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# Variability in lidar-derived particle properties over West Africa due to changes in absorption: towards an understanding

Igor Veselovskii<sup>1</sup>, Qiaoyun Hu<sup>2</sup>, Philippe Goloub<sup>2</sup>, Thierry Podvin<sup>2</sup>, Mikhail Korenskiy<sup>1</sup>, Yevgeny Derimian<sup>2</sup>, Michel Legrand<sup>2</sup>, and Patricia Castellanos<sup>3</sup>

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Abstract. Measurements performed in western Africa (Senegal) during the SHADOW field campaign are analyzed to show that spectral dependence of the imaginary part of the complex refractive index (CRI) of dust can be revealed by lidar-measured particle parameters. Observations in April 2015 provide good opportunity for such study, because, due to high optical depth of the dust, exceeding 0.5, the extinction coefficient could be derived from lidar measurements with high accuracy and the contribution of other aerosol types, such as biomass burning, was negligible. For instance, in the second half of April 2015, AERONET observations demonstrated a temporal decrease in the imaginary part of the CRI at 440 nm from approximately 0.0045 to 0.0025. This decrease is in line with a change in the relationship between the lidar ratios (the extinction-to-backscattering ratio) at 355 and 532 nm ( $S_{355}$  and  $S_{532}$ ). For instance in the first half of April,  $S_{355}/S_{532}$  is as high as 1.5 and the backscatter Ångström exponent,  $A_{\beta}$ , is as low as -0.75, while after 15 April  $S_{355}/S_{532} = 1.0$  and  $A_{\beta}$  is close to zero. The aerosol depolarization ratio  $\delta_{532}$  for the whole of April exceeded 30 % in the height range considered, implying that no other aerosol, except dust, occurred. The performed modeling confirmed that the observed  $S_{355}/S_{532}$  and  $A_{\beta}$  values match the spectrally dependent imaginary part of the refractive index as can be expected for mineral dust containing iron oxides. The second phase of the SHADOW campaign was focused on evaluation of the lidar ratio of smoke and estimates of its dependence on relative humidity (RH). For five studied smoke episodes the lidar ratio increases from  $44 \pm 5$ to  $66 \pm 7$  sr at 532 nm and from  $62 \pm 6$  to  $80 \pm 8$  sr at 355 nm, when RH varied from 25 % to 85 %. Performed numerical simulations demonstrate that observed ratio  $S_{355}/S_{532}$ , exceeding 1.0 in the smoke plumes, can indicate an increase in the imaginary part of the smoke particles in the ultraviolet (UV) range.

### 1 Introduction

Atmospheric dust has significant impact on the Earth's climate system, and this impact remains highly uncertain (IPCC, 2013). In modeling the direct aerosol effect, the vertical profile of the aerosol extinction is one of the basic input parameters, and when this profile is derived from the elastic backscatter lidar observations, knowledge of the extinctionto-backscatter ratio (so-called lidar ratio) is essential (Klett, 1985). Although the desert dust in source regions is sometimes qualified as "pure dust", it is always a mixture of various elements, e.g., iron oxides, clays, quartz and calciumrich species, whose proportions can vary (Sokolik and Toon, 1999; Kandler et al., 2011; Wagner et al., 2012; Di Biagio et al., 2017, 2019, and references therein). Thus, the dust optical properties, and hence the lidar ratio (S), can vary, depending on the relative abundance of various minerals in emission sources. The imaginary part of the complex refractive index (CRI) of different minerals can vary spectrally and often exhibits an increase in the UV spectral region for dust containing iron oxides. Therefore, the retrieval of the dust extinction profiles from elastic backscatter lidar observations should account for the spectral variation in the lidar ratio.

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Raman and HSRL lidars are capable of providing independent profiling of aerosol backscattering and extinction coefficients (Ansmann et al., 1992) and therefore are widely used to measure the lidar ratios of dust from different origins (e.g., Sakai et al., 2003; Papayannis et al., 2008, 2012; Xie et al., 2008; Ansmann et al., 2011; Mamouri et al., 2013; Burton et al., 2014; Nisantzi et al., 2015; Giannakaki et al., 2016; Hofer et al., 2017, 2020; Soupiona et al., 2018, 2019). The African deserts are the largest sources of mineral dust, and numerous studies have been conducted to quantify the particle intensive parameters (parameters independent of concentration) during dust transport from this source region to Europe and over the Atlantic Ocean (Mattis et al., 2002; Amiridis et al., 2005; Mona et al., 2006; Papayannis et al., 2008; Preißler et al., 2013; Groß et al., 2015; Rittmeister et al., 2017; Haarig et al., 2017). The dust properties are, however, modified during this transport, experiencing mixing and aging processes; thus the characterization of the dust properties near the source regions is highly important for the evaluation of the parameters of pure dust.

The lidar ratios at 355 and 532 nm ( $S_{355}$  and  $S_{532}$ ) were measured during the SAMUM-1 and SAMUM-2 experiments in Morocco and Cabo Verde, respectively (Esselborn et al., 2009; Tesche et al., 2009, 2011; Groß et al., 2011; Ansmann et al., 2011), as well as during the more recent SHADOW experiment in Senegal (Veselovskii et al., 2016, 2018). The lidar ratios  $S_{355}$  and  $S_{532}$  measured during the SAMUM experiments did not present significant spectral dependence. For example, for the SAMUM-2 campaign, the averaged values of  $S_{355}$  and  $S_{532}$  were  $53 \pm 10$  and  $54 \pm 10$  sr, respectively (Tesche et al., 2011). During SHADOW, however,  $S_{355}$  significantly exceeded  $S_{532}$  in many dust episodes, which was linked to an increase in the imaginary part of the CRI of dust at 355 nm (Veselovskii et al., 2016).

The dust backscattering coefficient, in contrast to the extinction coefficient, is sensitive to the imaginary part of the CRI (Perrone et al., 2004; Veselovskii et al., 2010; Gasteiger et al., 2011). Recall that the particle extinction is the sum of absorption and scattering, and an increase in absorption is accompanied by a decrease in scattering, leading to a weak dependence of the extinction on the imaginary part (Veselovskii et al., 2010, Fig. 6). Thus, it is expected that enhanced absorption in the UV should increase the lidar ratio. In turn, the ratio  $S_{355}/S_{532}$  should characterize the spectral variation in the imaginary part of the CRI. The latest version of AERONET products (3.0) provides inversions of the lidarrelated properties, including the lidar ratio, from almucantar scans with ground-based sun photometers. For these products, the shortest available wavelength is 440 nm. Despite the imaginary part at 440 nm (Im<sub>440</sub>) being lower than at 355 nm(Im<sub>355</sub>), AERONET observations still show an increase in absorption at 440 nm with respect to 675 nm which yields a ratio of  $S_{440}/S_{675}$  close to 1.4 for Saharan dust (Shin et al., 2018). The goal of this work is to analyze the correlation of variations in Im<sub>440</sub> from AERONET with measured values from lidar to reveal the effect of dust absorption on lidarderived aerosol properties. We focus on height and day-today variations in the dust intensive properties, such as  $S_{355}$ and  $S_{532}$  and the depolarization ratio ( $\delta$ ), as well as the extinction and backscatter Ångström exponents ( $A_{\alpha}$  and  $A_{\beta}$ , respectively) measured during several strong dust episodes in April 2015 during the SHADOW campaign.

The smoke aerosol particles, typically originated from biomass burning, can also have a pronounced spectral dependence of absorption (Nicolae et al., 2013). This is generally due to the presence of carbonaceous particles with organic compounds, so-called brown carbon (BrC; Sun et al., 2007; Kirchstetter, et al., 2004). The Sahel region is known for seasonal biomass burning caused by human activity through combustion of agricultural waste that can produce an abundant amount of BrC. The smoke can also be mixed with mineral dust during long-range transport or in the emission origin (Haywood et al., 2008). During SHADOW the observation period included the biomass burning season; thus an additional effort was dedicated to the examination of the spectral lidar ratio variability in transported biomass burning aerosol under different environmental conditions and presents a supplementary subject of the current study.

The paper is organized as follows. Section 2 describes the lidar system and provides the main expressions used for the data analysis. Several strong dust episodes, in April 2015, are analyzed in Sect. 3. In Sect. 4, the smoke episodes occurring from December 2015 to January 2016 are used to evaluate the variation in the smoke lidar ratio with relative humidity. The paper is finalized with a conclusion.

