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## To cite this version:

N. Carjan, M. Rizea. Acceleration induced neutron emission in heavy nuclei. Phys.Lett.B, 2020, 807, pp.135610. 10.1016/j.physletb.2020.135610 . hal-02911954

HAL Id: hal-02911954
https://hal.science/hal-02911954
Submitted on 18 Jul 2022

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# Acceleration induced neutron emission in heavy nuclei 

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#### Abstract

The effect of the acceleration of a nucleus on the neutron states is studied in the frame of the independent-particle nuclear shell model. For this we solve numerically the timedependent Schrodinger equation, with a moving mean-field of Woods-Saxon type. The time evolution of a neutron states at the Fermi level is calculated for ${ }^{236} U$ and acceleration parameter $\mathrm{A}=0.5$ (in $10^{44}\left[\mathrm{fm} / \mathrm{sec}^{2}\right]$ ). It is the acceleration during the Coulomb repulsion of two ${ }^{236} U$ nuclei when they are 20 fm apart. We keep this acceleration constant for $10^{-21} \mathrm{sec}$ before we switch it off $(\mathrm{A}=0)$ and follow the wave packet for another $10^{-21} \mathrm{sec}$. During the acceleration, the wave function oscillates with increasing amplitude until it escapes, mainly in the direction opposite to the motion of the nucleus. The mean value of its energy (in the nuclear system) increases from -4.80 MeV to -3.15 MeV and $12 \%$ of the wave packet leave the nucleus. During the uniform motion, the wave packet continues to oscillate and to escape at a lower rate: an extra $2 \%$. We repeated the calculations for two neighbouring states and found the emission rate to depend strongly on the position of the neutron state with respect to the Fermi energy. Finally, the effect of the nuclear deformation on the acceleration induced neutron emission is studied. In this case the period of oscillation is larger and the amplitude smaller. The angular distribution with respect to the direction of motion is more complex : it has, in the nuclear system, an intense component almost perpendicular to the deformation axis.


## Keywords:

neutron emission, acceleration, moving potential, time-dependent Schrödinger equation

What if, like a fully-filled water tank, a nucleus will spill its less bound nucleons when accelerated? Most probable this could happen to neutrons since, contrary to protons, clei.

July 1, 2020

Acceleration of heavy nuclei is an integral part of nuclear fission and reactions induced by heavy-ions. If neutrons are emitted during the acceleration of the fission fragments or when the projectile moves closer to or away from the target, one has to take into account this emission in the analysis of the data.

To give a first answer to this captivating question we choose a simple framework: the independent particle shell model [1] and a simple scenario borrowed from the operating mode of a linear accelerator [2]. We therefore solve the time-dependent Schrödinger equation with a moving mean-field of WoodsSaxon type and study the change in the neutron eigenstates during $10^{-21} \mathrm{sec}$ of constant acceleration followed by $10^{-21} \mathrm{sec}$ of constant velocity.

As in the independent particle shell model, we look at one neutron at a time. It moves in a mean field created by all other nucleons. At the same time it contributes to the potential seen by the others. If the nucleons are displaced, the nuclear density is displaced and, by selfconsistency, the potential follows. As for the shell model, our approach is expected to work for weakly bound nucleons.

The motion of a quantum particle in a moving one-dimensional potential well has been already studied in the field of control systems $[3,4,5]$.

In our case, we consider a neutron in a moving nuclear potential that has axial symmetry. It is represented by a wave function solution of the Schrödinger equation

$$
\begin{equation*}
i \hbar \frac{\partial \Theta(\rho, z, t)}{\partial t}=\mathcal{H}(\rho, z, t) \Theta(\rho, z, t) \tag{1}
\end{equation*}
$$

where $\mathcal{H}$ is the single-particle Hamiltonian.

$$
\begin{aligned}
& \mathcal{H}=-\frac{\hbar^{2}}{2 m} {\left[\frac{1}{\rho} \frac{\partial}{\partial \rho}+\frac{\partial^{2}}{\partial \rho^{2}}+\frac{\partial^{2}}{\partial z^{2}}-\frac{\Lambda^{2}}{\rho^{2}}\right] } \\
&+V(\rho, z-\alpha(t)) .
\end{aligned}
$$

To cover nuclei from spherical to strongly deformed, the formalism is developed in cylindrical coordinates. $\alpha(t)$ describes the displacement of the potential in time along the $z$ axis. $\Lambda$ is the projection of the orbital angular momentum on the symmetry axis. For simplicity the spin-orbit term is neglected.

