

# Depositional evolution of an extinct sinter mound from source to outflow, El Tatio, Chile

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1	Title: Depositional evolution of an extinct sinter mound from source to outflow, El
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#### 20 Abstract:

Siliceous sinter deposits from El Tatio, Chile, preserve a wide variety of depositional 21 environments and biosignatures, from high-temperature vent-proximal facies to distal 22 deposits dominated by silicified microbial mats. Four cores were drilled into an El Tatio 23 sinter mound and associated distal apron to investigate changes in hydrothermal 24 25 environments over geologic timescales. Sedimentary and geochemical analysis of multiple sinter cores records the initiation and accretion of diverse depositional features still observed 26 today in El Tatio. y. cal. BPFacies adjacent to hydrothermal vents are dominated by 27 28 laminated sinter crusts on the steep margins of a high-temperature pool, with sparse microbial 29 preservation. Outer margins of the same pool contain extensive sinter columns up to ten 30 centimeters in length, precipitated during repeated cycles of pool overflow and subsequent 31 evaporation. Low-relief hydrothermal pools also form minor deposits within distal debris aprons, and analogous pools are still active close to sampling locations. Debris aprons are 32 dominated by palisade, tufted, and arborescent microbial fabrics, with distinct mat textures 33 revealing well preserved microfossils. Surficial deposits in all cores feature detrital-rich and 34 microbially-influenced sinters overlying higher-temperature facies, indicating a relative 35 36 decrease in hydrothermal activity over time. Geochemical proxies for hydrothermal fluids and detrital input match depositional interpretations based on sedimentary structures. <sup>14</sup>C ages 37 38 from core deposits extend the mound's history by 11,000 years, recording at least three 39 thousand years of sinter deposition on top of glacial sandstones (13,337 - 10,232 y. cal. BP). Importantly, this work provides a detailed depositional model unavailable through surficial 40 41 sedimentology alone.

42 <u>Keywords:</u>

43 Siliceous sinter, hot springs, El Tatio, microbialites, microfossils

#### 44 **1. Introduction**

45 Siliceous sinter deposits form when amorphous silica (opal-A) precipitates in 46 terrestrial hydrothermal environments, such as geysers and hot springs. Water temperatures in hot springs decrease from near-boiling vents to distal outflow channels (< 30 °C), often 47 within several meters. Thermal gradients along geothermal springs produce lateral variations 48 49 in aqueous geochemistry, biological communities, and physical depositional processes. The complex relationships between these parameters produce an extraordinarily diverse range of 50 micro-environments within localized areas. Silica precipitation within sinters preserves 51 52 microscopic features in exquisite detail, especially microbial cells and their extracellular polymeric substances (EPS) (Walter, 1976a; Schultze-Lam et al., 1995; Cady & Farmer, 53 1996; Konhauser & Ferris, 1996; Jones et al., 1997, 1998, 2004, 2005; Konhauser et al., 54 55 2001, 2003; Handley et al., 2005; Campbell et al., 2015; Gong et al., 2019). Hydrothermal siliceous sinters therefore have the potential to record fine-scale transitions in sedimentology, 56 biology, and geochemistry over relatively short distances. Early sedimentary deposits (> 3 57 billion-years-old (Ga)) on Earth and Mars contain the oldest observed sinters (Ruff and 58 Farmer, 2016; Djokic et al., 2017; Ruff et al., 2019). On Earth, such deposits include the 59 60 earliest evidence for life in non-marine environments (Djokic et al., 2017), making work on 61 modern hot springs highly relevant for investigating ancient sedimentology and biosignatures 62 in paleo- and astrobiology.

Lateral gradients in sinter architecture are preserved in modern and ancient
hydrothermal systems (Cady and Farmer, 1996; Walter et al., 1996; Guidry and Chafetz,
2003a; Lowe and Braunstein, 2003; Hinman and Walter, 2005; ).Despite variations in local
geochemistry and climate, consistent patterns of biogenic and abiogenic textures are present
across multiple localities, forming a well-established model for sinter depositional facies. The
stratigraphic evolution of hydrothermal sinters over geologic timeframes has also been

69 studied in deposits of various ages (Trewin, 1993; Walter et al., 1996; Campbell et al., 2001), 70 including drill cores through recently active localities (Jones et al., 1998; Guidry and Chafetz, 2003b; Lynne et al., 2008, Campbell et al., 2020). However, compared with clastic and 71 72 carbonate environments, several aspects of sinter stratigraphy remain poorly constrained. For example, recent sinter cores only record cooler vent-distal environments such as outflow 73 74 channels and debris aprons, with primary water temperatures interpreted around 30 °C (Jones et al., 1998; Guidry and Chafetz, 2003b; Lynne et al., 2008, Campbell et al., 2020). In 75 contrast, only one study in Devonian strata interprets long-term stratigraphy of higher-76 77 temperature sinter facies (> 60 °C) closer to vent sources (Walter et al., 1996). The study of vent-proximal sinters is especially relevant to the field of geobiology, as higher-temperature 78 79 zones are often inhabited by different microbial communities than cooler outflow deposits. 80 Temperatures extremely close to vents in certain environments can be too hot for life, and 81 distinguishing which sinters are actively mediated by microbes, passively preserving traces of life, or are completely abiogenic, has been a continuing topic of research for decades (Walter, 82 83 1976a,b; Jones et al., 1997; Konhauser et al., 2001, 2003; Lowe & Braunstein, 2003; Handley & Campbell, 2011; Campbell et al., 2015). 84

85 Here, we present sedimentary, geochemical, and geochronological data from four 86 cores drilled within an extinct sinter mound in El Tatio, Chile, including the first stratigraphic 87 profiles of cores from vent-proximal sinters. Facies changes and the preservation of 88 biological communities were observed on various scales using stratigraphic logs, petrographic analyses, and scanning electron microscopy (SEM). Shifts in siliciclastic input 89 90 and primary silica precipitation were studied using geochemical signatures of sinters and 91 underlying sandstone. The detailed sedimentology and stratigraphy of multiple cores provides 92 the basis for a depositional model of the sinter mound, spanning thousands of years in an arid, high elevation geothermal setting. 93

94 **2. Geological setting** 

The El Tatio geothermal field is located in the Andes Mountains of northern Chile, 10 95 km west of the Chile-Bolivia border and 80 km north of San Pedro de Atacama (Fig. 1A). El 96 Tatio contains the world's highest geyser field (~4300 m a.s.l.), and the largest geyser field 97 (~30 km<sup>2</sup>) in the Southern Hemisphere (Zeil, 1959; Trujillo et al., 1969; Jones & Renaut, 98 99 1997; Glennon and Pfaff, 2003; Tassi et al., 2005; Munoz-Saez et al., 2018). The extreme altitude of El Tatio produces a different environment compared with many sinter locations. 100 At 4300 m elevation, atmospheric pressure is 0.58 atm, oxygen partial pressure is 0.12 atm, 101 102 and the boiling temperature of water is 86 °C. Daily temperatures have been measured between -5 and 25 °C during spring (Munoz-Saez et al., 2015), and drop to -30 °C in winter 103 (Fernandez-Turiel et al., 2005), resulting in frequent freeze-thaw cycles (Nicolau et al., 104 105 2014). Wind speeds also follow daily cycles, increasing to 7 km/h several hours after sunrise (Slagter et al., 2019). El Tatio is located in the arid Chilean Altiplano, receiving a mean 106 annual precipitation of 44 mm/yr, with a maximum of 160 mm/yr over the last 40 years 107 (DGA, 2017). Ultraviolet radiation levels are also elevated at the extreme altitudes of El 108 Tatio, with maximum UV-A and UV-B levels at 33 and 6  $W/m^2$ , respectively (Phoenix et al., 109 110 2006). Many environmental factors at El Tatio are analogous to conditions modelled on early Earth and Mars, including high ultra-violet radiation and lower atmospheric pressure 111 112 (Cockell, 2000; Som et al., 2012, 2016; Ruff & Farmer, 2016).

El Tatio occurs above the Altiplano-Puna Magma Body (Ward et al., 2014), which had an extensive explosive phase of eruption during the Miocene-Pleistocene (10-1 Ma) (De Silva, 1989; Leidig and Zandt, 2003). Surficial hydrothermal activity in El Tatio primarily occurs in the hanging wall of a northeast trending half graben, which is filled with ~1 km of sub-horizontal Miocene and Pliocene ignimbrites, tuffs, and lavas, and covered by Pleistocene/Holocene alluvial, glacial and sinter deposits (Healy, 1974; Lahsen and Trujillo,

119 1975). Hydrothermal waters are sourced from meteoric precipitation in the Bolivian Andes at >5,000 m elevation, ~15 km east of El Tatio (e.g., Giggenbach, 1978; Cortecci et al., 2005; 120 Munoz-Saez et al., 2018). Meteoric waters enter the subsurface and are heated by the 121 122 Altiplano-Puna Magma Body as they flow westward (Healy and Hochstein, 1973); heated water interacts with the surrounding host rock to generate the fluids in the El Tatio reservoir, 123 at an equilibrium temperature >200°C (e.g., Tassi et al., 2010; Munoz-Saez et al., 2018). The 124 residence time of reservoir water is still unknown, but the lack of tritium sets a boundary >60 125 years (e.g., Cortecci et al., 2005; Munoz-Saez et al., 2018). Surface geothermal activity 126 127 initiated within the peripheral margins of El Tatio in the late Pleistocene (~27 ka, Munoz-Saez et al., 2020), with continuous activity during the Holocene (~11 ka, Slagter et al., 2019). 128

#### 129 **3. Methods**

130 3.1 Sedimentology & petrography

Four cores (A, C, D, E) were drilled along a vent-proximal (A, C) to vent-distal 131 terrace (D, E) transect on an extinct sinter mound (22° 19' 40.41" S, 68° 0' 23.81" W, Fig. 132 1B-E, 2) in the Upper Basin of El Tatio, referred to in Slagter et al., 2019 as "Mound 502". 133 The cores were drilled vertically, with the exception of core D which was drilled at 27° from 134 vertical, using a modified portable Shawtool Drill machine (diamond bit, 41 mm diameter) 135 lubricated by cooled spring water from an adjacent spring. The cores were sawed in half 136 using a diamond-edged tungsten carbide rotary saw to measure sedimentary features and 137 stratigraphy; the working half of each core was used for petrographic and geochemical 138 sample collection, while the corresponding archival halves were used for non-destructive 139 140 analyses and stored at the National Museum of Natural History of Paris. Thin-sections were made from 27 core samples (Institut Universitaire Européen de la Mer, Brest, France; Thin 141 Section Lab, Toul, France), with standard petrographic techniques used to interpret primary 142

micro-scale fabrics and secondary alteration. Detailed photomosaics of all thin-sections were
created using a Zeiss Axio Zoom V16 Microscope equipped with Zeiss Zen imaging software
(Institut de Biologie Paris-Seine).

146 3.2 Scanning electron microscopy

Scanning electron microscopy (SEM) was performed on rough, unpolished surfaces of 14 core samples fractured into 3-5 mm chunks. Samples were placed directly on aluminium stages with carbon glue, air-dried, and sputter-coated with gold before imaging. SEM images were obtained at the IPGP PARI analytical platform with a Zeiss EVO MA-10, at 15 kV and 12-mm working distance on SE mode.3.3<sup>14</sup>C dating

152 Five sinter samples (10 g each) were cut, cleaned, and ground down: two samples from Core C and four samples from Core E (Table 1, Fig. 2). The fine-grained samples were 153 later treated with 1 N HCl for 24 h and rinsed with ultrapure water. All samples were 154 subsequently digested in 40% hydrofluoric acid (HF). The remaining carbonaceous (non-155 soluble) material was rinsed with ultrapure water and centrifuged. All samples, along with 156 <sup>14</sup>C-free wood and coal blanks, were dried, combusted, and cryogenically purified before 157 being converted to graphite and measured by accelerator mass spectrometry at the UC Irvine 158 Keck-CCAMS facility, using OX-I as the primary standard. The ages are reported in Table 1 159 as years before present (y. BP), where present is AD 1950 (Stuiver and Polach, 1977). 160 Calibrated ages obtained on CALIB (http://calib.qub.ac.uk) using the SHCal13 dataset (Hogg 161 et al., 2013) are also reported as mean ages according to a  $2\sigma$  confidence interval. Ages in the 162 text are reported as calibrated years before present (y. cal. BP).  $\delta^{13}$ C values for samples were 163 164 measured to a precision of <0.1‰ relative to PDB, using a Thermo Finnigan Delta Plus IRMS with Gas Bench input at UC Irvine (Table 1). 165

166 3.4 Major and trace element analysis

Powders from 76 core samples (~150 mg each) were analyzed at the Institut
Universitaire Européen de la Mer (IUEM, Brest). All weighing and preparation of sample
powders for solution mode ICP-MS and ICP-OES was carried out in a class 1000 clean lab at
IUEM. PFA vials were acid washed internally and externally overnight in near-boiling,
concentrated nitric acid (HNO<sub>3</sub>), rinsed three times with 18.2 MΩ-cm ultrapure water, then
dried at room temperature and stored under clean conditions.

