soufrière craters are located west and north of Dolomieu crater, respectively. Seismic stations are marked by green triangles and cameras by blue dots. The red zone marks a rockfall trajectory estimated from the video. (b) Seismograms (left) and corresponding spectrograms (right) of vertical velocity generated by a rockfall on December 13, 2016, corresponding to the red trajectory indicated in a). The signals were recorded at the four seismic stations surrounding Dolomieu crater. The red curves are the seismic energy proxy  $E_{iz}$  (according to Eq. 1), calculated from a sliding time window in steps of 2 s and of width d = 4 s. Ambient seismic noise can only be detected in the spectrogram at furthest station SNE below 3 Hz. (c) Inter-station energy ratios from vertical ground velocity. The beginning of the seismic signal emitted by the rockfall is marked by an abrupt change of the ratios.

#### 3.1 SEM simulations

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The propagation of the seismic wave field is simulated using the 3D Spectral Element Method (SEM, e.g. Komatitsch & Vilotte, 1998; Chaljub et al., 2007) in a numerical domain of dimensions x = 2100 m (easting), y = 1800 m (northing), and z = 600 m (depth) as shown in Figure 2a, identical to the domain of the simulations presented by Kuehnert et al. (2020).

The domain is meshed in the top  $150 \,\mathrm{m}$  with hexahedral elements of  $10 \,\mathrm{m}$  side length 179 to correctly accommodate the surface topography of Dolomieu crater provided by a 2009 180 Digital Elevation Model (DEM) of 10 m resolution. Further below in depth, the element 181 size is increased to 30 m to reduce computational costs. A Zone of refinement connects 182 the two different element sizes, while a smooth Buffer layer filters out short wavelength 183 variations of the fine mesh that cannot be represented in the coarse mesh. A polynomial 184 order of 5 is used in all elements. To simulate an unbounded domain, absorbing PMLs 185 (Perfectly Matched Layers, e.g. Festa & Vilotte, 2005) of 160 m thickness are attached 186 on the sides and the bottom of the domain. 187

The subsurface is parametrized using the generic velocity model proposed by Lesage 188 et al. (2018) for the shallow structure of volcanoes. It is characterized by a velocity pro-189 file gradually increasing with depth as illustrated in Figure 2b. It is implemented on the 190 3D domain by laterally following the surface topography (i.e. each point at the surface 191 is defined by depth z = 0 m), deforming the horizontal layers of equal velocity. Kuchnert 192 et al. (2020) validated that this velocity model represents reasonably well the present study 193 site at Piton de la Fournaise volcano by comparing simulated and recorded seismic sig-194 nals of different rockfalls. The rock density is set to  $\rho = 2000 \,\mathrm{kg \, m^{-3}}$  and quality fac-195 tors are set to  $Q_P = 80$  and  $Q_S = 50$  for P- and S-wave velocity, respectively. These 196 values are based on previous studies on Piton de la Fournaise and similar volcanoes (Battaglia 197 & Aki, 2003; O'Brien & Bean, 2009; Hibert et al., 2011). 198

Seismic sources are represented by a point force and a Ricker wavelet of 7 Hz dom-199 inant frequency. This corresponds to a frequency range between 2 Hz and 20 Hz, in agree-200 ment with the predominantly observed seismic spectrum associated with rockfalls at the 201 Dolomieu crater (see Fig. 1b and e.g. Hibert et al., 2014). The source-time function as 202 well as its spectral content is displayed in Figure 2c. A typical wave simulation is shown 203 in Figure 2d with a snapshot at time  $t = 3.2 \,\mathrm{s}$  illustrating the wave field radiated by 204 a vertical point source located at the southern crater wall (yellow arrow). It can be ob-205 served that the majority of seismic energy is located close to the surface as a result of 206 the shallow low seismic velocity. The surface topography causes a highly scattered wave 207 field. Synthetic seismograms recorded at the surface along the cross-section are shown 208 in a time-offset representation. The wave field originates at the source location (0 km off)209 set) and travels outwards in all directions. Wave scattering caused by the topography 210 is detectable here, especially close to the crater rim at around 0.6 km offset with reflec-211 tions back-propagating towards the bottom of the crater. 212

<sup>213</sup> Concerning computational efforts, it takes a CPU time of around 10,000 CPU hours <sup>214</sup> (10 cores per CPU) for one simulation on the presented domain (i.e. duration: 6 s, num-<sup>215</sup> ber of elements: 915,704, number of GLL points: 6, max. frequency: approx. 20 Hz, min. <sup>216</sup> velocity:  $320 \text{ m s}^{-1}$ ). We run the simulations in parallel on 200 CPUs, leading to 2.3 days <sup>217</sup> per simulation.

To efficiently explore different potential positions of the rockfall source without performing a simulation for each of them, the reciprocity principle is used (Aki & Richards, 2002): the synthetic source is located at the station location and the wave field is recorded at the source location. Potential rockfall source positions are confined within a rectangular area at Dolomieu crater, shown in Figure 3a. The area is sampled by a grid of measurement points (in blue) with 10 m spacing.

The principle of reciprocity is illustrated in Figure 3b. It is shown that performing reciprocal simulations by interchanging source and receiver (and their corresponding directions) results in identical synthetic seismograms. In order to collect all necessary information, simulations for each component of all seismometers are carried out, i.e. manuscript submitted to JGR: Earth Surface



Figure 2. (a) Cross-section through the meshed domain with 10 m resolution surface topography from Dolomieu crater. The color map depicts the seismic velocity model. The elements have a side lengths of 10 m in the top 150 m, increasing to 30 m below the *Zone of refinement*. Absorbing PMLs (Perfectly Matched Layers) of 160 m width are attached on the sides and bottom of the domain. (b) Generic velocity-depth profile for S-wave velocity  $v_S$  (solid line) and P-wave velocity  $v_P$  (dashed line) as proposed by Lesage et al. (2018) for the shallow structure of volcanoes. (c) Ricker wavelet of 7 Hz dominant frequency: source-time function (top) and corresponding spectrum (below). (d) Simulation of the wave field (vertical velocity) from a vertical source on the southern crater wall (yellow arrow). On the top, a snapshot of the wave field is shown at time t = 3.2 s, where positive amplitudes are denoted in red, negative in blue. The graph below shows the seismic traces recorded at the surface along the same cross-section. The green trace corresponds to the signal recorded at station BON.

200 Carried out 5 × 5 sindlations for the time-component sets inductors Dory, Borr, and

<sup>231</sup> SNE and one simulation for the single-component seismometer DSO. This is done for

<sup>&</sup>lt;sup>228</sup> a point source is placed at the position of a given seismometer while the input force di-<sup>229</sup> rection is aligned with the component of the seismometer. In total, ten simulations are <sup>230</sup> carried out:  $3 \times 3$  simulations for the three-component seismometers BON, BOR, and



Figure 3. (a) Grid of receivers (in blue) for reciprocal simulations. The yellow star and red triangle are used to illustrate reciprocity in b). The sampled area measures  $1200 \text{ m} \times 1000 \text{ m}$  (east  $\times$  north). Sample spacing is 10 m, resulting in  $121 \times 101 = 12221$  grid points. (b) Illustration of the principle of reciprocity. *Top:* Forward simulation where source (vertical) and receiver (horizontal) are placed at the position of the true source and the true receiver, respectively. *Bottom:* Reciprocal simulation, where a synthetic horizontal source replaces the true horizontal receiver and a synthetic vertical receiver replaces the true vertical source, resulting in identical synthetic seismograms.

both the model with Dolomieu crater topography and for a model with a flat surface forcomparison.

