

Dawsonite and ankerite formation in the LDX-1 structure, Yinggehai basin, South China sea: An analogy for carbon mineralization in subsurface sandstone aquifers

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| 3 | subsurface sandstone aquifers |
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| 12 | |
| 13 | Abstract |
| 14 | The geochemistry and petrology of the LDX-1 structure of the Yinggehai basin, a |
| 15 | natural analog of a sedimentary carbon storage site, was investigated to understand the |
| 16 | consequences of the charging of CO_2 gas in this system. The rocks in this structure are |
| 17 | dominated by subarkose and sublitharenite sandstones. The authigenic minerals formed |
| 18 | after CO ₂ injection are dawsonite, microcrystalline quartz, kaolinite and ankerite. |
| 19 | Dawsonite and ankerite are formed just beneath a CO ₂ bearing anticlinal structure due |
| 20 | to the reactions between silicate minerals (feldspars and clay minerals) and the fluid |
| 21 | phase. Carbon and oxygen isotopic analyses indicate that the main carbon source for |
| 22 | dawsonite and ankerite formation was mantle magmatic CO ₂ . The aqueous activities of |
| 23 | sodium and calcium, the partial pressure of CO ₂ , pH and temperature are the key factors |
| 24 | influencing the stability of the dawsonite and ankerite. The presence of the anticlinal |
| 25 | structure maintaining a locally high CO ₂ partial pressure in the waters beneath this |
| 26 | structure is likely responsible for the observed long-term persistence of dawsonite and |
| 27 | ankerite in this system. |
| | |

Keywords: authigenic carbonate minerals, carbon and oxygen isotopes, dawsonite
stability, carbon capture and storage (CCS), Yinggehai Basin

31

32 1 Introduction

33 Carbon-dioxide emissions to the atmosphere by anthropogenic industrial 34 activities are a key factor causing global climate change (IPCC, 2005, 2007; Manabe 35 and Stouffer, 1993). The injection of anthropogenic CO_2 into deep geologic formations has been advocated to slow increasing atmospheric CO2 concentrations 36 37 and its effect on global climate (Bachu et al., 1994; Bachu, 2003; IPCC, 2005, 2007; 38 Oelkers and Cole, 2008). There are four mechanisms of carbon storage in the 39 subsurface: structural and stratigraphic trapping, residual CO₂ trapping, solubility 40 trapping and mineral trapping (IPCC, 2005; Johnson et al., 2001). Of these, mineral 41 trapping, and in particular in-situ carbon mineralization, is the most stable and safe 42 carbon storage solution (Lackner et al 1995; Johnson et al., 2001; Matter et al., 2016; 43 Oelkers et al., 2008; Power et al. 2013).

44 Sandstone aquifers with impermeable cap-rocks are commonly considered for 45 carbon storage as they are have adequate permeability and pore space, and are 46 abundant in the subsurface (Benson and Cole, 2008; IPCC, 2005; Xu et al., 2005). A 47 number of studies have shown that calcite, siderite, magnesite, dolomite, ankerite and 48 dawsonite could form in various sandstones in response to CO₂ injection. (Baker et al., 49 1995; Fischer et al., 2010; Gao et al., 2009; Gaus et al., 2005; Hellevang et al., 2013; 50 Higgs et al., 2013, 2015; Hitchon, 1996; Klajmon et al., 2017; Liu et al., 2011; Moore 51 et al., 2005; Pauwels et al., 2007; Pearce et al., 2016; Pham et al., 2011; Uysal et al., 52 2011; Van Pham et al., 2012; Worden, 2006; Xu et al., 2004, 2005, 2007, 2010; Yu et 53 al., 2015; Zhao et al., 2018; Zhou et al., 2014). Of these minerals, dawsonite 54 (NaAl(OH)₂CO₃) is particularly attractive as a carbon storage host as it contains 55 sodium and aluminum, which can be abundant in sedimentary silicates. This mineral, 56 however, is more reactive and less stable than many of the aforementioned divalent

57 metal carbonate minerals (Bénézeth et al., 2007; Hellevang et al., 2005, 2011).

58 The LDX-1 structure in the Yinggehai basin of South China Sea provides an 59 insightful analogy for the study of mineral carbonation in sandstone aquifers (Fig. 1). 60 The gas phase in the LDX-1 structure contains as much as 93 volume percent CO_2 61 and this CO_2 has a distinct inorganic carbon isotopic signature (He et al., 2005; Huang 62 et al., 2003). In addition, past studies concluded that the CO_2 gas injection into this 63 structure occurred after that of the hydrocarbon gases (Huang et al., 2004; Sun and 64 Wang, 2000), which is similar to scenerios that might be encountered by injecting into former oil field reserviors. 65

The purpose of this study is to characterize the mineralogy and paragenesis in two wells drilled into the LDX structure to investigate sandstone diagenesis and carbonate mineral formation in response to the influx of CO_2 . Results will be used to gain further insight into the potential for mineral carbonation in sedimentary sandstones as part of subsurface carbon storage efforts.

71

72 2 Geological setting

73 The Yinggehai basin, which developed during the Cenozoic, is a natural gas-rich 74 rift basin. It is located between the Indian and South China plate, formed by the 75 combination of lithosphere extension and the deep, strike-slip Red River fault. The 76 direction of this basin is Northwest-Southeast, with a total area of approximately 11×10^4 km² (Huang et al., 2009; Wan et al., 2007; Fig. 1a, b). Two northwest-trending, 77 78 deep-to-basement faults divide the Yinggehai basin into three sub-structural units: The 79 Central Depression, the Ying Dong Slope and the Ying Xi Slope (see Fig 1). The central 80 depression consists of two tectonic units: the Lin Gao uplift and the diapir structure belt 81 (Fig 1b; Zhu et al., 2004). During the Miocene, the Yinggehai basin was filled with 82 argillaceous sediments along the fault and formed a series of diapir structures in the 83 central depression zone (Huang et al., 2009). He et al. (2005) reported that most of the 84 natural gas reservoirs with high CO₂ contents are concentrated in these diapir 85 structures.

86 The basement of the Yinggehai basin is dominated by palaeozoic magmatic rocks 87 (Hao et al., 1998). According to the drilling strata, the Cenozoic sediments include the 88 Eocene Lingtou Formation, the Oligocene Yacheng and Lingshui Formations, the 89 Miocene Sanya, Meishan, and Huangliu Formations, the Pliocene Yinggehai Formation 90 and the Quaternary Ledong Formation. The thickness of the Cenozoic sediments in this 91 basin is up to 16-17 km (Gong et al., 1997; Huang et al., 2004, 2009; Meng et al., 2012; 92 Zhu et al., 2007) (Fig. 2). The Lington Formation mainly consists of sandstone and 93 mudstone. The Yacheng and Lingshui Formations are composed of mudstone and silty 94 mudstone interbeded with mudstone. Both the Sanya and Meishan Formations are 95 dominated by calcareous mudstone. The Huangliu, Yinggehai and Ledong Formations 96 are mainly siltstone, sandstone and interbedded mudstone. The Miocene mudstone is 97 the main source rock of natural gas in this basin, while the interbedded mudstone and 98 sandstone of the overlying Huangliu and Yinggehai Formations provide a good 99 reservoir and cap rock system for the basin (Wang and Huang, 2008). According to the 100 logging data, the Yinggehai and Huangliu formations contain the main CO₂ gas 101 reservoirs in the basin.

102 The LDX-1 Structure is a dome structure located in the southeast of the Ledong 103 area in the central depression of the Yinggehai Basin (Fig. 1b). The gas reservoir in the 104 LDX-1 structure is an anticline formed by a mud diapir. The rocks containing the 105 CO₂-rich gas reservoirs consist mainly of the marine sandstones of the Pliocene 106 Yangehai Formation (Fig. 2) and are located at a depth of 392 m to 2300 m (Huang et al., 107 2004; Li et al., 2005). According to gas measurements, the LDX-1 reservoirs are 108 dominated by hydrocarbon gas and CO₂; the CO₂ content ranges from 41.9% to 78.7% 109 (Huang et al., 2004).

110 There were two major episodes of gas emplacement into the Yinggehai basin, the 111 first from an organic source and dominated by CH_4 and the second from an inorganic 112 sourced and dominated by CO_2 . According to the Xie et al., (2014), the arrival of the 113 CH₄ dominated gas was at approximately 3.7 Ma. The CO₂ dominated gas arrived 114 later. Estimates of its arrival time range from 0.3 to 1.9 Ma (Fu et al., 2016; Huang et 115 al., 2005; Wang et al., 2004; Xie et al., 2014). As such this basin provides a close 116 natural analogue to the filling of a former hydrocarbon reservoir with CO₂ as part of 117 subsurface carbon storage efforts.