For major element analyses, 150-250 mg of powder was weighed into 30ml PFA vials 173 to which 3 ml of analytical-grade concentrated HF, 3 ml of concentrated HNO<sub>3</sub>, and 1 ml of 174 175 concentrated HCl were added. Samples were digested completely overnight at 90°C, with kerogen as the only residue, and upon cooling in closed beakers, were diluted rapidly with 97 176 ml of concentrated H<sub>3</sub>BO<sub>3</sub> solution to neutralize HF and prevent loss of Si as SiF<sub>4</sub>, following 177 178 the method of Cotten et al. (1995). Diluted samples were analyzed for major elements using a Yvon Joriba Ultima 2 ICP-OES calibrated against 8 similarly prepared international 179 geostandards (CB, BELC, JB, BEN, ACE, WSE, ME, and GSN), with pure quartz as a high-180 SiO<sub>2</sub> standard, at the Pôle Spectrométrie Océan (PSO) at IUEM in Brest, France. 181

For trace element analyses, powders were completely digested with a three-step, 182 heated HF-HNO<sub>3</sub>-HCl attack, with kerogen as the only residue, using ultra-pure acids 183 distilled in sub-boiling stills under clean lab conditions. First, approximately 100-150 mg of 184 185 powder was placed in an acid-cleaned PFA vial. 1.5 mL of 16 M HNO<sub>3</sub> and 1.5 mL of 22 M HF were added. Samples were then capped and left to react overnight at 80°C. Samples were 186 then uncapped and left to evaporate to dryness for 24 hours. 3 mL of freshly mixed aqua regia 187 was added to each residue, samples were capped and allowed to react for 6 hours at room 188 temperature to avoid excessive gas buildup. After 6 hours, caps were removed, and the 189 sample was evaporated overnight at 80°C to dryness. Finally, 3 mL of 6 M HCl was added to 190 each residue and samples were heated for ~24h at 80°C. Prior to analysis, 100 mL of each 191

192	sample was diluted to 5 mL	with $2\%$ HNO <sub>3</sub> , and $1.5$ m	L of each sample was
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archived. International geostandards IF-G, BCR-2, and BHVO-2 were prepared in duplicatein the same batch and treated as unknowns to monitor precision.

Diluted samples were measured in solution mode on a Thermo Element2 HR-ICP-MS at the PSO. Samples were calibrated by gravimetrically-prepared multi-element standards at concentrations of 0.5, 5, and 50 ng/g that were measured repeatedly throughout the session. Additionally, a 5 ng/g In (Indium) standard was added to each sample, standard, and rinse, and was used to monitor signal stability and to correct for instrumental drift across the session. Each sample and standard were bracketed by a rinse for which data was also acquired to determine the method detection limit.

Rare earth element (REE) concentrations were normalized against background clastic deposition at Mound 502, as well as traditional standards such as chondritic meteorites and post-Archean Australian shale (PAAS). REE concentrations in Mound 502 clastic deposits were measured from a sinter-free massive sandstone sample from the lowest half-meter of Core E (Fig. 2). REE data shown and discussed below are clastic-normalized, for easier interpretation of clastic vs. hydrothermal influence in sinter deposits. Chondrite- and PAASnormalized REE values are present in Supplementary Table 1.

Eu anomalies were measured to compare the relative influence of hydrothermal fluids between samples (Michard, 1989; Lewis et al., 1997; Feng et al., 2014). Eu anomalies represent enrichments or depletions of Eu compared with Sm and Gd within samples, and are typically calculated in a given sample using the equation:

Eu anomaly = 2\*Eu/(Sm+Gd)

Anomaly values greater than 1 represent Eu enrichment, and are typically interpreted as
evidence for higher hydrothermal influence (Michard, 1989; Lewis et al., 1997; Feng et al.,
2014).

#### 217 **4. Depositional facies**

Ten sedimentary facies were identified during core analysis, each representing a different depositional environment (Fig. 2, Table 2). All facies described have modern equivalents within active El Tatio hydrothermal systems. Figure 2illustrates the progression of sedimentary facies and interpreted environments over time, and Table 2 contains diagnostic features for each facies.

223 Mound 502 rises 2-3 meters above surrounding topography, and is 20 meters in 224 diameter, with a roughly circular outline (Fig. 1C). Western margins are fairly steep, while eastern margins gradually slope into a debris apron. The mound's interior is dominated by a 225 debris-filled pit 1.5 meters deep and five meters wide which represents the inactive former 226 vent, surrounded by a one-meter wide flat-topped sinter rim. As with exterior mound slopes, 227 inner pit margins are steeper on the western face, surrounding the pit's deepest surface in the 228 229 south-western corner (Fig. 1C, D). Mound 502 is bordered to the north by a low-lying marsh with geothermal and meteoric water sources, bordered by vegetated hills further north and 230 west (Fig. 1B). Relatively flat, barren plains border the mound to the south and east (Fig. 1B). 231 While the mound itself contains minimal visible hydrothermal activity, adjacent plains within 232 ten meters to the south and east contain several active pools and widespread fumaroles. 233 Boiling pool margins contain small columnar and spicular sinters (see Section 4.4), but do not 234 235 form mounds above the surrounding flat topography.

Core A (3.38 m long) was drilled within the eastern sinter rim, adjacent to the inner
pit margin (Fig. 1C, D). Cores C (1.94 m long) and D (0.98 m long) were drilled ~1.5 meters

below the top of Core A along the outer eastern and southern margins, respectively. Core D
was drilled 27° from vertical (Fig. 1E). Core E (5.57 m long) was drilled in a flat debris apron
15 meters away from the sinter margin.

241 4.1 Inclined laminated sinter crust

Laminated sinter crusts are defined by dense, surface-normal fabrics of amorphous silica, forming mm-scale laminae of consistent thickness (Fig. 3). Similar deposits in Yellowstone siliceous sinters have been labeled "vent-proximal facies" and "laminated crusts" (Guidry & Chafetz, 2003a, c), as well as "dendritic sinter facies" (Lowe & Braunstein, 2003). The term "dendritic" is avoided in this study to prevent confusion with branching arborescent sinter elsewhere in El Tatio cores (see Section 4.5).

248 Laminated crusts are divided into two separate facies in Mound 502 based on bed orientation and depositional interpretation: inclined crusts and planar crusts (Section 4.2). 249 Inclined crusts are only observed in Core A, between 34 and 217 cm depth, with a shift in 250 textural preservation at 164 cm depth (see below). The inclination of laminae and bedding 251 varies between 45° and nearly vertical, with most areas around 70° or higher inclination (Fig. 252 253 3). Well-preserved zones are composed of silica laminae 1-2 mm thick, alternating between pale surface-normal fabrics and thinner (< 1 mm), darker layers of more massive silica (Fig. 254 3A-C). No microbial textures or cells are observed, and detrital grains are rare or absent. 255

Below 164 cm depth in Core A, inclined crusts are gray or light green and have little to no porosity. Detailed fabrics are only visible in thin sections, where individual laminae contain surface-normal clusters of amorphous silica, which flare and diverge away from bedding planes (Fig. 3D). Preservation varies between individual laminae, with more altered layers only exhibiting faint layering (Fig. 3D). Samples also contain dark, linear cross-cutting

features occurring 1-3 cm apart (Fig. 3B). Cross-cutting textures disrupt and propagate
through several centimeters of inclined crusts, independent of textural preservation.

Above 164 cm, inclined crusts are more well-preserved, and surface-normal fabrics 263 become increasingly visible in hand samples (Fig. 3A-C). In the uppermost facies (34-51 cm 264 depth), inclined crusts are white, extremely friable, and contain very few detrital grains (Fig. 265 3C). Inclination decreases upward from 70 to  $45^{\circ}$ , with rare cross-cutting textures present. In 266 SEM images of the uppermost inclined facies, surface-normal fabrics are composed of 267 elongated tubes of silica ~50 µm wide and up to 1 mm in length (Fig. 3E, F). Closer 268 269 examination of individual tubes reveals irregularly porous cores with scalloped circular 270 textures, surrounded by µm-scale laminae around the outermost tube margins (Fig. 3F). Rare silicified filaments are also present within individual laminae, comprising less than 5% of 271 272 inclined textures (Fig. 3G, H). Filaments are smaller than surrounding surface-normal tubes (10-20 µm wide, <200 µm long), and are not consistently oriented (Fig. 3G, H). Instead, 273 filaments more closely resemble preserved cells in microbial facies (Figs. 8, 10, 11, see 274 Sections 4.5-4.7). 275

Laminated sinter crusts are observed in subaqueous hydrothermal environments 276 where silica precipitation typically outpaces microbial mat formation, such as vent-proximal 277 pools (Guidry & Chafetz, 2003a, Lowe & Braunstein, 2003). SEM images most closely 278 resemble "shrub columns" with abundant porous silica structures described by Jones & 279 Renaut (2003) from vent-proximal deposits in New Zealand. Internal porosity within 280 branches of shrub columns are attributed to secondary dissolution, which can replace 281 microbial cells or abiogenic mineral dendrites (Jones & Renaut, 2003). While porous surface-282 normal textures are ambiguously biogenic, the occasional presence of filamentous 283 microfossils indicates that microbial communities were present. 284

The high angle of lamination over nearly 2 meters suggests extensive precipitation on 285 steep surfaces, most likely the inner subaqueous surfaces of a large high-temperature pool. 286 The interpretation fits with Mound 502's current near-vent architecture, which consists of a 287 288 5-meter-wide steep-sided pit. The darker surface-normal lineations which disrupt laminae are potentially the propagations of corrugated or "rilled" bedforms commonly seen on vent walls 289 (Braunstein & Lowe, 2001; Guidry & Chafetz, 2003a; Lowe & Braunstein, 2003). Therefore, 290 291 the inclined fibrous facies in Core A are interpreted as the margins of a steep-walled pool immediately above a hydrothermal vent, most likely above 70 °C. 292

293 4.2 Planar laminated sinter crusts

Planar laminated crusts contain thin layers of surface-normal silica fabrics (Fig. 4), 294 similar to inclined sinter crusts (see Section 4.1), but exhibit consistently horizontal bedding. 295 296 Planar crusts are only present in Core A from 11 to 34 cm depth, immediately above inclined crust facies. Planar and inclined crusts are separated from each other by a sharp 297 298 unconformity. As with inclined crusts, laminae in planar facies are 1-2 mm thick, containing pale surface-normal fabrics separated by thin, darker massive laminae (Fig. 4A). Boundaries 299 between light and dark layers are more distinct in planar than inclined fibrous facies. Detrital 300 301 sinter pieces and siliciclastic material are rare, but more abundant in darker laminae (Fig. 4A). Porosity in both light and dark layers is low on mm- to cm-scales. 302

SEM imaging of planar sinter crusts reveals two distinct styles of preservation (Fig.
4B,C). Light laminae are composed of adjacent vertical chains of silica spherules ~100 µm in
diameter (Fig. 4B). Vertical textures in light laminae superficially resemble surface-normal
textures within inclined sinters, but notably lack porous, scalloped cores (Figs. 3E, F, 4B). By
contrast, dark laminae contain linear highly-porous zones 20 µm wide, surrounded by rinds of

amorphous silica (Fig. 4C). Pores in planar crusts lack consistent orientation, and are
typically smaller and less interconnected than in inclined facies.

