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Unraveling the exhumation history of high-pressure ophiolites using 1 magnetite (U-Th-Sm)/He thermochronometry 2 3 Stéphane Schwartz^{1,2,*}, Cécile Gautheron², Richard A. Ketcham³, Fabrice Brunet¹, Marianna 4 5 Corre¹, Arnaud Agranier⁴, Rosella Pinna-Jamme², Frédéric Haurine², Gael Monvoin², Nicolas 6 Riel⁵ 7 ¹ ISTerre, Univ. Grenoble Alpes, Univ. Savoie Mont Blanc, CNRS, IRD, IFSTTAR, 38041 Grenoble, 8 9 France 10 ² GEOPS, Univ. Paris-Sud, CNRS, Université Paris-Saclay, 91405 Orsay, France 11 ³ Jackson School of Geosciences, The University of Texas at Austin, Austin, TX, USA 12 ⁴ Laboratoire Domaines Océaniques, CNRS, IUEM, Université de Bretagne Occidentale, 29280 13 Plouzané, France ⁵ Institute of Geosciences Johannes Gutenberg, University Mainz, D-55128 Mainz, Germany 14 15 16 * corresponding author: stephane.schwartz@univ-grenoble-alpes.fr 17 18 19 20 To be submitted to EPSL 21 22 Keywords: Thermochronometry; (U-Th-Sm)/He; Magnetite; Modeling; High-pressure 23 24 ophiolites; Exhumation

25 Abstract

26 Magnetite is a ubiquitous oxide in ultramafic and mafic rocks, which is present in a large 27 range of geological and tectonic settings. In the case of high-pressure ultramafic rocks, 28 exhumation timing is commonly constrained by geochronometers in nearby lithologies. The 29 development of the magnetite (U-Th-Sm)/He method (MgHe) has opened new perspectives 30 to refine the exhumation history of such rocks. However no thermal history has been deduced using MgHe data so far. Here we applied MgHe dating to magnetite from an HP-LT 31 32 alpine ophiolite body (Rocher Blanc) from the Schistes lustrés units (Western Alps, France) 33 where part the P-T-t history is constrained to further refine the thermal history and 34 demonstrate validity of the method. Textural and geochemical characterizations of the 35 magnetite grains reveal two crystallization environments leading to distinct grain morphologies, euhedral and pseudo-euhedral, with different amounts of mineral inclusions 36 37 (titanite, chlorite and augite-aegirine). Both magnetite types crystallized during the same 38 retrograde path at T> 250°C, under hydrothermal conditions shown by Si and V oscillatory 39 zoning and low Ti content. MgHe ages derived from euhedral magnetite crystals range from 14.8±2.1 to 20.9±3.0 Ma, between zircon and apatite fission track (ZFT) and (AFT) ages 40 41 deduced from the literature. Pseudo-euhedral grains display older ages interpreted as 42 resulting from the implantation of He from U-rich neighboring minerals such as titanite. For 43 the first time, MgHe data are inverted along with AFT and ZFT data to refine the thermal 44 history of an ophiolite related to their exhumation. The MgHe data suggest that the Rocher 45 Blanc sample has undergone an exhumation in ductile conditions related to subduction 46 dynamics and underplating processes, followed by an exhumation reflecting denudation by 47 erosion under brittle conditions. This exhumation regime change is found to take place at 48 ~20 Ma, as opposed to a less constrained ~26 Ma cooling as would be inferred from the ZFT

and AFT ages alone. Considering published thermochronometric data along the Western Alps, we propose that the exhumation is controlled by the upward indentation of the Apulian mantle producing regional tilting and associated erosion. With the present Schistes lustrés ophiolite example, we demonstrate that the MgHe method can be used as a lowtemperature thermochronometer (closure temperature in the 200-250°C range) that aids in unraveling the cooling history associated with exhumation of high-pressure ultramafic and mafic rocks at the brittle to ductile transition.

56

57 1. Introduction

58 Ultramafic and mafic rocks are widely distributed on the Earth's surface and are associated 59 with a large range of geological and geodynamic settings, from oceanic accretion to orogenic 60 zones (e.g., Guillot et al., 2015). In contrast to their abundance, mafic and ultramafic rocks 61 only yield a few minerals suitable for geo-thermochronometry such as U-Th-bearing 62 accessory minerals (zircon, titanite or apatite). However, recent advances in the analysis of 63 radiogenic isotopes have extended the application of (U-Th)/He geo-thermochronology to minerals incorporating U and Th at the ppb level such as magnetite (Blackburn et al., 2007; 64 65 Cooperdock and Stockli, 2016) or spinel (Cooperdock and Stockli, 2018). Magnetite (U-66 Th)/He geo-thermochronometry, abbreviated MgHe below, opens new perspectives with 67 respect to the use of basic and ultrabasic rocks for dating geodynamic processes.

Indeed, magnetite (Fe²⁺Fe³⁺₂O₄) is ubiquitous in these rock types where it can either crystallize from the parent magma or precipitate later in the rock history from hydrothermal fluids at lower temperature. The MgHe method is particularly attractive for dating geodynamic processes involving serpentinite formation, metamorphism or exhumation. Magnetite is produced during serpentinization of the lithospheric mantle at slow-spreading

ridges for T>200°C (Klein et al., 2013; Klein et al., 2014) and controls thereby hydrogen production (e.g., see Brunet, 2019 for a review). Magnetite in subducted serpentinite can survive temperature conditions up to the antigorite breakdown (Debret et al., 2014; Merkulova et al., 2017). Therefore, in addition to constraining the timing of mafic and ultramafic rocks exhumation, dating of magnetite by (U-Th-Sm)/He geochronometry can also put time constraints on the functioning of fossil hydrothermal systems and on serpentinite metamorphism in subduction-collision zones.

The MgHe method is based on radiogenic ⁴He production and accumulation in the crystal 80 lattice during alpha decay of U, Th, and Sm isotopes (²³⁸U, ²³⁵U, ²³²Th and ¹⁴⁷Sm) that lead to 81 stable isotopes (²⁰⁶Pb, ²⁰⁷Pb, ²⁰⁸Pb and ¹⁴³Nd respectively). The alpha particle is ejected with 82 83 a high kinetic energy of several MeV and travels a mean distance of ~15 µm in the magnetite 84 lattice (Ketcham et al., 2011). A correction for He loss due to alpha ejection (called F_T) needs 85 to be applied which can range from few percent to 20-30% for smaller grains with sizes of ~100 μ m (Farley et al., 1996). Otherwise, abrasion of the first ~15 μ m of the crystals or 86 87 selection of the inner parts of a large crystal can be done to avoid ejection or implantation from neighboring minerals (Blackburn et al., 2007; Cooperdock and Stockli, 2016). The 88 89 remaining He atoms in the crystal can diffuse outside of the crystal with a rate of diffusion 90 that is a function of the temperature and crystal volume. Based on experimental He diffusion 91 coefficients obtained on magmatic magnetite by Blackburn et al. (2007), the closure 92 temperature Tc (Dodson, 1973) is as 250±40°C for a crystal size of 500 μm and a cooling rate 93 of 10°C/Ma. The Tc for He in magnetite falls in the same range as for the zircon fission track 94 (ZFT) system, which is broadly estimated to be ~240±50°C (Gallagher et al., 1998). Ketcham (2019) derived a Tc value of 281°C based on fitting experimental data sets from relatively 95

96 young zircons (Murakami et al., 2006), and Rahn et al., (2004) showed that radiation damage
97 may lower annealing resistance.

