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Fast and efficient preparation of 1D chains and dense cold atomic clouds

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We report the efficient and fast (~ 2Hz) preparation of randomly loaded 1D chains of individual ⁸⁷Rb atoms and of dense atomic clouds trapped in optical tweezers using a new experimental platform. This platform is designed for the study of both structured and disordered atomic systems in free space. It is composed of two high-resolution optical systems perpendicular to each other, enhancing observation and manipulation capabilities. The setup includes a dynamically controllable telescope, which we use to vary the tweezer beam waist. A D1 A-enhanced gray molasses enhances the loading of the traps from a magneto-optical trap. Using these tools, we prepare chains of up to ~ 100 atoms separated by ~ 1 µm by retro-reflecting the tweezer light, hence producing a 1D optical lattice with strong transverse confinement. Dense atomic clouds with peak densities up to $n_0 \sim 10^{15}$ at/cm³ are obtained by compression of an initial cloud. This high density results into interatomic distances smaller than $\lambda/(2\pi)$ for the D2 optical transitions, making it ideal to study light-induced interactions in dense samples.

I. INTRODUCTION

The optical response of an ensemble of atoms illuminated by near-resonant light can be significantly different from the one of a single atom due to light induced dipole-dipole interactions [1]. They give rise to collective behaviors such as modified decay rates or spectral linewidths [2–5], or resonance shifts [6–8]. Recently these effects have drawn an increasing interest, for they can be relevant in fundamental optics and have possible applications ranging from optical lattice atomic clocks [9–11] to quantum technologies [12, 13].

In order to enhance the collective optical response of an atomic ensemble, two different paths can be followed. The first one consists in using high-density samples, so that the effect of light-induced dipole interactions is large. This requires the preparation of atomic clouds with densities n fulfilling $n/k^3 \sim 1$ where $k = 2\pi/\lambda_0$ with λ_0 the atomic resonance wavelength. Fundamental questions arise concerning disordered ensembles, such as the existence of Dicke superradiance in small samples [14] or the saturation of the index of refraction for high densities [15]. In disordered clouds, the field radiated by each emitter acquires a random propagation phase that renders difficult the pristine control of interaction effects. The second path thus consists in spatially structuring the cloud at the sub-wavelength scale [16, 17]. In this way, the interferences can be tailored, making it possible to enhance or suppress the effect of dipole interactions. This second route could pave the way to several applications: for example, mirrors made by an atomic layer [16-18], as recently realized using a 2D Mott insulator [5], controlled transport of excitations [19, 20] and light storage [13, 21] or in quantum metrology [12, 13, 22]. The investigation of collective effects in ordered ensembles is also relevant for optical lattice clocks [9, 10, 23], as they could limit their accuracy.

In this paper, we follow the two paths introduced above, relying on a new experimental platform, which we describe and characterize. This platform makes it possible to prepare 1D arrays [24] of ⁸⁷Rb atoms, and disordered atomic ensembles with peak densities reaching $n_0/k^3 \sim 1$. This apparatus is an upgrade of our previous experimental setup [7]. It consists of two high-resolution optical systems with axes perpendicular to one another in a "maltese cross" geometry similar to [25]. These two optical axes used together allow for the simultaneous observation of the fluorescence light emitted by the atoms (incoherent response [26]) and the transmission through the cloud (coherent part [27]). One of the axes is used to focus a tight optical dipole trap (tweezer) to confine the atoms. We have placed in the tweezer beam path a telescope made of two lenses with tunable focal length to dynamically control the tweezer waist. We use this control to prepare chains of atoms with variable length when retro-reflecting the tweezer laser beam, and dense elongated samples after compressing an initially loaded atomic cloud. The loading of the traps from a cloud of laser cooled atoms is enhanced by implementing Λ -enhanced gray molasses.

The paper is organized as follows. Section II describes the optical setup and its alignment, the imaging system, and the *OptoTelescope* that allows to produce optical tweezers with tunable waists. Section III presents the realization of a 1D chain with controllable length and its characterization. Section IV details the enhancement of the trap loading using gray molasses. Section V introduces a new protocol to prepare dense clouds using the tools described before.