As with inclined sinter crusts, planar facies are observed within vent-proximal hightemperature deposits (Guidry & Chafetz, 2003a, Lowe & Braunstein, 2003). In Mound 502, planar laminated crusts are only observed within the flat-topped inner rim which lines the inner pit (Fig. 1D). While inclined crusts were most likely deposited on the steep subaqueous margins of the main inner pit, planar crusts potentially represent adjacent, shallower flatbottomed pools.

Lowe & Braunstein, 2003 described scattered microfossil filaments in Yellowstone 316 sinter crusts ("dendritic sinter facies"), though many more samples contained no trace of 317 microbial presence. Similarly, planar crusts in El Tatio cores contain no definitive 318 319 microfossils, instead sharing similarities with ambiguously biogenic "shrub columns" described by Jones & Renaut, 2003. The lack of biosignatures does not preclude a microbial 320 321 presence in planar sinter crusts, but supports a silica precipitation patterns less conducive to microfossil preservation (Walter, 1976b; Lowe & Braunstein, 2003; Hinman & Walter, 322 2005). 323

324 4.3 Particulate sinter

Particulate sinters are defined as beds of "botryoidal grains and chains" over 5 μm in diameter, frequently associated with detrital grains (Braunstein & Lowe, 2001; Lowe & Braunstein, 2003; Watts-Henwood et al., 2017). Particulate sinter facies occur at the top of Core A, and in two zones within Core E (103-125 cm, 165-195 cm depth). Particulate beds are typically gray with brown detrital material (Fig. 5A), and are closely associated with laminated sinter crusts, arborescent sinters (see Section 4.5), and massive sinters (see Section 4.8).

Particulate sinters are defined by diffuse planar to wavy laminae between 1 and 5 mm thick, comprised of rounded 50-100 µm-wide agglomerates of opal spherules surrounded by pore space (Fig. 5B-E). Well-preserved hand samples of particulate facies show similar textures to associated thin sections (Fig. 5A), though many others are difficult to discern from massive sinter facies prior to microscopic analysis (see section 4.8). Detrital material is abundant in specific layers as individual sand grains (Fig. 5C). No microfossils or preserved microbial mat textures are visible using petrographic or scanning electron microscopy.

Particulate facies are interpreted as high-temperature deposits along the bases of non-339 boiling hydrothermal pools, which occur both proximal and distal to vent sources. Spherical 340 silica particles precipitate onto various substrates such as microbial cells and wind-blown 341 detrital sediments, forming a sticky, laminar ooze comprised of amorphous, wetted silica and 342 343 thin biofilms (Rimstidt & Cole, 1983; Lowe & Brownstein, 2001; Braunstein & Lowe, 2003; Watts-Henwood et al., 2017). Microbial cells can form a major component of particulate 344 sinters, preserving silicified cells or cell fragments of microbes from various temperature 345 regimes (Jones & Renaut, 1996; Braunstein & Lowe, 2003). Sediments are agitated by wind 346 and hydrothermal currents, forming sinuous, ripple-like bedforms on pool floors (Lowe & 347 348 Brownstein, 2001; Watts-Henwood et al., 2017). Similar textures are observed in extinct 349 sinter mounds elsewhere in El Tatio, bearing irregular surfaces of low-relief ridges (Sanchez-350 Garcia et al., 2019). Modern high-temperature pools are present within debris aprons several 351 meters downhill from Core E, providing a useful analogue for distal pool environments indicated by particulate facies. 352

353 4.4 Columnar and spicular sinter

Cores C and E contain several distinct facies dominated by clusters of elongate sinter protrusions with convex laminae, which can be classified as columns and spicules (Figs. 6,

7). *Columns* are between 0.1 and 1 cm wide, and can exceed 5 cm in length (Walter et al.,
1976b). *Spicules* are less than 1 mm in diameter, reaching up to 1 cm long (Walter et al.,
1976b). The internal structure of spicules and columns are nearly identical, with definitions
for each based on size (Walter, 1976b), though terminology varies slightly between studies
(Jones & Renaut, 1997, 2003, 2006; Braunstein & Lowe 2001).

361 Columnar sinter is most prevalent in Core C, from 70 to 167.5 cm. Individual columns are 0.5-1 cm wide, and 1-5.5 cm long (Fig. 6C-E). Core C often truncates column 362 lengths, but surface exposures of columnar facies verify the observed size ranges (Fig. 6A, 363 B). In transverse cross-section, columns are typically circular to oval in shape (Fig. 6E). 364 Columns frequently exhibit branching morphologies, with individual branches reaching no 365 more than 30° of inclination from parent columns (Fig. 6C, D). The only change in column 366 morphology occurs at 123.5 cm in Core C, where column orientation switches from 367 horizontal to 60° inclination. The change in angle occurs above an unconformity with planar 368 laminated sinter (Fig. 6E). 369

370 Sinter columns in Core C are internally laminated, with laminations visible in longitudinal and transverse cross-sections (Figs. 6, 7). Laminae are convex in the direction of 371 growth, either upwards in angled columns, or laterally in horizontal columns (Fig. 6C, D, 7A, 372 C, D). The synoptic relief of individual layers within columns is relatively low, between 0.5 373 374 and 2.0 mm (Fig. 7A, C, D). Layers have low porosity, with sub-mm pores scattered throughout columns. SEM imaging also reveal column layers (Fig. 7F), with the rare 375 presence of hollow silica tubes on the outer surfaces of columns (Fig. 7G). Interstitial gaps 376 between columns are porous, with coarse or fine sediments on the bottom of pores serving as 377 geopetal, or way-up, indicators (Fig. 7D, E). Fine geopetal sediments are often cross-cut with 378 secondary silica veins (Fig. 7E). Other mineral phases include rims of yellow and brown 379 oxidized crusts pervasively surrounding columnar edges (Fig. 7A, C, D). The presence of 380

crusts on the ends of broken columns indicates secondary precipitation after initial columngrowth (Fig. 7C).

383 Columnar sinter is also present at 364 cm depth in Core E (Figs. 6F, 7B), within a 3 cm thick bed. Core E columns are vertical features ~0.5 cm in width and 1 cm in height, 384 occurring in isolated clusters of one to three structures. Internal morphology closely 385 386 resembles Core C columns, bearing convex-upwards laminae with 0.5 mm relief (Figs. 6F, 7B). Laminae in Core E columns are thicker than Core C (0.25 - 0.5 mm), and contain higher 387 amounts of detrital material Lower columnar margins have minor yellow mineral crusts, 388 389 lacking dendritic textures observed in Core C. Columns are immediately overlain by tufted microbial sinters, but no evidence of significant microbial communities is present at the time 390 of column deposition. 391

Spicular sinter occurs in a 5 cm thick bed immediately below columnar sinters in Core
C, at 175 cm depth. Spicules are 1 mm wide and up to 1 cm long, with individual structures
curving upwards from 30 to 70° inclination (Fig. 6G). As with Core C columnar sinters,
spicules are tightly packed and bordered by yellow and red mineral crusts. Laminae are not
observed in spicule cross-sections. The spicule layer is immediately and unconformably
overlain by horizontal columnar sinters.

Columnar and spicular sinters are commonly observed around boiling pools and
geysers, in zones which exhibit frequent wetting and drying (Walter, 1976b; Jones et al.,
1997, 1998; Jones & Renaut, 1997, 2003, 2006; Braunstein & Lowe, 2001; Guidry &
Chafetz, 2003a, Lowe & Braunstein, 2003; Boudreau & Lynne, 2012). Splashes and surges of
hot, silica -rich water evaporate on nearby surfaces, precipitating opal-A. Over time, silica
precipitation produces micro-scale topography that progresses into columnar morphologies,
with capillary motion of water within columns providing silica over time (Walter, 1972,

405 1976b; Braunstein & Lowe, 2001; Jones & Renaut 1997, 2003, 2006; Mountain et al., 2003). Splash zones close to pools and vents form spicular textures, while larger columns occur 406 further away in surge and overflow environments (Walter, 1972, 1976b; Braunstein & Lowe 407 408 2001). The upward increasing angle of columnar growth in Core C is likely due to local topography, with horizontal columns growing on nearly-vertical mound faces, and increasing 409 angles of growth on progressively flatter surfaces closer to the tops of sinter rims. Incipient 410 411 growth of vertical columns in El Tatio occurs around modern boiling pools several meters south of Core E (Fig. 6B). The modern pool-margin textures are nearly identical in size and 412 413 orientation to Core E columns. Therefore, columnar sinters from Cores C and E are interpreted as silica precipitates that formed in the former splash and surge zones of boiling 414 pools. 415

416 Columnar sinters frequently preserve microbial communities, both within columnar laminae and interstitial spaces between columns (Cady & Farmer, 1996; Jones et al., 1997, 417 1998; Guidry & Chafetz, 2003; Jones & Renaut, 2003, 2006; Handley et al., 2005, 2008). In 418 other columnar sinter locations, biosignatures are less apparent (Walter, 1976b; Braunstein & 419 Lowe, 2001; Lowe & Braunstein, 2003). Most columnar facies are interpreted as forming via 420 421 biogenic and abiogenic silicification, with capillary forces, surge events, and precipitation 422 within microbial communities dominating at different points in time (Jones et al., 1997; Jones 423 & Renaut, 2003, 2006). In El Tatio cores, microbial textures in columnar sinter are obscured 424 by diagenesis, with only isolated, putative microfossils within inter-columnar spaces (Fig. 7G). The columnar sinters examined are conservatively interpreted as forming through 425 capillary forces after periodic surge events, with microbial communities likely present in 426 427 cooler, moist areas between columns.

428 4.5 Arborescent sinter

429 Arborescent sinter is defined in this study as horizontal beds composed of surfacenormal, extensively branching siliceous textures up to 1 cm in height (Fig. 8). The same 430 textures are nearly identical to "siliceous shrubs" described in detail by Guidry & Chafetz, 431 432 2003c. However, the term "shrub" has been used to describe various different sinter textures such as laminated sinter crusts ("dendritic sinter facies" in Lowe & Braunstein, 2003), 433 microbial palisade fabrics (Cady & Farmer, 1998; Walter et al., 1998; Campbell et al., 2001; 434 see Section 4.6), and geyser-associated facies (Jones & Renaut, 2003). Similarly, the term 435 "dendritic" is commonly applied to branching amorphous or mineral phases, and has been 436 437 used in various sinter fabrics, such as laminated crusts, spicules, and microbial mats (Jones & Renaut, 2003; Lowe & Braunstein, 2003; Hinman & Walter; 2005), as well as phases other 438 than amorphous silica, including quartz, calcite, and manganese (Bargar & Beeson 1981; 439 440 Jones et al., 1998; Walter et al., 1998; Guidry & Chafetz, 2003b). By contrast, the term "arborescent" is only applied to siliceous shrubs as described in Guidry & Chafetz, 2003a & 441 c, the previously published descriptions which most closely resemble the El Tatio sinter 442 443 facies presented here.

Various horizons in Cores C and E (Fig. 2) contain arborescent sinter. In Core C (0-3, 444 445 30-34 cm depth) and deeper in Core E (310-320 cm depth), arborescent sinters form minor 446 units of 1 cm-thick beds interspersed between sandstones, microbial, and massive sinter 447 facies (Fig. 8A-B). Arborescent layers higher in Core E form persistent beds between 142-448 187 cm (Fig. 8C) and are closely associated with particulate sinters (see Section 4.3). In Cores C and E, arborescent sinter contains individual branching structures that reach up to 1 449 cm high, with individual branches less than 1 mm in diameter (Fig. 8A-C). Arborescent zones 450 451 form highly porous areas compared with over- and underlying facies, and are often separated 452 by mm-scale interstitial layers of less porous sinter (Fig. 8B, C). Interstitial sinter layers also

453 contain higher abundances of detrital grains, often appearing darker in color than white or454 light gray arborescent layers (Fig. 8B, C).