- 118
- 119 3 Sampling and research methods

120 A total 159 sandstone samples were collected from the gas reservoir layers of the 121 LDX-1 structure of the Yinggehai Formation. The samples were collected from the 122 depths of 1841.62 to 1869.02 m (from well LDX-1-1) and of 1958.07 to 1905.13 m 123 (from well LDX-1-4). The compositions of the sandstone samples were determined by 124 counting 300 points in each thin section following the approach of Dickson (1965) 125 using an OLYMPUS BX51 polarizing microscope. To identify the carbonate minerals, 126 all the thin sections were stained with Alizarin Red S and K-ferricyanide solutions 127 following the procedure of Dickson (1965). Authigenic mineral observations and 128 energy spectrum analyses were performed at the Research Center of Paleontology and 129 Stratigraphy of Jilin University, using a JEOL JSM-6700F scanning electron 130 microscope (SEM), equipped with energy dispersive X-ray spectrometer (EDS) 131 analyzer manufactured by Oxford, UK; mineral surfaces were gold coated prior to this 132 analysis.

133 Based on a systematic petrographic study of the sandstone in the gas reservoirs, 10 134 dawsonite-bearing samples were selected for carbon and oxygen isotope analyses using 135 MAT253 isotopic ratio mass spectrometers at the Analytical Laboratory of The China 136 National Nuclear Corporation (CNNC) Beijing Research Institute of Uranium Geology. 137 The analytical method for obtaining stable isotopic compositions is based on past 138 studies such as Liu et al. (2011). Samples were grounded into a powder using an agate 139 mortar and pestle. The ground powers were sieved and the $> 74 \mu m$ fraction was 140 collected, then dried for 2 h at 105 °C under atmospheric pressure, to remove potential

141 surface volatile phases, before analyzing. For carbon and oxygen isotope analyses, 99% 142 phosphoric acid (produced by Merck KGaA, Darmstadt, Germany) was reacted with crushed whole rock powders at 25 °C or 50 °C (McCrea, 1950). For dawsonite analysis, 143 144 CO₂ released from samples during the first 4 h reactions at 25 °C was discarded, as this 145 gas likely originated from the decomposition of calcite. After that, the reaction was 146 continued for another 24h and the CO₂ was collected for the analysis of dawsonite 147 isotopic compositions (Becker, 1971; Rosenbaum and Sheppard, 1986; Sharma and 148 Clayton, 1965). For ankerite analysis, the CO₂ released during the first 24 h reaction at 149 25 °C was discarded, as it likely originated from calcite and dawsonite decomposition, 150 then the reaction temperature was raised to 50 °C followed by the collection of the CO₂ produced by the decomposition reaction during subsequent 48 h (Uysal et al., 2011). 151 The precision of the analyses is $\pm 0.1\%$ for δ^{13} C and $\pm 0.2\%$ for δ^{18} O. All reported 152 isotope values in this study are relative to the Belemnitella americana from the Peedee 153 154 Formation (PDB) standard for carbon and Standard Mean Ocean Water (SMOW) for 155 oxygen.

156 All the speciation and solubility calculations were conducted using PHREEQC v. 157 3 (Parkhurst and Appelo, 2013) with its Lawrence Livermore National Laboratory 158 (LLNL) thermodynamic database after adding to it the stability constant for ankerite. The stability of ankerite in this study is based on the formula $(Ca(Fe_{0.5}Mg_{0.5})(CO_3)_2)$ 159 160 below). thermodynamic properties were estimated from the (see Its 161 stoichiometrically weighted sum of the calcite, siderite and magnesite stability 162 constants (Xu et al., 2004). The thermodynamic properties required for this estimate 163 were calculated using the SUPCRTBL database (Zimmer et al., 2016). The resulting 164 ankerite equilibrium constant and enthalpy of reaction are 0.642 and -69.326 kJ/mol at 165 70 °C and -0.087 and -77.803 kJ/mol 100 °C. Thermodynamic properties in the LLNL 166 database are consistent with those of dawsonite reported by Bénézeth et al. (2007). 167 The Truesdell-Jones equation was used to calculate aqueous activity coefficients and 168 the fugacity for CO₂ gas was calculated using the ThermoSolver 1.0 software (Barnes,

169 2006).

170

171 *4 Results*

172 *4.1 Petrographic characteristics*

The dawsonite in the LDX-1 structure is hosted in fine-grained subarkose and lithic quartz sandstones. The observed detrital minerals are mainly quartz (16.7% to 64.7%), feldspar (up to 12.0%) and rock fragments (up to 23.0%). The modal compositions of the detrital minerals are shown schematically in Fig 3. The feldspar is mainly comprised of plagioclase and also contains small amounts of orthoclase and perthite. The rock fragments mainly consist of metamorphic debris (up to 19%), granite debris (up to 8%), and sedimentary debris (up to 7%).

180

181 4.2 Authigenic mineralogy and paragenesis

The authigenic minerals observed in the rock samples are dawsonite, quartz,
siderite, dolomite, ankerite, mixed-layer clays consisting of illite and smectite, and
kaolinite.

185

186 *4.2.1 Dawsonite*

187 Dawsonite is principally found at depths of 1841.27 to 1867.23 m in well LDX-1-1 and 1863 to 2018 m in well LDX-1-4 with its content ranging from 0.2 to 6 188 189 Vol% as determined by point counting (see above). Dawsonite aggregates nucleate as microcrystalline grains then grow radially outward into the sandstone pores. 190 191 Dawsonite growth usually terminates at the edge of the detrital minerals (eg. quartz 192 and feldspar). Dawsonite could also be observed growing on dolomite (Fig. 4a, b), 193 potassium feldspar and plagioclase (Fig 4c, d). The EDS results show that the 194 dawsonite mainly consists of carbon, oxygen, sodium and aluminum having weight 195 percents of 7.25 to 22.23, 50.43 to 66.93, 3.58 to 14.69, and 7.27 to 30.26 respectively 196 (see Table 1).

198 4.2.2 Authigenic quartz

199 Two kinds of authigenic quartz were observed: quartz overgrowths and 200 microcrystalline quartz. The content of authigenic quartz is low, generally less than 201 0.5%. Quartz overgrowths are the among the earliest authigenic silicate minerals 202 observed and are evident due to clay rims between detrital quartz and these overgrowths. The thickness of the quartz overgrowths ranges from 0.01 to 0.06 mm. 203 204 Furthermore, we observed some radial dawsonite surrounded by quartz overgrowths 205 (see Fig. 4e), suggesting some of the quartz precipitated concurrently with or after the 206 dawsonite. In addition, SEM images show that the microcrystalline quartz was 207 distributed in sandstone pores sporadically and is associated with kaolinite. This 208 kaolinite is also associated with dawsonite in some places (Fig. 4f). Such observations 209 further suggest that the microcrystalline quartz formed during same time period as the 210 dawsonite.

211

212 *4.2.3 Siderite*

213 Most siderite was distributed in pores as subhedral intergranular crystals or 214 aggregates along the boundaries of the detrital grains (Fig. 4g). The content of the 215 siderite ranges from 0.5% to 13.5% in the bulk rock and the crystal sizes range from 216 0.01 mm to 0.03 mm. The color of the siderite is yellow brown to dark brown under 217 single polarized light. The siderite crystals were distributed either in the inner side of 218 quartz overgrowth or near the clay rims surrounding the original quartz grain 219 boundaries, indicating that siderite formation was earlier than or concurrent with the 220 quartz overgrowths.

221

222 *4.2.4 Dolomite*

The dolomite content (0.5 to 43 Vol%) varied substantially in the studied cores. As dolomite is not stained by Alizarin Red S and K-ferricyanide solutions, it is readily distinguished from other carbonates. Under the polarizing microscope, some dolomite was observed being replaced by dawsonite aggregates (Fig. 4h). SEM analysis revealed that the ends of some needle-like dawsonite crystals terminate at the faces of dolomite crystals exhibiting a "bridged phenomenon" (Figure 4a, b). Both sets of observations imply that the dolomite formed earlier than the dawsonite, yet the presence of some detrital dolomite cannot be discounted.