II. OPTICAL SETUP, IN-VACUUM LENSES ALIGNMENT AND IMAGING SYSTEM

A. Optical setup

Trapping individual atoms or preparing dense atomic samples requires the waist of the dipole trap beam to be on the order of a few micrometers [28, 29]. This imposes to work with high numerical aperture (NA), diffraction-limited lenses [30]. As represented in Fig. 1, our apparatus is composed of



FIG. 1. Schematic of the experimental setup. Two orthogonal highresolution (NA=0.44) optical systems based on 4 in-vacuum aspheric lenses (AL) create an optical dipole trap, which can be retroreflected to realize a chain of single atoms in a 1D-optical lattice, and collect the scattered light on an electron-multiplying CCD (EMCCD) in two perpendicular directions. On the tweezer axis, the fluorescence is separated from the trapping light using a dichroic mirror (DM). The trap radial size is dynamically controlled with the *OptoTelescope* (OT). All light enters and exits the vacuum chamber through CF40 viewports (Vp). Insert : The x-axis is rotated by an angle of 45° with respect to the plane containing the horizontal beams of the MOT and the z-axis. It is therefore not superimposed to the vertical beam of the MOT, which is in the direction of gravity g.

four in-vacuum aspheric lenses, forming two orthogonal axes in a quasi-confocal configuration. The lenses are manufactured by Asphericon[®][31] and feature effective NA = 0.44. Their working distance (15 mm) is sufficiently large to allow for large optical access, in particular for the six counterpropagating magneto-optical trap (MOT) beams. The plane containing the optical axes of the lenses makes an angle of 45° with respect to the one containing horizontal MOT beams (see Insert Fig. 1): this gives an extra (vertical) access for the atomic beam entering the trapping region. This configuration allows the six MOT beams to be orthogonal, which facilitates alignment and the overlapping with the dipole trap. This also reduces the stray light scattered in the chamber and collected by the imaging system.

The conjugated planes has been optimized using an optical design software to minimize the aberrations of the two crossed optical systems, at both the trapping wavelength $\lambda_{trap} = 940 \text{ nm}$ and the ⁸⁷Rb D2 line ($\lambda_0 = 780 \text{ nm}$), the numerical aperture being fixed. Due to the dispersion properties of the glass of the aspheric lenses, the best performances at λ_{trap} and λ_0 are achieved at different focusing positions for initially collimated beams. For this reason, we work in a tradeoff configuration where the optical performances of the lenses are similar for the two different wavelengths. More precisely, we impose that the wavelength-dependent Strehl ratio (*S*) [32] is the same at λ_{trap} and λ_0 . In our specific case, we calculate S = 0.93, at a distance $d = +285 \,\mu\text{m}$ away from the focal point of a lens at λ_0 . For this configuration, we calculate that



FIG. 2. Sketch of the alignment procedure. A CCD camera is placed at a fixed position d_{780} while we shine a $\lambda_0 = 780$ nm laser beam onto a pinhole acting as a point source for the aspheric lens A. By moving the pinhole with respect to lens A, we optimize the Strehl ratio on the camera and have access to the best focus of this lens. The pinhole is then rotated to face the other lenses.

the image of an object emitting in vacuum at λ_0 is located at $d_{780} \simeq 1119 \,\text{mm}$ in air (see Fig. 2). This distance is used for the alignment procedure of the lenses described in the next section.

