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# **To cite this version:**

J. Deuzé, F.M. Bréon, C. Devaux, P. Goloub, M. Herman, et al.. Remote sensing of aerosols over land surfaces from POLDER-ADEOS-1 polarized measurements. Journal of Geophysical Research: Atmospheres, 2001, 106 (D5), pp.4913-4926.  $10.1029/2000JD900364$ . hal-03120974

# **HAL Id: hal-03120974 <https://hal.science/hal-03120974v1>**

Submitted on 26 Jan 2021

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# **Remote sensing of aerosols over land surfaces from POLDER-ADEOS-1 polarized measurements**

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**Abstract. The polarization measurements achieved by the POLDER instrument on ADEOS-! are used for the remote sensing of aerosols over land surfaces. The key advantage of using polarized observations itheir ability to systematically correct for the**  ground contribution, whereas the classical approach using natural light fails. The **estimation of land surface polarizing properties from POLDER has been examined in a**  previous paper. Here we consider how the optical thickness  $\delta_0$  and Angström exponent  $\alpha$ **of aerosols are derived from the polarized light backscattered by the particles. The inversion scheme is detailed, and illustrative results are presented. Maps of the retrieved optical thickness allow for detection of large aerosol features, and in the case of small**  aerosols, the  $\delta_0$  and  $\alpha$  retrievals are consistent with correlative ground-based **measurements. However, because polarized light stems mainly from small particles, the results are biased for aerosol distributions containing coarser modes of particles. To**  overcome this limitation, an aerosol index defined as the product  $AI = \delta_0 \alpha$  is proposed. Theoretical analysis and comparison with ground-based measurements suggest that AI is approximately the same when using  $\delta_0$ , and  $\alpha$  is related to the entire aerosol size **distribution or derived from the polarized light originating from the small polarizing particles alone. This invariance is specially assessed by testing the continuity of AI across coastlines, given the unbiased properties of aerosol retrieval over ocean. Although reducing the information concerning the aerosols, this single parameter allows a link between the POLDER aerosol surveys over land and ocean. POLDER aerosol index global maps enable the monitoring of major aerosol sources over continental areas.** 

# **1. Introduction**

**Recognition of the climatic influence of aerosols [Hansen and Lacis, 1990; Charlson et al., 1992] has led to large efforts for improving their global monitoring by satellite. Over the ocean, where the aerosol remote sensing is tractable from measurements at near-infrared wavelengths, aerosol surveys using satellite imagery have been conducted from decades [Fraser, !976; Quentzel and Koepke, 1984; Rao et al., 1989; Nakafirna and Higurashi, 1998]. Presently, aerosol maps are derived routinely from NOAA/AVHRR data [Stowe et al., 1992; Husar et al., 1997], and improved characterization of the aerosols is or should be derived from new remote sensing instruments like OCTS (Ocean Color Temperature Scanner)**  [Nakajima et al., 1999]. SeaWiFS (Sea-viewing Wide Field-of**view Sensor) [Gordon and Wang, !994], POLDER (Polariza**tion and Directionality of Earth's Reflectances) [Deuzé et al., **1999], MISR (Multiangle Imaging Spectroradiometer) [Kahn et at., 1998], or MODIS (Moderate-Resolution Imaging Spec**troradiometer) [Tanré et al., 1997].

By contrast, the aerosol remote sensing over land surfaces,

**3Centre National d'Etudes Spatiales, Toulouse, France.** 

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**Paper number 2000JD900364. 0! 48-0227/01/2000JD900364509.00**  **where the main aerosol sources are located, is much less advanced because of the difficulty in discriminating the aerosol contribution from the ground in top of the atmosphere (TOA) measurements. This problem has been approached in several different ways, for example, from the reflectance of dark targets [Kaufman and Sendra, 1988], the adjacency effect of the**  aerosols [Kaufman and Joseph, 1982; Tanré et al., 1988], or **their effect on the apparent temperature of desertic areas [Legrand et ai., 1988]. Recently, first monitoring of UVabsorbing aerosols, both over land and ocean, has been achieved by using the ultraviolet measurements of TOMS/ Nimbus 7 [Herman et al., 1997; Torres et al., I998]. Here again, the new generation of remote sensing instruments should im**prove the characterization of the aerosols from dark dense **vegetation reflectance [Kaufman et al., 1997a, 1997b], mul**tiangle imagery [Martonchik, 1997] or polarization measure**ments [Herman et al., 1997]. Here we examine the capacity of directional polarization measurements, using the 8 month data acquired by POLDER aboard ADEOS-1. The remote sensing of aerosols is based on the polarized light that they backscatter and the expectation that any correction in TOA measurements of the polarized light reflected by surfaces should be more tractable than the correction of the total reflected light.** 

**The wide field-of-view imaging spectroradiometer POLDER, on ADEOS-1 from November !996 to the end of ADEOS on June 1997, offered multispectral, multidirectional, and polarized measurements of the Earth's reflectances. The wide field-of-view optics and the matrix array detector of POLDER provide bidimensional pictures of the Earth's surface which allow daily observations with a resolution of 6 km x** 

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**Figure 1. In situ measurements of the polarized phase func**tions  $q_a(\Theta)$  for different kinds of aerosols as a function of the scattering angle  $\Theta$ ;  $q_a(\Theta)$  is derived from ground-based measurements of the polarized skylight at wavelength  $\lambda = 865$  nm **[Vermeulen et al., 2000]. Results are shown for desertic dusts (Hapex campaign, Maidives), aerosols from the Indian continent (Indoex campaign), and urban-industrial aerosols (La Crau campaign).** 

**7 km, from 13 different viewing geometries during the same orbit. A rotating filter wheel provides radiance measurements in nine channels ranging from 443 to 910 nm. Moreover, the channels centered at 443, 670, and 865 nm are equipped with polaroids, therefore permitting measurements of the polariza**tion [*Deschamps et al.,* 1994].

**The aerosol retrieval from these measurements is detailed in section 2. The polarized light scattered by the aerosols is firstly estimated by correcting the molecular and surface influences, which requires modeling of the POLDER measurements and surface bidirectional polarization distribution function (BPDF). By processing multitemporal series of POLDER**  data, Nadal and Bréon [1999] have developed BPDF models of **the land surfaces which have been entered in the inversion scheme in order to derive the aerosol contribution. Presently, the 443 nm data are not used. At this wavelength the molecular contribution dominates largely the measurements, so because of the uncertainty in the calibration factor of the blue channel, there is very little accuracy in the remaining aerosol contribution to use in an aerosol inversion scheme. Finally, comparison of the aerosol polarization measurements in the 670 and 865 nm channels along with calculations for a set of aerosol models**  provides the optical thickness of the particles  $\delta_0$  at  $\lambda_0 = 865$  nm and their Ångström exponent  $\alpha$  between wavelengths  $\lambda_0 = 865$ **nm** and  $\lambda_1 = 670$  nm such that  $\delta_1 = \delta_0(\lambda_1/\lambda_0)^{-\alpha}$ , where  $\delta_1$  is the optical thickness at  $\lambda_1$ .