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#### **3.2** Optimization method for source location

A source location probability estimate is associated with each point of the grid in Figure 3a by considering the inverse of the misfit between the synthetic energy ratio  $E_{ij}^{\text{simu}}/E_{\text{ref},j}^{\text{simu}}$ , computed when the source is actually located at that specific grid point, and the observed energy ratio  $E_{ij,\text{tw}}^{\text{obs}}/E_{\text{ref},j,\text{tw}}^{\text{obs}}$ . Here 'ref' refers to the reference station, while *i* designates the station and *j* the component considered. Given that the rockfall source is moving, the observed energy ratio is evaluated over a time window 'tw'. The misfit  $e_{\text{tw}}$  for each time window is defined as follows:

$$e_{\rm tw} = \frac{1}{N_{\rm Sta}} \sum_{ij}^{N_{\rm Sta}} \left| \log_{10} \left( \frac{E_{ij}^{\rm simu}}{E_{\rm ref,j}^{\rm simu}} \div \frac{E_{ij,\rm tw}^{\rm obs}}{E_{\rm ref,j,\rm tw}^{\rm obs}} \right) \right|,\tag{2}$$

where  $N_{\text{Sta}}$  is the number of station-channel pairs to be considered, with each compo-242 nent counted separately. Zero misfit is achieved when simulated and observed energy ra-243 tios are equal. Using the logarithm in equation (2) distributes their relative values equally 244 around zero. This, combined with the absolute value, ensures that the misfit estimation 245 is not biased by the reference station. The probability of the source location is calculated 246 by the inverse of misfit  $e_{tw}$  and scaled to a probability density function (PDF) with rel-247 ative values between 0 and 1. Alternative formulas were investigated, for example the 248 relative difference between simulated and observed energy ratios or approaches with con-249 ditional statements. However, the estimate in equation (2) was evaluated to be the most 250 appropriate for the location problem, notably because it results in spatially smooth vary-251 ing probability values and because it is not biased towards the reference station. 252

In order to consider the frequency dependency of the energy ratios, location is carried out in different frequency bands, namely at 3-7 Hz, 8-12 Hz, and 13-17 Hz. This selection is defined to cover a large part of the available frequency content from the simulations and the observations. A bandwidth of 4 Hz is assumed to be narrow enough to respect the dispersive character of the energy ratios and broad enough to average over fluctuations of the spectral ratios. Noise levels at Dolomieu crater are very low at frequencies above 3 Hz and can be ignored in the tests. Tests of the location method with added synthetic withe noise are performed in section 5.3.

Selecting the width of time window 'tw' over which the observed seismic energy is estimated requires special attention. On the one hand, the width has to be chosen as small as possible in order to sample the moving source. On the other hand, as the same time window is used for all stations, most of the seismic signal generated by a given rockfall source has to arrive at each of the stations within the time window. In order to respect both criteria, a window width of 4 s is defined, and confirmed by simulations to be an appropriate compromise.

To allow comparison between observed and simulated energy ratios, the recorded signals must first be corrected for local site amplification, not considered in the simulations. Therefore, site amplification functions were estimated for each station channel using thirty-six volcano-tectonic (VT) events that were centered around 2 km below Dolomieu crater. Station BON is used as the reference station given its low spectral amplitudes from VT recordings as well as low spectral H/V noise ratios. The resulting site amplification functions are shown in Figure 4. The site effect correction is performed prior to locating by deconvoluting the recorded signals with the spectral amplification functions.



Figure 4. Spectral amplification functions estimated from volcano-tectonic (VT) events relative to reference station BON. Amplification is calculated from smoothed spectral ground velocity recorded by: a) vertical components, b) north components, and c) east components. Smoothing is performed as proposed by Konno and Ohmachi (1998). The blue-shaded area indicates the standard deviation of the amplification as calculated from all VTs. Figure adapted from Kuehnert et al. (2020).

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To test the influence of the above parameters and site effects on the location method, a hands-on Jupyter notebook (Kluyver et al., 2016) is published on https://github.com/ Jubeku/RF\_localization (Kuehnert et al., 2019).

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#### 3.3 The influence of topography on inter-station energy ratios

The relative amplitudes recorded at various stations can be influenced by the topography (e.g. Kuehnert et al., 2020), thus modifying the energy ratios in equation (2). Having built databases of the simulated energies  $E_{ij}^{simu}$  for a domain with a flat surface and a domain with topography, we can gain a first insight into the influence of topography by comparing the resulting synthetic energy ratios.

This is done in Figure 5, where the energy ratio between station pair BOR and BON at each grid point of potential source locations is shown for a flat surface (top) and for the Dolomieu topography (bottom). The energy ratios  $E_{ij}^{\text{simu}}/E_{\text{ref},j}^{\text{simu}}$  are calculated respectively from all three components j = Z, N, E and reference station i = ref is chosen to be BON.

For the domain with a flat free surface, when vertical signal component j = Z is measured, we can observe a bipolar pattern of the energy ratios with values > 1 towards station BOR and values < 1 towards reference station BON, while unity is reached at equidistant positions between the station-pair. Values are determined purely by the sourcereceiver distances. The energy ratios from the horizontal signal components i = N, Eresult in a more complicated spatial pattern. This is because the radiation pattern from the vertical source in the horizontal plane is not radially isotropic.



Figure 5. Seismic energy ratios between station pair BOR and BON (in green) from simulations on a domain with a flat surface (top) and a domain with topography (bottom). At each grid point (see Fig 3a), representing a potential source position, the ratio is computed from vertical-component seismic energy  $E_{i,Z}$  (left), north-component seismic energy  $E_{i,N}$  (middle), and east-component seismic energy  $E_{i,E}$  (right). Unfiltered synthetic seismograms were used for the calculation.

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For the domain with surface topography, in the case of vertical component i = Z, 298 the pattern of energy ratios becomes distorted because of the signature of the topogra-299 phy in the wave field. In general, a bipolar spatial distribution of the energy ratios per-300 sists, indicating that the decay of seismic amplitude remains influenced by the source-301 receiver distance. The pattern of energy ratios from the horizontal signal components 302 i = N, E is comparable to the vertical pattern, indicating that the wave propagation 303 along the topography dominates over source-characteristic radiation patterns. This to-304 pographic path effect (e.g. Kumagai et al., 2011; Kuehnert et al., 2020) is similar to the 305 distortion of radiation patterns by the scattering of the wave field by small-scale soil het-306 erogeneities, validating the assumption of an isotropic radiation at high frequencies above 307

around 5 Hz (e.g. Takemura et al., 2009; Kumagai et al., 2010). As a consequence, the
presented method can be implemented independently of the source impact direction used
in the simulations (here we have chosen a vertical surface traction), whereby both vertical and horizontal component signals can be used for location. We show here that leveraging horizontal component signals can improve the locating results. Typically, only vertical component signals are used in rockfall location methods, except for polarization approaches such as proposed by Vilajosana et al. (2008).

### <sup>315</sup> 4 Application to rockfalls at Dolomieu crater

The proposed formalism to evaluate the relative probability of potential source locations on a predefined grid of positions is now applied to rockfall seismic signals recorded at Piton de la Fournaise volcano. After analyzing individual time windows, all probabilities derived from a sliding time window are combined in the attempt to reconstruct the full rockfall trajectory.

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## 4.1 Rockfall location at given time steps

The location method is initially tested for a rockfall that occurred on December 13, 2016, corresponding to the event presented in Figure 1. The analysis is carried out at six different times i) to vi) as indicated on the seismogram in Figure 6. Above the seismogram, the whole trajectory is shown as well as camera snapshots of times ii) to v).

