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Climate sensitivity and geomorphological response of cirque

glaciers from the Late Glacial to the Holocene

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Abstract

Through a detailed geomorphological study, including thourough mapping of the 15 geomorphic features as well as ¹⁰Be Cosmic-Ray Exposure (CRE) dating, the 16 17 geomorphological evolution of the Mulhacén cirque since the maximum ice extent of the 18 last glacial cycle until nowadays was determined. This glacial cirque is shaped on the northern face of the Mulhacén peak (3479 m a.s.l., 37°03′12″N / 3°18′41″W), Sierra 19 20 Nevada, southern Spain. It includes several depositional and erosional glacial landforms 21 that allowed reconstructing its environmental evolution since the last glacial cycle. 22 Furthermore, the sequence of glacial oscillations from this site was compared to that of 23 other cirques of the massif, evidencing that: (i) new glaciers formed in these cirques 24 during the Younger Dryas (YD), and (ii) disappeared at 11.7 ± 1.0 ka. Depending on the 25 altitude, orientation and height of the cirque walls, the final deglaciation of the cirques 26 generated a diversity of landscapes, including a wide range of glacial and periglacial 27 landforms, such as polished surfaces, sequences of moraines, proto-rock glaciers or large 28 rock glacier systems. No glaciers existed in the Sierra Nevada during the Middle 29 Holocene. Only the circues whose summits exceed 3300 m, are north-exposed and whose 30 walls exceed 300 m high (i.e. Mulhacén and Veleta) hosted glaciers during Neoglacial 31 phases, including the Little Ice Age (LIA) (approx. 1300-1850 CE). During these periods, 32 climate oscillations favoured the formation of small glaciers in these cirques, which 33 generated large moraine systems with either one polygenic ridge or a sequence of spaced 34 frontal arcs. The existence of glaciers impeded the formation of permafrost-related

started to form. These results are compared with the deglacial evolution in 55 cirques from Iberian mountains as well as from glacial cirques from other mid-latitude mountains and subpolar regions. The chronology of their deglaciation as well as the landforms generated during glacial retreat followed similar patterns, with no significant differences

landforms, such as rock glaciers and protalus lobes until the end of the LIA, when they

- 40 at regional scale. For each mountain range, the geomorphological diversity existing in
- 41 each cirque depends on the local topographic characteristics although they formed during
- 42 the same climatic phases.

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- 43 Key words: Glacial Cirque, Rock Glacier, Sierra Nevada, Cosmic-Ray Exposure Dating,
- 44 Late Glacial, Holocene.

1. Introduction

- 46 Glacial cirques are defined as armchair-shaped erosional hollows, typified by steep
- 47 headwalls, typically arcuate and with lateral spurs, and with overdeepened floors, often
- occupied by a lake or bog (Evans and Cox, 1974, 1995; Barr and Spagnolo, 2015). The
- 49 typology of cirques is extremely diversified and, therefore, the limits of this concept are
- 50 difficult to define (Mîndrescu and Evans, 2014). Their origin is associated with the first
- steps of glaciation (Benn and Evans, 2010), although the mechanisms involved in their
- formation and upgrowth are still under debate (Sanders et al., 2012, 2013). Since the first
- 53 studies, the importance of palaeoclimatic conditions driving the glacial and periglacial
- processes that shaped the heads of the valleys is highlighted (Benedict, 1973; Delmas et
- al., 2015; Barr et al., 2017). Consequently, several parameters such as the distribution,
- aspect, floor elevation and morphometry of the glacial cirques have been used to infer
- 57 palaeotemperatures, precipitation gradients, cloud-cover and prevailing wind directions
- during glaciation (e.g. Dahl and Nesje, 1992; Barr and Spagnolo, 2015; Ipsen et al., 2018).
- However, other factors such as the geological structure, post-glacial erosion and the
- 60 uncertainty regarding their origin make even more challenging to reconstruct past climate
- 61 regimes using cirque morphometry (Barr and Spagnolo, 2015).
- Moreover, it must be taken into account that enhanced by paraglacial dynamics, glacial
- cirques continue in many cases to evolve during the interglacial periods following the last
- phase of glaciation (Kleman and Stroeven, 1997; Ballantyne, 2002, 2013). This is
- especially important in the mountain ranges under Mediterranean influence, where glacial
- cirques are one of the most abundant glacial landforms, affected by the last glacial cycle

and subsequent deglaciation (Hughes et al., 2006; 2007). Over recent decades, in particular two palaeoenvironmental techniques, namely lake sediment record studies and Cosmic-Ray Exposure (CRE) dating of glacial landforms using cosmogenic nuclides, have provided evidence of the potential palaeoclimatic information preserved in alpine cirques.

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Many mountain glaciers reached their maximum extent between 26.5 and 19 ka (Clark et 72 73 al., 2009) or a few thousand years before (Hughes et al., 2013). This period is the so-74 called Last Glacial Maximum (LGM), which coincided with the minimum sea level at 75 global scale (Clark et al., 2009; Hughes et al., 2013). However, this maximum extension 76 was reached before or after the LGM in many other mountains, which is known as the 77 Local Last Glacial Maximum (LLGM). In the Mediterranean region, particular attention 78 has been paid to the reconstruction of the sequence of glacial phases recorded on cirque 79 floors in the form of moraine systems revealing either (i) phases of stabilization during 80 the long-term retreat since the LLGM, or (ii) short periods of glacial readvance. This is 81 the case in the Rila mountains, Balkan Peninsula, where 2 km-long cirques include a large 82 number of moraines that formed between the LGM and the Younger Dryas (YD) (12.9– 83 11.7 ka, Walker et al. 2009; GS-1 Rasmussen et al., 2014), i.e. within the approx. 24-12 84 ka period (Kuhlemann et al., 2013). One of the best examples of the geomorphological 85 interest of these small glacial circues comes from Mount Olympus, southern Balkan 86 Peninsula, where moraines from the Late Glacial to the latest Neoglacial advances (from 87 15.6 ± 2.0 to 0.64 ± 0.08 ka) are distributed over a distance of less than 500 m (Styllas et 88 al., 2018). Similarly, in the Dinaric mountains, Žebre et al. (2019) summarized previous 89 studies of numerous small circues with a sequence of moraine systems dating back from 90 Oldest Dryas (OD) to the YD (14.9 ka to 11.7 ka). Small cirques in the Apennines also 91 retain geomorphic evidence in the form of moraine arcs formed between the LLGM and 92 present-day (Giraudi, 2012; Baroni et al., 2018). There are many other examples of the 93 potential of small cirques, including information about glacial oscillations occurred 94 during Termination-1 (19 to 11.7 ka), such as in the Anatolian peninsula. Here, in Mount 95 Geyikdag (Sarıkaya et al., 2017) and Mount Uludag (Zahno et al., 2010), ca. 1 km long 96 cirques contain moraines from the LLGM to the Holocene. A similar pattern is found in 97 small cirgues of the High Atlas, Morocco, with moraine systems from different phases 98 encompassing the entire deglaciation until the early Holocene (Hughes et al., 2019).

Small glacial circues with rich palaeoclimatic information are not exclusive of the Mediterranean region. Cirques are also frequent glacial features in continental ranges as, for example, those in Central Europe. Similar glacial chronologies as those described above have been reported from circues in the Krkonoše Mountains, Sudetes range (Engel et al., 2014), or in circues of larger formerly glaciated ranges, such as the Tatra Mountains, Northern Carpathians (Engel et al., 2015; Makos et al., 2018; Zasadni et al., 2020), and Parang Mountains, Southern Carpathians (Gheorghiu et al, 2015), where moraines from the OD to the Holocene have been dated. Small cirques in the Alps also provided detailed information on glacial oscillations and the climatic evolution of the Late Holocene (Ribolini et al., 2007; Ivy-Ochs et al., 2009; Hippolyte et al., 2009; Moran et al., 2016; Ivy-Ochs, 2015; Le Roy et al., 2017). Well-known glacial oscillations from the last glacial phases in many circues of several mountain ranges of the Western United States confirming previous results obtained in certain ranges such as the Sierra Nevada (Clark and Gillespie, 1997) have been recently published (Marcott et al., 2019; Laabs et al., 2020) Glacial cirques in mountains that protruded from ice-sheet surfaces, such as those in nunataks, also include geomorphic evidence of glacial fluctuations during the last glacial cycle. This is the case of the Macgillycuddy's Reeks (South Ireland), where small cirques host moraines from the LLGM and even younger (Barth et al., 2016). Moreover, glacial cirques can also preserve accurate data on the last deglaciation phases of ice-sheets, as well as on Neoglacial advances during the current interglacial. For example, up to four glacial advances that occurred after the complete disappearance of the Scandinavian Icesheet have been evidenced in numerous cirques in Norwegian mountains (Dahl and Nesje, 1992; Paasche et al., 2007). In Svalbard, small cirques contain moraines revealing that the YD was only a minor advance, smaller than those of the Little Ice Age (from 1300 to about 1850 CE, LIA) (Mangerud and Landvik, 2007). One of the best examples of the palaeoclimatic significance of these cirques can be found in the Tröllaskagi Peninsula, northern Iceland, formerly covered by the Icelandic Ice-Sheet (Ipsen et al., 2018). In this peninsula, small cirques a few km long host up to 12 Holocene moraine complexes corresponding to various stages of the LIA, and even different glacial advances over the last decades (Fernández-Fernández et al., 2019). Moreover, other cirques in the Tröllaskagi Peninsula include a set of erratic boulders, moraines, and several generations

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132 the OD to the present (Tanarro et al., 2019; Fernández-Fernández et al., 2020). 133 In the Iberian Peninsula, small glacial cirques including several moraine systems have 134 been studied in the highest mountain ranges (Oliva et al., 2019). In the Eastern Pyrenees, 135 very small cirques preserve erratic boulders, complex moraine systems and rock glaciers 136 that reveal a complex glacial evolution from the LLGM to the Holocene (Pallàs et al., 137 2006, 2010; Delmas, 2015; Andrés et al., 2019; Jomelli et al., 2020). In the Central 138 Pyrenees, small cirques include moraine complexes from the OD to the Holocene 139 (Delmas, 2015; Palacios et al., 2017a; Crest et al., 2017; Tomkins et al., 2018), with only 140 a few sites recording Neoglacial advances (García-Ruiz et al., 2014). In the Cantabrian 141 Mountains, the heads of some circues host rock glaciers and moraines from the OD to the 142 Early Holocene (Rodríguez-Rodríguez, 2016, 2017). In the Iberian Range, cirques < 1 143 km long preserved moraines from the LLGM, OD and remnants of fossil debris-covered 144 glaciers, some of which were active until the Middle-Holocene due to their favourable 145 orientation (Fernández-Fernández et al., 2017). Other cirques contain moraines from the 146 LLGM to the OD, when the formation of rock glaciers occurs, and even subsequent 147 deglaciated cirque steps, coetaneous with the YD (García-Ruiz et al., 2020a). In the 148 Central Range, small circues in the massifs of Guadarrama (Palacios et al., 2011, 2012a, 149 2012b; Carrasco et al., 2016) and Gredos (Domínguez-Villar, 2013; Carrasco et al., 2013, 150 2015) include moraine complexes from the LLGM to the OD, YD and even the Holocene, 151 depending on the altitude and prevailing aspect of the cirques. Besides, a detailed study 152 of the peatbogs from a glacial cirque in Gredos has provided valuable palaeoclimatic 153 information from the deglaciation from the OD to the present (López-Sáez, et al., 2020). 154 The Sierra Nevada (South Iberia) contains 65 glacial cirques with geomorphic evidence 155 indicative of the glacial and palaeoclimatic evolution from the LLGM to the recent times 156 (Palacios et al., 2016; Palma et al., 2017) and, in some of these cirques, even from the 157 previous glacial cycle (Palacios et al., 2019). Regarding the Holocene, several 158 palaeoenvironmental archives such as lake records and peatbogs inside glacial cirques 159 testify to a remarkable climatic variability (Oliva, 2009; Oliva and Gómez-Ortiz, 2012), 160 as well as the impact of human activity (Anderson et al., 2011; García-Alix et al., 2013, 161 2017). Two of these circues held also small glaciers during the LIA, which were the 162 southernmost in Europe (Gómez-Ortiz et al., 2009, 2015, 2018). One of them, the Veleta 163 cirque, shaped on the northern slope of the Veleta peak, contains valuable information on

of rock glaciers and debris-covered glaciers showing a complex glacial evolution from

environmental dynamics from the first glacial phases of the LIA to the present-day, with evidence of: (i) the succession of several glacial advances and the gradual melting of the glacier, (ii) rock avalanches, (iii) landslides and, (iv) permafrost degradation. These processes, highly sensitive to small climatic variations, led to intense geomorphic readjustments typical of the paraglacial phase (Oliva et al., 2016; Palacios et al., 2019; Serrano et al., 2018; Gómez-Ortiz et al., 2019 and references therein). As a result, landforms from this circue are considered key geoindicators for monitoring the impact of climate change on mountain geomorphological processes in southern Europe (Gómez-Ortiz et al., 2019). The other cirque that hosted a glacier during the LIA in the Sierra Nevada is located under the northern rock wall of Mulhacén peak. To date, although it contains a wide range of glacial depositional and erosional features, this cirque has been little studied. The only palaeoclimatic records from this circue are those provided by the sediments of La Mosca Lake, which led to a detailed picture of the Mid-Late Holocene palaeoecological evolution as well as to the evidence of several glacial advances during the Late Holocene (Manzano et al., 2019; Oliva, 2009; Oliva et al., 2019).

On the other hand, the interpretation of the palaeoenvironmental records preserved in the interior of glacial cirques can come up against important difficulties, such as the alteration of glacial reliefs depending on the intensity of slope processes and the frequency of rock avalanches altering glacial landforms (Deline et al., 2015; Mercier et al., 2017; Knight et al., 2018) or the effects of neotectonic activity in cirque evolution (Oskin and Burbank, 2005). Regarding CRE dating, these processes can hinder the presumed undisturbed exposure of bedrock surfaces, rock glacier and moraine boulders to be dated, or potentially lead to nuclide inheritance in these surfaces (Li et al., 2016; Çiner et al., 2017; Köse et al., 2019),

Therefore, the current state of knowledge still presents some open questions, such as:

- (i) What is the origin of the complex moraine formations frequently found inside the cirques and covering a wide chronological range?
- (ii) Why do glaciers, once confined within the cirques, tend to evolve in many different ways, forming, in some cases, debris-covered glaciers or rock glaciers, and, in other cases, multiple moraine ridges developing in short periods of time?

(iii) Is it possible to circumvent the impact of geomorphological processes occurring
 inside the glacial circumstates to correctly extract the palaeoenvironmental significance
 preserved in their records?

In order to address these questions, the glacial landforms existing in one of the Mediterranean cirques housing the most detailed sequence of the environmental evolution since the LLGM, the Mulhacén cirque, have been analysed. To investigate the potential common behaviour and infer its palaeoclimatic implications on different spatial scales, the palaeoclimatic and geomorphologicalal evolution of this cirque is compared to that of others in the whole massif of Sierra Nevada, the Iberian Peninsula and other mountains of the Northern Hemisphere for which geomorphological and CRE data are available.

