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# 1 **Seismic evidence for partial melt below tectonic plates**

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5

6 **The seismic low velocity zone (LVZ) of the upper mantle is generally**  
7 **associated with a low-viscosity asthenosphere that plays a key role for the**  
8 **dynamics of plate tectonics<sup>1</sup>. However, its origin remains enigmatic, some**  
9 **authors attributing the reduction in seismic velocity to a small amount of**  
10 **partial melt<sup>2,3</sup>, others invoking solid-state mechanisms near the solidus<sup>4-6</sup>, or**  
11 **the effect of volatile contents<sup>6</sup>. Observations of shear attenuation provide**  
12 **additional constraints to unravel the origin of the LVZ<sup>7</sup>. Here, we report the**  
13 **discovery of partial melt within the LVZ from the simultaneous**  
14 **interpretation of global 3D shear attenuation and velocity models. We**  
15 **observe that partial melting down to 150-200 km depth beneath mid-ocean**  
16 **ridges, major hotspots and back-arc regions feeds the asthenosphere. A small**  
17 **part of this melt (<0.30%) remains trapped within the oceanic LVZ. The**  
18 **amount of melt is related to plate velocities and increases significantly**  
19 **between 3 and 5 cm yr<sup>-1</sup>, similar to previous observations of mantle crystal**  
20 **alignment underneath tectonic plates<sup>8</sup>. Our observations suggest that by**  
21 **reducing viscosity<sup>9</sup>, melt facilitates plate motion and large-scale crystal**  
22 **alignment in the asthenosphere. Melt is absent under most of the continents.**

23 Our finding results from the simultaneous analysis of two upper mantle  
24 tomographic models of shear wave velocity ( $V_s$ ) and attenuation (parameterised

25 with  $Q_s$  the quality factor, a measure of energy dissipation). Until now, most global  
26 tomographic studies of the upper mantle and their thermochemical interpretation  
27 have focused on shear velocity<sup>4,7</sup>. Recent experiments on olivine suggest that wave  
28 speed and attenuation are insensitive to water, which would imply that elevated  
29 water contents are not responsible for the LVZ<sup>10</sup>. However,  $V_s$  is sensitive to  
30 temperature, composition and melt content, and deciphering the causes of its  
31 variations represents a strongly non-unique inverse problem, severely limiting our  
32 understanding of the Earth's interior. Shear attenuation has a different sensitivity  
33 to these quantities, and therefore provides complementary constraints on the  
34 origin of seismic heterogeneities<sup>11</sup>. Shear attenuation is negligibly dependent on  
35 major element chemistry<sup>12</sup>, and exponentially dependent on temperature<sup>13</sup>. The  
36 relation between attenuation and melt is debated, with some authors arguing for a  
37 weak dependence based on experiments, models and seismological observations<sup>12</sup>,  
38 while others suggest a larger effect<sup>3</sup>. Measuring shear attenuation is nevertheless  
39 difficult. For this reason, only a few global  $Q_s$  models have been published in the  
40 last 20 years<sup>14</sup>, and the only recent joint interpretation of 3D  $Q_s$  and  $V_s$   
41 tomographic models at global scale is based on models built from different  
42 datasets and modeling approaches<sup>15</sup>.

43 The novelty of our study is to simultaneously interpret two recent global  $V_s$  and  $Q_s$   
44 models that are consistent as obtained from the same Rayleigh wave dataset, at the  
45 same resolution and using the same modelling approach. These  $V_s$  (DR2020s) and  
46  $Q_s$  ( $Q_s$ ASR17) models are displayed in **Figure 1**. Details of our tomographic  
47 procedure can be found in Methods. Before interpreting these two models in the  
48 light of laboratory experiments, a few words are needed to emphasize in what  
49 temperature and pressure range our interpretation is pertinent. The attenuation  
50 models derived by mineral physicists<sup>4,13</sup> are valid for temperatures  $T$  larger than  
51 900°C, which correspond to the base of the lithosphere and to the asthenosphere.

52 They consider thermally activated processes varying exponentially with  $1/T$  that  
53 would imply quality factor in excess of 2,000 in the upper 100 km of the  
54 lithosphere and reaching several million near the surface (see Methods). However,  
55 finite  $Q_s$  is observed in the crust and in the cold mantle lithosphere where  
56 attenuation is most likely related to non-thermal processes such as scattering or  
57 fluid-fracture interactions<sup>14</sup>. Furthermore, the attenuation observed in seismology  
58 accumulates along the seismic ray and the observation of  $Q_s$  is only possible when  
59 the amplitude of a wave is measurably smaller than in a pure elastic model. Given  
60 the uncertainties on amplitude data, it appears impossible to resolve quality  
61 factors larger than  $\sim 2,000$  with long period Rayleigh waves (at periods of 100s,  
62 assuming velocities of  $4.5 \text{ km s}^{-1}$  and ray lengths of 10,000 km, the amplitude  
63 reduction with respect to a pure elastic model, would be less than 3.5%). Our  
64 inversion leads indeed to a  $Q_s$  model with strong lateral variations, by two orders  
65 of magnitudes, but with a maximum  $Q_s$  of  $\sim 1,750$ . Our attenuation model is  
66 therefore mostly adapted to depths greater or equal to  $\sim 100$  km, where  $Q_s$  values  
67 between fifty and a few hundreds are expected<sup>15</sup>. Therefore, our interpretation  
68 applies to the oceanic asthenosphere and the mantle structure at depths greater or  
69 equal to  $\sim 100$  km, where our tomographic models are accurate and where  
70 conditions similar to those used in laboratory experiments prevail.

71 **Figure 2** displays  $Q_s$  as a function of  $V_s$  for each pixel of the maps, at different  
72 depths in the upper mantle. The curves in dark and light blue represent the  
73 theoretical relations due to temperature variations for a meltless pyrolytic  
74 mantle<sup>16</sup>, given by two anelasticity models<sup>4,13</sup> based on laboratory experiments,  
75 appropriate for asthenospheric conditions (see Methods for details). These models  
76 only explain a limited part of the velocities and attenuations of our dataset. We  
77 consider that the first theoretical curve<sup>4</sup> is compatible with a given  $V_s$ - $Q_s$   
78 observation if it falls within the typical uncertainties of the  $V_s$  and  $\ln(Q_s)$

79 observations, 1% and 10% respectively (ivory colour). We show in Methods and  
80 supplementary **Figures S1-10** that reconciling the low values of  $V_s$  with  $Q_s$  cannot  
81 be done by invoking radial anisotropy, elastically accommodated grain boundary  
82 sliding, or the effect of composition or water. However, this can be done by adding  
83 partial melt, thus reducing  $V_s$ , and shifting the theoretical curves to the left in  
84 **Figure 2**, since as discussed below, adding partial melt reduces  $V_s$  but has little effect  
85 on  $Q_s$ . Warm colours in **Figure 2** indicate the amount of melt ( $<0.7\%$ ) needed to  
86 reconcile observations with the first theoretical model<sup>4</sup> (dark blue curve), which  
87 requires the smallest amount of melt. On the right side of the ivory region, points in  
88 grey are those for which  $V_s$  is too high compared to the theoretical value. They are  
89 associated with the lithospheric depletion of cratonic roots. In this case, the grey  
90 intensity quantifies the departure in percent from the theoretical curve assuming a  
91 meltless pyrolitic mantle. Results using the second theoretical model<sup>13</sup> are shown  
92 in **Figures S11-12**. They lead essentially to the same conclusions, but require  
93 larger melt fractions (up to 1%), which are more difficult to reconcile with the very  
94 small melt fraction ( $\sim 0.1\%$ ) suggested by geochemical studies<sup>4</sup>.

95 The effect of melt on  $V_s$  has been estimated to 7.9% reduction per percent of melt  
96 based on model calculations<sup>17</sup>. Recent experiments<sup>3</sup> are in qualitative agreement  
97 but require a slightly larger  $V_s$  reduction (**Figure S13**). The effect of melt on  
98 attenuation is not well constrained and depends on the mechanisms of  
99 attenuation. We show in Methods that melt may have a large effect on  $Q_s$  at short  
100 period (1 s), but not in the period range of surface waves (50-250s). We therefore  
101 neglect the effect of melt on  $Q_s$  and we model its effect on  $V_s$  based on recent  
102 experiments<sup>3</sup>. **Figure 2** shows that the slowest shear velocities require less than  
103 0.7 % of melt.

104 **Figure 3** displays global maps of melt content at different depths, using the same  
105 colour-coding as in **Figure 2**. The associated mantle temperatures, derived by our  
106 approach are on average slightly above the solidus in oceanic regions, between  
107 100 and 200 km depth (**Figure S14, panel g**). The differences between the maps  
108 of temperature (**Figure S14 panels a-f**) and those of melt content are attributed to  
109 a variable amount of volatile<sup>9</sup>. A higher amount of volatiles above subduction  
110 zones may lower the solidus and favor melting, while a dryer mantle in other  
111 regions may impede melting. The heterogeneities in **Figure 3** display a strong  
112 correlation with surface tectonics. Regions where  $V_s$  is too high (in grey) are  
113 located beneath continents down to 150 km depth. The discrepancy is likely due to  
114 our assumption of a homogeneous and pyrolitic mantle. This assumption is  
115 reasonable in the well-mixed convective mantle. However, beneath cratons,  
116 compositional heterogeneities and depletion of incompatible elements are  
117 contributing to the high seismic velocities. Our observations under cratons, at 150  
118 km depth, are on average 2.4% faster than a pyrolitic mantle which is compatible  
119 with compositional effects<sup>18</sup>.

120 At 100 km depth, melt is required below mid-ocean ridges, some hotspots near the  
121 Atlantic ridge and in the south Pacific Ocean, back-arc basins around the Pacific  
122 Ocean including the eastern margin of Asia and some other active tectonic regions  
123 (Afar, Tibet, West of North-America and Southwest Pacific, including the North  
124 Fidji basin and the Northfolk ridge south of the Vanuatu arc). In these regions, the  
125 amount of melt exceeds 0.3% and can reach 0.7%. Melt is not required beneath  
126 most of the remaining oceanic and continental Phanerozoic lithosphere where  
127 temperature variations alone explain our observations.

128 The depth range 150-200 km corresponds to the oceanic LVZ, where a small  
129 amount of melt (<0.3%) is required over broad regions. The largest amount is

130 beneath hotspots, ridges, and volatile-rich back-arc regions and can extend deeper  
131 than the LVZ, suggesting that these deep regions feed the asthenosphere with  
132 partial melt. For example, partial melt is observed down to 250 km beneath Hawaii  
133 and 300 km beneath the Afar and East-African rift, the hotspots located on the  
134 western part of North-America, the region of the Balleny Islands in Antarctica, the  
135 western Pacific and the Indian ocean near the Ninety-East ridge. A melting  
136 anomaly near 300 km on the eastern part of the Tibetan plateau is also observed.

137 The depth extent and amount of melt beneath oceanic regions are consistent with  
138 local studies. The NoMELT experiment<sup>19</sup> was performed beneath a location  
139 relatively far away from Pacific hotspots (blue star in **Figure 3**) where we too  
140 confirm the absence of melt in what could be a volatile-poor mantle. We observe  
141 melting beneath the East Pacific Rise (green star in **Figure 3**) down to 250 km  
142 depth, in agreement with the MELT experiment<sup>20</sup>. The depth extent of melting  
143 beneath Hawaii down to 250 km is supported by measurements of water  
144 abundance in the deep region that feeds the plume<sup>21</sup>. Our results agree with  
145 observations of significant melting beneath the Philippine Sea plate and the  
146 Western Pacific<sup>2</sup>. Finally, melting down to at least 300 km beneath the Balleny  
147 Islands hotspot is also consistent with previous observations<sup>22</sup>.

