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## To cite this version:

Elise Chenot, Emmanuelle Pucéat, Nicolas Freslon, Jean-François Deconinck, Mohammad Javad Razmjooei, et al.. Late Cretaceous changes in oceanic currents and sediment sources in the eastern Tethys: insights from Nd isotopes and clay mineralogy. Global and Planetary Change, 2021, 198, pp.103353. 10.1016/j.gloplacha.2020.103353 . hal-03148725

HAL Id: hal-03148725
https://hal.science/hal-03148725
Submitted on 3 Feb 2023

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Late Cretaceous changes in oceanic currents and sediment sources in the eastern Tethys: insights from Nd isotopes and clay mineralogy

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#### Abstract

The Late Cretaceous is marked by geodynamical changes including Africa-Eurasia convergence that resulted in the narrowing of the Tethys Ocean and in ophiolite obduction along the southern margin of the various continental blocks in eastern Tethys. Geochemical and mineralogical analyses were performed on the Shahneshin section (Zagros Basin - eastern


Tethys), to explore the consequences of this evolution on oceanic circulation, and to estimate the role of ophiolite weathering that can impact climate through atmospheric $\mathrm{CO}_{2}$ drawdown. [ Ni ] and $[\mathrm{Cr}]$ sharply increase in the uppermost Campanian - Maastrichtian interval and are not coeval with an increase in the Nd isotope composition $\left(\varepsilon_{\mathrm{Nd}}\right)$ of the detrital fraction of the sediments. This is interpreted as reflecting weathering of the mantellic, ultramafic part of the ophiolite sequence at that time, implying that ophiolites were exposed on the continents as soon as the Coniacian in the vicinity of the Zagros Basin. Hence their weathering could have contributed to the Late Cretaceous climatic cooling. Clay mineralogy reveals an alternation of kaolinite-rich humid periods (Coniacian/Santonian and late early Campanian to latest Campanian) and kaolinite-depleted arid periods (early Campanian and the latest Campanian to late Maastrichtian). The $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ of bathyal waters (carbonate leached fraction) appears quite radiogenic, in the range of -3.4 to $-5.2 \varepsilon$-units, about 2 to $4 \varepsilon$-units higher than the detrital fraction. Changes in the $\varepsilon_{N d(t)}$ of detrital material delivered to the Zagros Basin appear to control the $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ of the local bottom waters from the Coniacian to the middle Campanian. By contrast, decoupling of detrital $\varepsilon_{\mathrm{Nd}(t)}$ from that of the local bottom waters from the middle Campanian to the Maastrichtian points to increasing fluxes of westward flowing Pacific radiogenic water masses into the eastern Tethys at that time. These results support an intensification of the Tethyan Circumglobal Current, likely related to the narrowing of the Tethyan seaway.

Keywords

Ocean circulation, Late Cretaceous, continental weathering, climate, Shahneshin section, Iran

1. Introduction

The Late Cretaceous was characterized by a long-term global cooling (Huber et al., 1995, 2002, 2018; Pucéat et al., 2003; Friedrich et al. 2012; Linnert et al., 2014) which has been linked to a decrease in atmospheric $\mathrm{pCO}_{2}$ (Royer et al., 2012; Franks et al., 2014; Wang et al., 2014; Mills et al., 2019). During the last decades, many studies have investigated the processes involved in this atmospheric $\mathrm{CO}_{2}$ evolution. A reduction in $\mathrm{CO}_{2}$ mantle degassing, linked to a reduction in seafloor production rates and continental arc magmatism, has been repeatedly invoked to explain a decrease in $\mathrm{pCO}_{2}$ during the Late Cretaceous (Cogné and Humler, 2006; Van der Meer et al., 2014; McKenzie et al., 2016). Enhanced $\mathrm{CO}_{2}$ drawdown by continental weathering has also been suggested as a driver of this climate evolution (Jagoutz et al., 2016; Chenot et al., 2018), with a potentially important role of ophiolite obduction and weathering onto the southern Tethyan margin recently pointed out by Jagoutz et al. (2016). Whilst some constraints have been brought on the timing of the initiation of ophiolite obduction, notably using radiometric geochemical dating (Delaloye and Desmons, 1980; Lanphere and Pamić, 1983; Shahidi and Nazari, 1997; Tilton et al., 1981; Babaie et al., 2006; Shafaii Moghadam and Stern, 2011; Agard et al., 2011; Ao et al., 2016), information is still largely lacking on the timing of effective subaerial exposure of ophiolitic rocks to weathering processes, that can only be inferred from biostratigraphy of the overlying sediments (Shafaii Moghadam and Stern, 2015). This lack of information impedes our knowledge of the links between climate and obduction during the Late Cretaceous.

The ophiolite obduction that affected a large part of the southern Tethyan margin has been related to the dynamics of Africa-Eurasia convergence (Guiraud and Bosworth, 1997; Şengör and Stock, 2014; Jolivet et al., 2016). This convergence also resulted in a narrowing of the Tethys Ocean, with potential implications on oceanic circulation, a major component of the climate system. A major reorganization in the global oceanic circulation has been depicted
during the Late Cretaceous, evidenced during the Campanian by a decrease of $2 \varepsilon$-units of the Nd isotope composition of bottom waters, a tracer of modern and ancient ocean circulation, in the Atlantic and Indian oceans (Robinson et al., 2010; Macleod et al., 2011; Murphy and Thomas, 2012, 2013; Robinson and Vance, 2012; Voigt et al., 2013; Moiroud et al., 2016). This $2 \varepsilon$-units decrease could have marked the onset or intensification of deep-water formation in the southern Atlantic (Robinson et al., 2010; Robinson and Vance, 2012; Murphy and Thomas, 2012, 2013; Voigt et al., 2013) or in the northern Atlantic (Macleod et al., 2011). The progressive narrowing of the Tethys during the Late Cretaceous, related to the northward motion of Africa, may have played a role in this reorganization of the oceanic circulation. The Tethys Ocean linked the proto-Indian and Atlantic oceans during the Cretaceous, and formed, along with the open Caribbean seaway linking the Atlantic and Pacific Oceans, a low-latitude seaway where seawater could flow unimpeded throughout the globe. A large circum global current, called the Tethyan Circumglobal Current (TCC), is thought to have flowed westward throughout this seaway (Föllmi et al., 1992; Pucéat et al., 2005; Soudry et al., 2006). Like the modern Antarctic Circum global Current that flows unimpeded around Antarctica, the TCC could have played a major role in the global oceanic circulation, impacting the density balance between oceanic water masses through the transport of salty waters. The development of intensive upwelling systems along the southern Tethyan margin during the Campanian and Maastrichtian, which is marked by large phosphorite deposits and high productivity of floro-faunal assemblages (e.g. Notholt, 1985; Almogi-Labin et al., 1993; Eshet and Almogi-Labin, 1996; Ashckenazi-Polivoda et al., 2011), points to potentially important modifications in the intensity of the TCC at that time.

In this context, we aim in this study at exploring the potential changes in the TCC during the narrowing of the Tethys Ocean in the Late Cretaceous. We present a new dataset of the $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ of seawater inferred from the acetic acid-leached fraction of sediments collected
along the Shahneshin section (Zagros Basin, Iran), following recent development for the acquisition of a seawater signal from different sediment fractions (Wilson et al., 2013). The Shahneshin outcrop is located in the eastern part of the Tethys, close to the eastern entrance into the Tethyan seaway. The $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ of seawater at this site has thus the potential to track changes in the TCC intensity. The evolution of clay mineral assemblages, of the $\varepsilon_{\mathrm{Nd}(t)}$ of the insoluble residual fraction of the sediment, and elemental ( Ni and Cr ) concentrations of the sediments are also determined, to discuss the possible impact of changes in the source of material eroded from the nearby continent on local seawater $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ and to discuss the potential links between weathering and erosion processes in this region, ophiolite obduction, and climate.
2. Geodynamical framework and palaeogeography

The Middle East region presents highly deformed sedimentary structures in the orogenic areas (Zagros, Caucasus) and metamorphic rocks around the obducted areas (Cyprus, Northern Turkey, Syria, south of Iran, Oman; Fig. 1). This modern geological framework is linked to the anticlockwise movement of Africa initiated during the Cretaceous, which ultimately led to the closure of the Tethyan Realm and to the subduction of the NeoTethys Ocean beneath central Iran (Bosworth et al., 1999; Jolivet et al., 2016). In the eastern part of the Tethys, this tectonic event is expressed by a $5,000 \mathrm{~km}$-long region of relief developed across the north of Africa and the Middle East, called the Ayyubid orogeny. It is associated to the establishment of a widespread ophiolitic belt along the Afro-Arabian passive continental margin during the Turonian to Campanian interval (Agard et al., 2005; Homke et al., 2009; Shafaii Moghadam and Stern, 2011; Şengör and Stock, 2014; Jolivet et al., 2016; Fig. 1). This compressional tectonic regime has induced the development of highly subsiding
basins and inversion reliefs (e.g. Zagros folded zone; Koop and Stoneley, 1982; Agard et al., 2005; Homke et al., 2009; Şengör and Stock, 2014). Weathering of newly created nearby continental masses most likely provided detrital particles transported and deposited into adjacent highly subsiding basins (Saura et al., 2011).
2.1. Zagros basin and the "Gurpi" Formation

The Zagros folded zone encompasses south-eastern Turkey, northern Syria, northern Iraq and western to southern parts of Iran (Alavi, 1980, 1994; Berberian and King, 1981; Dercourt et al., 1986; Fig. 1a). In the west and southwestern parts of Iran, this NW-SE folded structure is composed of four main mountains provinces called Lurestan, Khuzestan, Internals Fars and Coastal Fars at about 26 to $35^{\circ}$ N (Agard et al., 2005; Fig. 1a). Palaeogeographical reconstructions suggest a palaeolatitude of 5 to $10^{\circ} \mathrm{N}$ for the Zagros Basin during the Campanian and Maastrichtian (Philip and Floquet, 2000; Barrier and Vrielynck, 2007; Fig. $1 \mathrm{~b}-\mathrm{c}$ ). This basin is located within the equatorial Tethyan corridor, bordered to the south by the Arabian Shield, and to the north by an active subduction zone with ophiolitic nappes, possibly above sea-level during the early Campanian and immersed during the late Maastrichtian (Barrier and Vrielynck, 2007; Fig 1b-c). The ages of the sediments overlying the mantellic and crustal rocks, that provide an age range of the sub-aerial exposure of the Iranian Zagros ophiolites, are Triassic to Cretaceous for Kermanshah ophiolites (Lurestan; Saccani et al., 2013; Whitechurch et al., 2013; Shafaii Moghadam and Stern, 2015), Cenomanian-Turonian to early Santonian for Neyriz ophiolites (Babaei et al., 2006; Shafaii Moghadam and Stern, 2011), undifferentiated Late Cretaceous for Haji Abad ophiolites (Ghasemi et al., 2002), and late Turonian to Maastrichtian for Baft ophiolites (Fars Provinces; Shafaii Moghadam and Stern, 2011; Shafaii Moghadam et al., 2013).