B. In-vacuum lenses alignment

The alignment procedure is detailed in [33]. It is experimentally challenging as it involves intersecting two optical axes with a precision much smaller than their field of view $(\pm 50 \,\mu\text{m})$ [34]. We did the alignment in air, correcting for the difference in index of refraction with respect to vacuum. The barrels holding the aspheric lenses are placed inside a metallic lens holder and separated from it with glass spacers. The lens holder is designed such that the angle formed between the two axes is 90° with a tolerance of $\pm 0.1^{\circ}$. The only degree of freedom for each lens is its on-axis position. It is set by tuning the thickness of the glass spacers with a precision of $\pm 1 \,\mu$ m. As represented in Fig. 2, a CCD camera is first placed at a distance d_{780} away from one lens. A pinhole of diameter (1.0 ± 0.5) µm is then mounted on an XYZ translation stage and a rotation stage and placed inside the lens holder. This pinhole is not small enough to be considered as a point source when illuminated by a laser beam at λ_0 . We have taken its finite size into account for the characterization of the performance of the lenses [33]. The pinhole is first moved parallel to the lens axis to minimize the size of its image on the CCD. Once the pinhole is in the targeted object plane, we move it in the transverse plane to maximize the Strehl ratio, thus placing it on the lens optical axis. The pinhole is then rotated by 90° to face another lens. This procedure is performed for each lens and by keeping track of the pinhole motion, we obtain a mapping of the best foci. Finally, the spacers thickness is adjusted to bring all the foci at the same point. After the procedure, we obtain a satisfying alignment of the lenses and the optical axes cross with a residual offset $\lesssim 5 \,\mu\text{m}$, smaller than the field of view of the lenses.

C. Imaging system

The atoms held in the dipole trap are imaged with the two high-resolution axes (Fig. 1), with a diffraction-limited resolution of $1.22\lambda_0/(2NA) \simeq 1\,\mu\text{m}$. Along the trapping axis \hat{z} , the fluorescence or the transmitted light is separated from the trap light using a dichroic mirror and interferometric filters, and is collected by an electron-multiplying CCD (EMCCD) with pixel size $16\mu m \times 16\mu m$ [35]. The magnification of the imaging system along this axis is 6.4, leading to an effective pixel size in the object plane of 2.5 µm: this allows focusing the light emitted by a single trapped atom onto a single pixel, maximizing the signal-to-noise ratio, albeit at the cost of a lowered resolution with respect to the diffraction limit. The fluorescence emitted along the \hat{x} -axis is collected on the same camera, allowing for the observation of the atoms in two orthogonal directions in a single image. The magnification on the transverse axis is \sim 16, leading to an effective pixel size of 1 µm in the object plane. Both resolutions were verified using calibrated pinholes in planes conjugate to the atoms plane. The magnification was confirmed by measuring simultaneously the displacement of trapped atoms on both axes when moving the trapping beam by a known distance. The estimated collection efficiency of both imaging systems is $\sim 4\%$, taking into account the collection of the aspheric lens (5%), the transmission of the optical elements (90%) and the camera quantum efficiency (85% at $\lambda_0 = 780$ nm). This value is confirmed by the measurement of the fluorescence at saturation of a single ⁸⁷Rb atom in a tight dipole trap. As detailed below, we use this atom as a probe to characterize the trap (size and depth), as was done in [30].

D. The OptoTelescope

Our apparatus includes a telescope with tunable magnification, which we name here *OptoTelescope* (OT). This telescope is composed of a pair of 1 inch lenses with voltage-controlled focal lengths, manufactured by OptoTune[®] [36], and placed in an afocal configuration. Tunable lenses allow for the dynamical manipulation of dipole traps [37]. Here, using the OT, we dynamically change the size of the trapping beam before the aspherical lens and thus the optical tweezer waist. To limit aberrations from the OT, we use a beam diameter of $\simeq 1 \text{ mm}$ at its entrance. Also, we minimize gravity-induced aberrations by positioning the optical axis of the lenses vertically. In order to achieve small waists on the atoms, the beam after the OT is magnified by a ×4 telescope before being focused by the aspherical lens. The OT is designed for a magnification ranging from 1 to 3.5.

We characterized the *OptoTelescope* by performing *in situ* measurements on a single atom trapped in the tweezer. For a given magnification, the waist of the trap w_0 is measured as follows. For a fixed power *P*, the peak intensity and thus the light-shift induced by the trap (proportional to the trap depth *U*) are obtained by using a push-out beam expelling the atom from the trap. The light shift is measured from the detuning of this beam for which the push-out effect is the largest, recorded



OptoTelescope magnification

FIG. 3. Dipole trap waist at $1/e^2$ as a function of the OT magnification. The diffraction limit $1.22\lambda_{\rm trap}/2NA \simeq 1.15\,\mu{\rm m}$ is indicated as the smallest trap achievable with the apparatus. The dashed line corresponds to the expected size.

for various trap depths. The trap waist is then extracted using $U \propto P/w_0^2$. The results were checked by independent measurements of the oscillation frequencies of individual trapped atoms [30]. We are able to dynamically change the size of the trap between about 1.6 µm and 4.3 µm, in agreement with the theoretical values calculated using gaussian optics, as shown in Fig. 3.