148 The origin of partial melt within most of the oceanic LVZ is uncertain. Melting at  
149 mid ocean-ridges can exceed 1% in the depth range 40-80 km<sup>20,23</sup>, but smaller  
150 volume melting may occur down to depths of 150-250 km<sup>23</sup>. After melting at mid-  
151 ocean ridges, a small amount of melt may remain unextractable from the mantle  
152 peridotite<sup>9,24,25</sup>. Our observations suggest that decompression melting with at least  
153 0.3 to 0.7% of melt occurs also beneath some hotspots and back arc basins down to  
154 about 200 km. Most of the melt produced in these regions is extracted and  
155 incorporated to the oceanic crust, but a small amount remains in the oceanic LVZ

156 as it ages. The quantity of melt, if any, decreases close to continents below 0.1%,  
157 where a simple model without melt explains most of our observations. This may be  
158 due to the difficulty of melting the depleted continental lithosphere. However under  
159 tectonically active regions, beneath the Afar and East-African rifts, Tibet, Western  
160 North-America and Transantarctic Mountains, the asthenosphere contains a small  
161 amount of melt (**Figure 3**).

162 The quantity of melt needed to reconcile Vs and Qs under large swaths of the  
163 oceanic LVZ is larger than usual estimates of unextractable melt, which range from  
164 very small values<sup>9</sup> to a maximum of 0.1%<sup>24</sup>. Although melt is likely connected and  
165 able to percolate even at very small porosity<sup>9</sup>, surface tension resists the phase  
166 separation<sup>9,26</sup>. The ability of the melt to rise depends on various parameters,  
167 surface tension, buoyancy, permeability, melt and matrix viscosities, which are all  
168 known with large uncertainties<sup>25</sup>. The melt fractions greater than 0.1% that we  
169 obtain are therefore plausible. This melt concentration is in overall agreement  
170 with the range of estimates derived from electromagnetic studies<sup>27</sup> of the LVZ,  
171 which often propose even larger values (see Methods). A structure made of  
172 magma-rich sills embedded in a meltless mantle might also be mapped by  
173 tomography as an average medium with moderate melt content. This partially  
174 molten layered model has been proposed for the northwest Pacific and Philippine  
175 plates<sup>2</sup>, but it could extend more generally to the entire oceanic LVZ. This layering  
176 would explain both the radial anisotropy observed within the oceanic LVZ<sup>28</sup> and  
177 the sharp velocity and viscosity contrasts at the lithosphere-asthenosphere  
178 boundary<sup>2</sup>.

179 Finally, the amount of melt exhibits a very peculiar relation with the plate  
180 velocity<sup>29</sup> expressed in a no-net-rotation reference frame (**Figure 4**). The melt  
181 fraction in the asthenosphere, abruptly increases by a factor close to 2 when the



182 velocity is larger than 4 cm yr<sup>-1</sup>. This variation is very similar to the observed  
 183 variation of azimuthal anisotropy with present-day plate motion<sup>8</sup>. Our results  
 184 suggest that plate-scale crystal alignment beneath fast-moving plates is associated  
 185 with a greater amount of melt. This requires either that melt facilitates  
 186 deformation<sup>9</sup> or that deformation favours melt retention in the LVZ<sup>2</sup>, or both. In  
 187 any case, the small amount of melt observed beneath large swaths of the oceanic  
 188 LVZ is likely to significantly decrease viscosity (by one to two orders of  
 189 magnitude<sup>9</sup>, see Methods) and to play a significant role in the decoupling of  
 190 tectonic plates from the mantle.

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270

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 272 paper. E.D. wrote the codes for the interpretation of the seismic models and wrote  
 273 the drafts of the manuscript. E.D. wrote the tomography code for  $V_s$ , Y.R. adapted  
 274 this code for  $Q_s$ . T.B. contributed to the design of the figures and to the writing of  
 275 the manuscript. Y.R. developed preliminary codes for interpreting the seismic  
 276 models, contributed to all mineralogical aspects and to the writing of the  
 277 manuscript. S.D realized the tests of the effect of composition and contributed to  
 278 the writing of the revised manuscript.

279

280 **Fig. 1: Shear velocity and attenuation in the upper mantle.** Panels a-c-e-g:  
 281 perturbations in shear wave velocities from DR2020s shown in percent with respect to a  
 282 mean value  $V_{\text{ref}}$  given in  $\text{km s}^{-1}$  above the colour scale. Panels b-d-f-h: maps of our  $Q_s$   
 283 tomographic model QsADR17 at different depths in the upper mantle.  $Q_s$  is plotted with a  
 284 logarithmic scale. Its geometric average is given above the colour scale. Hotspot locations  
 285 are shown with black triangles.

286

287

288

289 **Fig. 2: Scatter plot depicting the observed shear attenuation as a function of shear**  
 290 **velocity compared with theoretical predictions.** The velocities (from DR2020s) and  
 291 attenuation values (from QsADR17) are plotted at 100 km (panel a), 150 km (panel b) and  
 292 200 km depth (panel c). Each dot corresponds to a geographical location. The dark blue  
 293 and light blue curves are the theoretical predictions assuming a pyrolitic composition<sup>34</sup> in  
 294 the absence of melt using the anelasticity models of Takei<sup>4</sup> and Jackson et al.<sup>13</sup>,  
 295 respectively. The upper colour scale indicates the amount of melt in percent required to  
 296 explain our observations using Takei<sup>4</sup>'s model. Adding melt enables to lower the predicted  
 297 velocity without changing  $Q_s$ . The lower grey scale indicates the misfit in percent between  
 298 theory and observations, in regions where  $V_s$  is too high and cannot be reconciled with  
 299 model predictions assuming a pyrolitic mantle.

300

301

302 **Fig. 3: Melt content at different depths in the upper mantle.** This melt content is  
 303 derived from the joint interpretation of QsADR17 and DR2020s. The colour coding is  
 304 identical to Figure 2. Melt content in percent is indicated with warm colour from ivory  
 305 (0% melt) to brown (0.4 to 0.7% melt). The grey scale indicates the misfit in percent  
 306 between theory and observations, in regions where  $V_s$  is too high compared with  
 307 predictions. Hotspots locations are indicated with black triangles. The blue and green stars  
 308 indicate the location of the NoMELT and MELT experiments, respectively.

309

310

311

312 **Fig. 4: Percentage of melt at different depths as a function of absolute plate velocity.**  
313 (The velocities<sup>29</sup> are expressed in a no-net-rotation reference frame. The percentage of  
314 melt is averaged for all geographical points with similar plate velocities, using a sliding  
315 window of  $\pm 2$  cm yr<sup>-1</sup>. The amount of melt increases significantly in the asthenosphere  
316 (100-200 km) for plate velocities between 3 and 5 cm yr<sup>-1</sup>. This result links with previous  
317 observations that only plates moving faster than 4 cm yr<sup>-1</sup> can organize the flow at large  
318 scale in the underlying asthenosphere<sup>8</sup>, suggesting that melt reduces viscosity<sup>9</sup> and  
319 facilitates large-scale crystal alignment.  
320

## 321 **Methods**

322 **Qs and Vs tomographic models.** Figure 1 presents maps of DR2020s, a new  
323 global Vs model and QsADR17 a recent global Qs model<sup>30</sup>. Both models are built  
324 from the same massive Rayleigh wave measurements<sup>31</sup>, and are obtained from a  
325 similar tomographic procedure. The first step is an automated waveform inversion  
326 approach which was applied to approximately 375,000 Rayleigh seismograms<sup>31</sup>.  
327 From a single surface wave seismogram, the waveform inversion extracts  
328 simultaneously a path-average depth-dependent shear velocity profile, Vs, and  
329 quality factor, Qs. By jointly interpreting the amplitude and phase of each  
330 waveform, we ensure that the interplay between Vs and Qs is accounted for, and  
331 that the shear quality factor and velocity profiles are constrained within the same  
332 period range. The waveform analysis is performed in the period range 50-250 s  
333 and accounts for the fundamental and up to the 5th higher mode of Rayleigh  
334 waves, thus ensuring a good depth resolution for Vs and Qs from 50 km depth  
335 down to the transition zone. It is a non-linear iterative process, which also  
336 produces frequency-dependent phase velocity and attenuation curves compatible  
337 with the recorded waveform. The effect of physical dispersion due to attenuation is  
338 accounted for in the modelling.

339 The second step is a regionalization of the 1D path-average models. DR2020s is  
340 obtained from the regionalization at each depth of the path-average Vs models  
341 using a continuous regionalization approach<sup>31</sup>. This tomographic inversion yields  
342 3D absolute velocities. QADR17, a global model of Rayleigh-wave attenuation was  
343 obtained after adapting the same regionalization approach to our dataset of  
344 Rayleigh wave attenuation curves, parameterized as  $\ln(Q)$ <sup>32</sup>. The regionalisation of  
345 the path-average attenuation curves accounts for frequency-dependent effects like  
346 focussing-defocussing, which can have important effects on the amplitude of

347 Rayleigh waves. The logarithmic parameterization brings the distribution of the  
348 quality factor dataset close to a Gaussian, allows the large variations of  $Q$   
349 documented by local seismic studies and guarantees to avoid negative attenuation  
350 values in the inverted model. The horizontal smoothing in DR2020s and QADR17  
351 is determined by a Gaussian a priori covariance function controlled by an angular  
352 correlation length. Adenis et al.<sup>32</sup> chose a conservative value of  $10^\circ$  (meaning that  
353 the  $Q$  model is resolved accurately up to spherical harmonic 12). Debayle and  
354 Ricard<sup>31</sup> used a shorter correlation length of  $3.6^\circ$  for their  $V_s$  model DR2012. We  
355 re-inverted their dataset using a correlation length of  $10^\circ$  in order to obtain a  $S_v$ -  
356 wave tomographic model at the same horizontal resolution and vertical smoothing  
357 than QsADR17. To minimize biases due to un-modelled radial anisotropy, we  
358 computed the isotropic Voigt average of our  $S_v$  model and the  $S_h$  model obtained  
359 by adding the radial anisotropy of PREM to our  $S_v$  model. The resulting isotropic  
360 model DR2020s is plotted in **Figure 1**. Using other 1D or 3D models of radial  
361 anisotropy does not affect our conclusions, as discussed below. QsADR17 was  
362 obtained from the inversion at depth of QADR17<sup>30</sup>, using the same vertical  
363 smoothing as for DR2020s. DR2020s and QsADR17 are therefore consistent since  
364 they are derived from the simultaneous inversion of the same waveforms and  
365 inverted using the same regionalization approach with the same Gaussian filtering.

366 **Prediction of  $Q_s$  and  $V_s$  using a temperature-dependent model**<sup>4</sup>. The dark blue  
367 curve in **Figure 2** is obtained by predicting  $V_s$  and  $Q_s$  at each depth for a range of  
368 temperatures, using a temperature-dependent anelasticity model<sup>4</sup> and assuming a  
369 pyrolitic composition<sup>16</sup>. This model parameterizes the relaxation spectrum with a  
370 small number of variables determined from experimental data. It accounts for a  
371 monotonic background spectrum plus a broad temperature-dependent absorption  
372 peak in the seismic frequency band, whose amplitude and width increase below  
373 the solidus. This produces a significant enhancement of polycrystal anelasticity

374 before melting. This pre-melting effect induces large Vs reductions under the  
 375 solidus. We choose the solidus according to Hirschman<sup>33</sup>. We use the model  
 376 parameters given in Tables 1 and 2 of Takei<sup>4</sup>, except for the unrelaxed shear  
 377 modulus  $\mu_U$ . We estimate  $\mu_U$  for a pyrolitic mantle using the mineralogic phase  
 378 diagram computed by the Perple X software<sup>34</sup>. In **Figures S15-S16** we show the  
 379 results obtained with the shear modulus of Takei<sup>4</sup>, which reduces the amount of  
 380 melt in oceanic regions but fails to explain our  $\ln(Q_s)$ -Vs observations beneath old  
 381 oceanic basins at 100 km depth. It also significantly increases the misfit beneath  
 382 continents.