455 Petrographic analysis of arborescent structures reveals branching chains of amalgamated siliceous spheres up to ~100 µm in diameter (Fig. 8D). Individual arborescent 456 features and branches are often separated by pore space, with more extensive silica 457 458 precipitation resulting in smaller interstitial voids. SEM images show similar branching textures as observed in thin sections (Fig. 8E), as well as the presence of silicified 459 microfossils and higher quantities of detrital material at the top of individual arborescent beds 460 (Fig. 8E-G). Microbial filaments are 10-20 µm wide and over 200 µm long, resembling cells 461 preserved in palisade mat and tufted mat facies (see Sections 4.6 & 4.7). Microfossils are 462 found in clusters within and around larger branching features at lower densities than other 463 464 microbial facies in El Tatio cores (Figs. 10, 11).

In active hydrothermal systems, arborescent sinters form within discharge channels as 465 hydrothermal waters flow away from vent sources (Guidry & Chafetz, 2003a,c). Within such 466 depositional environments, arborescent facies appear to be temperature-independent, forming 467 in subaqueous zones between 15 and 80 °C (Guidry & Chafetz, 2003c). Microbial 468 communities preserved in El Tatio arborescent facies corroborate findings from similar 469 Yellowstone sinters (Guidry & Chafetz, 2003c). Guidry & Chafetz, 2003c noted the presence 470 of branching "siliceous shrubs" in discharge channels with higher outflow volume, while 471 laminated microbial mats were abundant in lower outflow zones. Based on these 472 observations, arborescent sinters in El Tatio cores are interpreted as silicified microbial 473 communities in high-volume discharge channels, both in vent-proximal zones and distal 474 aprons. 475

476 4.6 Fenestral palisade sinter

477 Palisade sinters are defined by silicified filamentous microfossils alternating between layers of surface-normal and surface-parallel orientation (Cady & Farmer, 1996; Walter et al., 478 1996; Campbell et al., 2001; Gong et al., 2019). Similar textures are observed, but not 479 480 specifically labeled as palisades, in hydrothermal stromatolites (Konhauser et al., 2004; Jones et al., 1997, 1998, 2005; Berelson et al., 2011), with a variety of terms applied to microfossil 481 orientation (erect vs. prostrate, U- vs. P-laminae, and upright vs. flat-lying). Palisades occur 482 483 in the top 50 centimeters of Cores C, D, and E, and between 215 and 244 cm depth in Core E. Most El Tatio core palisades contain high abundances of fenestrae (Figs. 9, 10), defined as 484 485 primary pore spaces preserved within microbial textures (Tebbutt et al., 1965; Mata et al., 2012). 486

Fenestral porosity alternates between individual laminae, resulting in light-dark 487 488 layering within hand samples and thin sections. Lighter, more porous laminae are 2-5 mm thick, while darker laminae are less than 2 mm thick and contain fewer fenestrae. Detrital 489 grains are present but not abundant throughout light and dark palisade laminae (Fig. 9E, G). 490 Both styles of lamination contain filamentous microfabrics visible in thin-section and SEM 491 492 imaging (Figs. 9D-G, 10A-E). Filaments in less-porous laminae are oriented parallel to 493 bedding planes. In contrast, filaments in porous laminae contort around curved fenestral 494 margins in various orientations.

Palisades in Cores C, D, and E differ in fenestral size and morphology (Fig. 9A-C).
Fenestrae in Cores C and E are sub-mm in scale, with sub-circular margins in cross-section
(Fig. 9A, C). By contrast, Core D fenestrae have elongated oval or lobate morphologies
between 0.5 and 5 mm in diameter (Fig. 9B, D). Elongate fenestrae are especially prevalent
in thicker porous layers of Core D. Long axes of such fenestrae are oriented normal to
bedding planes, separated by thin vertical columns of silicified microbial filaments (Fig. 9D).

501 Darker layers of Core D contain smaller, more circular fenestrae comparable to morphologies
502 observed in Cores C and E.

Filamentous microfossils are abundant in all fenestral palisades (Fig. 10A-C). In SEM 503 504 images, Cores C and D exhibit round voids lined with 5-µm wide filaments up to 100 µm in length (Fig. 10A, B). Filaments retain equal diameter throughout their length, terminating 505 506 with rounded, untapered ends. Individual microfossils frequently form bridges across fenestral pore spaces. While most sinters between fenestrae only preserve microfossils on 507 exterior surfaces, exquisitely preserved regions reveal columnar bundles of filaments 508 509 stretching across pore spaces (Fig. 10D, E). Secondary overprinting is especially prevalent in Core E, where microfossils are still abundant, but rounded fenestral textures are less distinct 510 (Fig. 10C). 511

512 Palisades and similar microbial fabrics are globally present in hydrothermal sinter deposits of various ages (Walter, 1976a; Walter et al., 1996, 1998; Jones et al., 1997, 1998, 513 2005; Campbell et al., 2001, 2015; Konhauser et al., 2004; Fernandez-Turiel et al., 2005; 514 Berelson et al., 2011; Gong et al., 2019). Variations in filament orientation are interpreted 515 differently in various locations, including seasonal growth patterns (Walter, 1976a; Jones et 516 517 al., 1998, 2005), pulses in hydrothermal sheet-flow (Campbell et al., 2015), and quasi-daily shifts in water temperatures (Berelson et al., 2011). On steeper surfaces such as the edges of 518 sinter mounds, palisade formation is driven by cycles of wetting and drying (Gong et al., 519 2019). 520

Palisade sinters closely resemble cyanobacterial communities living at or below 40-45
°C, typically interpreted as *Calothrix* (Walter, 1976a; Cady & Farmer, 1996; Campbell et al.,
2001, 2015), though similar textures have been noted in communities of *Fischerella*, *Phormidium*, and *Leptolyngbya* (Jones et al., 2005; Pepe-Ranney et al., 2012; Gong et al.,

2019). In active El Tatio geyser terraces, morphological and molecular evidence show a
predominance of *Leptolyngbya*, and an absence of *Calothrix* or its more halotolerant sister
genus, *Rivularia* (Gong et al., 2019). Due to the robust structures of Calothrix-type
organisms, their absence is not likely due to a preservation bias, but rather interpreted as
absence from the characterized terraces.

530 In many sinter localities, erect palisade filaments are often tightly packed together, leaving no space for mm-scale fenestrae. However, several locations contain more porous 531 fenestral palisades in distal sinter aprons and terraces, including other zones in El Tatio 532 533 (Jones et al., 1998; Munoz-Saez et al., 2016). One environmental exception occurs in Obsidian Pool Prime in Yellowstone National Park (Berelson et al., 2011; Mata et al., 2012), 534 where fenestral textures occur in stromatolites around higher-temperature pool margins (>45 535 536 °C). However, Obsidian Pool Prime fenestrae have distinctly different morphologies from El Tatio cores, forming sub-spherical voids less than 1 mm in diameter. Instead, fenestrae in 537 core samples more closely resemble elongate pores observed in cooler, distal settings 538 elsewhere in El Tatio and New Zealand (Jones et al., 1998; Munoz-Saez et al., 2016). 539

540 4.7 Tufted mat sinter

Tufted mat sinters are defined as filamentous microbial mat textures with sharp peaks 541 and cones irregularly separated by 1-2 mm concave depressions (Cady & Farmer, 1996; 542 Walter et al., 1996, 1998; Campbell et al., 2001; Jones et al., 2005), most notable in thin-543 section (Fig. 11A, B). Tufted facies occur within two zones of Core E (0-103 cm, 289-367 cm 544 depth), interspersed as individual beds up to 15 cm thick. The top meter of Core E contains 545 546 tufted beds interspersed between palisade fabrics, while lower units briefly occur between 547 thicker units of arborescent sinters and laminated sandstone deposits. Individual tufted layers are ~100 µm thick, and are difficult to discern in hand samples. Laminae laterally increase in 548

thickness to 200 µm at tufted peaks, and diminish to 50 µm in interstitial depressions.
Textures immediately below tufts vary between layers. Most laminae do not preserve
additional microbial textures below tufts (Fig. 11B), while some laminae contain relict
microbial filaments (Fig. 11A). Occasionally, both highly circular and elongate fenestrae are
preserved, bounded by thin pillars of filamentous sinter (Fig. 11C).

SEM imaging of tufted sinters shows exquisite preservation of microbial mat textures.
Three-dimensional tufts are composed entirely of microbial filaments between 2-5 μm in
diameter (Fig. 11D-F). Filaments form diffuse meshwork textures at the base of tufts, while
sharp peaks exhibit denser accumulations, reflecting differences in thickness and density
observed in thin-sections (Fig. 11A, B). Large rounded fenestrae are also present, bounded
and separated from each other by silicified filaments (Fig. 10F, G). Fenestrae are less
frequent in tufted facies than palisades, but are larger and more spherical.

Tufted facies closely resemble modern mats dominated by the cyanobacterial taxa 561 Phormidium and Leptolyngbya (Walter, 1976a; Campbell et al., 2001; Fernandez-Turiel et 562 al., 2005; Lau et al., 2005; Bosak et al., 2012; Reyes et al., 2013, Bradley et al., 2017). Many 563 tufted mats grow between 45 and 60 °C (Walter, 1976a; Campbell et al., 2001; Lau et al., 564 2005; Bradley et al., 2017), though tufted *Leptolyngbya* mats are also reported between 17 565 and 45 °C (Bosak et al., 2012; Reves et al., 2013). Large rounded fenestrae below tufts are 566 also distinct from fenestral palisade textures, more closely resembling spherical gas bubbles 567 preserved within tufted hydrothermal mats (Lynne & Campbell, 2003; Bosak et al., 2008, 568 2010; Mata et al., 2012). Tufted facies are interspersed with palisade fabrics in the top meter 569 of Core E, with changes in mat morphology potentially facilitated by changes in water 570 temperatures or hydrochemistry. 571

572 4.8 Massive sinter

Massive sinters exhibit either obscured or no primary features such as laminae, 573 bedding, detrital grains, or microfossils (Fernandez-Turiel et al., 2005; Walter et al., 1996). 574 Massive sinter facies occur at various depths in Cores A, D, and E. The lowest meter of Core 575 576 A consists of green and gray low-porosity sinter with pale yellow and orange minerals along secondary fissures (Fig. 12A). While massive Core A sinter contains vertically oriented 577 lineations in hand sample similar to inclined sinter crusts, all traces of vertical fabric are 578 obscured in thin section (Fig. 12B). Sinters at the base of Core D (72 - 97 cm) have similarly 579 low porosity, and lack bedding, vertical lineation, or clear laminae in hand sample or thin 580 581 section (Fig. 12C). Massive sinters occur in 5-cm thick beds at various depths in Core E (130-135 cm, 320-325 cm, 330-335 cm). Core E facies are lighter in color than in Cores A and D, 582 and have gradational contacts with particulate and arborescent sinters. Microscopy of massive 583 584 sinters and adjacent facies also shows extensive growth of secondary silica phases within pore spaces (Fig. 12D, E). 585

Massive sinter fabrics are present in modern and ancient sinter deposits (Walter, 586 1976a; Walter et al., 1996; Guidry & Chafetz 2003a). The lack of primary fabrics in hand 587 samples and thin-sections implies secondary alteration, such as recrystallization and 588 589 overgrowth of original sinter deposits, most likely by high temperature silica-rich subsurface 590 fluids. Adjacent primary facies in Cores A and E support a high-temperature alteration 591 hypothesis. Massive sinter in Core A is overlain by more than 2 meters of inclined and planar 592 sinter crusts, representing the sides and floors of vent-proximal pools respectively. In Core E, massive sinters are associated with particulate sinters, also interpreted as subaqueous deposits 593 near secondary hydrothermal vents distal to Mound 502. In contrast, facies at the base of 594 595 Core D are not associated with high-temperature pool or vent facies. Based on high-596 temperature associations elsewhere, massive sinters in Core D are interpreted as former

597 microbial mat sinters which were secondarily altered by subsurface fluid flow and598 recrystallization.

