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## **REFLECT – Research Flight of EURADOS and CRREAT: Intercomparison of various radiation dosimeters onboard aircraft**

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1 **REFLECT – Research Flight of EURADOS and CRREAT**  
2 **Intercomparison of various radiation dosimeters onboard**  
3 **aircraft**

4 Running head: Dosimeters' intercomparison onboard aircraft

5

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36 **Abstract**

37 Aircraft crew are one of the groups of radiation workers which receive the highest  
38 annual exposure to ionizing radiation. Validation of computer codes used routinely  
39 for calculation of the exposure due to cosmic radiation and the observation of  
40 nonpredictable changes in the level of the exposure due to solar energetic particles,  
41 requires continuous measurements onboard aircraft. Appropriate calibration of  
42 suitable instruments is crucial, however, for the very complex atmospheric radiation  
43 field there is no single reference field covering all particles and energies involved.  
44 Further intercomparisons of measurements of different instruments under real  
45 flight conditions are therefore indispensable.

46 In November 2017, the REFLECT (**RE**search **FL**ight of **EUR**ADOS and **CR**REAT) was  
47 carried out. With a payload comprising more than 20 different instruments,  
48 REFLECT represents the largest campaign of this type ever performed. The  
49 instruments flown included those already proven for routine dosimetry onboard  
50 aircraft such as the Liulin Si-diode spectrometer and tissue equivalent proportional  
51 counters, as well as newly developed detectors and instruments with the potential  
52 to be used for onboard aircraft measurements in the future. This flight enabled  
53 acquisition of dosimetric data under well-defined conditions onboard aircraft and  
54 comparison of new instruments with those routinely used.

55 As expected, dosimeters routinely used for aircraft dosimetry (such as a tissue  
56 equivalent proportional counter (TEPC) or a silicon detector device like Liulin  
57 agreed reasonable with each other as well as with model calculations. Conventional  
58 neutron rem counters underestimated neutron ambient dose equivalent, while

59 extended-range neutron rem counters provided results comparable to routinely  
60 used instruments. Although the response of some instruments, not primarily  
61 intended for the use in a very complex mixed radiation field such as onboard  
62 aircraft, was as somehow expected to be different, the prove of their suitability was  
63 one of the objectives of the REFLECT. This campaign comprised a single short flight.  
64 For further testing of instruments, additional flights as well as comparison at  
65 appropriate reference fields are envisaged. The REFLECT provided valuable  
66 experience and feedback for future experiments.

67

68 **Keywords**

69 cosmic radiation; aircraft; dosimeter; intercomparison; research flight

70

71

## 72 **1. Introduction**

73 Aircraft crew and airline passengers are exposed to elevated dose rates due to  
74 cosmic radiation onboard aircraft; aircraft crew is considered as a group of workers  
75 receiving one of the highest annual effective doses (ICRP, 1991; ICRP, 2007; ICRP,  
76 2016; IAEA 2003). Radiation protection for aircraft crew has been regulated in the  
77 European Union since 1996 by the EU-Directive 29/96/EURATOM (EURATOM  
78 1996). Since then, this directive was updated with the EU-Directive  
79 2013/59/EURATOM (EURATOM 2013). The EU member states were obliged to  
80 comply with the new regulations by updating their national legislations by February  
81 2018. Annual personal doses from galactic cosmic radiation (GCR) to aircraft crew  
82 members are routinely calculated by various computer codes that are validated  
83 preferably by measurements but also by code intercomparisons. Ongoing  
84 validations of such codes need in-flight measurements with appropriately calibrated  
85 instruments.

86 An assessment of aircraft crew radiation exposure is a complex task. Radiation field  
87 at civil flight altitudes is formed by interactions of mainly GCR (and sporadically  
88 solar energetic particles – SEP) with the atoms of the atmosphere of the Earth. All  
89 types of particles and electromagnetic component such as protons, muons, pions,  
90 electrons, neutrons, gamma rays and others of a wide range of energies covering  
91 several orders of magnitude are present as primary or secondary radiation  
92 (Schraube 2000; ISO 2001; Lindborg et al. 2004). Depending on altitude and  
93 geomagnetic latitude, about 40 % - 70 % of ambient dose equivalent  $H^*(10)$  is due

94 to neutrons, 20 % - 30 % due to electrons, 10 % due to protons and 10 % due to  
95 photons and muons (Schraube et al. 2002a; Lindborg et al. 2004). In addition,  
96 radiation field in the atmosphere is not constant in time and space due to solar  
97 modulation of the GCR, strong variations of particle fluences and energies in  
98 occasional SEPs, latitude effects caused by the geomagnetic field and build-  
99 up/absorption effects resulting from nuclear reactions with the atmospheric nuclei.

100 An assessment of the radiation exposure of aircraft crew requires a determination  
101 of the radiation protection quantity effective dose  $E$  (ICRP 2007). Since the effective  
102 dose is not a measurable quantity, for operational radiation protection purposes, an  
103 operational quantity, the ambient dose equivalent  $H^*(10)$  was introduced (ICRU  
104 1993).  $H^*(10)$  should be a conservative estimate of  $E$ . An empirical determination of  
105  $H^*(10)$  onboard aircraft requires accurate measurements using radiation detectors  
106 sensitive to the different particles and energy ranges. The most important species  
107 are neutrons (from few hundred keV up to few GeV) as they deliver the largest  
108 fraction of dose. The  $H^*(10)$  can be measured with an instrument suitably calibrated  
109 for this quantity what is not a trivial task for instruments to be used in atmospheric  
110 radiation field. For the very complex atmospheric radiation field, with its broad  
111 range of different particles and energies, there exists no single reference field  
112 covering all those radiation components. ISO reference radiation fields do not fully  
113 cover the whole particle and energy range of interest (ISO 2012). Additionally, for  
114 proper calibration, instrument responses for all particles and energies shall be taken  
115 into account. To simulate a cosmic radiation field or some of its components at  
116 aviation altitude, an accelerator-produced field such as provided at CERN EU High

117 Energy Reference Field (CERF) facility (Silari and Pozzi 2017; Pozzi et al. 2017,  
118 Pozzi and Silari 2019) or fields at high-mountains could be also used. However, the  
119 composition and spectra of these fields are not exactly the same as the one present  
120 onboard aircraft. Today, well calibrated Tissue Equivalent Proportional Counters  
121 (TEPC) are considered as the instruments that reasonably well approximate the  
122 operational dose quantity ambient dose equivalent in atmospheric radiation field  
123 (ISO 2012, Lindborg 1999). Other instruments need to be calibrated in appropriate  
124 reference fields or in situ against a TEPC.

125 Many in-flight measurements with different instruments were performed in the past  
126 and an overview of the most important research projects in aviation dosimetry  
127 during 1997-2007 was given in (Beck 2009). Further descriptions and results from  
128 various measurement campaigns onboard aircraft between 1992 and 2003 have  
129 been summarized in (Lindborg et al. 2004). Such measurements were usually done  
130 on single flights with changing altitude and cut-off rigidity (Bottollier-Depois et al.  
131 2004; Kubančák et al. 2015). For constant flight conditions, measurements have  
132 been conducted with only a limited number of instruments, such as TEPC and silicon  
133 spectrometers (Meier et al. 2016; Lindborg et al. 2007; Latocha et al. 2007; Lillhök  
134 et al. 2007). Recently several new detectors that are potentially suitable for onboard  
135 aircraft dosimetry have been developed, but not yet fully tested in the field  
136 (Bottollier-Depois et al. 2019; Yasuda et al. 2020; Kakona et al. 2019).

137 Despite the measurements performed so far, there is still need for continuous  
138 measurements onboard aircraft especially for observing short-term variations of  
139 radiation levels associated with SEP. The silicon spectrometer Liulin has been used



140 onboard aircraft for many years. Several Liulin detectors are permanently installed  
141 onboard aircraft of Air France and Czech airlines (Ploc et al. 2013) although their  
142 sensitivity to neutrons is rather low and they are not tissue-equivalent. A TEPC (e.g.  
143 like Hawk-type) is typically not used for long-term measurements due to its rather  
144 large dimensions and relatively high power consumption. A unique exception is  
145 long-term TEPC measurements reported by (Beck et al, 2005) where the  
146 “Halloween Storms” between October and November 2003 were recorded.

147 Intercomparisons with different types of instruments, which are usually calibrated  
148 in different ways, are necessary. A comparison exercise employing different  
149 instruments conducted in regular time intervals (e.g. every few years) represents an  
150 independent form of a quality control for participating groups. In addition, in a view  
151 of a growing demand for increasing the quality of dosimetric measurements at  
152 aviation altitudes by the space weather community (Tobiska et al. 2015; Meier et al.  
153 2018) measurement campaigns onboard aircraft are necessary.