The Late Cretaceous sedimentation of the Zagros Basin (Coniacian to Maastrichtian) consists of organic matter-rich marls and limestones which compose the Gurpi Formation (Fm. - average of $3 \%$ of total organic carbon; Ala et al., 1980). This formation is largely outcropping in the four main Zagros mountain provinces (Fig. 1c). Several hiatuses are recorded in the Gurpi Fm., which have been related to maximum flooding surfaces and/or tectonic events (Razmjooei et al., 2014, 2018, 2020a; Zarei and Ghasemi-Nejad, 2015). The dinoflagellate diversity associated to the microfossil assemblages (planktonic and benthic foraminifera) highlight the existence of lateral variability of facies in the Gurpi Fm. The Fars province displays external ramp environments whereas the Lurestan and the Khuzestan Province record shallower environments characterized by two bioturbated and fossiliferous limestones units called the "Lopha limestone member" (late Campanian, Globotruncana aegyptica biozone) and the "Emam Hassan member" (late Campanian to early Maastrichtian, Gansserina gansseri biozone; Zarei and Ghasemi-Nejad, 2015). This facies lateral variability recorded in the Gurpi Fm., associated to regional gaps, has been linked to compressive tectonic movements generating different subsidence rates between provinces.

### 2.2. Shahneshin section

The investigated Late Cretaceous deposits of the Gurpi Fm. are located in the central part of the Zagros folded zone, west of the Fars province, to the north east of Kazerun city and the north west of the Shahneshin anticline (N $29^{\circ} 44^{\prime} 47^{\prime \prime}$ and E $51^{\circ} 46^{\prime} 31^{\prime \prime}$ for the base and $29^{\circ} 44^{\prime} 40.69^{\prime \prime}$ and E $51^{\circ} 46^{\prime} 26.87^{\prime \prime}$ for the top of the studied interval at Shahneshin). At Shahneshin, the sediments are composed of dark grey marls and light yellow argillaceous limestones alternations (Fig. 2). The marls and marly limestones of the Gurpi Fm. have been
deposited in a deep sedimentary basin (Motiei, 1993). Hemmati-Nasab et al. (2008) and Darabi et al. (2018) respectively suggest rough estimations of 800 to 1200 m and 450 to 600 m paleo-water depth for the Gurpi Fm. in Lurestan area. Occurrence of a very high abundance of benthic foraminifers in the upper Campanian of Shahneshin suggests a much shallower condition at this time, which could be due to a tectonically-induced regression caused by a phase of uplift (Razmjooei et al., 2018, 2020a).

The age-model relies on the integrated stratigraphic study of the Shahneshin section by Razmjooei et al. (2018; Fig. 3). Slight modifications were however proposed later on by Razmjooei et al. (2020a) for the basal part of the section based on carbon-isotope stratigraphy. In particular, the authors applied in Razmjooei et al. (2018) the biostratigraphic concept of the Coniacian/Santonian boundary based on the first occurrence of nannofossil L. cayeuxii at 20m. In Razmjooei et al. (2020a), the authors argued for the level of the Kingsdown carbon isotope event at 40 m being the best possible marker for the Coniacian/Santonian boundary, assuming some degree of plankton diachronism. This interpretation is kept here. Subsequently, the age-model applied here considers eight anchor points (Table 1; Suppl. Fig. 1). Out of these, the base Maastrichtian at 72.15 Ma is at the level identified in Razmjooei et al. (2018) but also accounts immediately for a ca. 2 Myr hiatus truncating the late Campanian (Suppl. Fig. 1). The FO of G. aegyptiaca, LO of M. furcatus and LO of D. asymetrica are given as in Razmjooei et al. (2018) age-model, the Kingsdown event is considered as the level of the Coniacian/Santonian boundary at 86.49 Ma and the basal 40 m of the section are considered to follow a similar sedimentation rate of $2.36 \mathrm{~cm} / \mathrm{kyr}$ as the interval above between 40 and 94 m , thus giving an age of 88.19 Ma for the base of the section. This well-constrained age model reveals important changes in the sedimentation rate along the section, from approximately $2.4 \mathrm{~cm} / \mathrm{kyr}$ in the upper Coniacian to Santonian, around $0.6 \mathrm{~cm} / \mathrm{kyr}$ in the lowermost Campanian, to $2.5 \mathrm{~cm} / \mathrm{kyr}$ in the remaining of the Campanian. A significant gap,
possibly of up to 2 Ma , has been suggested to occur also in the uppermost Campanian to lowermost Maastrichtian, at 280 to 290m (Razmjooei et al., 2018; Suppl. Fig. 1).
3. Material and methods
3.1. Clay mineralogy

For detailed fieldwork and sampling strategy, we refer here to Razmjooei et al. (2018). Our sampling resolution for clay mineralogy varies between 1 m and 5 m . Mineralogical analyses have been performed at the Biogéosciences Laboratory (University of BourgogneFranche Comté, Dijon, France). Clay mineral assemblages of 120 samples have been identified by X-ray diffraction (XRD) on oriented mounts of non-calcareous clay-sized particles $(<2 \mu \mathrm{~m})$. The procedure described by Moore and Reynolds (1997) has been used to prepare the samples. Diffractograms have been obtained using a Bruker D4 Endeavor diffractometer with $\mathrm{CuK}_{\alpha}$ radiations with LynxEye detector and Ni filter, under $40-\mathrm{kV}$ voltage and $25-\mathrm{mA}$ intensity. Three preparations have been analysed, after air-drying, after ethyleneglycol solvation and after heating at $490^{\circ} \mathrm{C}$ for 2 h (Fig. 4). The goniometer scanned from $2.5^{\circ}$ to $28.5^{\circ}$ for each run. Clay minerals have been identified by the position of their main diffraction peaks on the three XRD runs, while semi-quantitative estimates have been produced in relation to their area (Moore and Reynolds, 1997). Areas have been determined on diffractograms of glycolated runs with MacDiff 4.2.5. Software (Petschick, 2000). Beyond the evaluation of the absolute proportions of the clay minerals, the aim was to identify their relative fluctuations along the section.
3.2. Rare earth elements concentrations (REE) and neodymium (Nd) isotope analyses
3.2.1. Sample preparation: carbonate and insoluble fraction extraction

Eighteen sedimentary bulk rock samples from the Shahneshin section were finely crushed using an agate mortar and pestle. Aliquots of about 1.5 g of powdered sediments were leached with 20 ml of acetic acid solution ( $2 \%$ ) during 30 minutes in Teflon vial (Savillex®). This short leaching time using quite diluted acetic acid aims to dissolve part of the carbonates while limiting extraction of $\mathrm{Fe}-\mathrm{Mn}$ oxides and/or reactive minerals (Wilson et al., 2013). The solution was filtered $(0.45 \mu \mathrm{~m}, \mathrm{PTFE})$, evaporated to dryness and then re-dissolved with 4 M hydrochloric acid ( HCl ).

Another aliquot of about 2 g of each powdered sediment sample went through the 3step leaching procedure derived from Bayon et al. (2002). Carbonate was removed using 20ml acetic acid $(10 \%, \mathrm{AA})$ that reacted overnight (about 18 h ), then $\mathrm{Fe}-\mathrm{Mn}$ oxides were removed using a 20 ml reductive solution ( $25 \%$ acetic acid + 1M Hydroxylammonium Hydrochloride, $\mathrm{AA}+\mathrm{HH}$ ) during 6 hours, and organic matter was removed using a $5 \%$ hydrogen peroxide solution $\left(\mathrm{H}_{2} \mathrm{O}_{2}, \% \mathrm{wt} / \mathrm{v}\right)$ that reacted overnight (Tessier et al., 1979; Bayon et al., 2002). About 100 mg of the residue was then digested using alkaline fusion following Bayon et al. (2009) and the precipitated Fe -oxy-hydroxides dissolved in 4 M hydrochloric acid ( HCl ).

### 3.2.2. Determination of the REE concentrations

Trace elements abundances in the carbonate leachates and the residue were determined using an aliquot of the final mother solutions diluted in $10 \mathrm{ml} \mathrm{HNO}_{3} 2 \%(\mathrm{v} / \mathrm{v})$. The REE concentrations were measured with a quadrupole inductively coupled plasma mass spectrometer (Thermo X-Series II) at the "Pôle Spectrométrie Océan" (IUEM Brest, France),
using the Tm addition method (Barrat et al., 1996; Bayon et al., 2009). The REE were measured in low resolution mode, and polyatomic oxide and hydroxide interferences were corrected using oxide formation rates determined by analyzing solutions of ultrapure water, $\mathrm{Ba}+\mathrm{Ce}, \mathrm{Pr}+\mathrm{Nd}$ and $\mathrm{Sm}+\mathrm{Eu}+\mathrm{Gd}+\mathrm{Tb}$ at the beginning of the measurement cycle. The uncertainty on all measurements was better than $5 \%$ for REE. The accuracy of the method was monitored by repeated analyses of two certified reference materials from USGS (BHVO2 and BCR-2); deviation from the certified values (Jochum et al., 2016) was better than 5\% for all replicates.
3.2.3. Determination of the Nd isotopic composition

The REE were isolated using common ion exchange chromatography, including a first ion exchange resin to isolate the lanthanides (AG-50W-X8, Bio-Rad) and a second one (Ln resin, Eichrom) to isolate Nd (Bayon et al., 2012). The Nd isotopic compositions of the residues were carried on a Multi-Collector MC-ICP-MS (Neptune, Thermo Scientific) at the "Pôle Spectrométrie Océan" (IFREMER Brest, France) with instrumental bias corrected by a sample-standard bracketing method, using JNdi-1 (Tanaka et al., 2000). The carbonate leachate Nd isotope compositions were determined at the LGLTPE laboratory in Lyon (Ecole Normale Supérieure de Lyon, France) on a Thermo Scientific Neptune and a Nu Plasma 1700 MC-ICP-MS, using the Rennes in-house Nd standard (Chauvel and Blichert-Toft, 2001) as internal standard. All ${ }^{143} \mathrm{Nd} /{ }^{144} \mathrm{Nd}$ ratios were normalized to a value of 0.7219 for ${ }^{146} \mathrm{Nd} /{ }^{144} \mathrm{Nd}$ (Table 2). The 2s external reproducibility of our measurements was estimated using repeated analyses of the JNdi-1 ( $0.512115 \pm 0.000016$; average $\pm 2 \mathrm{~s})$ and the Rennes in-house Nd standard $(0.511961 \pm 0.000009)$. The overall reproducibility of the chemistry is estimated by repeated analyses of the BHVO-2 certified reference material. During the course of the study,
the measured ${ }^{143} \mathrm{Nd} /{ }^{144} \mathrm{Nd}$ ratios of BHVO-2 were $0.512987 \pm 0.000005$ (average $\pm 2 \mathrm{~s} ; \mathrm{n}=5$ ) and $0.512988 \pm 0.000007(\mathrm{n}=12)$ by using the JNdi-1 and the Rennes in-house Nd as internal standard, respectively. Those values agree well with the reference values for MC-ICP-MS and TIMS measurements (respectively $0.512990 \pm 0.000010$ and $0.512984 \pm 0.000011$; Weis et al., 2006).