2. Study area

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- 205 Located in the southern fringe of Iberia at latitude of 37°N, the Sierra Nevada is the 206 highest massif of the Betic Range (Fig. 1). It includes the highest peaks in western Europe 207 outside the Alps, with elevations exceeding 3000 m a.s.l. in its western fringe, such as 208 Mulhacén (3479 m, 37°03′12″N 3°18′41″W) and Veleta (3396 m, 37°03′22″N, 209 3°21'56"W). The landscape of the massif results from its bordering location between 210 different climatic influences: continental (Europe/Africa), maritime 211 (Atlantic/Mediterranean) and subtropical high pressure belt/mid-latitude westerlies 212 (Oliva et al., 2011). Currently, the mean annual air temperature at 2500 m is 4.4 °C (1965-213 1992), whereas at the highest summits of the massif at 3400 it is 0 °C (2001-2016) 214 (Gómez-Ortiz et al., 2019). Annual precipitation reaches 710 mm, mostly as snow 215 between October and April (Oliva, 2009; Oliva et al., 2008). The bedrock of the summit 216 area, the study site, is composed of micaschists which are intensely affected by periglacial 217 processes. The vegetation cover is very sparse, typical of the high semi-arid 218 Mediterranean mountains, and is mostly distributed across the valley floors near lakes 219 and wetlands (Oliva et al., 2011).
 - The contemporary landscape of the Sierra Nevada is mostly a consequence of past glaciations. A wide range of depositional and erosional landforms of glacial origin are distributed in elevations above 2000 (northern slopes) and 2500 m (southern slopes). They result from the sequence of glacial phases that shaped the highest lands of the massif since the maximum ice extent of the penultimate glacial cycle (Palacios et al., 2019). The

225 LIA was the last period with the presence of glaciers in the Sierra Nevada (Gómez-Ortiz 226 et al., 2009; 2012, 2015, 2018, Palacios et al., 2016). 227 The deglaciation of the circues favoured the appearance of many mountain lakes and bogs 228 which have been sampled for palaeoenvironmental and palaeoclimatic purposes 229 (Anderson et al., 2011; García-Alix et al., 2013; 2017; Jiménez-Moreno and Anderson, 230 2012; Oliva et al., 2011, 2010). At the northern foot of the Mulhacén peak, a glacial lake 231 dammed by a moraine ridge is located next to the cirque mouth at 2892 m a.s.l.: the Mosca 232 Lake (ca. 90 m wide, 160 m long, 1 ha, 3.2 m deep) (Oliva and Gómez-Ortiz, 2012) (Fig. 233 1). A sequence of recessional moraines is connected to the large talus cones generated by 234 frost shattering from the steep northern rock wall of the Mulhacén peak, almost 400 m 235 high. These moraines show multiple collapses and subsidence features (Fig. 1) which 236 evidence the degradation of the small isolated permafrost patches still existing today in 237 the massif (Oliva et al., 2018; Serrano et al., 2018). The large talus cones are also being 238 affected by widespread debris flows transporting sediments downslope during the melting 239 season or during extreme rainfall events in late summer and early autumn. At the foot of 240 the Mulhacén rock wall, there are a few semi-permanent snow-patches as well as a 241 protalus lobe formed following the final deglaciation of the cirque (Fig. 2), which is also 242 indicative of the existence of permafrost at this site (Serrano et al., 2018; Gómez-Ortiz et 243 al., 2019). 244 As revealed by historical documents and lacustrine sedimentary records, the cirque was 245 partly glaciated several times during the Holocene, namely at 2.8-2.7, 1.4-1.2 and 0.51-246 0.24 ka cal BP (Oliva and Gómez-Ortiz, 2012, Oliva et al., 2020). Lake sediment phases 247 recorded peaks with large concentrations of sand and very low organic carbon content, 248 which are attributed to colder and wetter conditions and the development of a glacier in 249 the catchment. The last of these periods coincided with the LIA, which promoted the 250 formation of the largest glacier over the three last millennia in the Mulhacén cirque (Oliva 251 and Gómez-Ortiz, 2012). There are other longer sedimentary cores collected from this 252 lake with radiocarbon dates up to 8.4 cal ka BP, which however do not come from the 253 basal sediments, but give a minimum age of lake formation following glacial retreat 254 (Manzano et al., 2019). No chronological evidence is reported about earlier glacial

phases. Therefore, the moraines located upstream of the La Mosca Lake have probably

developed throughout the Holocene, whereas the moraines closing the lake must be older.

3. Methodology

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258 3.1 Geomorphological mapping

- 259 In order to reconstruct the palaeoenvironmental evolution of the Mulhacén cirque since 260 the last glacial cycle, a detailed geomorphological analysis and mapping of the study area 261 was first performed. Although we consider the last glacial cycle as a synonymous of the 262 Late Pleistocene (approx. 123-14 ka; Hughes and Gibbard, 2018), most of the preserved 263 glacial landforms formed during the LGM or later. The different geomorphological units 264 were classified following the criteria used in previous studies of other glacial cirques in 265 this massif focusing on the distribution of moraines, rock glaciers, protalus lobes, rock 266 avalanches, debris flows and rock falls (Gómez-Ortiz et al., 2012; Palacios et al., 2016; 267 Palma et al, 2017). The spatial distribution of the different features of glacial origin 268 revealed in this cirque suggests a sequence of glacial events until the recent disappearance 269 of the LIA glacier. The geomorphological map was also used for planning the sampling 270 strategy for CRE dating in order to establish the chronology of the cirque deglaciation. In 271 addition, historical sources provided information concerning the periods during which 272 glaciers were present during the LIA in the Sierra Nevada (Gómez-Ortiz et al., 2009, 273 2012, 2015, 2018). These were also used to interpret the final deglaciation phases of the 274 Mulhacén cirque.
- 275 *3.2 CRE sampling strategy*
- 276 The sampling strategy for CRE dating is based on the sequence of the mapped landforms,
- 277 from the bottom to the upper part of the cirque. In late summer 2016, when the cirque
- was snow-free and geomorphic features were clearly visible in the field, eight samples
- from the Mulhacén cirque at elevations between 2890 and 2950 m were collected. Two
- of the samples were taken from glacially polished outcrops (MOSCA-2 and MOSCA-3),
- 281 while the others were taken from flat-topped surfaces of boulders located on moraine
- 282 ridges. The characteristics of the sampling sites and complementary field data are
- summarized in Table 1.
- 284 *3.3 CRE laboratory procedures and exposure age calculation*
- 285 The samples were crushed and sieved to the 0.25-1 mm fraction at the 'Physical
- 286 Geography Laboratory' of the Universidad Complutense de Madrid. The following steps
- of the sample preparation process were carried out at the 'Laboratoire National des
- 288 Nucléides Cosmogéniques' (LN2C) of the CEREGE (Centre Européen de Recherche et
- 289 d'Enseignement des Géosciences de l'Environnement, Aix-en-Provence, France). To

extract the cosmogenic ¹⁰Be isotope, a first rough isolation of the quartz mineral fraction of the sieved samples was performed by means of a Frantz LB-1 magnetic separator removing the magnetic minerals. Subsequently, the remaining non-quartz minerals were dissolved through successive acid attacks (a mixture of concentrated hydrochloric (HCl) and hexafluorosilicic (H₂SiF₆) acids in a proportion 1:2). Three consecutive partial dissolutions using concentrated hydrofluoric acid (HF) were performed to ensure removal of any non-quartz mineral from the treated samples and to decontaminate the pure quartz mineral fraction from meteoric ¹⁰Be. For the following steps, pure quartz masses ranged between 6 and 40 g. 150 µL of an accurately weighted ⁹Be carrier solution manufactured in-house from a phenakite crystal ($[^9Be] = 3025 \pm 9 \mu g g^{-1}$; Merchel et al., 2008) were added, and the quartz was totally dissolved in 48% HF (3.6 mL per g of quartz + 30 mL in excess). After the total dissolution, the resulting solution was evaporated until dryness and the solid residues were recovered in 7.1 M HCl. Samples were precipitated with ammonia before successive separations, first through an anion exchange column (Dowex 1X8) to remove iron, and then through one or several cation exchange columns (Dowex 50WX8) to discard boron (isobar) and to separate the Be from other elements (Merchel and Herpers, 1999). All samples contained a lot of muscovite, which was difficult to remove completely and complicated the treatment of some of the samples. Therefore 17 cation exchange columns were performed for samples MOSCA-2, 3, 5 and 7.

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- The eluted Be was precipitated to beryllium hydroxide (Be(OH)₂) with ammonia and oxidized to BeO at 700 °C. As a final step, BeO were mixed with niobium powder in an approximate 1:1 proportion, and then loaded in cathodes for the subsequent measurement of the ¹⁰Be/⁹Be ratio at the ASTER ('Accélérateur pour les Sciences de la Terre, Environnement et Risques') AMS (Accelerator Mass Spectrometry) national facility at CEREGE. A chemical blank was prepared along with the seven samples.
- Sample 10 Be/ 9 Be ratios were calibrated against the in-house standard STD-11, using an assigned 10 Be/ 9 Be nominal ratio of $(1.191 \pm 0.013) \times 10^{-11}$ (Braucher et al., 2015).
- Analytical 1σ uncertainties include uncertainties associated with AMS counting statistics, the standard 10 Be/ 9 Be ratio, an external AMS error of 0.5% (Arnold et al., 2010) and the chemical blank measurement. A 10 Be half-life of (1.387 ± 0.0012) × 10^{6} years was used (Chmeleff et al., 2010; Korschinek et al., 2010). More details of analytical data are given in Table 1.

- 322 ¹⁰Be surface CRE ages were calculated using the CREp online calculator (Martin et al.,
- 323 2017; available online at: http://crep.crpg.cnrs-nancy.fr/#/). The parameters used were:
- 324 the LSD elevation latitude scaling scheme (Lifton et al., 2014), the ERA40 atmospheric
- model (Uppala et al., 2005) and the geomagnetic database based on the LSD framework
- 326 (Lifton et al., 2014). This setting yielded a sea-level high latitude (SLHL) ¹⁰Be production
- rate of 3.98 ± 0.22 atoms g⁻¹ yr⁻¹. Rock density was assumed to be 2.7 g cm⁻³. Partial
- 328 shielding of the surrounding topography from the cosmic-ray flux was calculated through
- 329 the Topographic Shielding Calculator v.2 of CRONUS-Earth Web Calculators (consulted
- 330 2020). CRE ages are reported for the zero denudation scenario, and uncertainties included
- hereafter include analytical and production rate errors (Table 2).
- The results of the analysis carried out in the Mulhacén cirque have been confronted with
- previous studies involving CRE dating, some using ¹⁰Be and others ³⁶Cl, conducted in
- other cirques of the Sierra Nevada and in the rest of the Iberian mountains. In order to
- 335 compare the results obtained from the two cosmogenic nuclides, all the CRE ages
- mentioned in the text were recalculated in accordance with the protocols proposed for the
- 337 Iberian Peninsula in Oliva et al. (2019). Also, to compare our results with those obtained
- 338 in cirques from other mountains outside Iberian Peninsula where CRE dating was
- performed, we have checked all the publications of the last five years that include a
- detailed description of the CRE age calculation protocols and only those with similar
- protocols to those applied to Iberian mountains were considered.
- In addition, the information on the glacial chronological sequences provided by the new
- 343 CRE ages were compared with the palaeoenvironmental evidences previously inferred
- from the sedimentary studies of La Mosca Lake (Oliva and Gómez-Ortiz, 2012; Manzano
- et al., 2019) as well as from other lakes and peatbogs of the massif (Anderson et al., 2011;
- García-Alix et al., 2013, 2017; Jiménez-Moreno and Anderson, 2012; Oliva et al., 2011,
- 347 **2010**).

- 4. Results
- 349 4.1 Geomorphological setting and CRE sample selection
- 350 The results of the geomorphological analysis of the Mulhacén cirque are summarized in
- 351 the geomorphological map (Fig. 2) as well as along a transect including the most
- remarkable glacial and periglacial features (Fig. 3).

353 For CRE dating, we focused on rocky outcrops with glacially polished surfaces that 354 protrude above the bottom of the cirque and had probably not been covered with sediment 355 after deglaciation; however, few suitable sites were found as debris cover is very abundant 356 in the Mulhacén cirque floor. In addition, due to an efficient weathering of the micaschist 357 bedrock, the original glacial surface of the few outcrops standing in the cirque floor is 358 often not preserved and, therefore, only two polished surfaces were considered suitable 359 for CRE dating. One sample was taken from a bedrock step below La Mosca Lake 360 (MOSCA-2) and another one from a bedrock step above it (MOSCA-3) (Figs. 4 and 5). 361 The Mulhacén cirque ends downvalley in a large step that descends vertically from 2920 362 to 2600 m to the Valdecasillas gorge. The sample MOSCA-2 is located just on the edge 363 of this step, next to two moraine ridges damming the La Mosca Lake. Due to the highly 364 weathered surface of the rocks, only three boulders from these ridges appeared suitable 365 for CRE dating (not affected by denudation, and standing out above the cirque floor): one 366 from the outermost moraine (MOSCA-1) and two from the ridge closest to the lake 367 (MOSCA-6 and MOSCA-7). The quartz content in sample MOSCA-1 was too low for 368 ¹⁰Be extraction and was considered non-suitable for further analyses. 369 A bedrock step above the lake was sampled (MOSCA-3). Large boulders within fine-370 grained matrix are distributed on this step. In contrast to the majority of moraines in the 371 Sierra Nevada, composed of abundant fine sediments due to the weathering of the 372 micaschists (Gómez-Ortiz et al., 2012; Palacios et al., 2016), this ridge preserved metric-373 size boulders. However, most of them were intensely fractured by frost shattering after 374 their deposition, and only one sample from this landform was collected (MOSCA-5). 375 Overlapping this ridge, there is a moraine including both boulders and fine sediments. 376 One boulder from this ridge was sampled, whose surface showed glacial striations 377 ensuring that it has not undergone major degradation since its deposition (MOSCA-4). 378 Above this ridge, there is a large area with a massive accumulation of boulders that 379 corresponds to a large moraine system composed of several ridges. Due to the steep slope, 380 the shape of some of these ridges has been altered by rock falls and slope readjustment, 381 though they still show an arcuate morphology. All these moraines seem to be affected by 382 slow flow of the surface sediments towards the base of the cirque, likely due to the presence of buried ice and permafrost patches such as those detected and monitored in 383 384 the neighbouring Veleta cirque (Gómez-Ortiz et al., 2015; 2019). There are several