154 In November 2017, the research campaign REFLECT (**RE**search **FL**ight of **EURADOS**  
155 and **CRREAT**) was carried out by Nuclear Physics Institute CAS. The response of  
156 more than 20 different detectors was investigated during a flight onboard a small  
157 aircraft. The instruments’ ensemble included those already proved for dosimetry  
158 onboard aircraft such as Liulin and TEPCs, as well as newly developed detectors and  
159 instruments with the potential to be used for onboard aircraft measurements in  
160 future. Dosimetric data under well-defined conditions, including constant altitude  
161 and constant space weather conditions, were acquired. Sixteen institutes  
162 participated, several of them representing the leading research groups in aviation

163 dosimetry in their respective countries. As a result, REFLECT is the largest campaign  
164 of this type ever performed. This campaign was part of the research activities of  
165 Working Group 11 of EURADOS (EURADOS 2020) and of the CRREAT (Research  
166 Center of Cosmic Rays and Radiation Events in the Atmosphere) project (CRREAT  
167 n.d.).

## 168 **2.Instruments**

169 Radiation detectors included in the REFLECT campaign embraced instruments  
170 routinely used for cosmic radiation monitoring (TEPC, Liulin), newly developed  
171 radiation detectors as well as detectors with future potential for cosmic radiation  
172 monitoring onboard aircraft. With one exception, all instruments were active  
173 radiation detectors, i.e. electronic instruments capable of making time-resolved  
174 measurements.

175 An overview of the detectors used listing instruments, measured quantities,  
176 typically used radiation fields and participating institutes is given in Table 1. The  
177 detectors routinely used are underlined. Others are various neutron rem-counters,  
178 Si-detectors, recombination chamber or scintillation detectors.

179 Table 1. Instruments used during REFLECT

<b>Instrument</b>	<b>Quantity measured/ provided</b>	<b>Typical radiation field</b>	<b>Institute</b>
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<u>TEPC Hawk</u>	$H^*(10)$	Mixed radiation	Institute for Radiological Protection and Nuclear Safety, France (IRSN)  Seibersdorf Laboratories, Austria (SL)
<u>Sievert instrument</u>	$H^*(10)$	Mixed radiation	Swedish Radiation Safety Authority, Sweden (SSM)
<u>Liulin</u>	$D(Si), H^*(10)$	Mixed radiation	Nuclear Physics Institute of the CAS, Czech Republic (NPI)  National Institute of Radiological Sciences, QST, Japan (QST)  German Aerospace Center, Germany (DLR)
REM-2 recombination chamber	$H^*(10)$	Mixed radiation	National Centre for Nuclear Research, Poland (NCBJ)
LB 6419	$H^*(10)$	Mixed radiation  Neutrons (thermal – 300 MeV), photons	Deutsches Elektronen-Synchrotron, Germany (DESY)
TTM low-level neutron and gamma	$H^*(10)$	Mixed radiation	National Centre for Nuclear Research, Poland (NCBJ)

monitoring station			
Airdos	$D(Si)$	Mixed radiation	Nuclear Physics Institute of the CAS, Czech Republic (NPI)
Minipix	$D(Si)$	Mixed radiation	Nuclear Physics Institute of the CAS, Czech Republic (NPI)
NM2B-495Pb	$H^*(10)$	Neutrons (up to 10 GeV)	Helmholtz Zentrum München, Germany (HMGU)
LINUS	$H^*(10)$	Neutrons (up to 2 GeV)	European Council for Nuclear Research, Switzerland (CERN)
LB6411	$H^*(10)$	Neutrons (up to 20 MeV)	Nuclear Physics Institute of the CAS, Czech Republic (NPI)
Passive REM counter	$H^*(10)$	Neutrons	Politecnico di Milano, Italy (Polimi)
ELDO	$H_p(10)$	Neutrons (up to 200 MeV)	Helmholtz Zentrum München, Germany (HMGU)
HammerHead HH	$H^*(10)$	Photons (50 keV – 8 MeV), electrons, protons, muons, pions	HHtec for HHtec Association, Czech Republic (HHtec)

FH 40 G-10 with FHZ- 612B probe	$H^*(10)$	Photons	National Centre for Nuclear Research, Poland (NCBJ)
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180

## 181 **2.1 Tissue equivalent proportional counters (TEPC)**

182 A TEPC has the ability to provide values of the dose equivalent in tissue-equivalent  
 183 material from most radiation components reasonably well. It is therefore  
 184 particularly useful in comparisons of cosmic radiation measurements onboard  
 185 aircraft (EURADOS 1996). Several different TEPCs were used to measure the dose  
 186 equivalent during the REFLECT.

### 187 **2.1.1 Hawk environmental Monitoring System FW-AD**

188 The Hawk environmental Monitoring System FW-AD is a tissue equivalent  
 189 proportional counter from Far West Technology Inc. (Goleta, California, USA),  
 190 composed of a spherical chamber (127 mm diameter) with a wall from A-150 tissue  
 191 equivalent plastic (2 mm thick) and filled with pure propane gas at low pressure  
 192 (about 9.33 hPa) simulating of 2  $\mu\text{m}$  site size (Conroy 2004). The outer container is  
 193 made of 6.35 mm thick stainless steel. The dose equivalent is calculated from a  
 194 spectrum of single energy deposition events and a radiation quality factor  $Q$ ,  
 195 determined by the  $Q(L)$  relation given in (ICRP 60), where  $L$  denotes the  
 196 unrestricted linear energy transfer (LET) in the exposed material (ICRP 2007).

197 Both IRSN and SL used Hawk type 1 systems using two linear multichannel analyzers  
198 working in parallel with low and high gains. The low-gain analogue to digital  
199 converter (ADC) measures LET spectra up to  $1024 \text{ keV}\cdot\mu\text{m}^{-1}$  with  $1 \text{ keV}\cdot\mu\text{m}^{-1}$   
200 resolution. The high-gain channel uses an ADC measuring up to a lineal energy of  $25.6$   
201  $\text{keV}\cdot\mu\text{m}^{-1}$  with a resolution of  $0.1 \text{ keV}\cdot\mu\text{m}^{-1}$ . The energy deposition of the low and high  
202 LET components and the associated quality factor are stored in an output file once  
203 per minute. The separation between the low and the high LET component is set at  $10$   
204  $\text{keV}\cdot\mu\text{m}^{-1}$  according to the  $Q(L)$  relationship (ICRP 2007). Events, encountering  
205 significant electronic noise, below the so-called low energy threshold ( $0.3 \text{ keV}\cdot\mu\text{m}^{-1}$   
206 for IRSN and  $0.5 \text{ keV}\cdot\mu\text{m}^{-1}$  for SL) are not recorded. An extrapolation function based  
207 on  $^{60}\text{Co}$  gamma-rays LET spectrum was chosen for the SL Hawk. For the IRSN Hawk  
208 data analysis, a simple coefficient (the average of correction factor determined for  
209  $^{60}\text{Co}$  and  $^{137}\text{Cs}$  gamma-rays) was applied (Farah et al. 2017). No compensation of the  
210 counting loss due to dead time is included in the analysis software.

211 Correction factors,  $N_{low}$  and  $N_{high}$  to ambient dose-equivalent for the low and high LET  
212 components of the dose equivalent are used.  $N_{low}$  was determined in photon radiation  
213 fields with  $^{60}\text{Co}$  and  $^{137}\text{Cs}$  sources.  $N_{high}$  was defined using the neutron reference  
214 sources of  $^{241}\text{Am-Be}$  or  $^{252}\text{Cf}$  neutron sources. The values of  $N_{low}$  are  $1.11\pm 0.02$  and  
215  $1.34 \pm 0.03$  and the values of  $N_{high}$  are  $0.80\pm 0.09$  and  $0.84 \pm 0.10$  for IRSN and SL,  
216 respectively. Correction coefficients for neutrons were also evaluated for between  $0.5$   
217 and  $19 \text{ MeV}$  and were found similar to Am-Be or  $^{252}\text{Cf}$  neutron sources (Trompier et  
218 al. 2007).

### 219 **2.1.2 Sievert instrument**

220 The Sievert instruments are microdosimetric detectors developed by SSM (Kyllönen  
221 et al. 2001a; Lillhök et al. 2017). The detectors are TEPCs with 5 mm A-150 walls  
222 housed in vacuum containers of 2 mm aluminum. The detector volume has a  
223 diameter and length equal to 11.54 cm and a volume of 1207 cm<sup>3</sup>. The detectors are  
224 working at a gas pressure of 1.3 kPa of propane based tissue-equivalent gas with  
225 (volume fractions) 55% C<sub>3</sub>H<sub>8</sub>, 39.6% CO<sub>2</sub> and 5.4% N<sub>2</sub>, to simulate an object size  
226 with a mean chord length of 2 μm.

227 The electric charge is integrated for an integration time of typically 0.1 to 0.3 s. The  
228 absorbed dose to detector gas during this time interval is calculated from the  
229 average charge, the mass of the detector gas, the mean energy required to create an  
230 ion pair (an average value of 27.2 eV was used in the analysis), and the detector gas  
231 multiplication factor.

232 Characterization of the radiation quality is based on the variance-covariance  
233 method (Kellerer 1968; Bengtsson 1970; Lindborg and Bengtsson 1971; Kellerer  
234 and Rossi 1984).

