

Dawsonite and ankerite formation in the LDX-1 structure, Yinggehai basin, South China sea: An analogy for carbon mineralization in subsurface sandstone aquifers

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1 Introduction

 Carbon-dioxide emissions to the atmosphere by anthropogenic industrial activities are a key factor causing global climate change (IPCC, 2005, 2007; Manabe 35 and Stouffer, 1993). The injection of anthropogenic $CO₂$ into deep geologic 36 formations has been advocated to slow increasing atmospheric $CO₂$ concentrations and its effect on global climate (Bachu et al., 1994; Bachu, 2003; IPCC, 2005, 2007; Oelkers and Cole, 2008). There are four mechanisms of carbon storage in the 39 subsurface: structural and stratigraphic trapping, residual $CO₂$ trapping, solubility trapping and mineral trapping (IPCC, 2005; Johnson et al., 2001). Of these, mineral trapping, and in particular in-situ carbon mineralization, is the most stable and safe carbon storage solution (Lackner et al 1995; Johnson et al., 2001; Matter et al., 2016; Oelkers et al., 2008; Power et al. 2013).

 Sandstone aquifers with impermeable cap-rocks are commonly considered for carbon storage as they are have adequate permeability and pore space, and are abundant in the subsurface (Benson and Cole, 2008; IPCC, 2005; Xu et al., 2005). A number of studies have shown that calcite, siderite, magnesite, dolomite, ankerite and 48 dawsonite could form in various sandstones in response to $CO₂$ injection. (Baker et al., 1995; Fischer et al., 2010; Gao et al., 2009; Gaus et al., 2005; Hellevang et al., 2013; Higgs et al., 2013, 2015; Hitchon, 1996; Klajmon et al., 2017; Liu et al., 2011; Moore et al., 2005; Pauwels et al., 2007; Pearce et al., 2016; Pham et al., 2011; Uysal et al., 2011; Van Pham et al., 2012; Worden, 2006; Xu et al., 2004, 2005, 2007, 2010; Yu et al., 2015; Zhao et al., 2018; Zhou et al., 2014). Of these minerals, dawsonite (NaAl(OH)₂CO₃) is particularly attractive as a carbon storage host as it contains sodium and aluminum, which can be abundant in sedimentary silicates. This mineral, however, is more reactive and less stable than many of the aforementioned divalent

metal carbonate minerals (Bénézeth et al., 2007; Hellevang et al., 2005, 2011).

 The LDX-1 structure in the Yinggehai basin of South China Sea provides an insightful analogy for the study of mineral carbonation in sandstone aquifers (Fig. 1). 60 The gas phase in the LDX-1 structure contains as much as 93 volume percent $CO₂$ 61 and this $CO₂$ has a distinct inorganic carbon isotopic signature (He et al., 2005; Huang 62 et al., 2003). In addition, past studies concluded that the $CO₂$ gas injection into this structure occurred after that of the hydrocarbon gases (Huang et al., 2004; Sun and Wang, 2000), which is similar to scenerios that might be encountered by injecting into former oil field reserviors.

 The purpose of this study is to characterize the mineralogy and paragenesis in two wells drilled into the LDX structure to investigate sandstone diagenesis and carbonate 68 mineral formation in response to the influx of $CO₂$. Results will be used to gain further insight into the potential for mineral carbonation in sedimentary sandstones as part of subsurface carbon storage efforts.

2 Geological setting

 The Yinggehai basin, which developed during the Cenozoic, is a natural gas-rich rift basin. It is located between the Indian and South China plate, formed by the combination of lithosphere extension and the deep, strike-slip Red River fault. The direction of this basin is Northwest-Southeast, with a total area of approximately 11×10^4 km² (Huang et al., 2009; Wan et al., 2007; Fig. 1a, b). Two northwest-trending, deep-to-basement faults divide the Yinggehai basin into three sub-structural units: The Central Depression, the Ying Dong Slope and the Ying Xi Slope (see Fig 1). The central depression consists of two tectonic units: the Lin Gao uplift and the diapir structure belt (Fig 1b; Zhu et al., 2004). During the Miocene, the Yinggehai basin was filled with argillaceous sediments along the fault and formed a series of diapir structures in the central depression zone (Huang et al., 2009). He et al. (2005) reported that most of the 84 natural gas reservoirs with high $CO₂$ contents are concentrated in these diapir structures.

 The basement of the Yinggehai basin is dominated by palaeozoic magmatic rocks (Hao et al., 1998). According to the drilling strata, the Cenozoic sediments include the Eocene Lingtou Formation, the Oligocene Yacheng and Lingshui Formations, the Miocene Sanya, Meishan, and Huangliu Formations, the Pliocene Yinggehai Formation and the Quaternary Ledong Formation. The thickness of the Cenozoic sediments in this basin is up to 16-17 km (Gong et al., 1997; Huang et al., 2004, 2009; Meng et al., 2012; Zhu et al., 2007) (Fig. 2). The Lingtou Formation mainly consists of sandstone and mudstone. The Yacheng and Lingshui Formations are composed of mudstone and silty mudstone interbeded with mudstone. Both the Sanya and Meishan Formations are dominated by calcareous mudstone. The Huangliu, Yinggehai and Ledong Formations are mainly siltstone, sandstone and interbedded mudstone. The Miocene mudstone is the main source rock of natural gas in this basin, while the interbedded mudstone and sandstone of the overlying Huangliu and Yinggehai Formations provide a good reservoir and cap rock system for the basin (Wang and Huang, 2008). According to the 100 logging data, the Yinggehai and Huangliu formations contain the main $CO₂$ gas reservoirs in the basin.

 The LDX-1 Structure is a dome structure located in the southeast of the Ledong area in the central depression of the Yinggehai Basin (Fig. 1b). The gas reservoir in the LDX-1 structure is an anticline formed by a mud diapir. The rocks containing the CO₂-rich gas reservoirs consist mainly of the marine sandstones of the Pliocene Yangehai Formation (Fig. 2) and are located at a depth of 392 m to 2300 m (Huang et al., 2004; Li et al., 2005). According to gas measurements, the LDX-1 reservoirs are 108 dominated by hydrocarbon gas and CO_2 ; the CO_2 content ranges from 41.9% to 78.7% (Huang et al., 2004).

 There were two major episodes of gas emplacement into the Yinggehai basin, the first from an organic source and dominated by CH⁴ and the second from an inorganic 112 sourced and dominated by CO₂. According to the Xie et al., (2014), the arrival of the

113 CH₄ dominated gas was at approximately 3.7 Ma. The $CO₂$ dominated gas arrived later. Estimates of its arrival time range from 0.3 to 1.9 Ma (Fu et al., 2016; Huang et al., 2005; Wang et al., 2004; Xie et al., 2014). As such this basin provides a close 116 natural analogue to the filling of a former hydrocarbon reservoir with $CO₂$ as part of subsurface carbon storage efforts.