Time i) is just before the start of the rockfall. Time ii) is after the detachment, when 327 movements can be detected on the video. At time iii), the rockfall appears from behind 328 a small valley at the top of the crater wall. Thereafter, the rockfall accelerates, which 329 leads to stronger impacts and thus to the highest signal amplitudes. A total of three boul-330 ders are detected at time iv) on their way down towards the crater bottom. At time v), 331 the third boulder arrives at the bottom. No movement is detected later on the video at 332 time vi). Nevertheless, it can be assumed that smaller granular material is still active 333 on the flank, causing small amplitude seismic signals. 334

Rockfall location is performed here in the highest frequency band (13-17 Hz) with simulated energy ratios from the model with Dolomieu topography and using all available station-channel pairs, i.e.  $N_{\text{Sta}} = 7$ , adding up three station pairs for the vertical component and two station pairs each for the north and east component (DSO contains only the vertical component). Figure 7 shows the resulting source location probability maps at the six successive time steps i) to vi).

At time i), most probable seismic sources are located at the center and the south-341 eastern side of the crater. As the rockfall has not yet started at that time, the distribu-342 tion must be related to ambient seismic noise. At time ii), the source probability abruptly 343 moves southwest, marking the beginning of the rockfall. The position of maximum prob-344 ability is around 100 m from the estimated location of detachment. Then the area of prob-345 able source locations moves north at time iii) with the most probable source location ap-346 proaching the estimated rockfall position. The predicted source location continues to move 347 along the rockfall trajectory at time iv). At time v) it arrives at the bottom of the crater. 348 At this time, the position of maximum probability is at around 200 m from the estimated 349 location. However, the distribution also shows densely populated high probabilities close 350 to the estimated rockfall location. At time vi), after the last boulder visible on the video 351 has reached the crater floor, a zone of probable source positions remains in the lower part 352 of the trajectory. This may be explained by the late movement of granular material that 353 is not visible on the video. 354

The spatially scattered distribution of the predicted sources and the discrepancy between position of maximum probability and actual rockfall location imply a lack of resolution that can have several reasons. Firstly, the source positions are somewhat ambiguous, i.e. different locations can explain the observed seismic energy ratios equally



Figure 6. Camera images and seismic signal of a rockfall on December 13, 2016. Top: Images taken by camera DOEC. The full rockfall trajectory on the left is reconstructed from differences between successive images. Towards the right, snapshots at times ii) to v) are displayed. Rockfall positions are indicated by red circles and the direction of movement by red arrows. Bottom: Vertical ground velocity  $v_Z$  recorded at closest station BOR. Time steps i) to vi) are marked by red vertical dashed lines. Rockfall location is performed in time windows of  $\pm 2$  s around these time steps. The signal is bandpass filtered at 13-17 Hz.

well. Secondly, as observed on the video, the rockfall does not consist of a single boul-359 der. The resulting seismic signal is hence a superposition of multiple sources shifted in 360 time and space. Given that it is based on the assumption of a single source, the loca-361 tion method is flawed, a problem that will be studied in section 5.1 using synthetic sig-362 nals. The general southward shift of the predicted source locations compared to the true 363 trajectory may also be caused by soil heterogeneities that affect seismic wave propaga-364 tion and are not considered by either the simulations or the site impact correction. An-365 other cause could be an inaccurate representation of the topography, which is possible 366 since the DEM used here is from 2009 and the rockfall analyzed occurred in 2016. 367

#### **4.2** Spatio-temporal rockfall evolution

In order to reconstruct the full rockfall trajectory, the location method of equation (2) is used with a sliding time window. Results from all time windows are combined at each potential source position by selecting the maximum probability over time. For each grid point, the minimum misfit *e* between observed and simulated energy ratios is defined by

$$e = \min_{\mathrm{tw}} e_{\mathrm{tw}},\tag{3}$$



Figure 7. Location of seismic source at time steps i) to vi) as defined in Figure 6. The colorscale represents the source location probability. Black stars denote the position of maximum probability. Red shaded zone marks actual rockfall trajectory as estimated from the video and the red arrows approximate the current rockfall location. Locating is carried out in frequency band of 13-17 Hz using simulations from the model with Dolomieu topography. All stations and components are used, i.e.  $N_{\text{Sta}} = 7$ .

where  $e_{tw}$  is the misfit in each time window 'tw' defined in equation (2). The maximum probability is the inverse of misfit e and can be plotted for each spatial point. In this way, the temporal evolution of the rockfall trajectory can be displayed on a single graph.

Figure 8a shows the resulting location map of the previously analyzed rockfall, using the same method configuration (i.e. at high frequency 13-17 Hz and with all available station-channel pairs). Thanks to the color sequence, we can track how the rockfall moves from top to bottom of the crater over time in agreement with the observed rockfall trajectory from the video. Black stars denote the positions of maximum probability at time steps ii) to v), identical to those shown in Figure 7. Again, a general southward shift of around 100 m with respect to the video-estimated trajectory is observed.

Rockfall location is performed in different frequency bands. In the intermediate fre-384 quency band at 8-12 Hz, Figure 8b, the predicted source locations follow the movement 385 of the actual trajectory and the positions of maximum probability are at distances com-386 parable to those in Figure 8a. However, the resolution decreases as the spatial distribu-387 tion of probable sources becomes much wider, covering large parts of the crater. The res-388 olution is even worse in the lowest frequency band at 3-7 Hz, Figure 8c, where the gen-389 eral downward movement of the rockfall is hardly noticeable, with large discrepancies 390 of the maximum probability positions in the time steps ii) to v). The observed decrease 391 of resolution towards lower frequencies can be explained by the increase of the seismic 392 wavelength. Assuming mainly fundamental-mode Rayleigh waves (Kuehnert et al., 2020), 393 the wavelength for the velocity model used here increases from 26 m at 15 Hz by a fac-394 tor of 1.7 to 44 m at 10 Hz and by a factor of 4.5 to 116 m at 5 Hz; resolution can be ex-395 pected to decrease accordingly. 396

A reduction in location error when higher frequencies are used in the location process is also reported by Lacroix and Helmstetter (2011). When analyzing single impact signals with frequency contents up to 30 Hz, they achieve locating accuracies of 50 m us-



Figure 8. Spatio-temporal rockfall evolution. Color represents time of seismic record and intensity representing probability of source location. Black shaded zone indicates rockfall trajectory from video. Black stars denote positions of maximum probability at time steps ii) to v). The seismogram underneath was recorded at closest station BOR at 13-17 Hz, associating time with color using a 2D colorbar (MMesch, 2016). Signals recorded at all stations for all components are shown with scales in the supplementary material. Using the results from the same simulation including topography, the pre-processing of synthetic and observed signals for location changes as follows: (a) Location at high frequency band 13-17 Hz using all available station channels (ENZ, i.e. east, north, and vertical components), thus  $N_{\text{Sta}} = 7$ . (b) Location at intermediate frequency band 8-12 Hz, with all components ENZ, i.e.  $N_{\text{Sta}} = 7$ . (c) Location at low frequency band 3-7 Hz, with all components ENZ, i.e.  $N_{\text{Sta}} = 7$ . (d) Location at 13-17 Hz, using only vertical component Z, i.e.  $N_{\text{Sta}} = 3$ . (f) Location at 13-17 Hz, using vertical component Z, i.e.  $N_{\text{Sta}} = 3$ , and using simulations from a model with a flat surface.

ing beamforming and a priori measured seismic velocities. Similarly, when analyzing sin-

cation accuracies of 81 m, comparing the results of back-projecting seismic envelopes with

ing window to track snow avalanches and slush flows. When comparing the locations from

gle impacts and frequency contents of 5-25 Hz, Dietze et al. (2017) achieves average lo-

those of TLS-based measurements. For continuous and distributed sources, the reported

location accuracies decrease. Pérez-Guillén et al. (2019) use the ASL technique in a slid-

seismic signals to numerical flow simulations, they report mean locating accuracies be-

tween 85 m and 271 m, which is of a similar order of magnitude to the results presented here.

Figure 8d shows the location results without prior correction of the recorded signals from site amplification. The results fail to predict a clear rockfall trajectory and a large spatial discrepancy is observed between probable source positions and actual rockfall location.