- depressions of collapse and subsidence features (Fig. 2) related to the degradation of the
- frozen mass located below the debris cover, which also affected the stability of the
- boulders. In addition, there are also abundant rocks fallen from the steep north wall of the
- 388 Mulhacén, located only at less than 300 m from these deposits. Consequently, it was
- difficult to find boulders suitable for CRE dating and indicative of the age of stabilization
- of this moraine system. In fact, only one likely appropriate boulder in an external ridge
- was sampled (MOSCA-8).
- 392 According to historical sources (Gómez-Ortiz et al., 2009, 2015, 2018), the LIA glacier
- did not reach these outermost moraines and only occupied the innermost moraine systems
- 394 closer to the northern Mulhacén rock wall, which is very active supplying debris to the
- 395 talus cones. Currently, an active protalus lobe at the foot of the rock wall indicates the
- 396 occurrence of permafrost conditions at this site.
- *4.2 CRE results*
- As usual, we highly recommend using only the internal (analytical) errors, when the ¹⁰Be
- ages are compared amongst them (at the study site and at the other sites), because they
- are all equally impacted by the production rate uncertainty the production rate
- 401 uncertainty has to be considered when the ¹⁰Be ages are compared to other chronological
- 402 data.
- The CRE results are presented in Table 2, and Figs. 6 and 7. The two samples from
- 404 polished bedrock lead to indistinguishable CRE ages: 13.5 ± 0.8 ka (MOSCA-2) and 14.2
- ± 0.8 (MOSCA-3). Assuming that the age of the bedrock samples reveals the age of the
- 406 final ice retreat in the mouth of the cirque, this leads to a mean CRE age of 13.8 ± 0.8 ka
- 407 (n=2).
- The two samples from the moraine just below La Mosca Lake yielded to CRE ages of
- 409 12.6 ± 0.9 ka (MOSCA-6) and 16.4 ± 1.1 ka (MOSCA-7). Therefore, although from the
- same moraine, one boulder leads to an age 1.6 ka younger than the age of the bedrock
- 411 surface where it rests, whereas the other leads to an age 2.2 ka older, respectively.
- 412 Considering that the exposure age of the sample MOSCA-7 is older than that of the
- bedrock where it rests, it can be considered an outlier.
- The samples taken above the lake one located in a moraine of one of the lower ridges
- 415 (MOSCA-4, 9.8 ± 0.8 ka) and the other in a moraine of one of the highest ridges

- 416 (MOSCA-8, 10.2 ± 0.7 ka) lead to indistinguishable CRE ages whose mean value is
- $417 10.0 \pm 0.7$ ka (n = 2). The deposition of these boulders resulted from different glacial
- events separated by only a few hundred years. The sample taken from the ridge formed
- by large boulders, located above the lake, leads to a CRE age of 23.4 ± 1.5 ka (MOSCA-
- 5). Significantly older than the landforms located more distant from the cirque wall (mean
- 421 CRE age of 14.1 ± 0.9 ka (n=4)), this age is inconsistent with the geomorphological
- sequence, probably due to nuclide inheritance.

5. Discussion

- The results obtained in this work provide an approximate chronology of glacial retreat in
- 425 the Mulhacén cirque, located at the foot of the highest peak in the Iberian Peninsula. The
- scarce number of landforms suitable to be sampled for the application of CRE dating
- 427 impeded providing an accurate sequence of glacial oscillations for the transition from the
- 428 Late Glacial to the Early Holocene, when most of the cirque became ice-free. This is due
- 429 to the very intense geomorphological processes of the northern wall of the Mulhacén,
- with intense rock fall activity generating large talus cones that have covered some of the
- most recent moraines. In addition, the micaschist bedrock fractures very easily and it is
- therefore very difficult to find original glacial surfaces in boulders or bedrock outcrops.
- 433 5.1 Sequence of geomorphological phases in the Mulhacén cirque
- Despite the small number of available samples to which the CRE dating method can be
- applied with confidence, the combination of the results obtained in this study with those
- from previous studies points to the occurrence of a number of geomorphological phases
- inside the Mulhacén cirque (Table 2, and Fig. 6 and 7):
- 438 (i) The deglaciation of most of the circue culminated approximately at 14 ka including
- 439 the area where La Mosca Lake is located –, which is related to the Bølling-Allerød
- Interstadial (B-A; 14.6-12.9 ka; GI-1 Greenland ice cores; Rasmussen et al., 2014). These
- results are compatible with the previous study that suggested that lake sedimentation
- started before 8.4 cal ka BP (Manzano et al., 2019).
- 443 (ii) Other published works focusing on La Mosca Lake sediments reported glacial
- oscillations during the Late Holocene. Three coarse-grained layers deposited in the lake
- bottom were interpreted as indicating the existence of a glacier in the catchment at 2.8-
- 446 2.7, 1.4-1.2 and 0.51-0.24 ka cal BP (Oliva and Gómez-Ortiz, 2012). The presence of a
- 447 glacier during the LIA in this cirque was also confirmed by historical sources (Gómez-

- Ortiz et al., 2009, 2015, 2018). Our results show that the extent of these palaeoglaciers
- was spatially limited, with small features only located at the base of the wall in the
- 450 concavities that are currently occupied by talus cones. According to our results, the last
- Early Holocene glacial advance that generated moraines reached a distance of only 300
- m from the foot of the wall.
- 453 (iii) The set of moraines located between the talus cones and the lake, with up to 6
- different ridges in less than 400 m, is indicative of repeated glacier advances and retreats
- in the early stages of the Holocene, towards ca. 10 ka. These Early Holocene glaciers had
- a significant debris cover provided by the northern rock wall of the Mulhacén cirque.
- Indeed, the abundance of boulders across this moraine system shows evidence of the very
- 458 high debris supply on palaeoglaciers from the rock wall. During their retreat, the ice
- 459 gradually melted leaving multiple collapse depressions.
- The age difference between the samples that mark the deglaciation of the mouth of the
- 461 cirque during the B-A Interstadial, and those of the moraines deposited at the bottom of
- 462 the cirque during the Holocene, was also shown in a previous work (Oliva et al., 2015)
- by the application of Schmidt Hammer Exposure Dating, with higher rebound (R values)
- indicative of less weathered surfaces and thus younger ages. The moraine that led to a
- ¹⁰Be CRE age of 14 ka coincided with an R value of 55 ± 4 , whereas the R values in the
- 466 10 ka ridge were higher, 64 ± 6 (Oliva et al., 2015).
- 467 (iv) There are some samples that show evidence of cosmogenic nuclide inheritance. In
- 468 fact, the outlier sample MOSCA-5 is related to the unique moraine formed by large
- boulders with no fine-grained sediments. The overestimated exposure age may be related
- 470 to boulders fallen from the wall and transported supraglacially on the ice surface with no
- or very little surface readjustment. This process has earlier been inferred in other
- 472 mountain circues, especially in small ones with limited distance between the headwall
- and the moraines (Li et al. 2016; Ciner et al. 2017; Köse et al., 2019).
- 474 (v) According to our results, there is no evidence of glacial activity during the Holocene
- Thermal Maximum (HTM: 9-5 ka, Renssen et al., 2009) in this cirque.
- 476 (vi) The current intense geomorphological activity of the rock wall is shown by the
- formation of large rock fall cones and talus since the disappearance of the LIA glacier.
- These rock fall cones are underlain by permafrost, as it is demonstrated by the recent
- development of a protalus lobe at the foot of such rock wall (Serrano et al., 2018).

- 480 5.2 Common and diverse geomorphological phases in the Sierra Nevada cirques and
- 481 their palaeoclimatic and topographic significance.
- In the Sierra Nevada, there is remnant evidence of two Late Pleistocene glacial cycles
- that are preserved in the form of moraine complexes in the southern valleys of the massif.
- During the last glacial cycle, the maximum ice extent slightly predated the LGM, as it has
- been shown by CRE dating of boulders from lateral moraines from the largest glacial
- 486 valleys (Gómez-Ortiz et al., 2012; 2015; Palacios et al., 2016; 2019; Oliva et al., 2019).
- 487 Moreover, the sequence of glacial phases inferred from several other cirques in the Sierra
- 488 Nevada (Gómez-Ortiz et al., 2012; 2015; Palacios et al., 2016; 2019; Oliva et al., 2019)
- show a similar sequence of the last deglaciation to the one presented here from the
- 490 Mulhacén cirque, whose main phases have been summarized in Table 3. It is important
- 491 to consider that glacial cirques do not always form at the head of the valleys, such as in
- 492 the Sierra Nevada where many develop on lateral slopes of the main valley more
- favourable for snow accumulation, especially on E- or NE-facing slopes.
- 494 Several common spatio-temporal patterns can be inferred with regards to the deglaciation
- evolution in cirques of the Sierra Nevada (Table 3 and Fig. 7):
- 496 (i) All the cirques studied in the Sierra Nevada were deglaciated at the beginning of the
- 497 B-A, although glaciers either remained as small features or regenerated during the
- 498 following cold phases. Six of the eight cirques studied Sierra Nevada show
- 499 indistinguishable deglaciation ages based on samples obtained from polished bedrock
- surfaces, with an average of 14.0 ± 1.2 ka (n=6) (samples from Palacios et al., 2016 and
- present work) (Fig. 7 and 8A). Therefore, these datasets suggest that shortly after the
- onset of the B-A the valley floors were ice-free and glaciers were confined within the
- 503 cirques as small features at the foot of the highest rock walls and probably disappeared
- from many others (Fig. 8B and C).
- 505 (ii) There are only two cirques (Río Seco and Caldereta) including samples of polished
- bedrock surfaces that suggest a more recent deglaciation (12.0 \pm 1.1 ka, n= 3) (Palacios
- et al., 2016). They do not preserve evidence of later moraine formations or the
- development of large rock glaciers. These younger ages would reveal the presence of
- glaciers after the main deglaciation at the bottom of the cirque, as in other cirques, but
- 510 they did not generate moraines or rock glaciers when the ice melted, most likely due to
- 511 the low debris supply from relatively stable rock walls.

512 (iii) In most cases, after the most important deglaciation phase following the LLGM,

glaciers persisted in the circues, or they regenerated during the YD (Fig. 8.D). Some inner

moraines in the cirques have a stabilization age similar (considering their uncertainties)

to polished bedrock surfaces, with an average age of 11.4 ± 1.0 ka (n= 6) (Palacios et al.,

516 2016 and present work).

- 517 (iv) The importance of topographical constraints controlling the geomorphological
- evolution in each cirque of the Sierra Nevada after its final deglaciation at the end of the
- YD. Moreover, the topographical conditions of each cirque favoured the generation of
- 520 different type of landforms:
- a) The cirques with summit altitudes >3300 m and floors > 2950 m, north exposed
- and with cirque walls > 300 m high were the most climatically sensitive considering
- the geomorphological setting. They include evidence of glacier changes from the YD
- to the present. This is the case of the Veleta and Mulhacén cirques. But it is
- noteworthy that each one evolved differently: in the case of the Veleta cirque, a large
- polygenic moraine was formed (Gómez-Ortiz et al., 2009; 2012, 2015; Palacios et al.,
- 527 2019) (Fig. 8E.4), whereas in the Mulhacén cirque a sequence of moraine ridges
- developed (Fig. 8E.5). During the HTM glaciers probably melted away in these
- cirques, but regenerated under the cold Neoglacial conditions prevailing during the
- Late Holocene, namely during the LIA (Gómez-Ortiz et al., 2009; 2012, 2015, 2018;
- Oliva and Gómez-Ortiz, 2012; Palacios et al., 2016). These Neoglacial advances
- enlarged the polygenic moraine in the Veleta cirque, and must have generated new
- moraine ridges in the Mulhacén cirque. However, the very intense geomorphological
- processes of the Mulhacén wall accumulating large masses of debris at the foot of the
- steep slope of the cirque and the readjustment of the moraine boulders have hidden
- the associated evidence.
- b) Cirques with summit altitudes between 3000 and 3200 m, floors located above
- 538 2800 m, with small headwalls (elevation range < 200 m) and east-facing produced
- only small proto-rock glaciers. This is the case of the Peñón Colorado, Río Seco and
- Caldereta cirques. In these cirques, where proto-rock glaciers developed, their fronts
- stabilized shortly after they formed (Fig. 8E.1).
- c) In circues with summit altitudes < 3000 m and floors below 2800 m, the retreating
- glaciers left one or more moraines or even none depending on the intensity of the

paraglacial readjustment of their walls rather than the climate oscillations, and no rock

glaciers formed. This is the case of the Hoya de la Mora, Mojón de Trigo and Moro

cirques (Fig. 8.E.3). These moraines are located between 400-700 m away from the

- cirque walls, showing evidence of the small size of the YD glaciers.
- 548 (v) Except in the case of the Veleta and Mulhacén cirques, the landscape of the rest of the
- 549 cirques is similar to that existing during the HTM. No new glacial landforms have formed
- since and they have been affected only by limited periglacial and slope processes during
- Late Holocene cold phases (Oliva, 2009; Oliva and Gómez-Ortiz, 2012; Oliva et al., 2010;
- 552 2011 and 2020).
- 553 (vi) Based on the evolution and chronology of the Sierra Nevada glacial cirques, we
- assume that the origin of the rock glaciers they host are related to the process of
- 555 deglaciation that occurred at the beginning of warm periods in connection with
- paraglacial processes which triggered an intense geomorphological activity of the rock
- walls (Kleman and Stroeven, 1997; Ballantyne, 2002; 2013, Knight et al., 2018; Serrano
- 558 et al., 2018).
- (vii) As in other cirques of the Sierra Nevada at the end of the Late Pleistocene glacial
- 560 phases, the end of Neoglaciation with the disappearance of the LIA glaciers has favoured
- the development of a small rock glacier in the Veleta cirque (Gómez-Ortiz et al., 2015;
- 562 2019) and a protalus lobe in the Mulhacén cirque (Serrano et al., 2018), where small
- isolated permafrost patches exist under the debris cover (Oliva et al., 2018; Serrano et al.,
- 564 2018). These landforms are geoindicators of the end of cold periods and are out of balance
- with current climate conditions, and therefore, tend to become inactive (Oliva et al., 2018;
- 566 Gómez-Ortiz et al., 2019).
- 5.3 The typology of the cirques in Sierra Nevada in the context of the Iberian Peninsula
- 568 Studies similar to those conducted in the Sierra Nevada have been carried out in other
- 569 Iberian mountain ranges, although most of them focused on individual cirques, with the
- exception of some studies in the Pyrenees. The chronology of formation of different
- landforms in the circues of the Sierra Nevada is highly similar to the periods of formation
- of analogous geomorphological features in other Iberian cirques (Table 4).
- 573 (i) Based on the available literature, we have analyzed a total of 26 Iberian cirques, where
- 574 the maximum and minimum CRE ages of deglaciation at their mouth vary from $16.3 \pm$

