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1 **What can we learn from observed temperature and salinity isopycnal anomalies at eddy generation sites?**

2 **Application in the Tropical Atlantic Ocean**

3

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15 **Key points:**

- 16 ● At their generation site, 50% of observed eddies have non-significant isopycnal temperature/salinity
- 17 (θ/S) anomalies in the TAO.
- 18 ● On density-coordinates, both CEs and AEs can exhibit significant positive, negative or non-significant
- 19 isopycnal θ/S anomalies.
- 20 ● We discuss the relationship between isopycnal θ/S and PV anomalies and how they can inform us on
- 21 their generation mechanisms.

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27 **Abstract:**

28 Based on 18 years of satellite altimetry data and temperature/salinity (θ/S) profiles from Argo floats, we
29 analyze the isopycnal θ/S anomalies (θ'/S') within new-born eddies in the tropical Atlantic Ocean (TAO). Our
30 results show that on density-coordinates, both anticyclonic eddies (AEs) and cyclonic eddies (CEs) can exhibit
31 positive, negative or non-significant θ'/S' . Almost half of the sampled eddies do not have significant θ'/S' at their
32 generation site. The other half exhibits significant positive or negative θ' , up to ± 0.5 °C typically. More than
33 70% of them have a subsurface signature. Refined analyses of the vertical structure of new-born eddies in three
34 selected subregions show the dominance of cold (warm) subsurface AEs (CEs) likely due to isopycnal advection
35 of large scale potential vorticity (PV) and θ . PV is a key parameter to analyze eddy generation and dynamics.
36 Isopycnal advection, friction or diapycnal mixing can be involved in the generation of PV anomalies from which
37 vortices can then emerge. However, it is difficult to evaluate PV at mesoscale. In this study we propose to
38 combine θ'/S' and PV anomalies to analyze which process could be involved in its transformation in the ocean.
39 We argue that eddies created by diapycnal mixing or isopycnal advection of water-masses are associated with
40 PV anomalies and significant θ'/S' . In contrast, a frictional forcing also creates PV anomalies but without
41 modifying θ/S . Even though our results remain qualitative, the proposed diagnostics can be of interest to validate
42 realistic models and then use them to analyze the PV anomaly sources.

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44 **Keywords:** mesoscale eddies; isopycnal temperature/salinity anomalies; eddy generation mechanisms; potential
45 vorticity; tropical Atlantic Ocean.

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53 **Plain Language Summary**

54 Mesoscale vortices are common features in the global ocean, having typical length scales of 10 to 100
55 km and lifespans from several weeks to several months or even years for the more energetic structures. The
56 vorticity of these quasi-circular rotating structures emerging from large-scale currents can be formed by several
57 physical processes: isopycnal advection of a water mass, friction (wind stress, bottom stress) or diapycnal
58 mixing. In this study, we first analyze the vertical structure of temperature/salinity in the eddies at their
59 generation site. Based on the analysis of satellite and in-situ data in the tropical Atlantic Ocean, we show that
60 ~50% of the eddies do not present significant isopycnal temperature anomalies when they are generated. In
61 contrast, the remaining ~50% of eddies present significant isopycnal temperature anomalies, but with
62 unexpected numerous cold anticyclones and warm cyclones. We then discuss the link between the potential
63 vorticity and isopycnal temperature anomalies, showing combinations of the previous physical processes
64 involved have to be invoked to explain the observed eddy characteristics.

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78 **1. Introduction**

79 The Tropical Atlantic Ocean (TAO) is a key region for the inter-hemispheric exchange of heat, salt and
80 mass by thermohaline circulation, large-scale currents and mesoscale eddies (e.g. Thomas and Zhai, 2013;
81 Saenko et al., 2018). In the TAO, the upper ocean circulation is mainly composed of i) equatorial limbs of the
82 North and South Atlantic anticyclonic subtropical gyres, ii) zonal equatorial currents and iii) near-coastal
83 current systems (Fig. 1). In the surface layer of the North and South Atlantic subtropical gyres, the salty North
84 and South Atlantic Waters (NAW and SAW, respectively; Fig. 1a) have maximum salinities exceeding 37 in
85 their formation region with densities of $\sim 25.0 - 25.5 \text{ kg m}^{-3}$ (Tsuchiya et al., 1992; Bourlès et al., 1999a;
86 Stramma and Schott, 1999; Stramma et al., 2005a-b Kirchner et al. 2009). In contrast, in the surface layer of the
87 equatorial Atlantic, is found the relatively warm and fresh Tropical Surface Water (TSW; Fig. 1a) with densities
88 lower than 24.5 kg m^{-3} (e.g. Tsuchiya et al., 1992, Tomczak and Godfrey, 1994 Stramma and Schott, 1999 ;
89 Stramma et al., 2005a-b). Fraction of NAW and SAW is advected west-equatorward by the North Equatorial
90 Current (NEC) and the southern South Equatorial Current (sSEC), respectively (Fig. 1a). In the equatorial
91 region this fraction subducts and spreads below the lighter TSW forming a subsurface S maximum in the upper
92 thermocline, known as Subtropical Underwater (STUW; Fig 1b) within the density range $\sigma_{\theta} \sim 24.5-26.3 \text{ kg m}^{-3}$
93 (Snowden and Molinari, 2003; Tsuchiya et al., 1992; Stramma and Schott, 1999; Stramma et al., 2005b). In the
94 upper thermocline of the eastern TAO, another salinity maximum water, known as East Atlantic Water (EAW),
95 is observed, but noticeably fresher than the STUW (Wilson et al., 1994; Bourlès et al., 1999a; Urbano et al.
96 2008; Kirchner et al. 2009).

97 Underneath the surface waters are found the North and South Atlantic Central Waters (NACW and
98 SACW, respectively; Fig 1b). They are characterized by a linear temperature-salinity (θ -S) relationship in the
99 density range $\sigma_{\theta} \sim 25.8 - 27.1 \text{ kg m}^{-3}$ (e.g. Emery, 2003; Liu and Tanhua, 2019) and are connected at around
100 15°N (Sverdrup et al. 1942; Emery and Meincke, 1986; Stramma and Schott, 1999; Stramma et al., 2005b).
101 Below the base of the pycnocline ($\sigma_{\theta} > 26.0 \text{ kg m}^{-3}$), water-masses present θ -S properties close to the Atlantic
102 Subarctic Intermediate Water (ASIW) in the Northern TAO and Antarctic Intermediate Water (AAIW) in the
103 Southern TAO (e.g. Emery, 2003).

104 Although the relatively complex large-scale circulation shown in Fig. 1 is important for the
105 redistribution of water-masses, mesoscale eddies are also known to play a key role in the transfer and
106 redistribution of energy, heat, salt and physical/biogeochemical properties from their generation regions to their
107 dissipation sites (e.g. Chaigneau et al., 2011; Gaube et al., 2014; McGillicuddy 2016). They contribute to the
108 mixing and redistribution of water-masses through several mechanisms such as eddy horizontal stirring, eddy-
109 induced upwelling/downwelling, subduction or trapping and self-advection over long distances across the basin
110 (e.g. McWilliams and Flierl, 1979; Herbette et al, 2004; Chelton et al., 2011), being able to connect eastern and
111 western boundaries (e.g., Laxenaire et al., 2018).

112 Eddy evolutions also strongly depend on their dynamical properties, in particular their potential vorticity
113 (PV) signature which is a key quantity to analyze the formation and dynamics of eddies. Indeed, the core of an
114 eddy in geostrophic equilibrium is necessarily associated with a significant isopycnal PV anomaly, from which
115 the vorticity or stratification of the eddy core can be determined (Hoskins, 1985; Morel and McWilliams 1997;
116 Herbette et al., 2003). Since PV is also conserved for fluid particles in adiabatic and frictionless evolution, the
117 formation of a PV anomaly (and of an eddy subsequently) is necessarily associated with either adiabatic
118 displacement of particles in a background PV gradient, or friction or diapycnal mixing (including the diabatic
119 transformation associated with atmospheric heat and salt fluxes). Many studies have shown how vortices can be
120 formed by the displacement of particles in a background PV gradient. Meridional advection on the planetary
121 beta-plane (Wang, 2005), interaction of currents with sea-mounts, islands, or continental shelves (Aristégui et
122 al., 1994; Herbette et al, 2004) can also be interpreted as creation of PV anomalies by isopycnal advection of
123 particles. These processes are thought to play a significant role in the formation of mesoscale eddies at least in
124 some regions of the TAO, such as east of the North Brazil Current (NBC) retroflexion (e.g. Aguedjou et al.,
125 2019). Some recent studies have shown that non-conservative forcings could also lead to the generation of strong
126 PV anomalies and vortices. Indeed, theoretical and numerical studies have shown that diapycnal mixing (Haynes
127 and McIntyre, 1987; 1990; Morel and McWilliams, 2001) and frictional effects, associated with lateral viscous
128 layers (D'Asaro, 1988; Morel and McWilliams 2001; Akuetevi and Wirth, 2015), the wind (Thomas, 2005;
129 Morel et al, 2006; Holmes et al, 2014; Holmes and Thomas, 2016) or with the bottom boundary layer

130 (Benthuyssen and Thomas, 2012; Gula et al, 2015; 2016; Morvan et al, 2019), are all efficient mechanisms that
131 can modify PV, create PV anomalies and subsequently vortical structures (Morel et al, 2019; Assene et al, 2020).
132 Finally, also note that the well-known formation of eddies by barotropic and/or baroclinic instabilities of mean
133 currents is related to the existence of isopycnal PV gradient of opposite signs and can be interpreted as the result
134 of the creation of opposite sign PV anomalies or dipolar vortical structures (Charney and Stern, 1962; Morel and
135 McWilliams, 2001). The mechanisms responsible for the determination of the PV structure of the unstable mean
136 current is thus of fundamental importance too.

137 If, as noted above, there now exist several studies analyzing the origin of the processes (adiabatic,
138 frictional or diapycnal mixing) involved in the generation of PV anomalies and subsequent formation of eddies
139 in numerical simulations, it remains to be evaluated in nature. This is important because friction and diapycnal
140 mixing are parameterized in numerical models, so numerical results have to be taken cautiously as long as they
141 are not confirmed by observations. However, this is very challenging since the calculation of PV and its
142 evolution require a three-dimensional description of currents and stratification. As such, the detailed diagnostics
143 of PV evolution done in numerical simulation are impossible to reproduce for the real ocean. However, in the
144 present study we propose to separate mechanisms depending on their combined effects on the transformation of
145 PV and transformation of θ/S . Indeed, it has been shown that diapycnal mixing or advection of water masses can
146 lead to significant PV modification accompanied by isopycnal θ/S anomalies (θ'/S'), whose sign and strength can
147 be evaluated from in-situ data. In contrast, the effect of frictional forcings is more difficult to evaluate but
148 friction is expected to create PV anomalies too with very weak isopycnal θ'/S' . We thus argue that isopycnal θ'/S'
149 in vortex cores, where we know there exists significant PV anomalies, can be used to analyze some mechanisms
150 at the origin of the transformation of PV in the real ocean, and represent a first step for the validation of
151 numerical models in this respect.

152 In this study, based on observations, we thus propose to combine in-situ θ/S measurements with satellite
153 altimetry data to estimate the isopycnal temperature/salinity anomalies (θ'/S') of eddy cores at their generation
154 sites in the TAO. In Section 2 we describe the datasets and methods used to characterize θ'/S' inside eddies. In
155 Section 3 we first present θ/S characteristics of the large-scale water-masses over selected isopycnal levels.

156 Second, we show case studies of θ/S' inside individual eddies to illustrate that θ/S' computed from isopycnal
157 levels can strongly differ from the ones obtained along isobaric levels. Third, we characterize the θ/S isopycnal
158 structure of eddies in the entire TAO and focus in some particular areas of the northern TAO. These diagnostics
159 help i) to estimate the fraction of eddies that can participate to the trapping and redistribution of heat and salt
160 through isopycnal advection in the TAO and ii) to depict the mean isopycnal θ/S' in surface and subsurface new-
161 born eddies. Finally, the possibility to infer where diapycnal mixing, isopycnal advection and/or frictional effects
162 could play a significant role in the generation of PV anomalies forming eddy cores and determining their
163 dynamical characteristics is discussed in Section 4, as well as the coherency of this information with the known
164 dynamical features of the TAO.

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166 2. Data and methods

167 2.1 Altimetry data and eddy tracking

168 Mesoscale eddies are identified and tracked in the TAO from daily maps of the Salto/Duacs absolute
169 dynamic topography (ADT) gridded product. This multimission satellite altimetry product, was optimally
170 interpolated onto a $0.25^\circ \times 0.25^\circ$ longitude/latitude daily grid (Ducet et al., 2000; Le Traon et al., 1998;
171 Duacs/AVISO+, 2018; Pujol et al., 2016) and is freely distributed by the Copernicus Marine Environment
172 Monitoring Service (<http://marine.copernicus.eu/>).

173 Eddies were identified from January 2000 to December 2017, using the widely used algorithm
174 developed by Chaigneau et al. (2008; 2009). An eddy is identified by its center, corresponding to a local
175 extremum in ADT, being maximum for an anticyclonic eddy (AE) and minimum for a cyclonic eddy (CE), and
176 its external edge which corresponds to the outermost closed ADT contour around each detected eddy center.
177 Eddy trajectories are constructed according to their polarity (cyclonic or anticyclonic) using the algorithm
178 developed by Pegliasco et al. (2015). Briefly, this algorithm considers as part of the same trajectory, overlapping
179 eddies with the same polarity detected at time t and $t + 1$ day. If several eddies overlap, a cost function is
180 computed to determine the most similar eddy at the time $t + 1$ day. When no overlapping eddy is found neither

181 at time $t + 1$ day nor $t + 2$ day, the trajectory is stopped and the eddy is considered as dissipated. This eddy
182 tracking algorithm is also able to successfully identify eddy-merging or eddy-splitting events (16% of the long
183 lived tracked eddies) that frequently occur in the ocean (e.g. Pegliasco et al., 2015; Laxenaire et al., 2018). As in
184 Aguedjou et al. (2019), in order to consider only long-lived and coherent structures, we retained eddies lasting
185 more than 30 days and having amplitudes and radii greater than 1 cm and 30 km, respectively. A total of ~7800
186 long-lived AEs and ~8100 long-lived CEs were detected in the TAO between 2000 and 2017. Readers interested
187 in more detail regarding the main characteristics and seasonality of these eddies are invited to refer to Aguedjou
188 et al. (2019).

189 **2.2 Argo data**

190 **2.2.1. Argo data processing**

191 The vertical/isopycnal structure of mesoscale eddies is investigated using θ/S profiles acquired by Argo
192 floats in the TAO during the 2000-2017 period. These data were collected and made freely available by the
193 Coriolis project and programs that contribute to it (<http://www.coriolis.eu.org>). In the TAO, a total of 144033
194 θ/S profiles, flagged as good or probably good by the Coriolis Data Center, were quality controlled following
195 Chaigneau et al. (2011) and Pegliasco et al. (2015). First, we retained profiles for which: i) the shallowest data is
196 not deeper than 15 m depth and the deepest acquisition is below 950 m depth, ii) at least 30 data are available
197 between the surface and 950 m depth and iii) the depth difference between two consecutive data does not exceed
198 a given threshold (25 m for the 0–150 m layer, 50 m for the 150–300 m layer, 75 m for the 300–500 m layer and
199 100 m below 500 m depth). Second, each profile was visually inspected and was systematically discarded if it
200 presented a suspicious θ/S profile. The remaining 114440 profiles (Fig. 2a), representing ~80% of the original
201 database, were then linearly interpolated every 5 m from the surface to 300 m depth, and every 10 m below. We
202 here assumed that the shallowest values (within 0-15 m depth) were representative of the surface values. The
203 spatial distribution of the retained Argo profiles is rather homogeneous in the TAO, except along continental
204 shelves or around islands where Argo floats cannot reach their nominal parking depths (Fig. 2a). About 55% of
205 these profiles were acquired in the Northern Hemisphere, due to the higher number of international cruises and

206 Argo float deployments in this region. The temporal distribution of valid Argo profiles in the TAO shows a
207 linear increase from ~200 yearly profiles in 2000 to ~12000 yearly profiles after 2017 (Fig. 2b).

208 **2.2.2. Argo data within new-born eddies**

209 The 114440 retained θ/S profiles were then classified into three categories depending on whether Argo
210 floats surfaced within AEs or CEs (detected from altimetry) or outside eddies. However, along their trajectories,
211 TAO eddies exhibit a three-step evolution composed of growth, mature and decay phases (Pegliasco et al., 2015;
212 Sun et al., 2018; Aguedjou et al., 2019). During the growth phase of the eddies, which extends over the first
213 ~30% of their lifespan, their radius, amplitude, and kinetic energy strongly increase (Pegliasco et al., 2015; Sun
214 et al., 2018; Aguedjou et al., 2019). Also during this phase, eddies traveled on average 200 km from their
215 generation sites (See Appendix A3 and Fig. A4d). Thus, in order to investigate the isopycnal structure of new-
216 born eddies, close to their generation sites, and infer the physical mechanisms involved in their generation, we
217 only retained θ/S profiles acquired by Argo floats that surfaced within eddies located closer than 200 km from
218 their formation sites (Fig. 2c). Around ~25% of the ~16000 long-lived eddy trajectories were sampled within
219 200 km from their birth sites and are thus retained for the analysis of the thermohaline structure of new-born
220 eddies presented in Section 3.4. The spatial distribution of the Argo floats that sampled these new-born eddies is
221 shown in Fig. 2c. Note that varying the distance between 100 km and 300 km do not change the main results and
222 discussion drawn in this study (not shown). We thus selected 200 km, which is a good compromise between
223 being close enough from the eddy formation sites and having a satisfactory number of eddies sampled by Argo
224 floats.

225 Figure 2d shows the mean distribution of Argo floats that surfaced within new-born eddies, as a function
226 of the normalized radial distance from eddy centers. A normalized distance equals to 1 corresponds to the
227 equivalent eddy radius inferred from the area enclosed by the ADT contour of the eddy-edge. Within these new-
228 born eddies, 50% (75%, respectively) of the Argo floats surfaced within a normalized distance of 2/3 (0.9) the
229 eddy equivalent radius from the eddy center (Fig. 2d). About 20% surfaced at a normalized radial distance

230 higher than 1. This occurs when eddies are deformed and the profilers surface between the equivalent radii and
231 eddy edges.

232 **2.3 Isopycnal temperature anomalies in new-born eddies**

233 These depth-dependent θ/S profiles were first projected onto density-coordinates (here density refers to
234 potential density referenced to the sea-surface). Second, isopycnal θ' and S' , were inferred for each profile by
235 removing a local climatological profile representative of the large-scale background, also computed on density-
236 coordinates. In a similar way to Castelao et al. (2014) and Pegliasco et al. (2015), local climatological profiles (
237 \bar{P}), were obtained by weighted arithmetic means of all the available profiles (P_i) acquired outside eddies,
238 within a radius of 200 km and separated by less than ± 30 days (independently of the year) from the date of the
239 considered profile within the eddy (see Appendix A1). Selecting outside-eddy profiles collected no more than 30
240 days apart of the day of the year in which the eddy profile was collected, allows us to filter out unwanted θ/S
241 seasonal variations.

242 Vortices were then further classified into three main categories, depending on whether isopycnal θ' were
243 i) significant (positive or negative) in the surface layer, extending from the surface to the base of the pycnocline,
244 ii) significant (positive or negative) in the subsurface layer below the base of the pycnocline, or iii) not
245 significant neither above nor below the pycnocline (see Appendix). The corresponding eddies are then qualified
246 as surface eddies, subsurface eddies or eddies with non-significant anomalies, respectively. Eddies having both
247 surface and subsurface significant isopycnal θ' are considered as subsurface eddies. Note that the nature (surface
248 or subsurface) of the eddies is often defined based on kinematic properties such as geostrophic velocity (e.g.
249 Chaigneau et al., 2011; Pegliasco et al., 2015), vorticity (e.g. Assene et al., 2020), or Okubo-Weiss parameter
250 (e.g. Xu et al., 2019). Here in contrast, it is the vertical position of the significant isopycnal θ' within the eddies
251 (above or below the pycnocline, respectively) that defines their nature (surface or subsurface).

252 In order to determine whether θ' within eddies are significant or not, an isopycnal θ' threshold was
253 defined in a $1^\circ \times 1^\circ$ longitude/latitude grid at seasonal scale, from Argo profiles that surfaced outside eddies (see
254 Appendix). Figures 3a-b show the annual mean of θ' thresholds for the surface and subsurface layers,

255 respectively. In general, a given θ' profile acquired within an eddy is significant within the surface layer when
256 the square root of its quadratic mean values (see Eq. E4 in Appendix) integrated over the surface layer is greater
257 than 0.2-0.5 °C, except for some regions such as the frontal zone separating NAW from TSW, where θ' threshold
258 reaches up to 0.8 °C (Fig. 3a). In contrast, threshold values of θ' within the subsurface layer are much lower and
259 generally less than 0.3 °C. However, around the frontal zone along which the NEC is flowing, high θ' threshold
260 values are still noticed reaching up to 0.8 °C (Fig. 3b). Note that a classification based on Student's t-tests was
261 also performed and led to similar results at a 99% confidence level. For a given density, θ' and S' are
262 proportional and of the same sign, so that isopycnal maps of θ' or S' are similar by definition. We thus
263 hereinafter only focus on θ' along isopycnal levels.

264 It is important to note that the ocean mixed layer is the site of intense ocean-atmosphere interactions that
265 impact the mixed-layer properties from diurnal to inter-annual timescales (e.g. Guemas et al., 2011; Rugg and
266 Foltz, 2016). For instance, diurnal variations of the solar heating can modify the daily mean SST by about a 0.3°-
267 0.5 °C (even greater for extreme events) in the TAO (e.g. Guemas et al., 2011; Gentemann et al., 2008;
268 Wenegrat and McPhaden, 2015). Similarly, interannual variations of ± 0.5 °C are commonly observed in the
269 TAO (e.g. Carton et al., 1996; Rugg and Foltz, 2016; Foltz et al., 2019). Thus, using Argo floats' data within the
270 mixed-layer induces relatively strong positive or negative θ/S anomalies that are not directly related to eddy
271 dynamics and are difficult to interpret. Also, this high variability, can lead to a misclassification of eddies as the
272 eddy signal in the mixed layer can be more variable than below the mixed layer. Thus in this study, surface
273 eddies refer to eddies having significant θ' between the base of the mixed layer to the base of the pycnocline.

274 3. Results

275 3.1 Large-scale distribution of isopycnal temperature in the TAO

276 In order to better understand the eddy signature on θ along isopycnal surfaces, we first briefly depict the
277 large-scale water-mass θ distribution in the TAO along two isopycnal levels. As such, Fig. 4 presents the mean θ
278 on $\sigma_\theta = 25.5 \text{ kg m}^{-3}$ and $\sigma_\theta = 27.2 \text{ kg m}^{-3}$ levels, obtained from Argo floats that surfaced outside eddies. These
279 two particular σ_θ levels were retained because the maximum isopycnal θ' within surface/subsurface eddies are

280 mostly observed close to $\sigma_\theta = 25.5/27.2 \text{ kg m}^{-3}$ (see Fig. 7). On the shallower/lighter density-level, we can
281 observe the noticeable warmer areas located within the subtropical gyres where NAW and SAW are located
282 (Fig. 4a,d). On $\sigma_\theta = 25.5 \text{ kg m}^{-3}$ isopycnal, these water-masses θ ranges from 24 to 24.5 °C and 22 to 23 °C,
283 respectively (Fig. 4a,d). The along-isopycnal θ decreases equatorward from the gyre centers. In the eastern TAO,
284 where EAW is originated, θ decreases to ~ 20 °C and S is lower than 36 (see also on Fig. 4d). Along the equator
285 a relative warm core ($\theta \sim 22$ °C) water-mass is advected eastward by the EUC (see Fig. 1 and Fig. 4a). The θ of
286 this water-mass slightly decreases eastward and the water-mass is also characterized by relatively high salinities
287 (not shown but see, e.g. Hormann and Brandt, 2007; Kolodziejczyk et al., 2009; Da-Allada et al., 2017). Finally,
288 in the eastern boundary upwelling systems (Canary and Benguela), the mean θ is lower than 18 °C on $\sigma_\theta = 25.5$
289 kg m^{-3} due to the presence of the wind-forced coastal upwelling (Fig. 4a). The $\sigma_\theta = 25.5 \text{ kg m}^{-3}$ isopycnal layer
290 deepens from ~ 40 m in the eastern TAO to ~ 140 m in the western TAO at latitudes of $\pm 20^\circ$ (black lines in Fig.
291 4a). The westward deepening of this isopycnal layer, which is associated with the lower part of the
292 thermocline/pycnocline, is reduced along the equator where its depth varies from 50 m in the Gulf of Guinea to
293 ~ 100 m off Brazil.

294 On the $\sigma_\theta = 25.5 \text{ kg m}^{-3}$ ($\sigma_\theta = 27.2 \text{ kg m}^{-3}$, respectively) density layer, θ distribution shows a strong
295 isopycnal θ front that separates the warmer and saltier NAW (NACW) from the colder and fresher EAW
296 (SACW) and extends across the basin. East of $\sim 30^\circ\text{W}$, this front is known as the Cape Verde Frontal Zone
297 (CVFZ) (Zenk et al; 1991; Pérez-Rodríguez and Marrero-Díaz, 2001; Martínez-Marrero et al., 2008; Tiedemann
298 et al., 2018), which is an active area of water-mass exchange associated with the formation of mesoscale eddies
299 (e.g. Dadou et al., 1996; Schütte et al., 2016). However, we hereafter simply refer to the frontal zone to indicate
300 the connection area between NACW and SACW. The NEC flows along this frontal zone, which is oriented
301 southwestward from 20°N in the eastern TAO to 10°N in the western TAO (see also Fig. 1). This front is clearly
302 visible down to $\sigma_\theta = 27.2 \text{ kg m}^{-3}$ (Fig. 4b) where it is more diffuse and the strongest θ gradients are observed
303 southward, between 10°N and the equator. At this density level, the frontal zone is much more zonal, and
304 NACW and SACW have typical θ of 10-11 °C and 5-6 °C, respectively (Fig. 4b,d). The mean depth of this

305 isopycnal level is of ~ 650 m in a large part of the TAO (black lines in Fig. 4b), and deepens to 750-900 m
306 poleward of latitudes $\pm 20^\circ$.