The number of studies dealing with the MgHe method remains limited. MgHe has been applied to volcanic rocks (Blackburn et al., 2007), ore deposits (Fanale and Kulp, 1962) as well as serpentinites and chloritoschists (Cooperdock and Stockli, 2016). These studies have clearly demonstrated the relevance of the MgHe thermochronometer to obtain cooling or crystallization ages. However, MgHe data have never previously been used to infer the thermal history of rocks using inverse modeling, which is the main purpose of collecting thermochronometric data.

Here, the MgHe method is applied to an ophiolite from the Schistes lustrés units (Western
Alps, France), for which the cooling history is relatively well constrained, as the mineralogy
and P-T history of the area have been characterized (Pognant and Kienast, 1986; Tricart and
Schwartz, 2006) and ZFT and AFT data are available (Schwartz et al., 2007). The obtained
MgHe data are critically evaluated in the framework of the ophiolite thermal history
constrained by the ZFT and AFT literature data.

111

112 2. Geological setting

The studied samples were collected in the internal domain of the western Alps (Fig. 1A) that underlies the boundary between the European and African plates (e.g., Lemoine et al., 1986). This domain (Fig. 1B) derives from a part of the Tethyan ocean (Piedmont zone) and its northwestern European margin (Briançonnais zone). In the study area, the Piedmont zone is composed of the Schistes lustrés units and the Monviso ophiolite, which rest on the internal crystalline massif of Dora Maira (Fig. 1C). These units were juxtaposed during subduction and collision processes in Late Cretaceous to Late Tertiary times. The Piedmont zone 120 represents different levels of a paleo-subduction zone, preserving the low thermal gradient (5-8°C/km) associated with the subduction dynamics (Schwartz et al., 2001; Agard et al., 121 122 2002). In this context, the Schistes lustrés units correspond to a fossiliferous sedimentary 123 accretionary wedge metamorphosed under high pressure and low temperature (HP-LT) 124 conditions during the late Cretaceous-early Eocene subduction of the Tethyan ocean (Lafay 125 et al., 2013). This domain is derived from Mesozoic oceanic sediments that were originally 126 composed of marls, clays, and limestones associated with fragments of oceanic lithosphere. 127 These sediments were strongly deformed and metamorphosed during subduction, and they 128 crop out today as foliated and polydeformed calcschists enclosing boudinaged meter- to 129 kilometer-sized Jurassic ophiolites. The P-T conditions increase toward the east, from low-130 temperature blueschist (LT-BS, Fig. 1C) to high-temperature blueschist facies conditions (HT-131 BS, Fig. 1C) (Tricart and Schwartz, 2006). The blueschist domain is structurally located above 132 the Monviso eclogitic ophiolite (Fig. 1C). In this massif, the metasedimentary component is 133 small (<20 vol.%) relative to the voluminous oceanic lithosphere (Schwartz et al., 2001).

134 The P-T evolutions of the Schistes lustrés units are well documented and show clockwise loop paths, keeping a record of the early prograde evolution, and the retrograde 135 136 metamorphism stages have remained in the field typical of greenschist paragenesis (Fig. 1D), 137 with no temperature increase (Tricart and Schwartz, 2006). For the studied Rocher Blanc 138 ophiolite located in the central part of the Schistes lustrés units (Fig. 1C), the metamorphic 139 peak conditions are estimated at 10±2 kbar and >350°C (Tricart and Schwartz, 2006, 140 Schwartz et al., 2013), with laser probe Ar-Ar ages ranging from 62 to 55 Ma (Agard et al., 141 2002). The final exhumation is constrained by apatite and zircon fission track (AFT-ZFT) 142 analysis (Schwartz et al., 2007) (Fig. 1C), which reveals a diachronous exhumation of the HP-143 LT metamorphic rocks from east to west during Oligocene to Miocene time. Taking into

account the cooling rates from Schwartz et al. (2007), the Rocher Blanc ophiolite is
considered to have passed through the ~240-280°C isotherm at 24.8±1.6 Ma and the ~110°C
isotherm at 12.3±2.0 Ma (Fig. 1C).

147

148 **3. Samples and methods**

149 **3.1. Sampling and petrologic characterization**

150 Sample RB1 was collected from the Rocher Blanc massif (RB; WGS84 UTM31N, 6°57'08.01"E, 151 44°39'53.25"N; elevation: 2810 m a.s.l.) that corresponds to a km-scale ophiolitic body 152 composed of mafic and ultramafic rocks embedded in the Schistes lustrés units (Fig. 1C) that 153 record middle-temperature blueschist facies conditions (MT-BS) (Pognant and Kienast, 1986; 154 Tricart and Schwartz, 2006). The high-pressure blueschist paragenesis in mafic rocks 155 (glaucophane, zoisite) has been partially to totally retrogressed at greenschist facies 156 conditions (chlorite, titanite and augite-aegirine). In ultramafic rocks, high pressure is 157 characterized by the presence of antigorite (Schwartz et al., 2013). The studied sample (RB1) 158 corresponds to a magnetite-bearing chloritized Fe-Ti meta-gabbro (chloritoschist), some tens of centimeters thick, located at the interface between meter-size lenses of Fe-Ti gabbro and 159 160 serpentinite (Fig. 2A). The sample contains visible mm-scale euhedral magnetite crystals as 161 shown in Figure 2B, which are superimposed over the main schistosity defined by the 162 chlorite – titanite assemblage (Fig. 2C).

163

164 **3.2. Mineralogical and geochemical analysis**

165 The chemistry of both euhedral and pseudo-euhedral magnetite crystals of the RB1 sample 166 were characterized at the microscale. Semi-quantitative analyses and imaging of five 167 magnetite and their inclusions were performed with quantitative analyses and mapping of minor and trace elements in magnetite (Si, Ti, Cr, V, Mn and Mg). Several SEM images are shown in Figures 2D, E, F, and concentrations of Si, Ti, Sr, V, Mn, and Mg obtained on magnetite crystals are reported in Table 1. Figure 3 presents the Si and V contents of one euhedral and one pseudo-euhedral magnetite crystal identified also in Figure 2C. Additional information can be found in the supplementary section.

173 In-situ uranium and thorium concentrations in chlorites, titanites and magnetites from 174 sample RB1 were quantified by laser ablation inductively coupled plasma mass spectrometry 175 (LA-ICPMS). Results are presented in Table 2 and on Figure 4. Nevertheless, 2σ internal 176 errors are reported in Table 2 and additional details can be found in the supplementary 177 section.

178

179 **3.3. Crystal selection using CT scanning**

180 Magnetite grains scans were performed following guidelines by Cooperdock and Stockli 181 (2016, 2018), and addition details can be found in the supplementary section. The blob3D 182 software (Ketcham, 2005a; Cooperdock et al., 2019) was used to determine precisely the 183 volume, surface/volume ratio, sphere-equivalent radius and alpha-ejection factor FT 184 (Ketcham et al., 2011) for each grain, and the presence of large inclusions (>5 µm). 3D 185 images of the magnetite crystals were generated to aid in selecting crystals for (U-Th-Sm)/He 186 dating and documenting their euhedral and pseudo-euhedral shapes. 3D grain shape images 187 of MgHe-analyzed magnetite crystals are reported in Figures 5A and B.