### 2 Experimental setup and data analysis

The observations were performed with a LILAS multiwavelength Raman lidar during the SHADOW campaign at Mbour, Senegal. Information related to the SHADOW and observation site is presented in Veselovskii et al. (2016). The LILAS is based on a tripled Nd:YAG laser with a 20 Hz repetition rate and pulse energy of 90/100/100 mJ at 355/532/1064 nm. The aperture of the receiving telescope is 400 mm. During the campaign, the LILAS configuration  $(3\beta + 2\alpha + 1\delta)$  allowed the measurement of three particle backscattering coefficients ( $\beta_{355}$ ,  $\beta_{532}$ ,  $\beta_{1064}$ ), two extinction coefficients ( $\alpha_{355}$ ,  $\alpha_{532}$ ) and the depolarization ratio at 532 nm ( $\delta_{532}$ ). To improve the performance of the system at 532 nm the rotational Raman channel was used instead of the vibrational one (Veselovskii et al., 2015). The measurements were performed at a 47° angle to the horizon. The backscattering coefficients and depolarization ratios were calculated with a 7.5 m range resolution (corresponding to a 5.5 m vertical resolution), while the range resolution of the extinction coefficients varied from 50 (at 1000 m) to 125 m (at 7000 m). Particle extinction and backscattering coefficients at 355 and 532 nm are calculated from elastic and Raman backscatter

signals, as described in Ansmann et al. (1992), and corresponding uncertainties are shown in the plots. An additional Raman reception channel at 408 nm was set up for profiling the water vapor mixing ratio (WVMR; Whiteman et al., 1992).

The particle depolarization ratio  $\delta$ , determined as a ratio of cross- and copolarized components of the particle backscattering coefficient, was calculated and calibrated in the same way as described in Freudenthaler et al. (2009). The relative uncertainty in depolarization measurements due to calibration is estimated to be  $\pm 10$  %. To analyze the complex aerosol mixtures, containing dust (d) and smoke (s), we can write  $\beta = \beta^d + \beta^s$  and  $\alpha = \alpha^d + \alpha^s$ . The depolarization ratio of such a mixture is therefore

$$\delta = \frac{\left(\frac{\delta^{d}}{1+\delta^{d}}\right)\beta^{d} + \left(\frac{\delta^{s}}{1+\delta^{s}}\right)\beta^{s}}{\frac{\beta^{d}}{1+\delta^{d}} + \frac{\beta^{s}}{1+\delta^{s}}}.$$
 (1)

Here  $\delta^d$  and  $\delta^s$  are the particle depolarization ratios of dust and smoke components, respectively.

To characterize the spectral dependence of the extinction  $(\alpha)$  and backscattering  $(\beta)$  coefficients, the corresponding Ångström exponents are introduced as

$$A_{\alpha} = \frac{\ln\left(\frac{\alpha_{\lambda_1}}{\alpha_{\lambda_2}}\right)}{\ln\left(\frac{\lambda_2}{\lambda_1}\right)} \text{ and } A_{\beta} = \frac{\ln\left(\frac{\beta_{\lambda_1}}{\beta_{\lambda_2}}\right)}{\ln\left(\frac{\lambda_2}{\lambda_1}\right)},\tag{2}$$

where  $\alpha_{\lambda_1}$ ,  $\alpha_{\lambda_2}$  and  $\beta_{\lambda_1}$ ,  $\beta_{\lambda_2}$  are the extinction and backscattering coefficients at wavelengths  $\lambda_1$  and  $\lambda_2$ . For the mixture of smoke and dust, the extinction Ångström exponent (EAE) can be calculated from the ratio  $\frac{\alpha_{\lambda_1}}{\alpha_{\lambda_2}}$ :

$$\frac{\alpha_{\lambda_{1}}}{\alpha_{\lambda_{2}}} = \frac{\alpha_{\lambda_{1}}^{d} + \alpha_{\lambda_{1}}^{s}}{\alpha_{\lambda_{2}}^{d} + \alpha_{\lambda_{2}}^{s}} = \frac{\alpha_{\lambda_{1}}^{d}}{\alpha_{\lambda_{2}}^{d}} \frac{\left(1 + \frac{\alpha_{\lambda_{1}}^{s}}{\alpha_{\lambda_{1}}^{d}}\right)}{\left(1 + \frac{\alpha_{\lambda_{2}}^{s}}{\alpha_{\lambda_{2}}^{d}}\right)}$$

$$= \frac{\alpha_{\lambda_{1}}^{d}}{\alpha_{\lambda_{2}}^{d}} \frac{\left(1 + \frac{\alpha_{\lambda_{2}}^{s}\left(\frac{\lambda_{2}}{\lambda_{1}}\right)^{A_{\alpha}^{s}}}{\alpha_{\lambda_{2}}^{d}\left(\frac{\lambda_{2}}{\lambda_{1}}\right)^{A_{\alpha}^{d}}}\right)}{\left(1 + \frac{\alpha_{\lambda_{2}}^{s}}{\alpha_{\lambda_{2}}^{d}}\right)}$$

$$= \frac{\alpha_{\lambda_{1}}^{d}}{\alpha_{\lambda_{2}}^{d}} \frac{\left(1 + \frac{\alpha_{2}^{s}}{\alpha_{\lambda_{2}}^{d}}\left(\frac{\lambda_{2}}{\lambda_{1}}\right)^{(A_{\alpha}^{s} - A_{\alpha}^{d})}\right)}{\left(1 + \frac{\alpha_{\lambda_{2}}^{s}}{\alpha_{\lambda_{2}}^{d}}\right)}.$$
(3)

Here  $A_{\alpha}^{\rm d}$  and  $A_{\alpha}^{\rm s}$  are the Ångström exponents of dust and smoke. The Ångström exponent of the mixture is obtained from Eq. (3):

$$A_{\alpha} = \frac{\ln \frac{\alpha_{\lambda_1}}{\alpha_{\lambda_2}}}{\ln \frac{\lambda_2}{\lambda_1}} = A_{\alpha}^{d} + \frac{1}{\ln \frac{\lambda_2}{\lambda_1}} \ln \left[ \frac{\left(1 + \frac{\alpha_{\lambda_2}^{s}}{\alpha_{\lambda_2}^{d}} \left(\frac{\lambda_2}{\lambda_1}\right)^{\left(A_{\alpha}^{s} - A_{\alpha}^{d}\right)}\right)}{\left(1 + \frac{\alpha_{\lambda_2}^{s}}{\alpha_{\lambda_2}^{d}}\right)} \right]. \tag{4}$$

The backscattering Ångström exponent (BAE) can be calculated in a similar way. And finally, the lidar ratio of the aerosol mixture is calculated as

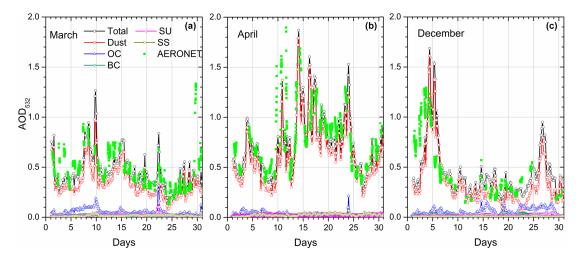
$$S = \frac{S^{\mathrm{d}}\beta^{\mathrm{d}} + S^{\mathrm{s}}\beta^{\mathrm{s}}}{\beta^{\mathrm{d}} + \beta^{\mathrm{s}}} = S^{\mathrm{d}} + \frac{\beta^{\mathrm{s}}}{\beta} \left( S^{\mathrm{s}} - S^{\mathrm{d}} \right), \tag{5}$$

where  $S^{d}$  and  $S^{s}$  are the lidar ratios of dust and smoke.

### 3 Dust observations in March and April 2015

The aerosol over West Africa presents strong seasonal variations. The spring is characterized by strong dust emissions, while, during the winter season, intense forest fires occurring in the equatorial regions emit smoke particles that are transported over Senegal (Veselovskii et al., 2018). The SHADOW campaign included the following periods of measurements: 13 March-25 April 2015, 8-25 December 2015 and 5-24 January 2016, so numerous dust and smoke episodes were observed. In our analysis of lidar-derived aerosol properties, we also considered aerosol columnar properties provided by AERONET (Holben et al., 1998) and aerosol profiles predicted by the Modern-Era Retrospective analysis for Research and Applications, Version 2 (MERRA-2) aerosol reanalysis (Gelaro et al., 2017; Randles et al., 2017). MERRA-2 is the first long-term global reanalysis to assimilate space-based aerosol observations and include their radiative coupling with atmospheric dynamics. MERRA-2 is driven by the Goddard Earth Observing System (GEOS) model version 5 that includes the Goddard Chemistry, Aerosol, Radiation and Transport (GOCART) module. GOCART models the sources, sinks and transformation of the following five aerosol species as external mixtures: dust, organic carbon (OC), black carbon (BC), sulfates (SU) and sea salt (SS). Dust and sea salt are represented by five noninteracting size bins and have wind-speed-dependent emissions. The MERRA-2 reanalysis assimilates aerosol optical depth (AOD) observations from the twin Moderate Resolution Imaging Spectroradiometer (MODIS) instruments, MODIS Terra and MODIS Aqua, as well as the AERONET ground-based sun photometer network. In addition, the profiles of meteorological variables (P, T, RH), provided by radiosondes at the Léopold Sédar Senghor International Airport, located  $\sim 70 \, \text{km}$  from the Mbour site, were also available. The relative humidity (RH) profiles over the Mbour site were calculated from a combination of lidar-derived WVMR and temperature profile from radiosounding.

