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Deuteration of ammonia with D atoms on oxidized partly ASW covered graphite surface

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Abstract

The deuteration of ammonia by D atoms has been investigated experimentally in the sub-monolayer regime on realistic analogues of interstellar dust grain surfaces. About 0.8 monolayer of solid NH₃ was deposited on top of an oxidized graphite surface held at 10 K, partly covered with ASW ice. Ammonia ice is subsequently exposed to D atoms for different exposure times using a differentially pumped beam-line. The deuteration experiments of ammonia were carried out by mass spectroscopy and temperature programmed desorption (TPD) technique. The experimental results showed the formation of three isotopologue ammonia species by direct exothermic H-D substitution surface reactions: NH₃ + D → NH₂D + H, NH₂D + D → NHD₂ + H, and NHD₂ + D → ND₃ + H. The formation of the deuterated isotopologues NH₂D, NHD₂, and ND₃ at low surface temperature (10 K) is likely to occur through quantum tunneling process on the oxidized graphite surface. A kinetic model taking into account the diffusion of D atoms on the surface is developed to estimate the width and the height of the activation energy barriers for the successive deuteration reactions of ammonia species by D atoms. Identical control experiments were performed using CH₃OH and D atoms. The deuteration process of solid methanol is ruled by H abstraction and D addition mechanism, and is almost five orders of magnitude faster than ammonia deuteration process.

1. Introduction

Ammonia molecules are detected in the gas phase of molecular clouds: Taurus Molecular Cloud-1 (TMC1-N) [1], with relatively high abundances 10⁻⁷-10⁻⁸ respective to H₂ molecules [2]. Solid NH₃ has been detected through infrared absorption in different astrophysical environments: high-mass protostars [3], low-mass protostars [4], comets [5], and in dense molecular clouds [6]. The interstellar grain mantles in dense molecular clouds are predominantly composed of H₂O ice, combined with other molecules such as (CO, CO₂, NH₃, H₂CO, and CH₃OH) [7, 4]. The abundance of ammonia (NH₃) in the icy mantles is 1 to 15 % with respect to water (H₂O) ice [8, 3], while in the cold dust envelopes of young stellar objects, the ammonia ice fraction is 5 % or less [6]. In comets, ammonia is present at the 1 % level relative to water ice [5]. Deuterated ammonia NH₂D was first detected by Rodriguez Kuiper et al [9] in high temperature molecular clouds such as Orion-KL Nebula region (T=50-150 K). NH₂D molecules have been also observed in many sources towards dark molecular clouds [10], galactic protostellar cores [11], and interstellar dense cores (L134N) [12]. The [NH₂D]/[NH₃] ratio in gas phase varies from 0.02 to 0.1. These abundance ratios are larger than the cosmic abundance of elemental deuterium relative to hydrogen

(D/H), which is expected from the Big-Bang nucleosynthesis to be 1.5 × 10⁻⁵ [13]. Observations in low mass protostellar cores showed the highest [NH₂D]/[NH₃] ratios (0.3), indicating that deuterium fractionation of ammonia increases towards protostellar regions [14]. Chemical models explained this high fractionation ratio by the gas-phase ion-molecule chemistry with depletion of C, O and CO from the gas phase [12, 14]. Doubly deuterated ND₂H ammonia has been also detected for the first time in cold, 10 K, dense cores L134N by Roueff et al. [15]. The expected fractionation ratio [ND₂H]/[NH₃] from models is 0.03 [14]. The Caltech Submillimeter Observatory (CSO) has detected the triply deuterated ammonia, ND₃, through its J_K emission transition near 310 GHz [16] in cold clouds (10 K). The observed [ND₃]/[NH₃] ratio in very cold clouds of gas and dust is found to be closer to 0.001. Such a high isotopic ratio between ND₃ and NH₃ suggested that the deuteration of NH₃ is likely to occur by ion-molecule reactions in the gas phase, in which deuterium transfer reactions are much faster than proton transfer [16, 17]. Theoretical models of pure gas-phase chemistry [18, 10, 19] explained relatively well the abundances of simply and multiply deuterated ammonia molecules in dense cores. According to Tielens et al. [20], grain surface chemistry would also build deuterated molecules by deuteration process on grain mantles with D atoms. The trapped deuterated species on grains are eventually released into the gas phase due to the heating of a close star in the formation stage. Recent chemical models of cold dark clouds [21] have shown that desorption of species into the gas phase via thermal evaporation is negligible

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for dark clouds with temperatures of 10 K.

The deuteration experiments of solid ammonia by D atoms has been already performed by two astrophysical groups; the Watanabe group (Nagaoka et al. [22]), and the Leiden group (Fedoseev et al. [23]) using mainly infrared spectroscopy. The experimental studies of Nagaoka et al. [22] have shown an efficient deuteration of CH₃OH ice by D atoms addition at low surface temperature. The deuterated methanol species are formed by direct H-D exchange and the combination of abstraction and addition reactions, through tunneling quantum process. These authors have reported that no deuterated species of ammonia are observed in the exposure of pre-deposited NH₃ ice to D atoms at temperatures below 15 K. Even the experimental results of Fedoseev et al. [23] have also shown that the deuteration of solid NH₃ by D atoms did not take place at temperature lower than 15 K, by depositing D atoms on ammonia ice, or by performing co-deposition experiments of NH₃ molecules with D atoms on gold cold surface.

Based on the previous experimental results of Fedoseev et al. [23] and Nagaoka et al. [22], someone wonders about the dramatic difference observed in the deuteration of ammonia and methanol by D atoms in the solid phase. If these authors [23, 22] did not observe the deuteration of the NH₃ by D atoms in their experiments, this is probably because of the very high activation energy barrier of the reaction NH₃+D in the solid phase in comparison to that of CH₃OH+D. The value of the activation energy barrier of the reaction (NH₃ + D → NH₂D + H) has been estimated from earlier experimental [24] and theoretical [25] works in the gas phase to be 11 kcal · mol⁻¹ or 46 kJ · mol⁻¹. While the activation energy barrier of the abstraction reaction (CH₃OH + D → CH₂OH + HD), has been reported from gas phase estimations to be lower than that of ammonia (27 kJ · mol⁻¹) [2]. But up to now, there is no laboratory studies providing activation energy barriers for the reaction CH₃OH + D and that of NH₃ + D in the solid phase.

It is obvious that laboratory experiments are important for understanding the deuteration reactions occurring on the cold grain surfaces between condensed molecules and the impinging D atoms. However, some factors related to the gas flux of the deuterium atoms, the thickness of the ices on the grains, and the fluences of the atomic species on the surface, may affect the progress and the evolution of these reactions. In the case of the the previous works of Nagaoka et al. [22], and Fedoseev et al. [23], the authors have performed experiments in the multi-layer regime by covering the aluminium surface with 10 ML of solid NH₃ [22], and the the gold surface by 50 ML of ammonia ice [23], and irradiating the corresponding ices with D-flux of 1.4×10^{13} atoms · cm⁻² · s⁻¹, and $\times 10^{14}$ atoms · cm⁻² · s⁻¹, respectively (see Table 1). First of all, the use of a high flux of D atoms in their experiments favors the recombination reactions D+D on the surface or in the bulk of the ices, and reduces therefore the reaction efficiency of D atoms with the adsorbed CH₃OH and NH₃ species. However, in the experiments of Nagaoka et al. [22], the deuteration reaction (CH₃OH + D) seems to be not affected by the high flux of D atoms. This is probably because the activation energy barrier of H-D exchange reaction between D and CH₃OH is lower than that between D and NH₃.

On the other hand, as reported by Fedoseev et al. [23], the use of a thick layer of ammonia ice favors the formation of hydrogen bonds (N-H), that can strength the interaction NH₃-NH₃ molecules, and prevent the D-H exchange between D atoms and adsorbed NH₃ molecules.

In this work, we performed deuteration experiments of solid NH₃ by D atoms in the sub-monolayer regime, and with low D-flux, on an oxidized highly oriented pyrolytic graphite (HOPG) surface, partly covered with an amorphous solid water (ASW) ice held at 10 K of less than 0.5 ML of thickness, simulating water ice contaminations. We deposited only a fraction of one monolayer of solid ammonia (0.8 ML) on the substrate to study the effect of the grain surface on the efficiency of the deuteration reaction between D atoms and the adsorbed NH₃ molecules. In this work, we considered that the physisorption of species on the oxidized HOPG surface dominates the chemisorption process. We also used low D-flux in comparison to the previous works [23, 22] to reduce the recombination efficiency of D atoms on the surface, and increase the probability of the H-D substitution reaction. As shown in the Table 1, at our experimental conditions, even by reducing the D atoms Fluence by factors 100 and 10, with respect to those of Nagaoka et al. [22] and Fedoseev et al. [23], respectively, the total amount of D atoms (53.5 ML) sent on the surface seems to be sufficient for the D-fractionation of solid ammonia, and the formation of the deuterated species NH₂D, NHD₂, and ND₃.

For comparison, similar D atoms addition experiments have been performed with CH₃OH molecules to corroborate the findings of Nagaoka et al. [26], and validate the deuteration method governed by the abstraction-addition mechanism. D atom addition and H atom abstraction may not be the only mechanism to deuterate molecules on ices. Direct H-D substitution reactions could also proceed at low temperatures to fractionate the astrophysical molecules.

The paper is organized as follows: in section 2, we describe the experimental setup and explain the procedures of the deuteration experiments; section 3 presents the experimental results for NH₃ + D and CH₃OH + D reactions, and in the section 4, we propose a kinetic model to estimate the activation energy barriers of the successive H-D substitution reactions. We make some concluding remarks in the final section.

2. Experimental

The experiments were performed with the FORMOLISM (FORMATION of MOLECULES in the InterStellar Medium) apparatus. The set-up is dedicated to studying the reactivity of atoms and molecules on surfaces of astrophysical interest, under the conditions of temperatures and pressures similar to those encountered in the ISM. The experimental setup is briefly described here and more details are given in a previous work [27]. The apparatus is composed of an ultra-high vacuum (UHV) stainless steel chamber with a base pressure of about 10⁻¹¹ mbar. The sample holder is located in the center of the main chamber. It is thermally connected to a cold finger of a closed-cycle Helium cryostat. The temperature of the sample is measured in the range of 6 K-350 K by a calibrated silicon diode

Table 1: Comparison between the experimental conditions and the results of the system (NH₃ + D) for different works and references.

Article	D Fluence	D Thickness	D-Flux	NH ₃ Thickness	results
	atoms · cm ⁻²	ML	atoms · cm ⁻² · s ⁻¹	ML	
Nagaoka et al. [22] (2015)	≤10 ¹⁸	≤1000	10 ¹⁴	10	no deuteration
Fedoseev et al. [23] (2015)	8 × 8 ¹⁶ -3 × 10 ¹⁷	≤100	1-4 × 10 ¹³	50	no deuteration
This work	≤5.35 × 10 ¹⁶	≤53.5	3.7 × 10 ¹²	0.8	deuteration

clamped to the sample holder. This one is made of a 1 cm diameter copper block which is covered with a highly oriented pyrolytic graphite (HOPG, ZYA-grade) substrate. The HOPG is a model of an ordered carbonaceous material mimicking interstellar dust grains analogues in astrophysical environments. It is characterized by an arrangement of carbon atoms in a hexagonal lattice. The HOPG grade (10 mm diameter × 2 mm thickness) was firstly dried in an oven at about 100 °C during two hours, and then cleaved several times using "Scotch tape" method at room temperature to yield several large terraces (micron scale) that contain limited defects and step edges. The HOPG was cleaved in air immediately prior to being inserted into the vacuum chamber. It was mounted directly onto the copper finger by means of a glue (ARMECO Product INC CERAMA BOND 571-P). In chamber, the HOPG sample was annealed to 300 K under UHV to remove any contaminants. In this work, we used an oxidized HOPG sample, which has been preliminary exposed to oxygen atomic beam under UHV for several exposure doses, and then warmed-up from 10 K to 300 K to desorb oxygen and other species from the substrate, mainly water molecules. The oxidation phase was achieved after the saturation bonds of the surface, the defects and the step edges of the sample. This behavior was deduced when there is no modification in the Thermally Programmed Desorption profiles of the adsorbates. Prior oxidation of the HOPG is expected to give stable surface, where the structure cannot be modified by other adsorbates. However, defects and/or some functional groups may be created on the surface during the oxidation phase. These active sites have not been identified by our diagnostics using mass spectrometry and infrared spectroscopy.