### 3.2.4. Post-sedimentation decay correction

The present-day Nd isotopic ratios measured in the samples were corrected for the radioactive decay of ${ }^{147} \mathrm{Sm}$ to ${ }^{143} \mathrm{Nd}$ using the measured Sm and Nd concentrations determined for each sample $\left({ }^{147} \mathrm{Sm} /{ }^{144} \mathrm{Nd}=0.6049 \times[\mathrm{Sm}] /[\mathrm{Nd}]\right)$ and a radioactive decay constant $\lambda=$ 6.539.10-12 $\mathrm{yr}^{-1}$ for the $\mathrm{Sm}-\mathrm{Nd}$ system (Begemann et al., 2001). The absolute ages of the Shahneshin samples were calculated using the age model described in the Part 2.2 above and in Table 1. The age-corrected isotopic ratios are expressed using the standard $\varepsilon_{N d(t)}$ notation, which corresponds to the relative deviation in part per thousand from the Chondritic Uniform Reservoir (CHUR) values at the deposition age, using the present-day ${ }^{143} \mathrm{Nd} /{ }^{144} \mathrm{Nd}$ ratio of 0.512630 and ${ }^{147} \mathrm{Sm} /{ }^{144} \mathrm{Nd}$ ratio of 0.1960 for the CHUR (Bouvier et al., 2008). The analytical uncertainties expressed under the $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ notation, estimated by propagating external errors introduced by the ${ }^{143} \mathrm{Nd} /{ }^{144} \mathrm{Nd}( \pm 0.000016$ or $\pm 0.000009, \pm 2 \sigma)$ and ${ }^{147} \mathrm{Sm} /{ }^{144} \mathrm{Nd}( \pm 0.004, \pm 2 \sigma)$, are $\pm 0.35$ for the insoluble fractions and $\pm 0.21$ for the carbonate leachates (Table 2).
3.3. Determination of major and trace element concentrations

Twenty-two bulk sediment samples were additionally analysed for their major (Al) and trace-element ( $\mathrm{Ni}, \mathrm{Cr}$ ) concentrations at the Service d'Analyse des Roches et des

Minéraux of the Centre National de la Recherche Scientifique (SARM, Vandœuvre-lèsNancy, France). Complete analytical procedure developed at the SARM is detailed in Carignan et al. (2001). About 300 mg of powdered rock samples are fused in Pt crucibles with ultrapure $\mathrm{LiBO}_{2}$ at $980^{\circ} \mathrm{C}$ in an oven for 60 min , and after cooling the fusion glass is dissolved in a $\mathrm{HNO}_{3}(1 \mathrm{~mol} / \mathrm{l})-\mathrm{H}_{2} \mathrm{O}_{2}(0.5 \% \mathrm{v} / \mathrm{v})$ - glycerol $(10 \% \mathrm{v} / \mathrm{v})$ mixture. The samples were analysed on an ICP-OES (Thermo Fisher iCap 6500) and an ICP-MS (Thermo Fisher iCapQ), respectively, at the Service d'Analyse des Roches et des Minéraux of the Centre National de la Recherche Scientifique (SARM, Vandœuvre-lès-Nancy, France). Uncertainties in measurements, based on analyses of international reference standards (BR, AN-G, UB-N, DR-N, and GH; Govindaraju, 1995) and replicates are lower than $15 \%$ for $\mathrm{Al}_{2} \mathrm{O}_{3}$, and $5 \%$ for Ni and Cr (Table 2).

In order to discriminate changes in nickel $(\mathrm{Ni})$ and chromium $(\mathrm{Cr})$ concentration of the detrital material from variations in detrital supply, trace element concentrations have been normalized to Al , as aluminum is considered as a robust tracer of detrital input, and are expressed as element/Al (X/Al) ratios in this study (Fig. 3).
4. Results
4.1. Clay mineralogical assemblages

Clay mineral assemblages of the Shahneshin section are composed of illite-smectite mixed-layers (I-S R1), illite, kaolinite, chlorite, and likely a mixture of random mixed-layers including chlorite-smectite (C-S) chlorite-vermiculite (C-V) and finally vermiculite whose presence is suggested by the relatively narrow reflection at $14 \AA$ which is not modified after saturation with ethylene-glycol and which decreases sharply after heating (Fig. 4). From the
base to 260 m , I-S R1 dominates the clay fraction ( 35 to $40 \%$ ) and decreases to $20 \%$ from 260 m to the top of the section. Four mineralogical intervals are identified:

- Interval 1 (Coniacian-Santonian):

From the base to 80 m , illite ( 20 to $30 \%$ ) is associated to I-S R1 and abundant kaolinite reaching $50 \%$ at the base of the section and showing a mixture of $\mathrm{C}-\mathrm{S}$ and $\mathrm{C}-\mathrm{V}$ (and vermiculite?; 10\%) decreasing proportions upward.

- Interval 2 (lowermost Campanian pars):

From 80 to 130 m , the interval is characterized by increasing proportions of illite and I-S R1 and by the absence of kaolinite. Chlorite is occasionally abundant (up to 30\%).

- Interval 3 (lower Campanian pars to upper Campanian):

From 130 m to 250 m , illite and I-S R1 decrease up section to $15 \%$ and less than $40 \%$ respectively. The proportions of chlorite never exceed $10 \%$, while the proportions of kaolinite progressively increase up to $40 \%$ from the base to the top of the interval.

- Interval 4 (uppermost Campanian and Maastrichtian):

From 250m to the top of the section, chlorite and illite increase again (respectively 35\% and 20\%). In the same interval, the mixture of C-S, C-V and vermiculite increases significantly around $30 \%$.
4.2. REE concentrations and Nd isotopes

REE concentrations for the insoluble fractions and the acetic acid leachates are reported on Table 2 and are presented on Fig. 5 normalized to PAAS (Post Archean Australian Shale; Taylor and McLenann, 1985). The insoluble fractions all display quite similar REE patterns, characterized by depletion in Middle Rare Earth Element (MREE) and a very small or absent cerium (Ce) anomaly. By contrast, the acid acetic leachates of all
samples display a negative Ce anomaly and enrichment in middle and heavy REEs, with lower REEs concentrations compared to the insoluble fraction. The acetic acid leachates yield systematically more radiogenic Nd isotopic values than the corresponding insoluble fractions, of about $3 \varepsilon$-units on average, but up to about $4 \varepsilon$-units (Table 2; Fig. 3). The insoluble fraction $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ ranges from -9.0 to -6.0 $\varepsilon$-units and shows a positive excursion encompassing an interval from the Coniacian to the uppermost Santonian, with $\varepsilon_{\mathrm{Nd}(t)}$ reaching maximum values of about -6 $\varepsilon$-units during the Santonian. Insoluble fraction $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ remains around $-8 \varepsilon$-units, within the exception of one sample (ShG E07) that yield a value of - $6.8 \varepsilon$-units, until the lower Campanian where $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ values increase again by about $2 \varepsilon$-units, from -8.1 to $-6.2 \varepsilon$ units in the uppermost lower Campanian. $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ values then decrease by about $3 \varepsilon$-units, down to $-9 \varepsilon$-units in the upper Maastrichtian at the top of the section. The acetic leachate $\varepsilon_{N d(t)}$ displays a similar evolution than that of the insoluble fraction, with an offset of about 2-3 $\varepsilon$ units toward more radiogenic values, from the Coniacian to the uppermost Santonian. By contrast, acetic leachate $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ values remain quite stable, around $-4 \varepsilon$-units, until the lower Maastrichtian and then decline slightly from the lower to the upper Maastrichtian down to values of about -4.5 $\varepsilon$-units.

### 4.3. Major and trace element concentrations

Analytical results for the twenty-two samples are shown in Table 3 and Fig. 3. The chemical data of trace elements show that the $\mathrm{CaO}, \mathrm{SiO}_{2}$ and $\mathrm{Al}_{2} \mathrm{O}_{3}$ contents ranges are 36-48 $\mathrm{wt} . \%, 7-18 \mathrm{wt} . \%$ and $2-7 \mathrm{wt} . \%$ respectively, and the chemical data of trace elements show that Ni and Cr contents ranges are 21-83 ppm and 42-101 ppm range. Thus, the bulk rocks of the Shahneshin section are rich in $\mathrm{CaO}, \mathrm{SiO}_{2}$ and $\mathrm{Al}_{2} \mathrm{O}_{3}$, while depleted in $\mathrm{Fe}_{2} \mathrm{O}_{3}, \mathrm{~K}_{2} \mathrm{O}, \mathrm{TiO}_{2}$, and MgO . An important characteristic of these carbonates is the highest proportions of Ni and

Cr from 250 m to the top of the section (upper Campanian to upper Maastrichtian), showed by markedly higher ratios of $\mathrm{Ni} / \mathrm{Al}$ and $\mathrm{Cr} / \mathrm{Al}$ above 250 m (Fig. 3).

## 5. Discussion

5.1. Diagenesis of the clay minerals

To interpret clay mineral assemblages in terms of palaeoenvironments, we have to ensure that post depositional transformations including burial diagenesis did not modify significantly the clay minerals and that their initial sedimentary signal is preserved. In this relatively thick section from a highly tectonised area, thermal influences may be expected to have played a significant role on clay mineral assemblages. The influence of a thermal diagenesis is even more probable that the clay assemblages are devoid of I-S R0 comparable to smectites (Fig. 3). This absence is surprising because in the Tethyan Realm and even all over the globe, the Late Cretaceous is marked by a clay sedimentation largely dominated by smectites (Chamley, 1989; Chamley et al., 1990; Thiry and Jacquin, 1993; Deconinck and Chamley, 1995; Chenot et al., 2018). Moreover, there is an abundance of I-S R1 and a complex mixture of C-S and C-V mixed-layering which are respectively all intermediate phases of the transformation of smectites into illite and chlorite. However, from the top to the bottom of the section, the proportions of illite, chlorite and mixed-layers do not increase steadily, as would have been expected under a strong influence of burial diagenesis. The preservation of kaolinite, which is also sensitive to burial diagenesis, suggests that maximum burial temperature was not high enough $\left(<150^{\circ} \mathrm{C}\right.$, Beaufort et al., 2015) to transform this mineral, whose fluctuations seem to correspond to a primary signal (Fig. 3).