Figure 1 shows the aerosol optical depth at  $532 \, \text{nm}$  (AOD<sub>532</sub>) for March, April and December 2015 recalculated from AERONET AOD at 500 nm using the 440–675 nm Ångström exponent. The same figure shows the AODs for the five aerosol species used in the MERRA-2 model, such as dust, organic carbon (OC), black carbon (BC), sulfates (SU) and sea salt (SS). The optical depths provided by MERRA-2 and AERONET are in a good agreement. Dust is the pre-



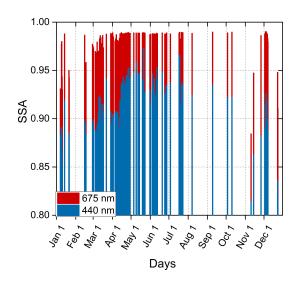
**Figure 1.** The aerosol optical depth (AOD) at 532 nm (open circles) and AODs of the main aerosol components, such as dust, organic carbon (OC), black carbon (BC), sulfates (SU) and sea salt (SS), provided by the MERRA-2 for (a) March, (b) April and (c) December 2015 over Mbour. Green squares show AOD<sub>532</sub> provided by AERONET.

dominant aerosol component for all 3 months with the highest values of AOD in April. The contribution of organic carbon (the main component of the biomass burning products) is significant in December, when the forest fire season starts in equatorial regions, though a noticeable amount of OC is also predicted for March and for the beginning of April. The contribution of BC and SU to the total AOD is low: the sum of the corresponding AODs is below 0.1 for all 3 months.

The single-scattering albedo (SSA) over the Mbour site in 2015 provided by AERONET at 440 and 675 nm is shown in Fig. 2. The SSA<sub>675</sub> is above 0.97 for the March–April period, but at 440 nm dust absorption is stronger, and, in March, SSA<sub>440</sub> is about 0.9. However, in the middle of April, SSA<sub>440</sub> increases up to 0.95, indicating that aerosols become less absorbing at shorter wavelengths. We can thus expect that variation in SSA at 355 nm between April and March should be even stronger. In our study we consider two groups of observations. The first group corresponds to the beginning of April, when SSA at 440 nm was lower. The second group covers the second half of April, when SSA at 440 nm increased. By analyzing these two groups we expect to reveal the effect of aerosol absorption on lidar-derived aerosol properties.

### 3.1 Dust episode on 1–4 April 2015

In the beginning of April the dust was transported by continental trades (Harmattan) from the northeastern and eastern drylands. For the period 1–4 April, as follows from Fig. 1b, the  $AOD_{532}$  over Dakar increased up to 1.0. Figure 3 shows the spatiotemporal distributions of the aerosol backscattering coefficient  $\beta_{532}$ , particle depolarization ratio  $\delta_{532}$ , and water vapor mixing ratio for the nights 1–2, 2–3 and 3–4 April 2015. The corresponding air mass back trajectories, shown in Fig. 4, demonstrate that on 1–2 and 2–



**Figure 2.** Aerosol single-scattering albedo (SSA) at 675 and 440 nm provided by AERONET for Mbour site in 2015.

3 April air masses at all heights arrive from the northeast, whereas on 3–4 April the air masses above 2500 m are advected from the east. These air masses are characterized by higher humidity and may contain biomass burning products. During these three nights, the linear particle depolarization ratio and WVMR present some evolution. On 1–2 April  $\delta_{532}$  exceeds 30% and does not change significantly within the dust layer, even if some decrease is observed above 2000 m after 03:00 UTC. By 3–4 April the depolarization ratio above 2500 m decreases below 25%, simultaneously with an increase in the WVMR. During the dust episode, the relative humidity did not exceed 20% on 1–3 April, but on 3–4 April it increased up to 40% above 2500 m.

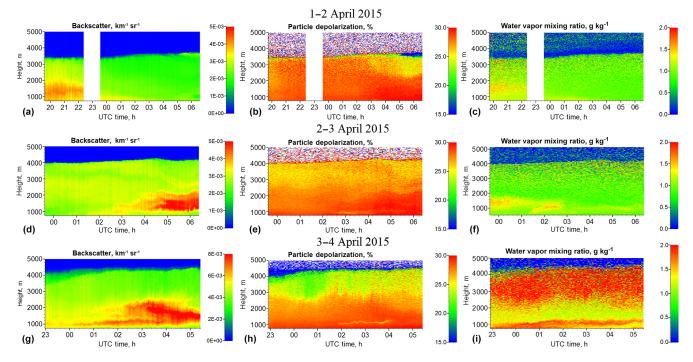


Figure 3. Spatiotemporal distributions of aerosol backscattering coefficient  $\beta_{532}$  (**a**, **d**, **g**), particle depolarization ratio  $\delta_{532}$  (**b**, **e**, **h**) and water vapor mixing ratio (**c**, **f**, **i**) for the nights 1–2 April (**a**, **b**, **c**), 2–3 April (**d**, **e**, **f**) and 3–4 April 2015 (**g**, **h**, **i**).

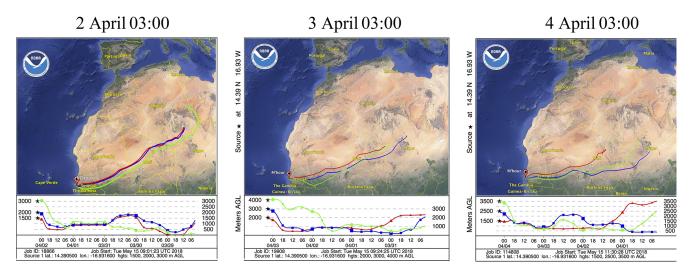


Figure 4. For the air mass in Mbour, 3 d backward trajectories on 2, 3, and 4 April 2015 at 03:00 UTC obtained with the HYSPLIT model.

Vertical profiles of dust particle properties such as aerosol extinction coefficients  $\alpha_{355}$  and  $\alpha_{532}$ , particle depolarization ratio  $\delta_{532}$ , and lidar ratios  $S_{355}$  and  $S_{532}$  are shown in Fig. 5 for the three observation periods on 1, 2–3 and 3–4 April 2015. The profiles of backscattering coefficients for 2–3 and 3–4 April are given in Fig. 6, while the extinction and backscatter Ångström exponents, calculated at 355 and 532 nm wavelengths for three temporal intervals from Fig. 5, are presented in Fig. 7. During all three observation periods  $A_{\alpha}$  is slightly negative ( $A_{\alpha} = -0.1 \pm 0.1$ )

up to 2000 m. For the dust component, MERRA-2 provides value of  $A_{\alpha}=-0.14$ , which agrees with observations. Above 2000 m,  $A_{\alpha}$  exhibits some increase, which is most significant on 3–4 April, when  $A_{\alpha}$  reaches  $0.3\pm0.1$  at 4000 m height. A simultaneous decrease in  $\delta_{532}$  indicates the possible presence of smoke particles above 2000 m. The backscatter Ångström exponent  $A_{\beta}$ , in contrast with  $A_{\alpha}$ , is sensitive to the spectral dependence of the imaginary part of CRI, thus yielding complicated vertical variability in  $A_{\beta}$ . In particu-

lar, on 2–3 April  $A_{\beta}$  decreases from -0.5 to -0.7 within a 1500–2500 m height range, when  $A_{\alpha}$  remains stable.