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- *3 Sampling and research methods*

 A total 159 sandstone samples were collected from the gas reservoir layers of the LDX-1 structure of the Yinggehai Formation. The samples were collected from the depths of 1841.62 to 1869.02 m (from well LDX-1-1) and of 1958.07 to 1905.13 m (from well LDX-1-4). The compositions of the sandstone samples were determined by counting 300 points in each thin section following the approach of Dickson (1965) using an OLYMPUS BX51 polarizing microscope. To identify the carbonate minerals, all the thin sections were stained with Alizarin Red S and K-ferricyanide solutions following the procedure of Dickson (1965). Authigenic mineral observations and energy spectrum analyses were performed at the Research Center of Paleontology and Stratigraphy of Jilin University, using a JEOL JSM-6700F scanning electron microscope (SEM), equipped with energy dispersive X-ray spectrometer (EDS) analyzer manufactured by Oxford, UK; mineral surfaces were gold coated prior to this analysis.

 Based on a systematic petrographic study of the sandstone in the gas reservoirs, 10 dawsonite-bearing samples were selected for carbon and oxygen isotope analyses using MAT253 isotopic ratio mass spectrometers at the Analytical Laboratory of The China National Nuclear Corporation (CNNC) Beijing Research Institute of Uranium Geology. The analytical method for obtaining stable isotopic compositions is based on past studies such as Liu et al. (2011). Samples were grounded into a powder using an agate mortar and pestle. The ground powers were sieved and the > 74 μm fraction was 140 collected, then dried for 2 h at 105 °C under atmospheric pressure, to remove potential

 surface volatile phases, before analyzing. For carbon and oxygen isotope analyses, 99% phosphoric acid (produced by Merck KGaA, Darmstadt, Germany) was reacted with 143 crushed whole rock powders at 25 °C or 50 °C (McCrea, 1950). For dawsonite analysis, 144 CO₂ released from samples during the first 4 h reactions at 25 °C was discarded, as this gas likely originated from the decomposition of calcite. After that, the reaction was 146 continued for another 24h and the $CO₂$ was collected for the analysis of dawsonite isotopic compositions (Becker, 1971; Rosenbaum and Sheppard, 1986; Sharma and 148 Clayton, 1965). For ankerite analysis, the $CO₂$ released during the first 24 h reaction at 25 ºC was discarded, as it likely originated from calcite and dawsonite decomposition, 150 then the reaction temperature was raised to 50 \degree C followed by the collection of the CO₂ produced by the decomposition reaction during subsequent 48 h (Uysal et al., 2011). 152 The precision of the analyses is $\pm 0.1\%$ for δ^{13} C and $\pm 0.2\%$ for δ^{18} O. All reported isotope values in this study are relative to the Belemnitella americana from the Peedee Formation (PDB) standard for carbon and Standard Mean Ocean Water (SMOW) for oxygen.

 All the speciation and solubility calculations were conducted using PHREEQC v. 3 (Parkhurst and Appelo, 2013) with its Lawrence Livermore National Laboratory (LLNL) thermodynamic database after adding to it the stability constant for ankerite. 159 The stability of ankerite in this study is based on the formula $(Ca(Fe_{0.5}Mg_{0.5})(CO₃)₂)$ (see below). Its thermodynamic properties were estimated from the stoichiometrically weighted sum of the calcite, siderite and magnesite stability constants (Xu et al., 2004). The thermodynamic properties required for this estimate were calculated using the SUPCRTBL database (Zimmer et al., 2016). The resulting ankerite equilibrium constant and enthalpy of reaction are 0.642 and -69.326 kJ/mol at 165 70 °C and -0.087 and -77.803 kJ/mol 100 °C. Thermodynamic properties in the LLNL database are consistent with those of dawsonite reported by Bénézeth et al. (2007). The Truesdell-Jones equation was used to calculate aqueous activity coefficients and 168 the fugacity for CO_2 gas was calculated using the ThermoSolver 1.0 software (Barnes, 2006).

4 Results

4.1 Petrographic characteristics

 The dawsonite in the LDX-1 structure is hosted in fine-grained subarkose and lithic quartz sandstones. The observed detrital minerals are mainly quartz (16.7% to 64.7%), feldspar (up to 12.0%) and rock fragments (up to 23.0%). The modal compositions of the detrital minerals are shown schematically in Fig 3. The feldspar is mainly comprised of plagioclase and also contains small amounts of orthoclase and perthite. The rock fragments mainly consist of metamorphic debris (up to 19%), granite debris (up to 8%), and sedimentary debris (up to 7%).

4.2 Authigenic mineralogy and paragenesis

 The authigenic minerals observed in the rock samples are dawsonite, quartz, siderite, dolomite, ankerite, mixed-layer clays consisting of illite and smectite, and kaolinite.

4.2.1 Dawsonite

 Dawsonite is principally found at depths of 1841.27 to 1867.23 m in well LDX-1-1 and 1863 to 2018 m in well LDX-1-4 with its content ranging from 0.2 to 6 Vol% as determined by point counting (see above). Dawsonite aggregates nucleate as microcrystalline grains then grow radially outward into the sandstone pores. Dawsonite growth usually terminates at the edge of the detrital minerals (eg. quartz and feldspar). Dawsonite could also be observed growing on dolomite (Fig. 4a, b), potassium feldspar and plagioclase (Fig 4c, d). The EDS results show that the dawsonite mainly consists of carbon, oxygen, sodium and aluminum having weight percents of 7.25 to 22.23, 50.43 to 66.93, 3.58 to 14.69, and 7.27 to 30.26 respectively (see Table 1).

4.2.2 Authigenic quartz

 Two kinds of authigenic quartz were observed: quartz overgrowths and microcrystalline quartz. The content of authigenic quartz is low, generally less than 0.5%. Quartz overgrowths are the among the earliest authigenic silicate minerals observed and are evident due to clay rims between detrital quartz and these overgrowths. The thickness of the quartz overgrowths ranges from 0.01 to 0.06 mm. Furthermore, we observed some radial dawsonite surrounded by quartz overgrowths (see Fig. 4e), suggesting some of the quartz precipitated concurrently with or after the dawsonite. In addition, SEM images show that the microcrystalline quartz was distributed in sandstone pores sporadically and is associated with kaolinite. This kaolinite is also associated with dawsonite in some places (Fig. 4f). Such observations further suggest that the microcrystalline quartz formed during same time period as the dawsonite.

