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Arnaud Cuisset, Francis Hindle, Gaël Mouret, Robin Bocquet, Jonas Bruckhuisen, et al.. Terahertz Rotational Spectroscopy of Greenhouse Gases Using Long Interaction Path-Lengths. Applied Sciences, 2021, 11 (3), pp.1229. 10.3390/app11031229. hal-03171300

HAL Id: hal-03171300 https://hal.science/hal-03171300

Submitted on 27 Oct 2021

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Article

Terahertz rotational spectroscopy of greenhouse gases using long interaction path-lengths

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Version January 18, 2021 submitted to Appl. Sci.

- Abstract: Even if on-board mm-wave/THz heterodyne receivers have been developed to measure
- ² greenhouse gases (GHGs) atmospheric profiles, rotational spectroscopy rests under-exploited for
- their monitoring unlike IR rovibrational spectroscopy. The present study deals with the ability of THz
- spectroscopy using long interaction path-lengths for GHG laboratory investigations. High-resolution
- ⁵ THz signatures of non-polar greenhouse molecules may be observed by probing very weak centrifugal
- $^{\circ}$ distortion induced rotational transitions. To illustrate, new measurements on CH₄ and CF₄ have been
- ⁷ carried out. For CH₄, pure rotational transitions, recorded by cw-THz photomixing up to 2.6 THz
- in a White type cell adjusted to 20 m, have allowed to update the methane line list of atmospheric
- databases. Concerning CF₄, Fabry-Perot THz absorption spectroscopy with a km effective pathlength
- was required to detect line intensities lower than 10^{-27} cm⁻¹/(molec cm⁻²). Contrary to previous
- synchrotron based FT-FIR measurements, the tetrahedral splitting of CF₄ THz lines is fully resolved.
- ¹² Finally, quantitative measurements of N₂O and O₃ gas traces have been performed in an atmospheric
- simulation chamber using a submm-wave amplified multiplier chain coupled to a Chernin type
- multi-pass cell on a 200 m path-length. The THz monitoring of these two polar GHGs at tropospheric
- ¹⁵ and stratospheric concentrations may be now considered.

Keywords: greenhouse gases; terahertz; long interaction path-lengths; rotational spectroscopy; trace gas
 monitoring.

18 1. Introduction

The main discussion on climate change tends to focus on carbon dioxide (CO₂), the most dominant Greenhouse Gas (GHG) (65% of the global emission) produced by the burning of fossil fuels, industrial

²¹ processes and changes in land use (deforestation, intensification of agriculture, flooding...) [1]. But

 $_{22}$ CO₂ is not the only GHG driving the global climate change: methane (CH₄), nitrous oxide (N₂O),

tropospheric ozone (O_3) and fluorinated gases have a direct contribution to the greenhouse effect. In

particular, the radiative forcing (warming influence) of long-lived GHGs such as tetrafluoromethane 24 (CF_4) with a 50000 years lifetime in earth's atmosphere [2], is undergoing accelerating growth (almost 25 a third of the increase attributed to the industrial age has been over the past 30 years) [3,4]. Several 26 currently-operating satellite and air-borne missions aim to estimate emissions and absorptions of the 27 main GHGs. By monitoring the column densities in our atmosphere, they accumulate new knowledge 28 on their global distribution and temporal variation. Fourier Transform Spectrometers (FTS) are the 29 most popular instruments used in these missions analysing the IR radiations reflected from the Earth's 30 surface and emitted from the atmosphere and the surface. We can mention: (i) the IASI interferometer in the European MetOp satellite which monitors CO₂, O₃, CH₄, N₂O GHG column amounts in 32 several IR channels of detection [5]; (ii) ACE on board the Canadian Satellite SCISAT where CF_4 and 33 others CFCs/HCFCs are retrieved routinely in addition to the most abundant GHGs [6]; (iii) TES on 34 board the Aura spacecraft which retrieves the O_3 and CH_4 profiles in the lower atmosphere from 35 their rovibrational signatures [7]. Amongst the Aura's instruments, there is also a passive microwave 36 limb-sounding (MLS) radiometer/spectrometer used to measure the pollution in the upper troposphere 37 in the presence of ice clouds and volcanic aerosols, which prevent measurements by IR/UV techniques. 38 The MLS instrument monitors O_3 and N_2O by measuring pure rotational lines in the mm-wave (240 39 GHz) and in the submm-wave (640 GHz) domains, respectively [8]. More recently, the National Space 40 Science Center of the Chinese Academy of Sciences has developed a THz Atmospheric Limb Sounder 41 (TALIS) for atmospheric vertical resolved profile observations: as for MLS/Aura, O₃ and N₂O are 42 the targeted GHGs and their rotational transitions are probed with four heterodyne radiometers (LO 43 frequencies: 118 GHz, 190 GHz, 240 GHz and 643 GHz) and several FFT spectrometers of 2 GHz 44 bandwidth with 2 MHz resolution [9]. With this instrument, they are able to retrieve the N_2O and O_3 45 profiles with average precisions lower than 10 % and 5 % from 10 to 42 kms and from 10 to 70 kms 46 altitude, respectively. Accurate THz laboratory measurements are required in order to provide the best set of GHGs 48 rotational line parameters (frequencies, widths, intensities...) and to optimize the inversion processes 49 in GHGs atmospheric concentrations retrieval. The quality of the fits depends directly on the SNR of 50 high resolution THz spectra and consequently on the sensitivity of the spectrometers [10]. According 51 to the Beer-Lambert law, longer interaction THz path-lengths are required in the situation of a weak 52 absorption. With longer wavelengths and larger beamsizes, the control of the propagation of THz 53 radiations on long-range beam paths is challenging compared to that carried out in the IR/UV 54

