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1 **Optimal configuration of a far-infrared radiometer to**
2 **the study the Arctic winter atmosphere**

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6 **Key Points:**

- 7 • FIR channels add information for UTLS water vapor compared to standard MIR
8 channels
9 • IC is used to optimize the channels frequencies and widths of a FIR radiometer
10 • A high DFS is reached with only a few channels of an optimized FIR radiometer

Abstract

Several FIR satellite missions are planned for the next decade, with a special interest for the Arctic region. A theoretical study is performed to help about the design of a FIR radiometer, whose configuration in terms of channels number and frequencies is optimized based on information content analysis. The problem is cast in a context of vertical column experiments (1D) to determine the optimal configuration of a FIR radiometer to study the Arctic polar night. If only observations of the FIR radiometer were assimilated, the results show that for humidity, 90 % of the total information content is obtained with 4 bands whereas for temperature 10 bands are needed. When the FIR measurements are assimilated on top of those from the Advanced Infrared Sounder (AIRS), the former bring in additional information between the surface and 850 hPa and from 550 hPa to 250 hPa for humidity. Moreover, between 400 hPa and 200 hPa, the FIR radiometer is better than AIRS at reducing the analysis error variance for humidity. This indicates the potential of FIR observations for improving water vapor analysis in the Arctic.

1 Introduction

Since the beginning of meteorological satellites, temperature profiling has been performed with sounders in the infrared (IR) (Wark & Hilleary, 1969). The state-of-the-art instruments that probe the mid-infrared (MIR) are the Infrared Atmospheric Sounding Interferometer (IASI) (Blumstein et al., 2004), the Atmospheric Infrared Sounder (AIRS) (Aumann et al., 2003) and the Cross-track Infrared Sounder (CrIS) (Bloom, 2001). Those instruments use the $15 \mu\text{m}$ CO_2 absorption band to probe atmospheric temperature and the water vapor vibrational band at $6.3 \mu\text{m}$ to retrieve humidity profiles (Rizzi et al., 2002). This humidity profiling capability results from the strong spectral variations of the water vapor absorption in that band.

Water vapor also exhibits an extended rotational absorption band as well as a continuum in the far-infrared (FIR; $15 \mu\text{m} < \lambda < 100 \mu\text{m}$). This absorption band is broader than the vibrational band, hence there is more energy in this region. Previous studies have pointed out the potential of FIR for atmospheric profiling, particularly in cold regions. First, a direct consequence of the temperature dependence of the Planck function in the radiative transfer equation, emission is shifted to the FIR as the temperature of the scene decreases, offering a greater capability for cold scenes (M. Mlynczak et al., 2007), in contrary to the traditional $6.3 \mu\text{m}$ band quickly loses energy as the scene gets colder (Susskind et al., 2003). Practically, more than half of the radiation is lost to space in the FIR in the polar regions (M. Mlynczak et al., 2007). Second, the water vapor rotational absorption band in the FIR has many absorption lines with larger optical depth than the MIR, which leads to an increased sensitivity to small water vapor variations (Harries et al., 2008). The increased sensitivity is especially important in the upper troposphere, where the water vapor concentration is scarce (Clough et al., 1992). Thus, the FIR region can be valuable for profiling the atmosphere and particularly in the stratosphere and the upper troposphere (Shahabadi & Huang, 2014).

Despite these acknowledged advantages of the FIR over the MIR for water vapor profiling, no direct spectrally resolved measurements of the atmospheric radiation have been made recently from space. The last measurements in the FIR, up to $25 \mu\text{m}$, were made 40 years ago on two Russian *Meteor* spacecrafts and 45 years ago by the IRIS (Infrared Interferometer Spectrometer and Radiometer) instruments on the NASA Nimbus III and IV (M. G. Mlynczak et al., 2002), data that has been used to identify changes in spectral outgoing longwave radiation (?). However at the time, the spectral and spatial resolutions of the observations, along with the large noise, prevented from getting much geophysical information out of the data. Since then, low noise liquid helium cooled bolometers operating in the far-IR have been developed, and used for instance in the Far-Infrared Spectrometer of the Troposphere (FIRST) instrument (M. G. Mlynczak et al.,

2006). Such systems are however too delicate, massive and expensive to be but on a satellite. This, in combination with the intrinsic higher sensitivity of MIR sensors compared to FIR sensors, explains why no FIR satellite has been flying for decades now.

Only with the recent advent of uncooled systems operating at room temperature that have space missions in the FIR seen a renewal of interest. The series of satellite missions Climate Absolute Radiance and Refractivity Observatory (CLARREO) (Wielicki et al., 2013) is intended to measure spectrally-resolved Earth emission spectrum between 5 - 50 μm with a spectral resolution of 0.5 cm^{-1} to determine small changes in the spectral outgoing radiation (infrared and reflected solar). The mission aims at detecting decadal changes in climate forcings, responses and feedbacks and to serve for reference intercalibration in space. It thus focuses on global or regional averages and their variations on annual timescales. As a consequence the noise-equivalent temperature difference (NETD), has the requirement to be smaller than 10 K in the FIR since averaged over a year it will be reduced to 0.01 K globally. The requirements on the absolute accuracy are on the contrary much more stringent. ESA candidate mission, called FORUM (Far Infrared Outgoing Radiation Understanding and Monitoring) (Palchetti et al., 2016) focuses on studying the forcings and the feedbacks of atmospheric water vapor and of ice clouds on the climate. The recently funded NASA PREFIRE (Polar Radiant Energy in the Far-InfraRed Experiment) CubeSat, to be launched in 2022, intends to measure in the 0-45 μm range to measure spatial and temporal variations in spectral fluxes on hourly to seasonal timescales (L'Ecuyer, 2019). Merrelli and Turner (2012) used the technical characteristics of CLARREO to compare two interferometers for remote sensing of temperature and humidity, with a spectral resolution of 0.5 cm^{-1} , one measuring in the MIR and another measuring in the FIR even though CLARREO was not designed to measure temperature and water vapor. They showed that there is more information content in the FIR compared to the MIR when the noise is equal in both spectral regions. However, if the uncertainty of the actual CLARREO is used for the FIR region, the advantage of the FIR is lost (Shahabadi et al., 2015).

