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REFLECT – Research Flight of EURADOS and CRREAT: Intercomparison of various radiation dosimeters onboard aircraft

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1 **REFLECT – Research Flight of EURADOS and CRREAT**
2 **Intercomparison of various radiation dosimeters onboard**
3 **aircraft**

4 Running head: Dosimeters' intercomparison onboard aircraft

5

6 Iva Ambrožová*¹, Peter Beck², Eric R. Benton^{1,3}, Robert Billnert⁴, Jean-Francois
7 Bottollier-Depois⁵, Marco Caresana⁶, Nesrine Dinar⁷, Szymon Domański⁸, Michał A.
8 Gryziński⁸, Martin Kákona^{1,9}, Antonín Kolros^{10,11}, Pavel Krist¹, Michał Kuć⁸, Dagmar
9 Kyselová^{1,9}, Marcin Latocha², Albrecht Leuschner¹², Jan Lillhök⁴, Maciej Maciak⁸,
10 Vladimír Mareš¹³, Łukasz Murawski⁸, Fabio Pozzi⁷, Guenther Reitz^{1,14}, Kai
11 Schennetten¹⁴, Marco Silari⁷, Jakub Šlegl^{1,9}, Marek Sommer^{1,9}, Václav Štěpán¹,
12 Francois Trompier⁵, Christoph Tscherne², Yukio Uchihori¹⁵, Arturo Vargas¹⁶,
13 Ladislav Viererbl^{10,11}, Marek Wielunski¹³, Mie Wising⁴, Gabriele Zorloni⁶, Ondřej
14 Ploc¹

15

16 ¹ Nuclear Physics Institute CAS, Czech Republic

17 ² Seibersdorf Labor GmbH, Austria

18 ³ Oklahoma State University, USA

19 ⁴ Swedish Radiation Safety Authority, Sweden

20 ⁵ Institute for Radiological Protection and Nuclear Safety, France

- 21 ⁶ Politecnico di Milano, Italy
- 22 ⁷ CERN, Switzerland
- 23 ⁸ National Centre for Nuclear Research, Poland
- 24 ⁹ Faculty of Nuclear Sciences and Physical Engineering CTU, Prague, Czech Republic
- 25 ¹⁰ Research Centre Rez, Czech Republic
- 26 ¹¹ HHtec Association, Czech Republic
- 27 ¹² Deutsches Elektronen-Synchrotron, Germany
- 28 ¹³ Helmholtz Zentrum München, Germany
- 29 ¹⁴ German Aerospace Center, Germany
- 30 ¹⁵ National Institute of Radiological Sciences / QST, Japan
- 31 ¹⁶ Technical University of Catalonia, Spain
- 32
- 33 * ambrozova@ujf.cas.cz, Nuclear Physics Institute CAS, Department of Radiation
- 34 Dosimetry, Na Truhlarce 39/64, Prague, 18000, Czech Republic
- 35

36 **Abstract**

37 Aircraft crew are one of the groups of radiation workers which receive the highest
38 annual exposure to ionizing radiation. Validation of computer codes used routinely
39 for calculation of the exposure due to cosmic radiation and the observation of
40 nonpredictable changes in the level of the exposure due to solar energetic particles,
41 requires continuous measurements onboard aircraft. Appropriate calibration of
42 suitable instruments is crucial, however, for the very complex atmospheric radiation
43 field there is no single reference field covering all particles and energies involved.
44 Further intercomparisons of measurements of different instruments under real
45 flight conditions are therefore indispensable.

46 In November 2017, the REFLECT (**RE**search **FL**ight of **EUR**ADOS and **CR**REAT) was
47 carried out. With a payload comprising more than 20 different instruments,
48 REFLECT represents the largest campaign of this type ever performed. The
49 instruments flown included those already proven for routine dosimetry onboard
50 aircraft such as the Liulin Si-diode spectrometer and tissue equivalent proportional
51 counters, as well as newly developed detectors and instruments with the potential
52 to be used for onboard aircraft measurements in the future. This flight enabled
53 acquisition of dosimetric data under well-defined conditions onboard aircraft and
54 comparison of new instruments with those routinely used.

55 As expected, dosimeters routinely used for aircraft dosimetry (such as a tissue
56 equivalent proportional counter (TEPC) or a silicon detector device like Liulin
57 agreed reasonable with each other as well as with model calculations. Conventional
58 neutron rem counters underestimated neutron ambient dose equivalent, while

59 extended-range neutron rem counters provided results comparable to routinely
60 used instruments. Although the response of some instruments, not primarily
61 intended for the use in a very complex mixed radiation field such as onboard
62 aircraft, was as somehow expected to be different, the prove of their suitability was
63 one of the objectives of the REFLECT. This campaign comprised a single short flight.
64 For further testing of instruments, additional flights as well as comparison at
65 appropriate reference fields are envisaged. The REFLECT provided valuable
66 experience and feedback for future experiments.

67

68 **Keywords**

69 cosmic radiation; aircraft; dosimeter; intercomparison; research flight

70

71

72 **1. Introduction**

73 Aircraft crew and airline passengers are exposed to elevated dose rates due to
74 cosmic radiation onboard aircraft; aircraft crew is considered as a group of workers
75 receiving one of the highest annual effective doses (ICRP, 1991; ICRP, 2007; ICRP,
76 2016; IAEA 2003). Radiation protection for aircraft crew has been regulated in the
77 European Union since 1996 by the EU-Directive 29/96/EURATOM (EURATOM
78 1996). Since then, this directive was updated with the EU-Directive
79 2013/59/EURATOM (EURATOM 2013). The EU member states were obliged to
80 comply with the new regulations by updating their national legislations by February
81 2018. Annual personal doses from galactic cosmic radiation (GCR) to aircraft crew
82 members are routinely calculated by various computer codes that are validated
83 preferably by measurements but also by code intercomparisons. Ongoing
84 validations of such codes need in-flight measurements with appropriately calibrated
85 instruments.

86 An assessment of aircraft crew radiation exposure is a complex task. Radiation field
87 at civil flight altitudes is formed by interactions of mainly GCR (and sporadically
88 solar energetic particles – SEP) with the atoms of the atmosphere of the Earth. All
89 types of particles and electromagnetic component such as protons, muons, pions,
90 electrons, neutrons, gamma rays and others of a wide range of energies covering
91 several orders of magnitude are present as primary or secondary radiation
92 (Schraube 2000; ISO 2001; Lindborg et al. 2004). Depending on altitude and
93 geomagnetic latitude, about 40 % - 70 % of ambient dose equivalent $H^*(10)$ is due

94 to neutrons, 20 % - 30 % due to electrons, 10 % due to protons and 10 % due to
95 photons and muons (Schraube et al. 2002a; Lindborg et al. 2004). In addition,
96 radiation field in the atmosphere is not constant in time and space due to solar
97 modulation of the GCR, strong variations of particle fluences and energies in
98 occasional SEPs, latitude effects caused by the geomagnetic field and build-
99 up/absorption effects resulting from nuclear reactions with the atmospheric nuclei.

100 An assessment of the radiation exposure of aircraft crew requires a determination
101 of the radiation protection quantity effective dose E (ICRP 2007). Since the effective
102 dose is not a measurable quantity, for operational radiation protection purposes, an
103 operational quantity, the ambient dose equivalent $H^*(10)$ was introduced (ICRU
104 1993). $H^*(10)$ should be a conservative estimate of E . An empirical determination of
105 $H^*(10)$ onboard aircraft requires accurate measurements using radiation detectors
106 sensitive to the different particles and energy ranges. The most important species
107 are neutrons (from few hundred keV up to few GeV) as they deliver the largest
108 fraction of dose. The $H^*(10)$ can be measured with an instrument suitably calibrated
109 for this quantity what is not a trivial task for instruments to be used in atmospheric
110 radiation field. For the very complex atmospheric radiation field, with its broad
111 range of different particles and energies, there exists no single reference field
112 covering all those radiation components. ISO reference radiation fields do not fully
113 cover the whole particle and energy range of interest (ISO 2012). Additionally, for
114 proper calibration, instrument responses for all particles and energies shall be taken
115 into account. To simulate a cosmic radiation field or some of its components at
116 aviation altitude, an accelerator-produced field such as provided at CERN EU High

117 Energy Reference Field (CERF) facility (Silari and Pozzi 2017; Pozzi et al. 2017,
118 Pozzi and Silari 2019) or fields at high-mountains could be also used. However, the
119 composition and spectra of these fields are not exactly the same as the one present
120 onboard aircraft. Today, well calibrated Tissue Equivalent Proportional Counters
121 (TEPC) are considered as the instruments that reasonably well approximate the
122 operational dose quantity ambient dose equivalent in atmospheric radiation field
123 (ISO 2012, Lindborg 1999). Other instruments need to be calibrated in appropriate
124 reference fields or in situ against a TEPC.