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575
       3.3 ka to 13.2 \pm 0.7 ka, considering either the time of stabilization of external moraines,
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       or the outcropping of the bedrock due to the disappearance of the ice (Table 4). The
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       average age of deglaciation taking into account all these circues is 15.1 \pm 1.3 ka (n = 21)
578
       (Table 4 and Fig.9.A). According to marine and terrestrial records from the Iberian
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       Peninsula and surrounding areas (Fletcher et al., 2010a, 2010b, Moreno et al., 2014;
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       López-Sáez, et al., 2020), a sudden temperature increase that led to values similar to
581
       present is recorded at the beginning of B-A. The accelerated deglaciation of the cirques
582
       probably during the B-A has been detected in many other cirques in Iberian mountains
583
       (Palacios et al., 2017b), such as the Central Range (Palacios et al., 2011; Palacios et al.,
584
       2012a, 2012b; Carrasco et al., 2015), Iberian Range (Fernández-Fernández et al., 2017;
585
       García-Ruiz et al., 2020a), Central Cantabrian Mountains (Rodríguez-Rodríguez et al.,
586
       2017) and the Pyrenees (Pallàs et al., 2006, 2010; Delmas, 2015; Palacios et al., 2017a;
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       Crest et al., 2017; Tomkins et al., 2018; Andrés et al, 2019; Jomelli et al., 2020).
588
       (ii) Although they still need to be studied in much more detail, YD glaciers existed within
589
       the cirques of other Iberian mountain ranges (García-Ruiz et al., 2016). There are few
590
       studies on YD related moraines in the in the Pyrenees (Pallàs et al., 2006; Crest et al.,
591
       2017; Palacios et al., 2017a; Andrés et al., 2019) and one in the Central Range (Carrasco
592
       et al., 2015). However, dating of bedrock surfaces in nine other cirques shows that
593
       glaciers occupied the bottom of the cirques until the beginning of the Holocene (Palacios
594
       et al., 2011, 2012b, 2017a; García-Ruiz et al., 2020a; Crest et al., 2017 Andrés et al.,
595
       2019; Jomelli et al., 2020). Although the age of the bedrock surfaces and the moraines
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       resting on them suggest the occurrence of different glacial phases, in this global analysis
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       the mean of the two types of ages can be interpreted as indicative of the final ice
598
       disappearance from the cirques. The average age of the deglaciation of all these cirques
599
       is 11.2 ka (n=17), which fits within the onset of the Holocene, once temperatures
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       increased following the YD (Table 4 and Fig.9.A). In the rest of the cirques, there is a
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       lack of information on the glacial impact of the YD due to lack of dating or because the
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       bottoms are still occupied by rock glaciers. According to the 9 cirques with dated
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       moraines, the glaciers in the YD never exceeded the limits of the cirque, with lengths
604
       between 500 and 1200 m (Table 4). The abrupt warming after the YD cold period is
605
       evidenced in marine and terrestrial records of the regional Iberian context (Fletcher et al.,
606
       2010a, 2010b, Moreno et al., 2014; García-Ruiz et al., 2016; López-Sáez, et al., 2020).
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- 607 (iii) In some cases, boulder exposure ages from the front of the rock glaciers are provided.
- 608 It is assumed that these ages represent the melting of the interstitial ice, i.e. the
- stabilization of the rock glacier fronts (Zasadni et al., 2020). In six cirques, rock glaciers
- 610 likely stabilized somewhat before or at the start of the B-A Interstadial (Rodríguez-
- Rodríguez et al., 2016, 2017; Andrés et al., 2019; García-Ruiz et al., 2020a), but the roots
- of the large rock glacier complexes stabilized during the Middle to Late Holocene
- 613 (Palacios et al., 2017a; Andrés et al, 2019; Jomelli et al., 2020). The stabilization periods
- 614 (onset, duration and ending) of two debris-covered glaciers from the Iberian Range were
- determined by the aspect, and hence by solar radiation. They range from the beginning of
- the B-A to the Middle Holocene (Fernández-Fernández et al., 2017) (Table 4).
- 617 (iv) Neoglacial landforms are located in cirque floors only at the foot of the northern face
- of peaks above 3000 m in the Central Pyrenees (García-Ruiz et al., 2014; 2020b).
- (v) In summary, in all highest Iberian mountains geomorphological features similar to
- 620 those in the Sierra Nevada cirques have been observed, regarding both type and
- development period. The deglaciation of the cirques started at the beginning of the B-A
- and new glaciers developed during the YD. Following the glacial retreat at the onset of
- 623 the B-A, as well as during the Early Holocene, shrinking glaciers triggered the formation
- of rock glaciers in areas of high sediment supply (e.g. at the foot of steep rock walls),
- with their fronts stabilizing shortly after forming. The sediment supply of the cirques
- depends on the rock types (see Table 4), but also on the persistence of weathered substrate
- of the cirque rock wall, especially in granitic areas (Palacios et al., 2017a; Andrés et al.,
- 628 **2019**).
- 629 5.4 Diversity of landforms associated with cirque deglaciation in the Northern
- 630 Hemisphere
- In order to explore if the geomorphological evolution within glacial circues in the Sierra
- Nevada and other Iberian ranges represents a local or a more widespread climatic pattern,
- we have examined dozens of glacial cirques in the Northern Hemisphere where CRE data
- are available. To this end, we have analysed chronological data from cirques from the
- Mediterranean region, Central Europe, the British Isles and subpolar regions. We have
- also included the analysis of numerous cirques in the western United States where precise
- chronologies are available (Marcott et al., 2019).

638 (i) Located between 31 and 37° N, 11 Mediterranean cirques from the Anatolian and 639 Balkan peninsulas and north Morocco benefitting from new detailed CRE chronology 640 were analysed (Table 5). Eight of the 11 circues include chronological data on their initial 641 deglaciation. All of them became ice-free at an averaged CRE age of 14.1 ± 1.4 ka (n=8), 642 in consistency with the findings from Iberia (Table 5, Fig. 9.B) (Köse et al., 2019; 643 Sarıkaya et al., 2017; Styllas et al., 2018; Hughes et al., 2018; Žebre et al., 2019). Seven cirques include moraine ridges inside the cirques, five of which formed during the YD 644 645 (Köse et al., 2019; Sarıkaya et al., 2017; Styllas et al., 2018; Hughes et al., 2018; Žebre 646 et al., 2019). Only a few limestone cirques in the Anatolian peninsula reported 647 significantly younger moraine ages (Sarıkaya et al., 2017). With the exception of these 648 excluded cases, the average CRE age of these moraines is 12.0 ± 0.8 ka (n=8) (Table 5, 649 Fig. 9.B) and the glacier lengths ranged from 700 to 2500 m. There are only some CRE-650 dated rock glaciers in this region. Evidencing cosmogenic nuclide inheritance, they were 651 not included in our analysis (Ciner et al. 2017). Neoglacial landforms appear only in 652 cirques with peaks near 3000 m (Styllas et al., 2018), with the exception of Anatolia 653 (Sarıkaya et al., 2017). 654 (ii) 17 cirques from Central Europe mountains located between 45° and 50° N were 655 analysed (Table 6). Most of them are situated in the Carpathian Mountains, only two 656 being low-altitude cirques from the Alps. In the Central Alps, the YD glaciers were 657 beyond the cirque area and, hence, they are not considered in our analysis. Up to 15 658 circues include moraines with CRE ages ranging from 15.5 ± 0.4 to 13.2 ± 1.0 ka, leading 659 to an averaged CRE age of 14.7 ± 0.9 ka (n=14) (Table 6, Fig. 9C) (Engel et al., 2014, 660 2017; Gheorghiu et al., 2015; Makos et al., 2018; Zasadni et al., 2020). Their stabilization 661 might be related to the beginning of the B-A. Only 6 cirques included moraine formation 662 during the YD/Early Holocene transition, whose CRE ages range between 12.0 ± 1.1 and 663 10.9 ± 0.6 ka, leading to an averaged CRE age of 11.6 ± 0.7 ka (n=6) (Table 6, Fig. 9.C) 664 (Engel et al., 2017; Gheorghiu et al, 2015; Makos et al., 2018). The YD glaciers had 665 limited areas, with lengths between 500 and 800 m. 6 cirques do not host YD moraines, 666 but include rock glaciers, whose fronts stabilized at CRE ages ranging from 12.9 ± 0.7 to 667 10.9 ± 1.0 ka, also during the YD/Early Holocene transition (Moran et al., 2016; Zasadni 668 et al., 2020) (Table 6). There is no evidence of Neoglacial landforms.

(iii) Six cirques from the British Isles and three from northern Iceland were analyzed

(Table 7). Only three British cirques provide information regarding their initial

669

- deglaciation. They became ice-free prior to or during the B-A at an average age of $15.2 \pm$
- 1.1 ka (n=3) (Table 7, Fig. 9D) (Barth et al., 2018). Four British and two Icelandic cirques
- host moraines whose stabilization ages range from 12.0 ± 0.9 to 10.7 ± 1.0 ka, leading to
- an average age of 11.6 ± 1.1 ka (n=6) (Table 7, Fig. 9D). Therefore, they probably formed
- during the YD (Barth et al., 2018; Hughes et al., 2019; Fernández-Fernández et al., 2020).
- The British YD cirque glaciers were between 300 and 950 m long while the Icelandic
- ones were longer, between 1500 and 3000 m. The stabilization of one rock glacier has
- been dated in Britain at 12.5 ± 1.1 ka (n=5; Barth et al., 2018) and two others in Iceland
- lead to a reported average age of 10.1 ± 1.0 ka (n=4; Fernández-Fernández et al., 2020).
- 680 (iv) Until very recently, chronological data about the formation of landforms inside
- 681 glacial cirques from North America was very limited, but a recent study provides
- information about cirgues from the Western United States (Marcott et al., 2019). Located
- between 35° to 47° N, 12 cirgues were examined (Table 8). Six of these cirgues provide
- deglaciation ages based on the chronology of the stabilization of their moraines that
- occurred during the B-A, their ages between 15.2 ± 0.7 and 13.4 ± 0.7 ka lead to an
- average age of 14.4 ± 0.8 ka (n=6) (Table 8, Fig. 9E) (Marcott et al., 2019). Eight of the
- American cirques have internal moraines whose ages between 12.8 ± 0.7 and 10.7 ± 0.6
- ka lead to an average age of 12.0 ± 0.7 ka (n=8) likely corresponding to the YD (Table 8,
- 689 Fig. 9.E) (Marcott et al., 2019). Theses YD glaciers were between 700 and 2800 m long.
- Three of these circues have rock glaciers whose ages of stabilization range between 10.5
- 691 \pm 0.6 and 9.6 \pm 0.5 ka (Marcott et al., 2019).
- 692 (v) The comparison of the different deglaciation chronologies and of the development of
- 693 rock glaciers allows inferring if there is a synchronic pattern across the Northern
- 694 Hemisphere. The comparison of all the averaged ages of the rock glacier front
- stabilization reveals that they chronologically coincide with phases following the
- deglaciation of these cirques, i.e. fifteen of them (average age: 10.8 ± 0.7 ka) after the
- YD, and 5 of them (average age: 15.4 ± 1.5 ka) with the end of the OD (Table 3, 4, 5, 6,
- 7 and 8 and Fig. 9.F). Since most of the analysed circues were deglaciated after the YD,
- most of the rock glacier fronts stabilized shortly thereafter, around 1 ka after deglaciation
- 700 (Palacios et al., 2016; Moran et al., 2016; Palacios et al., 2017a; Fernández-Fernández et
- 701 al., 2020; Marcott et al., 2019; Zasadni et al., 2020).
- 702 (vi) Based on the typology and age of the landforms observed inside the cirques in the
- 703 mid-latitude mountains of the Northern Hemisphere, we concluded that the

- geomorphological dynamics of all these cirgues were similar to those studied in the Sierra
- Nevada and in the rest of the Iberian Peninsula (Fig. 10). As in the Sierra Nevada, the
- diversity of geomorphic landforms in each cirque depends on the local topographic
- characteristics of the altitude of its summits and floors as well as on its orientation and
- 708 prevailing lithology. However, most of the moraine systems and rock glaciers distributed
- inside glacial circues formed during the same climatic phases throughout the Hemisphere,
- 710 from the Last Termination to the present.
- 711 In fact, two distinct geomorphological phases are clearly linked to the warm phases during
- 712 the Last Termintion in the Northern Hemisphere: (i) The glacier margins receded from
- 713 the mouth of the cirques generating in many cases moraines during the climate
- oscillations of the OD and the B-A transition; (ii) in most cirques, the glaciers expanded
- or remained during the YD and generated moraines within the limits of the cirques. These
- glaciers retreated up-valley and abandoned their moraines at the onset of the Holocene,
- 717 when rock glacier fronts stabilized (Fig. 10).

718 **6. Conclusions**

- 719 This work provides robust evidence that glacial cirques reflect the climatic conditions
- prevailing during their formation (Benedict, 1973; Dahl and Nesje, 1992; Delmas et al.,
- 721 2015; Barr et al., 2017; Barr and Spagnolo, 2015; Ipsen et al., 2018), and also preserve
- very valuable information on the climatic and geomorphological evolution that occurred
- 723 during their deglaciation:
- 724 (i) All the examined circues were deglaciated during the OD and the B-A transition. There
- are no geomorphic remnants in none of the examined cirques of the Northern Hemisphere
- supporting the existence of glaciers during the B-A. Although survival of glaciers through
- 727 the B-A in the cirques is fully possible, it is not proven. While glaciers retreated during
- 728 the OD and B-A transition in many cirques, in others the glaciers evolved into rock
- 729 glaciers but their fronts stabilized soon after their formation.
- 730 (ii) There is widespread evidence of the existence of glaciers during the YD in many
- cirques, as shown by: (i) the age of glacier retreat revealed by the age of polished bedrock
- surfaces, and (ii) the age of moraine boulders stabilization. In both cases, results provide
- ages spanning the late YD and early Holocene.
- 734 (iii) The retreat of YD glaciers generated a wide variety of landforms within each cirque:
- 735 (a) Sometimes, after the glaciers disappeared, only polished surfaces remained. This is a

widespread pattern in Iberian mountains, and probably in other mid-latitude mountains, although in other regions there are very few studies that have dated bedrock surfaces located at the bottom of the cirques. (b) In other cases, following the melting of glaciers, moraines rapidly stabilized. (c) Moraines did not form in many cirgues, but the retreat of YD glaciers favoured the development of rock glaciers. In many cases, their fronts stabilized shortly after their formation, even in less than 1 ka (Zasadni et al., 2020). (d) In other cases, debris-covered glaciers could also form. In fact, many of the analysed rock glaciers may have originated from debris covered glaciers formed in previous phases of the long-term deglaciation process (Anderson et al., 2018). The detailed study conducted on the wide variety of cirques in the Sierra Nevada provided information about the topographical setting that determined every type of evolution. Similar studies should be done in other mountain systems to better determine every single type of geomorphological evolution and infer the role of climate, topography and lithology in their evolution.

(iv) In most cirques, no major changes have occurred on geomorphological dynamics since the HTM. Landscapes have only undergone minor transformations related to periglacial and slope processes. Neoglacial landforms only developed in the highest mountains, such as in the case of the Sierra Nevada, where glaciers developed in cirques with summits exceeding 3300 m. The diversity of cirque landforms depends on local conditions and, above all, on the geomorphological activity of their headwalls. In some cases, the retreat of small Neoglacial ice masses triggered the formation of incipient rock glaciers.

(v) This study highlights that the formation of rock glaciers inside the cirques is triggered by the onset of warm phases that accelerate the retreat of glaciers. Therefore, they are more associated with the paraglacial readjustment following glacial retreat (Ballantyne, 2002) rather than with a periglacial origin driven by very cold temperatures. This explains the rapid stabilization of their fronts once the internal frozen body disappears (Gómez-Ortiz et al., 2014).

Glacial cirques have been shown to be highly climatically sensitive areas where small changes favoured the development or disappearance of glaciers. In addition, the activity of their walls is critical to generate a great variety of landforms, depending on their topographic characteristics. Glacial cirques are, therefore, a valuable source of

- 768 palaeoenvironmental information to better understand the landscape evolution of
- mountain systems since the last deglaciation to recent times.