regions. Nevertheless, recent progress has allowed gas phase THz spectroscopy experiments with

⁵⁶ molecules/beam interaction distances from several tens to several hundreds of meters [11] to be

performed. This article highlights recent advances of THz spectroscopy performed by our group, using
 long-interaction path-lengths for the measurement of weak GHG rotational absorptions. The ability to

 $_{10}$ measure fully resolved centrifugally induced rotational transitions of non polar GHGs such as CH₄

and CF_4 will be demonstrated and the first THz measurements in an atmospheric simulation chamber

⁶¹ will be presented towards the detection of polar stable N₂O and unstable O₃ GHGs at trace levels.

62 2. Materials and Methods

With the exception for the production of ozone, commercially available high purity CH_4 , N_2O , and CF_4 GHGs with natural isotopic abundances were used throughout. The ozone was produced by corona discharge generation with pure oxygen (Air Tree Ozone Technology, C-Lasky C-L010-DTI). Various THz sources and detectors have been used in the studies presented here. In particular, three commonly used approaches, optical, optoelectronic and electronic, have been employed for the generation of the THz radiation.

69 2.1. Synchrotron based FT THz spectroscopy

The AILES beamline of the SOLEIL synchrotron was used for high-resolution broadband FT measurements in the 0.6 to 19.5 THz, (20-650 cm⁻¹) spectral range on CH₄ and CF₄. The very

- ⁷² bright light extracted from the synchrotron bending magnet has the advantage of displaying a lower
- ⁷³ divergence compared with other sources aiding its use with long path interaction White type cells.
- The FT IFS125 spectrometer, used in this case, has a 5 m delay stage allowing a resolution of 30 MHz to
- ⁷⁵ be achieved [17].

76 2.2. Cw-THz spectroscopy by photomixing

Alternatively higher resolution measurements can be made using monochromatic THz sources. 77 The generation of THz radiation by photomixing is a frequency down-conversion technique which 78 was used for the measurements of individual rotational lines of CH_4 up to 2.6 THz [12]. Two extended 79 cavity laser diodes operating around 780 nm are mixed together in a LTG-GaAs device. The lasers 80 are detuned to the desired THz frequency inducing a corresponding current in the photomixer device 81 which is coupled to an integrated log spiral antenna [21]. The radiation is pre-colimated by a silicon hyper-hemispherical lens and propagated in free space. This type of source is monochromatic with 83 excellent spectral purity and can be used in the range of 100 GHz to 3.3 THz. The available power 84 is limited to 0.1 nW from 2.5 THz. By locking onto a frequency comb, the generated THz frequency 85 is determined with a kHz accuracy [13]. Both amplitude and frequency modulation schemes can be 86 easily implemented with this source, it does however suffer from available low power levels at the highest frequencies.

89 2.3. Submm-wave spectroscopy with amplified multiplier chains

We have used two AMC from the Virginia Diodes Inc. as electronic sources for the submm-wave measurements coupled to CHARME (see section 3.2) and to the THz FP cavity (see section 3.3). They are based on the up-conversion of a synthesized microwave frequency using a cascade of Schottky diode frequency multipliers mounted in rectangular wave-guide blocks. The guided radiation is launched into free space using a horn antenna and propagated through the interaction cell. Frequencies from 100 GHz to 1 THz can be conveniently produced and easily modulated in either amplitude or frequency as desired [27]. A typical power of 50 μ W is available at 600 GHz.

