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Article

# Terahertz rotational spectroscopy of greenhouse gases using long interaction path-lengths

Arnaud Cuisset <sup>1,\*</sup> , Francis Hindle <sup>1</sup> , Gaël Mouret <sup>1</sup> , Robin Bocquet <sup>1</sup> , Jonas Bruckhuisen <sup>1</sup> , Jean Decker <sup>1</sup> , Anastasiia Pienkina <sup>1,2</sup> , Cédric Bray <sup>1,†</sup> , Éric Fertein <sup>1</sup>  and Vincent Boudon <sup>3,\*</sup> 

<sup>1</sup> Université du Littoral Côte d'Opale, UR 4493, LPCA, Laboratoire de Physico-Chimie de l'Atmosphère, F-59140 Dunkerque, France; cuisset@univ-littoral.fr(A.C.); hindle@univ-littoral.fr(F.H.); mouret@univ-littoral.fr(G.M.); bocquet@univ-littoral.fr(R.B.); jonas.bruckhuisen@univ-littoral.fr(J.B.); jean.decker@univ-littoral.fr(J.D.); fertein@univ-littoral.fr(E.F.)

<sup>2</sup> SATT-Nord, 25, Avenue Charles Saint-Venant, 59800 Lille, France; Anastasiia.Pienkina@satt nord.fr(A.P)

<sup>3</sup> Laboratoire Interdisciplinaire Carnot de Bourgogne, UMR 6303 CNRS, Université de Bourgogne Franche-Comté, 9 Av. A. Savary BP47870, F-21078, Dijon Cedex, France; vincent.boudon@u-bourgogne.fr(V.B.)

\* Correspondence: cuisset@univ-littoral.fr(A.C.);vincent.boudon@u-bourgogne.fr(V.B.)

† Laboratoire Charles Coulomb, UMR 5221 CNRS, Université de Montpellier, Campus Triolet, Place Eugène Bataillon, F-34095 Montpellier Cedex 5, France; cedric.bray@umontpellier.fr(C.B.)

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**Abstract:** Even if on-board mm-wave/THz heterodyne receivers have been developed to measure greenhouse gases (GHGs) atmospheric profiles, rotational spectroscopy rests under-exploited for their monitoring unlike IR rovibrational spectroscopy. The present study deals with the ability of THz spectroscopy using long interaction path-lengths for GHG laboratory investigations. High-resolution THz signatures of non-polar greenhouse molecules may be observed by probing very weak centrifugal distortion induced rotational transitions. To illustrate, new measurements on CH<sub>4</sub> and CF<sub>4</sub> have been carried out. For CH<sub>4</sub>, pure rotational transitions, recorded by cw-THz photomixing up to 2.6 THz in a White type cell adjusted to 20 m, have allowed to update the methane line list of atmospheric databases. Concerning CF<sub>4</sub>, Fabry-Perot THz absorption spectroscopy with a km effective pathlength was required to detect line intensities lower than 10<sup>-27</sup> cm<sup>-1</sup>/(molec cm<sup>-2</sup>). Contrary to previous synchrotron based FT-FIR measurements, the tetrahedral splitting of CF<sub>4</sub> THz lines is fully resolved. Finally, quantitative measurements of N<sub>2</sub>O and O<sub>3</sub> gas traces have been performed in an atmospheric simulation chamber using a submm-wave amplified multiplier chain coupled to a Chernin type multi-pass cell on a 200 m path-length. The THz monitoring of these two polar GHGs at tropospheric and stratospheric concentrations may be now considered.

**Keywords:** greenhouse gases; terahertz; long interaction path-lengths; rotational spectroscopy; trace gas monitoring.

## 1. Introduction

The main discussion on climate change tends to focus on carbon dioxide (CO<sub>2</sub>), the most dominant Greenhouse Gas (GHG) (65% of the global emission) produced by the burning of fossil fuels, industrial processes and changes in land use (deforestation, intensification of agriculture, flooding...) [1]. But CO<sub>2</sub> is not the only GHG driving the global climate change: methane (CH<sub>4</sub>), nitrous oxide (N<sub>2</sub>O), tropospheric ozone (O<sub>3</sub>) and fluorinated gases have a direct contribution to the greenhouse effect. In

particular, the radiative forcing (warming influence) of long-lived GHGs such as tetrafluoromethane ( $\text{CF}_4$ ) with a 50000 years lifetime in earth's atmosphere [2], is undergoing accelerating growth (almost a third of the increase attributed to the industrial age has been over the past 30 years) [3,4]. Several currently-operating satellite and air-borne missions aim to estimate emissions and absorptions of the main GHGs. By monitoring the column densities in our atmosphere, they accumulate new knowledge on their global distribution and temporal variation. Fourier Transform Spectrometers (FTS) are the most popular instruments used in these missions analysing the IR radiations reflected from the Earth's surface and emitted from the atmosphere and the surface. We can mention: (i) the IASI interferometer in the European MetOp satellite which monitors  $\text{CO}_2$ ,  $\text{O}_3$ ,  $\text{CH}_4$ ,  $\text{N}_2\text{O}$  GHG column amounts in several IR channels of detection [5]; (ii) ACE on board the Canadian Satellite SCISAT where  $\text{CF}_4$  and others CFCs/HFCs are retrieved routinely in addition to the most abundant GHGs [6]; (iii) TES on board the Aura spacecraft which retrieves the  $\text{O}_3$  and  $\text{CH}_4$  profiles in the lower atmosphere from their rovibrational signatures [7]. Amongst the Aura's instruments, there is also a passive microwave limb-sounding (MLS) radiometer/spectrometer used to measure the pollution in the upper troposphere in the presence of ice clouds and volcanic aerosols, which prevent measurements by IR/UV techniques. The MLS instrument monitors  $\text{O}_3$  and  $\text{N}_2\text{O}$  by measuring pure rotational lines in the mm-wave (240 GHz) and in the submm-wave (640 GHz) domains, respectively [8]. More recently, the National Space Science Center of the Chinese Academy of Sciences has developed a THz Atmospheric Limb Sounder (TALIS) for atmospheric vertical resolved profile observations: as for MLS/Aura,  $\text{O}_3$  and  $\text{N}_2\text{O}$  are the targeted GHGs and their rotational transitions are probed with four heterodyne radiometers (LO frequencies: 118 GHz, 190 GHz, 240 GHz and 643 GHz) and several FFT spectrometers of 2 GHz bandwidth with 2 MHz resolution [9]. With this instrument, they are able to retrieve the  $\text{N}_2\text{O}$  and  $\text{O}_3$  profiles with average precisions lower than 10 % and 5 % from 10 to 42 kms and from 10 to 70 kms altitude, respectively.