235 In cosmic radiation applications where the high-LET events are rare and the  
236 absorbed dose rate is relatively low, a mixed single-event and multiple-event  
237 analysis can be used (Kyllönen et al. 2001b). The measured spectrum will in such  
238 situations have a region dominated by multiple events, and another region  
239 dominated by single high-LET events. The regions are chosen to be separated at  
240 150 keV·μm<sup>-1</sup>. The quality factor in the multiple-event region (<150 keV·μm<sup>-1</sup>) is

241 calculated from the dose-average lineal energy by using a linear  $Q(y)$  relation. In the  
242 region above  $150 \text{ keV}\cdot\mu\text{m}^{-1}$ , the events are treated as single events (after correction  
243 for a multiple-event contribution),  $y$  was set equal to  $L$  and the corresponding  
244 absorbed dose fraction multiplied by the quality factor defined in ICRP 103 (ICRP  
245 2007). In addition, a correction factor  $c_{Q, \text{high}} = 1.25$  for the high-LET component  
246 below  $150 \text{ keV}\cdot\mu\text{m}^{-1}$  for the difference between the  $Q(y)$ -function used and  $Q(L)$   
247 according to ICRP60 is obtained from a previous comparison of the two approaches  
248 on aircraft measurements (Lillhök et al 2007).

249 From Monte Carlo (MC) simulations of the neutron detector response (Lillhök 2007)  
250 using a simulated atmospheric neutron spectrum (Roesler et al 1998) the detector  
251 absorbed dose and the ambient absorbed dose  $D^*(10)$  agree within 3%.

252 The low-LET and high-LET components are defined as the contribution with dose-  
253 average lineal energy  $1.6 \text{ keV}\cdot\mu\text{m}^{-1}$  measured with these detectors in a  $^{60}\text{Co}$  gamma  
254 radiation field, and  $94 \text{ keV}\cdot\mu\text{m}^{-1}$  simulated for these detectors in a simulated  
255 atmospheric neutron spectrum (Lillhök 2007).

## 256 **2.2 Other detectors for mixed radiation fields**

### 257 **2.2.1 Liulin**

258 The Mobile Dosimetry Unit (MDU) Liulin is a silicon semiconductor spectrometer  
259 that has been used for cosmic radiation measurements (Dachev 2009) as well as  
260 aircraft crew dosimetry for many years (Ploc et al. 2013; Meier et al. 2009). Liulin is  
261 equipped with a Hamamatsu S2744-08 PIN diode ( $10 \times 20 \times 0.3 \text{ mm}^3$ ), low noise



262 hybrid charge-sensitive preamplifier AMPTEK Inc. type A225, fast 12-bit analogue-  
263 digital converter (ADC), 2 or 3 microcontrollers and flash memory. Liulin detects  
264 energy imparted to its active volume in a single energy deposition event. Pulse  
265 amplitudes are stored in a 256-channel spectrum (only 8 most significant bits are  
266 used from ADC), from which the absorbed dose in silicon is then calculated. The  
267 energy calibration of Liulin was obtained at HIMAC (Uchihori et al. 2002).

268 Liulin can also be used to estimate  $H^*(10)$  onboard aircraft, using the absorbed dose  
269 in silicon and an appropriate conversion factor, which can be determined by various  
270 means (Ploc et al. 2011; Wissmann and Klages 2018). In this experiment, several  
271 MDU models were used; however,  $H^*(10)$  is given only for Liulin MDU7 (NPI). MDU7  
272 was recently calibrated at CERF, which enabled to obtain calibration coefficient  
273 converting  $D_{Si}$  to  $H^*(10)$  as described for example in Ploc et al. (2011).

### 274 **2.2.2 Airdos**

275 Airdos is detector with similar design and sensitivity as Liulin. It has been designed  
276 as open source instrument for measurement in mixed radiation fields with low  
277 intensity such as those encountered onboard aircraft (Kakona et al., 2019). It is  
278 composed of a silicon PIN diode (Hamamatsu S2744-09) of the same type used in  
279 the Liulin MDU, electronics for converting the signal to the pulse-height spectra, a  
280 GPS module, an SD memory card and batteries. Full documentation is freely  
281 available at (GitHub, 2020). In this measurement campaign, a version Airdos 01 was  
282 used. Energy range of Airdos 01 is from 0.2 to 12.5 MeV of deposited energy in  
283 silicon with energy resolution 49.4 keV per channel. Accumulated pulse amplitudes

284 are stored in 250 channels spectra every 15 seconds. The detector was calibrated  
285 using heavy charged particle beams at HIMAC (NIRS, Japan) and at the U-120M  
286 cyclotron (NPI, Czech Republic).

### 287 **2.2.3 Timepix**

288 Timepix (Llopart et al. 2002; Llopart et al. 2007) is a hybrid semiconductor pixel  
289 detector which consists of matrix of 256 x 256 pixels (total of 65536 pixels).  
290 Timepix was developed by Medipix2 collaboration (Campbell 2011). The pixel pitch  
291 is 55  $\mu\text{m}$  and total sensitive area is nearly 2  $\text{cm}^2$ . For this flight the silicon with  
292 thickness of 500  $\mu\text{m}$  was used as a semiconductor sensor chip. The Timepix chip  
293 was readout by compact electronics which is in MiniPIX interface (Granja 2018,  
294 Granja and Pospisil 2014). The detector was operated in per-pixel energy mode  
295 which allows to measure the time that the signal spent over threshold. The  
296 calibration between time over threshold to deposited energy was done by method  
297 described in (Jakubek 2011). Due to high granularity provided by Timepix  
298 architecture the detector can measure single particle energy deposition events  
299 (Granja and Pospisil 2014). The configuration of Timepix device and data  
300 acquisition (including the pre-processing of data) was performed in PIXET software  
301 (Turecek 2015) which was run on standard Windows laptop.

### 302 **2.2.4 REM-2 recombination chamber**

303 The REM-2 is a cylindrical parallel-plate recombination chamber with an active  
304 volume of about 1800  $\text{cm}^3$  and total mass of 6.5 kg. The chamber has 25 tissue-

305 equivalent electrodes and it is filled with the gas mixture consisting of methane and  
306 5% nitrogen, with high pressure up to 1 MPa. The effective wall thickness (Al) of the  
307 chamber is equivalent to about 1.8 cm of tissue. The REM-2 chamber approximates  
308 the dosimetric parameters of the ICRU sphere in such a way, that the dose  
309 contribution and energy spectrum of secondary charged particles in the chamber  
310 active cavity are similar to those in the ICRU sphere at the depth of 10 mm (Maciak  
311 2018). Therefore it can be used for the determination of  $H^*(10)$  in mixed radiation  
312 fields (Zielczyński et al. 2008; Caresana et al. 2014; Murawski et al. 2018).

313 The chamber is designed in such a way that the initial recombination of ions occurs  
314 when the chamber operates at polarizing voltages below saturation and, for a  
315 certain range of gas pressure and dose rates, the initial recombination exceeds  
316 volume recombination. Measuring methods are based on the determination of the  
317 dose rate from the saturation current and the radiation quality from the amount of  
318 initial recombination. By means of recombination methods it is possible to estimate  
319 the radiation quality factor (Zielczyński et al. 1994; Golnik et al. 2004; Golnik 2018).

320 The method used for the determination of the radiation quality involves  
321 measurements of two ionization currents  $i_S$  and  $i_R$  at two properly chosen polarizing  
322 voltages  $U_S$  and  $U_R$ . A certain combination of these two currents is called  
323 recombination index of radiation quality  $Q_4$  and may serve as a measurable quantity  
324 that depends on LET in a similar way as the radiation quality factor does (Golnik  
325 2018). The polarizing voltage  $U_S$  is the high voltage, the same as for the  
326 measurements of the absorbed dose. The lower voltage  $U_R$ , called the recombination  
327 voltage, has been determined during calibration of the chamber in a reference

328 gamma radiation field of air kerma from  $^{137}\text{Cs}$  source.  $U_R$  ensures 96% of ion  
329 collection efficiency in such reference field. The ambient dose equivalent is  
330 calculated as the product of absorbed dose and  $Q_4$ .

331 The detector was calibrated at CERF in 2016, and twice in monoenergetic neutron  
332 reference fields: at PTB (Golnik et al. 1997) and in 2018 at NPL. Before the REFLECT  
333 measurements, the chamber was calibrated at 990 V saturation voltage in the  
334 accredited (AP 070) Radiation Protection Measurements Laboratory (LPD, NCBJ)  
335 according to the Operational procedure M-1 (2017) with a  $^{137}\text{Cs}$  reference photon  
336 source and PuBe reference neutron source.

### 337 **2.2.5 TTM low-level neutron and gamma monitoring station**

338 The low-level neutron and gamma monitoring station registers photons and  
339 neutrons in separate 'pulse-height' windows (Pszona et al. 2014). The detector is  
340 based on an 8 inch Leake neutron area survey instrument (Leake et al. 2004). It uses  
341 a Centronic SP9  $^3\text{He}$  spherical proportional counter, surrounded by an inner  
342 polyethylene layer, a spherical shell of natural cadmium and a further outer  
343 polyethylene moderator. The cadmium shield is composed of two hemispherical  
344 shells, 0.91 mm thick, with 25 holes. The areas covered by the holes are the same in  
345 both hemispheres, except for a 12.5 mm hole used by the SP9 connector (Tagziria et  
346 al. 2004). Discrimination between photons and neutrons is based on the analysis of  
347 the pulse-height spectrum, defining the photon and neutron windows (Pszona et al.  
348 2014). The neutron response function is shown in Figure 1. The TTM station was  
349 calibrated with  $^{137}\text{Cs}$  and AmBe reference sources in the accredited (AP 070)

350 Radiation Protection Measurements Laboratory (LPD, NCBJ). Calibration factors  
351 used for the measurement were 0.55 nSv and 1.28 nSv per count for the photon and  
352 neutron windows, respectively.