307 Figure 4c presents the meridional θ section at 35°W . Warmer waters with θ greater than 20°C are
308 located above the $\sigma_\theta = 26.0 \text{ kg m}^{-3}$ isopycnal layer and are associated with the previously described NAW, SAW
309 and TSW. TSW is the warmest water-mass, with θ reaching 28°C around the equator, but exhibits the lowest
310 surface salinity of 35.5-36 (not shown) due to the excess of precipitation to evaporation in this area. Below the σ_θ
311 $= 26.0 \text{ kg m}^{-3}$ density layer, there are the distinguishable warmer NACW and cooler SACW, separated by the θ
312 front that becomes more diffuse below $\sigma_\theta = 27.0 \text{ kg m}^{-3}$.

313 θ/S diagrams and the main water-mass characteristics found in the TAO are depicted in Figure 4d. They
314 were constructed from Argo floats that surfaced outside eddies within the northern, southern and eastern parts of
315 the TAO (see boxes delimited in white in Fig 4a). This Figure confirms that TSW is much fresher and lighter
316 than NAW or SAW and that NACW is warmer and saltier than SACW along isopycnal levels.

317 **3.2 Case studies of iso-depth versus isopycnal temperature anomalies in mesoscale eddies**

318 Estimates of θ and S anomalies within an eddy can strongly differ whether we use a depth-coordinate
319 system, where the local vertical displacement of isopycnal levels strongly impacts θ'/S' , or a density-coordinate
320 system. In order to better familiarize the reader with this concept, and to help the interpretation of the results
321 described in the following Sections, we here describe θ' observed in three individual mesoscale AEs at their
322 generation sites, using both the depth- and density-coordinate systems. These three eddies (AE1, AE2 and AE3,
323 respectively), detected by their ADT signature, were sampled by Argo floats in the western (AE1, Fig. 5a),
324 central (AE2, Fig. 5b) and eastern parts (AE3, Fig. 5c) of the northern TAO. As typically observed in the TAO
325 (Aguedjou et al., 2019), they have amplitudes of $\sim 2\text{-}4$ cm, radii of $\sim 70\text{-}100$ km, eddy kinetic energies of $85\text{-}100$
326 $\text{cm}^2 \text{ s}^{-2}$, and mean vorticities of $2\text{-}5 \times 10^{-5} \text{ s}^{-1}$.

327 Although the 3 eddies were sampled by Argo floats within their core, AE1 and AE2 were sampled
328 relatively close to their edge (Fig. 5a-b), whereas A3 was sampled in the vicinity of its center (Fig. 5c). As
329 expected for AEs, θ' observed using a depth-coordinate system are mainly positive for the 3 case-study eddies

330 (Fig. 4d-f, magenta lines). Indeed, in depth-coordinates, the isopycnal deepening (rising, respectively) occurring
331 inside AEs (CEs), is the main driver of the observed positive (negative) θ' (e.g. Assassi et al., 2016; Keppler et
332 al., 2018).

333 The θ' computed in depth-coordinates (Fig. 5d, magenta line) suggest that AE1 is a subsurface eddy with
334 a core extending from 100 m to 600 m depth. These anomalies are in part due to the observed ~ 14 m deepening
335 of the subsurface isopycnal levels in AE1 (not shown). However, θ' computed in density-coordinates (but plotted
336 in Fig 5d as a function of depth by associating each isopycnal level to its depth in the analyzed profile), are also
337 significantly positive in the $\sigma_\theta \sim 25.0 - 27.4 \text{ kg m}^{-3}$ isopycnal layer located between 100 and 800 m depth, with
338 the maximum of ~ 1 °C reached slightly below $\sigma_\theta = 27.0$ (Fig. 5d, red line). These positive isopycnal θ' suggest
339 that diapycnal mixing or isopycnal advection from remote regions occurred during the formation of AE1. Fig. 5g
340 compares the θ/S diagrams for the Argo profile acquired inside the vortex (red) and the corresponding
341 climatological profile obtained from profiles outside eddies (green). It can be misleading since there exists
342 strong discrepancies between both profiles in low density ranges ($\sigma_\theta = 24.0-25.5 \text{ kg m}^{-3}$) (Fig. 5d). However,
343 given the curvatures of isopycnal lines and of the vertical profiles, the strongest isopycnal eddy θ/S anomalies
344 are obtained in the subsurface layer containing NACW. Thus, for this particular AE1 case-study, θ' are located in
345 subsurface and of the same sign (positive) for both coordinate systems.

346 In AE2, θ' computed in depth-coordinates are positive between the surface and 250 m depth, and slightly
347 negative below 250 m (magenta line in Fig. 5e). The maximum anomaly is of ~ 1.5 °C at ~ 150 m depth, thus
348 AE2 is likely a surface eddy. In A2, these positive anomalies are associated with a deepening of the near-surface
349 isopycnal levels by ~ 45 m (not shown). However, when computing θ' in density-coordinates, AE2 is
350 characterized by negative θ' in the surface layers reaching maximum negative anomalies of -1.5 °C for $\sigma_\theta \sim 26.0$
351 kg m^{-3} at ~ 170 m depth (red line in Fig. 5e). Again, these strong negative isopycnal θ' suggest that AE2 contains
352 water from a remote region having distinct $\theta-S$ characteristics, or that diapycnal mixing, likely associated with
353 air-sea fluxes, locally modified the thermohaline structure of the water column. AE2 clearly illustrates the
354 differences that can exist when using depth- versus density-coordinates to compute θ/S anomalies.

355 Finally, in AE3, significant positive θ' computed from depth-coordinates are observed from the surface
356 to 200 m depth (Fig 5f) with maximum anomalies of ~ 3.5 °C observed at ~ 50 m depth. However, isopycnal θ'
357 show that this eddy does not contain significant water-mass anomalies relative to the background large-scale
358 environment (Fig. 5f). Thus, θ' observed in depth-coordinates only result from the local deepening of ~ 12 m
359 observed for the near-surface isopycnal layers (not shown). The isopycnal θ - S structure within AE3 is similar to
360 the one usually observed in this region, with EAW and SACW in the surface and subsurface layers, respectively
361 (Fig. 5i).

362 To summarize, based on 3 case-study AEs, we have shown that isopycnal θ' anomalies can strongly
363 differ from anomalies computed from the more commonly used depth-coordinate system. Obviously, similar
364 conclusions hold for CEs that generally depict negative anomalies in depth-coordinate system, but that can show
365 positive, negative or non-significant anomalies in density-coordinate system. These discrepancies are
366 problematic, in particular for the estimates of anomalous heat or salt eddy contents, and their associated
367 transports, which are commonly computed in depth-coordinates instead of using density-coordinates.

368 **3.3 Spatial distribution of isopycnal temperature anomalies in TAO new-born eddies**

369 Figure 6 shows the spatial distribution of surface and subsurface eddies, as well as eddies with non-
370 significant anomalies at their generation sites in the TAO. These sampled new-born eddies have a mean
371 amplitude, radius and lifespan of ~ 3 cm, ~ 90 km, and 150 days respectively. These mean values are associated
372 with standard errors of ~ 0.2 cm, ~ 2 km and ~ 7 days, respectively. A first striking result is that about half of the
373 analyzed new-born AEs and CEs have a weak and non-significant isopycnal θ' relatively to their local
374 environment at their generation site, with their mean θ' ranging mostly between ± 0.1 °C. They cover all areas
375 and their spatial distribution is quite identical for AEs and CEs (Fig. 6e-f). Although not associated with
376 significant isopycnal θ' , these eddies are associated with an average vertical displacement of ± 10 m of their
377 isopycnal levels, similar to those observed in eddies with significant anomalies, which shows that both type of
378 vortices are associated with dynamical signature of similar strength. Likewise, Chen et al., (2021) reported that
379 eddies induce an average vertical displacement of the pycnocline of the same order of magnitude in tropical

380 regions. As a result of the pycnocline vertical displacement, more than 90% of the new-born eddies that do not
381 have significant isopycnal θ' anomalies, exhibit significant θ' in depth coordinates.

382 Thus, about half of the new-born eddies have significant θ' anomalies at their generation sites (Fig. 6a-
383 d). Around 25% of these eddies are surface eddies whereas the remaining (~75%) exhibit significant subsurface
384 θ' . Note however, that ~30% of these subsurface eddies also exhibit significant surface anomalies. The spatial
385 distribution of AEs and CEs having significant isopycnal θ' is almost similar in surface and subsurface layers
386 (Fig. 6a-b and 6c-d). In the Southern Hemisphere, eddies with significant θ' are very few and their maximum
387 isopycnal θ' (± 0.3 °C) are observed around the Benguela upwelling system. In contrast, in the Northern
388 Hemisphere, positive and negative θ' are more similarly distributed with magnitude reaching up to ± 0.8 °C along
389 the frontal zone (Fig. 6a-b). Within the subsurface layer, AEs and CEs in the Northern Hemisphere are mostly
390 characterized by negative and positive θ' , respectively, contrary to what is observed in the Southern Hemisphere
391 (Fig. 6c-d). Again, this observed dominance of relatively cold AEs and warm CEs in subsurface is unusual when
392 working in depth-coordinates where the sign of the anomalies is largely driven by the deepening/rising of
393 isopycnal layers (see Section 3.2).

394 As described in Section 1, different mechanisms can create PV anomalies and generate eddies. However,
395 depending on the involved mechanisms, the newly formed eddies either have isopycnal θ'/S' or not. Frictional
396 mechanisms such as wind, lateral friction, or current shear, produce PV anomalies without creating isopycnal
397 θ'/S' . In contrast, the formation of PV anomalies by advection of water-masses with different characteristics or
398 by diapycnal mixing is associated with isopycnal θ'/S' . Thus, from the results obtained above, we can conclude
399 that about 50% of the TAO eddies are likely formed by frictional mechanisms, whereas 50% are likely formed
400 by diapycnal mixing or lateral advection of distinct water-masses. The majority of these new-born eddies have
401 distinct θ'/S' in subsurface isopycnal layers with cold anomalies in AEs and warm anomalies for CEs in the
402 Northern TAO. We decided to further study this area in the following section.

403 3.4 Eddy isopycnal structures in selected northern TAO regions

404 Three sub-regions (R1 to R3, see Fig. 3-6) were defined according to their large-scale dynamics and
405 characteristics (Fig. 1, Fig. 3 and Fig. 4a-b). R1 extends along the frontal zone separating NACW from SACW
406 (Fig. 3). The westward NEC which flows along the frontal zone is mainly fed by the eastern branch of the North
407 Atlantic subtropical gyre but also by the northern branch of the Guinean Dome (Fig. 1). Further west, a part of
408 the NEC retroflects cyclonically and feeds both the eastward North Equatorial Counter Current (NECC) and the
409 eastward North Equatorial Undercurrent (NEUC) (Fig. 1a-b; e.g. Stramma and Schott, 1999; Bourlès et al.
410 1999a; Schott et al., 2004). The large-scale isopycnal θ distribution (Fig. 4) shows that the NEC flows along the
411 frontal zone that separates relatively warm and salty waters of the North Atlantic subtropical gyre from cooler
412 and fresher waters of the equatorial Atlantic Ocean (Fig. 4). This frontal zone is known to exhibit strong
413 meanderings and eddy generations (e.g. Dadou et al., 1996; Shütte et al., 2016; Aguedjou et al., 2019). In R2,
414 that extends from north-east Brazil to West Africa between 0°N and 10°N (Fig. 6), is found the zonal equatorial
415 dynamics (Fig. 1). In this region, strong instabilities are frequently observed, in particular due to the horizontal
416 shear between the nSEC and NECC (Fig. 1) (e.g. Weisberg and Weingartner, 1988; Kelly et al., 1995; Athié and
417 Marin, 2008; Von Schuckmann et al., 2008, Aguedjou et al., 2019). Finally, subregion R3 includes the NBC
418 retroflexion (Fig. 1a-b), and is populated by relatively large and energetic eddies whose surface properties (size,
419 amplitude, eddy kinetic energy) exhibit a strong seasonal variability (e.g. Aguedjou et al., 2019).

420 Figure 7 shows for these 3 subregions the mean vertical θ' profiles inside new-born AEs and CEs using
421 density-coordinates. As already discussed, the 3 sub-regions show a high number of non-significant anomaly
422 profiles representing 40- 60% of observed eddies (Fig. 7, in black).

423 In R1, $\sim 75\%$ of the new-born eddies (AEs and CEs) having a significant θ' are subsurface eddies (Fig.
424 7a,d). Surface AEs (solid lines in Fig. 7a) and CEs (solid lines in Fig. 7d) are characterized by average θ' of
425 $\pm 0.5\text{-}0.7^{\circ}\text{C}$ between $\sigma_{\theta} \sim 25.5$ and 26.0 kg m^{-3} . Subsurface AEs (dotted lines in Fig. 7a) and CEs (dotted lines in
426 Fig. 7d) show maximum anomalies of $\pm 0.4^{\circ}\text{C}$ between $\sigma_{\theta} \sim 27 \text{ kg m}^{-3}$ and $\sigma_{\theta} \sim 27.2 \text{ kg m}^{-3}$ except subsurface
427 cold CEs whose mean anomaly of -0.5°C is found between $\sigma_{\theta} \sim 26\text{-}26.5 \text{ kg m}^{-3}$. Interestingly, around 73% (65%,
428 respectively) of these subsurface eddies are cold (warm) for AEs (CEs).

429 In R2, the surface eddies have strong anomalies of ± 0.5 °C just below the mixed layer, with the
430 maximum located in the density range of the eastward NECC and the westward nSEC, which carry relatively
431 warm and cold water, respectively (Fig. 1a and Fig. 4a). In the subsurface layer, similarly to what is observed in
432 R1 with a lower contrast with the surface layer, $\sim 65\%$ of the CEs and $\sim 70\%$ of the AEs that present a significant
433 θ' are of subsurface (Fig. 7b,e). Around 82% (65%, respectively) of the subsurface sampled AEs (CEs) have cold
434 (warm) anomalies and their maximum anomalies of $\sim \pm 0.3$ °C are located at $\sigma_\theta \sim 27.2$ kg m⁻³ (Fig. 7b,e).

435 In R3 (Fig. 7c,f), unlike in the two previous areas, very few new-born eddies were sampled by Argo
436 floats. Moreover $\sim 55\%$ of the AEs and $\sim 65\%$ of the CEs sampled at their generation sites do not exhibit
437 significant θ' . The remaining sampled new-born eddies are characterized by relatively large anomalies, reaching
438 ± 1 °C for CEs and $\sim \pm 0.7$ °C for AEs. Note, that the relatively small number of sampled eddies is due both to the
439 very few number of Argo floats in R3 and to the reduced size of this area. AEs characterized by positive θ' (red
440 lines in Fig. 7c) show anomalies of ~ 0.5 °C, observed at $\sigma_\theta = 26$ kg m⁻³ for surface AEs, and $\sigma_\theta = 26.8$ kg m⁻³ for
441 subsurface AEs. Surface CEs predominantly show strong positive and negative anomalies that are maximum at
442 $\sigma_\theta = 25.5$ kg m⁻³ (solid lines in Fig 7). In contrast, subsurface CEs mainly show positive θ' of 0.5 °C at the base of
443 the pycnocline or at $\sigma_\theta \sim 27$ kg m⁻³ (dotted lines in Fig. 7f). Note that the 3 profiles classified as cold subsurface
444 CEs (blue dotted line in Fig. 7e) are characterized by strong positive anomalies between $\sigma_\theta = 24.5$ kg m⁻³ and $\sigma_\theta =$
445 26.7 kg m⁻³ and weak negative anomalies in deeper levels. Integrated in the water column, these deeper negative
446 anomalies, that occupy a thicker layer, prevail and lead to a negative heat content anomaly that explains the
447 classification of these profiles as cold subsurface CEs. In R3, the sampled new-born eddies were mainly formed
448 on the eastern flank of the NBC retroflection (not shown), where relatively strong θ gradients exist. In the
449 subsurface layer, isopycnal mixing of NACW and SACW takes place (eg. Kirchner et al., 2009).

450 4. Qualitative analysis and discussion

451 4.1 Eddy vertical structure and potential implications for heat and salt transports

452 The first striking result of our analysis, underlined in Fig. 6, is that about half of the analyzed eddies
453 have non-significant isopycnal θ'/S' at their generation site. The other half is highly dominated by subsurface

454 eddies, with the strongest anomalies within the frontal zone in the Northern Hemisphere. The high number of
455 subsurface eddies obtained in our study have been also previously reported in different areas of the TAO, from
456 both numerical simulations and in-situ data. For instance, in the Canary and Benguela upwelling systems,
457 Pegliasco et al. (2015) estimated that subsurface eddies represent between 40 and 60% of the total number of
458 eddies. Although their analysis was based on depth-coordinates and did not specifically focused on new-born
459 eddies, these authors also showed that 40-60% of the sampled eddies do not have statistically significant θ/S' , in
460 agreement with our results. Other studies also revealed the presence and persistence of subsurface structures in
461 various parts of the TAO, in the eastern and western boundary currents regions and in the subtropical gyres (e.g.
462 Schütte et al., 2016; Garraffo et al., 2003; Assene et al., 2020; Amores et al., 2017; Laxenaire et al., 2020).
463 However, as already discussed by Pegliasco et al. (2015), the high number of subsurface eddies could be
464 influenced by a slight sampling bias in Argo data. Indeed, Argo floats drift for 10 days at a nominal parking
465 depth of 500-1000 m and may therefore be preferentially trapped within subsurface eddies that reach these
466 depths.

467 In terms of tracer transport, our results would suggest that only half of the analyzed eddies (eddies with
468 significant anomalies) might potentially contribute to the heat and salt transport within the basin. However, the
469 efficiency of the eddies to redistribute heat and salt through isopycnal advection from their formation region
470 strongly depends on the ambient θ/S characteristics at their dissipation sites and whether or not the trapped and
471 transported water-masses exhibit similar properties than the ambient ones. In the same way, other significant
472 processes such as diapycnal mixing along eddy trajectories, local eddy-stirring transport associated with the
473 lateral advection of large-scale thermohaline gradients (e.g. Sun et al., 2019) were not considered in our
474 estimation. Making a precise quantification of the eddy heat/salt transport exceeds the scope of this study.
475 However, it is important to recall that θ/S' computed in depth-coordinates are largely influenced by the vertical
476 displacement of isopycnal layers and do not reflect the net effect of eddy heat and salt contents that participate to
477 the redistribution of heat and salt in the global ocean. As such, isopycnal coordinates must be considered when
478 evaluating heat/salt transport by eddies.

479 **4.2 Effect of isopycnal advection, friction and diapycnal mixing on the generation of PV anomalies**

480 The equation of conservation of momentum and density can be written as

481
$$\frac{d}{dt} \vec{U} + \vec{f} \times \vec{U} = \frac{-\vec{\nabla} P}{\rho_o} - \vec{g} \frac{\rho}{\rho_o} + \vec{F}$$
 (1)

482
$$\nabla(\vec{U})=0$$
 (2)

483
$$\frac{d}{dt} \rho = \dot{\rho}$$
 (3)

484 where $\vec{U}=(u, v, w)$ is the velocity field, \vec{f} is the Earth rotation vector, whose projection on the local vertical
 485 axis defines the Coriolis parameter f , P is the pressure, ρ the density of seawater and ρ_o a mean density, \vec{g} is the
 486 gravity, $\frac{d}{dt} = \partial_t + u \partial_x + v \partial_y + w \partial_z$ is the Lagrangian derivative. Finally \vec{F} and $\dot{\rho}$ are non-conservative
 487 terms respectively representing frictional effects (acting on momentum) and diapycnal mixing (acting on θ
 488 and/or S and finally modifying the particle potential density). As shown by Ertel (1942; see also Muller, 2006;
 489 Morel et al, 2019; Assene et al, 2020), we can define a generalized PV as:

490
$$PV = -(\vec{\nabla} \times \vec{U} + \vec{f}) \cdot \vec{\nabla} Z(\rho)$$
 (4)

491 The traditional Ertel PV (obtained with $Z(\rho) = \rho$) is dominated by the signature of the pycnocline, so that the
 492 dynamical signal associated with isopycnal variations of PV is difficult to identify. In the following, we will use
 493 the rescaled PV, as proposed in Delpech et al (2020) and Assene et al (2020), determined by choosing $Z(\rho^*) = z$
 494 for a specific location, where the density profile $\rho^*(z)$ is typical of the stratification of the area. The rescaled PV
 495 is close to the quasi-geostrophic PV and scales as a vorticity with a reference PV value at rest close to f . Its
 496 evolution equation is (see Muller, 2006):

497
$$\frac{\partial PV}{\partial t} = -\vec{\nabla} \cdot (\vec{U} \cdot \nabla PV + \vec{F} \cdot \vec{\nabla} Z(\rho)) + (\vec{\nabla} \times \vec{U} + \vec{f}) \cdot \dot{Z}(\rho)$$
 (5)

498 This equation shows that PV anomalies can be formed by adiabatic displacement of water parcels in a
 499 background PV gradient, frictional effects (associated with the forcing term \vec{F}) and diapycnal mixing

500 (associated with the forcing terms $\dot{\rho}$ and $\dot{Z}(\rho) = \partial_\rho Z(\rho) \dot{\rho}$). We assume the modification of PV and θ takes
501 place in water masses of larger scale currents, and that vortices emerging from these currents keep similar
502 characteristics. Figure 8 qualitatively describes the expected effects of the identified three main mechanisms on
503 both the formation of PV and θ anomalies. First, since a water mass has specific isopycnal θ (and S) but
504 generally also PV signatures, the adiabatic displacement of a water mass to a region with different characteristics
505 potentially leads to the generation of eddies whose cores contain both PV and θ anomalies (Fig. 8 a-b). Second,
506 if both diapycnal fluxes $\dot{\rho}$ and frictional effects \bar{F} can generate significant PV anomalies and potentially
507 strong eddies (Fig. 8c,e) (see Morel et al, 2019 and references therein), only diapycnal fluxes (induced by mixing
508 or atmospheric fluxes) are also associated with a modification of the θ/S properties and the generation of
509 isopycnal θ anomalies (Fig. 8c-d). It is interesting to notice that interior mixing (Fig. 8c-d) generates new
510 isopycnal characteristics (green patch in Fig. 8d) that necessarily remain in the convex envelope of the initial
511 background θ/S profile (light green area, see Penney et al., 2020), whereas atmospheric fluxes can substantially
512 alter the initial θ/S properties of waters (red characteristics). In contrast, the frictional term acts on momentum
513 and thus on PV (Fig. 8e), but does not affect other tracers, so that no isopycnal θ anomalies are expected (Fig.
514 8f). To conclude, among the mechanisms that can lead to the formation of PV anomalies and eddies, friction is
515 the only one which does not create isopycnal θ anomalies.

516 If a rigorous evaluation of the PV production rate requires 3D fields, as explained in Assene et al (2020),
517 we can simplify Eq. (5) to estimate order of magnitudes for some terms. First, in a background PV gradient β_{bg} ,
518 isopycnal advection of water parcel will generate a PV anomaly

$$519 \quad \delta PV = \beta_{bg} \delta Y \quad (6)$$

520 where δY represents the displacement of the fluid parcels along the PV gradient. Similarly, for this adiabatic
521 process, an isopycnal temperature anomaly is created with

$$522 \quad \delta \theta = \gamma_{bg} \delta Y \quad (7)$$

523 where γ_{bg} is the background temperature gradient (assumed to be in the same direction as the PV gradient for
524 simplicity).

525 The effect of diapycnal mixing on PV is controlled by the impermeability theorem (Haynes and
 526 McIntyre, 1987). This powerful principle allows to evaluate the modification of PV within a layer subject to
 527 diapycnal fluxes. We get (see and McIntyre, 1987; Morel and McWilliams, 2001)

$$528 \quad \delta PV = -f \frac{\delta V}{V} \quad (8)$$

529 where V is the initial volume of the region undergoing mixing and δV is its variation associated with mixing. In
 530 contrast, the evolution temperature, in a layer subject to mixing, depends on values of temperature in adjacent
 531 layers. It can be roughly estimated considering a volume V of water in a layer, having an initial temperature θ_i
 532 mixes with a volume δV of an adjacent layer having a temperature θ_a . The modification of temperature in the
 533 initial layer is then:

$$534 \quad \delta\theta = \frac{\delta V}{V + \delta V} (\theta_a - \theta_i) \quad (9)$$

535 Finally, the effect of friction is more complicated to evaluate, as it depends on details of the wind or
 536 bottom stress structure and stratification close to the boundary but it is known to yield important PV changes too
 537 (see Thomas, 2005; Morel et al, 2006; Benthuisen and Thomas, 2012; Gula et al, 20015, 2016), but the
 538 frictional forcing term does not modify temperature or salinity.

539 In order to analyze the effect of these processes, we computed a mean large-scale rescaled PV,
 540 calculated using Eq. (4), where \bar{U} is the geostrophic velocity field computed from the World Ocean Atlas θ/S
 541 climatology (Locarnini et al., 2018; Zweng et al., 2018) with a level of no motion at 1000 m depth and where the
 542 reference profile to define $Z(\rho)$ is chosen at 27°W, 7.5°N, a dynamically less intense area corresponding also to a
 543 lower surface density. Figure 9a,c show the spatial distribution of the obtained rescaled PV, averaged within
 544 both a near-surface ($\sigma_\theta = 25.75 - 26.5 \text{ kg m}^{-3}$) and a subsurface layer ($\sigma_\theta = 26.9 - 27.4 \text{ kg m}^{-3}$). For each layer, the
 545 mean θ distribution is also shown by black contours (Fig. 9a,c).

546 In R1, a reservoir of relatively strong (weak, respectively) positive PV, associated with relatively cold (warm)
 547 water is observed on the southern (northern) edge of the thermal front in the surface layer (Fig. 9a). Thus in this
 548 layer, an isopycnal PV advection tends to generate either i) positive PV anomalies associated with negative θ' ,

549 leading to the formation of cold core surface CEs, or ii) negative PV anomalies associated with positive θ' ,
 550 leading to the generation of warm-core surface AEs. This PV/ θ relationship is better depicted in Fig. 9b. In
 551 contrast, in R2, maximum positive PV are associated with warmer θ , suggesting that isopycnal advection tends
 552 to generate cold AEs and warm CEs in the surface layer of this region (Fig. 9a-b). In R3, as shown in Fig. 9b, the
 553 PV/ θ relationship is more complex and not strictly monotonic, suggesting that both warm and cold CEs and AEs
 554 can be generated by isopycnal advection, although the general tendency is closer to the R2 region (cold AEs and
 555 warm CEs). From Figure 7, we indeed observed a higher number of cold AEs in the surface layer in R2 and R3,
 556 representing $\sim 66\%$ of the significant surface eddies.