In Figure 8e, the rockfall is located using seismic signals of only vertical compo-413 nent Z, leading to  $N_{\text{Sta}} = 3$  station-channel pairs. A narrow corridor of high proba-414 bilities can be seen, indicating a well resolved rockfall trajectory. However, compared to 415 the results in Figure 8a, a larger discrepancy with the actual rockfall location is observed. 416 This suggest that adding additional measurements may reduce the resolution as it be-417 comes harder to keep the misfit, as defined in equation (2), small, however, the predic-418 tions potentially improve as noisy or malfunctioning measurements can be compensated, 419 which is in agreement with the network resolution study in section 5.2. 420

In Figure 8f, locating is carried out using simulated energy ratios from a model with 421 a flat surface. In this case, only vertical components can be used as energy ratios from 422 horizontal components lead to values that are strongly modulated by radiation patterns, 423 as shown in Figure 5. The resulting source probability distribution consists of patch-like 424 areas that do not show smooth transitions over time (i.e. color), leading to a coarse rock-425 fall location that is not well resolved in time. This is because the energy ratios from the 426 flat model are dominated by the source-receiver distances and only these localized patches 427 can explain the observed energy ratios. Typically, rockfall location methods attempt to 428 account for the effect of the topography during the location process by considering a map 429 of elongated travel times (e.g. Hibert et al., 2014; Levy et al., 2015; Dietze et al., 2017), 430 assuming straight wave paths along the surface. The method proposed here allows the 431 high-resolution topography and its influence on the wave field to be fully accounted for 432 by 3D numerical modeling of the seismic wave field. The influence on the location by im-433 plementing a slightly coarser resolution DEM is demonstrated in section 5.3. 434

#### 4.3 Locating other rockfalls

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The comparison in the previous section suggests that the best locating results for the present study site can be achieved in the high frequency range (13-17 Hz), using both vertical and horizontal components (ENZ), removing site effects from the observed signals, and simulating energy ratios on the model with topography. With this configuration, the location method will now be evaluated using four rockfalls of different types and from different locations within the Dolomieu crater. The observed trajectories as well as the locations are shown in Figure 9 and described below.

Rockfall (a) is again located in the southwestern region, with an initial detachment further north compared to the previously analyzed event. Rockfall (b) is in the southeast and rockfall (c) is in the north of the Dolomieu crater. Rockfall (d) occurred in the same region as rockfall (c) but consisted of fine granular material in contrast to the other events which consisted of individual boulders.

For rockfall (a), Figure 9a, the most probable sources are all inferred at locations 448 close to the observed trajectory in the southwestern region of Dolomieu crater. In par-449 ticular, the location of the detachment phase (in purple) and the observed trajectory to-450 wards the east are well represented. However for the last stage of the rockfall, inferred 451 sources are located too far south, at the wall of the crater, and not at the bottom of the 452 crater as observed. This might be interpreted as the signature of possible superposition 453 of the seismic signals generated by subsequent boulders or an incorrect topography rep-454 resentation. Note that the resolution of the location method makes it possible to iden-455 tify the trajectory and the detachment of this rockfall event distinctly with respect to 456 the trajectory and detachment of the event analyzed in the previous section, Figure 8. 457 for which the detachment phase is located 100 m further south. 458

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Figure 9. Four rockfalls used for locating evaluation. Left: trajectory reconstructed from successive camera images (outline marked in orange and north-direction and camera indicated at the bottom). Right: map of predicted spatio-temporal source evolution (black-shaded video estimated trajectory). Underneath: seismograms recorded at closest station (vertical ground velocity, bandpass filtered at 13-17 Hz). Signals recorded at all stations for all components are shown with scales in the supplementary material. a) Rockfall consisting of individual boulders occurring on December 7, 2015 in the southwest with detachment position just beneath Bory crater. b) Rockfall consisting of individual boulders occurring on February 28, 2016 in the south. c) Rockfall consisting of individual boulders occurring on January 22, 2017 in the north. d) Rockfall consisting of fine granular material occurring on June 14, 2016; traces from dust clouds extend beyond the outlines of the sketched event location and station DSO was malfunctioning and could not be used for locating.

For rockfall (b), Figure 9b, the inferred sources are correctly located in the south-459 ern region of the Dolomieu crater, but with strongly deteriorated resolution in space and 460 time. The inferred source locations 30s after the start of the event (in green), are located 461 at the bottom of the crater. The video shows that the first boulder arrives at the crater 462 bottom at this time, but other boulders are still moving at the top of the crater wall. 463 Multiple sources hamper the ability of the location method to determine the trajectory 464 of a single source. As a result, the inferred sources at later times (yellow and red colors) 465 are located half-way down at the crater wall. Another explanation for the poor resolu-466

tion is the station network configuration which will be studied in section 5.2 using syn thetic signals.

For rockfall (c), Figure 9c, the inferred sources are well-located at the beginning of the event, while locations become more and more scattered in space at later times. This time-deterioration of the resolution can be analyzed with the help of the video that shows that at beginning the event initially involves a single boulder impacting the crater wall, with subsequent distribution of boulders originating from the fragmentation of the original boulder or from the mobilisation of basal rock deposits.

Finally for rockfall (d), Figure 9d, which occurred in the same region as event (c) 475 but consisted of fine granular material flowing down the steep crater wall, the method 476 is able to locate the event with high-resolution, in particular the initial activation loca-477 tion. This is quite remarkable given that station DSO was not functioning properly and 478 was disregarded for the analysis. Moreover, given that the source is parametrized as a 479 single force, this high-resolution of the source locations in the case of granular flow is not 480 intuitive as it generates a complex extended source. This might suggest that recorded 481 signals are dominated by a localized high-energy radiating source area, which we will fur-482 ther discuss in section 5.4. Using a similar approach based on analysing the seismic sig-483 nals in a sliding time window, Pérez-Guillén et al. (2019) are also able to track the dis-484 tributed and moving seismic sources generated by snow avalanches and slush flows. 485

### <sup>486</sup> 5 Evaluation of the presented location method using synthetic signals

Rockfall events generate complex and extended seismic sources, and the resolution 487 and the limitations of the proposed location methods need to be assessed through tests 488 with synthetic seismograms for which the seismic sources can be controlled, e.g., with 489 known source time functions and locations. In this way, we study the problem of the su-490 perposition of spatially distributed sources as well as the performance of the location method 491 in different frequency bands and the error introduced when topography is not consid-492 ered. The study with synthetic seismograms additionally offers the possibility to eval-493 uate the influence of the network geometry on the location resolution. 494

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## 5.1 Single-point sources mimicking boulder impacts

In order to assess the frequency-dependent location error introduced when topog-496 raphy effects are ignored, synthetic seismic signals are generated for two controlled point 497 sources (e.g. two distinct boulders impacting at the same time with the same force, lo-498 cated in the southwestern part of Dolomieu crater) through 3D wave simulation in the 499 model including topography. Locations of the sources are then inferred from different 500 frequency bands of the synthetic signals using the location method with and without to-501 pography effects. For better comparability between these two cases, only vertical com-502 ponent signals are used for location. Figure 10 summarizes the main results. The true 503 locations of the two point sources are located in the center of the red circles. 504

The left and middle columns show the inferred locations when considering sources at position P1 and P2 separately and the right column when the two sources are acting simultaneously. Rows a-d show the inferred locations when considering two frequency bands of the synthetic signals (i.e. 13-17 Hz and 3-7Hz) and propagating models with topography or a flat surface.

For the high-frequency band, Figure 10a shows the results when using propagating models including topography effects. The true positions of the sources P1 and P2, when considered acting separately, are well reconstructed as would be expected, corresponding to a point with high-probability (i.e. dark purple) in the center of the red circles. In contrast, when using a flat propagating model, Figure 10b, sources are reconstructed with a 100 m-200 m location error and with a spatial ambiguity (multiple source positions with similar probability).



