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# 1 **3D morphology and timing of the giant fossil pockmark of Beauvoisin, SE Basin of**

## 2 **France**

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18 Abbreviated title: Morphology and timing of a fossil giant pockmark

## 19 **1. Abstract**

20 The resolution of data acquired over modern seafloors does not allow imaging of the inner features of a  
21 fluid seep structure, particularly in the shallow subsurface. In the South-East Basin of France (Drôme), fossil  
22 cold seep structures comprising fossil-rich carbonate lenses were identified about 30 years ago within the  
23 Oxfordian (Late Jurassic) Terres Noires Formation. These structures were first interpreted as pseudo-  
24 bioherms related to hydrothermal activity, but comparison with active seep sites on modern margins,  
25 together with isotopic analyses led to a re-interpretation involving cold fluids instead. To date, all seep sites  
26 have generally been studied individually without considering any link to neighboring or more distant sites.  
27 Based on 23 high-resolution stratigraphic logs within the structure coupled to mosaicked aerial  
28 photographs from a drone survey, the 19 fluid seep events were correlated in the area, including two new  
29 sites exposed due to weathering. We have shown that each identified sub-site is composed of sub-vertically

30 stacked fossil-rich carbonate lenses interbedded with marls, which developed in smooth, 4 to 6 m deep  
31 depressions beneath the local seabed. The nodules present within the depressions are of primary  
32 importance as they mark the area of active seeping. This general organization is very similar to the modern  
33 Regab giant pockmark in the Lower Congo Basin where only a few sub-sites are active at the same time. A  
34 spatio-temporal 3D reconstruction of the position of these sub-sites shows that the carbonate lenses are  
35 organized into clusters comprising up to 7 sub-sites grouped together in the same stratigraphic interval and  
36 the same geographic zone. A sandbox experiment where gas is injected at constant flow rate at the base of  
37 a box filled with a matrix of water-saturated grains displays a pattern consisting of disturbed sediments  
38 inside a parabolic-shaped area. This parabolic shape was also identified on a seismic profile across the  
39 Regab giant pockmark, suggesting that the processes are similar for the Beauvoisin and Regab seep areas.  
40 The laboratory experiments also show that the seeping conduit is stable during a given period of time and  
41 suddenly shifts laterally. It is mainly due to the collapse of the conduit, the lateral migration and the  
42 reopening at a new position. The general log obtained in the Beauvoisin seep area suggests a similar  
43 pattern with periods of seeping alternating with periods of quiescence, each of which is approximately 200  
44 ka. Even if a pockmark seems to have been inactive for a long period of time, it could be due to the lateral  
45 shift of the feeder conduit meaning that the sub-seafloor is still charged in gas. This is of primary  
46 importance for risk assessment, hydrocarbon exploration and general understanding of geobiology at  
47 seafloor seeps.

#### 48 **200 words short abstract:**

49 In the SE Basin of France (Drôme), fossil cold seep structures made of fossil-rich carbonate lenses were  
50 identified in the Oxfordian (Late Jurassic) Terres Noires Formation about 30 years ago. To date, all seep  
51 sites have generally been studied individually without consideration of any link to neighboring or more  
52 distant sites. Based on a detailed fieldwork in the Beauvoisin area and comparison with the modern, active  
53 giant Regab pockmark and sandbox experiments, we suggest that this site can now be considered as a 800  
54 m wide fossil analogue of a giant pockmark. It comprises several 4-6 m deep, 80-120 m wide coalesced sub-  
55 seep sites. A spatio-temporal 3D reconstruction of the position of these sub-sites shows that the carbonate  
56 lenses are organized in clusters with only one cluster active at a given period. Two periods of seep activity  
57 are separated by a period of quiescence due to the lateral shift of the feeder conduit beneath the  
58 structure. This suggests that even if a seafloor pockmark appears to be inactive, this could be due to the  
59 lateral shift of the feeder conduit meaning that the sub-seafloor is still charged in gas.

## 60 **2. Introduction**

61 Hydrogen sulfide and methane that sustain highly specialized chemosynthetic-based ecosystems in cold  
62 seep environments have been recognized in many modern basins worldwide (*Campbell, 2006; Amano et*  
63 *al., 2010; Teichert and Van de Schrootbrugge, 2013*) and in the fossil record in the early Cenozoic -  
64 onwards (*Campbell et al., 1995b; Kiel and Little, 2006; Kiel, 2010; Vrijenhoek, 2013*). Sulfide and methane

65 release results in carbonate precipitation forming lenses encased in surrounding limestones or marlstones  
66 ([Campbell et al., 2002](#)). These fossil organism-rich, authigenic, carbonate rocks developed at the seafloor  
67 where oil and/or gas migrated from an underlying network of focused fluid flow in sedimentary basins  
68 ([Bohrmann et al., 1998](#); [Aloisi et al., 2000](#); [Aloisi et al., 2002](#)). The fossils dominated by lucinid bivalves  
69 generally show larger dimensions than at other sites as they rely on chemotrophic symbionts for nutrition  
70 ([Paull et al., 1992](#)). This biogeochemical process mainly relies on anaerobic oxidation of methane (AOM)  
71 ([Boetius et al., 2000](#)), where the symbiotic consortia of sulfate-reducing bacteria and methanotrophic  
72 archaea use both sulfate and methane for their metabolism and biomass synthesis, thus supplying energy  
73 and nutrients to their hosts ([Knittel and Boetius, 2009](#); [Deusner et al., 2014](#)). However, fossil-rich  
74 carbonate assemblages have always been considered as individual seep sites without comparing their  
75 distribution to the geometrical organization of modern clusters of seep sites. This lack of connection  
76 between isolated seep sites can be due 1) to poor outcrop conditions or stratigraphic uncertainty  
77 ([Natalicchio et al., 2015](#)), or 2) to the large size of modern pockmarks that are several hundred meters in  
78 diameter and a few meters only in depth (typically 500-800 m wide and 5-80 m deep) leading to difficult or  
79 incorrect correlations in the field ([Gay, 2002](#); [Stewart, 2015](#)). To date, a fossil giant pockmark has never  
80 been identified nor characterized at outcrop, although some microbial assemblages have been  
81 reinterpreted in the light of examination of modern hot vents at mid-ocean ridges and cold seeps at  
82 continental margins ([Whiticar, 1999](#); [Gay et al., 2006](#)). This is the case for the fossil seep sites in the  
83 Jurassic limestones of Beauvoisin in the SE Basin of France, which were first interpreted as hydrothermal  
84 vents ([Rolin, 1987](#); [Gaillard et al., 1996](#)) and then reinterpreted as cold seeps ([Peckmann et al., 1999](#); [Kiel,  
85 2013](#)). The Beauvoisin seep site is of high importance as it is one of the few worldwide examples that is  
86 equivalent to Cenozoic chemosymbiotic-related fossil sites ([Campbell et al., 1995a](#); [Louis-Schmid et al.,  
87 2007](#)). All seep sites have been generally studied individually without consideration of any link to  
88 neighboring sites even those situated in more distant locations ([Louis-Schmid et al., 2007](#)). Furthermore, in  
89 the fossil record isolated carbonate lenses or -groups of small-scale carbonate lenses, about 10 m wide and  
90 5 m high, have often been considered as individual pockmarks or as a field of pockmarks respectively  
91 ([Agirrezabala et al., 2013](#)).

92 The aim of this paper is to characterize the 3D architecture of a 800 m-wide cluster of fossil seep sites at  
93 Beauvoisin in the SE Basin of France, in the vicinity of Buis les Baronnies and Propiac (Drôme) ([Fig. 1](#)), based  
94 on the correlation of detailed sedimentological logs, sampling of key seep features such as carbonate  
95 lenses, marls, nodules and related petrographical and geochemical analyzes. The Beauvoisin architecture is  
96 then compared to that of the Regab giant pockmark in the Lower Congo Basin and to sandbox and  
97 numerical models simulating fluid injection at the base of a box filled with a matrix of water-saturated  
98 grains.

### 99 3. Methods and data

100 Since the first studies of the Beauvoisin site conducted in the 80's (*Gaillard et al., 1985; Rolin, 1987, Rolin*  
101 *et al., 1990*), recent erosion has exposed new sites and to date, 19 embedded carbonate masses (sub-sites  
102 A to T, **see Fig. 2 for location**) have been identified at various positions in the stratigraphic record, using the  
103 initial nomenclature of *Rolin (1987)*. This suggests that all sites are not cropping out due to vegetal cover  
104 and agricultural fields and some sub-sites are probably still to be found. In order to define precisely the  
105 stratigraphical relationship between the encasing fine-grained marls and the fossil-rich carbonate units,  
106 twenty-three stratigraphical logs were obtained through the entire area using a Jacob's stick, between 2002  
107 and 2016. The precision is about 2 cm for each log. Field analysis of the Beauvoisin unit comprises  
108 identification of major stratigraphic and sedimentological features. The general correlation is related to the  
109 Tethyan sequences (*Gradstein et al., 2012*) and some key layers were easily identified (**Fig. 3**), such as the  
110 few cm-thick bentonite layers documented by *Pellenard, (2003)*, or the sub-continuous 10 cm to 1 m thick  
111 layers of massive carbonate layers R1 to R20 of the Argovian sequence identified by *Gaillard et al. (1992)*.  
112 This study is then based on the analysis of the macroscopic fabric of carbonate lenses related to the  
113 reference layers in the surrounding marls of the "Terres Noires" Formation. Sampling for petrographical,  
114 mineralogical and geochemical analyses of the carbonate deposits was carried out on steep slopes in marls.  
115 Where thalwegs were not accessible for sampling or mapping, a drone was used for high-resolution  
116 photography. The obtained mosaics were coupled to aerial photographs and then georeferenced to  
117 topographic maps (IGN map #3140ET) in order to establish the areal distribution of outcrop (**Fig. 2**).

118 Analyses were conducted at the University of Barcelona for C and O isotopes. Forty-eight microsamples  
119 were prepared after the petrographic study to determine the carbon and oxygen stable isotope ratio of the  
120 different cements using the standard technique of Craig and Gordon (1965) and Claypool et al., (1980). The  
121 CO<sub>2</sub> was extracted from 60±10 µg of powdered carbonate samples which were reacted with 103%  
122 phosphoric acid for 2 min at 70° C for calcites. The CO<sub>2</sub> was analyzed using an automated Kiel Carbonate  
123 Device attached to a Thermal Ionization Mass Spectrometer Thermo Electron (Finnigan) MAT-252. The  
124 results are precise to ±0.02‰ for δ<sup>13</sup>C and ±0.04‰ for δ<sup>18</sup>O.

125 The general organization of the fossil seep sites of Beauvoisin was then compared to the modern and active  
126 Regab pockmark in the Lower Congo Basin (*Ondreas et al., 2005; Gay, 2002; Gay et al., 2006; Marcon et*  
127 *al., 2014*). Bathymetry and imaging data were acquired during various surveys between 1999 (ZAIANGO)  
128 and 2011 (WACS) using IFREMER's remotely operated vehicle (ROV) *Victor 6000*. The first survey was  
129 conducted during a site exploration and mosaicking at about 2 m above the sea-bottom. The main survey  
130 was then conducted from 30 m above the seafloor using a Reson Seabat 7125 multibeam echosounder  
131 (MBES) running at 400 kHz.

132 The sandbox model consists of a Plexiglas cylindrical tank (24 cm diameter) for the 3D and a vertical Hele-  
133 Shaw cell (glass plates 40 × 30 cm, separated by a 2 mm gap) for the 2D laboratory experiments. Both

134 systems are filled with particles immersed in water and a constant flow of air,  $\Phi$ , is injected from a single  
135 inlet centered at the base of the cell. The particles are polydispersed, spherical glass beads (Sovitech glass  
136 spheres), sieved to obtain four batches with particle diameters  $d = 218 \pm 17$ ,  $318 \pm 44$ ,  $631 \pm 37$ , and  $802 \pm$   
137  $68 \mu\text{m}$ . The grain size distribution for the different batches is measured by means of a microscope (Optika  
138 B-163) and roughly displays a Gaussian shape (*Ramos et al., 2015*). In the 2D experiment the grain  
139 dynamics are tracked by analyzing the absolute difference in the intensity of two consecutive images,  
140 which gives access to regions where motion occurred in the immersed granular layer, due to the ascending  
141 gas flow. In order to quantify the generation of the fluidized zone, which corresponds to a cumulative  
142 process due to the continuous gas emission, we define the normalized flow density, computed as the  
143 cumulation of successive image differences (*Ramos et al., 2015; Varas et al., 2015; Poryles et al., 2016*).  
144 This variable makes it possible to quantify the regions where motion occurred (disturbed particles).

#### 145 4. Geological setting

146 The Beauvoisin seep site is located in the northern part of the South East Basin of France (**Fig. 1**) which is  
147 related to the Jurassic opening of the Liguro-tethyan Ocean (*Lemoine 1985*). From the Hettangian to the  
148 Bathonian the basin was a shallow-water carbonate platform with frequent subaerial exposure or erosion  
149 indicating a stable margin with a low rate of subsidence. From the Bathonian, sedimentation compensated  
150 for the moderate subsidence of the southern platform and environments remained shallow, whereas along  
151 the northern platform cherts and organic-rich marls indicate a general deepening of at least 600 m  
152 (*Dardeau, 1988*). From the Oxfordian, general subsidence affected the entire platform, moderate in the  
153 southern part (reefal facies and confined environments) and increasing in the northern part (deep marl  
154 facies and submarine slides).