As follows from Fig. 5, on 1 April the lidar ratio  $S_{355} =$  $70 \pm 6$  sr does not change with height, while  $S_{532}$  gradually decreases from  $60 \pm 5$  sr at 1000 m to  $50 \pm 4$  sr at 3000 m in height. In sessions that followed (Fig. 5b, c) the lidar ratios at both 355 and 532 nm decreased. Thus, the range of lidar ratio variation for the dust episode on 1-4 April is 60-70 sr at 355 nm and 45–60 sr at 532 nm. The lidar ratios ( $S_{355}$  and  $S_{532}$ ) modeled by MERRA-2 for the dust component are also shown in Fig. 5. The corresponding lidar ratio values are 70 and 42 sr, respectively, and do not vary with altitude as the model optical properties of all dust size bins based on spectral complex refractive indices from the Optical Properties of Aerosols and Clouds (OPAC) tables (Hess et al., 1998) and the spheroidal-shape models developed by Meng et al. (2010) are the same and fixed, as dust is treated as hydrophobic. The modeled  $S_{355}$  value is near the top of the range of observed values, while modeled  $S_{532}$  underestimates the observations.

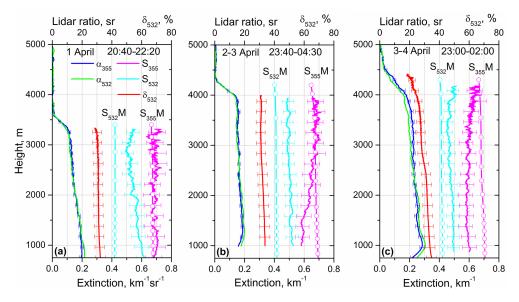
The gradual decrease in  $S_{532}$  with height in Fig. 5a and c is however unusual. There are, at least, two possible reasons to explain  $S_{532}$  height variation. The first one is the presence of nondust particles, for example, smoke. The second reason is that the properties (composition) of dust change with height. If nondust particles are present, the particle intensive properties, such as S,  $\delta$  and  $A_{\alpha}$ , should vary with height in a consistent way. The MERRA-2 modeling reported in Fig. 1 shows that in the beginning of April the organic carbon is the second main contributor to the AOD, after dust. We should recall, however, that while the model can provide a realistic range of OC variation, it does not necessarily reproduce the exact spatiotemporal distribution of the OC extinction coefficient.

In the dust episode considered, the most significant smoke contribution was observed on 3-4 April. Figure 8a shows the profiles of measured  $\alpha_{355}$  and  $\alpha_{532}$  together with MERRA-2 modeled extinction coefficients at 532 nm for five aerosol components. The extinction Angström exponents measured by lidar and modeled by MERRA-2 for dust components are given in Fig. 8b. The same figure also shows the lidar-derived water vapor mixing ratio profile together with the relative humidity. At low altitudes (below 2500 m), where aerosol is represented by pure dust, the measured and modeled values of extinction coefficients are close. Above 2500 m the measured value of  $\alpha_{355}$  exceeds that of  $\alpha_{532}$ , indicating the presence of smoke particles, while the modeled contribution of OC to the total extinction is very low. The measured extinction Ångström exponent is about -0.1 below 2000 m, which agrees well with modeling results for pure dust. An increase in WVMR and RH above 2000 m coincides with growth in the  $A_{\alpha}$ . For the considered case, the model reproduces correctly the dust loading but underestimates the smoke contribution. At 3500 m, the difference between measured and modeled  $\alpha_{532}$  is about  $0.045 \, \mathrm{km}^{-1}$  which can be attributed to the smoke contribution.

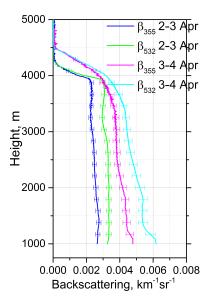
Dust and smoke particles contributions to the total backscattering coefficient can be also separated on the basis of the depolarization measurements, assuming that depolarization ratios of these particles are known (Tesche et al., 2009). The results of such decomposition are presented in Fig. 8c, assuming 35 % and 7 % for the dust and smoke depolarization ratio, respectively. The  $\delta_{532} = 7\%$  was the lowest value that we observed in elevated smoke layers during the SHADOW experiment (Veselovskii et al., 2018); however, due to a large difference in smoke and dust depolarization ratios, the choice of the exact value for the smoke did not influence significantly the results. The contribution of smoke to the total  $\beta_{532}$  at 3500 m is 0.0009 km<sup>-1</sup> sr<sup>-1</sup>. For the smoke lidar ratio of 50 sr at 532 nm (validity of this choice will be discussed in Sect. 3.3), the smoke extinction coefficient is about  $0.045 \,\mathrm{km}^{-1}$ . This value agrees well with the smoke contribution obtained from Fig. 8a at 3500 m and thus can be used for estimating the smoke effect on the intensive aerosols properties derived from lidar measurements.

The depolarization ratio of the dust-smoke mixture, calculated with Eq. (1), matches the observed value since decomposition in Fig. 8c is based on depolarization measurements. The Ångström exponent at 3500 m computed with Eq. (4) for  $\alpha_{532}^{s} = 0.045 \text{ km}^{-1}$ ,  $\alpha_{532}^{d} = 0.147 \text{ km}^{-1}$ ,  $A_{\alpha}^{d} = -0.1 \text{ and}$  $A_{\alpha}^{\rm s} = 0.9$  yields  $A_{\alpha} = 0.28$ , which is close to the observed value  $0.26\pm0.08$ . Hence, the observed variation in  $A_{\alpha}$  above 2000 m on 3-4 April is well explained by the smoke contribution. In a similar way, using Eq. (5) we can estimate the smoke lidar ratio  $(S_{532}^{s})$  that would match the observed decrease in  $S_{532}$ . To explain the decrease in the lidar ratio at 3500 m from 50 to 45 sr, the smoke lidar ratio should be about 25 sr, which is unrealistically small (Burton et al., 2012). Such a small lidar ratio could be attributed to the maritime aerosol, but then the lidar ratios at both wavelengths should decrease simultaneously. Recall that on 1–2 April the smoke contribution was significantly lower, while the decrease in S<sub>532</sub> was about 10 sr. Thus, presence of smoke particles cannot explain the observed decrease in  $S_{532}$  and it should probably be attributed to changes in dust composition (and so the imaginary part) with height.

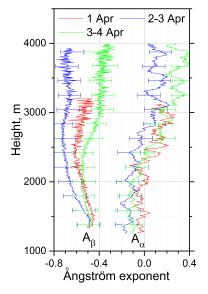
The smoke lidar ratio is usually assumed to be higher than that of dust (Tesche et al., 2011; Burton et all., 2012); meanwhile in Fig. 5c the lidar ratio  $S_{532}$  is not increased in the presence of the smoke particles. It should however be noticed that our results were obtained at low RH. The smoke particles are hygroscopic, and the lidar ratio should increase with RH. The way to characterize  $S_{532}^{\rm s}$  over the Dakar site can be based on the analysis of the lidar measurements during smoke episodes within a height range where smoke contribution becomes predominant. The results of such analysis will be discussed later in Sect. 3.3.



**Figure 5.** Vertical profiles of extinction coefficients ( $\alpha_{355}$ ,  $\alpha_{532}$ ) and lidar ratios ( $S_{355}$ ,  $S_{532}$ ) at 355 and 532 nm together with the particle depolarization ratio  $\delta_{532}$  measured on 1 April (20:40–22:20 UTC), 2–3 April (23:40–04:30 UTC) and 3–4 April 2015 (23:00–02:00 UTC). Symbols show the lidar ratios of dust provided by the MERRA-2 model ( $S_{355}M$ ,  $S_{532}M$ ).



**Figure 6.** Backscattering coefficients for observations presented in Fig. 5 for 2–3 and 3–4 April.

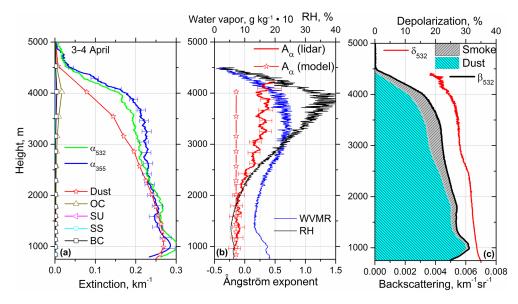


**Figure 7.** Vertical profiles of the extinction and backscattering Ångström exponents  $(A_{\alpha} \text{ and } A_{\beta})$  at 355–532 nm for three temporal intervals from Fig. 5.