231

232 *4.2.5 Ankerite*

233 Ankerite (0.5 to 36 Vol%) mainly surrounds dolomite crystals as rim-like growths 234 (Fig. 4i) or as euhedral to subhedral intergranular crystals (Fig. 4j, k). Under the Back 235 Scattered Electron (BSE) mode, the ankerite is seen to form of nearly equal ring-like 236 shapes along the dolomite surfaces as it replaced this mineral. Some ankerite is seen to 237 have grown on dawsonite (Fig 4j, k). Such observations suggest that ankerite formation 238 occurred later than both the dolomite and dawsonite. EDS analyses results for ankerite 239 showed that its Ca, Mg, and Fe content ranged from 0.39 to 9.22, 2.4 to 6.52, and 1.47 240 to 7.82 atomic percent respectively (see Table 2), consistent with the formula 241 $Ca(Fe_{0.5}Mg_{0.5})(CO_{3})_{2}$.

242

243 4.2.6 Clay minerals

244 The XRD results indicate that the clay minerals in the CO₂-containing reservoirs 245 of the Yinggehai Basin mainly consist of mixed-layer illite and smectite, and kaolinite 246 (Huang et al., 2004; Jiang et al., 2015; Tong et al., 2012). The mixed layer illite and 247 smectite commonly surround the detrital minerals as "clay rims", which could be 248 identified by polarizing microscope. Mixed layer illite and smectite are not found in the 249 sandstone pores, suggesting that it is the earliest formed secondary mineral in the 250 diagenetic sequence. Honeycomb-like authigenic clay is seen to grow on K-feldspar 251 (see Fig. 4l, m). According the EDS results shown in Table 3, most of the smectite and 252 mixed layer illite are Ca depleted. This suggests that the Ca in the smectite and mixed

layer of illite and smectite may have been removed over time through water-mineralreactions.

255 Authigenic kaolinite grows as vermicular assemblages in the sandstone pores. 256 Hexagonal kaolinite crystals are found attached to dawsonite crystals (Fig. 4n), at times 257 adjacent to small "etch pits" on the dawsonite surface (Fig. 40), suggesting that some 258 kaolinite post-dates the dawsonite. However, vermicular kaolinite is commonly 259 interspersed between acicular dawsonite crystals, suggesting that this kaolinite formed 260 slightly earlier than or concurrently with dawsonite (Fig. 4p). In some places ankerite 261 grew on the top of kaolinite aggregates (Fig. 4q). This suggests that the formation of 262 kaolinite might have spanned from the period from when dawsonite grew through that 263 of ankerite formation.

Chlorite was not observed in the samples considered in this study. Its presence in the rocks of the LDX system was reported by Fu et al. (2016). Moreover, the presence of chlorite within and below the CO_2 gas reservoirs is inferred as a potential source of the iron leading to ankerite precipitation (see below).

268

269 4.3 Vertical characteristics of authigenic carbonates in the reservoirs

270 The dawsonite and ankerite contents across the study area, measured on 159 thin sections, are shown in Fig. 5. Over the depths of this study (from 1841.6 m to 1868.89 271 272 m), dawsonite (0.7 to 8.3%) and ankerite (1.7 to 26.0%) in the LDX-1-1 well mainly 273 developed near the gas water contact (GWC) (Fig. 5a). Owing to lack of samples, few 274 observations of these minerals were made above the GWC in this well. In well 275 LDX-1-4, the limited measurements of mineral content in the gas reservoirs show the 276 dawsonite and ankerite content ranging from 0.7 to 2.3% and 2.0 to 14.0% above the 277 GWC, whereas the corresponding content results for dawsonite and ankerite under the 278 GWC in this well are 1.0 to 6.3% and 1.0 to 24.0 % (Fig. 5b).

279

280 4.4 C and O isotope compositions

The stable isotopic compositions of dawsonite and ankerite are shown in Table 4 and Fig. 6. The carbon isotope (δ^{13} C) values for dawsonite vary from -4.3‰ to -0.6‰ with an average value of -2.0‰. The oxygen isotope (δ^{18} O) values for dawsonite range from +22.4‰ to +24.0‰. The δ^{13} C values for ankerite range from -2.6‰ to -0.8‰ with an average value of -1.4‰ and the δ^{18} O values range from +22.8‰ to +23.7‰.

The measured δ^{13} C values of dawsonite and ankerite have been used to estimate 286 the corresponding equilibrium δ^{13} C of the CO₂ (δ^{13} C_{CO2}) from which they precipitated. 287 The $\delta^{13}C_{CO2}$ for the CO₂ that formed the dawsonite, was estimated assuming dawsonite 288 289 exhibits an identical CO₂ carbon isotope fractionation as that of calcite, using the equation proposed by Ohmoto and Rye (1972) (Baker et al., 1995). The calculated $\delta^{13}C$ 290 for the CO₂ in equilibrium with ankerite was estimated assuming its fractionation was 291 292 identical to the dolomite- CO_2 fractionation equation proposed by Horita (2014). The δ^{13} C of the CO₂ (δ^{13} C_{CO2}) calculated to be in isotopic equilibrium with dawsonite range 293 from -9.66‰ to -6.69‰ with an average value of -7.90‰. The equilibrated $\delta^{13}C_{CO2}$ for 294 ankerite varies from -7.27‰ to -5.94‰ with a mean value of -6.4‰. 295

296 The formation temperature for dawsonite and ankerite were calculated from their 297 oxygen isotope compositions. This was done by assuming oxygen fractionation 298 between dawsonite and H₂O is identical to the calcite-H₂O fractionation equation proposed by O'Neil et al. (1969) and that oxygen fractionation between ankerite and 299 H_2O is identical to the dolomite- H_2O fractionation equation of Horita (2014). The ¹⁸O 300 value for water used in these calculations was $\delta^{18}O_{H2O} = +2.7\%$ (Deng and Chen, 1994), 301 302 which was measured in the formation waters of the DFY-1 structure in the 303 Yinggehai-Huangliu Formation of the Yinggehai basin (Fig. 1). The calculated 304 temperature for dawsonite formation varies from 67.8 °C to 79.6 °C and the calculated 305 temperature of ankerite formation ranges from 89.6 °C to 96.4 °C (see Table 4). These temperatures are consistent with 1) the earlier formation of dawsonite and 2) the 306 307 increasing temperature of the study site over time as the basin subsided. Note that the 308 current temperature of the study site is 92 to 103 °C.

310 5 Discussion

311 5.1 The origin of the carbon in the dawsonite and ankerite

The $\delta^{13}C_{CO2}$ of the gas forming dawsonite and ankerite reported above and in 312 Table 4 are similar to the current δ^{13} C values in the CO₂-rich reservoirs of the LDX 313 system; the current $\delta^{13}C_{CO2}$, are reported to range from -8.2% to -4.2% (He et al., 314 315 2005). Moreover, the hydrogen and oxygen isotopic compositions of the formation 316 water in the Yinggehai Basin reservoirs are significantly different from that of 317 meteoric water, which implies the lack of exchange between meteoric water and that 318 present in the LDX system (Deng and Chen, 1994). Such evidence suggests that the CO_2 in the gas reservoir is likely a major carbon source for the both dawsonite and 319 320 ankerite formation.

321 Carbon dioxide in the LDX gas reservoirs at the present time is likely dominated by magmatic derived CO₂. Firstly, the $\delta^{13}C_{CO2}$ value in the gas reservoir ranges from 322 -8.2‰ to -4.2‰, which is significantly different from that obtained from kerogen 323 degradation; $\delta^{13}C_{CO2}$ from kerogen degradation ranges from -25% to -20% (Irwin et 324 325 al., 1977; Tissot and Welte, 1984) and CO₂ from bacterial fermentation typically has a $\delta^{13}C_{CO2} \sim +15\%$ (Irwin et al., 1977). Secondly, the carbon isotope values of CO_2 in 326 the LDX system exhibits a $\delta^{13}C_{CO2}$ range similar to the -8% to -3% range observed 327 for CO₂ sourced from mantle-derived magma (Baker et al., 1995; Clayton et al., 1990; 328 329 Dai, 1996; Ohmoto and Rye, 1972; Rollinson, 2014;). Taking account of the 330 observations summarized above, we suggest that origin of the CO₂ in the LDX-1 331 structures is mantle-derived magma degasing.

The timing of dawsonite and ankerite formation can be constrained by the timing of the basaltic magmatism related CO_2 charging of the LDX system. The Yinggehai Formation, where the dawsonite and ankerite developed, was formed between 5.5 Ma and 1.9Ma (Fig. 2). The Honghe fault zone, which controls the formation and distribution direction of Yinggehai basin is still active (Zhu et al., 2009) (Fig. 1). The ages of alkaline basalt in this region range from 4 to 6 Ma and from 1 Ma up to
present (Zhu and Wang, 1989). Several ages of the CO₂ charging into the Yinggehai
Formation have been reported in the literature; these ages range from 0.3 to 1.9 Ma
(Fu et al., 2016; Huang et al., 2005; Wang et al., 2004; Xie et al., 2014). Based on
these ages, it seems reasonable to conclude the dawsonite and ankerite in the
Yinggehai Formation is formed during the last 1.9 Ma.