557 In R1, the background gradient of PV is $\beta_{bg} \sim 2 \times 10^{-11} \text{ m}^{-1} \text{ s}^{-1}$, similar to the planetary beta (but note the
 558 gradient has opposite orientation here). The background gradient of temperature has the same direction and we
 559 can estimate $\gamma_{bg} \sim 4 \times 10^{-6} \text{ }^\circ\text{C m}^{-1}$. From Eq. (6) and (7) we can thus estimate that to create an eddy with significant
 560 vorticity and PV anomaly, say $\delta PV \sim \pm 10^{-5} \text{ s}^{-1}$, by isopycnal advection, fluid parcels must be displaced by $\delta Y \sim$
 561 500 km. This yields a temperature anomaly $\delta\theta \sim \pm 2 \text{ }^\circ\text{C}$ (positive for AEs and negative for CEs) which is 3 to 4
 562 times the commonly observed anomalies in the area for these types of vortices. Thus, even for vortices whose
 563 PV and temperature anomalies are coherent with the background fields and could possibly be generated by
 564 isopycnal advection, other processes probably play a role in their generation. Evaluating the possible influence
 565 of diapycnal mixing of the near surface layer with the deeper subsurface one, Eq. (8) shows that a relative
 566 increase of volume $\delta V/V \sim 0.25$ is necessary to reach $\delta PV \sim 1 \cdot 10^{-5} \text{ s}^{-1}$ (using $f \sim 4 \cdot 10^{-5} \text{ s}^{-1}$). Taking a temperature
 567 difference of $\theta_i - \theta_a \sim 5 \text{ }^\circ\text{C}$ between adjacent layers, we get $\delta\theta \sim 1 \text{ }^\circ\text{C}$, about 2 times the observed anomalies in
 568 the area, so atmospheric fluxes, or other processes, also probably have to be taken into account if diapycnal
 569 mixing is to explain the characteristics of observed eddy structures. As mentioned above it is difficult to evaluate
 570 the effect of a frictional term, but this mechanism creates PV anomalies without modifications of the isopycnal
 571 temperature, so it cannot explain alone the formation of eddies with significant temperature anomalies. We can
 572 conclude that in region R1, in the near surface layer, the PV and temperature structure of most eddies are
 573 probably determined by a combination of processes to explain the observed limited temperature anomalies while
 574 still leading to significant PV anomaly, and vorticity.

575 PV/ θ relationships were also examined in the subsurface layer where we note a general northward
576 increase of both the PV and θ (Fig. 9c). This distribution, as well as the main PV/ θ relationship shown in Fig. 9d,
577 suggests that large-scale isopycnal advection leads to the formation of cold subsurface AEs and warm subsurface
578 CEs in the TAO. Thus, isopycnal advection likely explains the dominance (65-80% of the subsurface structures)
579 of warm CEs and cold AEs observed from the Argo profiles in R1 and R2 (Fig. 7). For other eddies, similar
580 analysis as proposed above can be performed and yield the same conclusions: a combination of processes is
581 necessary to explain some of the observed eddy structures.

582 Note that the PV (and temperature) structure of eddies without significant isopycnal θ'/S' signature could
583 be explained by frictional effects alone. Since they are present everywhere and represent about half the eddy
584 population, this would suggest that wind and bottom stress are important mechanisms in the transformation of
585 PV and creation of PV anomalies feeding the core of eddies in the TAO. Likewise, it is tempting to argue that
586 friction could also explain the generation of PV anomalies for vortices exhibiting significant but unstructured
587 θ'/S' signature. Note that even a constant wind (with weak Ekman pumping effects) is able to modify PV along a
588 front (Thomas, 2005), a process that also leads to the destabilization of upwelling currents (Morel et al., 2006),
589 strong modification of Ekman drift (Morel and Thomas, 2009) or the reinforcement of preexisting vortical
590 structures (Holmes et al., 2014). However, this remains hypothetical, and specific combinations of advection and
591 mixing can also be imagined to explain the observed structures. In fact, up to now, we did not find a way to
592 evaluate the influence of friction in the transformation of PV from observations alone.

593 **4.3 Limitations of the study**

594 Diagnostics proposed in this study remain mostly qualitative, but, to our knowledge, it is the first attempt
595 to analyze the importance of non-conservative effects in the generation of eddies and their associated θ/S and PV
596 properties using only observations. The respective influence of the wind-stress and diapycnal mixing in the
597 formation of vortices can serve as reference for realistic numerical models, for which these processes are
598 parameterized. Given the reasonable number of eddies sampled in this study, the statistics calculated here are

599 thought to be significant, but could obviously be refined in the future, when more observations become
600 available.

601 Although our classification of significant versus non-significant θ/S' is robust, it is important to point
602 out some limitations of our diagnostics. First, the fact that eddies must be sampled at their generation (± 200 km)
603 sites strongly reduces the number of analyzed eddies.

604 Second, some of Argo profiles classified as outside eddy profiles might have sampled vortices which
605 were not detectable by altimetry, as it has been highlighted by previous studies (e.g. Garraffo et 2003; Assene et
606 al., 2020). The isopycnal θ climatology, obtained from profiles supposedly acquired outside eddies and used to
607 evaluate eddy anomalies could be slightly spoiled by this effect. However, given the very large number of Argo
608 profiles available, we believe this remains marginal.

609 Third, the exact location, relative to the eddy-centers, of Argo floats that surfaced within eddies were not
610 considered when calculating the eddy θ/S' (Fig. 2d). However, both theoretically and practically, Argo vertical
611 profiles are on average acquired at a distance of $2/3$ of the equivalent eddy radius from the eddy center (e.g.
612 Chaigneau and Pizarro, 2005; Pegliasco et al., 2015; Fig 2d). Although it may slightly impact our results, eddy
613 cores can generally be considered as homogeneous in θ/S and the general discussion on the mechanisms
614 involved in the generation of θ/S' , that shape the whole eddy structure from the eddy center to the eddy edge,
615 remains valid.

616 Fourth, in Fig. 8 we assumed the PV and θ characteristics of eddies are determined by the properties of
617 the water-mass forming their cores, so the generation process itself is adiabatic (such as baroclinic/barotropic
618 instabilities, overshoot of a current at a cape, ...). Friction or mixing could act during the formation process, and
619 we can even imagine that they directly create eddies. This is the case for the formation of eddies by local wind or
620 bottom stress in the lee of islands (e.g. Pullen et al, 2008; Ioannou et al, 2020) or local convection (McWilliams,
621 1985). However, we think these events do not modify our interpretation as even at local scales advection,
622 friction or mixing have the same qualitative effects on PV and θ/S' . One may also wonder how the complex
623 eddy dynamics, characterized by frequently observed merging or splitting events, may influence our results. In

624 our study, which focused on the eddy isopycnal structure close to the eddy generation sites, only splitting events
625 can be problematic because the eddy tracking algorithm considers that a new eddy is formed at the separation
626 site of the parent trajectory (e.g. Pegliasco et al., 2015; Laxenaire et al., 2018). However, among the ~3700 new-
627 born eddies sampled by Argo floats in the TAO, only ~590 (16%) were associated with splitting events.
628 Excluding these eddies from the analyses did not alter the main results and conclusions drawn in this study.
629 However, this point could deserve special consideration in future investigations.

630 Fifth, although satellite altimetry has been proven to be an efficient approach to identify and track
631 subsurface eddies (e.g. Ciani et al., 2017), the exact position of the birth location of subsurface eddies could be
632 slightly biased using altimeter data. Indeed, subsurface eddies do not necessarily have a clear signature on sea
633 surface height which prevents their identification in altimetry maps (e.g. Assene et al., 2020). Therefore, in our
634 study, some subsurface eddies could have been formed a few days/weeks before being detected by altimetry.
635 However, quantifying such biases is not possible without using three-dimensional numerical simulations (e.g.
636 Ciani et al., 2017), what is beyond the scope of the present study.

637 Sixth, we argued that the effect of a frictional forcing term \bar{F} (Eq. 1) on PV modification (Eq. 5) is
638 difficult to assess. An additional difficulty is that frictional processes are generally also accompanied with
639 diapycnal mixing. This is indeed generally the case for the wind stress, which can thus be expected to lead to
640 modification of θ/S too. However, strong mixing is usually associated with convection generated by atmospheric
641 heat fluxes, or strong vertical shear in ocean currents. Without such strong preconditioning, diapycnal mixing
642 associated with the wind or bottom stress is expected to be weak, with θ'/S' remaining non-significant.

643 Finally, Classification of eddies with significant vs non-significant temperature anomalies relies
644 on thresholds defined according to the local background variability and computed from Argo profiles that
645 surfaced outside eddies (Appendix A2). This variability represents the typical anomaly a particle can get when
646 displaced adiabatically over a typical vortex diameter. Thus a vortex whose core has an isopycnal θ' below the
647 threshold must have gained its PV anomaly from processes that do not significantly alter the local water masses,
648 which is typical of friction. We used the 80th percentile of this variability to define θ' threshold,
649 since the isopycnal θ' profiles are distinguishably different from those of eddies with significant θ' (see Fig. 5d-f

650 and 7). The number of identified eddies obviously depend on the choice of the percentile but the proportion of
651 eddies with non-significant θ' remains important as long as the chosen percentile remains reasonable (the
652 proportion is still 40% when choosing the 70th percentile, it reaches 70% when choosing the 90th percentile).
653 Likewise, results are also sensitive the confidence level chosen for the performed Student's t-test. Indeed, the
654 more we decrease the confidence level of the Student's t-test, the more we increase the number of significant
655 anomalies.

656 5. Summary and perspectives

657 Combining 18 years of satellite altimetry and θ/S data acquired by Argo floats in the TAO, we first
658 showed that isopycnal θ'/S' can strongly differ from the ones obtained using depth-coordinates. Indeed, although
659 AE (CE, respectively) mostly induced positive (negative) θ'/S' in depth-coordinates, both AE and CE can exhibit
660 positive, negative or non-significant isopycnal θ'/S' . We then focused on the vertical structure of eddies close to
661 their generation site, to investigate the proportion and distribution of new-born eddies and found that half of the
662 total analyzed new-born eddies in the TAO are characterized by non-significant θ' . In contrast, the second half,
663 composed of 70% of subsurface, and of 30% of near-surface eddies, likely involve diapycnal mixing and/or
664 lateral advection. As a consequence, in terms of tracer transport, our results would suggest that eddies that are
665 generated with a significant θ'/S' mostly contribute to heat and salt transport in the TAO.

666 Refined diagnostics in three selected subregions in the northern TAO, were proposed. Along the frontal
667 zone and in the northern equatorial subregions, the mean vertical structure of eddies is dominated by subsurface
668 eddies with maximum anomalies reaching up to ± 0.5 °C mostly found between $\sigma_\theta = 27$ and 27.2 kg m^{-3} isopycnal
669 layers. For these subsurface eddies, $\sim 75\%$ of AEs exhibit a negative maximum θ' whereas $\sim 65\%$ of CEs
670 maximum θ' are positive. In the third subregion, within the NBC retroflexion, results are questionable because
671 of the reduced number of eddies sampled by Argo profiles at their generation sites.

672 We then proposed qualitative diagnostics based on the links between θ'/S' and PV anomalies under the
673 influence of isopycnal advection, friction or diapycnal mixing. For eddies with significant anomalies, in the
674 surface layer, PV/ θ relationships suggest that isopycnal water-mass advection can explain the generation of PV

675 anomalies of cold core AEs observed in R2 and R3 areas. In contrast, in R1, isopycnal advection would
676 preferentially lead to the generation of cold core CEs and warm AEs. But observations show no preference in θ' .
677 Thus isopycnal advection alone cannot explain our diagnostics and, to modify PV or θ'/S' structures of water-
678 masses, other processes must be involved too. Both the wind-stress or diapycnal mixing can explain the
679 observations and, even though we think friction is probably a key process to explain the observed characteristics,
680 it seems difficult to prove that one process is more important than the other. Additional diagnostics can be
681 imagined to further analyze the relative influence of each process. For instance, friction associated with the wind
682 stress is probably sensitive to the position of the ITCZ in the studied region, leading to weaker stress in summer
683 in R1, when the ITCZ migrates from the equator to this region. Preliminary analysis has shown that the
684 production rate of eddies with non-significant θ'/S' drastically drops during summer, which could be a sign that
685 friction is indeed a major mechanism for the generation of PV and vortices in this region. However, this drop
686 also exists for some other vortex types, and the atmospheric heat fluxes also vary seasonally, so the
687 interpretation is not easy and deeper analyses are necessary before concluding. In the subsurface-layer, water-
688 mass advection is also suggested to explain the formation of PV anomalies of warm (cold, respectively) core
689 CEs (AEs) especially in R1 and R2, which indeed corresponds to the distribution dominantly observed.

690 As far as perspectives are concerned, this study provide a new diagnostics that can be useful for
691 numerical models. As mentioned above, frictional and diapycnal mixing effects are parameterized in the models,
692 and thus imperfectly represented. Given the qualitative link between these processes and the transformation of
693 both PV and isopycnal θ'/S' fields, combined diagnostics involving isopycnal θ'/S' and vorticity of eddies is thus
694 challenging for numerical results. In particular, since a major fraction of the observed eddies are associated with
695 non-significant θ/S anomalies, frictional effects are thought to play a major role in the generation of their PV
696 structure. Parameterizations of frictional effects are very difficult to evaluate and remain one of the Achille's
697 heel of circulation models at mesoscale. The proposed diagnostics can thus be very useful to compare different
698 parameterizations. Further, it would be interesting to combine isopycnal θ/S structures with trajectories to infer
699 the long term behavior of specific long-lived eddies, having been sampled at different times by Argo floats, to
700 analyze the evolution of their heat/salt contents. The present general approach can also be applied to other

701 regions to infer the relative contribution of friction and diapycnal mixing or isopycnal advection in the PV
 702 sources forming the core of eddies, using only observations, particularly in energetic areas such as the Gulf
 703 Stream and Kuroshio or in specific regions where wind-stress curl lead to the eddy generation (Canary,
 704 Ierapetra, etc.). Finally, when a significant number of Argo floats will be equipped with biogeochemical sensors
 705 (dissolved oxygen, nutrients, chlorophyll-a, ...), it will be very interesting to evaluate if there exists some
 706 structuring in the transport of biogeochemical tracers by AEs and CEs, in particular in the TAO. Indeed,
 707 combining different tracers with different sources and sinks can help to more precisely identify the physical
 708 mechanisms responsible for the transformation of the water characteristics, including their PV.

709
 710

711 **Appendix A: Determination of the significance of eddy temperature anomalies**

712 **A1. Weighted arithmetic means**

713 Isopycnal θ/S' were computed for each Argo profile acquired within a new-born eddy by removing a
 714 local climatological profile representative of the large-scale background, also computed on density-coordinates.
 715 These local climatological profiles (\bar{P}) were obtained by weighted arithmetic means of all the available profiles
 716 (P_i) acquired outside eddies, within a radius of 200 km and separated by less than ± 30 days (independently of the
 717 year) from the date of the considered profile. Weighted arithmetic mean profiles (\bar{P}) were computed by
 718 equation E1:

$$719 \quad \bar{P} = \frac{1}{\sum \Omega_i(r_i, t_i)} \sum \Omega_i(r_i, t_i) \cdot P_i(\rho) \quad (E1)$$

720 where Ω_i are the weights, assigned to each profile, which depend on the distance (r_i) and time (t_i) of
 721 the profile P_i (outside eddies) from the considered profile (located inside the eddy):

$$722 \quad \Omega_i(r_i, t_i) = e^{-\frac{1}{2} \left[\left(\frac{r_i}{\Delta R} \right)^2 + \left(\frac{t_i}{\Delta t} \right)^2 \right]} \quad (E2)$$

723 where $\Delta R = 100$ km and $\Delta t = 15$ days are the typical length and time scales. The choice of ΔR and Δt does not
 724 significantly alters P and the results presented in the study.

725 **A2. Temperature anomaly threshold**

726 In order to determine whether θ'/S' obtained within new-born eddies are significant or not, a climatology
 727 of monthly isopycnal θ' thresholds was computed on a $1^\circ \times 1^\circ$ longitude/latitude grid within surface and
 728 subsurface layers. The surface layer extends from the base of the mixed layer to the base of the seasonal
 729 pycnocline and the subsurface layer extends from the base of the pycnocline to the deepest isopycnal level (Fig.
 730 A1). As in de Boyer Montégut et al. (2004), the mixed layer depth was defined as the depth where the density
 731 increased by 0.03 kg m^{-3} from its value at 10 m depth. The reference depth was chosen at 10 m in order to avoid
 732 the strong diurnal cycle that occurs in the first few meters of the ocean (de Boyer Montégut et al. 2004). The
 733 base of the pycnocline was determined as the depth of fluid layer possessing one half of the maximum of the
 734 squared buoyancy frequency (see Cheng and Hsu, 2014). The square buoyancy is defined as :

$$735 \quad N^2 = \frac{-g}{\rho} \frac{d\rho}{dz} \quad (\text{E3})$$

736 where g , ρ and z , are gravity, density, and depth, respectively.

737 Typically, the surface layer extends from $\sigma_\theta \sim 22 \text{ kg m}^{-3}$ to $\sigma_\theta \sim 26 \text{ kg m}^{-3}$ density layers in equatorial regions
 738 (15°N - 15°S), and from $\sigma_\theta \sim 25 \text{ kg m}^{-3}$ to $\sigma_\theta \sim 26.5 \text{ kg m}^{-3}$ at mid-latitudes (15 - 30°). In both areas, the subsurface
 739 layer extend from the base of the surface layer to $\sigma_\theta \sim 27.5 \text{ kg m}^{-3}$. In both the surface and subsurface layers
 740 (Fig. A₁), the monthly climatological θ' thresholds were computed at each grid point as follows: first, we
 741 selected all the profiles within 200 km around the grid point, that surfaced outside eddies during the
 742 corresponding month (regardless of the year). Second, we computed for each profile, the square root of the
 743 quadratic mean of θ' integrated over the layer thickness using equation E4.

$$744 \quad M_1 = \sqrt{\frac{1}{\sum h(\rho_i)} \cdot \sum h(\rho_i) \cdot (\theta'(\rho_i))^2} \quad (\text{E4})$$

745 where $h(\rho_i)$ and $\theta'(\rho_i)$ are the thickness and the θ' of the isopycnal layer ρ_i . Third, we retained the 80th percentile
 746 of M1 values as the monthly θ' threshold at the considered grid-point. This percentile was carefully chosen after
 747 a sensitivity study and is considered as significantly different from noise. For accuracy reasons, at least 30
 748 monthly profiles are required around the grid-point to compute the threshold. Figures A2 and A3 show the
 749 monthly θ' threshold distributions for the surface and subsurface layers. Note the relatively high θ' threshold
 750 within the frontal zone throughout the year in both layers. No strong seasonality was observed, except a slight
 751 increase of the θ' threshold value in February in the surface layer (Fig. A2).

752 Isopycnal θ' observed in an eddy is considered as significant if its M1 values in surface and/or
 753 subsurface are higher than the corresponding thresholds shown in Fig. A2-A3. In contrast if M1 values in the
 754 eddy are lower than the corresponding thresholds in both the surface and subsurface layers, eddy anomalies are
 755 classified as non-significant. When an anomaly is significant in a given layer, its sign is given by that of the
 756 mean θ' computed over the layer, as follow:

$$757 \quad M_2 = \frac{1}{\sum h(\rho_i)} \cdot \sum h(\rho_i) \cdot \theta'(\rho_i) \quad (E5)$$

758 A similar method was used by Itoh and Yasuda (2010) to identify warm and cold eddies in the northwestern
 759 Pacific ocean.

760 **A3. Mean eddy dynamic properties**

761 The mean evolution of the study eddy properties (amplitude, radius and kinetic energy), similarly to
 762 what has been already observed by several observational (e.g. Pegliasco et al., 2015; Sun et al., 2018; Aguedjou
 763 et al., 2019) and numerical (Kurian et al., 2011) studies, follows three main phases during eddy life-cycle: the
 764 growth phase, the maturity phase and the dacaying phase (Fig A4a-c). The first 20-30% of eddy lifespan
 765 represents eddy growth phase during which eddy properties stongly increase, whereas the maturity phase
 766 corresponds to a plateau-like evolution of the mean properties representing 20-30% to ~ 80 % of eddy lifespan.

767 The decaying phase is the last phase of eddy life-cycle during which eddy properties decrease rapidly. During
768 the growth phase, TAO eddies travel on average over 200 km (Fig A4d).

769

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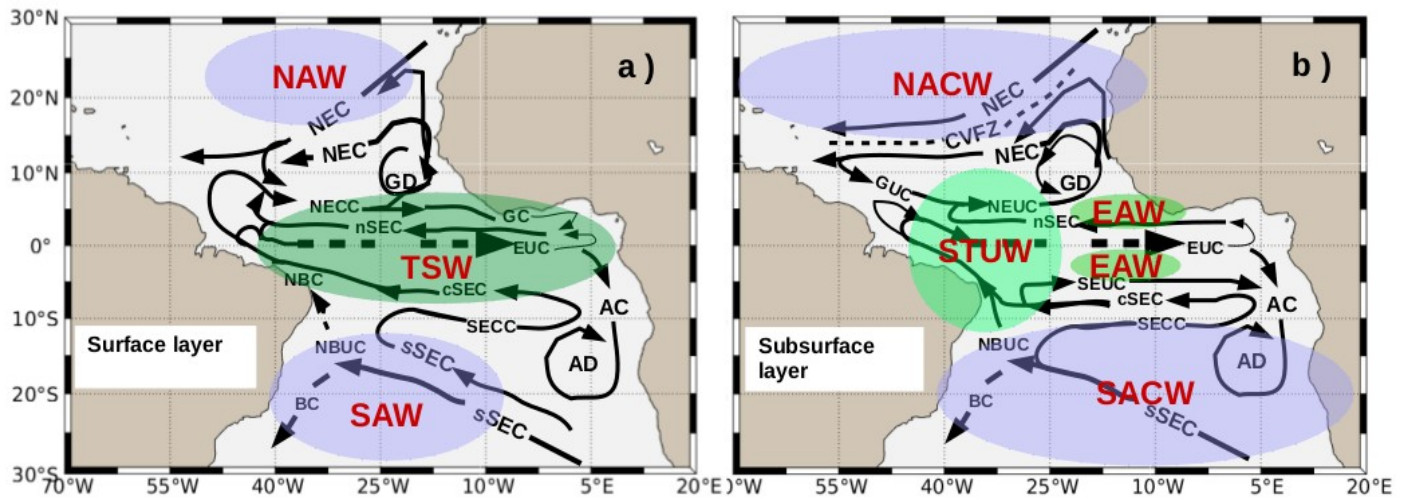
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1030 **Figures**

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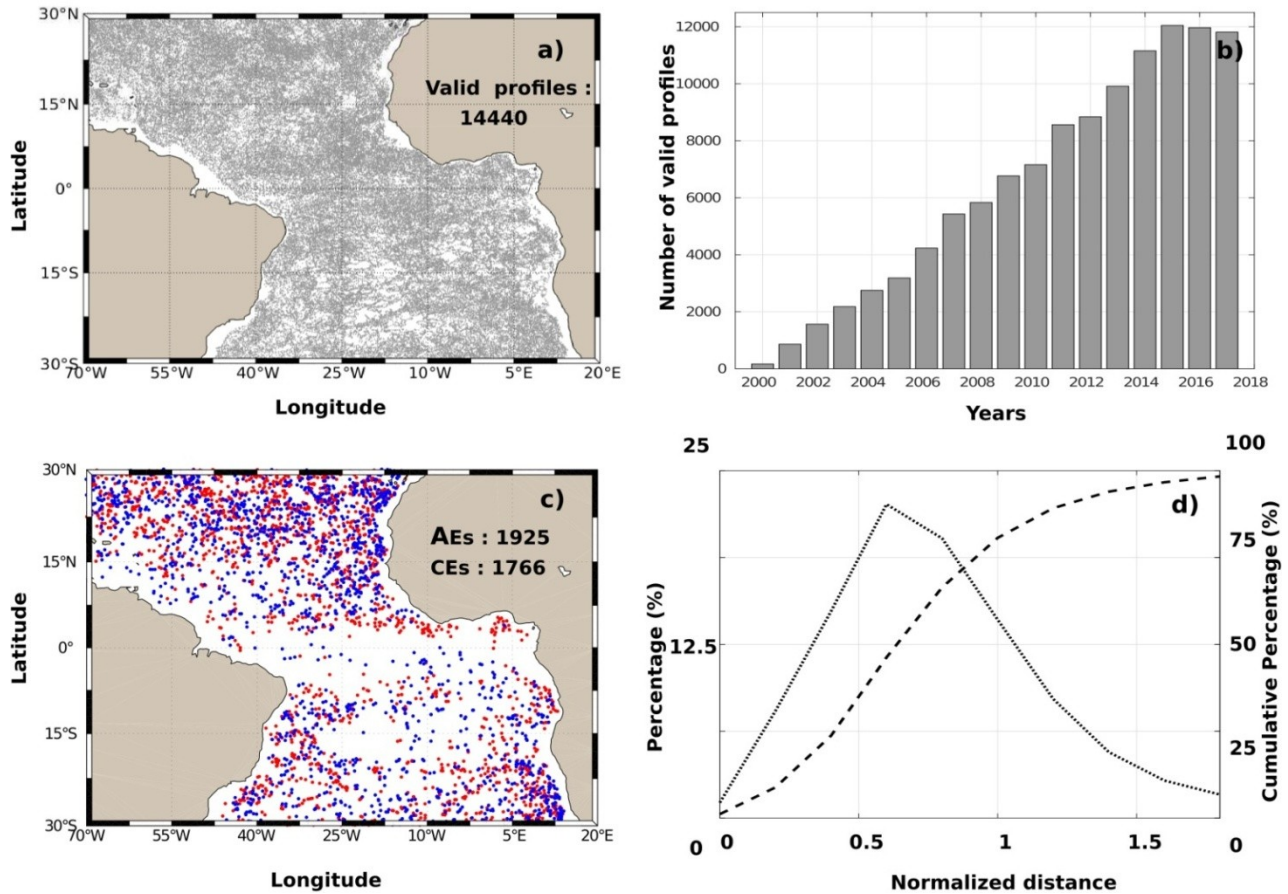


1032 Figure 1. Schematic view of a) surface and b) subsurface circulation in the TAO (adapted from Stramma and
 1033 Schott, 1999). Main currents are: Brazil Current (BC), North Brazil Current (NBC), North Brazil Undercurrent
 1034 (NBUC), North Equatorial Current (NEC), North Equatorial Countercurrent (NECC), northern, central and
 1035 southern branches of South Equatorial Current (nSEC, cSEC, sSEC), Guinea Current (GC), Angola Current
 1036 (AC), Angola Dome (AD), Guinea Dome (GD), South Equatorial Countercurrent (SECC), Equatorial
 1037 Undercurrent (EUC), North Equatorial Undercurrent (NEUC), South Equatorial Undercurrent (SEUC), Guiana
 1038 Undercurrent (GUC). Tropical surface water-masses and central water-masses are also indicated: Tropical
 1039 Surface water (TSW), Northern, southern and eastern tropical Atlantic water (NAW, SAW, EAW), Subtropical
 1040 Underwater (STUW), North and South Atlantic Central Water (NACW, SACW).