Only a limited selection of channels among those in the water vapor absorption band of interferometers at high resolution is used in data assimilation (Fourrié & Thépaut, 2003). Supposedly, adjacent bands could thus be merged into larger bands to refine the remote sensing capability. The Arctic was selected since around 60 % of the outgoing longwave radiation is in the FIR region. Also, there is a need for precise water vapor measurements in the troposphere (Müller et al., 2016) and especially in the Arctic (Boullot et al., 2016) and the FIR upwelling spectrum contains a large amount of potential profiling information. The objective of the present study is to design an optimal FIR radiometer to study the Arctic polar night by examining different configurations, noise levels and the trade-off between spectral resolution and noise level. Thus, this study considers a radiometer in the Arctic region using different filters or a gratings to allow different bandwidths within a spectral region. As in Observing System Simulation Experiments (OSSEs), synthetic measurements are created for different configurations of the FIR radiometer. The radiometric noise of the Far InfraRed Radiometer (FIRR) is used as a baseline to constrain the detector's performance (Libois et al., 2016). The optimal configuration is selected with information content as a metric to lead to the best temperature and humidity analyses. The impact of FIR measurements is also evaluated in terms of their added value when assimilated on top of currently assimilated AIRS data. The experiments are done under the assumption that AIRS and the FIR radiometer are collocated and assimilated in a simple 1D assimilation system.

The paper is organized as follows. Section 2 presents the information content framework, the characteristics of the instrument, and the context of the experiments. Section 3 presents the results of the evaluation of different instrument configurations. Section 4 compares the impact of measurements of the FIR radiometer with that of AIRS measurements. A discussion and conclusions are presented in section 5.

2 Methods

This section first presents the characteristics of the FIR radiometer and the atmospheric conditions used in this study. Finally, we present the method used to evaluate the information content of measurements which is based on the reduction of analysis error obtained in the context of data assimilation using a numerical weather forecast as an *a priori* background state.

2.1 Instrument characteristics

A synthetic spaceborne FIR radiometer is considered in this study. The goal is not to investigate a particular instrument, but to explore the potential of such a novel instrument, in the framework of the preparation of the TICFIRE mission. Nevertheless, the characteristics of this radiometer are based on the FIRR instrument. The characteristics of the optics are fixed (field of view, spatial resolution, F-number, etc), only those of the detector are changed. The detector performance explore a realistic range, although the feasibility study for such performance is left to the industry. The two principal characteristics considered are its number of bands and its noise-equivalent radiance (NER). The bands are adjacent and fully cover the range of 15 - 100 μm . The transmittance is one in the bands and zero outside. The bandwidths are set in three different ways hereafter referred as 'equi-energetic', 'constant wavelength' and 'constant wavenumber'. Equi-energetic means that each band receives the same amount of energy at the top of the atmosphere (TOA), this is calculated for each atmospheric profile used. This implies that the spectral widths of the bands are not constant. Constant wavelength and constant wavenumber bands means that each bands has the same spectral width in microns or cm^{-1} respectively. The NER is varied through the experiments, but two specific NER will be highlighted, called baseline NER and target NER. The baseline NER is equal to $0.01 \text{ Wm}^{-2}\text{sr}^{-1}$, according to the findings of Libois et al. (2016). Those detector characteristics are consistent with a microbolometer sensor coated with gold black for a integration time of 1 s (Proulx et al., 2009) The other specific value of NER used, target NER, is equal to $0.002 \text{ Wm}^{-2}\text{sr}^{-1}$, which is the expected NER in a few years from now, expected from efforts by the industry, mainly on the electronics and on the analog to digital conversion. Also, band splitting is achievable with a grating or filters. It was chosen to use NER instead of NETD for the radiometric resolution in order to work at the sensor level. This allows to evaluate the gain of changing the radiometric resolution and the spectral width of the bands independently. NER remains constant independently of the instrument spectral configuration, while NETD would change. Since the NER is constant, this results in less energy per band when the bandwidth is reduced. It needs to be noted the correlation between radiometric and spectral resolutions, when the number of bands increases, the signal-to-noise ratio decreases as the energy per band decreases.

Figure 1 shows the NETD for a blackbody at 250 K for a constant NER of $0.01 \text{ Wm}^{-2}\text{sr}^{-1}$ for an instrument with 10, 15, 20, 25 and 40 equi-energetic bands. The NETD is not constant for the bands of a configuration. This allows to compare this experiment with other studies using NETD. It shows that the NETD, for a configuration with 10 bands, is comparable to the NETD of AIRS, below 0.5 K (Garand et al., 2007), and of MODIS, less than 0.35 K (Xiong et al., 2008).

2.2 Atmospheric profiles

The radiosonde profiles are from the Integrated Global Radiosonde Archive (IGRA) database (<http://www.ncdc.noaa.gov/oa/climate/igra/>) (Durre et al., 2006). Figure 2 shows the locations of the eight stations where the different vertical profiles were taken. Those stations are the same as in Serreze et al. (2012) and were selected to represent the various atmospheric conditions in the Arctic. It needs to be noted that the Arctic region was chosen, but those results would be similar for the Antarctic region. For each station, 6 profiles were selected randomly from the months of January or February of 2015

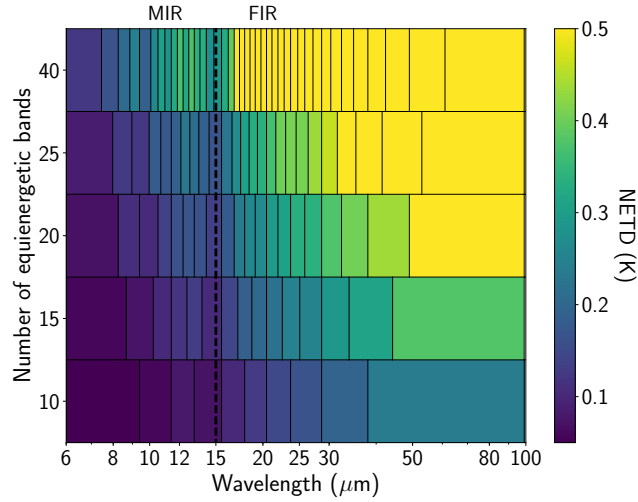


Figure 1. NETD for different configurations of equienergetic bands for a blackbody at 250 K with a constant NER of $0.01 \text{ Wm}^{-2}\text{sr}^{-1}$. The vertical lines represent the widths of equi-energetic bands.

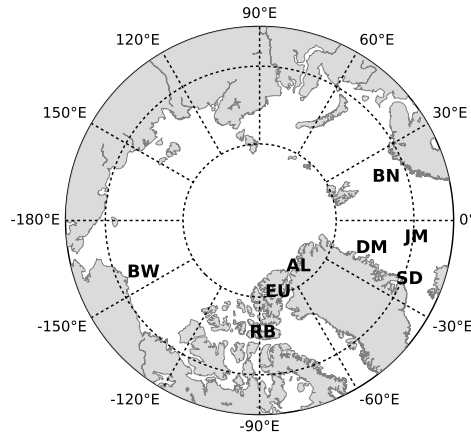


Figure 2. Locations of the eight Arctic stations. The letter codes are: JM- Jan Mayen; BN- Bjornoya; SD- Scoresbysunde; DM- Danmarkshavn; BW- Barrow, AL- Alert; EU- Eureka; RB- Resolute Bay.