Figure 10. Location of point source at positions P1 and P2 and after simultaneously activating P1 and P2. The exact source position is located in the center of the corresponding red circle. Location is performed using vertical component Z, i.e.  $N_{\text{Sta}} = 3$ . In each map, the color is normalized by the maximum probability. (a) Location of signals in frequency band 13-17 Hz using simulations from the domain with topography. (b) Location of signals in frequency band 13-17 Hz using simulations from the domain with a flat surface. (c) Location of signals in frequency band 3-7 Hz using simulations from the domain with topography. (d) Location of signals in frequency band 3-7 Hz using simulations from the domain with a flat surface.

For the low-frequency band, when including topographic effects, Figure 10c, the location of the source P1 is well reconstructed, while the location is more ambiguous for the source P2 and spatially scattered within an area of size up to 300 m. The imperfect location could be caused by the 4s time window cutting a part of the low frequency signal. As expected, when topographic effects are not included (flat model), Figure 10d,

the inferred locations become more blurred. With longer probing wavelengths (i.e. for 522 the low frequency band), we would expect reduced location resolution for both models. 523 The good reconstruction of source P1 in this low-frequency band for the model with to-524 pography is therefore puzzling. This cannot be explained by the proximity of station BOR, 525 since better resolution of the source P1 would in that case also be observed when using 526 the flat model. This might be the signature of topography effects, since the source P1 527 is located just below the crater rim, one of the steepest regions of the crater geometry, 528 leading to generated signal characteristics quite distinct from those of neighboring po-529 tential locations. 530

When both sources are acting simultaneously (Figure 10, right column), positions 531 of the individual sources can no longer be determined for all the test cases. Taking into 532 account topography effects, Figures 10a and c, the probability distribution of source lo-533 cation inferred from the high-frequency band (Figure 10a) is spatially scattered with rel-534 atively high-probability patches of around 300 m size in the neighborhood of the indi-535 vidual sources, while the distribution inferred from the low-frequency band is focused 536 in a single region that seems to best explain the superposed signal from the two impacts. 537 In other words, a single source in this region would result in similar relative energy mea-538 surements at the stations as the superimposed signal from the two sources. Interestingly, 539 a small shift to the south is observed, similar to what occurs when locating real rock-540 falls in this area (see Fig. 8a and 9a), suggesting that the observed shift could partly be 541 caused by the superposition of impacts from several boulders at different locations. 542

When ignoring topographic effects with the flat wave propagating model, Figures 10b and d, the inferred probability distribution of source location becomes smoother and less spatially resolved, since recorded seismic signals then contain only information on the source-receiver distances. For the low-frequency band, the relatively high-probability areas are loosely defined and shift further away from the actual source positions.

To summarize, the individual sources can be well located only from high-frequencies 548 and when taking into account topography effects. When the two sources are acting si-549 multaneously, i.e. impacting at the same time and with same force, the individual sources 550 can no longer be distinguished and the location probabilities are concentrated somewhere 551 in the vicinity between the sources. Similar results were reported by Kumagai et al. (2009) 552 who numerically tested the amplitude source location (ASL) method with two simulta-553 neous, spatially separated sources, resulting in the best location being between the two 554 sources. Nonetheless it is important to bear in mind that for real rockfalls, radiating sources 555 are non-uniform in space and time. This means that recorded signals are dominated by 556 the signature of the most strongly radiating sources at a given time. This makes it pos-557 sible to locate the strongest sources in space at each time and to reconstruct rockfall tra-558 jectories reasonably well. 559

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### 5.2 The influence of the network geometry on the location resolution

The previous analysis is extended by quantifying the decay of the location probabilities as a function of distance from the actual source. In this way, a proxy for the spatial resolution is defined for each grid point, which can serve as an array response function for single-impact sources.

Considering each grid point as the actual source position, the location probability of all other grid points is calculated (resulting in probability maps as in Figure 10). Then, assuming a circular symmetry of the probability as a first order approximation, the source probability p as a function of distance d from the actual source is approximated by an exponential decay of the following form:

$$p(d) = a \exp(-dk) + b, \tag{4}$$

where a, k, and b are fitting parameters. The fit is performed without considering the probability value at the actual source position to avoid influence from singularities (such as the high probability at the source position in Figure 10a). Finally, a proxy R for the spatial resolution is defined by the *half-life* of the exponential decay using rate constant k:

$$R = \frac{\ln(2)}{k}.$$
(5)

The shift b can be ignored, since only relative probability variations are important.

Figure 11 presents the analysis performed for a network using the vertical components Z of all four stations. The decay of probabilities with distance is shown for three points as example in Figure 11a, where points P1 and P2 correspond to the points analyzed in Figure 10. The fitted exponential function is shown in red and half-decay is marked with black dashed lines. The uncertainty  $\delta R$  on resolution proxy R has been propagated from the fitting error  $\delta k$  in the rate constant k.



Figure 11. Resolution proxy R from a network of four stations with vertical components. (a) Source location probability as a function of distance from the actual source at P1, P2, and P3, respectively. Blue circles in the distribution correspond to all the other grid points of potential source locations. The fitted exponential function is shown in red. Black dashed lines mark the position of resolution proxy R, defined by the half-decay. The plots are normalized so that p(0) = 1 and  $p(d \rightarrow \infty) = 0$ . (b) Map of the resolution proxy R, constructed by calculating the half-decay for each grid point in case it is the actual source, as shown for points P1, P2, and P3.

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The map of resolution proxy R in Figure 11b shows a median resolution of  $R_{\text{med}} = 107.4 \text{ m}$  and generally indicates values below 100 m in the northwest (as for points P1 and P2) and values above 100 m in the southeast (as for point P3). The southeastern region is not enclosed by the network geometry which may explain the poorer resolution. The poorer resolution in the vicinity of station DSO may be caused by the proximity of this station to the crater rim, suggesting that this is not an optimal position for locating seismic sources.

To evaluate the influence of the network geometry on the resolution, the analysis was performed with one of the stations removed alternately. Figure 12a shows the resulting maps of resolution proxy R using a reduced network of three stations with vertical components only. Generally, the resolution becomes poorer in the direction of the removed station while the three remaining stations form a triangle that spans an area of enhanced resolution, best seen in i) and iii). This triangle is not so clearly visible in ii) and iv) because of the generally poorer resolution in the southeast (see Fig. 11b).

The median resolution  $R_{\text{med}}$  decreases slightly compared to the previous analysis with four stations (except for case iv). If fewer stations are involved, the accumulated



Figure 12. Influence of the network geometry on the resolution proxy R. (a) Map of resolution when using three stations of vertical components Z. Removed station is shown in grey. Station combinations consist of i) BON-BOR-SNE, ii) BON-DSO-SNE, iii) BON-BOR-DSO, and iv) BOR-DSO-SNE. (b) Map of resolution when using four stations and all available channels, i.e. three component ENZ of stations BON, BOR, and SNE; vertical component Z of station DSO.

misfit at positions in the vicinity of the actual source is lower, resulting in higher probabilities and steeper decay of the exponential curve. However, the relative median uncertainty  $R_{\text{med}}$  increases by a factor between 2 and 3 (from 1.7% to 3.4%-4.8%), indicating a more scattered probability distribution.