4.2.3 Siderite

 Most siderite was distributed in pores as subhedral intergranular crystals or aggregates along the boundaries of the detrital grains (Fig. 4g). The content of the siderite ranges from 0.5% to 13.5% in the bulk rock and the crystal sizes range from 0.01 mm to 0.03 mm. The color of the siderite is yellow brown to dark brown under single polarized light. The siderite crystals were distributed either in the inner side of quartz overgrowth or near the clay rims surrounding the original quartz grain boundaries, indicating that siderite formation was earlier than or concurrent with the quartz overgrowths.

4.2.4 Dolomite

 The dolomite content (0.5 to 43 Vol%) varied substantially in the studied cores. As dolomite is not stained by Alizarin Red S and K-ferricyanide solutions, it is readily distinguished from other carbonates. Under the polarizing microscope, some dolomite was observed being replaced by dawsonite aggregates (Fig. 4h). SEM analysis revealed that the ends of some needle-like dawsonite crystals terminate at the faces of dolomite crystals exhibiting a "bridged phenomenon" (Figure 4a, b). Both sets of observations imply that the dolomite formed earlier than the dawsonite, yet the presence of some detrital dolomite cannot be discounted.

4.2.5 Ankerite

 Ankerite (0.5 to 36 Vol%) mainly surrounds dolomite crystals as rim-like growths 234 (Fig. 4i) or as euhedral to subhedral intergranular crystals (Fig. 4j, k). Under the Back Scattered Electron (BSE) mode, the ankerite is seen to form of nearly equal ring-like shapes along the dolomite surfaces as it replaced this mineral. Some ankerite is seen to have grown on dawsonite (Fig 4j, k). Such observations suggest that ankerite formation occurred later than both the dolomite and dawsonite. EDS analyses results for ankerite showed that its Ca, Mg, and Fe content ranged from 0.39 to 9.22, 2.4 to 6.52, and 1.47 to 7.82 atomic percent respectively (see Table 2), consistent with the formula 241 Ca(Fe_{0.5}Mg_{0.5})(CO₃)₂.

4.2.6 Clay minerals

244 The XRD results indicate that the clay minerals in the $CO₂$ -containing reservoirs of the Yinggehai Basin mainly consist of mixed-layer illite and smectite, and kaolinite (Huang et al., 2004; Jiang et al., 2015;Tong et al., 2012). The mixed layer illite and smectite commonly surround the detrital minerals as "clay rims", which could be identified by polarizing microscope. Mixed layer illite and smectite are not found in the sandstone pores, suggesting that it is the earliest formed secondary mineral in the diagenetic sequence. Honeycomb-like authigenic clay is seen to grow on K-feldspar (see Fig. 4l, m). According the EDS results shown in Table 3, most of the smectite and mixed layer illite are Ca depleted. This suggests that the Ca in the smectite and mixed layer of illite and smectite may have been removed over time through water-mineral reactions.

 Authigenic kaolinite grows as vermicular assemblages in the sandstone pores. Hexagonal kaolinite crystals are found attached to dawsonite crystals (Fig. 4n), at times 257 adjacent to small "etch pits" on the dawsonite surface $(Fig. 40)$, suggesting that some kaolinite post-dates the dawsonite. However, vermicular kaolinite is commonly interspersed between acicular dawsonite crystals, suggesting that this kaolinite formed slightly earlier than or concurrently with dawsonite (Fig. 4p). In some places ankerite 261 grew on the top of kaolinite aggregates ($Fig. 4q$). This suggests that the formation of kaolinite might have spanned from the period from when dawsonite grew through that of ankerite formation.

 Chlorite was not observed in the samples considered in this study. Its presence in the rocks of the LDX system was reported by Fu et al. (2016). Moreover, the presence 266 of chlorite within and below the $CO₂$ gas reservoirs is inferred as a potential source of 267 the iron leading to ankerite precipitation (see below).

4.3 Vertical characteristics of authigenic carbonates in the reservoirs

 The dawsonite and ankerite contents across the study area, measured on 159 thin sections, are shown in Fig. 5. Over the depths of this study (from 1841.6 m to 1868.89 m), dawsonite (0.7 to 8.3%) and ankerite (1.7 to 26.0%) in the LDX-1-1 well mainly developed near the gas water contact (GWC) (Fig. 5a). Owing to lack of samples, few observations of these minerals were made above the GWC in this well. In well LDX-1-4, the limited measurements of mineral content in the gas reservoirs show the dawsonite and ankerite content ranging from 0.7 to 2.3% and 2.0 to 14.0% above the GWC, whereas the corresponding content results for dawsonite and ankerite under the GWC in this well are 1.0 to 6.3% and 1.0 to 24.0 % (Fig. 5b).

4.4 C and O isotope compositions

281 The stable isotopic compositions of dawsonite and ankerite are shown in Table 4 282 and Fig. 6. The carbon isotope (δ^{13} C) values for dawsonite vary from -4.3‰ to -0.6‰ 283 with an average value of -2.0‰. The oxygen isotope $(\delta^{18}O)$ values for dawsonite range 284 from +22.4‰ to +24.0‰. The δ^{13} C values for ankerite range from -2.6‰ to -0.8‰ with 285 an average value of -1.4‰ and the δ^{18} O values range from +22.8‰ to +23.7‰.

286 The measured δ^{13} C values of dawsonite and ankerite have been used to estimate 287 the corresponding equilibrium δ^{13} C of the CO₂ (δ^{13} C_{CO2}) from which they precipitated. 288 The $\delta^{13}C_{CO2}$ for the CO₂ that formed the dawsonite, was estimated assuming dawsonite 289 exhibits an identical $CO₂$ carbon isotope fractionation as that of calcite, using the 290 equation proposed by Ohmoto and Rye (1972) (Baker et al., 1995). The calculated $\delta^{13}C$ 291 for the $CO₂$ in equilibrium with ankerite was estimated assuming its fractionation was 292 identical to the dolomite- CO_2 fractionation equation proposed by Horita (2014). The 293 δ^{13} C of the CO₂ (δ^{13} C_{CO2}) calculated to be in isotopic equilibrium with dawsonite range 294 from -9.66‰ to -6.69‰ with an average value of -7.90‰. The equilibrated $\delta^{13}C_{CO2}$ for 295 ankerite varies from -7.27‰ to -5.94‰ with a mean value of -6.4‰.