125 Many in-flight measurements with different instruments were performed in the past
126 and an overview of the most important research projects in aviation dosimetry
127 during 1997-2007 was given in (Beck 2009). Further descriptions and results from
128 various measurement campaigns onboard aircraft between 1992 and 2003 have
129 been summarized in (Lindborg et al. 2004). Such measurements were usually done
130 on single flights with changing altitude and cut-off rigidity (Bottollier-Depois et al.
131 2004; Kubančák et al. 2015). For constant flight conditions, measurements have
132 been conducted with only a limited number of instruments, such as TEPC and silicon
133 spectrometers (Meier et al. 2016; Lindborg et al. 2007; Latocha et al. 2007; Lillhök
134 et al. 2007). Recently several new detectors that are potentially suitable for onboard
135 aircraft dosimetry have been developed, but not yet fully tested in the field
136 (Bottollier-Depois et al. 2019; Yasuda et al. 2020; Kakona et al. 2019).

137 Despite the measurements performed so far, there is still need for continuous
138 measurements onboard aircraft especially for observing short-term variations of
139 radiation levels associated with SEP. The silicon spectrometer Liulin has been used

140 onboard aircraft for many years. Several Liulin detectors are permanently installed
141 onboard aircraft of Air France and Czech airlines (Ploc et al. 2013) although their
142 sensitivity to neutrons is rather low and they are not tissue-equivalent. A TEPC (e.g.
143 like Hawk-type) is typically not used for long-term measurements due to its rather
144 large dimensions and relatively high power consumption. A unique exception is
145 long-term TEPC measurements reported by (Beck et al, 2005) where the
146 “Halloween Storms” between October and November 2003 were recorded.

147 Intercomparisons with different types of instruments, which are usually calibrated
148 in different ways, are necessary. A comparison exercise employing different
149 instruments conducted in regular time intervals (e.g. every few years) represents an
150 independent form of a quality control for participating groups. In addition, in a view
151 of a growing demand for increasing the quality of dosimetric measurements at
152 aviation altitudes by the space weather community (Tobiska et al. 2015; Meier et al.
153 2018) measurement campaigns onboard aircraft are necessary.

154 In November 2017, the research campaign REFLECT (**RE**search **FL**ight of **EURADOS**
155 and **CRREAT**) was carried out by Nuclear Physics Institute CAS. The response of
156 more than 20 different detectors was investigated during a flight onboard a small
157 aircraft. The instruments’ ensemble included those already proved for dosimetry
158 onboard aircraft such as Liulin and TEPCs, as well as newly developed detectors and
159 instruments with the potential to be used for onboard aircraft measurements in
160 future. Dosimetric data under well-defined conditions, including constant altitude
161 and constant space weather conditions, were acquired. Sixteen institutes
162 participated, several of them representing the leading research groups in aviation

163 dosimetry in their respective countries. As a result, REFLECT is the largest campaign
164 of this type ever performed. This campaign was part of the research activities of
165 Working Group 11 of EURADOS (EURADOS 2020) and of the CRREAT (Research
166 Center of Cosmic Rays and Radiation Events in the Atmosphere) project (CRREAT
167 n.d.).

168 **2.Instruments**

169 Radiation detectors included in the REFLECT campaign embraced instruments
170 routinely used for cosmic radiation monitoring (TEPC, Liulin), newly developed
171 radiation detectors as well as detectors with future potential for cosmic radiation
172 monitoring onboard aircraft. With one exception, all instruments were active
173 radiation detectors, i.e. electronic instruments capable of making time-resolved
174 measurements.

175 An overview of the detectors used listing instruments, measured quantities,
176 typically used radiation fields and participating institutes is given in Table 1. The
177 detectors routinely used are underlined. Others are various neutron rem-counters,
178 Si-detectors, recombination chamber or scintillation detectors.

179 Table 1. Instruments used during REFLECT

Instrument	Quantity measured/ provided	Typical radiation field	Institute
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<u>TEPC Hawk</u>	$H^*(10)$	Mixed radiation	Institute for Radiological Protection and Nuclear Safety, France (IRSN) Seibersdorf Laboratories, Austria (SL)
<u>Sievert instrument</u>	$H^*(10)$	Mixed radiation	Swedish Radiation Safety Authority, Sweden (SSM)
<u>Liulin</u>	$D(Si), H^*(10)$	Mixed radiation	Nuclear Physics Institute of the CAS, Czech Republic (NPI) National Institute of Radiological Sciences, QST, Japan (QST) German Aerospace Center, Germany (DLR)
REM-2 recombination chamber	$H^*(10)$	Mixed radiation	National Centre for Nuclear Research, Poland (NCBJ)
LB 6419	$H^*(10)$	Mixed radiation Neutrons (thermal – 300 MeV), photons	Deutsches Elektronen-Synchrotron, Germany (DESY)
TTM low-level neutron and gamma	$H^*(10)$	Mixed radiation	National Centre for Nuclear Research, Poland (NCBJ)

monitoring station			
Airdos	$D(Si)$	Mixed radiation	Nuclear Physics Institute of the CAS, Czech Republic (NPI)
Minipix	$D(Si)$	Mixed radiation	Nuclear Physics Institute of the CAS, Czech Republic (NPI)
NM2B-495Pb	$H^*(10)$	Neutrons (up to 10 GeV)	Helmholtz Zentrum München, Germany (HMGU)
LINUS	$H^*(10)$	Neutrons (up to 2 GeV)	European Council for Nuclear Research, Switzerland (CERN)
LB6411	$H^*(10)$	Neutrons (up to 20 MeV)	Nuclear Physics Institute of the CAS, Czech Republic (NPI)
Passive REM counter	$H^*(10)$	Neutrons	Politecnico di Milano, Italy (Polimi)
ELDO	$H_p(10)$	Neutrons (up to 200 MeV)	Helmholtz Zentrum München, Germany (HMGU)
HammerHead HH	$H^*(10)$	Photons (50 keV – 8 MeV), electrons, protons, muons, pions	HHtec for HHtec Association, Czech Republic (HHtec)

FH 40 G-10 with FHZ- 612B probe	$H^*(10)$	Photons	National Centre for Nuclear Research, Poland (NCBJ)
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181 **2.1 Tissue equivalent proportional counters (TEPC)**

182 A TEPC has the ability to provide values of the dose equivalent in tissue-equivalent
 183 material from most radiation components reasonably well. It is therefore
 184 particularly useful in comparisons of cosmic radiation measurements onboard
 185 aircraft (EURADOS 1996). Several different TEPCs were used to measure the dose
 186 equivalent during the REFLECT.

187 **2.1.1 Hawk environmental Monitoring System FW-AD**

188 The Hawk environmental Monitoring System FW-AD is a tissue equivalent
 189 proportional counter from Far West Technology Inc. (Goleta, California, USA),
 190 composed of a spherical chamber (127 mm diameter) with a wall from A-150 tissue
 191 equivalent plastic (2 mm thick) and filled with pure propane gas at low pressure
 192 (about 9.33 hPa) simulating of 2 μm site size (Conroy 2004). The outer container is
 193 made of 6.35 mm thick stainless steel. The dose equivalent is calculated from a
 194 spectrum of single energy deposition events and a radiation quality factor Q ,
 195 determined by the $Q(L)$ relation given in (ICRP 60), where L denotes the
 196 unrestricted linear energy transfer (LET) in the exposed material (ICRP 2007).

197 Both IRSN and SL used Hawk type 1 systems using two linear multichannel analyzers
198 working in parallel with low and high gains. The low-gain analogue to digital
199 converter (ADC) measures LET spectra up to $1024 \text{ keV}\cdot\mu\text{m}^{-1}$ with $1 \text{ keV}\cdot\mu\text{m}^{-1}$
200 resolution. The high-gain channel uses an ADC measuring up to a lineal energy of 25.6
201 $\text{keV}\cdot\mu\text{m}^{-1}$ with a resolution of $0.1 \text{ keV}\cdot\mu\text{m}^{-1}$. The energy deposition of the low and high
202 LET components and the associated quality factor are stored in an output file once
203 per minute. The separation between the low and the high LET component is set at 10
204 $\text{keV}\cdot\mu\text{m}^{-1}$ according to the $Q(L)$ relationship (ICRP 2007). Events, encountering
205 significant electronic noise, below the so-called low energy threshold ($0.3 \text{ keV}\cdot\mu\text{m}^{-1}$
206 for IRSN and $0.5 \text{ keV}\cdot\mu\text{m}^{-1}$ for SL) are not recorded. An extrapolation function based
207 on ^{60}Co gamma-rays LET spectrum was chosen for the SL Hawk. For the IRSN Hawk
208 data analysis, a simple coefficient (the average of correction factor determined for
209 ^{60}Co and ^{137}Cs gamma-rays) was applied (Farah et al. 2017). No compensation of the
210 counting loss due to dead time is included in the analysis software.

211 Correction factors, N_{low} and N_{high} to ambient dose-equivalent for the low and high LET
212 components of the dose equivalent are used. N_{low} was determined in photon radiation
213 fields with ^{60}Co and ^{137}Cs sources. N_{high} was defined using the neutron reference
214 sources of $^{241}\text{Am-Be}$ or ^{252}Cf neutron sources. The values of N_{low} are 1.11 ± 0.02 and
215 1.34 ± 0.03 and the values of N_{high} are 0.80 ± 0.09 and 0.84 ± 0.10 for IRSN and SL,
216 respectively. Correction coefficients for neutrons were also evaluated for between 0.5
217 and 19 MeV and were found similar to Am-Be or ^{252}Cf neutron sources (Trompier et
218 al. 2007).