On the contrary, adding measurements increases the median resolution as can be 602 seen in Figure 12b with  $R_{\rm med} = 131.8 \,\mathrm{m}$ , where all available station components have 603 been combined (i.e. three components of BON, BOR, and SNE, and vertical-component 604 of DSO). Nevertheless, the relative median uncertainty decreases to 1.3% and the spa-605 tial variation of the resolution is smoother compared to the response from vertical com-606 ponents only (Figure 11b). This suggests that the location method is more stable with 607 an increased number of measurements which help to better determine the solution space. 608 This is analogous to findings of Kraft et al. (2013) whose optimal network design algo-609 rithm, which takes into account laterally variable noise levels, often extends an estab-610 lished network with stations near existing station locations to further enhance the seis-611 mic source resolution. Their algorithm is based on the linearized earthquake location prob-612 lem (D-criterion), first implemented by Kijko (1977). Toledo et al. (2020) use the same 613 theory to develop a network design tool for seismic sources in geothermal and volcanic 614 contexts. Their study shows how the first four stations can significantly improve the cost-615 benefit given optimal locations, while the added value decreases with each additional sta-616 tion in a power-law like manner. 617

For the given network at Dolomieu crater, the above tests indicate enhanced res-618 olution in the area which is enclosed by the network geometry. This is an effect which 619 can be observed in previous rockfall location studies (e.g. Lacroix & Helmstetter, 2011; 620 Gracchi et al., 2017) and agrees with findings from optimal network design studies (e.g. 621 Rabinowitz & Steinberg, 1990). In the present case, the southeast part of the crater shows 622 lower resolution, explaining the poor location of the rockfall on February 28, 2016, in Fig-623 ure 9b. Adding additional measurements, such as horizontal channels, can increase the 624 stability of the solution with only a slight loss of resolution, which is especially impor-625

tant in the field when measurements can be contaminated by noise or when the site amplification functions are poorly known, which is in agreement to the rockfall location results in Figure 8a and e using three components (ENZ) and the vertical component (Z), respectively.

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## 5.3 Multiple-point sources mimicking a down-slope moving rockfall

Synthetic rockfall seismic signals are generated here from a downward moving seismic point source, i.e. parametrized as a single vertical traction, kinematically constrained by the boulder trajectory observed during one rockfall event at the Dolomieu crater (December 13, 2016) already discussed in section 4.1. The space-and-time positions of the seismic point-source is mapped in Figure 13. Another representation of the space-time trajectory of the seismic source is shown in the graph at the top of Figure 14a.

To construct the source space-and-time trajectory of the point-source, the position 637 and time of seven markers (Figure 14a) during the rockfall were determined manually 638 from the analysis of the video images of the rockfall event on December 13, 2016. The 639 time and space positions of the source between the selected markers are interpolated in-640 cluding small fluctuations using the  $10 \times 10$  m spatial grid covering the observed rockfall 641 trajectory, leading to a total of 200 impacts. The source can be activated at the same 642 spatial position a number of times as the trajectory is spatially discretized by only 60 643 cells. 644

Synthetic seismic signals, hereafter designated the reference signals, are generated at the different stations from all the source positions and activation times, using the wave propagating model including topography. The source-time function (7 Hz Ricker wavelet) and the amplitude of the vertical traction are the same for all the impact sources. The corresponding generated signals can be seen in Figure A1b.

In the test, the location method is applied using: 1) the same topographic model 650 used for the generated reference signals, in a high-frequency (13-17 Hz, defined as base-651 line model) and a low-frequency (8-12 Hz) band of the reference signals (Figures 13 a and 652 b, respectively); 2) a wave propagating model including a low-pass filtered topography 653 with a  $30 \,\mathrm{m}$  corner wavelength reducing the topography resolution from  $10 \,\mathrm{m}$  to  $20 \,\mathrm{m}$  (Fig-654 ure 13c); and 3) a wave propagating model including the original topography but a ve-655 locity model increased by 10% (Figure 13d). Further, the influence of the assumption 656 of vertical rockfall impacts is tested by synthesizing reference signals generated by source 657 impacts normal to the slope and locating them using the same wave propagation model 658 as above from vertical sources (Figure 13e) and a wave propagation model from slope-659 normal sources (Figure 13f). Finally the method is tested after adding different levels 660 of white noise to the reference signal (Figure 13g to i). 661

Results in Figure 13a are expected to be best because the synthetic signals are an-662 alyzed using the same model used for their generation and because the best spatial res-663 olution is expected at high frequencies as already seen before. Nevertheless, in contrast 664 to a single-point source which could be located exactly (Fig. 10a), we can observe a cor-665 ridor of high probability that extends to up to 200 m. This is related to the superposi-666 tion of signals from temporally overlapping sources, which compromises the predictions. 667 Still, the predictions correctly follow the progressively downhill moving active source re-668 gion. 669

The second low-frequency test, Figure 13b, demonstrates that valuable information is also contained in the low-frequency band, even though larger-wavelengths result in lower spatial resolution than when using high frequencies, extending the high probability corridor especially at later times to up to 300 m. The contained information still suggests that developing methods that can exploit information across different frequency bands would be a major improvement, but is beyond the scope of this study.

<sup>676</sup> In the third test, Figure 13c, in which the forward modeling part of the location <sup>677</sup> method includes a smoother representation of the topography, inferred source locations <sup>678</sup> are shifted compared to those inferred from first step, e.g. the source positions between manuscript submitted to JGR: Earth Surface



Figure 13. Location of a synthetic signal for a single rock moving down-slope. Color-filled circles mark the space-and-time positions of the vertical point impacts. The bottom graph shows the generated seismogram (vertical velocity) recorded at BOR and filtered at 13-17 Hz, resulting from a total of 200 impacts (the low amplitude gap in the signal is random and corresponds to the gap at around 12s in the offset-delay distribution in Figure 14a). Signals generated at all stations for all components are shown with scales in Figure A1. (a) Baseline configuration for location at 13-17 Hz using a wave propagation model with a topography resolution of 10 m and all station-channel pairs ( $N_{\text{Sta}} = 7$ ). (b) Location in a lower frequency band at 8-12 Hz. (c) Location using a wave propagation model with a topography resolution of 20 m. (d) Location using a wave propagation model with a 10% faster medium velocity. (e) Location of a synthetic rockfall signal from sources that are directed normal to the topography, referred to as  $S_n$ , using a wave propagation model with vertical sources S<sub>Z</sub>. (f) Location of a synthetic rockfall signal from sources which are directed normal to the topography, referred to as  $S_n$ , using a wave propagation model with normal sources  $S_n$ . (g) Location of a reference signal contaminated by a median white noise level of 3.4%, corresponding to a median signal-to-noise ratio SNR  $\approx$  30, comparable to the noise level observed in this frequency band for the previously analyzed and relatively small rockfall on December 13, 2016 (see Appendix A). (h) Location of reference signal with median SNR  $\approx 2.4$ (Figure A2b). (i) Location of reference signal with median SNR  $\approx 1.4$  (Figure A2c).

20 s and 30 s are shifted by around 50 m towards the south. This stresses the importance 679 of properly resolving the topographic effects at the scale of the frequency bands that are 680 analyzed. Besides the already discussed superposition of multiple sources at different lo-681 cations, the observed southwards shift of the predicted trajectory in section 4.1 can partly 682 be interpreted as possibly resulting from an inaccurate outdated DEM given that the 683 surface topography of the Dolomieu crater is continuously reshaped by high rockfall ac-684 tivity (e.g. Hibert, Mangeney, et al., 2017; Durand et al., 2018; Derrien et al., 2019). Which 685 of the two effects is stronger is inherently dependent on the location of the rockfall and 686 the relative positions and magnitudes of the inferring sources, and cannot be predicted 687 in a general way. 688