167 or 2016 in order to sample Arctic winter conditions. The profiles were truncated at 20
 168 km altitude. Table 1 shows the vertical resolution of the atmospheric profiles selected.
 169 Figure 3 shows the averaged 48 temperature and humidity profiles selected with the red
 170 and blue lines respectively and the shaded area of the same color shows the correspond-
 171 ing standard deviation. The natural variability in the profiles seen through the standard
 172 deviation can be associated with different meteorological situations. The larger spread
 173 near the surface is expected since there is more variability in that region. Also, the peak
 174 seen at 4 km in the standard deviation is due to the averaged water vapor mixing ratio
 175 being almost equal to the standard deviation at that point.

Altitude interval (km)	Vertical resolution (km)
0 - 0.1	0.01
0.1 - 1	0.025
1 - 3	0.1
3 - 5	0.2
5 - 8	0.5
8 - 12	0.5
12 - 20	2

Table 1. Vertical resolution of the atmospheric profiles from IGRA

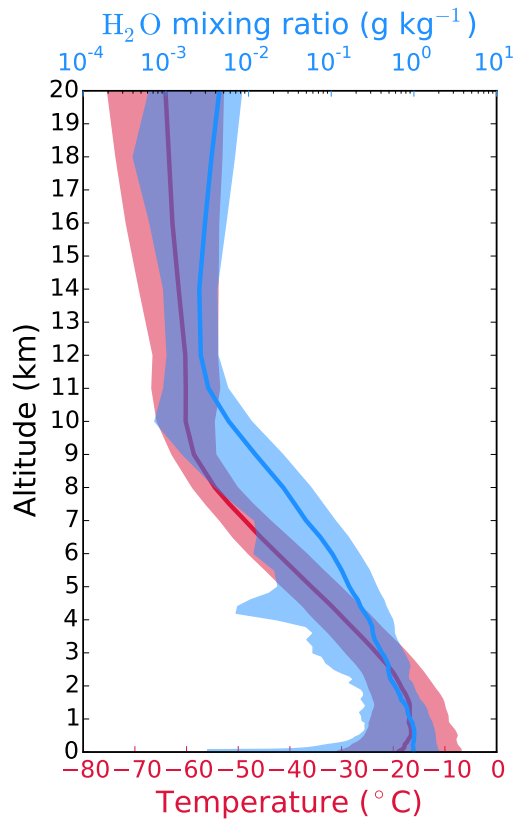


Figure 3. Temperature and humidity profiles averaged for the 48 radiosoundings at eight Arctic stations shown with the red and blue lines respectively. The shaded area shows the standard deviation associated with the variables.

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2.3 Theoretical framework

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In this section, the theoretical framework used is explained. This study is based on linear statistical estimation theory in the context of numerical weather prediction (NWP) (Rodgers, 2000). The different notations, definitions, approximations and data used are described in this section (Lewis et al., 2006).

181 The atmospheric state is represented by a vector \mathbf{x} and the satellite radiance mea-
 182 surements at different wavelengths at the TOA are represented by the vector \mathbf{y} .
 183 The observation is related to the atmospheric state through the equation

$$\mathbf{y} = H(\mathbf{x}) + \epsilon_o, \quad (1)$$

184 where H is the forward model linking the observation to the atmospheric profile
 185 and ϵ_o is the observation error. In this case, the state corresponds to atmospheric pro-
 186 files of temperature, \mathbf{T} , and logarithm of specific humidity $\ln \mathbf{q}$ defined on \mathbf{k} vertical lev-
 187 els on which the model state is defined. The dimension of the model state \mathbf{x} is thus $2\mathbf{k}$.
 188 The ozone and other trace gases are kept constant. The assimilation seeks to correct an
 189 *a priori* estimate of the state of the atmosphere, \mathbf{x}_b , also referred to as the *background*
 190 *state*, using the information contained in the observations. It takes into account the rela-
 191 tive accuracies of \mathbf{x} and \mathbf{y} to obtain a minimum variance estimate, \mathbf{x}_a , called the *anal-*
 192 *ysis*.

193 A linearization of the forward model around the atmospheric profile, \mathbf{x}_b is done,
 194 which gives, assuming that the radiative-transfer equation is weakly nonlinear near the
 195 background state

$$H(\mathbf{x}) \cong H(\mathbf{x}_b) + \mathbf{H}(\mathbf{x} - \mathbf{x}_b), \quad (2)$$

196 where $H(\mathbf{x}_b)$ is the background state in the observations space and $\mathbf{H} = \left. \frac{\partial H(x)}{\partial x} \right|_{x_b}$
 197 is the linearized observation operator with respect to \mathbf{x} evaluated at $\mathbf{x} = \mathbf{x}_b$, referred
 198 to as the Jacobian.

199 The analysis, \mathbf{x}_a , which represent the corrected atmospheric state after the obser-
 200 vations and the model are taken into consideration, is given by

$$\mathbf{x}_a = \mathbf{x}_b + \mathbf{K}(\mathbf{y} - \mathbf{H}\mathbf{x}_b) \quad (3)$$

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 202 with $\mathbf{K} = \mathbf{B}\mathbf{H}^T(\mathbf{R} + \mathbf{H}\mathbf{B}\mathbf{H}^T)^{-1}$ being the gain matrix. \mathbf{B} is the background er-
 203 ror covariance matrix and \mathbf{R} , the observation error covariance matrix (Rodgers, 2000).
 204 The superscript T and -1 denote respectively the transpose and inverse of a matrix.

205 2.3.1 Jacobians

206 For each band, the Jacobian indicates how temperature and humidity variations
 207 in each band impact the radiance measured at the TOA. The Jacobians were obtained
 208 by finite difference with the radiative-transfer model MODTRAN v 5.4 (Berk et al., 2005)
 209 by perturbing the background state \mathbf{x}_b , in this case a temperature and humidity profile.
 210 The Jacobians were computed for each band of the FIR radiometer and AIRS. A sub-
 211 set of 142 channels were used for AIRS. It is assumed that the FIR radiometer and AIRS
 212 are collocated on a pixel with the same spatial response and that the lag between the
 213 instruments is negligible. More specifically, for the temperature Jacobians, \mathbf{H}_{T_i} , at the
 214 level i , a perturbation of ± 0.5 K was done (Garand et al., 2001). Perturbations of 1 K
 215 have been deemed sufficiently small for this experiment. This gives the variation of ra-
 216 diance seen at the TOA for a variation of 1 K in the atmospheric profile at each atmo-
 217 spheric level. In the same manner, the humidity Jacobians in logarithm of specific hu-
 218 midity, $s = \ln q$, at the level i , were obtained by perturbations of ± 0.05 q, where q is the
 219 specific humidity and $s = \ln(q)$. As shown with the following equation, in order to ob-
 220 tain a Jacobian with respect to a logarithm, the perturbations are done on the profile
 221 in q. Thus by multiplying the difference of perturbations by 10, this results in Jacobians
 222 with the units of $\text{Wm}^{-2} \text{sr}^{-1} \log(\text{L L}^{-1})^{-1}$.