To sum up, it is likely that the initial clay sedimentation included a significant
proportion of I-S R0 (smectite) but that these minerals under mild thermal influence turned into I-S R1, C-S and C-V mixed-layers. In addition, this is in agreement with the moderate to poor preservation of the nannofossils (Razmjooei et al., 2020b). However, the evolution of the relative proportions of illite, chlorite and kaolinite seem to be mainly a primary signal that can be interpreted in terms of palaeoenvironment and palaeoclimate.
5.2. Climatic control of the clay assemblages of the Shahneshin section

The most striking feature of the clay mineralogy of the Shahneshin section is the alternation of kaolinite-rich intervals with intervals devoid of kaolinite (Fig. 6). The formation of kaolinite in soils requires a high level of hydrolysis and consequently a significant humidity and runoff favoring water renewal. A high temperature accelerates the hydrolysis reactions but humidity is the key parameter to form kaolinite in weathering profiles, (Chamley, 1989; Ruffell et al., 2002). Currently, kaolinite is abundant in weathering profiles that developed mainly at low latitudes. Fluctuations in kaolinite proportion in clay assemblages, especially when they are linked to opposite trends in illite and chlorite that derive from weathering and erosion of the basement in more arid environments, is often interpreted in terms of changes in humidity/aridity (eg. Vanderaveroet et al., 1999; Adatte et al., 2002; Moiroud et al., 2012). Kaolinite can also be reworked from old non-metamorphosed rocks. In this case, it is generally accompanied by primary minerals (illite and chlorite), often abundant in ancient rocks that underwent significant diagenesis (Chenot et al., 2016). On the Shahneshin section, percentages of illite and chlorite do not show the same trends as kaolinite but rather display opposite trends. This suggests that kaolinite is most likely reworked from pedogenic blankets developed at the time of sedimentation, which points to the existence, at the time of kaolinite deposition, of a relatively warm and humid climate. Such climatic
conditions are supported by the occurrence of bauxites on emerged peri-Tethyan continental masses during the Late Cretaceous (Bárdossy and Dercourt, 1990), and by the compilation of lithological and paleontological indicators, including vegetation (Otto-Bliesner and Upchurch, 1997; Upchurch et al., 2015). By contrast, the intervals with little or no kaolinite, that present higher proportions of illite and chlorite in the clay mineral assemblages, would then point to a drier climate during these intervals. Interestingly, the locally drier climatic conditions recorded during the late Campanian to Maastrichtian at Shahneshin (mineralogical interval 4, Fig. 3) do not reflect the overall humid conditions (kaolinite up to $60-80 \%$ ) recorded at other tethyan sites with paleolatitudes comprised between approximately 0 and $25^{\circ} \mathrm{N}$ (Tunisia, Li et al., 2000; Adatte et al., 2002; Egypt, Tantawy et al., 2001; Keller et al., 2002; Iran, Rostami et al., 2018; Italy; Chenot et al., 2018). This feature may be related to the specific paleogeography of the Zagros region, inducing different local climatic conditions in the Zagros studied area in this interval.

It is thus likely that an evolution in local climate could have contributed to the observed changes in clay mineral assemblages (Fig. 6). More humid conditions in intervals 1 (Coniacian to latest Santonian) and 3 (middle Campanian to earliest late Campanian) could have favoured kaolinite formation, while more arid environments in intervals 2 (early Campanian) and 4 (late Campanian to late Maastrichtian) rather favoured formation of illite and chlorite. Surface marine temperatures inferred from TEX86 values at a comparable latitude in Israel highlight a cooling of a few degrees from the Coniacian to the early Campanian, but remain relatively stable up to the end of early Maastrichtian and do not show any correspondence with the evolution of clay mineralogy on the remaining of the section (Fig. 6; Alsenz et al., 2013). Yet surface temperatures recorded in Israel (Adheret 1 borehole and PAMA Quarry) are likely affected by the development of a large upwelling system during the Campanian and Maastrichtian, depicted by the occurrence of large phosphorite deposits
and floro-faunal assemblages of high productivity (Nothold, 1985; Almogi-Labin et al., 1993; Ashckenazi-Polivoda et al., 2011). At a global scale, climate fluctuations are identified within the Campanian-Maastrichtian interval, with a warming episode evidenced in benthic foraminifera $\delta^{18} \mathrm{O}$, in TEX86, and in planktonic foraminifera during the mid-late Campanian (Friedrich et al., 2012; Linnert et al., 2014; Falzoni et al., 2016: Fig. 6). These climate fluctuations may have modified to some extent the hydrological cycle and changed the local hydrolysing conditions in the vicinity of the Zagros Basin. Such changes in sea-surface temperatures of the central part of the Zagros Basin are also suggested from calcareous nannofossil assemblages (Razmjooei et al., 2020b).

Variations in detrital sources are also susceptible to drive changes in clay mineralogical assemblages through time (Fig. 6). The main geodynamical evolution evident in this area is the $5,000 \mathrm{~km}$ long obduction-driven orogeny that affected the southern Tethyan margin from Lybia to Oman, linked to Africa-Eurasia convergence during the Late Cretaceous (Agard et al., 2011; Şengör and Stock, 2014). On both sides of the intra-oceanic subduction zone in the Zagros area, two ophiolitic belts were obducted, parallel to the southwestern part of Iran. These complete ophiolites sequences contain well-preserved mantellic (harzburgite) and crustal (gabbro and basalt) rocks (Shafaii Moghadam and Stern, 2011). Because of their mineralogical composition, oceanic crust rocks are very sensitive to chemical weathering, and plagioclase feldspar (basalt and gabbro's major components) degradation could provide the aluminum required to form kaolinite, incorporated in lateritic profiles widespread at low latitudes (Schopka et al., 2011; Giorgis et al., 2014). Illite and chlorite are primary minerals deriving from weathering and erosion of felsic, chlorite and mica-rich metamorphic rocks from the basement.

However, if the evolution of clay assemblages along the Late Cretaceous Shahneshin section was controlled by changes in the nature and provenance of the weathered and eroded
rocks as sediment sources, a co-variation of kaolinite proportions with $\mathrm{Ni} / \mathrm{Al}$ and $\mathrm{Cr} / \mathrm{Al}$ ratios would be expected, that is not observed here (Fig. 6). Indeed, mafic and ultramafic rocks are notably concentrated in Ni and Cr (Azizi et al., 2013; MacKay and Ansdell, 2013), and detrital material eroded from such rocks and deposited nearby is expected to be enriched in Ni and Cr compared to material derived from weathering of crystalline basement. On the contrary, bulk rock $\mathrm{Cr} / \mathrm{Al}$ and $\mathrm{Ni} / \mathrm{Al}$ evolution along the section appears to be decoupled from the evolution in kaolinite proportions. The $\mathrm{Cr} / \mathrm{Al}$ and $\mathrm{Ni} / \mathrm{Al}$ ratios both display a sharp increase in the middle and late Campanian followed by a milder increase in the Maastrichtian, while kaolinite proportions decrease to values close to zero at the end of the late Campanian and remain close to zero afterwards, after the ca. 2Myr hiatus.

The evolution of Ni and Cr concentrations in the sediments do not support a change in the source of the detrital material as the main driver of clay mineral assemblage evolution in the studied section, between an ophiolite source generating kaolinite and a crystalline basement source generating illite and chlorite. More likely, evolution of clay mineral assemblages is controlled by changes in hydrolysing conditions in the studied area, with kaolinite-rich intervals reflecting more humid conditions and illite and chlorite-rich intervals reflecting drier conditions. The increase in $\mathrm{Ni} / \mathrm{Al}$ and $\mathrm{Cr} / \mathrm{Al}$ ratios from the middle Campanian to the Maastrichtian may either mark the initiation of ophiolite weathering in this region, or of the weathering of the ultramafic, mantellic part of the ophiolite sequence that is over 10 times more enriched in Ni and Cr than the overlying basalt and gabbros (Azizi et al., 2013; Shafaii Moghadam and Stern, 2015; Nouri et al., 2019; Monsef et al., 2018). Weathering of mafic and ultramafic rocks can also provide the high magnesium content required to form chlorite and vermiculitic clays at the expense of smectites.
5.3. Evolution of the insoluble fraction $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ : insights on ophiolite weathering in Iran?

Detrital particles derived from the weathering and erosion of the bedrock retain the initial Nd isotope signature of the source rock (Goldstein et al., 1984). The Nd isotope signature of the detrital fraction recovered from marine sediments has thus been widely used to track sediment provenance (e.g. Frank et al., 2002; Goldstein and Hemming, 2003; Revel et al., 2010; van der Lubbe et al., 2016). Mantellic material, characterized by high $\mathrm{Sm} / \mathrm{Nd}$ ratios, yield high $\varepsilon_{\mathrm{Nd}}$ values reflecting a radiogenic composition, while ancient crustal material, characterized by lower $\mathrm{Sm} / \mathrm{Nd}$ ratios, yield negative $\varepsilon_{\mathrm{Nd}}$ values reflecting an unradiogenic composition.

The insoluble fraction analyzed in this work represents the residue obtained after removal of carbonates, $\mathrm{Fe}-\mathrm{Mn}$ oxides, and organic matter from bulk sediment samples using acetic acid, hydroxyl-ammonium hydrochloride, and hydrogen peroxide respectively (Fig. 6, see the Material and Method part). In distal oceanic sites, authigenic clays can in some case contribute significantly to the insoluble fraction of the sediments which in that case may partly integrate the Nd isotope composition of bottom seawater (Chamley and BonnotCourtois, 1981; Dubinin, 2004; Karpoff et al., 2007). At Shahneshin, the sediments have been deposited in a bathyal environment within an oceanic channel during the Late Cretaceous (Fig. 1), relatively close to the continent. This proximity to land masses is also highlighted by the quite high sedimentation rate recorded throughout the section (Razmjooei et al., 2018; Suppl. Fig. 1). Therefore, we think it unlikely for authigenic clays to represent a significant part of the insoluble fraction, and interpret the $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ of this fraction as mostly reflecting the Nd isotope composition of local detrital inputs (Fig. 6). This view is supported by the REE pattern of the residue that is quite different from that of seawater characterized by a negative Ce anomaly, by depletion in light REE and enrichment in heavy REE (Fig. 5).