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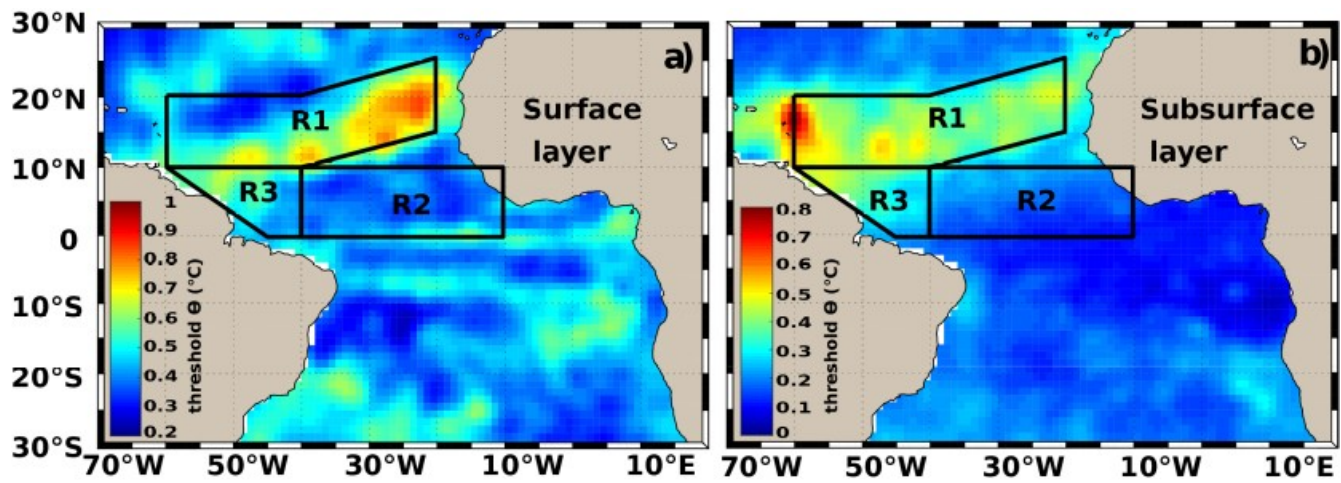


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1046 Figure 2. Spatio-temporal distribution of Argo floats in the TAO: a) Spatial and b) temporal distribution of valid
 1047 Argo profiles, c) spatial distribution of Argo profiles that surfaced within anticyclonic eddies (AEs, red dots) and
 1048 cyclonic eddies (CEs, blue dots), d) Histogram (dotted line, left axis) and cumulative distribution (dashed line,
 1049 right axis) of the normalized radial distance of Argo floats to eddy centers. Normalized distances were obtained
 1050 dividing the distance (in km) by the equivalent eddy radius.

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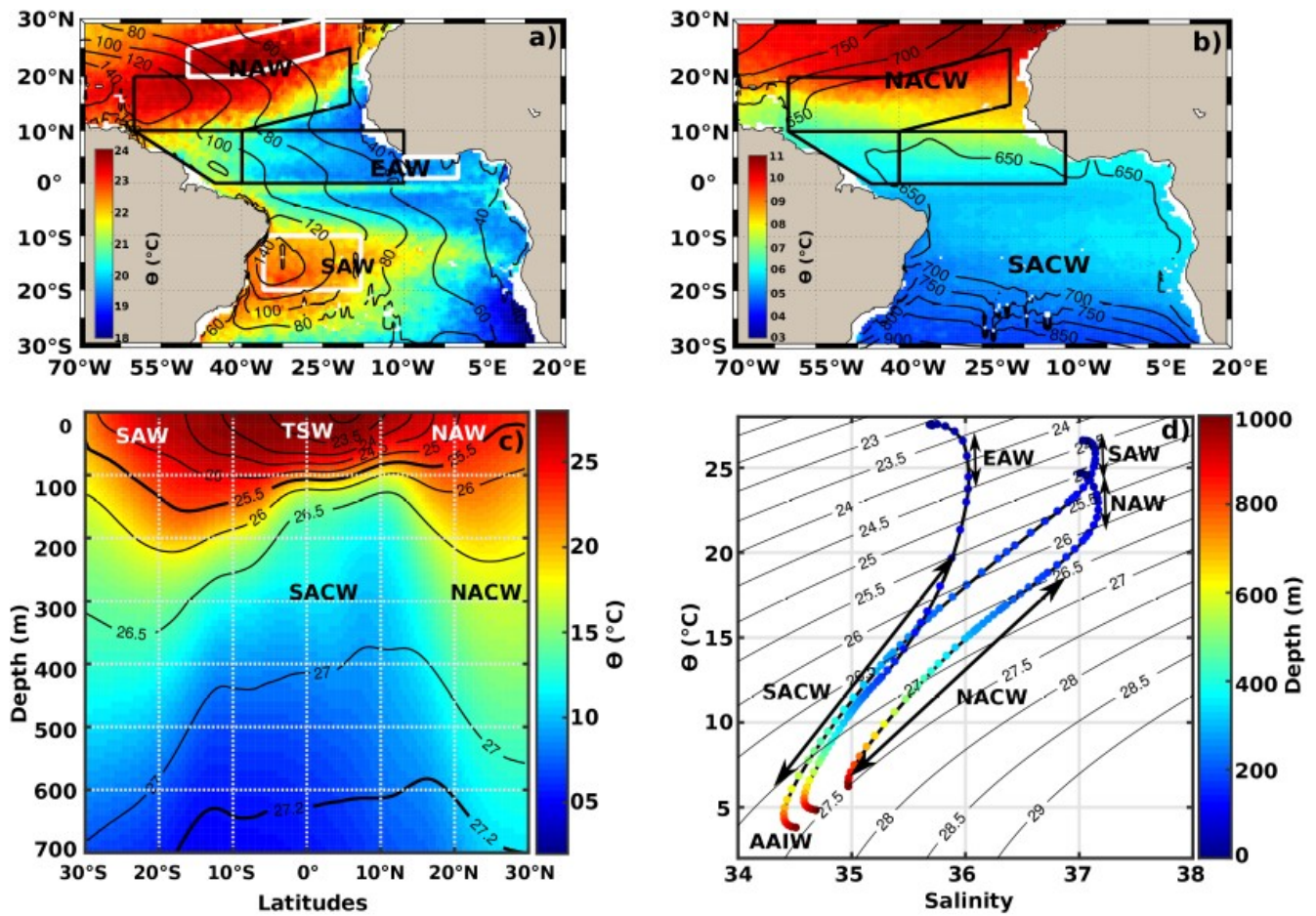


1054 Figure 3. Annual mean of the isopycnal temperature anomaly threshold within the a) surface and b) subsurface
 1055 layers. Three dynamically different subregions (R1, R2, R3) used in this study are delimited in black (see in-text
 1056 description of these regions). Surface layer extends from the mixed layer depth to the base of the pycnocline,
 1057 whereas subsurface layer extends below the pycnocline to 1000 m depth.

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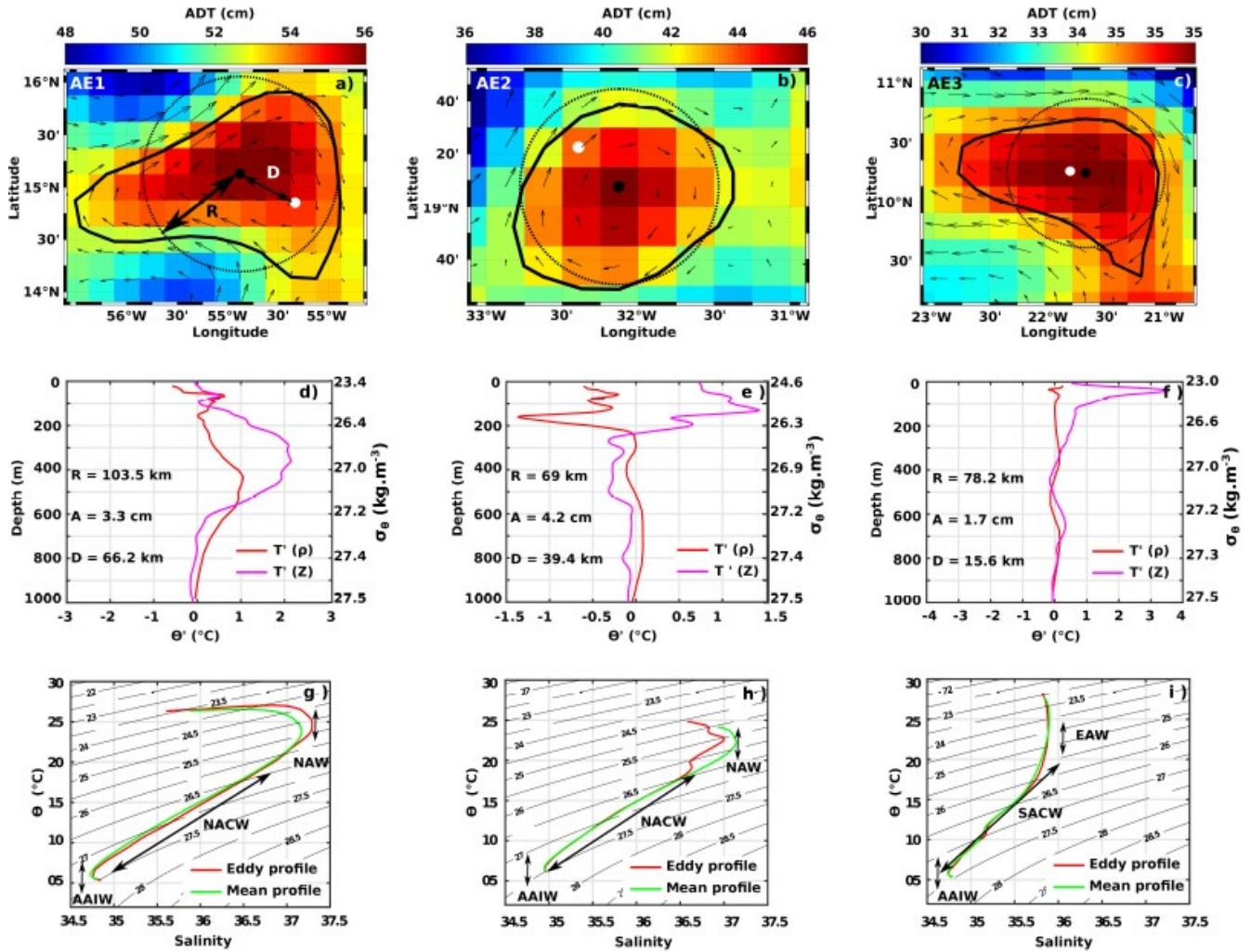
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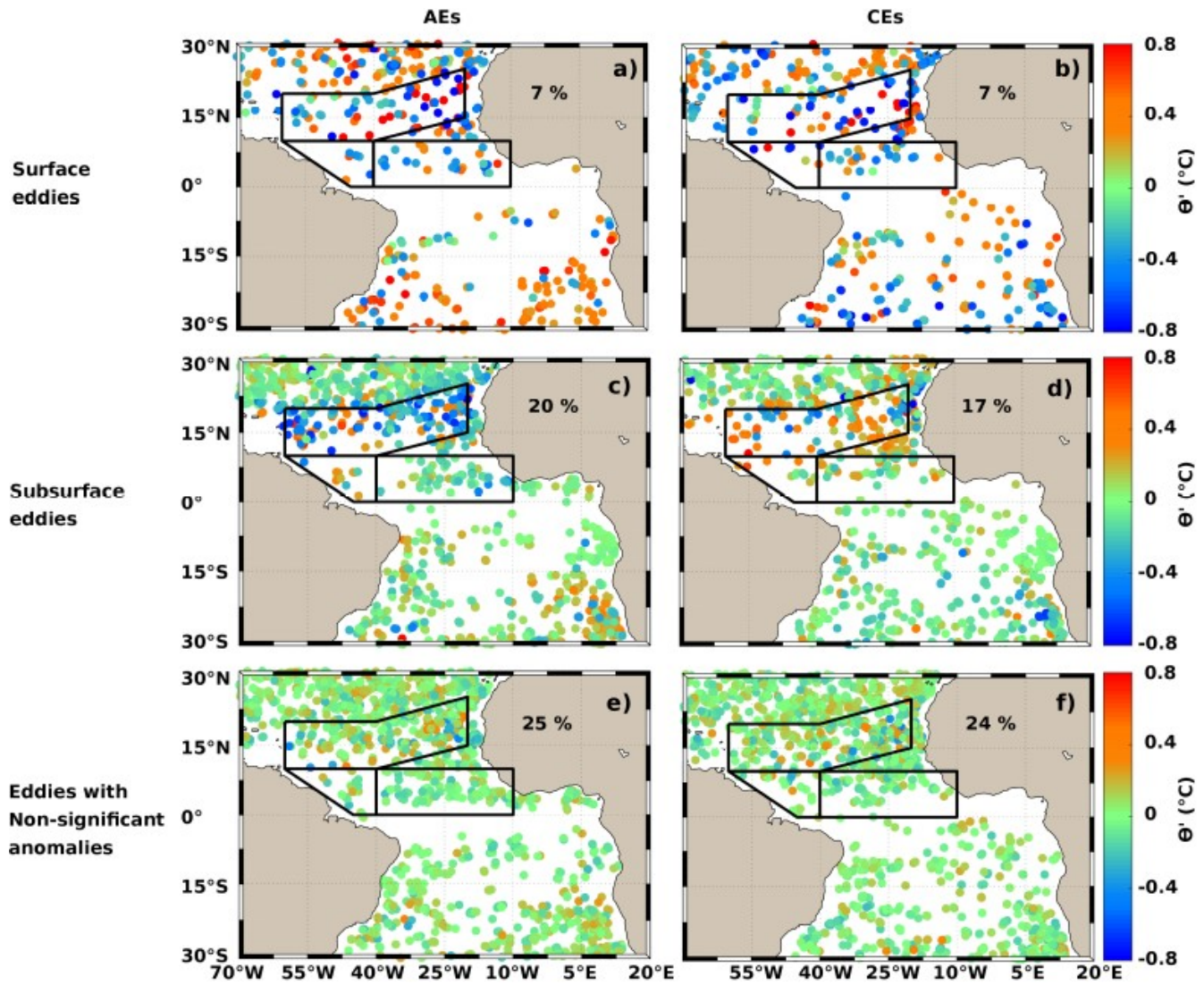
1062 Figure 4. Large-scale temperature distribution (in °C) on a) $\sigma_\theta = 25.5 \text{ kg m}^{-3}$ and b) $\sigma_\theta = 27.2 \text{ kg m}^{-3}$. Black
 1063 contours correspond to the depth (in m) of the corresponding σ_θ layer, whereas black boxes delimit the R1-R3
 1064 subregions. c) Mean meridional temperature section along 35°W ; black contours depict σ_θ levels. d) Mean θ/S
 1065 diagram and main water-masses observed within the 3 white boxes delimited in a). Water-masses: NAW: North
 1066 Atlantic Water; SAW: South Atlantic Water; TSW: Tropical Surface Water; EAW: Eastern Atlantic Water;
 1067 NACW: North Atlantic Central Water; SACW: South Atlantic Central Water; AAIW: Antarctic Intermediate
 1068 Water.

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1073 Figure 5. Density-coordinate anomalies compared to depth-coordinate anomalies for three case-study
 1074 anticyclonic eddies. a-c) Eddy characteristics in AVISO maps. Black and white dots correspond to the eddy
 1075 centers and the location of surfaced Argo floats, respectively, black contours delimit eddy edges whereas dotted
 1076 circles represent the contours of the disc corresponding to eddy areas. d-f) Temperature anomalies observed
 1077 within eddies in depth-coordinates (magenta lines) and in density-coordinates (red lines). R, A and D indicate
 1078 the eddy radius, amplitude and the distance of the surfaced Argo float to the eddy center. g-i) θ -S diagram
 1079 obtained within eddies (red lines) and for the mean climatology at the same location (green curves). Water-
 1080 masses: NAW: North Atlantic Water; EAW: Eastern Atlantic Water; NACW: North Atlantic Central Water;
 1081 SACW: South Atlantic Central Water; AAIW: Antarctic Intermediate Water.

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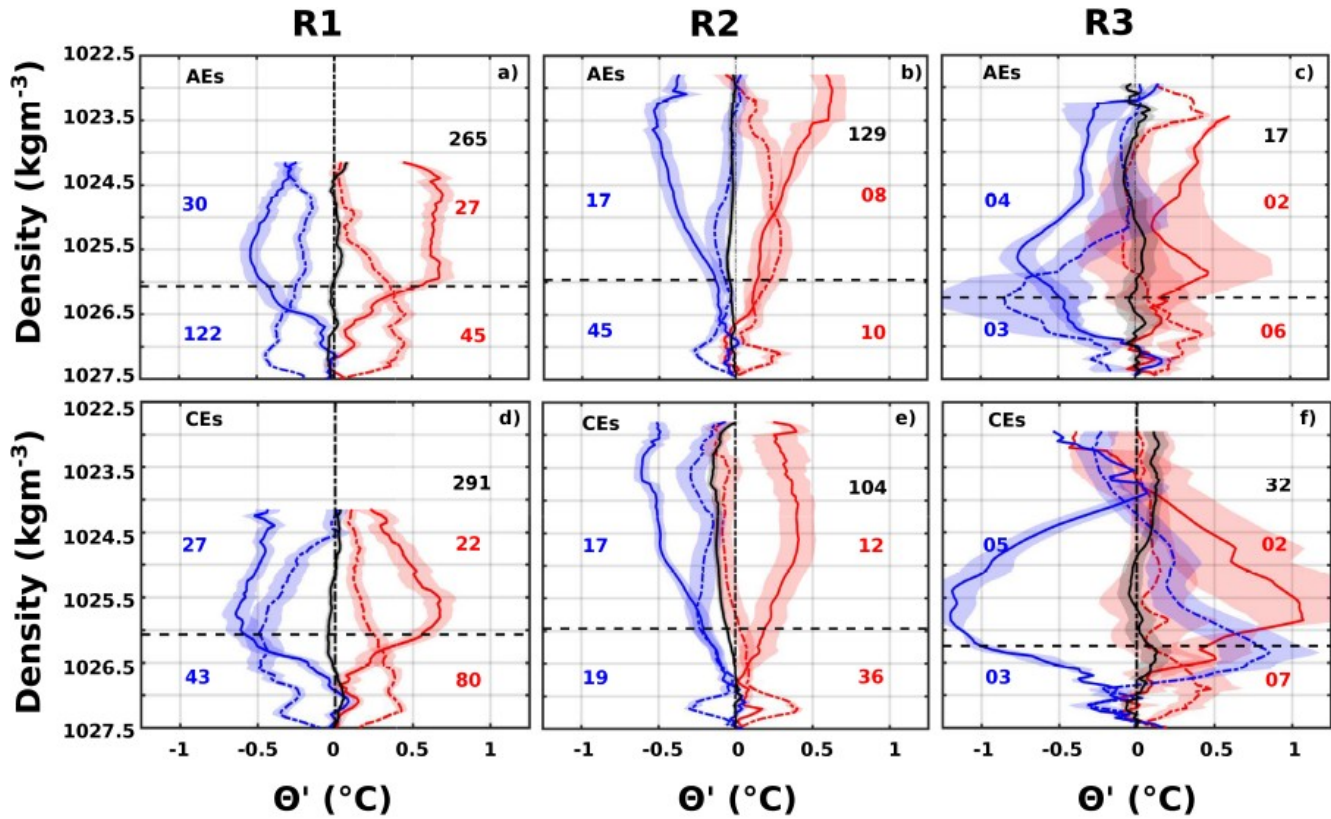


1083 Figure 6. Spatial distribution of isopycnal temperature anomalies within new-born eddies. a-b) Significant
1084 surface θ' within AEs and CEs. c-d) Significant subsurface θ' within AEs and CEs. e-f) non-significant θ' within
1085 AEs and CEs. Black boxes delimit the R1-R3 subregions. Numbers indicate the percentage of sampled eddies
1086 for each category.

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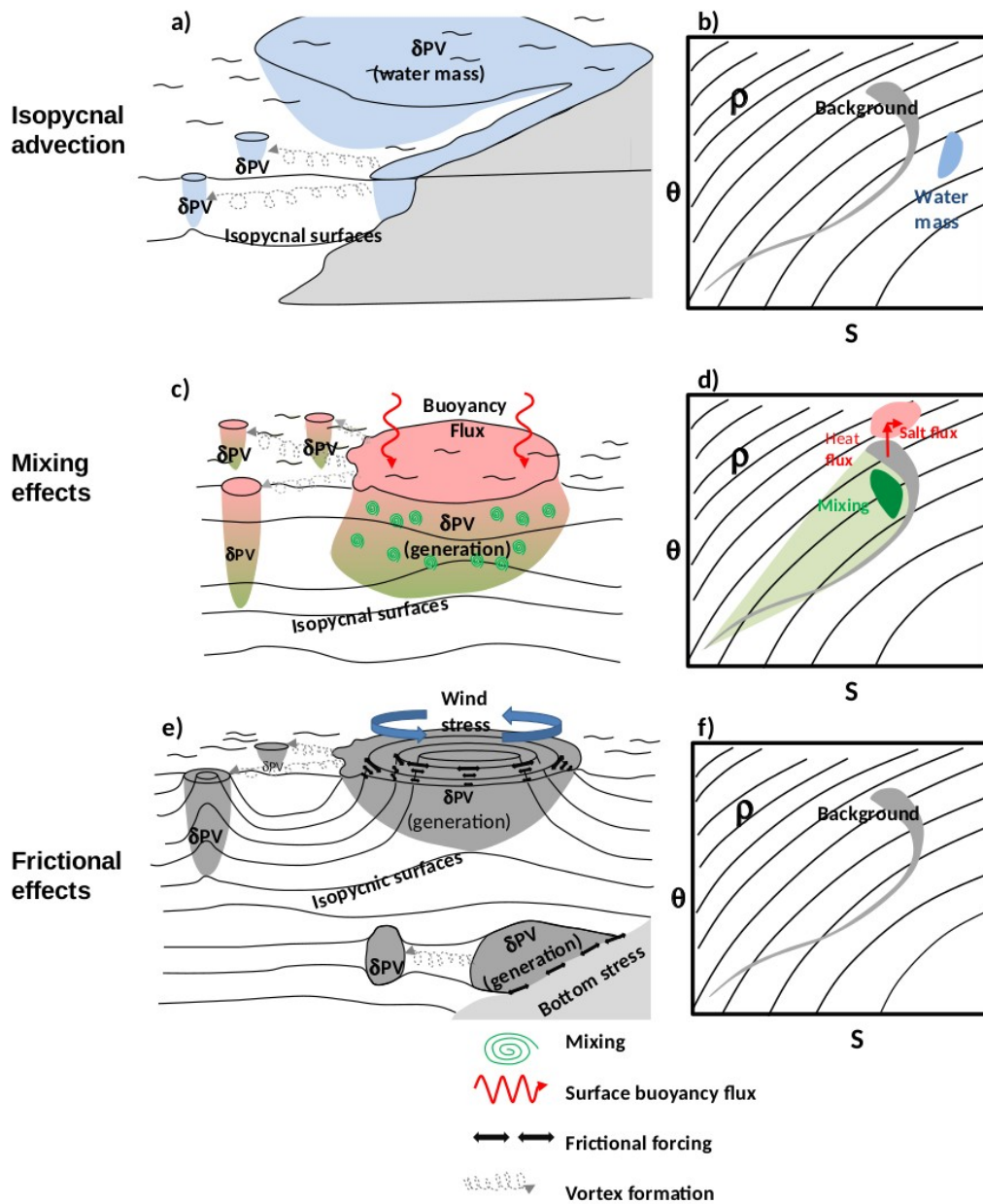


1090 Figure 7. Mean isopycnal temperature structures of newborn a-c) anticyclonic (AEs) and d-f) cyclonic (CEs)
1091 eddies within 3 sub-regions (R1-R3.) Significant surface and subsurface anomalies are represented in solid and
1092 dashed lines, respectively, whereas positive anomalies are in red and negative anomalies in blue. Non-significant
1093 anomalies are represented in black. For each subplot, black numbers indicate the number of sampled eddies
1094 having non-significant θ' , the right red top (bottom, respectively) numbers indicate the number of surface
1095 (subsurface) eddies with positive anomalies and the left blue top (bottom, respectively) numbers indicate the
1096 number of surface (subsurface) eddies with negative anomalies. Shading areas correspond to standard errors of
1097 the means, evaluated using Student's t test with a significance level of 95%. For each subregion, the horizontal
1098 dashed line corresponds to the base of the pycnocline that delimits the surface and subsurface layers.