Accurate THz laboratory measurements are required in order to provide the best set of GHGs rotational line parameters (frequencies, widths, intensities...) and to optimize the inversion processes in GHGs atmospheric concentrations retrieval. The quality of the fits depends directly on the SNR of high resolution THz spectra and consequently on the sensitivity of the spectrometers [10]. According to the Beer-Lambert law, longer interaction THz path-lengths are required in the situation of a weak absorption. With longer wavelengths and larger beamsizes, the control of the propagation of THz radiations on long-range beam paths is challenging compared to that carried out in the IR/UV regions. Nevertheless, recent progress has allowed gas phase THz spectroscopy experiments with molecules/beam interaction distances from several tens to several hundreds of meters [11] to be performed. This article highlights recent advances of THz spectroscopy performed by our group, using long-interaction path-lengths for the measurement of weak GHG rotational absorptions. The ability to measure fully resolved centrifugally induced rotational transitions of non polar GHGs such as  $\text{CH}_4$  and  $\text{CF}_4$  will be demonstrated and the first THz measurements in an atmospheric simulation chamber will be presented towards the detection of polar stable  $\text{N}_2\text{O}$  and unstable  $\text{O}_3$  GHGs at trace levels.

## 2. Materials and Methods

With the exception for the production of ozone, commercially available high purity  $\text{CH}_4$ ,  $\text{N}_2\text{O}$ , and  $\text{CF}_4$  GHGs with natural isotopic abundances were used throughout. The ozone was produced by corona discharge generation with pure oxygen (Air Tree Ozone Technology, C-Lasky C-L010-DTI). Various THz sources and detectors have been used in the studies presented here. In particular, three commonly used approaches, optical, optoelectronic and electronic, have been employed for the generation of the THz radiation.

### 2.1. Synchrotron based FT THz spectroscopy

The AILES beamline of the SOLEIL synchrotron was used for high-resolution broadband FT measurements in the 0.6 to 19.5 THz, ( $20\text{-}650\text{ cm}^{-1}$ ) spectral range on  $\text{CH}_4$  and  $\text{CF}_4$ . The very

72 bright light extracted from the synchrotron bending magnet has the advantage of displaying a lower  
73 divergence compared with other sources aiding its use with long path interaction White type cells.  
74 The FT IFS125 spectrometer, used in this case, has a 5 m delay stage allowing a resolution of 30 MHz to  
75 be achieved [17].

## 76 2.2. *Cw-THz spectroscopy by photomixing*

77 Alternatively higher resolution measurements can be made using monochromatic THz sources.  
78 The generation of THz radiation by photomixing is a frequency down-conversion technique which  
79 was used for the measurements of individual rotational lines of CH<sub>4</sub> up to 2.6 THz [12]. Two extended  
80 cavity laser diodes operating around 780 nm are mixed together in a LTG-GaAs device. The lasers  
81 are detuned to the desired THz frequency inducing a corresponding current in the photomixer device  
82 which is coupled to an integrated log spiral antenna [21]. The radiation is pre-collimated by a silicon  
83 hyper-hemispherical lens and propagated in free space. This type of source is monochromatic with  
84 excellent spectral purity and can be used in the range of 100 GHz to 3.3 THz. The available power  
85 is limited to 0.1 nW from 2.5 THz. By locking onto a frequency comb, the generated THz frequency  
86 is determined with a kHz accuracy [13]. Both amplitude and frequency modulation schemes can be  
87 easily implemented with this source, it does however suffer from available low power levels at the  
88 highest frequencies.

## 89 2.3. *Submm-wave spectroscopy with amplified multiplier chains*

90 We have used two AMC from the Virginia Diodes Inc. as electronic sources for the submm-wave  
91 measurements coupled to CHARME (see section 3.2) and to the THz FP cavity (see section 3.3). They are  
92 based on the up-conversion of a synthesized microwave frequency using a cascade of Schottky diode  
93 frequency multipliers mounted in rectangular wave-guide blocks. The guided radiation is launched  
94 into free space using a horn antenna and propagated through the interaction cell. Frequencies from  
95 100 GHz to 1 THz can be conveniently produced and easily modulated in either amplitude or frequency  
96 as desired [27]. A typical power of 50  $\mu$ W is available at 600 GHz.

## 97 2.4. *Detection schemes*

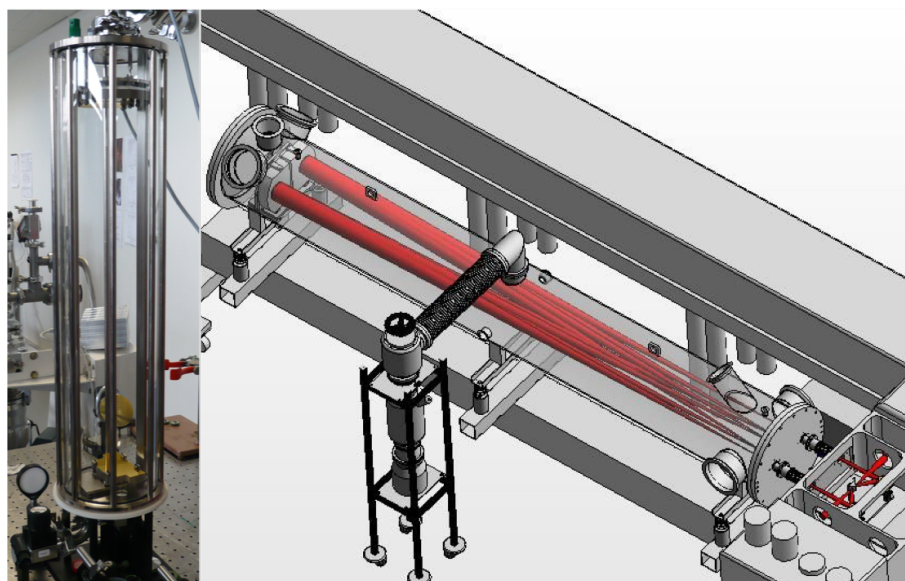
98 For all experiments, two types of detector have been employed. Zero Biased Detector (ZBD) is  
99 an unbiased Schottky diode mounted in a wave-guide operating in detection mode. This type of  
100 detector is uncooled and can provide a typical NEP of 10 pW/ $\sqrt{\text{Hz}}$ . When greater sensitivity was  
101 required a Helium cooled bolometer was used with a NEP of 1 pW/ $\sqrt{\text{Hz}}$ . For the measurements of  
102 very weakly intense CF<sub>4</sub> lines, we have coupled a very high finesse FP cavity with an AMC in order to  
103 reach very long equivalent interaction paths (see section 3.3). A control loop is used to lock the cavity  
104 to the frequency of the AMC with a frequency modulation providing an error signal. The second  
105 harmonic of the frequency modulation is simultaneously exploited to detect the presence of molecular  
106 absorption present in the cavity. The AMC frequency and cavity are thus swept together allowing  
107 sensitive absorption measurements to be made [34].