By the Liouville transformation $\Phi=\rho^{1 / 2} \Theta$, the first derivative with respect to $\rho$ from $\mathcal{H}$ is removed, resulting a simplified Hamiltonian $H$ of the form:

$$
\begin{gathered}
H=-\frac{\hbar^{2}}{2 m}\left[\frac{\partial^{2}}{\partial \rho^{2}}+\frac{\partial^{2}}{\partial z^{2}}-\frac{\Lambda^{2}-1 / 4}{\rho^{2}}\right] \\
+V(\rho, z-\alpha(t))
\end{gathered}
$$

To solve the corresponding equation, a transformation of both the variable and the function from the non-inertial (laboratory) to the inertial (nuclear) system is convenient. It avoids an interpolation of the potential between the grid points at each time step.

We will explain this transformation on the 1-D TDSE:
$i \hbar \frac{\partial \Phi(t, z)}{\partial t}=-\frac{\hbar^{2}}{2 m} \frac{\partial^{2} \Phi(t, z)}{\partial z^{2}}+V(z-\alpha(t)) \Phi(t, z)$
We go in the nuclear frame by the following changes of the variable $z \rightarrow q$ and of the function $\Phi \rightarrow \Psi[6]$ :

$$
\begin{equation*}
q=z-\alpha(t), \quad \Phi(t, z)=\exp (u) \Psi(t, q) \tag{3}
\end{equation*}
$$

where

$$
u=i b\left(z \dot{\alpha}-\alpha \dot{\alpha}+\frac{1}{2} \int_{0}^{t} \dot{\alpha}^{2}\left(t^{\prime}\right) d t^{\prime}\right) .
$$

By taking $b=\frac{m}{\hbar}$, it can be shown that Eq.(2) will be transformed in

$$
\begin{align*}
i \hbar \frac{\partial \Psi(t, q)}{\partial t}= & -\frac{\hbar^{2}}{2 m} \frac{\partial^{2} \Psi(t, q)}{\partial q^{2}}+V(q) \Psi(t, q) \\
& +m q \ddot{\alpha}(t) \Psi(t, q) \tag{4}
\end{align*}
$$

To eliminate the linear term in $q$ (which tends to $\infty$ as $q \rightarrow \infty$ ), a further function transformation is performed [7]

$$
\begin{equation*}
\Psi(t, q)=\exp \left(-i \frac{\lambda}{\hbar}\right) \chi(t, q) \tag{5}
\end{equation*}
$$

$$
\begin{equation*}
i \hbar \frac{\partial \chi(t, q)}{\partial t}=-\frac{\hbar^{2}}{2 m}\left(\frac{\partial^{2} \chi(t, q)}{\partial q^{2}}+\frac{2}{\hbar i} B t \frac{\partial \chi(t, q)}{\partial q}-\frac{1}{\hbar^{2}} B^{2} t^{2} \chi(t, q)\right)+V(q) \chi(t, q) \tag{6}
\end{equation*}
$$

The advantage of these transformations is that they lead to equations in which the potential depends on a time-independent variable, the dependence on $\alpha(t)$ being transferred in the coefficients of Eqs.(4),(6).