97 2.4. Detection schemes

For all experiments, two types of detector have been employed. Zero Biased Detector (ZBD) is 98 an unbiased Schottky diode mounted in a wave-guide operating in detection mode. This type of 99 detector is uncooled and can provide a typical NEP of 10 pW/ \sqrt{Hz} . When greater sensitivity was 100 required a Helium cooled bolometer was used with a NEP of 1 pW/ \sqrt{Hz} . For the measurements of 101 very weakly intense CF₄ lines, we have coupled a very high finesse FP cavity with an AMC in order to 102 reach very long equivalent interaction paths (see section 3.3). A control loop is used to lock the cavity 103 to the frequency of the AMC with a frequency modulation providing an error signal. The second 104 harmonic of the frequency modulation is simultaneously exploited to detect the presence of molecular 105 absorption present in the cavity. The AMC frequency and cavity are thus swept together allowing 106 sensitive absorption measurements to be made [34]. 107

3. Results and discussion

3.1. THz measurements of CH₄ in a White type multi-pass cell

THz long-path spectroscopic experiments were initially conducted in the 2000's using home-made or commercial White type gas cells. The White cell was first described in 1942 by John U. White [14] and was a significant improvement over previous long path measurement techniques used in optical spectroscopy. A White cell is constructed using three spherical, concave mirrors having the same radius of curvature. The mirrors are separated by a distance equal to their radii of curvature. FT Michelson based interferometers have been connected by different groups to White type gas cells

allowing to improve the detection sensitivity in the Far-IR domain. The THz pathlengths reached vary 116 from several tens of meters with classical sources such as Hg lamps [15,16] up to 180 m with weakly 117 divergent and high brightness Far-IR synchrotron radiations [17,18]. As example, the White type cell connected to the Bruker IFS125 high-resolution interferometer coupled to the AILES Far-IR/THz 119 beamline of the SOLEIL synchrotron is shown in Fig. 1 (right part). Optoelectronic THz sources have 120 been also coupled to White type cells to attempt THz rotational transitions measurements with long 121 pathlengths (see Fig. 1, left part). Ppm and subppm LOD of strongly polar compounds were reached 122 in pulsed THz-TDS and in cw-THz spectroscopy with maximal THz path-lengths of 5 m [20] and 20 m 123 [21], respectively. 124



Figure 1. Picture/Scheme of the two White-type gas cells used for the CH₄ rotational spectra measurements. Left: standard White-type cell (Infrared Analysis, 35-V) used in the cw-THz photomixing measurement [21]. Right: White-type cell located in the AILES beamline of the SOLEIL synchrotron[17].

To date, the spectroscopic investigation of pure rotational THz transitions of CH₄ provides 125 probably the best results in terms of sensitivity obtained by coupling cw-THz sources with White-type 126 cells. Even if CH₄ is not polar, centrifugal distortion and vibrational effects induce a weak dipole (\simeq 127 $1 \mu D$) which allows transitions between rotational energy levels. So very weak rotational absorptions in 128 the THz domain may be measured with an optimized level of sensitivity. Using the HITRAN database 129 and its graphical tool "HITRAN on the web" [22,23], the rotational absorbance of pure methane has 130 been modelled in Fig. 2 (main panel) in the 2.0 - 2.7 THz frequency range for a THz path length of 131 20 m and a pressure of 10 mbar. The THz spectroscopy of CH_4 requires the measurements of very 132 weak rotational absorptions lower than 2×10^{-5} cm⁻¹ in a THz frequency region difficult to access. 133 FT-Far-IR synchrotron based spectroscopy [18,19] and cw-THz by photomixing [12] succeeded to 134 measure these weak rotational lines (see insets Fig. 2) with 150 m and 20 m pathlengths, respectively. 135 The White type multi-pass gas cells shown in Fig. 1 were used to reach these long interaction distances 136 and to increase the THz absorbances. While the FT-Far-IR rotational lines are limited by the maximal 137 resolution of the high-resolution interferometer (around 30 MHz), the THz CH₄ lines measured by 138 cw-THz spectroscopy are fully resolved and Doppler limited. These THz frequencies resulting from 139 the gaussian fit of the low-pressure line profile have improved the accuracies of the ground state and 140 the $v_4 \leftarrow v_4$ hot band molecular parameters allowing a better global modelling of the CH₄ emission 141 and absorption used to determine the molecular abundance of this major GHG in Earth's atmosphere 142 and also in various (exo)planetary upper atmospheres [18,26]. 143