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1276 Table captions

- Table 1. Field data of sampling sites, topographic shielding factor, sample thickness and
- distance from the headwall.
- Table 2. Analytical data and cosmic-ray exposure (CRE) ages. ¹⁰Be/⁹Be ratios were
- measured at the ASTER AMS facility. The numbers in italics correspond to the internal
- 1281 (analytical) uncertainty at 1σ level.
- Table 3. Location, main topographic characteristics, geomorphologicalal units, and
- 1283 average CRE ages from the cirques studied in the Sierra Nevada compared to the
- Mulhacén cirque.
- Table 4. Location, main topographic characteristics, geomorphologicalal units, and
- average CRE ages of the circues studied in the rest of Iberian Peninsula (outside Sierra
- Nevada). All CRE ages are updated following Oliva et al. (2019).
- Table 5. Location, main topographic characteristics, geomorphologicalal units, and
- average CRE ages of the cirques studied in the Mediterranean region (Iberian Peninsula
- not included). All CRE ages are updated.
- Table 6. Location, main topographic characteristics, geomorphologicalal units, and
- average CRE ages of the cirques studied in the Central European region (outside Sierra
- 1293 Nevada). All CRE ages are updated.
- 1294 Table 7. Location, main topographic characteristics, geomorphologicalal units, and
- average CRE ages of the cirques studied in the British Isles and Iceland. All CRE ages
- are updated.
- Table 8. Location, main topographic characteristics, geomorphologicalal units, and
- average CRE ages of the cirques studied in the Western North America. All CRE ages
- 1299 are updated.

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1301 Figure captions

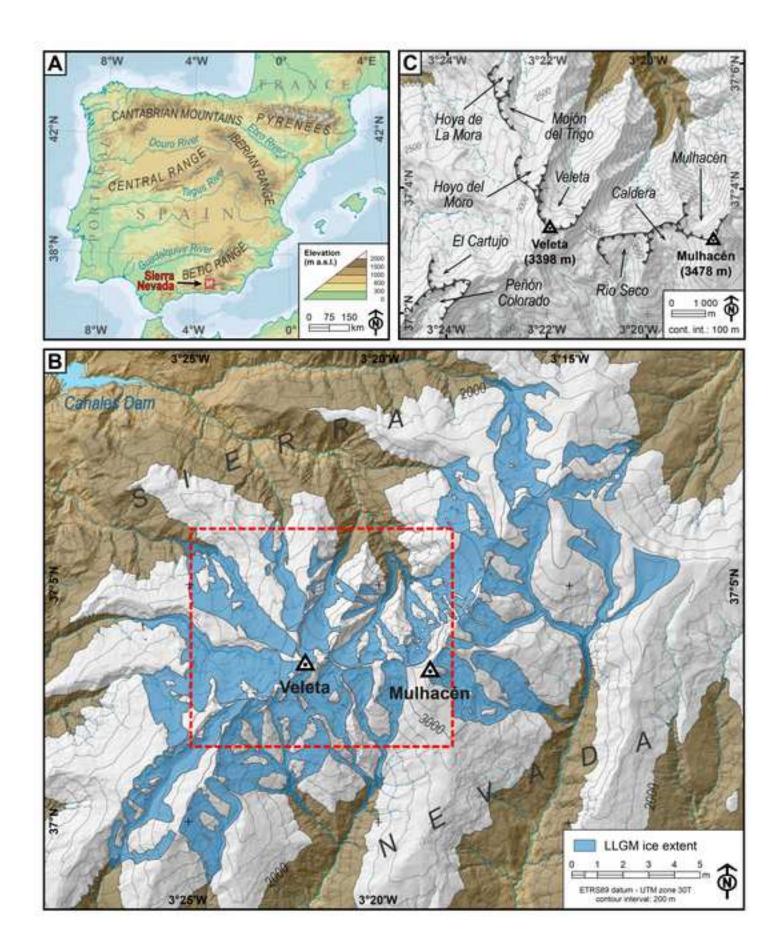
- Figure 1. Location map of the study area. A) Location of the Sierra Nevada in the context
- of the Iberian Peninsula. B) Glacier extent of the Sierra Nevada glaciers during their
- 1304 LLGM. C) Western sector of the Sierra Nevada where the main peaks and the cirques
- cited throughout the text are distributed.
- Figure 2. Mulhacén cirque and location of the CRE sampling sites and ¹⁰Be ages. A)
- 1307 Aerial orthophoto. B) Geomorphological map.
- 1308 Figure 3. Geomorphological transect along the Mulhacén cirque (S-N direction) showing
- 1309 the 10 Be CRE ages (in ka).
- Figure 4. Photos of the Mulhacén cirque floor, with the location of the CRE sampling
- sites and ¹⁰Be ages. Panels A) and B) are oblique views of La Mosca Lake taken from the
- 1312 SE and E, respectively.
- Figure 5. Photos of the Mulhacén cirque floor, location of the CRE sampling sites and
- 1314 ¹⁰Be CRE ages. A) Oblique view of La Mosca Lake from the W. B) Blocky moraine
- where the sample MOSCA-5 (with nuclide inheritance) was collected.
- Figure 6. CRE age correlation with: A) elevation, and B) distance to the Mulhacén
- headwall. In the panel B) we have removed the sample MOSCA-5 as being an outlier
- 1318 (cosmogenic nuclide inheritance). In the panel A) the correlation with altitude is very
- weak as all the samples are located on the flat floor of the cirque.
- Figure 7. Probability density plots of CRE ages for different chronostratigraphic units in
- the A) Mulhacén cirque (data shown in Table 3) and B) in the Sierra Nevada cirques (data
- shown in Table 4). Panels correspond to different stages according to the
- chronostratigraphy: a) Deglaciation of the Mulhacén cirque floor after the Younger
- Dryas. b) Deglaciation of the cirque mouth at the beginning of the Bølling-Allerød
- interstadial. We have included two outlier samples, probably with cosmogenic nuclide
- inheritance (MOSCA-5 and MOSCA-7). The number associated with each averaged CRE
- age refers to the number of CRE ages from each geomorphological unit considered. We
- assume that this average is indicative of the age for each geomorphological phase detected
- in each cirque.
- Figure 8. Synthetic evolution model of the deglaciation phases in the cirques of the Sierra
- Nevada. A) The glaciers filled the cirgues during the OD, until just before the onset of

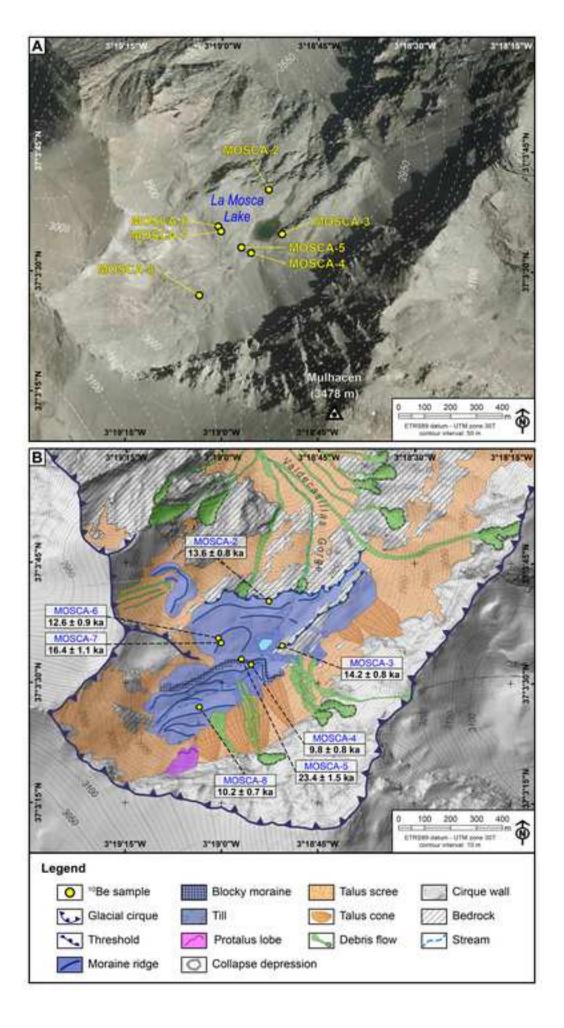
1332 the B-A Interstadial. B) The mouth of most of the cirques was deglaciated during the B-1333 A Interstadial. C) Glaciers might have disappeared from the Sierra Nevada cirques. D) 1334 Most of the cirques hosted new small glaciers during the YD. After this phase, the cirques 1335 evolved differently, according to their topographic characteristics. E.1) Cirques with 1336 summits at 3000-3200 m and floors at >2800 m, with small headwalls (< 200 m of altitude 1337 range) and east-facing produced only small proto-rock glaciers. E.2) Cirques with 1338 summits at 3000-3200 m and floors at > 2800 m, with steep and long headwalls (> 300 m 1339 of altitude range) developed larger rock glacier complexes, especially if they are north facing. E.3) Cirques with summits at < 3000 m and floors at < 2800 m. The shrinking 1340 1341 glaciers left one or more moraines - or even none - depending on the intensity of the 1342 paraglacial readjustment of their surrounding walls. Cirques with summits at > 3300 m and floors at > 2950 m, north-facing and with > 300-m-high walls recorded geomorphic 1343 1344 evidence from glacier changes from the YD to present. This is the case of the Veleta and 1345 Mulhacén cirques, but their evolution was different. E.4) Veleta palaeoglacier formed 1346 only a large polygenic moraine. E.5) Mulhacén palaeoglacier developed a sequence of 1347 moraine ridges. 1348 Figure 9. Probability density plots of CRE ages for different chronostratigraphical units 1349 in A) The Iberian Peninsula cirques (data shown in Table 4); B) The Mediterranean 1350 cirques (outside Iberian Peninsula) (data shown in Table 5); C) The Central European 1351 cirques (data shown in Table 6); D) The British Isles and Iceland cirques; E) The Western 1352 North America cirques (data shown in Table 8); F) Rock glacier front stabilization in the cirques analysed in the North Hemisphere (data shown in Tables 3, 4, 5, 6, 7 and 8). For 1353 1354 Figure 9, B), C), D) and E) the plots are clustered in a) Deglaciation of the cirque floor 1355 after the YD (ages of the stabilization of the youngest moraine and polished bedrock 1356 surfaces); b) Deglaciation of the mouth of the cirque, at the beginning of the B-A 1357 Interstadial (ages of the stabilization of the oldest moraine inside the cirque and polished 1358 bedrock surfaces). For the Panel F) the plots are clustered in a) Rock glacier front 1359 stabilization after Neoglacial advances; b) Rock glacier front stabilization after the Younger Dryas; c) Rock glacier front stabilization after at the beginning of the B-A 1360 1361 Interstadial. The number associated with each averaged CRE age refers to the number of 1362 CRE ages from each geomorphological unit considered. 1363 Figure 10. Probability density plots of CRE ages for different chronostratigraphical units,

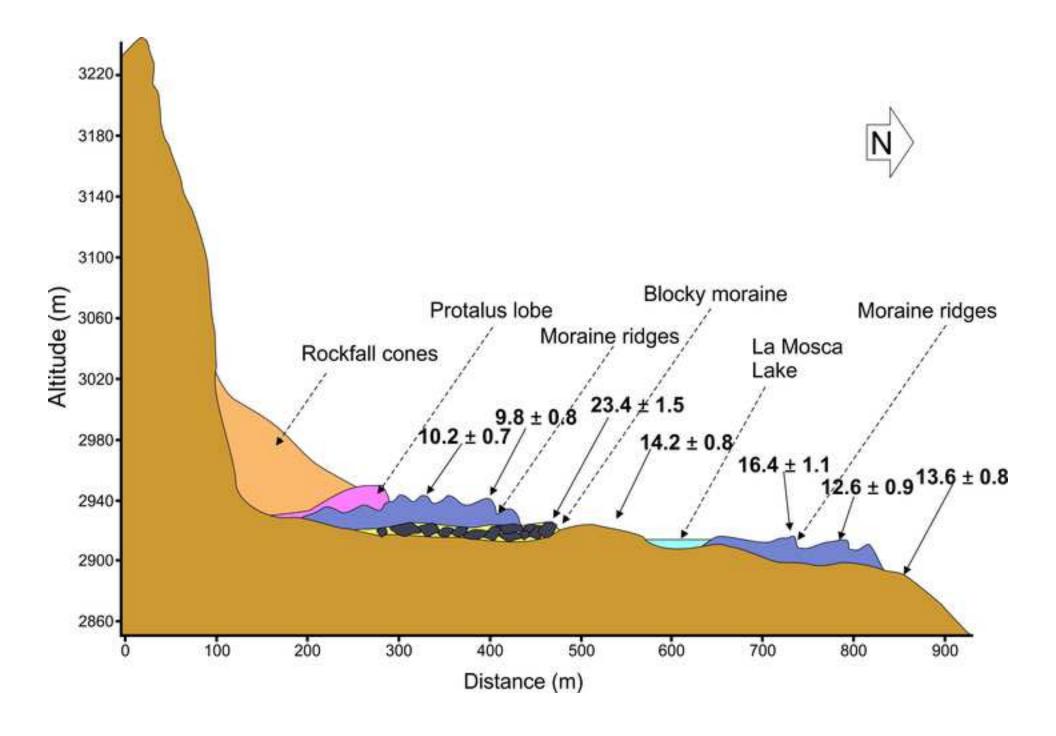
comparing all the results from the six regions. The plots are clustered in a) Rock glacier