188

189 **3.4. (U-Th-Sm)/He on magnetite**

190 Non-abraded single mm-size euhedral and pseudo-euhedral magnetite crystals were
191 degassed under-vacuum in order to determine the He content using the protocol described

in Allard et al. (2018). He content is corrected from the blank level, which is from 40 to 160 times lower than the He signal. Helium content error is estimated to 2% accordingly. The wet chemistry digestion protocol was adapted from Blackburn et al. (2007) allowing the determination of the ²³⁸U, ²³²Th, ¹⁴⁷Sm content. Additional details are placed in the supplementary section. Analytical uncertainties in the U, Th and Sm content translate into a 15% error on alpha-ejection-corrected MgHe ages. Sample mass, volume, He, U, Th, Sm content, raw and alpha-ejection-corrected (U-Th-Sm)/He ages are reported in Table 3.

199

200 **4. Results**

4.1. Magnetite phase relationships and chemical composition

202 At the thin section scale, magnetite crystallizes in two zones of distinct mineralogy, which 203 are likely inherited from the initial distribution of magmatic minerals in the Fe-Ti gabbro (Fig. 204 2C). The mineralogy of these two zones is either dominated by a chlorite + magnetite ± relict 205 magmatic pyroxene (Zone 1) or titanite + magnetite ± chlorite ± relict magmatic ilmenite 206 (Zone 2) assemblages. In Zone 1, the augite-aegirine pyroxene may stably coexist with 207 titanite and chlorite contains a 10-20 % jadeitic component. In Zone 2, titanite, magnetite 208 and chlorite appear texturally as destabilization products of magmatic ilmenite and 209 generates local reaction coronas (Fig. 2D). Plagioclase has not been identified in either of the 210 two zones. Magnetite crystal habits differ between the two zones. In Zone 2, magnetite is 211 pseudo-euhedral to anhedral, possibly polycrystalline, and contains numerous inclusions (Fig. 212 2D-E) of titanite and chlorite. In Zone 1, magnetite is euhedral, and can form larger grains 213 exceeding 500 µm with rare inclusions (Fig. 2F) of augite-aegirine, chlorite and titanite.

Chemical characterization of euhedral magnetite crystals from Zone 1 and pseudo-euhedral
magnetite crystals from Zone 2 are reported in Table 1. Significant Mn, Cr and Ti contents

216 are present but remain below 500 ppm (Table 1). Mg is below detection limit (i.e., <25 ppm), 217 while Si and V are the two most concentrated cations after Fe, with mean concentrations of 218 809±513 and 1821±265 ppm, respectively (Table 1). The Si and V distribution patterns are 219 the same in euhedral and in pseudo-euhedral magnetite grains indicating similar growth 220 process and history. It is characterized by oscillatory zoning with a relatively good correlation 221 between Si and V-rich growth zones (Fig. 3), although vanadium zoning is sharper. The 222 nature of the inclusions and their distribution seem independent of the chemical variations 223 recorded by these two elements.

224

4.2. Magnetite, chlorite and titanite in-situ U content

226 The U content of magnetite derived from laser ablation ranges from 5±0.4 to 24±16 ppb and 227 is ~five times higher in the rim compared to the core (Table 2, Figure 4). In addition, the U 228 content is similar to the U value obtained by wet chemistry for the dated crystals with value 229 of 8±1 to 29±2 ppb (Table 3). The U content is also similar between magnetite from Zone 1 230 (euhedral) and Zone 2 (pseudo-euhedral) (Table 2 and 3). The U content in chlorite from 231 Zone 1 of ~16±8 ppb is lower than the content in chlorite from Zone 2 which ranges from 232 42±5 to 68±10 ppb (Table 2, Figure 4). Although variable, the U content of titanite is by far 233 the highest with values ranging from 27±5 to 1569±141 ppb. In the RB1 sample titanite is the 234 U-bearing phase.

235

236 4.3. Magnetite (U-Th-Sm)/He

Five euhedral magnetite crystals (Fig. 5A) and two pseudo-euhedral grains (Fig. 5B) were selected for MgHe dating, utilizing the CT scan data to select grains with the fewest (for pseudo-euhedral magnetite) or no inclusions. The selected crystals have equivalent-sphere radii (Rs) ranging from 362 to 483 μm (Table 3). Helium content of these seven magnetite
grains ranges from 0.9 to 3.3 pmol/g and effective uranium (U+0.24Th) ranges from 9 to 34
ppb (Table 3). Ejection-corrected MgHe ages for euhedral and pseudo-euhedral grains range
from 14.8±2.1 to 20.9±3.0 Ma and 30.4±4.4 to 32.0±4.6 Ma, respectively (Fig. 5C).

244

245 **5. Discussion**

246 **5.1. Conditions of magnetite formation**

247 Magnetite is a common mineral in ultramafic and mafic rocks and appears over a large range 248 of P-T and fO₂ conditions (Nadoll et al., 2014). In the case of magnetite created during the 249 oceanic serpentinization process, it is classically assumed to form at temperatures above 250 ~200°C (e.g., Alt et al., 2013; Klein et al., 2013; Klein et al., 2014), even if magnetite could 251 potentially crystallize at temperatures even below 120°C (Alt et al., 2013). In the case of 252 subducted ophiolites (serpentinite and gabbro lithologies), magnetite is observed from 253 eclogite paragenesis (Li et al., 2004; Debret et al., 2014) up to greenschist facies conditions 254 related to the retrograde stage for temperatures ranging from 550 to 250°C (Schwartz et al., 255 2013).

256 The petrological and geochemical studies of the RB1 sample indicate that the euhedral and 257 pseudo-euhedral magnetite grains crystallized during the retrograde path under greenschist 258 facies conditions assisted by metamorphic hydrothermal fluids and are not derived from 259 magmatic origin. Indeed, the Ti content of magnetite crystals is under 100 ppm (Table 1), too 260 low to be of magmatic origin (Nadoll et al., 2014), and the host inclusions (titanite, chlorite, 261 augite-aegirine) that are representative of HP-LT paragenesis demonstrate that magnetite 262 grains crystallized during the greenschist retrograde path. In this context, the oscillatory V 263 and Si chemical zoning, with a Si-richer rim observed for euhedral and pseudo-euhedral 264 magnetite grains (Figure 3) record the changing conditions and compositions of the 265 metamorphic fluid related to magnetite growth (Dare et al., 2014). The Si and V are the main 266 minor elements at the ~1000 to 2000 ppm level, compared to Ti, Cr, Mn, and Mg (Table 1); 267 these Si and V contents reveal high solubility indicating crystallization from a high 268 temperature fluid (Dare et al., 2014). Such high Si content has already been encountered in 269 hydrothermal systems (skarns, metamorphic veins; e.g., Shimazaki, 1998) and Si oscillatory 270 zoning has been described in hydrothermal magnetite (Deditius et al., 2018). Hydrothermal 271 conditions (250-350°C, Fig. 6A) in the fluid-circulation zone marked by the magnetite-bearing 272 chloritoschists may have favored Si and V incorporation, the latter element being supplied 273 locally by the Fe-Ti gabbro itself.

Based on previous thermobarometric studies of the host metamorphic Fe-Ti gabbro (e.g., Pognante and Kienast, 1986), and petrological investigation in this study, the magnetite begins to crystallize most probably near the blueschist – greenschist transition, as residual augite-aegirine is found in inclusions, and magnetite crystallization continues during the entire greenschist retrograde path (Fig. 6B).