383 **Prediction of  $Q_s$  and Vs from the experimental results of Jackson et al.**<sup>13</sup>. The  
 384 light blue curve in **Figure 2** is obtained by a two-step process. First, we use  
 385 experimental results on melt-free polycrystalline olivine to predict a theoretical  
 386 quality factor  $Q_t$ . Computations are performed in the temperature range 800-1800  
 387 K and the pressure range 1.49-12.99 GPa, corresponding to the depth range 50-  
 388 390 km. We use the following relation<sup>13</sup> :

$$389 \quad Q_t^{-1}(T_0, T, P, d) = A (T_0 d^{-1} \exp(-(E + PV)/RT))^{\alpha} \quad \text{Eq. A.1}$$

390 In this formulation,  $Q_t^{-1}$  at period  $T_0$  depends on the temperature  $T$ , the pressure  $P$   
 391 and to a lesser extent the grain size  $d$ .  $R$  is the gas constant,  $E$  is the activation  
 392 energy and  $V$  is the activation volume. **Table S1** summarizes the values of the  
 393 different parameters of Eq. A.1. In the range of values compatible with  
 394 experiments<sup>13</sup>, we choose  $T_0=100$  s, the average period of our long period Rayleigh  
 395 waves and  $d=0.01$ m. The theoretical curves corresponding to Eq. A1 are shown in  
 396 **Figure S9**, with the effect of changing grain size.

397 Second, we use Perple X<sup>34</sup> to estimate an isotropic Vs for a pyrolitic model<sup>16</sup>. We  
 398 compute Vs for the same temperatures and pressures as  $Q_t$ . Perple\_X produces

399 unrelaxed S-wave velocities at infinite frequency, above the absorption band of  
 400 seismic attenuation, while our long period (>50 s) seismic surface waves are likely  
 401 to see relaxed velocities. We therefore correct for the effect of anelasticity on  $V_s^{35}$  :

$$402 \quad V_s^t = V_{s\infty} \left( 1 - \frac{Q_t^{-1}(T_0, T, P, d)}{2 \tan(\pi\alpha/2)} \right) \quad \text{Eq. A.2}$$

403 where  $V_{s\infty}$  is the unrelaxed velocity and  $V_s^t$  is the velocity corrected from the effect  
 404 of attenuation. According to Eq. A2, the stronger the attenuation, the stronger the  
 405 velocity reduction is. The light blue curve in **Figure 2** is finally obtained by plotting  
 406 at a given depth the obtained  $\ln(Q_t)$  as a function of  $V_s^t$  for each temperature.

407 **Comparison between our observations and the theoretical relations.** At each  
 408 geographical location and depth, we extract the observed value of  $Q_s$  in QsADR17  
 409 and  $V_s$  in DR2020s and we compare our observations with the chosen theoretical  
 410 relation. We consider that a theoretical curve is compatible with a given  $V_s$ - $Q_s$   
 411 observation if it falls within  $\pm 1\%$  and  $\pm 10\%$  of the  $V_s$  and  $\ln(Q_s)$  observations,  
 412 respectively (**Figure 2**). These errors are on the conservative side of recent  
 413 estimates<sup>36,37</sup>. If the theoretical relation cannot explain our observations, two  
 414 situations can arise: 1) the observed  $V_s$  is lower than the theoretical  $V_s^t$ . In this case  
 415 melt can be added to reduce  $V_s^t$ . This is done assuming melt has no effect on  $Q_s$ ; 2)  
 416 the observed  $V_s$  is higher than  $V_s^t$ . This is mostly the case under continents and can  
 417 be explained by the depletion of the pyrolytic mantle increasing the theoretical  
 418  $V_s^{t18}$ .

419 **Effect of grain size.** Grain sizes in the shallow upper mantle are likely 1-20 mm as  
 420 observed in upper mantle xenoliths, harzburgite and dunite bodies<sup>13</sup>. The effect of  
 421 grain size on Eq. A1 is shown in **Figure S9**. The blue areas around the theoretical  
 422 curves cover the influence of grain size from 1 mm (bottom dot curve) to an upper  
 423 bound of 100 mm (upper dashed curve). For a given value of  $Q_s$ , increasing the



424 grain size increases the temperature and therefore decreases the velocity. As melt  
425 is required for explaining slow velocities, increasing grain size will decrease the  
426 amount of melt required to explain our observations. Using the equation A1  
427 proposed by Jackson, we show in **Figure S12** melt distributions obtained for  
428  $d=0.01$  m and in **Figure S17** those for the extreme value  $d=0.1$  m, which minimizes  
429 the amount of melt required in the LVZ. Maps in **Figure S17** are similar to those of  
430 **Figure S12** with a smaller amount of melt. However, in both cases, amounts of  
431 melt are larger than for our preferred model using the equations of Takei<sup>4</sup> (**Figure**  
432 **3**).

433 **Amounts of melt.** Although the amount of melt present in the LVZ depends on the  
434 choice of the anelasticity model and on the effect of melt on  $V_s$ , we show in **Figures**  
435 **S12, S16, S17, S18** that different choices of parameters yields similar maps and  
436 that melt fractions away from source regions always stay below 1%, in agreement  
437 with the partially molten asthenosphere model<sup>2</sup>. We note however that our  
438 preferred anelasticity model<sup>4</sup> generally explains our  $V_s$  and  $\ln(Q_s)$  observations  
439 without the need of partial melt above and below the asthenosphere (**Figure 3**).  
440 This is therefore a conservative choice, able to reconcile our  $V_s$  and  $\ln(Q_s)$   
441 observations in most of the upper mantle, while minimizing the amount of melt  
442 required in the LVZ.

443 **Effect of radial anisotropy.** We used the anisotropic parameter  $\xi=(V_{sh}/V_{sv})^2$  of  
444 PREM to convert our  $S_v$  observations into an isotropic model, DR2020s. We  
445 checked that 3D models of  $\xi$  obtained from anisotropic models like SEMUM2<sup>38</sup> or  
446 S362ANI<sup>39</sup> would not affect our main conclusions. As an example, **Figure S1** is  
447 similar to **Figure 3** but based on our  $S_v$  observations corrected with the  
448 anisotropic  $\xi$  parameters of SEMUM2 (**Figure S2**). Only minor changes are  
449 observed (for example in the Pacific or Indian oceans at 150 and 200 km), but

450 none that would change the interpretations of this paper. The lateral variations of  $\xi$   
 451 in SEMUM2 (**Figure S2**) are too small to affect our results. To confirm this point,  
 452 we computed the anisotropy  $\xi$  that would be needed to account for the discrepancy  
 453 between observed Qs and Vs, without invoking mantle depletion or partial melt  
 454 (**Figure S3**). Comparison of this figure with **Figure S2** shows that neither the  
 455 patterns nor the amplitude would fit (a 3 to 4 times larger anisotropy than in  
 456 SEMUM2 would be needed to explain our observations). The anisotropy therefore  
 457 does not affect our conclusions.

458 **Elastically accommodated grain boundary sliding (EAGBS) hypothesis.**

459 Anelastic relaxation caused by EAGBS can produce a sharp velocity reduction and  
 460 may explain the LAB and the mid-lithosphere discontinuity observed beneath  
 461 continents<sup>40</sup>. We tested the EAGBS hypothesis as an alternative explanation to  
 462 partial melting for reconciling Vs and Qs in the LVZ. EAGBS is characterized by a  
 463 sharp attenuation peak at a characteristic frequency  $\omega_{EAGBS}$ , followed by a diffused  
 464 absorption band at lower frequencies. We first compute the characteristic  
 465 frequency for the transition between unrelaxed and relaxed shear moduli. It is  
 466 given by Karato et al.,<sup>40</sup> :

$$467 \quad \omega_{EAGBS} = A \cdot d^{-1} \cdot \exp\left(-\frac{E^* + PV^*}{R.T}\right) \cdot \left[1 + \left(\frac{C_w}{C_{w0}}\right)^r\right] \quad \text{Eq. A3}$$

468 where  $C_w$  is the water content in wt%,  $A=2.3 \times 10^{-14} \text{ m s}^{-1}$ ,  $C_{w0}=10^{-4} \text{ wt}\%$ ,  $r=1$ ,  
 469  $E^*=350 \text{ kJ mol}^{-1}$ ,  $V^*=10^{-6} \text{ m}^3$ ,  $d=5 \times 10^{-3} \text{ m}$ . Following Cobden et al.<sup>7</sup> we assume a dry  
 470 mantle with  $C_w=10^{-6} \text{ wt}\%$ .

471 Surface waves analysed at frequencies lower than  $\omega_{EAGBS}$  should see relaxed  
 472 velocities within the absorption band (see Fig. SI-2 of Karato et al.<sup>40</sup>), providing  
 473 depth and temperature are within the activation ranges of EAGBS. EAGBS is  
 474 activated between 60 and 160 km depth in the temperature range 900-1,350 K. We

475 consider that EAGBS may affect the interpretations of our long period (>40 s) Qs  
 476 and Vs maps when  $\omega_{EAGBS} > 0.025$  Hz and, from figure 4 of Karato et al.<sup>40</sup>, in the  
 477 following temperature ranges for the depths of our tomographic inversion:

- 478 • 70 km: 930-1,140 K
- 479 • 90 km: 980-1,180 K
- 480 • 100 km: 1,010-1,220 K
- 481 • 125 km: 1,060-1,270 K
- 482 • 150 km: 1,130-1,330 K

483 Regions where EAGBS can apply are displayed in green in **Figure S4**. These  
 484 regions correspond mostly to continental areas. Assuming that EAGBS correction  
 485 applies, Karato et al.<sup>40</sup> list three alternative formulations for the velocity reduction.  
 486 The Ghahremani<sup>41</sup> equation, which produces the smallest velocity reduction is :

$$487 \quad V_s^{eagbs} = V_s^t \cdot \sqrt{\frac{0.86-0.83\nu}{1-\nu} \cdot \frac{1+\nu}{1.14+0.83\nu}} \quad \text{Eq. A.4}$$

488 where  $\nu$  is the Poisson's ratio ( $\sim 0.3$  for upper mantle minerals). Regions where  
 489 this can reconcile our Vs and Qs observations are displayed in red in **Figure S4**.  
 490 The two other formulations give too large velocity reduction. **Figure S4** shows that  
 491 EAGBS fails at reconciling our long period observations, excepted in a few localized  
 492 regions (in red). This does not preclude EAGBS to play a significant role in  
 493 explaining the sharp velocity reduction associated with the LAB, which cannot be  
 494 resolved by long period surface waves. However, the strong velocity reduction  
 495 observed within the LVZ requires another mechanism.

496 **Effect of composition.** Compositional heterogeneities exist in the upper mantle,  
 497 likely in the domain limited by pyrolite, harzburgite and pyroxenite<sup>42</sup>. Here, we  
 498 check the effect of changing the composition from pyrolite to these end-members  
 499 using Perple X<sup>34</sup> to compute the velocities. Pyroxenites have variable bulk

500 compositions<sup>43</sup> and we consider both a silica-deficient composition (see, Stixrude  
 501 and Lithgow-Bertelloni<sup>44</sup>, Table 1) and a silica-excess composition similar to that  
 502 of a subducted basalt (see, Xu et al.,<sup>16</sup>, Table 1). A silica-excess pyroxenite would  
 503 only be faster than pyrolite after eclogitization, below 150 km depth (**Figure S5**,  
 504 orange curve). At 100 km depth, a mantle made entirely of silica-excess pyroxenite  
 505 would be too slow to explain the observed velocities (**Figure S6**). At larger depth,  
 506 after eclogitization, it would lead to large melt fractions (**Figure S6**). The silica-  
 507 deficient pyroxenite and the harzburgite are characterized at asthenospheric  
 508 conditions by much faster or comparable velocities than pyrolite (**Figure S5**) and  
 509 hence cannot explain the low velocities that we attribute to the presence of melt.  
 510 Harzburgite would give melt fractions and distribution similar to pyrolite (**Figure**  
 511 **S7**), while the silicate-deficient pyroxenite would lead to higher melt fractions (up  
 512 to 1.7%) and a ubiquitous melt presence below 150 km depth (**Figure S8**).  
 513 Pyroxenite, whether enriched or depleted in silica, should only be present in a  
 514 small proportion (~5%)<sup>42</sup> that hardly affects the seismic velocities of the  
 515 peridotitic mantle. At any rate, large-scale chemical heterogeneities (> 1000 km)  
 516 are not likely to exist in the LVZ, where viscosity is small (below 10<sup>20</sup> Pa s) and the  
 517 convective mantle should be well stirred and mixed. Therefore, our observations  
 518 cannot be explained by reasonable variations of compositions in the LVZ.