## 108 3. Results and discussion

### 109 3.1. *THz measurements of CH<sub>4</sub> in a White type multi-pass cell*

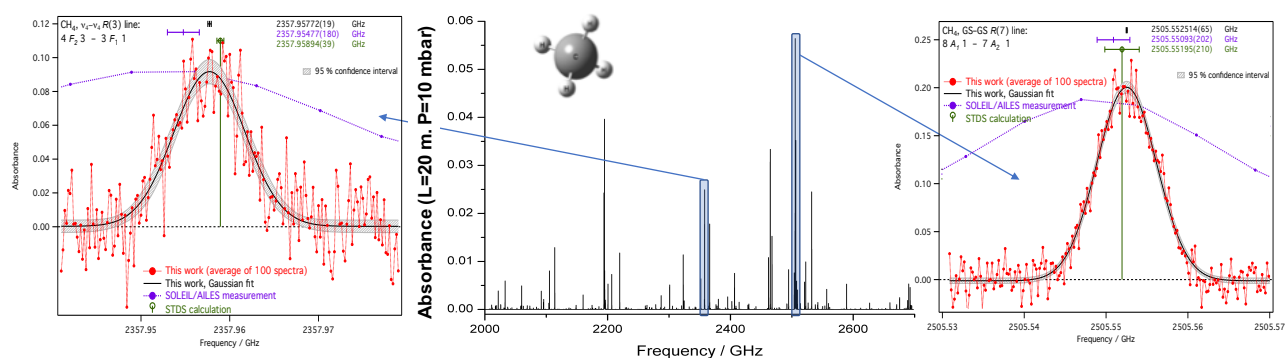
110 THz long-path spectroscopic experiments were initially conducted in the 2000's using home-made  
111 or commercial White type gas cells. The White cell was first described in 1942 by John U. White [14]  
112 and was a significant improvement over previous long path measurement techniques used in optical  
113 spectroscopy. A White cell is constructed using three spherical, concave mirrors having the same  
114 radius of curvature. The mirrors are separated by a distance equal to their radii of curvature. FT  
115 Michelson based interferometers have been connected by different groups to White type gas cells

116 allowing to improve the detection sensitivity in the Far-IR domain. The THz pathlengths reached vary  
117 from several tens of meters with classical sources such as Hg lamps [15,16] up to 180 m with weakly  
118 divergent and high brightness Far-IR synchrotron radiations [17,18]. As example, the White type  
119 cell connected to the Bruker IFS125 high-resolution interferometer coupled to the AILES Far-IR/THz  
120 beamline of the SOLEIL synchrotron is shown in Fig. 1 (right part). Optoelectronic THz sources have  
121 been also coupled to White type cells to attempt THz rotational transitions measurements with long  
122 pathlengths (see Fig. 1, left part). Ppm and subppm LOD of strongly polar compounds were reached  
123 in pulsed THz-TDS and in cw-THz spectroscopy with maximal THz path-lengths of 5 m [20] and 20 m  
124 [21], respectively.



**Figure 1.** Picture/Scheme of the two White-type gas cells used for the CH<sub>4</sub> rotational spectra measurements. Left: standard White-type cell (Infrared Analysis, 35-V) used in the cw-THz photomixing measurement [21]. Right: White-type cell located in the AILES beamline of the SOLEIL synchrotron [17].

125 To date, the spectroscopic investigation of pure rotational THz transitions of CH<sub>4</sub> provides  
126 probably the best results in terms of sensitivity obtained by coupling cw-THz sources with White-type  
127 cells. Even if CH<sub>4</sub> is not polar, centrifugal distortion and vibrational effects induce a weak dipole ( $\approx$   
128  $1 \mu D$ ) which allows transitions between rotational energy levels. So very weak rotational absorptions in  
129 the THz domain may be measured with an optimized level of sensitivity. Using the HITRAN database  
130 and its graphical tool "HITRAN on the web" [22,23], the rotational absorbance of pure methane has  
131 been modelled in Fig. 2 (main panel) in the 2.0 - 2.7 THz frequency range for a THz path length of  
132 20 m and a pressure of 10 mbar. The THz spectroscopy of CH<sub>4</sub> requires the measurements of very  
133 weak rotational absorptions lower than  $2 \times 10^{-5} \text{ cm}^{-1}$  in a THz frequency region difficult to access.  
134 FT-Far-IR synchrotron based spectroscopy [18,19] and cw-THz by photomixing [12] succeeded to  
135 measure these weak rotational lines (see insets Fig. 2) with 150 m and 20 m pathlengths, respectively.  
136 The White type multi-pass gas cells shown in Fig. 1 were used to reach these long interaction distances  
137 and to increase the THz absorbances. While the FT-Far-IR rotational lines are limited by the maximal  
138 resolution of the high-resolution interferometer (around 30 MHz), the THz CH<sub>4</sub> lines measured by  
139 cw-THz spectroscopy are fully resolved and Doppler limited. These THz frequencies resulting from  
140 the gaussian fit of the low-pressure line profile have improved the accuracies of the ground state and  
141 the  $\nu_4 \leftarrow \nu_4$  hot band molecular parameters allowing a better global modelling of the CH<sub>4</sub> emission  
142 and absorption used to determine the molecular abundance of this major GHG in Earth's atmosphere  
143 and also in various (exo)planetary upper atmospheres [18,26].