Figure 2. Central part: Absorption of ground state rotational THz lines of CH₄ induced by centrifugal distortion simulated for a pathlength of 20 m. and a pressure of 10 mbar. Left and right parts: two examples of experimental lines measured at 30 MHz resolution, P = 10 mbar and L = 150 m by FT-FIR spectroscopy using the AILES beamline of the SOLEIL synchrotron (blue solid line) and at 30 kHz resolution, P = 1.5 mbar and L = 20 m with the cw-THz photomixing spectrometer [12]. Calculations were performed using the STDS software [24], which is par of the XTDS [25] package that implements the tensorial formalism developed for spherical-top molecules in the Dijon group.

3.2. Trace gases THz measurements of N_2O and O_3 in an atmospheric simulation chamber equipped with a Chernin type multi-pass cell

A ppm LOD is required to monitor CH_4 in our atmosphere. Due to the weak intensities of pure 146 rotational lines, THz spectroscopic monitoring of methane at atmospheric concentrations would prove 147 serious difficulties even after a strong improvement of the spectrometers sensitivity. On the contrary, 148 the THz trace gas detection of polar GHGs already proved its capability by the monitoring of N₂O and 149 O_3 in the stratosphere with the MLS and TALIS sounders [8,9]. These THz heterodyne receivers are 150 able to determine atmospheric subppm concentration profiles in the submm- and in the mm-wave 151 domains. A possibility to get close to the detection levels of N₂O and O₃ is demonstrated in this study 152 by performing the first THz absorption measurements in an atmospheric simulation chamber. 153

Compared to CH_4 , rather strong rotational transitions of N_2O and O_3 , with intensities 154 around 10^{-22} cm⁻¹/(molecule cm⁻²) [22] were probed in the submm-wave domain using versatile 155 spectrometers based on AMC [27]. Nevertheless, with increasing wavelengths, the THz beam 156 divergence rises and long path THz absorption experiments becomes challenging. Recently, promising 157 results have been obtained with compact circular multi-pass cells able to refocus the beam at each 158 internal reflection:(i) in Ref. [28], Rothbart et al., using a AMC, have measured traces of acetaldehyde in 159 methanol down to 110 ppm with an optical path of 1.9 m in the 250 GHz region; (ii) in Ref. [11], Kim et 160 *al.*, using THz-TDS, have recorded the pure rotational spectrum of N_2O at atmospheric pressure with a 161 dilution of 1% by reaching a THz path of 18.61 m. Even if the results of these studies are interesting 162 since they were obtained without pre-concentration in a very compact and easily transportable gas cell, 163 the LOD reached is clearly not sufficient for trace gas detection in the atmosphere. Yet it is possible to reach lower LOD with longer THz path-lenghts using a multi-pass cell based on the clever Chernin 165 arrangement [29]. Compared to the White type arrangement discussed in the previous section, a 166 Chernin multi-pass cell optimizes the recirculating of the beam over many focused lines on the field 167 mirrors [30]. In particular longer wavelengths are very critical due to the overlapping between adjacent 168 refocusing points within the cell what needs to be avoided or at least limited in order to prevent the 169 presence of stationary waves yielding to baseline variations. Moreover the optical path-length in the 170 Chernin arrangement is easy to adjust with a matrix distribution of the refocusing points and a variable 171 number of rows and columns. As shown in Fig.3, we have dimensionned a Chernin multi-pass cell in 172 order to integrate it in CHARME (Chamber for the Atmospheric Reactivity and the Metrology of the 173 Environment), a 9.2 m³, evacuable cylinder used to simulate the physical-chemistry of the atmosphere 174 in a controlled environment [31]. The Chernin type optical setup called hereafter "MultiCHARME" 175 is composed of 5 mirrors (2 rectangular field mirrors and 3 spherical objective mirrors) each with 176

5 m radius of curvature corresponding to the 5 m baselength of CHARME. All mirror holders are

equipped with computer-controlled micrometric screws for optical adjustments and path length

changes. MultiCHARME was designed to be coupled with different spectrometers covering a very

¹⁸⁰ large spectral range from visible to sub-mm wavelengths.