III. REALIZATION OF A CHAIN OF ATOMS WITH CONTROLLABLE LENGTH

In this section, we present the preparation and characterization of one-dimensional atomic chains of cold Rb atoms, using the tools described in the previous section.

As represented in Fig. 1, we produce the chain by retroreflecting the tweezer beam using the second aspherical lenses placed on the same axis, thus forming a 1D optical lattice with an inter-site spacing $\lambda_{trap}/2 = 470 \text{ nm}$ [38]. The small beam waist of the tweezer ensures a tight transverse confinement. This 1D array in then loaded from the MOT with a filling fraction averaged along the chain of $\simeq 0.25$. We will show in the next section that the loading can be improved up to ~ 0.5 using gray molasses. We collect the fluorescence emitted by the chain in the transverse direction under a 20ms excitation by the MOT beams. A typical example of the atomic chain is shown in Fig. 4(a) (the resolution being about twice the intertrap separation, we do not resolve individual sites). The length of the atomic chain is given by the range around the focal point where the beam intensity is high enough to trap atoms, which is set by the Rayleigh distance $z_{\rm R} = \pi w_0^2 / \lambda_{\rm trap}$. Experimentally, we realize atomic chains with different lengths (and atom number) by tuning the waist of the trapping beam using the OT. As changing the waist also modifies the trap depth, we adapt its power to keep the depth at the center of the chain at $\sim 1 \,\mathrm{mK}$. In Fig. 4(b) we present cuts of the fluorescence along the chain for various lengths. Our longest chains have lengths of $\sim 100 \,\mu\text{m}$ (hence $\sim 200 \,\text{sites}$).



FIG. 4. (a) Averaged image of the fluorescence collected by the transverse imaging axis. (b) Cuts of the fluorescence along the chain for various chain lengths.



FIG. 5. Atom losses as a function of the position in the chain and the modulation frequency of the trapping beam. The dashed lines are the calculated axial oscillation frequencies. The multiple resonances correspond to $2\omega_z/p$ with *p* integer.

To characterize the chain, we measure the local transverse and longitudinal trapping frequencies ω_r and ω_z along the chain axis. To do so, we rely on parametric heating by modulating the intensity of the trapping beam at a given frequency, inducing losses at $2\omega_r$ or $2\omega_z$. Since the trap depth varies along the chain, the oscillation frequencies depend on the position, and so do the resonant frequencies of the parametric heating. Experimentally, we first load a chain from the MOT and take a first reference fluorescence image. The trap beam power is then set at a value of 140 mW while, for this measurement, the waist is set to $3.3 \mu m$. With these parameters we expect $\omega_z \simeq 2\pi \times 1 \text{ MHz}$ and $\omega_r \simeq 2\pi \times 70 \text{ kHz}$ at the center of the chain. The beam intensity is then modulated with a relative amplitude of 5% during 100 ms using an arbitrary waveform generator. A second fluorescence image of the chain is then taken and compared to the reference image to evaluate the atom losses. This sequence is repeated 50 times to average over the chain filling.

Figure 5 shows the atom losses due to the axial excitation. The resonance frequencies extracted with this method are in good agreement with the calculated oscillation frequencies (dashed lines), confirming the expected value of the waist. The different dashed lines reported in Fig. 5, are given by $2\omega_z/p$ with *p* integer. We observe losses at these frequencies since the amplitude modulation is not perfectly sinusoidal and thus contains tones at multiples *p* of the driving frequency. We also observe losses on the chain edges where the trap is the shallowest: these are due to the reference imaging light expelling atoms from the shallow traps, which are thus not recovered in the second fluorescence image. The same experiment was done for radial oscillation frequencies, obtaining also in this case a good agreement between the measured trapping frequencies and the predicted ones.