519 **Possible effect of water.** Although, it has been recently shown that in olivine,  
 520 wave-speed and attenuation are insensitive to water<sup>10</sup>, some previous studies had  
 521 suggested an effect of water quantified following the equation<sup>45</sup>:

$$522 \quad Q_t^{-1}(T_0, T, P, C_{OH}, d) = A \left( T_0 d^{-1} \left( \frac{C_{OH}}{C_{OH(ref)}} \right)^r \exp(-(E + PV)/RT) \right)^\alpha \quad \text{Eq. A.5}$$

523 where  $C_{OH}$  is the water concentration,  $r$  is a dimensionless constant and all other  
 524 parameters are as in Eq. A1. The average H<sub>2</sub>O content of Earth's upper mantle is  
 525 estimated in the range 50-200 ppm<sup>46</sup>. We test a value close to the upper bound

526 ( $C_{OH}=1000H/10^6$  Si, corresponding to  $125 \pm 75$  ppm by weight of water) with  $r=1$ , a  
 527 standard value for moderate water content and  $C_{OH(ref)}=50H/10^6$  SI<sup>7</sup>. The  
 528 corresponding theoretical curves are shown in blue on **Figure S9**. At fixed  
 529 temperature, water increases attenuation. However to keep Q constant, an  
 530 increase of water content must be balanced by a decrease of temperature and  
 531 therefore by a higher expected velocity  $V_s^t$ . Reconciling this higher  $V_s^t$  with our  
 532 observations requires larger amounts of melt (**Figure S10**). Therefore, if there were  
 533 an effect of water on Vs and Qs, it would increase the required amount of melt.

534 **Effect of melt on attenuation.** A small amount of melt might have a large effect on  
 535 attenuation<sup>47</sup> in the case of grain boundary sliding, where a broad dissipation peak  
 536 is observed in the seismic frequency band. Body wave studies have reported low  
 537 Qs values at short period (1s) with  $Q_s \leq 25$  beneath the Juan de Fuca and Gorda  
 538 ridges<sup>48</sup> and  $25 \leq Q_s \leq 80$  in the back-arc mantle of Central America, the Marianas  
 539 and the Lau Basin<sup>49</sup>. A recent experimental study suggests that 0.2% of melt  
 540 produces  $Q_s=48$ <sup>3</sup>. However such a large effect of melt on Qs is difficult to reconcile  
 541 with long period (>30 s) surface waves observations of  $Q_s=80$  beneath the fast  
 542 spreading southern East Pacific Rise<sup>20</sup> where up to 1-2% of melt is expected<sup>50</sup>. If  
 543 the attenuation mechanism is melt squirt, then the dissipation peak may lie outside  
 544 the seismic frequency band and cause little attenuation. Model calculations<sup>51</sup>, long-  
 545 period surface waves<sup>20</sup> and experiments<sup>12</sup> favour such interpretation. This is also  
 546 supported by our long period (100 s) global surface wave tomographic models,  
 547 which suggest that strong Vs and Qs reductions are not necessarily correlated<sup>30,32</sup>.  
 548 For these reasons, we assume that in the period range of surface waves (50-250s),  
 549 melt does not significantly affect attenuation and we neglect its effect on Qs.

550 **Temperature in the upper mantle.** Our Qs and Vs seismic models are interpreted  
 551 using mineral physics results<sup>4,13</sup>. These experimental results are valid at temperatures

552 larger than 900°C or 1000°C, which are reached at the base of the lithosphere and in the  
553 asthenosphere. The predictions of our model can be checked with respect to the well-  
554 known thermal behaviour of the cooling oceanic lithosphere<sup>52,53</sup> although this is  
555 certainly pushing the model outside its applicability range. **Figure S19 (panel a)**,  
556 displays the temperature variations beneath the Pacific as a function of sea-floor age  
557 predicted from our seismic models using experimental results<sup>4</sup>. It demonstrates that  
558 even at depths shallower than 100 km, we retrieve the well-known age-dependence of  
559 temperature in oceanic regions (we also pick up the cold signal of the West Pacific  
560 subductions at old ages). However, our temperature variations between the ridge and an  
561 old lithosphere (~100°) are lower than the predictions of the plate-cooling model (250°  
562 at 75 km depth, **panel b**). **Panel c** displays the quality factors that would be deduced  
563 from the plate-cooling model using the experimental results<sup>4</sup> used in this paper. In the  
564 oceanic lithosphere,  $Q_s$  values much greater than 2000 ( $\ln(Q_s) > 7.6$ ) would be  
565 predicted, that cannot be retrieved by long period seismology. The limitations of  
566 experimental data at low temperatures together with the inability of surface wave  
567 seismology to quantify precisely negligible attenuations, explain why we do not  
568 interpret results at depths shallower than 100 km. Below the oceanic plates, **Figure S14**,  
569 **panel g** shows that our predicted temperatures beneath oceans are perfectly compatible  
570 with geodynamic and petrologic expectations. In agreement with our findings, under  
571 oceans, the average 1D temperature profile is above the solidus<sup>4</sup> and the adiabat<sup>54</sup>  
572 in the depth range 100-200 km, where we predict partial melting. This overshoot  
573 of the temperature, above the adiabat, is indeed found in all numerical simulations  
574 of mantle convection<sup>55</sup>. Under continents, the temperature appears to be below the  
575 solidus<sup>4</sup> and the adiabat<sup>54</sup>. It reaches the adiabat around 250 km depth (**Figure**  
576 **S14, panel g**).

577 **Compatibility of our results with electrical conductivity.** Our results are  
578 consistent with the interpretation of a number of recent studies of electrical

579 conductivity. In the depth range 100-150 km, the joint interpretation of electrical  
580 and seismic data<sup>56</sup> requires 0.3-2.5% of melt beneath the mid-Atlantic ridge, 1% or  
581 less melt beneath Hawaiï and less than 5% of melt beneath the East-pacific rise, in  
582 good agreement with **Figure 3**. We also confirm the absence of melt in the region  
583 of the NoMELT experiment<sup>19</sup>, in agreement with conductivity data in this region<sup>25</sup>.  
584 Previous experimental results suggest that between 0.3 and 2% of hydrous  
585 basaltic melt can account for the observed electrical conductivity in the LVZ<sup>27</sup>. A  
586 more recent study<sup>57</sup> has refined these results by simultaneously measuring wave  
587 velocity and electrical conductivity on a simplified partial melt analogue. They  
588 conclude that the low velocity zone away from spreading ridges can be explained  
589 by 0.3-0.8% volatile-bearing melt, the upper bound of our observations.

590 **Implications for viscosity.** The variation of viscosity as a function of melt content  
591 occurs in two steps<sup>9</sup>. The onset of melting brings already a significant effect on  
592 viscosity when a connected network of melt tubules is formed. The viscosity  $\eta$  is  
593 expected to decrease by one or two orders of magnitude before the melt fraction  $\phi$   
594 reaches 0.1%. For larger melt content,  $\eta$  decreases further<sup>9</sup> with  $d\ln(\eta)/d\phi = -26$   
595 but this effect is minor for the low melt content that we observe. Under oceanic  
596 plates where  $\phi$  is around 0.3%,  $\eta$  should be one or two orders of magnitude lower  
597 than under continents.

598 **Data availability:** The dataset generated during this study (3D Vs, Qs models and  
599 melt fraction models) is available as IRIS data products at  
600 <https://doi.org/10.17611/dp/emc.2020.dbrdnature.1>.

601 **Code availability:** Numerical modelling codes related with this paper can be  
602 downloaded from <https://doi.org/10.17611/dp/emc.2020.dbrdnature.1>. Requests  
603 about the numerical modelling codes associated with this paper should be sent to

604 Eric.Debayle@ens-lyon.fr. Most figures were created with open software GMT

605 4.5.13.

606 **Competing interest:** The authors declare no competing interests.

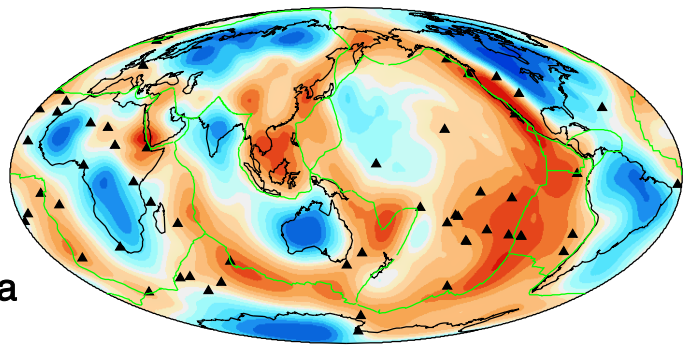
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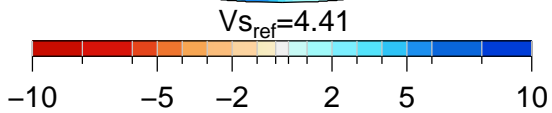