FORMOLISM is equipped with a quadrupole mass spectrometer (QMS) which allows routinely the simultaneous detection of several species in the gas phase by their masses. The QMS can be placed either in front of the surface for the detection of species desorbed into the gas phase during the warming-up of the sample, or in front of the beam-line for the characterization and the calibration of the NH₃, CH₃OH, and D atoms beams. The experimental setup is also equipped with a Fourier transform infrared spectrometer (FTIR) for the in-situ solid phase measurements by reflection absorption infrared spectroscopy (RAIRS) in the spectral range 4000-700 (cm⁻¹ [28].

The D atomic jet is prepared in a triply differentially pumped beam-line aimed at the sample holder. Its is composed of three vacuum chambers connected together by tight diaphragms of

3 mm diameters. The beam-line is equipped with a quartz tube with inner diameter of 4 mm, which is surrounded by a microwave source cavity for the dissociation of D₂ molecules. When the microwave source (Sairem) is turned on, the cavity is cooled down with a pressurized air jet, and D atoms are produced from the D₂ molecular plasma. The D₂ plasma is generated by a microwave power supply coupled into a Surfatron cavity operating at 2.45 GHz and producing up to 300 W. The warm D atoms undergo several collisions with the inner walls of the tube, and finally thermalize at the room temperature of about 350 K before they reach the surface. However, the charged particles composed of excited atoms, ions and electrons, produced in the plasma quickly recombine within the tight quartz tube [29, 30]. Because of the high micro-wave frequency, the hot energetic particles cannot leave the discharge pipe, as reported in some astrophysical laboratory works [29].. With our direct beam-line configuration, even if some residual photons can be sent on the surface during the exposure of the D-beam, the fraction of these particles is negligible, and they cannot affect the adsorbed species.

The deuterium beam dissociation rate, measured with the quadrupole mass spectrometer from the D₂ signals (m/z=4) during the discharge (ON) and the discharge (OFF) of the microwave source is calculated from the following relationship $\tau = \frac{D_2(\text{OFF}) - D_2(\text{ON})}{D_2(\text{OFF})}$. In this work, the dissociation rate τ of D₂ beam reaches a high value of 85 % with an effective microwave power of 50 W.

The flux of the dissociated D atoms coming from the gas phase and hitting the surface is $\Phi_{D,\text{ON}} = (3.7 \pm 0.5) \times 10^{12}$ atom · cm⁻² · s⁻¹. It is defined as $\Phi_{D,\text{ON}} = 2\tau \Phi_{D_2,\text{OFF}}$, where $\Phi_{D_2,\text{OFF}} = (2.2 \pm 0.4) \times 10^{12}$ molecule · cm⁻² · s⁻¹ is the flux of D₂ beam before running the microwave discharge. The flux of D₂ molecules coming from the beam-line is determined by the so-called the King and Wells method [31], which is generally used to evaluate the sticking coefficient of particles incident on a cold surface. This method consists to measure with the mass spectrometry the indirect D₂ signal in the vacuum chamber in the real time during the exposure of D₂ on the oxidized HOPG surface. Φ_{D_2} is calculated from the ratio between the exposure dose of D₂ molecules that saturate the graphitic surface, expressed in (molecule · cm⁻²) by the corresponding exposure time of D₂ on the surface, expressed in (second). According to the estimation made by Amiaud et al. [27], a compact ice layer begins to saturate after an exposure to 0.45 ML of D₂ (i.e. 0.45×10^{15} molecule · cm⁻²). More detail description and er-

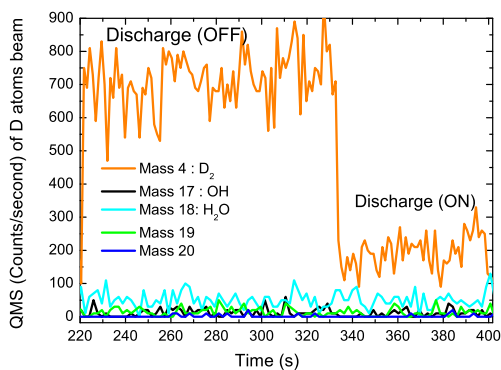


Figure 1: The QMS signals (in counts/seconds) as a function of the time (s) of $m/z=2$ (D), $m/z=4$ (D_2), $m/z=17$ (NH_3), $m/z=18$ (NH_2D , ND_2H), $m/z=19$ (ND_2H), and $m/z=20$ (ND_3), given by the QMS, placed in front of the D beam after the deposition of the NH_3 molecules on the oxidized HOPG surface using the same beamline.

rors about the estimation of D-flux are given in the reference [27].

In this study, we have deposited all species (NH_3 , CH_3OH , H_2O molecules, and D atoms) by using only one beam-line, oriented at 45° relatively to the surface of the sample. That guarantees a quasi-perfect match between the effective areas on which particles are deposited.

In our NH_3+D experiments, the beam-line is pumped off to evacuate the residual gas of ammonia species, after the deposition phase of NH_3 ice on the cold surface. Then the D atoms are generated in the same beam-line by the microwave dissociation of D_2 molecules. We have checked with the QMS, placed in front of the beam-line, that no deuterated species (NH_2D , ND_2H , HDO , ND_3 , D_2O) and radicals (ND_2 , OD , OH) contaminants are coming from the D beam. Figure 1 shows the signal of ($m/z=4$) before the dissociation of deuterium molecules (discharge OFF) and during the dissociation phase (discharge ON). We notice that there is no increase in the signal of mass 18 during the discharge ON, which may correspond to NH_2D and ND_2 species formed from NH_3 and D atoms within the beam-line. The small signal of mass 18 is the background signal of H_2O molecules contaminants in the main chamber. Moreover, the absence of the signals ($m/z=19$) and ($m/z=20$) excludes any possible formation of NHD_2 and ND_3 species in the D beam.

Ammonia and methanol ice films, with a thickness of 0.8 ML were grown on the oxidized HOPG surface held at 10 K by beam-line vapor deposition of NH_3 molecules (from Eurisotop bottle with 99.9 % purity) and CH_3OH molecules (from liquid methanol with 99.5 % purity). The monolayer surface coverage corresponds to the number density of molecules that populate 10^{15} sites on the surface. It is defined as 1 ML = 10^{15} molecules \cdot cm $^{-2}$. In this work, the fluxes of ammonia and methanol species that hit the surface, are defined as the amounts of these species that saturate the surface per unit time ($\Phi = \frac{\text{exposure dose for 1 ML}}{\text{exposure time for 1 ML}}$). The values of the fluxes are found to be $\Phi_{NH_3} = 2.1 \times 10^{12}$ molecules \cdot cm $^{-2} \cdot$ s $^{-1}$ and Φ_{CH_3OH}

$$= 1.7 \times 10^{12} \text{ molecules} \cdot \text{cm}^{-2} \cdot \text{s}^{-1}.$$

Because water is always present as contaminant in the ultra-high vacuum chamber, and can be condensed on the cold surface at 10 K, we have performed experiments to study the effect of the water ice on the deuteration of solid ammonia. In order to simulate the small amount of the water ice that can be condensed on the surface during the exposure phase of the reactants at 10 K, we have deposited a very thin film of porous amorphous solid water (ASW) ice with ~ 0.5 ML of thickness on the oxidized graphite surface at 10 K, by H_2O vapor deposition during 5 minutes, using the same beam-line as that for ammonia and D atoms. The water vapor was obtained from deionized water which had been purified by several pumping cycles under cryogenic vacuum. We have estimated the thickness in ML of the amorphous water ice film grown on the surface at 10 K by beam-line vapor (H_2O) deposition using the reflection absorption infrared spectroscopy (RAIRS). We have deposited water ice on the surface at 10 K, for different exposure times, then we have recorded the RAIR spectra, and measured the integrated areas $\int \nu d\nu$ in (cm^{-1}) below the IR absorption bands of water ice at about 3200 cm^{-1} . Using the formulae $S = \frac{\text{Ln}10 \int \nu d\nu}{N}$ [32], where N is the column density of water ice in ($\text{molecule} \cdot \text{cm}^{-2}$), and S is the strength band of H_2O at 3200 cm^{-1} in ($\text{cm} \cdot \text{molecule}^{-1}$), we have estimated the exposure time required to form 1 ML surface coverage of ASW ice at 10 K. The first monolayer of amorphous water ice covering the surface at 10 K is reached after about 10 minutes of water deposition time with a flux of 10^{12} molecules \cdot cm $^{-2} \cdot$ s $^{-1}$. The absorbance value of the corresponding H_2O infrared band at 3200 cm^{-1} is very low ($AB=0.0005$). Because of the low surface coverage of solid NH_3 (0.8 ML) on oxidized graphite, and the low absorbance ($AB=0.0002$) of the infrared band of ammonia at 1106 cm^{-2} for the vibrational mode ν_2 , the deuteration experiments of ammonia by D atoms on the oxidized graphite surface has been analyzed in this work only by TPD-QMS spectroscopic method.

For the deuteration experiments of solid NH_3 (or CH_3OH), we firstly deposited 0.8 ML of NH_3 (or CH_3OH) ices on the oxidized HOPG surface held at 10 K, and then we exposed the films of ammonia (or methanol) to D atomic beam at the same surface temperature. After the exposure phases, we used TPD technique by warming-up the sample from 10 K to 210 K with a linear heating rate of $0.17 \text{ K} \cdot \text{s}^{-1}$, until the sublimation of the ices from the surface. The species desorbed into the gas phase are then detected and identified through mass spectrometry.

3. Results

3.1. Co-deposition of H_2O-NH_3

In order to study the effect of the water on the adsorption-desorption of ammonia molecules, two kinds of experiments have been performed on the graphite surface held at 10 K using H_2O and NH_3 molecules. In the first experiment, we grow (~ 0.5 ML) of amorphous solid water (ASW) ice on the oxidized graphite surface by exposing the sample held at 10 K to H_2O water beam during 5 minutes. In the second experiment,

we deposited 0.8 ML of solid NH_3 on top of 0.5 ML surface coverage of water (H_2O) ice grown on the oxidized graphite surface at 10 K.

Figure 2, top panel shows the TPD signals of masses ($m/z=17$) for pure H_2O and mixed ($\text{H}_2\text{O} - \text{NH}_3$) films. These signals are compared to the masses ($m/z=16$), ($m/z=15$), and ($m/z=14$) of the film ($\text{H}_2\text{O} - \text{NH}_3$). Figure 2, bottom panel shows the TPD signals of masses ($m/z=18$) for both H_2O and ($\text{H}_2\text{O} - \text{NH}_3$) films.

The TPD peak of the cracking pattern (OH^+) ($m/z=17$) by the electron impact of the ionized H_2O molecules is centered at 147 K (Figure 2, top panel), and the maximum of the TPD peak of pure H_2O ($m/z=18$) appears at 147 K (Figure 2, top panel). For the film ($\text{H}_2\text{O} - \text{NH}_3$), Figure 2, top panel shows a strong TPD peak ($m/z=17$) at 106 K, which corresponds to the desorption of pure ionized NH_3 molecules from the surface, and a small TPD peak at 147 K ($m/z=17$), which is attributed to the cracking pattern OH^+ of the ionized H_2O molecules by the electron impact. For the same mixed film, the TPD peak at 147 K displayed in the Figure 2, bottom panel is that of ionized H_2O molecules ($m/z=18$). We have checked that no cracking pattern NH_2^+ ($m/z=16$), NH^+ ($m/z=15$) and N^+ ($m/z=14$) of the ionized NH_3 molecules desorb at 147 K (Figure 2, top panel).

For the two H_2O and ($\text{H}_2\text{O} - \text{NH}_3$) ices, the Figure 2, bottom panel shows only one desorption peak at around 147 K, which is slightly more intense for $\text{H}_2\text{O} - \text{NH}_3$ ice (in magenta line) than pure H_2O ice (in black line). The small increase in the area below the desorption curve of $m/z=18$ at 147 K for the $\text{H}_2\text{O} - \text{NH}_3$ ice results from the instability of the H_2O flux.