**The applicability of the method is examined in section 3, and**  examples of level 2 aerosol products,  $\alpha$  and  $\delta_{0}$ , over land are **shown. Some large aerosol features are recognizable in daily maps of the retrieved optical thickness. Section 4 presents comparisons of the POLDER results with correlative ground**based measurements. The retrieved parameters  $\alpha$  and  $\delta_0$  are **significant in the case of small aerosols, such as in biomass burning events, but are generally biased. Because polarized light is mainly sensitive to scattering by small particles and**  cannot detect coarser modes of aerosols,  $\alpha$  and  $\delta_0$  are probably **related to the smaller mode of particles, leading to an overes**timation of the **Angström exponent and an underestimation of the optical thickness in the case of bimodal size distribution of**  the particles. In a first approach it is proposed, in section 5, to combine the parameters  $\alpha$  and  $\delta_0$  into a simple aerosol index,

 $AI = \alpha \delta_0$ , which should be approximately the same when **calculated either with the biased parameters derived from po.**  larization or with those corresponding to the whole size distri**bution. This aerosol index makes it possible to link the two**  different aerosol surveys performed by POLDER over land **and over the ocean, and global maps of AI provide a new monitoring opportunity of major aerosol sources over conti. nental areas, including desertic particles.** 

# **2. Aerosol Retrieval From Polarized Light**

#### **2.1. Scheme of Aerosol Retrieval**

**The principle of the method has been discussed previously**  [Herman et al., 1997]. Ground-based measurements suggest **that the polarized light reflected by ground targets is small and stable enough to allow for correction in TOA measurements, and we assume that these surface BPDFs have been deter. mined. Then, the POLDER data can provide the polarized**  light scattered by the aerosols at several wavelengths  $\lambda$ , and for several scattering angles  $\Theta$ , this is about  $\delta_q q_q(\Theta)$ , where  $\delta_q$  is the optical thickness of the aerosols and  $q_a(\Theta)$  is their polar**ized phase function; that is,** 

$$
q_a(\Theta) = p_a(\Theta) P_a(\Theta), \qquad (1)
$$

where  $p_a(\Theta)$  is the phase function and  $P_a(\Theta)$  is the polariza**tion ratio for single scattering of the aerosols. The derivation of**   $\delta_a q_a(\Theta)$  from the measurements is detailed in section 2.2. We **examine here how the aerosol retrieval may be achieved from**   $\delta_a q_a(\Theta)$ , instead of  $\delta_a p_a(\Theta)$ , used in the classical approach **of total radiance measurements.** 

**The inversion seems more delicate because the polarized**  phase function  $q_a(\Theta)$  is more versatile than the normalized phase function  $p_a(\Theta)$ . Ground-based measurements, however, suggest that at near-infrared wavelengths  $(\lambda = 850 \text{ nm})$  $\lambda = 1650$  nm),  $q_a(\Theta)$  is reasonably defined for many terrestrial aerosols [Vermeulen et al., 2000]. Figure 1 shows examples of **q, (©) functions measured from the ground for different kinds of aerosols. Most aerosols exhibit rather stable polarized phase functions, except desertic particles whose polarization is typically 5-10 times lower. POLDER measurements now taken**  over the ocean will enable estimations of  $q_a(\Theta)$  for a broader **class of terrestrial aerosols. The present scheme is based on monomodal aerosol models whose polarization exhibitsimilar behavior to that of the data set in Figure 1. Lognormal size distributions have been retained (see section 2.3), with a fixed**  value for the standard deviation  $\sigma$ , and several values of the **modal radius ?, and refractive index m, of the aerosols. Figure 2 shows examples of the resulting polarized phase functions,**  for  $m = 1.40$ . They are in reasonable agreement with the **observations.** 

To select the best model from the measurements, one may look at the directional behavior of  $q_a(\Theta)$  within the different **viewing directions of POLDER, or/and to the spectral behav**ior of  $\delta_a q_a(\Theta)$  in the three polarized channels. The interval of scattering angles varies from a few degrees to more than 80<sup>e</sup> **depending on the target latitude and localization within the**  POLDER swath [Bréon et al., 1997]. As explained previously, **the present results are obtained by using only the 670 and 865 nm data. In these conditions, Figures 3 and 4 show that retrieval of the aerosol model is rather delicate.** 

For a few models, Figure 3 shows how  $q_a(\Theta)/q_a(100^\circ)$ **varies according to the particle refractive index and Angstr6m** 

**On the other hand, for the same aerosol models, Figure 4 shows how the polarized light varies as a function of the observation wavelength. The spectral dependance of the polarized light is estimated here by** 

$$
X(\lambda) = \delta_a(\lambda) q_a(\lambda, \Theta_0) / \delta_a(\lambda_0) q_a(\lambda_0, \Theta_0), \qquad (2)
$$

with  $\Theta_0 = 100^\circ$  and  $\lambda_0 = 850$  nm. The spectral information is **much more reliable than the directional one, although restricting to channels at 670 and 865 nm limits the sensitivity of the method.** 

### **2.2. Modeling of the Measurements**

**We consider cloud-free POLDER pixels, selected according**  to the cloud-screening algorithm of Bréon and Colzy [1999]. Preliminary correction of the influences of the gaseous absorption [Hagolle et al., 1999; Vesperini et al., 1999] and of the **stratospheric aerosol contribution [Lafrance and Herman, 1998] is performed. To reduce the noise level, the aerosol algorithm is applied to 3 x 3 POLDER pixels (i.e., about 20**  km  $\times$  20 km areas). Moreover, viewing directions corresponding to specular reflection within 3° are eliminated in order to **avoid contamination by the Sun glint from subpixel water surfaces (e.g., lakes, bogs).** 

**The partially polarized light at TOA is assumed to be linearly polarized, and the three measurements performed in a POLDER-polarized channel are converted [Hagolle et al.,**  1996] into the Stokes parameters  $(I_0, Q_0, U_0)$  with

$$
I_0 = L
$$
,  $Q_0 = L_p \cos(2\psi_0)$ ,  $U_0 = L_p \sin(2\psi_0)$ , (3)

where  $L$  and  $L<sub>p</sub>$  are the normalized radiance and polarized **radiance;** that is,  $L(L_p) = \pi L^*(L_p)/E$  where E is the solar **irradiance at TOA and**  $L^*(L_p)$  **is the radiance (polarized**  $\overline{C}$ **) radiance). Here the Stokes parameters are referred to axes linked to the meridian plane, which contains the viewing di**rection; hence in (3),  $\psi_0$  is the angle between the polarization direction and the meridian plane  $(-\pi/2 < \psi_0 < +\pi/2)$ . The **accuracy of the polarization measurements has been controlled**  by using observations of the Sun glint over the ocean [*Toubbé* 



**Figure 2.** Examples of the polarized phase functions  $q_a(\Theta)$ at  $\lambda = 865$  nm for the monomodal models used in the inversion scheme, according to the corresponding modal radius  $\bar{r}$  or the corresponding Angström exponent  $\alpha$ . The assumed refractive index of the particles is  $m = 1.40$ .