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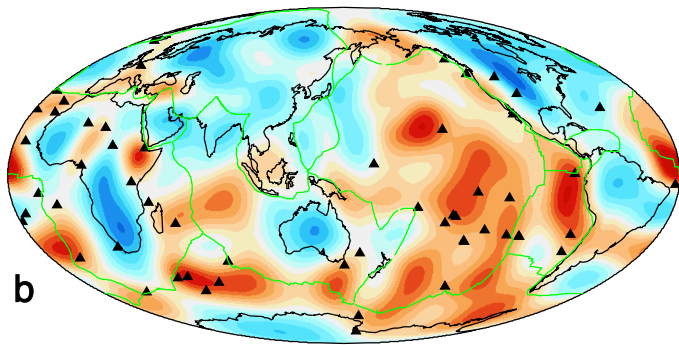
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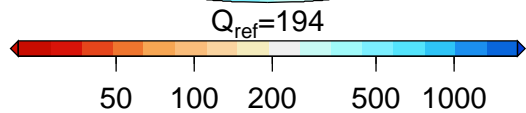
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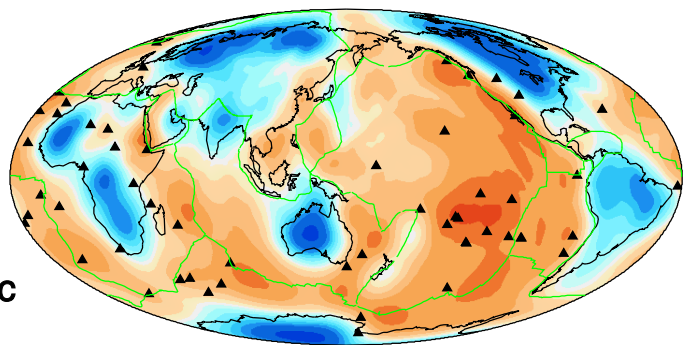
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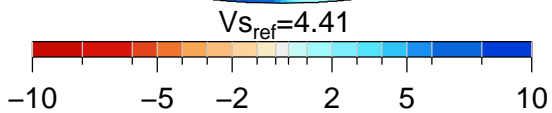
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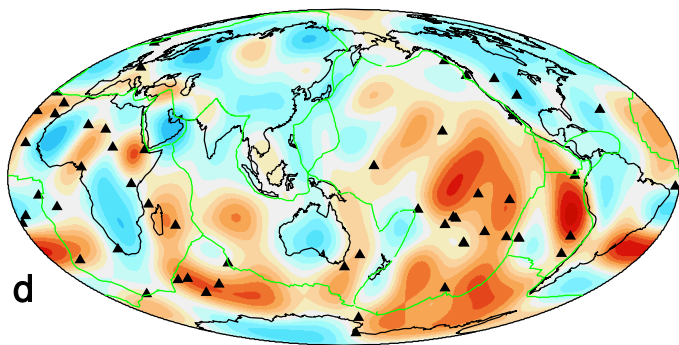
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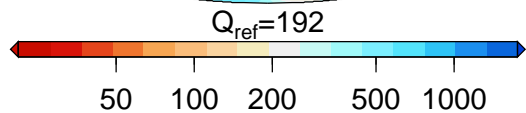
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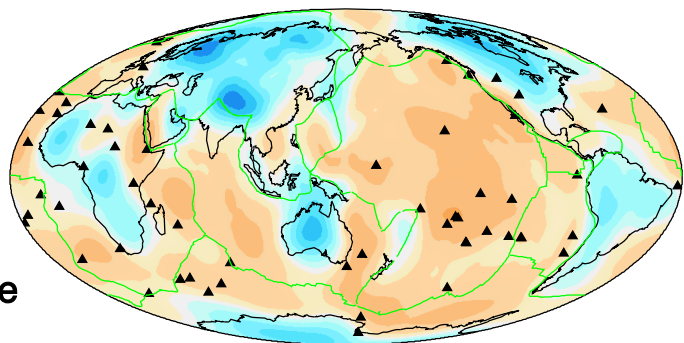
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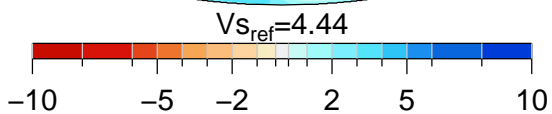
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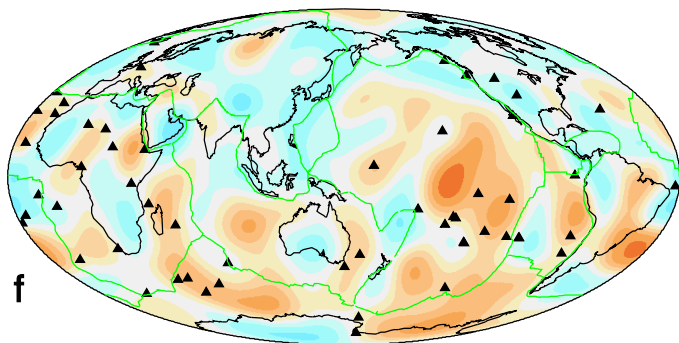
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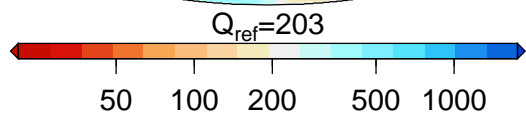
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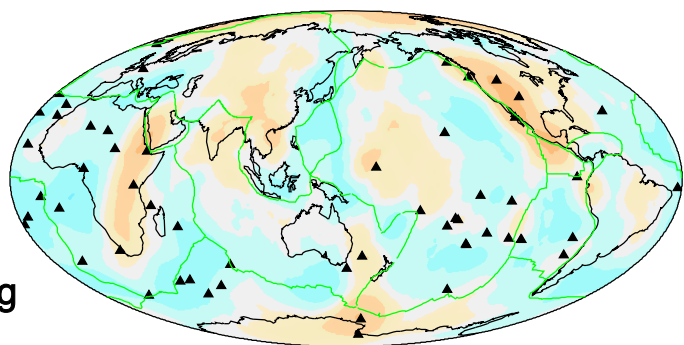
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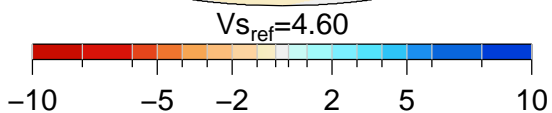
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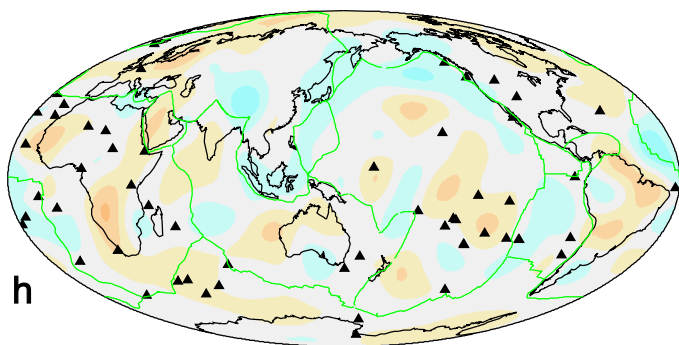
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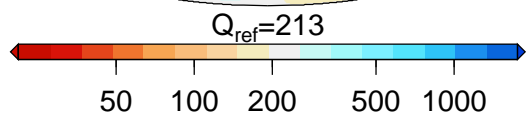
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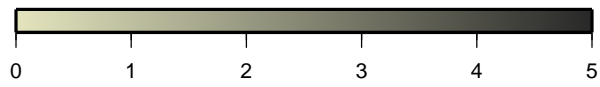
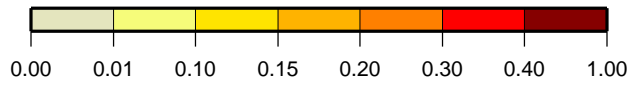
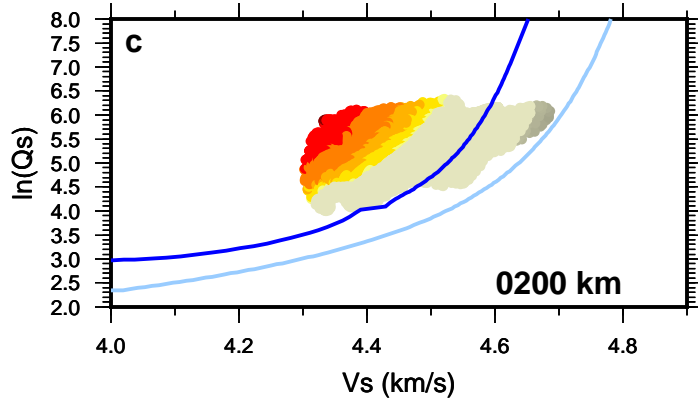
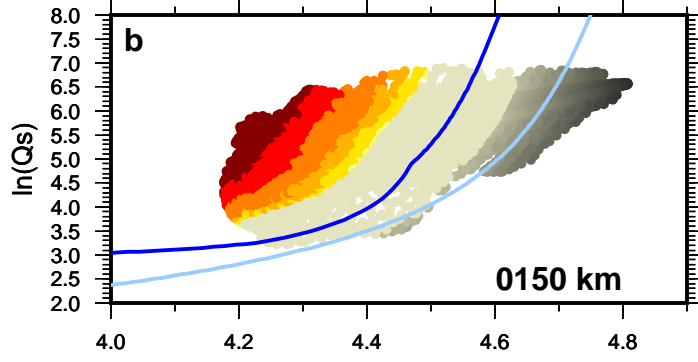
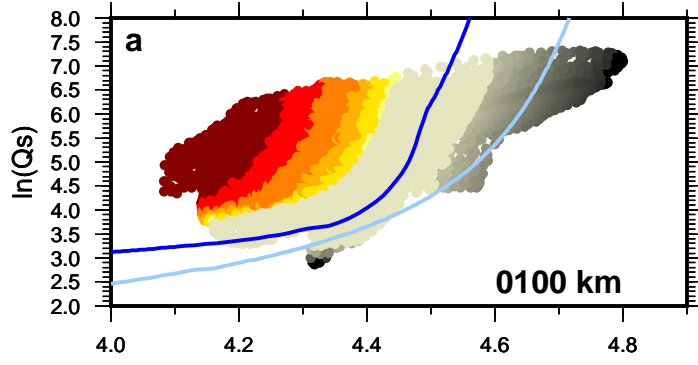


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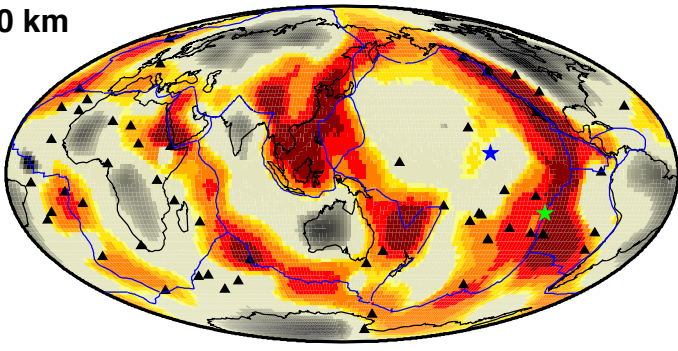


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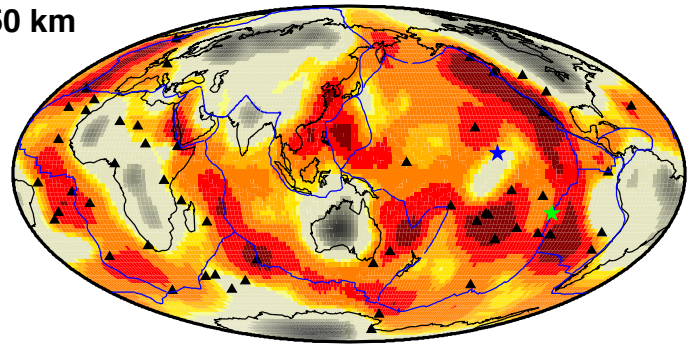