3.2. Exposure of NH_3 and D atoms on graphite surface

In the first experiment, we prepared a film 1 (15.5 ML (D) + 0.8 ML (NH_3)), by exposing firstly the oxidized HOPG surface at 10 K, to D beam for 15.5ML surface coverage, and then to 0.8 ML of NH_3 ice at the same surface temperature 10 K. In the second experiment, the film 2 (0.8 ML (NH_3) + 15.5 ML (D)), is prepared by exposing 15.5 ML of D atoms on top of 0.8 ML of NH_3 ice pre-deposited on the oxidized HOPG surface at 10 K. Two TPD control experiments were also performed in addition to the previous ones, by depositing separately 0.8 ML of solid NH_3 , and 15.5 ML of D atoms on the oxidized HOPG surface. The TPD curves of all the experiments are displayed in the Figure 4 for the masses 17, 18, 19 and 20. The TPD curves of the two films of NH_3 (0.8 ML) and that of D atoms (15.5 ML) deposited separately on the surface, showed small peaks at around 152 K, for masses 18, 19 and 20 (see panels (b), (c) and (d) of Figure 3). These peaks correspond to the desorption of water contaminations, such as H_2O ($m/z=18$), HDO ($m/z=19$), and D_2O ($m/z=20$). These water impurities came either from the beam-line during the exposure phases of NH_3 and D atoms, or from the ultra-high vacuum chamber. In the case of the film 1, where D atoms are deposited before NH_3 on the surface, the TPD curve of mass 18 in the Figure 3b shows a small desorption peak at around 96 K overlapping a large peak at 130 K-150 K. While in the panels (c) and (d) of the Figure 3, we observe only the desorption peaks at 152 K for the masses

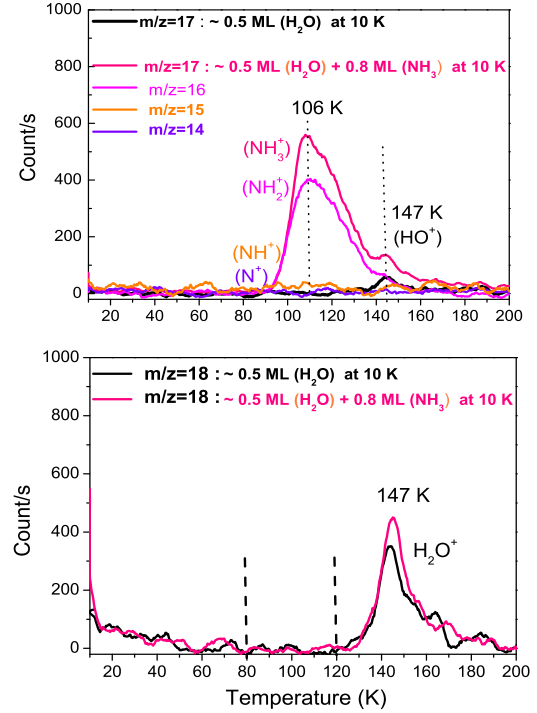


Figure 2: TPD signals of mass $m/z=17$ (Top panel) and mass $m/z=18$ (bottom panel) between 10 K and 200 K. Black curve: (~ 0.5) ML of pure porous amorphous water (H_2O) ice deposited on the oxidized graphite surface at 10 K. Magenta curve: 0.8 ML of NH_3 ice deposited on top of (~ 0.5) ML of H_2O ice pre-deposited on the oxidized graphite surface at 10 K.

19 and 20, respectively. The desorption peaks at the higher temperatures 150 K and 152 K for the masses 18, 19 and 20 are likely to be originated from water impurities H_2O , HDO and D_2O , respectively. Based on the computational results of Burke et al. [33], the sticking coefficients of the impinging D atoms coming from the gas phase at room temperature onto graphite and ASW ice (held at 10 K) is 90 % and 60 %, respectively. However, the experimental studies of Matar et al. [34] for the sticking coefficient of D atoms on the non-porous ASW ice is estimated to be 30 %. Since our substrate is constituted of an oxidized HOPG, partly covered with ASW ice contaminants (H_2O , HDO and D_2O), most of the D atoms exposed on the substrate for the (film 1), will stick both on the oxidized graphite and on the water surface adsorption sites. These atoms promptly form D_2 molecules by $\text{D} + \text{D}$ surface recombination, either by Langmuir-Hinshelwood mechanism based on the diffusion of two adsorbed D atoms on the surface, or via Eley-Rideal abstraction reactions between adsorbed D atoms and incoming D atoms from the gas phase [35, 36]. Moreover, the experimental and the theoretical studies of Horneker et al. [37] have revealed a possible route for D_2 formation on the HOPG surface through D adsorbate clusters. The D_2 molecules formed on the graphitic surface cannot react with the NH_3 molecules adsorbed on the surface, and cannot be therefore involved in the formation of the new isotopic species of ammonia. We suggested that water contaminants present on the surface, such as HDO , D_2O , may react with the deposited

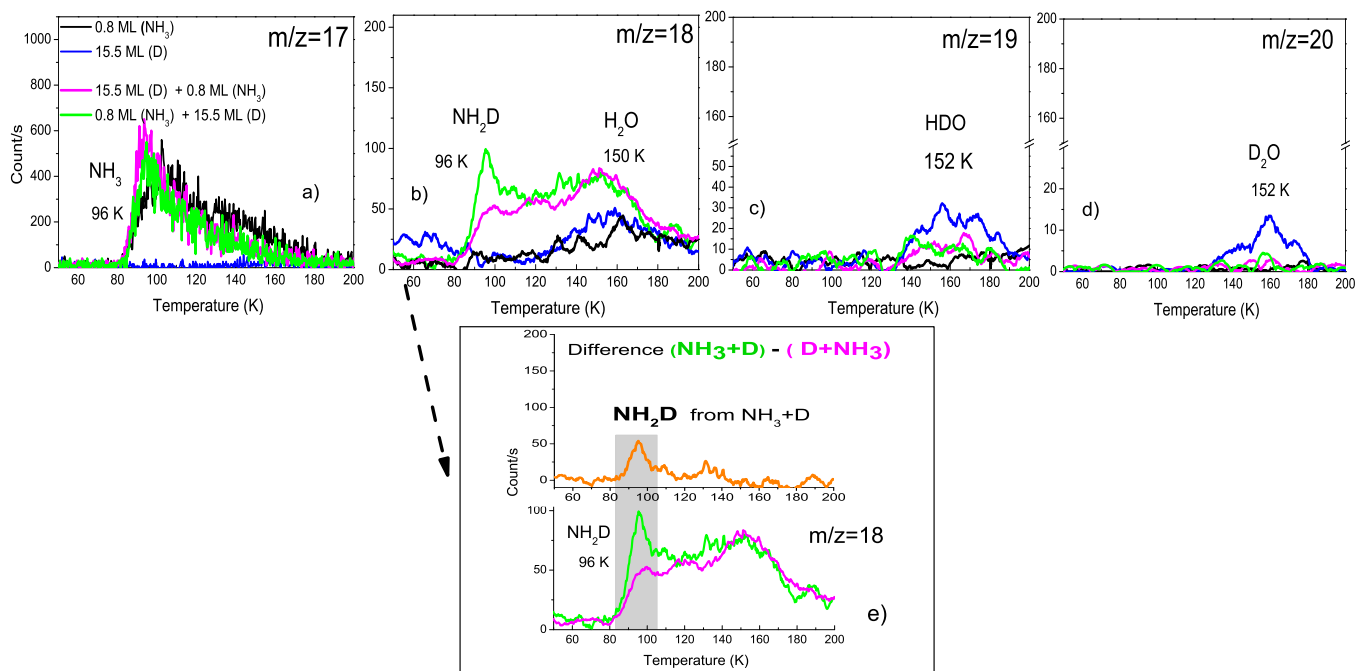
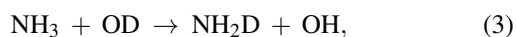


Figure 3: TPD signals between 50 K and 200 K of masses: a) $m/z=17$, b) $m/z=18$, c) $m/z=19$, and d) $m/z=20$. Black curve: deposition of 0.8 ML of NH_3 ice, blue curve: deposition of 15.5 ML of D atoms, magenta curve: deposition of the film 1 (15.5 ML D atoms + 0.8 ML NH_3), green curve: deposition of the film 2 (0.8 ML NH_3 + 15.5 ML D atoms). The deposition of the species is performed on an oxidized HOPG surface held at 10 K. The Figure e) gives the difference between the green curve (Film 2) and the magenta curve (Film 1) for mass 18. It illustrates the desorption peak of NH_2D ($m/z=18$) at around 96 K, really formed by the surface reaction $\text{NH}_3 + \text{D}$.

NH_3 molecules and form NH_2D species ($m/z=18$) through the following exothermic reactions (1) and (2), provided by Nist web-book [38].



The presence of a very small desorption peak ($m/z=20$) for the deuterated water at 150 K following exposure to ammonia (magenta traces in Figure 3d) may support our suggestion. The NH_2D molecules that can be formed by isotopic exchange reaction between NH_3 molecules and HDO and D_2O species on the oxidized graphite surface may desorb between 50 K and 120 K. Moreover, the exposure of D atoms on the oxidized graphite surface may create new functional groups or intermediates, such as (-OD). These reactive intermediate species may interact with NH_3 and form NH_2D following the exothermic reaction (3.2), provided by Nist web-book [38].



All these suggestions for the formation of NH_2D by heavy water contaminants or by -OD intermediates, could explain the observed small desorption peak (in magenta), for mass 18, in the Figure 3b, at around 96 K (film 1), where NH_3 molecules are deposited on top of D atoms on the oxidized graphite surface. In the case of the film 2 (0.8 ML NH_3 +15.5ML of D),

where D atoms are deposited on top of the solid NH_3 film, the desorption peak (in green) at 96 K for mass 18 becomes larger than previously (see Figure 3b). The increase in the TPD area of the peak at 96 K for mass 18 is likely to result from the reaction of D atoms with NH_3 molecules on the surface. The TPD peak at 96 K of the NH_2D molecules produced by the reaction ($\text{NH}_3 + \text{D}$) is shown in the Figure 3e. The desorption curve (in orange) is the difference (film 2-film 1), between the TPD signal of NH_2D ($m/z=18$), expected to be formed from the reaction ($\text{NH}_3 + \text{D}$), and the TPD signal of NH_2D ($m/z=18$), produced from the reaction of NH_3 with -OD, HDO and/or D_2O contaminants on the surface.

3.3. Kinetics of $\text{NH}_3 + \text{D}$ reaction

The kinetic reaction $\text{NH}_3 + \text{D}$ has been studied by exposing 0.8 ML of solid NH_3 to different doses of D atoms (0 ML, 1.0 ML, 6.6 ML, 15.5 ML, 31.0 ML and 53.2 ML). TPD curves of species for masses 17, 18, 19 and 20 are shown in Figure 4 between 50 K and 210 K, for each film of NH_3 and D atoms. As shown in the Figure 4a, the maximum of the TPD peak of NH_3 shifts slightly toward lower temperatures with the D-exposure time from 104 K to 96 K, and in parallel we observe the disappearance of a second desorption peak, as a shoulder at about 150 K. These desorption temperatures differences can be explained in terms of reaction sites and/or surface contamination, such as water molecules. In these experiments the amount of water ices contaminants on the surface is negligible.

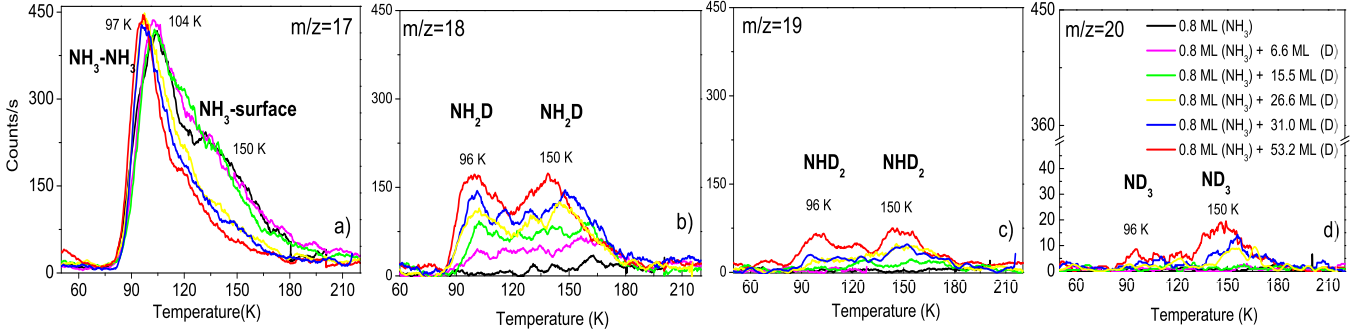


Figure 4: TPD desorption curves of ammonia species between 60 K and 220 K as a function of D-atoms exposure doses (0 ML, 6.6 ML, 15.5 ML, 26.6 ML, 31.0 ML and 53.2 ML) on 0.8 ML of solid NH_3 ice grown on the oxidized HOPG surface held at 10 K with amorphous water ice contaminants: a) $m/z=17$ (NH_3), b) $m/z=18$ (NH_2D), c) $m/z=19$ (NHD_2), and d) $m/z=20$ (ND_3).