155 In the Beauvoisin area the identified seep sites occur in the deepest, central part of the basin where the  
156 subsidence was at a maximum from the Bathonian to the Oxfordian (*Gaillard et al. 1985*). This led to  
157 deposition of up to 2000-2500 m of organic-rich marls called the "Terres Noires" Formation, limited to the  
158 North by the Jura high, to the West by the Cevennes domain and to the East by the Briançonnais domain of  
159 the Western Alps (**Fig. 1**). The "Terres Noires" Formation has also been identified in boreholes in the south  
160 in Camargue and Provence (*Dardeau, 1988*). Subsidence was mostly due to basement faults with  
161 kinematics controlled by salt withdrawal in the extensional domain of the margin (*Masclé et al, 1988*). The  
162 fluid seep sites are bounded to the west by major salt-rooted faults which facilitated growth of the main  
163 salt diapirs into the marls, as at Propiac and Condorcet (**Fig. 1**).

164 The "Terres Noires" Formation is divided into three main intervals corresponding to the three major  
165 geodynamical episodes that affected the margin (*Gaillard et al., 1985; Gaillard et al., 1988; Rolin, 1987;*  
166 *Rolin et al., 1990*). The Bathonian sequence is characterized by marls alternating with thin mudstone  
167 layers. The Callovian to Middle Oxfordian sequence is dominated in the lower part by cm to pluri-cm thick  
168 dolomudstones and in the middle and upper part by marls containing isolated red to grey nodules

169 ("Nodules Chocolat" by [Artru, 1972](#)). The upper Oxfordian is composed of marls alternating with pluri-cm  
170 to m thick dolomudstones layers in a general thickening upward sequence. The transition to the  
171 Kimmeridgian is marked by the Argovian sequence composed of thick amber dolomudstone layers  
172 alternating with light brown marls.

173 To date the origin of focused fluids has never been clearly elucidated. They were hypothetically related to  
174 crustal faults ([Lemoine, 1988](#)), to halokinesis ([Masclé et al., 1988](#)) or to biogenic origin derived from the  
175 decomposition of organic matter contained in the "Terres Noires" Formation ([Gaillard et al., 1996](#)).  
176 Fluorescence conducted on fluid inclusions coupled to stable carbon isotope analyses have shown the  
177 presence of oil containing n-alkanes generated from thermal maturation of organic matter from the  
178 sedimentary pile ([Peckman et al., 1999](#)). The latter is consistent with the total thickness of the "Terres  
179 Noires" Formation which has clearly reached a burial depth through the oil window in the lower part.

## 180 5. Log correlations

181 The fossil-rich carbonate suite of Beauvoisin crops out within the thick succession of well-stratified Jurassic  
182 marls of the "Terres Noires" Formation in the Lower to Upper Oxfordian interval ([Gaillard et al. 1985](#)).  
183 Between 2002 and 2016, 23 stratigraphical logs were collected allowing a detailed correlation of the 19  
184 sites (A to T) related to the general stratigraphy of the Oxfordian marls ([Fig. 2](#)). All sites are concentrated  
185 west of Beauvoisin and south of the "Col de la Taillade" ([Fig. 2](#)) where they are exposed in very steep  
186 valleys between 660 m and 870 m altitude. This work is based on the initial nomenclature established by  
187 [Gaillard et al. \(1985\) and Rolin \(1987\)](#). They grouped together carbonate lenses in sites from A to R ([Rolin  
188 et al., 1990](#)). The intense weathering in spring and fall in the region for the last 30 years allowed 2 new sites  
189 to be exposed in the Terres Noires Formation. The sites S and T were integrated into the general  
190 stratigraphy of Beauvoisin in Logs 11 and 15, respectively ([Fig. 2](#)).

191 The 23 stratigraphical logs were constrained using key layers, R1 to R20 following the nomenclature of  
192 [Gaillard et al. \(1985\) and Rolin \(1987\)](#). However, additional layers (not displayed in this study) were also  
193 used following [Gaidon \(1988\), Pellenard et al. \(1999\) and Pellenard et al. \(2003\)](#). This included the  
194 identification of a bentonite layer, located at about 220 m above the bottom of Log 1 ([Fig. 3](#)), which is the  
195 reference for all measurements in this study. Some cm-thick ochre and gypsum-rich layers were also  
196 recognized in the area. They can be correlated over all the basin and mark some regional events ([Pellenard  
197 et al., 1999](#)). The correlation is centered and flattened on the R1 level corresponding to the base of the  
198 Argovian sequence ending on top with the R6 level. It is characteristic of the upper part of the Oxfordian in  
199 the SE Basin of France and it is composed of dm to m, yellowish to light brown, limestones alternating with  
200 light grey marls and claystones.

201 It is worth noting that the thicknesses of the Argovian sequence varies by a factor of 3 to 4 over a distance  
202 of 700 m which is the longest distance between two logs in the area. For instance, this is the case for the



203 R2-R6 interval which is about 110 m thick in log 16 (North), 60 m thick in log 8 (centre of the area), 30 m  
204 thick in log 9 (West) and 90 m thick in log 13 (South West) (see Fig. 2 for location). In particular, log 9  
205 shows evidence of intense deformations due to slumps at 260 m and 275 m (Fig. 3). The thickness  
206 variations in the Argovian sequence could be due to syn-to-post depositional erosion or slump processes  
207 which increase towards the centre of the area.

208 Based on the log correlation it is then possible to generate a composite log taking into account the eroded  
209 missing sequences, slump deposits and faults (Fig. 4). When varying laterally, the thickest part of a  
210 sequence was kept in order to correspond to the 360 m total thickness of the studied Oxfordian interval.  
211 On this composite log, the oldest site A is located within the *Cordatum* Zone (Lower Oxfordian) whereas the  
212 youngest site T is in the *Bifurcatis* Zone (Upper Oxfordian). Given the uncertainty of absolute dating in the  
213 Jurassic, the Tethysian sequences and age model from Gradstein et al. (2012) are reported on the right side  
214 as an indication only. It shows that only two sites (A and S) developed in the *Cordatum* Zone corresponding  
215 to the Lower Oxfordian whereas most sites developed in the *Plicatilis* Zone (B, C, D, E, F, G, H, I, J, K) and in  
216 the *Transversarium* Zone (L, M, N, O, P, Q, R, T) that are Middle Oxfordian in age and only 2 sites (R and T)  
217 developed in the *Bifurcatis* Zone corresponding to the Upper Oxfordian (Fig. 4). Given an average rate of  
218 sedimentation and a constant compaction rate within the interval, 310 m were deposited in 3.1 Ma  
219 between 157,4 and 160,5 Ma. The base of site A1 at 15 m and the top of site T at 359 m can be estimated  
220 at 157.14 Ma and 160.58 Ma respectively as a first approximation. Despite 15 years of fieldwork in the  
221 area, no more sites have been discovered beneath site A or above site T. However, it does not mean that  
222 they do not exist as they may have not been exposed yet, like the newly discovered sites S and T. To date,  
223 the sites A to T can be estimated for a total of 3.44 My in duration with an average rate of 10 cm per 1000  
224 years of sedimentation.

## 225 6. Vertical organization of seep facies

226 The 19 sites A to T correspond to carbonate lenses encased in marls or claystones and forming local  
227 unconformities in the Terres Noires Formation. Each site taken individually is organized as sub-vertically  
228 stacked, 2-15 m wide, lenses that are in contact (i.e., the top of an underlying lens is in sharp contact with  
229 the bottom of the overlying lens) or interbedded with marls. As for sub-site F (Fig. 5A), the three basal  
230 carbonate lenses, 1 to 1.5 m thick and 5 m wide, are in contact with each other and the two top carbonate  
231 lenses, 0.5 to 1 m thick and 3 m wide, are interlayered with a 0.8 to 1 m thick interval of nodule-rich marls.  
232 Sub-site F forms a 7 m high edifice composed of sub-vertically stacked carbonate lenses, lenticular in shape.  
233 The two first basal lenses are brecciated showing various-size, cemented, sub-angular clasts (Fig. 5B). In  
234 detail, two major types of breccia were identified:

235 > A matrix-supported cemented macro-breccia made up of cm to dm sub-angular and microsparite clasts  
236 within a yellowish or light brown microbial carbonate matrix.



237 > A dark brown cement-supported micro-breccia made up of mm to pluri-mm clasts within a slightly oil (or  
238 bitumen) impregnated matrix.

239 This brecciated facies is in close association with sub-vertical mineralized veins (Fig. 5B). The veins are filled  
240 by several generations of carbonate cement. The first generation is an iron-rich drusy calcite, 20-300 mm  
241 thick followed by bladed, high Mg-calcite or aragonite and then by 550 mm of botryoidal aragonite forming  
242 the final cement generation (Peckman et al., 1999). The matrix-supported cemented macro-breccia  
243 comprises clasts, the shape of which is concordant with the massive matrix of the carbonate lenses  
244 suggesting a puzzle-like structure (Fig. 5B&C). The dark brown cement-supported micro-breccia is more  
245 concentrated towards the veins and at some places surrounding the mineralization.

246 The carbonate lenses are dominated by mollusk macrofossils mostly represented by lucinid bivalves  
247 *Beauvoisina carinata* (Gaillard et al. 1985, Rolin et al., 1990; Gaillard et al., 1992), gastropods including  
248 *Paskentana umbilicata*, *Hokkaidoconcha novacula* (Kiel, 2013) and cephalopods such as ammonites (Rolin  
249 et al., 1990). Lucinids are usually restricted to carbonate lenses and they are not found in surrounding  
250 marls suggesting that they were endofauna. They can form dense groups with some lucinid specimens  
251 reaching 18 cm in diameter (Rolin, 1987). The shells provide average positive  $\delta^{13}\text{C}$  values as high as +5 ‰  
252 PDB (Peckman et al., 1999), probably due to symbiosis with chemosynthetic bacteria (Rolin et al. 1990;  
253 Gaillard et al. 1992). However, the  $\delta^{13}\text{C}$  values range between -26.5 ‰ to +13 ‰ PDB in carbonate lenses  
254 (Peckman et al., 1999), which is consistent with the more recent values measured at site F by Tribovillard  
255 et al. (2013) (-18.8 ‰ <  $\delta^{13}\text{C}$  < +12.7 ‰ PDB). In the basal lenses of site F, some lucinid specimen, 12 to 15  
256 cm wide, were observed very close to the brecciated facies and mineralized veins (Fig. 5C). Some other  
257 species have been identified, such as crustacean exoskeletons, fragments and coprolites, belonging to the  
258 form-genus *Favreina* and attributed to the anomuran superfamilies *Thalassinoidea* and *Galattheoidea*, as  
259 well as fish teeth (reflecting the presence of additional predators or scavengers), sponge spicules and the  
260 irregular echinoid *Tithonia oxfordiana* (Gaillard et al., 1985; Gaillard et al., 1992; Senowbari-Daryan et al.,  
261 2007; Kiel et al., 2010; Gaillard et al., 2011). Probable deposit feeders, such as holothuroids (*Sclerites*), and  
262 suspension feeders, such as crinoids (ossicles), are found in minor abundance (Gaillard et al., 2011).  
263 Benthic foraminifers (such as *Spirillinidae*, *Nodosariidae*, *Textulariidae* and *Ophthalmidium*), ostracods,  
264 planktonic foraminifers (proto-*globigerinids*), radiolarians and dinoflagellates are present but less  
265 numerous. Sponges are also frequent, mainly *Lyssacid hexactinellids*, accompanied by *Lithistid*  
266 demosponges, unidentified demosponges and lychniscid hexactinellids (Gaillard et al., 2011). These biota  
267 are common in the Jurassic seafloor in the area suggesting a bathyal environment estimated around 600 m  
268 water depth (Tribovillard et al, 2013). Ovoid to irregularly shaped fecal pellet concentrations locally  
269 scattered in the micrite indicate benthic activity (Gaillard et al., 1985, Gaillard et al., 1992; Gaillard et al.,  
270 2011).

271 The Oxfordian seep carbonates are very rich in micritic nodules, mm to 15 cm in diameter depending on  
272 their position relative to carbonate lenses (**Fig. 5D**). These concretions formed around body fossils, such as  
273 ammonites, bivalves, spicules and any biotritus or burrows (*Gaillard et al., 1985; Rolin, 1987*). Nodules  
274 are darker than the micritic matrix and are densely packed within the basal carbonate lenses where they  
275 form mm to 2 cm aggregates encased in a dark micrite (**Fig. 6A**). They contain framboidal pyrite, 40 to 250  
276 mm in diameter, and are often lined by an outer rim of pyrite (*Peckman et al., 1999; Gay, 2002*). The  $\delta^{13}\text{C}$   
277 values in nodules range between -26.5 ‰ and -22.8 ‰ PDB (*Peckman et al., 1999*), with minimum values  
278 similar to the encasing carbonate lenses. They become scattered towards the margins and they can form 5  
279 to 10 cm-long twin nodules where the carbonate lenses thin (**Fig. 6B**). Twin nodules are separated from  
280 each other over a distance between a few dm to about 1 m. These nodules appear at the same  
281 stratigraphical level as 15 cm-large nodules separated from each other by a few dm to a few meters (**Fig.**  
282 **6B**). Other small nodules, 15 cm to mm in diameter can be found at the same stratigraphical level over a  
283 distance of 30 to 60 m from the centre of the carbonate lens (**Fig. 6B**). However, their diameter decreases  
284 as the distance to each other increases with distance from the carbonate lens. Beyond 60 m from the  
285 carbonate lens, they fully disappear. Bitumen is frequent within the nodules encased in marls and not in  
286 nodules encased in carbonate lenses. Other types of concretions related to sponge taphonomy and  
287 burrowing were locally described by *Gaillard et al. (2011)*.