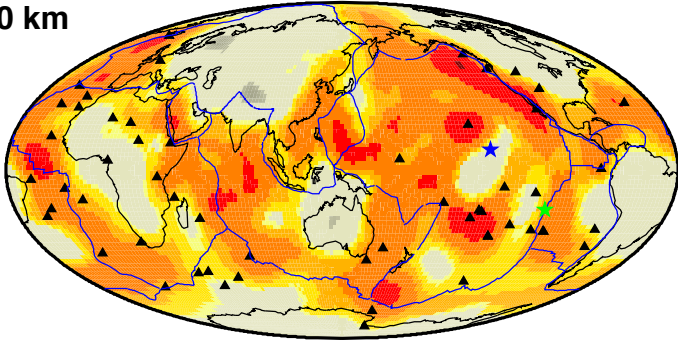
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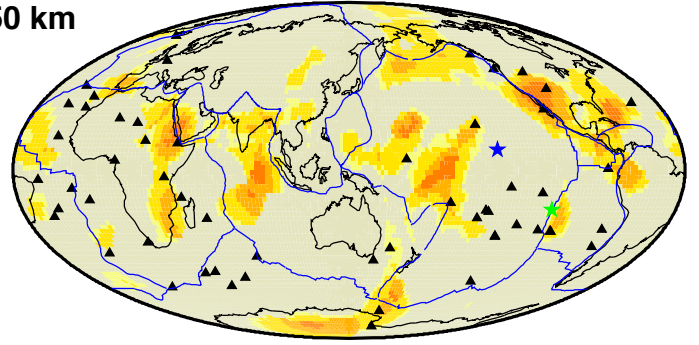
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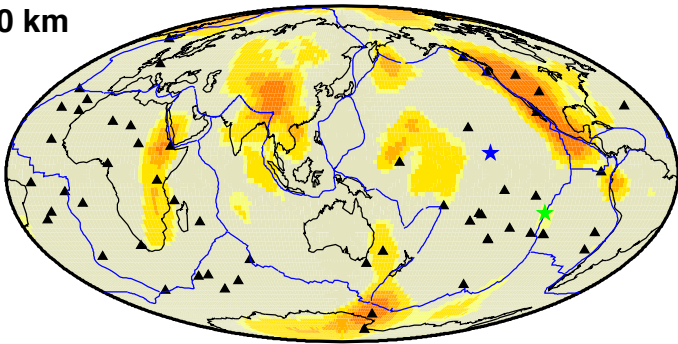
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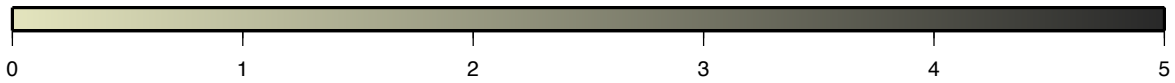
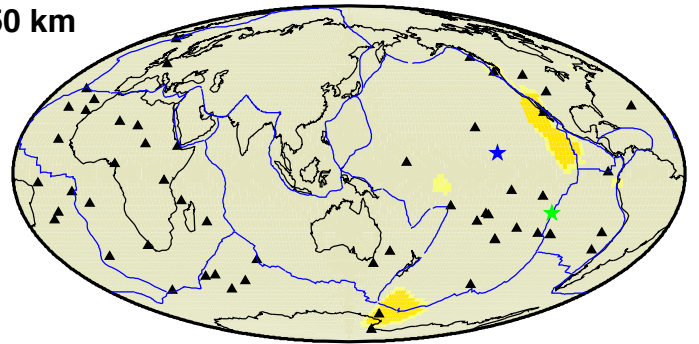
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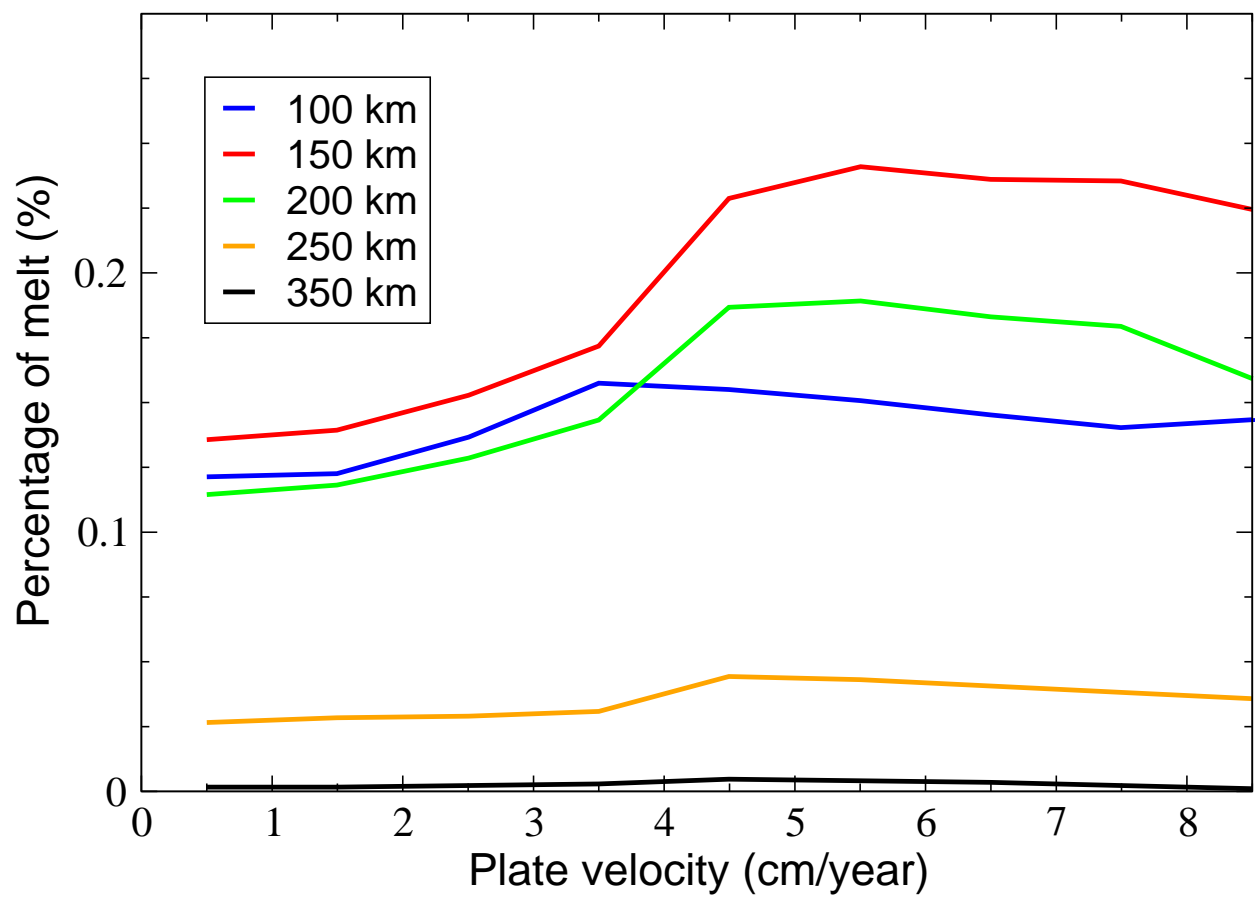


0300 km



0350 km





<i>Symbol</i>	<i>Description</i>	<i>Value</i>	<i>Units</i>
$\alpha$	exponent	0.26	
A	prefactor	$7.5 \times 10^{-2}$	$\text{s}^{-\alpha} \mu\text{m}^{\alpha}$
d	grain size	1-100	mm
$T_0$	period	100	s
E	activation energy	424	kJ/mol
V	activation volume	$6 \times 10^{-6}$	$\text{m}^3/\text{mol}$

Table S1: Reference parameters for Eq.A.1 after Jackson et al.<sup>1</sup>

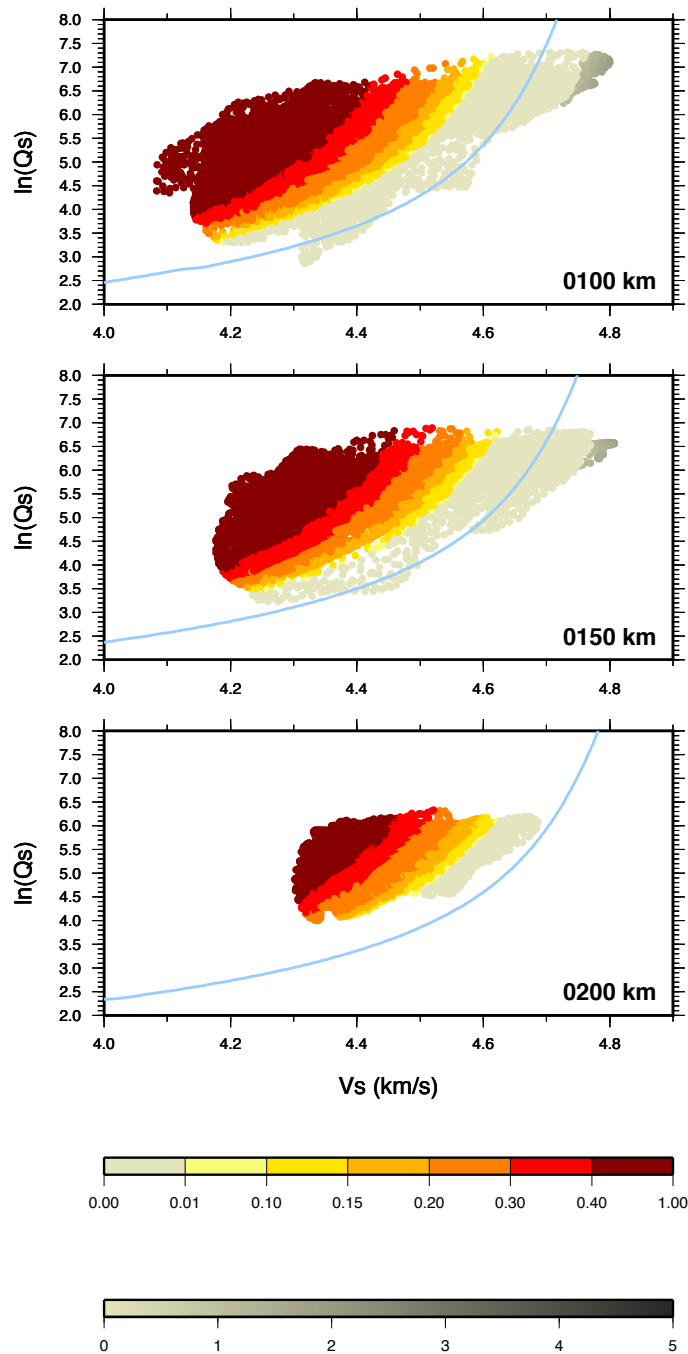


Fig. S1: same as Figure 2 but the color scales indicate the departure from the model of Jackson et al.<sup>1</sup> (light blue curve), for a grain size of 10 mm. The upper color scale indicates the amount of melt in percent required to explain our observations (ivory color from 0 to 0.01% of melt underlines data for which the model can reconcile our  $Q_s$  and  $V_s$  observations). The lower grey scale indicates the misfit in percent between theory and observations, in regions where  $V_s$  is too high and cannot be reconciled with model predictions.

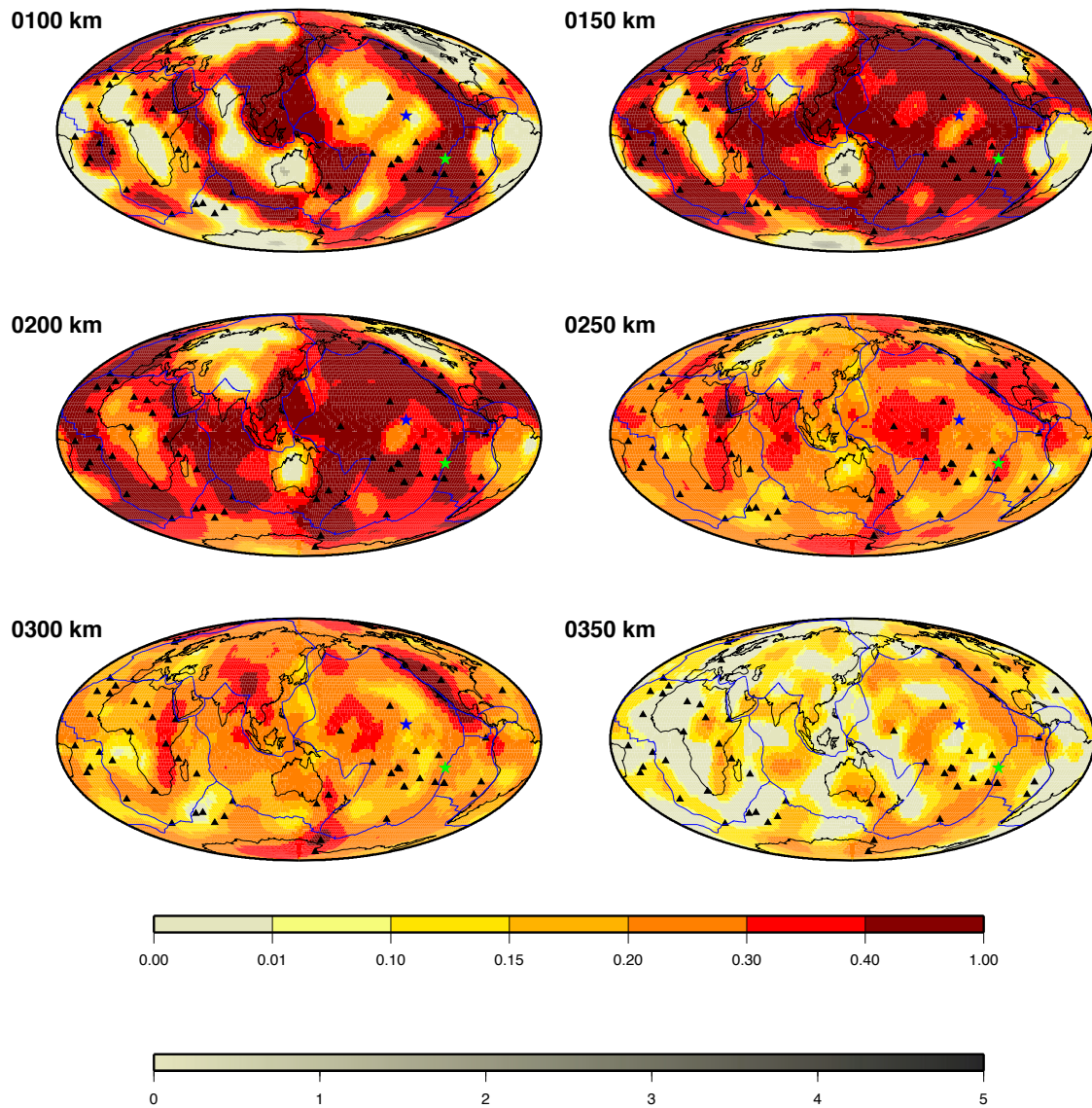


Fig. S2: same as Figure 3 but color scales indicate the departure from the model of Jackson et al.<sup>1</sup>, for a grain size of 10 mm. Melt content in percent is indicated with warm colors (upper color scale) from ivory (0% melt) to brown (0.4 to 0.7% melt). The lower grey scale indicates the misfit in percent between the theory and observations, in regions where  $V_s$  is too high compared with predictions.