### 3.2 Dust episodes on 14 and 24 April 2015

In the second part of April 2015, dust AOD<sub>532</sub> exceeded 1.0 (Fig. 1b) and contributions of other aerosol components were insignificant. Meanwhile, as follows from Fig. 2, SSA<sub>440</sub> increased after 15 April; thus dust became less absorbing in the UV, which should influence the lidar-derived aerosol intensive properties. Figure 9 shows the values of the extinction coefficients and lidar ratios at 355 and 532 nm, together with the depolarization ratio  $\delta_{532}$  and the Ångström exponents  $A_{\alpha}$ 

and  $A_{\beta}$  observed on 14 April (00:00–05:00 UTC) and 23–24 April (23:00–06:00 UTC). The first case is a transition day when SSA<sub>440</sub> starts to increase. The aerosol extinction profiles presented in Fig. 9a show that two dust layers can be distinguished. In the first layer (below 2.5 km), aerosol intensive properties are similar to that of 1–4 April with  $S_{355} > S_{532}$ , slightly negative  $A_{\alpha} = -0.1$  and  $A_{\beta}$  as low as -0.35. In the second layer,  $S_{355}$  and  $S_{532}$  coincide and both  $A_{\alpha}$  and  $A_{\beta}$  are



**Figure 8.** Vertical profiles of (a) extinction coefficients at 355 and 532 nm ( $\alpha_{355}$ ,  $\alpha_{532}$ ) measured by lidar (lines) and modeled by MERRA-2 (line + symbol) for five aerosol components at 532 nm; (b) extinction Ångström exponents at 355–532 nm obtained from lidar observations and modeled by MERRA-2 for pure dust (stars) together with water vapor mixing ratio (WVMR) and the relative humidity; (c) contribution of dust and smoke particles to  $\beta_{532}$  together with particle depolarization ratio  $\delta_{532}$ . Values of WVMR are multiplied by a factor of 10. Lidar measurements were performed on 3–4 April 2015 for the period 23:00–02:00 UTC. Modeling results are given for 4 April 00:00 UTC.

close to zero. The depolarization ratio in the second layer is about 31 %, slightly lower than in the first one. Thus, we can assume that an increase in the imaginary part in UV in the first layer is more significant than in the second one. From the analysis of air mass back trajectories given in Fig. 10, we can conclude that the air masses in the first layer originate from the northeastern and eastern drylands, while in the second layer the air masses arrive from the east. After 14 April,  $S_{355}$  and  $S_{532}$  coincided for the whole height range, and results obtained on 23-24 April (Fig. 8c, d) are the example of such observations. Air mass back trajectories show that the air masses at both 2.0 and 3.0 km height are transported from the east. The ratio  $S_{355}/S_{532}$  is close to 1.0 within the whole dust layer, and both Ångström exponents  $A_{\alpha}$ ,  $A_{\beta}$  are close to zero. Thus, the results from Figs. 9 and 10 indicate that lidarderived aerosol properties depend on the dust source origin.

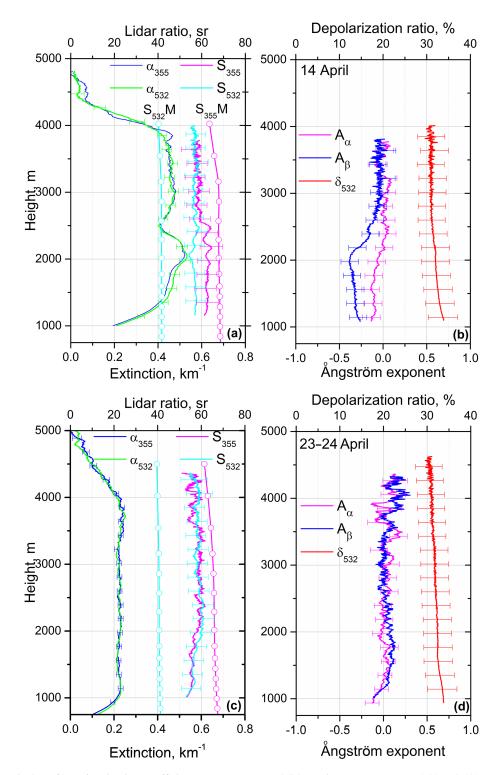
# 3.3 Analysis of lidar ratio variations in March–April 2015

Figure 11 summarizes the lidar ratio measurements for the period from 29 March to 24 April 2015 (first phase of SHADOW ended on 25 April). Here we focus on the properties of pure dust and thus do not show results before 29 March, when AOD was lower and the contribution of other aerosol types could be significant (Fig. 1). For Fig. 11 we have chosen height intervals where the S value is stable and  $\delta$  exceeds 30%. For example, on 14 and 24 April lidar ratios are averaged inside 2.7–3.7 and 2.0–4.0 km layers, respectively. For the period considered,  $S_{355}$  and  $S_{532}$  vary in

the ranges of 50–80 and 45–60 sr, respectively, with mean values of 62 and 51 sr. Enhanced variability in  $S_{355}$  compared to in  $S_{532}$  can be explained by variation in the imaginary part at 355 nm. At the beginning of the 29 March and 8 April dust episodes, the  $S_{355}/S_{532}$  ratio is as high as 1.5 and then gradually decreases. After 14 April, the  $S_{355}/S_{532}$  ratio becomes close to 1.0; thus S presents no spectral dependence.

The day-to-day variation in the aerosol column properties, including the spectrally dependent complex refractive index, can be obtained from AERONET (Holben et al., 1998). Figure 12 shows the imaginary part of the aerosol refractive index at 440 and 675 nm (Im<sub>440</sub>, Im<sub>675</sub>) provided by AERONET for the same period of time as in Fig. 11. The Im<sub>440</sub> strongly decreases after 14 April, correlating with the decrease in the  $S_{355}/S_{532}$  ratio in Fig. 11, which corroborates the suggestion that variations in the  $S_{355}/S_{532}$  ratio are related to variation in dust absorption in the UV range. The retrieved real part (Re) of the complex refractive index oscillates around Re = 1.45 and shows no significant spectral dependence. Correlation between enhancement of Im<sub>440</sub>, with respect to  $Im_{675}$ , and an increase in lidar-derived  $S_{355}/S_{532}$ is clearly seen in Fig. 13, showing time series of the Im<sub>440</sub>- $Im_{675}$  difference and  $S_{355}/S_{532}$  ratio.

To analyze the variations in the observed lidar ratios and the Ångström exponents, a simplified numerical simulation has been performed. For a realistic modeling of the dust lidar ratio, various mixtures of different mineral components and particles shapes should be considered. Sensitivity of the modeling results to the dust mixture parameters was demonstrated in the study of Gasteiger et al. (2011). Such detailed



**Figure 9.** (**a**, **c**) Vertical profiles of extinction coefficients ( $\alpha_{355}$ ,  $\alpha_{532}$ ) and lidar ratios ( $S_{355}$ ,  $S_{532}$ ) at 355 and 532 nm, together with (**b**, **d**) the particle depolarization ratio  $\delta_{532}$  and extinction and backscattering Ångström exponents ( $A_{\alpha}$ ,  $A_{\beta}$ ) measured on (**a**, **b**) 14 April 2015 (00:00–05:00 UTC) and (**c**, **d**) the night 23–24 April (23:00–06:00 UTC). Open symbols on plots (**a**, **c**) show the lidar ratios  $S_{355}$ M and  $S_{532}$ M provided by the MERRA-2 model on 14 and 24 April at 00:00 UTC.

# 14 April 03:00

# The Gamma Superstants Service Service

# 24 April 03:00

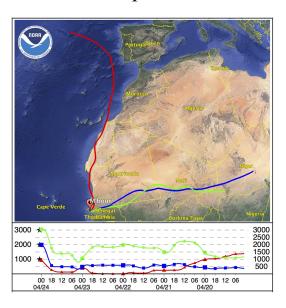
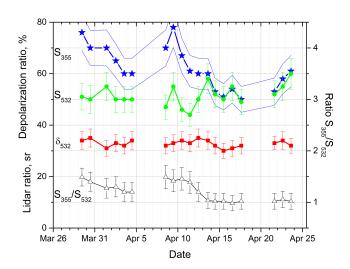


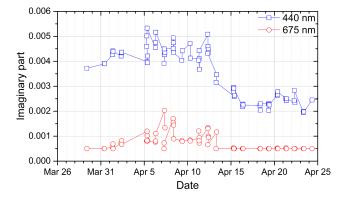
Figure 10. For 14 April (03:00 UTC) and 24 April (00:00 UTC) 2015, 4 d backward trajectories obtained with the HYSPLIT model.



**Figure 11.** Lidar ratios  $S_{355}$  and  $S_{532}$  and the particle depolarization ratio  $\delta_{532}$  for dust episodes in March–April 2015. Open triangles show the ratio  $S_{355}/S_{532}$ .

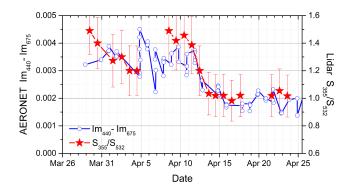
modeling, however, is outside the scope of the present paper. Here we only intend to evaluate the main impact when the imaginary part of CRI is modified.