### 353 **2.2.6 LB 6419**

354 The LB 6419 was designed by DESY and Berthold Technologies to measure the  
355 ambient dose equivalent  $H^*(10)$  of pulsed and continuous neutron and photon  
356 radiation at high-energy accelerators (Leuschner et al. 2017). The LB 6419  
357 comprises a cylindrical moderated rem-counter with a  $^3\text{He}$  proportional counter  
358 and a plastic scintillator.

359 The response to low-energy neutrons  $H_{LEN}$  is obtained from the proportional  
360 counter by counting the reaction products of the nuclear reaction  $^3\text{He}(n,p)\text{T}$ . Its  
361 moderator is made of polyethylene and contains neither any response-shaping  
362 absorbers like Cd or B nor converters like Pb. So it measures  $H_{LEN}$  with a calibration  
363 factor of 0.1 nSv per count. Its neutron response function is shown in Figure 1.

364 The response to high-energy neutrons  $H_{HEN}$  is obtained from the scintillator by  
365 collecting scintillator light above 20 MeV, a threshold where any response from  
366 electro-magnetic radiation such as  $\gamma$ ,  $e^\pm$ ,  $\mu^\pm$  can be discriminated. The response  
367 comes from the energy deposition of charged products from neutron scattering on  
368 hydrogen nuclei of the scintillator  $\text{H}(n,n)p$  and on carbon nuclei as  $\text{C}(n,p)$  and  
369  $\text{C}(n,\alpha)$ . As this response is based on the measurement of absorbed energy rather  
370 than counting it cannot be shown in Figure 1. The corresponding calibration factor  
371 was measured and validated at CERF in 2010, 2012 and 2017.

372 The total neutron dose  $H_N$  is obtained by summing up the doses of the two energy  
373 ranges  $H_{LEN}$  and  $H_{HEN}$ .

374 Electromagnetic radiation ( $H_{ELM}$ ) can be separated from the neutron response  
375 because it shows up in the energy spectrum as the so called “muon peak”. In the  
376 cylindrical scintillator with its dimension of 4.1 cm these minimum ionizing  
377 particles lose about 8 MeV ( $2 \text{ MeV}\cdot\text{cm}^{-2}$ ). The calibration is done by means of the  
378 Compton edges of radioactive sources such as  $^{137}\text{Cs}$  and  $^{60}\text{Co}$ .

379 Finally the total dose  $H_{TOT}$  is obtained by summing up the neutron dose  $H_N$  and the  
380 dose of the electro-magnetic radiation  $H_{ELM}$ .

### 381 **2.2.7. HammerHead HH**

382 The HammerHead HH (HHtec Association, Czech Republic) is a wide-range  
383 scintillation detector designed for high-precision  $H^*(10)$  measurements. The  
384 ambient dose equivalent rate range is from  $5 \text{ nSv}\cdot\text{h}^{-1}$  to  $10 \text{ mSv}\cdot\text{h}^{-1}$  for a photon  
385 energy range from 50 keV to 8 MeV. The typical type A uncertainty is 12 % for 1 s  
386 measuring interval ( $1\sigma$  and  $H^*(10)_{\text{terrestrial}} = 130 \text{ nSv}\cdot\text{h}^{-1}$ ). The HH meter is a portable  
387 detector with dimensions of  $\varnothing 80 \text{ mm} \times 340 \text{ mm}$  and mass of 1.6 kg.

388 The HammerHead HH has been designed in order that the measured value best  
389 corresponds to the physical definition of  $H^*(10)$  for photons and meet the strict  
390 criteria required by IEC 60846 for ambient dose equivalent meters. As detector, a  
391  $\text{CaF}_2:\text{Eu}$  scintillator with low atomic number is used. Its shape is close to a sphere of  
392 64 mm diameter, therefore the meter has excellent  $-135^\circ$  to  $+135^\circ$  angular  
393 response. The HH meter works in current mode, therefore the measurement is not

394 influenced by dead time. The unique time-energy analysis of the measured signals  
395 makes it possible to distinguish the contribution  $H^*(10)_L$  from particles with energy  
396 below 4 MeV and  $H^*(10)_H$  from particles with energy above 4 MeV. When measuring  
397 on the Earth's surface, the  $H^*(10)_L$  value represents the terrestrial component of the  
398 radiation field, whereas the  $H^*(10)_H$  value allows estimating the secondary cosmic  
399 ray component but without the influence of neutrons. The typical duration of a  
400 measurement is 9 hours when the instrument is connected to a tablet for data  
401 transfer.

402 HammerHead HH was calibrated in the accredited calibration laboratory at Czech  
403 Metrology Institute in Prague with a X-ray device (40 keV - 250 keV) and  $^{137}\text{Cs}$ ,  $^{60}\text{Co}$   
404 reference sources in terms of  $H^*(10)$ .

### 405 **2.2.8 FH 40 G-10 with FHZ-612B probe**

406 The FH 40 G-10 is a portable dose rate meter based on an internal energy filtered  
407 proportional counter. Without any external probe connected, this device is sensitive  
408 to photons only. During this experiment, an additional FHZ-612B Beta Gamma  
409 probe was connected. Even with the external FHZ-61B connected, the detector was  
410 used as gamma detector since the beta detector cap was installed. The  $H^*(10)$   
411 measuring ranges of the FH 40 G-10 and FHZ-612B are  $10 \text{ nSv}\cdot\text{h}^{-1} - 1 \text{ Sv}\cdot\text{h}^{-1}$  and  $100$   
412  $\text{ nSv}\cdot\text{h}^{-1} - 10 \text{ Sv}\cdot\text{h}^{-1}$ , respectively. The energy range is 20 keV – 4.4 MeV for the FH 40  
413 G-10 and 82 keV – 1.3 MeV for the FHZ-612B.

414 The instrument was calibrated in the accredited (AP 070) Radiation Protection  
415 Measurements Laboratory (LPD, NCBJ) with a  $^{137}\text{Cs}$  reference source in terms of  
416  $H^*(10)$  (ISO 1999).

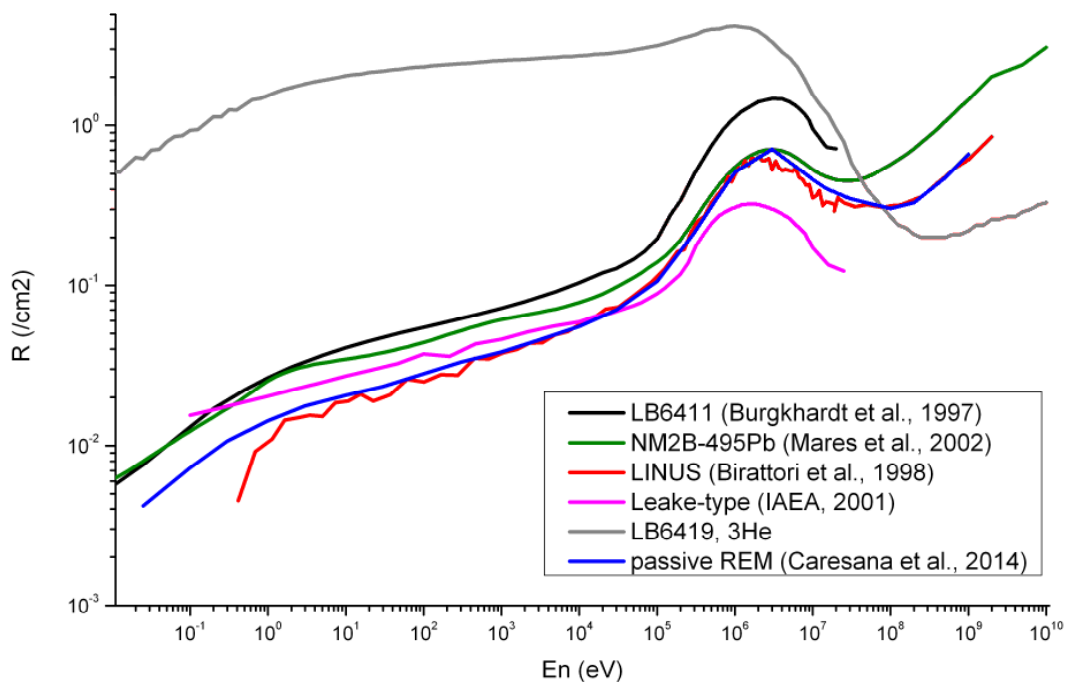
### 417 **2.3 Neutron rem-counters and dosimeters**

418 Several neutron dosimeters and rem-counters were used. The design of neutron  
419 rem counters is based mostly on the Andersson-Braun type (Andersson and Braun  
420 1963) or Leake type (Leake 1966) and they measure the neutron ambient dose  
421 equivalent,  $H^*(10)$ .

422 Neutron fluence response functions of the used neutron detectors are shown in  
423 Figure 1. Response functions are usually calculated with MC codes; several energy  
424 points are validated through measurements in monoenergetic neutron fields.

