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► To cite this version:

Stéphane Schwartz, Cécile Gautheron, Richard Ketcham, Fabrice Brunet, Marianna Corre, et al.. Unraveling the exhumation history of high-pressure ophiolites using magnetite (U-Th-Sm)/He thermochronometry. *Earth and Planetary Science Letters*, 2020, 543, pp.116359 (IF 5,273). <10.1016/j.epsl.2020.116359>. <hal-02911613>

HAL Id: hal-02911613

<https://hal.science/hal-02911613v1>

Submitted on 3 Jun 2022

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1 **Unraveling the exhumation history of high-pressure ophiolites using**
2 **magnetite (U-Th-Sm)/He thermochronometry**

3
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19
20 To be submitted to EPSL

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22
23 **Keywords:** Thermochronometry; (U-Th-Sm)/He; Magnetite; Modeling; High-pressure
24 ophiolites; Exhumation

25 **Abstract**

26 Magnetite is a ubiquitous oxide in ultramafic and mafic rocks, which is present in a large
27 range of geological and tectonic settings. In the case of high-pressure ultramafic rocks,
28 exhumation timing is commonly constrained by geochronometers in nearby lithologies. The
29 development of the magnetite (U-Th-Sm)/He method (MgHe) has opened new perspectives
30 to refine the exhumation history of such rocks. However no thermal history has been
31 deduced using MgHe data so far. Here we applied MgHe dating to magnetite from an HP-LT
32 alpine ophiolite body (Rocher Blanc) from the Schistes lustrés units (Western Alps, France)
33 where part the P-T-t history is constrained to further refine the thermal history and
34 demonstrate validity of the method. Textural and geochemical characterizations of the
35 magnetite grains reveal two crystallization environments leading to distinct grain
36 morphologies, euhedral and pseudo-euhedral, with different amounts of mineral inclusions
37 (titanite, chlorite and augite-aegirine). Both magnetite types crystallized during the same
38 retrograde path at $T > 250^{\circ}\text{C}$, under hydrothermal conditions shown by Si and V oscillatory
39 zoning and low Ti content. MgHe ages derived from euhedral magnetite crystals range from
40 14.8 ± 2.1 to 20.9 ± 3.0 Ma, between zircon and apatite fission track (ZFT) and (AFT) ages
41 deduced from the literature. Pseudo-euhedral grains display older ages interpreted as
42 resulting from the implantation of He from U-rich neighboring minerals such as titanite. For
43 the first time, MgHe data are inverted along with AFT and ZFT data to refine the thermal
44 history of an ophiolite related to their exhumation. The MgHe data suggest that the Rocher
45 Blanc sample has undergone an exhumation in ductile conditions related to subduction
46 dynamics and underplating processes, followed by an exhumation reflecting denudation by
47 erosion under brittle conditions. This exhumation regime change is found to take place at
48 ~ 20 Ma, as opposed to a less constrained ~ 26 Ma cooling as would be inferred from the ZFT

49 and AFT ages alone. Considering published thermochronometric data along the Western
50 Alps, we propose that the exhumation is controlled by the upward indentation of the
51 Apulian mantle producing regional tilting and associated erosion. With the present Schistes
52 lustrés ophiolite example, we demonstrate that the MgHe method can be used as a low-
53 temperature thermochronometer (closure temperature in the 200-250°C range) that aids in
54 unraveling the cooling history associated with exhumation of high-pressure ultramafic and
55 mafic rocks at the brittle to ductile transition.

56

57 **1. Introduction**

58 Ultramafic and mafic rocks are widely distributed on the Earth's surface and are associated
59 with a large range of geological and geodynamic settings, from oceanic accretion to orogenic
60 zones (e.g., [Guillot et al., 2015](#)). In contrast to their abundance, mafic and ultramafic rocks
61 only yield a few minerals suitable for geo-thermochronometry such as U-Th-bearing
62 accessory minerals (zircon, titanite or apatite). However, recent advances in the analysis of
63 radiogenic isotopes have extended the application of (U-Th)/He geo-thermochronology to
64 minerals incorporating U and Th at the ppb level such as magnetite ([Blackburn et al., 2007](#);
65 [Cooperdock and Stockli, 2016](#)) or spinel ([Cooperdock and Stockli, 2018](#)). Magnetite (U-
66 Th)/He geo-thermochronometry, abbreviated MgHe below, opens new perspectives with
67 respect to the use of basic and ultrabasic rocks for dating geodynamic processes.

68 Indeed, magnetite ($\text{Fe}^{2+}\text{Fe}^{3+}_2\text{O}_4$) is ubiquitous in these rock types where it can either
69 crystallize from the parent magma or precipitate later in the rock history from hydrothermal
70 fluids at lower temperature. The MgHe method is particularly attractive for dating
71 geodynamic processes involving serpentinite formation, metamorphism or exhumation.

72 Magnetite is produced during serpentinization of the lithospheric mantle at slow-spreading

73 ridges for $T > 200^\circ\text{C}$ (Klein et al., 2013; Klein et al., 2014) and controls thereby hydrogen
74 production (e.g., see Brunet, 2019 for a review). Magnetite in subducted serpentinite can
75 survive temperature conditions up to the antigorite breakdown (Debret et al., 2014;
76 Merkulova et al., 2017). Therefore, in addition to constraining the timing of mafic and
77 ultramafic rocks exhumation, dating of magnetite by (U-Th-Sm)/He geochronometry can also
78 put time constraints on the functioning of fossil hydrothermal systems and on serpentinite
79 metamorphism in subduction-collision zones.

80 The MgHe method is based on radiogenic ^4He production and accumulation in the crystal
81 lattice during alpha decay of U, Th, and Sm isotopes (^{238}U , ^{235}U , ^{232}Th and ^{147}Sm) that lead to
82 stable isotopes (^{206}Pb , ^{207}Pb , ^{208}Pb and ^{143}Nd respectively). The alpha particle is ejected with
83 a high kinetic energy of several MeV and travels a mean distance of $\sim 15\ \mu\text{m}$ in the magnetite
84 lattice (Ketcham et al., 2011). A correction for He loss due to alpha ejection (called F_T) needs
85 to be applied which can range from few percent to 20-30% for smaller grains with sizes of
86 $\sim 100\ \mu\text{m}$ (Farley et al., 1996). Otherwise, abrasion of the first $\sim 15\ \mu\text{m}$ of the crystals or
87 selection of the inner parts of a large crystal can be done to avoid ejection or implantation
88 from neighboring minerals (Blackburn et al., 2007; Cooperdock and Stockli, 2016). The
89 remaining He atoms in the crystal can diffuse outside of the crystal with a rate of diffusion
90 that is a function of the temperature and crystal volume. Based on experimental He diffusion
91 coefficients obtained on magmatic magnetite by Blackburn et al. (2007), the closure
92 temperature T_c (Dodson, 1973) is as $250 \pm 40^\circ\text{C}$ for a crystal size of $500\ \mu\text{m}$ and a cooling rate
93 of $10^\circ\text{C}/\text{Ma}$. The T_c for He in magnetite falls in the same range as for the zircon fission track
94 (ZFT) system, which is broadly estimated to be $\sim 240 \pm 50^\circ\text{C}$ (Gallagher et al., 1998). Ketcham
95 (2019) derived a T_c value of 281°C based on fitting experimental data sets from relatively

96 young zircons (Murakami et al., 2006), and Rahn et al., (2004) showed that radiation damage
97 may lower annealing resistance.

98 The number of studies dealing with the MgHe method remains limited. MgHe has been
99 applied to volcanic rocks (Blackburn et al., 2007), ore deposits (Fanale and Kulp, 1962) as
100 well as serpentinites and chloritoschists (Cooperdock and Stockli, 2016). These studies have
101 clearly demonstrated the relevance of the MgHe thermochronometer to obtain cooling or
102 crystallization ages. However, MgHe data have never previously been used to infer the
103 thermal history of rocks using inverse modeling, which is the main purpose of collecting
104 thermochronometric data.

105 Here, the MgHe method is applied to an ophiolite from the Schistes lustrés units (Western
106 Alps, France), for which the cooling history is relatively well constrained, as the mineralogy
107 and P-T history of the area have been characterized (Pognant and Kienast, 1986; Tricart and
108 Schwartz, 2006) and ZFT and AFT data are available (Schwartz et al., 2007). The obtained
109 MgHe data are critically evaluated in the framework of the ophiolite thermal history
110 constrained by the ZFT and AFT literature data.

111

112 **2. Geological setting**

113 The studied samples were collected in the internal domain of the western Alps (Fig. 1A) that
114 underlies the boundary between the European and African plates (e.g., Lemoine et al., 1986).

115 This domain (Fig. 1B) derives from a part of the Tethyan ocean (Piedmont zone) and its
116 northwestern European margin (Briançonnais zone). In the study area, the Piedmont zone is
117 composed of the Schistes lustrés units and the Monviso ophiolite, which rest on the internal
118 crystalline massif of Dora Maira (Fig. 1C). These units were juxtaposed during subduction
119 and collision processes in Late Cretaceous to Late Tertiary times. The Piedmont zone

120 represents different levels of a paleo-subduction zone, preserving the low thermal gradient
121 (5–8°C/km) associated with the subduction dynamics (Schwartz et al., 2001; Agard et al.,
122 2002). In this context, the Schistes lustrés units correspond to a fossiliferous sedimentary
123 accretionary wedge metamorphosed under high pressure and low temperature (HP-LT)
124 conditions during the late Cretaceous–early Eocene subduction of the Tethyan ocean (Lafay
125 et al., 2013). This domain is derived from Mesozoic oceanic sediments that were originally
126 composed of marls, clays, and limestones associated with fragments of oceanic lithosphere.
127 These sediments were strongly deformed and metamorphosed during subduction, and they
128 crop out today as foliated and polydeformed calcschists enclosing boudinaged meter- to
129 kilometer-sized Jurassic ophiolites. The P-T conditions increase toward the east, from low-
130 temperature blueschist (LT-BS, Fig. 1C) to high-temperature blueschist facies conditions (HT-
131 BS, Fig. 1C) (Tricart and Schwartz, 2006). The blueschist domain is structurally located above
132 the Monviso eclogitic ophiolite (Fig. 1C). In this massif, the metasedimentary component is
133 small (<20 vol.%) relative to the voluminous oceanic lithosphere (Schwartz et al., 2001).