219 **2.1.2 Sievert instrument**

220 The Sievert instruments are microdosimetric detectors developed by SSM (Kyllönen
221 et al. 2001a; Lillhök et al. 2017). The detectors are TEPCs with 5 mm A-150 walls
222 housed in vacuum containers of 2 mm aluminum. The detector volume has a
223 diameter and length equal to 11.54 cm and a volume of 1207 cm³. The detectors are
224 working at a gas pressure of 1.3 kPa of propane based tissue-equivalent gas with
225 (volume fractions) 55% C₃H₈, 39.6% CO₂ and 5.4% N₂, to simulate an object size
226 with a mean chord length of 2 μm.

227 The electric charge is integrated for an integration time of typically 0.1 to 0.3 s. The
228 absorbed dose to detector gas during this time interval is calculated from the
229 average charge, the mass of the detector gas, the mean energy required to create an
230 ion pair (an average value of 27.2 eV was used in the analysis), and the detector gas
231 multiplication factor.

232 Characterization of the radiation quality is based on the variance-covariance
233 method (Kellerer 1968; Bengtsson 1970; Lindborg and Bengtsson 1971; Kellerer
234 and Rossi 1984).

235 In cosmic radiation applications where the high-LET events are rare and the
236 absorbed dose rate is relatively low, a mixed single-event and multiple-event
237 analysis can be used (Kyllönen et al. 2001b). The measured spectrum will in such
238 situations have a region dominated by multiple events, and another region
239 dominated by single high-LET events. The regions are chosen to be separated at
240 150 keV·μm⁻¹. The quality factor in the multiple-event region (<150 keV·μm⁻¹) is

241 calculated from the dose-average lineal energy by using a linear $Q(y)$ relation. In the
242 region above $150 \text{ keV}\cdot\mu\text{m}^{-1}$, the events are treated as single events (after correction
243 for a multiple-event contribution), y was set equal to L and the corresponding
244 absorbed dose fraction multiplied by the quality factor defined in ICRP 103 (ICRP
245 2007). In addition, a correction factor $c_{Q, \text{high}} = 1.25$ for the high-LET component
246 below $150 \text{ keV}\cdot\mu\text{m}^{-1}$ for the difference between the $Q(y)$ -function used and $Q(L)$
247 according to ICRP60 is obtained from a previous comparison of the two approaches
248 on aircraft measurements (Lillhök et al 2007).

249 From Monte Carlo (MC) simulations of the neutron detector response (Lillhök 2007)
250 using a simulated atmospheric neutron spectrum (Roesler et al 1998) the detector
251 absorbed dose and the ambient absorbed dose $D^*(10)$ agree within 3%.

252 The low-LET and high-LET components are defined as the contribution with dose-
253 average lineal energy $1.6 \text{ keV}\cdot\mu\text{m}^{-1}$ measured with these detectors in a ^{60}Co gamma
254 radiation field, and $94 \text{ keV}\cdot\mu\text{m}^{-1}$ simulated for these detectors in a simulated
255 atmospheric neutron spectrum (Lillhök 2007).

256 **2.2 Other detectors for mixed radiation fields**

257 **2.2.1 Liulin**

258 The Mobile Dosimetry Unit (MDU) Liulin is a silicon semiconductor spectrometer
259 that has been used for cosmic radiation measurements (Dachev 2009) as well as
260 aircraft crew dosimetry for many years (Ploc et al. 2013; Meier et al. 2009). Liulin is
261 equipped with a Hamamatsu S2744-08 PIN diode ($10 \times 20 \times 0.3 \text{ mm}^3$), low noise

262 hybrid charge-sensitive preamplifier AMPTEK Inc. type A225, fast 12-bit analogue-
263 digital converter (ADC), 2 or 3 microcontrollers and flash memory. Liulin detects
264 energy imparted to its active volume in a single energy deposition event. Pulse
265 amplitudes are stored in a 256-channel spectrum (only 8 most significant bits are
266 used from ADC), from which the absorbed dose in silicon is then calculated. The
267 energy calibration of Liulin was obtained at HIMAC (Uchihori et al. 2002).

268 Liulin can also be used to estimate $H^*(10)$ onboard aircraft, using the absorbed dose
269 in silicon and an appropriate conversion factor, which can be determined by various
270 means (Ploc et al. 2011; Wissmann and Klages 2018). In this experiment, several
271 MDU models were used; however, $H^*(10)$ is given only for Liulin MDU7 (NPI). MDU7
272 was recently calibrated at CERF, which enabled to obtain calibration coefficient
273 converting D_{Si} to $H^*(10)$ as described for example in Ploc et al. (2011).

274 **2.2.2 Airdos**

275 Airdos is detector with similar design and sensitivity as Liulin. It has been designed
276 as open source instrument for measurement in mixed radiation fields with low
277 intensity such as those encountered onboard aircraft (Kakona et al., 2019). It is
278 composed of a silicon PIN diode (Hamamatsu S2744-09) of the same type used in
279 the Liulin MDU, electronics for converting the signal to the pulse-height spectra, a
280 GPS module, an SD memory card and batteries. Full documentation is freely
281 available at (GitHub, 2020). In this measurement campaign, a version Airdos 01 was
282 used. Energy range of Airdos 01 is from 0.2 to 12.5 MeV of deposited energy in
283 silicon with energy resolution 49.4 keV per channel. Accumulated pulse amplitudes

284 are stored in 250 channels spectra every 15 seconds. The detector was calibrated
285 using heavy charged particle beams at HIMAC (NIRS, Japan) and at the U-120M
286 cyclotron (NPI, Czech Republic).

287 **2.2.3 Timepix**

288 Timepix (Llopart et al. 2002; Llopart et al. 2007) is a hybrid semiconductor pixel
289 detector which consists of matrix of 256 x 256 pixels (total of 65536 pixels).
290 Timepix was developed by Medipix2 collaboration (Campbell 2011). The pixel pitch
291 is 55 μm and total sensitive area is nearly 2 cm^2 . For this flight the silicon with
292 thickness of 500 μm was used as a semiconductor sensor chip. The Timepix chip
293 was readout by compact electronics which is in MiniPIX interface (Granja 2018,
294 Granja and Pospisil 2014). The detector was operated in per-pixel energy mode
295 which allows to measure the time that the signal spent over threshold. The
296 calibration between time over threshold to deposited energy was done by method
297 described in (Jakubek 2011). Due to high granularity provided by Timepix
298 architecture the detector can measure single particle energy deposition events
299 (Granja and Pospisil 2014). The configuration of Timepix device and data
300 acquisition (including the pre-processing of data) was performed in PIXET software
301 (Turecek 2015) which was run on standard Windows laptop.

302 **2.2.4 REM-2 recombination chamber**

303 The REM-2 is a cylindrical parallel-plate recombination chamber with an active
304 volume of about 1800 cm^3 and total mass of 6.5 kg. The chamber has 25 tissue-

305 equivalent electrodes and it is filled with the gas mixture consisting of methane and
306 5% nitrogen, with high pressure up to 1 MPa. The effective wall thickness (Al) of the
307 chamber is equivalent to about 1.8 cm of tissue. The REM-2 chamber approximates
308 the dosimetric parameters of the ICRU sphere in such a way, that the dose
309 contribution and energy spectrum of secondary charged particles in the chamber
310 active cavity are similar to those in the ICRU sphere at the depth of 10 mm (Maciak
311 2018). Therefore it can be used for the determination of $H^*(10)$ in mixed radiation
312 fields (Zielczyński et al. 2008; Caresana et al. 2014; Murawski et al. 2018).

313 The chamber is designed in such a way that the initial recombination of ions occurs
314 when the chamber operates at polarizing voltages below saturation and, for a
315 certain range of gas pressure and dose rates, the initial recombination exceeds
316 volume recombination. Measuring methods are based on the determination of the
317 dose rate from the saturation current and the radiation quality from the amount of
318 initial recombination. By means of recombination methods it is possible to estimate
319 the radiation quality factor (Zielczyński et al. 1994; Golnik et al. 2004; Golnik 2018).

320 The method used for the determination of the radiation quality involves
321 measurements of two ionization currents i_S and i_R at two properly chosen polarizing
322 voltages U_S and U_R . A certain combination of these two currents is called
323 recombination index of radiation quality Q_4 and may serve as a measurable quantity
324 that depends on LET in a similar way as the radiation quality factor does (Golnik
325 2018). The polarizing voltage U_S is the high voltage, the same as for the
326 measurements of the absorbed dose. The lower voltage U_R , called the recombination
327 voltage, has been determined during calibration of the chamber in a reference

328 gamma radiation field of air kerma from ^{137}Cs source. U_R ensures 96% of ion
329 collection efficiency in such reference field. The ambient dose equivalent is
330 calculated as the product of absorbed dose and Q_4 .