The decrease in $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ of the detrital fraction during the sharp rise in $\mathrm{Ni} / \mathrm{Al}$ and $\mathrm{Cr} / \mathrm{Al}$
ratios recorded in the middle Campanian to Maastrichtian interval argues against an initiation of ophiolite weathering at that time, that should rather be associated by a large increase in $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ values. The Nd isotope data rather support weathering of the mantellic ultramafic material (mantellic rocks) underlying basalt and gabbro (volcanic and crustal rocks) of the ophiolite sequence as a driver for the observed increase in $\mathrm{Ni} / \mathrm{Al}$ and $\mathrm{Cr} / \mathrm{Al}$ ratios. Ultramafic peridotites are much more enriched in Ni and Cr (Azizi et al., 2013; Shafaii Moghadam et al., 2014; Monsef et al., 2018; Nouri et al., 2019) but are not expected to markedly increase the $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ value of detrital material (Blusztajn et al., 1995), compared to that derived from the weathering of the overlying basalt and gabbros of the ophiolite sequence. If this interpretation is correct, our data then imply that ophiolite weathering was already initiated in the vicinity of the Zagros region during the Coniacian, at the base of the section, at least for the upper parts of the ophiolite sequences.

The decrease in Nd isotope values of the detrital fraction recorded from the middle Campanian to the Maastrichtian requires however the existence of an increasing input of an additional material with an unradiogenic composition (Fig. 6). The Nd isotope composition of the detrital fraction, in the range of -9 to $-6 \varepsilon$-units, actually appears quite unradiogenic all along the section considering the age and nature of the potentially exposed terranes in the vicinity of the Zagros Basin at that time, that are expected to deliver detrital material with a more radiogenic composition. The Nd isotope composition of obducted ophiolites both in the Outer and Inner Zagros Ophiolitic Belts is radiogenic, with $\varepsilon_{N d(t)}$ values at 80 Ma in the range of +4.9 to $+9.8 \varepsilon$-units (Shafaii Moghadam et al., 2012; Azizi et al., 2013; Shafaii Moghadam et al., 2014). Granitoid bodies formed prior to the Late Cretaceous and outcropping at present in Iran also tend to present relatively radiogenic $\varepsilon_{\mathrm{Nd}}$ values (e.g. -3.5 to - $2.9 \varepsilon$-units calculated at 80 Ma for the Tuyeh-Darvar Granitoid, Naderi et al., 2018, Suppl. Table 1 ; or -4 to $+1.2 \varepsilon$ units at 80 Ma for the granitoid body of the Misho region in northwest Iran, although two
samples yield more unradiogenic values of -6.4 and -9.7 $\varepsilon$-units; Ahankoub et al., 2013, Fig. 7; Suppl. Table 1).

The relatively low $\varepsilon_{\mathrm{Nd}}$ values of the insoluble fraction, down to $-9 \varepsilon$-units, thus require mixing with an additional, unradiogenic source, with inputs that could have varied along the section. In a recent study, Nouri et al. (2019) has shown that some parts of the Harsin-Sahneh serpentinized peridotites in the mantle sequence of the Zagros ophiolite complex, exposed along the Zagros suture zone in western Iran, yield more unradiogenic values, down to -7.8 عunits at 80 Ma (Fig. 7; Suppl. Table 1). These low values have been interpreted by these authors as reflecting interaction of peridotites with seawater during metasomatism in a suprasubduction zone system. Alteration of serpentinized peridotites may have partly contributed to the relatively low $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ values observed in the insoluble fraction of Shahneshin. Other sources of unradiogenic material, farther away from the studied area, are the Khida terrane within the Arabian Shield, or the East Saharan Craton on the other side of the modern Red Sea. The Arabian-Nubian Shield, that represents the closest area with unradiogenic material, is composed of amalgamated juvenile oceanic arc terranes and ophiolite remnants, and of reworked older crust and metasediments in the flanks of the shield, that formed the suture between East and West Gondwana (Stern, 1994; Stein and Goldstein, 1996; Whitehouse et al., 1998, 2001; Stern, 2002; Kusky and Li., 2003). The juvenile oceanic arc terranes and ophiolite remnants, that compose the largest part of the shield, have $\varepsilon_{\mathrm{Nd}}$ values calculated at 80Ma typically between -5 to +18 e-units (Duyverman et al., 1982; Johnson and Woldehaimanot, 2003; Stoeser and Frost, 2006; Fig. 7; Suppl. Table 1). On the western flank of the shield, a mix of juvenile and more ancient crust yield more variable and less radiogenic $\varepsilon_{\mathrm{Nd}}$ values, down to $-35 \varepsilon$-unit at 80 Ma (Harms et al., 1990, 1994; Stern, 1994; Johnson and Woldehaimanot, 2003). The Khida terrane, that represents a small portion of a more ancient continental crust, outcrops at present in the eastern part of the shield and yields the most
unradiogenic Nd isotope values of the Arabian Shield ( $\varepsilon_{\mathrm{Nd}}$ values calculated at 80 Ma ranging between -23.8 and -6.3 $\varepsilon$-units; Fig. 7; Suppl. Table 1). Although quite far away from the studied section, detrital material derived from these unradiogenic parts of the Arabian-Nubian Shield may have been brought by rivers, in proportions that could have evolved in the context of Africa-Eurasia convergence and associated tectonic movements, susceptible to modify river catchment areas.

Interestingly, variations in detrital fraction $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ present a striking temporal coincidence with the evolution of clay mineralogical assemblages. More radiogenic $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ values are recorded in the intervals rich in kaolinite while unradiogenic values are recorded during intervals devoid of kaolinite and rich in chlorite and illite (Fig. 6). Wind-blown dust can contribute significantly to the Nd isotope composition of the detrital material deposited in oceanic basins (Grousset et al., 1988; Grousset and Biscaye, 2005). Along the western African margin, markedly lower $\varepsilon_{\mathrm{Nd}}$ values (by up to $+4 \varepsilon$-units) have been recorded in the detrital fraction of deep-sea core sediments deposited during the drier last glacial maximum, compared to present-day values measured in core-tops at the same sites (Grousset et al., 1988). These authors have interpreted these more unradiogenic values as deriving from increasing inputs of Saharan wind-blown dust at these off-shore sites during the overall drier climate interval corresponding to the Last Glacial Maximum. Similarly, the relatively unradiogenic $\varepsilon_{N d}$ values recorded in the detrital fraction at Shahneshin may reflect to some extent the presence of unradiogenic wind-blown dust mixed with radiogenic detrital material derived from the weathering of nearby ophiolites and granitoids. The origin of this unradiogenic dust would be difficult to determine, as it would be dependent of both atmospheric circulation and of the location of deserts on ancient continental crust, but it may have come from the East Saharan Craton, the Khida terrane, or, given the likely position of the site relative to the trade-winds, from the Eurasian craton as much of south-east Asia is
thought to have been very dry during the late Cretaceous (Hasegawa et al., 2012). If this interpretation is correct, then the co-variations of kaolinite proportions in clay mineral assemblages with the $\varepsilon_{N d}$ values of the insoluble fraction would derive from significant variations in the contribution of this unradiogenic dust material. This contribution would have been higher in drier intervals marked by higher proportions of illite and chlorite, and lower in more humid intervals marked by higher proportions of kaolinite. The marked decrease in Nd isotope values of the insoluble fraction from the middle Campanian to the Maastrichtian, associated with high proportions of illite and chlorite, could thus reflect drier climatic conditions and an increased contribution of wind-blown unradiogenic material to the sediments of the section at that time.

### 5.4. An oceanic signal partially preserved?

The carbonate leachates all present a REE pattern typical of non-skeletal marine carbonates formed in shallow oxic seawater, with a negative Ce anomaly, a positive La anomaly, and an enrichment in heavy rare earth elements (Webb and Kamber, 2000; Tostevin et al., 2016). These REE patterns along with the significant difference, from 2 to over $4 \varepsilon$ units recorded between the carbonate leachate and the residues $\varepsilon_{\mathrm{Nd}}$, strongly suggest that the carbonate leachates record the $\varepsilon_{\mathrm{Nd}}$ of the local bottom seawater at Shahneshin.

It has been shown that on continental margins and proximal settings, local bottom seawater $\varepsilon_{\mathrm{Nd}}$ can be impacted by local detrital inputs, through processes called boundary exchanges (Lacan and Jeandel, 2001; Jeandel, 2016). Therefore, in such settings, the evolution of local seawater $\varepsilon_{N d}$ can be governed by changes in the source of the material eroded nearby. At Shahneshin, variations in local bottom seawater $\varepsilon_{\mathrm{Nd}}$ mirror that of the detrital fraction from the Coniacian to the middle Campanian. This co-variation strongly
suggests that changes in the Nd isotope composition of the sediments transported to the Zagros Basin have likely driven the Nd isotope composition of the local seawater during this interval. By contrast, a decoupling of both isotopic signals is depicted from the late Campanian onward, highlighted by the increasing difference recorded between $\varepsilon_{\mathrm{Nd}(\mathrm{t}) \mathrm{bottom}}$ seawater and $\varepsilon_{\mathrm{Nd}(\mathrm{t}) \text { insoluble residue }}$ from about 2 to over $4 \varepsilon$-units $\left(\Delta \varepsilon_{\mathrm{Nd}(t)}\right.$ boot. Sweater - insol resi; Fig. 6). While the $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ of the residue displays a marked decrease of about $3 \varepsilon$-units from the middle Campanian to the late Maastrichtian, the $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ of the local bottom waters only shows a mild decrease of less than $1 \varepsilon$-unit on average. This difference points to increasing inputs of a radiogenic water mass in the Zagros Basin during this interval, driving the Nd isotope composition of the local seawater further away from the unradiogenic composition of the nearby detrital supply. At that time, the only deep or intermediate water masses reported in the literature with radiogenic compositions above $-5 \varepsilon$-units are water masses from the Pacific Ocean ( $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ ranging between -7 and $-3 \varepsilon$-units during the Late Cretaceous; Thomas, 2004; Hague et al., 2012; Thomas et al., 2014; Moiroud et al., 2016; Haynes et al., 2020).

During the Late Cretaceous, the large wind-driven westward flowing TCC flowed all around the globe at low latitudes, from the Indian and Pacific Oceans into the Atlantic Ocean through the Tethys, and then back to the Pacific through the Caribbean seaway (Winterer, 1991; Föllmi and Delamette, 1991; Otto-Bliesner et al., 2002; Donnadieu et al., 2006). The development of upwellings along the southern Tethyan margin during the Campanian and Maastrichtian, inferred from high productivity assemblages as radiolarites orrganic-matter layers, and the occurrence of large phosphorite deposits (Salaj and Nairn, 1987 ; Föllmi et al., 1992 ; Almogi-Labin et al., 1993 ; Pufahl et al., 2003 ; Lüning et al., 1998 ; Meilijson et al., 2014 ; El-Shafeiy et al., 2014 ; Bou Daher et al., 2014, 2015), has been linked to an intensification of the TCC at that time (Pucéat et al., 2005; Soudry et al., 2004, 2006). Our new results thus support a more intense TCC in the late Campanian to Maastrichtian interval,
that would have resulted in enhanced inputs of radiogenic Pacific waters into the Zagros Basin, driving the $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ of the local seawater further away from that of the more unradiogenic detrital inputs. Our work additionally suggests that the TCC could have extended to bathyal depths during the Late Cretaceous, at least within the Zagros Basin.