**Figure 3. (a) Directional dependence of the polarized light.**  Behavior of  $q_a(\Theta)/q_a(100^\circ)$  as a function of  $\Theta$  for mono**modal models with refractive index m = 1.33. Large particles exhibit characteristic polarization features but for small parti**cles ( $\alpha > 1.35$ ) no information on  $\bar{r}$  and m (see Figures 3b and 3c), hence on  $q_a(\Theta)$ , can be derived from the angular depen**dence of polarized light. (b) Same as Figure 3a but for particle**  refractive index  $m = 1.40$ . (c) Same as Figure 3a but for **particle refractive index m = 1.50.** 

et al., 1999] and is estimated at approximately  $\pm 2\%$  polarization [Hagolle et al., 1999].

**Given the uncertainties, the measurements are modeled in a simplified way by considering only the polarized light corresponding to single scattering by aerosols, single scattering by molecules, and direct reflection on the surface. Because primary scattering largely dominates the polarized light, this ap**proximation is sufficient to take into account most of the po**larized light in channels centered at 670 and 865 nm, where molecular scattering is small, and for aerosol optical thickness**  smaller than about 0.25 [Bréon et al., 1997; Lafrance, 1997]. **Single scattering by molecules or disordered particles gives polarized light perpendicular or parallel to the scattering plane, which contains the solar and viewing directions. Similarly, BPDFs are modeled by assuming that the polarized light comes from specular reflection on surface elements, whose normal bisects the solar and viewing directions. The resulting polarization therefore is perpendicular to the scattering plane** 



**Figure 4. (a) Spectral dependence of the polarized light. Be**havior of  $(\delta_a q_a)_{\lambda}$  $(\delta_a q_a)_{\lambda_0}$  as a function of  $\lambda$  for the mono**modal models of Figure 3a. The scattering angle is fixed (© =**  100°) and  $\lambda_0 = 850$  nm. The spectral behavior of polarized light is linked mainly to the behavior of  $\delta_{a,\lambda}/\delta_{a,\lambda_0}$  and allows the retrieval of the particle dimension or Angström exponent. (b) Same as Figure 4a but for particle refractive index  $m = 1.40$ . **(c) Same as Figure 4a but for particle refractive index m = 1.50.** 

**also. Thus by considering the Stokes parameters with reference to axes perpendicular and parallel to the scattering plane (i.e.,**  with  $\psi = 0$  or  $\pi/2$  for the angle between the polarization **direction and the normal to the scattering plane), each polar**ized term reduces into its Q parameter,  $Q = L_p \cos(2\psi)$  =  $\pm L_p$ , with

$$
Q_a = \delta_a q_a(\Theta) / (4 \cos \theta_v), \qquad (4a)
$$

$$
Q_m = \delta_m(z) q_m(\Theta) / (4 \cos \theta_v), \qquad (4b)
$$

$$
Q_g = L_p^g(\Omega), \tag{4c}
$$

where  $\theta_{\nu}(\theta_{s})$  is the zenith viewing (solar) angle, and  $\Omega$  is the Sun incident angle on the surface element (i.e.,  $(\pi - \Theta)/2$ ); the molecular polarized phase function  $q_m(\Theta)$  is positive, while the aerosol-polarized phase function  $q_a(\Theta)$  may be positive or **negative for polarization direction, respectively, perpendicular**  or parallel to the scattering plane; the BPDF modeling gives  $L_n^g$  positive; note that the molecular optical thickness  $\delta_m(z)$  is

**scaled according to the altitude of the target. Finally, the polarized radiance at TOA, referred to this way, is written as** 

$$
Q_{\text{cal}} = Q_m + e^{-M\delta_m} (Q_a + e^{-M\beta\delta_a} Q_g), \qquad (5)
$$

**where M is the air mass factor, and the exponential transmis. sion terms account for the screening effects of the upper molecular layer and aerosol layer on the polarized light from**  lower layers;  $\beta$  accounts for the large forward scattering of the aerosols, a property that reduces their screening effect on the polarized light from the ground [Bréon et al., 1995; Lafrance, 1997]. For typical surface BPDFs and aerosol models, simula**tions of the TOA-polarized radiance have been performed with vertical mixing of the molecules and aerosols. The aerosol and molecular densities were given exponential profiles with scale heights of 2 and 8 km, respectively. The results show that**  with  $\beta \approx 0.50$ , (5) is accurate within about 0.0001 at 865 nm **and 0.0002 at 670 nm for aerosol optical thickness smaller than**  0.25 and viewing angles smaller than 60° [Lafrance, 1997]. The **accuracy of (5) decreases for larger aerosol optical thicknesses.** 

On the other hand, given  $(Q_0, U_0)$  and the observation geometry, the angle  $\psi_{\text{meas}}$  between the observed polarization **direction and the normal to the scattering plane is evaluated. Of course, because of the multiple scattering, scattering/**  reflection processes, and measurement errors,  $\psi_{\text{meas}}$  is not exactly zero or  $\pi/2$ . Numerical simulations, however, show that for the 865 and 670 nm wavelengths,  $\psi_{\text{meas}}$  should depart significantly from zero or  $\pi/2$  only when the polarized light is **nearly null. To be consistent with the simplified modeling,**  given  $L_p^{\text{meas}} = \sqrt{Q_0^2 + U_0^2}$  and  $\psi_{\text{meas}}$ , the measured Stokes **parameters with reference to the scattering plane are approximated into the forrn** 

$$
Q_{\text{meas}} = \varepsilon_{\text{meas}} \sqrt{Q_0^2 + U_0^2} \quad \varepsilon_{\text{meas}} = +1 \text{ for } |\psi_{\text{meas}}| < 45^\circ
$$
\n
$$
\varepsilon_{\text{meas}} = -1 \text{ for } |\psi_{\text{meas}}| > 45^\circ. \tag{6}
$$