331 The detector was calibrated at CERF in 2016, and twice in monoenergetic neutron
332 reference fields: at PTB (Golnik et al. 1997) and in 2018 at NPL. Before the REFLECT
333 measurements, the chamber was calibrated at 990 V saturation voltage in the
334 accredited (AP 070) Radiation Protection Measurements Laboratory (LPD, NCBJ)
335 according to the Operational procedure M-1 (2017) with a ^{137}Cs reference photon
336 source and PuBe reference neutron source.

337 **2.2.5 TTM low-level neutron and gamma monitoring station**

338 The low-level neutron and gamma monitoring station registers photons and
339 neutrons in separate 'pulse-height' windows (Pszona et al. 2014). The detector is
340 based on an 8 inch Leake neutron area survey instrument (Leake et al. 2004). It uses
341 a Centronic SP9 ^3He spherical proportional counter, surrounded by an inner
342 polyethylene layer, a spherical shell of natural cadmium and a further outer
343 polyethylene moderator. The cadmium shield is composed of two hemispherical
344 shells, 0.91 mm thick, with 25 holes. The areas covered by the holes are the same in
345 both hemispheres, except for a 12.5 mm hole used by the SP9 connector (Tagziria et
346 al. 2004). Discrimination between photons and neutrons is based on the analysis of
347 the pulse-height spectrum, defining the photon and neutron windows (Pszona et al.
348 2014). The neutron response function is shown in Figure 1. The TTM station was
349 calibrated with ^{137}Cs and AmBe reference sources in the accredited (AP 070)

350 Radiation Protection Measurements Laboratory (LPD, NCBJ). Calibration factors
351 used for the measurement were 0.55 nSv and 1.28 nSv per count for the photon and
352 neutron windows, respectively.

353 **2.2.6 LB 6419**

354 The LB 6419 was designed by DESY and Berthold Technologies to measure the
355 ambient dose equivalent $H^*(10)$ of pulsed and continuous neutron and photon
356 radiation at high-energy accelerators (Leuschner et al. 2017). The LB 6419
357 comprises a cylindrical moderated rem-counter with a ^3He proportional counter
358 and a plastic scintillator.

359 The response to low-energy neutrons H_{LEN} is obtained from the proportional
360 counter by counting the reaction products of the nuclear reaction $^3\text{He}(n,p)\text{T}$. Its
361 moderator is made of polyethylene and contains neither any response-shaping
362 absorbers like Cd or B nor converters like Pb. So it measures H_{LEN} with a calibration
363 factor of 0.1 nSv per count. Its neutron response function is shown in Figure 1.

364 The response to high-energy neutrons H_{HEN} is obtained from the scintillator by
365 collecting scintillator light above 20 MeV, a threshold where any response from
366 electro-magnetic radiation such as γ , e^\pm , μ^\pm can be discriminated. The response
367 comes from the energy deposition of charged products from neutron scattering on
368 hydrogen nuclei of the scintillator $\text{H}(n,n)p$ and on carbon nuclei as $\text{C}(n,p)$ and
369 $\text{C}(n,\alpha)$. As this response is based on the measurement of absorbed energy rather
370 than counting it cannot be shown in Figure 1. The corresponding calibration factor
371 was measured and validated at CERF in 2010, 2012 and 2017.

372 The total neutron dose H_N is obtained by summing up the doses of the two energy
373 ranges H_{LEN} and H_{HEN} .

374 Electromagnetic radiation (H_{ELM}) can be separated from the neutron response
375 because it shows up in the energy spectrum as the so called “muon peak”. In the
376 cylindrical scintillator with its dimension of 4.1 cm these minimum ionizing
377 particles lose about 8 MeV ($2 \text{ MeV}\cdot\text{cm}^{-2}$). The calibration is done by means of the
378 Compton edges of radioactive sources such as ^{137}Cs and ^{60}Co .

379 Finally the total dose H_{TOT} is obtained by summing up the neutron dose H_N and the
380 dose of the electro-magnetic radiation H_{ELM} .

381 **2.2.7. HammerHead HH**

382 The HammerHead HH (HHtec Association, Czech Republic) is a wide-range
383 scintillation detector designed for high-precision $H^*(10)$ measurements. The
384 ambient dose equivalent rate range is from $5 \text{ nSv}\cdot\text{h}^{-1}$ to $10 \text{ mSv}\cdot\text{h}^{-1}$ for a photon
385 energy range from 50 keV to 8 MeV. The typical type A uncertainty is 12 % for 1 s
386 measuring interval (1σ and $H^*(10)_{\text{terrestrial}} = 130 \text{ nSv}\cdot\text{h}^{-1}$). The HH meter is a portable
387 detector with dimensions of $\varnothing 80 \text{ mm} \times 340 \text{ mm}$ and mass of 1.6 kg.

388 The HammerHead HH has been designed in order that the measured value best
389 corresponds to the physical definition of $H^*(10)$ for photons and meet the strict
390 criteria required by IEC 60846 for ambient dose equivalent meters. As detector, a
391 $\text{CaF}_2:\text{Eu}$ scintillator with low atomic number is used. Its shape is close to a sphere of
392 64 mm diameter, therefore the meter has excellent -135° to $+135^\circ$ angular
393 response. The HH meter works in current mode, therefore the measurement is not

394 influenced by dead time. The unique time-energy analysis of the measured signals
395 makes it possible to distinguish the contribution $H^*(10)_L$ from particles with energy
396 below 4 MeV and $H^*(10)_H$ from particles with energy above 4 MeV. When measuring
397 on the Earth's surface, the $H^*(10)_L$ value represents the terrestrial component of the
398 radiation field, whereas the $H^*(10)_H$ value allows estimating the secondary cosmic
399 ray component but without the influence of neutrons. The typical duration of a
400 measurement is 9 hours when the instrument is connected to a tablet for data
401 transfer.

402 HammerHead HH was calibrated in the accredited calibration laboratory at Czech
403 Metrology Institute in Prague with a X-ray device (40 keV - 250 keV) and ^{137}Cs , ^{60}Co
404 reference sources in terms of $H^*(10)$.

405 **2.2.8 FH 40 G-10 with FHZ-612B probe**

406 The FH 40 G-10 is a portable dose rate meter based on an internal energy filtered
407 proportional counter. Without any external probe connected, this device is sensitive
408 to photons only. During this experiment, an additional FHZ-612B Beta Gamma
409 probe was connected. Even with the external FHZ-61B connected, the detector was
410 used as gamma detector since the beta detector cap was installed. The $H^*(10)$
411 measuring ranges of the FH 40 G-10 and FHZ-612B are $10 \text{ nSv}\cdot\text{h}^{-1} - 1 \text{ Sv}\cdot\text{h}^{-1}$ and 100
412 $\text{ nSv}\cdot\text{h}^{-1} - 10 \text{ Sv}\cdot\text{h}^{-1}$, respectively. The energy range is 20 keV – 4.4 MeV for the FH 40
413 G-10 and 82 keV – 1.3 MeV for the FHZ-612B.

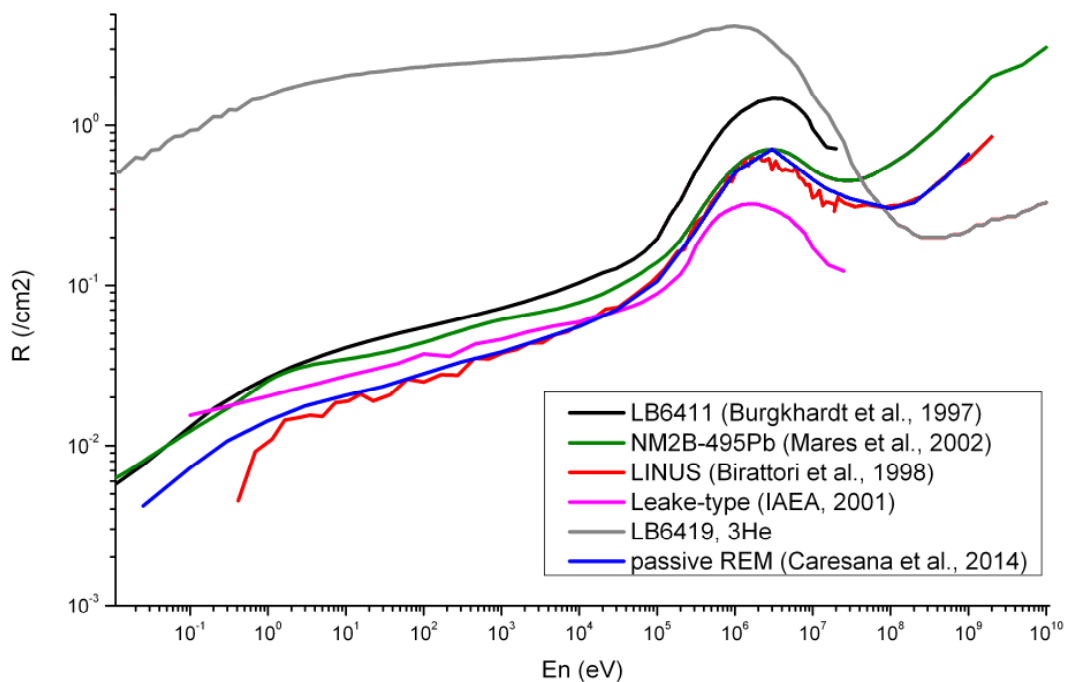
414 The instrument was calibrated in the accredited (AP 070) Radiation Protection
415 Measurements Laboratory (LPD, NCBJ) with a ^{137}Cs reference source in terms of
416 $H^*(10)$ (ISO 1999).