Our results also highlight the quite radiogenic composition of bathyal waters in the Zagros Basin, that display $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ values in the range of -3.5 to $-5.5 \varepsilon$-units, about 2 to $3 \varepsilon$-units above the $\varepsilon_{N d(t)}$ values of shallow waters of the Negev (Soudry et al., 2004, 2006). If such a radiogenic composition is confirmed as a feature of bathyal waters in the eastern Tethys, the development of an intense upwelling system could then have contributed to the increase in shallow water $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ values recorded in the Negev from the Turonian to the Campanian (Soudry et al., 2004; 2006), along with the intensification of the TCC. More data from other bathyal sites in the Tethys are however required to confirm the latter hypothesis.

## 6. Conclusions

Our new clay mineralogical data coupled to sediment $\mathrm{Ni} / \mathrm{Al}$ and $\mathrm{Cr} / \mathrm{Al}$ ratios suggest that the clay sedimentation of the eastern Tethys Shahneshin section was mainly driven by changes in local hydrolysing conditions, potentially linked to global climatic change as suggested by global records in benthic foraminifer $\delta^{18} \mathrm{O}$. Humid intervals in the Coniacian to Santonian and middle Campanian to early late Campanian periods could have favoured kaolinite formation while more arid intervals (early Campanian and latest Campanian to late Maastrichtian) would have inhibited kaolinite formation and promoted a higher proportion of illite and chlorite in clay mineral assemblages.

The large increase recorded in sediment $\mathrm{Ni} / \mathrm{Al}$ and $\mathrm{Cr} / \mathrm{Al}$ from the upper Campanian onward, that is not associated to a concomitant increase in the $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ of the detrital fraction, is
interpreted as reflecting the alteration and erosion of the mantellic ultramafic material of the ophiolite sequence, that underlies basalt and gabbro. Following this interpretation, our data set imply that the ophiolite sequence was exposed to continental alteration and erosion in this area as soon as the Coniacian. Our results thus support a possible role of ophiolite obduction along the southern Tethyan margin in the Late Cretaceous long-term cooling through weathering of the exposed material, a process that draws down atmospheric $\mathrm{CO}_{2}$.

The increasing difference recorded between the $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ of the local bottom seawater and of the residue from the middle Campanian to the Maastrichtian points to a growing influence of a radiogenic water mass entering the Zagros Basin during this period. These results support an intensification of the westward flowing TCC during the Campanian and Maastrichtian that would have brought larger fluxes of radiogenic Pacific waters into the eastern Tethys. Our new data suggest that the TCC would have extended down to bathyal depth at that time, at least in the Zagros Basin. They also provide the first $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ values of intermediate to deep waters within the Tethyan area revealing quite radiogenic values in this part of the eastern Tethys, in the range of -3.5 to $-5.5 \varepsilon$-units. Upwelling of such radiogenic waters onto the Negev platform during the Campanian and Maastrichtian could have partly contributed to the recorded increase in local seawater $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ values at that time.

## Acknowledgements

This research was funded by the ANR Anox-Sea coordinated by E. Pucéat. We would like to thank Marie-Laure Rouget, Emmanuel Ponzevera, Yoan Germain, Alexis de Prunelé and Philippe Telouk for their technical support in the geochemical analyses. M.J. Razmjooei and N . Thibault were funded by Carlsbergfondet CF16-0457. We are grateful to Thierry Adatte and an anonymous reviewer for constructive comments on the manuscript.

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Figures and Tables

Figure 1: Structural map of the western Iranian domain (modified from Homke et al., 2009) replaced on a google earth picture (A), palaeogeographical maps of the eastern Tethyan Realm during the early Campanian (B) and late Maastrichtian (C; modified from Barrier and Vrielynck, 2007). Abbreviations: $\mathrm{TCC}=$ Tethyan Circum global Current; Zag. Bas. $=$ Zagros Basin.

Figure 2: Field view of the Shahneshin section (Zagros Basin, Iran) (A) uppermost Santonian to lowemost Campanian stratigraphic succession and (B) upper Campanian stratigraphic succession.

Figure 3: Clay mineralogy of the Shahneshin section compared with the $\varepsilon_{N d(t)}$ of the acetic acid leachates and of the insoluble fractions, and with the $\mathrm{Ni} / \mathrm{Al}[(\mathrm{Ni}(\mathrm{ppm}) / \mathrm{Al}(\%)) \times 10000]$ and $\mathrm{Cr} / \mathrm{Al}[(\mathrm{Cr}(\mathrm{ppm}) / \mathrm{Al}(\%) \times 10000]$ ratios. These data have been replaced along the lithological succession, the calcareous nannofossils and planktonic foraminifera biostratigraphy and the $\delta^{13} \mathrm{C}$ bulk rock curve of Razmjooei et al. (2018). The $2 \sigma$ error are 0.35 $\varepsilon$-units for the insoluble fraction and $0.21 \varepsilon$-units for the acetic acid leachate fraction.

Abbreviations: $\mathrm{C}=$ Coniacian ; Maast. $=$ Mastrichtian $; \mathrm{Sa}=$ Santonian
Figure 4: X-ray diffractograms of the clay fraction of some samples from the Shahneshin section, at $81 \mathrm{~m}, 234 \mathrm{~m}$ and 313 m . The comparison of the X-ray diffractograms below 250 m (at 81 m and 234 m ) and above 250 m (at 313 m ) highly suggests the presence of vermiculite with high and narrow $14 \AA$ peak in the upper part of the section. This difference reflects the probable occurrence of vermiculite.

Figure 5: PAAS-normalized REE patterns for acetic acid leachates (at the top) and insoluble fractions (at the bottom) of the Shahneshin samples.

Figure 6: Evolution of the proportions of illite, chlorite and kaolinite compared with $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ of
the local bottom seawater and of the insoluble fraction at Shahneshin, and the $\mathrm{Ni} / \mathrm{Al}$ and $\mathrm{Cr} / \mathrm{Al}$ ratios, during the Late Cretaceous. The relative ages have been calculated using the new agemodel established for the Shahneshin section (see Table 1 and Suppl. Fig. 1., modified from Razmjooei et al., 2018, based on GTS 2016 from Ogg et al., 2016). The $2 \sigma$ error are $0.35 \varepsilon-$ units for the insoluble fraction and $0.21 \varepsilon$-units for the local bottom seawater. $\Delta \varepsilon_{\mathrm{Nd}(t)}$ represent the difference between $\varepsilon_{\mathrm{Nd}(\mathrm{t})}$ of the seawater and the insoluble fraction. TEX86 palaeotemperature data are from Pama quarry (Israel), Adrehet 1 borehole (Israel; Alsenz et al., 2013) and Shulqualak-Evans borehole (Linnert et al., 2014) and $\delta^{18} \mathrm{O}_{\text {planktonic foraminifera }}$ palaeotemperature data from ODP Site 1259 (Bornemann et al., 2008), ODP Site 1050 (Huber et al., 2002; O'Brien et al., 2017) and ODP Site 762 (Falzoni et al., 2016). These palaeo sea surface temperature are replaced along age model from O'Brien et al. (2017), based on GTS 2012 (Gradstein et al., 2012)

Figure 7: Nd isotopic composition $\left(\varepsilon_{\mathrm{Nd}}\right)$ calculated at 80 Ma (references and calculation in Suppl. Table 1) of different magmatic complexes outcropping in the general area of the Zagros Basin. Present-day extension of the Arabian-Nubian Shield (dark grey), terrane delimitation, and Precambrian deposition basins (dashed outline) from Johnson and Woldehaimanot (2003).

Abbreviations: EG = Eastern Gulf Basin; F = Fahud Basin; G = Ghudun Basin; Gh = Ghaba Basin; $\mathrm{SO}=$ South Oman Basin; WG $=$ Western Gulf Basin. The hatched area corresponds to the western oceanic arc, and the dotted area to the eastern oceanic arc, as defined in Stoeser and Frost (2006).

Table 1: Eight anchor points used to build the age-model for the Shahneshin section (updated from Razmjooei et al., 2018 taking into account the new results on Bowan section presented in Razmjooei et al., 2020a).

Table 2: REE concentrations and Nd isotopic data of the insoluble fraction, the acetic acid
leachate of the Shahneshin section and the certified reference material used in the analytic series. $\varepsilon_{\mathrm{Nd}(t)}$ values of the Shahneshin section's samples have been calculated using the agemodel of Table 1.

Table 2: Bulk rock major and trace element concentrations of the samples from the Shahneshin section.
Deep marine
Deep shelf, slope
Carbonate platform
Terrigenous platform
Exposed land and
fluviatile domains Volcanos

Accretionary complex

Active subduction
Shahneshin section - site studied

## AfB - Afghan Block

AIB - Alcapa Block
ArS - Arabian Shield
BPN - Bisitoune Platform Nappe
BPN - Bisitoune Platform Nap
BPM - Bitlis-Puturge Massif
IAM - Izmir-Ankara-Ercinkan Massif
IAM - Izmir-Ankara-Ercinkan Massif
ION - Izmir-Ankara-Ercinkan Ophiolite Nappes
KCN - Kizildag-Cilo Ophiolites Nappe
KKC - Karodut-Kocali Complexes
KRN - Khermansha Radiolarite Nappe
KsM - Kirsehir Massif
LuB - Lut Block