343

344 5.2 Cation source of dawsonite and ankerite

345 The dissolution of feldspars in the LDX-1 structure provides the cations for 346 dawsonite. The formation of dawsonite from feldspar has been observed in numerous 347 systems. For example, feldspars were observed to be altered to dawsonite in the Lam 348 Formation of Shabwa basin Yemen (Worden, 2006), in the reservoirs of the Castillo 349 structure of the Golfo San Jorge basin (Comerio et al., 2014), and in the Songliao basin, 350 the Hailaer basin, and the East Sea basin in China (Gao et al., 2009; Liu et al., 2011; 351 Zhao et al., 2018; Zhou et al., 2014). The reaction forming dawsonite from albite can be 352 written (Bénézeth et al., 2007; Hellevang et al., 2005; Ryzhenko, 2006; Xi et al., 2016): 353 $NaAlSi_{3}O_{8} + CO_{2} + H_{2}O \rightleftharpoons NaAlCO_{3}(OH)_{2} + 3SiO_{2(a_{1})}$ (Eq. 1) 354 Albite Dawsonite

- 355 where some of the liberated $SiO_{2(aq)}$ could provoke quartz precipitation. The equivalent
- 356 reaction forming dawsonite from K-feldspar:

357
$$\operatorname{KAlSi_3O_8} + \operatorname{Na^+} + \operatorname{CO_2} + \operatorname{H_2O} \rightleftharpoons \operatorname{NaAlCO_3(OH)_2} + 3\operatorname{SiO_{2(aq)}} + \operatorname{K^+}$$
 (Eq. 2)

358 K-feldspar Dawsonite

has been discounted by Hellevang et al. (2011) due to the likely build-up of K^+ in the aqueous phase arresting this reaction.

Evidence for the coupling of feldspar dissolution to dawsonite precipitation reactions could be found throughout the LDX system (Fig. 4c, d, r). If dawsonite originates from K-feldspar by Eq. 2, however, its growth also requires Na⁺ from the fluid phase. This would require the system to be relatively open to fluid transport, which may be limited in this system. Note that according to mass conservation, for every 1 cm³ of dawsonite formed from albite by reaction 1, there should be 1.16 cm^3 of SiO₂ precipitation (Zhou et al., 2014). However, far less microcrystalline quartz is observed in our system than dawsonite (Fig. 4b), which suggests that some of the SiO₂ may have been consumed by other water-mineral reactions.

370 Geochemical calculations results suggest that chlorite dissolution provoked by the charging of the system by CO_2 could provide Fe^{2+} for ankerite formation (Xu et al., 371 372 2005, 2007, 2014; Yu et al., 2014). The dissolution of smectite also could provide 373 some of the iron for ankerite growth (Hellevang et al., 2013; Pham et al., 2011; 374 Worden and Morad, 2003; Worden, 2006), as well as some aluminum promoting 375 dawsonite formation. The dissolution of chlorite and smectite coupled to ankerite 376 formation has been observed in the sandstone reservoirs of the Otway basin of Austrailia., in the Songliao Basin of China, the Jurassic and late Cretaceous reservoirs 377 378 of North Sea oil fields, and in the Eocene strata from Wilcox of Texas, (Awwiller, 379 1993; Hendry et al., 2000; Liu et al., 2011; Thyberg et al., 2010; Watson et al., 2004). 380 The chlorite content of the Yinggehai Basin CO₂ bearing reservoirs has been reported 381 to be significantly lower than that in CO₂-free reservoirs (Tong et al., 2012). Similar 382 observations have been reported in wells of the Yinggehei Basin (Huang et al., 2004; Jiang et al., 2015). Such observations suggest that the dissolution of iron bearing clay 383 minerals such as chlorite is the main Fe^{2+} source for ankerite. The Mg²⁺ and some 384 Ca^{2+} in the ankerite crystals could come from dolomite dissolution, as fresh ankerite is 385 observed growing within a dolomite matrix (Fig. 4i). Additional Ca²⁺ can be sourced 386 387 from the transformation of clays in the system; a number of studies have reported Ca 388 release as smectite transforms to mixed layered clays and illite (e.g. Boles and Frank, 389 1979; McHargue and Price, 1982). A close relation between ankerite and smectite is 390 evident in the studied core samples (see Fig. 7). Taking account of these observations, 391 a representative reaction coupling chlorite dissolution to ankerite precipitation could be expressed as (modified from Kirste et al., 2004; see also Luquot et al., 2012): 392

$$393 \quad \text{Fe}_2\text{Al}_2\text{SiO}_5(\text{OH})_4 + 2 \text{ Ca}(\text{MgCO}_3)_2 + \text{SiO}_2 + 2 \text{ Ca}^{2+} + 4 \text{ CO}_2 + 2 \text{ H}_2\text{O} =$$

394 Chamosite Dolomite Ouartz

395

396

Kaolinite

 $4 \operatorname{Ca}(\operatorname{Fe}_{0.5}\operatorname{Mg}_{0.5})(\operatorname{CO}_3)_2 + \operatorname{Al}_2\operatorname{Si}_2\operatorname{O}_5(\operatorname{OH})_4 + 4 \operatorname{H}^+$

397 Note this reaction, balanced on Fe and Mg, requires substantial magmatic CO₂ for 398 ankerite formation, consistent with its carbon isotope composition. This reaction also 399 consumes SiO₂, which may account, at least in part, for the low mass of authigeic 400 quartz in the present day system.

(Eq. 3)

Ankerite

401

402 5.3 CO₂ partial pressure and carbonate mineral stability

403 It is likely that both ankerite and dawsonite are formed after the arrival of mantle 404 sourced CO₂, providing a clear analogy for modern carbon capture and storage 405 systems. The observations described in this study indicate that dawsonite is first 406 formed at ~70 °C, followed by ankerite at ~ 90 °C, in response to the arrival of CO₂ 407 into the overlying reservoirs. The increase in temperature in the system likely stems 408 from its increased burial depth over time. Ankerite is also reported to be the last 409 forming mineral in the sandstone reservoirs in Songliao (Liu et al., 2011; Yu et al., 410 2015) and Hailaer basins (Gao et al., 2009). Some kaolinite is seen to precipitate 411 coupled to the dissolution of dawsonite; some dissolution etch pits on dawsonite crystal were evident near the kaolinite (Fig. 4m). Kaolinite precipitation continues as 412 ankerite precipitates, consistent with Eq. (3). The formation of ankerite also appears 413 414 to be linked to dolomite dissolution.

415 All of these observations are consistent with the phase diagrams shown in Figs. 8 416 to 10. In Fig. 8, the relative stability of dawsonite in the Na₂O-Al₂O₃-SiO₂-CO₂-H₂O 417 system balanced on aluminum and saturated with quartz is shown. Dawsonite and 418 quartz become stable relative to albite at log (f_{CO2}) > 0.15 and 0.9 at 70 and 100 °C, 419 respectively. This diagram thus suggests that albite would be favored to transform to 420 dawsonite as CO₂ fills the reservoirs in accord with reaction 1. As suggested by Fig. 9, 421 some of the precipitated dawsonite and quartz will tend to react to form kaolinite due 422 to an increase in temperature, which tends to shrink the dawsonite stability field. The 423 value of log (f_{CO2}) in equilibrium with the assemblage albite = dawsonite + quartz at 424 70 and 100 °C plot as vertical lines on the Fig. 10, which also shows the relative 425 stability field of ankerite + kaolinite relative to that of chlorite, dolomite and quartz consistent with reaction 3. It can be seen that for all $a_{Ca2+}\!/\!a^2_{~H^+}\!>$ ~9.7 at 70 °C and >426 ~7.3 at 100 °C, dawsonite would first tend to form before ankerite as the CO₂ partial 427 428 pressure in the reservoirs increase. The partial pressure of CO₂ in the LDX system 429 would increase as the CO₂ fraction of the gas progressively increased in the reservoirs 430 located just above the studied rock samples (Fig. 5). The persistence of dawsonite and 431 ankerite in this system is assured by the continued presence of the CO₂-rich gas 432 reservoir, which maintains f_{CO2} at relatively high levels. Note the current temperature 433 of the gas reservoirs are 92.3 and 102.9 °C for well LDX-1-4 and well LDX-1-1 434 respectively, and the corresponding gas pressure in reservoirs, which is dominated by 435 CO₂, are between 18.9 and 43.2 Mpa (China National Offshore Oil Company, 436 personal communication). Hellevang et al. (2005), noted, that dawsonite, would likely 437 rapidly dissolve once it became thermodynamically unstable as the f_{CO2} decreased in 438 the such systems.