To better understand the influence of the velocity model for the source predictions, 689 the location method is applied using a wave propagation model including the original 690 topography but with a modified velocity model in which velocities are globally increased 691 by 10%, Figure 13d. This also influences intrinsic attenuation by decreasing the velocity-692 dependent absorption coefficient (e.g. Aki & Richards, 2002). The inferred source loca-693 tions don't differ significantly from the best reference test with the original velocity model, 694 Figure 13a, in the same frequency band. This might appear to be surprising as in this 695 test the forward modeling part of the location method is computed using the same to-696 pography resolution but with a different velocity model. However in this modified ve-697 locity model, seismic velocities are uniformly increased by 10%, which does not signif-698 icantly alter the energy ratios between the different stations. More systematic scenar-699 ios, including possible spatially localized velocity perturbations and local site effects at 700 the stations, need to be investigated in the future to properly assess the influence of the 701 a priori uncertainties in the seismic velocity model on the performance of the location 702 method. In the case that information about the subsurface properties is available, it can 703 be considered in the spectral-element based 3D propagation model and can therefore taken 704 into account in the proposed location method. This is in contrast to other locating meth-705 ods where the seismic velocity is used as a free parameter and optimized during the lo-706 cating process to maximize the correlation between stations (e.g. Burtin et al., 2013; Hi-707 bert et al., 2014; Dietze et al., 2017; Pérez-Guillén et al., 2019), therefore three-dimensional 708 velocity models cannot be considered. 709

The energy ratios from the wave propagation model are generally computed un-710 der a vertical source assumption, even though a rockfall can generate forces normal and 711 tangential to the slope. Kuehnert et al. (2020) showed that wave propagation along the 712 topography dominates over the source direction, suggesting that the source direction and 713 the resulting radiation patterns are a second order effect for location. To support this 714 assumption and verify that non-vertical forces generated from the rockfall on the ground 715 can indeed be ignored, a new reference signal is generated from sources directed normal 716 to the slope, which we refer to as the  $S_n$ -signal. In the test, the location method is then 717 applied using a wave propagation model with vertical sources, referred to as the  $S_{Z}$ -model, 718 Figure 13e; and using a wave propagation model with normal sources, referred to as the 719 S<sub>n</sub>-model, Figure 13f. 720

Results from the first test with the vertical source assumption, Figure 13e, don't
differ significantly from the previous test with a reference signal from vertical sources,
Figure 13a, suggesting that the direction of the rockfall source impact does not influence
the performance of the location method.

This conclusion is further supported by the second test using a wave propagation model with normal sources, Figure 13f, where the outline of the predicted source distribution is displaced by a maximum of 50 m compared to the previous distribution in Figure 13e. Given the unpredictability of the source field in the case of real rockfalls, we conclude that the vertical source assumption is very reasonable and the most straightforward solution when predicting rockfall trajectories with the proposed location method.

Finally, noise contaminated reference signals were located. In Figure 13g, the noise causes the probability distribution to be slightly blurred with location probabilities reducing by around 10% and the width of the spatial distribution increasing by around

100 m compared to the noise-free test in Figure 13a. Here, the added white noise is sim-734 ilar in amplitude as observed at 13-17 Hz in the signals from Dolomieu crater, i.e. a me-735 dian noise level of 3.4% or median signal-to-noise ratio of SNR = 30.0 for the rockfall 736 on December 13, 2016, which does not comprise large volumes, as can be seen in Fig-737 ure 6. Higher noise levels, Figure 13h and i, increasingly blur the predicted source prob-738 ability distribution with location probabilities reducing by around 30% and 70%, and 739 the width of the spatial distribution increasing by around 200 m and 400 m, respectively, 740 compared to the noise-free test in Figure 13a. The tests suggest that the location method 741 is robust to noise levels many times higher than those observed at Dolomieu crater, and 742 that at an average SNR of 2.4, the rockfall trajectory can still be tracked reasonably well 743 with an error of about 200 m. 744

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## 5.4 Distributed point sources mimicking complex rockfalls and granular flows

To increase complexity, the downward moving seismic source is activated twice with a respective time shift of 20 s, as shown in Figure 14a where the top graph shows the spacetime trajectory of two successive boulders. As a consequence of the respective time shift, the first boulder arrives at the bottom of the crater while the second boulder is still located in the top half of the crater wall, visible by the red-filled circles for times > 40 s in the map of Figure 14a.

Location results show a high-probability corridor of around 200 m width compa-753 rable with the probability distribution from the single-boulder test in Figure 13a. How-754 ever, the superposition of the two simultaneously acting sources compromises the time 755 resolution of the method, i.e. the color sequence is mixed so that, for example, the red 756 color for times > 40 s is strongly scattered and spread along almost the whole crater wall. 757 This loss of spatio-temporal resolution due to superposition of multiple sources was al-758 ready observed in the previous test with single sources P1 and P2, Figure 10, and partly 759 explains the poor spatio-temporal location of the rockfall on February 28, 2016, located 760 on the southeastern crater wall, Figure 9b. 761

In a second test, a large distributed source with variable impact amplitudes was 762 constructed, aiming to synthesize the characteristics of a granular flow. The space-time 763 distribution of a total of 10,000 source impacts, presented in Figure 14b, is constructed 764 by defining a minimum and a maximum velocity curve as well as a third curve in between 765 where impact amplitudes are maximum. The total number of impacts reaches its maximum at around 10s and decays subsequently towards zero after 45s. The source area, 767 marked by color-filled circles on the map in Figure 14b, is spatially discretized by 87 cells 768 (selected from the  $10 \times 10$  m spatial grid), each of which can be activated multiple times 769 to simulate a total of 10,000 impacts. The corresponding generated signals are shown 770 in Figure A1d. 771

Despite the superposition of the numerous distributed seismic sources, high-probability
predictions are correctly located on the northwestern crater wall. The width of the probability distribution of up to 300 m is very similar to the one from the real granular event
on June 14, 2016, Figure 9d, and the global downward movement of the sources is well
captured and can be followed by means of the correctly ordered color sequence.

## 6 Conclusion

We propose a new rockfall location method based on seismic energy ratios between stations. In an optimization routine, observed energy ratios are compared to a database of simulated energy ratios in a region of interest. The benefit of the method is that once the database has been created, locations can be estimated quickly without the need for complicated analyses of the seismic signal such as precise picking of arrival times. The rockfall seismic signals are analyzed in sliding time windows, making it possible to follow the rockfall trajectory over time. The method can therefore potentially be used for



**Figure 14.** (a) The graph on top shows space-time distribution of two source trajectories mimicking two successive boulder tracks of 200 impacts each. Red asterisks mark source locations of the rockfall on December 13, 2016 as estimated from video images that serve as interpolation points. Under the map, the generated reference signal is shown, recorded at BOR and filtered at 13-17 Hz. The map shows picked source positions as well as location results at 13-17 Hz using a wave propagation model with a topography of 10 m resolution. (b) Space-time distribution of 10,000 sources mimicking a granular flow. The sources are distributed within two velocity curves. An additional curve inbetween defines sources of maximum impact amplitude. The amplitude is represented in arbitrary units with a minimum amplitude of 1. Under the map, the generated reference signal is shown, recorded at BON and filtered at 13-17 Hz. Signals generated at all stations for all components are shown with scales in Figure A1. Note that the signal amplitude is controlled by both the amplitude of each individual impact and the number of sources that act simultaneously. The map shows picked source positions as well as location results at 13-17 Hz using a wave propagation model with a topography of 10 m resolution.

continuous monitoring in real time, in parallel with existing methods that detect and classify rockfall seismic signals (e.g. Dammeier et al., 2016; Dietze et al., 2017; Provost et al., 2017; Maggi et al., 2017; Hibert, Provost, et al., 2017; E. J. Lee et al., 2019).

By direct numerical modeling of the wave field on a domain representing the study
site, no assumptions about the wave type of the recorded signal are required, high-resolution
surface topography and its influence on the wave field can be accounted for, and a priori information about subsurface properties and a corresponding three-dimensional seis-

mic velocity model can be considered and are not required to be estimated during the
 location process.