$$\mathbf{H}_{s,i} = \frac{\partial R}{\partial \ln q} = \frac{\partial q}{\partial \ln q} \frac{\partial R}{\partial q} = q \frac{\partial R}{\partial q} \quad (4)$$

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2.3.2 Background error covariance matrix

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The matrix \mathbf{B} is the background error covariance matrix associated with the background state \mathbf{x}_b . Figure 4 represents the \mathbf{B} matrix used for temperature (left) and humidity (right). The \mathbf{B} matrices are the stationary components of the background term of the Environment Canada assimilation system (Buehner et al., 2015). Those matrices were evaluated for a latitude of $79^\circ 59' 20''$, which corresponds to Eureka, Canada for the month of February. The units used are K^2 and $\log(\text{L L}^{-1})^2$ for temperature and humidity respectively. The cross-terms \mathbf{B}_{Ts} and \mathbf{B}_{sT} are considered equal to zero, and thus only the components \mathbf{B}_{TT} and \mathbf{B}_{ss} of the \mathbf{B} matrix are considered. Making this approximation allows to calculate the DFS and analysis error for temperature and humidity separately. In this study, the matrices \mathbf{B}_{TT} and \mathbf{B}_{ss} are kept constant.

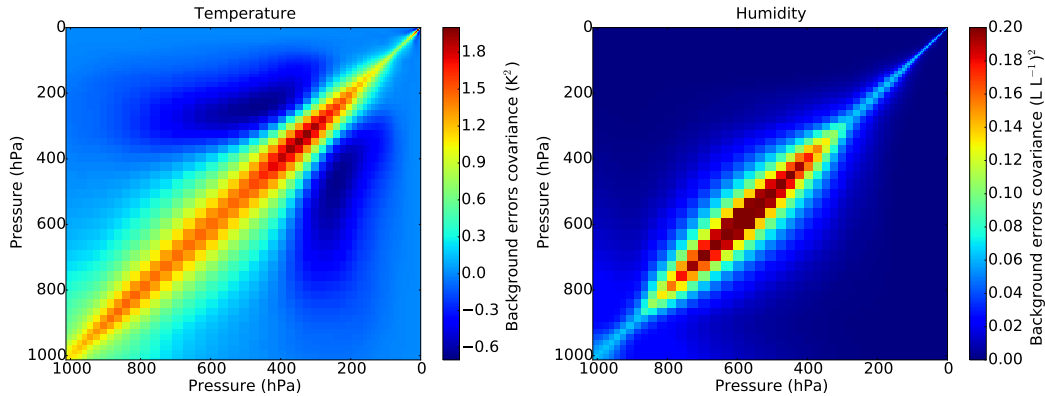


Figure 4. Background error covariances matrices \mathbf{B} for temperature (left) and logarithm specific humidity (right) at a latitude of $79^\circ 59' 20''$ for the month of February

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2.3.3 Observation error covariance matrix

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The matrix \mathbf{R} is the observation error covariance matrix. Normally, the matrix \mathbf{R} takes into consideration the measurement error, the forward-model error, the representativeness error and the error associated with quality control to name a few (Bormann et al., 2010) but for this study, only the measurement error was considered. This approximation was taken to be consistent with previous studies in the FIR region such as Merrelli and Turner (2012), Shahabadi and Huang (2014) and Mertens (2002). They show that, in the thermal IR region, the main contribution to the observation error is the measurement error and also, the interchannel correlation error is small (Garand et al., 2007). The measurement error is assumed to be gaussian and unbiased, assumptions used especially in data assimilation (Rodgers, 2000). Therefore, the matrix \mathbf{R} is assumed to be diagonal with the NER values on the diagonal. The spectral NER for the instrument AIRS was taken from the AIRS website (<https://airs.jpl.nasa.gov/index.html>) version 5 L1B data. The NER for the synthetic instrument is assumed to be constant for each configuration and each band since it comes from the sensor.

2.3.4 Information content

The impact of measurements is estimated from the analysis error covariance and the degrees of freedom per signal (DFS). The analysis error, assumed here to be unbiased, is $\varepsilon_a = \mathbf{x}_a - \mathbf{x}_t$ where \mathbf{x}_t is the true state of the atmosphere. So, $\mathbf{A} = \langle \varepsilon_a \varepsilon_a^T \rangle$, with $\langle \dots \rangle$ being the statistical average, is the analysis error covariance matrix and can be shown to be

$$\mathbf{A} = (\mathbf{I} - \mathbf{KH})\mathbf{B}. \quad (5)$$

The reduction of analysis error due to the assimilation of observations is measured by

$$\text{tr}(\mathbf{AB}^{-1}) = N - \text{tr}(\mathbf{KH}), \quad (6)$$

where $\text{tr}(\mathbf{KH}) = \text{tr}(\mathbf{HK})$. The gain in information, or the DFS is defined as

$$\text{DFS} = \text{tr}(\mathbf{HK}). \quad (7)$$

The DFS can then be viewed in two ways, in the observation space and in the model space. In the observation space, the DFS measures the independent degrees of freedom measured by the observations and take into account redundancy. In the model space, it measures the reduction of analysis error with respect to the background error.

It is an evaluation technique based on the relative errors between the observations and the prior information (Purser & Huang, 1993). It has also been used to quantify the added value of a new set of observations by comparison with other types of measurements and also on top of measurements already assimilated (McNally et al., 2006; Lupu et al., 2011)

Thus, the analysis error variance matrix \mathbf{A} and the DFS depend on the background error covariance matrix \mathbf{B} , the observation error covariance matrix \mathbf{R} and the Jacobian matrix \mathbf{H} . The DFS will be used as a metric for obtaining the optimal configuration of the FIR radiometer and to discuss the trade-off between spectral resolution and noise level. The analysis error variance matrix will be used to see the vertical impact when the FIR radiometer is assimilated. Those calculations for the DFS and analysis error variance were done for the 48 atmospheric profiles individually, and are then averaged. The calculations for 48 atmospheric profiles show the added value on average and also its variability for the different possible atmospheric situations in the Arctic. The standard deviation spread for the DFS and the analysis error variance will be shown in figures 7 and 10 respectively.

3 Evaluation of configurations

3.1 Optimization under constraints

In this section, the DFS is used to discuss the trade-off between spectral resolution and noise level.