417 **2.3 Neutron rem-counters and dosimeters**

418 Several neutron dosimeters and rem-counters were used. The design of neutron
419 rem counters is based mostly on the Andersson-Braun type (Andersson and Braun
420 1963) or Leake type (Leake 1966) and they measure the neutron ambient dose
421 equivalent, $H^*(10)$.

422 Neutron fluence response functions of the used neutron detectors are shown in
423 Figure 1. Response functions are usually calculated with MC codes; several energy
424 points are validated through measurements in monoenergetic neutron fields.

425



426

427 Figure1. Neutron response function R (counts per unit neutron fluence) of the neutron detectors

428 used in REFLECT.

429

430 2.3.1 NM2B-495Pb Rem Counter

431 The NM2B-495Pb Rem Counter is based on the conventional Andersson-Braun rem-

432 counter (NE Technology Ltd.) with a cylindrical BF_3 proportional counter

433 surrounded by an inner polyethylene moderator, a boron-doped synthetic rubber

434 absorber, and an outer polyethylene moderator. To extend the detection range to

435 higher energy neutrons, a 1 cm thick lead shell is added around the boron rubber.

436 For this experimental flight, pulse height spectra were registered to control the

437 photon background and properly set up the region of interest (ROI). This procedure

438 enables an appropriate evaluation of the number of counts which are then

439 converted to $H^*(10)$ through the calibration coefficient. The fluence response
440 function from thermal to 10 GeV neutrons was calculated by means of different
441 Monte Carlo codes (Mares et al. 2002). The rem counter calibration was performed
442 using a 185 GBq ^{241}Am -Be neutron source following the ISO recommendations (ISO
443 2001). The rem counter was also used and calibrated in 100 and 300 MeV quasi-
444 mono-energetic neutron fields at RCNP in Osaka (Mares et al. 2017) and at CERF.
445 The response function of the detector is shown in Figure 1.

446 **2.3.2 LINUS**

447 The LINUS (Birattari et al. 1990; Birattari et al. 1992; Birattari et al. 1993; Birattari
448 et al. 1998) is the original extended-range rem counter. It consists of a ^3He
449 proportional counter embedded in a spherical polyethylene moderator, which
450 incorporates a boron-doped rubber absorber and a 1 cm thick lead shell so that its
451 response function extends up to several hundred MeV. The signal is treated with a
452 standard counting chain (pre-amplifier, amplifier, single channel analyzer and
453 counter) and the TTL output is analyzed by a custom LabVIEW interface. The
454 response function of the detector is shown in Figure 1. Neutron detectors are
455 sensitive to some extent to gamma rays, which can transfer energy to the system
456 through Compton scattering in the walls or fill gas. The gamma rejection for the
457 LINUS is obtained by setting a discriminator below the low energy neutron signal to
458 reject counts due to gamma rays and electronic noise. The threshold was
459 determined by analyzing the pulse height spectrum of the ^3He counter.

460 The LINUS was calibrated with an AmBe source (Dinar et al. 2017) in the CERN
461 CALibration LABoratory (CALLAB) (Pozzi et al. 2015). The calibration provided a
462 calibration factor of 0.89 nSv per count with an overall uncertainty of 3.2% at one
463 sigma.

464 **2.3.3 LB6411**

465 The LB 6411 neutron probe (Burgkhardt et al. 1997), connected to the universal
466 monitor LB 123, is designed for measurement of neutron ambient dose equivalent
467 $H^*(10)$ in accordance with ICRP 60 (BERTHOLD n.d.). The LB 6411 consists of a
468 cylindrical ^3He proportional counter centered in a polyethylene sphere with
469 diameter 25 cm. The neutron energy range is from thermal to 20 MeV. The spectrum
470 from the bare ^{252}Cf neutron source has been used as the calibration spectrum. The
471 numerical calibration factor is 0.32 nSv per count (Burgkhardt et al. 1997). The
472 response function over the whole energy range was calculated with MCNP. For
473 several energies the results were crosschecked with monoenergetic neutron
474 measurements. The response function of the detector is shown in Figure 1. The
475 response to gamma radiation is approx. 10^{-3} counts per nSv, which means a
476 discrimination factor of 3×10^3 .

477 **2.3.4 Passive REM counter**

478 The neutron contribution to $H^*(10)$ was also measured with a system consisting of
479 two CR-39 detectors $3 \times 4 \text{ cm}^2$ in dimension coupled to a ^{10}B enriched converter,
480 positioned inside a sphere made with polyethylene, lead, and cadmium. The ^{10}B is

481 contained in boron carbide (B_4C) deposited on an aluminum plate. The thickness of
482 the boron carbide is about $10\ \mu\text{m}$. The instrument is an extended range rem counter
483 and the response function is shown in Figure 1. The full description of the
484 instrument is in (Caresana et al. 2014) while previous experience in measuring
485 onboard aircraft is described in (Federico et al. 2015).

486 The plug, hosting the two CR39 detectors assembled with the boron converter, was
487 removed from the moderating sphere during the shipment, inserted into the sphere
488 immediately before take-off and removed immediately after landing.

489 A check of the calibration coefficient was performed at CERF in August 2017 and
490 resulted in $10.6\ \text{cm}^{-2}\cdot\mu\text{Sv}^{-1}$ with an uncertainty equal to 14% ($k=1$). The sensitivity is
491 about 3 times higher than the one reported in (Caresana et al. 2014). This is because
492 the boron converter used in the above cited work is the Enriched Converter Screen
493 BE10 by Dosirad (France) whose thickness is about $100\ \mu\text{m}$. Using this converter,
494 only a layer of about $10\ \mu\text{m}$ directly facing the CR39 detector contributes to the
495 signal, while interactions occurring at longer distance generate reaction products
496 that are self-absorbed in the converter. The effect is a depression of the thermal
497 neutron flux, resulting in a reduced sensitivity.

498 **2.3.5 Electronic neutron dosimeter ELDO**

499 The ELDO is an individual dosimeter developed at the Helmholtz Zentrum München
500 (HMGU), sensitive to neutrons from thermal energies up to about 200 MeV
501 (Wielunski et al. 2004). It is a small ($160\ \text{g}$, $115\times 60\times 16\ \text{mm}^3$) personal dosimeter
502 with a dose measurement range between $1\ \mu\text{Sv}$ and $10\ \text{Sv}$. Its operational lifetime is

503 about 400 hours. It consists of four Si PIN-diodes with LiF or polyethylene (PE)
504 converters encapsulated in lead or cadmium. The combination of diodes and
505 converter enables separate measurements of neutrons with one fast-sensor (PE)
506 operating in the 1–200 MeV neutron energy range, two delta-sensors (LiF)
507 functioning between 50 keV and 2 MeV, and one albedo-sensor (LiF) sensitive to
508 low-energy neutrons (<50 keV). Each sensor is sensitive to a certain neutron energy
509 range and has its own calibration factor. The measured dose and dose rate in terms
510 of the personal dose equivalent, $H_p(10)$ (an operational quantity for individual
511 monitoring for the assessment of effective dose), are also shown on its LCD display.
512 Calibration of the ELDO was done at PTB Braunschweig, Germany, in mono-
513 energetic neutron fields with energies between 138 keV and 14.8 MeV (Bergmeier
514 et al. 2013). Additionally, the ELDO was also tested in the reference field of CERN-
515 CERF providing high-energy fields similar to that of secondary cosmic rays at flight
516 altitudes (Wielunski et al., 2018) and at the Environmental Research Station "UFS
517 Schneefernerhaus" (2,650 m above sea level) close to the summit of the Zugspitze
518 Mountain, Germany (Volnhals, 2012). In these experiments, an excess in the
519 measured counts was observed which are due to protons. Although the fluence of
520 protons is only 6% of the neutron fluence, protons caused around 12% of the
521 measured counts. The sensor response to protons and muons could be calculated
522 with GEANT4 calculations (Volnhals 2012) which support the observation.

523 **2.4 Calculations**

524 Ambient dose equivalent rates for different particles can be calculated using various
525 models; the overview of codes assessing radiation exposure of aircraft crew is given
526 in (Bottollier-Depois et al. 2012). All these codes provide calculations for the GCR
527 induced radiation field in aircraft flight altitudes agreeing within 20 % with
528 reference measurements (Bottollier-Depois et al. 2012). In this publication, the
529 EPCARD.Net code (Mares et al. 2009) was used for comparison.