DOMAINS
MaN - Mamoni Nappe
NOM - North-Anatolian Ophiolitic Melange
NON - Neyriz Ophiolites Nappe
PAM - Peri-Arabian Massif
PiN - Pichakun Nappe
RSM - Rhodopes-Standja Massif
SAM - Sevan-Akera Ophiolitic Massif
SHO - Sahneh-Harsin Ophiolites Nappes
TbB-Tabas Block
TDB - Tisza-Dacia Block
TON - Troodos Ophiolite Nappe



| Sample Depth (meters) | Raw <br> Diffractogram | Interpreted Diffractogram |
| :---: | :---: | :---: |
| 81 |  |  |
| 234 |  |  |
| 313 |  |  |






| Stratigraphic anchors | Height (m) | Age (Ma) | Sedimentation rate |
| :---: | :---: | :---: | :---: |
| K-Pg Boundary | 345,85 | 66,04 | - |
| Base Maastrichtian | 283,8 | 72,15 | 1,02 |
| Base hiatus in Razmjooei agemodel | 283,8 | 74,01 | 0,00 |
| FO G. aegytiaca | 274 | 74,42 | 2,39 |
| LO M. furcatus | 113 | 81,00 | 2,45 |
| LO D. asymetrica/Base Campanian | 94 | 84,20 | 2,59 |
| Kingsdown event/Base Santonian | 40 | 86,49 | 2,36 |
| Base section with constant sedimentation rate | 0 | 88,19 |  |


| Sample ID | Depth (m) | Age (Ma) | La (ppm) | $\begin{gathered} \mathrm{Ce} \\ (\mathrm{ppm}) \end{gathered}$ | Pr (ppm) | $\begin{gathered} \begin{array}{c} \mathrm{Nd} \\ (\mathrm{ppm}) \end{array} \\ \hline \end{gathered}$ | $\begin{gathered} \mathrm{Sm} \\ (\mathrm{ppm}) \end{gathered}$ | $\begin{gathered} \mathrm{Eu} \\ (\mathrm{ppm}) \end{gathered}$ | $\begin{gathered} \begin{array}{c} \mathrm{Gd} \\ (\mathrm{ppm}) \end{array} \\ \hline \end{gathered}$ | $\begin{gathered} \mathrm{Tb} \\ (\mathrm{ppm}) \end{gathered}$ | $\begin{gathered} \hline \mathrm{Dy} \\ (\mathrm{ppm}) \end{gathered}$ | $\begin{gathered} \mathrm{Ho} \\ (\mathrm{ppm}) \end{gathered}$ | Er (ppm) | $\begin{gathered} \begin{array}{c} \mathrm{Tm}^{*} \\ (\mathrm{ppm}) \end{array} \\ \hline \end{gathered}$ | $\begin{gathered} \mathrm{Yb} \\ (\mathrm{ppm}) \end{gathered}$ | Lu (ppm) | ${ }^{147} \mathrm{Sm} /{ }^{144} \mathrm{Nd}$ | ${ }^{143} \mathrm{Nd} /{ }^{144} \mathrm{Nd}$ | 2 s | $\varepsilon \mathrm{Nd}(\mathrm{t})$ | 2s |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Insoluble fractions ( $n=17$ ) |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| ShG E18 | 333.0 | 67.31 | 17.5 | 27.1 | 2.85 | 9.77 | 1.67 | 0.333 | 1.54 | 0.291 | 2.3 | 0.586 | 2.12 | 0.33 | 2.51 | 0.38 | 0.103 | 0.512128 | 0.000006 | -9.01 | 0.15 |
| ShG E17 | 311.0 | 69.47 | 11.8 | 19.1 | 2.07 | 7.24 | 1.32 | 0.254 | 1.2 | 0.228 | 1.77 | 0.429 | 1.53 | 0.24 | 1.83 | 0.28 | 0.110 | 0.512200 | 0.000008 | -7.63 | 0.19 |
| ShG E16 | 292.5 | 71.29 | 14.1 | 20.6 | 2.11 | 6.98 | 1.12 | 0.234 | 1.04 | 0.186 | 1.44 | 0.369 | 1.37 | 0.21 | 1.61 | 0.25 | 0.097 | 0.512145 | 0.000005 | -8.56 | 0.13 |
| ShG E15 | 271.0 | 74.54 | 12 | 17.7 | 1.81 | 6.26 | 1.12 | 0.210 | 1.05 | 0.188 | 1.52 | 0.392 | 1.43 | 0.23 | 1.77 | 0.26 | 0.108 | 0.512217 | 0.000011 | -7.22 | 0.25 |
| ShG E13 | 232.0 | 76.14 | 16.6 | 23.8 | 2.46 | 8.13 | 1.46 | 0.292 | 1.17 | 0.223 | 1.66 | 0.411 | 1.47 | 0.23 | 1.77 | 0.27 | 0.109 | 0.512254 | 0.000009 | -6.49 | 0.21 |
| ShG E12 | 211.7 | 76.97 | 16.4 | 21.3 | 2.11 | 6.71 | 1.05 | 0.198 | 0.89 | 0.161 | 1.23 | 0.318 | 1.15 | 0.18 | 1.39 | 0.22 | 0.094 | 0.512222 | 0.000005 | -6.95 | 0.14 |
| ShG E11 | 192.0 | 77.77 | 17.6 | 25 | 2.53 | 8.29 | 1.53 | 0.276 | 1.25 | 0.239 | 1.74 | 0.425 | 1.5 | 0.24 | 1.85 | 0.29 | 0.112 | 0.512268 | 0.000008 | -6.23 | 0.20 |
| ShG E10 | 171.1 | 78.63 | 17 | 24.8 | 2.57 | 8.6 | 1.54 | 0.318 | 1.34 | 0.241 | 1.76 | 0.418 | 1.46 | 0.23 | 1.68 | 0.26 | 0.108 | 0.512250 | 0.000011 | -6.53 | 0.25 |
| ShG E09 | 152.0 | 79.41 | 15.2 | 21.5 | 2.26 | 7.69 | 1.3 | 0.234 | 1.04 | 0.186 | 1.45 | 0.368 | 1.25 | 0.19 | 1.46 | 0.24 | 0.103 | 0.512192 | 0.000013 | -7.60 | 0.29 |
| ShG E08 | 132.9 | 80.19 | 20.6 | 27 | 2.79 | 9.19 | 1.46 | 0.287 | 1.18 | 0.210 | 1.61 | 0.406 | 1.42 | 0.23 | 1.72 | 0.27 | 0.096 | 0.512174 | 0.000006 | -7.87 | 0.17 |
| ShG E07 | 113.0 | 81.00 | 19.5 | 26.2 | 3.02 | 10.73 | 1.99 | 0.456 | 1.97 | 0.325 | 2.33 | 0.554 | 1.84 | 0.27 | 1.99 | 0.31 | 0.112 | 0.512237 | 0.000008 | -6.80 | 0.20 |
| ShG E06 | 91.8 | 84.29 | 8.9 | 12.3 | 1.21 | 3.96 | 0.71 | 0.157 | 0.73 | 0.157 | 1.29 | 0.332 | 1.15 | 0.17 | 1.23 | 0.19 | 0.108 | 0.512167 | 0.000006 | -8.09 | 0.16 |
| ShG E05 | 74.2 | 85.04 | 17.4 | 22.2 | 2.23 | 7.2 | 1.19 | 0.228 | 1.01 | 0.177 | 1.34 | 0.327 | 1.17 | 0.18 | 1.38 | 0.22 | 0.100 | 0.512245 | 0.000006 | -6.47 | 0.17 |
| ShG E04 | 60.0 | 85.64 | 13.6 | 18 | 1.81 | 5.96 | 1 | 0.208 | 0.85 | 0.146 | 1.06 | 0.261 | 0.91 | 0.14 | 1.08 | 0.17 | 0.102 | 0.512273 | 0.000006 | -5.95 | 0.16 |
| ShG E03 | 39.8 | 86.50 | 21.5 | 29.9 | 3.21 | 11.01 | 2.02 | 0.418 | 1.77 | 0.297 | 2.16 | 0.501 | 1.74 | 0.27 | 2.03 | 0.32 | 0.111 | 0.512267 | 0.000014 | -6.14 | 0.32 |
| ShG E02 | 22.0 | 87.26 | 22.8 | 29.1 | 3.04 | 9.95 | 1.57 | 0.309 | 1.25 | 0.194 | 1.38 | 0.332 | 1.17 | 0.18 | 1.36 | 0.21 | 0.095 | 0.512185 | 0.000006 | -7.56 | 0.16 |
| ShG E01 | 1.0 | 88.15 | 24.8 | 31.8 | 3.46 | 11.62 | 1.96 | 0.402 | 1.61 | 0.253 | 1.73 | 0.401 | 1.33 | 0.21 | 1.54 | 0.24 | 0.102 | 0.512251 | 0.000004 | -6.34 | 0.13 |
|  |  |  |  |  |  |  |  |  |  |  |  |  |  | 25 external reproducibility: |  |  | $\pm 0.004$ |  | $\pm 0.000016$ |  | $\pm 0.35$ |
| Acetic acid leachate ( $n=18$ ) |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| ShG E18 | 333.0 | 67.31 | 5.1 | 7.2 | 1.02 | 3.89 | 0.75 | 0.188 | 0.78 | 0.116 | 0.71 | 0.149 | 0.42 | 0.057 | 0.38 | 0.059 | 0.117 | 0.512361 | 0.000007 | -4.57 | 0.17 |
| ShG E17 | 311.0 | 69.47 | 4.9 | 6.5 | 0.89 | 3.45 | 0.70 | 0.182 | 0.76 | 0.114 | 0.70 | 0.150 | 0.43 | 0.060 | 0.41 | 0.064 | 0.122 | 0.512365 | 0.000006 | -4.52 | 0.14 |
| ShG E16 | 292.5 | 71.29 | 3.6 | 4.9 | 0.68 | 2.68 | 0.54 | 0.132 | 0.58 | 0.087 | 0.56 | 0.118 | 0.35 | 0.048 | 0.32 | 0.048 | 0.122 | 0.512408 | 0.000008 | -3.66 | 0.19 |
| ShG E15 | 271.0 | 74.54 | 4.8 | 6.3 | 0.86 | 3.41 | 0.69 | 0.174 | 0.78 | 0.116 | 0.75 | 0.163 | 0.49 | 0.067 | 0.45 | 0.070 | 0.122 | 0.512420 | 0.000007 | -3.39 | 0.17 |
| ShG E14 | 252.2 | 75.31 | 5 | 5.6 | 0.90 | 3.59 | 0.74 | 0.185 | 0.82 | 0.127 | 0.85 | 0.187 | 0.56 | 0.077 | 0.52 | 0.082 | 0.124 | 0.512385 | 0.000005 | -4.10 | 0.14 |
| ShG E13 | 232.0 | 76.14 | 4.1 | 6.6 | 0.87 | 3.41 | 0.68 | 0.162 | 0.68 | 0.104 | 0.65 | 0.138 | 0.41 | 0.056 | 0.37 | 0.056 | 0.120 | 0.512419 | 0.000007 | -3.38 | 0.18 |
| ShG E12 | 211.7 | 76.97 | 6.1 | 7.8 | 1.02 | 3.82 | 0.73 | 0.180 | 0.83 | 0.126 | 0.82 | 0.185 | 0.56 | 0.079 | 0.54 | 0.084 | 0.115 | 0.512371 | 0.000007 | -4.25 | 0.17 |
| ShG E11 | 192.0 | 77.77 | 4.9 | 8 | 1.06 | 4.12 | 0.80 | 0.192 | 0.79 | 0.119 | 0.75 | 0.159 | 0.46 | 0.064 | 0.43 | 0.066 | 0.118 | 0.512403 | 0.000007 | -3.65 | 0.18 |
| ShG E10 | 171.1 | 78.63 | 5.8 | 8.5 | 1.12 | 4.42 | 0.90 | 0.222 | 0.97 | 0.142 | 0.87 | 0.189 | 0.56 | 0.077 | 0.52 | 0.080 | 0.123 | 0.512374 | 0.000006 | -4.26 | 0.16 |
| ShG E09 | 152.0 | 79.41 | 4.2 | 4.6 | 0.70 | 2.73 | 0.53 | 0.126 | 0.60 | 0.088 | 0.58 | 0.129 | 0.40 | 0.053 | 0.35 | 0.052 | 0.117 | 0.512355 | 0.000006 | -4.57 | 0.16 |
| ShG E08 | 132.9 | 80.19 | 4.8 | 4.9 | 0.78 | 3.07 | 0.59 | 0.148 | 0.70 | 0.104 | 0.69 | 0.157 | 0.47 | 0.064 | 0.43 | 0.067 | 0.117 | 0.512348 | 0.000005 | -4.70 | 0.15 |
| ShG E07 | 113.0 | 81.00 | 7.8 | 7.8 | 1.19 | 4.75 | 0.91 | 0.227 | 1.10 | 0.158 | 1.02 | 0.231 | 0.69 | 0.094 | 0.62 | 0.098 | 0.115 | 0.512339 | 0.000005 | -4.85 | 0.14 |
| ShG E06 | 91.8 | 84.29 | 7.2 | 5.7 | 0.96 | 4.61 | 0.96 | 0.292 | 1.44 | 0.171 | 1.14 | 0.269 | 0.84 | 0.108 | 0.68 | 0.117 | 0.125 | 0.512325 | 0.000012 | -5.19 | 0.28 |
| ShG E05 | 74.2 | 85.04 | 9.9 | 14.3 | 1.79 | 6.94 | 1.34 | 0.347 | 1.44 | 0.214 | 1.34 | 0.290 | 0.87 | 0.120 | 0.81 | 0.125 | 0.117 | 0.512394 | 0.000005 | -3.74 | 0.13 |
| ShG E04 | 60.0 | 85.64 | 8.7 | 13.8 | 1.71 | 6.66 | 1.28 | 0.314 | 1.30 | 0.193 | 1.17 | 0.246 | 0.71 | 0.096 | 0.64 | 0.098 | 0.116 | 0.512403 | 0.000005 | -3.56 | 0.14 |
| ShG E03 | 39.8 | 86.50 | 7.1 | 7.3 | 1.14 | 5.51 | 1.16 | 0.371 | 1.70 | 0.208 | 1.26 | 0.304 | 0.91 | 0.119 | 0.75 | 0.111 | 0.128 | 0.512377 | 0.000016 | -4.19 | 0.36 |
| ShG E02 | 22.0 | 87.26 | 5.1 | 6.9 | 1.01 | 4.01 | 0.78 | 0.191 | 0.87 | 0.128 | 0.79 | 0.172 | 0.50 | 0.065 | 0.42 | 0.064 | 0.118 | 0.512345 | 0.000005 | -4.70 | 0.14 |
| ShG E01 | 1.0 | 88.15 | 6.1 | 8.1 | 1.21 | 4.82 | 0.95 | 0.233 | 0.99 | 0.148 | 0.92 | 0.197 | 0.58 | 0.079 | 0.52 | 0.080 | 0.119 | 0.512365 | 0.000006 | -4.31 | 0.16 |
|  |  |  |  |  |  |  |  |  |  |  |  |  |  | 2 ext | nal repro | ducibility: | $\pm 0.004$ |  | $\pm 0.000009$ |  | $\pm 0.21$ |
| REE concentrations in Certified reference material |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |  |
| USGS BHVO-2 Average ( $\mathrm{n}=10$ ) reference values (Jochum et al., 2016): |  |  | 14.8 | 36.2 | 5.15 | 23.59 | 5.88 | 2.000 | 6.07 | 0.905 | 5.23 | 0.970 | 2.53 | 0.320 | 2.009 | 0.278 | 0.1508 |  |  |  |  |
|  |  |  | 15.2 | 37.53 | 5.339 | 24.27 | 6.023 | 2.043 | 6.207 | 0.939 | 5.28 | 0.989 | 2.51 | 0.335 | 1.994 | 0.275 | 0.1501 |  |  |  |  |
| RSD (\%) |  |  | -2.7 | -3.6 | -3.6 | -2.8 | -2.4 | -2.1 | -2.1 | -3.6 | -0.9 | -1.9 | 0.6 | -3.3 | 0.8 | 0.8 | 0.5 |  |  |  |  |
| USGS BCR-2 Average ( $\mathrm{n}=7$ ) reference values (Jochum et al., 2016): RSD (\%) |  |  | 25.4 | 53.4 | 6.81 | 28.5 | 6.54 | 1.945 | 6.81 | 1.049 | 6.48 | 1.32 | 3.74 | 0.520 | 3.50 | 0.520 | 0.139 |  |  |  |  |
|  |  |  | 25.08 | 53.12 | 6.827 | 28.26 | 6.547 | 1.989 | 6.811 | 1.077 | 6.424 | 1.313 | 3.67 | 0.534 | 3.392 | 0.505 | 0.138-10 |  |  |  |  |
|  |  |  | 1.4 | 0.6 | -0.2 | 0.9 | -0.1 | -2.2 | 0.0 | -2.6 | 0.9 | 0.5 | 1.9 | -3.2 | 3.2 | 3.0 |  |  |  |  |  |