The lidar ratio depends not only on the complex refractive index but also on the dust particle size distribution (PSD). The PSDs provided by AERONET on 2 and 23 April 2015 (three distributions for each day) are shown in Fig. 14. The PSDs are similar, and the effective radii for both days are about  $0.75 \, \mu m$ ; thus, differences in S observed for 2 and 23 April should be related mainly to the complex refractive

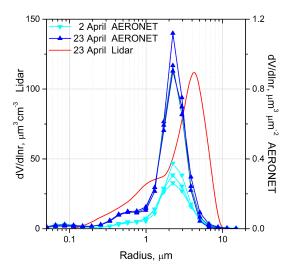


**Figure 12.** Imaginary part of the refractive index at 440 and 675 nm provided by AERONET for March–April 2015.

index. Figure 15a presents modeled  $S_{355}$  and  $S_{532}$  lidar ratios together with the extinction and backscattering Ångström exponents  $A_{\alpha}$  and  $A_{\beta}$  as a function of the imaginary part. Computations were performed for the AERONET-derived size distribution on 23 April from Fig. 14 using the assembly of randomly oriented spheroids (Dubovik et al., 2006) with the real part Re = 1.55.  $S_{355}$  and  $S_{532}$  increase with the imaginary part, and the ratio  $S_{355}/S_{532}$  is about 1.1. The extinction coefficient is slightly sensitive to the imaginary part; thus an increase in S in Fig. 15 is due to a decrease in the backscattering coefficient with Im. The modeled  $A_{\alpha}$  is about  $A_{\alpha}=0.1$ , while  $A_{\beta}$  decreases with Im to  $A_{\beta}=-0.2$ . To estimate the influence of a spectrally dependent imaginary part Im( $\lambda$ ) on  $A_{\beta}$ , we have also performed computations assuming a fixed Im $_{532}=0.002$  and only Im $_{355}$  is free to vary. Corresponding



**Figure 13.** Difference of  $Im_{440}$ – $Im_{675}$  from Fig. 12 together with lidar measured values  $S_{355}/S_{532}$  from Fig. 11 for days in April 2015.



**Figure 14.** The particle size distributions provided by AERONET on 2 and 23 April 2015 (three PSDs for each day). Red line shows the PSD derived from  $3\beta + 2\alpha$  lidar measurements on 23–24 April within a 2.0–3.0 km height range.

results are shown in Fig. 15a with open stars. Spectral dependence of the imaginary part significantly decreases  $A_{\beta}$ : for Im<sub>355</sub> = 0.005 (Im<sub>355</sub>–Im<sub>532</sub> = 0.003),  $A_{\beta}$  decreases to -0.75.

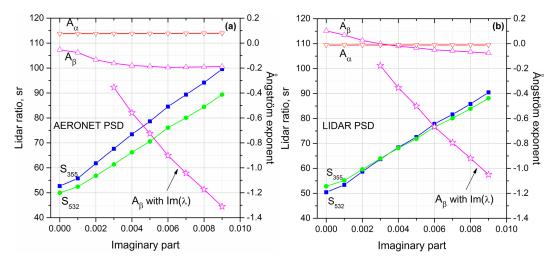
We should recall, however, that for the second half of April the observed ratio  $S_{355}/S_{532}$  was about 1.0 and both extinction and backscatter Ångström exponents were close to zero. To figure out the kind of PSD that would reproduce those observations, we retrieved the PSD from  $3\beta + 2\alpha$  measurements, as described in Veselovskii et al. (2002, 2010). For that purpose, data from 23–24 April (Fig. 9), averaged within a 2–3 km layer, were inverted, and the corresponding PSD is shown in Fig. 14 with a red line. Inversion was performed for the assembly of randomly oriented spheroids, assuming a spectrally independent refractive index. Due to the limited number of input data (five) we are able to reproduce only the main features of the PSD. The maximum of this lidar-

derived PSD is shifted towards larger radii, with respect to the AERONET size distribution, but at the same time, the retrieved PSD contains a significant contribution from the fine particles. The simulation results for this lidar-derived PSD, are given in Fig. 15b. The lidar ratios  $S_{355}$  and  $S_{532}$  for all values of the imaginary part are close. The backscatter and extinction Ångström exponents are close to zero, matching the observations of the second half of April 2015. Thus simulation results demonstrate dependence on the PSD chosen, but in both cases these lead to the same conclusion: observed low values of  $A_{\beta}$  cannot be reproduced without accounting for spectral dependence of the imaginary part.

To compare computations and observations, information on Im<sub>355</sub> and Im<sub>532</sub> values is needed. The recently measured refractive indices of dust, sampled in different regions of Africa, are presented by Di Biagio et al. (2019). In particular, for the countries located north and east of Senegal, the aerosol imaginary parts at 370, 470, 520 and 660 nm are of 0.0043, 0.0033, 0.0026 and 0.0013 for Mauritania and 0.0048, 0.0038, 0.0030 and 0.0024 for Mali, respectively. The highest values of lidar ratios observed in our measurements are about 60 and 80 sr at 532 and 355 nm, respectively. Corresponding imaginary parts of the CRI from Fig. 15 can be estimated as  $Im_{532} = 0.002-0.003$  and  $Im_{355} = 0.005-$ 0.006, which agrees with results presented by Di Biagio et al. (2019). Assuming  $Im_{355} = 0.005$  and  $Im_{532} = 0.002$ , the modeled ratio  $S_{355}/S_{532}$  is about 1.44 and  $A_{\beta}$  is about -0.75 for both AERONET and lidar-derived PSDs, which again agrees reasonably with observations. The modeling performed is very simplified; still it confirms that the observed values of the  $S_{355}/S_{532}$  ratio and  $A_{\beta}$  can be explained by the spectral dependence of the imaginary part of the CRI.

Thus, based on our measurement results, two types of dust can be distinguished. The first type has a high  $S_{355}/S_{532}$  ratio (up to 1.5). Such a kind of dust is characterized by an increase in the imaginary part in the UV range, and it was observed, for example, during 29 March and 10 April episodes. For the second type, the ratio  $S_{355}/S_{532}\approx 1.0$ , so variation in the imaginary part of the refractive index between 532 and 355 nm wavelengths should be smaller than for the first type. Such dust was observed in the second half of April 2015. Both types are characterized by a high depolarization ratio,  $\delta_{532}$  values, exceeding 30 %, so depolarization measurements at 532 nm are not capable of discriminating between these two types of dust.

The difference in the observed dust properties is probably related to the mineralogical characteristics in the source region. From the back trajectories analysis presented in Figs. 4 and 10 one can suppose that the first type of dust was transported from the north–east, while the second type from the east. In order to verify if a difference in the dust emission source region and transport takes place, we also analyzed the Infrared Difference Dust Index (IDDI) derived from the Meteosat Second Generation (MSG) geostationary satellite imagery in the thermal infrared (TIR) range. The IDDI was



**Figure 15.** Lidar ratios  $S_{355}$  and  $S_{532}$  together with the extinction and backscattering Ångström exponents  $A_{\alpha}$  and  $A_{\beta}$  calculated for (a) AERONET PSD on 23 April from Fig. 14 and (b) lidar-derived PSD from Fig. 14 as a function of the imaginary part. Open stars show  $A_{\beta}$  for spectrally dependent imaginary part Im( $\lambda$ ), assuming that Im<sub>532</sub> = 0.002 is fixed and only Im<sub>355</sub> is free to vary. Computations are performed for the assembly of randomly oriented spheroids with the real part Re = 1.55.

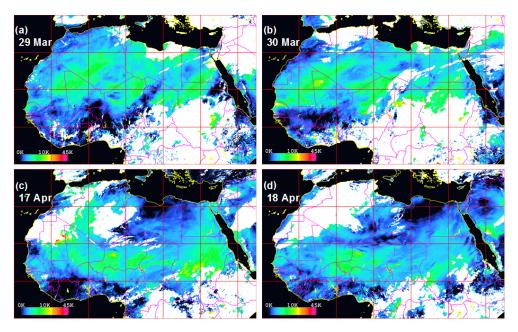
originally developed by Legrand et al. (1989, 2001) for the Meteosat First Generation (MFG) and is based on the impact of the airborne mineral dust on the TIR radiation emitted by the terrestrial surface. The physical principle of the IDDI derivation is in the thermal contrast between terrestrial surface and atmosphere, and the best sensitivity is found at around noontime when the surface temperature is maximal (Legrand et al., 1989). The IDDI product shows that the brightness temperature of terrestrial surface observed by satellite can be reduced by up to about 50° K in the presence of airborne mineral dust, and a reduction by about 10° K already indicates a major dust event (Legrand et al., 2001). A direct relationship between the IDDI and aerosol optical thickness in the solar spectrum and visibility was also found (Legrand et al., 2001). It should be mentioned here that the IDDI was initially developed for MFG and the absolute consistency with the IDDI values from MSG should be examined due to differences in spatial and spectral resolutions between the two sensors. However, the physical principles used for the IDDI determination are the same, and a direct application of the MFG IDDI algorithm to MSG was found to be possible. Moreover, tests showed that the absolute values of the IDDI for a coincident overlapping period of MFG and MSG are very close. Nevertheless, employment of the IDDI from MSG is indeed applicable for the required purpose in the current analysis of solely dust spatial pattern detection.