530 **2.4.1 EPCARD.Net**

531 The European Program package for the Calculation of Aviation Route Doses
532 (EPCARD) is a widely used program for estimating the exposure of aircraft crew.
533 This code was developed at the Helmholtz Zentrum München (Schraube et al.
534 2002b) and further improved in a new object-oriented code EPCARD.Net (Mares et
535 al. 2009). In 2010, EPCARD.Net ver. 5.4.3 Professional was approved for official use
536 for assessing radiation exposure from secondary cosmic radiation at aviation
537 altitudes by the German Aviation Authority (LBA) and the National Metrology
538 Institute, Physikalisch-Technische Bundesanstalt (PTB).

539 EPCARD.net is based on the results of extensive FLUKA Monte Carlo (Ferrari et al.
540 2005; Böhlen et al. 2014) calculations of particle energy spectra of neutrons,
541 protons, photons, electrons and positrons, muons, and pions at various depths in the
542 atmosphere down to sea level for all possible values of solar activity and
543 geomagnetic shielding conditions (Roesler et al. 2002). The primary particle spectra

544 used in the FLUKA calculations as well as the modulation potential describing solar
545 activity were based on the model of Badhwar and O'Neill (Badhwar 1997; Badhwar
546 et al. 2000).

547 To determine the dose rates at specific locations in the atmosphere during a flight,
548 the cut-off rigidity, the solar deceleration potential and the barometric altitude are
549 calculated to quantify geomagnetic shielding, solar activity and atmospheric
550 shielding. The EPCARD.Net parameter database includes energy-averaged dose
551 conversion coefficients, calculated by folding each single-particle fluence spectrum
552 with the appropriate dose conversion function (Mares et al. 2004; Mares and
553 Leuthold 2007), which depends on barometric altitude, cut-off rigidity, and solar
554 activity, since the shape of the particle energy spectra also depends on these
555 parameters. Ambient dose equivalent, $H^*(10)$, and effective dose, E , are calculated
556 separately for each particle, i.e. the dose contributions from neutrons, protons,
557 photons, electrons, muons, and pions are assessed individually.

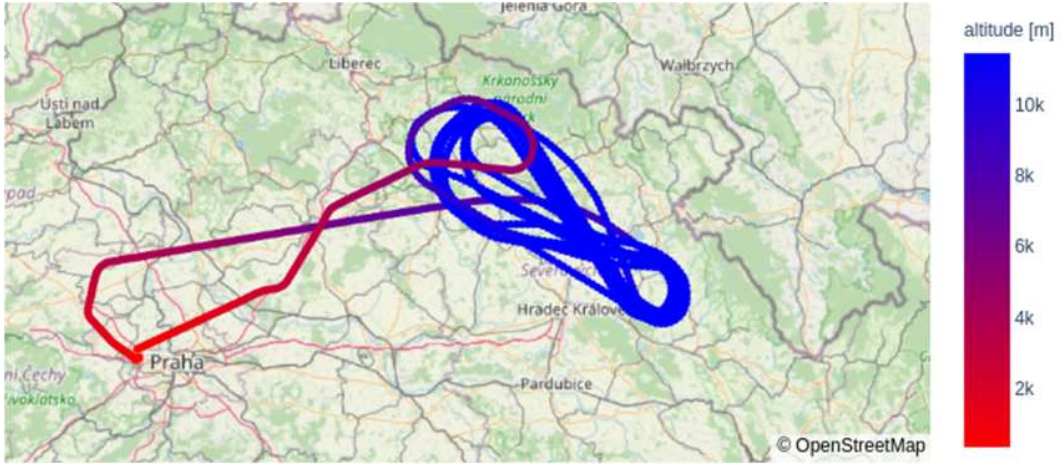
558 More general information about EPCARD is available on the web site (EPCARD
559 2020), where a simplified on-line version of the EPCARD calculator for public use
560 can also be found.

561 **3. Experiment**

562 **3.1 Flight**

563 The radiation detectors were exposed aboard an Embraer Legacy 600 aircraft
564 operated by ABS Jets. The aircraft, together with 250 kg of equipment and eight
565 scientific staff, flew from Vaclav Havel Airport in Prague (50.1°N, 14.2° E) to the
566 FL390 flight level, on the 29th November 2017. The flight took off at the airport at
567 13:06 UTC and reached stable flight conditions (barometric altitude 11871 ± 8 m,
568 range from 11853 to 11893), latitude 50.41 ± 0.14 °N (range from 50.18 to 50.58),
569 longitude 15.80 ± 0.27 ° E (range from 15.26 to 16.24) at 13:38 UTC. At this level,
570 the aircraft circled over the northern part of the Czech Republic (the area is a
571 reserved airspace that is commonly used for operating test flights) for 90 minutes
572 and landed back at Prague Airport at 15:34 UTC. Navigation data (barometric
573 altitude, latitude, and longitude) were taken from the aircraft record and GPS. The
574 flight route and flight profile are shown in Figure 2 and Figure 3, respectively. The
575 space weather conditions were stable during the whole flight and no short-term
576 solar activity affected the results. Space weather situation can be assessed e.g. by
577 neutron monitors (nmdb.eu). During the flight the variation in the count rates of the
578 neutron monitor at Lomnický štít (the nearest neutron monitor) was below 0.5%,
579 which indicates stable space weather conditions.

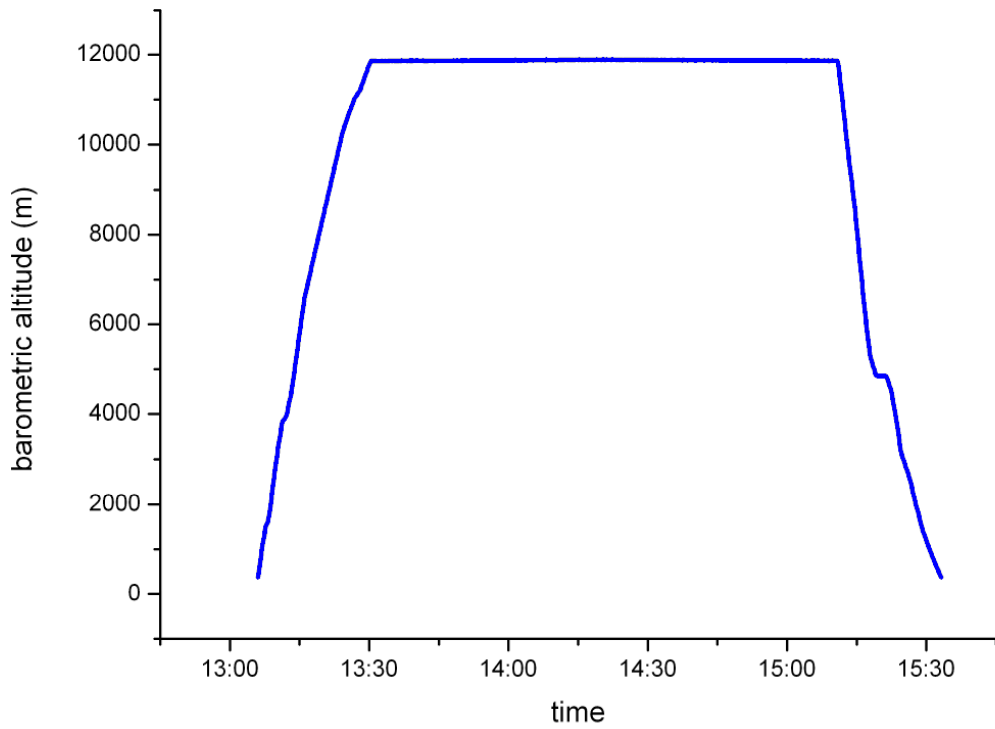
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581

582 Figure 2. Flight route

583



584

585 Figure 3. Flight profile

586

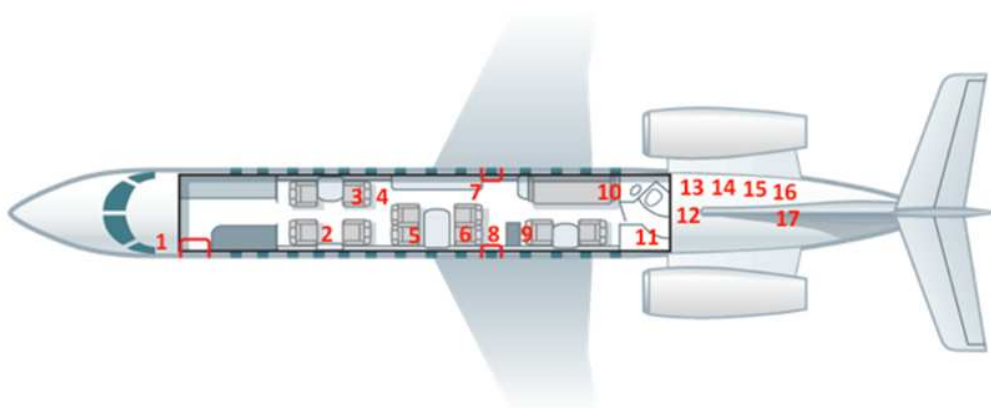
587 **3.2 Location of detectors inside the aircraft**

588 The detectors were placed at various locations inside the aircraft (Fig. 4).

589 Equipment that needed power and manual control were installed on or behind the
590 seats. Smaller devices like Liulins or Airdos were distributed in various locations
591 inside the aircraft. The rest of the instruments were stored in the baggage
592 compartment.