In addition, panels (b), (c) and (d) of Figure 4 show the growth of three double desorption peaks at 96 K and 150 K, for masses 18, 19 and 20, respectively. The desorption peak at about 96 K in Figure 4b is likely to be attributed to NH_2D ($m/z=18$) species, mainly produced from the reaction between NH_3 and D atoms. Similarly, the desorption peaks observed at 96 K in Figure 4c, and d are attributed to the doubly deuterated species NHD_2 ($m/z=19$), and triply deuterated ammonia ND_3 ($m/z=20$), formed mainly by the reaction $\text{NH}_2\text{D}+\text{D}$, and NHD_2+D , respectively. We neglected the contribution of NH_2D , NHD_2 , ND_3 formed from the contaminants on the surface, mainly water ices in these experiments.

Furthermore, the cracking pattern of the ionized ammonia molecules ND_3^+ ($m/z=20$), NHD_2^+ ($m/z=19$), NH_2D^+ ($m/z=18$), and NH_3^+ ($m/z=17$) by electron impact, in the ion source of the QMS, are ND_2^+ , NHD^+ , NH_2^+ , ND^+ , NH^+ , D_2^+ , H_2^+ , D^+ , H^+ , and N^+ . The ion fragments NHD^+ ($m/z=17$) and ND_2^+ ($m/z=18$) provided by the QMS in the vacuum chamber during the warming-up phase, can be added to the TPD signals of ionized NH_3^+ ($m/z=17$) and NH_2D^+ ($m/z=18$) molecules, respectively. This means that the TPD curves in the Figure 4a ($m/z=17$) peaking at 96 K and 150 K are the mixture of the ionized NH_3^+ , and the cracking pattern NHD^+ of the ionized NHD_2^+ ($m/z=19$) and NH_2D^+ ($m/z=18$). Similarly, the TPD curves in the Figure 4b ($m/z=18$), peaking at 96 K and 150 K are the mixture of the ionized NH_2D^+ molecules, and the cracking pattern ND_2^+ of ionized NHD_2^+ ($m/z=19$) and ND_3^+ ($m/z=20$) of the deuterated molecules.

In our experimental conditions, the electron's energy of the QMS ion source is 32 eV. With this energy, only 30 % of molecules desorbing from the surface are ionized in the head of the QMS. So we can not determine the precise contribution of species having the same mass m/z to the QMS data, but we can assume that most of ammonia molecules desorbing from the surface are not fragmented in the QMS head but only ionized.

As previously discussed in section 3.2, TPD peaks observed at 150 K in figure 4, panels b), c) and d) match well with the desorption of water impurities H_2O , HDO , and D_2O , respec-

tively. The observed deuterated ammonia species in the TPD spectra are likely to be formed by H-D substitution reaction between the impinging D atoms and the ammonia adsorbed on the oxidized graphite surface. We excluded any energetic particles (photons, electrons and ions) in the formation of the deuterated ammonia species. Previous control experiments realized in the laboratory exclude any possible interaction of the electrons with the surface. The energetic particles produced in the microwave plasma of D atoms inside the beam-line can not reach the cold surface of the sample during the D exposure phase, and dissociate therefore the adsorbed NH_3 molecules and cause their deuteration.

In Figure 4a, the strong TPD peak at 96 K (in black line) has the behavior of multilayer desorption of NH_3 ice, where NH_3 is probably bound to adsorbed NH_3 molecules. While the TPD peak (in black line) at 150 K corresponds to NH_3 molecules physisorbed on the sites of the oxidized HOPG surface and contaminants (OD, CD...). In Figure 4b ($m/z=18$), the TPD peaks at 96 K that increase with the increase of the exposure dose of D atoms, are likely to be attributed to NH_2D molecules ($m/z=18$), formed from the reaction ($\text{NH}_3 + \text{D}$) on ammonia ice deposited on the surface. In the same Figure 4b, the growth of the TPD peaks at 150 K with D exposure doses seems to be coherent with the decrease of the TPD peaks at 150 K in the Figure 4a. The desorption peaks at 150 K Figure 4b are likely to be attributed to NH_2D formed from the reaction ($\text{NH}_3 + \text{D}$) on the surface of the oxidized HOPG. The maximum of these TPD peaks shifts toward lower temperatures with the increase of the peak height, and the NH_2D coverage on the surface. This means that the interaction of the clusters of NH_3 molecules with D atoms leads to the formation of deuterated species NHD_2 , NHD_2 , and ND_3 , desorbing into the gas phase at 96 K, as seen in the TPD curves of Figure 4b, c, and d, respectively. While the successive deuteration of adsorbed NH_3 on the oxidized graphite surface, produces deuterated ammonia molecules desorbing from the stronger binding sites of the surface at 150 K (see Figure 4b, c, and d).

The astrophysical group of Watanabe et al. [22] have demonstrated experimentally the efficient formation of the deuterated

isotopologue species of methanol at low surface temperature (10 K) by the D atoms exposure on CH₃OH ice. The isotopic species were observed and detected by infrared spectroscopy during the exposure of the adsorbates at 10 K. Similarly to methanol molecules, we believed that the deuteration reaction NH₃ + D proceeds during the exposure phase of NH₃ and D atoms on the oxidized HOPG substrate at 10 K, thanks to the tunneling process. At 10 K, D atoms are mobile [39] and can diffuse on the surface to react with solid ammonia molecules. However, since in our experiments the deuterated species of ammonia are detected by TPD measurements from 10 K to 200 K, it is possible that the formation of the deuterated species of ammonia proceeds during the warming-up phase of the sample, rather than during the exposure of the reactants on the surface at 10 K. This assumption for the deuteration of ammonia by D atoms at higher surface temperatures is not taken into account in our experiments.

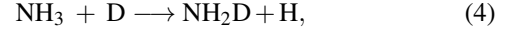
3.4. Kinetics of CH₃OH+D reaction

In this section, we would like to compare the kinetic reaction NH₃ + D to that of CH₃OH + D molecules in the submonolayer regime. We investigated similar deuteration experiments of solid CH₃OH by D atoms as for ammonia molecules. The experiments were performed under the same conditions: same low surface coverage (~ 0.8 ML), same D atomic flux $\phi(D) = 3.7 \times 10^{12}$ molecules \cdot cm⁻² \cdot s⁻¹, and same surface temperature (10 K). Firstly, we deposited 0.8 ML of solid CH₃OH on the oxidized HOPG surface at 10 K, and then we added 6.6 ML of D atoms for the first experiment, and 15.5 ML of D atoms for the second one. After the D-addition phase, each film of CH₃OH + D was heated linearly from 10 K to 210 K using the same heating rate of 0.17 K \cdot s⁻¹. Figure 5 shows the TPD desorption curves of CH₃OH ($m/z=32$), and the newly formed isotopic species CH₂DOH ($m/z=33$), CHD₂OH ($m/z=34$) and CD₃OH ($m/z=35$) between 100 K and 200 K. According to Nagaoka et al. [26] and Hiraoka et al. [40], the H-abstraction of CH₃OH by D atoms is likely to occur in the methyl -CH₃ group rather than the hydroxyl -OH group of the CH₃OH ($m/z=32$) molecules. We have thus attributed the TPD signals of masses $m/z=33$, $m/z=34$ and $m/z=35$ in the Figure 5 to the newly formed deuterated species CH₂DOH, CHD₂OH, and CD₃OH, respectively, which are deuterated in the methyl group. The formation of the deuterated species in the hydroxyl group, such as CH₃OD ($m/z=33$), CH₂DOD ($m/z=34$), and CHD₂OD ($m/z=35$), by the reaction system (CH₃OH + D) is expected to be negligible in this work. However, these deuterated methanol species in the hydroxyl group can be formed by D-H isotopic exchange between the species (CH₃OH, CH₂DOH, and CHD₂OH) with the deuterated D₂O water ice contaminants, during the transition phase from the amorphous to the crystalline state of the water ice at 120 K [41].

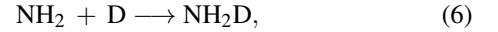
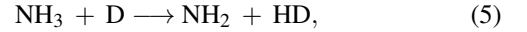
4. Analysis

4.1. Rate equations of NH₃+D system reactions

We suggested that the reaction between NH₃ and D atoms on the oxidized graphite surface held at 10 K, proceeds through direct Hydrogen-Deuterium substitution process by H-abstraction and D-addition mechanism, as proposed by Nagaoka et al. [22, 42] for H₂CO + D and CH₃OH + D reactions. In fact, the direct H-D substitution reaction (4) leading to the formation of NH₂D species is slightly exothermic with a formation enthalpy $\Delta H^0 = -781.8$ K.

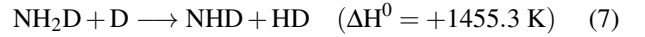


In the case of the H-abstraction and D-addition mechanism of NH₃, the indirect H-D substitution process is described by the following reactions (5) and (6).

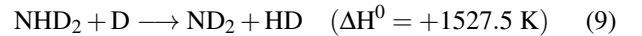


The first H-atom abstraction reaction (5) of NH₃ molecule by D atom leads to the formation of HD molecule and the NH₂ radical. This reaction is endothermic with a reaction enthalpy of $\Delta H^0 = +1527.5$ K, and needs an excess thermal energy to be produced. While the second D-addition reaction (6) leading to the formation of the first isotopologue NH₂D is exothermic with higher heat of formation $\Delta H^0 = -54480$ K. All the standard reaction enthalpies involving ammonia species and D atoms are provided by NIST database [38].

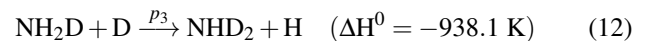
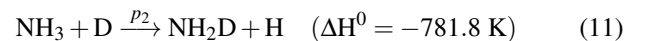
The same endothermic behavior takes place in the H-abstraction reactions (7) and (9) of NH₂D and NHD₂ species by D atoms, respectively.



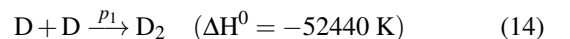
and



In order to fit the TPD experimental data of NH₃, NH₂D, NHD₂ and ND₃ species, given in the Figure 7, we used a kinetic model described by the following exothermic system reactions (11-13) for the three direct H-D substitution reactions.



These reactions are in competition with the exothermic D+D surface reaction leading to the formation of D₂ molecules.



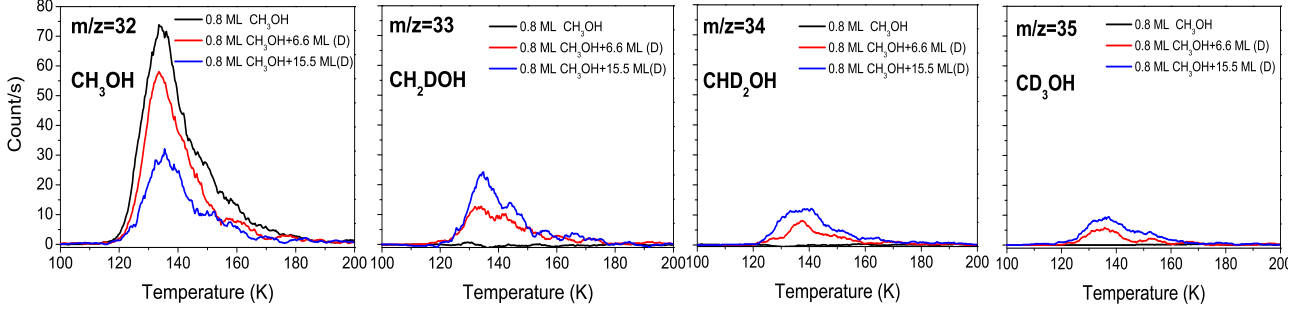


Figure 5: TPD curves of CH_3OH ($m/z=32$), CH_2DOH ($m/z=33$), CHD_2OH ($m/z=34$) and CD_3OH ($m/z=35$) between 100 K and 200 K. Black curve: 0.8 ML of CH_3OH ice pre-deposited on the oxidized graphite surface held at 10 K; Red curve: after the exposure of 6.6 ML of D atoms at 10 K on 0.8 ML of CH_3OH ice; Blue curve: after the exposure of 15.5 ML of D atoms at 10 K on 0.8 ML of CH_3OH ice.