Here, location was performed for rockfalls at Dolomieu crater, Reunion Island. All 794 analyzed rockfall events could be located in the correct area of the crater. Generally, the best spatial resolution (below 100 m) is observed in the beginning of the rockfall when 796 the seismic source is very confined in space. Thereafter, the predicted source locations 797 become more scattered. This is linked to the spatial distribution of the seismic source, 798 comprising multiple simultaneous impacts at different positions. The superposition of 799 multiple sources is not considered in the method and hence compromises the location 800 results. Nonetheless, the method performs remarkably well in this regard and is even able 801 to locate a downward moving granular flow, likely because the signals are dominated by 802 the signature of the most radiating sources at a given time. 803

It is shown that the influence of the assumed source impact direction on the loca-804 tion is of the second order, since propagation along the Dolomieu crater topography dom-805 inates over source-characteristic radiation patterns for the investigated frequencies above 806 about 3 Hz. Thus, a vertical surface traction can be assumed, even though the actual source 807 field of real rockfalls remains unknown. Furthermore, the insignificance of the source-808 characteristic radiation patterns makes it possible to use all vertical and horizontal com-809 ponent signals for location, which makes the method more robust against ambient noise 810 or poorly known site amplifications. 811

Experiments with synthetic rockfall sources confirmed that the best spatial resolution is achieved at high frequencies (here in the frequency band of 13-17 Hz). For future development of the method, a combination of multiple and possibly also narrower frequency bands should be considered. The synthetic tests also revealed that a precise representation of the surface topography is crucial to the quality of the location results.

Investigations on the influence of the network geometry on the resolution suggests 817 that best resolution (below 100 m) is achieved when the source area is triangulated by 818 the seismic stations. The method currently assumes that the signal of a seismic source 819 arrives fully within the defined time window at all stations. This is possible because of 820 the positions of the seismometers with respect to the rockfalls at Dolomieu crater but 821 might be a limitation for other source-receiver geometries. In order to overcome this lim-822 itation, a time shift can potentially be introduced at each station with respect to the re-823 gion of interest after estimating the approximate arrival times. 824

Comparisons with other location methods that are able to track moving seismic surface sources, as for example the approach of amplitude source location (ASL, e.g. Pérez-Guillén et al., 2019), need to be carried out at the same study site and using the same station network to assess the benefits of each method and compare their resolution.

Noise levels at Dolomieu crater are very low at the here studied frequencies above 3 Hz and could be ignored when locating the observed rockfalls. Tests with added white noise showed that the location method is robust against noise levels that are considerably higher than those observed at Dolomieu crater, and that using signals in the 13-17 Hz frequency band with an average SNR of 2.4, rockfall trajectories can still be tracked with an error of about 200 m.

No significant effects on the location results were found when modifying the subsurface velocity model. However, seismic velocities in the test were uniformly increased by 10%, which does not significantly alter the energy ratios between the different stations. More systematic scenarios, including possible spatially localized velocity perturbations and local site effects at the stations, will need to be investigated in future studies to properly assess the influence of the a-priori uncertainties in the seismic velocity model on the performance of the location method.

#### Appendix A Comparison of real, synthetic and noise-contaminated syn-842 thetic rockfall signals 843

Seismograms generated by the above analyzed rockfall events on December 13, 2016 and on June 14, 2016 are shown in Figures A1a and c, respectively. They are shown together with synthetic rockfall signals, mimicking the real events, in Figure A1b for the boulder-type event synthesized and analyzed in section 5.3 and in Figure A1d for the granulartype event synthesized and analyzed in section 5.4.



Figure A1. Seismograms of all station channels for real rockfalls at Dolomieu crater and synthetic signals mimicking these events. Signals of each event are normalized and real seismograms are bandpass filtered at 1-35 Hz. (a) Signals of boulder-type event on December 13, 2016, analyzed in Figures 6 to 8. (b) Synthetic signals generated by 200 point impacts mimicking a down-slope moving rock, analyzed in Figure 13. (c) Signals of granular-type event on June 14, 2016, analyzed in Figure 9d. (d) Synthetic signals generated by 10,000 point impacts mimicking a granular flow, analyzed in Figure 14.

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To test the location method with noise-contaminated synthetic signals, the noise levels on the observed signals at Dolomieu crater are analyzed using the rockfall event 850 on December 13, 2016. Observed signals filtered at 13-17 Hz and corresponding signalto-noise ratios (SNR) are shown in Figure A2a. The minimum SNR is 15.1 while the me-

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dian SNR is 30.0, which corresponds to a median noise level of 3.4%. For the relatively small rockfall, the SNR is high, which indicates a low noise level at the Dolomieu crater at these high frequencies.

v (m/s) (a) Real rockfall SNR=30	(b) Synth SNR=30	(c) Synth SNR=2.4	(d) Synth SNR=1.4
BON_E	BON_E	— BON_E	- BON_E
SNR=15.	9 SNR=29.0	SNR=2.2	SNR=1.4
BON_N	BON_N	— BON_N	- BON_N
0	7 SNR=27.8	SNR=2.3	SNR=1.4
1 BON_Z	— BON_Z	— BON_Z	BON_Z
0 SNR=27	1 SNR=29 7	SNR=2.5	SNR=1 4
1 BOR_E	BOR_E	BOR_E	BOR_E
0 - SNR=40.	7 SNR=235.0	SNR=16.2	SNR=7.9
1 BOR_N	BOR_N	- BOR_N	- BOR_N
0 - SNR=79.	6 SNR=128.7	SNR=8.7	SNR=4.3
1 BOR_Z	BOR_Z	BOR_Z	BOR_Z
0 - SNR=65.	8 SNR=279.8	SNR=19.3	SNR=9.6
1 DSO_Z	DSO_Z	DSO_Z	DSO_Z
0	5 SNR=95.4	SNR=6.8	SNR=3.2
1 SNE_E	SNE_E	SNE_E	SNE_E
0	5 SNR=10.9	SNR=1 4	SNR=1 1
1 SNE_N			
0	9 SND-11 0	SND_1 1	
1			
0			
0 20 40 60	0 20 40 60	0 20 40 60	0 20 40 60
Times (s)	Times (s)	Times (s)	Times (s)

Figure A2. Comparison of signal-to-noise ratios (SNR) for real and synthetic signals, filtered at 13-17 Hz and with normalized amplitudes. (a) Observed signal at Dolomieu crater generated by rockfall on December 13, 2016, with minimum SNR of 15.1 and median SNR of 30.0, which corresponds to a median noise level of 3.4%. (b) Synthetic signal from the test in section 5.3, contaminated with a median noise level of 3.4% (i.e. median SNR  $\approx$  30), similar to the level on the observed rockfall signal. (c) Synthetic signal contaminated with a median noise level of 41.5 % (i.e. median SNR  $\approx$  2.4), 12 times higher than the noise level on the observed rockfall signal. SNR values are close to 1 at station SNE, hiding the signals almost entirely. (d) Synthetic signal contaminated with a median noise level of 71.0% (i.e. median SNR  $\approx 1.4$ ), 21 times higher than the level on the observed rockfall signal. SNR values are close to 1 at both station BON and SNE, hiding the signals almost entirely.

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White noise is now added to the synthetic rockfalls in Figure A1b. Figure A2a shows 856 the synthetic signals contaminated with a noise level so that the median SNR  $\approx 30$ , 857 comparable to the SNR observed at Dolomieu crater. Noise levels are then increased by 858 a factor of 12 and a factor of 21, resulting in the synthetic signals shown in Figures A2b 859 and c, respectively. A factor of 12 increases the median noise level to 41.5% (i.e. me-860 dian SNR  $\approx 2.4$ ), hiding almost entirely the signal at station SNE. A factor 21 increases 861 the median noise level to 71.0 % (i.e. median SNR  $\approx$  1.4), hiding not only signals at 862 SNE, but also at BON. 863

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