288 The carbonate lenses are composed of a micritic matrix rich in Mg-calcite, aragonite and dolomite, which  
289 are the common authigenic minerals at cold seeps (*Roberts et al. 1993*). Previous analyses conducted on  
290 marls show  $\delta^{13}\text{C}$  values ranging between -19 ‰ to -17.7 ‰ far from the carbonate lenses (*Peckman et al.,*  
291 *1999*). These values increase close to the carbonate lenses ( $-0.7\text{ ‰} < \delta^{13}\text{C} < +1.1\text{ ‰}$ ), which was also reported  
292 by *Tribovillard et al. (2013)*. In addition, different types of carbonate cements were identified, such as  
293 splayed calcite, yellow calcite and botryoidal aragonite and calcite (*Beauchamp and Savard 1992*). In  
294 particular, the botryoidal fabric is presumed to be of bacterial origin (*Roberts et al. 1993*), with values of  
295  $\delta^{13}\text{C}$  between -14.8 ‰ to -12 ‰ (*Peckman et al., 1999*) reinforcing the role of microbial mediation in the  
296 building of carbonate lenses.

297 The top of the upper lens of many sub-sites is affected by pervasive corrosion creating vugs and irregular  
298 surfaces. Such as on top of sub-site G (**see Fig. 2 for location**), the corrosion formed a pluri-cm thick Mn-  
299 and Fe-rich crust (**Fig. 6C**), which is often associated with pyrite. Remnant pyrite-encrusted micrite nodules  
300 can be found in overlying marls for a few cm to a few dm above the oxidized crust (**Fig. 6D**).

## 301 7. Seeping sub-sites

302 The sub-site A is one of the more recognizable sub-seep sites of the Beauvoisin area (**See Fig. 2 for**  
303 **location**). It was first described as a columnar structure, different from other sub-sites B to T considered as  
304 lenticular (*Gaillard et al., 1985; Rolin, 1987*). However, the outcrop conditions with steep flanks of the  
305 Terres Noires did not give access to sub-site A until we used climbing ropes and harnesses. Sub-site A is

306 composed of lenticular carbonate lenses that are currently intensively eroded by present-day weathering  
307 (Fig. 7). All carbonate lenses that are in sharp contact with each others are considered as one unit. In  
308 consequence, 3 main units can be defined, forming sub-sites A1, A2 and A3, respectively from base to top.  
309 In terms of stratigraphic position, sub-sites A1 to A3 are Lower Oxfordian in age (See Fig. 3 for the log  
310 correlation), belonging to the Cordatum Zone at about 160.5 Ma (See Fig. 4 for age estimation). Three  
311 high-resolution stratigraphic sections were obtained in the area. Logs 1 and 2 are located respectively 60 m  
312 and 30 m south-east of the sub-sites A1 to A3, whereas log 3 follows the crest descending through the 3  
313 sub-sites A1 to A3 (Fig. 7). At this scale, relevant levels 1 to 5 can be visually correlated from log to log  
314 following the general stratigraphy ( $S_0$ ). Levels 1 to 4 correspond to aligned nodules following the general  
315 stratigraphy whereas level 5 is a pluri-cm thick layer of mudstone. All sections were flattened on level 5  
316 located on top of sub-site A2 in log 3 (Fig. 8). The first remarkable feature is that all levels are at the same  
317 topographic level in logs 1 and 2, but levels 1 to 3 are located 4 to 6 m beneath the general stratigraphy at  
318 the base of log 3. Despite our efforts, level 4 cannot be identified in log 3 directly beneath sub-site A1. All  
319 levels, including the flat level 5, are marked by a decrease in nodule concentrations from log 3 to log 1. The  
320 number of nodules per volume of marls is markedly reduced beyond log 1 (i.e., beyond 60 m) and the  
321 nodules almost fully disappear beyond 70-80 m, e.g. at Site G (see Fig. 6).

## 322 8. Comparison with an active seep site

323 Sub-seeps were previously identified at the bottom of the Regab giant pockmark in the Lower Congo Basin  
324 (LCB) where a 800 m wide and 15 m deep depression at about 3150 m water depth has been documented  
325 (Gay, 2002; Ondreas et al., 2005). The average slope is about 2 to 3° compared to the surrounding seafloor.  
326 The detailed microbathymetry was obtained onboard the ROV-*Victor 6000* during the ZAIROV 2 and  
327 BIOZAIRE cruises in 1999-2000. The first observations show that the giant pockmark is composed of several  
328 depressions, each about 100 m wide and 6 m deep (Ondreas et al., 2005). In detail the small-scale  
329 depressions are not homogeneously distributed at the bottom of the pockmarks (Fig. 9A). The deepest two  
330 depressions are located near the centre of the Regab pockmark field. They are marked by steep flanks with  
331 an angle of about 5° compared to the horizontal plane, locally reaching 7° if compared to the regional slope  
332 whereas other small-scale depressions have smoother slopes with an angle of about 1 to 2° (Fig. 9B). They  
333 all have a flat bottom, meaning that they look more like a plate than a bowl structure. This is consistent  
334 with the spider structures identified at shallower depths in the Lower Congo Basin (Casenave et al., 2017).  
335 However, only a few concretionary carbonates have been recovered in gravity cores from the Regab giant  
336 pockmark field (Gay et al., 2006). This apparent lack of nodules or concretions may be due to the inability  
337 of modern coring techniques to adequately sample such carbonates (Lloyd et al., 2015).

338 The studies conducted in the Regab pockmark field have shown that the highest concentrations of methane  
339 (up to 150  $\mu\text{l/l}$ ) in bottom waters was measured near the centre of the main depression, suggesting that  
340 this area is the actual and active methane seepage point (Gay et al., 2006). More precisely, a recent study

341 showed that only two zones are actually harboring tubeworms, mussels and clams at the same time, all  
342 living in close association with 4 m high carbonate build-ups (see Fig. 2B in [Marcon et al., 2014](#)), also a  
343 common feature of other giant structures such as in the Hydrate Ridge ([Teichert et al., 2005](#)) and in the  
344 Mediterranean Sea ([Ingrassia et al., 2015](#)). These two zones are annotated with green arrows on **Fig. 9B**  
345 and they correspond to the two deepest small-scale depressions identified on the dip map of the Regab  
346 pockmark field. In addition, sea-bottom pictures and videos taken within the other small-scale depressions  
347 displayed only relict features such as carbonate pavement oxidized on top and surrounding dead fauna or  
348 fields of clams on reduced black sediments (red arrows on **Fig 9B**). This may indicate a "fossil" site as  
349 remnant methane seepage (<20 µl/l) in bottom waters coupled to dead or dying fauna and oxidized crusts  
350 suggesting that fluid fluxes may have operated at higher rates in the past ([Gay et al. 2006](#)), as observed in  
351 other active and modern seep sites ([Greinert et al., 2001](#), [Han et al., 2004](#), [Haas et al., 2010](#)). This also  
352 suggests that only one or two points of focused fluid emission can act in a pockmark field, such as for the  
353 Hydrate Hole further North in the Lower Congo Basin ([Wenau et al., 2017](#)) and in other giant pockmark  
354 fields such as in the Zannone area ([Ingrassia et al., 2015](#)). This is probably due to the pressure gradient  
355 allowing only one path at a time to be open as in any plumbing system ([Gay et al., 2007](#)).

## 356 9. Clustering of fluid venting

357 Predicting the pattern of gas invasion and rise in liquid-saturated sediments is still a challenge. From a  
358 theoretical point of view, the relative importance of the dominant forces in the system, i.e., buoyancy and  
359 capillary forces, have to be taken into account. The Bond (or Eötvös) number, defined as the ratio between  
360 gravity over capillarity forces, is traditionally used to characterize multiphase flow in porous media. However, this  
361 number does not account for the different spatial scales which govern the dominant forces. Indeed, the  
362 largest capillary overpressure is controlled by the narrowest space between grains ("pore neck") while the largest  
363 buoyancy force is controlled by the wider pore space ("pore body"). In order to take into account  
364 this so-called porous media aspect ratio, a modified Bond number has been introduced by [Brooks et al.](#)  
365 [\(1999\)](#). This dimensionless parameter allows a physical classification of the flow patterns which strongly  
366 depend on the porous media aspect ratio ([Selker et al., 2007](#)). This latter, however, is difficult to estimate,  
367 in particular in geological settings where porosity, permeability and grain distribution may drastically  
368 change over short distances.

369 Laboratory experiments were therefore developed to investigate gas invasion into liquid-saturated grains in  
370 polydispersed systems. They bring the advantage of being closer to field situations, while still allowing a  
371 precise control of the imposed parameters (geometry, grain size distribution, injected gas flow-rate, and so  
372 on). These sandbox models have been used to illustrate different regimes of gas invasion in the liquid-filled  
373 granular medium (see for instance [Selker et al., 2007](#), [Varas et al., 2015](#) and references therein): formation  
374 of gas channels, dendritic invasion, fracturing, etc. Sandbox experiments were recently conducted to  
375 investigate the temporal variations of gas emission due to gas invasion into liquid-saturated grains ([Vidal et](#)

376 [al., 2010](#); [Varas et al., 2011](#); [Varas et al., 2015](#)). In a 3D experiment ([Fig. 10a](#)), a constant air flow-rate ( $Q$ ) is  
377 injected at the bottom of a layer of grains immersed in water. The gas is injected via a dry chamber (volume  
378  $V$ ), through a nozzle of inner diameter 1mm, localized at the bottom centre of the cell ([Fig. 10a](#)). The  
379 pressure variations in the chamber are monitored with a pressure sensor. Surface imaging makes it possible  
380 to quantify the gas emission location and dynamics. The analysis of the position of bubble emissions shows  
381 that they can migrate laterally through time ([Fig. 10b](#)). Indeed, the distance from the centre (i.e. vertical  
382 axis of the injection point, red cross in [Fig.10b](#)) varies from 0 to a few cm in any direction of the cell, for an  
383 initial grain height  $h_g=14$ cm. No specific pattern appeared at this stage, probably because grain supply  
384 simulating sedimentation has not been implemented yet in the experiment. However, two stable states of  
385 gas emission were reported during the experiment. For a significant period of time, bubbles were expelled  
386 at roughly the same position, with a minor lateral migration, defining the clusters C1 and C2 at a distance of  
387 about 2.5 and 4 cm from the centre, respectively ([Fig. 10b and 10c](#)). This study points out that, due to the  
388 strong vertical pressure gradient, gas channels cannot co-exist at the same time. In other words, only one  
389 gas channel can be active and expel gas at the surface at a given time.

## 390 10. Discussion

### 391 a) Seep site identification

392 The carbonate lenses of Beauvoisin were initially considered as pseudobioherms in the 80's due to their  
393 bioherm-like structure and composition ([Gaillard et al. 1985](#)). Although they contain reef-building  
394 organisms it is now considered that they did not form any significant relief on the seafloor ([Gaillard et al.](#)  
395 [1985](#); [Rolin, 1987](#); [Peckman et al., 1999](#); [Gay, 2002](#)) even if some examples on modern sites show that  
396 they can form a positive relief above the seabed ([Teichert et al., 2005](#); [Himmler, et al., 2015](#)). It has been  
397 shown that carbonate pavements formed in situ into the sediments ([Bayon et al., 2009](#)). The presence of  
398 autochthonous biological communities, together with a high density of one or very few taxa of  
399 chemotrophic megafauna, and lateral faunal gradients, are recognizable features of seep sites ([Kauffman](#)  
400 [et al., 1996](#); [Barbieri et al., 2005](#)). Moreover, the relation of fossil-rich carbonate lenses within the general  
401 stratigraphy and structural framework, their mineralogy and geochemical features, including the biomarker  
402 signatures, are indicative of chemosynthetic organism activity ([Campbell et al., 2002](#)). Lucinid bivalves are  
403 the most prominent biota associated with the Oxfordian carbonate lenses, with some giant specimens, up  
404 to 15-20 cm in diameter, suggesting that they fed on abundant food and nutrients ([Rolin, 1987](#); [Gay, 2002](#)).  
405 Detailed paleoecology studies indicate that the communities are similar to those of recent/present  
406 hydrothermal vents and cold seeps ([Rolin et al. 1990](#); [Gaillard et al. 1992](#)) and to other fossil seep sites  
407 ([Taviani, 1994](#); [Clari et al., 1994](#); [Taylor et al., 2009](#)). More recently, geochemical analyses conducted on  
408 several Beauvoisin sub-sites have led to the conclusion that they were dependent on chemosynthesis  
409 ([Peckman et al., 1999](#); [Tribovillard et al., 2013](#)). The microbial oxidation of methane under anoxic  
410 conditions, referred to as sulfate-dependent anaerobic oxidation of methane (AOM), is a major biochemical

411 process occurring at cold seeps. At sites where AOM is active, methane concentration varies strongly  
412 depending on spatial and temporal availability of methane supply and microbial turnover rates (*Knittel and*  
413 *Boetius, 2009; Deusner et al., 2014*). Sulfate depletion is driven by methane oxidation in sediment pore-  
414 waters through bacterially-mediated reactions, inducing a change in redox conditions (*Feng et al., 2011*)  
415 and favoring carbonate precipitation (*Bayon et al., 2007; Ge et al., 2010; Vanneste et al., 2012, and*  
416 *references therein*). At the Beauvoisin sub-sites the activity of sulfate-reducing bacteria is firstly indicated  
417 by the dispersively-distributed framboidal pyrite within the nodules and by an outer rim of pyrite on the  
418 nodules (*Gaillard et al. 1992; Gay, 2002*). Then, biomarkers of archaea were identified in Middle Oxfordian  
419 carbonate lenses, where methanogens produced heavy CO<sub>2</sub> leading to <sup>13</sup>C-enriched carbonate precipitation  
420 in the methanogenic zone with δ<sup>13</sup>C values as high as +13 ‰ and 12,7 ‰ PDB (*Peckmann et al., 1999;*  
421 *Tribovillard et al., 2013*). The stable isotopic composition of carbonate lenses and nodules coupled with  
422 taxonomic endofauna show that all sub-sites A to T located in the Beauvoisin anticline structure can be  
423 considered as cold seeps that developed within the sediment, near to the sediment water interface in a  
424 similar fashion to modern seep sites (*Sibuet and Olu 1998, Bayon et al., 2006; Campbell et al., 2006*). This  
425 conclusion is also supported by specific enrichments in Mo, As and Sb in carbonate lenses, echoing the  
426 bacterially-mediated formation of authigenic carbonate crusts through AOM (*Tribovillard et al., 2013*) as  
427 previously shown in the modern mud-volcano of the Malta Plateau (*Cangemi et al., 2010*).