Figure 3. Left: Scheme of the MULTICHARME Chernin cell with pictures, inside CHARME, of the 6×6 arrangement highlighted with a He-Ne laser on the 2 fields and 3 objective mirrors. Right: external view of CHARME: CHamber for Atmospheric Reactivity and Metrology of the Environment located in the LPCA, Dunkirk, France.

N₂O was chosen as a test molecule to characterize the performances of MultiCHARME on three 181 frequency decades. In particular, the linearity of the absorption was checked with rovibrational 182 measurements up to 480 m in the near-IR and with rotational measurements up to 240 m in the THz 183 domain [32]. In Figure 4, a rotational absorbance of 32 % at 577.58 GHz (P=0.7 mbar) is obtained by 184 measuring in CHARME 400 ppm of residual N₂O traces with a THz path-length adjusted to 200 m 185 (configuration 4×5 on the field mirror). Measurements of ozone at a trace level of 200 ppm were 186 also performed in the same configuration with an absorbance of 25 % at 577.58 GHz (P=1 mbar). We 187 benefited from the possibility to simultaneously modulate the amplitude and frequency of the AMC. It 188 was of particular utility to minimise the effects of the standing waves disturbing the baseline spectra 189 measured in MultiCHARME. A rapid frequency modulation was applied with a depth corresponding 190 to the FSR of the interaction length. The absorption profiles of the targeted lines displayed a collisional 191 broadening in excess of this FSR. The profiles were measured using a significantly slower amplitude 192 modulation and a lock-in amplifier to extract the correct modulation frequency. In the right part of fig.4, 193 the absorbances of N₂O and O₃, respectively at typical tropospheric and stratospheric concentrations 194 have been simulated with HITRAN on the web [23]. Absorption levels lower than $1.0 \times 10^{-8} \text{ cm}^{-1}$ and 195 2.5×10^{-7} cm⁻¹ should be reached to measure, respectively, tropospheric N₂O and stratospheric O₃. 196 With a THz path-length of 200 m, the associated absorbances are respectively 2000 times and 50 times 197 weaker that those measured with MultiCHARME. The capability to monitor N₂O and O₃ with typical 198 atmospheric concentrations depend now on a correct modelisation of the subsisting baseline variations 199 due to long path FP effects. By removing these baseline oscillations, we can reasonably hope to reach 200 sufficient SNR to monitor O_3 in CHARME at stratospheric concentration with the same setup. The 201 measurement of tropospheric N_2O requires to improve by at least two orders of magnitude the LOD. 202 Finally, we can notice that the LOD actually reached using THz spectroscopy in MultiCHARME is 203 already sufficient to monitor N₂O collected in specific polluted atmosphere such as dental office where 204 the waste anesthetic N₂O exposure could reach several hundreds of ppm without adapted extractors 205 [33]. 206



Figure 4. Left: Simulation of the rotational absorption in the 575-625 frequency region of N₂O and O₃ at tropospheric (330 ppb) and stratospheric (7.5 ppm) concentrations, respectively, for a THz pathlength of 200 m. Right: 400 ppm of N₂O at P = 0.7 mbar and 200 ppm of O₃ at P = 1.0 mbar measured in CHARME with the MultiCHARME setup in a 4 × 5 Chernin configuration providing a THz pathlength of 200 m (black open circles). Both lines have been measured with an amplitude modulated at 4.5 kHz. An additional frequency modulation at 50 kHz with a 1.1 MHz depth was added to limit the baseline oscillations. The red solid lines correspond to Voigt profiles simulated with HITRAN2016 tabulated frequencies, linewidths and intensities.

207 3.3. Intra-cavity THz measurements of CF₄

In the previous sections, we have shown that multi-pass cells provide higher sensitivity for the 208 measurements of weak rotational resonances or small number of absorbing molecules leading to THz 209 line with low intensities. But these cells are limited by a significant attenuation of the input THz power 210 and generally require large volumes to reach distances exceeding 100 m. An alternative approach 211 is to adapt the intra-cavity techniques developed in the IR domain to the longer wavelengths of the 212 THz/submm-wave spectral domains. In 2019, our group challenges this feat by developing a THz 213 resonator based on a low-loss oversized corrugated waveguide closed with highly reflective photonic 214 mirrors made from silicon discs [34]. The length of the cavity is finely adjusted using piezo-electric 215 actuators to move the mirrors by up to 250 μ m, a control loop locks the cavity length to the THz 216 frequency. With Fabry-Perot THz absorption spectroscopy (FP-TAS), we were able to perform gas 217 phase measurements with a finesse better than 3000 in the 620 GHz frequency range. With such 218 high-finesse, an equivalent interaction length of 1 km is estimated with a cavity of 48 cm long. It was 219 demonstrated by measurements on a low abundant OCS isotopologue that FP-TAS is clearly able to 220 measure transitions with intensities of about 10^{-27} cm⁻¹/(molec cm⁻²). A scheme and a picture of 221 the THz FP cavity is given in Fig. 5. We can notice that the quality of the cavity should be optimised 222 and must remain stable to perform quantification. The measured signal is not directly the molecular 223 absorption it is the second harmonic of a frequency modulated THz. It indicates the strength and 224 width of the cavity mode as it scans across the molecular line. The absorption by the molecule causes a 225 reduction in the signal allowing the line centre to be reliably determined and quantification undertaken 226 for a given cavity configuration. For all these reasons, a step of calibration is still required. 227