439

440 5.4 The paragenetic sequence

441 Based on the observations described above, the proposed paragenetic sequence for 442 the studied dawsonite-bearing sandstones is shown in Fig. 11. Secondary clays (mainly 443 illite and smectite mixed layer clays) began forming before compaction during the early 444 stage of diagenesis and coated detrital minerals as clay rims. Sideite is found to form 445 next; siderite precipitation is readily observed in reducing, non-sulphidic pore waters in 446 the suboxic and methanogenesis zones (Morad, 1998). Quartz overgrowths are then 447 observed to form, perhaps due to the recrystallization of biogenic SiO₂, but also 448 accompanied with the formation of minor kaolinite in response to feldspar dissolution.

Subsequently, dolomite is observed to form, potentially promoted by Ca and Mg release from the dissolution or transformation of the earlier formed clays. After the influx of CO_2 into the reservoir. dawsonite and additional quartz forms from the dissolution of feldspar, and subsequently ankerite and kaolinite forms from the dissolution of dolomite, quartz, chlorite and potentially some clay minerals..

454

455 5.5 Implications for carbon storage in sandstone reservoirs

456 The formation and long-term stability of dawsonite and ankerite in the LDX 457 system are likely good analogies for some modern carbon storage systems such as the 458 Shenhua CCS demonstration project in China (Yang et al., 2017). The distribution of 459 dawsonite and ankerite in the LDX reservoirs (Fig. 5) indicates that these minerals 460 can form in the water layers beneath gas water contact (GWC). Some dawsonite and 461 ankerite were also observed above the present day GWC, which may indicate that 462 either 1) the position of the GWC has changed over time or 2) that these phases grew 463 in the gas filled section of the reservoir, potentially on water-wet surfaces. This later, 464 process, however, is likely limited as the dawsonite and ankerite forming reactions 1 465 and 3 consume H_2O . The formation of such carbonate minerals adjacent to the CO_2 466 gas reservoir is also supported by the results of TOUGHREACT calculations reported 467 by Pan et al. (2017) that suggest that when CO_2 is injected into sandstone reservoirs, 468 carbonate minerals would form on both upper and lower layers of the reservoir. 469 Carbonate mineral precipitation in the water layers below the GWC has also been 470 observed in New Zealand sandstone reserviors (Higgs et al., 2015). As such, we suggest 471 that the CO_2 in the gas reservior maintains a sufficient CO_2 partial pressure in the 472 formation waters below the GWC to promote the growth and long term stability of 473 dawsonite and ankerite. However, according to the paragenetic sequence determined 474 in this study, kaolinite precipitated from the fluid concurrently with the precipitation 475 of dawsonite and ankerite. Numerous petrographic observations and experimental 476 simulations suggest that kaolinite and dawsonite formation could significantly reduce

477 sandstone permeability (Luquot et al., 2012; Shiraki and Dunn, 2000; Yu et al., 2012). 478 Kaolinite also commonly grows in response to CO₂-feldspar or CO₂-clay mineral 479 interaction (Luquot et al., 2012; Moore et al., 2005; Shiraki and Dunn, 2000; Yu et al., 480 2012). Shiraki and Dunn (2000) concluded that the precipitation and growth of 481 kaolinite in the pores could block pore throats leading to the reduction of the 482 permeability of reservoir. Similarly, Worden (2006) suggested that the fiber-like crystal 483 morphology of dawsonite would also reduce sandstone permeability. The reduced 484 permeability in the sandstone could slow chemical transport, and including potentially 485 the transport of dissolved CO_2 out of the storage system. Such processes could thereby 486 promote the stability of carbonate minerals in this system. The decrease in 487 permeability could also lead to fluid pressure build up in the gas reservoir as further 488 minerals form in this system.

489

490 6 Conclusions

491 This study focused on the mineral reactions provoked by the influx of CO_2 into the 492 subarkose and sublitharenite sandstones of the Yinggehai basin. This system is likely a 493 good analogy for subsurface carbon storage in sandstone hosted hydrocarbon reservoirs 494 as the influxed CO₂ replaces a methane-rich gas phase in an anticlinal structure. This 495 leads to the presence of a CO_2 -rich gas phase stable over a ~1-2 million year time frame. 496 The influx of provoked the CO₂ provoked the formation of dawsonite from plagioclase 497 and ankerite and kaolinite likely from chlorite, quartz, and dolomite, at temperatures of 68 to 80 °C. and 90 to 96° C, respectively, largely just below the CO₂-gas water 498 499 interface. The formation of such phases could provide for the mineral storage of some 500 of the CO₂ injected during subsurface carbon storage efforts.

501 The key to maintaining dawsonite and ankerite stability in such systems appears to 502 be preserving a high partial pressure of CO_2 . As such, although the formation of these 503 minerals might lead to the storage of CO_2 injected into sandstones as part of 504 engineered geologic storage efforts, such minerals may not provide for the 'permanent' 505 mineral storage of this gas that one might anticipate from the formation of carbonate 506 minerals such as calcite, which can be stable at far lower CO_2 partial pressure, unless 507 the system tends to become impermeable as due to the formation of various secondary 508 minerals.

509

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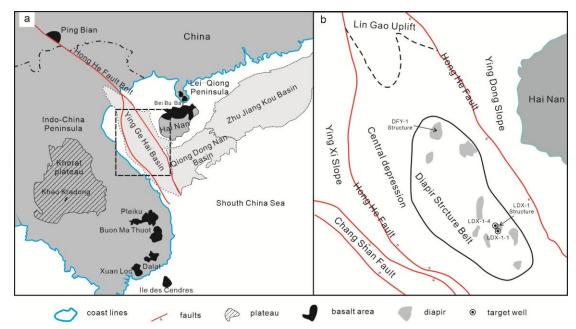
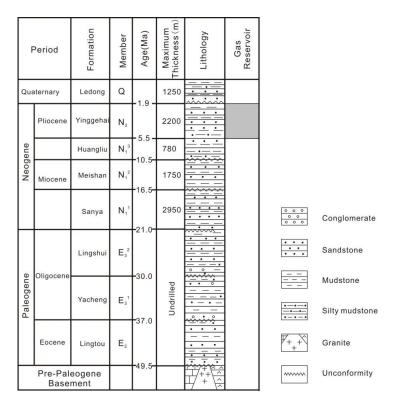


Figure 1. Location map of LDX-1 structure in Yinggehai Basin, south China Sea area. a) The location of Yinggehai
basin and basalt area modified after Shi and Yan (2011), basalt areas are modified from Hoang et al. (1996, 1998), Li
et al. (2005), Wang et al. (2001); Zhou et al. (1997) and Zhu et al. (1989), b) Geotectonic units modified after Luo et

837 al., (2003) and Meng et al. (2012).

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840 Figure 2. Simplified stratigraphic column of the Yinggehai basin modified after Gong et al. (1997) and Huang et al.

841 (2003).