Figure 5 shows the total DFS for temperature of different configurations for a synthetic FIR radiometer with constant wavenumber configuration when the NER error level varies between 0.0003 and 0.02 $\text{Wm}^{-2}\text{sr}^{-1}$ and a spectral range of 15 to 100 μm . The total number of bands varies between 1 and 200 bands. The color represents the value of the DFS for this configuration. Hence, this figure shows that for a fixed NER, the DFS increases and then decreases as the number of bands increases. This figure can be useful when there are technological constraints for example. If the NER is imposed by the available technology, taking a horizontal line on the top panel of figure 5 highlights the

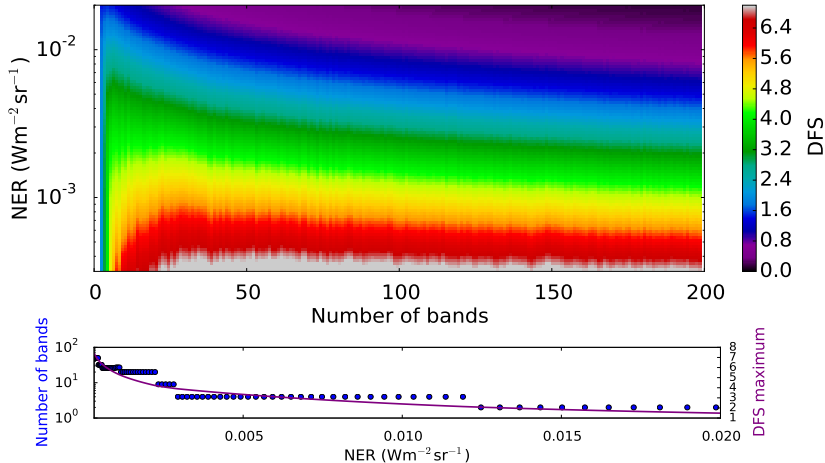


Figure 5. Top panel: The averaged total DFS is shown for variations of the NER level (y-axis) and variations of the total number of bands for temperature of the constant wavenumber bands configuration. The DFS for each configuration is shown with the colorbar. Bottom panel: The number of bands that maximizes DFS as a function of the NER (blue dots) and the DFS maximum as a function of the NER (purple line).

290 available spectral configurations. The maximum DFS is not with smaller bands (right
 291 side of the figure) but always with a configuration which has less than 50 bands. This
 292 is due to the constant NER. By having smaller bands, the amount of energy per band
 293 decreases and hence the signal-to-noise ratio decreases. This is also shown in the bot-
 294 tom panel, which shows with blue dots the number of bands of the configuration with
 295 the maximal DFS for a variation of the NER. This shows that having more bands is not
 296 always the best configuration, since the DFS is not increasing as the number of bands
 297 increases. The number of bands with the peak DFS gets larger as the NER decreases.
 298 Another interesting way to analyze this map is by having a constraint on the number
 299 of bands an instrument can have. A radiometer can be operated with a filter-wheel and,
 300 as the number of bands increases, the rate of repetition decreases and also the cost in-
 301 creases.

302 Figure 6 is similar to figure 5 but for humidity, but was cut off at 100 bands to bet-
 303 ter see the shift in the DFS peak. It has a lot more variability in the DFS for an hori-
 304 zontal line compared to temperature. This variability is partly due to spectral features
 305 of transmittance. For humidity, the maximum DFS is always obtained with a configu-
 306 ration which has less than 55 bands.

307 3.2 Maximisation of the total DFS

308 In this section, the DFS is used to find an optimal configuration for the FIR ra-
 309 diometer considered in this study. Three different splitting of the bands are considered,
 310 equi-energetic, constant wavelength bands and constant wavenumber bands. The instru-
 311 ment will be split between 1 and 250 bands.

312 Considering those three configurations, figure 7 shows the total DFS averaged over
 313 the 48 profiles for temperature and humidity. Two values of the NER were used for the
 314 level of error, namely $0.01 \text{ Wm}^{-2}\text{sr}^{-1}$ and $0.002 \text{ Wm}^{-2}\text{sr}^{-1}$. To show the variability, the
 315 standard deviation is shown with the shaded area. Note that several papers have inves-
 316 tigated the channel selection and information content of AIRS (e.g. Fourrié and Thépaut

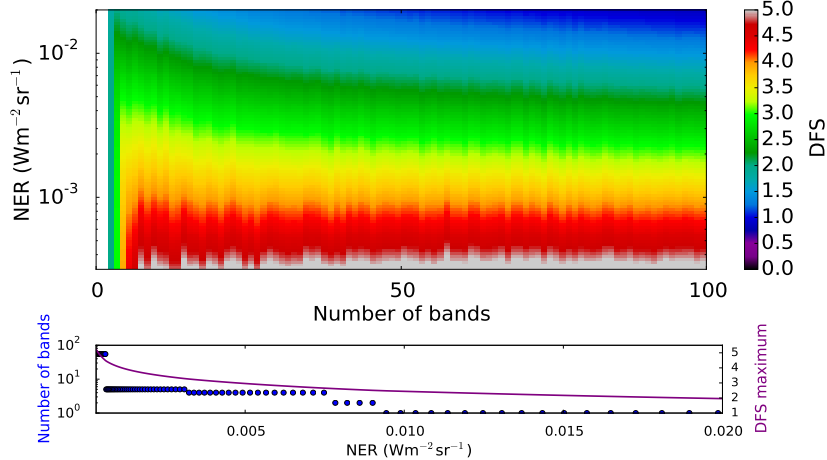


Figure 6. Top panel: The averaged total DFS is shown for variations of the NER level (y-axis) and variations of the total number of bands for humidity of the constant wavenumber bands. The DFS for each configuration is shown with the colorbar. Bottom panel: The number of bands that maximizes DFS as a function of the NER (blue dots) and the DFS maximum as a function of the NER (purple line).

(2003); Divakarla et al. (2006); Garand et al. (2007)). However, different \mathbf{B} and \mathbf{R} matrices were used, which prevents direct comparison with this previous work.

For temperature, for the three cases, as the number of bands increases the total DFS decreases. This is due to the constant NER, which results in less energy when the bandwidth is reduced. For the configuration having constant wavelength bands, the peak in DFS is reached with more bands compared to the other configurations. The bands being too wide near $15 \mu\text{m}$, there is not enough resolution to capture the variation in the transmittance in this part of the spectrum. Table 2 shows the maximum DFS for each configuration with the corresponding number of bands for both temperature and humidity for the target noise. When the NER is reduced, there is a large increase in the DFS as expected since it means that the measurements are more accurate. For example, there is an increase by a factor of 2.00 in the DFS for the constant wavenumber band configuration when the NER goes from $0.01 \text{ Wm}^{-2}\text{sr}^{-1}$ to $0.002 \text{ Wm}^{-2}\text{sr}^{-1}$. For the different configurations, the standard deviation varies between 0.209 to 0.247 which is less than for AIRS which is equal to 0.513 for temperature. The highest DFS is with the constant wavenumber band configuration with 22 bands for the NER level of $0.002 \text{ Wm}^{-2}\text{sr}^{-1}$. This is the configuration that will be used for temperature for the remainder of this study. The individual DFS of the optimal configuration of the 22 constant wavenumber bands is 20 % smaller than AIRS. However, when those bands are assimilated on top of AIRS, (table 2), the value of the DFS increases by 13.2 % compared to when AIRS is assimilated alone. This means that even after the information in the thermal IR is assimilated, there is still value in assimilating data in the FIR.