Our model includes both Eley-Rideal (ER) and Langmuir-Hinshelwood (LH) mechanisms for the reactions of D atom either with another D atom on the surface or with an ammonia species already adsorbed on the surface at 10 K. The Eley-Rideal mechanism occurs when one of the species already adsorbed on the surface promptly reacts with a particle coming from the gas phase, before being adsorbed on the surface. The Langmuir-Hinshelwood mechanism describes the formation of molecules on the surface when two adsorbed reaction-partners diffuse on the surface. D atoms are thermalized with the surface and they react with ammonia molecules thanks to surface diffusion. The ER mechanism is independent of the temperature of the surface T_s , while LH mechanism is very sensitive to T_s since it depends on diffusion coefficients. Moreover, LH is more efficient than ER mechanism at low surface coverage [43]. In our experiment, a D atom coming from the gas phase can hit an ammonia species already adsorbed on the surface, react and form a newly isotopic species of ammonia through ER mechanism. If the adsorbed D atom does not react through ER mechanism, it can diffuse on the surface from one site to a neighboring one. The diffused D atom can react either with another D atom on the surface to form D_2 molecule, or with an adsorbed NH_3 , NH_2D or NHD_2 molecules to form NHD_2 , NH_2D , or ND_3 species, respectively, through the LH mechanism. All species, except D atoms, are not mobile on the surface at 10 K.

4.1.1. Kinetic model

The model used to fit our experimental data is very similar to the one described by Minissale et al. [43, 44]. It is composed of six differential equations, one for each of the species considered: D atoms, coming exclusively from the beam; NH_3 molecules, deposited on the surface; NH_2D , NHD_2 , and ND_3 , formed on the surface; and finally D_2 , coming both from the beam and formed on the surface. Each differential equation is composed of positive and negative terms, indicating respectively an increase (i.e. species arriving from the gas phase or formed on the surface), or a decrease (i.e. species reacting on the surface) in the surface coverage of the species. The terms involving the ER and LH mechanisms are independent

of one another, thus we are able to determine the amount of a species formed (or consumed) via ER or LH mechanism. Below, we present the list of differential equations governing the NH_3 deuteration:

$$\begin{aligned} \frac{d[\text{D}]}{dt} = & \phi_{\text{D}} \left(1 - 2p_{1\text{ER}}[\text{D}] - p_{2\text{ER}}[\text{NH}_3] - \right. \\ & \left. p_{3\text{ER}}[\text{NH}_2\text{D}] - p_{4\text{ER}}[\text{NHD}_2] \right) - \\ & k_{\text{diff}} \left(4p_{1\text{LH}}[\text{D}][\text{D}] + p_{2\text{LH}}[\text{D}][\text{NH}_3] + \right. \\ & \left. p_{3\text{LH}}[\text{D}][\text{NH}_2\text{D}] + p_{4\text{LH}}[\text{D}][\text{NHD}_2] \right) \end{aligned} \quad (15)$$

$$\begin{aligned} \frac{d[\text{D}_2]}{dt} = & \phi_{\text{D}_2} + 2\phi_{\text{D}}(1 - e_1)p_{1\text{ER}}[\text{D}] + \\ & 2(1 - e_1)k_{\text{diff}} \cdot p_{1\text{LH}}[\text{D}][\text{D}] \end{aligned} \quad (16)$$

$$\frac{d[\text{NH}_3]}{dt} = -\phi_{\text{D}}p_{2\text{ER}}[\text{NH}_3] - k_{\text{diff}} \cdot p_{2\text{LH}}[\text{D}][\text{NH}_3] \quad (17)$$

$$\begin{aligned} \frac{d[\text{NH}_2\text{D}]}{dt} = & \phi_{\text{D}} \left(p_{2\text{ER}}[\text{NH}_3] - p_{3\text{ER}}[\text{NH}_2\text{D}] \right) + \\ & k_{\text{diff}} \left(p_{2\text{LH}}[\text{D}][\text{NH}_3] - p_{3\text{LH}}[\text{D}][\text{NH}_2\text{D}] \right) \end{aligned} \quad (18)$$

$$\begin{aligned} \frac{d[\text{NHD}_2]}{dt} = & \phi_{\text{D}} \left(p_{3\text{ER}}[\text{NH}_2\text{D}] - p_{4\text{ER}}[\text{NHD}_2] \right) + \\ & k_{\text{diff}} \left(p_{3\text{LH}}[\text{D}][\text{NH}_2\text{D}] - p_{4\text{LH}}[\text{D}][\text{NHD}_2] \right) \end{aligned} \quad (19)$$

$$\frac{d[\text{ND}_3]}{dt} = \phi_{\text{D}} \cdot p_{4\text{ER}}[\text{NHD}_2] + k_{\text{diff}} \cdot p_{4\text{LH}}[\text{D}][\text{NHD}_2] \quad (20)$$

The $[D]$, $[D_2]$, $[NH_3]$, $[NH_2D]$, $[NHD_2]$, and $[ND_3]$ quantities are the surface coverages of species. $[X]$ is dimensionless and represents the percentage of surface covered with the X species. For each species $[X]$ ranging between 0 and 1. This condition is true for all species except for D_2 , whose surface coverage can be bigger than one. We stress that it does not represent a problem to evaluate activation barrier, since D_2 is an inert species and does not have any effect on kinetics of reactions. The initial reaction conditions at $t=0$ simulate the experimental conditions: $[NH_3](t=0)=0.8$ and $[NH_2D]=[NHD_2]=[ND_3]=0$ for $t=0$. Furthermore, we impose that at any time:

$$[NH_3](t) + [NH_2D](t) + [NHD_2](t) + [ND_3](t) = [NH_3](t=0) \quad (21)$$

Dimensionless surface coverage is then converted in ML (or molecule \cdot cm $^{-2}$) by multiplying $[X]$ for the amount of adsorption sites of our surface (10^{15} sites \cdot cm $^{-2}$) and compared with experimental results. ϕ_X represents the part of surface covered per second by the X species coming from the gas phase. We know that in our experimental conditions the total number density of the impinging D-atoms coming from the gas phase and hitting the surface is given by the flux of D atoms in the beamline: $\phi_D = 3.7 \times 10^{12}$ atoms \cdot cm $^{-2}$ \cdot s $^{-1}$. If we consider again that a surface contains 10^{15} sites \cdot cm $^{-2}$, the flux of D atoms landing the surface is $\phi_D=3.7 \times 10^{-3}$ s $^{-1}$.

The terms concerning the chemical desorption of ammonia species NH_2D , NHD_2 , and ND_3 (formed by the reaction with D atoms) are not considered in the model, since the thermal desorption of these considered species is negligible at 10 K. Despite the various heats of formation of NH_3 isotopologues (ΔH less than 1100 K), no desorption of the newly formed ammonia species has been observed experimentally at 10 K from the graphite surface. This is because the local heats of formation of these species through the exothermic reactions (11-13) do not exceed the desorption energy of these ammonia species ($E_{des}=2300$ K) [45]. So once the isotopologue ammonia species are formed, they stay in the solid phase on the graphite surface at 10 K, because their binding energy of about 2300 K is higher than the excess energies of formation. The non chemical desorption of the corresponding molecules at 10 K is also confirmed experimentally by using the DED (During Exposure Desorption) method [46], which consists of monitoring with the QMS placed in front of the sample, the signal of the deuterated molecules released into the gas phase during the deposition phase. However, the parameter e_1 characterizing the prompt desorption of some D_2 molecules, upon formation on the surface at 10 K, through the very exothermic reaction (14), is expected to be non negligible. We point out that this term e_1 does not influence ammonia species surface coverage. In fact, as we have already said D_2 is a non-reactive species and it cannot consume neither D atoms nor NH_xD_y species on the surface. However, it has been demonstrated by Amiaud et al. [27] that the presence of D_2 molecules already adsorbed on the water ice increases the recombination efficiency of D atoms through the barrierless $D + D \rightarrow D_2$ reaction, by increasing the sticking co-

efficient of the deuterium atoms on the surface. This behaviour may have an important impact in the deuteration experiments, since the presence of condensed inert D_2 species may separate D and NH_3 , resulting in the decrease of the recombination efficiency of D atoms with adsorbed NH_3 , and the reduction of the H-D substitution reaction.

p_{1ER} and p_{1LH} parameters are the reaction probabilities (dimensionless) of the D+D surface reaction (14), and we fixed to one their values. Similarly, p_{2ER} , p_{3ER} , p_{4ER} , and p_{2LH} , p_{3LH} and p_{4LH} parameters represent the probabilities of the reactions (11-13) to be occurred via ER and LH mechanisms, respectively. Ammonia and D atoms species can react together by overcoming the activation barrier (Arrhenius term), or by crossing the barrier (tunneling term) as expressed by the following equations (22) and (23):

$$p_{iER} = e^{-E_{a_i}/(k_B \times T_{eff})} + e^{-2 Z_r \times \sqrt{2 \mu \times E_{a_i} \times k_B/h}} \quad (22)$$

and

$$p_{iLH} = e^{-E_{a_i}/(k_B \times T_s)} + e^{-2 Z_r \times \sqrt{2 \mu \times E_{a_i} \times k_B/h}} \quad (23)$$

Where k_B is the Boltzmann constant, h the Planck constant, E_{a_i} ($i=2-4$) are the activation energy barriers of the reactions (11-13), Z_r is the width of the (rectangular) activation barrier, and μ is the tunneling mass which is described by the reduced mass of the system involved in bi-molecular atom transfer reaction. This tunneling mass is defined as:

$$\mu = \frac{m_{NH_xD_y} \times m_D}{m_{NH_xD_y} + m_D}, \text{ with } x,y=0-3 \text{ and } x+y=3 \quad (24)$$

T_{eff} is the effective temperature of the reaction between NH_3 , NH_2D , NHD_2 and D atoms given by:

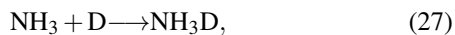
$$T_{eff} = \mu \left(\frac{T_{solid}}{m_{NH_3}} + \frac{T_{gas}}{m_D} \right) = 314 \text{ K} \quad (25)$$

T_s (=10 K) is the surface temperature. The parameter k_{diff} is the diffusion coefficient of D atoms between sites on the surface. It represents the amount of surface sites scanned in one second by D atoms. It is defined by the following equation (26)

$$k_{diff} = \nu \left(e^{-E_{diff}/k_B \times T_s} + e^{-2 Z_d \times \sqrt{2 \mu \times E_{diff} \times k_B/h}} \right) \quad (26)$$

Where $\nu = 10^{12}$ s $^{-1}$ is the attempt frequency for overcoming the diffusion barrier of D atoms and E_{diff} is the energy barrier for the diffusion of D atoms on cold surfaces held at 10 K. Bonfant et al. [47] have reported an extremely low diffusion barriers of 4 meV for hydrogen atoms on graphite surface, meaning that hydrogen atoms physisorbed on graphite is highly mobile at low surface temperatures. However, for irregular surfaces such as ASW, the diffusion energy barrier of D atoms does not have a single value but follows a distribution, because there are several potential sites of different depths. Since our substrate used in our experiments is composed of an oxidized HOPG mixed with ASW ice deposits, the diffusion energy value of D atoms used in this model is that estimated on ASW ice for low surface coverage, $E_{diff} = (22 \pm 2)$ meV or (255 ± 22) K [39]. Even if this

diffusion energy is different from that calculated on graphite surface, its high value does not affect the modeling results. The parameter Z_d is the width of the (rectangular) diffusion barrier. We have fixed the width of the diffusion barrier of D atoms (Z_d) to 1 Å, a value commonly used to describe H or D atoms diffusion on the surface. In our kinetic model, the term of the tunneling probability $e^{-2 Z_r \sqrt{2 \mu E_a} k_B / \hbar}$ for crossing the rectangular activation barrier depends on the tunneling mass of the reaction. This tunneling mass is described by the reduced mass μ of the system involving ammonia species NH_xD_y and D atom with $x, y=0-3$ and $x+y=3$. The value of μ for each reaction system is equal to 1.8 amu, and is close to the mass of the deuterium particle D ($m/z=2$ amu). This means that for direct D-H exchange reaction between D atom and NH_3 molecules, the D atom is considered to be the tunneling particle that conduct the system $\text{NH}_3 + \text{D}$ to across the rectangular barrier through quantum tunneling. According to Hidaka et al. [48], the tunneling mass significantly depends on the reaction mechanism. For the addition reaction ($\text{AX} + \text{B} \rightarrow \text{AXB}$), the tunneling mass in the reaction coordinate is simply described by the reduced mass of the two-body system. However, for the abstraction reaction ($\text{AX} + \text{B} \rightarrow \text{A} + \text{BX}$), which involves three free particles in the reaction system, the tunneling mass is described by the effective mass, defined in the papers of Hidaka et al. [48]. For the direct H-D exchange reaction (11) between NH_3 and D atoms, the description of the tunneling mass is not straightforward according to the reference [48]. However, if we assume that the H-D substitution reaction (11) occurs via an intermediate NH_3D , having a tetrahedral geometry as demonstrated by ab-initio calculations [25], we can thus apply the reduced mass μ to describe the tunneling mass of the addition (27) reaction.