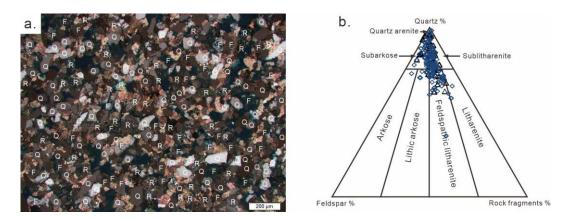
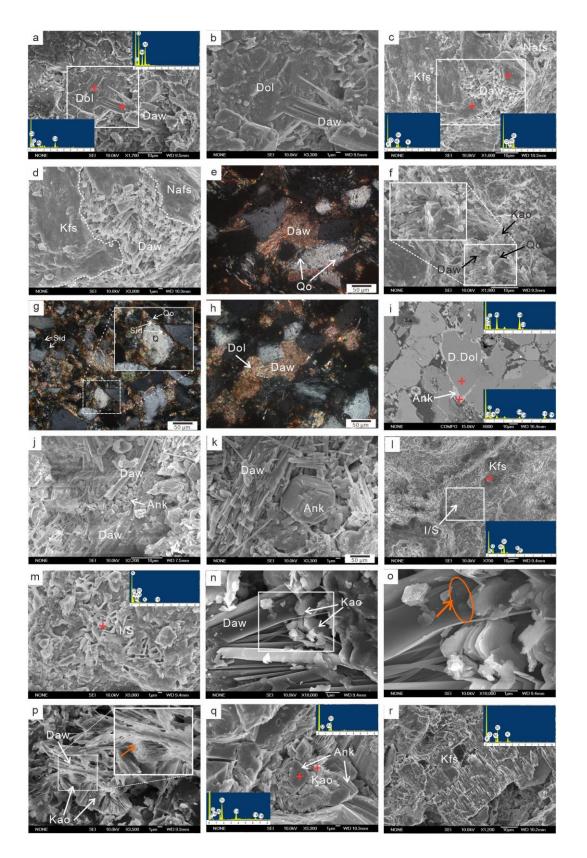


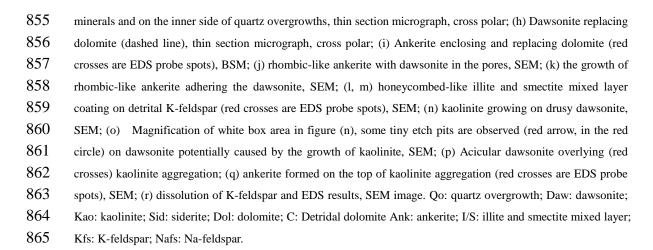
Figure 3. a) Image of a thin section of a sandstone sample collected from a depth of 1860.72 m in the LDX-1-1 well
showing its detrital minerals: Q=quartz, F=feldspar and R=rock fragment. b) Mineral composition diagram, after
Folk (1968), of the dawsonite-bearing samples from well LDX-1-1 (blue rhombuses) and well LDX-1-4 (black

- triangles). All the data were obtained through point-counting see text.
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Figure 4. Micrographs of diagenetic minerals. (a, b) Acicular dawsonite growing on top of dolomite (red crosses
are EDS probe spots), SEM; (c, d) Dawsonite replacing feldspar and plagioclase with EDS analysis results (red
crosses are EDS probe spots); (e) Dawsonite replacing quartz overgrowths, thin section micrograph, cross polar; (f)
Microcrystalline quartz grows with kaolinite on the top of drusy dawsonite, SEM; (g) Siderite surrounding detrital



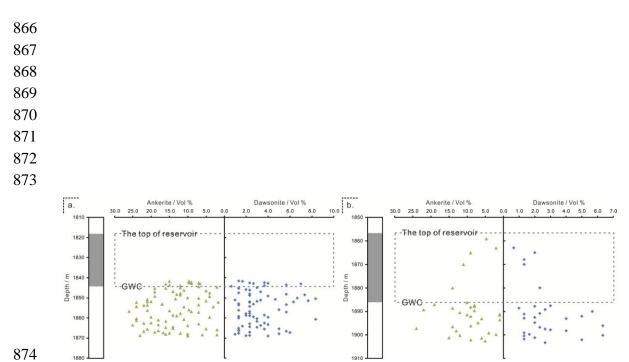
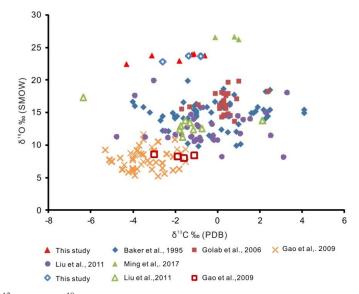
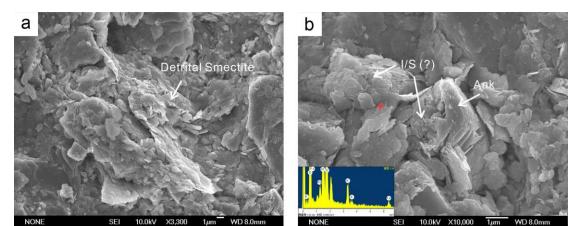


Figure 5. Plot of volume percent versus depth for dawsonite (blue rhombuses) and ankerite (green triangles) in
well LDX-1-1 (Fig 3a) and well LDX-1-4 (Fig 3b). The filled gray areas and dashed boxes indicate the location of
gas layers, the water layers in the reservoir system are below the gas layers. The gas-water-contact (GWC), as
reported by the China National Offshore Oil Company are located at 1844.1m and 1886.0m depth in wells
LDX-1-1 and LDX-1-4, respectively.



883 Figure 6. Plot of δ^{13} C versus δ^{18} O values for dawsonite and ankerite samples. Filled symbols represent the 884 compositions of dawsonite, and open symbols for ankerite The filled red triangles and open blue rhombus 885 correspond to compositions measured in this study from the LDX-1 structure, Yinggehai basin, China. 886 Compositions of dawsonite from the Bowen-Gunnedah-Sydney basin system in Australia reported by Baker et al. 887 (1995) are shown as filled blue rhombuses, those from the late Permian coal measures of the Upper Hunter Valley 888 in Australia reported by Golab et al. (2006) are shown as filled dark red squares, those of the Cretaceous Hailaer 889 basin in China reported by Gao et al. (2009) are shown as filled yellow crosses, those of the upper Cretaceous 890 Songliao Basin in China reported by Liu et al. (2011) are shown as filled purple circles, and those of the Yaojie 891 coalfield, Minhe basin in China reported by Ming et al. (2017) are shown as filled green triangles. Compositions 892 of ankerite from the Cretaceous Hailaer basin in China reported by Gao et al. (2009) are shown as open red 893 squares, and those of the upper Cretaceous Honggang anticline of the Songliao Basin in China reported by Liu et 894 al. (2011) are shown as open green triangles.





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Figure 7. Images of clay minerals observed in this study. (a) detrital smectite in the sandstone and (b) image showing
the relationship between smectite and ankerite. Rhombic ankerite crystals grew and adhered to the mixed layer of
illite/smectite, suggesting the transformation of smectite could provide some of the materials for ankerite. I/S: mixed

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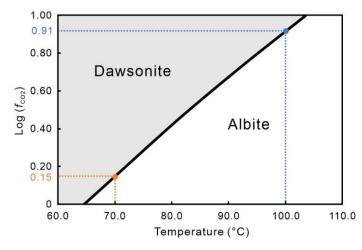
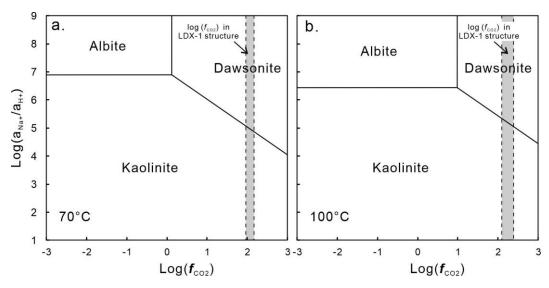


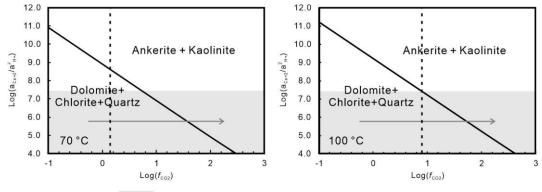
Figure 8 Log fugacity of CO₂ (log(f_{CO2})) versus temperature (°C) diagram depicting mineral stability fields in the system Na₂O-Al₂O₃-SiO₂-CO₂-H₂O balanced on Al and at quartz equilibrium. The grey area indicates the stability field of dawsonite, and the white area indicates the stability field for albite. The orange and blue lines highlight the log(f_{CO2}) values in equilibrium with dawsonite and albite at 70 and 100 °C, respectively. This figure was generated using PHREEQC together with its llnl database – see text.

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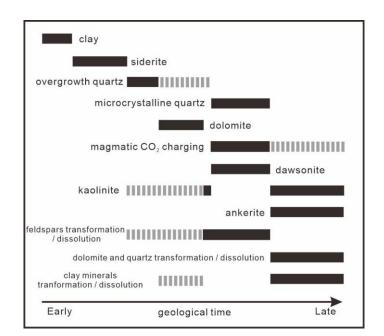
911Figure 9. Log fugacity of CO2 ($\log(f_{CO2})$) versus log a_{Na+}/a_{H+} diagrams depicting mineral stability fields in the912system Na2O-Al2O3-SiO2-CO2-H2O balanced on Al and at quartz equilibrium at 70 °C and 100 °C. The dashed913area represents the current log (f_{CO2}) range in the reservoirs considered in this study. This figure was generated914using PHREEQC together with its llnl database – see text.