134 The P-T evolutions of the Schistes lustrés units are well documented and show clockwise
135 loop paths, keeping a record of the early prograde evolution, and the retrograde
136 metamorphism stages have remained in the field typical of greenschist paragenesis (Fig. 1D),
137 with no temperature increase (Tricart and Schwartz, 2006). For the studied Rocher Blanc
138 ophiolite located in the central part of the Schistes lustrés units (Fig. 1C), the metamorphic
139 peak conditions are estimated at 10 ± 2 kbar and $>350^\circ\text{C}$ (Tricart and Schwartz, 2006,
140 Schwartz et al., 2013), with laser probe Ar-Ar ages ranging from 62 to 55 Ma (Agard et al.,
141 2002). The final exhumation is constrained by apatite and zircon fission track (AFT-ZFT)
142 analysis (Schwartz et al., 2007) (Fig. 1C), which reveals a diachronous exhumation of the HP-
143 LT metamorphic rocks from east to west during Oligocene to Miocene time. Taking into

144 account the cooling rates from [Schwartz et al. \(2007\)](#), the Rocher Blanc ophiolite is
145 considered to have passed through the ~240-280°C isotherm at 24.8±1.6 Ma and the ~110°C
146 isotherm at 12.3±2.0 Ma (Fig. 1C).

147

148 **3. Samples and methods**

149 **3.1. Sampling and petrologic characterization**

150 Sample RB1 was collected from the Rocher Blanc massif (RB; WGS84 UTM31N, 6°57'08.01"E,
151 44°39'53.25"N; elevation: 2810 m a.s.l.) that corresponds to a km-scale ophiolitic body
152 composed of mafic and ultramafic rocks embedded in the Schistes lustrés units (Fig. 1C) that
153 record middle-temperature blueschist facies conditions (MT-BS) ([Pognant and Kienast, 1986](#);
154 [Tricart and Schwartz, 2006](#)). The high-pressure blueschist paragenesis in mafic rocks
155 (glaucofan, zoisite) has been partially to totally retrogressed at greenschist facies
156 conditions (chlorite, titanite and augite-aegirine). In ultramafic rocks, high pressure is
157 characterized by the presence of antigorite ([Schwartz et al., 2013](#)). The studied sample (RB1)
158 corresponds to a magnetite-bearing chloritized Fe-Ti meta-gabbro (chloritoschist), some tens
159 of centimeters thick, located at the interface between meter-size lenses of Fe-Ti gabbro and
160 serpentinite (Fig. 2A). The sample contains visible mm-scale euhedral magnetite crystals as
161 shown in Figure 2B, which are superimposed over the main schistosity defined by the
162 chlorite – titanite assemblage (Fig. 2C).

163

164 **3.2. Mineralogical and geochemical analysis**

165 The chemistry of both euhedral and pseudo-euhedral magnetite crystals of the RB1 sample
166 were characterized at the microscale. Semi-quantitative analyses and imaging of five
167 magnetite and their inclusions were performed with quantitative analyses and mapping of

168 minor and trace elements in magnetite (Si, Ti, Cr, V, Mn and Mg). Several SEM images are
169 shown in Figures 2D, E, F, and concentrations of Si, Ti, Sr, V, Mn, and Mg obtained on
170 magnetite crystals are reported in Table 1. Figure 3 presents the Si and V contents of one
171 euhedral and one pseudo-euhedral magnetite crystal identified also in Figure 2C. Additional
172 information can be found in the supplementary section.

173 In-situ uranium and thorium concentrations in chlorites, titanites and magnetites from
174 sample RB1 were quantified by laser ablation inductively coupled plasma mass spectrometry
175 (LA-ICPMS). Results are presented in Table 2 and on Figure 4. Nevertheless, 2σ internal
176 errors are reported in Table 2 and additional details can be found in the supplementary
177 section.

178 179 **3.3. Crystal selection using CT scanning**

180 Magnetite grains scans were performed following guidelines by Cooperdock and Stockli
181 (2016, 2018), and additional details can be found in the supplementary section. The blob3D
182 software (Ketcham, 2005a; Cooperdock et al., 2019) was used to determine precisely the
183 volume, surface/volume ratio, sphere-equivalent radius and alpha-ejection factor F_T
184 (Ketcham et al., 2011) for each grain, and the presence of large inclusions ($>5 \mu\text{m}$). 3D
185 images of the magnetite crystals were generated to aid in selecting crystals for (U-Th-Sm)/He
186 dating and documenting their euhedral and pseudo-euhedral shapes. 3D grain shape images
187 of MgHe-analyzed magnetite crystals are reported in Figures 5A and B.

188

189 **3.4. (U-Th-Sm)/He on magnetite**

190 Non-abraded single mm-size euhedral and pseudo-euhedral magnetite crystals were
191 degassed under-vacuum in order to determine the He content using the protocol described

192 in Allard et al. (2018). He content is corrected from the blank level, which is from 40 to 160
193 times lower than the He signal. Helium content error is estimated to 2% accordingly. The wet
194 chemistry digestion protocol was adapted from Blackburn et al. (2007) allowing the
195 determination of the ^{238}U , ^{232}Th , ^{147}Sm content. Additional details are placed in the
196 supplementary section. Analytical uncertainties in the U, Th and Sm content translate into a
197 15% error on alpha-ejection-corrected MgHe ages. Sample mass, volume, He, U, Th, Sm
198 content, raw and alpha-ejection-corrected (U-Th-Sm)/He ages are reported in Table 3.

199

200 **4. Results**

201 **4.1. Magnetite phase relationships and chemical composition**

202 At the thin section scale, magnetite crystallizes in two zones of distinct mineralogy, which
203 are likely inherited from the initial distribution of magmatic minerals in the Fe-Ti gabbro (Fig.
204 2C). The mineralogy of these two zones is either dominated by a chlorite + magnetite \pm relict
205 magmatic pyroxene (Zone 1) or titanite + magnetite \pm chlorite \pm relict magmatic ilmenite
206 (Zone 2) assemblages. In Zone 1, the augite-aegirine pyroxene may stably coexist with
207 titanite and chlorite contains a 10-20 % jadeitic component. In Zone 2, titanite, magnetite
208 and chlorite appear texturally as destabilization products of magmatic ilmenite and
209 generates local reaction coronas (Fig. 2D). Plagioclase has not been identified in either of the
210 two zones. Magnetite crystal habits differ between the two zones. In Zone 2, magnetite is
211 pseudo-euhedral to anhedral, possibly polycrystalline, and contains numerous inclusions (Fig.
212 2D-E) of titanite and chlorite. In Zone 1, magnetite is euhedral, and can form larger grains
213 exceeding 500 μm with rare inclusions (Fig. 2F) of augite-aegirine, chlorite and titanite.

214 Chemical characterization of euhedral magnetite crystals from Zone 1 and pseudo-euhedral
215 magnetite crystals from Zone 2 are reported in Table 1. Significant Mn, Cr and Ti contents

216 are present but remain below 500 ppm (Table 1). Mg is below detection limit (i.e., <25 ppm),
217 while Si and V are the two most concentrated cations after Fe, with mean concentrations of
218 809 ± 513 and 1821 ± 265 ppm, respectively (Table 1). The Si and V distribution patterns are
219 the same in euhedral and in pseudo-euhedral magnetite grains indicating similar growth
220 process and history. It is characterized by oscillatory zoning with a relatively good correlation
221 between Si and V-rich growth zones (Fig. 3), although vanadium zoning is sharper. The
222 nature of the inclusions and their distribution seem independent of the chemical variations
223 recorded by these two elements.

224

225 **4.2. Magnetite, chlorite and titanite in-situ U content**

226 The U content of magnetite derived from laser ablation ranges from 5 ± 0.4 to 24 ± 16 ppb and
227 is ~five times higher in the rim compared to the core (Table 2, Figure 4). In addition, the U
228 content is similar to the U value obtained by wet chemistry for the dated crystals with value
229 of 8 ± 1 to 29 ± 2 ppb (Table 3). The U content is also similar between magnetite from Zone 1
230 (euhedral) and Zone 2 (pseudo-euhedral) (Table 2 and 3). The U content in chlorite from
231 Zone 1 of $\sim 16\pm 8$ ppb is lower than the content in chlorite from Zone 2 which ranges from
232 42 ± 5 to 68 ± 10 ppb (Table 2, Figure 4). Although variable, the U content of titanite is by far
233 the highest with values ranging from 27 ± 5 to 1569 ± 141 ppb. In the RB1 sample titanite is the
234 U-bearing phase.