279

280 **5.2. Magnetite He age distribution**

The MgHe ages for all crystals are younger than the high-pressure ages recorded for the Schistes lustrés units (62-54 Ma; Agard et al., 2002), and potentially represent exhumation ages, i.e., related to the retrograde path. Interestingly, the euhedral grains are quite homogeneous with a mean age of 18.1±2.4 Ma that falls between the AFT and ZFT ages as deduced from Schwartz et al. (2007) for the Rocher Blanc area (Fig. 5C), and are thus compatible with cooling ages associated with the exhumation history of the area. Indeed, the MgHe ages are younger than ZFT ages, and older than AFT ages, in agreement with the 288 published thermal sensitivity of MgHe (250±50°C; Blackburn et al., 2007), ZFT (240±50°C, Gallagher et al., 1998; ~281°C, Ketcham, 2019) and AFT (110±10°C, Gallagher et al., 1998). 289 290 Conversely, the ages obtained on pseudo-euhedral crystals are significantly older than the 291 euhedral crystal and ZFT ages (Fig. 5C). This is not consistent with petrological data which 292 show that euhedral and pseudo-euhedral magnetite all crystallized simultaneously under 293 greenschist conditions at T > 250°C. Thus, we interpret these ages as related to the presence 294 of excess He compared to euhedral crystals with similar size and eU content (Table 3); they 295 cannot be considered as cooling ages. Excess of He may be due to (i) the impact of alpha 296 implantation from neighboring minerals or circulating rich U-Th fluids, and/or (ii) the impact 297 of inclusions with higher U-Th content on the global He budget of magnetite.

298 Alpha implantation has been frequently observed or modeled for the (U-Th)/He method, 299 especially for apatite (e.g., Spiegel et al., 2009; Gautheron et al., 2012) and magnetite 300 (Blackburn et al., 2007; Cooperdock and Stockli, 2016). When minerals contain lower or 301 equivalent U and Th content compared to neighboring phases, the alpha implantation effect 302 can be significant. This is particularly the case for magnetite, which typically contains low U-303 Th content (<30 ppb) compared to minerals with high U-Th content of >100 ppb like titanite 304 in the present case (Table 2). In Zone 1, magnetite is embedded in chlorite which is U-poor 305 with contents ~15 ppb. It must be noted that the higher U content in Zone 2 results in a 306 higher U content, from 40 to 60 ppb, in chlorite as well (Table 2, Figure 4). In the RB1 sample, 307 for the chlorite area (Zone 1) alpha implantation in magnetite should be negligible whereas 308 in the titanite area (Zone 2) alpha implantation in magnetite could be significant. As the RB1 309 grains are large (Rs >400 μm) with small alpha ejection corrections (F_T>0.97; Ketcham et al., 310 2011), a 20-40 ppb matrix would simply cancel out the F_T correction, and only thousands of 311 ppb U-Th will have an age effect approaching 10% (Gautheron et al., 2012). For euhedral

magnetite, MgHe ages younger than ZFT imply no significant alpha implantation and reinforce the hypothesis that euhedral magnetites crystallized in Zone 1 (chlorite-rich environment). Similarly, MgHe ages older than ZFT observed for the pseudo-euhedral magnetite may reflect implantation from U-Th-rich neighboring minerals such as titanite, consistent with the pseudo-euhedral magnetites crystallizing in Zone 2, in agreement with petrological observations.

318 Inclusions richer in U-Th than the host minerals can also impact (U-Th)/He ages by adding He 319 to the crystal lattice, as already reported for apatite (Farley, 2002) and magnetite (Blackburn 320 et al., 2007; Cooperdock and Stockli, 2016). An excess of He can occur if the total U-Th 321 content of host mineral and inclusions is incompletely retrieved by the chemical digestion 322 protocol that is usually designed to target the host mineral. For magnetite, the digestion 323 protocol is not adapted to dissolve completely silicates (e.g., chlorite, titanite, zircon) or 324 other oxide minerals (e.g., titanite). A small fraction of the U-Th budget coming from 325 inclusions may be retrieved, but the total U-Th budget is incomplete, leading to spuriously 326 older MgHe ages with no geological significance. For the two pseudo-euhedral magnetites 327 from RB1, the MgHe ages older than the euhedral magnetites with similar U-Th content 328 (Table 2) would be consistent with the absence of U-Th-rich inclusions in the euhedral 329 crystals. However, because the CT scan analysis indicated minimal detectable inclusions, and 330 because the primary U-Th-bearing phase, titanite, has eU compositions only ~100x higher 331 than the magnetite (Table 2, 3), we do not believe inclusions can explain the magnitude of 332 excess He. We thus interpret that MgHe ages older than ZFT reflect implantation across 333 large areas of grain boundaries shared with titanite (Fig. 4D), and for the rest of the 334 discussion we do not use or further discuss the two pseudo-euhedral crystals, as their 335 measured ages are not geologically meaningful. However, we can conclude that the MgHe

ages from the euhedral crystals of the RB1 sample are not significantly impacted byimplantation and reflect cooling due to exhumation along the retrograde path.

338 Finally, the results of this study give some methodological insight into MgHe dating. First, CT 339 imaging is essential for detecting the presence of mineral inclusions larger than \sim 5 μ m, as 340 already stated by Cooperdock and Stockli (2016). Comparison of U-Th content for crystals 341 with older MgHe ages can provide indirect evidence of the impact of rich U-Th inclusions 342 below the CT-scan detection limit (<1 µm). Second, the removal of the 20 µm outer margin 343 of the magnetite crystals, as recommended by Blackburn et al. (2007), is not always 344 necessary. In fact, the result of this study shows that implantation is not always significant 345 and depends on the petrological environment of magnetite (i.e. chlorite versus titanite 346 zones), as well as the size of the crystal. Furthermore, alpha implantation along long grain 347 boundaries shared with titanite, as documented in our zone 2 non-euhedral magnetites, if 348 combined with diffusional migration of helium into the magnetite grain (Gautheron et al. 349 2012) may render abrasion an insufficient remedy, depending on cooling history. Thorough 350 petrological characterization of mineral textures and compositions thus appears to be a 351 necessary step of the MgHe method.

352

353 **5.3. Origin of MgHe age dispersion**

Significant MgHe age dispersion is encountered for euhedral RB1 magnetites (MgHe age = 18.1±2.4 Ma, 13% at 1 σ). MgHe age dispersion has also been reported in previous work in the HP-LT subduction complex on the island of Syros, Greece (Cooperdock and Stockli, 2016). This age dispersion is higher than typical analytical error (i.e. ~5-10 %, for apatite and zircon) but similar or lower than frequently found in practice for apatite or zircon (U-Th)/He methods (e.g., Ault et al., 2019; Ketcham et al., 2018). Using investigations already published

360 on the origin of apatite or zircon (U-Th)/He age dispersion, some factors affecting MgHe 361 ages related to the impact of (i) U-Th zonation, (ii) slight alpha-implantation, or (iii) crystal-362 dependent He diffusion coefficient can be anticipated.