The IDDI calculations, applied to the MSG images taken during the field campaign, clearly show a major dust event in northern and central Africa. The elevated IDDI values over Senegal are also visible. The IDDI images show distinguishable changes in the emission sources and transport features during the different phases of the observations. For instance, Fig. 16 shows that the dust emissions during the first

phase of the event originated in south Algeria, Mauritania and Mali (examples of images from 29 and 30 March 2015). Weeks later, spatial patterns of the elevated IDDI are shifted to the south and show source regions in the south of Niger (Fig. 15c, d). Of course, attribution of emission sources mineralogy to aerosol spectral absorption is a complex task (Alfaro et al., 2004; Lafon et al., 2006; Di Biagio et al., 2017, 2019), and it is difficult to point to a specific source that could clearly explain the change in the aerosol absorbing properties observed in this study. However, the IDDI images clearly suggest a change in the dust transport regime that is consistent with the change in the dust optical properties.

### 4 Smoke episodes in December 2015–January 2016

During the SHADOW campaign, we had several strong smoke episodes in December 2015-January 2016, when air mass transported the products of biomass burning from the areas of intensive forest fires in the equatorial region. The relative humidity in the advected smoke layers varied from episode to episode, allowing evaluation of the RH influence on the smoke lidar ratios  $S_{355}$  and  $S_{532}$ . We should keep in mind, however, that for different days the smoke particles could have different chemical compositions, so evaluated RH dependence can be considered as semiquantitative only. The spatiotemporal evolution of the particle backscattering coefficient and depolarization ratio at 532 nm, during the 14–15 December 2015 smoke episode, is given in Fig. 17. The same figure also shows the water vapor mixing ratio, a convenient tracer to identify wet air mass arriving from the equatorial region. The smoke particles are usually contained in elevated layers, being mixed with dust (Veselovskii et al., 2018). The height ranges where the smoke particles are pre-



**Figure 16.** Infrared Difference Dust Index (IDDI) derived from MSG geostationary satellite at noontime. Panels (a) and (b) show elevated IDDI values, representing airborne dust emission and transport, over central and northern Sahara on 29 and 30 March 2015. The dust transport regime is visibly changed a few days later (17, 18 April 2015; panels **c**, **d**); the elevated IDDI values are shifted to the south. The areas in white are cloud-screened pixels; the IDDI is derived only over land due to the algorithm physical principle.

dominant can be identified by a low depolarization ratio and enhanced WVMR. For the event considered, the smoke particles are predominant above 1500 m after midnight.

The vertical profiles of  $\alpha_{355}$ ,  $\alpha_{532}$ ,  $S_{355}$ ,  $S_{532}$ ,  $A_{\alpha}$  and  $A_{\beta}$ together with the water vapor mixing ratio and the relative humidity, for 15 December (04:00-06:00 UTC), are shown in Fig. 18. The same figure presents decomposition of  $\beta_{532}$ into the dust and smoke contributions, based on depolarization measurements (Tesche et al., 2011). The smoke episodes are characterized by different relative humidity within the elevated layer. On 15 December, RH is about 40 % in the 1500-2100 m range and the ratio  $\frac{\beta_{532}^8}{\beta_{532}}$  is about 0.57 at 2000 m. The lidar ratio  $S_{532}$  decreases from 50 to 44 sr in a 1000–2000 m range, while  $S_{355}$  rises from 58 to 67 sr; thus  $S_{355}$  significantly exceeds  $S_{532}$ . We should recall that lidar ratios presented in Fig. 18 are attributed to a dust-smoke mixture. In principle, we can estimate  $S_{532}^{s}$  using Eq. (5), because the ratio  $\frac{\beta_{532}^2}{\beta_{532}}$  is available. The corresponding  $S_{532}^{s}$  profile obtained for assumed  $S_{532}^{d} = 50$  sr is shown in Fig. 18a (black line).  $S_{532}^{\rm s}$  is about 40 sr at 2000 m, and it is close to the measured  $S_{532}$  value. In the smoke layer, the extinction Ångström exponent  $A_{\alpha}$  can exceed  $A_{\beta}$ , due to the negative contribution of  $A_{\beta}^{d}$ . In particular, on 15 December  $A_{\alpha}$  is about 1.1, while  $A_{\beta}$  is close to zero.

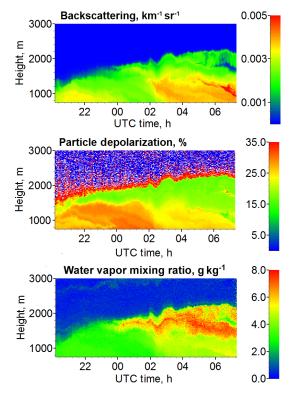
To estimate the dependence of smoke lidar ratios  $S_{355}$  and  $S_{532}$  on RH, five smoke episodes on 14–15, 15–16, 22–23 and 24–25 December 2015 and 19–20 January 2016 were analyzed.  $S_{532}$  and  $S_{355}$ , together with relative humidity and

the  $\frac{\beta_{532}^2}{\beta_{532}}$  ratio, are summarized, for these episodes, in Table 1. The heights chosen correspond to the values of relative humidity close to maximum. The calculated values of RH are characterized by high uncertainties, because lidar and sonde measurements are not collocated. Estimations of the corresponding uncertainties are also given by Table 1. The lidar ratio values from Table 1 are plotted in Fig. 19 as a function of RH. These plots, however, should be taken with care, because, due to variation in chemical composition and the aging processes, results may depend not only on RH. Moreover, the dust particles occurring in the elevated layers, as discussed, can introduce an additional ambiguity in the results. Nevertheless, Fig. 19 demonstrates a clear increasing trend of S with RH, at both wavelengths. From this figure, one can also conclude that  $S_{355}$  always exceeds  $S_{532}$  and that  $S_{532}$  for smoke can be as small as  $44 \pm 5$  sr at low humidity. The small values of  $S_{532}$  for the fresh smoke (about 40 sr) were also reported by Burton et al. (2012).

To compare our observations with the lidar ratios used in the MERRA-2 model, we have also performed the simulation of  $S_{532}^{\rm OC}$  (RH) and  $S_{355}^{\rm OC}$  (RH) dependence for organic carbon (OC) based on the particle parameters and hygroscopic growth factor from the MERRA-2 model. In MERRA-2 the organic carbon is the main component of the biomass burning products. The imaginary part of the OC increases in the UV due to the presence of brown carbon (BrC), which is a subset of organic carbon with strong absorption in the UV region (Bergstrom et al., 2007; Torres et al., 2007). The majority of BrC is emitted into the atmosphere through

**Table 1.** Lidar ratios  $S_{355}$  and  $S_{532}$  for five smoke episodes in December 2015–January 2016 and the corresponding relative humidity RH. The table also provides the height and temporal interval of observations. The contribution of the smoke particles to the total backscattering  $\frac{\beta_{532}^s}{\beta_{532}}$  is derived from depolarization measurements.

Date	Height (m)	Time (UTC)	$\frac{\beta_{532}^{\mathrm{s}}}{\beta_{532}}$	RH (%)	S <sub>355</sub> (sr)	S <sub>532</sub> (sr)
15 Dec	2000	04:00-06:00	0.57	$42\pm8$	$67 \pm 7$	$44 \pm 5$
15 Dec	1850	19:20-20:30	0.57	$25 \pm 6$	$62 \pm 6$	$50 \pm 5$
23 Dec	2250	05:00-07:00	0.65	$65 \pm 13$	$76 \pm 8$	$56 \pm 6$
24 Dec	3200	19:00-23:00	0.66	$75 \pm 14$	$76 \pm 8$	$62 \pm 6$
20 Jan	4500	01:00-07:00	0.8	$85 \pm 15$	$80 \pm 8$	$66 \pm 7$



**Figure 17.** Spatiotemporal distributions of aerosol backscattering coefficient  $\beta_{532}$ , particle depolarization ratio  $\delta_{532}$  and water vapor mixing ratio during smoke episode on the night 14–15 December 2015.