Figure 5. Left: Scheme of the Fabry-Perot THz cavity based on a low-loss oversized corrugated waveguide with highly reflective photonic mirrors. Successive cavity modes are shown inside the cavity scheme and an achieved finesse above 3500 in the 626-629 GHz frequency range is highlighted ($L_{eff.} > 1.1 \text{ km}$). Right: Picture of the FP-THz absorption spectrometer with some legends.

In the present article, we present new measurements performed with this cavity on CF_4 a very 228 stable GHG, with a high greenhouse warming potential (> 6500) [2]. As for CH₄, CF₄ is not polar 229 and only rotational transitions induced by centrifugal distortion or vibrational anharmonicity may be 230 observed. Compared to CH_4 , the heavier CF_4 molecule exhibits ground state pure rotational transitions 231 at lower frequencies with a maximum intensity around 650 GHz. According to ab initio parameters 232 [35], the distortion-induced dipole is estimated to only 0.324 μ D. Therefore, all the THz lines are very 233 weak with intensities never exceeding 10^{-27} cm⁻¹/(molec cm⁻²). At room temperature, rotational 234 transitions belonging to the hot band $v_3 = 1$ are slightly more intense compared to the ground state 235 transitions. Rotational clusters in the R branch of this band were measured by FT-THz spectroscopy 236 based on synchrotron source with a path-length of 150 m in the White-type cell shown in Fig. 1 (right 237 part) [36]. In this study, the observation of CF₄ rotational clusters has required to co-add more than 238 5000 spectra measured at middle resolution (0.01 cm^{-1}) and pressure-broadened at P=100 mbar. Due 239 to a lack of sensitivity and resolution, it was not possible to measure at lower pressure and to fully 240 resolve the tetrahedral splitting of CF_4 rotational lines. Moreover, the ground state transitions with 241 absorbances lower than 1% (Fig. 6, upper panel) were not observed. The first measurements of the 242 tetrahedral split components of distortion-induced rotational lines of CF₄ are presented in this work 243 with the FP-TAS measurements performed in the 625 GHz region. The middle panel of Fig. 6 targets the 244 frequency range accessible by FP-TAS. The THz spectrum of CF_4 is simulated at a resolution of 100 kHz 245 with fully resolved and Doppler limited rotational lines at $P=100 \ \mu$ bar. With a 1 km path-length, 246 absorbances above 1 % are reached for several transitions in the ground state and in the $v_3 = 1$ hot 247 band. Finally, four examples of measured lines by FP-TAS in the R(20) cluster are presented in the 248 lower panel. The lines of the multiplet were measured individually with finesses between 2800 to 3250 249 and pressures between 80 μ bar and 150 μ bar. The frequency step was fixed to few kHz and the time 250 constant to 200 ms. The SNR for the four lines of Fig.6 are measured between 20 to 120, which shows 251 that the measurement of very weak THz line intensities below 10^{-28} cm⁻¹/(molec cm⁻²) may be for 252 now considered in THz spectroscopy. Considering this level of intensity now reachable, we can expect 253 trace gas detection of N₂O and O₃ (Fig. 4) by FP-TAS in the THz cavity at a subppm LOD, enough for 254 a detection at atmospheric concentrations. 255