425





426

427 Figure1. Neutron response function  $R$  (counts per unit neutron fluence) of the neutron detectors

428 used in REFLECT.

429

### 430 2.3.1 NM2B-495Pb Rem Counter

431 The NM2B-495Pb Rem Counter is based on the conventional Andersson-Braun rem-

432 counter (NE Technology Ltd.) with a cylindrical  $\text{BF}_3$  proportional counter

433 surrounded by an inner polyethylene moderator, a boron-doped synthetic rubber

434 absorber, and an outer polyethylene moderator. To extend the detection range to

435 higher energy neutrons, a 1 cm thick lead shell is added around the boron rubber.

436 For this experimental flight, pulse height spectra were registered to control the

437 photon background and properly set up the region of interest (ROI). This procedure

438 enables an appropriate evaluation of the number of counts which are then

439 converted to  $H^*(10)$  through the calibration coefficient. The fluence response  
440 function from thermal to 10 GeV neutrons was calculated by means of different  
441 Monte Carlo codes (Mares et al. 2002). The rem counter calibration was performed  
442 using a 185 GBq  $^{241}\text{Am}$ -Be neutron source following the ISO recommendations (ISO  
443 2001). The rem counter was also used and calibrated in 100 and 300 MeV quasi-  
444 mono-energetic neutron fields at RCNP in Osaka (Mares et al. 2017) and at CERF.  
445 The response function of the detector is shown in Figure 1.

### 446 **2.3.2 LINUS**

447 The LINUS (Birattari et al. 1990; Birattari et al. 1992; Birattari et al. 1993; Birattari  
448 et al. 1998) is the original extended-range rem counter. It consists of a  $^3\text{He}$   
449 proportional counter embedded in a spherical polyethylene moderator, which  
450 incorporates a boron-doped rubber absorber and a 1 cm thick lead shell so that its  
451 response function extends up to several hundred MeV. The signal is treated with a  
452 standard counting chain (pre-amplifier, amplifier, single channel analyzer and  
453 counter) and the TTL output is analyzed by a custom LabVIEW interface. The  
454 response function of the detector is shown in Figure 1. Neutron detectors are  
455 sensitive to some extent to gamma rays, which can transfer energy to the system  
456 through Compton scattering in the walls or fill gas. The gamma rejection for the  
457 LINUS is obtained by setting a discriminator below the low energy neutron signal to  
458 reject counts due to gamma rays and electronic noise. The threshold was  
459 determined by analyzing the pulse height spectrum of the  $^3\text{He}$  counter.

460 The LINUS was calibrated with an AmBe source (Dinar et al. 2017) in the CERN  
461 CALibration LABoratory (CALLAB) (Pozzi et al. 2015). The calibration provided a  
462 calibration factor of 0.89 nSv per count with an overall uncertainty of 3.2% at one  
463 sigma.

### 464 **2.3.3 LB6411**

465 The LB 6411 neutron probe (Burgkhardt et al. 1997), connected to the universal  
466 monitor LB 123, is designed for measurement of neutron ambient dose equivalent  
467  $H^*(10)$  in accordance with ICRP 60 (BERTHOLD n.d.). The LB 6411 consists of a  
468 cylindrical  $^3\text{He}$  proportional counter centered in a polyethylene sphere with  
469 diameter 25 cm. The neutron energy range is from thermal to 20 MeV. The spectrum  
470 from the bare  $^{252}\text{Cf}$  neutron source has been used as the calibration spectrum. The  
471 numerical calibration factor is 0.32 nSv per count (Burgkhardt et al. 1997). The  
472 response function over the whole energy range was calculated with MCNP. For  
473 several energies the results were crosschecked with monoenergetic neutron  
474 measurements. The response function of the detector is shown in Figure 1. The  
475 response to gamma radiation is approx.  $10^{-3}$  counts per nSv, which means a  
476 discrimination factor of  $3 \times 10^3$ .

### 477 **2.3.4 Passive REM counter**

478 The neutron contribution to  $H^*(10)$  was also measured with a system consisting of  
479 two CR-39 detectors  $3 \times 4 \text{ cm}^2$  in dimension coupled to a  $^{10}\text{B}$  enriched converter,  
480 positioned inside a sphere made with polyethylene, lead, and cadmium. The  $^{10}\text{B}$  is

481 contained in boron carbide ( $B_4C$ ) deposited on an aluminum plate. The thickness of  
482 the boron carbide is about  $10\ \mu\text{m}$ . The instrument is an extended range rem counter  
483 and the response function is shown in Figure 1. The full description of the  
484 instrument is in (Caresana et al. 2014) while previous experience in measuring  
485 onboard aircraft is described in (Federico et al. 2015).

486 The plug, hosting the two CR39 detectors assembled with the boron converter, was  
487 removed from the moderating sphere during the shipment, inserted into the sphere  
488 immediately before take-off and removed immediately after landing.

489 A check of the calibration coefficient was performed at CERF in August 2017 and  
490 resulted in  $10.6\ \text{cm}^{-2}\cdot\mu\text{Sv}^{-1}$  with an uncertainty equal to 14% ( $k=1$ ). The sensitivity is  
491 about 3 times higher than the one reported in (Caresana et al. 2014). This is because  
492 the boron converter used in the above cited work is the Enriched Converter Screen  
493 BE10 by Dosirad (France) whose thickness is about  $100\ \mu\text{m}$ . Using this converter,  
494 only a layer of about  $10\ \mu\text{m}$  directly facing the CR39 detector contributes to the  
495 signal, while interactions occurring at longer distance generate reaction products  
496 that are self-absorbed in the converter. The effect is a depression of the thermal  
497 neutron flux, resulting in a reduced sensitivity.

### 498 **2.3.5 Electronic neutron dosimeter ELDO**

499 The ELDO is an individual dosimeter developed at the Helmholtz Zentrum München  
500 (HMGU), sensitive to neutrons from thermal energies up to about 200 MeV  
501 (Wielunski et al. 2004). It is a small ( $160\ \text{g}$ ,  $115\times 60\times 16\ \text{mm}^3$ ) personal dosimeter  
502 with a dose measurement range between  $1\ \mu\text{Sv}$  and  $10\ \text{Sv}$ . Its operational lifetime is

503 about 400 hours. It consists of four Si PIN-diodes with LiF or polyethylene (PE)  
504 converters encapsulated in lead or cadmium. The combination of diodes and  
505 converter enables separate measurements of neutrons with one fast-sensor (PE)  
506 operating in the 1–200 MeV neutron energy range, two delta-sensors (LiF)  
507 functioning between 50 keV and 2 MeV, and one albedo-sensor (LiF) sensitive to  
508 low-energy neutrons (<50 keV). Each sensor is sensitive to a certain neutron energy  
509 range and has its own calibration factor. The measured dose and dose rate in terms  
510 of the personal dose equivalent,  $H_p(10)$  (an operational quantity for individual  
511 monitoring for the assessment of effective dose), are also shown on its LCD display.  
512 Calibration of the ELDO was done at PTB Braunschweig, Germany, in mono-  
513 energetic neutron fields with energies between 138 keV and 14.8 MeV (Bergmeier  
514 et al. 2013). Additionally, the ELDO was also tested in the reference field of CERN-  
515 CERF providing high-energy fields similar to that of secondary cosmic rays at flight  
516 altitudes (Wielunski et al., 2018) and at the Environmental Research Station "UFS  
517 Schneefernerhaus" (2,650 m above sea level) close to the summit of the Zugspitze  
518 Mountain, Germany (Volnhals, 2012). In these experiments, an excess in the  
519 measured counts was observed which are due to protons. Although the fluence of  
520 protons is only 6% of the neutron fluence, protons caused around 12% of the  
521 measured counts. The sensor response to protons and muons could be calculated  
522 with GEANT4 calculations (Volnhals 2012) which support the observation.

## 523 **2.4 Calculations**

524 Ambient dose equivalent rates for different particles can be calculated using various  
525 models; the overview of codes assessing radiation exposure of aircraft crew is given  
526 in (Bottollier-Depois et al. 2012). All these codes provide calculations for the GCR  
527 induced radiation field in aircraft flight altitudes agreeing within 20 % with  
528 reference measurements (Bottollier-Depois et al. 2012). In this publication, the  
529 EPCARD.Net code (Mares et al. 2009) was used for comparison.

### 530 **2.4.1 EPCARD.Net**

531 The European Program package for the Calculation of Aviation Route Doses  
532 (EPCARD) is a widely used program for estimating the exposure of aircraft crew.  
533 This code was developed at the Helmholtz Zentrum München (Schraube et al.  
534 2002b) and further improved in a new object-oriented code EPCARD.Net (Mares et  
535 al. 2009). In 2010, EPCARD.Net ver. 5.4.3 Professional was approved for official use  
536 for assessing radiation exposure from secondary cosmic radiation at aviation  
537 altitudes by the German Aviation Authority (LBA) and the National Metrology  
538 Institute, Physikalisch-Technische Bundesanstalt (PTB).