Since the potential moves along the z-axis, the inclusion of the second coordinate $\rho$ is straightforward. We therefore work with the variables $\rho$ and $q$ on a finite numerical grid: $[0,160] \times[-256,256], \Delta \rho=\Delta q=1 / 8$ fm . For the time evolution we use the step $\Delta t=1 / 128 \times 10^{-22} \mathrm{sec}$. The equations (4) and (6) are solved numerically by the CrankNicolson method. One obtains a linear system which is solved by a routine based on the Strong Implicit Procedure [8]. As initial so-
with

$$
\lambda(t, q)=q m \int_{0}^{t} \ddot{\alpha}\left(t^{\prime}\right) d t^{\prime}=q \beta(t) .
$$

In our particular case $\alpha(t)=\frac{1}{2} A t^{2}$.

$$
\begin{gathered}
\dot{\alpha}=A t, \ddot{\alpha}=A, \lambda=B q t, B=m A \\
u=i b\left(z A t-\frac{1}{3} A^{2} t^{3}\right)
\end{gathered}
$$

and the equation for $\chi$ assumes the form:
lutions (at $t=0$ ) we consider eigenfunctions of the original Hamiltonian.

The propagation in time is done in two steps. We consider a quadratic $\alpha(\mathrm{t})$ and use Eq.(6) for $10^{-21}$ sec. Then we consider a linear $\alpha(\mathrm{t})$ and use Eq.(4) without the $\ddot{\alpha}$ term for another $10^{-21} \mathrm{sec}$. At any time, by performing the inverse tranformations, we can retrieve the solution $\Phi(\rho, z, t)$ of the original equation. For the 1st step we choose the constant acceleration $\mathrm{A}=0.5\left[10^{44} \mathrm{fm} / \mathrm{sec}^{2}\right]$ resulting a constant velocity $v=5\left[10^{22} \mathrm{fm} / \mathrm{sec}\right]=0.167 \mathrm{c}$ for the 2 nd step. c is the speed of light. As first example, we take spherical ${ }^{236} U$ and study 3 neutron states with $\Lambda=0$ around the Fermi level. The


Figure 1: (Color online) Time evolution of $\left|\Phi_{14}^{0}\right|^{2}$ during (left) and after (right) acceleration for spherical ${ }^{236} U . N_{i n}$ is the norm inside the dotted circle defined by $V_{0} / 100$. T is the time in $10^{-22} \mathrm{sec}$.
parameters of the Woods-Saxon potential are fitted to single-particle and single-hole states in the ${ }^{208} \mathrm{~Pb}$ region [9].

Fig. 1 shows the result for the 14th eigenstate laying at -4.8 MeV . For $\mathrm{T}>0$, the wave packet exhibits a changing asymmetry with respect to the $\rho$ axis indicative of an oscillation. In fact, we are dealing with an oscillation over the usual vibration of a quasistationary state. A similar vibration causes the barrier asaults in the Gamow picture of $\alpha$ decay [10]. The presence of the neutron inside the nucleus, measured by $N_{i n}$, decreases from 1.00 (at $\mathrm{T}=0$ ) to 0.88 (at $\mathrm{T}=10$ ). A neutron,
initially occupying this state, has therefore $12 \%$ chance to be emitted during the acceleration phase. For $\mathrm{T}>10$ the oscillatory motion continues. At $\mathrm{T}=20 N_{\text {in }}$ reaches an even lower value ( 0.86 ) because at $\mathrm{T}=10$ the wave packet has short unbound tails which will inevitably leave the nucleus.

The most probable direction of emission is $180^{\circ}$ with respect to the displacement of the nucleus as in classical mechanics. There is however a weaker branch at $\approx 150^{\circ}$ which has a quantum origin. It is known that the tunneling path of a metastable state is mainly dictated by its quantum numbers [11, 12]. In
other words the emission preserves the spatial distribution of the respective state. Hence we expect a 2 nd peak at $125^{\circ}$ in the nuclear system which translates into an emission at $150^{\circ}$ in the laboratory system.

Fig. 2 shows the time dependence of a neutron states with energy ( -3.4 MeV ) above the Fermi level. $49 \%$ of this wave function is emitted, in the same interval of time, through strong oscillations. Predictably, the emission starts earlier and is more intense. As for $\Phi_{14}^{0}$, there are also two directions of emission: one intense at $180^{\circ}$ and one weak at an angle between $180^{\circ}$ and $90^{\circ}$.