Reproducibility of the Isotopic measurements of the UGSS BHVO-2 certified reference material

| Sample ID | $\mathrm{SiO}_{2}$ (\%) | $\mathrm{Al}_{2} \mathrm{O}_{3}$ (\%) | $\mathrm{Fe}_{2} \mathrm{O}_{3}(\%)$ | MnO (\%) | MgO (\%) | CaO (\%) | $\mathrm{Na}_{2} \mathrm{O}$ (\%) | $\mathrm{K}_{2} \mathrm{O}$ (\%) | TiO ${ }_{2}$ (\%) | $\mathrm{P}_{2} \mathrm{O}_{5}$ (\%) | Ni (ppm) | Cr (ppm) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| G3 | 9.43 | 4.72 | 1.56 | < D.L. | 0.39 | 44.57 | 0.04 | 0.60 | 0.18 | < D.L. | 33.2 | 63.4 |
| G10 | 15.39 | 6.95 | 2.44 | < D.L. | 0.51 | 38.77 | 0.05 | 1.02 | 0.31 | < D.L. | 33.6 | 76.6 |
| G22 | 10.80 | 3.63 | 1.42 | < D.L. | 0.44 | 44.62 | 0.05 | 0.82 | 0.17 | 0.12 | 31.3 | 45.6 |
| G23 | 12.21 | 4.39 | 1.65 | < D.L. | 0.46 | 43.14 | 0.03 | 0.94 | 0.20 | 0.12 | 29.9 | 45.9 |
| G25 | 13.84 | 5.15 | 2.10 | 0.037 | 0.49 | 41.22 | 0.04 | 1.09 | 0.23 | 0.11 | 65.6 | 47.5 |
| G26 | 13.65 | 5.06 | 1.68 | < D.L. | 0.47 | 41.73 | 0.04 | 1.08 | 0.23 | 0.13 | 22.9 | 42.4 |
| G32 | 18.36 | 6.65 | 2.26 | 0.016 | 0.58 | 36.61 | 0.07 | 1.58 | 0.31 | 0.11 | 46.9 | 53.4 |
| G38 | 15.00 | 4.92 | 1.76 | < D.L. | 0.57 | 40.84 | 0.06 | 1.44 | 0.23 | 0.12 | 22.6 | 40.3 |
| G48 | 15.25 | 3.86 | 1.52 | < D.L. | 0.57 | 41.53 | 0.08 | 1.14 | 0.17 | 0.10 | 29.2 | 42.4 |
| G51 | 12.57 | 3.18 | 1.31 | < D.L. | 0.45 | 44.02 | 0.06 | 0.90 | 0.14 | < D.L. | 22.6 | 33.6 |
| G55 | 16.27 | 4.91 | 1.86 | < D.L. | 0.64 | 39.89 | 0.04 | 1.17 | 0.23 | < D.L. | 32.3 | 49.7 |
| G60 | 12.45 | 4.03 | 1.52 | < D.L. | 0.53 | 43.48 | 0.03 | 0.82 | 0.19 | < D.L. | 21.1 | 48.0 |
| G73 | 10.08 | 3.32 | 1.38 | < D.L. | 0.41 | 45.88 | 0.03 | 0.51 | 0.15 | < D.L. | 26.4 | 47.6 |
| G74 | 11.23 | 4.17 | 1.59 | < D.L. | 0.44 | 44.31 | 0.04 | 0.55 | 0.19 | < D.L. | 32.5 | 56.9 |
| G82 | 11.88 | 4.32 | 1.70 | < D.L. | 0.43 | 43.65 | 0.05 | 0.55 | 0.20 | 0.10 | 29.8 | 58.4 |
| G83 | 10.99 | 3.81 | 1.68 | < D.L. | 0.45 | 44.52 | 0.05 | 0.51 | 0.18 | 0.11 | 29.0 | 61.5 |
| G88 | 11.70 | 4.52 | 1.62 | < D.L. | 0.44 | 43.81 | 0.03 | 0.54 | 0.21 | < D.L. | 37.3 | 66.1 |
| G94 | 11.44 | 3.87 | 1.66 | < D.L. | 0.54 | 44.03 | 0.10 | 0.61 | 0.18 | 0.19 | 40.8 | 101 |
| G96 | 11.30 | 3.47 | 1.59 | < D.L. | 0.54 | 44.30 | 0.07 | 0.60 | 0.17 | < D.L. | 48.5 | 89.8 |
| G102 | 8.73 | 2.14 | 1.14 | 0.018 | 0.51 | 47.89 | 0.11 | 0.43 | 0.11 | < D.L. | 41.3 | 53.3 |
| G108 | 7.63 | 1.99 | 1.06 | < D.L. | 0.52 | 48.27 | 0.07 | 0.40 | 0.10 | < D.L. | 40.0 | 65.7 |
| G120 | 12.41 | 3.24 | 2.29 | 0.025 | 1.06 | 43.59 | 0.10 | 0.69 | 0.18 | 0.11 | 83.4 | 98.5 |

D.L. : detection limit