363 First, a slight impact of alpha-implantation could affect the MgHe age, as the U-Th content of 364 the neighboring minerals could be slightly different For example, U and Th content for 365 chlorite of ~15 ppb has been determined for Zone 1 (Table 2), but slightly higher U-Th values 366 are possible (Table 2). However, the large crystal size of the magnetite from this study 367 reduces the impact of implantation. Nevertheless, other parameters can account for 368 dispersion. For instance, Cooperdock and Stockli (2016) found similar MgHe age dispersion 369 on abraded crystals (i.e. mechanical removal of the 20 µm border of the crystal that can 370 have been affected by alpha implantation).

371 Second, due to the low U-Th content, analytical determination of zoning is difficult, and up 372 to a ~4x higher U concentration in the rim has been measured for three magnetite crystals 373 (Table 2). Because the ablation spot is large (150 μ m), the exact size of the U-Th-enriched 374 rim is not known but can be similar to the one observed by Si-V analysis, i.e. \sim 100 μ m (Fig. 3). 375 In this case, the presence of a \sim 100 μ m enriched U-Th rim will not have much impact on the 376 ejection factor as it is much larger than the mean stopping distance (15 μ m). If finer zoning 377 exists, modeling can be performed (e.g., Farley et al., 1996; Gautheron et al., 2012) assuming 378 that U-Th zoning follows Si and V zoning (Fig. 3C and D). Using the Monte Carlo code 379 associated with Gautheron et al. (2012), we modeled a 400-µm radius sphere with a 50 µm 380 rim depleted or enriched in U-Th by 50%, and as a worst-case scenario a rim containing all of 381 the crystal's U-Th (i.e., no U-Th in the core). We obtained F_T ejection factors of 0.99, 0.96, 382 and 0.76, respectively, implying a 25% MgHe age difference for the worst case. However, in 383 comparing our bulk and in situ core and rim analyses we find no evidence of such severe

zoning in U-Th, and we can estimate a maximum MgHe age dispersion component of only up
to a few percent. Given that implantation can be from minerals with much more than 50%
higher eU, it is likely to be a source of equal or greater dispersion.

387 Third, a defect- and/or composition-dependent He diffusion coefficient could also be 388 invoked to explain age dispersion, as has been proposed and modeled for other minerals 389 such as apatite and zircon (Gautheron et al., 2009; Flowers et al., 2009; Guenthner et al., 390 2013). Indeed, damage due to radioactivity, and also crystal defects could trap or facilitate 391 He diffusion within the crystal structure. A damage content of some 5% would be required 392 to change the He diffusion coefficient significantly, as modeled for apatite (e.g., Gerin et al., 393 2017), which is unlikely given the low eU content of magnetite (<<1%). Alternatively, 394 substitution by minor or trace elements could deform the crystal structure and thus modify 395 He diffusion, as shown in apatite at the atomic scale (Djimbi et al., 2015). The magnetite 396 crystals from RB1 have ~0.5 wt% of minor elements, and if we assume similar effect as for 397 apatite (Djimbi et al., 2015), no significant change in the closure temperature compared to 398 pure magnetite is expected. This is highly speculative, as the composition of the volcanic 399 magnetite used by Blackburn et al. (2007) is not known. The effect of minor elements in 400 magnetite on He volume diffusion still needs to be investigated. It is also possible that fluid 401 inclusions can act as traps for He, as has been proposed for apatite and zircon (Zeitler et al. 402 2017; Danišík et al., 2017); this possibility could be further pursued with higher-resolution 403 imaging and/or ramped heating experiments (McDannell et al., 2018).

Finally, we stress that there is currently only one experimental study documenting He diffusion kinetics in magnetite (Blackburn et al., 2007), and even that data set shows signs of complex release behavior. Further data are certainly required to build a more robust database of kinetic parameters and explore possible anomalous behaviors.

409 5.4. MgHe data inversion: resolving exhumation rate from the Rocher Blanc ophiolite 410 The petrological and geochemical arguments deduced from the RB1 sample analysis indicate 411 that magnetite crystalized at T> 250°C, as illustrated in Figure 6, at temperature higher than 412 the MgHe closure temperature (Tc), and thus that MgHe ages obtained from euhedral 413 magnetite of ca. 18.1±2.4 Ma correspond to cooling ages of the Rocher Blanc ophiolite (Fig. 414 6). Specifically, for the dated RB1 magnetites, which all have similar sizes (Table 3), Tc ranges 415 from 242 to 248°C for sphere-equivalent radii of 362 to 483 µm and a cooling rate of 416 10°C/Ma. MgHe ages derived from are sandwiched between ZFT and AFT ages as expected 417 by their thermal sensitivity, and in fact shed further light on the thermal sensitivity of the 418 ZFT method, supporting a ZFT Tc >260 °C. Figure 6 summarizes the P-T path for RB1 sample, 419 including the magnetite growth field during the retrograde path and the time constraint 420 obtained from AFT, ZFT and MgHe ages. 421 5.4.1. MgHe data inversion 422 The thermal history of the studied area was retrieved using the HeFTy software (Ketcham, 423 2005b) by inverse modeling of MgHe, AFT and ZFT low temperature thermochronometric 424 data. HeFTy uses a constrained Monte Carlo scheme that poses a large number of candidate 425 time-temperature (t-T) paths, and retains those that fit the thermochronometric data to a 426 "good" (closer than statistical expectation for a random sample) or "acceptable" (does not 427 fail a 95% confidence test) extent. For the inversions the fission-track annealing models by 428 Ketcham et al. (1999) and Ketcham (2019) were used for AFT and ZFT data, respectively. For 429 MgHe ages, the sphere-equivalent radii (Rs) of each crystal and He diffusion parameters

430 from Blackburn et al. (2007) were used. In detail, the HeFTy software models (i) the fission-

431 track production and annealing in apatite and zircon, and (ii) helium production, taking into

432 account ejection and diffusion, in magnetite. In this way, the data inversion identifies a 433 family of potential cooling histories by using 17 t-T nodes randomly placed to define a 434 cooling-only path with a maximum allowed cooling rate of 50°C/m.y. over any segment. In 435 addition to the MgHe, AFT and ZFT data, the peak HP-LT metamorphic conditions for the 436 Rocher Blanc ophiolite massif as reported in Figure 6, estimated to be 350±20°C at 58±4 Ma 437 (Agard et al., 2002) is used at a starting point. Further model inputs are listed in Table 4.

Thermal history modeling was performed with AFT-ZFT alone, with MgHe data alone and 438 439 with all three combined (AFT-ZFT and MgHe). First, modeling was done with interpolated 440 AFT and ZFT ages for RB1 sample assuming that ages vary linearly with the sample position 441 on a W-E transect (Schwartz et al., 2007) (Fig. 1), and with and without MgHe data. Second, inversions were also performed with AFT and ZFT data from RN and CG samples, that 442 443 represent the available data geographically framing the RB ophiolite (Fig. 1C), and with and 444 without MgHe data. Finally, for comparison, the AFT-ZFT data and the MgHe data separately 445 for RB1 sample were inverted. For modeling of MgHe data, all euhedral grains were 446 individually included. The computed thermal histories are reported on Figure 7. In the on-447 line supplement, Figure S1 presents the modeling inversion results obtained for AFT and ZFT 448 data of CG and RN samples.