593 Two fuel tanks are located in the wings, two fuel tanks are in the bottom part of the
594 body and two fuel tanks are in the rear part of the plane, behind the baggage
595 compartment. Because the flight was quite short (about 2.5 hours), only the tanks in
596 the wings were filled with fuel. The total amount of fuel before take-off was 5482
597 liters (4380 kg), 2530 kg were burned during the flight.

598



599

600 Figure 4. Placement of detectors inside the aircraft: 1 – Liulin MDU10; 2 – Sievert, 3 – Timepix; 4 –

601 HammerHead; 5 – LB6411; 6 – LB6419; 7 – Liulin MDU7; 8 – NM2B-495Pb Rem Counter; 9 – Liulin

602 MDU14; 10 – LINUS; 11 – AIRDOS; 12 – REM-2 recombination chamber; 13 – TTM + FH 40 G-10 +
603 FHZ-612B, 14 – ELDO; 15 – TEPC Hawk (IRSN); 16 – passive REM counter + Liulin MDU1; 17 – TEPC
604 Hawk (SL).

605 **4. Results and Discussion**

606 Not all devices operated during the whole flight, part of the instruments were
607 started when stable flight conditions were reached. To compare the results obtained
608 with the active detectors, we consider only data acquired at a constant flight
609 altitude.

610 Values of $\dot{H}^*(10)$ for various particles calculated with EPCARD are shown in Table 2.
611 It should be noted that the calculations were done in free air, whereas the
612 instruments measured inside the aircraft and therefore small differences can be
613 expected due to shielding effects (Ferrari et al. 2004). As can be seen from Table 2,
614 the most important contribution to $\dot{H}^*(10)$ is from neutrons (57% of the total
615 $\dot{H}^*(10)$), followed by electrons (20%) and protons (15%). The uncertainty on the
616 calculated values is estimated to be less than 20%, based on (Bottollier-Depois et al.
617 2012) who compared various codes used for assessing radiation exposure of aircraft
618 crew due to GCR with the conclusion that the agreement between the codes was
619 better than 20% from the median. The codes have also been previously validated by
620 measurements with an agreement better than $\pm 20\%$ (Lindborg, 2004).

621 Table 2. Calculated values of $H^*(10)$ rate during the REFLECT at FL390 for different particles using
622 EPCARD.Net ver. 5.5.0

	neutrons	photons	protons	electrons	muons	total
$\dot{H}^*(10)$ [$\mu\text{Sv/h}$]	3.8 ± 0.8	0.3 ± 0.1	1.0 ± 0.2	1.3 ± 0.3	0.2 ± 0.0	6.6 ± 1.3

623

624 Table 3. Dose-equivalent rate measured with the various detectors. Uncertainties given as combined
625 uncertainties with $k=1$ and with the contribution from measurement statistics in parenthesis.

	Instrument	$\dot{H}^*(10)$ [$\mu\text{Sv/h}$]	$\dot{H}^*(10)_{\text{Low-LET OR}}$ $\dot{H}^*(10)_{\gamma+e}$ [$\mu\text{Sv/h}$]	$\dot{H}^*(10)_{\text{High-LET OR}}$ $\dot{H}^*(10)_n$ [$\mu\text{Sv/h}$]
routinely used instruments	TEPC HAWK (SL)	7.1 ± 0.5 (0.3)	3.3 ± 0.1 (<0.1)	3.8 ± 0.6 (0.3)
	TEPC HAWK (IRSN)	7.9 ± 0.6 (0.3)	3.5 ± 0.1 (<0.1)	4.4 ± 0.5 (0.3)
	Sievert (SSM)	7.4 ± 0.6 (0.4)	2.9 ± 0.2 (0.1)	4.5 ± 0.6 (0.6)
	Liulin MDU7 (NPI)	7.1 ± 1.1	3.1 ± 0.5	4.0 ± 0.7
Neutron rem-counters	NM2B-495Pb (HMGU)			3.7 ± 0.4
	LINUS (CERN)			3.9 ± 0.1
	LB6411 (NPI)			2.1 ± 0.4
	Passive REM counter (Polimi)			7.5 ± 2.5

Others instruments	LB6419 (DESY)	9.1 ± 1.8	1.7 ± 0.4	7.4 ± 1.5
	REM-2 (NCBJ)	10.1 ± 8.9	–	–
	TTM (NCBJ)	4.8 ± 0.8	2.1 ± 0.4	2.8 ± 0.5
	ELDO (HMGU)			4.4 ± 0.9*
	HammerHead HH (HHtec Association)		2.7 ± 0.2	
	FH 40 G-10 (NCB)		3.2 ± 0.5	
	FHZ-612B (NCB)		4.7 ± 0.8	

626 * $H_p(10)$

627

628 Table 3 lists the measurement results (for all instruments except Si detectors
629 measuring only D_{Si}). The results are grouped in a low-LET component that
630 comprises the contribution of low ionizing radiation (photons, electrons, muons,
631 protons, pions) and in a high-LET component representing mostly contributions of
632 neutrons, stopping protons and higher Z ionizing particles. Even if the neutron
633 contribution can extend below $10 \text{ keV}\cdot\mu\text{m}^{-1}$ and for low energy photons it can be
634 above $10 \text{ keV}\cdot\mu\text{m}^{-1}$, the previous assumption (high-LET assimilated to neutrons) is
635 usually made when comparing results of various instruments and calculations. One
636 should also note that most detectors designed for low-LET measurements exhibit a
637 response to neutrons that is usually unknown, especially for high-energy neutrons.

638 In the table, the results provided are not corrected for this unknown neutron
639 response. Uncertainties are given as combined uncertainties with contributions
640 from calibration and measurement statistics and presented with coverage factor
641 $k=1$. For TEPC, the statistical uncertainty is given in parenthesis. The first group
642 includes instruments routinely used onboard aircraft, measuring both low and high-
643 LET components, the second group includes neutron rem-counters and the third
644 group includes the remaining instruments.

645 When comparing the results, an agreement of $\pm 20\%$ at a 95 % confidence level is
646 considered satisfactory. The recommendation on acceptable uncertainties in
647 radiation protection is given in (ICRP 1997) where it is stated: *"...overall uncertainty*
648 *at the 95 % confidence level in the estimation of effective dose around the relevant*
649 *dose limit may well be a factor of 1.5 in either direction for photons and may be*
650 *substantially greater for neutrons of uncertain energy and for electrons. Greater*
651 *uncertainties are also inevitable at low levels of effective dose for all qualities of*
652 *radiation."*

653 The measurements of the various TEPCs (HAWK, Sievert) agreed well with each
654 other, with Liulin, and with the EPCARD calculations, as it was during a previous
655 flight comparison (CAATER) (Lillök et al 2007). The differences in low- and high-
656 LET components between the Sievert instrument and the other TEPCs were likely
657 due to the fact that the Sievert instrument distinguished $H^*(10)$ contributions from
658 photons and neutrons rather than in terms of a low and high-LET threshold. The
659 differences could have been also due to different locations of the TEPCs (Hawks in
660 the baggage compartment, Sievert in the front of the plane). During the

661 approximately 90 minutes of the cruise, the TEPC experienced statistically low
662 number of high LET events, which resulted in higher uncertainties for this
663 component.

664 The LB6419 and REM-2 measured larger values of total $H^*(10)$ than the TEPCs. The
665 REM-2 is very sensitive to vibrations, which probably led to the very high value of
666 the uncertainty. In principle, there are recombination methods for separating the
667 dose according to LET, but in this flight, the measurement method was simplified
668 because of the short time and difficult conditions. Provision of the values for the low
669 and high-LET components would be helpful to better interpret the data. Improved
670 calibration is needed to make use of REM-2 for routine aircraft dosimetry.

671 The LB6419 also measured a higher value of total $H^*(10)$ than the TEPCs. In
672 addition, Table 3 shows a small measured value of the low-LET radiation
673 contribution and an increased contribution of the high-LET component as compared
674 to the TEPCs. There is a need to check the separation method of the two components
675 and the calibration, since the total ambient dose equivalent was also too high
676 compared to the TEPC.

677 The $H^*(10)$ values measured with the TTM monitoring station, especially the
678 neutron component, was lower than the other instruments. Since only polyethylene
679 was used as a moderator (no lead or other high atomic number material was
680 included in the shell), the neutron energy range of this instrument was limited to
681 20 MeV.

682 Except for the LB6411 and TTM, the instruments measuring only the neutron
683 component provided results comparable to the TEPC results and with the EPCARD
684 calculations.

685 According to the EPCARD model, neutrons contributed for more than 50% to the
686 total $H^*(10)$; neutrons can reach energies up to several hundreds of MeV
687 (Pazianotto et al. 2017). Conventional neutron REM counters have a detection range
688 usually limited to about 15–20 MeV, their response dropping sharply at higher
689 energies. To extend the range to higher energies (up to several hundreds of MeV), a
690 shell of high-Z material (like tungsten or lead) is usually added to the PE moderator.