Finally, given  $L_p^g$  and the known molecular term, the aerosol **parameters are derived from** 

$$
Q_{\text{meas}} = \frac{\delta_m q_m}{4 \cos \theta_v} + e^{-M \delta_m} \left( \frac{\delta_a q_a}{4 \cos \theta_v} + e^{-M \beta \delta_a} L_p^g \right). \tag{7}
$$

#### **2.3. Aerosol Inversion**

**Crude modelings only of surface BPDFs were known before the POLDER launch [Vanderbilt et al., 1990; Rondeaux and**  Herman, 1991; Bréon et al., 1995]. From a preliminary analysis **of time series measurements of the polarized radiance in the**  865 nm POLDER channel, Nadal and Bréon [1999] have developed improved BPDF models with  $L_p^g$  in the form

$$
L_p^g = \alpha_0 \cos \theta_s \{1 - \exp \left[ \beta_0 R_p(\Omega) / (\cos \theta_s + \cos \theta_v) \right] \}, \quad (8)
$$

where  $R_p(\Omega)$  is the Fresnel reflection coefficient for polarized light, calculated for a refractive index of  $m = 1.50$ ;  $\alpha_0$  and  $\beta_0$ **are empirical coefficients adjusted for the different classes of land surface according to the main International Geosphere-Biosphere Programme (IGBP) biotypes and the normalized**  difference vegetation index (NDVI) [Nadal and Bréon, 1999]. Equation (8) gives values of  $L_p^g$  similar to those of the expected polarized radiance scattered by the aerosols, which explains **why remote sensing of aerosols in polarized light should be feasible. We assume that the polarized light reflected by surfaces is independent on the wavelength [Vanderbilt et at., 1990;**  Rondeaux and Herman, 1991; Bréon et al., 1995], and so the **same BPDF is used in the 670 nm channel.** 

The aerosol modeling consists of tabulations of the  $q_a(\lambda)$ ,  $\Theta$ ) functions and scattering coefficient  $s_a(\lambda)$  of the aerosols at **670 and 865 nm. We consider spherical nonabsorbing particles with monomodal size distributions in the form** 

$$
\frac{dN(r)}{d \ln r} = \frac{1}{\sigma \sqrt{2\pi}} \exp\left[-\frac{1}{2}\left(\frac{\ln r - \ln \bar{r}}{\sigma}\right)^2\right].
$$
 (9)

**Following Shettle and Fenn [1979], a value of 0.864 is used for o'. Three values of the aerosol real refractive index are considered, m = 1.33, 1.40, and 1.50, because of possible different behaviors of**  $q_a(\Theta)$  **in backscattering directions depending on**  $m$  (see Figure 3). For each value of  $m$ , 12 values of  $\bar{r}$  are adjusted to provide Ångström exponents ranging from 0.30 to **2.50.** 

**Given the 2N (two wavelengths, N-viewing directions) mea**surements, for each aerosol model the error  $\eta$ ,

$$
\eta = \sqrt{\frac{1}{2N} \sum_{\lambda_0, \lambda_1} \sum_j [\mathcal{Q}_{\text{cal}}(\lambda, \Theta_j) - \mathcal{Q}_{\text{meas}}(\lambda, \Theta_j)]^2}, \quad (10)
$$

is minimized by adjusting  $\delta_0$ . The retrieved aerosol parameters are the Ångström exponent  $\alpha$  and optical thickness  $\delta_0$  corre**sponding to the model which gives the smallest residual error**   $\eta$ . Examination of the measurements shows that we never **observe the large polarization features corresponding to the rainbow of large particles, in Figures 2 and 3, which could provide valuable information on the particle refractive index. Therefore the refractive index corresponding to the selected model is not expected to be significant and will not be examined.** 

#### **3. Results**

#### **3.1. Preliminary Qualitative Results**

**The ability to control the land surface contribution from satellite is a key point of the method. A qualitative illustration of this problem is shown in Plate 1. For each ground pixel of a large geographic area, pictures A and B display the maximum and minimum values of the polarized radiance measured in the**  865 nm channel times cos  $\theta_v$ , in the viewing direction corresponding to a constant scattering angle,  $\Theta = 120^{\circ}$ , over a 30 **day period. The measurements are corrected from the molecular term according to (4b). By selecting aconstant ©, thus a**  constant  $\Omega$ , variations of the surface BPDF should be minimized. The 120<sup>°</sup> scattering angle provides a good repetition of **the measurements at the considered latitudes and is far enough**  from the cloud rainbow direction (about 140°) [Goloub et al., **1994] to prevent contamination of the polarized light by residual liquid clouds.** 

**The noiseless behaviors of the pictures give confidence in the results, and the mean levels are consistent with the ex**pected orders of magnitude. In picture A, with  $\delta_a \approx 4$  cos  $\theta_{\nu}L_{p}/q_{a}(120^{\circ})$  and  $q_{a}(120^{\circ}) \approx 0.06$  to 0.12 according to Figure 1, the maximum values cos  $\theta_{\nu}L_{p} \approx 0.0175$  correspond to a  $\delta_a \approx 0.35$  to 0.70. In picture B, minimum values ranging from cos  $\theta_{v}L_{p} \approx 0.002$  to 0.006 are consistent with previous **ground-based measurements of vegetation and bare soil polar**izations [Vanderbilt et al., 1990; Rondeaux and Herman, 1991; Bréon et al., 1995], and show that land surface polarization is **generally no larger than aerosol polarization. Moreover, the smoothness of the rninimum picture confirms that land surface polarization exhibits lower spatial variability than land surface**  **reflectance, hence making the correction of the measurements easier.** 

**On the other hand, comparison of pictures A and B shows that very clear conditions probably never prevail over polluted areas like China or the Ganga basin, therefore preventing the estimation of the surface polarization by a simple minimum technique. The applicability of the method clearly needs some general modeling of the land surface BPDF according to the**  land surface geotype, as achieved by *Nadal and Bréon* [1999].