235

236 **4.3. Magnetite (U-Th-Sm)/He**

237 Five euhedral magnetite crystals (Fig. 5A) and two pseudo-euhedral grains (Fig. 5B) were
238 selected for MgHe dating, utilizing the CT scan data to select grains with the fewest (for
239 pseudo-euhedral magnetite) or no inclusions. The selected crystals have equivalent-sphere

240 radii (Rs) ranging from 362 to 483 μm (Table 3). Helium content of these seven magnetite
241 grains ranges from 0.9 to 3.3 pmol/g and effective uranium ($\text{U}+0.24\text{Th}$) ranges from 9 to 34
242 ppb (Table 3). Ejection-corrected MgHe ages for euhedral and pseudo-euhedral grains range
243 from 14.8 ± 2.1 to 20.9 ± 3.0 Ma and 30.4 ± 4.4 to 32.0 ± 4.6 Ma, respectively (Fig. 5C).

244

245 **5. Discussion**

246 **5.1. Conditions of magnetite formation**

247 Magnetite is a common mineral in ultramafic and mafic rocks and appears over a large range
248 of P-T and $f\text{O}_2$ conditions (Nadoll et al., 2014). In the case of magnetite created during the
249 oceanic serpentinization process, it is classically assumed to form at temperatures above
250 $\sim 200^\circ\text{C}$ (e.g., Alt et al., 2013; Klein et al., 2013; Klein et al., 2014), even if magnetite could
251 potentially crystallize at temperatures even below 120°C (Alt et al., 2013). In the case of
252 subducted ophiolites (serpentinite and gabbro lithologies), magnetite is observed from
253 eclogite paragenesis (Li et al., 2004; Debret et al., 2014) up to greenschist facies conditions
254 related to the retrograde stage for temperatures ranging from 550 to 250°C (Schwartz et al.,
255 2013).

256 The petrological and geochemical studies of the RB1 sample indicate that the euhedral and
257 pseudo-euhedral magnetite grains crystallized during the retrograde path under greenschist
258 facies conditions assisted by metamorphic hydrothermal fluids and are not derived from
259 magmatic origin. Indeed, the Ti content of magnetite crystals is under 100 ppm (Table 1), too
260 low to be of magmatic origin (Nadoll et al., 2014), and the host inclusions (titanite, chlorite,
261 augite-aegirine) that are representative of HP-LT paragenesis demonstrate that magnetite
262 grains crystallized during the greenschist retrograde path. In this context, the oscillatory V
263 and Si chemical zoning, with a Si-richer rim observed for euhedral and pseudo-euhedral

264 magnetite grains (Figure 3) record the changing conditions and compositions of the
265 metamorphic fluid related to magnetite growth (Dare et al., 2014). The Si and V are the main
266 minor elements at the ~1000 to 2000 ppm level, compared to Ti, Cr, Mn, and Mg (Table 1);
267 these Si and V contents reveal high solubility indicating crystallization from a high
268 temperature fluid (Dare et al., 2014). Such high Si content has already been encountered in
269 hydrothermal systems (skarns, metamorphic veins; e.g., Shimazaki, 1998) and Si oscillatory
270 zoning has been described in hydrothermal magnetite (Deditius et al., 2018). Hydrothermal
271 conditions (250-350°C, Fig. 6A) in the fluid-circulation zone marked by the magnetite-bearing
272 chloritoschists may have favored Si and V incorporation, the latter element being supplied
273 locally by the Fe-Ti gabbro itself.

274 Based on previous thermobarometric studies of the host metamorphic Fe-Ti gabbro (e.g.,
275 Pognante and Kienast, 1986), and petrological investigation in this study, the magnetite
276 begins to crystallize most probably near the blueschist – greenschist transition, as residual
277 augite-aegirine is found in inclusions, and magnetite crystallization continues during the
278 entire greenschist retrograde path (Fig. 6B).

279

280 **5.2. Magnetite He age distribution**

281 The MgHe ages for all crystals are younger than the high-pressure ages recorded for the
282 Schistes lustrés units (62-54 Ma; Agard et al., 2002), and potentially represent exhumation
283 ages, i.e., related to the retrograde path. Interestingly, the euhedral grains are quite
284 homogeneous with a mean age of 18.1 ± 2.4 Ma that falls between the AFT and ZFT ages as
285 deduced from Schwartz et al. (2007) for the Rocher Blanc area (Fig. 5C), and are thus
286 compatible with cooling ages associated with the exhumation history of the area. Indeed,
287 the MgHe ages are younger than ZFT ages, and older than AFT ages, in agreement with the

288 published thermal sensitivity of MgHe ($250\pm 50^\circ\text{C}$; Blackburn et al., 2007), ZFT ($240\pm 50^\circ\text{C}$,
289 Gallagher et al., 1998; $\sim 281^\circ\text{C}$, Ketcham, 2019) and AFT ($110\pm 10^\circ\text{C}$, Gallagher et al., 1998).
290 Conversely, the ages obtained on pseudo-euhedral crystals are significantly older than the
291 euhedral crystal and ZFT ages (Fig. 5C). This is not consistent with petrological data which
292 show that euhedral and pseudo-euhedral magnetite all crystallized simultaneously under
293 greenschist conditions at $T > 250^\circ\text{C}$. Thus, we interpret these ages as related to the presence
294 of excess He compared to euhedral crystals with similar size and eU content (Table 3); they
295 cannot be considered as cooling ages. Excess of He may be due to (i) the impact of alpha
296 implantation from neighboring minerals or circulating rich U-Th fluids, and/or (ii) the impact
297 of inclusions with higher U-Th content on the global He budget of magnetite.

298 Alpha implantation has been frequently observed or modeled for the (U-Th)/He method,
299 especially for apatite (e.g., Spiegel et al., 2009; Gautheron et al., 2012) and magnetite
300 (Blackburn et al., 2007; Cooperdock and Stockli, 2016). When minerals contain lower or
301 equivalent U and Th content compared to neighboring phases, the alpha implantation effect
302 can be significant. This is particularly the case for magnetite, which typically contains low U-
303 Th content (< 30 ppb) compared to minerals with high U-Th content of > 100 ppb like titanite
304 in the present case (Table 2). In Zone 1, magnetite is embedded in chlorite which is U-poor
305 with contents ~ 15 ppb. It must be noted that the higher U content in Zone 2 results in a
306 higher U content, from 40 to 60 ppb, in chlorite as well (Table 2, Figure 4). In the RB1 sample,
307 for the chlorite area (Zone 1) alpha implantation in magnetite should be negligible whereas
308 in the titanite area (Zone 2) alpha implantation in magnetite could be significant. As the RB1
309 grains are large ($R_s > 400 \mu\text{m}$) with small alpha ejection corrections ($F_T > 0.97$; Ketcham et al.,
310 2011), a 20-40 ppb matrix would simply cancel out the F_T correction, and only thousands of
311 ppb U-Th will have an age effect approaching 10% (Gautheron et al., 2012). For euhedral

312 magnetite, MgHe ages younger than ZFT imply no significant alpha implantation and
313 reinforce the hypothesis that euhedral magnetites crystallized in Zone 1 (chlorite-rich
314 environment). Similarly, MgHe ages older than ZFT observed for the pseudo-euhedral
315 magnetite may reflect implantation from U-Th-rich neighboring minerals such as titanite,
316 consistent with the pseudo-euhedral magnetites crystallizing in Zone 2, in agreement with
317 petrological observations.

318 Inclusions richer in U-Th than the host minerals can also impact (U-Th)/He ages by adding He
319 to the crystal lattice, as already reported for apatite (Farley, 2002) and magnetite (Blackburn
320 et al., 2007; Cooperdock and Stockli, 2016). An excess of He can occur if the total U-Th
321 content of host mineral and inclusions is incompletely retrieved by the chemical digestion
322 protocol that is usually designed to target the host mineral. For magnetite, the digestion
323 protocol is not adapted to dissolve completely silicates (e.g., chlorite, titanite, zircon) or
324 other oxide minerals (e.g., titanite). A small fraction of the U-Th budget coming from
325 inclusions may be retrieved, but the total U-Th budget is incomplete, leading to spuriously
326 older MgHe ages with no geological significance. For the two pseudo-euhedral magnetites
327 from RB1, the MgHe ages older than the euhedral magnetites with similar U-Th content
328 (Table 2) would be consistent with the absence of U-Th-rich inclusions in the euhedral
329 crystals. However, because the CT scan analysis indicated minimal detectable inclusions, and
330 because the primary U-Th-bearing phase, titanite, has eU compositions only ~100x higher
331 than the magnetite (Table 2, 3), we do not believe inclusions can explain the magnitude of
332 excess He. We thus interpret that MgHe ages older than ZFT reflect implantation across
333 large areas of grain boundaries shared with titanite (Fig. 4D), and for the rest of the
334 discussion we do not use or further discuss the two pseudo-euhedral crystals, as their
335 measured ages are not geologically meaningful. However, we can conclude that the MgHe

336 ages from the euhedral crystals of the RB1 sample are not significantly impacted by
337 implantation and reflect cooling due to exhumation along the retrograde path.
338 Finally, the results of this study give some methodological insight into MgHe dating. First, CT
339 imaging is essential for detecting the presence of mineral inclusions larger than $\sim 5 \mu\text{m}$, as
340 already stated by [Cooperdock and Stockli \(2016\)](#). Comparison of U-Th content for crystals
341 with older MgHe ages can provide indirect evidence of the impact of rich U-Th inclusions
342 below the CT-scan detection limit ($< 1 \mu\text{m}$). Second, the removal of the $20 \mu\text{m}$ outer margin
343 of the magnetite crystals, as recommended by [Blackburn et al. \(2007\)](#), is not always
344 necessary. In fact, the result of this study shows that implantation is not always significant
345 and depends on the petrological environment of magnetite (i.e. chlorite versus titanite
346 zones), as well as the size of the crystal. Furthermore, alpha implantation along long grain
347 boundaries shared with titanite, as documented in our zone 2 non-euhedral magnetites, if
348 combined with diffusional migration of helium into the magnetite grain (Gautheron et al.
349 2012) may render abrasion an insufficient remedy, depending on cooling history. Thorough
350 petrological characterization of mineral textures and compositions thus appears to be a
351 necessary step of the MgHe method.