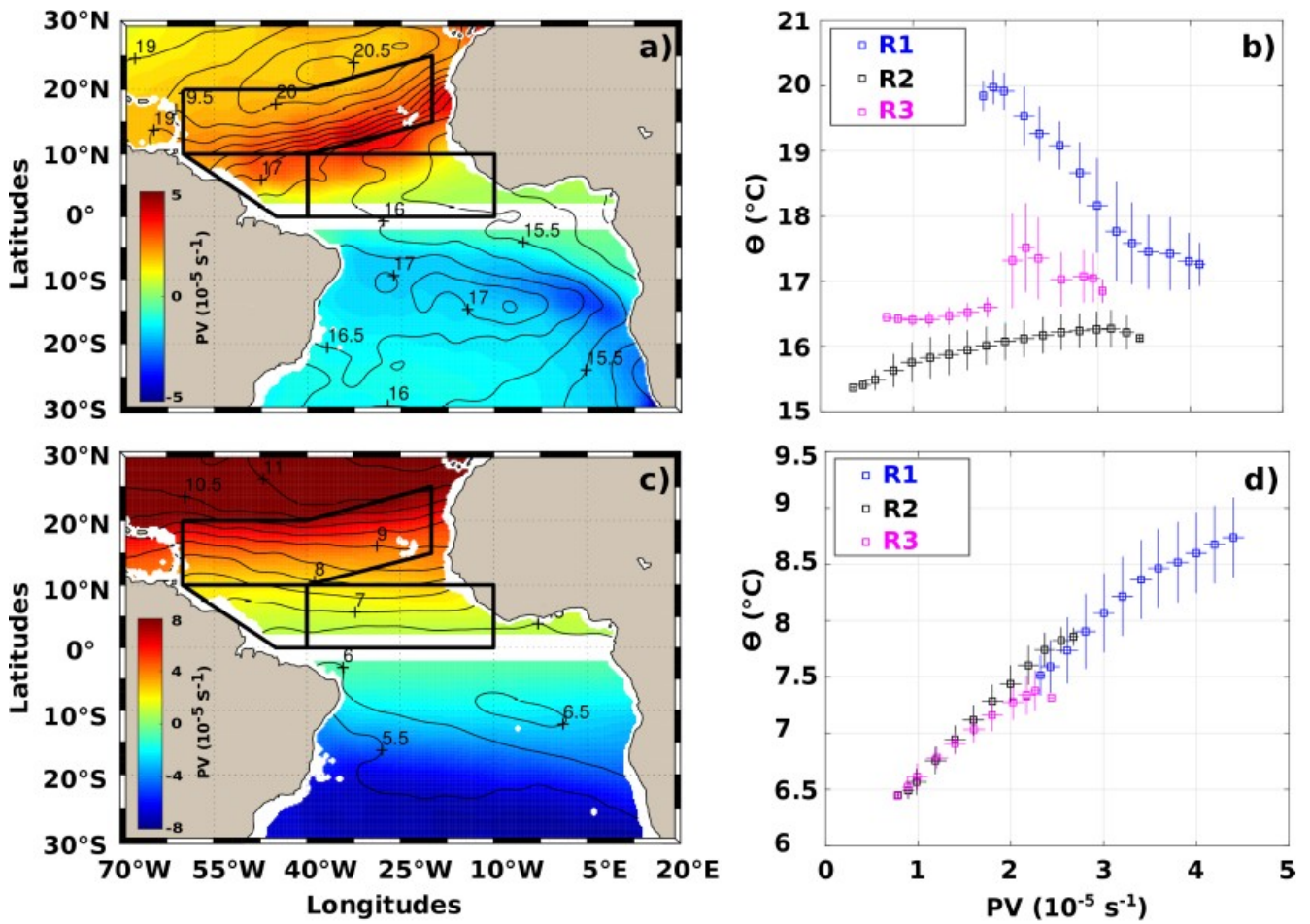
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1121 Figure 8. Qualitative analysis of the combined effects of specific mechanisms on the creation of PV (left
 1122 column) and corresponding θ/S diagrams (right column). a-b) Advection of specific water-mass (in blue) in a
 1123 region with different PV and θ/S characteristics (in white/black). c-d) Creation of PV anomalies by diapycnal
 1124 mixing (green) and atmospheric buoyancy fluxes (red); These mechanisms also create θ/S anomalies. e-f)
 1125 Frictional effects associated with surface wind or bottom stress; This mechanism is the only one that does not
 1126 create θ/S anomalies. θ/S diagrams representative of the background are in black.



1128 Figure 9. Large-scale rescaled potential vorticity (PV, color shading) and θ (in $^\circ\text{C}$, black contours), averaged
 1129 between a) $\sigma_\theta = 25.75 - 26.5 \text{ kg m}^{-3}$ and c) $\sigma_\theta = 26.9 - 27.4 \text{ kg m}^{-3}$. Black boxes delimit the R1-R3 subregions. b
 1130 & d) PV/ θ relationships in R1-R3 subregions. PV (and corresponding θ) were averaged in intervals of $0.2 \times 10^{-5} \text{ s}^{-1}$
 1131 (dots) and the corresponding standard deviations are indicated by solid lines.

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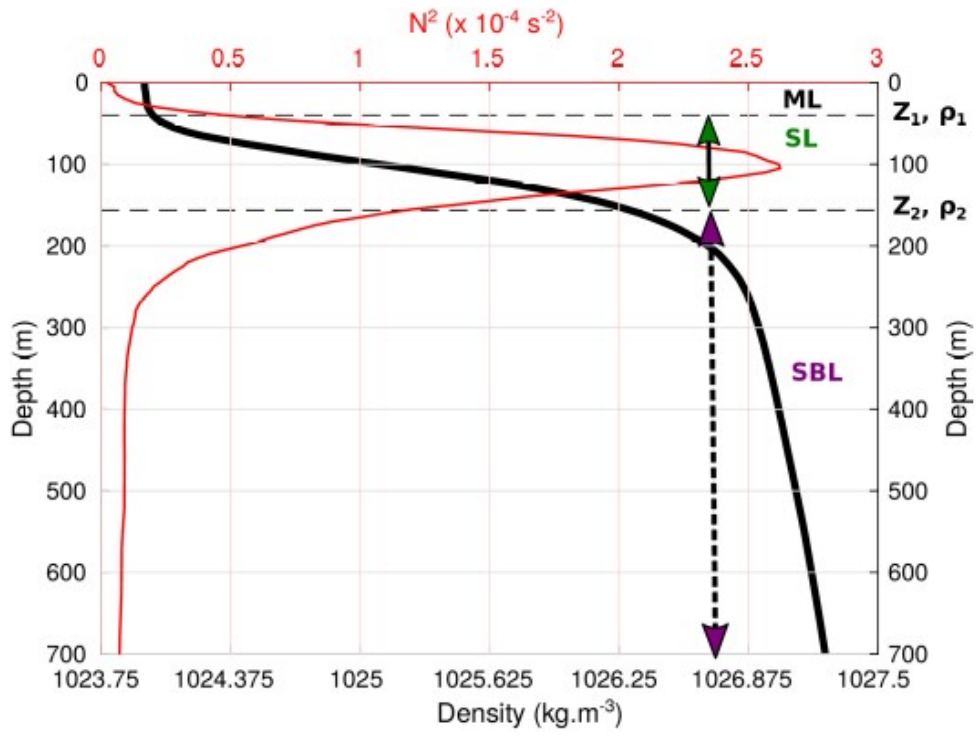
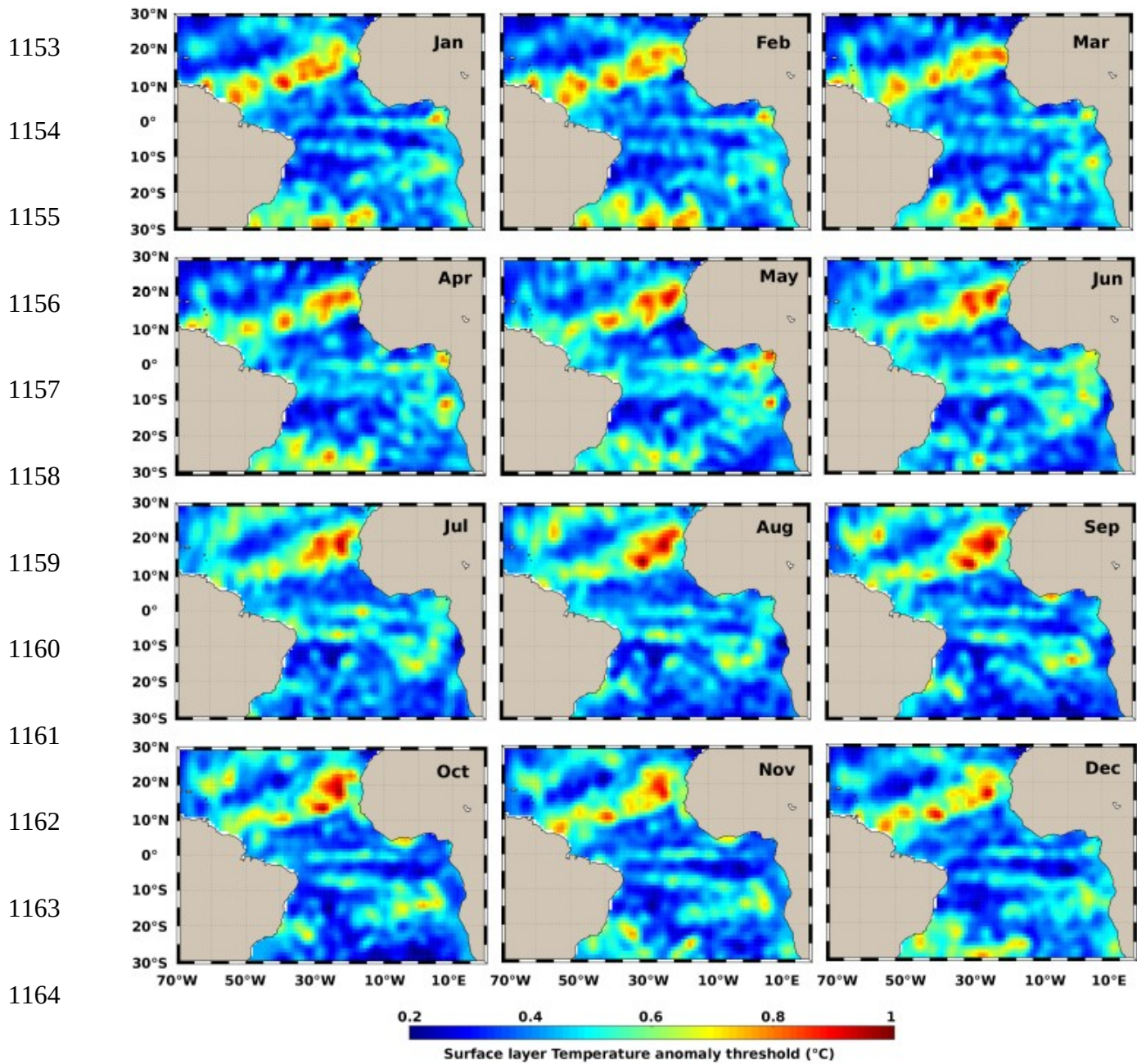
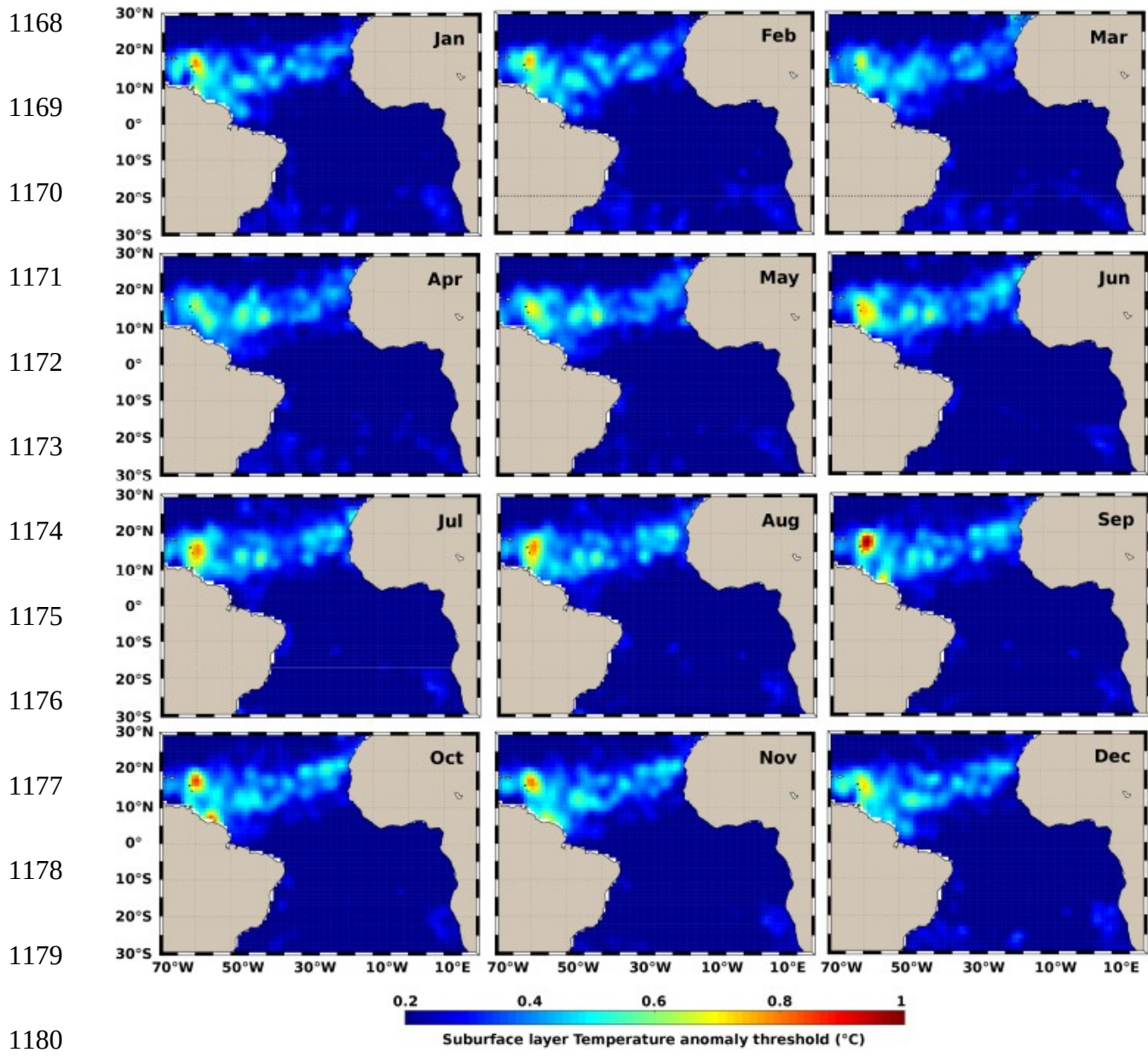


Figure A1. Mean density (black curve) and square buoyancy frequency (N^2 , red curve) profiles at a given grid location showing the mixed (ML), surface (SL) and subsurface (SBL) layers. (Z_1, ρ_1) and (Z_2, ρ_2) are the mixed layer and the base of the pycnocline depths and densities.



1166 Figure A2. Monthly mean of the temperature anomaly threshold within the surface layer.

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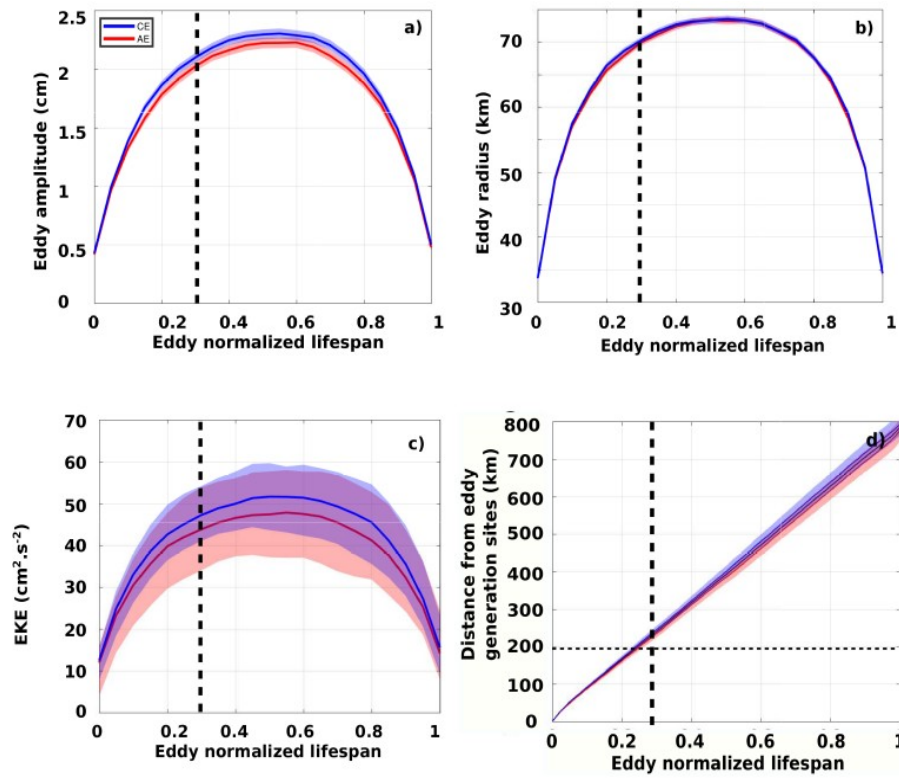
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Figure A3. Monthly mean of the temperature anomaly threshold within the subsurface layer.

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1194 Figure A4. Mean eddy a) amplitude, b) radius, c) kinetic energy (EKE) and d) propagation distance as a function
1195 of normalized lifespan. Vertical dashed lines delimit the eddy growth phase from the mature phase, and
1196 correspond to 30% of eddy lifespan and the horizontal dashed line indicates the 200 km distance. Shading areas
1197 correspond to standard errors around the mean.

Figure1.

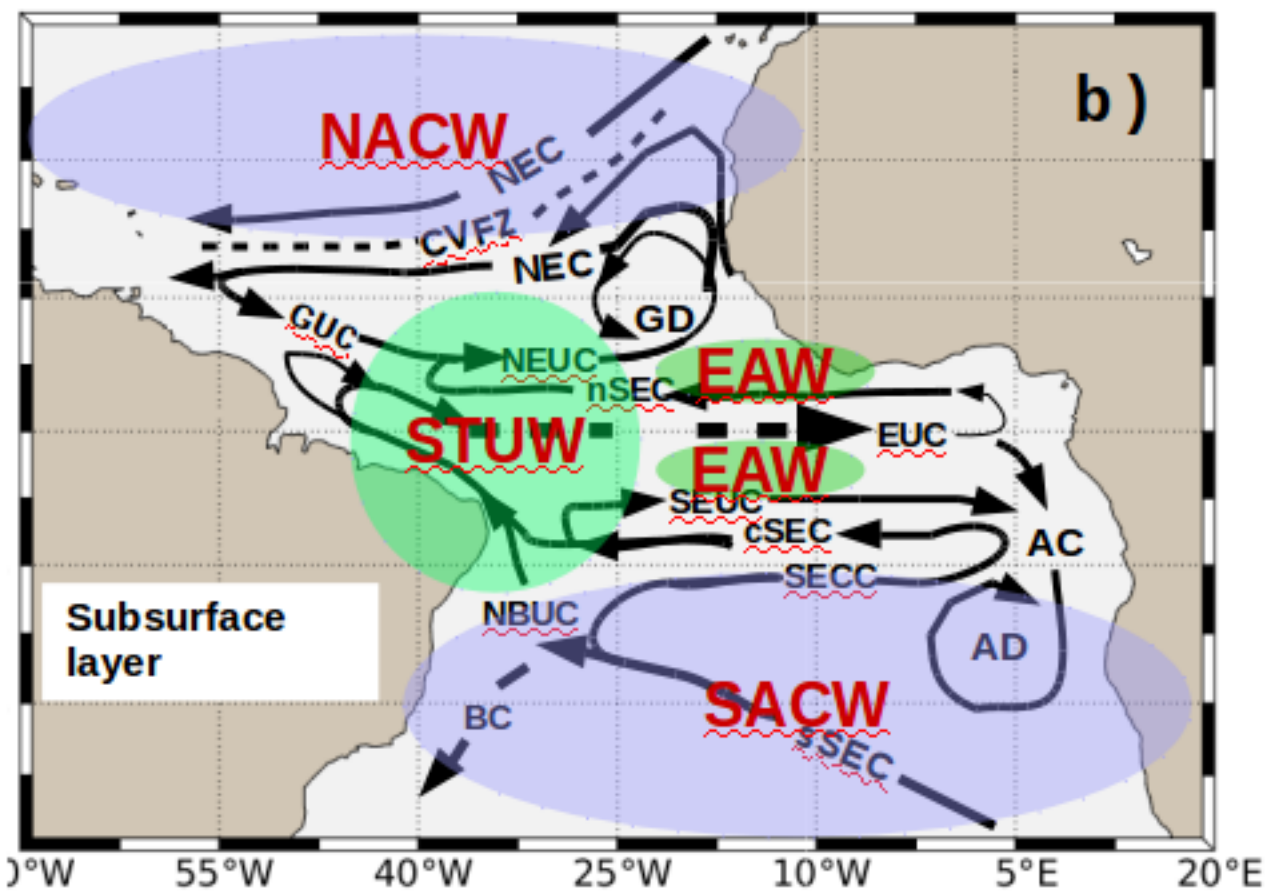
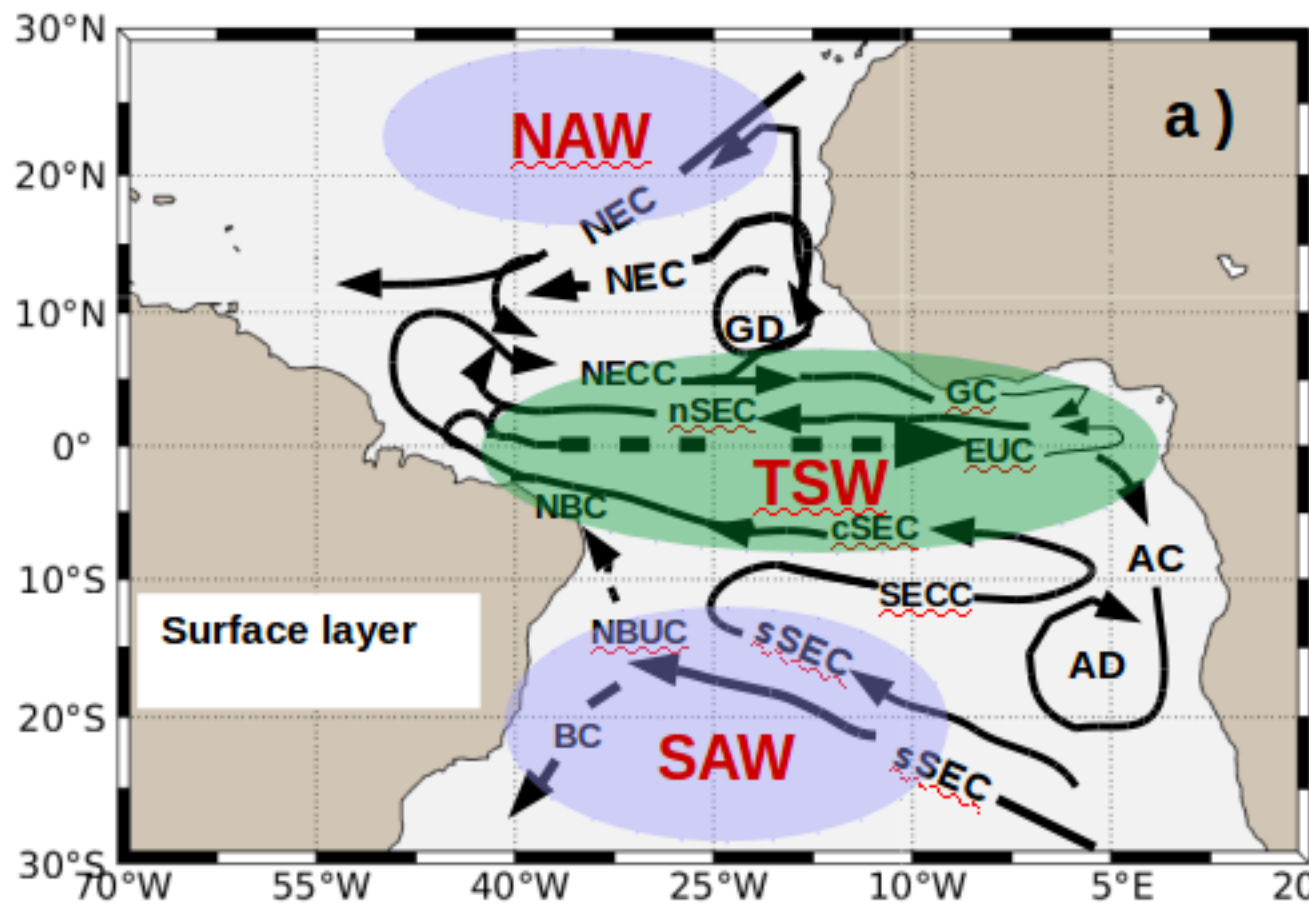