IV. OPTIMIZATION OF THE LOADING USING A-ENHANCED GRAY MOLASSES

Gray molasses (GM) are commonly used to achieve sub-Doppler cooling of atoms using dark states [39–43]. The use of GM in a tight optical tweezer offers two interesting prospects. First, the low photon scattering rate in dark states reduces light-induced collisions. This yields a higher density of the atomic cloud the tweezer is loaded from, and hence a larger number of atoms in the tweezer. Second, their blue detuning with respect to the atomic frequency should permit to tailor light-induced collisions to selectively remove a single atom out of a pair, resulting into exactly one atom per trap with high probability [44, 45].

We first consider the loading of a single atom in a small (non-retroreflected) tweezer, and apply Λ -enhanced gray molasses [46] on the ⁸⁷Rb D2 line $(\lambda_0 = 780$ nm) [47]. The cooling beam is blue-detuned from the $(5S_{1/2}, F = 2)$ to $(5P_{3/2}, F' = 2)$ transition and superimposed with the six MOT beams with intensity $I \sim I_{sat} = 1.67 \,\mathrm{mW \, cm^{-2}}$ per beam. The coherent repumper is created from the same laser using an electro-optical modulator with frequency equal to the ground state hyperfine splitting v = 6834.68 MHz. The intensity of the repumper is $I \sim I_{sat}/10$ per beam, given by the sideband amplitude. Since gray molasses rely on the blue detuning of the cooling lasers, the optimal detuning will depend on the light-shift induced by the tweezer beam. After the MOT beams are switched off, we study the loading of a single atom from the GM into the tweezer (waist $w_0 = 1.6 \,\mu\text{m}$) varying the detuning of the GM and the trap depth. For each set of parameters, we record the loading probability and the atom temperature, using a release and recapture method [30, 48]. We have found that using the GM on the D2 line does result into individual atoms in the tweezer being colder than when loaded directly from the MOT ($\sim 20 \mu K$ instead of $80 \mu K$), and for a much broader range of the tweezer depth. Also, when loading directly from the MOT, the atoms can be captured in traps with depth $U/k_{\rm B} \sim 1 \,\mathrm{mK}$ while applying the GM stage allows trapping for depth down to $U/k_{\rm B} \sim 200\,\mu {\rm K}$. Furthermore, we observe that the GM detuning does not significantly change the temperature or the loading over a wide range of parame-



FIG. 6. Histogram of the collected fluorescence of atoms in a trap loaded with the GM (blue), in comparison with a trap loaded with a single atom (red). The fluorescence is induced in both cases by 20 ms of MOT beams with detuning 3Γ , where Γ is the natural linewidth of the ⁸⁷Rb D2 line. A background image, without atom, has been subtracted.

ters for detunings between 50 and 120 MHz above the transition and depths $U/k_{\rm B}$ between 200 µK and 1 mK. For larger trap depths and small detunings, the GM frequency becomes resonant with the $(5S_{1/2}, F = 2)$ to $(5P_{3/2}, F' = 2)$ transition, resulting in heating of the atom. However, while we observe efficient cooling when applying the GM, we have not found loading probabilities significantly higher than 50% in a single tweezer, or 25% in chains of traps (retroreflected tweezers), similar to what we achieved with the MOT. This might be due to the fact that the blue-detuned beam on the $(5S_{1/2}, F =$ 2) to $(5P_{3/2}, F' = 2)$ transition is detuned to the red of the $(5S_{1/2}, F = 2)$ to $(5P_{3/2}, F' = 3)$ transition (267 MHz higher in frequency), causing light-induced collisions, which may limit the loading.