599 4.9 Laminated sandstone and sinter

Laminated sandstone and sinter (LSS) deposits are defined as interbedded siliciclastic and sinter layers between 0.1 and 5 cm thick. The LSS facies is only observed in the lower 2.5 meters of Core E. Sandstones are medium- to coarse-grained, and are composed of subrounded to sub-angular volcaniclastic grains and occasionally detrital sinter pieces (Fig. 13A, C, E). No evidence of cross-bedding is observed in sandstone beds, while sinters contain millimeter- scale planar laminae lacking fenestrae or filaments. Laminated sinters often have irregular, unconformable contacts with overlying sandstones.

The ratio of sinter to sandstone in LSS facies varies in different areas of Core E.
Below 4 m depth, sandstone layers are thicker, containing isolated, centimeter-scale sinter
deposits. Unconformities separate lower LSS facies from intervening massive sandstones
lacking sinter in Core E (described below). LSS facies above 4 meters in Core E are
dominated by planar sinter, punctuated with sub-centimeter sandy layers. Upper LSS facies in
Core E are typically less than 5 cm thick and are often under- or overlain by tufted microbial
textures, distinguished by filamentous fabrics.

LSS facies in Core E are interpreted as distal sheet-flow deposits, with sinter layers representing relative lulls in siliciclastic input. Beds below 4 m depth indicate higher-energy environments where clastic sedimentation prevented persistent sinter deposition. In contrast, LSS facies above 4 m depth exhibit minor pulses of siliciclastic deposition in a calmer, sinterdominated regime. In both zones, sinter deposits were consistently eroded by subsequent sandstone deposition, forming irregular upper contacts.

620 4.10 Massive sandstone

Massive sandstones contain no bedding textures or interstitial sinter deposits, and are only found in the lowest meter of Core E. Similar to LSS facies, sands are medium to coarsegrained, and are composed of sub-rounded and sub-angular volcaniclastic grains (Fig. 13B, D, F). The majority of siliciclastic material is composed of volcanic lithic fragments, feldspar, mica and minor portions of redeposited sinter material. The sandstone is notably free of detrital quartz, but largely cemented by siliceous material such as opal and chalcedony (Fig. 13F).

Massive sandstone units have unconformable boundaries with adjacent sinter and 628 629 sandstone beds, and occasionally contain unconformities within otherwise homogenous units. One notable unconformity occurs at 538 cm depth in Core E. Sandstones containing 1-2 cm 630 wide matrix-supported rounded green clasts are conformably overlain by planar beds of sands 631 632 fining upward from coarse to medium-grained. Both units are eroded by an unconformably overlying homogenous, coarse-grained sandstone. Vitreous brown cements follow the 633 unconformity, cross-cutting into over- and underlying sediments (Fig. 13B). No other 634 sedimentary structures are apparent in massive sandstones. 635

Massive sandstones are interpreted as debris flows with little to no hydrothermal
influence. The only hydrothermal silica precipitation present (538 cm) occurs as secondary
precipitation following unconformities and cross-cutting primary lithologies in a narrow
zone.

#### 640 **5. Core geochemistry**

641 Sinters in Cores A, C, D, and the upper 4 meters of Core E are largely composed of
642 siliceous material with more than 94 wt.% of SiO<sub>2</sub> (Fig. 14). Cores closer to the vent source
643 (A, C, D) are especially enriched in SiO<sub>2</sub>, with all measured values above 98.5 wt.%. SiO<sub>2</sub>
644 concentrations in Cores A and C decrease towards surface deposits, with higher values in

sinter crusts and columnar sinters, and lower values in mat and particulate sinter facies. Core D shows little variability between massive sinters and palisade mat facies, while sandstones at the base of Core E are relatively SiO<sub>2</sub>-depleted (70-85 wt.%), with a sharp increase to >90 wt.% at 361 cm depth. SiO<sub>2</sub> concentrations in Core E increase to ~99 wt.% in arborescent, particulate and massive sinters between 1-2 m depth before slightly decreasing within the uppermost meter of mat facies, in a similar fashion to Cores A and C (Fig. 14).

Other major elements are anti-correlated to SiO<sub>2</sub> concentrations in all cores (Fig. 14, 651 Supplementary Information). For example, Al<sub>2</sub>O<sub>3</sub>, Fe<sub>2</sub>O<sub>3</sub>, TiO<sub>2</sub>, etc. increase upwards in 652 Cores A and C as SiO<sub>2</sub> decreases. In Core E, sandstones below 4 m depth have different 653 geochemical signatures than overlying sinters. Many elements are several orders of 654 magnitude more abundant in Core E sandstone deposits (Fig. 14), including Al<sub>2</sub>O<sub>3</sub> 655 656 concentrations above 10 wt.%. In Core E sinter facies above 4 m depth, elemental abundances reach minimum values at ~2 m (e.g., Al<sub>2</sub>O<sub>3</sub> concentrations at 0.1 wt.%) before 657 slowly increasing within the top meter of mat sinter facies. While certain elements are 658 elevated in individual layers (e.g., P<sub>2</sub>O<sub>5</sub> in Core E, 103 cm), larger shifts are consistent 659 between elements in all cores. 660

Most trace elements exhibit similar stratigraphic trends to major elements in all four 661 cores. For example Li, Cs, Rb, Se, U, Cr, and many other elements are enriched in Core E 662 sandstones below 4 m depth, and depleted in overlying sinters, with slightly higher 663 concentrations of Li and Cs (Supplementary Information). In rare cases, certain elements are 664 enriched in sinters compared with clastic deposits. For example, Sb concentrations increase in 665 Core E sinters above 4 m, and represent the most abundant trace element in all sinter samples 666 (Supplementary Information). As is slightly more abundant in sinter than sandstone facies, 667 with elevated concentrations in the top 0.5 m of all cores (Supplementary Information). 668

Rare earth element (REE) core profiles are nearly identical to most major and trace
elements, with strong anti-correlation to SiO<sub>2</sub> concentration (Fig. 14, 15). When REE
concentrations of individual samples are normalized to clastic deposits of Mound 502, neither
light nor heavy rare earth elements (LREE and HREE) are enriched (Fig. 15, Supplementary
Information). Slight HREE enrichment occurs when core samples are normalized to
chondritic meteorites and post-Archean Australian shale (PAAS) (Supplementary
Information).

676 Total REE abundance increases towards the tops of Cores A and C, with a slight decrease around 1.2 m depth in Core A (Fig. 15). Core D has consistent REE concentrations 677 over a limited sampling distribution. As with major elements, REE concentration in Core E 678 varies greatly based on depth and facies. Sandstones below 4 m depth have concentrations 679 680 one to two orders of magnitude higher than overlying sinters, forming a sample population distinct from facies in any other core (Fig. 15). REE abundance drops drastically by 681 approximately one order of magnitude at 4 m depth, when sinter becomes more abundant 682 relative to sandstone. Between ca. 3.6 and 3.0 m depth, REE abundance decreases steadily by 683 one order of magnitude where sandstone bearing beds become less abundant (Fig. 15). 684 685 Arborescent and particulate sinters between 1 and 2 meters contain the lowest REE 686 concentrations in Core E, with values gradually increasing in overlying mat facies.

Different core facies also vary in REE patterns between individual samples (Fig. 15).
When normalized to values of clastic deposits within Mound 502, REE patterns are
essentially flat in facies with elevated concentrations, such as LSS and various microbial mat
facies. In samples with depleted REE, certain elements are relatively enriched, including La,
Eu, Y, and Lu (Fig. 15). Enrichment of such elements typically increases as total REE
concentration decreases, and is most prevalent in sinter crust, particulate, and columnar sinter
facies. For example, positive Eu anomalies (enrichments in Eu relative to standards) are

highest in sinter crust and columnar facies in Cores A and C, respectively (Fig. 15, 16). The
highest positive Eu anomalies in Core E occur in particulate sinters between 1.5 and 2.0 m
depth, with all other facies exhibiting slightly negative or no Eu anomalies (Fig. 16).

697

### 6. Discussion: Core depositional histories

Depositional interpretations based on core sedimentology are well-supported by 698 geochemical data, especially regarding proxies for siliciclastic material relative to mineral 699 precipitation in hydrothermal fluids. For example, Al-concentration in chemical sediments is 700 frequently used as a proxy for siliciclastic components, such as clay minerals and feldspar 701 702 (Calvert, 1976; Tribovillard et al., 2006). The high concentration of Al<sub>2</sub>O<sub>3</sub> and other elements in sandstone beds compared with sinters confirms such interpretations for El Tatio systems 703 704 (Fig. 14). Therefore, subtle differences in  $Al_2O_3$  and associated elements within sinters can 705 determine relative siliciclastic input and corroborate environmental interpretations based on sedimentary structures. Similarly, positive Eu anomalies are often applied as a proxy to 706 estimate the relative influence of hydrothermal fluids on lithologies (Michard, 1989; Lewis et 707 al., 1997; Feng et al., 2014). On broad lateral scales, positive Eu anomalies in El Tatio cores 708 are larger with proximity to the vent source (Fig. 16). Such trends confirm the utility of Eu 709 710 anomalies within cores to determine relative fluctuations in hydrothermal activity.

Al and Eu proxies appear to be broadly connected in El Tatio cores, with samples forming a spectrum between two geochemical end-member proxies. Facies with higher  $Al_2O_3$ concentrations such as sandstones and microbial mat sinters have little to no Eu anomaly, while samples with high positive Eu anomalies (sinter crust, particulate, columnar sinters) have relatively low  $Al_2O_3$  abundance (Figs. 14-16). This dichotomy is most likely due to differences in REE abundance between hydrothermal sinters and detrital sandstones. Sandstones have the highest  $\Sigma REE$  in any core facies, with values orders of magnitude above

associated sinters (Fig. 15). The abundance of REE in sandstone implies that even a relatively
minor siliciclastic component in sinter can overpower the REE signature of depleted sinters.
Separating ∑REE into LREE or HREE reveals no enrichment of either group in clasticnormalized samples (Fig. 15), while normalizations to other standards reveal only minor
enrichments in HREE (Supplementary Information).

723 Therefore, in El Tatio cores, positive Eu anomalies are conservatively interpreted as the relative influence of hydrothermal precipitation compared to siliciclastic influx. The 724 dichotomy between Al<sub>2</sub>O<sub>3</sub> concentrations and Eu anomalies persists across all four cores, and 725 726 matches interpretations based on sedimentary and geochemical evidence. Distal environments have higher abundances of siliciclastic material relative to hydrothermal silica precipitation 727 than vent-proximal zones. Microbial mats in particular are noted for trapping detrital material 728 729 with filamentous textures and adhesive extracellular polymers (Gebelein, 1969; Riding, 2000; Frantz et al., 2015). 730

731 Trace element proxies support previous observations of El Tatio water and sinter chemistry. Two primary types of hydrothermal fluids exist in the El Tatio geyser field: alkali-732 chloride and acid-sulfate (Giggenbach, 1978; Cortecci et al., 2005). Alkali-chloride waters in 733 El Tatio are enriched in As, Cs, and Li, while acid-sulfate waters are only enriched in B 734 compared to meteoric sources (Cortecci et al., 2005). While B was not examined in this 735 study, cored sinters are slightly enriched in Li, Cs, and As compared with most other trace 736 elements (Supplementary Information), indicating the presence of alkali-chloride waters 737 similar to modern Upper Basin environments (Giggenbach, 1978; Cortecci et al., 2005). In 738 addition, elevated concentrations of Sb in Mound 502 sinters are consistent with previous 739 sinter studies at El Tatio and other Chilean hydrothermal zones (Landrum et al., 2009; 740 Sanchez-Yanez et al., 2017). Combined sedimentary and geochemical data therefore provide 741 depositional histories for each core and form a local model for sinter mound growth (Fig. 17). 742

743 6.1 Core A

Core A was drilled on top of Mound 502, along the inner rim of the steep-walled
inner pit. Core A is the first sedimentary core described from vent-proximal sinter deposits,
and it provides several insights into high-temperature facies which corroborate observations
of active modern hydrothermal systems.