4.1.2. Activation energy barriers of the reactions

In our kinetic model, we have four free parameters: the activation barriers (E_{a2} , E_{a3} , E_{a4}) for reaction (reaction barrier) and the width of the reaction barrier Z_r . This last can be constrained between 0.7 and 0.9 Å. Actually, solid-state chemistry at low temperatures should be dominated by quantum tunneling, according to Harmony [49] and Goldanskii [50]. In particular, H-abstraction and D-substitution should be ruled by tunneling, as shown by Goumans et al. [51] in the case of CH_3OH deuteration. The reaction $\text{NH}_3 + \text{D}$ has been studied experimentally very long time ago in the gas phase by Kurylo et al. [24] over the temperature range 423-741 K. These authors found that the H-D exchange between NH_3 and D may proceed through an intermediate NH_3D following the reaction $\text{NH}_3 + \text{D} \rightarrow [\text{NH}_3\text{D}] \rightarrow \text{NH}_2\text{D} + \text{H}$. As mentioned previously, the $\text{NH}_3 + \text{D}$ system reaction has been also studied theoretically [25] using ab-initio interpolated potential energy surface calculations. In both papers [25, 24], the activation barrier of the H-D exchange reaction (11) is reported to be $E_a = 11$ kcal/mol or 5540 K. However, some works (i.e. Bell [52] and references therein, Chapter 6: Tunneling in molecular spectra, the inversion of ammonia and related processes, page 153) argue that in the case of ammonia inversion, tunneling should be the dominant process, with

a typical width of the reaction barrier (Z_r) is 0.7-0.8 Å. This reaction is considered the prototype of processes involving tunneling in a symmetrical (or quasi-symmetrical) potential energy curve. To the best of our knowledges no experimental and theoretical works have deal with the width of ammonia deuteration barrier. Thus we have used a value for width of deuteration similar to that of ammonia inversion, aware that the two reactions involve not identical chemical processes. For the sake of simplicity, we have used a common Z_r value for the deuteration reactions (11-13), instead of a value for each reaction. We suggest that tunneling is necessary for ammonia deuteration (in analogy with methanol deuteration). For this reason we use quantum tunneling in our model but we point out that our simple formulation of tunneling is useful only for a qualitative evaluation of our experimental results. Quantum tunneling refers to the quantum mechanical phenomenon where a particle tunnels through a barrier that it classically could not surmount. Quantum tunneling is known to be an important process for molecular synthesis on interstellar grains at very low temperatures [53]. A detailed description of tunneling falls outside the scope of this work.

Figure 6 shows the surface densities of NH_3 , NH_2D , NHD_2 and ND_3 species as a function of D atoms Fluences. These surface densities, expressed in fraction of monolayer (ML) are the normalized integrated areas below the TPD curves of NH_3 , NH_2D , NHD_2 and ND_3 peaked at 97 K for each D atom fluence, with respect to the TPD integrated area of NH_3 for one monolayer coverage. As previously explained in the section 3.3, we assumed that all the deuterated ammonia species are ionized by electron impact in the ion source of the QMS during the TPD experiments. However, it has been reported by Rejoub et al. [54] that the ionization cross-section of light NH_3 molecules is twice larger than that of ND_3 , meaning that the cross-sections for formation of ion fragments from heavy deuterated molecules, are much smaller than those from NH_3 . Because we neglected the contribution of the cracking patterns in the TPD experiments, we did not considered the different ionization cross-section values of the deuterated species in the measurements of the areas from TPD profiles. As shown in the Figure 6, there is a good correlation between the experimental data and the fits obtained by the model for the exponential decay of NH_3 , and the growth of NH_2D , NHD_2 and ND_3 species on the surface when increasing the amount of D atoms on the surface. The plots of the Figure 6 show that for the higher D-irradiation time of 240 minutes, and the higher D-fluence of 5.34×10^{16} atoms \cdot cm $^{-2}$, about 20 % of the adsorbed NH_3 molecules are mainly deuterated into NH_2D and NHD_2 species with traces of ND_3 . The formation yields of the single, the double and the triple deuterated ammonia species are approximately (14 %), (5 %), and (1 %), respectively.

Thanks to our model, we have tested different scenario: we used three values (150 K, 250 K, 350 K) for the diffusion barrier E_{diff} of D atoms, and for each value, we have varied Z_r from 0.7 to 0.9 Å (step of 0.01 Å). In the case of $E_{\text{diff}} = 250$ K, the activation energy barriers of the successive H-D substitution reactions of ammonia species by D atoms are found to be $E_{a2} = (1840 \pm 270)$ K for the reaction (11), $E_{a3} = (1690 \pm 245)$ K for the reaction (12), and E_{a4}

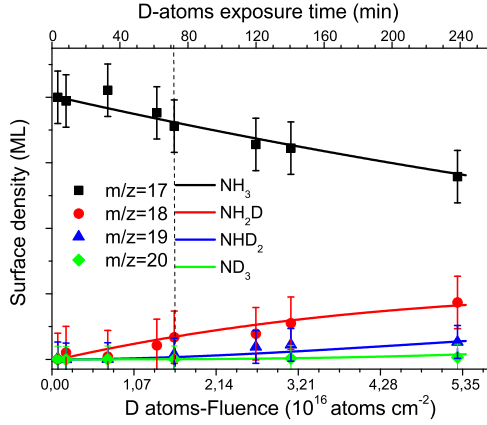


Figure 6: Kinetic evolutions of NH_3 , NH_2D , NHD_2 , and ND_3 species present on the surface as a function of D atoms exposure doses, and D fluences on 0.8 ML coverage of solid ammonia already deposited on the oxidized graphite substrate at 10 K. Black squares, red circles, blue triangles, green diamonds are the TPD experimental data of NH_3 , NH_2D , NHD_2 , and ND_3 , respectively. Solid lines are the fits obtained from the model. The uncertainties are given by the errors bars.

$= (1670 \pm 230)$ K for the reaction (13). In the Table 2, we list the width and the energy activation barrier for the H-D substitution reactions of NH_2D , NHD_2 and ND_3 species. The listed values of the activation energies E_{a_1} minimize the χ^2 value of our fit with respect to our experimental data. The statistical parameter χ^2 is obtained for each set of parameters by using the following formula:

$$\chi^2 = \frac{\sum_{\text{ML,mol}} [S_t(\text{ML, mol}) - S_e(\text{ML, mol})]^2}{S_e(\text{ML, mol})} \quad (28)$$

Where $S_t(\text{ML, mol})$ and $S_e(\text{ML, mol})$ are respectively the theoretical and experimental surface density for each molecule at a certain D-fluence.

In order to have a good correlation between the model and the experiments, Figure 7 shows how we can minimize the χ^2 value by setting a couple of activation energies values (E_{a_2} , E_{a_3}) for the deuteration reactions (11) and (12) and varying only the value of the third energy E_{a_4} for the reaction (13). Our activation energy barriers for the direct H-D substitution reactions (11-13) between ammonia species and D atoms (see Table 2) are smaller than the activation energy barrier $E_a = 5540$ K reported by the two references [24, 25], both in gas and solid phases. The low values of the activation energies obtained in this work can be explained by the catalytic effect of the ASW ice + oxidized HOPG on the deuteration reaction $\text{NH}_3 + \text{D}$. This substrate favors the diffusion of D atoms, and increases therefore the reactivity between NH_3 molecules and D atoms on the surface.

As shown in the Table 2, the width Z_r and the energy of the activation barriers E_a depend on the diffusion energy of D atoms on the surface E_{diff} . One can note that the higher the diffusion energy (E_{diff}) of D atoms, the lower the width (Z_r) of the energy barriers. The diffusion of D atoms on the cold surface increases the probability of the H-D substitution reactions of ammonia in the solid phase. Table 2 also shows that for each E_{diff} , the value of the activation energy barrier is al-

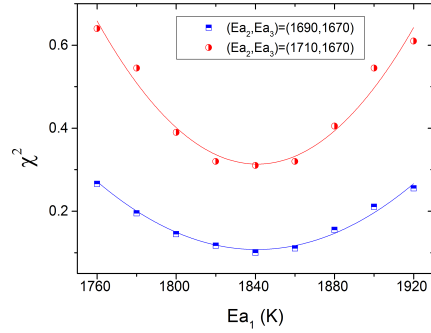


Figure 7: The curves minimizing the χ^2 value between the kinetic model and the experimental measurements for the reaction between NH_3 and D atoms, by setting a couple of activation energies barriers (E_{a_2} , E_{a_3}) for the deuteration reactions (11) and (12) and varying the value of third one (E_{a_4}) for the reaction (13). E_{diff} and Z_r are respectively 250 K and 0.83 Å.

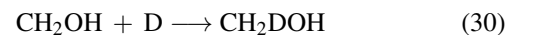
Table 2: The width Z_r and the height of the energy barriers E_a , expressed in (Å) and in kelvin (K), respectively, of the successive H-D substitution reactions of NH_3 molecules by D atoms on the oxidized, partly ASW covered graphite surface at 10 K, for a fixed value of D-atom diffusion energy E_{diff} . The minimum χ^2 value of the fits is found to vary between 0.3 and 0.1.

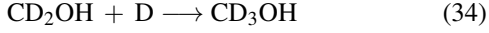
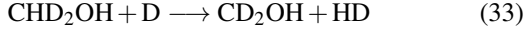
Reactions	E_{diff}	Z_r	E_a
units	K	Å	K
$\text{NH}_3 + \text{D} \xrightarrow{p_2} \text{NH}_2\text{D} + \text{H}$	150	0.86	1950 ± 250
	250	0.83	1840 ± 270
	350	0.81	1750 ± 320
$\text{NH}_2\text{D} + \text{D} \xrightarrow{p_3} \text{NHD}_2 + \text{H}$	150	0.86	1820 ± 220
	250	0.83	1690 ± 245
	350	0.81	1610 ± 290
$\text{NHD}_2 + \text{D} \xrightarrow{p_4} \text{ND}_3 + \text{H}$	150	0.86	1800 ± 210
	250	0.83	1670 ± 230
	350	0.81	1600 ± 250

ways high for the first deuteration reaction $\text{NH}_3 + \text{D}$, and then decreases by almost 10 % for the second $\text{NH}_2\text{D} + \text{D}$ and the third $\text{NHD}_2 + \text{D}$ deuteration reactions. Our activation energy barriers for the deuteration reaction $\text{NH}_3 + \text{D}$ in the solid phase is much lower than the value ($46 \text{ kJ} \cdot \text{mol}^{-1}$) given in the gas phase [25, 24]. This large difference can be explained by the catalytic effect of substrate composed of oxidized graphite and ASW ices deposits.

4.2. Rate equations of $\text{CH}_3\text{OH} + \text{D}$ system reactions

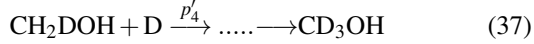
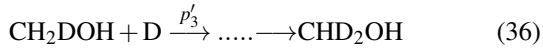
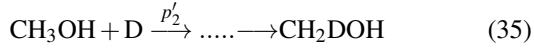
As has been suggested by Nagaoka et al. [26], the deuteration of CH_3OH , CH_2DOH and CHDOH species by D atoms on cold surfaces, occur through the successive H abstraction and D addition mechanism as follows:





Where the H-abstraction reactions (29, 31 and 33) are exothermic with small activation barriers in comparison to the direct H-D substitution reactions. While the D-addition reactions (30, 32 and 34) are exothermic with no barriers. In Hama et al's review [2], the direct H-D substitution reaction $\text{CH}_3\text{OH} + \text{D} \longrightarrow \text{CH}_2\text{DOH} + \text{H}$ has a very large activation energy barrier of 169 kJ/mol (or 20330 K), in comparison to the following H-abstraction reaction $\text{CH}_3\text{OH} + \text{D} \longrightarrow \text{CH}_2\text{OH} + \text{HD}$, which has an activation energy of 34 kJ · mol⁻¹ (or 4090 K), estimated from gas phase calculations.