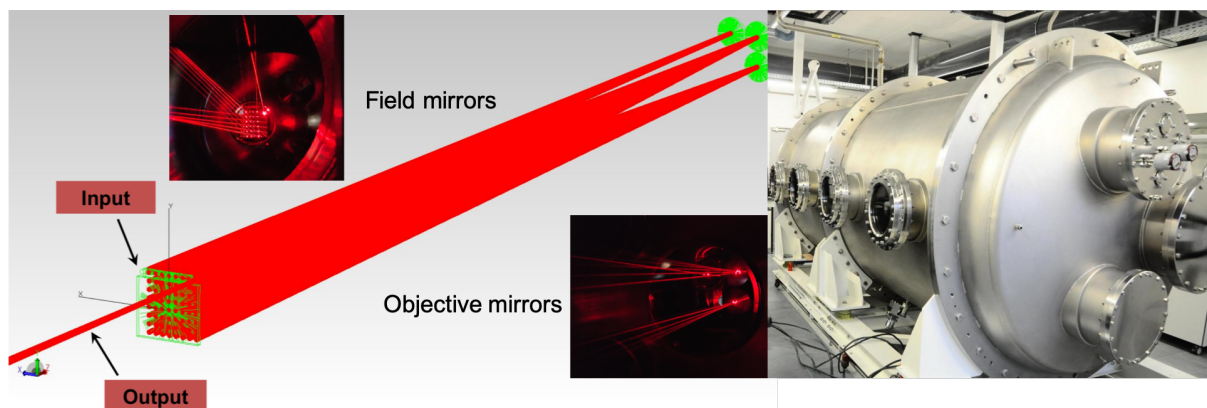
**Figure 2.** Central part: Absorption of ground state rotational THz lines of CH<sub>4</sub> induced by centrifugal distortion simulated for a pathlength of 20 m. and a pressure of 10 mbar. Left and right parts: two examples of experimental lines measured at 30 MHz resolution,  $P = 10$  mbar and  $L = 150$  m by FT-FIR spectroscopy using the AILES beamline of the SOLEIL synchrotron (blue solid line) and at 30 kHz resolution,  $P = 1.5$  mbar and  $L = 20$  m with the cw-THz photomixing spectrometer [12]. Calculations were performed using the STDS software [24], which is par of the XTDS [25] package that implements the tensorial formalism developed for spherical-top molecules in the Dijon group.

### 144 3.2. Trace gases THz measurements of N<sub>2</sub>O and O<sub>3</sub> in an atmospheric simulation chamber equipped with a 145 Chernin type multi-pass cell

146 A ppm LOD is required to monitor CH<sub>4</sub> in our atmosphere. Due to the weak intensities of pure  
147 rotational lines, THz spectroscopic monitoring of methane at atmospheric concentrations would prove  
148 serious difficulties even after a strong improvement of the spectrometers sensitivity. On the contrary,  
149 the THz trace gas detection of polar GHGs already proved its capability by the monitoring of N<sub>2</sub>O and  
150 O<sub>3</sub> in the stratosphere with the MLS and TALIS sounders [8,9]. These THz heterodyne receivers are  
151 able to determine atmospheric subppm concentration profiles in the submm- and in the mm-wave  
152 domains. A possibility to get close to the detection levels of N<sub>2</sub>O and O<sub>3</sub> is demonstrated in this study  
153 by performing the first THz absorption measurements in an atmospheric simulation chamber.

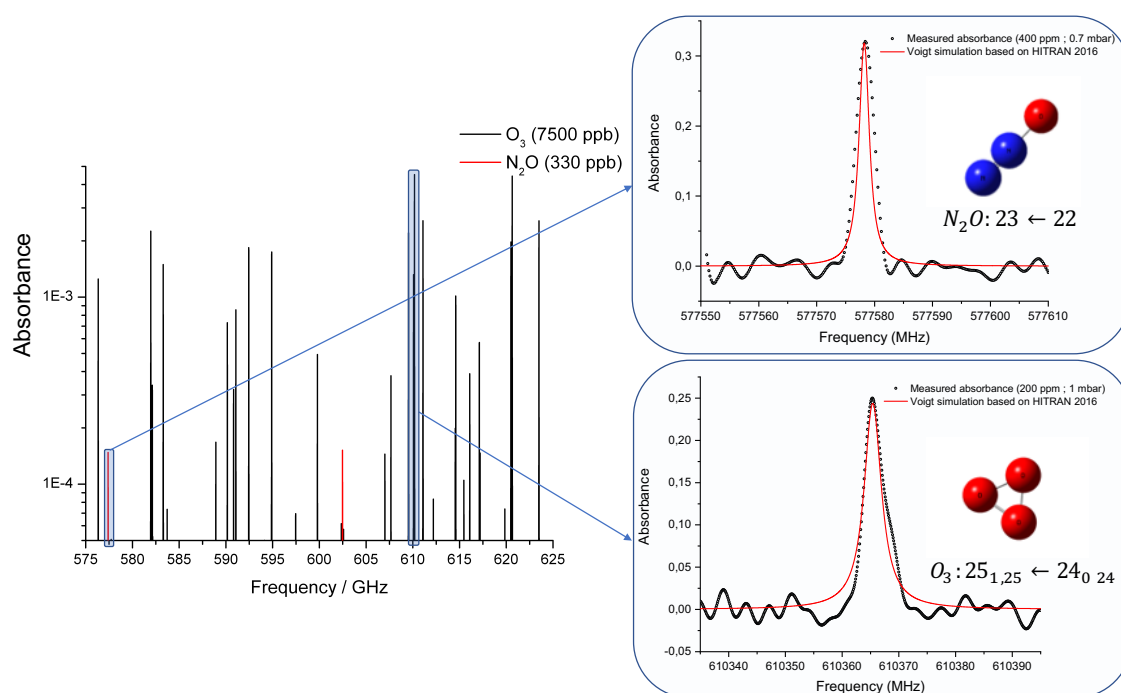
154 Compared to CH<sub>4</sub>, rather strong rotational transitions of N<sub>2</sub>O and O<sub>3</sub>, with intensities  
155 around  $10^{-22}$  cm<sup>-1</sup> / (molecule cm<sup>-2</sup>) [22] were probed in the submm-wave domain using versatile  
156 spectrometers based on AMC [27]. Nevertheless, with increasing wavelengths, the THz beam  
157 divergence rises and long path THz absorption experiments becomes challenging. Recently, promising  
158 results have been obtained with compact circular multi-pass cells able to refocus the beam at each  
159 internal reflection:(i) in Ref. [28], Rothbart *et al.*, using a AMC, have measured traces of acetaldehyde in  
160 methanol down to 110 ppm with an optical path of 1.9 m in the 250 GHz region; (ii) in Ref. [11], Kim *et*  
161 *al.*, using THz-TDS, have recorded the pure rotational spectrum of N<sub>2</sub>O at atmospheric pressure with a  
162 dilution of 1% by reaching a THz path of 18.61 m. Even if the results of these studies are interesting  
163 since they were obtained without pre-concentration in a very compact and easily transportable gas cell,  
164 the LOD reached is clearly not sufficient for trace gas detection in the atmosphere. Yet it is possible to  
165 reach lower LOD with longer THz path-lengths using a multi-pass cell based on the clever Chernin  
166 arrangement [29]. Compared to the White type arrangement discussed in the previous section, a  
167 Chernin multi-pass cell optimizes the recirculating of the beam over many focused lines on the field  
168 mirrors [30]. In particular longer wavelengths are very critical due to the overlapping between adjacent  
169 refocusing points within the cell what needs to be avoided or at least limited in order to prevent the  
170 presence of stationary waves yielding to baseline variations. Moreover the optical path-length in the  
171 Chernin arrangement is easy to adjust with a matrix distribution of the refocusing points and a variable  
172 number of rows and columns. As shown in Fig.3, we have dimensionned a Chernin multi-pass cell in  
173 order to integrate it in CHARME (Chamber for the Atmospheric Reactivity and the Metrology of the  
174 Environment), a 9.2 m<sup>3</sup>, evacuable cylinder used to simulate the physical-chemistry of the atmosphere  
175 in a controlled environment [31]. The Chernin type optical setup called hereafter "MultiCHARME"  
176 is composed of 5 mirrors (2 rectangular field mirrors and 3 spherical objective mirrors) each with