The lowermost meter of Core A is massive sinter, showing no primary texture in hand 748 sample or thin section (Fig. 12A, B). Massive facies contain low Al<sub>2</sub>O<sub>3</sub> concentrations and a 749 significant positive Eu anomaly (Fig. 14-16), indicating extensive silica precipitation in 750 751 hydrothermal waters with low siliciclastic input. Such precipitation can occur during primary vent-proximal deposition or secondary subsurface alteration. While specific depositional 752 features are absent, Al<sub>2</sub>O<sub>3</sub> concentrations in massive sinters as low as 0.16 wt.% indicating 753 754 very little to no siliciclastic input. Therefore, we interpret massive Core A sinters as ventproximal hydrothermal deposits secondarily altered by subsurface flow, consistent with 755 previous studies of massive sinter (Walter, 1976b; Walter et al., 1996; Guidry & Chafetz 756 2003a). 757

758 Inclined sinter crust is only observed in Core A, forming a nearly continuous unit  $\sim 2$ meters thick. Depositional angles and bed thicknesses correspond to the surficial topography 759 of Mound 502's inner pit, which descends ~1.5 meters below rim margins (Fig. 1D). Sinters 760 from 121 cm depth contain the lowest measured concentrations of  $Al_2O_3$  and  $\Sigma REE$  in all 761 core samples, as well as the highest percentages of SiO<sub>2</sub> and a strongly pronounced positive 762 Eu anomaly (Fig. 14-16). Geochemistry and sinter topography support the presence of a high-763 764 temperature pool sourced by subsurface hydrothermal fluids (Fig. 17), currently represented by the inner pit of the extinct mound. Microscopic surface-normal silica fabrics support a 765 hydrothermal pool interpretation for sinter crusts (Figs. 3, 4), closely resembling fabrics 766

observed in subaqueous conditions over 73 °C in Yellowstone pools (Lowe & Braunstein,
2003). While microbial textures are occasionally observed in SEM, the prevailing surfacenormal fabrics more closely resemble ambiguously biogenic "shrub columns" (Jones &
Renaut, 2003) or abiogenic "dendritic sinter fabrics" (Lowe & Braunstein, 2003), forming
when silica precipitation outpaces microbial growth.

772 Positive Eu anomalies and SiO<sub>2</sub> percentages decrease further up section while Al<sub>2</sub>O<sub>3</sub> concentrations increase, indicating an increase in siliciclastic input relative to sinter 773 precipitation. An angular unconformity separates inclined sinter crusts from overlying planar 774 775 facies. Planar sinter crusts match the flat-topped topography of Mound 502's inner rims, potentially signifying a shift from silica accretion within steep-sided pools to vertical build-776 up of pool-side margins (Fig. 17). No textural features in planar crusts indicate the presence 777 778 of microbial mats or indicate biological mediation of silica precipitation. The top ten centimeters of Core A are particulate sinter with abundant siliciclastic grains visible in hand 779 sample, signifying a decrease in relative hydrothermal activity compared to clastic 780 sedimentation (Fig. 17). 781

Vent-proximal sinter facies are frequently studied in active hydrothermal systems, 782 exhibiting a diverse range of environments including geysers and pools with varying degrees 783 of boiling, surging, and overflow (Braunstein & Lowe, 2001; Guidry & Chafetz, 2003; Lowe 784 & Braunstein, 2003; Hinman & Walter, 2005). However, the stratigraphic context of high-785 temperature vents is less well-understood. The most detailed vent-proximal stratigraphy 786 comes from Devonian sinters in the Drummond Basin of Australia (Walter et al., 1996, 787 1998). Extensive stratigraphic columns of Drummond exposures revealed several high-788 temperature facies: massive and brecciated sinters interpreted as sub-surface or altered vent 789 zones, as well as thin flat-bedded and "diffusely layered" sinters interpreted as pool-bottom 790 facies (Walter et al., 1996). Vent-proximal sinters are spread out in areas over  $> 100 \text{ m}^2$  in the 791

792 Drummond Basin, and are frequently intercalated with lower-temperature mat facies,

reaching no more than 50 cm thickness. In contrast, no sinter crusts or particulate facies are
described. The differences between Core A and Drummond strata highlight the diversity in
vent-proximal facies, and the need for further stratigraphic investigation.

796 6.2 Core C

Core C was drilled five meters to the east and 1.5 meters below the top of Core A, on 797 the outer margin of Mound 502. The lower meter of Core C is a nearly continuous unit of 798 distinctive columnar sinter (Figs. 6, 7). Despite the major morphological differences between 799 columnar sinters in Core C and sinter crusts in Core A, both have similar chemical signatures. 800  $Al_2O_3$ ,  $\Sigma REE$ , and many other elements increase in roughly equivalent concentrations up-801 section, while SiO<sub>2</sub> shows a steady decrease by ~1 wt.% (Figs. 14, 15). Positive Eu anomalies 802 803 near the base of Core C are the highest in any core, including sinter crust facies in Core A (Fig. 16). The geochemical signatures indicate a high-temperature depositional environment 804 with little to no siliciclastic input (Fig. 17). Such interpretations are strongly supported by 805 current observations of columnar and spicular sinter formation in splash and spray zones 806 close to hydrothermal vents and pools (Walter, 1976b; Jones et al., 1997, 1998; Braunstein & 807 Lowe, 2001; Lowe & Braunstein, 2003; Jones & Renaut 2003; Guidry & Chafetz, 2003a). 808

Despite the general consistency of columnar morphologies in Core C, several hiatuses and depositional changes are present. The transition from mm-scale spicules to cm-scale columns at the base of Core C is separated by a brief section of missing core. The sequence most likely represents a depositional hiatus between splash zones immediately adjacent to a hydrothermal vent, and surge-flow zones forming larger columns further away from hydrothermal activity. At 1.5 m depth, laminated arborescent sinter immediately overlies horizontal columns, with no evidence of intervening erosional activity (Fig. 6E). Overlying

816	columnar sinters maintain similar widths and branching morphologies, but form at increasing
817	depositional angles over the next 70 centimeters (Fig. 6C, D). The brief interval of
818	arborescent sinter formation indicates an outflow-dominated period with active microbial
819	communities colonizing the sinter (Fig. 6E).

Microbial cells and their growth textures are either extremely altered or absent in 820 821 columnar sinter facies (Fig. 7F, G), as opposed to abundant and well-preserved textures in palisade mat fabrics overlying columnar sinters (Figs. 9, 10). Columnar facies in various 822 locations appear to form with or without the presence of microbial communities (Walter, 823 824 1976b; Jones et al., 1997, 1998; Braunstein & Lowe, 2001; Jones & Renaut 2003; Lowe & Braunstein, 2003). As with sinter crusts, the paucity of microbial textures in well-preserved 825 columns does not preclude the earlier presence of microbes, but supports the dominance of 826 827 abiogenic processes of sinter growth.

Palisade mats in the top meter of Core C contain chemical signatures with no Eu 828 829 anomalies and elevated siliciclastic signatures (Fig. 14-16). Palisade mats are typically 830 interpreted as forming at water temperatures at or below 40-45 °C (Walter, 1976a; Campbell et al., 2015; Gong et al., 2019), though the lack of an Eu-anomaly combined with higher 831 concentrations of Al<sub>2</sub>O<sub>3</sub> in Core C palisades is more indicative of increased detritus than 832 relative temperature. Arborescent sinter appears as thin interstitial layers between palisade-833 834 mat sinters, and separate palisade mat facies from underlying columnar facies at 0.5 m depth. No chemical data was collected from arborescent sinters, but visual observations of detrital 835 grains, and comparisons with "siliceous shrubs" from Yellowstone imply outflow channels 836 with potentially higher discharge than palisade mat facies. The top half meter of Core C is 837 interpreted as a shift from a hot splashing zone to more clastic-dominated outflow 838 environments with occasional pulses of hydrothermal activity (Fig. 17). 839

As with vent-proximal facies in Core A, columnar and spicular sinters of Core C are 840 well-described from surface deposits in active and recently inactive hydrothermal settings 841 (Walter, 1976b; Jones et al., 1997, 1998; Braunstein & Lowe, 2001; Lowe & Braunstein, 842 843 2003; Jones & Renaut 2003; Guidry & Chafetz, 2003a), but this study is the first to observe columnar sinters in drill core. Similar laminated columns ( $\leq 20$  cm long) were observed in 844 stratigraphic studies of Devonian sinters from the Drummond Basin in Australia (Walter et 845 846 al., 1996, 1998), but poor exposure prevented correlation to adjacent facies. In contrast, the transition from columns to palisades in Core C from records changing sinter deposition over 847 848 time, from splash and surge-dominated forces to extensive microbial preservation.

849 6.3 Core D

Core D was drilled 6 m southeast of Core A, starting at a roughly equivalent elevation to Core C (Fig. 1E). Unlike the vertical cores of A, C, and E, D was drilled at a 27° angle from vertical alignment along the outer edge of Mound 502, with the drilling axis pointed towards the NW. Core D is the shortest core, and contains the lowest diversity in sedimentary facies and geochemical signatures.

Massive sinters at the base of Core D are not visibly associated with high-temperature facies, such as sinter crusts, particulate, or columnar sinters, although such deposits could underlie the core base (Fig. 2). However, the single chemistry data point within massive sinters is nearly identical to overlying palisade mat facies, with a minor positive Eu anomaly and a moderate siliciclastic signature (Fig. 14-16). Massive facies in Core D are therefore interpreted as relatively distal sinters secondarily altered by subsurface fluid flow (Fig. 17).

Fenestral palisade fabrics at the top of Core D are nearly identical to facies observed
in Cores C and E (Fig. 9B), preserving microbial communities visible in thin-section and
SEM imaging. Chemical analyses reveal slight differences between palisade sinters of

864 different cores, with Core D palisades bearing slightly higher positive Eu anomalies and lower siliciclastic signatures such as Al<sub>2</sub>O<sub>3</sub> than Cores C and E (Fig. 14-16). While it is 865 possible that Core D mats were deposited under warmer conditions than other palisade 866 textures, the relative lack of data keeps temperature estimates conservatively around 40-45 °C 867 as described in previous surficial and stratigraphic studies (Walter, 1976a, 1996, 1998; Jones 868 et al., 1998; Lynne et al., 2008; Campbell et al., 2001, 2015, 2020; Gong et al., 2019). In 869 either case, the top of Core D matches broad trends observed in surficial facies of all cores, 870 featuring strata deposited in cooler water temperatures with higher amounts of siliciclastic 871 872 material before termination of hydrothermal activity (Fig. 17).

873 6.4 Core E

Core E was drilled into a sinter debris apron 30 meters southeast of Core A, and 15 meters southeast of Mound 502 itself. The core begins approximately four meters below the top of Core A, and is the deepest of the four cores, reaching 560 cm depth. Core E records the highest diversity of sedimentary facies and geochemical signatures, including highertemperature columnar sinters, lower-temperature mat facies, and clastic sandstone beds.