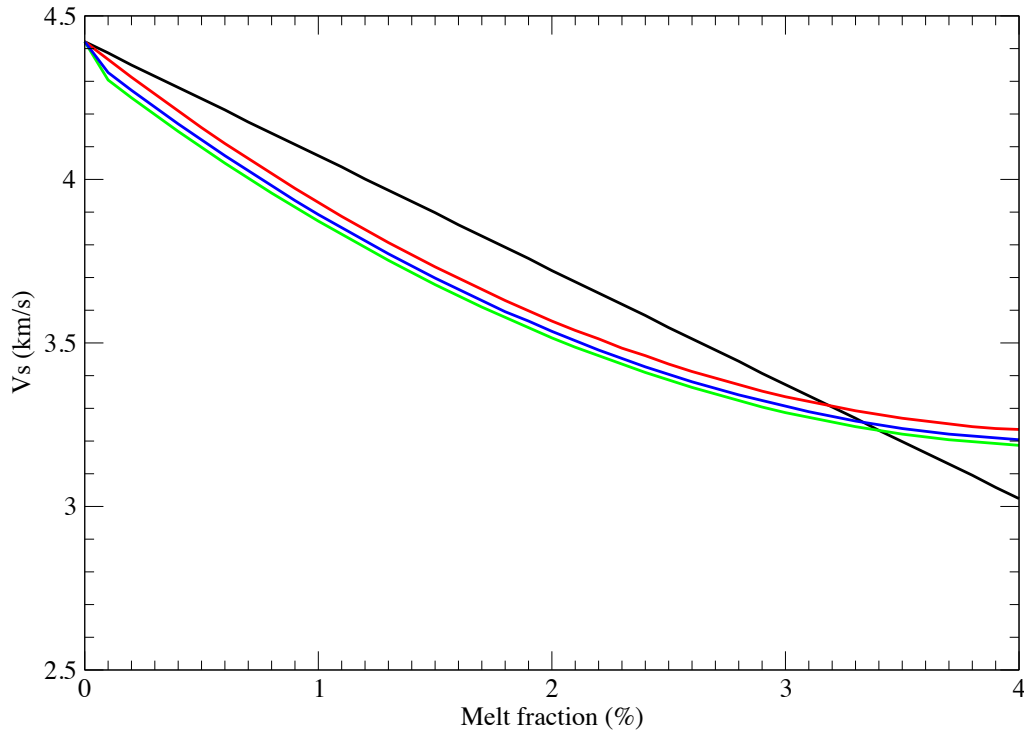


Fig. S3: Dependence of shear velocity on melt fraction  $\varphi$ . The linear  $V_s$  reduction of 7.9% per percent of melt<sup>2</sup> is shown in black from a reference velocity  $V_{ref}$ . The polynomial expression ( $V_s = 0.065\varphi^2 - 0.5565\varphi + V_{ref}$ ) derived from experimental results<sup>3</sup> is shown in red. The anelastic effect expected for seismic waves at high temperature<sup>4</sup> is shown assuming  $Q_s=80$ , the value of PREM<sup>5</sup> in the asthenosphere, for two values of  $\alpha$  (blue curve for  $\alpha=0.38$  and green curve for  $\alpha=0.26$ ). For small melt fractions ( $<1\%$ ), a stronger  $V_s$  reduction is obtained using the polynomial expression and the choice of  $\alpha$  has a small effect.

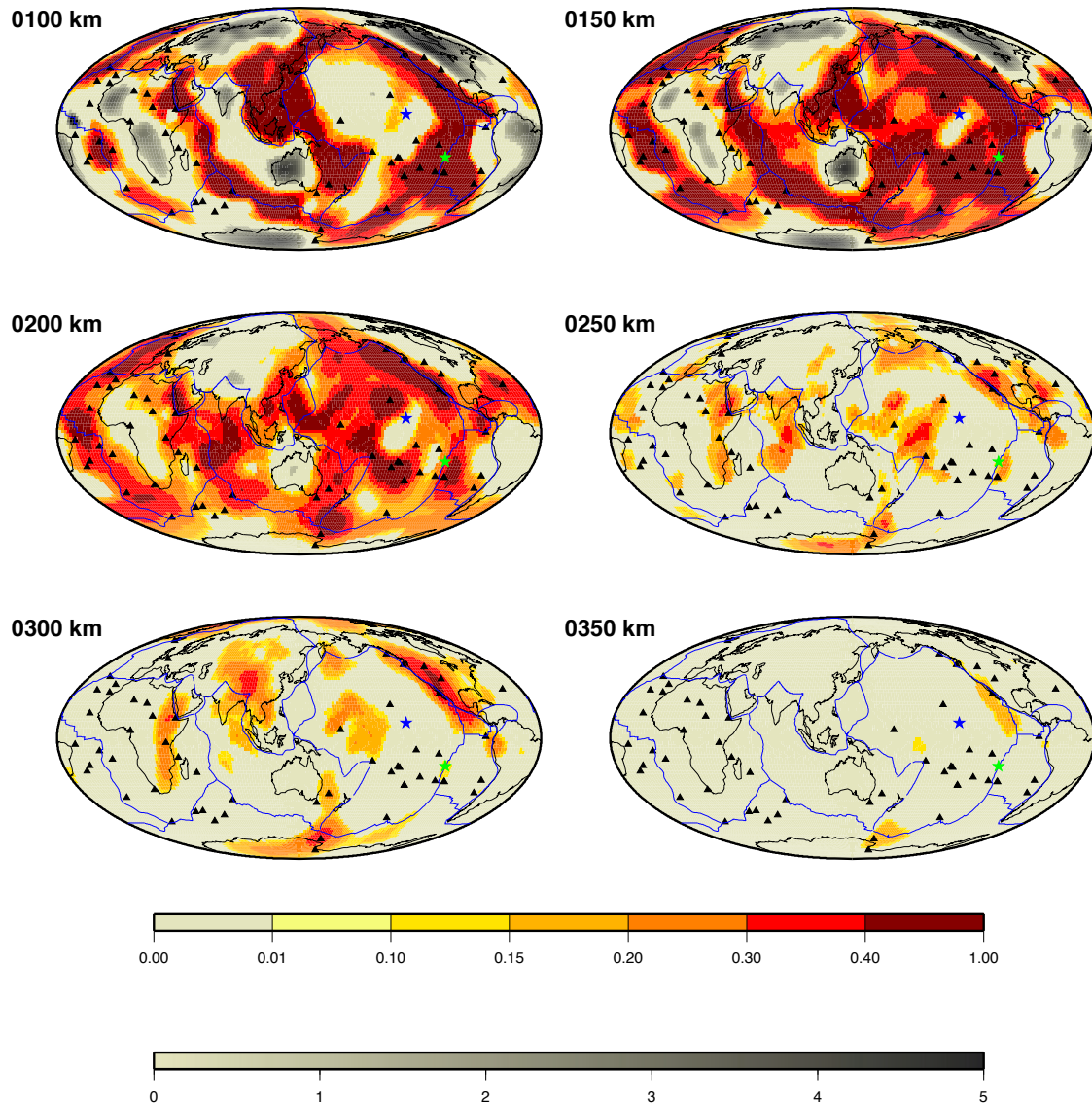


Fig. S4: same as Fig. 3 but using a linear 7.9%  $V_s$  reduction per percent of melt<sup>2</sup> instead of the polynomial expression derived from experimental results<sup>3</sup>. The upper color scale is slightly modified to allow melt content up to 1% (the maximum value at 100 km depth).

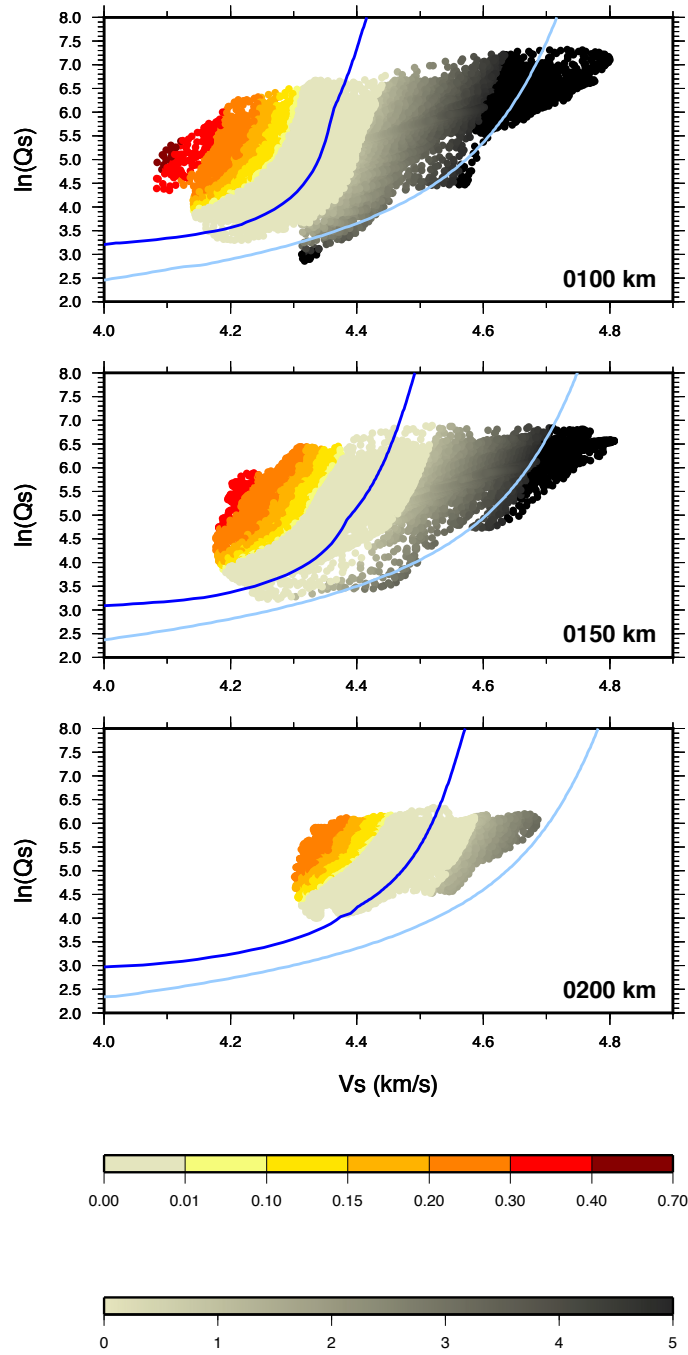


Fig. S5: same as Figure 2 but the unrelaxed shear modulus  $\mu_U$  needed to compute the temperature-dependent model (dark blue curve) is calculated using parameters proposed for of the temperature model of the Pacific<sup>6</sup>, instead of those deduced from Perple X assuming a pyrolitic composition.