#### 428 **b) Spatial organization**

429 To better understand the spatial organization of a sub-site, the centerpiece is the nodule distribution and  
430 concentration. The size, shape and spatial organization of nodules at sub-sites A1 to A3 is quite consistent  
431 with the observations made at sub-site F and G (**Fig. 6 and 8**), suggesting they have been formed through  
432 similar processes. These observations can be extended to almost all sub-sites, in spite of different outcrop  
433 conditions and forest cover. Only a few sub-sites seem not to be nodule-dependent, such as sub-sites L and  
434 M. These very close sub-sites (**See Fig. 2 for location**) mostly contain the irregular echinoid *Tithonia*  
435 *oxfordiana* (*Gaillard et al., 1985; Gaillard et al., 1992; Senowbari-Daryan et al., 2007; Kiel et al., 2010;*  
436 *Gaillard et al., 2011*) and may represent some kind of exception as they do not contain a large macrofauna  
437 either.

438 The general organization of a single sub-site can be summarized as a set of vertically stacked carbonate  
439 lenses located in the centre of a concentric area of nodules. The basal lens is composed of a breccia made  
440 of mineralized veins, clasts-rich conduits crosscutting limestones and embedded nodules. This facies is  
441 similar to the "brecciated limestone" from Italy (*Ladanza et al., 2013*), meaning that it may be due to an  
442 intense overpressured fluid flow impacting microbial carbonate and related macro-fauna in the sub-  
443 surface, such as what is observed in the injectite system of the Panoche Hills (*Blouet et al., 2016*). The  
444 carbonate lenses are interbedded with marls or claystones corresponding to varying fluxes of fluids such as  
445 in modern sites (*Feng et al., 2010*). The area affected by nodules is about 120 m wide forming a depression  
446 4 to 6 m deep at a maximum compared to the regional stratigraphy as shown for sub-site A1 and A2 (**Fig. 8**



447 **and 11**). The average calculated slope is about 6°, which is visually almost undetectable in the field. Only  
448 high resolution logs and the identification of marker horizons allow the identification of such low gradients.

#### 449 **c) Teachings from modern systems and sandbox modelling**

450 Based on previous studies conducted on modern seep sites, particularly on the Regab pockmark in the LCB,  
451 we have shown that the centre of an active pockmark is not the main area for focused fluid seepage and  
452 the main point of emission may migrate laterally through time forming new active sub-depressions at the  
453 seafloor. Consequently, a giant pockmark is the result of the coalescence of successive seepage sub-areas  
454 forming a wide depression through time. This behavior is confirmed in sandbox models in which the walls  
455 of a feeding channel must collapse prior to the development of a new point of emission (**Varas et al., 2015**  
456 and references therein). The main consequence is an intermittent fluid seep, although the fluid supply from  
457 deeper levels remains constant. This has also been documented in the Norway basin, through the Giant  
458 Gjallar Vent (**Gay et al., 2012; Dumke et al., 2012**). In the light of observations made on the modern giant  
459 Regab pockmark and sandbox models, an intermittent fluid seep can also be interpreted in the Beauvoisin  
460 area. With respect to the stratigraphy some sub-sites can be grouped together as clusters (**Fig. 12A**). Seven  
461 clusters C1 to C7 were defined in the Beauvoisin area. Except for clusters C3 and C5 containing 2 sub-sites  
462 occurring at different positions at the same time, all other clusters are grouped in 200 m wide areas. It  
463 means that the point of fluid emission on the seabed has remained at almost the same geographical  
464 position for a significant period of time. A 3D view of the 7 clusters shows that the system shifts in time  
465 from one cluster to the other (**Fig. 12B**). It also means that the occurrence of each cluster is separated from  
466 the next one by a period of relative quiescence during which no fluids (or only relict fluids) were expelled at  
467 the seabed. The durations of fluid seepage and fluid quiescence steps are variable. Given an estimated  
468 burial of 2000-2500 m for the Terres Noires Formation in the area (**Gaillard et al., 1985; Rolin, 1987**), the  
469 compaction rate at the base and bottom of the composite log can be considered equivalent. In a first  
470 approximation, 7 periods of active seafloor fluid seepage and 8 periods of fluid quiescence alternated for  
471 3.44 My, corresponding to periods of 200 ky each on average. However, at this stage the driving forces for  
472 focused fluid migration remain unclear. Major deep-rooted faults structuring the margin during the  
473 Oxfordian in the vicinity of the Beauvoisin area (in particular the Propiac fault related to halokinesis, **see**  
474 **Fig. 1 for location**) may have played an efficient driving role for a long period of time, as suggested for  
475 other fossil seep sites (**Aiello, 2005**). Finally, the Beauvoisin seep area can be considered as the result of  
476 clustered fluid seep sub-sites. This has led to the formation of a 800 m-wide area of fluid seepage active for  
477 over 3.44 My. In the light of this work, the Beauvoisin area can now be considered as a giant pockmark  
478 field.

#### 479 **d) Evolution model**

480 Based on the comparison of fossil outcrop, modern and active pockmark, and sandbox models data, we  
481 propose here a new model for the internal evolution of a fluid seep area. Previous work conducted on  
482 sandbox models and numerical simulations have pointed out that, independently of the initial air invasion



483 regime (percolation or fracture), similar systems develop a fluidized zone of parabolic shape characterized  
484 by a central air-channel (*Varas et al., 2011; Ramos et al., 2015; Varas et al., 2015, Poryles et al., 2016*).  
485 These authors show that particles are intensively reworked within the parabolic area whereas particles do  
486 not have any movement outside of the parabolic area (**Fig. 13A&B**). This behavior has been previously  
487 studied for hydraulic fractures showing that they can occur without tectonic stresses (*Mourgues et al.,*  
488 *2011*), which is compatible with shallow unconsolidated and fluidized sediments. During the vertical growth  
489 of the hydraulic fracture, representing the fluid-channelling conduit, a cone of vertical upward  
490 displacements can be observed just above the tip of the fracture and the cone is limited by two large shear  
491 bands, indicating that the fracture opened exclusively in pure tension mode (*Mourgues et al., 2012*). This  
492 kind of cone deformation was observed in sedimentary basins, such as in the Norway basin (*Gay et al.,*  
493 *2012*) or in the Northern Mediterranean Sea (*Gay et al., 2017*) where the cone structure is accompanied by  
494 an uplift of the seabed suggesting that this doming is the initial phase of fluid emission preceding a general  
495 collapse of the structure after which a pockmark forms. However, in the light of recent studies conducted  
496 on fluid flow through a water-saturated matrix of grains, such cone structures are most probably of  
497 parabolic shape (*Poryles et al., 2016*). A 2D seismic profile across the studied Regab giant pockmark field  
498 shows that reflections are very chaotic below the seabed (**Fig. 13C**). The disturbed zone is in fact parabolic  
499 in shape and is characterized by high amplitude reflections that were interpreted as carbonate structures  
500 and/or gas-charged sediments (*Gay, 2002; Ondreas et al., 2005; Gay et al., 2006, Marcon et al., 2014*). The  
501 base of the parabolic area is clearly located on top of a vertically fractured zone that is usually interpreted  
502 as a seismic pipe (*Gay et al., 2006; Ho et al., 2012; Løseth et al., 2011*), very similar to the structure found  
503 further North in the Lower Congo Basin at the Hydrate Hole where the deformed zone in the subsurface is  
504 located right above an underlying fault clearly visible on seismic sections (*Wenau et al., 2017*). Previous  
505 studies conducted in the Norway Basin have also documented V-shaped anomalies beneath pockmarks  
506 corresponding to the transition from focused to distributed fluid flow (*Betzler et al., 2011; Gay et al.,*  
507 *2012*). This seismic facies appears as a "flower" structure with a vertical narrow zone considered as the  
508 stem (i.e. the pipe) feeding a wide area considered as the corolla (i.e. the disturbed sediments) (*Gay et al.,*  
509 *2012*). It can be interpreted as deformation of cohesionless, unlithified sediments caused by fluid injection  
510 (*Gay et al., 2017*).

511 We propose here a simple model of a giant fossil pockmark considered as an outcrop analogue to modern  
512 giant pockmarks (**Fig. 13D**). Based on previous studies the V-shaped anomaly (*Gay et al., 2012*), or the  
513 cone-in-cone structure (*Mourgues et al., 2012*), is more probably parabolic in shape. However, to date in  
514 the fossil pockmark of Beauvoisin, neither the point of injection (i.e., the base of the parabolic area) nor the  
515 fluid pipes were clearly identified (**Fig. 13D**). This is probably due to outcrop conditions, the Beauvoisin area  
516 being located in the heart of an eroded E-W anticline structure with intense forest cover. The measured  
517 values of  $\delta^{13}\text{C}$  and the presence of n-alkanes in the hydrocarbon fraction of veins and nodules thus  
518 correspond to a contribution of different sources of carbon including a methane thermogenic origin coming

519 from thermal maturation of organic matter at the base of the Terres Noires Formation or below it  
520 (*Peckmann and Thiel, 2004*). This observation coupled with fluid migration being active for at least 3.4 My  
521 suggests that a significant amount of fluids migrated during this period. However, it does not mean that the  
522 fluid flow in the sub-surface is constant over time. On modern sites varying fluid flows were documented in  
523 the sub-surface, producing authigenic high Mg-Calcite or aragonite depending on the CH<sub>4</sub> flux rate and  
524 forming pavements or concretions (*Nöthen et al., 2011*). If the flux rate is not high enough, only  
525 disseminated concretions precipitated. At a specific point of emission, the fluid flux in the shallow sub-  
526 surface is not high enough to sustain massive carbonate precipitation, but it does not mean that the fluid  
527 flux is also reduced at depth. In the disturbed zone, carbonate lenses are organized in clusters that laterally  
528 migrate through time. At time  $t_0$  the first cluster is fed by an irregular conduit (green line on **Fig. 13D**). Time  
529  $t_1$  corresponds to the time necessary for the conduit to collapse, then to laterally shift and then to open at  
530 a different location. The time  $t_2$  corresponds to the birth of a new cluster of carbonate lenses. Based on the  
531 detailed stratigraphy we showed that the times  $t_0$ ,  $t_1$  and  $t_2$  are relatively equivalent in terms of duration,  
532 about 200 ky each. It means that even if a pockmark seems dead and inactive, it could be at stage  $t_1$   
533 corresponding to a shift of the feeding conduit in the disturbed zone. As observed in both modern and  
534 fossil fluid seep areas, a long lasting fluid migration is channelized through focused structures (*Ho et al.,*  
535 *2012; Løseth et al., 2011*) that could constantly feed the disturbed zone. The next challenge in the area will  
536 be to identify such underlying pipe structures, as the bitumen infilling of veins and nodules may correspond  
537 to a later stage of heavier fluid migration as shown for the "Brecciated Limestone" in Italy (i.e., secondary  
538 migration in the petroleum system; *Ladanza et al., 2015*). In the case of the Beauvoisin seep site, it would  
539 mean that the feeding pipes remained open for a long period of time.

## 540 **11. Conclusion**

541 For over 30 years, a fluid seep system has been known in the Beauvoisin area, located on the Jurassic  
542 passive margin of the SE Basin of France (*Gaillard et al., 1985*). Modern marine investigations coupled with  
543 sandbox models permitted comparison of the architecture of the Beauvoisin seep site with an active fluid  
544 seep area on the modern passive margin of the Lower Congo Basin. The main conclusion is that they have  
545 very similar features even if they do not occur at the same geological time. In the same way, the sandbox  
546 models of fluid injection structures have shown that the involved processes have led to a similar  
547 organization in 3D, in particular:

548 > **Stratigraphy:** based on 23 sedimentological sections in the Beauvoisin area, 19 sub-sites A to T have been  
549 reported in a composite log, representing a total thickness of about 370 m. Due to weathering and erosion  
550 since the site discovery, two new sub-sites S and T were identified and integrated in the general log, based  
551 on the initial nomenclature established by *Rolin (1987)*.

552 > **Sub-Site shape and organization:** Each identified sub-site is composed of sub-vertically stacked carbonate  
553 lenses. Detailed logs at various distances from the vertical axis of carbonate lenses have shown that the

554 sub-sites developed in smooth depressions, 4 to 6 m deeper than the surrounding seabed. Nodules are  
555 present within the depression only. They are encased in the carbonate lenses in the centre whereas their  
556 size varies depending on the distance from the centre. Coupled with the varying  $\delta^{13}\text{C}$  values, this indicates  
557 that nodules can be considered as markers of a sub-site, even if outcrop conditions do not allow  
558 identification of carbonate lenses buried in marls.

559 > **General organization:** The carbonate lenses are organized in clusters up to 7 sub-sites grouped together  
560 in the same stratigraphic interval and the same geographic zone within the seeping area. Only a few sub-  
561 sites seem active at the same time in a given cluster. The coalescence of all clusters leads to the formation  
562 of a wide depression. This general organization is very similar to that of the modern Regab giant pockmark  
563 field in the Lower Congo Basin. The seep site of Beauvoisin can be clearly considered as a fossil giant  
564 pockmark field, analogue to pockmarks in modern basins.