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Estimate fluid composition and its evolution direction due to CO₂ fugacity increases

917 Figure 10. Phase diagram of log CO₂ fugacity (log (f_{CO2})) versus log a_{Ca2+}/a^2_{H+} . The dashed vertical lines 918 correspond to equilibrium compositions of reaction 1, whereas the sloped lines correspond to equilibrium 919 compositions of reaction 3, The grey area is the likely aqueous activity ratio of a_{Ca2+}/a^2_{H+} in the studied system and 920 the arrow illustrates the direction of the CO₂ fugacity evolution in the system as the reservoir fills with CO₂. This 921 figure was generated using PHREEQC together with its llnl database after adding estimated values for the 922 thermodynamics properties of ankerite– see text. 923



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927 Figure 11. Paragenetic sequence for dawsonite-bearing sandstones in this study. The black bars are indicated by

928 petrologic observations, the grey dashed bars are estimated based on the mass balance considerations and reactions929 (1) and (3).

| Q | 3 | n |
|---|---|---|
| 7 | J | υ |

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| Table 1. EDS analysis data for dawsonite. | Table 1. | EDS | analysis | data for | dawsonite. |
|---|----------|-----|----------|----------|------------|
|---|----------|-----|----------|----------|------------|

| W 7-11 | | Weigh | t Percent | age, % | | Atomic Percentage, % | | | | | | |
|---------------|-------|-------|-----------|--------|-------|----------------------|-------|-------|-------|-------|--|--|
| Well | С | Na | Al | 0 | Si | С | Na | Al | 0 | Si | | |
| LDX-1-1 | 18.76 | 4.98 | 12.99 | 63.27 | n.d. | 25.14 | 3.48 | 7.75 | 63.63 | n.d. | | |
| LDX-1-1 | 15.88 | 11.13 | 14.19 | 58.8 | n.d. | 22.01 | 8.06 | 8.76 | 61.18 | n.d. | | |
| LDX-1-1 | 11.81 | 12.44 | 12.96 | 55.55 | 7.23 | 17.14 | 9.44 | 8.38 | 60.55 | 4.49 | | |
| LDX-1-4 | 14.88 | 13.8 | 14.24 | 57.09 | n.d. | 20.87 | 10.11 | 8.89 | 60.13 | n.d. | | |
| LDX-1-4 | 15.61 | 12.53 | 13.7 | 58.15 | n.d. | 21.71 | 9.11 | 8.48 | 60.7 | n.d. | | |
| LDX-1-4 | 12.62 | 14.69 | 17.98 | 54.72 | n.d. | 18.19 | 11.06 | 11.54 | 59.21 | n.d. | | |
| LDX-1-4 | 12.64 | 8.37 | 14.29 | 57.5 | 7.2 | 18.15 | 6.28 | 9.14 | 62 | 4.43 | | |
| LDX-1-4 | 7.83 | 6.83 | 8.83 | 55.28 | 21.23 | 11.88 | 5.41 | 5.96 | 62.97 | 13.78 | | |
| LDX-1-4 | 12.58 | 12.01 | 19.97 | 55.44 | n.d. | 18.13 | 9.04 | 12.82 | 60.01 | n.d. | | |
| LDX-1-4 | 9.14 | 14.68 | 24.74 | 51.45 | n.d. | 13.75 | 11.54 | 16.57 | 58.13 | n.d. | | |
| LDX-1-1 | 10.16 | 13.06 | 23.91 | 52.87 | n.d. | 15.09 | 10.13 | 15.81 | 58.96 | n.d. | | |
| LDX-1-1 | 11.59 | 9.4 | 23.75 | 55.26 | n.d. | 16.9 | 7.16 | 15.42 | 60.52 | n.d. | | |
| LDX-1-1 | 18.76 | 4.98 | 12.99 | 63.27 | n.d. | 25.14 | 3.48 | 7.75 | 63.63 | n.d. | | |
| LDX-1-1 | 11.97 | 10.19 | 18.04 | 55.91 | 3.88 | 17.36 | 7.72 | 11.64 | 60.86 | 2.41 | | |
| LDX-1-1 | 7.25 | 12.06 | 30.26 | 50.43 | n.d. | 11.18 | 9.71 | 20.76 | 58.35 | n.d. | | |
| LDX-1-1 | 13.42 | 13.99 | 16.92 | 55.67 | n.d. | 19.16 | 10.43 | 10.75 | 59.66 | n.d. | | |
| LDX-1-1 | 13.68 | 13.29 | 16.93 | 56.11 | n.d. | 19.46 | 9.88 | 10.72 | 59.94 | n.d. | | |
| LDX-1-1 | 18.6 | 7.9 | 11.21 | 62.28 | n.d. | 24.98 | 5.54 | 6.7 | 62.78 | n.d. | | |
| LDX-1-1 | 12.75 | 11.24 | 20.18 | 55.83 | n.d. | 18.34 | 8.45 | 12.92 | 60.29 | n.d. | | |
| LDX-1-1 | 10.17 | 12.49 | 24.3 | 53.04 | n.d. | 15.1 | 9.69 | 16.07 | 59.14 | n.d. | | |
| LDX-1-4 | 13.71 | 7.2 | 12.99 | 58.84 | 7.26 | 19.43 | 5.33 | 8.2 | 62.63 | 4.4 | | |
| LDX-1-4 | 11.98 | 13.11 | 16.51 | 55.02 | 3.38 | 17.39 | 9.93 | 10.66 | 59.92 | 2.1 | | |
| LDX-1-4 | 10.86 | 13.85 | 21.98 | 53.31 | n.d. | 16 | 10.65 | 14.41 | 58.94 | n.d. | | |
| LDX-1-4 | 22.23 | 3.58 | 7.27 | 66.93 | n.d. | 28.65 | 2.41 | 4.17 | 64.77 | n.d. | | |
| LDX-1-4 | 10.18 | 12.73 | 24.1 | 52.99 | n.d. | 15.12 | 9.88 | 15.93 | 59.07 | n.d. | | |
| LDX-1-4 | 12.26 | 13.58 | 19.45 | 54.7 | n.d. | 17.75 | 10.27 | 12.53 | 59.44 | n.d. | | |

933 n.d.: not be detected.

| Table 2. EDS | analysis data | for ankerite |
|--------------|---------------|--------------|
|--------------|---------------|--------------|

| | | | | | | Table 2. | EDS alla | ysis uata | ioi anisci | ite | | | | | | |
|---------|-------|-------|------|--------|---------|----------|----------------|-----------|------------|-------|------|------|------|------|------|------|
| | | | | Weight | Percent | | Atomic Percent | | | | | | | | | |
| Well | С | 0 | Na | Mg | Al | Si | Ca | Fe | С | 0 | Na | Mg | Al | Si | Ca | Fe |
| LDX-1-4 | 9.66 | 51.32 | n.d. | 6.54 | n.d. | 4.42 | 6.29 | 21.77 | 16.14 | 64.34 | n.d. | 5.4 | n.d. | 3.16 | 3.15 | 7.82 |
| LDX-1-4 | 13.15 | 52.72 | n.d. | 8.46 | n.d. | n.d. | 19.73 | 5.95 | 20.52 | 61.75 | n.d. | 6.52 | n.d. | n.d. | 9.22 | 2 |
| LDX-1-4 | 10.3 | 53.98 | n.d. | 7.09 | n.d. | 1.77 | 17.26 | 9.6 | 16.53 | 65.02 | n.d. | 5.62 | n.d. | 1.22 | 8.3 | 3.31 |
| LDX-1-4 | 12.96 | 57.98 | n.d. | 5.16 | 2.28 | 1.72 | 1.12 | 18.77 | 19.89 | 66.79 | n.d. | 3.92 | 1.56 | 1.13 | 0.52 | 6.2 |
| LDX-1-4 | 16.52 | 61.57 | n.d. | 7.64 | n.d. | n.d. | 0.92 | 13.35 | 23.71 | 66.35 | n.d. | 5.42 | n.d. | n.d. | 0.39 | 4.12 |
| LDX-1-4 | 12.79 | 61.42 | n.d. | 6.25 | 1.52 | 2.56 | 2.73 | 12.73 | 19 | 68.49 | n.d. | 4.59 | 1 | 1.62 | 1.21 | 4.07 |
| LDX-1-4 | 15.63 | 55.85 | n.d. | 3.64 | 1.34 | 4.16 | 11.59 | 7.79 | 23.37 | 62.69 | n.d. | 2.69 | 0.89 | 2.66 | 5.19 | 2.5 |
| LDX-1-1 | 13.97 | 59.87 | 1.84 | 5.04 | 1.69 | 0.84 | 12.09 | 4.66 | 20.51 | 66 | 1.41 | 3.66 | 1.1 | 0.53 | 5.32 | 1.47 |
| LDX-1-1 | 13.39 | 59.42 | n.d. | 5.56 | n.d. | 2.16 | 12.7 | 6.78 | 20.01 | 66.65 | n.d. | 4.1 | n.d | 1.38 | 5.69 | 2.18 |
| LDX-1-1 | 12.41 | 60.07 | n.d. | 3.29 | 2.27 | 11.6 | 3.16 | 7.2 | 18.36 | 66.71 | n.d. | 2.4 | 1.49 | 7.34 | 1.4 | 2.29 |
| LDX-1-1 | 10.25 | 45.23 | n.d. | 3.86 | 2.24 | 5.18 | 16.05 | 17.17 | 17.73 | 58.71 | n.d. | 3.3 | 1.73 | 3.83 | 8.32 | 6.39 |

n.d.: not be detected.