To circumvent this issue, we have thus implemented gray molasses on the D1 line $[(5S_{1/2}, F = 2)$ to $(5P_{1/2}, F' = 2)$ transition]. In the single non-retroreflected tweezer, after optimization, we were not able to obtain individual atoms with a probability significantly higher than 50%, whatever the detuning. This is in contrast to what was reported using a bluedetuned beam [44] or GM on the D1 line [45]. To explain this observation, we compare the volume of our tweezer to the one used in Ref. [45] and estimate ours to be a factor of > 10larger. Thus our collision rate is reduced by this factor and the time for blue-detuned light-induced collisions to induce selective losses and leave a single atom in the trap should be much longer than experimentally achievable timescales. We thus infer that more than one atom are left inside the trap. To confirm this, we compare the result of loading via the GM, with the direct loading from the MOT. In one case, we load directly the trap from the MOT: the collisional blockade mechanism operates [28, 30] and when sending near resonant light for imaging, we observe two clear fluorescence levels corresponding to either 1 or 0 atom in the trap. In the other case, we apply a 200 ms-long GM to load the trap and then image the atoms as before. Under this condition, we record a broad fluorescence histogram, as shown in Fig. 6. We explain it by the fact that the initial atom number is large. However the imaging light induces strong losses removing the atoms during the imaging time thus preventing us from counting precisely the *in-situ* atom number.

Finally, we have used D1 gray molasses to improve the loading of the atom chain. We are now able to load a chain of traps with a 50% probability. This is likely due to the fact that on average there are more than one atom per site following the gray molasses loading. The application of the MOT light for imaging then induces strong light-induced collisions, leaving either 0 or 1 atom. Further investigations will be necessary to unravel the loading mechanism of this chain of closely-spaced traps by D1 λ -enhanced gray molasses. We have also found that the loading using GM is more stable than the direct loading from the MOT in terms of daily fluctuations.

V. PREPARATION OF DENSE ATOMIC CLOUDS

As mentioned in the introduction, one of the motivations for our new set-up is the study of light scattering in dense ensembles. We present here a loading protocol based on the new tools of the setup that allows preparing dense enough samples. The main idea is to load as many atoms as possible into a large single-beam dipole trap using GM on the D1 line, and compress the cloud by dynamically reducing the beam waist [49] using the *OptoTelescope*.

We start from a 3D-MOT, which is compressed in 15 ms by red-detuning the MOT beams from -3Γ to -5Γ . We then decrease the magnetic field gradient by 50%. The MOT beams are then switched off and the GM is applied for 200ms, with the dipole trap on. At this stage, the trap depth is $U/k_{\rm B} \simeq$ 4.2 mK and the waist is $w_0 \simeq 2.5 \,\mu\text{m}$ [50]. In this starting configuration, we trap up to 6000 atoms at a temperature of 625 µK yielding a peak density $n_0 \approx 1.6 \times 10^{14} \text{ at/cm}^3$. The use of GM is the key ingredient here that allows for the loading of this large number of atoms. The cloud has an aspect ratio of about 12 along the trapping axis. The atom number is evaluated from the fluorescence collected during the illumination of the cloud with a 10µs-pulse of resonant light and dividing the signal by the same quantity measured with a single atom. To avoid effects caused by light-induced interactions, the imaging pulse in sent after a time-of-flight of 10µs during which the density drops by about an order of magnitude [51]. The temperature is measured by fitting the cloud size for a variable time-of-flight.

The trap is then compressed to a waist $w_0 = 1.8 \mu m$ by changing the magnification of the *OptoTelescope* in 30 ms, keeping the power constant. Next, the trap depth is increased in 10 ms up to 7.6 mK. The duration of the compression has been optimized to be short enough to minimize three-body losses but long enough compared to the response time of the OT lenses (2.5 ms). At this stage, we obtain a cloud of about 2500 atoms in the trap at a temperature of 700 μ K, which corresponds to a cloud peak density $n_0 \sim 10^{15}$ at/cm³ or equivalently to $n_0/k^3 = 1.7 \pm 0.3$. This density is three times larger than the one obtained in clouds of ~ 500 atoms [26, 29] with our previous apparatus which relied on a first, large dipole trap



FIG. 7. Time evolution in the final trap of atom number (a) and temperature (b). In (a) and (b), the solid lines correspond to the solutions of (1) and (2) with L_3 as single fit parameter. (c) Peak density n_0 in the trap, deduced from (a) and (b).

acting as a reservoir to load a second small tweezer.