**Another qualitative example is provided by (Plate 1, bottom) composites of POLDER pictures acquired over occidental South Africa on June 4 and 23, 1997, at the very beginning and at the height of the biomass burning activity, respectively, as confirmed by the ATSR-2 fire products [Goloub and Arino, 2000].** 

**The viewing direction selected for each pixel corresponds**  again to a constant scattering angle (here  $\Theta = 110^{\circ}$ ). More**over, the two orbits differ from a few POLDER subcycles (about 4 days), so the viewing geometries and thus the surface contribution should be nearly the same in the two pictures. For each day, images B and P display in true color the radiances and polarized radiances measured in the three channels. Because of the importance of the molecular polarization at 443 nm, the molecular contribution has been subtracted from the polarization data. Clearly, from June 4 to 23, the large increase of the extended plume of biomass burning aerosols is more discernible in the polarized radiance picture than in the radiance one. Note that the yellow color of the polarized picture is quite a persistent feature, which shows that most aerosols scatter less polarized light at short wavelengths.** 

#### **3.2. Level 2 Aerosol Products**

Let us first illustrate the algorithm functioning on typical **measurements. Figure 5 shows the case of a pixel in the Ganga**  valley, near New Delhi. Rather than level 1 data  $(Q_0, U_0)$ , **Figure 5 displays the polarization angle and polarized radiance**   $(\psi, L_n)$  as a function of the scattering angle. Figure 5a con**firms that the polarization direction, as expected, is nearly orthogonal to the scattering plane except in backscattering directions where the polarized light becomes very small. In these directions the aerosol single-scattering contribution vanishes and multiple-scattering processes result in more involved polarization behavior. In the present case, the aerosols exhibit probably negative polarization (parallel to the scattering plane)** in a small angular range near  $\Theta = 170^{\circ}$ , but analysis of **such effects is out the scope of our simplified modeling. Figures 5b and 5c show the polarized radiances. Different labels indicate the raw measurements**  $(L_p^{\text{meas}} = \sqrt{Q_0^2 + U_0^2})$ **, the** molecular contribution  $(L_p^m = Q_m)$ , the measurements corrected from the molecular effect  $((L_p^{\text{meas}} - L_p^m)e^{\delta_m M})$ , the surface contribution  $(L_p^g)$ , and the retrieved signal  $(|Q_{cal}|)$ , respectively. The retrieved optical thickness here is  $\delta_0 = 0.44$ , with  $\alpha$  = 1.38. Over this highly polluted pixel the aerosol **contribution is well above the molecular and ground contributions, and the inversion scheme allows a correct fit of the measurements.** 

**Examples of global results derived in this way are shown in Plate 2 as maps of the retrieved aerosol optical thickness (pic**tures iA and MA) and Ångström exponent (pictures iB and **MB). Pictures iA to iC (Plate 3) are for measurements over the whole Indian subcontinent on November 14, 1996 (two consecutive orbits), and pictures MA to MC are for measurements over Mexico on May 27, 1997. For comparison, the results** 



**Figure 5, POLDER observations over a pixel in the Ganga basin, near New Delhi. (a) Measurements and calculations of the angle between the polarization direction and the perpen**dicular to the scattering plane,  $\psi$ , as a function of the scattering **angie corresponding to the viewing direction. Solid circles and open circles: measurements at 670 and 865 nm, respectively. Solid and dashed curves: calculations at 670 and 865 nm, re**spectively. The polarized vibration is perpendicular to the scat**tering plane, except in backward directions where multiple scattering and/or changes in the aerosol polarization direction make the polarization behavior more complex. (b) As Figure**  5a but for polarized light at wavelength  $\lambda = 865$  nm. Measure**ments (solid circles) compare well with calculations for the retrieved aerosol model (solid line). The surface and molecular contributions are indicated. (c) As Figure 5b but for wave**length  $\lambda = 670$  nm.

obtained over the ocean in a very different way [Deuzé et al., **1999] are reported also. As these last results are derived from more confident radiance data and compare well with correla**tive ground-based measurements [Goloub et al., 1999], com**parison allows an indirect validation of the land products.** 

The global distribution of the aerosol optical thickness over **land is not unrealistic. In pictures iA to iC (Plate 3), for example, the known pollution prevailing within the Ganga basin contrasts with lower loading over the main Indian continent,**  and the orders of magnitude of  $\delta_0$  and  $\alpha$  are reasonable. In **pictures MA to MC the large aerosol loading ofsmall particles** 

with large Ångström exponents is consistent with biomass **burning activity in this region in May. However, while the results obtained over land and over the ocean are rather con. sistent in pictures MA to MC, they differ markedly in pictures iA to iC. Such a discrepancy between the results of the two processing lines over coastlines is a common feature of similar comparisons. The exception iswhen the results over the ocean**  indicate very small particles with a high Angström exponent. **Generally, the inversion scheme over land underestimates the optical thickness, while it overestimates the Angström exponent, which suggests some systematic bias in the aerosol prod. ucts over lands.** 

# **4. Comparison With Ground-Based Measurements**

**The reliability of the results has been assessed more directly**  from comparison with direct measurements of  $\delta_0$  and  $\alpha$ **achieved by ground-based Sun photometers. Data obtained during the POLDER overpass, with clear-sky conditions according to the POLDER cloud mask, have been selected in the data bank of AERONET (Aerosol Robotic Network) [Holben et al., 1998] or provided independently by T. Nakajima, R. Santer, and V. Souffiet. Table 1 lists the stations where these correlative measurements were performed. Note that a lot of the data concerns desertic sites.** 

**The methodology of the comparison is explained by Goloub et al. [1999]. Filtering of the Sun photometer data was performed in order to select measurements (1) simultaneous to the satellite overpass within half an hour, (2) corresponding to stable atmospheric conditions, (3) specially controlled to prevent undetected subpixel clouds, and (4) with a large enough**  aerosol content; only observations with  $\delta_0 > 0.05$  were retained. The POLDER results are averaged over a  $100 \times 100$  $km<sup>2</sup>$  area around the ground station (average of  $5 \times 5$  boxes of  $3 \times 3$  pixels each).

**The comparison between the POLDER and the Sun pho**tometer results is shown in Figures 6a and 6b, respectively, for  $\alpha$  and  $\delta_0$ . Both results are extremely noisy, as seen also in Plate **2 when comparing the very different qualities of the products over the land and over the ocean. This is not too surprising.**  The uncertainty on the surface contribution, typically  $\Delta L_n^g \cong$ 0.001 to 0.002 according to *Nadal and Bréon* [1999] is notice**able compared with the aerosol contribution, generally about**   $L_p^a \cong 0.010$ . However, the outstanding feature in Figures 6a and 6b is the systematic bias in the  $\alpha$  and  $\delta_0$  retrievals, which **confirms the previous results over coastlines. In most cases the**  retrieved Ångström exponent is overestimated, while correla**tively, the aerosol optical thickness is underestimated. For all measurements, Figure 6c shows how the optical thickness ratio**   $\delta_0^{\text{POLDER}}/\delta_0^{\text{SUNPH}}$  varies as a function of  $\alpha$ . The linear fit is **indicated. The aerosol optical thickness derived from POLDER is largely underestimated for large aerosols and tends progressively toward consistent figures for very small aerosols, as outlined previously.** 