539 EPCARD.net is based on the results of extensive FLUKA Monte Carlo (Ferrari et al.  
540 2005; Böhlen et al. 2014) calculations of particle energy spectra of neutrons,  
541 protons, photons, electrons and positrons, muons, and pions at various depths in the  
542 atmosphere down to sea level for all possible values of solar activity and  
543 geomagnetic shielding conditions (Roesler et al. 2002). The primary particle spectra

544 used in the FLUKA calculations as well as the modulation potential describing solar  
545 activity were based on the model of Badhwar and O'Neill (Badhwar 1997; Badhwar  
546 et al. 2000).

547 To determine the dose rates at specific locations in the atmosphere during a flight,  
548 the cut-off rigidity, the solar deceleration potential and the barometric altitude are  
549 calculated to quantify geomagnetic shielding, solar activity and atmospheric  
550 shielding. The EPCARD.Net parameter database includes energy-averaged dose  
551 conversion coefficients, calculated by folding each single-particle fluence spectrum  
552 with the appropriate dose conversion function (Mares et al. 2004; Mares and  
553 Leuthold 2007), which depends on barometric altitude, cut-off rigidity, and solar  
554 activity, since the shape of the particle energy spectra also depends on these  
555 parameters. Ambient dose equivalent,  $H^*(10)$ , and effective dose,  $E$ , are calculated  
556 separately for each particle, i.e. the dose contributions from neutrons, protons,  
557 photons, electrons, muons, and pions are assessed individually.

558 More general information about EPCARD is available on the web site (EPCARD  
559 2020), where a simplified on-line version of the EPCARD calculator for public use  
560 can also be found.

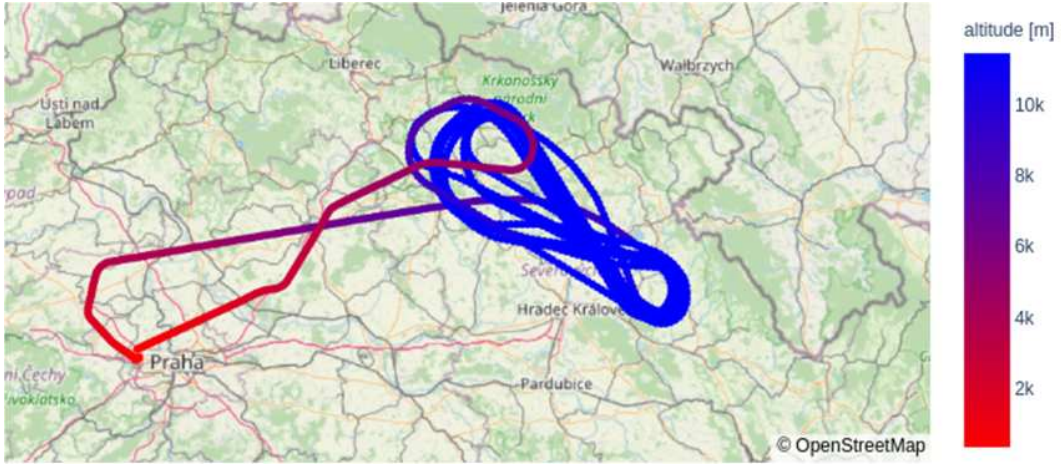
## 561 **3. Experiment**

### 562 **3.1 Flight**

563 The radiation detectors were exposed aboard an Embraer Legacy 600 aircraft  
564 operated by ABS Jets. The aircraft, together with 250 kg of equipment and eight  
565 scientific staff, flew from Vaclav Havel Airport in Prague (50.1°N, 14.2° E) to the  
566 FL390 flight level, on the 29<sup>th</sup> November 2017. The flight took off at the airport at  
567 13:06 UTC and reached stable flight conditions (barometric altitude  $11871 \pm 8$  m,  
568 range from 11853 to 11893), latitude  $50.41 \pm 0.14$  °N (range from 50.18 to 50.58),  
569 longitude  $15.80 \pm 0.27$  ° E (range from 15.26 to 16.24) at 13:38 UTC. At this level,  
570 the aircraft circled over the northern part of the Czech Republic (the area is a  
571 reserved airspace that is commonly used for operating test flights) for 90 minutes  
572 and landed back at Prague Airport at 15:34 UTC. Navigation data (barometric  
573 altitude, latitude, and longitude) were taken from the aircraft record and GPS. The  
574 flight route and flight profile are shown in Figure 2 and Figure 3, respectively. The  
575 space weather conditions were stable during the whole flight and no short-term  
576 solar activity affected the results. Space weather situation can be assessed e.g. by  
577 neutron monitors (nmdb.eu). During the flight the variation in the count rates of the  
578 neutron monitor at Lomnický štít (the nearest neutron monitor) was below 0.5%,  
579 which indicates stable space weather conditions.

580

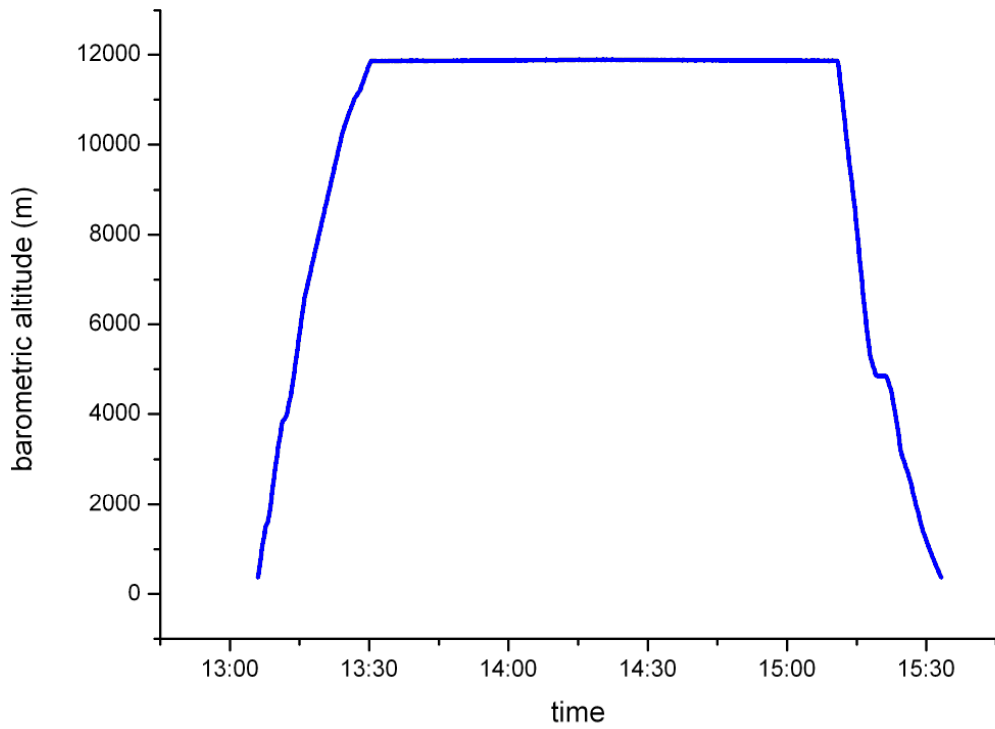




581

582 Figure 2. Flight route

583



584

585 Figure 3. Flight profile

586

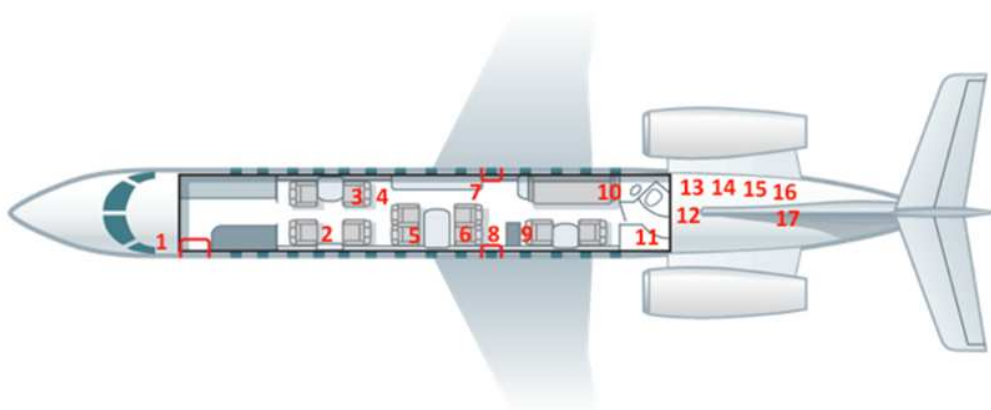
### 587 **3.2 Location of detectors inside the aircraft**

588 The detectors were placed at various locations inside the aircraft (Fig. 4).

589 Equipment that needed power and manual control were installed on or behind the  
590 seats. Smaller devices like Liulins or Airdos were distributed in various locations  
591 inside the aircraft. The rest of the instruments were stored in the baggage  
592 compartment.

593 Two fuel tanks are located in the wings, two fuel tanks are in the bottom part of the  
594 body and two fuel tanks are in the rear part of the plane, behind the baggage  
595 compartment. Because the flight was quite short (about 2.5 hours), only the tanks in  
596 the wings were filled with fuel. The total amount of fuel before take-off was 5482  
597 liters (4380 kg), 2530 kg were burned during the flight.