352

353 **5.3. Origin of MgHe age dispersion**

354 Significant MgHe age dispersion is encountered for euhedral RB1 magnetites (MgHe age =
355 $18.1 \pm 2.4 \text{ Ma}$, 13% at 1σ). MgHe age dispersion has also been reported in previous work in
356 the HP-LT subduction complex on the island of Syros, Greece ([Cooperdock and Stockli, 2016](#)).
357 This age dispersion is higher than typical analytical error (i.e. $\sim 5\text{-}10\%$, for apatite and zircon)
358 but similar or lower than frequently found in practice for apatite or zircon (U-Th)/He
359 methods (e.g., [Ault et al., 2019](#); [Ketcham et al., 2018](#)). Using investigations already published

360 on the origin of apatite or zircon (U-Th)/He age dispersion, some factors affecting MgHe
361 ages related to the impact of (i) U-Th zonation, (ii) slight alpha-implantation, or (iii) crystal-
362 dependent He diffusion coefficient can be anticipated.

363 First, a slight impact of alpha-implantation could affect the MgHe age, as the U-Th content of
364 the neighboring minerals could be slightly different. For example, U and Th content for
365 chlorite of ~15 ppb has been determined for Zone 1 (Table 2), but slightly higher U-Th values
366 are possible (Table 2). However, the large crystal size of the magnetite from this study
367 reduces the impact of implantation. Nevertheless, other parameters can account for
368 dispersion. For instance, [Cooperdock and Stockli \(2016\)](#) found similar MgHe age dispersion
369 on abraded crystals (i.e. mechanical removal of the 20 μm border of the crystal that can
370 have been affected by alpha implantation).

371 Second, due to the low U-Th content, analytical determination of zoning is difficult, and up
372 to a ~4x higher U concentration in the rim has been measured for three magnetite crystals
373 (Table 2). Because the ablation spot is large (150 μm), the exact size of the U-Th-enriched
374 rim is not known but can be similar to the one observed by Si-V analysis, i.e. ~100 μm (Fig. 3).
375 In this case, the presence of a ~100 μm enriched U-Th rim will not have much impact on the
376 ejection factor as it is much larger than the mean stopping distance (15 μm). If finer zoning
377 exists, modeling can be performed (e.g., [Farley et al., 1996](#); [Gautheron et al., 2012](#)) assuming
378 that U-Th zoning follows Si and V zoning (Fig. 3C and D). Using the Monte Carlo code
379 associated with [Gautheron et al. \(2012\)](#), we modeled a 400- μm radius sphere with a 50 μm
380 rim depleted or enriched in U-Th by 50%, and as a worst-case scenario a rim containing all of
381 the crystal's U-Th (i.e., no U-Th in the core). We obtained F_T ejection factors of 0.99, 0.96,
382 and 0.76, respectively, implying a 25% MgHe age difference for the worst case. However, in
383 comparing our bulk and in situ core and rim analyses we find no evidence of such severe

384 zoning in U-Th, and we can estimate a maximum MgHe age dispersion component of only up
385 to a few percent. Given that implantation can be from minerals with much more than 50%
386 higher eU, it is likely to be a source of equal or greater dispersion.

387 Third, a defect- and/or composition-dependent He diffusion coefficient could also be
388 invoked to explain age dispersion, as has been proposed and modeled for other minerals
389 such as apatite and zircon (Gautheron et al., 2009; Flowers et al., 2009; Guenther et al.,
390 2013). Indeed, damage due to radioactivity, and also crystal defects could trap or facilitate
391 He diffusion within the crystal structure. A damage content of some 5% would be required
392 to change the He diffusion coefficient significantly, as modeled for apatite (e.g., Gerin et al.,
393 2017), which is unlikely given the low eU content of magnetite (<<1%). Alternatively,
394 substitution by minor or trace elements could deform the crystal structure and thus modify
395 He diffusion, as shown in apatite at the atomic scale (Djimbi et al., 2015). The magnetite
396 crystals from RB1 have ~0.5 wt% of minor elements, and if we assume similar effect as for
397 apatite (Djimbi et al., 2015), no significant change in the closure temperature compared to
398 pure magnetite is expected. This is highly speculative, as the composition of the volcanic
399 magnetite used by Blackburn et al. (2007) is not known. The effect of minor elements in
400 magnetite on He volume diffusion still needs to be investigated. It is also possible that fluid
401 inclusions can act as traps for He, as has been proposed for apatite and zircon (Zeitler et al.
402 2017; Danišik et al., 2017); this possibility could be further pursued with higher-resolution
403 imaging and/or ramped heating experiments (McDannell et al., 2018).

404 Finally, we stress that there is currently only one experimental study documenting He
405 diffusion kinetics in magnetite (Blackburn et al., 2007), and even that data set shows signs of
406 complex release behavior. Further data are certainly required to build a more robust
407 database of kinetic parameters and explore possible anomalous behaviors.

408

409 **5.4. MgHe data inversion: resolving exhumation rate from the Rocher Blanc ophiolite**

410 The petrological and geochemical arguments deduced from the RB1 sample analysis indicate
411 that magnetite crystallized at $T > 250^{\circ}\text{C}$, as illustrated in Figure 6, at temperature higher than
412 the MgHe closure temperature (T_c), and thus that MgHe ages obtained from euhedral
413 magnetite of ca. 18.1 ± 2.4 Ma correspond to cooling ages of the Rocher Blanc ophiolite (Fig.
414 6). Specifically, for the dated RB1 magnetites, which all have similar sizes (Table 3), T_c ranges
415 from 242 to 248°C for sphere-equivalent radii of 362 to $483 \mu\text{m}$ and a cooling rate of
416 $10^{\circ}\text{C}/\text{Ma}$. MgHe ages derived from are sandwiched between ZFT and AFT ages as expected
417 by their thermal sensitivity, and in fact shed further light on the thermal sensitivity of the
418 ZFT method, supporting a ZFT $T_c > 260^{\circ}\text{C}$. Figure 6 summarizes the P-T path for RB1 sample,
419 including the magnetite growth field during the retrograde path and the time constraint
420 obtained from AFT, ZFT and MgHe ages.

421 **5.4.1. MgHe data inversion**

422 The thermal history of the studied area was retrieved using the HeFTy software (Ketcham,
423 2005b) by inverse modeling of MgHe, AFT and ZFT low temperature thermochronometric
424 data. HeFTy uses a constrained Monte Carlo scheme that poses a large number of candidate
425 time-temperature (t-T) paths, and retains those that fit the thermochronometric data to a
426 “good” (closer than statistical expectation for a random sample) or “acceptable” (does not
427 fail a 95% confidence test) extent. For the inversions the fission-track annealing models by
428 Ketcham et al. (1999) and Ketcham (2019) were used for AFT and ZFT data, respectively. For
429 MgHe ages, the sphere-equivalent radii (R_s) of each crystal and He diffusion parameters
430 from Blackburn et al. (2007) were used. In detail, the HeFTy software models (i) the fission-
431 track production and annealing in apatite and zircon, and (ii) helium production, taking into

432 account ejection and diffusion, in magnetite. In this way, the data inversion identifies a
433 family of potential cooling histories by using 17 t-T nodes randomly placed to define a
434 cooling-only path with a maximum allowed cooling rate of 50°C/m.y. over any segment. In
435 addition to the MgHe, AFT and ZFT data, the peak HP-LT metamorphic conditions for the
436 Rocher Blanc ophiolite massif as reported in Figure 6, estimated to be 350±20°C at 58±4 Ma
437 (Agard et al., 2002) is used as a starting point. Further model inputs are listed in Table 4.
438 Thermal history modeling was performed with AFT-ZFT alone, with MgHe data alone and
439 with all three combined (AFT-ZFT and MgHe). First, modeling was done with interpolated
440 AFT and ZFT ages for RB1 sample assuming that ages vary linearly with the sample position
441 on a W-E transect (Schwartz et al., 2007) (Fig. 1), and with and without MgHe data. Second,
442 inversions were also performed with AFT and ZFT data from RN and CG samples, that
443 represent the available data geographically framing the RB ophiolite (Fig. 1C), and with and
444 without MgHe data. Finally, for comparison, the AFT-ZFT data and the MgHe data separately
445 for RB1 sample were inverted. For modeling of MgHe data, all euhedral grains were
446 individually included. The computed thermal histories are reported on Figure 7. In the on-
447 line supplement, Figure S1 presents the modeling inversion results obtained for AFT and ZFT
448 data of CG and RN samples.

449

450 **5.4.2. Thermal history modeling**

451 The inversion of the MgHe data alone or with the AFT and ZFT literature data allows us to
452 refine the final t-T exhumation history of the Rocher Blanc ophiolite massif, while also
453 verifying good consistency with AFT and ZFT data. First, model outputs showing the set of
454 good and acceptable fits, and their weighted mean, demonstrate that the MgHe data
455 provide time constraints at ca. 230-240°C (Fig. 7A) whereas ZFT and AFT ages put constraints

456 on the Rocher Blanc ophiolite cooling history at ~ 280 and $\sim 100^\circ\text{C}$ (Fig. 7B) respectively. In
457 particular, the MgHe data refine the thermal history by requiring that the rapid phase of
458 cooling did not commence until after 20 Ma (Fig. 7C). The three systems are thus
459 complementary with respect to the temperature range that they cover. The apparent
460 consistency between MgHe, AFT and ZFT cooling ages suggests that the He diffusion data
461 obtained at 1 bar by diffusion experiments in vacuum at temperatures between 350 and
462 620°C over several days by Blackburn et al. (2007) can be extended to geological conditions
463 and timescales. As the AFT and ZFT data for RB1 sample are extrapolated from a different
464 locality, we also ran models with AFT and ZFT ages from samples CG and RN (Fig. S1). The
465 resulting cooling histories only slightly differ from those obtained with the extrapolated
466 Rocher Blanc ophiolite ages (Fig. S1) supporting the relevance of this thermal modeling
467 approach.