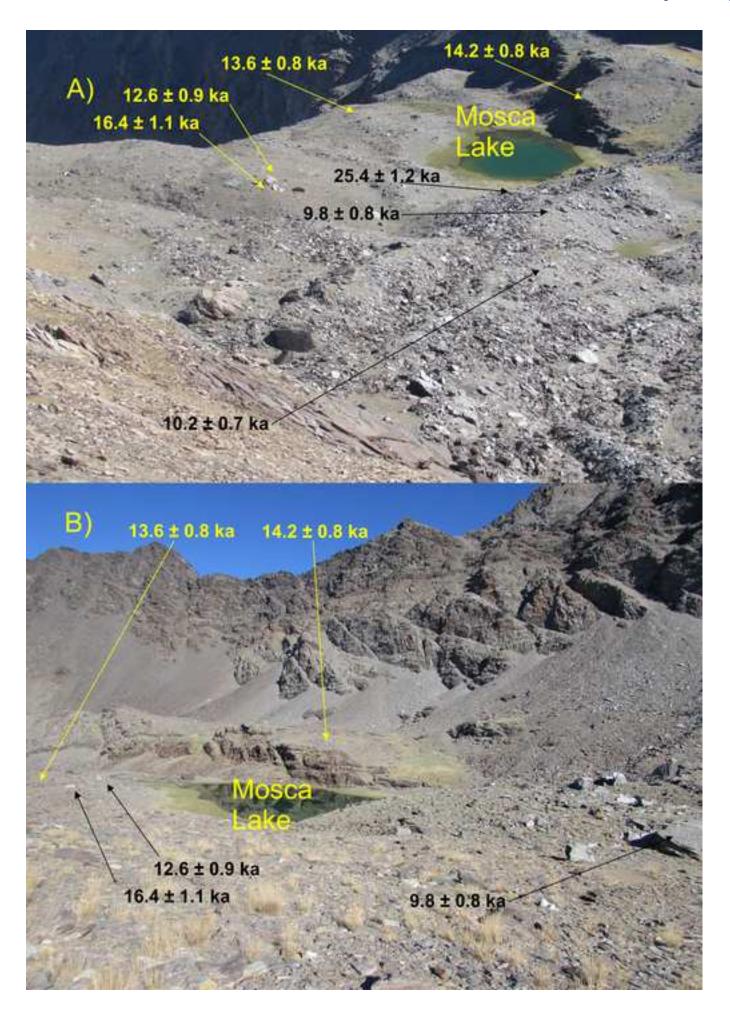
1364

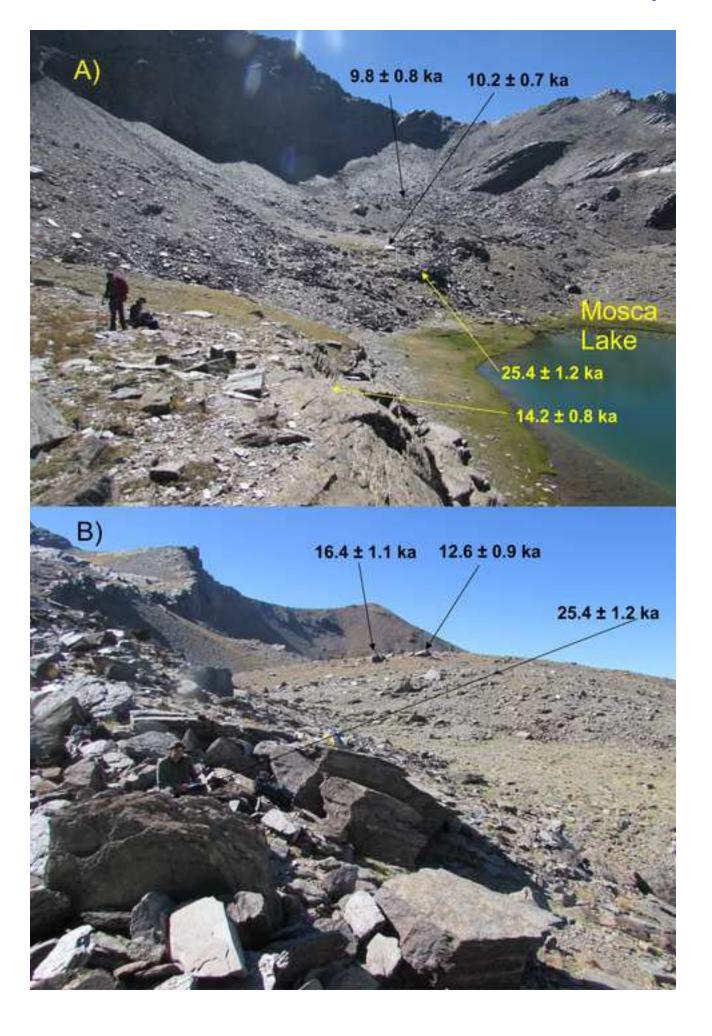
front stabilization after the YD; b) Deglaciation of the cirque floor after the YD (ages of the stabilization of the youngest moraine and polished bedrock surfaces); c) Deglaciation of the mouth of the cirque, at the beginning of the B-A Interstadial (ages of the stabilization of the oldest moraine inside the cirque and polished bedrock surfaces). The number associated with each averaged CRE age refers to the number of cirques with available CRE ages in a given region analysed. The average CRE age of each geomorphological unit is compared with the Greenland ice core chronology (Rasmussen et al., 2014).

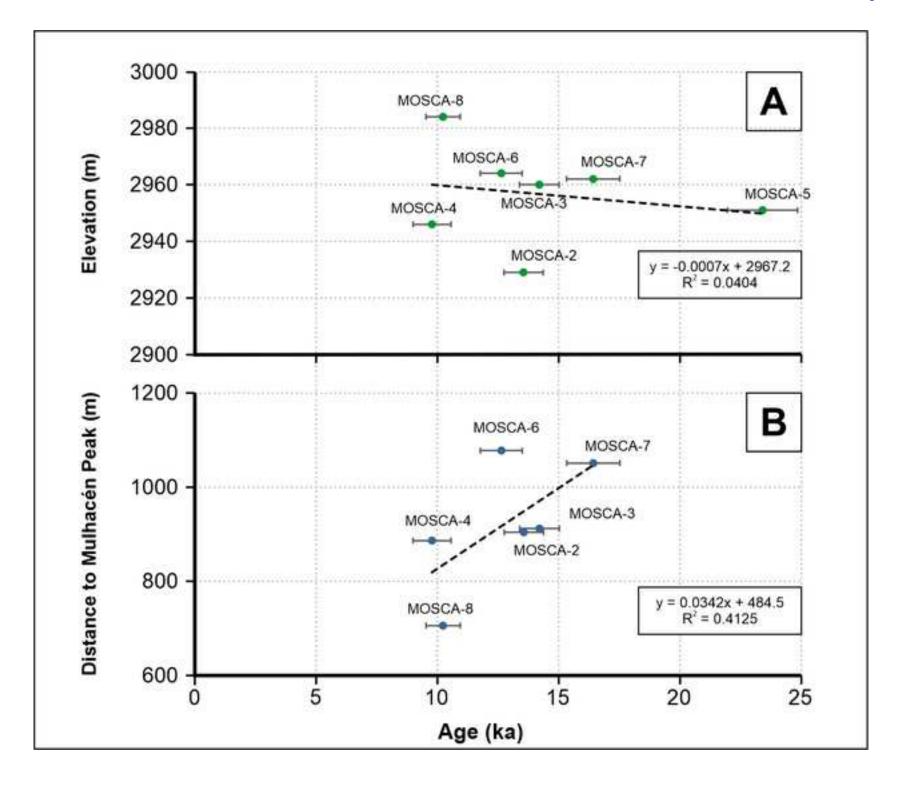


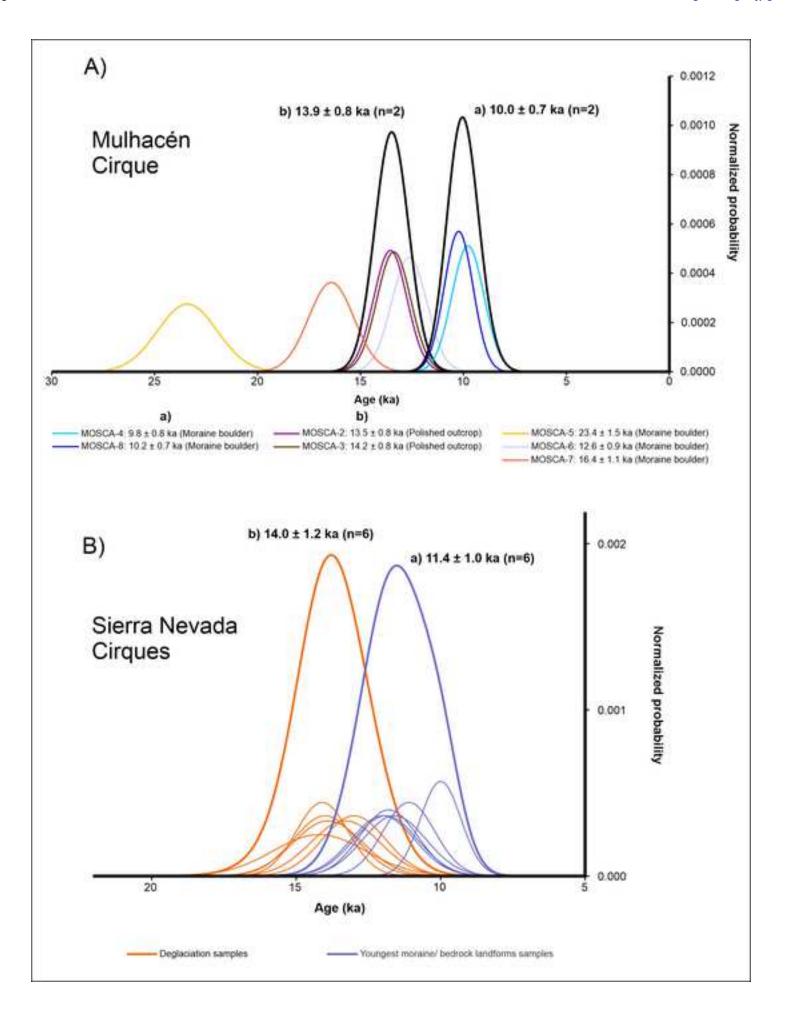


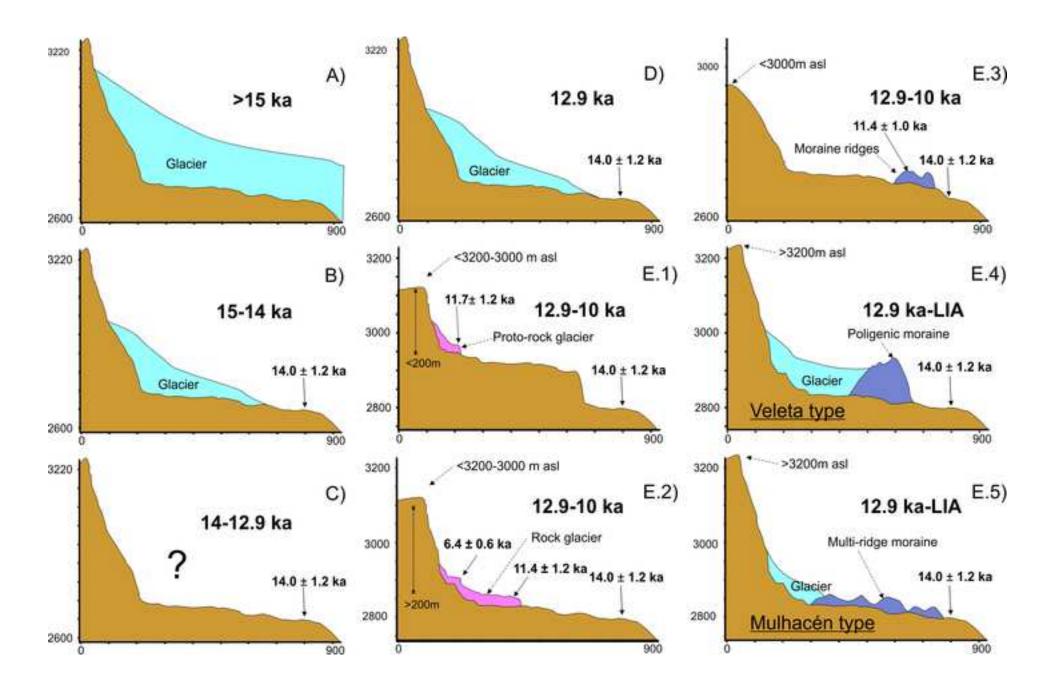


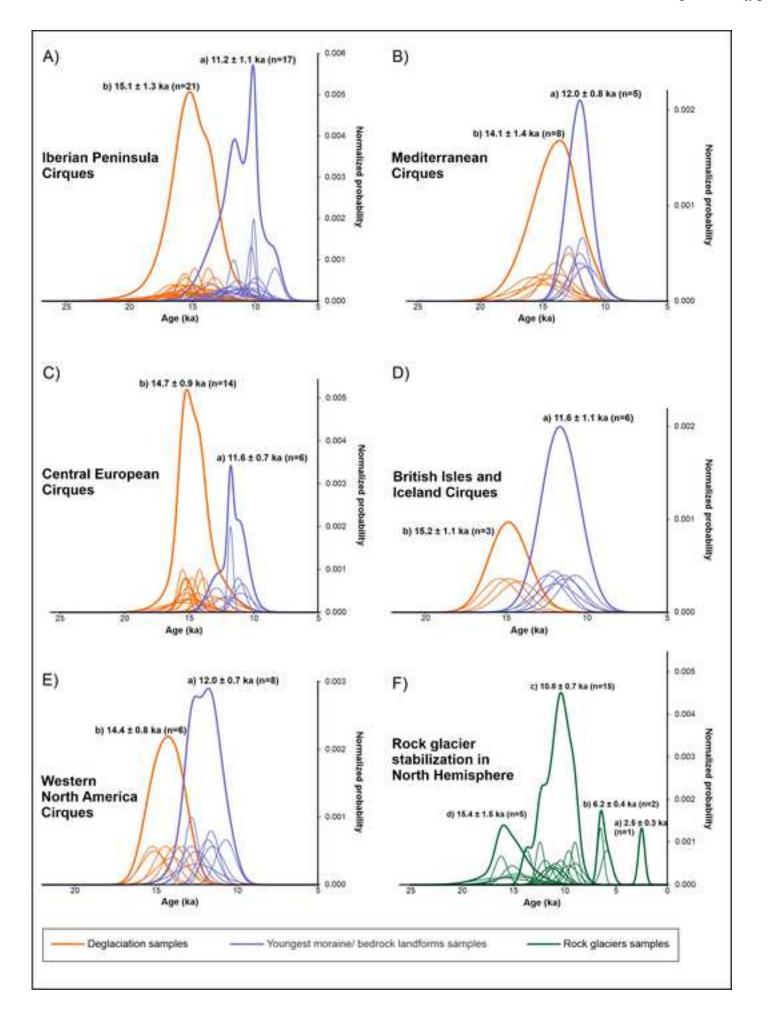


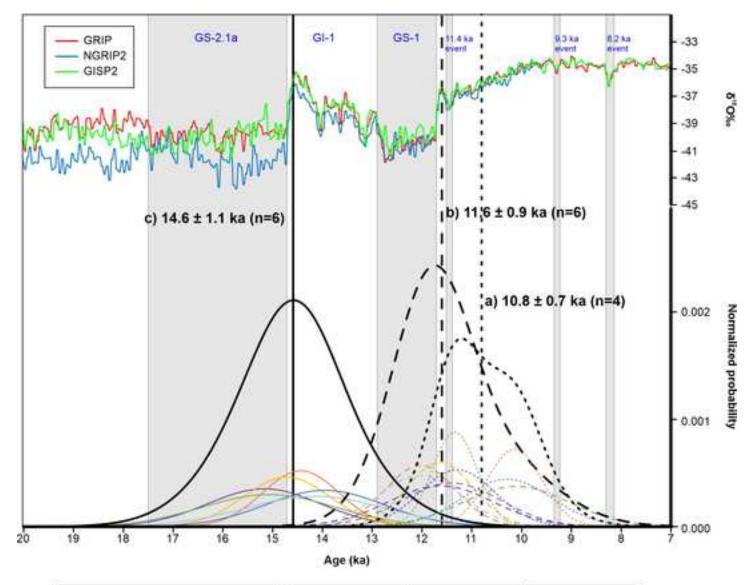












	Deglaciation of the cirque	Youngest moraine/ bedrock — —	Rock glacier landforms
Sierra Nevada			***************
Iberian Peninsula			***************************************
Mediterranean region			
Central European region			
British Isles and Subpolar regions			
Western North America			***************************************

Table 1. Field data of sampling sites, topographic shielding factor, sample thickness and distance from the headwall.