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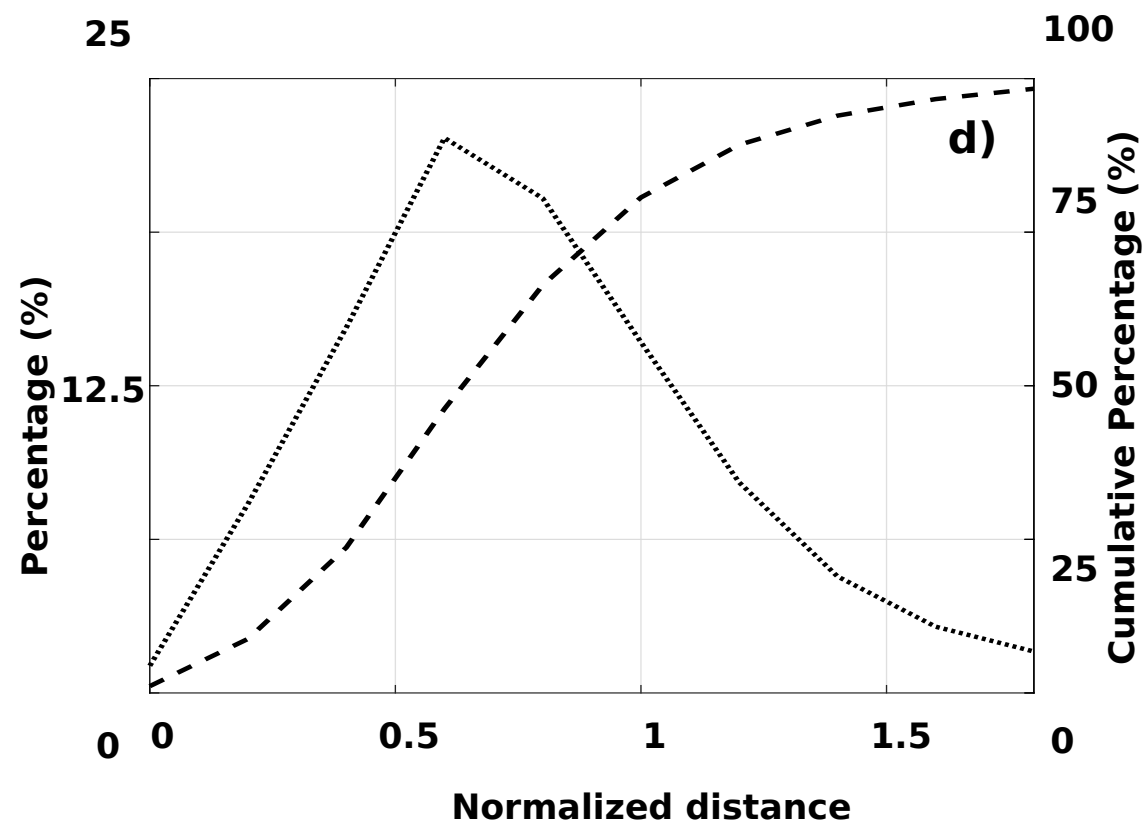
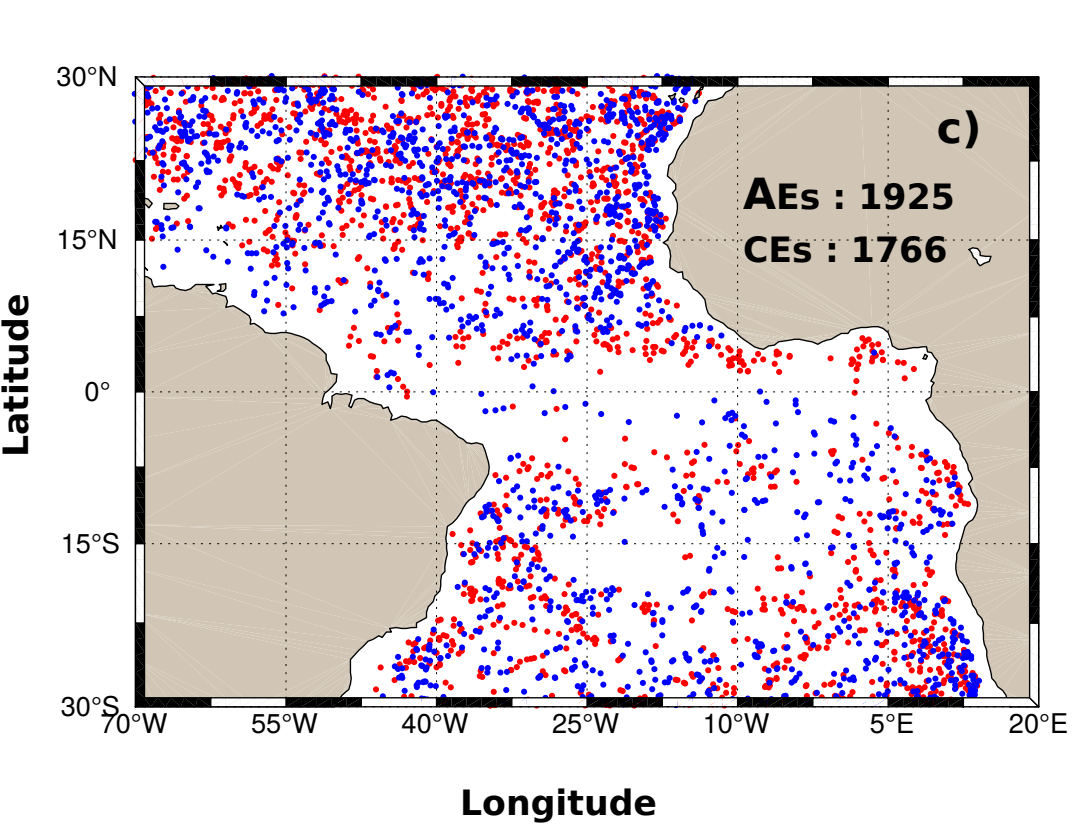
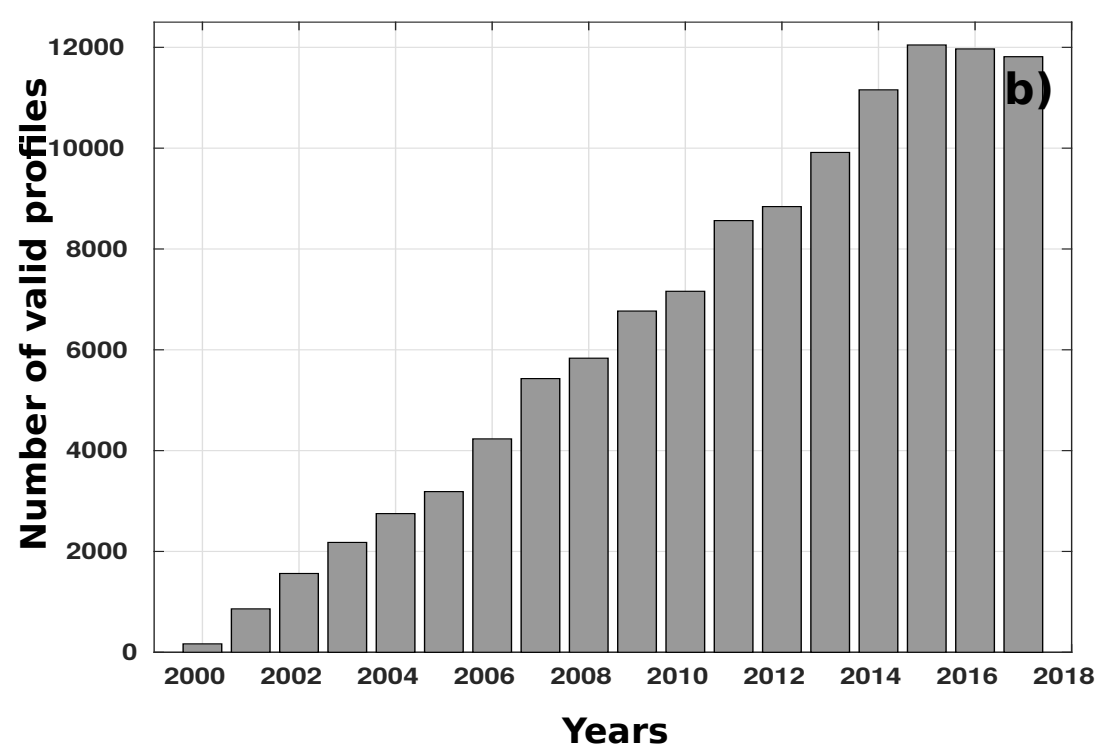
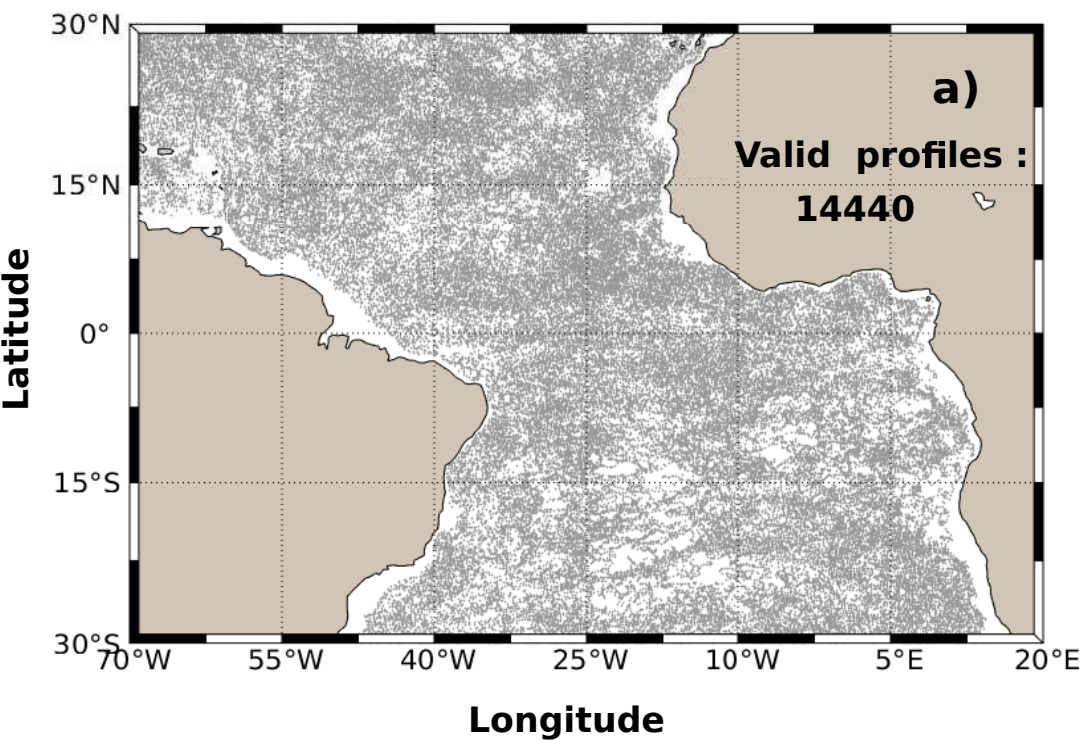


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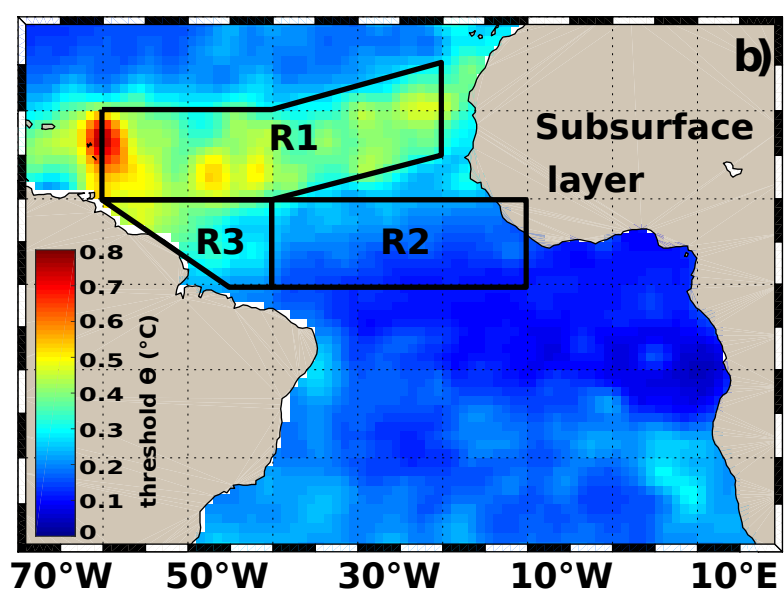
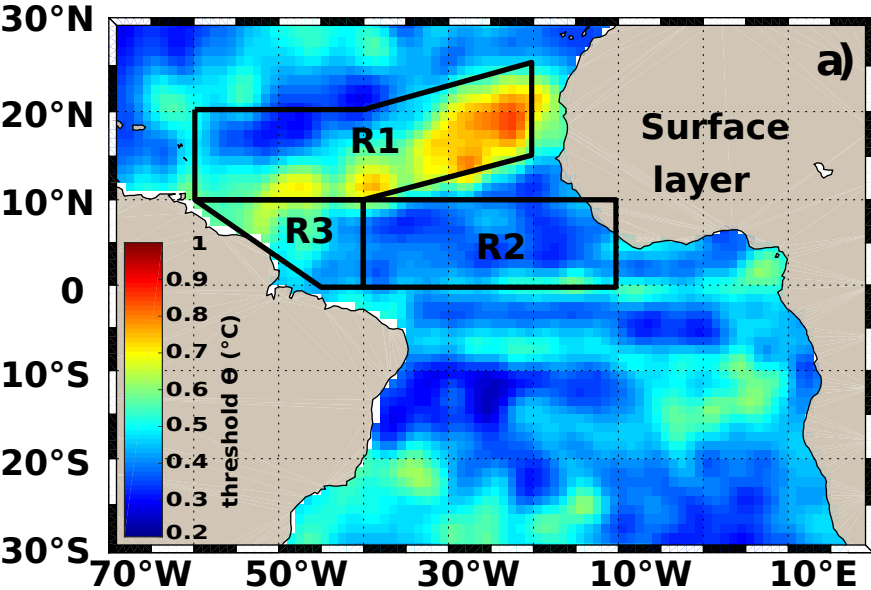


Figure4.

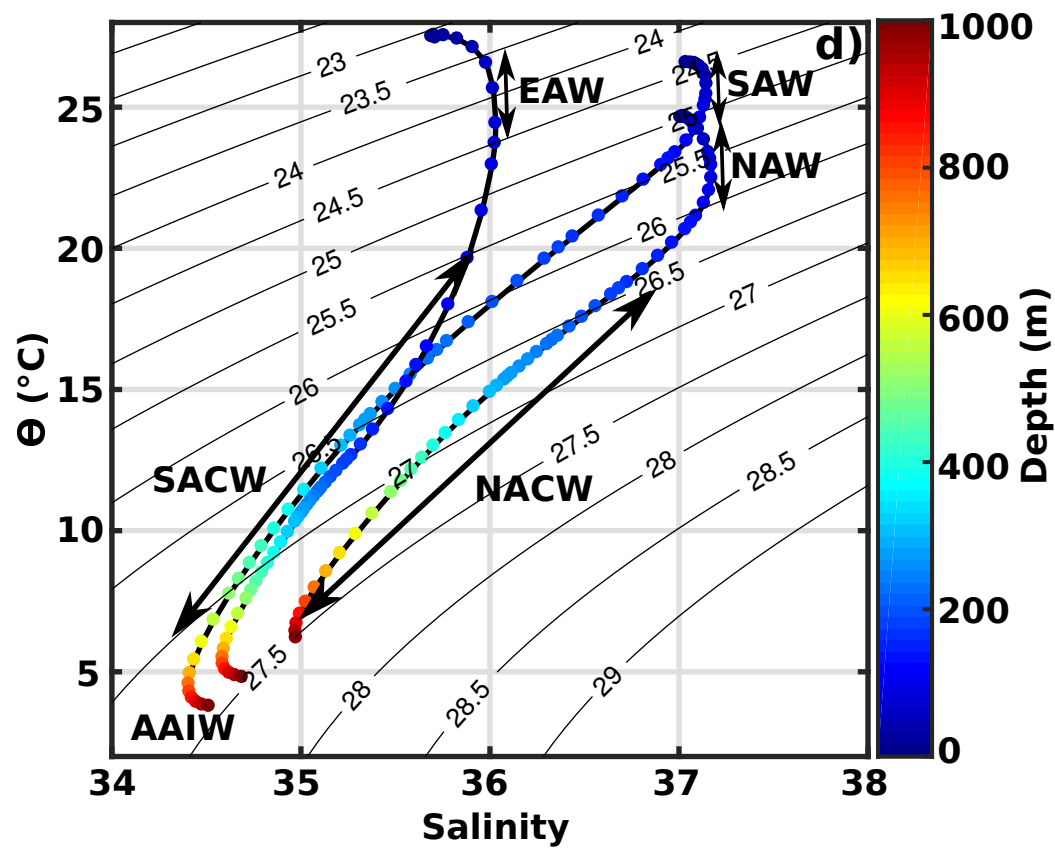
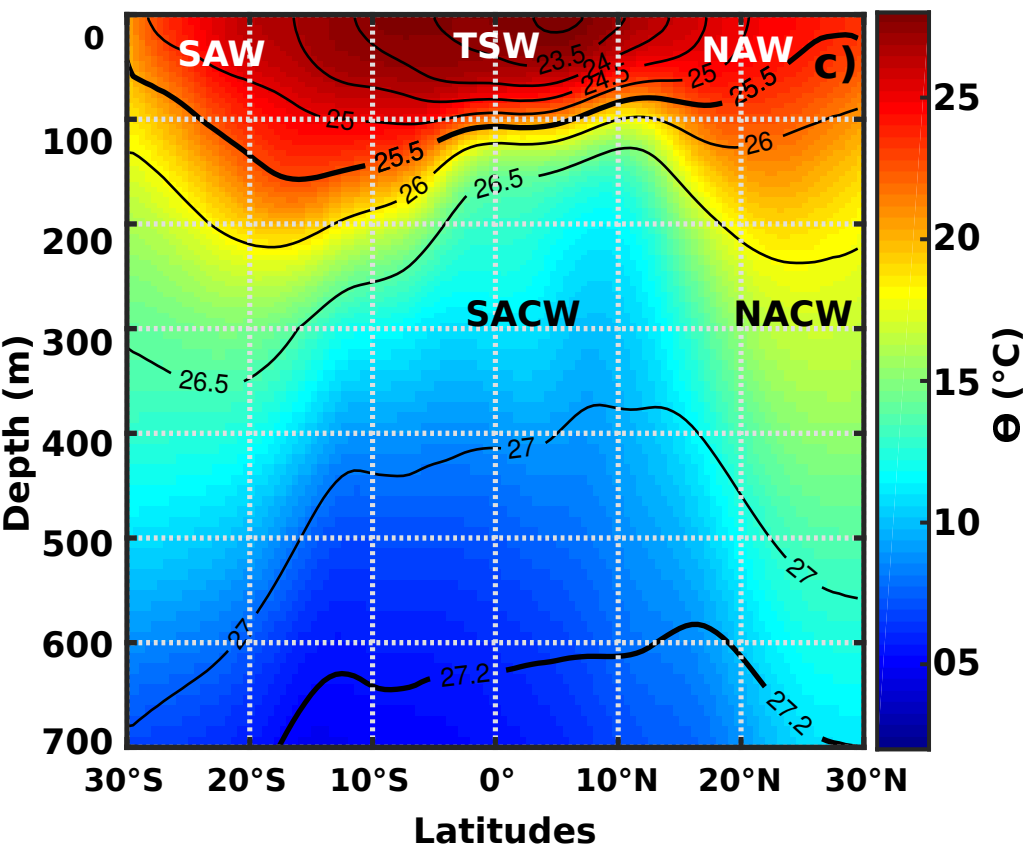
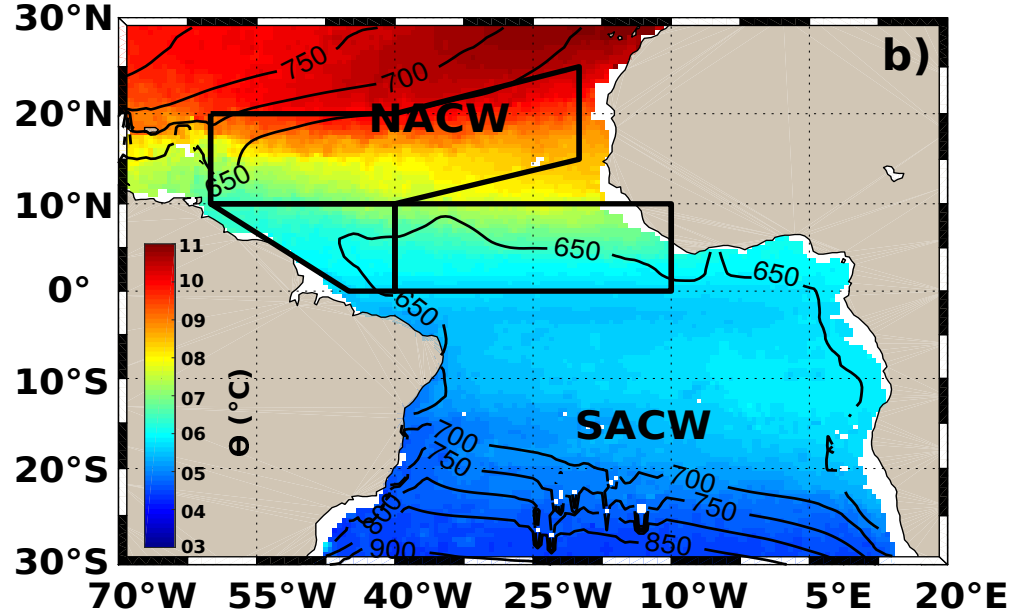
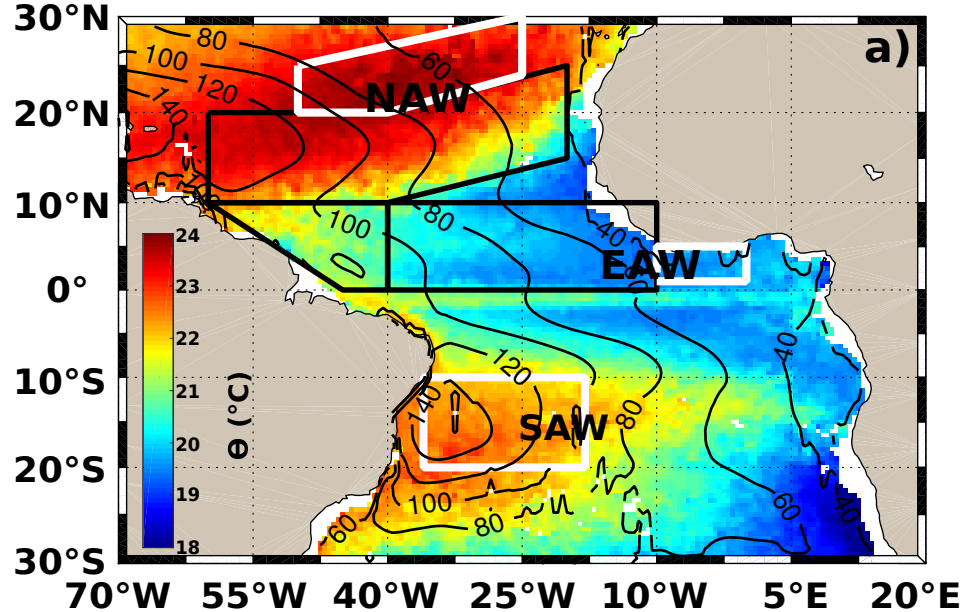