565 > **Timing:** A detailed stratigraphic correlation coupled with extensive mapping of fluid seep sub-sites in the  
566 Beauvoisin area has led to a spatio-temporal 3D reconstruction of the position of these sub-sites. It shows  
567 that periods of fluid seeping alternated with periods of apparent quiescence, about 200 ky each.

568 A pockmark is a seafloor depression hosting seep sites, both active and inactive. The study of the modern  
569 Regab pockmark shows that it is formed by the coalescence of smaller depressions, each about 100 m  
570 wide. However, submersible dives and gravity cores provide only the partial vision of the seafloor without  
571 integrating the sub-seafloor and the connection with fluid pipes at depth. A way to access the pockmark  
572 underground is studying outcrop analogues, such as the Beauvoisin giant pockmark. The study of this site  
573 provided a real breakthrough in the understanding of fluid activity in the disturbed zone below the  
574 seafloor. The main result is that a pockmark could seem inactive for a long period of time due to the lateral  
575 shift of the feeder conduit, meaning that the sub-seafloor remains charged in gas. This observation will be  
576 of great importance in the geobiology of fluid seeps ([Peckmann, 2005](#); [Judd et al., 2007](#)), in risk assessment  
577 for anthropic activities at the sea-bottom ([Gay et al., 2017](#)) and for hydrocarbon exploration ([Capozzi et al.,](#)  
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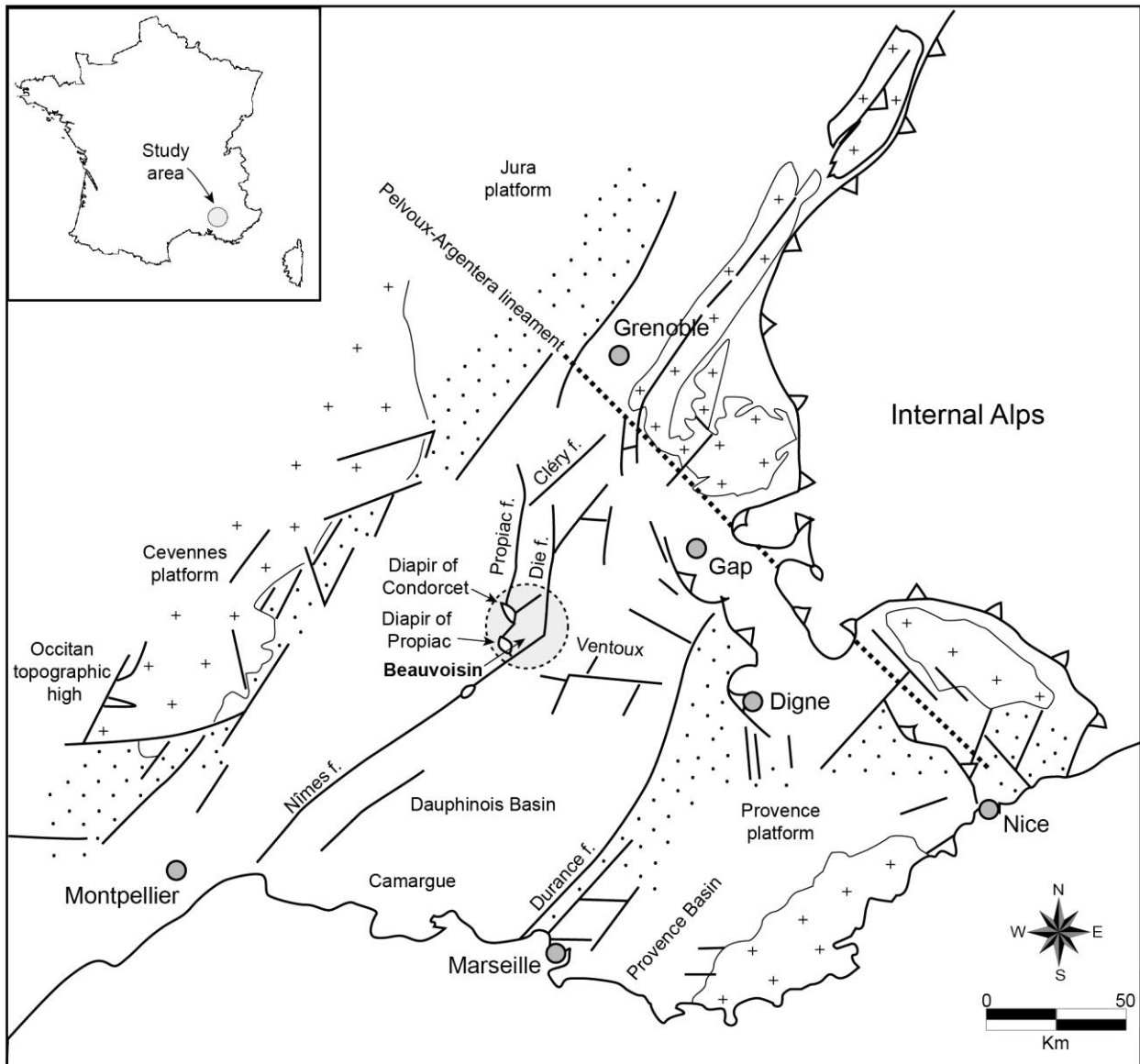
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832 methane. - *Chemical Geology*, **161**, 1, 291-314.

833 **List of figures**

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836

**Figure 1: Simplified structural map of the South East Basin of France. The study area is located between the Propiac**

837

**Fault and the Die Fault north of the Ventoux in a relay zone corresponding to the heritage of syn-sedimentary faults**

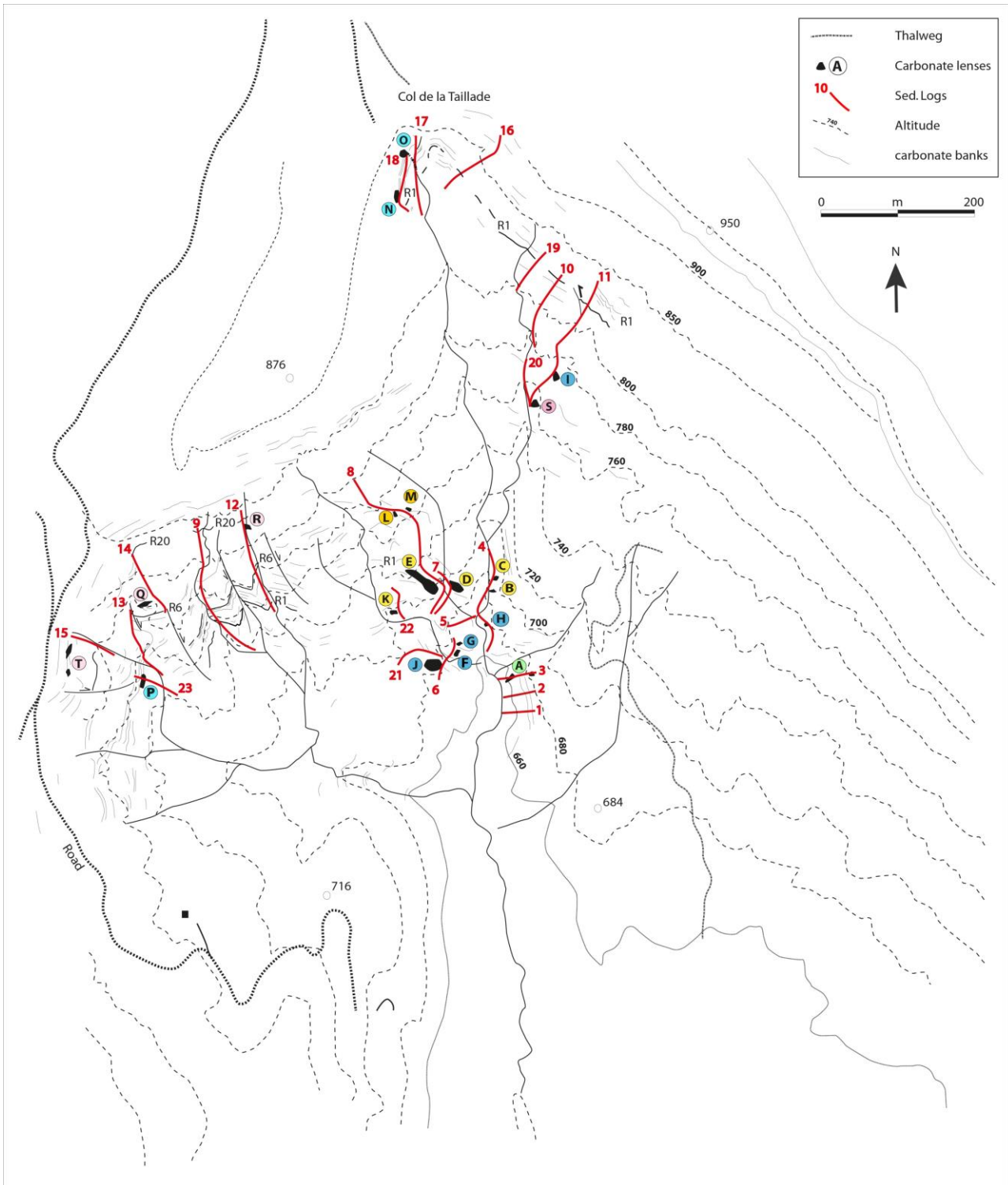
838

**that structured the Jurassic Tethyan margin. In the Beauvoisin area (grey circle on the map) the identified seep sites**

839

**developed in dark marls called the "Terres Noires" Formation (Oxfordian).**

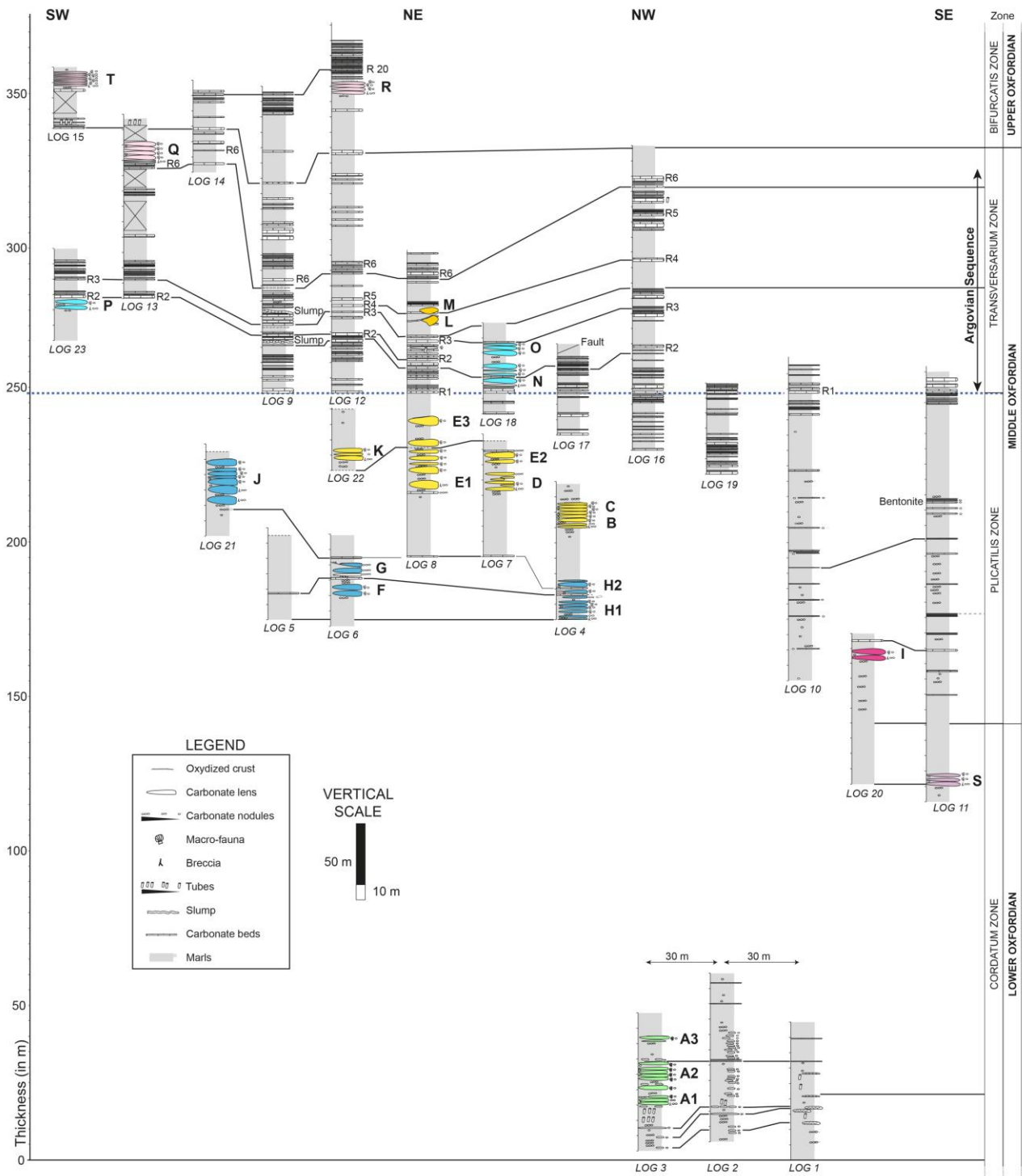
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842 **Figure 2: Detailed topographic map of the Beauvoisin area (modified after Gaillard et al., 1985). 23 sedimentological**  
 843 **logs have been recovered (red lines) between 2002 and 2016 in steep valleys. Almost all valleys have been explored,**  
 844 **depending on safety conditions. 19 sub-sites have been reported, including two new sites, S and T, that were**  
 845 **exposed due to weathering since the initial mapping.**