For humidity, the highest DFS is also with the constant wavenumber bands configuration with now 7 bands and a total DFS of 3.594 for the NER error of $0.002 \text{ Wm}^{-2}\text{sr}^{-1}$. For the remainder of the study, this configuration will be considered as the optimal configuration. Similarly to temperature, there is more variability for the constant wavelength configuration. For humidity, the standard deviation is larger compared to temperature, i.e. it varies between 0.318 and 0.354 depending on the configuration while it is 0.666 for AIRS. Also, when compared individually to AIRS (table 2), the DFS for the optimal configuration with 7 constant wavenumber bands is smaller than the DFS of AIRS by 14%.

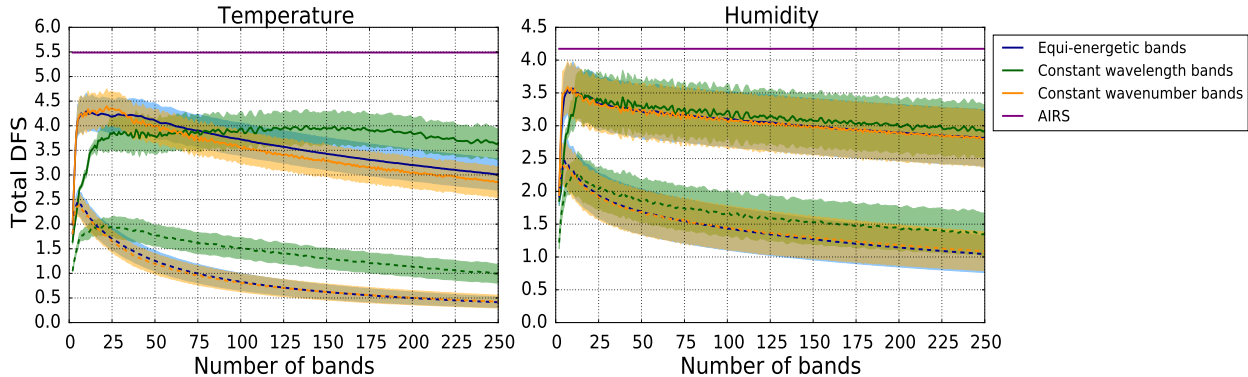


Figure 7. The averaged total DFS as a function of the total number of bands for three configurations which are equi-energetic bands (blue line), constant bandwidths in terms of wavelength (green line) and wavenumber (orange line) for temperature (left) and humidity (right). The dashed lines are for a NER of $0.01 \text{ Wm}^{-2}\text{sr}^{-1}$ whereas the full lines are for the target NER of $0.002 \text{ Wm}^{-2}\text{sr}^{-1}$. The shaded area represent the standard deviation of the 48 atmospheric profiles which are shown for all configurations except for AIRS. The standard deviation of AIRS is equal to 0.53 and 0.67 for temperature and humidity respectively. The purple line represents the averaged total DFS of AIRS for the 48 atmospheric profiles.

347 When those 7 bands in the FIR are assimilated on top of AIRS (table 2), the DFS in-
 348 creases by 11.5%, compared to assimilating only AIRS, which shows that measurements
 349 in the FIR add information when assimilated on top of AIRS data.

350 3.3 Selection of the bands with most information

351 Considering a fixed number of bands for the FIR radiometer, following Rabier et
 352 al. (2002), we now evaluate the DFS sequentially. First, for each atmospheric profile, the
 353 DFS is calculated for each band of the configuration and the one that maximises the DFS
 354 is selected. It is shown as the first position in figure 8. The next band selected is the one
 355 that, when added to the previous one, adds the largest information content. This pro-
 356 cess is done until all the bands are selected. Each new band thus optimally increases the
 357 DFS. This type of calculation was also done for ice cloud properties for AIRS by Chang
 358 et al. (2017). This selection was done for each of the 48 atmospheric profiles and for both
 359 temperature and humidity. Thus, figure 8 shows the frequency each band is selected at
 360 each position for the optimal configuration of the FIR radiometer with 22 constant wavenum-
 361 ber bands for temperature. It shows that 50 % of the time, the first band selected cor-
 362 respond to the first band (dark orange) in the splitting with the bandwidth $15.02\text{-}15.62$
 363 μm . In first and second position, the bands selected are always between $15.02 \mu\text{m}$ and
 364 $20.58 \mu\text{m}$. Also, in the last position, the last band (dark pink) with bandwidth 79.53-
 365 $100 \mu\text{m}$ is selected 66.67 % of the times whereas the second to last band in the splitting
 366 (bandwidth $66.03\text{-}79.53 \mu\text{m}$) is selected 33.33 % of the time.

367 Similarly for humidity, figure 9 shows the order of selection of the 7 constant wavenum-
 368 ber bands with respect to humidity through the atmospheric profiles. It shows that the
 369 first band selected is 58.3 % of the time the third band (mint green) which has bound-
 370 aries of $19.83 - 23.62 \mu\text{m}$ whereas the second band (yellow) is selected 22.92 % of the times.
 371 For the second band selected, 62.5 % of the time, it is the 5th band (dark blue) that is
 372 selected. The last band selected is always the first band (orange) which has boundaries
 373 between $15.02 - 17.09 \mu\text{m}$.

	Number of bands	DFS	Analysis error variance
Temperature			
Constant wavenumber bands	22	4.399	12.64
Constant wavelength bands	138	3.996	-
Equi-energetic bands	10	4.294	-
AIRS	-	5.488	10.58
Background error	-	-	19.62
AIRS + constant wavenumber bands	22	6.213	9.97
Humidity			
Constant wavenumber bands	7	3.594	1.09
Constant wavelength bands	15	3.482	-
Equi-energetic bands	10	3.569	-
AIRS	-	4.173	1.03
Background error	-	-	1.96
AIRS + constant wavenumber bands		4.714	0.95

Table 2. Total averaged DFS for a NER of $0.002 \text{ Wm}^{-2}\text{sr}^{-1}$ and analysis error variance. The units for the analysis error variance is K^2 and $\log(LL^{-1})^2$ for temperature and humidity respectively

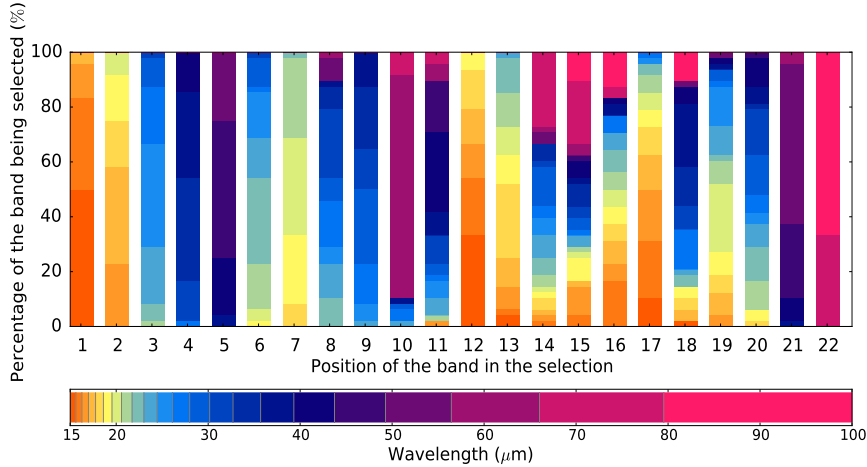


Figure 8. Probability of each band to be selected at each position for the 22 constant wavenumber bands with respect to temperature.