Such a high density results in large 3-body losses and high elastic collision rates. To characterize them and confirm the extracted value of the density, we study its dynamics. To do so, we have measured the cloud atom number N and temperature T as a function of the time after the end of the compression. The results are shown in Fig. 7(a). The temporal evolution of N and T is described by the following system of coupled equations that takes into account of 2- and 3-body losses [29, 52, 53]:

$$\frac{dN}{dt} = -\gamma_3 \frac{N^3}{T^5} - \gamma_2 \left(\boldsymbol{\sigma}(T), T \right) \frac{N^2}{T} \tag{1}$$

$$\frac{dT}{dt} = \frac{T}{3} \left[\frac{5}{3} \gamma_3 \frac{N^2}{T^5} - \tilde{\gamma}_2 \left(\sigma(T), T \right) \frac{N}{T} \right]$$
(2)

where the parameter γ_3 depends on the trap geometry and is proportional to the 3-body losses coefficient L_3 . The coefficients γ_2 and $\tilde{\gamma}_2$ depend on the temperature, the trap geometry and on the two-body elastic cross-section $\sigma(T)$, whose temperature dependence takes into account the *d*-wave resonance at 350 µK. We interpolate the data of [54] to find a functional form of $\sigma(T)$. We fit the decay of the atom number with the solution of Eq. (1), leaving solely L_3 as a fit parameter. We obtain $L_3 = (4 \pm 1) \times 10^{-28} \text{ cm}^6/\text{s}$. This value is larger [55] than those found in the literature [56, 57]. Note that there exists no prediction for the effect of the *d*wave resonance on 3-body losses, which could enhance L_3 at $T = 650 \mu$ K. We thus do not expect to find the literature values, which were measured deep in the *s*-wave regime. We also compare the model prediction of the temperature evolution to the data [see Fig. 7(b)], and find a very good agreement. The temperature is almost constant, which justifies the assumption of a temperature-independent L_3 (and hence γ_3) in the model. Combining the measurements of the atom number and of the temperature, we calculate the cloud density. Its evolution is shown in Fig. 7(c).

Our experiment is therefore able to efficiently produce microscopic clouds containing up to a few thousand atoms at densities $n_0 \sim k^3$. This corresponds to the regime where the atoms become strongly coupled by light-induced resonant dipole-dipole interactions (scaling as $\hbar\Gamma/(kr)^{\alpha}$ with $\alpha = 1,2,3$). Moreover the repetition rate of the experiment is high: about 2 Hz, limited by the MOT loading. Thanks to this, fast data acquisition is possible, which has allowed us to observe and control subradiance in the time domain [58]. It is in addition a strong asset when measuring, e.g., intensity correlations of the light emitted by the atomic ensemble.

VI. CONCLUSION

We have built an experimental setup that is well-suited for the study of light scattering in cold atom ensembles either in an ordered or disordered configuration. Our platform combines two high-resolution optical systems perpendicular to each other, an optical tweezer with a dynamically tunable waist and gray molasses on the D1 line. By retroreflecting the optical tweezer we create an optical lattice of controllable length, allowing for the preparation of atomic arrays with an average interatomic distance $1.2\lambda_0$. We recently used this feature to investigate a collective enhancement of lightinduced interactions in 1D arrays [8, 59]. The same strategy can be applied with an optical lattice of shorter wavelength (e.g. combining a repulsive optical lattice at 532 nm with the infrared tweezer for confinement). This would increase collective effects even further, enabling the observation of subradiant modes in ordered arrays [21, 60]. Furthermore, we presented a protocol for preparing dense clouds in a tightly focused optical tweezer that exploits the dynamical tunability of the OT. In this way we create clouds with a peak density larger than k^3 at a rate > 2 Hz. The short inter-atomic distances reached in this configuration also offers interesting prospects for investigations of superradiance in small ensembles and subradiance as we recently reported in [58], as well as the study of fundamental questions such as the saturation of the refractive index of dense media [15].

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