 The formation temperature for dawsonite and ankerite were calculated from their oxygen isotope compositions. This was done by assuming oxygen fractionation 298 between dawsonite and H_2O is identical to the calcite- H_2O fractionation equation proposed by O'Neil et al. (1969) and that oxygen fractionation between ankerite and H_2O is identical to the dolomite-H₂O fractionation equation of Horita (2014). The ¹⁸O 301 value for water used in these calculations was $\delta^{18}O_{\text{H2O}} = +2.7\%$ (Deng and Chen, 1994), which was measured in the formation waters of the DFY-1 structure in the Yinggehai-Huangliu Formation of the Yinggehai basin (Fig. 1). The calculated temperature for dawsonite formation varies from 67.8 ºC to 79.6 ºC and the calculated temperature of ankerite formation ranges from 89.6 ºC to 96.4 ºC (see Table 4). These temperatures are consistent with 1) the earlier formation of dawsonite and 2) the increasing temperature of the study site over time as the basin subsided. Note that the 308 current temperature of the study site is 92 to 103 $^{\circ}$ C.

5 Discussion

5.1 The origin of the carbon in the dawsonite and ankerite

312 The $\delta^{13}C_{CO2}$ of the gas forming dawsonite and ankerite reported above and in 313 Table 4 are similar to the current δ^{13} C values in the CO₂-rich reservoirs of the LDX 314 system; the current $\delta^{13}C_{CO2}$, are reported to range from -8.2‰ to -4.2‰ (He et al., 2005). Moreover, the hydrogen and oxygen isotopic compositions of the formation water in the Yinggehai Basin reservoirs are significantly different from that of meteoric water, which implies the lack of exchange between meteoric water and that present in the LDX system (Deng and Chen, 1994). Such evidence suggests that the CO₂ in the gas reservoir is likely a major carbon source for the both dawsonite and ankerite formation.

 Carbon dioxide in the LDX gas reservoirs at the present time is likely dominated 322 by magmatic derived CO_2 . Firstly, the $\delta^{13}C_{CO2}$ value in the gas reservoir ranges from -8.2‰ to -4.2‰, which is significantly different from that obtained from kerogen 324 degradation; $\delta^{13}C_{CO2}$ from kerogen degradation ranges from -25‰ to -20‰ (Irwin et 325 al., 1977; Tissot and Welte, 1984) and $CO₂$ from bacterial fermentation typically has a 326 $\delta^{13}C_{CO2} \sim +15\%$ (Irwin et al., 1977). Secondly, the carbon isotope values of CO₂ in 327 the LDX system exhibits a $\delta^{13}C_{CO2}$ range similar to the -8‰ to -3‰ range observed for CO² sourced from mantle-derived magma (Baker et al., 1995; Clayton et al., 1990; Dai, 1996; Ohmoto and Rye, 1972; Rollinson, 2014;). Taking account of the 330 observations summarized above, we suggest that origin of the $CO₂$ in the LDX-1 structures is mantle-derived magma degasing.

 The timing of dawsonite and ankerite formation can be constrained by the timing 333 of the basaltic magmatism related $CO₂$ charging of the LDX system. The Yinggehai Formation, where the dawsonite and ankerite developed, was formed between 5.5 Ma and 1.9Ma (Fig. 2). The Honghe fault zone, which controls the formation and distribution direction of Yinggehai basin is still active (Zhu et al., 2009) (Fig. 1). The

 ages of alkaline basalt in this region range from 4 to 6 Ma and from 1 Ma up to 338 present (Zhu and Wang, 1989). Several ages of the $CO₂$ charging into the Yinggehai Formation have been reported in the literature; these ages range from 0.3 to 1.9 Ma (Fu et al., 2016; Huang et al., 2005; Wang et al., 2004; Xie et al., 2014). Based on these ages, it seems reasonable to conclude the dawsonite and ankerite in the Yinggehai Formation is formed during the last 1.9 Ma.

5.2 Cation source of dawsonite and ankerite

 The dissolution of feldspars in the LDX-1 structure provides the cations for dawsonite. The formation of dawsonite from feldspar has been observed in numerous systems. For example, feldspars were observed to be altered to dawsonite in the Lam Formation of Shabwa basin Yemen (Worden, 2006), in the reservoirs of the Castillo structure of the Golfo San Jorge basin (Comerio et al., 2014), and in the Songliao basin, the Hailaer basin, and the East Sea basin in China (Gao et al., 2009; Liu et al., 2011; Zhao et al., 2018; Zhou et al., 2014). The reaction forming dawsonite from albite can be written (Bénézeth et al., 2007; Hellevang et al., 2005; Ryzhenko, 2006; Xi et al., 2016): NaAlSi₃O₈ + CO₂ + H₂O \Rightarrow NaAlCO₃(OH)₂ + 3SiO_{2(aq)} (Eq. 1)

Albite Dawsonite

355 where some of the liberated $SiO_{2(aq)}$ could provoke quartz precipitation. The equivalent

reaction forming dawsonite from K-feldspar:

$$
357 \quad \text{KAISi}_3\text{O}_8 + \text{Na}^+ + \text{CO}_2 + \text{H}_2\text{O} \rightleftharpoons \text{NaAlCO}_3(\text{OH})_2 + 3\text{SiO}_{2(aq)} + \text{K}^+ \tag{Eq. 2}
$$

K-feldspar Dawsonite

359 has been discounted by Hellevang et al. (2011) due to the likely build-up of K^+ in the aqueous phase arresting this reaction.

 Evidence for the coupling of feldspar dissolution to dawsonite precipitation reactions could be found throughout the LDX system (Fig. 4c, d, r). If dawsonite 363 originates from K-feldspar by Eq. 2, however, its growth also requires $Na⁺$ from the fluid phase. This would require the system to be relatively open to fluid transport, which may be limited in this system. Note that according to mass conservation, for 366 every 1 cm³ of dawsonite formed from albite by reaction 1, there should be 1.16 cm³ of SiO₂ precipitation (Zhou et al., 2014). However, far less microcrystalline quartz is 368 observed in our system than dawsonite (Fig. 4b), which suggests that some of the $SiO₂$ may have been consumed by other water-mineral reactions.

 Geochemical calculations results suggest that chlorite dissolution provoked by 371 the charging of the system by CO_2 could provide Fe²⁺ for ankerite formation (Xu et al., 2005, 2007, 2014; Yu et al., 2014). The dissolution of smectite also could provide some of the iron for ankerite growth (Hellevang et al., 2013; Pham et al., 2011; Worden and Morad, 2003; Worden, 2006), as well as some aluminum promoting dawsonite formation. The dissolution of chlorite and smectite coupled to ankerite formation has been observed in the sandstone reservoirs of the Otway basin of Austrailia,, in the Songliao Basin of China, the Jurassic and late Cretaceous reservoirs of North Sea oil fields, and in the Eocene strata from Wilcox of Texas, (Awwiller, 1993; Hendry et al., 2000; Liu et al., 2011; Thyberg et al., 2010; Watson et al., 2004). The chlorite content of the Yinggehai Basin $CO₂$ bearing reservoirs has been reported 381 to be significantly lower than that in $CO₂$ –free reservoirs (Tong et al., 2012). Similar observations have been reported in wells of the Yinggehei Basin (Huang et al., 2004; Jiang et al., 2015). Such observations suggest that the dissolution of iron bearing clay 384 minerals such as chlorite is the main Fe^{2+} source for ankerite. The Mg^{2+} and some $Ca²⁺$ in the ankerite crystals could come from dolomite dissolution, as fresh ankerite is 386 observed growing within a dolomite matrix (Fig. 4i). Additional Ca^{2+} can be sourced from the transformation of clays in the system; a number of studies have reported Ca release as smectite transforms to mixed layered clays and illite (e.g. Boles and Frank, 1979; McHargue and Price, 1982). A close relation between ankerite and smectite is evident in the studied core samples (see Fig. 7). Taking account of these observations, a representative reaction coupling chlorite dissolution to ankerite precipitation could be expressed as (modified from Kirste et al., 2004; see also Luquot et al., 2012):