177 5 m radius of curvature corresponding to the 5 m baselength of CHARME. All mirror holders are  
178 equipped with computer-controlled micrometric screws for optical adjustments and path length  
179 changes. MultiCHARME was designed to be coupled with different spectrometers covering a very  
180 large spectral range from visible to sub-mm wavelengths.



**Figure 3.** Left: Scheme of the MULTICHARME Chernin cell with pictures, inside CHARME, of the  $6 \times 6$  arrangement highlighted with a He-Ne laser on the 2 fields and 3 objective mirrors. Right: external view of CHARME: CHamber for Atmospheric Reactivity and Metrology of the Environment located in the LPCA, Dunkirk, France.

181  $N_2O$  was chosen as a test molecule to characterize the performances of MultiCHARME on three  
182 frequency decades. In particular, the linearity of the absorption was checked with rovibrational  
183 measurements up to 480 m in the near-IR and with rotational measurements up to 240 m in the THz  
184 domain [32]. In Figure 4, a rotational absorbance of 32 % at 577.58 GHz ( $P=0.7$  mbar) is obtained by  
185 measuring in CHARME 400 ppm of residual  $N_2O$  traces with a THz path-length adjusted to 200 m  
186 (configuration  $4 \times 5$  on the field mirror). Measurements of ozone at a trace level of 200 ppm were  
187 also performed in the same configuration with an absorbance of 25 % at 577.58 GHz ( $P=1$  mbar). We  
188 benefited from the possibility to simultaneously modulate the amplitude and frequency of the AMC. It  
189 was of particular utility to minimise the effects of the standing waves disturbing the baseline spectra  
190 measured in MultiCHARME. A rapid frequency modulation was applied with a depth corresponding  
191 to the FSR of the interaction length. The absorption profiles of the targeted lines displayed a collisional  
192 broadening in excess of this FSR. The profiles were measured using a significantly slower amplitude  
193 modulation and a lock-in amplifier to extract the correct modulation frequency. In the right part of fig.4,  
194 the absorbances of  $N_2O$  and  $O_3$ , respectively at typical tropospheric and stratospheric concentrations  
195 have been simulated with HITRAN on the web [23]. Absorption levels lower than  $1.0 \times 10^{-8} \text{ cm}^{-1}$  and  
196  $2.5 \times 10^{-7} \text{ cm}^{-1}$  should be reached to measure, respectively, tropospheric  $N_2O$  and stratospheric  $O_3$ .  
197 With a THz path-length of 200 m, the associated absorbances are respectively 2000 times and 50 times  
198 weaker than those measured with MultiCHARME. The capability to monitor  $N_2O$  and  $O_3$  with typical  
199 atmospheric concentrations depend now on a correct modelisation of the subsisting baseline variations  
200 due to long path FP effects. By removing these baseline oscillations, we can reasonably hope to reach  
201 sufficient SNR to monitor  $O_3$  in CHARME at stratospheric concentration with the same setup. The  
202 measurement of tropospheric  $N_2O$  requires to improve by at least two orders of magnitude the LOD.  
203 Finally, we can notice that the LOD actually reached using THz spectroscopy in MultiCHARME is  
204 already sufficient to monitor  $N_2O$  collected in specific polluted atmosphere such as dental office where  
205 the waste anesthetic  $N_2O$  exposure could reach several hundreds of ppm without adapted extractors  
206 [33].