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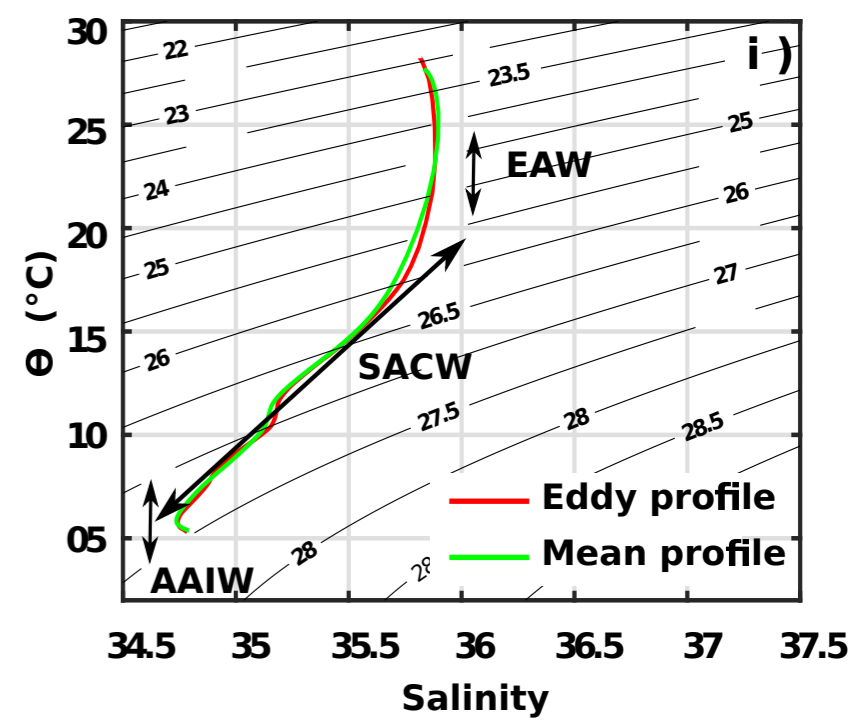
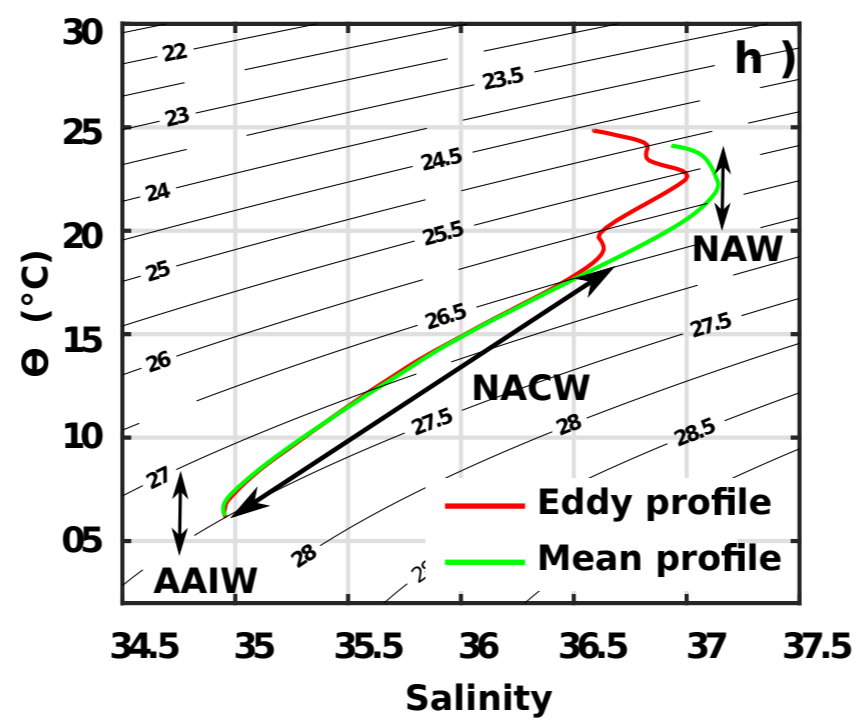
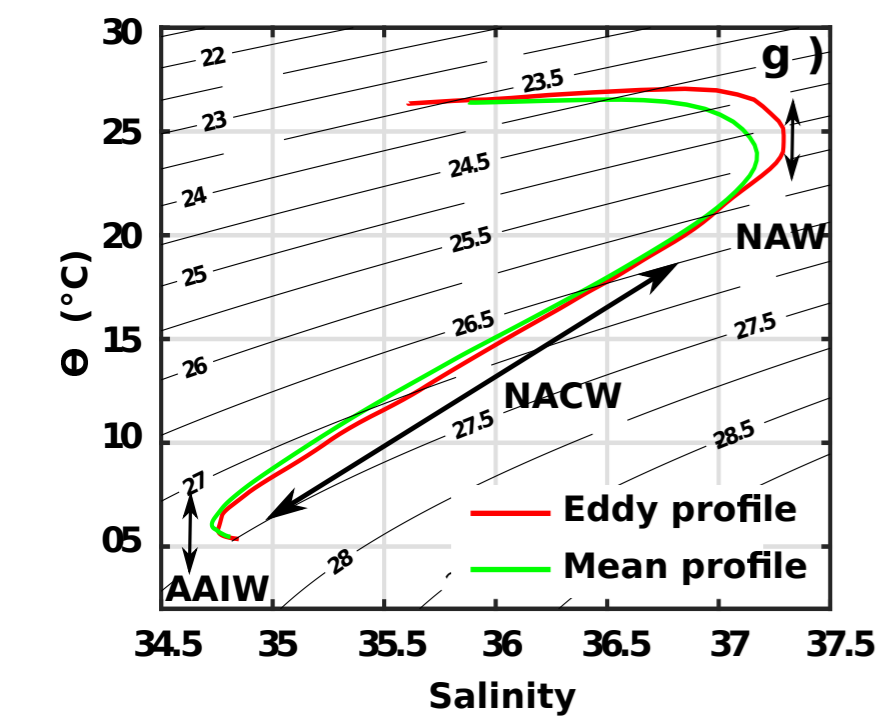
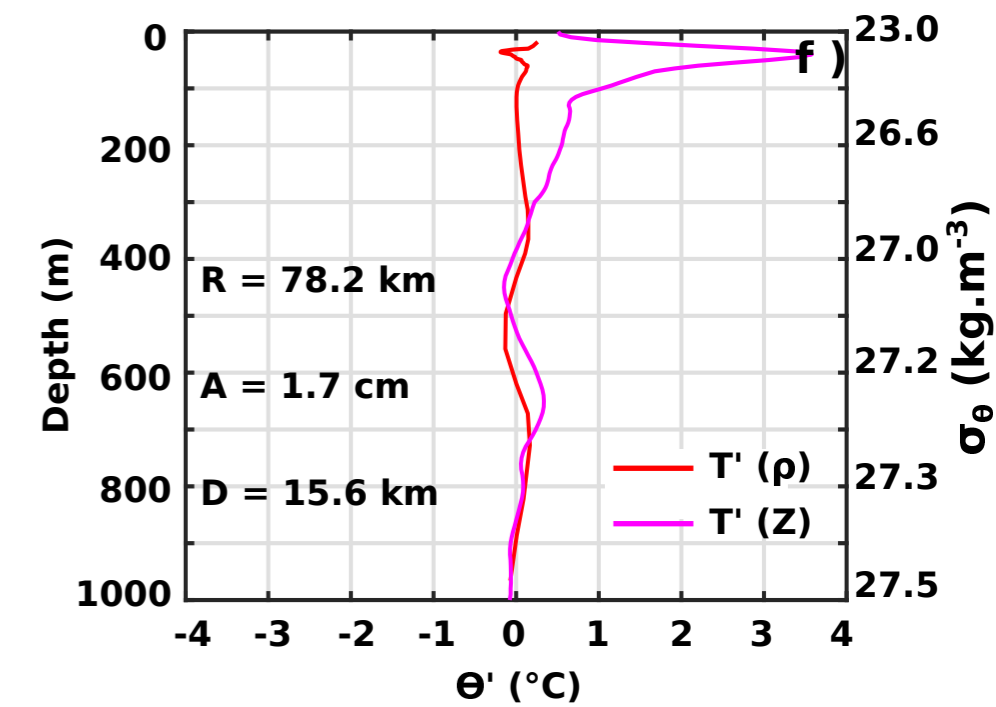
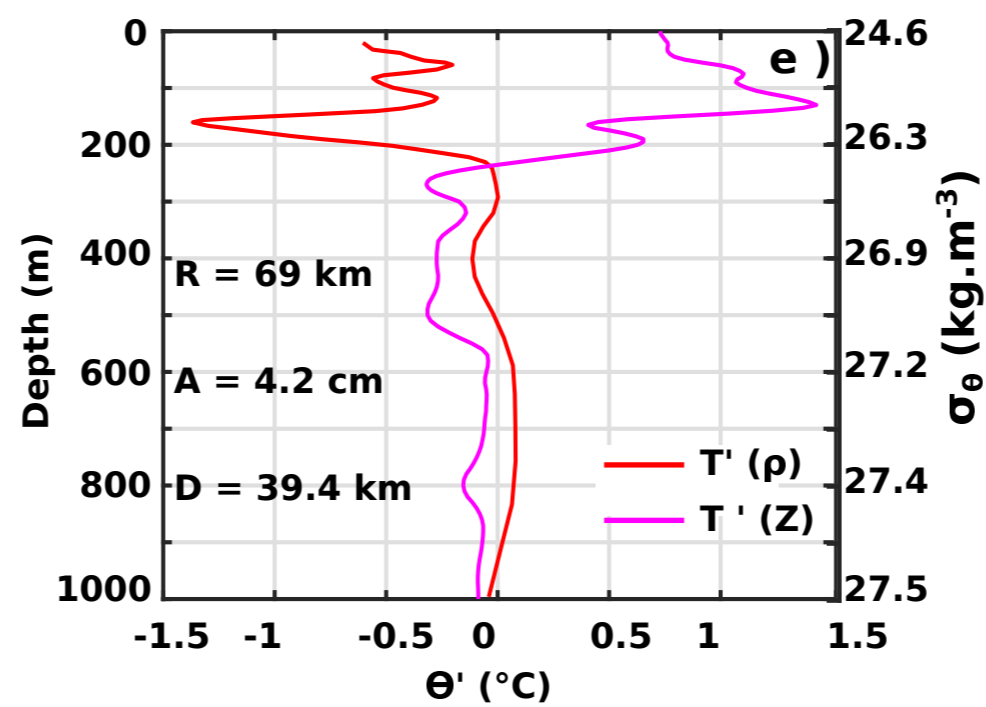
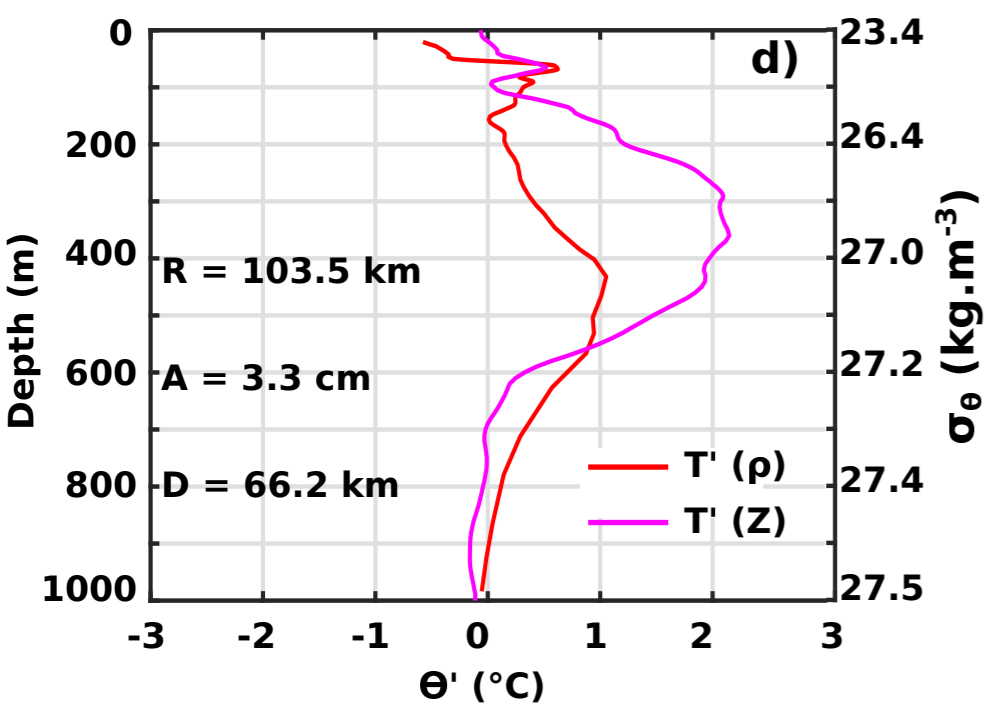
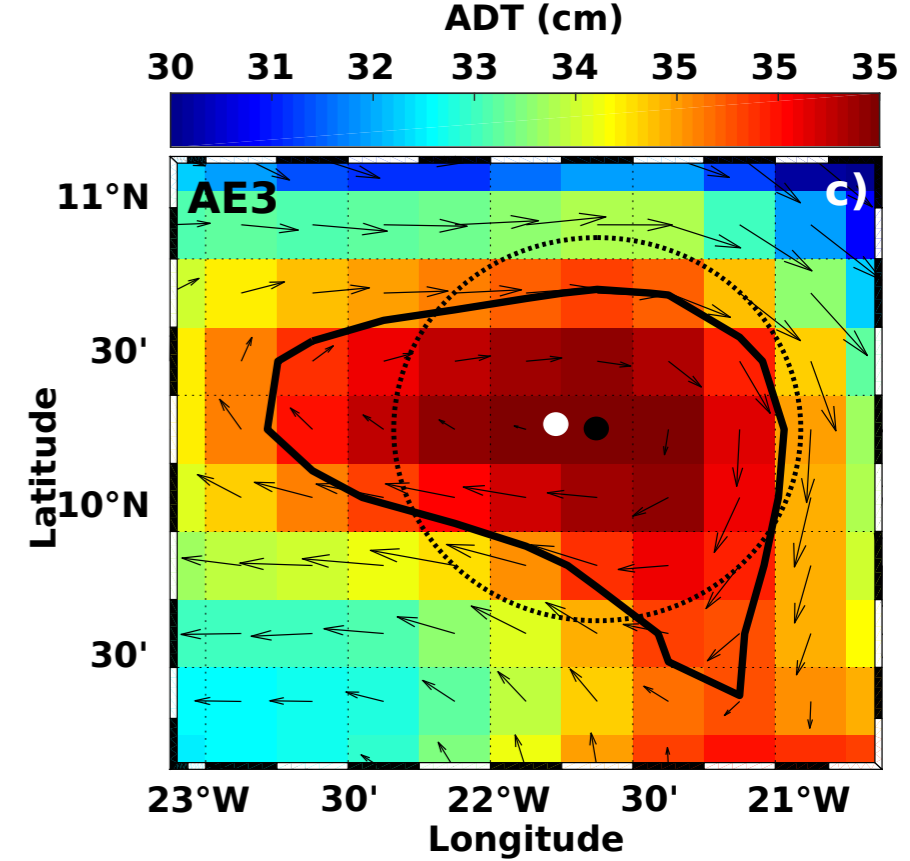
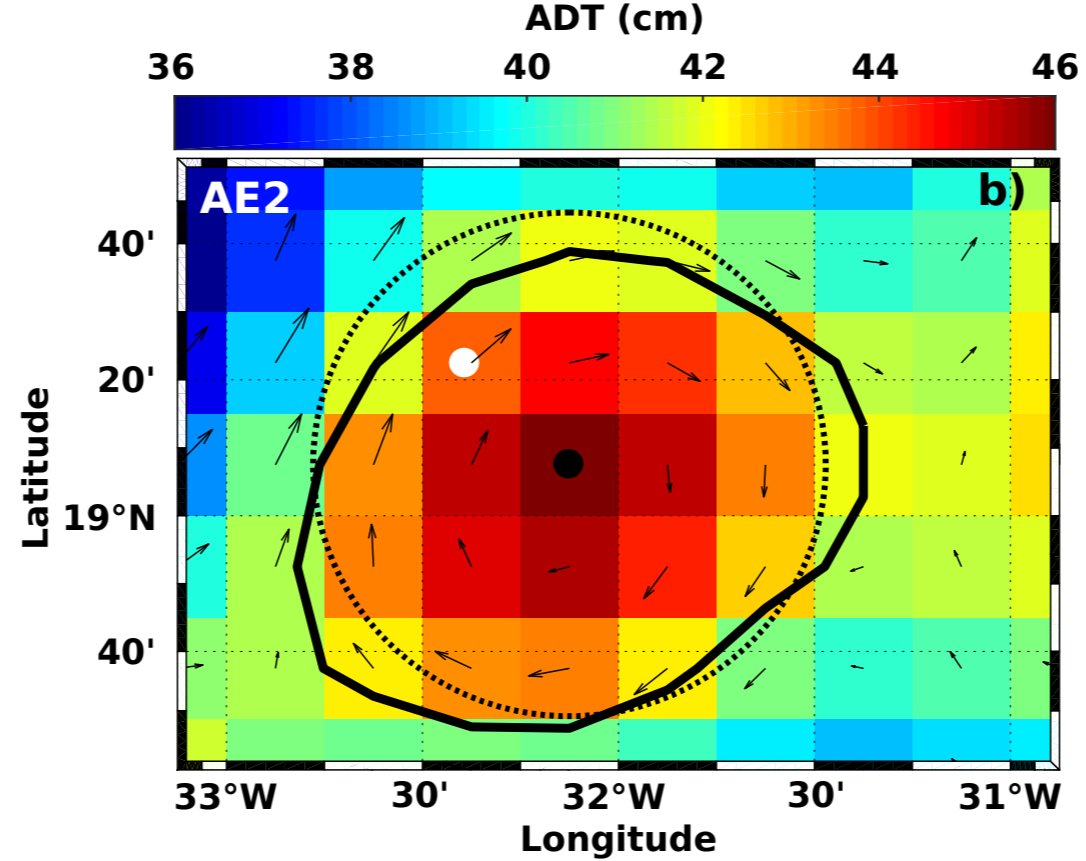
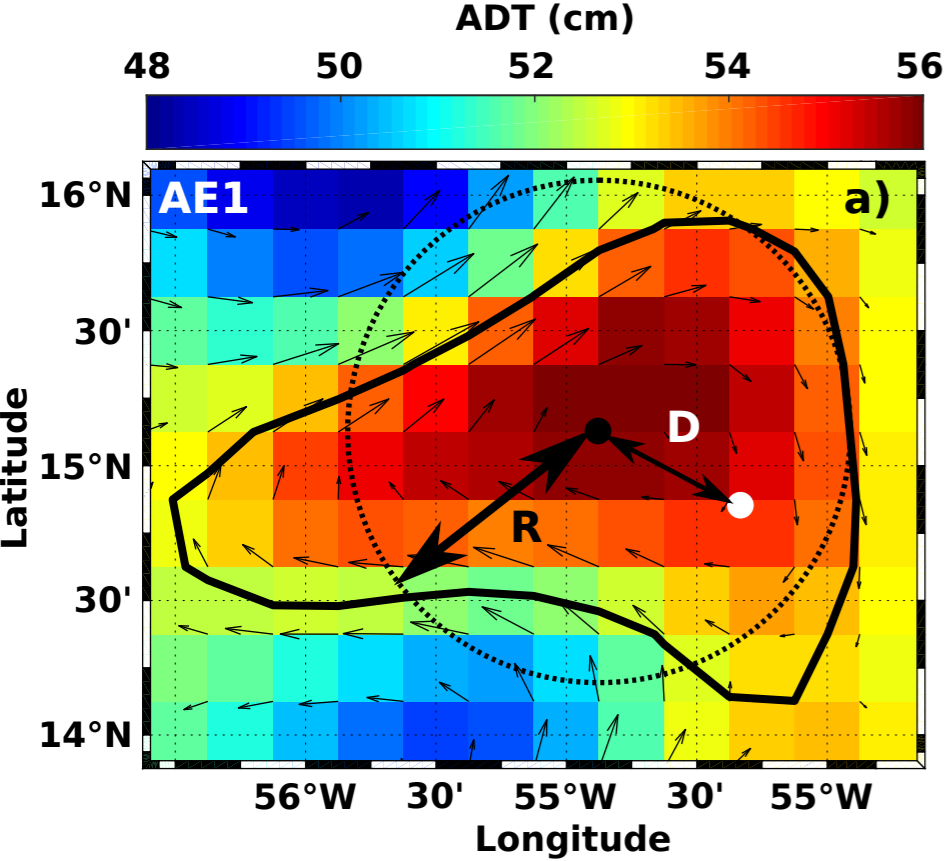


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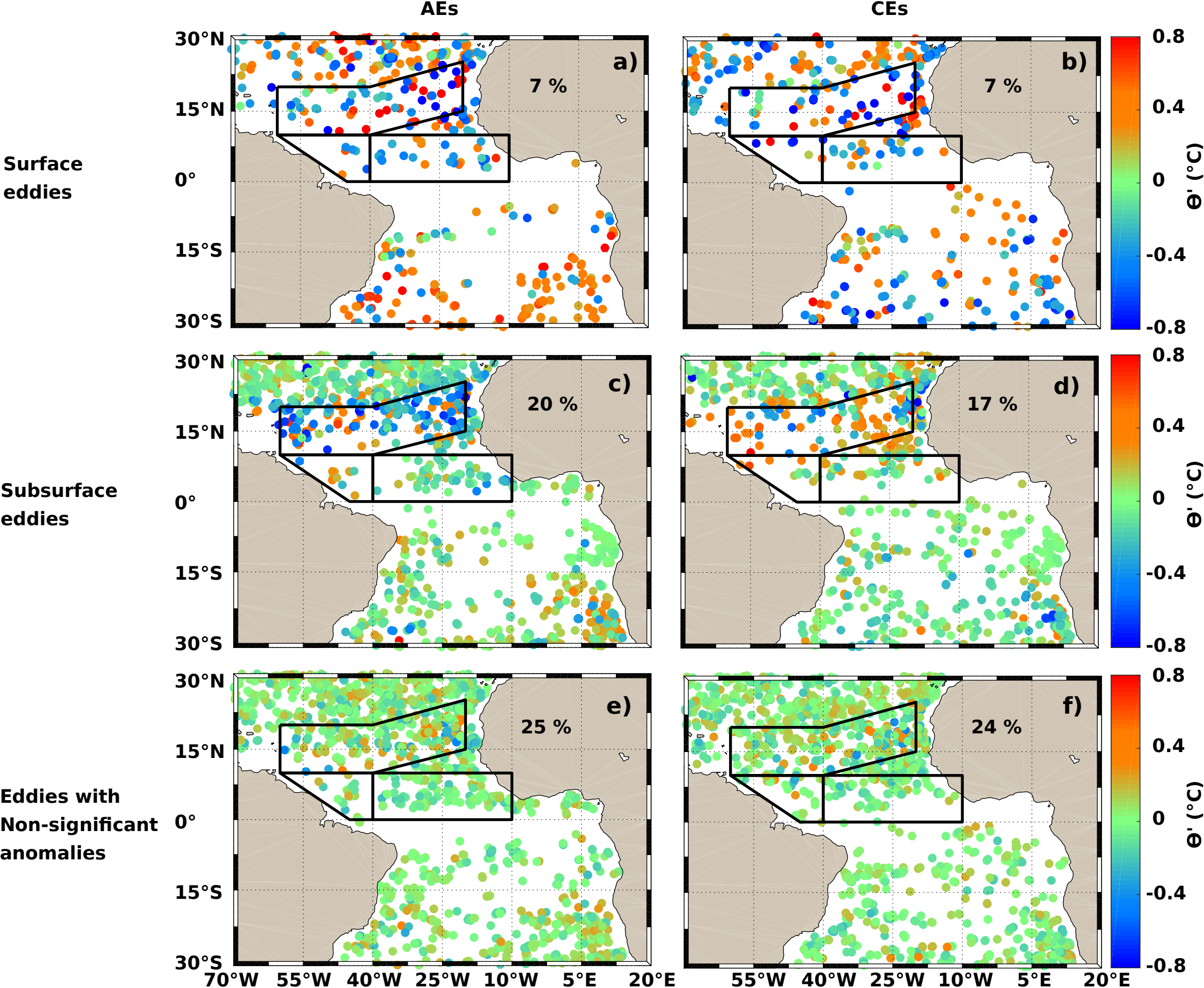


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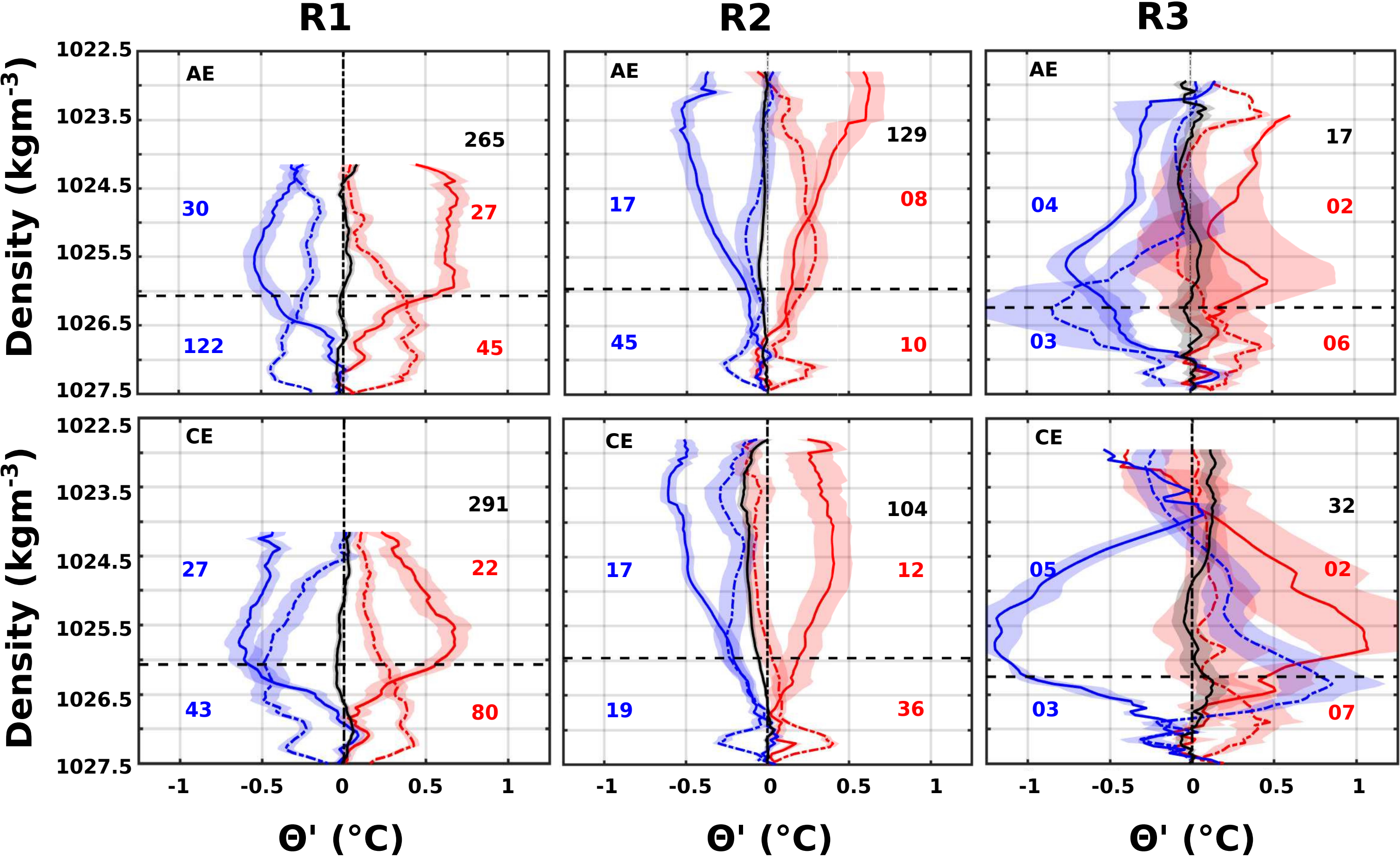
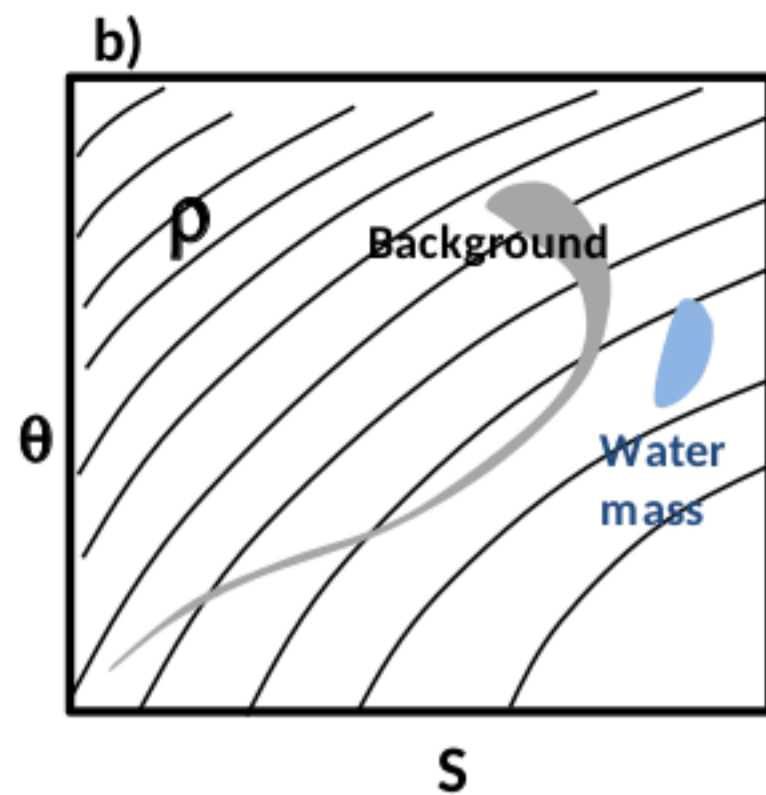
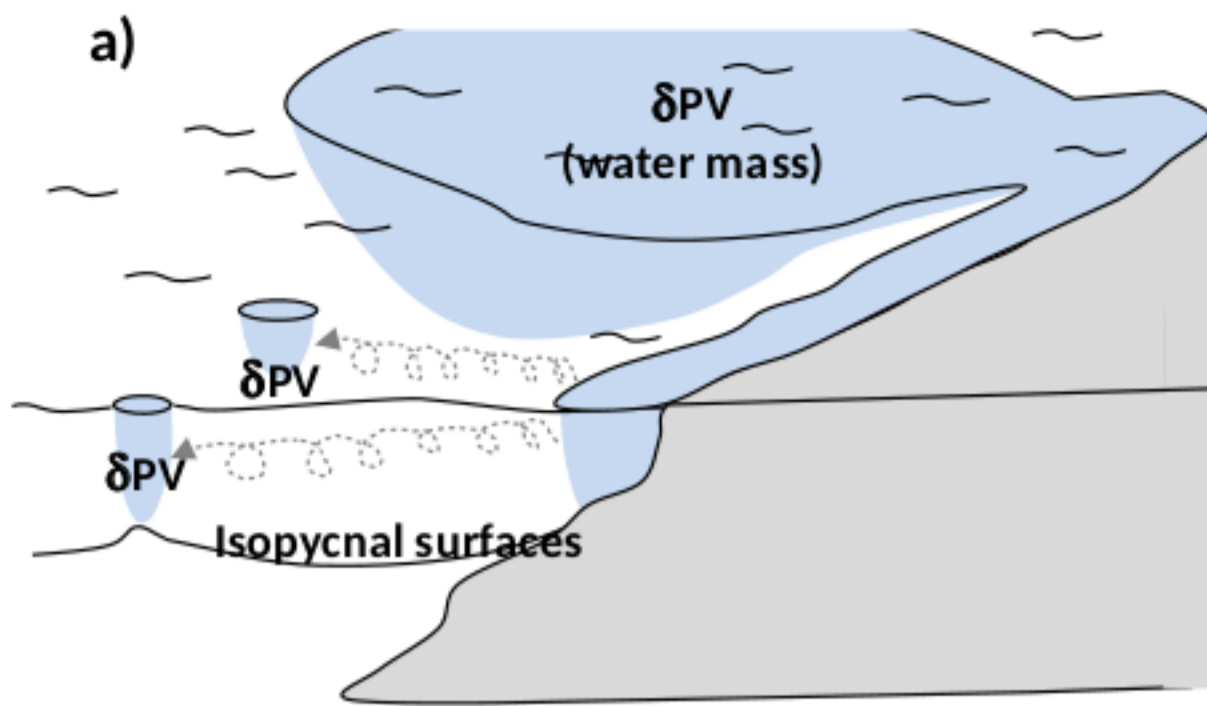
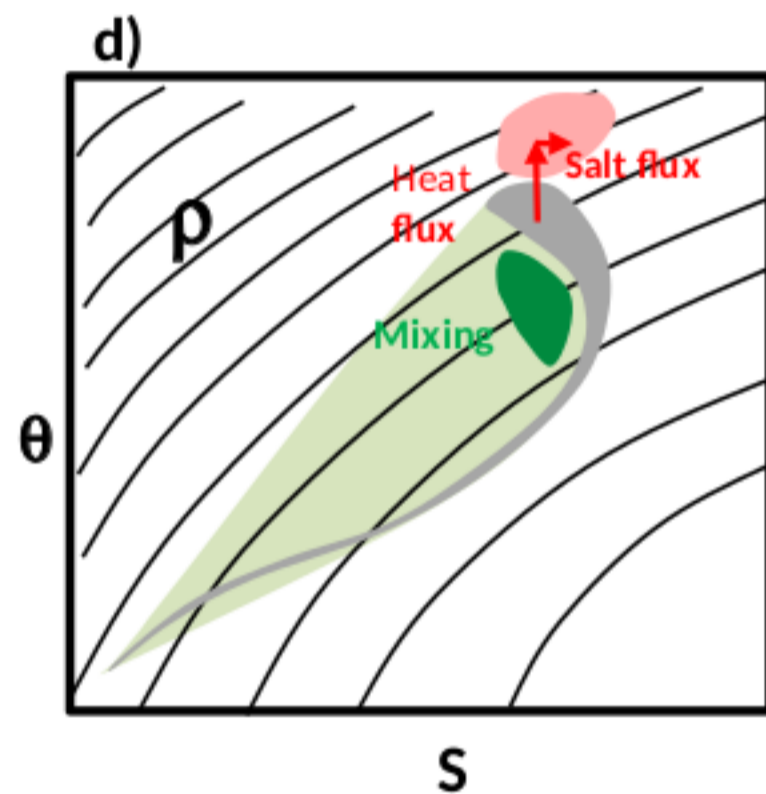
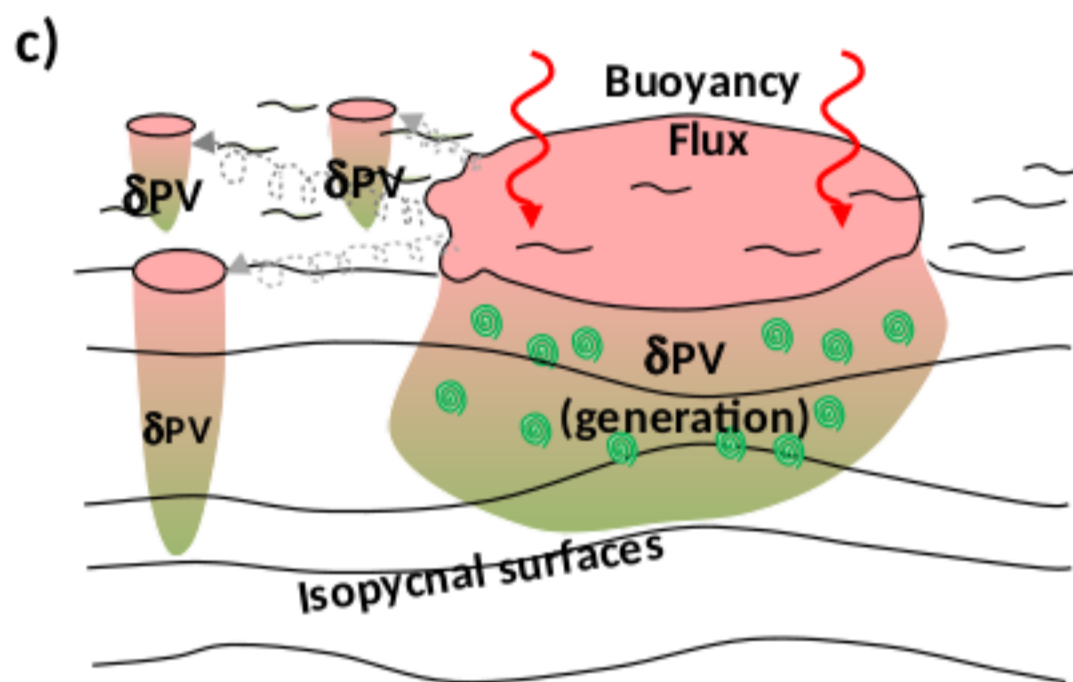