### **5. Discussion: Aerosol Index**

#### **5.1. Biases of Results and Bimodality of Aerosols**

**Let us focus on the polarized light scattered in side**scattering directions, in the range  $80^\circ < \Theta < 120^\circ$  where **polarization is large (see Figure 5). Polarization features like the rainbow should deserve a specific study. In side-scattering** 



**Maximum Minimum** 









**June 23-R June 23-P** 

Plate 1. (top) (maximum and minimum): For each ground pixel of the area, the images display the maximum and minimum value, respectively, of cos  $\theta_{\nu}L_{p}$  observed over a 1 month period with cloudless conditions. **To minimize variations in surface reflection and aerosol-polarized phase function, the data correspond to a**  viewing direction providing a constant scattering angle ( $\dot{\Theta} = 120^\circ$ ). Qualitatively, the images are expected to display the land surface and the largest aerosol contributions in polarized light. Minimum image confirms that the land surface contribution does not exhibit large spatial variability, and comparison with maximum image shows that the aerosol contribution may be well above the surface contribution. (bottom) True color com**posites (443 nm, blue; 670 nm, green; 865 nm, red) of POLDER pictures over occidental South Africa, on June**  4, 1997 and June 23, 1997, for radiance (R) and polarized radiance (P); polarized radiances are corrected for molecular scattering. The large plume of biomass burning aerosols, which appears on June 23, is more discernible in polarized light.

directions it is well known that the polarizing efficiency of scattering increases when the particle dimension decreases [*Hansen and Travis*, 1974], so polarization measurements allow detection of only the smaller particles within the aerosol size distribution. Therefore the bias in the optical thickness retrieval indicates that the implicit relationship between large/ **rampolarizing and small polarizing particles entailed bv our**  aerosol modeling is wrong. Two simple possibilities can be examined in a first approach. An obvious explanation is that some particles, because they are nonspherical, provide less

**India, Nov. 14, 1996** 



**MA-Optical thickness ••• MB-Ångström exponent** 

**MC-Aerosol index** 

**Plate 2.** (top) (IA to IC): Aerosol parameters derived from POLDER. Comparison between the results over land derived from polarized light (present paper), with the more confident results over ocean derived mainly from total light [Deuzé et al., 2000]. The observed area is over India; POLDER orbits 3161 and 3162, on **November 14, 1996. For each cloud-free pixel, image iA displays the retrieved optical thickness 8o, image iB,**  the retrieved Angström exponent  $\alpha$ , and image iC, the product  $\delta_0 \alpha$ . The color codes are indicated, increasing **values from blue to red. For these urban-industrial aerosols the results over land and over ocean contrast over**  coastlines, indicating overestimation of  $\alpha$  and underestimation of  $\delta_0$  over land. Image iC indicates compensation of the biases in the aerosol index  $AI = \delta_0 \alpha$ . (bottom) (MA to MC): Same legend but for an area over **Mexico; POLDER orbits 8012 and 8013 on May 27, 1997. For these small biomass burning aerosols the results are reasonably consistent on either sides of the coastlines, indicating that POLDER results over land are consistent.** 

**polarized light than expected according to Mic theory, which should explain the underestimation of the optical thickness. Another simple possibility isthat by assuming monomodal size distribution wc underestimate the number of largc/nonpolar**izing particles linked to the small/polarizing ones. This defect may appear for pronounced bimodality of the size distribution, **and ground-based obserwttions have shown that many' aerosols**  exhibit significant bimodality [Shettle and Fenn, 1979; Nakajima et al., 1996; Remer and Kaufman, 1998].

Let us consider the POLDER observations over the ocean, **which proxide both the total and the polarized light scattered**  by the aerosols. As explained by *Deuzé et al.* [2000], these observations are also analyzed with monomodal aerosol mod**els but the inversion starts with the total radiance in the 670** 



May 1997

$$
AI = 0.00
$$
 
$$
AI = 0.50
$$

**Plate 3.** One month average values of the aerosol index  $AI = \delta_0 \alpha$ . (top) November 1996, (bottom) May 1997. While the POLDER aerosol-processing lines over land and ocean lead to  $\delta_0$  and  $\alpha$  results, which may differ largely though coastlines, the aerosol index exhibits good continuity and allows for detecting large pollution sources (Indian continent, China), areas of biomass burning activity (Madagascar and South Africa in November, Mexico in May), and Saharan aerosols.

and 865 nm channels. These theoretical monomodal models enable a fit of the radiance and lead to values of  $\delta_0$  and  $\alpha$ consistent with ground-based observations [Goloub et al., 1999], but they tend to underestimate the polarized light mea-

Table 1. List of the Stations Where Correlative Aerosol Measurements Used for POLDER Validation Have Been Performed, With Number of Observations Available According to Different Filterings

	Site Name	Location	Number of <b>Filtered Data</b>
Europe/France	Aire/Adour	43.7°N, 0.2°E	24
Europe/France	Lille	50.5°N, 3°E	5
Europe/France	Wimereux	50.7°N, 1.6°E	30
Asia/Japan	Niigata	38°N, 139°E	21
Africa/Senegal	Dakar	14.4°N, 16.9°W	40
Africa/Burkina Fasso	Bidibahn	$14^{\circ}$ N. $0.5^{\circ}$ W	27
Africa/Burkina Fasso	Bondoukoui	$11.8^{\circ}$ N, $3.75^{\circ}$ W	39
Africa/Niger	Banizoumbou	3.5°N, 2.6°E	57

sured in side-scattering directions [Deuzé et al., 2000]. This trend is inconsistent with the hypothesis of nonpolarizing particles but is very understandable in the presence of bimodal size distribution of the aerosols. Let us consider according to Remer and Kaufman [1998] a size distribution consisting of a small accumulation mode and a large coarse mode. Mie calculations show that the coarse mode gives negligible polarized light in the range  $80^{\circ} < \Theta < 120^{\circ}$ , while its contribution in the total light is noticeable. Then, as shown qualitatively in Figure 7a and confirmed by numerical simulations [Deuzé et al., 2000], because the monomodal model has to take into account the contribution of the coarse mode in the total radiance, its adjustment is at the expense of the small particles, which leads to underestimation of the polarized light. Conversely, it can be seen that in the present algorithm, where the monomodal model only accounts for polarized light, it adjusts on the accumulation mode as shown qualitatively in Fig. 7b, at the expense of the large particles and their contribution in  $\delta_0$  and  $\alpha$ , thus leading to larger  $\alpha$  and smaller  $\delta_0$ .