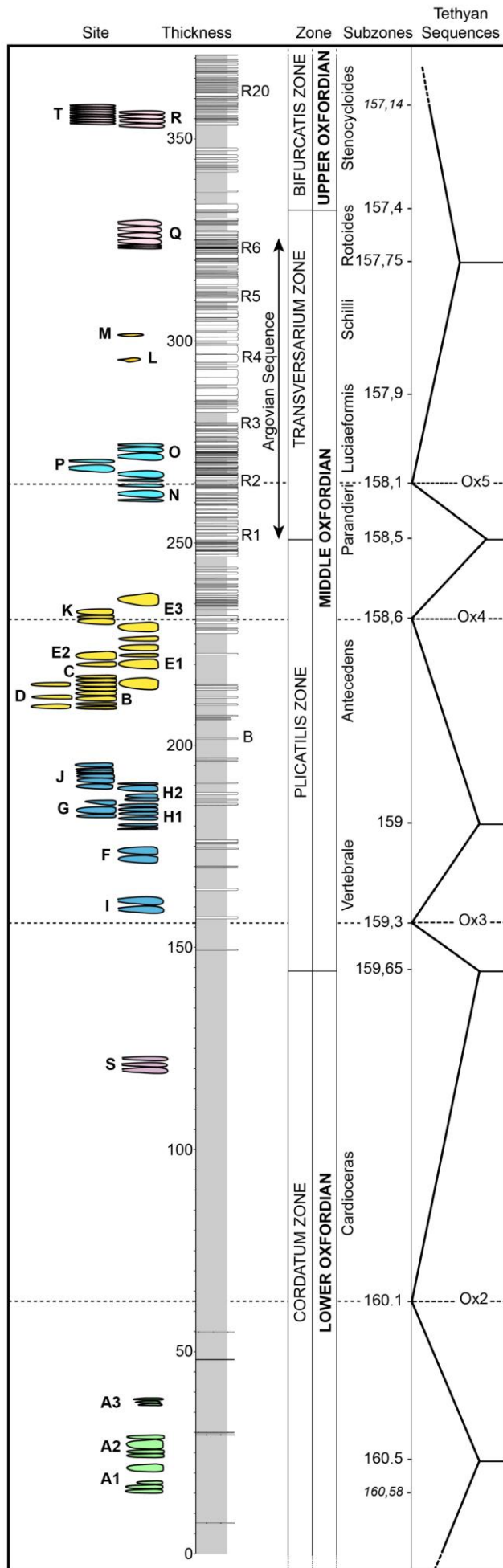




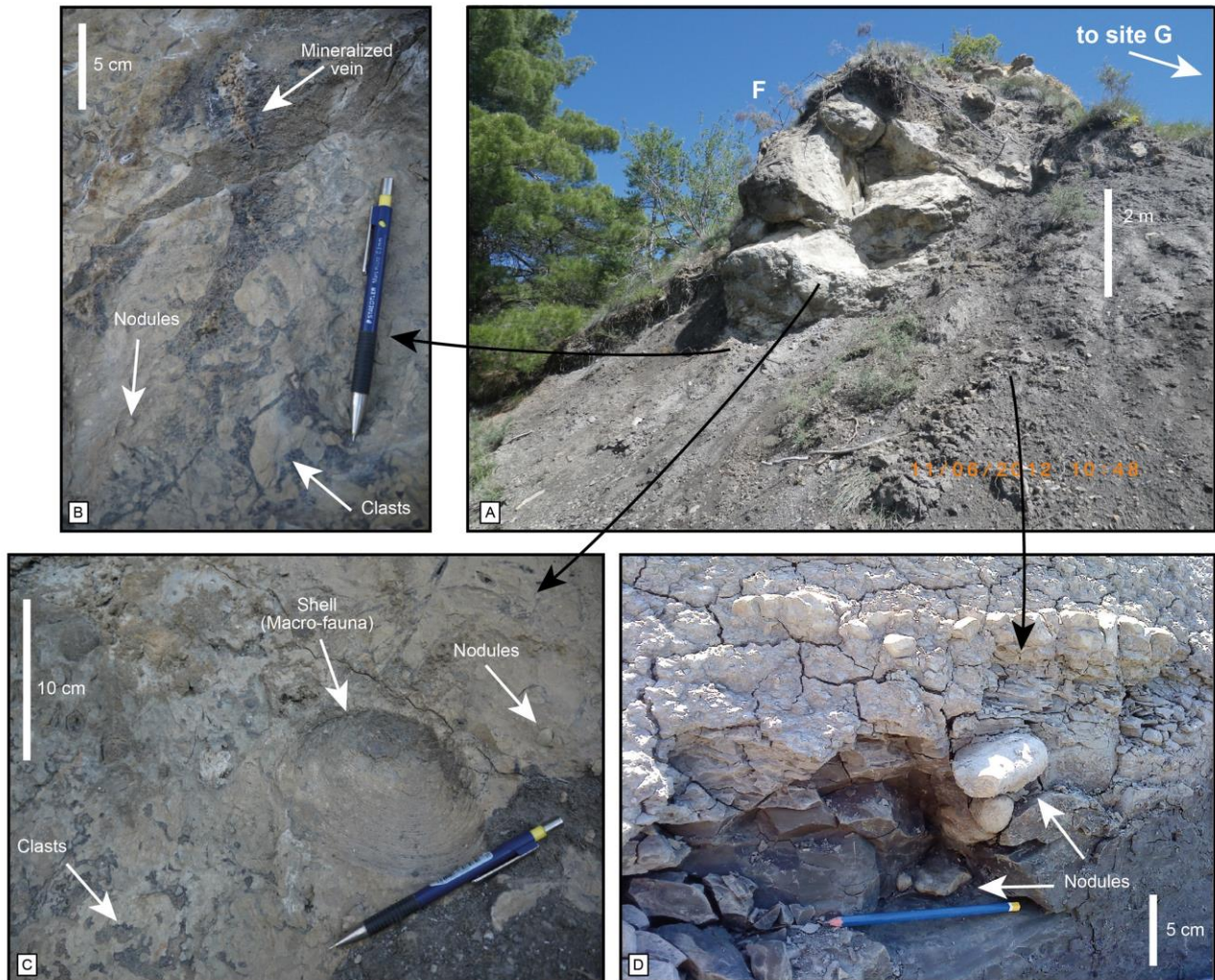
846

847 **Figure 3: Correlation of the 23 sedimentological logs in the Beauvoisin area regarding the general stratigraphy of**  
 848 **the Oxfordian. The 23 stratigraphical logs were obtained utilising key marker layers, R1 to R20 following the**  
 849 **nomenclature of Gaillard et al. (1985) and Rolin (1987). The correlation is centered and flattened on the R1 level**  
 850 **corresponding to the base of the Argovian sequence ending on top with the R6 level.**

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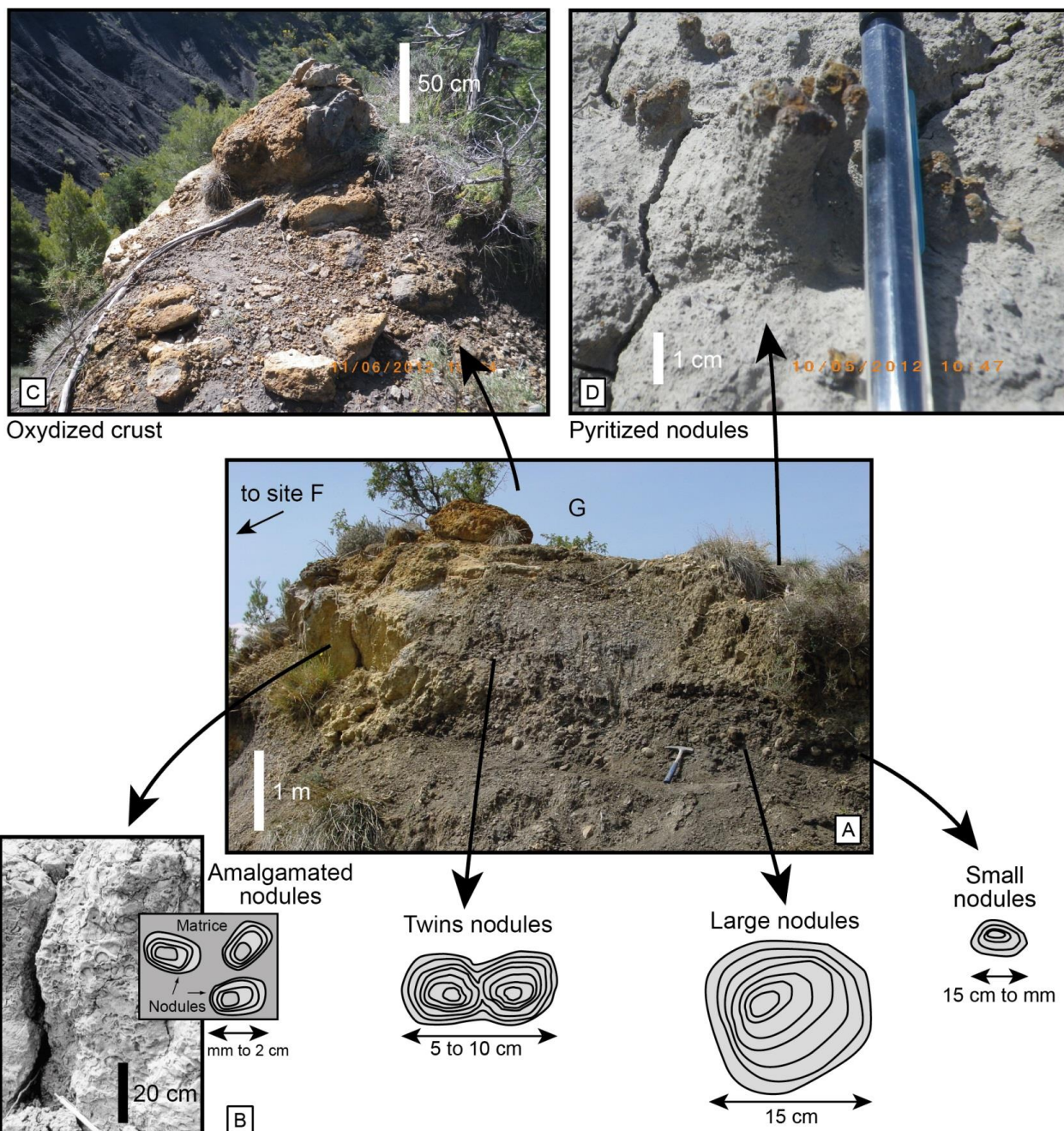
853 **Figure 4: Composite log of the Oxfordian in the Beauvoisin area taking into account eroded missing sequences,**  
 854 **slump deposits and faults (360 m in thickness). The oldest site A is located within the Cordatum Zone (Lower**  
 855 **Oxfordian) whereas the youngest site T is in the Bifurcatis Zone (Upper Oxfordian) according to the Tethysian**  
 856 **sequences and age model from [Gradstein et al. \(2012\)](#) .**



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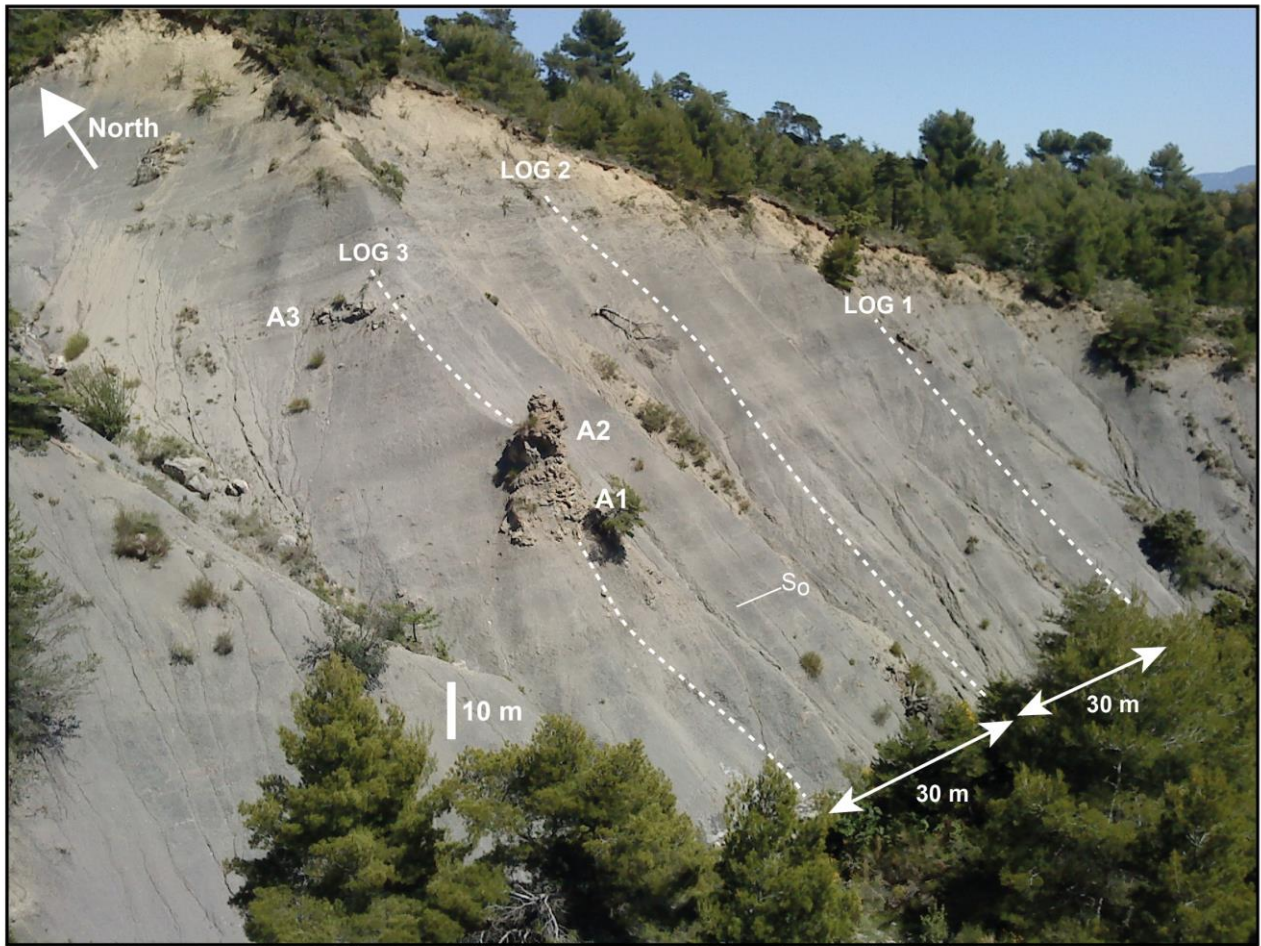
858 **Figure 5: Sub-site F located in the centre of the Beauvoisin area (See Fig. 2 for location). A) View to the West of site F**  
 859 **forming a 7 m high edifice composed of sub-vertically stacked carbonate lenses. B) Detailed photograph of the basal**  
 860 **carbonate lens showing the brecciated facies dominated by sub-vertical veins yielding mineralizations and**  
 861 **displaying a puzzle-like structure. C) Detailed photograph showing lucinid specimen, 12 to 15 cm wide, in close**  
 862 **association with the brecciated facies and mineralized veins. D) Detailed photograph showing micritic nodules, mm**  
 863 **to 15 cm in diameter.**