374 4 Analysis error

375 The impact of the observations can be seen with the analysis error variance since
 376 it provides information about the vertical distribution. The DFS previously discussed
 377 gives information integrated through the profile. Equation 2.3 is used to obtain the anal-
 378 ysis error variance. To show the added value of the FIR radiometer when other types
 379 of instruments are assimilated, the instrument AIRS is considered.

380 Figure 10 shows the analysis error variance profile for temperature and humidity
 381 when the optimal FIR radiometer and AIRS are assimilated. The dark green curve rep-
 382 represents the background error, the dark blue curve is the average over the 48 atmospheric
 383 profiles for the FIR radiometer whereas the shaded area represents the standard devi-
 384 ation associated with the variability obtained through the different profiles. The dark

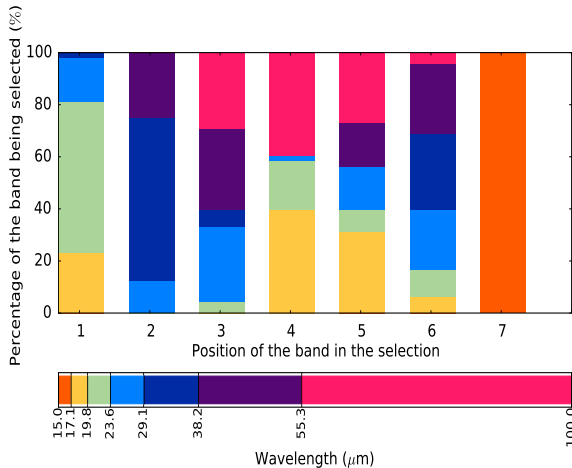


Figure 9. Probability of each band to be selected at each position for the 7 constant wavenumber bands with respect to humidity.

385 pink curve represents the averaged analysis error variance for AIRS whereas the peach
 386 curve shows when the FIR radiometer is assimilated with AIRS. For temperature, AIRS
 387 is better at reducing the error compared to the FIR radiometer except at 300 hPa. It
 388 is also at this height that the gain in assimilating the FIR radiometer on top of AIRS
 389 is seen. Also, the sum of the analysis error variance through the profile is shown in table
 390 2, which allows to see the impact of the observations in the model space. Hence, as-
 391 similating the FIR radiometer after AIRS allows to reduce by 5.65 % the analysis error
 392 variance. The main gain of the FIR radiometer is with respect to humidity. The FIR
 393 radiometer is better at reducing the error in the upper atmosphere, between 400 hPa and
 394 200 hPa, than AIRS. When the 7 bands of the FIR radiometer are assimilated on top
 395 of AIRS, there is a non negligible gain near the surface (between the surface and 850 hPa)
 396 and in the upper part of the atmosphere (between 400 hPa and 200 hPa). The gain near
 397 the surface is due to the FIR radiometer being sensitive to the temperature inversion layer.
 398 For the different profiles, there is some variability in the atmospheric conditions which
 399 is seen with the standard deviation, especially between the surface and 600 hPa which
 400 is expected. In the same way, with table 2, it shows that assimilating the FIR radiometer
 401 after AIRS allows to reduce by 12.84 % the analysis error variance for humidity. Mertens
 402 (2002) did a similar study on the ability of the FIR to improve water vapor retrievals.
 403 The conclusions are similar in the sense that both studies find that the main reduction
 404 in the analysis error variance is between 1000 hPa and 100 hPa and that there is a gain
 405 in using both the MIR and the FIR.

406 To show the impact of each individual band on the analysis error variance profile,
 407 the humidity Jacobians of the FIR radiometer (left panels) and the analysis error vari-
 408 ance associated with each of these bands when assimilated sequentially (right panels)
 409 is shown in figure 11 for two specific atmospheric cases with (top panel) and without (bot-
 410 tom panel) a temperature inversion. The colors of the bands of the Jacobians are asso-
 411 ciated with the colorbar. The top left panel of figure 11 shows the signature of two ef-
 412 fects: the greenhouse effect and the presence of a temperature inversion layer. The neg-
 413 ative peak is due to the greenhouse effect of water vapor. Increasing the humidity tends
 414 to reduce emission of radiance by masking the lower warmer layers. The positive part
 415 near the surface is due to an inversion. Increasing humidity elevates the effective emis-
 416 sion altitude, where the atmosphere is warmer due to the inversion (?). Moreover, as
 417 the wavelength of the band increases, the peak’s height increases from around 600 hPa

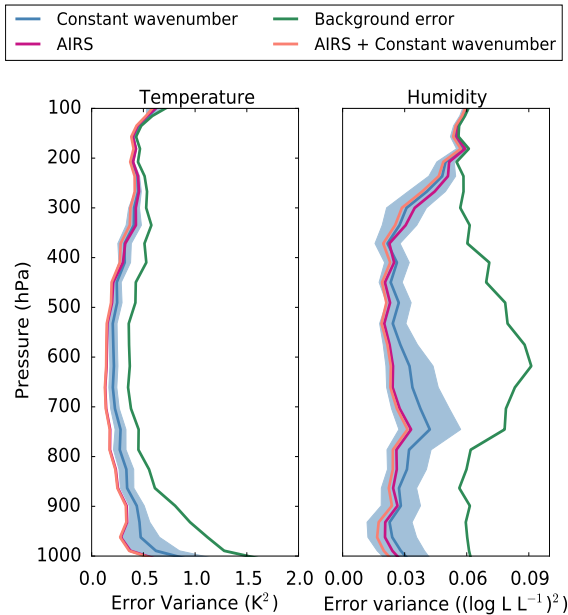


Figure 10. Analysis error variance profile for temperature (left) and humidity (right) for the optimized FIR radiometer. The green curve represents the background error **B**, the dark blue curve the average on the 48 atmospheric profile whereas the shaded area represents its associated standard deviation. The dark pink line is the analysis error variance of AIRS whereas the peach line represents when the optimized FIR radiometer is assimilated after AIRS.