**Figure 6.** (a) Comparison of the aerosol Angström exponent derived from POLDER with that derived from Sun photometer measurements. Polarized light systematically leads to an overestimation of  $\alpha$ . (b) As Figure 6a, but for the aerosol optical thickness at 865 nm. Polarized light systematically leads<br>to an underestimation of  $\delta_0$ . (c) The ratio  $\delta_0^{\text{Polder}}/\delta_0^{\text{Sun-Phot}}$  of the aerosol optical thicknesses derived from POLDER and Sun photometers is reported as a function of the Angström exponent of the particles. The solid line corresponds to linear regression. The POLDER results tend toward consistent values for small particles with a large Angström exponent.

#### 5.2. Aerosol Index

The discussion suggests that  $\delta_0$  and  $\alpha$  derived from polarization are characteristics of the small mode of aerosols and correspond to the aerosol global properties only when the size distribution reduces to this mode. A more precise assessment of the physical significance of  $\delta_0$  and  $\alpha$  requires further analysis of the POLDER data, but the previous analysis may be supported indirectly by examination of the reduced parameter

$$
AI = \alpha \delta_0, \tag{11}
$$

which, moreover, could provide a convenient aerosol indicator

Let  $(\alpha_A, \delta_A)$  and  $(\alpha_C, \delta_C)$  represent the parameters of the accumulation and coarse modes of a bimodal model, respectively. According to the previous considerations, let us assume crudely that the coarse mode corresponds to  $\alpha_c \approx 0$  and given negligible polarized light, so our inversion leads to  $(\alpha_A, \delta_A)$ As the Angström exponents are derived from nearby channels  $\lambda_0$  and  $\lambda_1$ , with  $\delta_{\lambda_1} \sim \delta_{\lambda_0} (1 + \alpha \Delta \lambda / \bar{\lambda})$ , the Angström exponent of the bimodal model is approximately

$$
\alpha = \frac{\delta_A \alpha_A + \delta_C \alpha_C}{\delta_A + \delta_C} = \frac{\delta_A \alpha_A + \delta_C \alpha_C}{\delta},
$$
 (12)

so with  $\alpha_C = 0$ ,

$$
AI = \delta \alpha = \delta_A \alpha_A. \tag{13}
$$

As shown in the Appendix, the previous crude derivation man be supported by a more refined analysis. Therefore (13) susgests that simple comparison of  $AI$  values based on the true aerosol parameters with  $AI$  estimates derived only from polarization measurements allows us to check the "bimodal" explanation of the observed biases in  $\delta_0$  and  $\alpha$ . If this explanation is correct, the aerosol index  $AI$  should provide partial characterization of the whole aerosol size distribution from data related only to the small polarizing particles.

#### 5.3. Validation-Aerosol Survey Over Lands and Ocean

The  $\delta_0$  and  $\alpha$  parameters reported in Figures 6a and 6b have been combined according to (11) in order to compare aerosal indices derived from direct Sun photometer measurements with those derived from the correlative POLDER data. Figure 8 shows that the results are reasonably consistent and that the systematic biases in  $\delta_0$  and  $\alpha$  are reduced in the resulting AI.



Figure 7. (a) To fit the radiance from bimodal aerosols,  $\bullet$ monomodal model of the ocean processing line has to take account the coarse mode of aerosols, at the expense of a lots small polarizing particles. (b) Inversely, to fit the polarion radiance, the monomodal model of the land surface proces line has mainly to take into account the accumulation mode particles, at the expense of the large particles.



Figure 8. Comparison of the aerosol index  $AI$  derived from POLDER with that derived from Sun photometer measurements. The biases in  $\delta_0$  and  $\alpha$  derived from POLDER (see Figures 6a and 6b) are reduced in  $AI = \alpha \delta_0$ .

The same transformation has been applied to the aerosol parameters  $\delta_0$  and  $\alpha$  reported in images iA, iB and MA, MB of Plate 2, and the resulting aerosol indices are reported in images iC and MC of the plate. The systematic discrepancies between the POLDER aerosol results of the two independent processing lines disappear largely when expressed in terms of aerosol index.

Therefore these comparisons support the possibility of combining the present POLDER aerosol results over land in a single aerosol indicator,  $AI$ , which despite a reduction in the information about the particles, is related to the entire aerosol distribution.

It is worthwhile to examine the behavior of  $AI$  for the particular case of desertic particles. These aerosols are known to consist partly of nonspherical and slightly polarizing particles, so the ability of polarization measurements to detect this important component is questionable. The results of Figure 8 have been dispatched in Figures 9a and 10a according to the Ångström exponent of the particles. Figures 9a and 10a correspond, respectively, to  $\alpha > 0.50$  and  $\alpha < 0.50$  according to the ground-based values as shown by the histograms of the Ångström exponent in Figures 9b and 10b. Most values of  $\alpha$  < 0.50 correspond to the three African stations of AERONET and are most likely associated with large dust events.

Strikingly, Figure 10a shows that the aerosol index is sensitive to desertic particles, with a large number of events exhibiting significant values of  $AI$ , larger than 0.05 to 0.10. Despite the systematic bias in Figure 10a between the POLDER and the Sun photometer results, which may be indicative of the inadequacy of Mie theory for such particles, the aerosol index derived from polarized light allows qualitative detection of desertic particles.

Finally, Plate 3 shows global maps of the aerosol index derived from the two aerosol-processing lines, over land and ocean, respectively. The results are average values of  $AI$  over one month, for November 1996 and May 1997. The consistency between the two kinds of results is impressive. More extensive and in depth analysis of these results should be developed in another paper, but Plate 3 shows that  $AI$  detects some large and well-known aerosol sources over land such as the Indian





Figure 9. (a) As Figure 8, but the comparison between the POLDER and the Sun photometer  $AI$  is shown for the case of rather small particles;  $\alpha > 0.50$  according to the ground-based values. (b) Histogram of the Angström exponent (from ground-based measurements) for observations reported in Figure 9a.