864

865 **Figure 6: Sub-site G located in the centre of the Beauvoisin area and just above the sub-site F (See Fig. 2 for**  
 866 **location). A) Outcrop photograph showing densely packed mm to 2 cm aggregates nodules encased in a dark**  
 867 **micrite within the basal carbonate lenses. B) Organization and shape of nodules depending on their distance to the**  
 868 **carbonate lens. C) Photograph of the top of sub-site G showing corrosion forming a pluri-cm thick Mn and Fe rich**  
 869 **crust. D) Detailed photograph showing pyrite-encrusted micrite nodules overlying marls for a few cm to a few dm**  
 870 **above the oxidized crust.**



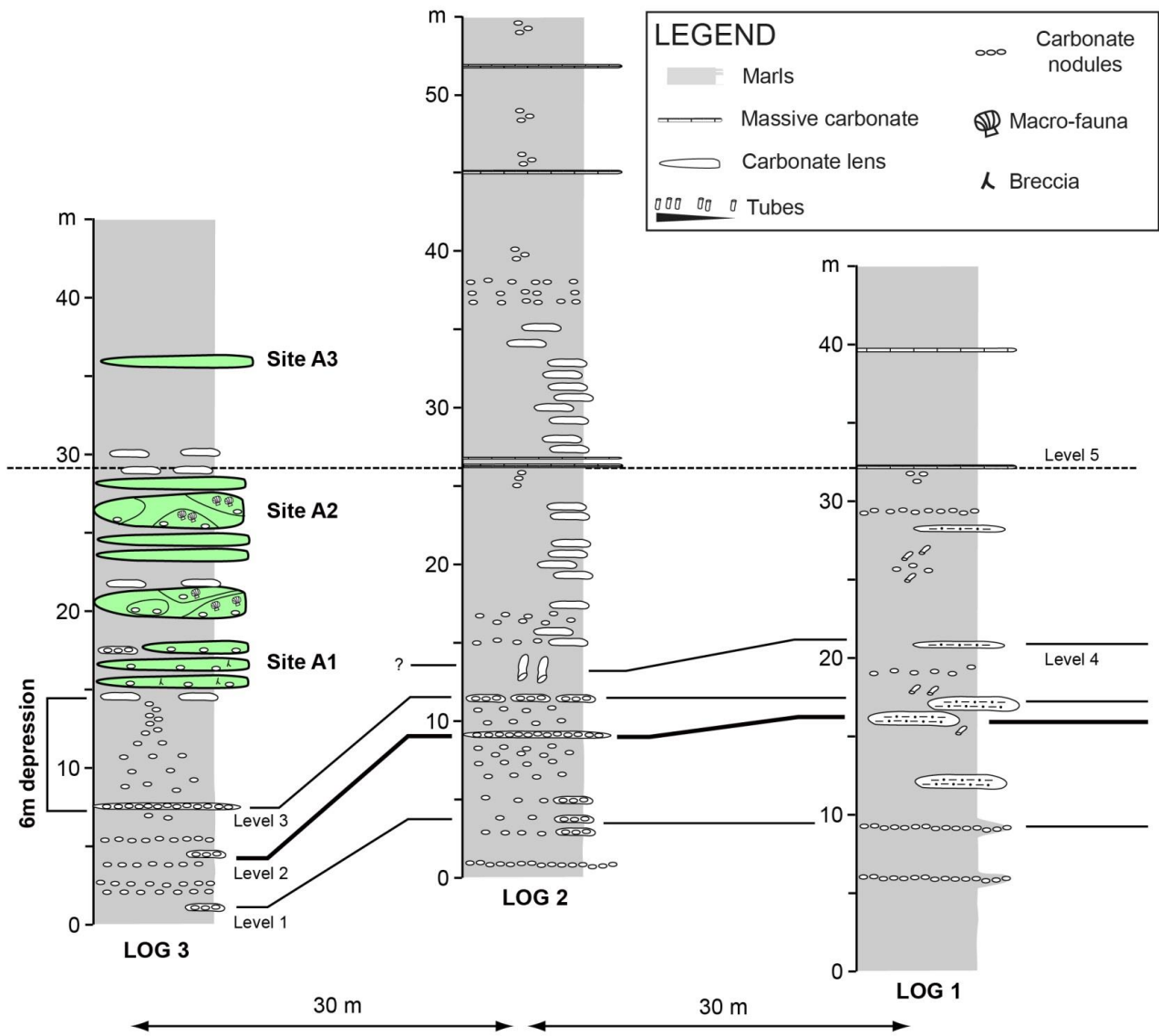
871

872 *Figure 7: Photograph to the NE taken from site G and showing site A in the Lower Oxfordian (See Fig. 2 for location).*

873 *Three sub-sites, A1 to A3, can be identified from to base to the top. Three high-resolution stratigraphic sections, logs*

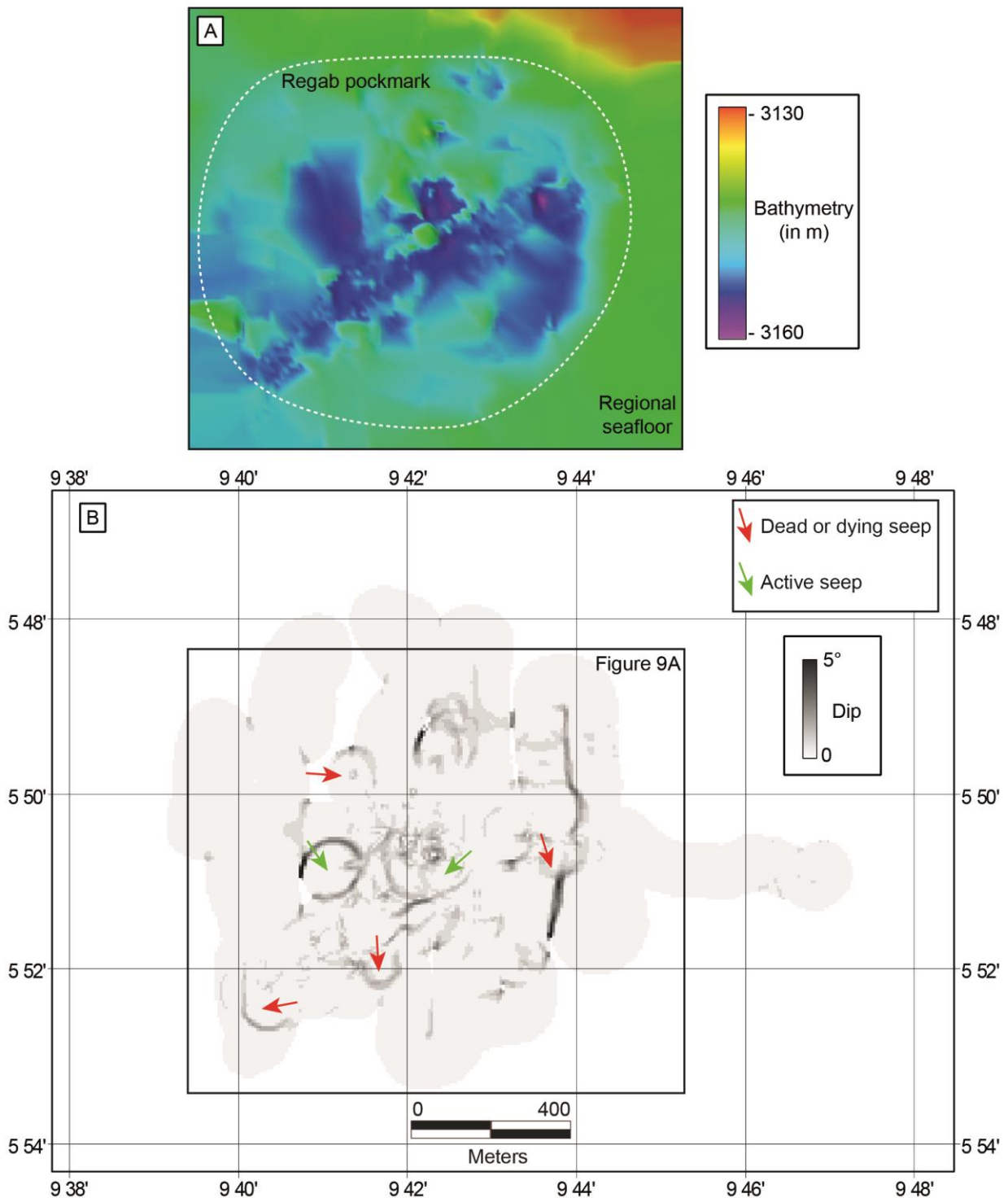
874 *1 to 3, were obtained at 60 m, 30 m and 0 m (vertical axis of the 3 sub-sites A1 to A3).*





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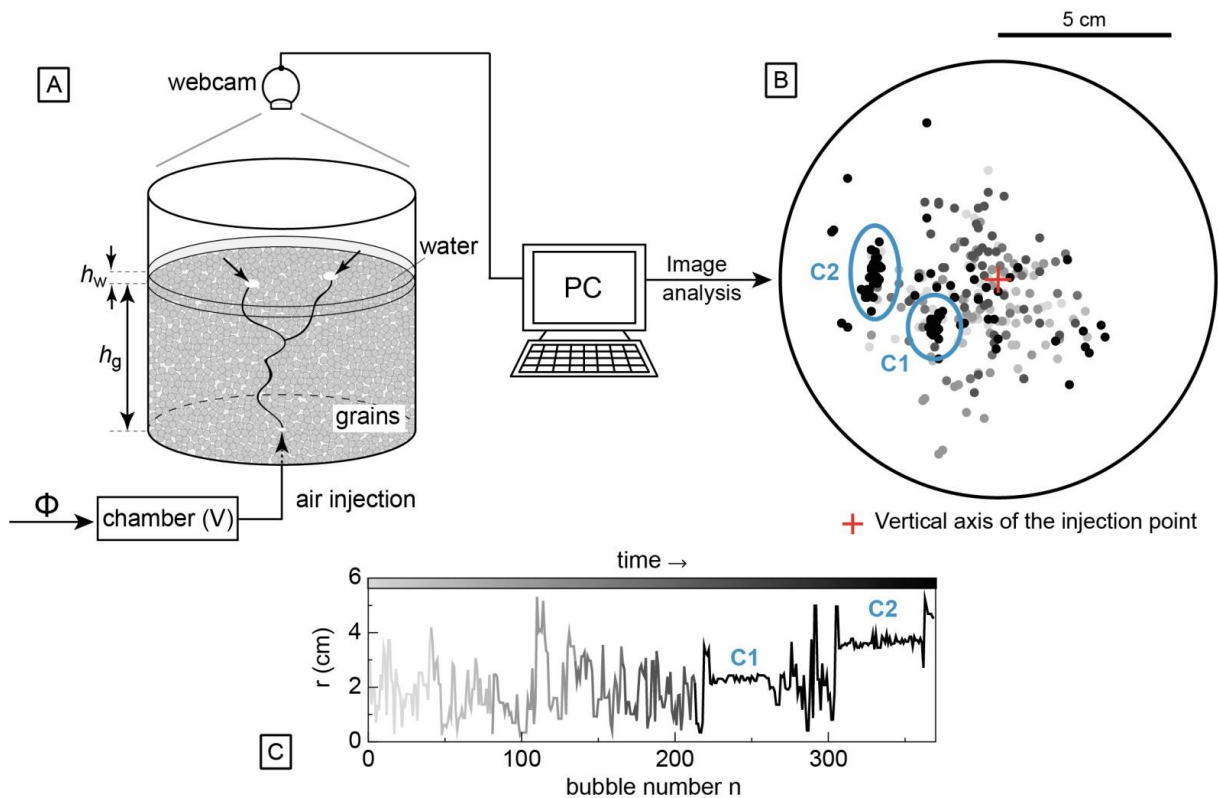
876 **Figure 8: Sedimentological logs 1 to 3 obtained at site A (see Fig. 2 for location, see Fig. 7 the detailed location). All**  
 877 **sections were flattened on level 5 located on top of sub-site A2 in log 3. Levels 1 to 3 are located 4 to 6 m beneath**  
 878 **the general stratigraphy at the base of log 3 defining a smooth depression. All levels are marked by a decrease in**  
 879 **nodule concentrations from log 3 to log 1.**