449

450 **5.4.2.** Thermal history modeling

The inversion of the MgHe data alone or with the AFT and ZFT literature data allows us to refine the final t-T exhumation history of the Rocher Blanc ophiolite massif, while also verifying good consistency with AFT and ZFT data. First, model outputs showing the set of good and acceptable fits, and their weighted mean, demonstrate that the MgHe data provide time constraints at ca. 230-240°C (Fig. 7A) whereas ZFT and AFT ages put constraints

456 on the Rocher Blanc ophiolite cooling history at ~280 and ~100°C (Fig. 7B) respectively. In 457 particular, the MgHe data refine the thermal history by requiring that the rapid phase of 458 cooling did not commence until after 20 Ma (Fig. 7C). The three systems are thus 459 complementary with respect to the temperature range that they cover. The apparent 460 consistency between MgHe, AFT and ZFT cooling ages suggests that the He diffusion data 461 obtained at 1 bar by diffusion experiments in vacuum at temperatures between 350 and 462 620°C over several days by Blackburn et al. (2007) can be extended to geological conditions 463 and timescales. As the AFT and ZFT data for RB1 sample are extrapolated from a different 464 locality, we also ran models with AFT and ZFT ages from samples CG and RN (Fig. S1). The 465 resulting cooling histories only slightly differ from those obtained with the extrapolated 466 Rocher Blanc ophiolite ages (Fig. S1) supporting the relevance of this thermal modeling 467 approach.

468

469 **5.5.** Final exhumation of the Rocher Blanc ophiolite: geodynamic implications

470 The thermal history obtained using the combination of thermochronometric methods (AFT-471 ZFT and MgHe data) brings strong constraint on the final exhumation of the Rocher Blanc 472 ophiolite and also on the Schistes lustrés units of the Piedmont zone (Western Alps). Our 473 thermal modeling reveals that the RB1 sample has undergone nonlinear cooling and records 474 a major change in the thermal pathway at 20 Ma (Fig. 7C). During the first exhumation stage, 475 from 55 to 20 Ma, the sample recorded slow cooling from 350 to 300 °C associated with an 476 extensive exhumation from 30 to 10 km (Fig. 6). The exhumation of ~20 km of rocks from 477 the Alpine orogenic wedge cannot be accounted for by surface denudation (i.e. erosion) only. 478 Large exhumation with limited cooling rather corresponds to exhumation in ductile 479 conditions related to subduction dynamics and underplating processes (Agard et al., 2001).

480 During ductile exhumation the sample passes from blueschist to greenschist conditions with a mean exhumation rate lower than 1 mm/yr, calculated using the depth and time 481 482 constraints (Fig. 6A and 7). Over the second exhumation stage, from 20 Ma to 10 Ma, the 483 RB1 sample records rapid cooling from 300 to 100°C (Fig. 7) associated with an exhumation 484 from 10 to 3 km (Fig. 6A). This exhumation and strong change of cooling rate reflect 485 denudation in the brittle conditions. A mean exhumation rate of 0.7 mm/yr can be deduced (Fig. 6A and 7). The last information of the thermal history is given by the final exhumation 486 487 stage recorded by AFT data and an exhumation of 3 km over 10 Ma leading to a mean lower 488 exhumation rate of 0.3 mm/yr (Fig. 7). Interestingly, the sample records similar exhumation 489 rate in ductile and brittle domains around <1 mm/yr, but significant cooling rate variations. 490 The change at ~20 Ma corresponds to the passage of the Rocher Blanc ophiolite below the 491 transition from ductile to brittle conditions in the lower grade of the greenschist facies 492 conditions (~250°C) during the retrograde path.

493 The final exhumation recorded by the thermal history reflects denudation at brittle 494 conditions (Fig. 7), where denudation is related to erosion linked to relief creation. The relief 495 creation is associated to a westward tilting of the Schistes lustrés units (Tricart et al., 2004; 496 Schwartz et al., 2009), as a response to the upward indentation of the Dora Maira massif. 497 Indeed, the crustal scale geometry of the internal zone is characterized by the presence of 498 Apulian mantle in a superficial position (Fig. 8), at 12 kilometers below the Dora Maira 499 massif (Lardeaux et al., 2006; Zhao et al., 2015). This portion of cold and rigid mantle acts as 500 an indenter that allowed the deformation of the thinned Briançonnais margin by crustal 501 underthrust process constituting a continental subduction wedge that forms the current 502 basement of the Piedmont zone (oceanic subduction wedge in Fig. 8). This continental

subduction wedge was overthrust on the European crust by the activation of the crustal
Penninic Frontal Thrust (PFT in Fig.8) at ~34 Ma (Dumont et al., 2012).

505 Taking in consideration the MgHe data of this study, the AFT-ZFT data of Piedmont zone 506 (Schwartz et al., 2007) and AFT data of the internal crystalline Massif of Dora Maira (Beucher 507 et al., 2012), we can illustrate the evolution of the area with their respective positions 508 though time in the orogenic wedge (Fig. 8). At present-day, all samples crop out at similar 509 elevations (>2000 m a.s.l.) for the Piedmont zone and at lower elevation (<1500 m a.s.l.) for 510 Dora Maira. At present-day the global shortening affects the front of the European foreland 511 (Schwartz et al., 2017), and allows the uplift and exhumation of the external crystalline 512 massifs (Pelvoux, Beucher et al., 2012) (Fig. 8). Using the AFT-ZFT and MgHe ages and 513 thermal sensitivity of the thermochronometric systems for all the samples, we can 514 reconstruct their position at depth ~20 Ma ago. Firstly, the samples CR from Schistes lustrés 515 and from Dora Maira present AFT age of 20-25 Ma indicating a shallow position around 3 km 516 (assuming an AFT closure temperature of 100°C and a thermal gradient of 25 °C). The 517 position at depth for the other samples from the Piedmont zone need to be progressively 518 deeper from east to west (Fig. 8). RB1 sample was passing the 250°C isotherm ~20 Ma based 519 on the thermal history obtained in this study (Fig. 7). From ~20 Ma to present-day, the 520 Piedmont zone has recorded a differential exhumation with a westward titling and uplift 521 that is recorded by the youngest ages to the east in the Monviso (Fig. 8). This exhumation is 522 controlled by the upward indentation of the Apulian mantle producing the global tilting of 523 the Piedmont zone (Fig. 8).

524

525 6. Conclusions

This study contributes to the understanding of the formation conditions of magnetite in HP-LT ophiolite and to the use of magnetite (U-Th-Sm)/He as a thermochronometer to quantify the exhumation history. We demonstrate that a comprehensive characterization of magnetite crystals and crystallization environment is needed in conjunction with MgHe dating in order to interpret the age data.