Figure 10. (a) As Figure 8, but the comparison between the POLDER and the Sun photometer  $AI$  is shown for the case of large particles;  $\alpha$  < 0.50 according to the ground-based values. Most results are from AERONET stations in Africa and correspond to Saharan dusts. (b) Histogram of the Ångström exponent (from ground-based measurements) for observations reported in Figure 10a.

continent, the Sahara desert, and China. Major areas of bio**mass burning activity are particularly well detected, over Madagascar and South Africa in November and over Mexico**  and China in May. A dedicated analysis of these specific aero**sols [Goloub and Arino, 2000] highlights a good correlation between the time series of AI with fire occurrences derived from the ATSR infrared channels. Moreover, the main geographical and seasonal aerosol features derived from the POLDER data over the oceans are consistent with the properties of the aerosols derived previously from the NOAA/**  AVHRR data with respect to their mean content by Stowe et al. **[1992] and Husar et al. [1997] and to their size by Higurashi and Nakafima [1999] [Deuz• et al., 1999]. Although resulting in a**  reduction in information about the aerosols, AI appears as a **powerful key in order to achieve a unified global survey of aerosols over both land and ocean.** 

## **6. Conclusion**

**Remote sensing of aerosols over land surfaces is examined by using the polarized light measured by POLDER at nearinfrared wavelengths, after due correction for the surface contribution. The polarized light scattered by the aerosols is analyzed through monomodal models of spherical nonabsorbing**  particles in order to derive their Angström exponent  $\alpha$  and optical thickness δ<sub>0</sub>.

**Comparisons with in situ measurements show that the retrieved parameters are meaningful for small aerosols, specially biomass burning aerosols, but otherwise are generally biased,**  with systematic overestimation of  $\alpha$  and correlative underestimation of  $\delta_0$ . Since polarized light is mainly sensitive to scat**tering by small particles, the biases are probably linked to bimodality of the aerosol size distribution, with the retrieved parameters related to the smaller accumulation mode of particles. POLDER measurements over the ocean, which allow for more in-depth analysis, will be processed with a bimodal inversion scheme. Further examination of these improved re**sults near coastlines should allow a more precise assessment of the significance of the individual parameters  $\alpha$  and  $\delta_0$  retrieved over land. In a first approach, an aerosol index  $AI = \alpha \delta_0$  is **proposed. It permits a link between the POLDER aerosol surveys over land and ocean, so the global maps of AI allow some monitoring of the main aerosol sources over continental areas, including desertic particles.** 

# **Appendix A**

**Equation (13) may be supported by a more refined analysis. In order to estimate more precisely how the present algorithm**  performs with bimodal aerosols, we have computed the Angström exponents  $\alpha$  and polarized phase functions  $q_{\alpha}(\lambda, \Theta)$  for **various monomodal and bimodal size distributions. Given the lack of directional information (see Figure 3), we have just**  calculated  $q_a(\lambda, 100^\circ)$  for the fixed scattering angle  $\Theta = 100^\circ$ .

**As the algorithm operates mainly with the ratio of polarized light,** 

$$
X=\frac{L_p^a(\lambda_1)}{L_p^a(\lambda_0)}=\frac{\delta_a(\lambda_1)q_a(\lambda_1)}{\delta_a(\lambda_0)q_a(\lambda_0)},
$$

Figure A1 firstly displays the relationship between X and  $\alpha$  for **monomodal and bimodal models, respectively. According to Figure Ala by fitting with a monomodal model, the X ratio** 



**Figure A1. For monomodal and bimodal aerosol models, calculated relationships between the ratio of polarized light X**  and, respectively  $(A1a)$  the Angström exponent  $\alpha$ ;  $(A1b)$  the polarized phase function  $q_a(\lambda_0, 100^\circ)$ ; and (A1c) the ratio  $\alpha/q_a(\lambda_0, 100^\circ)$  of the corresponding model. Given  $L^a_p(\lambda_0)$ and  $\ddot{X} = L_p^a(\lambda_1)/L_p^a\lambda_0$  corresponding to bimodal aerosols, **estimations based on monomodal models are shown to lead**  generally to an overestimation of  $\alpha$  (Figure A1a), underestimation of  $\delta_0$ , with  $\delta \approx L_p(\lambda_0)/q_a(\lambda_0)$  (Figure Alb), but **reasonable values of AI with**  $AI = \alpha \delta_0 \approx (\alpha/q(\lambda_0)) L_p^a(\lambda_0)$ **(Figure A!c). Monomodal models: lognormal size distribution**  with  $\bar{r}$  varied and  $\sigma$  fixed ( $\sigma = 0.8735$ ). Bimodal models [adapted from Tanré et al., 1997]: two lognormal size distributions with the percentages of the two modes varied; mode 1:  $\bar{r}$  varied  $(\bar{r}_1 = 0.050, (0.015), 0, 11 \mu m)$  and  $\sigma$  fixed  $(\sigma_1 = 0.6)$ ; mode 2:  $\bar{r}$  and  $\sigma$  fixed ( $\bar{r}_2 = 1.0 \mu \text{m}$ ,  $\sigma_2 = 0.8$ ). For the two **kinds of models: spherical nonabsorbing particles, with refrac**tive index  $m = 1.33, 1.40,$  and 1.50.

**corresponding to bimodal aerosols will generally lead to an overestimation of the 3mgstr6m exponent, which is the general trend of the present algorithm. Secondly, Figure Alb compares**  the relationship between X and  $q_a(\lambda_0, 100^\circ)$  for the two kinds of models. Figure A1b shows that  $q_a(\lambda_0, 100^\circ)$  is generally **smaller for bimodal than for monomodal models. Thus given**  the polarized light  $L_p^a(\lambda_0)$  scattered by bimodal aerosols, the optical thickness  $\delta_0 \approx L_p^a(\lambda_0)/q_a(\lambda_0)$  will be underestimated when assuming monomodal models. Finally, as  $AI = \alpha \delta_0$ 

 $(a/q_a(\lambda_0)) L_p^a(\lambda_0)$ , Figure A1c displays the direct relationships between X and  $\alpha/q_a(\lambda_0, 100^\circ)$  for monomodal and **bimodal models. Figure Ala confirms that, given X, the biases**  in  $\alpha$  and  $\delta_0$  compensate partly in the product  $\alpha\delta_0$ , so the **aerosol index AI derived from the monomodal algorithm will not differ significantly from its true value.** 

**Acknowledgments. Results presented in this paper were obtained using data from CNES's POLDER onboard NASDA's ADEOS. The authors would like to acknowledge A. Lifermann, the POLDER project scientist, for her constant support and helpful comments. The authors are grateful to T. Nakajima of the University of Tokyo, Japan,**  and to R. Santer and V. Soufflet for the Université du Littoral, France, **who kindly provided ground-based measurements. This study was**  funded by the Centre National d'Etudes Spatiales (CNES), the Centre National de la Recherche Scientifique (CNRS), the Conseil Régional du Nord-Pas de Calais, the Fonds Européen de Développement Régional (FEDER) and the Délégation Régionale à la Recherche et à la **Technologie du Nord-Pas de Calais. Informations on POLDER data can be found on the web site http://polder@www-projet.cnes.fr:8060.** 

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**(Received January 11, 2000; revised May 19, 2000; accepted May 28, 2000.)**