In this work, the successive deuteration reactions of methanol species by D atoms are described by the following simple reactions (35-37).



Where p'_2, p'_3 and p'_4 parameters are the reaction probabilities of the H-abstraction reactions (29, 31 and 33) of CH_3OH , CH_2DOH and CH_2DOH by D atoms, respectively. Using the same kinetic model previously described for $\text{NH}_3 + \text{D}$ system reactions in solid phase, we have estimated the activation energy barriers of the H-abstraction reactions (29, 31 and 33) for $\text{CH}_3\text{OH} + \text{D}$ system reactions. Table 3 summarizes the values of the width Z_r and the activation energy barriers E_a for each H-D substitution reaction of methanol species. Our kinetic model provides an activation energy barrier $E_a = (1080 \pm 180)$ K or (8.9 ± 1.5) kJ · mol⁻¹ for the first abstraction reaction ($\text{CH}_3\text{OH} + \text{D}$) given by the equation (29) with $E_{\text{diff}} = 250$ K. This value is more than a factor of four smaller than the activation energy value (34 kJ · mol⁻¹ or 4090 K), reported by [2] from theoretical estimations in the gas phase. Once again the catalytic role of the surface can explain the difference between gas phase and solid phase activation barriers. The activation energy barriers of the successive deuteration reactions of methanol species decrease significantly with the increase of the diffusion energy E_{diff} of D atoms on the surface. Nevertheless, the values of the activation energy barriers for $\text{CH}_3\text{OH} + \text{D}$ system reactions (see Table 3) are always smaller than those of the $\text{NH}_3 + \text{D}$ system reactions (see Table 2).

Figure 8 shows the best fit of the data for the exponential decay of CH_3OH , and the increase of the surface densities of CH_2DOH , CHD_2OH and CD_3OH with the increasing time and fluence of D atoms exposure on 0.8 ML of solid methanol CH_3OH pre-deposited on the oxidized HOPG surface. We can note that after 70 minutes of D atoms addition, about 0.44 ML of the adsorbed CH_3OH molecules, are deuterated into three isotopologue species, with the formation yields of ($\sim 22\%$) for CH_2DOH , ($\sim 15\%$) for CHD_2OH ,

Table 3: The width Z_r and the height of the energy barriers E_a , expressed in (Å) and in kelvin (K), respectively, of the successive H-D substitution reactions of CH_3OH molecules by D atoms on the oxidized, partly ASW covered graphite surface at 10 K, for a fixed value of D-atom diffusion energy E_{diff} . The χ^2 value of the fit vary between 0.3 and 0.1.

Reactions units	E_{diff} K	Z_r Å	E_a K
$\text{CH}_3\text{OH} + \text{D} \xrightarrow{p'_2} \text{CH}_2\text{DOH} + \text{H}$	150	0.86	1450 ± 210
	250	0.85	1080 ± 180
	350	0.84	860 ± 120
$\text{CH}_2\text{DOH} + \text{D} \xrightarrow{p'_3} \text{CHD}_2\text{OH} + \text{H}$	150	0.86	1330 ± 200
	250	0.85	990 ± 180
	350	0.84	770 ± 145
$\text{CHD}_2\text{OH} + \text{D} \xrightarrow{p'_4} \text{CD}_3\text{OH} + \text{H}$	150	0.86	1300 ± 205
	250	0.85	980 ± 170
	350	0.84	780 ± 150

and ($\sim 8\%$) for CD_3OH . The comparison with the previous kinetic results given in the Figure 7, shows that when 0.8 ML of solid NH_3 is irradiated with D atoms during the same exposure time of 70 minutes, only 10 % of NH_3 molecules are deuterated into NH_2D ($\sim 8\%$) with traces of NHD_2 ($\sim 2\%$) and NHD_3 ($< 1\%$). This means that during 70 minutes of D atoms exposure phase, 0.44 ML (or 55 %) of adsorbed CH_3OH molecules are consumed by D atoms, while only 0.08 ML (or 10 %) of solid ammonia can be destructed by D atoms. We defined the deuteration rate $\nu_X = \frac{d[X]}{dt}$ of an adsorbed species X by D atoms, as the amount [X] of this species (in molecules · cm⁻²), consumed per unit time (in min) of D atoms exposed on the surface. The deuteration rate value of CH_3OH is estimated to $\nu_{\text{CH}_3\text{OH}} \simeq 0.005 \times 10^{15}$ molecules · cm⁻² · min⁻¹, while that of NH_3 species is found to be $\nu_{\text{NH}_3} \simeq 0.001 \times 10^{15}$ molecules · cm⁻² · min⁻¹, and can slightly decrease for extended irradiation up to 240 minutes. The relationship between the two deuteration rates is $\frac{\nu_{\text{CH}_3\text{OH}}}{\nu_{\text{NH}_3}} \simeq 5$, meaning that the deuteration rate of CH_3OH molecules by D atoms on cold and oxidized graphite HOPG surfaces with ASW ice deposits, is five times higher than of NH_3 .

5. Discussion and conclusions

In this work, we demonstrated experimentally the possible deuteration of NH_3 molecules by D atoms on cold oxidized HOPG surface, partly covered with ASW ices. The deuteration experiments of solid ammonia were performed at low surface coverage and low temperature 10 K using mass spectroscopy and temperature programmed desorption (TPD). The isotopologue ammonia species NH_2D , NHD_2 and ND_3 desorbing from the surface at 96 K and 150 K are likely to be formed by direct exothermic H-D substitution reactions between the adsorbed ammonia species on the surface and the impinging D atoms. A kinetic model taking into account the diffusion of D atoms on the surface provides the activation energy barriers of the deuteration reactions ($\text{NH}_3 + \text{D}$) in the solid phase. We found that

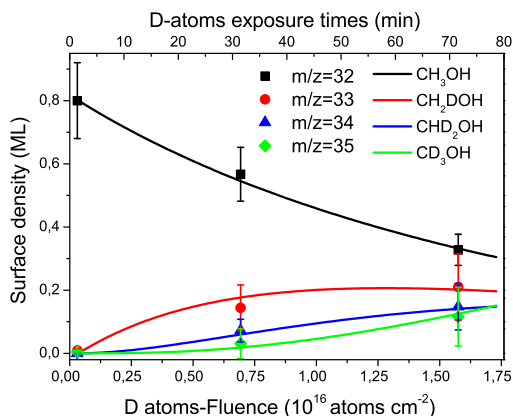


Figure 8: The variation of the surface densities of CH_3OH ($m/z=32$), CH_2DOH ($m/z=33$), CHD_2OH ($m/z=34$) and CD_3OH ($m/z=35$) as a function of D atoms exposure times and D fluences on 0.8 ML coverage of CH_3OH film pre-deposited on the surface of an oxidized graphite substrate held at 10 K. Full squares, circles and triangles are the TPD experimental data. Solid lines are the fits obtained from the model. The uncertainties are given by the errors bars.

the energy barrier for the D-H exchange reaction ($\text{NH}_3 + \text{D}$) is (1840 ± 270 K) or (15.4 ± 2.5 $\text{kJ} \cdot \text{mol}^{-1}$), three times lower than that predicted in the gas phase (5530 K or 46 $\text{kJ} \cdot \text{mol}^{-1}$) [24, 25]. Our results also show that the activation energy barrier for the first deuteration reaction ($\text{NH}_3 + \text{D}$) of ammonia is almost two times higher than that of the abstraction reaction (29) of CH_3OH (1080 ± 180 K) or (9.0 ± 1.2 $\text{kJ} \cdot \text{mol}^{-1}$). Our experimental results showed that the deuteration reaction ($\text{NH}_3 + \text{D}$) occurs through quantum tunneling, and it is five orders of magnitude slower than methanol (CH_3OH) deuteration process.

If our laboratory experiments lead to the formation of the deuterated ammonia species in comparison to the previous experiments of Nagaoka et al. [22] and Fedoseev et al. [23]. This is because our experimental conditions are different from the other works, and help to overcome the classical, the quantum tunneling, and the diffusion activation barriers of the reactions between NH_3 and D atoms on the oxidized graphite surface. It seems that the main factors that enhance the deuteration reactions between ammonia and D atoms in our experiments are the low D flux and the low thickness of solid NH_3 . Water impurities do not affect the formation of deuterated ammonia, including the effect as a catalyst.

By lowering significantly the D atoms-flux in this work with respect to the previous works of Fedoseev et al. [23] and Nagaoka et al. [22], we increase the density of D atoms available to diffuse on the surface, and interact efficiently with physisorbed ammonia species on the oxidized HOPG surface. However, the D-H exchange reaction between NH_3 and D atoms is almost in competition with the barrier-less recombination $\text{D} + \text{D} \rightarrow \text{D}_2$ reaction. We notice that the used D-flux could be suitable to have H-D substitution reactions in the experiments of Fedoseev et al. [23] and Nagaoka et al. [22], both for methanol and ammonia, if they have reduced the thickness of their ices to a fraction of one monolayer. These authors did not try to reduce simultaneously the thickness of the ices and

the flux of D atoms to study the effect of these two parameters on the the efficiency of the deuteration reactions at low temperatures.

Other factors related to a specific orientation of the graphitic surface, or a possible interaction of ammonia with the substrate by chemisorption process, at low surface temperature at 10 K, can induce the H-D exchange between NH_3 and D atoms. Some experimental [55] and theoretical [56] works reported in the literature have demonstrated the possible dissociative chemisorption of NH_3 molecules on the oxidized graphite surface at the epoxy functional groups, created by the oxidation of the HOPG surface at room temperature [57]. The dissociation of the adsorbed ammonia leads to the formation of $\text{C} - \text{NH}_2$ radicals, which can react with deuterium atoms. If the chemisorption occurs for NH_3 in our experiments, CH_3OH molecules will be also chemisorbed on the oxidized graphite surface, and leads to the formation of the radical CH_2OH . It may be the case, but all the molecules in our experiments desorb at the physisorption temperatures, at 140 K for methanol, 96 K for ammonia species, and 150 K for traces of water ices. In addition, the energy barrier for the dissociative chemisorption of NH_3 on the oxidized HOPG surface has been predicted to be 97.90 $\text{kJ} \cdot \text{mol}^{-1}$ [56]. This activation energy barrier is significantly high to be overcome, in comparison to the activation energy barriers found in this work (16 $\text{kJ} \cdot \text{mol}^{-1}$), and makes chemisorption improbable in our experiments.

Our results for the deuteration of small molecules, such as NH_3 by surface chemistry are important in the fields of astronomy, astrochemistry and low-temperature physics. The formation of the isotopologue ammonia species on cold grain surfaces can contribute to the D-enrichment of ammonia in the interstellar medium, and explain therefore the observed ratios of H-and D-ammonia bearing molecules in dark clouds. However, the amount of the NH_3 molecules expected to be deuterated in dense molecular clouds on a timescale of $10^5 - 10^6$ years is not enough important to reproduce the large gas-phase interstellar abundances of deuterated ammonia molecules. In space, the low reaction probability of NH_3 molecules with D atoms on interstellar grain mantles results from the competitive surface reactions of D atoms with the accreted species (D, N, O, O_2 ...) from the gas phase, leading to the formation of D_2 molecules through D+D recombination [37], NH_2D and ND_3 through D+N addition reactions [23], and heavy D_2O water ices through D+O and D+ O_2 chemical reactions [29, 46, 58].

Observational studies towards cold regions (dark molecular clouds and dense cores L134N) [15, 10, 16] show deuterated ammonia species in gas phase. In these cold environments of the interstellar medium, grain mantles are exposed to cosmic rays and UV irradiation fields originating from hot stars. These energetic UV photons may induce the desorption into the gas phase of the newly formed deuterated species NH_2D , NHD_2 and ND_3 on the grain surfaces by non thermal photo-desorption process [59]. The desorption of the deuterated ammonia species in cold regions may also result from exothermic reactions occurring on the grain, in particular the formation of molecular hydrogen (H_2) [60] by H bombardment. This reaction releases 4.5 eV of excess energy which can be transferred to the grain

surface, and causes the local heating of the deuterated species. It has been demonstrated that such local heating can reorganize the local structure of the ice mantle [61], although it can hardly induce indirect desorption [62].

Because the abundance of ammonia (NH₃) in icy mantles is nearly 15 % with respect to water (H₂O) ice [8, 3], it is interesting to study the efficiency of the deuteration reaction NH₃ + D on amorphous solid water ASW ice surfaces, and explore the role of the ice grain chemistry in the interstellar deuterium fractionation of ammonia molecules.