A wave function with energy lower than the Fermi level was found to oscillate during both regimes: quadratic and linear $\alpha(t)$. It doesn't however succeed to escape: its presence inside the nucleus doesn't change.

The percentage of the wave packet that leaves the nucleus, $\mathrm{N}_{\text {in }}(0)-\mathrm{N}_{\text {in }}(\mathrm{T})$, is a measure of the neutron emission probability. We
have plotted this quantity as a function of time for $\Phi_{14}^{0}$ and $\Phi_{15}^{0}$ in Fig.3. As expected, neutrons that are less bound have higher emission probability. One can observe the effect of the oscillation of the wave function in the mean-field and the plateau at large times (the emission is gradually reduced when the acceleration is turned off).

Summing up, we start with an eigenstate that, due to acceleration, becomes wave packet, i.e., it occupies with time dependent probabilities neighbouring states (oscillation) including some in the continuum (emission).

For a better understanding of the physics involved, we divide the total energy in the laboratory system in significant terms:

$$
\begin{aligned}
\langle\Phi| H|\Phi\rangle= & -\frac{\hbar^{2}}{2 m} \iint\left(\Psi^{*} \frac{\partial^{2} \Psi}{\partial \rho^{2}}+\Psi^{*} \frac{\partial^{2} \Psi}{\partial q^{2}}\right) d \rho d q \\
& +\frac{\hbar^{2}}{2 m} b^{2} \dot{\alpha}^{2} \iint|\Psi|^{2} d \rho d q
\end{aligned}
$$

$$
\begin{equation*}
-\frac{\hbar^{2}}{2 m} 2 i b \dot{\alpha} \iint \Psi^{*} \frac{\partial \Psi}{\partial q} d \rho d q+\frac{\hbar^{2}}{2 m}\left(\Lambda^{2}-1 / 4\right) \iint \frac{1}{\rho^{2}}|\Psi|^{2} d \rho d q+\iint V(\rho, q)|\Psi|^{2} d \rho d q \tag{7}
\end{equation*}
$$

The 1st term is the average kinetic energy in the nuclear frame. The 2nd term reduces to $m \dot{\alpha}^{2} / 2$; it is the extra kinetic energy due to the velocity of the potential. The 3rd term reduces to $\dot{\alpha}<p>$ where $<p>$ is the average momentum in the nuclear frame. It represents the variation of the neutron energy
due to the interaction with the moving wall of the potential; equivalent to the "one-body" dissipation [13]. The last two terms are the average centrifugal and nuclear potentials respectively.

The sum $\mathrm{E} 1+\mathrm{E} 4+\mathrm{E} 5$ i.e., the average neutron energy in the nuclear frame, slightly in-


Figure 2: (Color online) The same as in Fig. 1 but for $\Phi_{15}^{0}$.
creases up to $\mathrm{T}=10$ and stays constant afterwards. After $\mathrm{T}=10$ the oscillations of E are due only to E3 as clearly seen in Fig.4.

In the absence of correlations the oscillations last long time. In reality, with increasing excitation energy, the time between collisions decreases and the oscillations are slowly damped due to the onset of correlations.

Finally we study the effect of nuclear deformation on the neutron emission due to acceleration of the nucleus along the deformation axis. For this we describe the nuclear shape by a pure Cassini oval with a prolate deformation parameter $\epsilon=0.52$. It represents
the shape isomer of ${ }^{236} U$ [14]. The results for two neutron states, one below and one above the Fermi level, are presented in Figs. 5 and 6 respectively. The emission probability is again larger in the latter case: $N_{\text {in }}=$ 0.77 vs 0.39 . One can also notice that the emission at angles smaller than $180^{\circ}$ to the $z$-axis is much intenser than in the previous cases. It is because a quasi-stationary state in a deformed nucleus tunnels most probably pependicularly to the deformation axis [15], irrespective if the barrier is lower or higher along this direction. It is what happens here: the emission occurs at $90^{\circ}$ in the


Figure 3: (Color online) The neutron emission probability as a function of time for $\Phi_{14}^{0}$ and $\Phi_{15}^{0}$.