$$
393 \t Fe2Al2SiO5(OH)4 + 2 Ca(MgCO3)2 + SiO2 + 2 Ca2+ + 4 CO2 + 2 H2O =
$$

Chamosite Dolomite Quartz

395 $4 \text{Ca}(Fe_{0.5}Mg_{0.5})(CO_3)_2 + Al_2Si_2O_5(OH)_4 + 4 H^+$ (Eq. 3) Ankerite 396 Kaolinite

397 Note this reaction, balanced on Fe and Mg, requires substantial magmatic $CO₂$ for ankerite formation, consistent with its carbon isotope composition. This reaction also consumes SiO2, which may account, at least in part, for the low mass of authigeic quartz in the present day system.

5.3 CO² partial pressure and carbonate mineral stability

 It is likely that both ankerite and dawsonite are formed after the arrival of mantle sourced CO2, providing a clear analogy for modern carbon capture and storage systems. The observations described in this study indicate that dawsonite is first 406 formed at ~70 °C, followed by ankerite at ~90 °C, in response to the arrival of $CO₂$ into the overlying reservoirs. The increase in temperature in the system likely stems from its increased burial depth over time. Ankerite is also reported to be the last forming mineral in the sandstone reservoirs in Songliao (Liu et al., 2011; Yu et al., 2015) and Hailaer basins (Gao et al., 2009). Some kaolinite is seen to precipitate coupled to the dissolution of dawsonite; some dissolution etch pits on dawsonite crystal were evident near the kaolinite (Fig. 4m). Kaolinite precipitation continues as ankerite precipitates, consistent with Eq. (3). The formation of ankerite also appears to be linked to dolomite dissolution.

 All of these observations are consistent with the phase diagrams shown in Figs. 8 416 to 10. In Fig. 8, the relative stability of dawsonite in the $Na_2O-Al_2O_3-SiO_2-CO_2-H_2O$ system balanced on aluminum and saturated with quartz is shown. Dawsonite and 418 quartz become stable relative to albite at log $(f_{CO2}) > 0.15$ and 0.9 at 70 and 100 °C, respectively. This diagram thus suggests that albite would be favored to transform to 420 dawsonite as $CO₂$ fills the reservoirs in accord with reaction 1. As suggested by Fig. 9, some of the precipitated dawsonite and quartz will tend to react to form kaolinite due to an increase in temperature, which tends to shrink the dawsonite stability field. The 423 value of log (f_{C_Q2}) in equilibrium with the assemblage albite = dawsonite + quartz at 424 70 and 100 °C plot as vertical lines on the Fig. 10, which also shows the relative stability field of ankerite + kaolinite relative to that of chlorite, dolomite and quartz 426 consistent with reaction 3. It can be seen that for all $a_{Ca2+}/a_{H+}^2 > -9.7$ at 70 °C and $>$ 427 \sim 7.3 at 100 °C, dawsonite would first tend to form before ankerite as the CO₂ partial 428 pressure in the reservoirs increase. The partial pressure of $CO₂$ in the LDX system 429 would increase as the $CO₂$ fraction of the gas progressively increased in the reservoirs located just above the studied rock samples (Fig. 5). The persistence of dawsonite and 431 ankerite in this system is assured by the continued presence of the $CO₂$ -rich gas 432 reservoir, which maintains f_{CO2} at relatively high levels. Note the current temperature of the gas reservoirs are 92.3 and 102.9 ºC for well LDX-1-4 and well LDX-1-1 respectively, and the corresponding gas pressure in reservoirs, which is dominated by CO2, are between 18.9 and 43.2 Mpa (China National Offshore Oil Company, personal communication). Hellevang et al. (2005), noted, that dawsonite, would likely 437 rapidly dissolve once it became thermodynamically unstable as the f_{CQ2} decreased in the such systems.

5.4 The paragenetic sequence

 Based on the observations described above, the proposed paragenetic sequence for the studied dawsonite-bearing sandstones is shown in Fig. 11. Secondary clays (mainly illite and smectite mixed layer clays) began forming before compaction during the early stage of diagenesis and coated detrital minerals as clay rims. Sideite is found to form next; siderite precipitation is readily observed in reducing, non-sulphidic pore waters in the suboxic and methanogenesis zones (Morad, 1998). Quartz overgrowths are then 447 observed to form, perhaps due to the recrystallization of biogenic $SiO₂$, but also accompanied with the formation of minor kaolinite in response to feldspar dissolution.

 Subsequently, dolomite is observed to form, potentially promoted by Ca and Mg release from the dissolution or transformation of the earlier formed clays. After the 451 influx of $CO₂$ into the reservoir. dawsonite and additional quartz forms from the dissolution of feldspar, and subsequently ankerite and kaolinite forms from the dissolution of dolomite, quartz, chlorite and potentially some clay minerals..