Figure 6. Upper panel: Simulation of the CF₄ THz absorbance in the conditions of the synchrotron-based FT-Far-IR measurements (see Ref.[36]), namely a resolution of 0.01 cm⁻¹, a gas pressure of 100 mbar and a path-length of 150 m. Middle panel: Simulation of the CF₄ THz absorbance measured in the frequency range accessible by FP-TAS at a resolution of 100 kHz with a gas pressure of 100 μ bar and an effective path-length of 1 km. For both simulations, ground state transitions are in blue, $\nu_3 = 1$ hot band transitions are in red. Lower panel: Zoom on experimental individual rotational lines (*R*(20) cluster in the $\nu_3 \leftarrow \nu_3$ hot band) measured by FP-TAS. The experimental THz frequencies are determined by the microwave synthesiser used to drive the amplified multiplier chain. The microwave synthesiser is referenced to a time signal provided by a GPS receiver. The calculated frequencies are determined using the STDS software [24]. Dashed red line correspond to the FT-Far-IR synchrotron based measurements at a pressure of 100 mbar and a resolution of 0.01 cm⁻¹.

4. Conclusions

The present article highlights the best performances reached by rotational submm-wave/THz 257 long-path absorption spectroscopy of important atmospheric GHGs. The results obtained on CH_4 and 258 CF₄ demonstrate that the weakly intense centrifugal distortion-induced rotational lines are measurable 259 with a high degree of accuracy thanks to the progresses of THz gas phase high-resolution spectroscopy. 260 Several percents of methane absorbances with a path-length of 20 m may be measured above 2 THz 261 by a cw-THz photomixing source coupled to a White-type multi-pass cell. Thanks to a frequency 262 metrology based on a frequency comb, the rotational line centers are measured with an accuracy 263 competitive with those of electronic synthesizers and few cw-THz measurements are able to improve 264 the ground state low-order molecular parameters of the CH_4 spherical top [12]. The study of CH_4 has 265 been focused on THz line position analysis but absolute intensities measurements are required for 266 quantitative spectrosocopy of methane. Such analysis requires the improvement of the spectrometer 267 sensitivity. For this purpose, photomixers using a metallic mirror-based FP cavity could be tested at 268 short term in order to increase significantly the power of the source [37]. Next a new generation of 269 spectrometers has to be imagined allowing technological breakthroughs in the THz frequency gap. The 270 use of cavity enhanced techniques initially developed in the IR hold great promise, as demonstrated 271 by the results on CF₄ presented in this study where the tetrahedral splitting of CF₄ rotational lines 272 have been measured and resolved by FP-TAS in the submm-wave domain. This technique based on a 273 high finesse cavity allows to reach km effective path-length in a compact gas cell. According to the 274 CF₄ lines measured in Fig. 6 (lower panel), FP-TAS with a finesse >3200 should be able to measure rotational lines with SNR>3 for transitions with minimal intensities up to 10^{-28} cm⁻¹/(molec.cm⁻²). 276 The results presented in this article constitute a first step of a larger study of the CF₄ pure rotational 277 spectrum by FP-TAS: numerous lines in the 600-650 GHz frequency range (see Fig. 6, middle panel) 278 with absorbances > 1 % have to be measured, especially the ground state lines. The measured line 279 frequencies will be included in a global fit of the CF₄ ground state and $v_3 = 1$ molecular parameters in order to update the TFMeCaSDa database [38]. No doubt that the THz spectroscopy studies of 281 GHGs such as CH₄ and CF₄ have to be continued. In addition, new spectroscopic data may be also 282 obtained by considering the asymmetric isotopologues of non polar GHGs. In particular, with the 283 high sensitivity offered by our THz cavity, it will probably possible to measure the R(27) and R(28) 284 rotational transitions of asymmetric ${}^{16}O{}^{12}C{}^{18}O$ and to improve its ground state constants fitted only with low *J* value cm- and mm-wave transitions[39]. 286

On the other hand, the trace gas detection of non or weakly polar GHGs at atmospheric 287 concentration can not be seriously considered. In the case of the polar stable N_2O and unstable 288 O_3 GHG the situation is different since we demonstrated our ability to measure these molecules 289 at trace levels. For the first time, THz spectroscopy was used for a direct monitoring of molecular 290 species in an atmospheric simulation chamber. Thanks to a specially designed Chernin multi-pass cell 291 allowing to adjust a THz pathlength from 120 m to 240 m, quantitative rotational measurements of 292 N_2O and O_3 traces were performed. Right now, the accessible detection levels for both compounds 293 are limited to tens of ppm. We are working on a correct baseline modelisation in order to remove 294 its variations due to multiple interfering stationary waves in the Chernin cell. Anyway, the first 295 results obtained in CHARME with the coupling of a THz source to the MultiCHARME setup opens 296 new possibilities especially for the monitoring of stratospheric reaction processes at low-pressure. In 297 particular, the versatility of the electronic sources will allows to perform time-resolved quantitative 298 spectroscopies of reactants, oxidants and products involved in targeted reactions occurring in the 299 high altitude atmospheric layers. Moreover, it will be interesting to measure in CHARME, at different 300 pressures, the THz self-continuum of water absorption with the AMC coupled to MultiCHARME and 301 to compare our spectra with those recently performed with the coherent THz synchrotron radiation coupled to the White-type cell shown in Fig. 1 (right part) [40]. Finally, the high-finesse THz cavity 303 has also to be employed for the detection of stable N_2O and unstable O_3 GHG. Next studies will 304