691 As expected, due to their response functions (Fig. 1), the LB6411 and TTM measured
692 lower values. This was in agreement also with measurements by Yasuda et al.
693 (2018), who investigated neutron doses during long-haul flights with two neutron
694 monitors and compared their results with JISCARD EX calculations. They found that
695 the relative contribution to $H^*(10)$ of neutrons with energies above 15 MeV could
696 exceed 50%.

697 The passive REM counter only provided an integral value over the whole flight (the
698 detector was installed inside the moderator just before take-off and removed after
699 landing). The total measured $H^*(10)$ was $15 \pm 5 \mu\text{Sv}$; the dose rate at flight level
700 FL390 can be assessed assuming a taxi time of about 2 h and neglecting the small
701 contribution arising from the 0.5 h spent to reach the flight altitude and get back to
702 ground.

703 A possible explanation of the large measured $H^*(10)$ value is that the passive REM
704 counter – because of a misunderstanding with the shipping company – reached
705 Prague by airmail. Of course, the plug with the detector was not in the measuring
706 position, thus insensitive to fast neutrons. However, a small contribution from
707 thermal neutrons cannot be excluded.

708 Amongst the instruments measuring only the low-LET component, only the
709 HammerHead HH and the FH 40 G-10 obtained reasonable results. The results from
710 the other instruments disagreed with both the EPCARD calculation and the TEPC
711 measurements. It is difficult to compare the results because some photon detectors
712 are not only sensitive to photons and electrons, but also to protons, muons, pions
713 and to neutrons in some extent. For these instruments, their response to
714 components of the field other than that intended to be measured is not always
715 known.

716 Table 4 summarises the results of the silicon detectors (in terms of absorbed dose).
717 For these instruments, the dose in silicon was converted to dose in water using a
718 dose conversion factor of 1.23 (Ploc 2009). In a previous intercomparison flight
719 with Liulin MDUs (Meier et al. 2016) it was found that there could be differences in
720 the mass of the sensitive volume of the detectors (Si sensor size) considered in the
721 calculation of absorbed dose. In this experiment, we calculated the absorbed dose in
722 silicon using the same Si sensor mass (0.16597 g (Meier et al. 2016)) for all Liulin
723 units and for Airdos.

724 Table 4: Results (absorbed dose rate in silicon and in water) of the measurements with the Silicon
725 detectors

Instrument	MDU 7 Liulin (NPI)	MDU 10 Liulin (QST)	MDU14 Liulin (QST)	MDU1 Liulin (DLR)	Airdos T4 (NPI)	Minipix (NPI)
D_{Si} ($\mu\text{Gy}/\text{h}$)	2.8 ± 0.2	2.1 ± 0.3	2.7 ± 0.5	2.0 ± 0.3	1.8 ± 0.2	1.9 ± 0.3
D_{H_2O} ($\mu\text{Gy}/\text{h}$)*	3.4 ± 0.3	2.6 ± 0.4	3.3 ± 0.6	2.5 ± 0.3	2.2 ± 0.2	2.3 ± 0.3

726 * The dose D_{H_2O} was calculated from D_{Si} using a conversion factor 1.23

727 Although both Liulin and Airdos have similar sensitive volumes (mass, area,
728 thickness), they have different properties such as energy range of deposited energy
729 and width of the channel. Airdos has channel width of 49.4 keV whereas Liulin's is
730 81.4 keV. The energy range of Airdos (up to 12 MeV) is smaller than Liulin's (up to
731 24 MeV), in order to provide more detailed information on the lower part of the
732 energy spectrum. The significant part of the deposited energy when measuring
733 onboard aircraft comes from events depositing energy in the first several channels
734 (for Liulins, 65–83% of the absorbed dose was due to events with deposited energy
735 below 1 MeV, only 2–6% was due to events with deposited energy above 10 MeV).
736 To compare Airdos with Liulin, we considered only events within the energy range
737 of Airdos for the calculation of absorbed dose for Liulin. For MDU 7, D_{Si} was 2.6
738 $\mu\text{Gy}\cdot\text{h}^{-1}$, for MDU 10 D_{Si} was 1.9 $\mu\text{Gy}\cdot\text{h}^{-1}$, and for MDU 14 D_{Si} was 2.6 $\mu\text{Gy}\cdot\text{h}^{-1}$, to be
739 compared with 1.8 $\mu\text{Gy}\cdot\text{h}^{-1}$ measured by Airdos.

740 Some differences in the results could be due to the different shielding configurations
741 (for example, the aircraft fuel acts as a good neutron moderator) around the
742 locations in which the devices were installed (Fig. 4). The DLR Liulin was in the

743 baggage compartment, whereas the NPI Liulin and the Airdos were in the central
744 part of the aircraft or in the crew cabin. For Embraer Legacy 600, the baggage
745 compartment is located in the rear part of the aircraft, between the engines (Fig. 4).
746 Therefore, the baggage compartment, loaded with several larger instruments and
747 suitcases, is supposed to be more shielded than other areas of the aircraft. As was
748 shown in Ferrari et al. (2004), the shielding provided by the aircraft structure
749 (wings, engines, passengers, fuel) can cause a notable reduction in E or $H^*(10)$ for
750 most components of cosmic radiation. Differences in ambient dose equivalent for
751 various places inside the aircraft can be up to about 20% (Ferrari et al. 2004;
752 Battistoni et al. 2005; Kubancak et al. 2014). Nevertheless, the differences between
753 individual Liulin-type detectors seem to be too large to be explained only by
754 different locations in the aircraft. There appears to be some systematic differences.
755 The reason of these differences should be further investigated in comparison on
756 ground. Even for the detectors using the same Si diode, several factors can influence
757 the results, e.g. energy calibration, choice of the noise threshold, channel width
758 (Kakona et al. 2019).

759 Although the Minipix has larger energy range (from 5 keV), it showed a lower
760 absorbed dose than the MDUs. This is difficult to explain since the Minipix was
761 calibrated against the TEPC and showed a very good agreement to the TEPC Hawk
762 low-LET part (Ploc 2009). However, it should be mentioned that the setting of
763 Timepix, especially bias, could have been different for different experiments, which
764 might have caused some discrepancies. In this flight, the bias was set to 30 V. It is
765 important to use the same conditions (parameters) for all instruments.

766 Taking the MDU7 (MDU 7 has been calibrated at CERF) energy deposition spectra
767 and performing the calibration according to (Ploc 2009), the total $H^*(10)$ rate
768 arrived at $7.1 \mu\text{Sv}\cdot\text{h}^{-1}$, which agreed well with the TEPC results.

769 **5. Conclusions**

770 For the instruments with the potential to be used onboard aircraft, appropriate
771 calibration and determination of calibration/correction factor is crucial. A good
772 measurement of the atmospheric ambient dose equivalent requires that the
773 instrument response for all particles and energies is properly taken into account. As
774 there is no traceable reference field for the total radiation field in the atmosphere,
775 comparison of instruments onboard aircraft is necessary.

776 Various radiation detector systems were compared onboard aircraft under stable
777 flight conditions during the REFLECT measurement campaign – the largest
778 comparison of this type ever performed. As expected, the dosimeters routinely used
779 for aircraft dosimetry (TEPC, Liulin) worked adequately, the results agreed with
780 each other as well as with the EPCARD computer model calculations.

781 Since high-energy neutrons contribute significantly to $H^*(10)$, conventional neutron
782 rem counters (with energy range limited to about 20 MeV) underestimated neutron
783 $H^*(10)$ with standard calibration procedures using neutron sources such as AmBe
784 or Cf-252. Extended-range neutron rem counters (NM2B-495Pb, LINUS) provided
785 results comparable to $H^*(10)_n$ determined with routinely used instruments and the
786 EPCARD calculation.

787 The reading of some instruments (LB6419, FHZ-612B) was higher than expected
788 from the assumption that the detector is only sensitive to a specific component of
789 the radiation field. However, it should be noted that these instruments are not
790 primarily intended for use onboard aircraft in a very complex mixed radiation field.
791 Their response to the various components of the cosmic radiation field and the
792 energy dependence still need to be fully characterised.

793 The REFLECT campaign enabled measurement in uniform, well-defined conditions
794 onboard an aircraft and comparison of new instruments with those routinely used.
795 Although the response of some instruments, not primarily intended for the use in a
796 very complex mixed radiation field such as onboard aircraft, was different than as
797 somehow expected, more experiments are necessary too finally declare them as not
798 suitable. This campaign consisted only of one flight with one set of parameters
799 (vertical cut-off rigidity, altitude, phase of solar cycle) and it was relatively short.
800 For further testing of instruments showing some potential to be used for routine
801 dosimetry onboard aircraft, additional flights at different geomagnetic cut-offs and
802 altitudes as well as comparison at appropriate reference fields (CERF) are needed.
803 The REFLECT provided valuable experience enabling to discuss some issues
804 connected with the use of the dosimeters in such complex radiation field and it also
805 provided feedback for designing future in-flight campaigns.

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