468

469 **5.5. Final exhumation of the Rocher Blanc ophiolite: geodynamic implications**

470 The thermal history obtained using the combination of thermochronometric methods (AFT-
471 ZFT and MgHe data) brings strong constraint on the final exhumation of the Rocher Blanc
472 ophiolite and also on the Schistes lustrés units of the Piedmont zone (Western Alps). Our
473 thermal modeling reveals that the RB1 sample has undergone nonlinear cooling and records
474 a major change in the thermal pathway at 20 Ma (Fig. 7C). During the first exhumation stage,
475 from 55 to 20 Ma, the sample recorded slow cooling from 350 to 300°C associated with an
476 extensive exhumation from 30 to 10 km (Fig. 6). The exhumation of ~ 20 km of rocks from
477 the Alpine orogenic wedge cannot be accounted for by surface denudation (i.e. erosion) only.
478 Large exhumation with limited cooling rather corresponds to exhumation in ductile
479 conditions related to subduction dynamics and underplating processes (Agard et al., 2001).

480 During ductile exhumation the sample passes from blueschist to greenschist conditions with
481 a mean exhumation rate lower than 1 mm/yr, calculated using the depth and time
482 constraints (Fig. 6A and 7). Over the second exhumation stage, from 20 Ma to 10 Ma, the
483 RB1 sample records rapid cooling from 300 to 100°C (Fig. 7) associated with an exhumation
484 from 10 to 3 km (Fig. 6A). This exhumation and strong change of cooling rate reflect
485 denudation in the brittle conditions. A mean exhumation rate of 0.7 mm/yr can be deduced
486 (Fig. 6A and 7). The last information of the thermal history is given by the final exhumation
487 stage recorded by AFT data and an exhumation of 3 km over 10 Ma leading to a mean lower
488 exhumation rate of 0.3 mm/yr (Fig. 7). Interestingly, the sample records similar exhumation
489 rate in ductile and brittle domains around <1 mm/yr, but significant cooling rate variations.
490 The change at ~20 Ma corresponds to the passage of the Rocher Blanc ophiolite below the
491 transition from ductile to brittle conditions in the lower grade of the greenschist facies
492 conditions (~250°C) during the retrograde path.

493 The final exhumation recorded by the thermal history reflects denudation at brittle
494 conditions (Fig. 7), where denudation is related to erosion linked to relief creation. The relief
495 creation is associated to a westward tilting of the Schistes lustrés units (Tricart et al., 2004;
496 Schwartz et al., 2009), as a response to the upward indentation of the Dora Maira massif.
497 Indeed, the crustal scale geometry of the internal zone is characterized by the presence of
498 Apulian mantle in a superficial position (Fig. 8), at 12 kilometers below the Dora Maira
499 massif (Lardeaux et al., 2006; Zhao et al., 2015). This portion of cold and rigid mantle acts as
500 an indenter that allowed the deformation of the thinned Briançonnais margin by crustal
501 underthrust process constituting a continental subduction wedge that forms the current
502 basement of the Piedmont zone (oceanic subduction wedge in Fig. 8). This continental

503 subduction wedge was overthrust on the European crust by the activation of the crustal
504 Penninic Frontal Thrust (PFT in Fig.8) at ~34 Ma (Dumont et al., 2012).
505 Taking in consideration the MgHe data of this study, the AFT-ZFT data of Piedmont zone
506 (Schwartz et al., 2007) and AFT data of the internal crystalline Massif of Dora Maira (Beucher
507 et al., 2012), we can illustrate the evolution of the area with their respective positions
508 though time in the orogenic wedge (Fig. 8). At present-day, all samples crop out at similar
509 elevations (>2000 m a.s.l.) for the Piedmont zone and at lower elevation (<1500 m a.s.l.) for
510 Dora Maira. At present-day the global shortening affects the front of the European foreland
511 (Schwartz et al., 2017), and allows the uplift and exhumation of the external crystalline
512 massifs (Pelvoux, Beucher et al., 2012) (Fig. 8). Using the AFT-ZFT and MgHe ages and
513 thermal sensitivity of the thermochronometric systems for all the samples, we can
514 reconstruct their position at depth ~20 Ma ago. Firstly, the samples CR from Schistes lustrés
515 and from Dora Maira present AFT age of 20-25 Ma indicating a shallow position around 3 km
516 (assuming an AFT closure temperature of 100°C and a thermal gradient of 25 °C). The
517 position at depth for the other samples from the Piedmont zone need to be progressively
518 deeper from east to west (Fig. 8). RB1 sample was passing the 250°C isotherm ~20 Ma based
519 on the thermal history obtained in this study (Fig. 7). From ~20 Ma to present-day, the
520 Piedmont zone has recorded a differential exhumation with a westward tilting and uplift
521 that is recorded by the youngest ages to the east in the Monviso (Fig. 8). This exhumation is
522 controlled by the upward indentation of the Apulian mantle producing the global tilting of
523 the Piedmont zone (Fig. 8).

524

525 6. Conclusions

526 This study contributes to the understanding of the formation conditions of magnetite in HP-
527 LT ophiolite and to the use of magnetite (U-Th-Sm)/He as a thermochronometer to quantify
528 the exhumation history. We demonstrate that a comprehensive characterization of
529 magnetite crystals and crystallization environment is needed in conjunction with MgHe
530 dating in order to interpret the age data.

531 We specifically focused on one sample (RB1) from the Rocher Blanc massif that is an
532 ophiolitic body embedded in the Schistes lustrés units (Western Alps, France), which
533 contains abundant magnetite. This sample recorded middle-temperature blueschist facies
534 conditions and was overprinted by greenschist conditions during exhumation. We show that
535 for the RB1 sample, magnetite crystallized in two zones of distinct chemistry, that originated
536 from magmatic relictual zones, with euhedral and pseudo-euhedral grains with a dominant
537 chlorite or titanite mineralogy, respectively. In addition, the petrological and magnetite
538 geochemical analyses (e.g. V and Si) indicate that magnetite crystallized during the
539 retrograde metamorphic path under greenschist conditions ($T > 250^{\circ}\text{C}$). The Si and V zoning
540 observed in euhedral and pseudo-euhedral magnetite grains reinforce the fact that both
541 type of magnetite crystallized under same physical conditions. The pseudo-euhedral
542 magnetite, crystallizing in dominant titanite mineralogy, present older MgHe ages compared
543 to euhedral magnetite. The old MgHe ages are interpreted by the impact of alpha-
544 implantation from U-Th-rich neighbor minerals (e.g., titanite). On the contrary, alpha-
545 implantation in euhedral magnetite crystallizing in chlorite dominant zone is insignificant, as
546 chlorite contain similar U-Th content than magnetite. The MgHe ages derived from euhedral
547 magnetite range from 14.8 ± 2.1 to 20.9 ± 3.0 Ma, sandwiched between ZFT and AFT ages as
548 expected. The MgHe data along with AFT and ZFT data were inverted for the first time to
549 refine the local thermal history. The MgHe data suggest that fast cooling was delayed until

550 ~20 Ma, as opposed to the less constrained ~26 Ma that would be inferred from the ZFT and
551 AFT ages alone. The MgHe data suggest that the RB1 sample has undergone an exhumation
552 in ductile conditions related to subduction dynamics and underplating processes, followed
553 by an exhumation reflecting denudation by erosion at brittle conditions. Considering all
554 thermochronometric data along the Western Alps, we propose that the exhumation is
555 controlled by the upward indentation of the Apulian mantle producing the global tilting of
556 area and associated erosion.

557 Additional work on U-Th zoning in magnetite and He diffusion in magnetite will be necessary
558 in the future, as the thermal inversion interpretation is contingent on the [Blackburn et al.](#)
559 [\(2007\)](#) kinetic parameters being correct. Nevertheless, the present study further
560 demonstrates that MgHe can be used as a low-temperature thermochronometer to date HP
561 ultramafic final exhumation and the ductile-brittle transition opening new avenues of
562 research. This method effectively complements the available thermochronological tools,
563 making it a powerful tool for emerging applications.

564

565 **Acknowledgments**

566 This work has been funded thanks to TelluS-SYSTER INSU 2016 project and CNRS Delegation
567 of S. Schwartz. The UTCT lab is supported by NSF-EAR-IF grant EAR-1762458 to R. Ketcham.
568 Jessie Maisano is thanked for her help during CT scanning. V. Batanova, V. Magnien and N.
569 Findling are thanked for the technical support of microprobe and SEM analysis. B. Guéguen
570 is warmly thanked for her help during the magnetite chemical dissolution part. The three
571 anonymous reviewers and An Yin editor are warmly thanked for their constructive reviews.

572

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737

738 **Figure caption**

739 **Figure 1:** (A) Geographical setting of the Alps showing the location of Fig. 1B. (B) Simplified
740 structural and geological map of the Western Alps between France and Italy, with location of
741 Figure 1C. (C) Structural sketch-map of the Schistes Lustrés units of Piedmont zone, with
742 published apatite and fission tracks ages (modified from [Schwartz et al., 2007](#)) and location
743 of the Rocher Blanc ophiolite massif (RB), with extrapolated AFT and ZFT ages (in *). The
744 location of CG (Cirque de la Gavie) and RN (Rocca Nera) are also indicated. (D) Pressure-
745 Temperature path from the Rocher Blanc ophiolite (modified from [Schwartz et al., 2009](#)).