5.5 Implications for carbon storage in sandstone reservoirs

 The formation and long-term stability of dawsonite and ankerite in the LDX system are likely good analogies for some modern carbon storage systems such as the Shenhua CCS demonstration project in China (Yang et al., 2017). The distribution of dawsonite and ankerite in the LDX reservoirs (Fig. 5) indicates that these minerals can form in the water layers beneath gas water contact (GWC). Some dawsonite and ankerite were also observed above the present day GWC, which may indicate that either 1) the position of the GWC has changed over time or 2) that these phases grew in the gas filled section of the reservoir, potentially on water-wet surfaces. This later, process, however, is likely limited as the dawsonite and ankerite forming reactions 1 465 and 3 consume H₂O. The formation of such carbonate minerals adjacent to the $CO₂$ gas reservoir is also supported by the results of TOUGHREACT calculations reported 467 by Pan et al. (2017) that suggest that when $CO₂$ is injected into sandstone reservoirs, carbonate minerals would form on both upper and lower layers of the reservoir. Carbonate mineral precipitation in the water layers below the GWC has also been observed in New Zealand sandstone reserviors (Higgs et al., 2015). As such, we suggest 471 that the $CO₂$ in the gas reservior maintains a sufficient $CO₂$ partial pressure in the formation waters below the GWC to promote the growth and long term stability of dawsonite and ankerite. However, according to the paragenetic sequence determined in this study, kaolinite precipitated from the fluid concurrently with the precipitation of dawsonite and ankerite. Numerous petrographic observations and experimental simulations suggest that kaolinite and dawsonite formation could significantly reduce sandstone permeability (Luquot et al., 2012; Shiraki and Dunn, 2000; Yu et al., 2012). 478 Kaolinite also commonly grows in response to CO_2 -feldspar or CO_2 -clay mineral interaction (Luquot et al., 2012; Moore et al., 2005; Shiraki and Dunn, 2000; Yu et al., 2012). Shiraki and Dunn (2000) concluded that the precipitation and growth of kaolinite in the pores could block pore throats leading to the reduction of the permeability of reservoir. Similarly, Worden (2006) suggested that the fiber-like crystal morphology of dawsonite would also reduce sandstone permeability. The reduced permeability in the sandstone could slow chemical transport, and including potentially 485 the transport of dissolved $CO₂$ out of the storage system. Such processes could thereby promote the stability of carbonate minerals in this system. The decrease in permeability could also lead to fluid pressure build up in the gas reservoir as further minerals form in this system.

6 Conclusions

491 This study focused on the mineral reactions provoked by the influx of $CO₂$ into the subarkose and sublitharenite sandstones of the Yinggehai basin. This system is likely a good analogy for subsurface carbon storage in sandstone hosted hydrocarbon reservoirs 494 as the influxed CO_2 replaces a methane-rich gas phase in an anticlinal structure. This 495 leads to the presence of a CO_2 -rich gas phase stable over a ~1-2 million year time frame. 496 The influx of provoked the $CO₂$ provoked the formation of dawsonite from plagioclase and ankerite and kaolinite likely from chlorite, quartz, and dolomite, at temperatures of 498 68 to 80 °C. and 90 to 96 \degree C, respectively, largely just below the CO₂-gas water interface. The formation of such phases could provide for the mineral storage of some of the $CO₂$ injected during subsurface carbon storage efforts.

 The key to maintaining dawsonite and ankerite stability in such systems appears to 502 be preserving a high partial pressure of $CO₂$. As such, although the formation of these 503 minerals might lead to the storage of $CO₂$ injected into sandstones as part of engineered geologic storage efforts, such minerals may not provide for the 'permanent' mineral storage of this gas that one might anticipate from the formation of carbonate 506 minerals such as calcite, which can be stable at far lower $CO₂$ partial pressure, unless the system tends to become impermeable as due to the formation of various secondary minerals.

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 Figure 1. Location map of LDX-1 structure in Yinggehai Basin, south China Sea area. a) The location of Yinggehai 835 basin and basalt area modified after Shi and Yan (2011), basalt areas are modified from Hoang et al. (1996, 1998), Li

836 et al. (2005), Wang et al. (2001); Zhou et al. (1997) and Zhu et al. (1989), b) Geotectonic units modified after Luo et

al., (2003) and Meng et al. (2012).

840 Figure 2. Simplified stratigraphic column of the Yinggehai basin modified after Gong et al. (1997) and Huang et al.

(2003).

844 Figure 3. a) Image of a thin section of a sandstone sample collected from a depth of 1860.72 m in the LDX-1-1 well 845 showing its detrital minerals: Q=quartz, F=feldspar and R=rock fragment. b) Mineral composition diagram, after 846 Folk (1968), of the dawsonite-bearing samples from well LDX-1-1 (blue rhombuses) and well LDX-1-4 (black

- 847 triangles). All the data were obtained through point-counting see text.
- 848

 Figure 4. Micrographs of diagenetic minerals. (a, b) Acicular dawsonite growing on top of dolomite (red crosses are EDS probe spots), SEM; (c, d) Dawsonite replacing feldspar and plagioclase with EDS analysis results (red crosses are EDS probe spots); (e) Dawsonite replacing quartz overgrowths, thin section micrograph, cross polar; (f) Microcrystalline quartz grows with kaolinite on the top of drusy dawsonite, SEM; (g) Siderite surrounding detrital

 Figure 5. Plot of volume percent versus depth for dawsonite (blue rhombuses) and ankerite (green triangles) in well LDX-1-1 (Fig 3a) and well LDX-1-4 (Fig 3b). The filled gray areas and dashed boxes indicate the location of gas layers, the water layers in the reservoir system are below the gas layers. The gas-water-contact (GWC), as reported by the China National Offshore Oil Company are located at 1844.1m and 1886.0m depth in wells 879 LDX-1-1 and LDX-1-4, respectively.

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883 Figure 6. Plot of $\delta^{13}C$ versus $\delta^{18}O$ values for dawsonite and ankerite samples. Filled symbols represent the 884 compositions of dawsonite, and open symbols for ankerite The filled red triangles and open blue rhombus 885 correspond to compositions measured in this study from the LDX-1 structure, Yinggehai basin, China. 886 Compositions of dawsonite from the Bowen-Gunnedah-Sydney basin system in Australia reported by Baker et al. 887 (1995) are shown as filled blue rhombuses, those from the late Permian coal measures of the Upper Hunter Valley 888 in Australia reported by Golab et al. (2006) are shown as filled dark red squares, those of the Cretaceous Hailaer 889 basin in China reported by Gao et al. (2009) are shown as filled yellow crosses, those of the upper Cretaceous 890 Songliao Basin in China reported by Liu et al. (2011) are shown as filled purple circles, and those of the Yaojie 891 coalfield, Minhe basin in China reported by Ming et al. (2017) are shown as filled green triangles. Compositions 892 of ankerite from the Cretaceous Hailaer basin in China reported by Gao et al. (2009) are shown as open red 893 squares, and those of the upper Cretaceous Honggang anticline of the Songliao Basin in China reported by Liu et 894 al. (2011) are shown as open green triangles.

897
898 Figure 7. Images of clay minerals observed in this study. (a) detrital smectite in the sandstone and (b) image showing 899 the relationship between smectite and ankerite. Rhombic ankerite crystals grew and adhered to the mixed layer of 900 illite/smectite, suggesting the transformation of smectite could provide some of the materials for ankerite. I/S: mixed

904 Figure 8 Log fugacity of CO_2 ($log(f_{CO2})$) versus temperature (°C) diagram depicting mineral stability fields in the 905 system Na₂O-Al₂O₃-SiO₂-CO₂-H₂O balanced on Al and at quartz equilibrium. The grey area indicates the stability 906 field of dawsonite, and the white area indicates the stability field for albite. The orange and blue lines highlight the $907 \log(f_{CO2})$ values in equilibrium with dawsonite and albite at 70 and 100 °C, respectively. This figure was 908 generated using PHREEQC together with its llnl database – see text.