879 Core E is the only core examined to record non-hydrothermal sediments deposited prior to extensive sinter formation, appearing as sandstones below 4 m depth (Fig. 13). Most 880 sandstones present are massive and are likely the result of debris flows from surrounding 881 882 hillsides into valleys to the south (Fig. 1B). However, minor layers of laminated sinter occur throughout sandstone beds, indicating an influence from hydrothermal activity. The high 883 abundance of SiO<sub>2</sub> throughout the sandstone succession in combination with the lack of 884 885 detrital quartz (Fig. 13F) indicates that this hydrothermal influence persisted throughout the sandstone deposits. Minor variations in the SiO<sub>2</sub> concentration are most likely due to 886 differences in the amount of silica cement. 887

Layers between 2.9 and 3.7 m depth in Core E form the most diverse assemblage of 888 facies in any core measured in this study. Six different facies occur within 80 centimeters, 889 with no single unit exceeding ten centimeters in thickness (Fig. 2). Such disparate lithologies 890 891 imply relatively rapid environmental shifts on Mound 502's outer distal apron, including hydrothermal vent sources separate from the mound's central pool. Two sedimentary cycles 892 are noted, each shifting from subaqueous or pool-side facies to outflow environments. The 893 894 lower cycle (3.4-3.7 m depth) features a thin layer of small columnar sinter (Figs. 6F, 7B) overlain by tufted mat sinters and laminated sandstones. The upper cycle (2.9-3.4 m depth) 895 896 contains massive and arborescent sinters overlain by mixed tufted mats and sandstones. In other cores, columnar and massive sinters typically contain elevated positive Eu anomalies 897 indicative of extensive hydrothermal silica precipitation. However, between 2.9 and 3.7 cm 898 899 depth in Core E the relative concentration of Eu is similar to that of the underlying sandstone, 900 but increases only slightly up section (Fig. 15). By contrast, concentrations of Al<sub>2</sub>O<sub>3</sub> and other elements decrease by orders of magnitude (Fig. 14). Together, petrographic-, Eu-, and 901 902 Al<sub>2</sub>O<sub>3</sub>-data depict a location experiencing locally rapid depositional shifts, with an underlying trend of increasing hydrothermal sinter precipitation and decreasing siliciclastic input (Fig. 903 904 16).

905 The only section of Core E recovered between 2-3 meters is a 30-centimeter section 906 (2.2-2.5 cm depth) of palisade mat fabrics underlain by thin beds of massive sinter. The 907 abrupt transition from massive to palisade sinters is similar to Core D, with more detailed geochemical data sampling revealing distinct shifts in temperature and siliciclastic influence. 908 909 Positive Eu anomalies in massive sinters are equivalent to core facies at 3 m depth (Fig. 16). 910 In contrast, immediately overlying palisade mat sinters contain no Eu anomaly, but quickly return over the subsequent twenty centimeters. Al<sub>2</sub>O<sub>3</sub> and other siliciclastic constituents are 911 anti-correlated to relative Eu concentration, increasing sharply in lower palisade sinters and 912

decreasing up-section (Fig. 14, 15). Geochemical trends indicate hydrothermal environments
with a low siliciclastic component, succeeded by palisade mats potentially growing at ~40 °C
(Walter, 1976a; Campbell et al., 2015; Gong et al., 2019), though potentially persisting in
higher temperatures. No change in mat architecture is associated with changes in relative Eu
concentrations.

918 Between 1.0 and 2.0 m depth, Core E is entirely composed of particulate, massive, and arborescent sinters. Particulate and massive facies are frequently interpreted as 919 subaqueous deposits in high-temperature pools (Walter, 1976b; Braunstein & Lowe, 2001; 920 921 Lowe & Braunstein, 2003; Hinman & Walter, 2005), and Core E geochemistry supports similar conclusions. The highest positive Eu anomalies in Core E occur between 1.5 to 2.0 m 922 depth, similar to values observed in sinter crusts and columnar sinters in Cores A and C (Fig. 923 924 16). As in more vent-proximal cores, positive Eu anomalies gradually decrease up-section, eventually approaching similar relative Eu ratios observed in lower Core E sinters. 925 Concentrations of  $Al_2O_3$ ,  $\Sigma REE$  and other siliciclastic proxies reach minimum values in Core 926 E between 1.0 and 2.0 meters, and remain consistently low throughout dendritic and 927 associated sinters (Fig. 14, 16). Unlike inclined sinter crusts in Core A and columnar sinter in 928 929 Core C, high-temperature facies in Core E appear to have precipitated on flat-bottomed pools 930 on distal aprons, separate from the elevated sinter mound (Fig. 17). Such pools are still active 931 several meters away from Core E, supporting their potential presence during Core E 932 deposition. The presence of arborescent sinter between particulate facies most likely represents a shift from pool deposits to outflow channels with microbial communities (Guidry 933 934 & Chafetz, 2003c).

935 The top meter of Core E contains the thickest continuous package of microbial facies
936 in any core examined. Palisades comprise the majority of preserved mats, with nearly
937 identical fenestral morphologies to the upper beds of Cores C and D (Fig. 9C). The lower

938 boundary between particulate sinters and overlying palisades is associated with an immediate loss of the positive Eu anomaly, followed by fluctuations around the average for clastic 939 deposition around Mound 502 (Fig. 16). Unlike Cores C and D, palisades in Core E contain 940 941 distinct interstitial layers of tufted microbial sinters. Tufted sinters are typically assigned to higher-temperature communities (45-60 °C) than palisade mats (< 45 °C), based on 942 morphologies of modern hot spring communities (Walter, 1976a; Campbell et al., 2001; 943 Lynne & Campbell, 2003; Fernandez-Turiel et al., 2005). No clear difference in REE 944 signatures exists between tufted and palisade mat sinters in El Tatio cores, most likely due to 945 946 the entrapment of siliciclastic material on sticky microbial surfaces. The diversity of facies observed in Core E resembles previous stratigraphic studies 947 from distal sinter deposits around the world (Trewin, 1993; Walter et al., 1996, 1998; Jones et 948 949 al., 1998; Campbell et al., 2001, 2020; Guidry & Chafetz, 2003b; Lynne et al., 2008). Palisade and tufted microbial mats are ubiquitous facies in cooler environments (Walter et al., 950 1996, 1998; Jones et al., 1998; Campbell et al., 2001, 2020), while arborescent textures are 951 described less frequently (Walter et al., 1996, 1998; Campbell et al., 2001). The frequent 952 953 interbedding between microbial facies and more massive deposits in Core E (1-2, 3-3.5 m 954 depth) is also a common feature of distal sinter stratigraphy (Walter et al., 1996; Campbell et 955 al., 2001; Lynne et al., 2008). In contrast, Core E lacks extensive preservation of metazoan 956 communities such as plants or diatoms, which are often observed in distal sinters (Trewin, 957 1993; Walter et al., 1996, 1998; Campbell et al., 2001, 2020; Guidry & Chafetz, 2003b). While the climate at El Tatio is more extreme than many other sinter localities, scattered 958 areas with plant life are present. However, areas selected for coring in Mound 502 lack 959 960 extensive vegetation, and the absence of plant remains suggests similar environmental 961 conditions during sinter deposition. Core stratigraphy from Mound 502 therefore presents a closer environmental analogue to ancient sinters on Earth or Mars than other distal deposits. 962

963 6.5 Sinter ages

Five <sup>14</sup>C ages were measured from Cores C and E (Table 1, Fig. 2), ranging between 964 6,907 and 13,337 y. cal. BP. Core E contains a consistent progression from 13,337 to 10,028 965 966 y. cal. BP, underlain by an age of 6,907 y. cal. BP at 347 cm depth (Table 1, Fig. 2). Similar irregularities have been noted in previous sinter cores (Lynne et al., 2008; Lowenstern et al., 967 2016), and have several potential explanations. The presence of younger dates below a series 968 of consistently older data points can be explained by transport of younger material into 969 deeper sediments by subsurface infiltration through fractures and pores (Lynne et al., 2008; 970 971 Lynne, 2012). In support of this hypothesis, various core depths contain fracture zones associated with secondary mineral phases, including massive sandstones deeper in Core E 972 973 (Fig. 13B). While no specific fractures are apparent at 347 cm, the sinter bed is fairly porous, 974 and is ten centimeters below a massive, altered sinter (Fig. 2). We propose that younger material permeated through fissures and pores in older sinters via secondary fluid flow, 975 producing an anomalous radiocarbon date thousands of years younger than three succeeding 976 deposits. 977

The four remaining <sup>14</sup>C ages (one in Core C, three in Core E) record more than three 978 979 thousand years of sinter deposition in Mound 502 (Figs. 2, 17, Table 1). A transition from clastic sandstones to siliceous sinter occurred more than 13,337 y. cal. BP in distal debris 980 aprons (Core E, 310.5 cm depth). Between 13,337 and 10,931 y. cal. BP (Core E, 33 cm 981 depth), Core E experienced shifts between various sinter facies, including microbial mats and 982 detrital-poor vent proximal zones (Figs. 2, 17). After 10,931 y. cal. BP, palisade mat sinters 983 dominate deposition in Cores C and E (and presumably D) until at least 10,028 y. cal. BP 984 (Core E, 6 cm depth). The ages from Mound 502 supports previous timelines of hydrothermal 985 activity in El Tatio, which initiated at least 27 ka, and persisted until the present day without 986 major hiatuses (Slagter et al., 2019, Munoz-Saez et al., 2020). 987

The <sup>14</sup>C ages from Mound 502 cores are thousands of years older than surficial 988 deposits from the same location (Slagter et al., 2019). Four data points collected on a transect 989 between Cores A and C (Fig. 17B, C) record a period of deposition between 2,241 and 197 y. 990 991 cal. BP (Slagter et al., 2019). The age difference between the two datasets is interpreted as a depositional hiatus separating two phases of sinter deposition. The first phase formed the 992 majority of deposits in Cores C and E between ~14,000 and 10,000 y. cal. BP. After ~7,000 993 years, surficial hydrothermal activity resumed, forming surficial sinters between ~2,500 and 994 200 y. cal. BP (Slagter et al., 2019). Depositional hiatuses of similar length have previously 995 996 been reported from hydrothermal zones in Yellowstone and New Zealand (Foley, 2006; Drake et al., 2014; Hurwitz & Lowenstern, 2014), and are most likely due to changes in 997 998 subsurface hydrothermal systems.

999 The age of Core A remains uncertain relative to other cores. One unconformity exists 1000 in Core A at 35 cm depth, indicating a prolonged period of hydrothermal activity, followed by a hiatus and minor subsequent sinter deposition. The current set of <sup>14</sup>C ages from surficial 1001 and core samples presents two potential timelines (Fig. 17A). In one scenario, most of Core A 1002 was deposited during the initial period of hydrothermal activity, forming an elevated sinter 1003 1004 mound more than 10,000 y. cal. BP followed by a resurgence of activity after 2,000 y. cal. BP. Alternatively, the inner mound and Core A could have formed after 2,000 y. cal. BP, 1005 1006 with Cores C, D, and E representing relatively low-relief mound and outflow deposits (Fig. 1007 17A). In either case, depositional interpretations for all cores remain the same, with timing 1008 more well-constrained between more distal depositional facies.

## 1009 7. Conclusions

1010 The four cores examined provide the first complete stratigraphic transect of a recent1011 sinter mound from vent to outflow. Lateral trends of hydrothermal deposits record detailed

transitions in chemistry, sedimentology and biology over broad areas, while vertical profiles
provide similar records for local environments over geologic time. Initial depositional
interpretations from sedimentology and petrography match geochemical data from the same
facies. Combining all approaches within a single location provides a detailed environmental
history that includes the initiation, accretion, and eventual cessation of sinter deposition
within a single mound.