746

747 **Figure 2:** Photographic views of the Rocher Blanc ophiolite massif, polished thin section of
748 RB1 sample and SEM backscattered electron images. (A) Fe-Ti gabbro embedded in
749 serpentinite and developing magnetite-bearing chloritoschist at its interface. (B) Millimeter
750 scale euhedral magnetite crystals visible on outcrop indicated by white arrows. (C) Thin-
751 section of the RB1 magnetite-bearing chloritoschist sample. The magnetite crystal analyzed
752 is presented in Figure 3. (D), (E) and (F) are SEM-BSE images of the sample RB1. The sample
753 presents two mineralogical zones (2 and 1), rich in titanite and ilmenite (D and E) and
754 chlorite (F), respectively. Mag, Ttn, Ilm and Chl mean magnetite, titanite, ilmenite and
755 chlorite.

756

757 **Figure 3:** Geochemical characterization of magnetite crystal from the chlorite-rich Zone 1
758 and titanite-rich Zone 2 shown in Fig. 2C (RB1-Mag1 and RB1-Mag5, Table 1). (A,E) EMP-SE
759 image of the magnetite crystal with EMP elemental X-ray maps (1 μm beam size, 1.5 μm step
760 size and 1s dwell time) and profile. (B,F) Si and V concentration in ppm along the profile.

761 **(C,G)** and **(D,H)** map of the Si and V concentration in ppm. Inclusions of silicate minerals
762 appear in white on the chemical maps.

763

764 **Figure 4:** In-situ U concentration of magnetite, chlorite and titanite from Zone 1 and Zone 2.

765 **(A)** Thin-section of the RB1 magnetite-bearing chloritoschist sample, with location of the U
766 analysis zones **(B)**, **(C)** and **(D)**. Zoom of the **(B)**, **(C)** and **(D)** areas and U content in ppb.

767

768 **Figure 5:** Magnetite (U-Th-Sm)/He ages from Rocher Blanc ophiolite. **(A)** and **(B)** are 3D
769 reconstructions from CT-scan images of each euhedral and pseudo-euhedral crystals. **(C)**
770 MgHe age as a function of the sphere equivalent radius (Rs) estimated from CT-scan analysis
771 (Ketcham, 2005a; Cooperdock et al., 2019). The AFT and ZFT age ranges from Schwartz et al.
772 (2007) and shown in Figure 1C are shown.

773

774 **Figure 6:** Pressure-temperature-time pathway record by the RB ophiolite and magnetite
775 growth condition. **(A)** Pressure-Temperature-time path of the HP-LT Rocher Blanc ophiolite
776 modified from Schwartz et al. (2009; 2013). The AFT and ZFT estimated ages from RB (Fig.
777 1C), MgHe ages from euhedral magnetites are reported for their associated closure
778 temperature condition (Blackburn et al., 2007; Gallagher et al., 1998; Ketcham, 2019). The
779 age of HP conditions are taken from Agard et al. (2002). **(B)** Sketch of magnetite crystals
780 growth during the retrograde pathway in three main steps, from (I) blueschist conditions
781 (BS), (II) High-Pressure greenschist (HP-GS) and (III) Low-Pressure greenschist (LP-GS)
782 conditions. Zones 1 and 2 refer to the initial magmatic sites with augite-plagioclase and Fe-Ti
783 oxide (ilmenite) respectively. (I) The HP blueschist conditions are marked by the partial
784 mineralogical transformation of augite to aegirine, and the development of glaucophane and

785 zoisite in the plagioclase site (Pognante et al., 1986). In Zone 2, ilmenite is progressively
786 transformed to titanite. (II) The HP greenschist conditions are marked by the progressive
787 destabilization of augite-aegirine, glaucophane, and zoisite to chlorite and magnetite in Zone
788 1, whereas, ilmenite continues to be destabilized in titanite and \pm magnetite and chlorite.
789 The magnetite grains show low Si content whatever the considered zones. (III) Finally, the LP
790 greenschist facies conditions, show the pervasive development of the chlorite and magnetite
791 in the two zones. In the zone 2, only relics of ilmenite remain locally in the core of titanite
792 crystals. The magnetites present higher Si content.

793

794 **Figure 7:** Thermochronological data inversion modeling of the HP-LT Rocher Blanc RB
795 ophiolite using HeFTy software (Ketcham, 2005). **(A).** T-t pathways obtained for MgHe data
796 inversion only. **(B).** T-t pathways obtained for AFT and ZFT data inversion only. **(C).** T-t paths
797 obtained for joint AFT, ZFT and MgHe data inversion.

798

799 **Figure 8:** Schematic evolution of the Western Alps between France and Italy \sim 20 Ma ago and
800 at present-day. At \sim 20 Ma, the Penninic Frontal Thrust (PFT) is already active since 10 Myr
801 associated with the mantellic upward indentation, and the AFT-ZFT and MgHe dated
802 samples have started to be exhumed in the brittle condition (above 250°C). The samples
803 record differential exhumation related with the westward tilting and uplift. One
804 representative sample from Dora Maira massif is represented, where AFT ages of 20-25Ma
805 have recorded by Beucher et al. (2012).

806

807 **Figure S1:** Thermochronological data inversion modeling of the HP-LT Rocher Blanc RB
808 ophiolite using HeFTy software (Ketcham, 2005). T-t paths obtained using AFT and ZFT data

809 from sample CG **(A, B, C)**, sample RB **(D, E, F)**, RN sample **(G, H, I)**, combined without or with
810 MgHe data. **(J-K)** Weighted model Tt paths resumed results for CG, RB and RN samples,
811 without or with MgHe data. **(L)**. Tt pathway model obtained for MgHe data inversion only.

812

813 **Table 1:** Elemental concentration in ppm obtained on four representative magnetite crystals
814 from zone 1 of the RB1 sample.

815

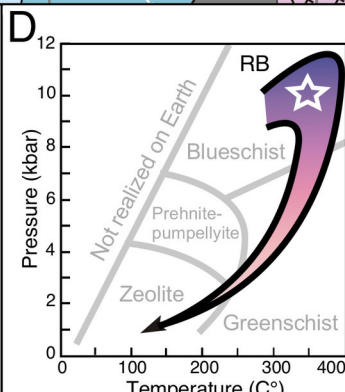
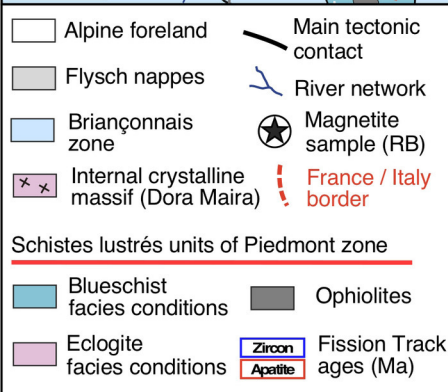
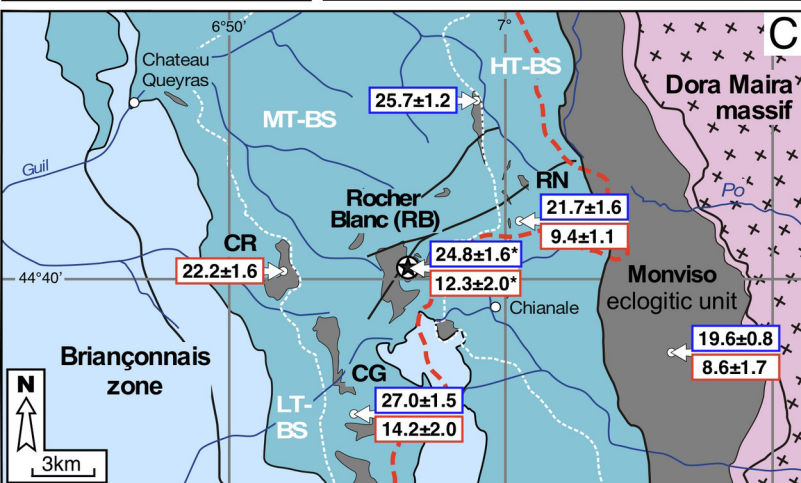
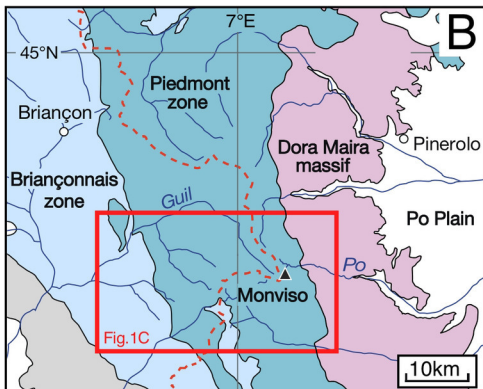
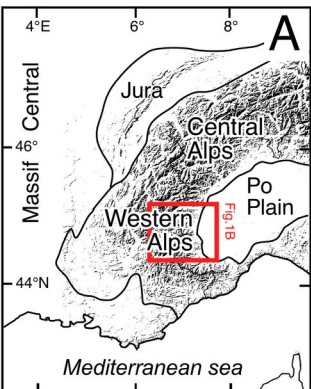
816 **Table 2:** In-situ U and Th concentrations obtained by LA-ICP-MS in sample RB1 (thick section).