Table 3. EDS analysis data for smectite and the mixed layer of illite and smectite

| | | | | Weig | ght Percen | tage, % | | | | | | | | | | | | |
|---------|------|-------|------|------|------------|---------|-------|-------|------|------|-------|------|------|-------|-------|------|------|------|
| Well | С | 0 | Na | Mg | Al | Si | K | Ca | Fe | С | 0 | Na | Mg | Al | Si | Κ | Ca | Fe |
| LDX-1-1 | 3.31 | 57 | 2.64 | 0.63 | 6.64 | 26.9 | 1.24 | 0.37 | 1.27 | 5.25 | 67.91 | 2.19 | 0.49 | 4.69 | 18.26 | 0.6 | 0.18 | 0.43 |
| LDX-1-1 | n.d. | 41.41 | n.d. | 1.22 | 9.69 | 35.81 | 6.72 | n.d. | 5.14 | n.d. | 57.05 | n.d. | 1.11 | 7.92 | 28.11 | 3.79 | n.d. | 2.03 |
| LDX-1-1 | n.d. | 37.5 | n.d. | 1.55 | 9.06 | 39.51 | 4.7 | n.d. | 7.67 | n.d. | 53.18 | n.d. | 1.44 | 7.62 | 31.92 | 2.73 | n.d. | 3.12 |
| LDX-1-1 | n.d. | 47.11 | 1.39 | 1.93 | 14.08 | 26.38 | 2.12 | n.d. | 6.99 | n.d. | 62.32 | 1.28 | 1.68 | 11.04 | 19.88 | 1.15 | n.d. | 2.65 |
| LDX-1-4 | n.d. | 57.17 | 0.7 | 0.99 | 15.27 | 22.26 | 2.53 | n.d. | 1.08 | n.d. | 70.24 | 0.6 | 0.8 | 11.13 | 15.58 | 1.27 | n.d. | 0.38 |
| LDX-1-4 | n.d. | 54.36 | n.d. | 4.31 | 8.23 | 11.71 | n.d. | 13.46 | 7.93 | n.d. | 71.16 | n.d. | 3.71 | 6.38 | 8.73 | n.d | 7.03 | 2.98 |
| LDX-1-4 | n.d. | 52.97 | n.d. | 1.18 | 13.35 | 26.82 | 4.82 | n.d. | 0.87 | n.d. | 66.92 | n.d. | 0.98 | 10 | 19.3 | 2.49 | n.d. | 0.31 |
| LDX-1-4 | n.d. | 55.21 | n.d. | 1.52 | 11.51 | 27.18 | 1.39 | n.d. | 3.18 | n.d. | 69.01 | n.d. | 1.25 | 8.53 | 19.35 | 0.71 | n.d. | 1.14 |
| LDX-1-4 | n.d. | 45.84 | n.d. | 1.44 | 14.38 | 28.31 | 3.13 | n.d. | 6.91 | n.d. | 61.37 | n.d. | 1.27 | 11.42 | 21.59 | 1.71 | n.d. | 2.65 |
| LDX-1-4 | n.d. | 46.01 | n.d. | 1.67 | 12.96 | 29.18 | 2.94 | n.d. | 7.24 | n.d. | 61.6 | n.d. | 1.47 | 10.29 | 22.25 | 1.61 | n.d. | 2.78 |
| LDX-1-4 | n.d. | 46.45 | n.d. | 1.52 | 15.13 | 26.85 | 4.9 | n.d. | 5.16 | n.d. | 61.77 | n.d. | 1.33 | 11.93 | 20.34 | 2.67 | n.d. | 1.96 |
| LDX-1-4 | n.d. | 52.5 | n.d. | 1.34 | 13.43 | 26.4 | 2.83 | n.d. | 3.49 | n.d. | 66.84 | n.d. | 1.12 | 10.14 | 19.15 | 1.48 | n.d. | 1.27 |
| LDX-1-4 | n.d. | 52.32 | 1.04 | 1.26 | 9.4 | 29.16 | 1.99 | n.d. | 3.28 | n.d. | 66.71 | 0.92 | 1.05 | 7.11 | 21.18 | 1.04 | 0.79 | 1.2 |
| LDX-1-4 | n.d. | 39.97 | n.d. | 1.14 | 10.55 | 39.83 | 2.53 | n.d. | 4.37 | n.d. | 55.06 | n.d. | 1.04 | 8.62 | 31.26 | 1.43 | 0.87 | 1.73 |
| LDX-1-4 | n.d. | 32.01 | n.d. | 0.97 | 12.81 | 34.05 | 12.63 | n.d. | 7.54 | n.d. | 47.8 | n.d. | 0.95 | 11.34 | 28.97 | 7.72 | n.d. | 3.23 |
| LDX-1-4 | n.d. | 53.67 | 0.98 | 0.99 | 11.99 | 28.2 | 2.38 | n.d. | 1.8 | n.d. | 67.37 | 0.86 | 0.82 | 8.92 | 20.16 | 1.22 | n.d. | 0.65 |
| LDX-1-1 | n.d. | 53.91 | n.d. | 1.68 | 12.75 | 25.32 | 4.2 | n.d. | 2.13 | n.d. | 67.96 | n.d. | 1.4 | 9.53 | 18.18 | 2.17 | n.d. | 0.77 |
| LDX-1-1 | n.d. | 61.56 | 1.25 | 0.99 | 9.68 | 22.71 | 2.14 | n.d. | 1.68 | n.d. | 74.07 | 1.04 | 0.79 | 6.9 | 15.56 | 1.05 | n.d. | 0.58 |
| LDX-1-1 | n.d. | 63.93 | 1.67 | 2.62 | 8.94 | 16.76 | 1.11 | n.d. | 4.97 | n.d. | 76.52 | 1.39 | 2.06 | 6.35 | 11.43 | 0.54 | n.d. | 1.7 |
| LDX-1-1 | n.d. | 35.98 | n.d. | 1.82 | 11.82 | 35.41 | 8.01 | n.d. | 6.96 | n.d. | 51.68 | n.d. | 1.72 | 10.07 | 28.97 | 4.7 | n.d. | 2.86 |

939 n.d.: not be detected.

| 940 | |
|-----|--|
| 941 | |

Table 4. Carbon and oxygen isotopic compositions for dawsonite and ankerite $\delta^{13}C$ δ^{18} O $\delta^{18}O$ $\delta^{13}C_{CO2}$ Depth Isotopic temp. Well Mineral ‰, PDB ‰, PDB ‰, SMOW °C ‰, PDB m LDX-1-4 Dawsonite -1.2 -6.8 23.9 1894.8 68.53 -7.39 LDX-1-4 1896.07 Dawsonite -1.1 -6.7 24.0 67.83 -7.34 Dawsonite LDX-1-4 69.94 1897.57 -0.6 -7.0 23.7 -6.69 LDX-1-4 1900.2 Dawsonite -1.8 -7.8 22.9 75.76 -7.44 LDX-1-1 1842.5 Dawsonite -4.3 -8.3 22.4 79.56 -9.66 LDX-1-1 1863.34 Dawsonite -3.1 -7.0 23.7 69.94 -9.17 LDX-1-4 1889.07 -2.6 96.44 -7.27 Ankerite -7.8 22.8 LDX-1-4 1894.8 Ankerite -0.8 -7.0 23.7 89.56 -5.94 LDX-1-4 1897.57 Ankerite -0.8 -7.1 23.6 90.31 -5.88 LDX-1-4 1900.2 Ankerite -1.4 -7.0 23.7 89.56 -6.50