Figure 8.

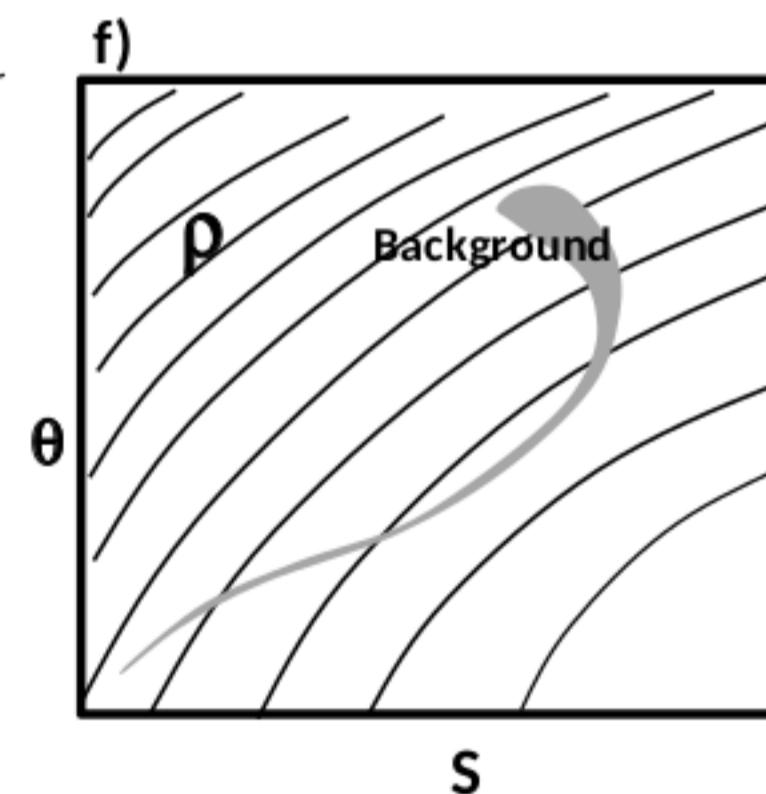
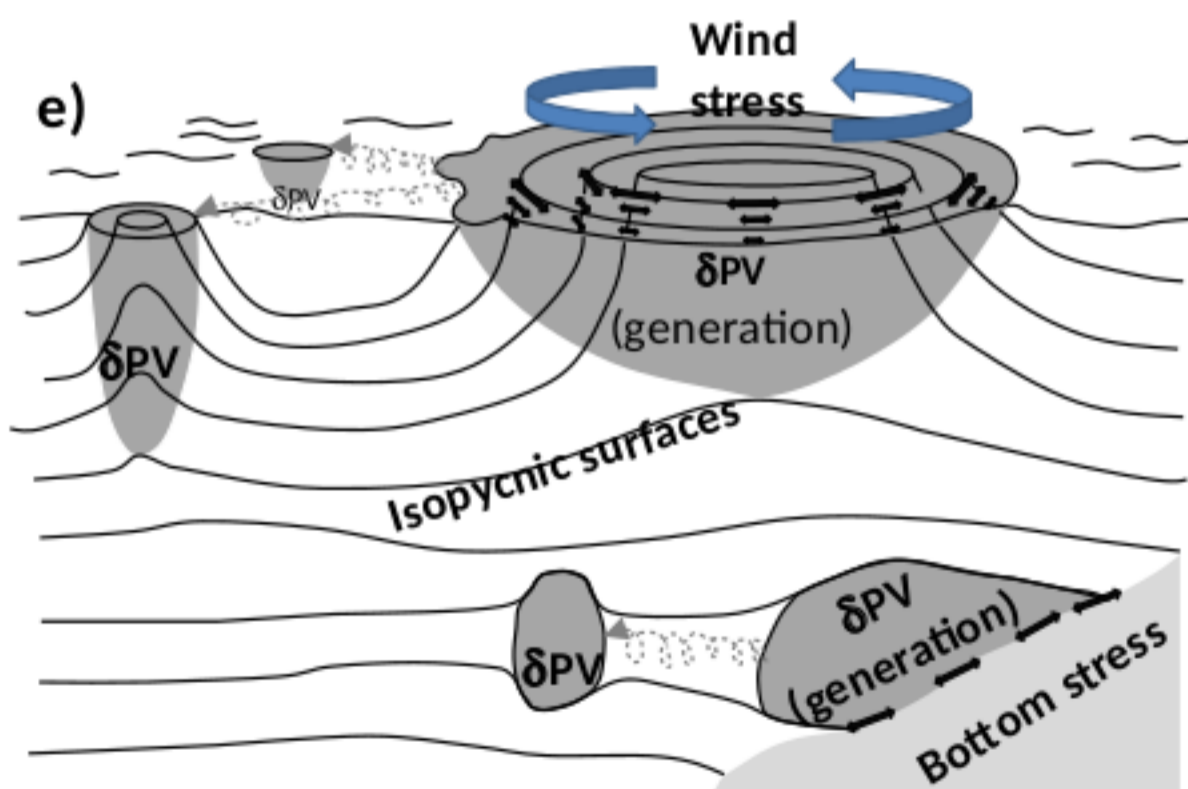
Isopycnal advection



Mixing effects



Frictional effects




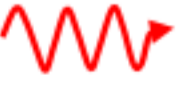

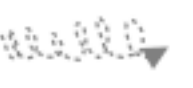
-  Mixing
-  Surface buoyancy flux
-  Frictional forcing
-  Vortex formation

Figure9.

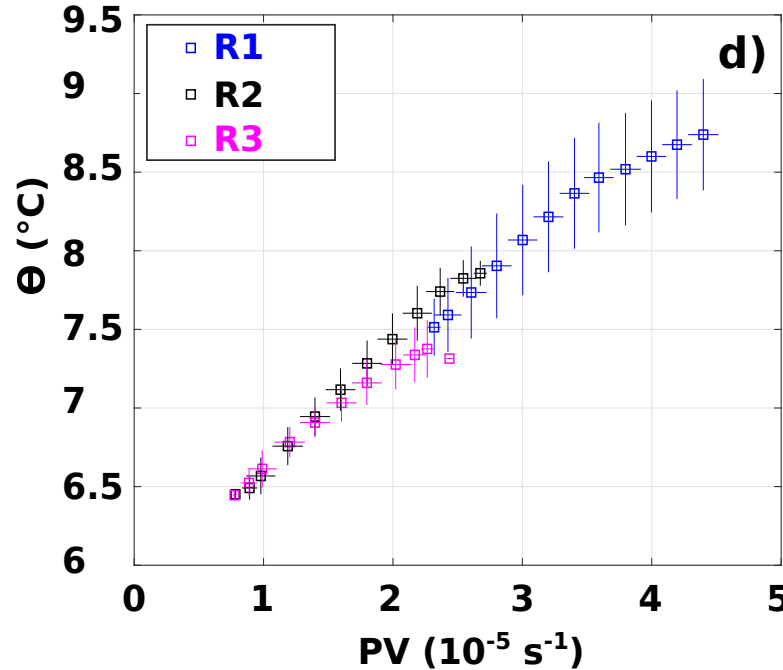
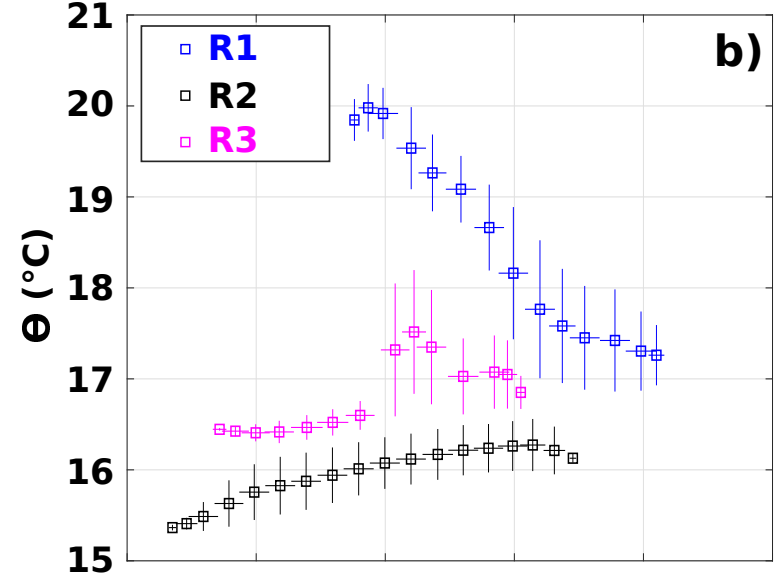
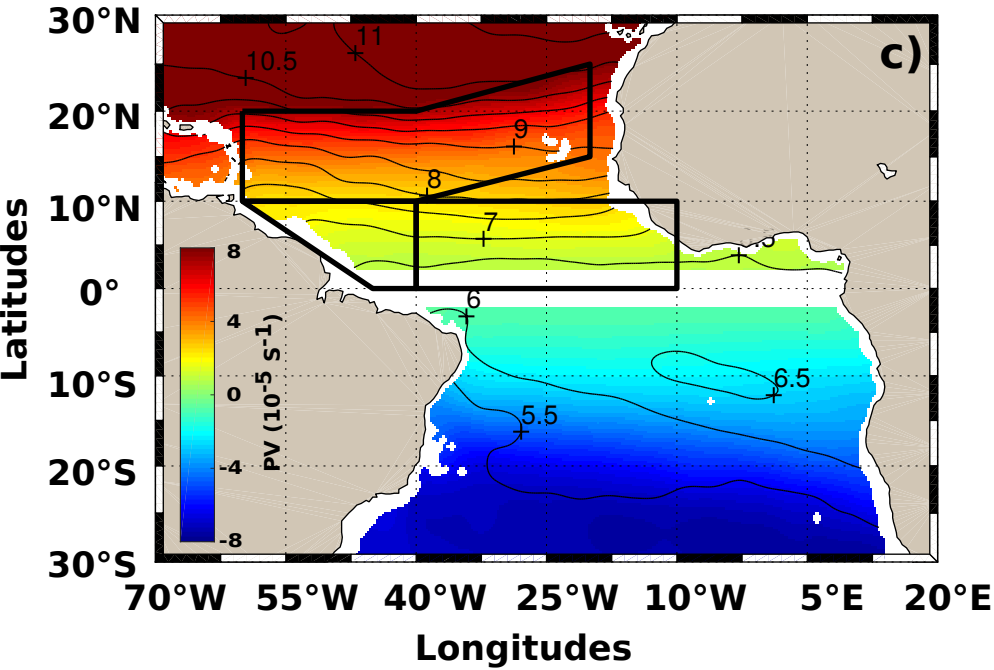
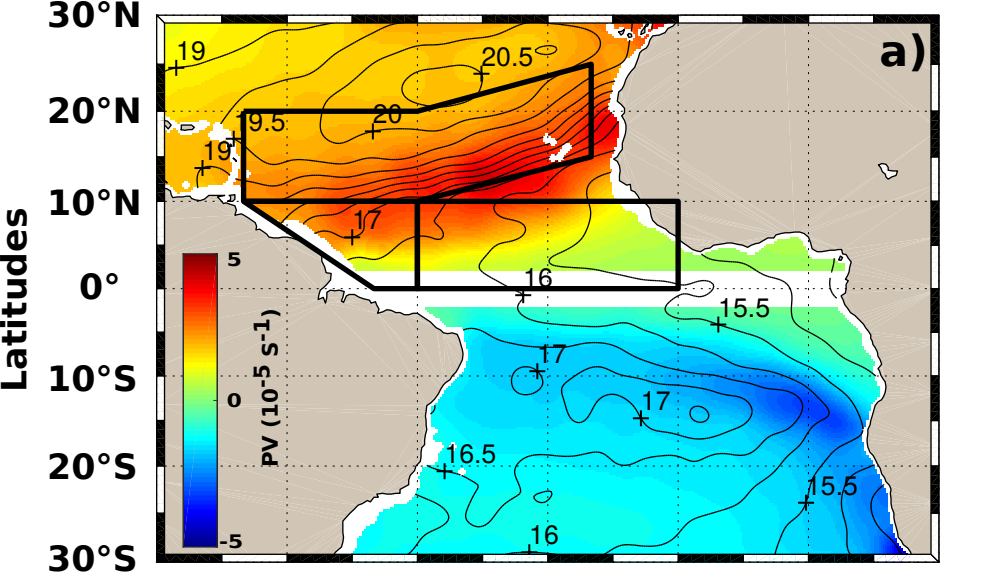


Figure A1.

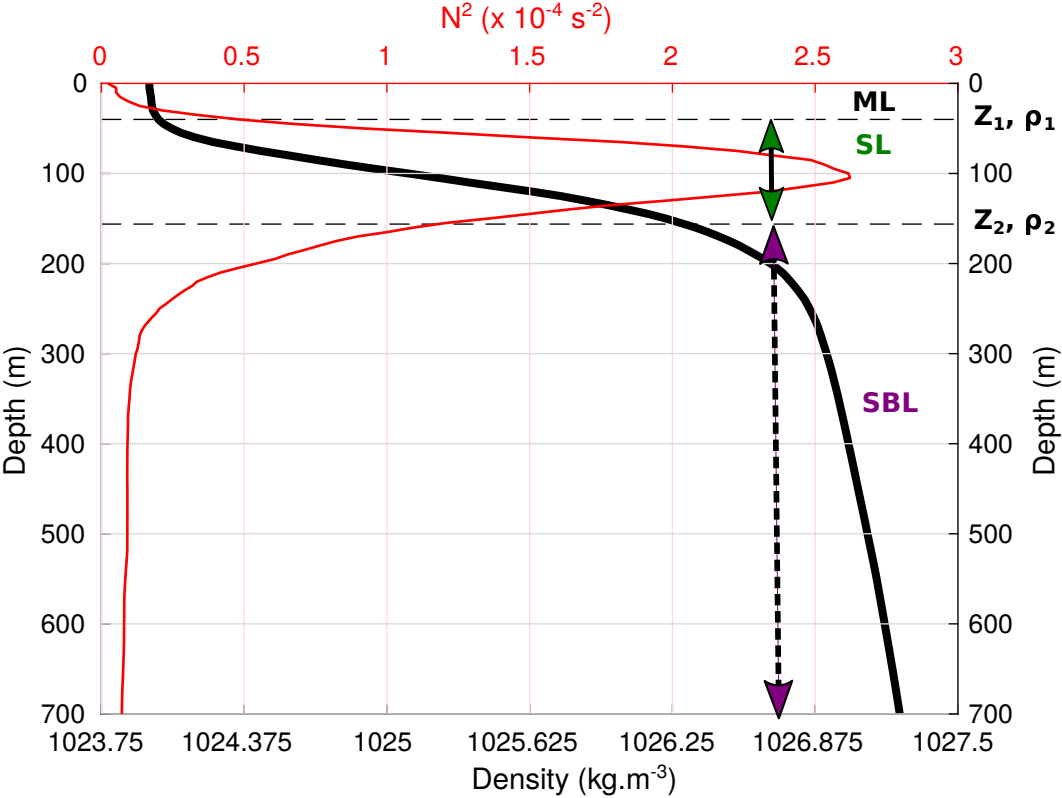


Figure A2.

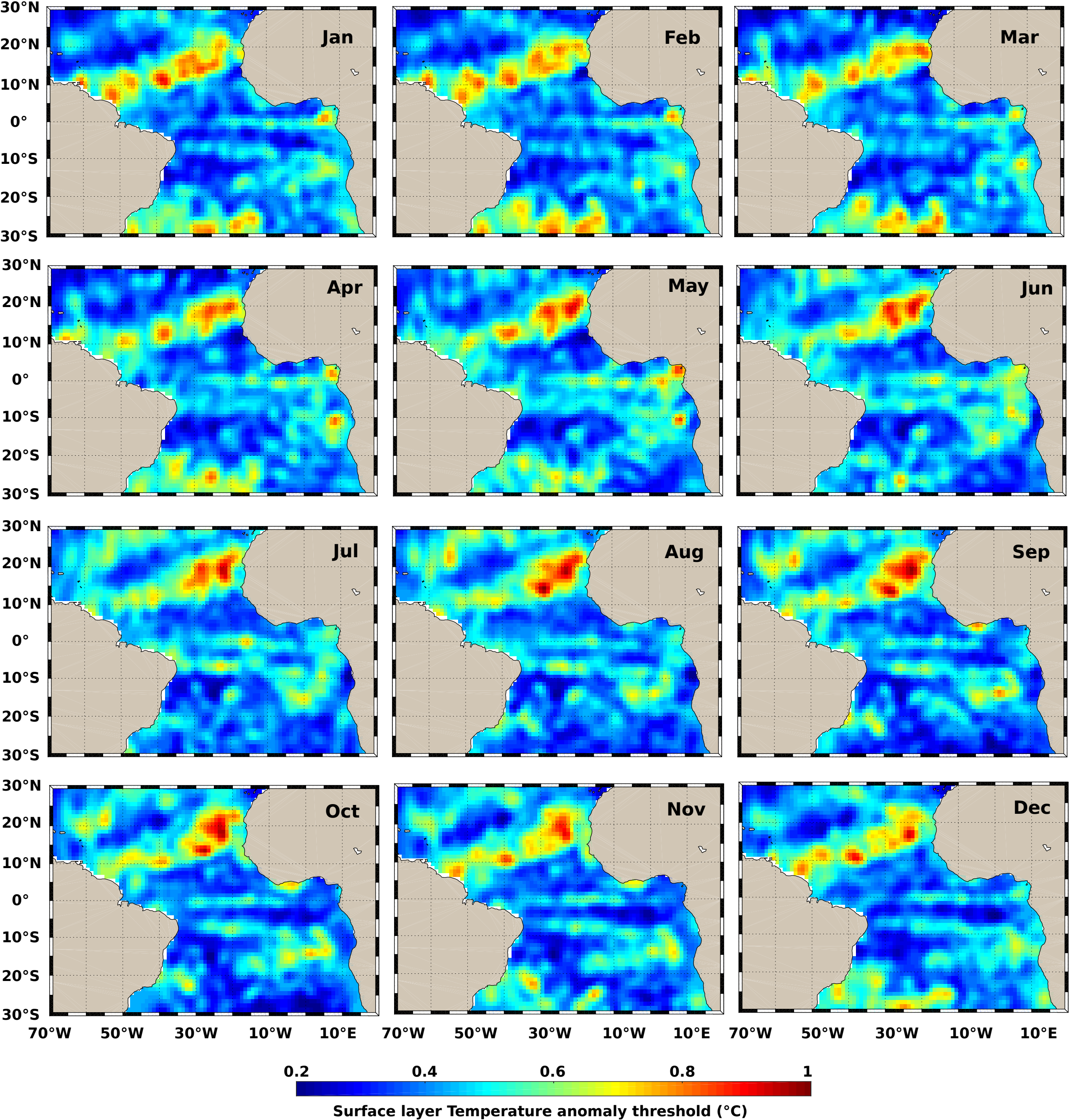


Figure A3.

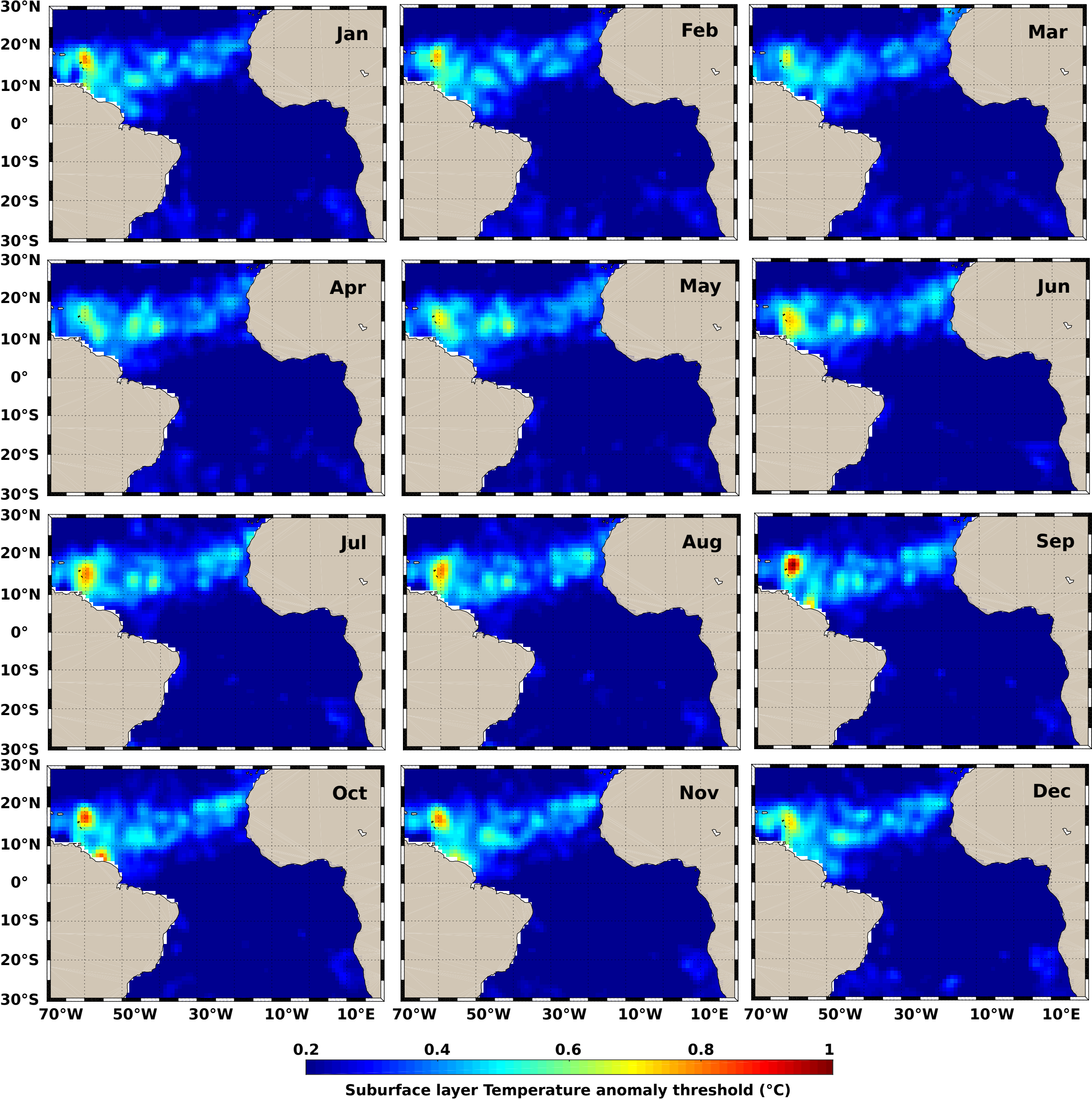


Figure A4.