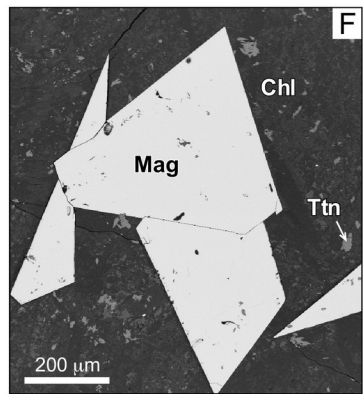
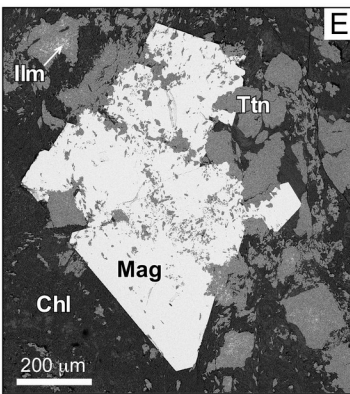
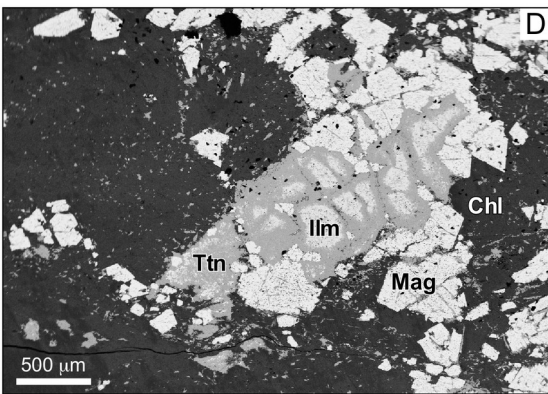
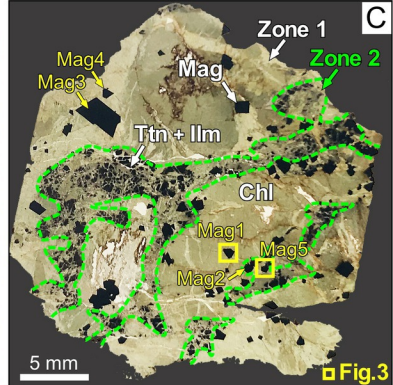
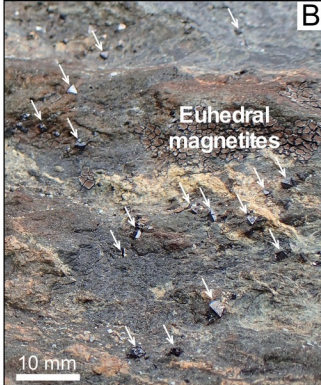
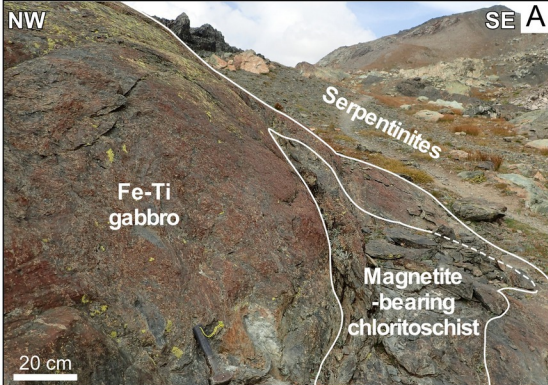
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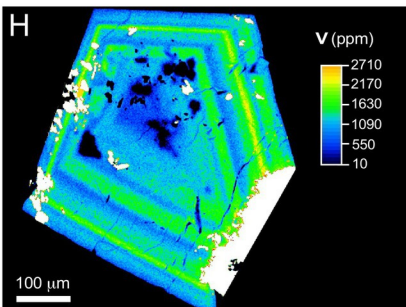
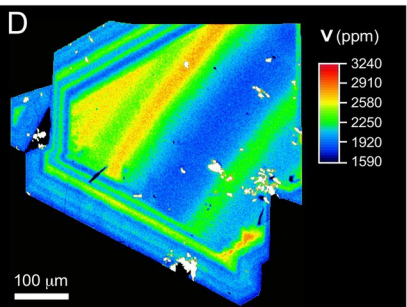
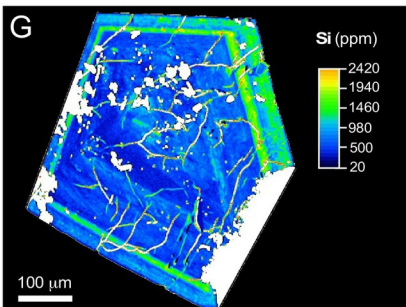
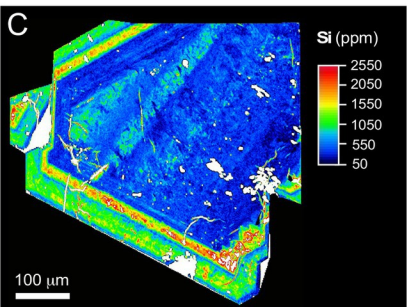
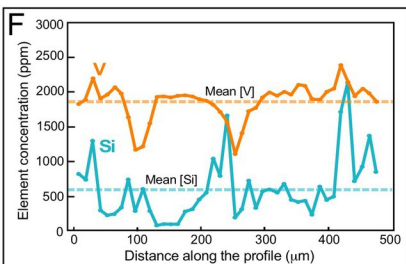
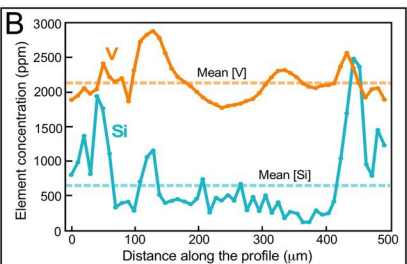
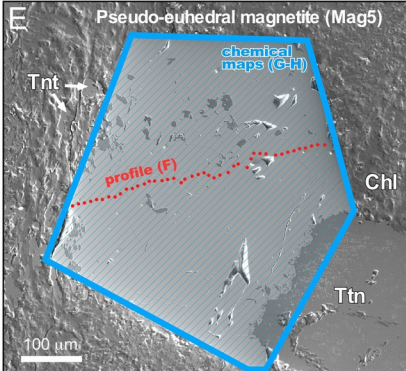
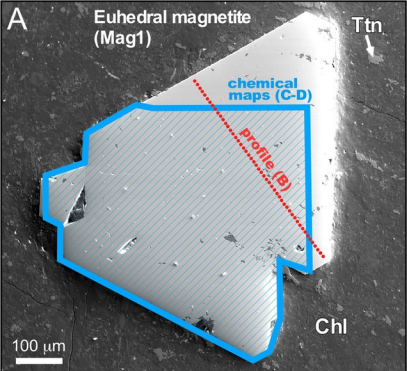
818 **Table 3:** CT-scan and (U-Th-Sm)/He results of the magnetite-bearing chloritoschist RB1
819 sample

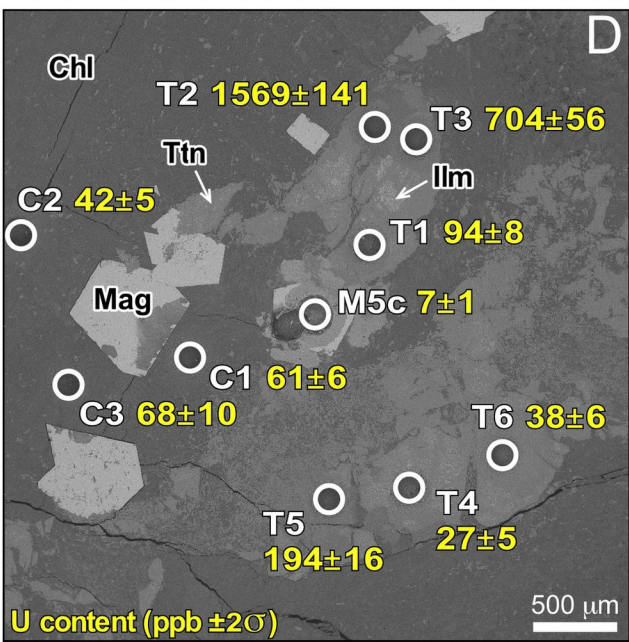
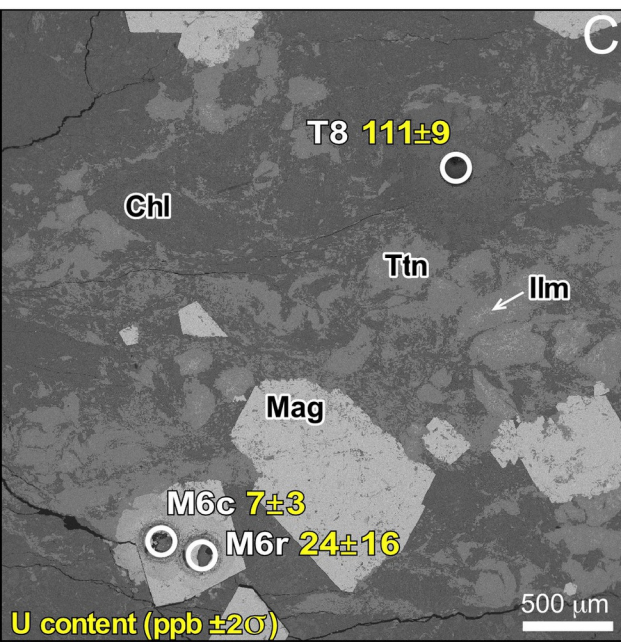
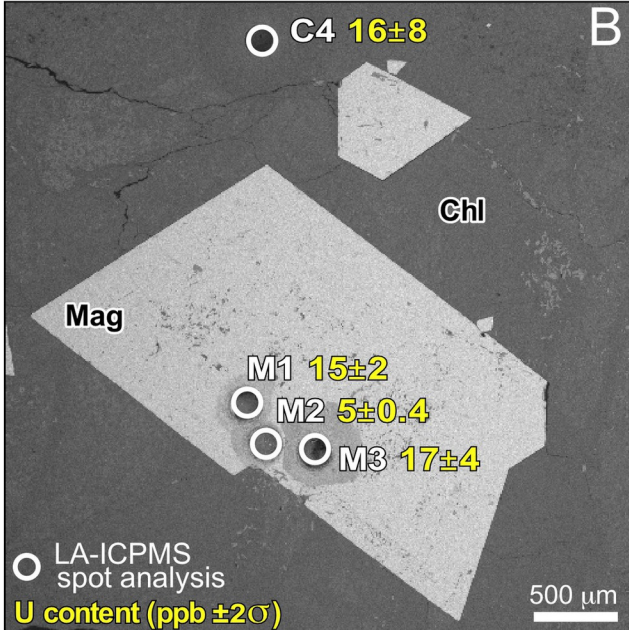
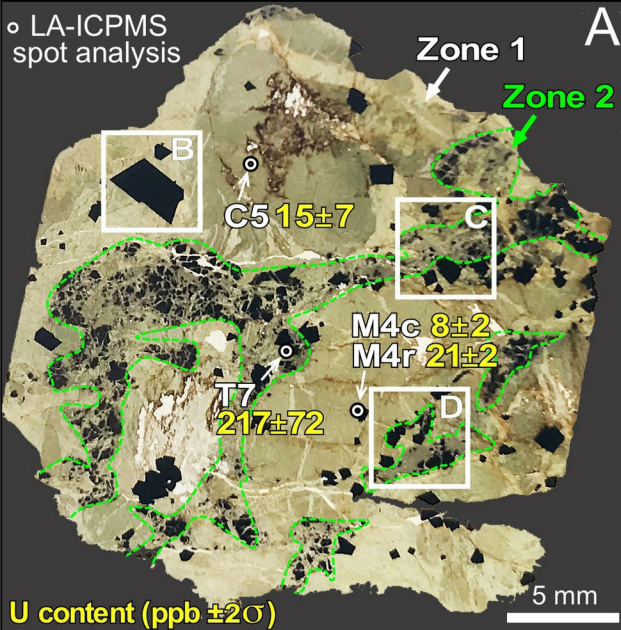
820