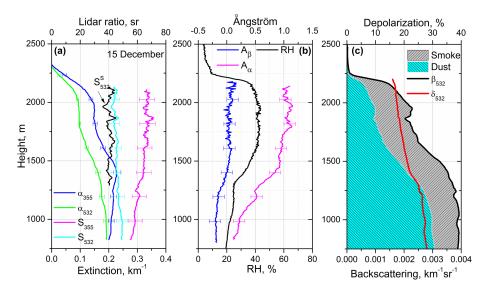
low-temperature, incomplete combustion of biomass. In the newest development of GEOS, biomass burning OC is now emitted as a new BrC tracer species that uses  $Im_{532} = 0.009$  and  $Im_{355} = 0.048$  values (Hammer et al., 2016). Thus, the spectral behavior of the imaginary part of the organic carbon refractive index depends on the contribution of the BrC fraction to the primary organic carbon and on the physical-chemical processes in the smoke layer during its transportation. As a result, the spectral dependence of Im can present strong variations. In our study, the computations at 355 nm were performed for four values of the imaginary part of dry particles,  $Im_{355} = 0.048, 0.03, 0.02, 0.01$ . At 532 nm two val-

ues,  $Im_{532} = 0.005$  and 0.009, were considered. The parameters of the dry particle size distribution, the real part of the CRI and the hygroscopic growth factor used in computations are given in Veselovskii et al. (2018). The particles are assumed to be homogeneous spheres, and an increase in the volume for every RH value (calculated from the growth factor) occurs due to water uptake. Thus both the real and the imaginary part of the CRI depend on RH.

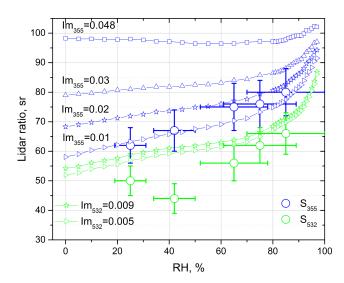
The results of the simulations, shown in Fig. 19, demonstrate a strong dependence of the organic carbon lidar ratio on the imaginary part of dry particles and on the relative humidity. For  $Im_{355} = 0.048$ , for all RH,  $S_{355}$  is above 95 sr, which strongly exceeds the observed values. For lower  $Im_{355}$ , the  $S_{355}$  (RH) dependence is more pronounced, and for  $Im_{355}$  within the range 0.01–0.02, computed  $S_{355}$  values are close to observed values. Computed  $S_{532}$  values at low RH exceed the measured ones, but for RH > 70 %, agreement between measurements and GEOS-assumed optical properties for OC becomes reasonable.

### 5 Summary and conclusion

Our study shows the impact of the imaginary-part variation on the lidar-derived dust properties. In contrast to extinction, the backscattering coefficient and so the lidar ratio are sensitive to the imaginary part of the CRI. Hence, the  $S_{355}/S_{532}$ ratio can be an indicator of the imaginary refractive index enhancement in the UV. Measurements performed during the SHADOW campaign, in dust conditions, show a correlation between the decrease in Im<sub>440</sub>, derived from AERONET observations, and the decrease in the lidar-derived  $S_{355}/S_{532}$ ratio. Namely, in the second half of April 2015,  $S_{355}/S_{532}$  decreased from 1.5 to 1.0, when Im<sub>440</sub> decreased from 0.0045 to 0.0025. Our numerical simulations confirm that observed  $S_{355}/S_{532}$  (ratio close to 1.5) and  $A_{\beta}$  (value close to -0.75) can be due to spectral variation in the imaginary part, attributed to iron oxides contained in dust particles. The simulations were performed for the model of randomly oriented spheroids; however, we should recall that an increase in the particle lidar ratio with the imaginary part should occur for any particle shape.



**Figure 18.** Vertical profiles of (a) extinction coefficients ( $\alpha_{355}$ ,  $\alpha_{532}$ ) and lidar ratios ( $S_{355}$ ,  $S_{532}$ ); (b) extinction and backscattering Ångström exponents ( $A_{\alpha}$ ,  $A_{\beta}$ ) at 355–532 nm and relative humidity RH; (c) contribution of dust and smoke to  $\beta_{532}$  together with particle depolarization ratio  $\delta_{532}$  on 15 December (04:00–06:00 UTC). Black line in plot (a) shows the lidar ratio of smoke  $S_{532}^{s}$  calculated from Eq. (5).



**Figure 19.** Modeled lidar ratios of organic carbon at 355 and 532 nm (line + symbol) as a function of the relative humidity for the particle parameters used in the MERRA-2 model. At 355 nm results are given for four values of the imaginary part of dry particles:  $Im_{355} = 0.048$ , 0.03. 0.02, 0.01. At 532 nm two values,  $Im_{532} = 0.009$  and 0.005, are considered. The scattered symbols (circles) show the lidar ratios ( $S_{355}$ ,  $S_{532}$ ) observed during five smoke episodes from Table 1.

Thus, April 2015 observations suggest the presence of different dust types, characterized by the distinct spectral dependence of  $Im(\lambda)$ . The analysis of backward trajectories and the IDDI derived from the MSG geostationary satellite confirms different air mass and dust particles transport features in the beginning and at the end of April. Hence, the

observed variations in  $S_{355}/S_{532}$  can be related to the source region mineralogy. During April, the particle depolarization systematically exceeded 30%; therefore no discrimination between different types of dust was possible. Dependence of the  $S_{355}/S_{532}$  ratio on dust origin, in particular, could explain, why during SAMUM experiments no significant spectral dependence of the lidar ratio was observed.

The results presented in this study demonstrate also that, for the selected temporal interval, the dust lidar ratios may present significant variation with height. Dust of different size and mineralogical composition can have different deposition rates; hence, the complex refractive index can be height-dependent. For instance, on 1 April, the  $S_{532}$  decreased with height from 60 to 50 sr within a 1000-3000 m range, while the depolarization ratio exceeded 30%. The analysis of this episode showed that variation in the lidar ratio is entirely attributed to variations in dust characteristics; the smoke aerosol contribution was insignificant. The data also demonstrate that a seemingly uniform dust layer may have quite a complex height variation. The results therefore suggest the relevance of including a varying mineralogy in radiative and climatic modeling of desert dust impacts. Dust mineralogy should also be taken into account when the possibility of the particle microphysical parameters' characterization on a base of multiwavelength lidar measurements is analyzed (Perez-Ramirez et al., 2019, 2020).

During December–January, the dry season in western Africa, our observations allowed in addition the analysis of biomass burning aerosol properties. These particles are a product of the seasonal forest fires and intensive agricultural waste combustion and can contain a substantial number of organic compounds, characterized by an enhanced imag-

inary part in the UV range (so-called BrC). For this aerosol type, the  $Im(\lambda)$  dependence should increase the lidar ratio at 355 nm and influence  $S_{355}/S_{532}$ . The smoke particles can also be hydrophilic, and the lidar ratio can therefore exhibit a strong dependence on RH. Several strong smoke episodes were observed during the SHADOW campaign. While we were able to evaluate the RH profiles, the dependence of the smoke lidar ratio with RH has been estimated. The results obtained should be taken as semiqualitative only, due to possible variation in smoke particle composition from episode to episode and due to the presence of dust particles. Nevertheless, the results clearly demonstrate an increase in  $S_{532}$  from  $44\pm5$  to  $66\pm7$  sr and of  $S_{355}$  from  $62\pm6$  to  $80\pm8$  sr, when the RH increased from 25% to 85%.

We would like to conclude that the multiwavelengths Raman and depolarization lidar measurements in western Africa enabled quite unique and comprehensive profiling of dust and smoke spectral absorption properties. The results demonstrated a high variability in the lidar ratio and the presence of its spectral dependence. Our study is one of the first attempts to track aerosol composition variability using lidar measurements and to understand the mechanism underlying the observed variations. However, the results presented were obtained for a single region in western Africa. It is important to repeat such studies at different locations around the world, including the Middle East, central and east Asia, Australia, and North America, in order to improve our knowledge on real-world dust optical properties needed in climate-relevant atmospheric modeling.

Data availability. Lidar measurements are available upon request (philippe.goloub@univ-lille.fr).

Author contributions. IV processed the data and wrote the paper. QH and TP performed the measurements. PG supervised the project and helped with paper preparation. MK developed software for data analysis. YD and ML analyzed the satellite data, and PC provided MERRA-2 simulations.

*Competing interests.* The authors declare that they have no conflict of interest.

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