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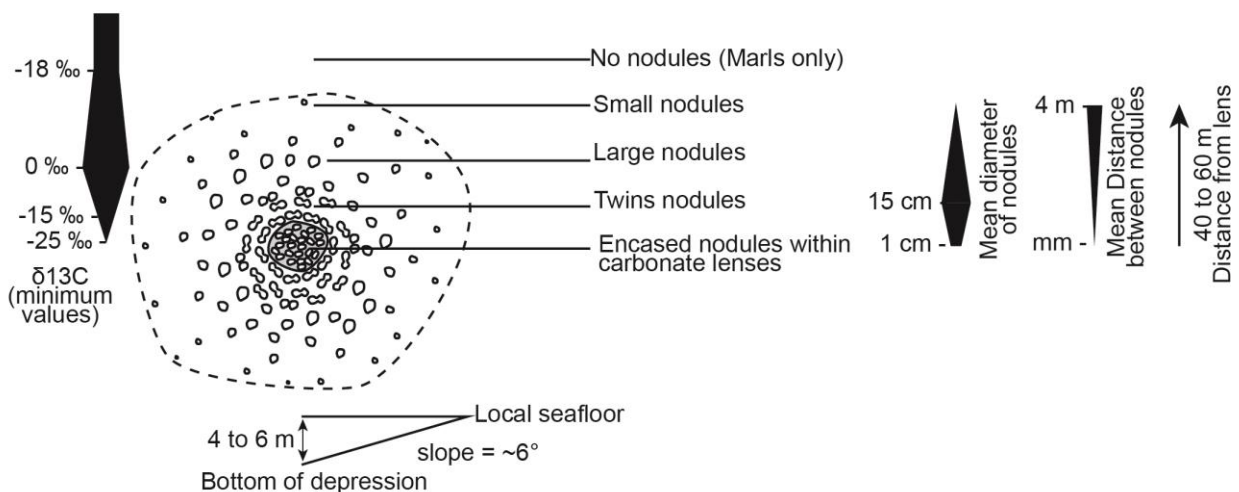
881 *Figure 9: A) High resolution bathymetric map of the Regab pockmark, acquired during the Zairov and Biozaire*  
 882 *scientific surveys in 2000-2002. B) Dip map of the Regab pockmark obtained by deriving the bathymetric map. It*  
 883 *shows well expressed 100 m wide sub-circular depressions. Green arrows represent the depressions with active and*  
 884 *high concentration methane seepage corresponding to the most active area (Gay et al. 2006). Red arrows represent*  
 885 *intermediate or low methane flow rates interpreted as dormant or dying sub-sites.*





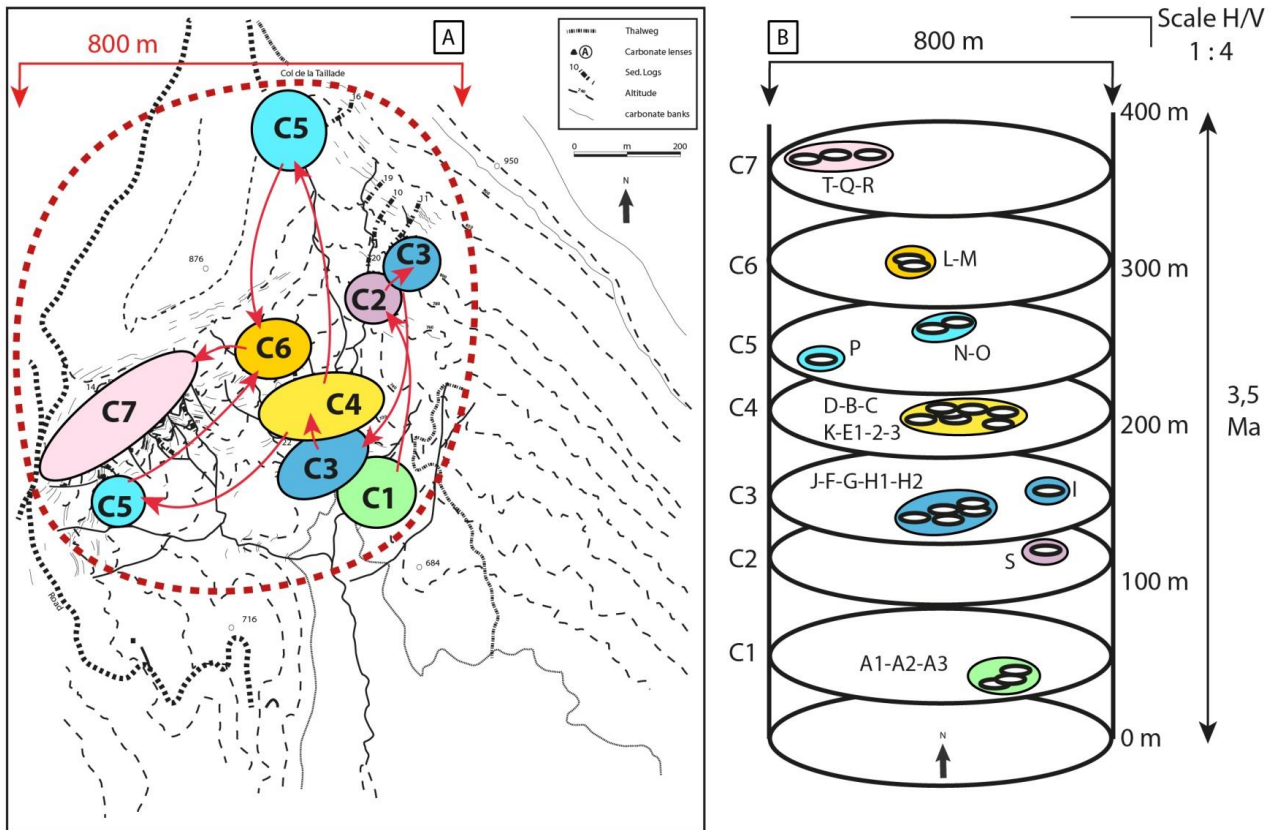
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887 **Figure 10: 3D sandbox experiment (Varas et al. 2015).** A) **Experimental setup.** Air is injected at a constant flow rate  
 888  $\Phi$  at the bottom of an immersed granular column ( $h_g$  is the height of grains and  $h_w$  the height of water above the  
 889 granular bed). Bubble emission at the free surface is recorded by a webcam positioned over the experimental cell.  
 890 B) **Detection of bubble emission at the free surface** [ $d = 318 \pm 44 \mu\text{m}$ ,  $h_g = 14 \text{ cm}$ ,  $h_w = 2 \text{ cm}$ ,  $\Phi = 4.3 \text{ mL/s}$ ; the dots  
 891 color from light gray to black is proportional to time]. The system releases bubbles in different (uncorrelated)  
 892 positions. For two significant periods of time, however, several successive bubbles are emitted in a narrow region  
 893 (clusters C1 and C2). It indicates the formation of a gas channel fixed at depth (region C1), which then collapses  
 894 before it can reopen in another region (C2). C) **Distance  $r$  between the bubble emission location and the centre of**  
 895 **the cell (red cross in B) as a function of the bubble number  $n$  for the sequence presented in B.**



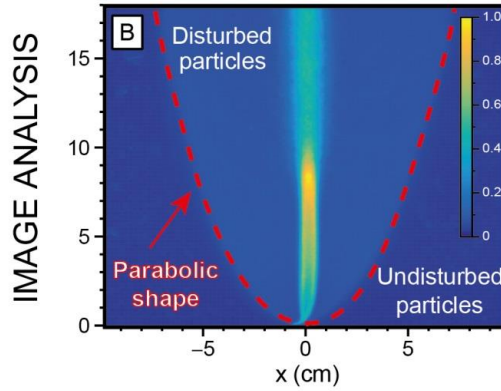
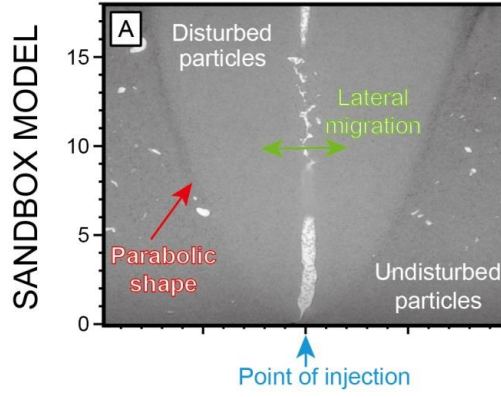
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897 **Figure 11: Schematic diagram of the organization of a single carbonate lens located in the centre of an 80 to 120 m**  
 898 **wide concentric area of nodules. Marker horizons on high resolution logs allowed the identification of a 4 to 6 m**  
 899 **deep depression with an average slope of about 6°, which is visually almost undetectable on field.**

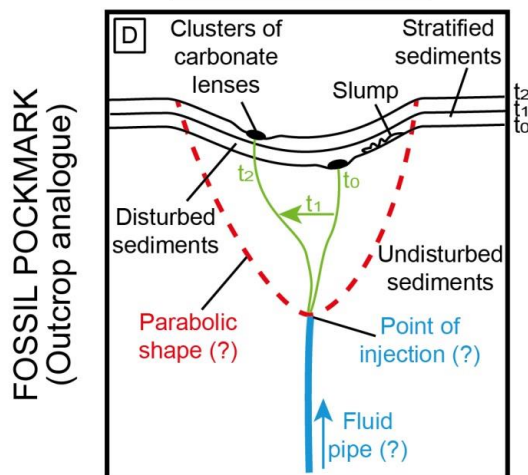
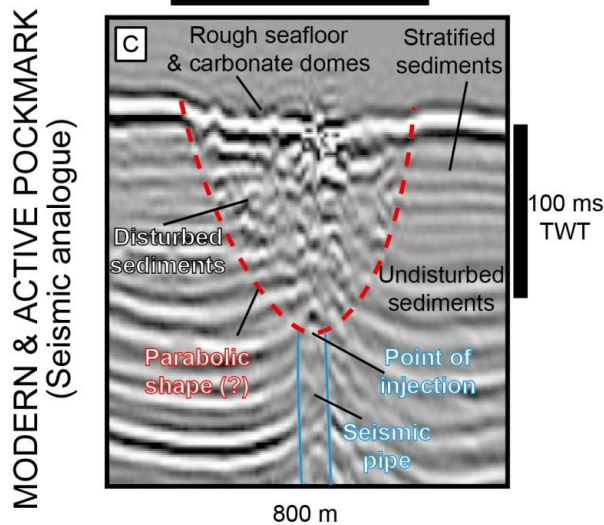


900  
 901 **Figure 12: A) Map of the sub-sites A to T grouped by clusters regarding their stratigraphical and geographical**  
 902 **positions in the Beauvoisin area. B) 3D schematic view of the 7 clusters showing that the point of emission at the**  
 903 **seafloor has laterally migrated through time. The 7 periods of active seafloor fluid seepage alternated with periods**  
 904 **of fluid quiescence for about 3.4 My, corresponding to periods of 200 ky each on average. It means that all sub-sites**  
 905 **A to T are genetically linked and they can be grouped in an 800 m wide giant pockmark.**

Clustering of the point of emission



800 m



907 *Figure 13: A) 2D sandbox experiment of air injection in a Hele-Shaw cell filled with water-saturated grains [ $d = 318 \pm$*   
908 *44  $\mu\text{m}$ ,  $h_g = 20 \text{ cm}$ ,  $h_w = 2 \text{ cm}$ ,  $\Phi = 0.66 \text{ mL/s}$ ]. The zone where the sediments are disturbed has a roughly parabolic*  
909 *shape (in red), and the central gas channel displays erratic lateral motion in time (in green), leading to lateral*  
910 *variations of bubbles emission at the surface. B) Normalized flow density (indicated by the colorbar) computed for*  
911 *the series of images from which A) is extracted. High values of the flow density indicate regions where motion*  
912 *occurred in the system. This analysis reveals the parabolic shape of the disturbed zone (in red). C) 2D seismic profile*  
913 *of the Regab giant pockmark showing a disturbed zone right beneath the main depression (modified after Gay et*  
914 *al., 2006). The limit between undisturbed reflections and the chaotic reflections is parabolic in shape (in red). The*  
915 *base of the parabolic area is located on top of a vertically fractured zone (in blue) that is usually interpreted as a*  
916 *seismic pipe, playing the role of a feeder conduit for fluids. This facies is called the "flower" structure with a vertical*  
917 *narrow zone considered as the stem (i.e. the seismic pipe) feeding a wide area considered as the corolla (i.e. the*  
918 *disturbed sediments). D) Model of the Beauvoisin seep site considered as an outcrop analogue for a giant*  
919 *pockmark. In the main depression, carbonate lenses are organized in clusters that have laterally migrated through*  
920 *time. At time  $t_0$  the first cluster is fed by an irregular conduit (green line). During time  $t_1$  the conduit has collapsed,*  
921 *laterally shifted and reopened at a different location feeding a new cluster of carbonate lenses during time  $t_2$ . It*  
922 *means that even if a pockmark seems dead or inactive, it could be at stage  $t_1$  corresponding to the shifting period of*  
923 *the feeding conduit in the disturbed zone. However, at present neither the parabolic shape of the disturbed zone (in*  
924 *red), nor the feeding conduits at depth (in blue) have been identified yet.*  
925