Hydrothermal activity started in the western margins of El Tatio's upper basin prior to 1018 13,337 y. cal. BP. Sinter deposits formed on underlying glacial sandstones, intermittently at 1019 1020 first, then forming a wide diversity of biogenic and abiogenic textures. Deposits closest to the 1021 original hydrothermal vent are dominated by subaqueous precipitation on the steep margins 1022 of a high-temperature pool. Several meters away, extensive columnar sinter formed on outer 1023 pool margins experiencing alternate periods of pool overflow and subaerial evaporation. Less extensive pools with subaqueous and columnar facies also formed in distal aprons. 1024 Microfossils are rare to absent in high-temperature facies (>70 °C), with silica precipitation 1025 most likely driven by physico-chemical factors such as silica concentrations, cooling, and 1026 1027 evaporation. On the other hand, extensive microbial deposits are present in cooler, distal 1028 environments with more siliciclastic input. Palisade, tufted, and arborescent sinters indicate 1029 the presence of multiple distinct microbial communities, potentially fluctuating with water 1030 temperature and flow dynamics. Mat deposits and detrital-rich facies eventually replace high-1031 temperature sinters in the upper sections of all cores, signifying cooling and eventual termination of local deposition no more than 10,232 y. cal. BP. In other areas of the same 1032 mound, hydrothermal activity persisted for several thousand years (Slagter et al., 2019). 1033

As seen within the El Tatio cores, sinter stratigraphy is complex, and varies both laterally and vertically on relatively fine scales. The deposits examined represent only one type of hydrothermal outflow, a topographically-elevated boiling, surging pool. Similar

1037 coring studies performed on extinct structures such as fountain geysers and low-relief, non-1038 surging pools are likely to reveal different facies patterns. Further sinter stratigraphy, especially in the harsh environmental conditions of El Tatio, also provides a comparison for 1039 1040 Archean and Martian sinters and their biosignature potential. Microfossils are present in all four cores, but are far more prevalent and well-preserved in outflow and distal deposits. 1041 1042 Many vent-proximal sinters are ambiguously biogenic, but the sparse remains of microbial 1043 filaments indicate a biological presence, and such deposits should not be overlooked in future 1044 stratigraphic studies.

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1053 The authors have no competing interests to declare.

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1353 Figure Captions:

1354 Figure 1

1355 Sampling location. A: Map showing general location of El Tatio in Chile. B: Aerial view of

1356 El Tatio's North Basin. Yellow dots represent significant hydrothermal outflow sources.

1357 Dashed outline represents the area shown in C and D. C: Aerial view of extinct sinter mound

and drill locations. D: Oblique view of extinct sinter mound, highlighting topography of the

sinter mound, interior pit, and drill locations. E: Profile of extinct sinter mound along transectin Fig. 1C.

1361 Figure 2

Generalized sedimentary facies in El Tatio cores. Scale represents core depth below surface,
see Fig. 1 for relative core positions within sinter mound. Asterisks by age dates represent
secondary influence on <sup>14</sup>C data. Table 1 lists specific depths for <sup>14</sup>C samples, Supplementary
Table 1 lists specific sample depths for all other chemistry samples.

1366 Figure 3

1367 Inclined sinter crust. A, B and C are hand specimen photographs. A, B: Nearly vertical sinter crusts in Core A at 1.4 m (A) and 1.6 m depth (B). Arrows in B highlight darker surface-1368 normal lineations which transect through sinter crusts. C: Core A, 0.4 m depth. Lower-angle 1369 laminations at the top of inclined beds. Note paler, more porous textures compared with 1370 deeper fabrics. D: Photomicrograph showing the boundary between well-preserved sinter 1371 crusts in the lower right and more altered textures in the upper left, marked by dashed line. 1372 Arrows highlight layers with surface normal fabrics flaring upward and to the left. Core A, 2 1373 m depth. E-H: SEM images of Core A, 0.4 m depth. E: Cross-section of surface-normal 1374 fabric, composed of silica tubes growing from right to left, highlighted by dashed lines. F: 1375 Detail of surface-normal silica tube, highlighting scalloped porous core surrounded by a 1376

thinly-laminated cortex of silica. G, H: Filamentous microfossils within inclined sinter. Note
the difference in size and preservation between microfossils and surface-normal fabrics in E
and F.

1380 Figure 4

- 1381 Planar sinter crust. A: Hand specimen photograph, Core A, 0.2 m depth. Note the presence of
- 1382 fine detrital grains in darker layers. B, C: SEM images of Core A, 0.25 m depth. B: Light
- 1383 layer, comprised of vertical chains of silica spheres separated by interstitial pore space. C:
- 1384 Dark layer, highlighting porous zones surrounded by massive silica rinds.

1385 Figure 5

- 1386 Particulate sinter. A, B: Hand specimen photographs of particulate facies in Core A, 0.1 m
- 1387 depth (A), and Core E, 1.1 m depth, highlighting diffuse, wavy layering. Orange areas are
- 1388 secondary mineral phases. C: Photomicrograph from Core E, 1.2 m depth. Particulate
- 1389 microfacies. Lighter, rounded textures are opal-A spherules, accompanied by larger detrital
- 1390 grains. D, E: SEM images from Core E, 1.2 m depth, showing a cross-section of particulate
- sinter layers. Note the absence of microbial filaments or well-defined laminae.

1392 Figure 6

- 1393 Macroscale features of columnar and spicular sinter. A: Surficial exposure of columnar sinter
- 1394 growing upwards. Exposure location is on the outer margin of the cored sinter, 5 m northwest
- 1395 of Core A (Fig. 1C). B: Recent columnar sinter fragments adjacent to active boiling pools in
- 1396 El Tatio, ~40 m southeast of Core E. C: Horizontal columns with various stages of branching,
- 1397 Core C, 1.6 m depth. D: Angled branching columns, Core C, 1.1 m depth. E: Cross-section
- through horizontal columns, Core C, 1.5 m depth. Planar sinter laminae immediately
- 1399 overlying columns mark the boundary between lower horizontal columns and upper angled
- 1400 columns. F: Vertical column showing convex laminae, overlain by tufted microbial sinter.

1401 Core E, 3.65 m depth. G: Curved spicules. Note the size difference between spicules and1402 columns. Core C, 1.7 m depth.

1403 Figure 7

1404 Petrography of columnar sinter. A: Rounded termination of laminated sinter column, Core C, 1 m depth. B: Two adjacent columns showing coarser, more porous laminae than Core C, 1405 from Core E, 3.65 m depth. C-E: Core C, 1 m depth. C: Broken termination of sinter column. 1406 Note the faint convex laminations within the column, as well as thin yellow minerals present 1407 around all margins, including broken surfaces. D: Boundary between finely-laminated 1408 1409 column and surrounding porous, detrital-rich interstitial sinter. A detrital-rich geopetal feature 1410 is highlighted by the white arrow. E: Fine geopetal silts located between columns, extensively 1411 cross-cut and separated by secondary silica veins. F, G: Core C, 1.8 m depth. F: Cross-section 1412 from the top down of an individual column, showing multiple generations of silica precipitation. G: Hollow silicified tube extending from a columnar surface, potentially 1413 1414 representing a rare preserved microfossil.

1415 Figure 8

1416 Arborescent sinter. A-C: hand specimen photographs. A: Arborescent facies at the top of Core C, with textures branching upward into open air. B: Pronounced cm-scale arborescent 1417 1418 facies, over- and underlain by layers of dense, less porous sinter. Core E, 3.15 m depth. Note 1419 the branching textures in individual "shrubs". C: Upper branches of arborescent structures extend into detrital-rich darker overlying sinter, noted by a white arrow. Core E, 1.5 m depth. 1420 1421 D: Photomicrograph of arborescent facies, showing branching chains of silica spherules separated by pore space. Core E, 1.4 m depth. E-G: SEM images from Core E, 3.15 m depth. 1422 1423 E: Individual branching structure, highlighting the presence of microbial filaments. F:

Magnified image of highlighted area in E, showing microbial filaments with associatedangular detrital material. G: Microbial filaments without detrital material present.

1426 Figure 9

1427 Palisade mat sinter. A-C : Hand specimen photographs. A: Core C, 0.1 m depth. B: Core D, 0.2 m depth. Note that fabrics at the top of the sample maintain fenestral qualities, but lose 1428 distinct lamination. C: Core E, 2.2 m depth. D-G: Photomicrographs of palisade sinters. D: 1429 Palisade fabric alternating between darker, less porous laminae and lighter, more porous 1430 zones with extensive fenestrae. Core D, 0.1 m depth. E: Palisade sinter with alternating low-1431 and high-porosity laminae, with fewer fenestrae compared to Cores C and D. Low-porosity 1432 laminae are lighter in thin-sections, but darker in hand samples (see Fig. 9C). Core E, 2.35 m 1433 1434 depth. F: Fenestrae within a light porous lamina, Core D, 0.2 m depth. G: Detrital sinter fragment (most likely columnar or spicular) embedded in palisade fabric. Core D, 0.1 m 1435 1436 depth.

1437 Figure 10

1438 SEM imagery of different microbial sinters. A-C: Palisade fabrics from (A) Core C, 0.1 m

1439 depth, (B) Core D, 0.2 m depth, (C) Core E, 2.2 m depth. Fenestrae are less abundant in C

1440 than A or B. D: Filamentous microfossils separating palisade fenestrae. Core D, 0.2 m depth.

1441 E: Detailed view of filaments in D. F: Rounded fenestrae in tufted mat sinters, Core E, 3.6 m

- 1442 depth. G: Detailed view of microbial sinter separating fenestrae in F, showing well-preserved
- 1443 microbial filaments on top of more silicified zones.

1444 Figure 11

1445 Tufted mat sinter. A-C: Photomicrographs of microbial tufts from (A) Core E, 0.3 m depth

1446 (B) Core E, 0.5 m depth. C: Core E, 3.6 m depth. Rounded fenestrae separated by filamentous

1447 hourglass structures (see Fig. 10 for SEM of similar features). D-F: SEM images from Core

1448 E, 3.6 m depth. Single microbial tuft, highlighting (D) the entire tuft, (E) the basal meshwork

1449 of filamentous microfossils, and (F) individual filaments near the tufted tip.

1450 Figure 12

- 1451 Massive sinter. A: Hand specimen photograph from Core A, 2.7 m depth. B-D:
- 1452 Photomicrographs of massive sinter in Core A, 2.4 m depth (B); Core D, 0.8 m depth (C);
- 1453 Core E, 1.5 m depth, cross-polarized light (D). Sample comes from an intercalation of
- 1454 massive sinter within dendritic sinter, highlighting extensive white rims of secondary silica
- 1455 precipitation. E: SEM image from Core E, 1.9 m depth.

1456 Figure 13

1457 Sandstone facies. A-B: Hand specimen photographs. A: Laminated sandstone and sinter,

1458 Core E, 3.5 m depth. B: Massive sandstone containing matrix-supported pebbles and

- 1459 secondary silica cements propagating along fracture zones, Core E, 5.4 m depth. C-E:
- 1460 Photomicrographs under plane polarized light. C: Laminated sandstone and sinter, Core E,
- 1461 3.5 m depth. D: Massive sandstone, Core E, 5.2 m depth. E: Laminated sandstone and sinter
- 1462 containing a fragment of columnar or spicular sinter, Core E, 3.5 m depth. F: Massive
- sandstone, photomicrograph in cross-polarized light, Core E, 5.2 m depth.

1464 Figure 14

Major element patterns in all cores. Due to the disparity in elemental concentrations between sandstones and overlying sinters in Core E, elements in the upper four meters are shown separately on different axes. Supplementary Table 1 contains detailed concentration and depth measurements for all major and trace elements.

1469 Figure 15

Rare earth element patterns in all cores, normalized to clastic sediments within Mound 502.
Depth profiles show total REE concentrations in each core, as well as profiles for Ce and Yb
representing light and heavy REE signatures. Note the nearly identical values for Ce and Yb
in all cores. Darker shades in REE patterns represent deeper samples in core, as shown in
profiles. Core E is further split into four sub-sections dominated by different facies. Specific
values are presented in the Supplementary Information.

1476 Figure 16

1477 Europium anomalies in all cores. Values are unit-less, and represent the relative enrichment

1478 or depletion of Eu compared to Sm and Gd (See Section 3.4 for calculations), in contrast with

1479 Eu concentrations (mg/kg) shown Figure 15 and the Supplementary Information. Values

above 1 are enriched in Eu, and are interpreted here as a proxy for the relative influence ofhigh-temperature hydrothermal fluid compared to clastic sedimentation.

1482 Figure 17

1483 Depositional history of the cored El Tatio sinter mound. <sup>14</sup>C dates interpreted as detrital

signatures or secondary alteration are not included. A: Schematic reconstruction of mound

1485 formation through time. B: Field photograph representing the sampling sites (yellow stars)

1486 from Slagter et al., 2019 compared to the position of drill cores presented in this study

1487 labelled with white arrows, modified from Slagter et al., 2019. C: Drone image showing the

1488 mound from above, representing the position of the features described in B.