Figure 4: (Color online) The neutron energy in the laboratory system (red), in the nuclear system (blue) and the energy due to the interaction with the moving potential (green).
nuclear system. Due to the displacement of the nucleus however, in the laboratory system, the emission occurs at slightly larger angles $\left(\approx 110^{\circ}\right)$. Hence there is a qualitative difference between spherical and strongly deformed nuclei: in the latter case, the direction imposed by quantum mechanics competes in intensity with the classically expected direction of emission.

Since in the present calculation the potential moves in the z-direction (Eq. (2)), one can study nuclei with the deformation axes aligned with the velocity vector (axial symmetry). In heavy-ion reactions however a deformed projectile approches a target with different orientations of the deformation axes. For a quantitative evaluation, an average over all orientations has to be performed.

This type of neutron emission doesn't occur only during the acceleration of a nucleus but also during the slowing down $(\mathrm{A}<0)$ of a projectile when it approaches a target. Among possible effects, an increase of the neutron-transfer cross section is expected.

The chosen value of A (0.5) is two times larger than the acceleration during the Coulomb repulsion of two equal fission fragments from ${ }^{236} \mathrm{U}$ separated by $D_{c m}=20 \mathrm{fm}$ ( $\mathrm{A}=0.249$ ) but comparable with that for the collision of two ${ }^{236} \mathrm{U}$ nuclei at the same distance of approach $(A=0.498)$. In conclusion, the excitation of the fission fragments during their acceleration (separation) is improbable. On the other hand, a heavy projectile could become excited (even emit neutrons) when it approaches a heavy target or moves away from it. For a quantitative answer a dedicated study is necessary. The acceleration


Figure 5: (Color online) Time evolution of $\left|\Phi_{17}^{0}\right|^{2}$ during (left) and after (right) acceleration for the shape isomer (deformation parameter $\epsilon=0.52$ ) of ${ }^{236} U . N_{i n}$ is the norm inside the dotted ellipse defined by $V_{0} / 100$. T is the time in $10^{-22} \mathrm{sec}$.
produced in Coulomb interactions is not constant, it depends on $D_{c m}$. A high level can be however maintained for $10^{-21} \mathrm{sec}$, a short time at the scale of the above mentioned processes ( $10^{-20} \mathrm{sec}$ ).

A heavy-ion of mass $m$ and charge $q$ in an electric field $E$ is accelerated along the field lines; $a=q E / m=0.376 \times 10^{23} \times E\left[\mathrm{fm} / \mathrm{s}^{2}\right]$ with $E$ in $[\mathrm{V} / \mathrm{m}]$. To attain an acceleration comparable with the one used above requires $E=10^{21}[\mathrm{~V} / \mathrm{m}]$ i.e., a much too large value.

A comment on the peculiarity of accelerating a heavy-ion is appropriate. Altough all
forces that produce the acceleration of heavy ions (such as the Coulomb force or the force $q E$ exerted by an electric field $E$ on a charge q) act only on protons, the neutrons follow. To decouple the two fields one needs extreme conditions. For instance to excite an isovector giant dipole resonance, a relativistic projectile is necessary [16].

The process of acceleration induced neutron emission has similarities with the realease of neutrons at scission i.e., during the last stage of nuclear fission $[17,18]$. They are both due to a fast change of the poten-


Figure 6: (Color online) The same as in Fig. 5 but for $\left|\Phi_{19}^{0}\right|^{2}$.
tial in which they move that transforms each neutron state into a wave packet with components in the continuum. In the case of scission neutrons however, it is the shape of the potential and not its position that changes.

A more quantitative evaluation of this process requires a time-dependent, selfconsistent microscopic approach [19]. At present it is a difficult task considering the large numerical grid that is necessary.

We are grateful to Gurgen Ter-Akopian and Jorgen Randrup for their interest in this work and illuminating discussions.

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