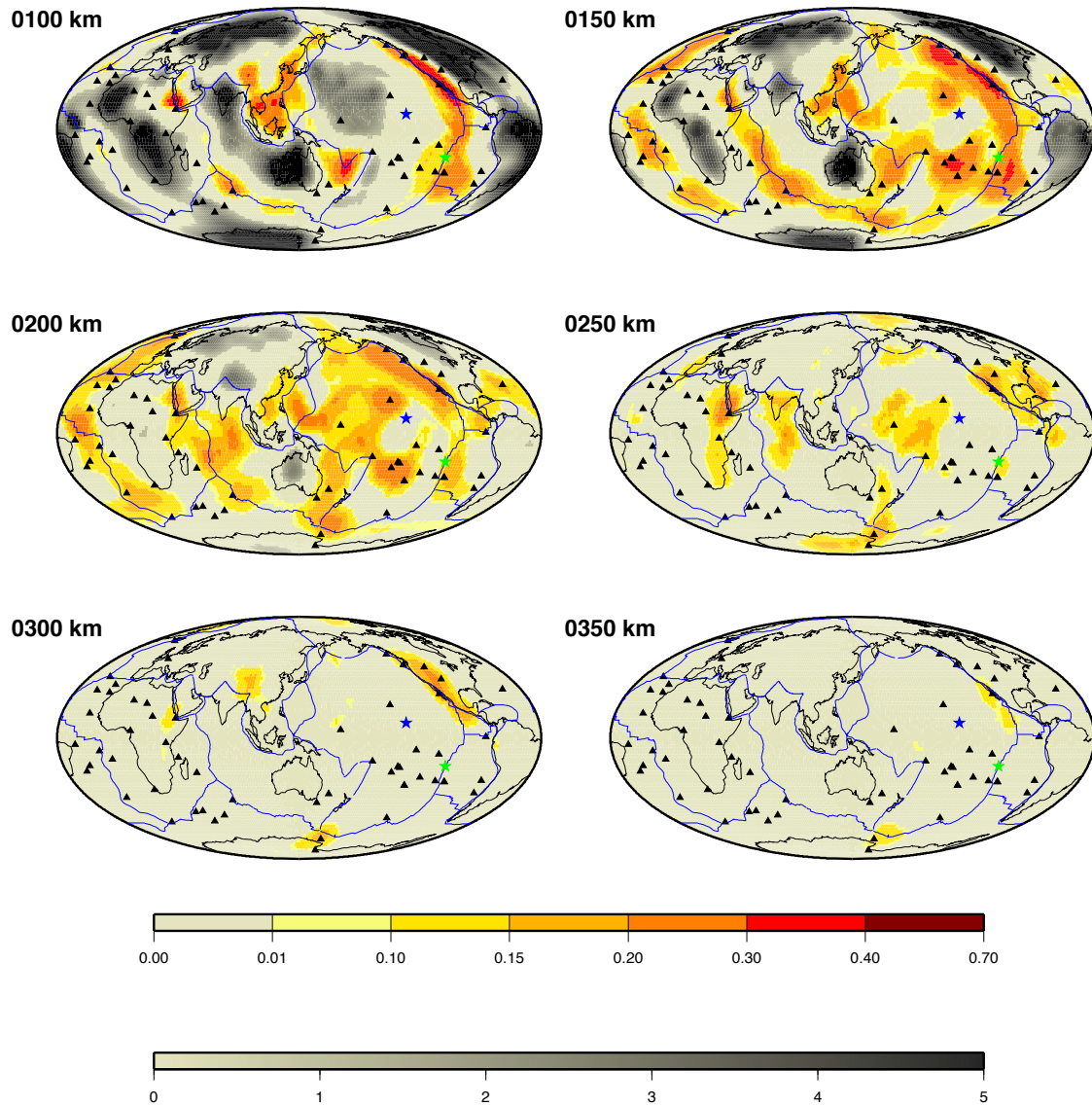


Fig. S6: same as Figure 3 but instead to estimate the unrelaxed shear modulus  $\mu_U$  using Perple X and a pyrolitic model, we use fitting parameters of the temperature model for the Pacific<sup>6</sup>.

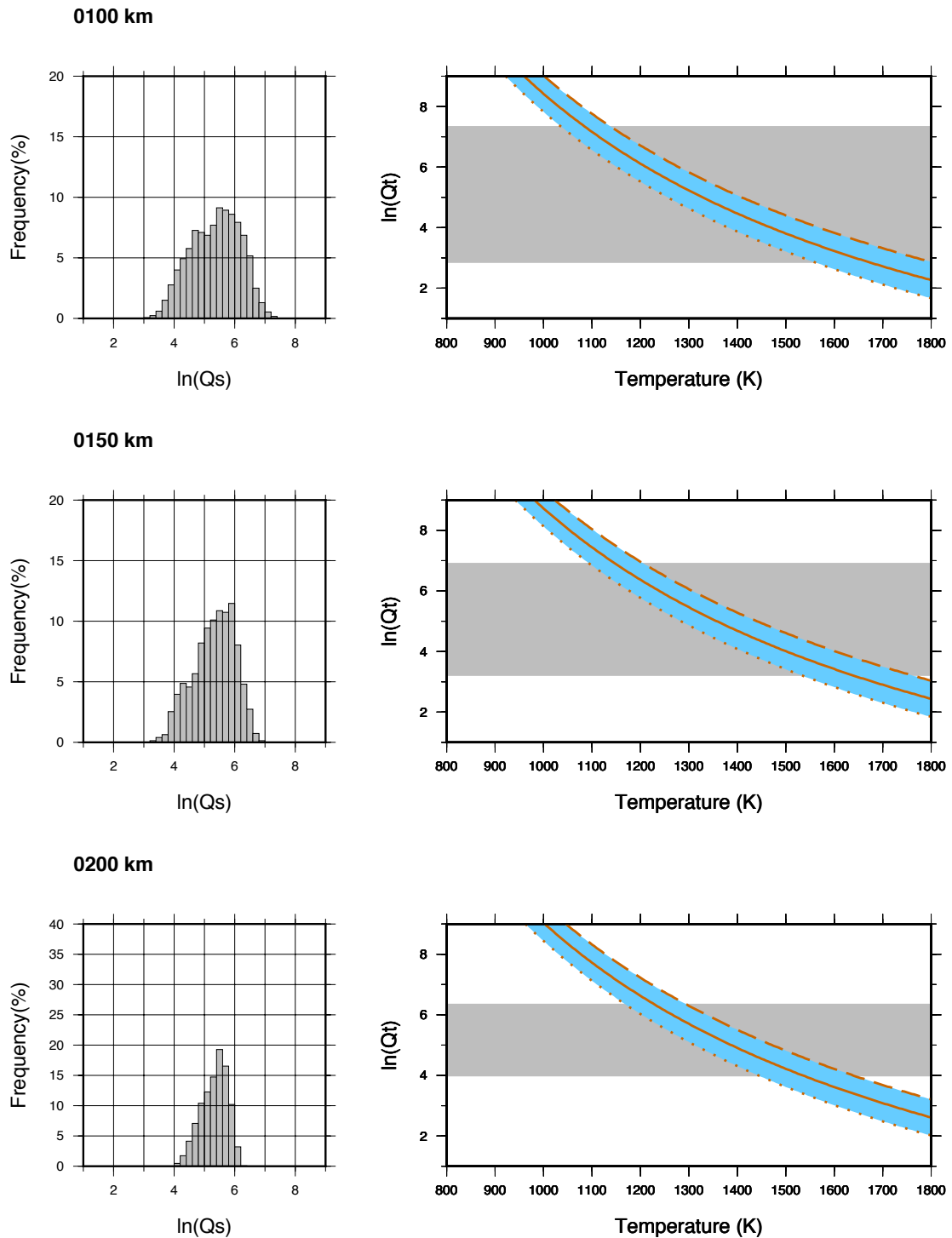


Fig. S7: Left column: Histograms of the distribution of  $\ln(Q_s)$  values extracted at 100 (top), 150 (middle) and 200 (bottom) km depth in QsADR17<sup>7</sup>. Right column: at the same depths, theoretical relation between  $\ln(Q_t)$  and temperature computed using Eq. A.1 for different grain sizes. The continuous lines in brown are the theoretical curves assuming a grain size of 10 mm. The light blue areas around this curve cover the influence of grain sizes from 1 mm (bottom dotted curve) to 100 mm (upper dashed curve). The shaded grey area shows the range of  $Q_s$  variations observed in QsADR17.

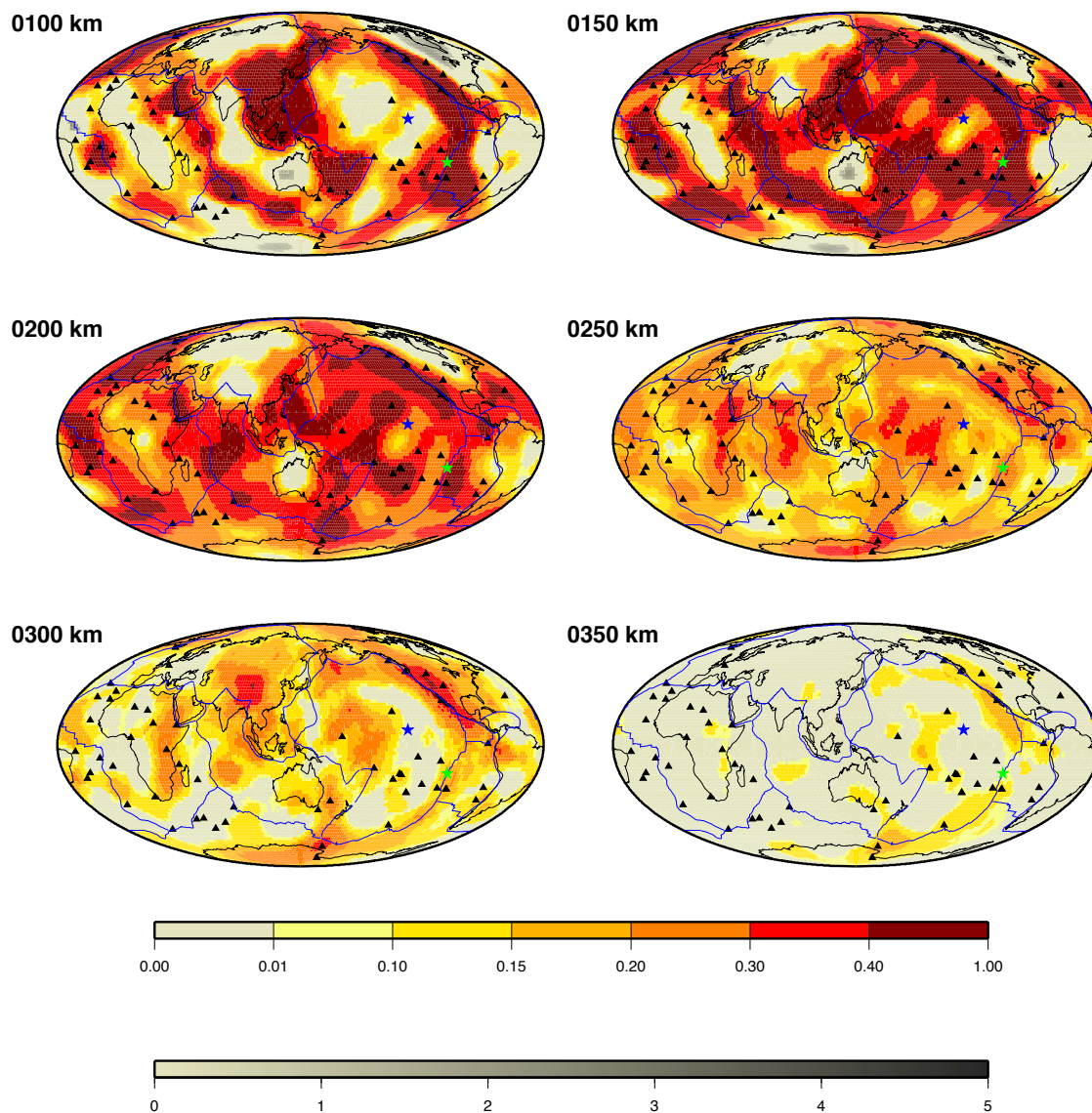


Fig. S8: same as Figure S2 but for a grain size of 100 mm.

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