418 to 350 hPa. The top right panel of figure 11 shows, as in figure 10, the analysis error vari-
 419 ance for the background error, when all the bands of the FIR radiometer and AIRS are
 420 assimilated separately which is represented by the dark green, blue and dark pink lines
 421 respectively. To show the impact on the analysis error of each band of the FIR radiome-
 422 ter, bands were sequentially assimilated and the analysis error variance was calculated
 423 after each new band of the FIR radiometer was assimilated. The order of assimilation
 424 of each band is the same as in section 3.3. It shows that the first band assimilated, the
 425 pale green one, reduces the error between 800 hPa and 300 hPa, which is where the Ja-
 426 cobian’s peak is. The second band, navy blue one, has a higher Jacobian’s peak (at around
 427 400 hPa) and it is mainly where the analysis is reduced, between 650 hPa and 200 hPa.
 428 The next two bands assimilated, yellow and purple lines respectively, are also shown in
 429 the figure. The reduction of the analysis error near the surface is mainly due to the yel-
 430 low band (17.1- 19.8 μm), which is interesting since it allows to restrain the uncertainty
 431 in the inversion layer. This illustrates the complementarity of these bands to obtain the
 432 best analysis over the whole vertical extent. Compared to AIRS (dark pink line), the FIR
 433 radiometer is better at reducing the error between 350 hPa and 250 hPa. To show the
 434 impact of the temperature inversion, the bottom panels of figure 11 shows the Jacobians
 435 and the analysis error variances for an atmospheric case without a temperature inver-
 436 sion. The left panel shows that the Jacobians are more spread out and sample the at-
 437 mosphere from near the surface up to 300 hPa. Also, compared to the other case, there
 438 is no positive peak in the Jacobians which is expected since it is due to a temperature
 439 inversion layer. For the analysis error variance (bottom right panel), with the Jacobians
 440 peak being lower, the analysis error variance is less reduced compared to the case with
 441 an inversion. Even though, the FIR radiometer reduces more the analysis error variance
 442 than AIRS between 350 hPa and 200 hPa. Also, near the surface, the FIR radiometer
 443 is better at reducing the analysis error variance when there is an inversion.

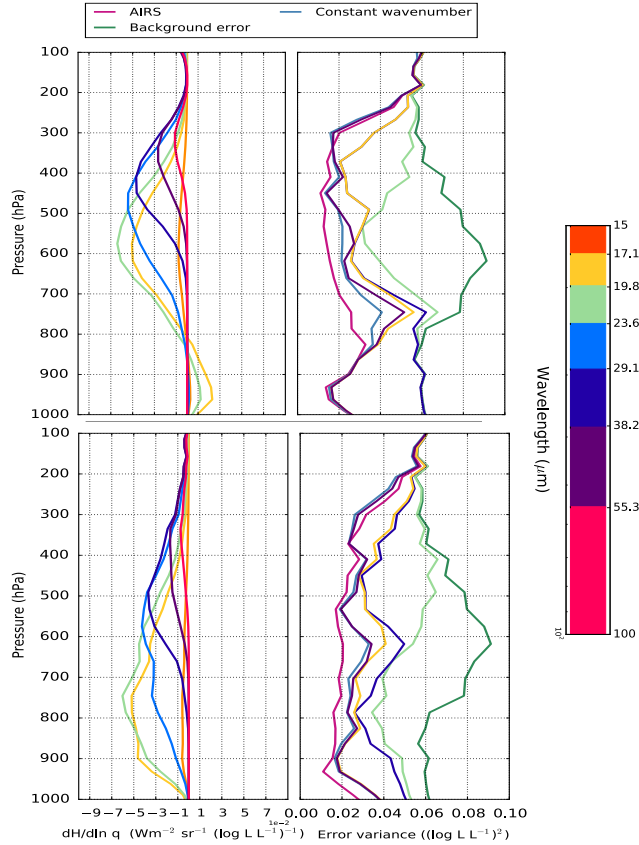


Figure 11. Left panels show the humidity Jacobians for the FIR radiometer associated with two atmospheric cases, which are with (top panel) and without (bottom panel) a temperature inversion. The different colors are the 7 associated bands of this configuration shown by the colorbar which is valid for both panels. The right panel show the analysis error variance for the two cases. The curves from right to left are the background error, when the bands of the FIR radiometer are assimilated one at the time from one band to four (light green, dark navy, yellow and purple), when all the 7 bands are assimilated and AIRS is assimilated.

5 Conclusions

The objective of the present study is to design an optimal FIR radiometer to study the Arctic polar night by examining different configurations, noise levels and the trade-off between spectral resolution and noise level. This was investigated through an information content analysis based on optimal estimation method. The optimal configuration for the synthetic FIR radiometer is with 22 constant wavenumber bands for temperature and 7 constant wavenumber bands for humidity. It was shown that too many bands with a large noise do not give enough information on the atmosphere. With a few bands, it was possible to get a DFS similar to AIRS when compared individually. Given that AIRS provides information also on both temperature and humidity, the impact of assimilating FIR measurements on top of AIRS data was evaluated by the reduction in analysis error variance. With respect to temperature, there is a small impact in assimilating the FIR radiometer measurements over AIRS between 400 and 250 hPa. On the contrary, for humidity, there is a non negligible gain near the surface (between the surface and 850 hPa) and in the upper part of the atmosphere (between 400 hPa and 200

459 hPa). Something else that is worth noting is that between 400 hPa and 200 hPa, taken
 460 individually, the FIR radiometer is better at reducing the humidity analysis error vari-
 461 ance than AIRS.

462 Measurements in the FIR are unlikely to be assimilated in the next few years in
 463 NWP systems, however the results shown in this paper highlight the potential of this
 464 new type of observations which may become available in the next decade. It is non neg-
 465 ligible to get results similar to AIRS in reducing the analysis error for humidity with only
 466 7 bands compared to a subset of 142 bands from AIRS. FIR measurements could be used
 467 in regions where there is still large uncertainties in water vapor retrieval or assimilation.
 468 It was shown to be useful for retrieval of water vapor in the 400 hPa to 200 hPa region
 469 for the Arctic, but FIR radiometry can be useful in other regions as well.

470 The results presented here, are based on a 1D assimilation of two collocated instru-
 471 ments. Another interesting aspect of this study is the method, which facilitates test-
 472 ing rapidly multiple configurations of an instrument. Also, it allowed to compare the rel-
 473 ative impact of measurements in the FIR and the MIR. However, there are limitations
 474 to this approach that need to be kept in mind. Because a satellite does provides mea-
 475 surements over the whole globe, it would be important to examine the impact one could
 476 expect in other regions such as the Tropics for instance. Finally, complex Observing Sim-
 477 ulated Systems Experiments (OSSEs) would be needed to evaluate the global impact in
 478 a context including all observations currently assimilated.

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 486 These simulations and the codes used to generate the figures are available from Laurence
 487 Coursol.

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