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911 Figure 9. Log fugacity of CO_2 ($log(f_{CO2})$) versus log a_{Na}/a_{H+} diagrams depicting mineral stability fields in the 912 system Na₂O-Al₂O₃-SiO₂-CO₂-H₂O balanced on Al and at quartz equilibrium at 70 °C and 100 °C. The dashed 913 area represents the current log (*f*_{CO2}) range in the reservoirs considered in this study. This figure was generated 914 using PHREEQC together with its llnl database – see text.

916

 \implies Estimate fluid composition and its evolution direction due to CO₂ fugacity increases

917 Figure 10. Phase diagram of log CO₂ fugacity (log (f_{CO2})) versus log a_{Ca2+}/a_{H+}^2 . The dashed vertical lines 918 correspond to equilibrium compositions of reaction 1, whereas the sloped lines correspond to equilibrium 919 compositions of reaction 3, The grey area is the likely aqueous activity ratio of a_{Ca2}/a_{H+}^2 in the studied system and 920 the arrow illustrates the direction of the CO_2 fugacity evolution in the system as the reservoir fills with CO_2 . This 921 figure was generated using PHREEQC together with its llnl database after adding estimated values for the 922 thermodynamics properties of ankerite– see text. 923

Figure 11. Paragenetic sequence for dawsonite-bearing sandstones in this study. The black bars are indicated by

928 petrologic observations, the grey dashed bars are estimated based on the mass balance considerations and reactions 929 (1) and (3).

931

932 **Table 1. EDS analysis data for dawsonite.**

| | | | Weight Percentage, % | | | Atomic Percentage, % | | | | | | | |
|-----------|-------------|-------|----------------------|-------------|------------|----------------------|-------|----------------|-------------|-------|--|--|--|
| Well | $\mathbf C$ | Na | \mathbf{A} l | $\mathbf O$ | Si | $\mathbf C$ | Na | \mathbf{A} l | $\mathbf O$ | Si | | | |
| $LDX-1-1$ | 18.76 | 4.98 | 12.99 | 63.27 | n.d. | 25.14 | 3.48 | 7.75 | 63.63 | n.d. | | | |
| $LDX-1-1$ | 15.88 | 11.13 | 14.19 | 58.8 | n.d. | 22.01 | 8.06 | 8.76 | 61.18 | n.d. | | | |
| $LDX-1-1$ | 11.81 | 12.44 | 12.96 | 55.55 | 7.23 | 17.14 | 9.44 | 8.38 | 60.55 | 4.49 | | | |
| $LDX-1-4$ | 14.88 | 13.8 | 14.24 | 57.09 | n.d. | 20.87 | 10.11 | 8.89 | 60.13 | n.d. | | | |
| $LDX-1-4$ | 15.61 | 12.53 | 13.7 | 58.15 | n.d. | 21.71 | 9.11 | 8.48 | 60.7 | n.d. | | | |
| $LDX-1-4$ | 12.62 | 14.69 | 17.98 | 54.72 | $\rm n.d.$ | 18.19 | 11.06 | 11.54 | 59.21 | n.d. | | | |
| $LDX-1-4$ | 12.64 | 8.37 | 14.29 | 57.5 | 7.2 | 18.15 | 6.28 | 9.14 | 62 | 4.43 | | | |
| $LDX-1-4$ | 7.83 | 6.83 | 8.83 | 55.28 | 21.23 | 11.88 | 5.41 | 5.96 | 62.97 | 13.78 | | | |
| $LDX-1-4$ | 12.58 | 12.01 | 19.97 | 55.44 | n.d. | 18.13 | 9.04 | 12.82 | 60.01 | n.d. | | | |
| $LDX-1-4$ | 9.14 | 14.68 | 24.74 | 51.45 | n.d. | 13.75 | 11.54 | 16.57 | 58.13 | n.d. | | | |
| $LDX-1-1$ | 10.16 | 13.06 | 23.91 | 52.87 | n.d. | 15.09 | 10.13 | 15.81 | 58.96 | n.d. | | | |
| $LDX-1-1$ | 11.59 | 9.4 | 23.75 | 55.26 | n.d. | 16.9 | 7.16 | 15.42 | 60.52 | n.d. | | | |
| $LDX-1-1$ | 18.76 | 4.98 | 12.99 | 63.27 | n.d. | 25.14 | 3.48 | 7.75 | 63.63 | n.d. | | | |
| $LDX-1-1$ | 11.97 | 10.19 | 18.04 | 55.91 | 3.88 | 17.36 | 7.72 | 11.64 | 60.86 | 2.41 | | | |
| $LDX-1-1$ | 7.25 | 12.06 | 30.26 | 50.43 | n.d. | 11.18 | 9.71 | 20.76 | 58.35 | n.d. | | | |
| $LDX-1-1$ | 13.42 | 13.99 | 16.92 | 55.67 | n.d. | 19.16 | 10.43 | 10.75 | 59.66 | n.d. | | | |
| $LDX-1-1$ | 13.68 | 13.29 | 16.93 | 56.11 | n.d. | 19.46 | 9.88 | 10.72 | 59.94 | n.d. | | | |
| $LDX-1-1$ | 18.6 | 7.9 | 11.21 | 62.28 | n.d. | 24.98 | 5.54 | 6.7 | 62.78 | n.d. | | | |
| $LDX-1-1$ | 12.75 | 11.24 | 20.18 | 55.83 | n.d. | 18.34 | 8.45 | 12.92 | 60.29 | n.d. | | | |
| $LDX-1-1$ | 10.17 | 12.49 | 24.3 | 53.04 | n.d. | 15.1 | 9.69 | 16.07 | 59.14 | n.d. | | | |
| $LDX-1-4$ | 13.71 | 7.2 | 12.99 | 58.84 | 7.26 | 19.43 | 5.33 | 8.2 | 62.63 | 4.4 | | | |
| $LDX-1-4$ | 11.98 | 13.11 | 16.51 | 55.02 | 3.38 | 17.39 | 9.93 | 10.66 | 59.92 | 2.1 | | | |
| $LDX-1-4$ | 10.86 | 13.85 | 21.98 | 53.31 | n.d. | 16 | 10.65 | 14.41 | 58.94 | n.d. | | | |
| $LDX-1-4$ | 22.23 | 3.58 | 7.27 | 66.93 | n.d. | 28.65 | 2.41 | 4.17 | 64.77 | n.d. | | | |
| $LDX-1-4$ | 10.18 | 12.73 | 24.1 | 52.99 | n.d. | 15.12 | 9.88 | 15.93 | 59.07 | n.d. | | | |
| $LDX-1-4$ | 12.26 | 13.58 | 19.45 | 54.7 | n.d. | 17.75 | 10.27 | 12.53 | 59.44 | n.d. | | | |