598



599

600 Figure 4. Placement of detectors inside the aircraft: 1 – Liulin MDU10; 2 – Sievert, 3 – Timepix; 4 –

601 HammerHead; 5 – LB6411; 6 – LB6419; 7 – Liulin MDU7; 8 – NM2B-495Pb Rem Counter; 9 – Liulin

602 MDU14; 10 – LINUS; 11 – AIRDOS; 12 – REM-2 recombination chamber; 13 – TTM + FH 40 G-10 +  
603 FHZ-612B, 14 – ELDO; 15 – TEPC Hawk (IRSN); 16 – passive REM counter + Liulin MDU1; 17 – TEPC  
604 Hawk (SL).

## 605 **4. Results and Discussion**

606 Not all devices operated during the whole flight, part of the instruments were  
607 started when stable flight conditions were reached. To compare the results obtained  
608 with the active detectors, we consider only data acquired at a constant flight  
609 altitude.

610 Values of  $\dot{H}^*(10)$  for various particles calculated with EPCARD are shown in Table 2.  
611 It should be noted that the calculations were done in free air, whereas the  
612 instruments measured inside the aircraft and therefore small differences can be  
613 expected due to shielding effects (Ferrari et al. 2004). As can be seen from Table 2,  
614 the most important contribution to  $\dot{H}^*(10)$  is from neutrons (57% of the total  
615  $\dot{H}^*(10)$ ), followed by electrons (20%) and protons (15%). The uncertainty on the  
616 calculated values is estimated to be less than 20%, based on (Bottollier-Depois et al.  
617 2012) who compared various codes used for assessing radiation exposure of aircraft  
618 crew due to GCR with the conclusion that the agreement between the codes was  
619 better than 20% from the median. The codes have also been previously validated by  
620 measurements with an agreement better than  $\pm 20\%$  (Lindborg, 2004).

621 Table 2. Calculated values of  $H^*(10)$  rate during the REFLECT at FL390 for different particles using  
622 EPCARD.Net ver. 5.5.0

	neutrons	photons	protons	electrons	muons	total
$\dot{H}^*(10)$ [ $\mu\text{Sv/h}$ ]	$3.8 \pm 0.8$	$0.3 \pm 0.1$	$1.0 \pm 0.2$	$1.3 \pm 0.3$	$0.2 \pm 0.0$	$6.6 \pm 1.3$

623

624 Table 3. Dose-equivalent rate measured with the various detectors. Uncertainties given as combined  
625 uncertainties with  $k=1$  and with the contribution from measurement statistics in parenthesis.

	Instrument	$\dot{H}^*(10)$ [ $\mu\text{Sv/h}$ ]	$\dot{H}^*(10)_{\text{Low-LET OR}}$ $\dot{H}^*(10)_{\gamma+e}$ [ $\mu\text{Sv/h}$ ]	$\dot{H}^*(10)_{\text{High-LET OR}}$ $\dot{H}^*(10)_n$ [ $\mu\text{Sv/h}$ ]
routinely used instruments	TEPC HAWK (SL)	$7.1 \pm 0.5$ (0.3)	$3.3 \pm 0.1$ (<0.1)	$3.8 \pm 0.6$ (0.3)
	TEPC HAWK (IRSN)	$7.9 \pm 0.6$ (0.3)	$3.5 \pm 0.1$ (<0.1)	$4.4 \pm 0.5$ (0.3)
	Sievert (SSM)	$7.4 \pm 0.6$ (0.4)	$2.9 \pm 0.2$ (0.1)	$4.5 \pm 0.6$ (0.6)
	Liulin MDU7 (NPI)	$7.1 \pm 1.1$	$3.1 \pm 0.5$	$4.0 \pm 0.7$
Neutron rem-counters	NM2B-495Pb (HMGU)			$3.7 \pm 0.4$
	LINUS (CERN)			$3.9 \pm 0.1$
	LB6411 (NPI)			$2.1 \pm 0.4$
	Passive REM counter (Polimi)			$7.5 \pm 2.5$

Others instruments	LB6419 (DESY)	9.1 ± 1.8	1.7 ± 0.4	7.4 ± 1.5
	REM-2 (NCBJ)	10.1 ± 8.9	–	–
	TTM (NCBJ)	4.8 ± 0.8	2.1 ± 0.4	2.8 ± 0.5
	ELDO (HMGU)			4.4 ± 0.9*
	HammerHead HH (HHtec Association)		2.7 ± 0.2	
	FH 40 G-10 (NCB)		3.2 ± 0.5	
	FHZ-612B (NCB)		4.7 ± 0.8	

626 \*  $H_p(10)$

627

628 Table 3 lists the measurement results (for all instruments except Si detectors  
629 measuring only  $D_{Si}$ ). The results are grouped in a low-LET component that  
630 comprises the contribution of low ionizing radiation (photons, electrons, muons,  
631 protons, pions) and in a high-LET component representing mostly contributions of  
632 neutrons, stopping protons and higher Z ionizing particles. Even if the neutron  
633 contribution can extend below  $10 \text{ keV}\cdot\mu\text{m}^{-1}$  and for low energy photons it can be  
634 above  $10 \text{ keV}\cdot\mu\text{m}^{-1}$ , the previous assumption (high-LET assimilated to neutrons) is  
635 usually made when comparing results of various instruments and calculations. One  
636 should also note that most detectors designed for low-LET measurements exhibit a  
637 response to neutrons that is usually unknown, especially for high-energy neutrons.

638 In the table, the results provided are not corrected for this unknown neutron  
639 response. Uncertainties are given as combined uncertainties with contributions  
640 from calibration and measurement statistics and presented with coverage factor  
641  $k=1$ . For TEPC, the statistical uncertainty is given in parenthesis. The first group  
642 includes instruments routinely used onboard aircraft, measuring both low and high-  
643 LET components, the second group includes neutron rem-counters and the third  
644 group includes the remaining instruments.

645 When comparing the results, an agreement of  $\pm 20\%$  at a 95 % confidence level is  
646 considered satisfactory. The recommendation on acceptable uncertainties in  
647 radiation protection is given in (ICRP 1997) where it is stated: *"...overall uncertainty*  
648 *at the 95 % confidence level in the estimation of effective dose around the relevant*  
649 *dose limit may well be a factor of 1.5 in either direction for photons and may be*  
650 *substantially greater for neutrons of uncertain energy and for electrons. Greater*  
651 *uncertainties are also inevitable at low levels of effective dose for all qualities of*  
652 *radiation."*

653 The measurements of the various TEPCs (HAWK, Sievert) agreed well with each  
654 other, with Liulin, and with the EPCARD calculations, as it was during a previous  
655 flight comparison (CAATER) (Lillök et al 2007). The differences in low- and high-  
656 LET components between the Sievert instrument and the other TEPCs were likely  
657 due to the fact that the Sievert instrument distinguished  $H^*(10)$  contributions from  
658 photons and neutrons rather than in terms of a low and high-LET threshold. The  
659 differences could have been also due to different locations of the TEPCs (Hawks in  
660 the baggage compartment, Sievert in the front of the plane). During the

661 approximately 90 minutes of the cruise, the TEPC experienced statistically low  
662 number of high LET events, which resulted in higher uncertainties for this  
663 component.

664 The LB6419 and REM-2 measured larger values of total  $H^*(10)$  than the TEPCs. The  
665 REM-2 is very sensitive to vibrations, which probably led to the very high value of  
666 the uncertainty. In principle, there are recombination methods for separating the  
667 dose according to LET, but in this flight, the measurement method was simplified  
668 because of the short time and difficult conditions. Provision of the values for the low  
669 and high-LET components would be helpful to better interpret the data. Improved  
670 calibration is needed to make use of REM-2 for routine aircraft dosimetry.

671 The LB6419 also measured a higher value of total  $H^*(10)$  than the TEPCs. In  
672 addition, Table 3 shows a small measured value of the low-LET radiation  
673 contribution and an increased contribution of the high-LET component as compared  
674 to the TEPCs. There is a need to check the separation method of the two components  
675 and the calibration, since the total ambient dose equivalent was also too high  
676 compared to the TEPC.

677 The  $H^*(10)$  values measured with the TTM monitoring station, especially the  
678 neutron component, was lower than the other instruments. Since only polyethylene  
679 was used as a moderator (no lead or other high atomic number material was  
680 included in the shell), the neutron energy range of this instrument was limited to  
681 20 MeV.

682 Except for the LB6411 and TTM, the instruments measuring only the neutron  
683 component provided results comparable to the TEPC results and with the EPCARD  
684 calculations.

685 According to the EPCARD model, neutrons contributed for more than 50% to the  
686 total  $H^*(10)$ ; neutrons can reach energies up to several hundreds of MeV  
687 (Pazianotto et al. 2017). Conventional neutron REM counters have a detection range  
688 usually limited to about 15–20 MeV, their response dropping sharply at higher  
689 energies. To extend the range to higher energies (up to several hundreds of MeV), a  
690 shell of high-Z material (like tungsten or lead) is usually added to the PE moderator.

691 As expected, due to their response functions (Fig. 1), the LB6411 and TTM measured  
692 lower values. This was in agreement also with measurements by Yasuda et al.  
693 (2018), who investigated neutron doses during long-haul flights with two neutron  
694 monitors and compared their results with JISCARD EX calculations. They found that  
695 the relative contribution to  $H^*(10)$  of neutrons with energies above 15 MeV could  
696 exceed 50%.