Sample name	Sample type	Latitude (DD)	Longitude (DD)	Elevation (m a.s.l.)	Topographic shielding factor	Thickness (cm)	Distance from the headwall (m): map vs ground length
MOSCA-2	Polished outcrop	37.06117	-3.31463	2929	0.9757	3.0	904 / 1058
MOSCA-3	Polished outcrop	37.05963	-3.31405	2960	0.9545	2.0	911 / 1049
MOSCA-4	Moraine boulder	37.05897	-3.31539	2946	0.9579	2.5	886 / 1034
MOSCA-5	Moraine boulder	37.05916	-3.31581	2951	0.9579	2.5	932 / 1071
MOSCA-6	Moraine boulder	37.05989	-3.31681	2964	0.9572	3.0	1077 / 1194
MOSCA-7	Moraine boulder	37.05972	-3.31669	2962	0.9579	4.0	1050/ 1170
MOSCA-8	Moraine boulder	37.05749	-3.31763	2984	0.9460	4.0	706 / 862

Table 2. Analytical data and cosmic ray exposure (CRE) ages. 10 Be/ 9 Be ratios were measured at the ASTER AMS facility. The numbers in italics correspond to the internal (analytical) uncertainty at 1σ level.

Sample name	Quartz weight (g)	Mass of carrier (9Be mg)	ASTER cathode ID	¹⁰ Be/ ⁹ Be (10 ⁻¹³) corrected of chemical blank	Blank correction (%)	[¹⁰ Be] (10 ⁴ atoms g ⁻¹)	Age (ka)
MOSCA-2	39.7779	155.9	ICYQ	4.7124 ± 0.1461	0.52	37.321 ± 1.163	$13.55 \pm 0.8 (0.4)$
MOSCA-3	29.2882	155.5	ICYR	3.6722 ± 0.1148	0.67	39.418 ± 1.238	$14.20 \pm 0.8(0.4)$
MOSCA-4	8.4979	158.3	ICYS	0.7182 ± 0.0350	3.31	27.047 ± 1.321	$9.78 \pm 0.8 \; (0.5)$
MOSCA-5	5.9343	156.3	ICYT	1.2924 ± 0.0571	1.87	68.786 ± 3.045	$23.41 \pm 1.5 (0.9)$
MOSCA-6	40.4572	153.2	ICYU	4.5547 ± 0.2008	0.54	34.864 ± 1.541	$12.63 \pm 0.9 (0.5)$
MOSCA-7	21.5507	155.5	ICYV	3.1261 ± 0.1501	0.78	45.598 ± 2.194	$16.42 \pm 1.1 (0.7)$
MOSCA-8	40.1676	156.3	ICYW	3.5672 ± 0.1130	0.68	28.061 ± 0.893	$10.23 \pm 0.7 (0.4)$
¹⁰ Be Blank							
MOSCA-BK	-	157.8	ICYP		-	-	-

Table 3. Location, main topographic characteristics, geomorphological units, and average CRE ages from the cirques studied in the Sierra Nevada compared to the Mulhacén cirque.

Cirque/ valley	Location ¹	Cirque elevation range ² (m a.s.l.) and (m)	Length ³ (m)	Aspect ⁴	Deglaciation (ka) ⁵	Youngest moraine/ bedrock (ka) ⁶	Distance from headwall to the youngest moraine (m) ⁷	Rock glacier and protalus lobe landforms (ka) ⁸	Neoglacial landforms (ka/CE) ⁹	References
La Mora/ San Juan	37°5'43"N 3°22'57"W	2609-2350 (259)	600	NE	13.9 ± 1.2 (n=2)	4 moraine ridges 11.6 ± 1.1 (n=2)	400	-		Palacios et al. (2016)
Mojón del Trigo/ San Juan	37°5'22"N 3°22'36"W	2609-2350 (259)	800	Е	14.5 ± 1.2 (n=2)	3 moraine ridges 11.8 ± 1.0 (n=2)	600	-		Palacios et al. (2016)
El Moro/ San Juan	37°4'28"N 3°22'22"W	2925-2790 (135)	700	Е	14.0 ± 1.1 (n=2)	1 ridge 11.1 ± 0.9 (n=1)	500			Palacios et al. (2016)
Cartujo/ Dílar	37°3'9"N 3°22'57"W	3152-2700 (452)	1400	N	14.2 ± 1.6 (n=2)	No data		Rock glacier From 11.4 ± 1.0 to 6.4 ± 0.6		Palacios et al. (2016)
Peñón Colorado/ Lanjarón	37°1'56"N 3°24'23"W	3113-2950 (163)	600	Е	13.0 ± 1.1 (n=4)	3 ridges No data		Proto-rock glacier 10.5 ± 0.9		Palacios et al. (2016)
Río Seco/ Río Seco	37°2'55"N 3°20'39"W	3141-3000 (141)	600	Е		Bedrock 11.9 ± 1.1 (n=1)		Proto-rock glacier 9.0 ± 0.8		Palacios et al. (2016)
Caldereta/ Mulhacén	37°3'5"N 3°19'32"W	3182-3000 (182)	900	E		Bedrock 12.0 ± 1.1 (n=2)		Rock glacier from 13.1 ± 1.2 to 6.3 ± 0.5		Palacios et al. (2016)
Corral del Veleta/ Guarmón	37°3'35"N 3°21'57"W	3396-3000 (396)	500	N	No data	Large poligenic ridge		Rock glacier from 1950 CE	Poligenic LIA ridge From 1355 to 1900 CE	Palacios et al. (2019)

Mulhacén cirque	37° 3'36"N 3°18'57"W	3479-2900 (579)	950	N	14.1 ± 0.9 (n=1) (a)	6 ridges 10.0 ± 0.7 (n=2)(a)	700	Protalus lobe from 1900 CE (b)		(a) Present work (b) Serrano et al. (2018)
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¹Geographic coordinates of the center point of the cirque floor.

²Maximum and minimum elevation of the cirque and elevation range.

³Cirque length. From the summit to the lower sector measured on the map. It is necessary to consider that many cirques are wider than they are long.

⁴Aspect. It is necessary to consider that many circues do not coincide with the head of the valley where they are located, but rather they are housed on one of its slopes with a more appropriate orientation for the accumulation of snow.

⁵Deglaciatoin age of the cirque according to the CRE ages obtained from bedrock outcrops on its bottom.

⁶Number of moraines and their CRE age in ka or CE located into the cirque. When no moraine ages are available, the age corresponds to dated bedrock surfaces.

⁸Rock glaciers and protalus lobe located in the cirque and the stabilization CRE age of its front and its root, when available.

⁹Neoglacial landforms present in the cirque.

Table 4. Location, main topographic characteristics, geomorphological units, and average CRE ages of the cirques studied in the rest of Iberian Peninsula (outside Sierra Nevada). All CRE ages are updated following Oliva et al. (2019).

Cirque/ Valley and rock type	Massif/ range ¹	Elevation (m a.s.l.) and geographic coordinates ²	Aspect ³	Deglaciation of the cirque (ka) ⁴	Youngest: moraine/ bedrock landforms (ka) ⁵	Distance from headwall to the youngest moraine (m)	Rock glacier/debris- covered glacier/other landforms (ka) ⁶	Neoglacial landforms (ka) ⁷	References
Cinco Lagunas/ Pinar (Granite)	Sierra de Gredos/ Central Range	2572 40°15'22"N 5°18'24"W	NNE	16.3 ± 3.3 (n=2)	Bedrock 10.3 ± 1.3 (n=3)		Protalus rampart,	-	Palacios et al. (2011)
Gredos (Granite)	Sierra de Gredos/ Central Range	2591 40°14'56"N 5°17'58"W	N	15.9 ± 1.0 (n=2)				-	Palacios et al. (2012a)
Cuerpo de Hombre (Granite)	Sierra de Gredos/ Central Range	2399 40°17'32"N 5°45'14"W	NW	15.1 ± 1.0 (n=3)	Moraine 12.1 ± 1.2 (n=1)	600	Rock avalanche	-	Carrasco et al. (2015)
Peñalara/ Lozoya (Gneis)	Sierra de Guadarrama/ Central Range	2428 40°50'60"N 3°58'1" W	ESE	15.9 ± 1.0 (n=2)	Bedrock 11.7 ± 04 (n=1)		Proto-rock glacier 16.1 ± 2.5 (n=1)	-	Palacios et al. (2012b)
Hoyo Grande/ Lozoya (Gneis)	Sierra de Guadarrama/ Central Range	2209 40°58'46"N 3°50'51"W	SE	15,6 ± 0,6 (n=2)					Carrasco et al. (2016)
San Lorenzo/ Najerilla (Conglomerat es)	Sierra de la Demanda/ Iberian Range	2271 42°14'33" N 2°58'22" W	SE	16.1 ± 1.0 (n=2)			Debris-covered glacier 9.0 ± 0.4 (n=3)	-	Fernández- Fernández et al. (2017)
Mencilla/ Arlanzón (Conglomerat es)	Sierra de la Demanda/ Iberian Range	1932 42°11'9"N 3°18'44"W	NNE				Debris-covered glacier 6.5 ± 0.3 (n=8)	-	Fernández- Fernández et al. (2017)
Peña Negra/ Mayor (Conglomerat es)	Sierra de la Cebollera/ Iberian Range	2023 m 42°2'39" 2°45'35"W	ENE	16.6 ± 1.0 (n=2)	Bedrock 13.5 ± 1.0 (n=1)		Rock glacier 15.1 ± 0.9 (n=3)	-	García-Ruiz et al. (2020a)

Monasterio (Quartzites)	Montaña Central/Cantabria n Mountains	2019 43°5'1"N 5°20'24"W	N	14.8 ± 0.5 (n=5)			Rock glacier 13.7 ± 0.5 (n=5)	-	Rodríguez- Rodríguez et al. (2017)
Silván (Quartzites)	Montaña Central/Cantabria n Mountains	1935 43°2'31"N 5°21'14"W	ENE				Rock glacier 16.2 ± 0.6 (n=5)	-	Rodríguez- Rodríguez et al. (2016)
Malniu-Guils cirques (Granite)	Cerdanya/ Eastern Pyrenees	2692 42°29'8"N 1°47'31"E	S	14.8 ± 2.0 (n=1)			Rock glacier 14.7 ± 1.6 (n=1)	-	Andrés et al. (2019)
Perafita/ Arànser (Granite)	Cerdanya/ Eastern Pyrenees	2761 42°26'59"N 1°34'53"E	Е	13.7 ± 0.5 (n=1)	Moraine 12.6 ± 1.7 (n=1)		Rock glacier From 14.7 ± 2.1 to 8.6 ± 1.1	-	Andrés et al, (2019)
Bassiès/ Escale (Granite)	Eastern Pyrenees	2763 42°37'0"N 1°57'34"E	N	15.8 ± 1.7 (n=2)	Moraine 11.7± 1.5(n=2)	1200		-	Crest et al. (2017); Tomkins et al. (2018)
Rec de la Grava/ (Granite)	Cerdanya/ Eastern Pyrenees	2763 42°37'7"N 1°57'16"E	S		Bedrock 11.5 ± 2.0 (n=3)	1100		-	Crest et al. (2017); Tomkins et al. (2018)
Picot Ariège valley (Granite)	Eastern Pyrenees	2797 42°40'34"N 1°28'55"E	NW	15.5 ± 0.7 (n=4)	Bedrock 8.4± 0.5 (n=1)		Rock glacier From 7.2 ± 0.4 to 1.4 ± 0.2	-	Jomelli et al. (2020)
Médécourbe/ Ariège valley (Granite)	Eastern Pyrenees	2914 42°36'13"N 1°26'31"E	N	13.2 ± 0.7 (n=2)	Bedrock /moraine 9.9± 0.7 (n=3)			-	Jomelli et al. (2020)
Mulleres* /Noguera Ribagorçana (Granite)	Central Pyrenees	3010 42°37'44"N 0°41'54"E	E	13.7 ± 0.9 (n=3)	Moraine 10.3 ± 0.3 (n=2)			-	Pallàs et al. (2006)
Bessiberri* /Noguera Ribagorçana (Granite)	Central Pyrenees	3017 42°35'41"N 0°49'13"E	W	16.3 ± 2.2 (n=1)	Moraine 10.1 ± 0.2 (n=2)			Rock glacier (active)	Pallàs et al. (2006)

Maladeta/ Esera (Granite)	Central Pyrenees	3323 42°38'47"N 0°38'25" E	N	13.7 ± 1.4 (n=2)	Bedrock /moraine 11.8± 1.1 (n=2)	1200		From 4.5 to LIA moraine	Crest et al. (2017); Tomkins et al. (2018)
Marboré/ Pineta (Sandstones)	Central Pyrenees	3348 42°40'32"N 0°2'4E	N					5.6 ± 0.6 3.6 ± 0.4 1.1 ± 0.1 and LIA moraines	García-Ruiz et al. (2014)
Arrémoulit/ Osseau (Granite)	Central Pyrenees	2821 42°50'5" N 0°19'51"W	N		Bedrock 10.2 ± 0.9 (n=2)			-	Palacios et al. (2017a)
Balaitus/ Aguas Limpias (Granite)	Central Pyrenees	3147 42°50'20"N 0°17'25"W	S	14.6 ± 2.0	Bedrock 11.0 ± 1.4 (n=2)			LIA moraine	Palacios et al. (2017a)
Bachimaña/ Caldarés (Granite)	Central Pyrenees	2728 42°47'54"N 0°13'2"W	S	12.9 ± 1.5 (n=3)				-	Palacios et al. (2017a)
Brazato/ Caldarés (Granite)	Central Pyrenees	2722 42°44'21"N 0°12'33"W	NNW	14.5 ± 1.2	Bedrock 11.4 ± 0.8		Rock glacier 6.1 ± 0.3 (n=2)	-	Palacios et al. (2017a)
Catieras/ Caldarés (Granite)	Central Pyrenees	2564 42°43'4"N 0°11'18"W	W	15.9 ± 1.6	Moraine 10.9 ± 1.2	600	Rock glacier 12.0 ± 1.3 (n=3)	-	Palacios et al. (2017a)
Piniecho/ Caldarés (Granite)	Central Pyrenees	2696 42°43'43"N 0°12'12"W	W	15.6 ± 2.3 (n=4)	Moraine 12.4 ± 1.9	500	Rock glacier 13.0 ± 1.3 (n=4)	-	Palacios et al. (2017a)

¹Name of the cirque within the mountain range.

² Elevation and geographic coordinates of the highest summit of the cirque.

³Main aspect of the cirque.

⁴CRE age showing the deglaciation of the cirque. Note that small glaciers may have remained at the foot of the cirque walls. All ages are related to the bedrock and/or moraines distributed at the mouth of the cirque.

⁵CRE ages indicating the final deglaciation of the cirque. All ages correspond to bedrock or moraines located in the highest parts of the cirque.

⁶CRE ages reporting the stabilization of the rock glacier fronts. If there are two ages, the second one corresponds to the stabilization of the rocks of the rock glacier. Other debris landforms (debris-covered glacier, protalus rampart, rock avalanches, etc.) located in the cirque floor are also included, with their age of stabilization.