References

- [1] A.C. Cheung, D.M. Rank, C.H. Townes, D.D. Thornton, W.J. Welch, *Physical Review Letters*. **21**, 1701 (1968). DOI 10.1103/PhysRevLett.21.1701
- [2] T. Hama, N. Watanabe, *Chemical Reviews* **113**, 8783 (2013). DOI 10.1021/cr4000978
- [3] E.L. Gibb, D.C.B. Whittet, A.C.A. Boogert, A.G.G.M. Tielens, *Astrophys. J. Soc.* **151**, 35 (2004). DOI 10.1086/381182
- [4] S. Bottinelli, A.C.A. Boogert, J. Bouwman, M. Beckwith, E.F. van Dishoeck, K.I. Öberg, K.M. Pontoppidan, H. Linnartz, G.A. Blake, N.J. Evans, II, F. Lahuis, *Astrophys. J.* **718**, 1100-1117 (2010). DOI 10.1088/0004-637X/718/2/1100
- [5] D. Bockelée-Morvan, J. Crovisier, M.J. Mumma, H.A. Weaver, ed. M. C. Festou. H. U. Keller, and H. A. Weaver; Univ. of Arizona Press p. 391 (2004)
- [6] E. Dartois, L. d'Hendecourt, *Astron.Astrophys.* **365**, 144 (2001). DOI 10.1051/0004-6361/20000174
- [7] A.G.G.M. Tielens, W. Hagen, *Astron. Astrophys.* **114**, 245 (1982)
- [8] E.L. Gibb, D.C.B. Whittet, W.A. Schutte, A.C.A. Boogert, J.E. Chiar, P. Ehrenfreund, P.A. Gerakines, J.V. Keane, A.G.G.M. Tielens, E.F. van Dishoeck, O. Kerkhof, *Astrophys. J.* **536**, 347 (2000)
- [9] E.N. Rodriguez Kuiper, T.B.H. Kuiper, B. Zuckerman, *Astrophys. J. Lett.* **219**, L49 (1978). DOI 10.1086/182604
- [10] S. Saito, H. Ozeki, M. Ohishi, S. Yamamoto, *Astrophys. J.* **535**, 227 (2000). DOI 10.1086/308818
- [11] R.Y. Shah, A. Wootten, *Astrophys. J.* **554**, 933 (2001). DOI 10.1086/321396
- [12] S. Tiné, E. Roueff, E. Falgarone, M. Gerin, G. Pineau des Forêts, *Astron.Astrophys.* **356**, 1039 (2000)
- [13] J.L. Linsky, *Space Sci. Rev.* **106**, 49 (2003)
- [14] J. Hatchell, *Astron.Astrophys.* **403**, L25 (2003). DOI 10.1051/0004-6361:20030297
- [15] E. Roueff, S. Tiné, L.H. Coudert, G. Pineau des Forêts, E. Falgarone, M. Gerin, *Astron.Astrophys.* **354**, L63 (2000)
- [16] F.F.S. van der Tak, P. Schilke, H.S.P. Muller, D.C. Lis, T.G. Phillips, M. Gerin, E. Roueff, *Astron. Astrophys.* **388**, L53 (2002). DOI 10.1051/0004-6361:20020647
- [17] D.C. Lis, E. Roueff, M. Gerin, T.G. Phillips, L.H. Coudert, F.F.S. van der Tak, P. Schilke, *Astrophys. J. Lett.* **571**, L55 (2002). DOI 10.1086/341132
- [18] H. Roberts, T.J. Millar, *Astron.Astrophys.* **364**, 780 (2000)
- [19] E. Roueff, D.C. Lis, F.F.S. van der Tak, M. Gerin, P.F. Goldsmith, *Astron. Astrophys.* **438**, 585 (2005). DOI 10.1051/0004-6361:20052724
- [20] A.G.G.M. Tielens, *Astron. Astrophys.* **119**, 177 (1983)
- [21] H. Roberts, E. Herbst, T.J. Millar, *Astrophys. J.* **424**, 905 (2004). DOI 10.1051/0004-6361:20040441
- [22] A. Nagaoka, N. Watanabe, A. Kouchi, *Astrophys. J. Lett.* **624**, L29 (2005). DOI 10.1086/430304
- [23] G. Fedoseev, S. Ioppolo, H. Linnartz, *Mon. Not. R. Astron. Soc.* **446**, 449 (2015). DOI 10.1093/mnras/stu1852
- [24] M.J. Kurylo, A.H. Gerard, F.L. Henry, B.T. Richard, *J. Chem. Phys.* **51**, 4497 (1969). DOI http://dx.doi.org/10.1063/1.1671817
- [25] G.E. Moyano, M.A. Collins, *Theor. Chem. Acc.* **113**, 225 (2005). DOI 10.1007/s00214-004-0626-8
- [26] A. Nagaoka, N. Watanabe, A. Kouchi, *J.Phys.Chem.A.* **111**(16), 3016 (2007). DOI 10.1021/jp068978r
- [27] L. Amiaud, F. Dulieu, J.H. Fillion, A. Momeni, J.L. Lemaire, *J. Chem. Phys.* **127**(14), 144709 (2007). DOI 10.1063/1.2746323
- [28] H. Chaabouni, H. Bergeron, S. Baouche, F. Dulieu, E. Matar, E. Congiu, L. Gavilan, J.L. Lemaire, *Astron.Astrophys.* **538**, A128 (2012). DOI 10.1051/0004-6361/201117409
- [29] S. Ioppolo, H.M. Cuppen, C. Romanzin, E.F. van Dishoeck, H. Linnartz, *Physical Chemistry Chemical Physics (Incorporating Faraday Transactions)* **12**, 12065 (2010). DOI 10.1039/C0CP00250J
- [30] P. Theule, F. Borget, F. Mispelaer, G. Danger, F. Duvernay, J.C. Guillemin, T. Chiavassa, *Astron. Astrophys.* **534**, A64 (2011). DOI 10.1051/0004-6361/201117494
- [31] D.A. King, M.G. Wells, *Surface Science* **29**, 454 (1972). DOI http://dx.doi.org/10.1116/1.1317816
- [32] S.E. Bisschop, G.W. Fuchs, E.F. van Dishoeck, H. Linnartz, *Astron. Astrophys.* **474**, 161 (2007). DOI 10.1051/0004-6361/201117494
- [33] J.R. Burke, D.J. Hollenbach, *Astrophys. J.* **265**, 223 (1983). DOI 10.1086/160667
- [34] E. Matar, H. Bergeron, F. Dulieu, H. Chaabouni, M. Accolla, J.L. Lemaire, *J. Chem. Phys.* **133**(10), 104507 (2010). DOI 10.1063/1.3484867
- [35] T. Zecho, A. Güttler, X. Sha, D. Lemoine, B. Jackson, J. Küppers, *Chemical Physics Letters* **366**, 188 (2002). DOI 10.1063/1.2746323
- [36] S. Cazaux, P. Caselli, A.G.G.M. Tielens, J. LeBourlot, M. Walmsley, *Journal of Physics Conference Series* **6**, 155 (2005). DOI 10.1088/1742-6596/6/1/016
- [37] L. Hornekaer, E. Rauls, W. Xu, Ž. Šljivančanin, R. Otero, I. Stensgaard, E. Lægsgaard, B. Hammer, F. Besenbacher, *Physical Review Letters* **97**(18), 186102 (2006). DOI 10.1103/PhysRevLett.97.186102
- [38] NIST. National institute of standards and technology (2014). <http://webbook.nist.gov/chemistry/>
- [39] E. Matar, E. Congiu, F. Dulieu, A. Momeni, J.L. Lemaire, *Astron. Astrophys.* **492**, L17 (2008). DOI 10.1051/0004-6361:200810434
- [40] K. Hiraoka, A. Wada, H. Kitagawa, M. Kamo, H. Unagiike, T. Ueno, T. Sugimoto, T. Enoura, N. Sogoshi, S. Okazaki, *Astrophys. J.* **620**, 542 (2005). DOI 10.1086/426958
- [41] A. Ratajczak, E. Quirico, A. Faure, B. Schmitt, C. Ceccarelli, *Astron. Astrophys.* **496**, L21 (2009). DOI 10.1051/0004-6361/200911679
- [42] A. Nagaoka, N. Watanabe, A. Kouchi, in *Astrochemistry - From Laboratory Studies to Astronomical Observations, American Institute of Physics Conference Series*, vol. 855, ed. by R.I. Kaiser, P. Bernath, Y. Osamura, S. Petrie, A.M. Mebel (2006), *American Institute of Physics Conference Series*, vol. 855, pp. 69–75. DOI http://dx.doi.org/10.1063/1.2359541
- [43] M. Minissale, E. Congiu, G. Manicò, V. Pirronello, F. Dulieu, *Astron. Astrophys.* **559**, A49 (2013). DOI 10.1051/0004-6361/201321453
- [44] M. Minissale, J.C. Loison, S. Baouche, H. Chaabouni, E. Congiu, F. Dulieu, *Astron. Astrophys.* **577**, A2 (2015). DOI 10.1051/0004-6361/201424342
- [45] A.S. Bolina, W.A. Brown, *Surface Science* **598**, 45 (2005). DOI 10.1016/j.susc.2005.08.025
- [46] F. Dulieu, E. Congiu, J. Noble, S. Baouche, H. Chaabouni, A. Moudens, M. Minissale, S. Cazaux, *Scientific Reports* **3**, 1338 (2013). DOI 10.1038/srep01338
- [47] M. Bonfant, R. Martinazzo, J.F. Tantardini, A. Ponti, *J. Phys. Chem. C* **111**(16), 5825 (2007). DOI 10.1021/jp070616b
- [48] H. Hidaka, M. Watanabe, A. Kouchi, N. Watanabe, *Astrophys. J.* **702**, 291 (2009). DOI 10.1088/0004-637X/702/1/291
- [49] M.D. Harmony, *Chem. Phys. Lett.* **10**, 337 (1971). DOI http://dx.doi.org/10.1016/0009-2614(71)80303-2
- [50] V.I. Goldanskii, *Ann. Rev. Phys. Chem.* **27**, 85 (1976). DOI 10.1146/annurev.pc.27.100176.000505
- [51] T.P.M. Goumans, J. Kästner, *Journal of Physical Chemistry A* **115**, 10767 (2011). DOI 10.1111/j.1365-2966.2011.18329.x
- [52] R.P. Bell, *The tunnel effect in chemistry: Chapter 6: Tunneling in molecular spectra, the inversion of ammonia and related processes; page 153.* (1980)
- [53] N. Watanabe, A. Kouchi, *Progress In Surface Science* **83**, 439 (2008). DOI 10.1016/j.progsurf.2008.10.001
- [54] R. Rejoub, B.G. Landsay, R.F. Stebbings, *J. Chem. Phys.* **115**, 5053 (2001). DOI 10.1063/1.1394748
- [55] T.F. Yeh, S.J. Chen, C.S. Yeh, H. Teng, *J. Phys. Chem.C* **117**, 6516 (2013). DOI 10.1021/jp312613

- [56] S. Tang, Z. Cao, *J. Phys. Chem. C* **116**, 8778 (2012). DOI 10.1021/jp212218w
- [57] R. Larciprete, P. Lacovig, S. Gardonio, A. Baraldi, S. Lizzit, *J. Phys. Chem. C* **116**, 9900 (2012). DOI 10.1021/jp2098153
- [58] H. Chaabouni, M. Minissale, G. Manicò, E. Congiu, S. Noble J. A, Baouche, M. Accolla, J. Lemaire, V. Pirronello, F. Dulieu., *J. Chem. Phys.* **137**, 234706 (2012). DOI <http://dx.doi.org/10.1063/1.4771663>
- [59] T.W. Hartquist, D.A. Williams, *Mon. Not. R. Astron. Soc.* **247**, 343 (1990)
- [60] W.W. Duley, D.A. Williams, *Mon. Not. R. Astron. Soc.* **260**, 37 (1993). DOI 10.1093/mnras/260.1.37
- [61] M. Accolla, E. Congiu, F. Dulieu, G. Manicò, H. Chaabouni, E. Matar, H. Mokrane, J.L. Lemaire, V. Pirronello, *Physical Chemistry Chemical Physics (Incorporating Faraday Transactions)* **13**, 8037 (2011). DOI 10.1039/c0cp01462a
- [62] M. Minissale, A. Moudens, S. Baouche, H. Chaabouni, F. Dulieu, *Mon. Not. R. Astron. Soc.* **458**, 2953 (2016). DOI 10.1093/mnras/stw373