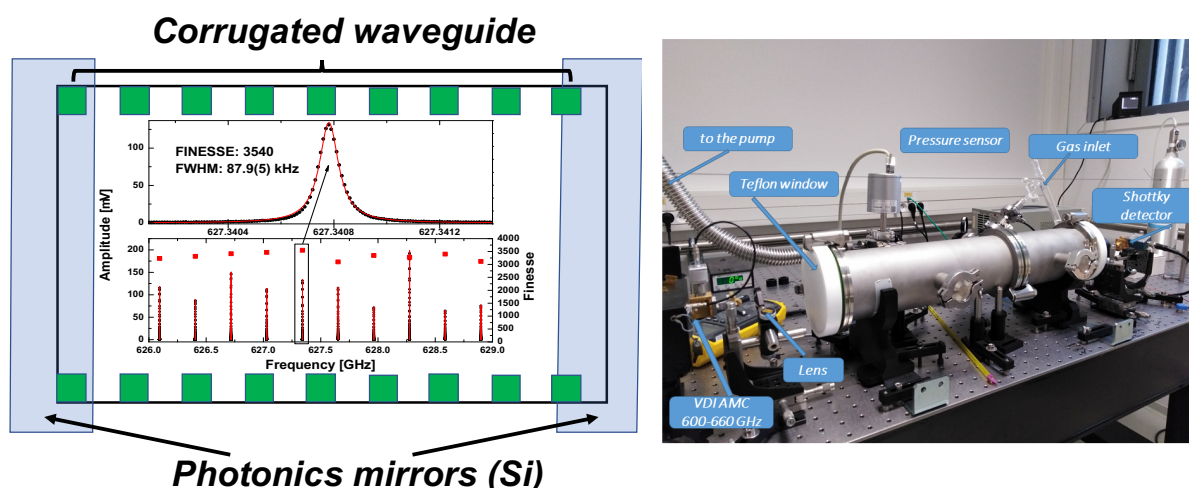


**Figure 4.** Left: Simulation of the rotational absorption in the 575-625 frequency region of  $N_2O$  and  $O_3$  at tropospheric (330 ppb) and stratospheric (7.5 ppm) concentrations, respectively, for a THz pathlength of 200 m. Right: 400 ppm of  $N_2O$  at  $P = 0.7$  mbar and 200 ppm of  $O_3$  at  $P = 1.0$  mbar measured in CHARME with the MultiCHARME setup in a  $4 \times 5$  Chernin configuration providing a THz pathlength of 200 m (black open circles). Both lines have been measured with an amplitude modulated at 4.5 kHz. An additional frequency modulation at 50 kHz with a 1.1 MHz depth was added to limit the baseline oscillations. The red solid lines correspond to Voigt profiles simulated with HITRAN2016 tabulated frequencies, linewidths and intensities.

### 207 3.3. Intra-cavity THz measurements of $CF_4$

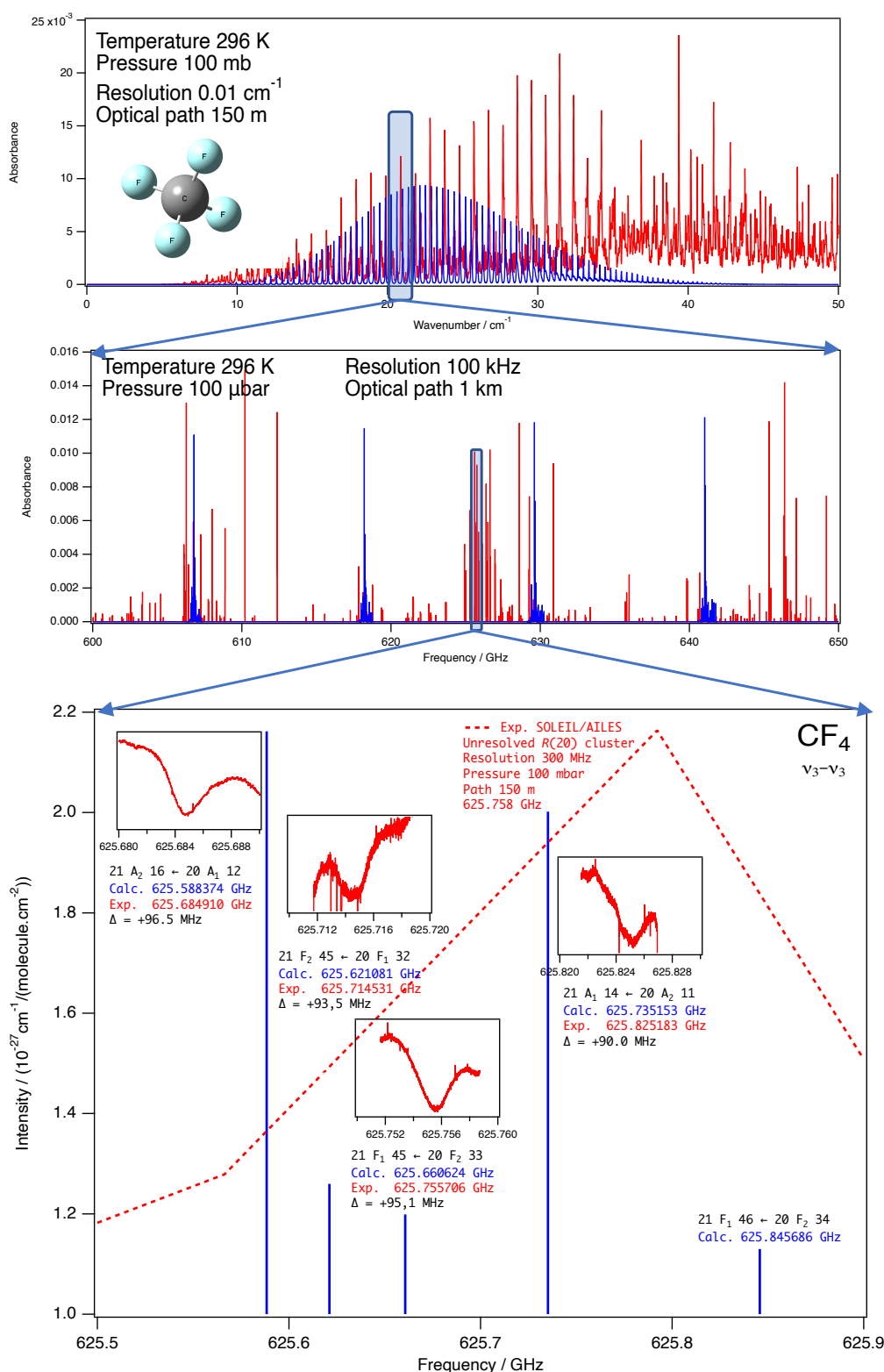
208 In the previous sections, we have shown that multi-pass cells provide higher sensitivity for the  
 209 measurements of weak rotational resonances or small number of absorbing molecules leading to THz  
 210 line with low intensities. But these cells are limited by a significant attenuation of the input THz power  
 211 and generally require large volumes to reach distances exceeding 100 m. An alternative approach  
 212 is to adapt the intra-cavity techniques developed in the IR domain to the longer wavelengths of the  
 213 THz/submm-wave spectral domains. In 2019, our group challenges this feat by developing a THz  
 214 resonator based on a low-loss oversized corrugated waveguide closed with highly reflective photonic  
 215 mirrors made from silicon discs [34]. The length of the cavity is finely adjusted using piezo-electric  
 216 actuators to move the mirrors by up to 250  $\mu m$ , a control loop locks the cavity length to the THz  
 217 frequency. With Fabry-Perot THz absorption spectroscopy (FP-TAS), we were able to perform gas  
 218 phase measurements with a finesse better than 3000 in the 620 GHz frequency range. With such  
 219 high-finesse, an equivalent interaction length of 1 km is estimated with a cavity of 48 cm long. It was  
 220 demonstrated by measurements on a low abundant OCS isotopologue that FP-TAS is clearly able to  
 221 measure transitions with intensities of about  $10^{-27} \text{ cm}^{-1} / (\text{molec cm}^{-2})$ . A scheme and a picture of  
 222 the THz FP cavity is given in Fig. 5. We can notice that the quality of the cavity should be optimised  
 223 and must remain stable to perform quantification. The measured signal is not directly the molecular  
 224 absorption it is the second harmonic of a frequency modulated THz. It indicates the strength and  
 225 width of the cavity mode as it scans across the molecular line. The absorption by the molecule causes a  
 226 reduction in the signal allowing the line centre to be reliably determined and quantification undertaken  
 227 for a given cavity configuration. For all these reasons, a step of calibration is still required.