We specifically focused on one sample (RB1) from the Rocher Blanc massif that is an 531 ophiolitic body embedded in the Schistes lustrés units (Western Alps, France), which 532 533 contains abundant magnetite. This sample recorded middle-temperature blueschist facies 534 conditions and was overprinted by greenschist conditions during exhumation. We show that 535 for the RB1 sample, magnetite crystallized in two zones of distinct chemistry, that originated 536 from magmatic relictual zones, with euhedral and pseudo-euhedral grains with a dominant 537 chlorite or titanite mineralogy, respectively. In addition, the petrological and magnetite 538 geochemical analyses (e.g. V and Si) indicate that magnetite crystallized during the 539 retrograde metamorphic path under greenschist conditions (T> 250°C). The Si and V zoning 540 observed in euhedral and pseudo-euhedral magnetite grains reinforce the fact that both type of magnetite crystallized under same physical conditions. The pseudo-euhedral 541 542 magnetite, crystallizing in dominant titanite mineralogy, present older MgHe ages compared to euhedral magnetite. The old MgHe ages are interpreted by the impact of alpha-543 544 implantation from U-Th-rich neighbor minerals (e.g., titanite). On the contrary, alpha-545 implantation in euhedral magnetite crystallizing in chlorite dominant zone is insignificant, as 546 chlorite contain similar U-Th content than magnetite. The MgHe ages derived from euhedral 547 magnetite range from 14.8±2.1 to 20.9±3.0 Ma, sandwiched between ZFT and AFT ages as 548 expected. The MgHe data along with AFT and ZFT data were inverted for the first time to 549 refine the local thermal history. The MgHe data suggest that fast cooling was delayed until ⁵⁵⁰ ~20 Ma, as opposed to the less constrained ~26 Ma that would be inferred from the ZFT and ⁵⁵¹ AFT ages alone. The MgHe data suggest that the RB1 sample has undergone an exhumation ⁵⁵² in ductile conditions related to subduction dynamics and underplating processes, followed ⁵⁵³ by an exhumation reflecting denudation by erosion at brittle conditions. Considering all ⁵⁵⁴ thermochronometric data along the Western Alps, we propose that the exhumation is ⁵⁵⁵ controlled by the upward indentation of the Apulian mantle producing the global titling of ⁵⁵⁶ area and associated erosion.

Additional work on U-Th zoning in magnetite and He diffusion in magnetite will be necessary in the future, as the thermal inversion interpretation is contingent on the Blackburn et al. (2007) kinetic parameters being correct. Nevertheless, the present study further demonstrates that MgHe can be used as a low-temperature thermochronometer to date HP ultramafic final exhumation and the ductile-brittle transition opening new avenues of research. This method effectively complements the available thermochronological tools, making it a powerful tool for emerging applications.

564

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738 Figure caption

Figure 1: (A) Geographical setting of the Alps showing the location of Fig. 1B. **(B)** Simplified structural and geological map of the Western Alps between France and Italy, with location of Figure 1C. **(C)** Structural sketch-map of the Schistes Lustrés units of Piedmont zone, with published apatite and fission tracks ages (modified from Schwartz et al., 2007) and location of the Rocher Blanc ophiolite massif (RB), with extrapolated AFT and ZFT ages (in *). The location of CG (Cirque de la Gavie) and RN (Rocca Nera) are also indicated. **(D)** Pressure-Temperature path from the Rocher Blanc ophiolite (modified from Schwartz et al., 2009).

746

Figure 2: Photographic views of the Rocher Blanc ophiolite massif, polished thin section of 747 748 RB1 sample and SEM backscattered electron images. (A) Fe-Ti gabbro embedded in 749 serpentinite and developing magnetite-bearing chloritoschist at its interface. (B) Millimeter 750 scale euhedral magnetite crystals visible on outcrop indicated by white arrows. (C) Thin-751 section of the RB1 magnetite-bearing chloritoschist sample. The magnetite crystal analyzed 752 is presented in Figure 3. (D), (E) and (F) are SEM-BSE images of the sample RB1. The sample 753 presents two mineralogical zones (2 and 1), rich in titanite and ilmenite (D and E) and 754 chlorite (F), respectively. Mag, Ttn, Ilm and Chl mean magnetite, titanite, ilmenite and 755 chlorite.

756

Figure 3: Geochemical characterization of magnetite crystal from the chlorite-rich Zone 1
and titanite-rich Zone 2 shown in Fig. 2C (RB1-Mag1 and RB1-Mag5, Table 1). (A,E) EMP-SE
image of the magnetite crystal with EMP elemental X-ray maps (1 μm beam size, 1.5 μm step
size and 1s dwell time) and profile. (B,F) Si and V concentration in ppm along the profile.

761 (C,G) and (D,H) map of the Si and V concentration in ppm. Inclusions of silicate minerals762 appear in white on the chemical maps.

763

Figure 4: In-situ U concentration of magnetite, chlorite and titanite from Zone 1 and Zone 2.
(A) Thin-section of the RB1 magnetite-bearing chloritoschist sample, with location of the U
analysis zones (B), (C) and (D). Zoom of the (B), (C) and (D) areas and U content in ppb.

767

Figure 5: Magnetite (U-Th-Sm)/He ages from Rocher Blanc ophiolite. (A) and (B) are 3D
reconstructions from CT-scan images of each euhedral and pseudo-euhedral crystals. (C)
MgHe age as a function of the sphere equivalent radius (Rs) estimated from CT-scan analysis
(Ketcham, 2005a; Cooperdock et al., 2019). The AFT and ZFT age ranges from Schwartz et al.
(2007) and shown in Figure 1C are shown.

773

774 Figure 6: Pressure-temperature-time pathway record by the RB ophiolite and magnetite 775 growth condition. (A) Pressure-Temperature-time path of the HP-LT Rocher Blanc ophiolite 776 modified from Schwartz et al. (2009; 2013). The AFT and ZFT estimated ages from RB (Fig. 777 1C), MgHe ages from euhedral magnetites are reported for their associated closure 778 temperature condition (Blackburn et al., 2007; Gallagher et al., 1998; Ketcham, 2019). The 779 age of HP conditions are taken from Agard et al. (2002). (B) Sketch of magnetite crystals 780 growth during the retrograde pathway in three main steps, from (I) blueschist conditions 781 (BS), (II) High-Pressure greenschist (HP-GS) and (III) Low-Pressure greenschist (LP-GS) 782 conditions. Zones 1 and 2 refer to the initial magmatic sites with augite-plagioclase and Fe-Ti 783 oxide (ilmenite) respectively. (I) The HP blueschist conditions are marked by the partial 784 mineralogical transformation of augite to aegirine, and the development of glaucophane and

785 zoisite in the plagioclase site (Pognante et al., 1986). In Zone 2, ilmenite is progressively transformed to titanite. (II) The HP greenschist conditions are marked by the progressive 786 787 destabilization of augite-aegirine, glaucophane, and zoisite to chlorite and magnetite in Zone 788 1, whereas, ilmenite continues to be destabilized in titanite and ± magnetite and chlorite. 789 The magnetite grains show low Si content whatever the considered zones. (III) Finally, the LP 790 greenschist facies conditions, show the pervasive development of the chlorite and magnetite 791 in the two zones. In the zone 2, only relics of ilmenite remain locally in the core of titanite 792 crystals. The magnetites present higher Si content.

793

794 **Figure 7:** Thermochronological data inversion modeling of the HP-LT Rocher Blanc RB

ophiolite using HeFTy software (Ketcham, 2005). (A). T-t pathways obtained for MgHe data
inversion only. (B). T-t pathways obtained for AFT and ZFT data inversion only. (C). T-t paths
obtained for joint AFT, ZFT and MgHe data inversion.

798

Figure 8: Schematic evolution of the Western Alps between France and Italy ~20 Ma ago and at present-day. At ~20 Ma, the Penninic Frontal Thrust (PFT) is already active since 10 Myr associated with the mantellic upward indentation, and the AFT-ZFT and MgHe dated samples have started to be exhumed in the brittle condition (above 250°C). The samples record differential exhumation related with the westward tilting and uplift. One representative sample from Dora Maira massif is represented, where AFT ages of 20-25Ma have recorded by Beucher et al. (2012).