933 **n.d.: not be detected.**

936 **n.d.: not be detected.**

938 **Table 3. EDS analysis data for smectite and the mixed layer of illite and smectite**

| Well | Weight Percentage, % | | | | | | | | Atomic Percentage, % | | | | | | | | | |
|-----------|----------------------|----------------|------|------|-------|-------|-------|-------|----------------------|--------------|----------|------|------|-------|-------|---------|------|------|
| | $\mathbf C$ | \overline{O} | Na | Mg | Al | Si | K | Ca | Fe | C | Ω | Na | Mg | Al | Si | $\bf K$ | Ca | Fe |
| LDX-1-1 | 3.31 | 57 | 2.64 | 0.63 | 6.64 | 26.9 | 1.24 | 0.37 | 1.27 | 5.25 | 67.91 | 2.19 | 0.49 | 4.69 | 18.26 | 0.6 | 0.18 | 0.43 |
| LDX-1-1 | n.d. | 41.41 | n.d. | 1.22 | 9.69 | 35.81 | 6.72 | n.d. | 5.14 | n.d. | 57.05 | n.d. | 1.11 | 7.92 | 28.11 | 3.79 | n.d. | 2.03 |
| $LDX-1-1$ | n.d. | 37.5 | n.d. | 1.55 | 9.06 | 39.51 | 4.7 | n.d. | 7.67 | n.d. | 53.18 | n.d. | 1.44 | 7.62 | 31.92 | 2.73 | n.d. | 3.12 |
| $LDX-1-1$ | n.d. | 47.11 | 1.39 | 1.93 | 14.08 | 26.38 | 2.12 | n.d. | 6.99 | n.d. | 62.32 | 1.28 | 1.68 | 11.04 | 19.88 | 1.15 | n.d. | 2.65 |
| LDX-1-4 | n.d. | 57.17 | 0.7 | 0.99 | 15.27 | 22.26 | 2.53 | n.d. | 1.08 | n.d. | 70.24 | 0.6 | 0.8 | 11.13 | 15.58 | 1.27 | n.d. | 0.38 |
| $LDX-1-4$ | n.d. | 54.36 | n.d. | 4.31 | 8.23 | 11.71 | n.d. | 13.46 | 7.93 | n.d. | 71.16 | n.d. | 3.71 | 6.38 | 8.73 | n.d | 7.03 | 2.98 |
| LDX-1-4 | n.d. | 52.97 | n.d. | 1.18 | 13.35 | 26.82 | 4.82 | n.d. | 0.87 | n.d. | 66.92 | n.d. | 0.98 | 10 | 19.3 | 2.49 | n.d. | 0.31 |
| $LDX-1-4$ | n.d. | 55.21 | n.d. | 1.52 | 11.51 | 27.18 | 1.39 | n.d. | 3.18 | n.d. | 69.01 | n.d. | 1.25 | 8.53 | 19.35 | 0.71 | n.d. | 1.14 |
| $LDX-1-4$ | n.d. | 45.84 | n.d. | 1.44 | 14.38 | 28.31 | 3.13 | n.d. | 6.91 | n.d. | 61.37 | n.d. | 1.27 | 11.42 | 21.59 | 1.71 | n.d. | 2.65 |
| LDX-1-4 | n.d. | 46.01 | n.d. | 1.67 | 12.96 | 29.18 | 2.94 | n.d. | 7.24 | n.d. | 61.6 | n.d. | 1.47 | 10.29 | 22.25 | 1.61 | n.d. | 2.78 |
| $LDX-1-4$ | n.d. | 46.45 | n.d. | 1.52 | 15.13 | 26.85 | 4.9 | n.d. | 5.16 | n.d. | 61.77 | n.d. | 1.33 | 11.93 | 20.34 | 2.67 | n.d. | 1.96 |
| LDX-1-4 | n.d. | 52.5 | n.d. | 1.34 | 13.43 | 26.4 | 2.83 | n.d. | 3.49 | n.d. | 66.84 | n.d. | 1.12 | 10.14 | 19.15 | 1.48 | n.d. | 1.27 |
| $LDX-1-4$ | n.d. | 52.32 | 1.04 | 1.26 | 9.4 | 29.16 | 1.99 | n.d. | 3.28 | n.d. | 66.71 | 0.92 | 1.05 | 7.11 | 21.18 | 1.04 | 0.79 | 1.2 |
| $LDX-1-4$ | n.d. | 39.97 | n.d. | 1.14 | 10.55 | 39.83 | 2.53 | n.d. | 4.37 | n.d. | 55.06 | n.d. | 1.04 | 8.62 | 31.26 | 1.43 | 0.87 | 1.73 |
| LDX-1-4 | n.d. | 32.01 | n.d. | 0.97 | 12.81 | 34.05 | 12.63 | n.d. | 7.54 | n.d. | 47.8 | n.d. | 0.95 | 11.34 | 28.97 | 7.72 | n.d. | 3.23 |
| $LDX-1-4$ | n.d. | 53.67 | 0.98 | 0.99 | 11.99 | 28.2 | 2.38 | n.d. | 1.8 | n.d. | 67.37 | 0.86 | 0.82 | 8.92 | 20.16 | 1.22 | n.d. | 0.65 |
| LDX-1-1 | n.d. | 53.91 | n.d. | 1.68 | 12.75 | 25.32 | 4.2 | n.d. | 2.13 | n.d. | 67.96 | n.d. | 1.4 | 9.53 | 18.18 | 2.17 | n.d. | 0.77 |
| LDX-1-1 | n.d. | 61.56 | 1.25 | 0.99 | 9.68 | 22.71 | 2.14 | n.d. | 1.68 | n.d. | 74.07 | 1.04 | 0.79 | 6.9 | 15.56 | 1.05 | n.d. | 0.58 |
| LDX-1-1 | n.d. | 63.93 | 1.67 | 2.62 | 8.94 | 16.76 | 1.11 | n.d. | 4.97 | n.d. | 76.52 | 1.39 | 2.06 | 6.35 | 11.43 | 0.54 | n.d. | 1.7 |
| LDX-1-1 | n.d. | 35.98 | n.d. | 1.82 | 11.82 | 35.41 | 8.01 | n.d. | 6.96 | n.d. | 51.68 | n.d. | 1.72 | 10.07 | 28.97 | 4.7 | n.d. | 2.86 |

939 **n.d.: not be detected.**

| 40 | |
|----|--|
| | |

941 **Table 4. Carbon and oxygen isotopic compositions for dawsonite and ankerite**