**Figure 5.** Left: Scheme of the Fabry-Perot THz cavity based on a low-loss oversized corrugated waveguide with highly reflective photonic mirrors. Successive cavity modes are shown inside the cavity scheme and an achieved finesse above 3500 in the 626-629 GHz frequency range is highlighted ( $L_{eff} > 1.1$  km). Right: Picture of the FP-THz absorption spectrometer with some legends.

228 In the present article, we present new measurements performed with this cavity on  $\text{CF}_4$  a very  
 229 stable GHG, with a high greenhouse warming potential ( $> 6500$ ) [2]. As for  $\text{CH}_4$ ,  $\text{CF}_4$  is not polar  
 230 and only rotational transitions induced by centrifugal distortion or vibrational anharmonicity may be  
 231 observed. Compared to  $\text{CH}_4$ , the heavier  $\text{CF}_4$  molecule exhibits ground state pure rotational transitions  
 232 at lower frequencies with a maximum intensity around 650 GHz. According to *ab initio* parameters  
 233 [35], the distortion-induced dipole is estimated to only  $0.324 \mu\text{D}$ . Therefore, all the THz lines are very  
 234 weak with intensities never exceeding  $10^{-27} \text{ cm}^{-1}/(\text{molec cm}^{-2})$ . At room temperature, rotational  
 235 transitions belonging to the hot band  $v_3 = 1$  are slightly more intense compared to the ground state  
 236 transitions. Rotational clusters in the *R* branch of this band were measured by FT-THz spectroscopy  
 237 based on synchrotron source with a path-length of 150 m in the White-type cell shown in Fig. 1 (right  
 238 part) [36]. In this study, the observation of  $\text{CF}_4$  rotational clusters has required to co-add more than  
 239 5000 spectra measured at middle resolution ( $0.01 \text{ cm}^{-1}$ ) and pressure-broadened at  $P=100$  mbar. Due  
 240 to a lack of sensitivity and resolution, it was not possible to measure at lower pressure and to fully  
 241 resolve the tetrahedral splitting of  $\text{CF}_4$  rotational lines. Moreover, the ground state transitions with  
 242 absorbances lower than 1% (Fig. 6, upper panel) were not observed. The first measurements of the  
 243 tetrahedral split components of distortion-induced rotational lines of  $\text{CF}_4$  are presented in this work  
 244 with the FP-TAS measurements performed in the 625 GHz region. The middle panel of Fig. 6 targets the  
 245 frequency range accessible by FP-TAS. The THz spectrum of  $\text{CF}_4$  is simulated at a resolution of 100 kHz  
 246 with fully resolved and Doppler limited rotational lines at  $P=100 \mu\text{bar}$ . With a 1 km path-length,  
 247 absorbances above 1 % are reached for several transitions in the ground state and in the  $v_3 = 1$  hot  
 248 band. Finally, four examples of measured lines by FP-TAS in the *R*(20) cluster are presented in the  
 249 lower panel. The lines of the multiplet were measured individually with finesesses between 2800 to 3250  
 250 and pressures between  $80 \mu\text{bar}$  and  $150 \mu\text{bar}$ . The frequency step was fixed to few kHz and the time  
 251 constant to 200 ms. The SNR for the four lines of Fig.6 are measured between 20 to 120, which shows  
 252 that the measurement of very weak THz line intensities below  $10^{-28} \text{ cm}^{-1}/(\text{molec cm}^{-2})$  may be for  
 253 now considered in THz spectroscopy. Considering this level of intensity now reachable, we can expect  
 254 trace gas detection of  $\text{N}_2\text{O}$  and  $\text{O}_3$  (Fig. 4) by FP-TAS in the THz cavity at a subppm LOD, enough for  
 255 a detection at atmospheric concentrations.



**Figure 6.** Upper panel: Simulation of the CF<sub>4</sub> THz absorbance in the conditions of the synchrotron-based FT-Far-IR measurements (see Ref.[36]), namely a resolution of 0.01 cm<sup>-1</sup>, a gas pressure of 100 mbar and a path-length of 150 m. Middle panel: Simulation of the CF<sub>4</sub> THz absorbance measured in the frequency range accessible by FP-TAS at a resolution of 100 kHz with a gas pressure of 100 μbar and an effective path-length of 1 km. For both simulations, ground state transitions are in blue, ν<sub>3</sub> = 1 hot band transitions are in red. Lower panel: Zoom on experimental individual rotational lines (R(20) cluster in the ν<sub>3</sub> ← ν<sub>3</sub> hot band) measured by FP-TAS. The experimental THz frequencies are determined by the microwave synthesiser used to drive the amplified multiplier chain. The microwave synthesiser is referenced to a time signal provided by a GPS receiver. The calculated frequencies are determined using the STDS software [24]. Dashed red line correspond to the FT-Far-IR synchrotron based measurements at a pressure of 100 mbar and a resolution of 0.01 cm<sup>-1</sup>.