806

Figure S1: Thermochronological data inversion modeling of the HP-LT Rocher Blanc RB
ophiolite using HeFTy software (Ketcham, 2005). T-t paths obtained using AFT and ZFT data

809	from sample CG (A, B, C), sample RB (D, E, F), RN sample (G, H, I), combined without or with
810	MgHe data. (J-K) Weighted model Tt paths resumed results for CG, RB and RN samples,
811	without or with MgHe data. (L). Tt pathway model obtained for MgHe data inversion only.
812	
813	Table 1: Elemental concentration in ppm obtained on four representative magnetite crystals
814	from zone 1 of the RB1 sample.
815	
816	Table 2: In-situ U and Th concentrations obtained by LA-ICP-MS in sample RB1 (thick section).
817	
818	Table 3: CT-scan and (U-Th-Sm)/He results of the magnetite-bearing chloritoschist RB1
819	sample
820	
821	Table 4: Thermal History Inverse Modeling Parameters

















Table 1: Elemental concentration in ppm obtained on representative magnetite crystalsfrom zone 1 and zone 2 of the RB1 sample.

Element	Si	Ti	Cr	V	Mn	Mg
Detection Limit (ppm)	24	24 30 30 30		30	35	25
		Euhe	dral magnetite			
RB1-Mag1 (n=76)						
Size 500 µm	680±474	138±102	55±11	2119±244	432±35	B.d.l.
Profile B – Fig. 3						
RB1-Mag3 (n=69)	1010+629	77±49	49±10	1660+267	420+24	D d l
Size 650 µm	10191020	72140	46±10	10091207	420±24	D.U.I.
RB1-Mag4 (n=61)	744+240	60+29	45+10	1902+220	420+10	DAL
Size 600 µm	744±349	0UI38	45±10	1802±230	429±19	в.и.і.
		Pseudo-e	uhedral magnet	ite		
RB1-Mag2 (n=53)	002+650	50+21	<i>1</i> 2+11	1652+220	A19+2A	Pdl
Size 550 µm	9951030	22721	45111	10331325	410104	D.u.i.
RB1-Mag5 (n=42)						
480 μm	608±456	n.a.	n.a.	1863±257	n.a.	n.a.
Profile F – Fig. 3						
RB1 (n=301)	809±513	82±55	48±10	1821±265	425±28	B.d.l.

B.d.l.: Below detection limits

n.a. not analyzed

Magnetite	M1	M2	М3	M4c ^{\$}	M4r ^{\$}	M5c	M6c	M6r
		Zone 1	- euhedra	al mag.		Zone 2 - pse	udo euhed	ral mag.
U ± 2σ (ppb)	15±2	5±0.4	17±4	8±2	21±2	7±1	7±3	24±16
Th $\pm 2\sigma$ (ppb)	3±3	bdl	6±3	1±0.7	6±9	3±0.6	4±1	bdl
Titanite	T1	T2	Т3	Τ7	Т8	T4*	T5*	T6*
				Zc	one 2			
$U \pm 2\sigma$ (ppb)	94±8	1569±141	704±56	217±72	111±9	27±5	194±16	38±6
Th $\pm 2\sigma$ (ppb)	28±22	628±546	109±9	163±117	4±1	4±1	36±3	6±1
Chlorite	C4	C5	C1	C2	C3			
	Z	one 1		Zone 2				
U ± 2σ (ppb)	16±8	15±7	61±6	42±5	68±10			
Th $\pm 2\sigma$ (ppb)	5±1	13±5	7±1	13±3	17±12			

 Table 2: In-situ U and Th concentrations obtained by LA-ICP-MS in sample RB1 (thick
 section).

c^{\$} and r^{\$} mean magnetite core and rim * Mn-rich titanite

bdl= below detection limit

Aliquot	Volume	Surface area	Rs	FT	Mass	⁴He	²³⁸ U	²³² Th	¹⁴⁷ Sm	⁴ He±s	²³⁸ U	²³² Th	¹⁴⁷ Sm	eU	Th/U	Raw date	Corrected date±s
	(mm³)	(mm²)	(µm)		(mg)	(pmol)	(ng)	(ng)	(ng)	(pmol/g)	(ppb)	(ppb)	(ppb)	(ppb)		(Ma)	(Ma)
Euhedra	l magnetit	e															
RB1-3	0.29	2.72	444	0.98	1.54	2.33×10 ⁻³	0.025	0.020	0.018	1.49±0.03	16.2	13.0	11.7	19	0.8	14.4±2.2	14.8±2.1
RB1-4	0.16	1.82	368	0.97	0.84	1.85×10 ⁻³	0.019	0.007	0.021	2.18±0.04	22.7	8.4	25.1	25	0.4	16.9±2.5	17.4±2.6
RB1-5	0.21	2.28	410	0.97	1.12	1.16×10 ⁻³	0.009	0.007	0.015	0.88±0.02	7.8	6.5	13.3	9	0.8	20.4±3.1	20.9±3.1
RB1-7	0.25	2.46	483	0.98	1.32	4.71×10 ⁻³	0.038	0.027	0.075	1.85±0.04	28.7	20.2	56.9	34	0.7	19.5±2.9	19.9±2.9
RB1-10	0.16	1.82	362	0.97	0.85	2.03×10 ⁻³	0.020	0.009	0.015	2.31±0.05	23.2	11.2	17.5	26	0.5	17.1±2.6	17.6±2.6
Pseudo-	euhedral r	nagnetite															
RB1-8	0.18	2.12	477	0.98	0.94	3.23×10 ⁻³	0.016	0.012	0.020	3.26±0.07	17.3	12.3	21.9	20	0.7	31.3±4.7	32.0±4.8
RB1-11	0.13	1.80	466	0.98	0.71	1.60×10 ⁻³	0.008	0.007	0.004	2.39±0.05	11.7	9.7	6.2	14	0.8	29.7±4.5	30.4±4.6

Table 3: CT-scan and (U-Th-Sm)/He results of the magnetite-bearing chloritoschist RB1 sample

Rs and F_T are the sphere equivalent radius and alpha ejection factor.

(U-Th-Sm)/He date has been corrected from alpha ejection factor (F_T) (Ketcham et al., 2011).

1. Thermochron	ologic data						
Complex and data used	AFT and ZFT data from Schwartz et al. (2007)						
samples and data used	MgHe data from this paper: Table 2, samples RB1-3, 4, 5, 7, and 10						
Data treatment: Dates, unce	rtainties, and other relevant constraints						
	Kinetics: Apatite assumed to have CI=0 apfu						
AFTUALA	Uncertainty model: Asymmetric 95% bounds						
	Initial mean track length: N/A; no track length data available						
<u>ZFT data</u>	Uncertainty model: Asymmetric 95% bounds						
<u>MgHe data</u>	Uncertainty model: Uniform 15% (1SE) uncertainty						
2. Additional ge	ologic information and constraint placement						
Present-day temperature	20°C						
Initial constraint	Assumed peak HP-LT conditions: 350±20°C at 58±4 Ma						
	HeFTy code E3/50: Paths episodic, halved 3 times, with maximum cooling rate						
I-t path complexity	50°C/m.y.						
3. System- and r	nodel-specific parameters						
AFT annealing model: Ketcha	m et al. (1999)						
ZFT annealing model: Ketcha	m (2019)						
MgHe diffusion parameters:	Blackburn et al. (2007)						
Modeling code: HeFTy v1.9.3							
Statistical fitting criteria: HeF	Ty default						
Ending condition: Models rur	n until 50 good fits found						

Table 4: Thermal History Inverse Modeling Parameters