³⁰⁵ be dedicated to verify if N_2O and O_3 submm-wave transitions could be measured at tro ³⁰⁶ (330 ppb) and stratospheric (7500 ppb) concentrations by FP-TAS.

Author Contributions: conceptualization, A.C. and V.B.; methodology, A.C., F.H., G.M., R.B., J.D., E.F. and V.B.;
software, F.H.; validation, A.C. and V.B.; investigation, C.B., J.B., J.D., E.F, A.P. and A.C.; formal analysis, C.B., J.B.,
J.D., A.P. and V.B.; writing–original draft preparation, A.C.; writing–review and editing, F.H., R.B., J.B. and V.B.;
supervision, A.C., G.M. and F.H.; project administration, A.C., V.B. and G.M.; funding acquisition, A.C., G.M. and
E.F.

Funding: This work was supported by the CaPPA project (Chemical and Physical Properties of the 312 Atmosphere) funded by the French National Research Agency (ANR-11-LABX-0005-01) and the CLIMIBIO 313 program supported by the Hauts-de-France Regional Council, the French Ministry of Higher Education and 314 Research and the European Regional Development Fund. C.B. and J.D. were also funded by CLIMIBIO. 315 The photomixing cw-THz spectrometer has been funded both by the Délégation Générale pour l'Armement 316 (DGA, ProjetANR-11-ASTR-0035), the Institut de Recherche en Environnement Industriel (IRENI) and the 317 European Commission via the Interreg IVA-2seas (Cleantech Project). MultiCHARME was founded by the 318 University of Littoral and its Research Quality Bonus. The THz cavity and A.P. were funded by the Satt-Nord 319 (project Teraspec-M0407), the French ANR (ANR-15-CE29-0017) and the European Regional Development Fund 320 (ÎNTERREG Ŷ FR-WA-VL 1.2.11). 321

Acknowledgments: We thank: (i) the AILES beamline team for the synchrotron based FT-THz experiments, especially the beamline manager P. Roy and O. Pirali in charge of the FT-FIR measurements in the White type cell; (ii) W. Zhao and B. Fang for the development of the Chernin cell; (iii) C. Coeur, N. Houzel and P. Kulinski for their assistance during the measurements in the CHARME chamber.

326 Conflicts of Interest: The authors declare no conflict of interest.

327 Abbreviations

³²⁸ The following abbreviations are used in this manuscript:

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ACE	Atmospheric Chemistry Experiment
AILES	Advanced Infrared beamLine Exploited for Spectroscopy
AMC	Amplified Multiplier Chains
CFCs/HCFCs	ChloroFluoroCarbons
CHARME	CHamber for the Atmospheric Reactivity and the Metrology of the Environment
cw-THz	continuous-wave Terahertz
FFT	Fast Fourier Transform
FP-TAS	Fabry-Perot Terahertz Absorption Spectroscopy
FSR	Free Spectral Range
FTS	Fourier Transform Spectrometer
GHG	Greenhouse gas
HITRAN	HIgh-resolution TRANsmission molecular absorption database
IASI	Infrared Atmospheric Sounding Interferometer
IR	InfraRed
LOD	Level of Detection
LTG-AsGa	Low-Temperature Grown - Arsenide Gallium
MLS	Microwave Limb Sounder
NEP	Noise Equivalent Power
SCISAT	SCIence Satellite
SNR	Signal to Noise Ratio
SOLEIL	Source Optimisée de Lumière d'Énergie Intermédiaire du LURE
STDS	Spherical-Top Data System
TALIS	THz Atmospheric Limb Sounder
TES	Tropospheric Emission Spectrometer
TFMeCaSDa	TetraFluoroMethane Calculated Spectroscopic Databases
THz-TDS	Terahertz-Time Domain Spectroscopy
LO	Local Oscillator
UV	Ultraviolet
ZBD	Zero Bias Detector

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