## 4. Conclusions

The present article highlights the best performances reached by rotational submm-wave/THz long-path absorption spectroscopy of important atmospheric GHGs. The results obtained on CH<sub>4</sub> and CF<sub>4</sub> demonstrate that the weakly intense centrifugal distortion-induced rotational lines are measurable with a high degree of accuracy thanks to the progresses of THz gas phase high-resolution spectroscopy. Several percents of methane absorbances with a path-length of 20 m may be measured above 2 THz by a cw-THz photomixing source coupled to a White-type multi-pass cell. Thanks to a frequency metrology based on a frequency comb, the rotational line centers are measured with an accuracy competitive with those of electronic synthesizers and few cw-THz measurements are able to improve the ground state low-order molecular parameters of the CH<sub>4</sub> spherical top [12]. The study of CH<sub>4</sub> has been focused on THz line position analysis but absolute intensities measurements are required for quantitative spectroscopy of methane. Such analysis requires the improvement of the spectrometer sensitivity. For this purpose, photomixers using a metallic mirror-based FP cavity could be tested at short term in order to increase significantly the power of the source [37]. Next a new generation of spectrometers has to be imagined allowing technological breakthroughs in the THz frequency gap. The use of cavity enhanced techniques initially developed in the IR hold great promise, as demonstrated by the results on CF<sub>4</sub> presented in this study where the tetrahedral splitting of CF<sub>4</sub> rotational lines have been measured and resolved by FP-TAS in the submm-wave domain. This technique based on a high finesse cavity allows to reach km effective path-length in a compact gas cell. According to the CF<sub>4</sub> lines measured in Fig. 6 (lower panel), FP-TAS with a finesse >3200 should be able to measure rotational lines with SNR>3 for transitions with minimal intensities up to 10<sup>-28</sup> cm<sup>-1</sup>/(molec.cm<sup>-2</sup>). The results presented in this article constitute a first step of a larger study of the CF<sub>4</sub> pure rotational spectrum by FP-TAS: numerous lines in the 600-650 GHz frequency range (see Fig. 6, middle panel) with absorbances > 1 % have to be measured, especially the ground state lines. The measured line frequencies will be included in a global fit of the CF<sub>4</sub> ground state and  $v_3 = 1$  molecular parameters in order to update the TFMecaSDa database [38]. No doubt that the THz spectroscopy studies of GHGs such as CH<sub>4</sub> and CF<sub>4</sub> have to be continued. In addition, new spectroscopic data may be also obtained by considering the asymmetric isotopologues of non polar GHGs. In particular, with the high sensitivity offered by our THz cavity, it will probably possible to measure the R(27) and R(28) rotational transitions of asymmetric <sup>16</sup>O<sup>12</sup>C<sup>18</sup>O and to improve its ground state constants fitted only with low  $J$  value cm- and mm-wave transitions[39].

On the other hand, the trace gas detection of non or weakly polar GHGs at atmospheric concentration can not be seriously considered. In the case of the polar stable N<sub>2</sub>O and unstable O<sub>3</sub> GHG the situation is different since we demonstrated our ability to measure these molecules at trace levels. For the first time, THz spectroscopy was used for a direct monitoring of molecular species in an atmospheric simulation chamber. Thanks to a specially designed Chernin multi-pass cell allowing to adjust a THz pathlength from 120 m to 240 m, quantitative rotational measurements of N<sub>2</sub>O and O<sub>3</sub> traces were performed. Right now, the accessible detection levels for both compounds are limited to tens of ppm. We are working on a correct baseline modelisation in order to remove its variations due to multiple interfering stationary waves in the Chernin cell. Anyway, the first results obtained in CHARME with the coupling of a THz source to the MultiCHARME setup opens new possibilities especially for the monitoring of stratospheric reaction processes at low-pressure. In particular, the versatility of the electronic sources will allows to perform time-resolved quantitative spectroscopies of reactants, oxidants and products involved in targeted reactions occurring in the high altitude atmospheric layers. Moreover, it will be interesting to measure in CHARME, at different pressures, the THz self-continuum of water absorption with the AMC coupled to MultiCHARME and to compare our spectra with those recently performed with the coherent THz synchrotron radiation coupled to the White-type cell shown in Fig. 1 (right part) [40]. Finally, the high-finesse THz cavity has also to be employed for the detection of stable N<sub>2</sub>O and unstable O<sub>3</sub> GHG. Next studies will

305 be dedicated to verify if N<sub>2</sub>O and O<sub>3</sub> submm-wave transitions could be measured at tropospheric  
 306 (330 ppb) and stratospheric (7500 ppb) concentrations by FP-TAS.

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## 327 Abbreviations

328 The following abbreviations are used in this manuscript:

329 ACE	Atmospheric Chemistry Experiment
AILES	Advanced Infrared beamLine Exploited for Spectroscopy
AMC	Amplified Multiplier Chains
CFCs/HCFCs	ChloroFluoroCarbons
CHARME	CHamber for the Atmospheric Reactivity and the Metrology of the Environment
cw-THz	continuous-wave Terahertz
FFT	Fast Fourier Transform
FP-TAS	Fabry-Perot Terahertz Absorption Spectroscopy
FSR	Free Spectral Range
FTS	Fourier Transform Spectrometer
GHG	Greenhouse gas
HITRAN	HIgh-resolution TRANsmission molecular absorption database
IASI	Infrared Atmospheric Sounding Interferometer
IR	InfraRed
330 LOD	Level of Detection
LTG-AsGa	Low-Temperature Grown - Arsenide Gallium
MLS	Microwave Limb Sounder
NEP	Noise Equivalent Power
SCISAT	SCIENCE Satellite
SNR	Signal to Noise Ratio
SOLEIL	Source Optimisée de Lumière d'Énergie Intermédiaire du LURE
STDS	Spherical-Top Data System
TALIS	THz Atmospheric Limb Sounder
TES	Tropospheric Emission Spectrometer
TFMeCaSDa	TetraFluoroMethane Calculated Spectroscopic Databases
THz-TDS	Terahertz-Time Domain Spectroscopy
LO	Local Oscillator
UV	Ultraviolet
ZBD	Zero Bias Detector

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