697 The passive REM counter only provided an integral value over the whole flight (the  
698 detector was installed inside the moderator just before take-off and removed after  
699 landing). The total measured  $H^*(10)$  was  $15 \pm 5 \mu\text{Sv}$ ; the dose rate at flight level  
700 FL390 can be assessed assuming a taxi time of about 2 h and neglecting the small  
701 contribution arising from the 0.5 h spent to reach the flight altitude and get back to  
702 ground.



703 A possible explanation of the large measured  $H^*(10)$  value is that the passive REM  
704 counter – because of a misunderstanding with the shipping company – reached  
705 Prague by airmail. Of course, the plug with the detector was not in the measuring  
706 position, thus insensitive to fast neutrons. However, a small contribution from  
707 thermal neutrons cannot be excluded.

708 Amongst the instruments measuring only the low-LET component, only the  
709 HammerHead HH and the FH 40 G-10 obtained reasonable results. The results from  
710 the other instruments disagreed with both the EPCARD calculation and the TEPC  
711 measurements. It is difficult to compare the results because some photon detectors  
712 are not only sensitive to photons and electrons, but also to protons, muons, pions  
713 and to neutrons in some extent. For these instruments, their response to  
714 components of the field other than that intended to be measured is not always  
715 known.

716 Table 4 summarises the results of the silicon detectors (in terms of absorbed dose).  
717 For these instruments, the dose in silicon was converted to dose in water using a  
718 dose conversion factor of 1.23 (Ploc 2009). In a previous intercomparison flight  
719 with Liulin MDUs (Meier et al. 2016) it was found that there could be differences in  
720 the mass of the sensitive volume of the detectors (Si sensor size) considered in the  
721 calculation of absorbed dose. In this experiment, we calculated the absorbed dose in  
722 silicon using the same Si sensor mass (0.16597 g (Meier et al. 2016)) for all Liulin  
723 units and for Airdos.

724 Table 4: Results (absorbed dose rate in silicon and in water) of the measurements with the Silicon  
725 detectors

Instrument	MDU 7 Liulin (NPI)	MDU 10 Liulin (QST)	MDU14 Liulin (QST)	MDU1 Liulin (DLR)	Airdos T4 (NPI)	Minipix (NPI)
$D_{Si}$ ( $\mu\text{Gy}/\text{h}$ )	$2.8 \pm 0.2$	$2.1 \pm 0.3$	$2.7 \pm 0.5$	$2.0 \pm 0.3$	$1.8 \pm 0.2$	$1.9 \pm 0.3$
$D_{H_2O}$ ( $\mu\text{Gy}/\text{h}$ )*	$3.4 \pm 0.3$	$2.6 \pm 0.4$	$3.3 \pm 0.6$	$2.5 \pm 0.3$	$2.2 \pm 0.2$	$2.3 \pm 0.3$

726 \* The dose  $D_{H_2O}$  was calculated from  $D_{Si}$  using a conversion factor 1.23

727 Although both Liulin and Airdos have similar sensitive volumes (mass, area,  
728 thickness), they have different properties such as energy range of deposited energy  
729 and width of the channel. Airdos has channel width of 49.4 keV whereas Liulin's is  
730 81.4 keV. The energy range of Airdos (up to 12 MeV) is smaller than Liulin's (up to  
731 24 MeV), in order to provide more detailed information on the lower part of the  
732 energy spectrum. The significant part of the deposited energy when measuring  
733 onboard aircraft comes from events depositing energy in the first several channels  
734 (for Liulins, 65–83% of the absorbed dose was due to events with deposited energy  
735 below 1 MeV, only 2–6% was due to events with deposited energy above 10 MeV).  
736 To compare Airdos with Liulin, we considered only events within the energy range  
737 of Airdos for the calculation of absorbed dose for Liulin. For MDU 7,  $D_{Si}$  was 2.6  
738  $\mu\text{Gy}\cdot\text{h}^{-1}$ , for MDU 10  $D_{Si}$  was 1.9  $\mu\text{Gy}\cdot\text{h}^{-1}$ , and for MDU 14  $D_{Si}$  was 2.6  $\mu\text{Gy}\cdot\text{h}^{-1}$ , to be  
739 compared with 1.8  $\mu\text{Gy}\cdot\text{h}^{-1}$  measured by Airdos.

740 Some differences in the results could be due to the different shielding configurations  
741 (for example, the aircraft fuel acts as a good neutron moderator) around the  
742 locations in which the devices were installed (Fig. 4). The DLR Liulin was in the

743 baggage compartment, whereas the NPI Liulin and the Airdos were in the central  
744 part of the aircraft or in the crew cabin. For Embraer Legacy 600, the baggage  
745 compartment is located in the rear part of the aircraft, between the engines (Fig. 4).  
746 Therefore, the baggage compartment, loaded with several larger instruments and  
747 suitcases, is supposed to be more shielded than other areas of the aircraft. As was  
748 shown in Ferrari et al. (2004), the shielding provided by the aircraft structure  
749 (wings, engines, passengers, fuel) can cause a notable reduction in  $E$  or  $H^*(10)$  for  
750 most components of cosmic radiation. Differences in ambient dose equivalent for  
751 various places inside the aircraft can be up to about 20% (Ferrari et al. 2004;  
752 Battistoni et al. 2005; Kubancak et al. 2014). Nevertheless, the differences between  
753 individual Liulin-type detectors seem to be too large to be explained only by  
754 different locations in the aircraft. There appears to be some systematic differences.  
755 The reason of these differences should be further investigated in comparison on  
756 ground. Even for the detectors using the same Si diode, several factors can influence  
757 the results, e.g. energy calibration, choice of the noise threshold, channel width  
758 (Kakona et al. 2019).

759 Although the Minipix has larger energy range (from 5 keV), it showed a lower  
760 absorbed dose than the MDUs. This is difficult to explain since the Minipix was  
761 calibrated against the TEPC and showed a very good agreement to the TEPC Hawk  
762 low-LET part (Ploc 2009). However, it should be mentioned that the setting of  
763 Timepix, especially bias, could have been different for different experiments, which  
764 might have caused some discrepancies. In this flight, the bias was set to 30 V. It is  
765 important to use the same conditions (parameters) for all instruments.

766 Taking the MDU7 (MDU 7 has been calibrated at CERF) energy deposition spectra  
767 and performing the calibration according to (Ploc 2009), the total  $H^*(10)$  rate  
768 arrived at  $7.1 \mu\text{Sv}\cdot\text{h}^{-1}$ , which agreed well with the TEPC results.

## 769 **5. Conclusions**

770 For the instruments with the potential to be used onboard aircraft, appropriate  
771 calibration and determination of calibration/correction factor is crucial. A good  
772 measurement of the atmospheric ambient dose equivalent requires that the  
773 instrument response for all particles and energies is properly taken into account. As  
774 there is no traceable reference field for the total radiation field in the atmosphere,  
775 comparison of instruments onboard aircraft is necessary.

776 Various radiation detector systems were compared onboard aircraft under stable  
777 flight conditions during the REFLECT measurement campaign – the largest  
778 comparison of this type ever performed. As expected, the dosimeters routinely used  
779 for aircraft dosimetry (TEPC, Liulin) worked adequately, the results agreed with  
780 each other as well as with the EPCARD computer model calculations.

781 Since high-energy neutrons contribute significantly to  $H^*(10)$ , conventional neutron  
782 rem counters (with energy range limited to about 20 MeV) underestimated neutron  
783  $H^*(10)$  with standard calibration procedures using neutron sources such as AmBe  
784 or Cf-252. Extended-range neutron rem counters (NM2B-495Pb, LINUS) provided  
785 results comparable to  $H^*(10)_n$  determined with routinely used instruments and the  
786 EPCARD calculation.

787 The reading of some instruments (LB6419, FHZ-612B) was higher than expected  
788 from the assumption that the detector is only sensitive to a specific component of  
789 the radiation field. However, it should be noted that these instruments are not  
790 primarily intended for use onboard aircraft in a very complex mixed radiation field.  
791 Their response to the various components of the cosmic radiation field and the  
792 energy dependence still need to be fully characterised.

793 The REFLECT campaign enabled measurement in uniform, well-defined conditions  
794 onboard an aircraft and comparison of new instruments with those routinely used.  
795 Although the response of some instruments, not primarily intended for the use in a  
796 very complex mixed radiation field such as onboard aircraft, was different than as  
797 somehow expected, more experiments are necessary too finally declare them as not  
798 suitable. This campaign consisted only of one flight with one set of parameters  
799 (vertical cut-off rigidity, altitude, phase of solar cycle) and it was relatively short.  
800 For further testing of instruments showing some potential to be used for routine  
801 dosimetry onboard aircraft, additional flights at different geomagnetic cut-offs and  
802 altitudes as well as comparison at appropriate reference fields (CERF) are needed.  
803 The REFLECT provided valuable experience enabling to discuss some issues  
804 connected with the use of the dosimeters in such complex radiation field and it also  
805 provided feedback for designing future in-flight campaigns.

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