⁷Existence of landforms generated by Neoglacial advances with the available CRE ages, if existing.

Table 5. Location, main topographic characteristics, geomorphological units, and average CRE ages of the cirques studied in the Mediterranean region (Iberian Peninsula not included). All CRE ages are updated.

Cirque/ Valley	Massif/ range ¹	Elevation (m a.s.l.) and coordinates ²	Aspect ³	Deglaciation of the cirque (ka) ⁴	Youngest moraine (ka) ⁵	Distance from headwall to the youngest moraine (m)	Rock glacier landforms (ka) ⁶	Neoglacial landforms (ka) ⁷	References
Kisbe/ Sayacak	Mt. Dedegol Mts Taurus Anatolia Pen.	2750 37°40'40"N 31°15'16"E	N		11.5 ± 1.1 (n=2)	2500			Köse et al. (2019)
Karagol	Mt. Dedegol Mts Taurus Anatolia Pen.	2900 37°38'35"N 31°17'25"E	Е	15.8 ± 1.6 (n=1)	12.0 ± 1.0	1800			Köse et al. (2019)
North Çimi	Central Taurus Anatolia Pen.	2411 36°57'21"N 31°59'32"E	Е	13.7 ± 0.8 (n=1)	8.1 ± 0.9* (n=2)	1500			Sarıkaya et al. (2017)*
South Çimi	Central Taurus Anatolia Pen.	2411 36°57'28"N 31°58'55"E	Е	14.9 ± 1.4 (n=5)	7.3 ± 0.6* (n=4)	2500			Sarıkaya et al. (2017)*
Güneycik	Central Taurus Anatolia Pen.	2440 37°1'41"N 31°59'24"E	Е	14.5 ± 1.7 (n=4)				5.9 ± 0.5 (n=4)	Sarıkaya et al. (2017)
Çündüre	Central Taurus Anatolia Pen.	2638 36°58'40"N 32°1'46"E	NE	14.9 ± 2.3 (n=3)					Sarıkaya et al. (2017)
Megala Kazania	Mount Olympus Balkans	2918 40°05'0"N 22°21'0"E	NNE	13.8 ± 1.4 (n=3)	12.0 ± 0.8 (n=2)	800			Styllas et al. (2018)
Throne of Zeus	Mount Olympus Balkans	2918 40°05'30"N 22°21'30"E	NNW	14.0 ± 1.0 (n=4)				2.5 ± 0.3 (n=4) and LIA	Styllas et al. (2018)
Velez Mountain	Dinaric Mts.	1965 43°19'02"N 18°2'6"E	N	14.9 ± 1.1 (n=2)					Zebre et al. (2019)

Irhzer n´Likemt	Akusal Atlas Mts. Morocco	3555 31°7'39"N 7°49'56"W	N	12.9 ± 0.7 (n=2)	1200		Hughes et al. (2018)
Azib Mzik	Akusal Atlas Mts. Morocco	3129 31°6'53"N 7°56'28"W	NE	11.8 ± 0.6 (n=3)	800		Hughes et al. (2018)

¹Name of the cirque within the mountain range.

⁴CRE age showing the deglaciation of the cirque, though small glaciers may have remained at the foot of the cirque walls. All ages are related to the bedrock and/or moraines distributed at the mouth of the cirque.

⁵CRE ages indicating the final deglaciation of the cirque. All ages correspond to bedrock or moraines located in the highest parts of the cirque.

⁶CRE ages reporting the stabilization of the rock glacier fronts. If there are two ages, the second one corresponds to the stabilization of the rock glacier.

⁷Existence of landforms generated by Neoglacial advances with the available CRE ages, if existing.

* These moraines are not included in the statistical analysis of the Fig. 11, as they are considered exceptional. Ages obtained in limestone under heavy erosion.

² Elevation and geographic coordinates of the highest summit of the cirque.

³Main aspect of the cirque.

Table 5. Location, main topographic characteristics, geomorphological units, and CRE ages of the cirques studied in the Central European region (outside Sierra Nevada). All CRE ages are updated.

Cirque/ Valley	Massif/ range ¹	Elevation (m a.s.l.) and coordinates ²	Aspect ³	Deglaciation of the cirque (ka) ⁴	Youngest moraine (ka) ⁵	Distance from headwall to the youngest moraine (m)	Rock glacier landforms (ka) ⁶	Neoglacial landforms (ka) ⁷	References
North Mohoru	Parâng Mts. Romanian Carpathians	2365 45°20'27"N 23°36'25"E	N	13.2 ± 1.1 ka (n=5)	11.8 ± 0.2 ka (n=2)	500			Gheorghiu et al. (2015)
Zanoaga Mare	Parâng Mts. Romanian Carpathians	2278 45°21'4"N 23°31'60"E	N	13.7 ± 1.2 ka (n=4)	11.0 ± 0.9 ka (n=2)	800			Gheorghiu et al. (2015)
Galcescu	Parâng Mts. Romanian Carpathians	2519 45°20'24"N 23°32'25"E	N	15.2 ± 1.3 (n=1)	12.0 ± 1.1 ka (n=3)	800			Gheorghiu et al. (2915)
Spitze Rumer Spitze	Karwendel Mts. North Alps	2454 47°19'13"N 11°25'35"E	N				From 12.4 ± 0.4 to 9.6 ± 0.6 (n=7)		Moran et al. (2016)
Mandlspitze	Karwendel Mts. North Alps	2370 47°19'7"N 11°24'25"E	N				10.9± 0.8 (n=4)		Moran et al. (2016)
Nefcerská	High Tatra Western Carpathians	2428 49°10'14"N 20°1'37"E	N	14.2 ± 0.4 (n=2)			11.3 ± 0.9 (n=3)		Zasadni et al. (2020)
Suchá važecká	High Tatra Western Carpathians	2350 49°9'47"N 20°0'48"E	S	14.9 ± 1.4 (n=2)			10.9 ± 1.0 (n=3)		Zasadni et al. (2020)
Mlynická	High Tatra Western Carpathians	2428 49°10'14"N 20°1'37"E	N				10.4 ± 0.7 (n=3)		Zasadni et al. (2020)
Hincova	High Tatra Western Carpathians	2438 49°11'11"N 20°3'38"E	SE	14.8 ± 0.5 (n=1)			11.9 ± 0.7 (n=3)		Zasadni et al. (2020)

Kasprowy Wierch Bystra	High Tatra Western Carpathians	1987 49°13'55"N 19°58'54"E	N	14.4±0.9 ka (n=4)				Makos et al. (2018)
Sucha Woda	High Tatra Western Carpathians	2235 49°13'6"N 20°1'45"E	N	15.5±0.4 ka (n=5)				Makos et al. (2018)
Biata Woda	High Tatra Western Carpathians	2300 49°13'10"N 20°0'34"E	N	15.2±0.9 ka (n=10)				Makos et al. (2018)
Piec Stawow Polskich valley	High Tatra Western Carpathians	2503 49°10'46"N 20°5'16"E	N	15.1±0.7 ka (n=5)	10.9±0.6 ka (n=5)	700		Makos et al. (2018)
Velká Studená	High Tatra Western Carpathians	2383 49°10'59"N 20°8'40"E	E	15.2 ± 0.5 (n=4)				Engel et al. (2015)
Malá Studená	High Tatra Western Carpathians	2627 49°12'9"N 20°11'46"E	SE	13.9 ± 0.5 (n=1)	11.2 ± 0.5 (n=1)	500		Engel et al. (2015)
Sněžka Úpa valley	Krkonoše Mountains	1602 50°44′10″N 15°44′25″E	S	15.3 ± 0.5 (n=1)				Engel et al. (2017)
Snowy cirque	Krkonoše Mountains	1509 50°46′39″N 15°34′03″E	NE	15.4 ± 1.8 (n=1)	12.9 ± 0.7 (n=1)	400		Engel et al. (2017)

¹Name of the cirque within the mountain range.

² Elevation and geographic coordinates of the highest summit of the cirque.

³Main aspect of the cirque.

⁴CRE age showing the deglaciation of the cirque, though small glaciers may have remained at the foot of the cirque walls. All ages are related to the moraines distributed at the mouth of the cirque.

⁵CRE ages indicating the final deglaciation of the cirque. All ages correspond to moraines located in the highest parts of the cirque.

⁶CRE ages reporting the stabilization of the rock glacier fronts. If there are two ages, the second one corresponds to the stabilization of the rock glacier.

⁷Existence of landforms generated by Neoglacial advances with the available CRE ages, if existing.

Table 7. Location, main topographic characteristics, geomorphological units, and average CRE ages of the cirques studied in the British Isles and Iceland. All CRE ages are updated.

Cirque/ Valley	Massif/ Range ¹	Elevation (m a.s.l.) and coordinates ²	Aspect ³	Deglaciatio n of the cirque (ka) ⁴	Youngest moraine/ bedrock (ka) ⁵	Distance from headwall to the youngest moraine (m)	Rock glacier landforms (ka) ⁶	Neoglacial landforms ⁷	References
Corranabinna Lough	Mayo Western Ireland	670 53°57'40"N 9°41'15"W	NNE	15.4 ± 1.1 (n=4)					Barth et al. (2018)
Glascairns Hill	Donegal NW Ireland	580 54°46'15"N 8° 1'42"W	NE	15.9 ± 1.1 (n=8)	12.0 ± 0.9 (n=4)	850			Barth et al. (2018)
Logaharry Lough	Mayo Western Ireland	620 53°37'38"N 9°42'10"W	NE	14.2 ± 1.2 (n=4)					Barth et al. (2018)
Sruhauncullin more	Mayo Western Ireland	803 53°38'29"N 9°47'31"W	NE		11.4 ± 1.0 (n=4)	950	12.5 ± 1.1 (n=5)		Barth et al. (2018)
Keskadale/ Newlands	Derwent Fells/ Lake District British Isles	734 54°32'30"N 3°14'8"W	N		12.4±1.0 ka (n=2)	600			Hughes et al. (2019)
Ling Comb	Derwent Fells/ Lake District British Isles	737 54°31'52"N 3°18'24"W	E		11.9±1.3 ka (n=4)	300			Hughes et al. (2019)
Fremri- Grjótárdalur West	Tröllaskagi Northern Iceland	1183 65°42'47"N 19°0'6.32"W	N		11.3 ± 1.1 (n=2)	1500	10.8 ± 1.0 (n=2)	Active rock glaciers	Fernández- Fernández et al. (2020)
Fremri- Grjótárdalur East	Tröllaskagi Northern Iceland	1183 65°42'47"N 19°0'6.32"W	N				9.4 ± 1.1 (n=2)	Active rock glaciers	Fernández- Fernández et al. (2020)
Hólajökull	Tröllaskagi Northern Iceland	65°42'7"N 18°57'2"W	N		10.7 ± 1.0 (n=2)	3000		Active debris covered glacier	Fernández- Fernández et al. (2020)

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² Elevation and geographic coordinates of the highest summit of the cirque.

³Main aspect of the cirque.

⁴CRE age showing the deglaciation of the cirque. Note that small glaciers may have remained at the foot of the cirque walls. All ages are related to the bedrock and/or moraines distributed at the mouth of the cirque.

⁵CRE ages indicating the final deglaciation of the cirque. All ages correspond to bedrock or moraines located in the highest parts of the cirque.

⁶CRE ages reporting the stabilization of the rock glacier fronts. If there are two ages, the second one corresponds to the stabilization of the rock glacier.

⁷Existence of landforms generated by Neoglacial advances with the available CRE ages, if existing.

Table 8. Location, main topographic characteristics, geomorphological units, and average CRE ages of the cirques studied in the Western North America. All CRE ages are updated.

Cirque/ valley	Massif/ range ¹	Elevation (m a.s.l.) and coordinates ²	Aspect ³	Deglaciation of the cirque (ka) ⁴	Youngest moraine/ bedrock (ka) ⁵	Distance from headwall to the youngest moraine (m)	Rock glacier landforms (ka) ⁶	Neoglacial landforms ⁷	References
Little Anapurna/ Inspiration Lake	Central Cascades Washington	2660 47°28'8"N 120°48'50"W	NE		10.7 ± 0.6 (n=5)	1200			Marcott et al. (2019)
Solicitude Lake	Teton Range Wyoming	3209 43°48'14"N 110°50'52"W	E		12.8 ± 0.4 (n=3)	1300			Licciardi et al. (2008); Marcott et al. (2019)
Roaring Fork Stough Creek	Wind River Range Wyoming	3720 42°38'2"N 109°1'29"W	NE		12.8 ± 0.7 (n=4)	700	Proto-rock glacier 9.6 ± 0.5 (n=7)	Active rock glacier	Marcott et al. (2019)
Temple lake	Wind River Range Wyoming	3953 42°41'55"N 109°10'15"O	NW	14.0 ± 1.0 (n=7)					Marcott et al. (2019)
Medicine Bow	Rocky Mts, Wyoming	3580 41°20'36"N106° 19'51"W	E	14.5 ± 0.7 (n=6)	11.5 ± 0.7 (n=6)	600	Proto-rock glacier 10.5 ± 0.6 (n=6)		Marcott et al. (2019)
Agassiz/ Blue Lake	Uinta Mts., Utah	3788 40°42'39"N 110°49'30"W	E	14.4 ± 0.7 (n=6)					Marcott et al. (2019)
Dead Horse Lake	Uinta Mts., Utah	3650 40°44'24"N 110°40'48"W	NE	13.4 ± 0.7 (n=6)	11.8 ± 0.6 (n=6)	600	Proto-rock glacier 10.5 ± 0.5 (n=6)		Marcott et al. (2019)
Arapahoe Cirque	Colorado Front Range	4115 40°01'35"N 105°39'01"W	Е		11.6 ± 0.5 (n=6)	1800	Inner moraine 10.0 ± 0.6 (n=6)		Marcott et al. (2019)

Warren Mt. Chicago lakes	Colorado Front Range	4055 39°36'19"N 105°37'59"w	NE	15.2 ± 0.7 (n=5)				Marcott et al. (2019)
Wheeler	South Snake Range Nevada	3982 38°59′10″N 114°18′48″W	NE		12.5 ± 1.3 (n=6)	1800	Active rock glacier	Marcott et al. (2019)
Mount Thompson Boon lakes	Sierra Nevada California	4112 37°8'35"N 118°36'48"W	N		12.5 ± 0.8 (n=5)	2800		Marcott et al. (2019)
Badly/ Katherine	Sangre de Cristo Mts, New Mexico.	3840 35°49'57"N105° 45'28"W	SE	15.1 ± 0.8 (n=10)				Marcott et al. (2019)

¹Name of the cirque within the mountain range.

⁴CRE age showing the deglaciation of the cirque. Note that small glaciers may have remained at the foot of the cirque walls. All ages are related to the bedrock and/or moraines distributed at the mouth of the cirque.

⁵CRE ages indicating the final deglaciation of the cirque. All ages correspond to bedrock or moraines located in the highest parts of the cirque.

⁶CRE ages reporting the stabilization of the rock glacier fronts. If there are two ages, the second one corresponds to the stabilization of the rock glacier.

⁷Existence of landforms generated by Neoglacial advances with, if existing, the available CRE ages.

² Elevation and geographic coordinates of the highest summit of the cirque.

³Main aspect of the cirque.