821 **Table 4:** Thermal History Inverse Modeling Parameters

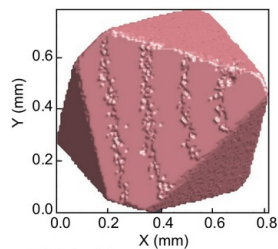




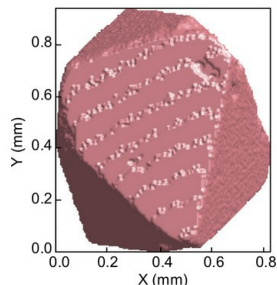




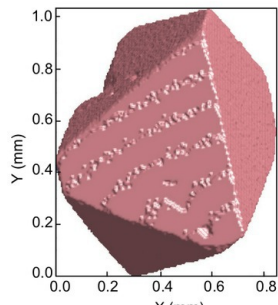
Euhedral magnetite



RB1-10
17.6 ± 2.6 Ma



RB1-7
19.9 ± 2.9 Ma

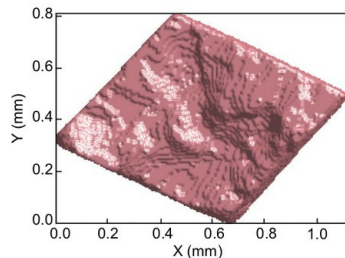


RB1-5
20.9 ± 3.0 Ma

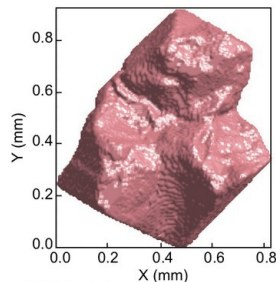
A

Pseudo-euhedral magnetite

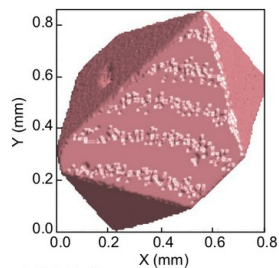
B



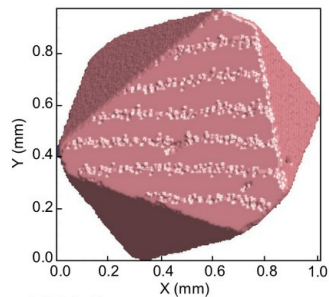
RB1-8
32.0 ± 4.6 Ma



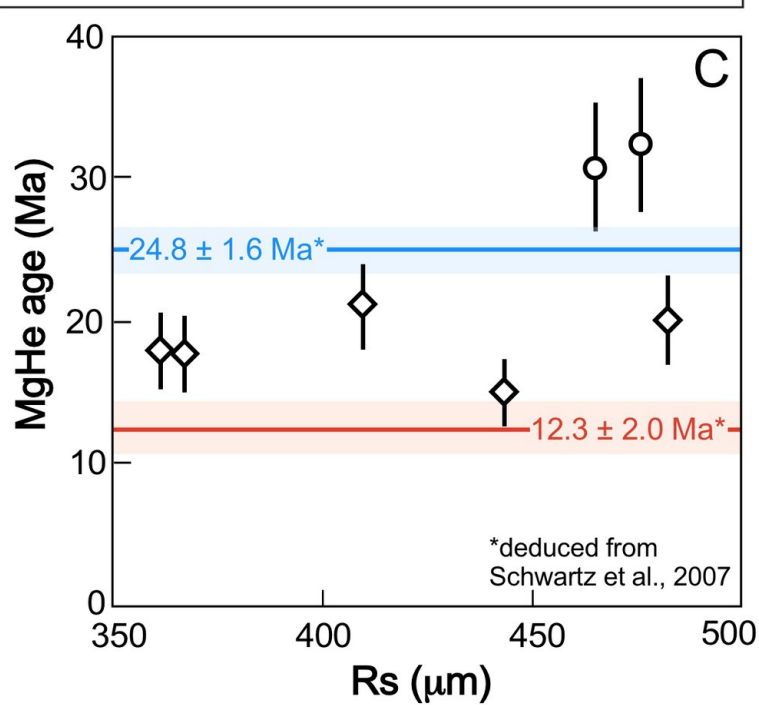
RB1-11
30.4 ± 4.4 Ma



RB1-4
17.4 ± 2.6 Ma



RB1-3
14.8 ± 2.1 Ma



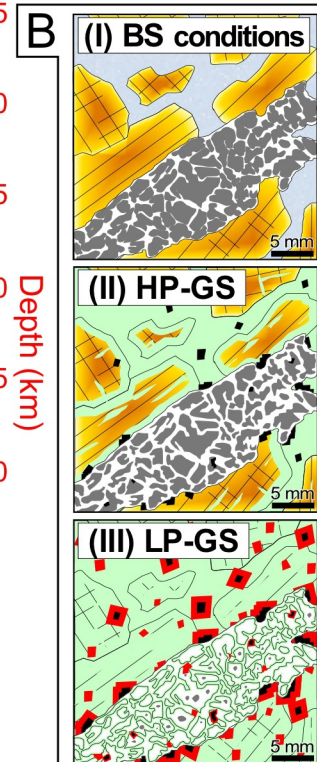
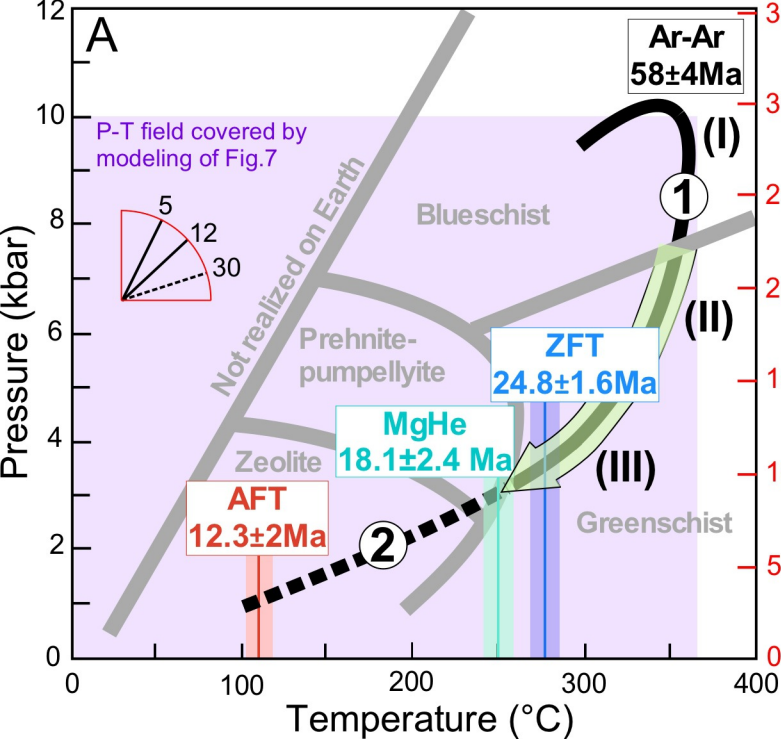
— Apatite FT ages

— Zircon FT ages

Magnetite He age (This study)

○ Pseudo-euhedral magnetite

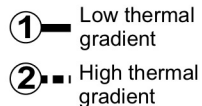
◇ Euhedral magnetite



Closure temperature (T_c):



Exhumation path of RB ophiolite:



Geothermal gradients ($^{\circ}\text{C}/\text{km}$)



Magnetite growth during retrograde path

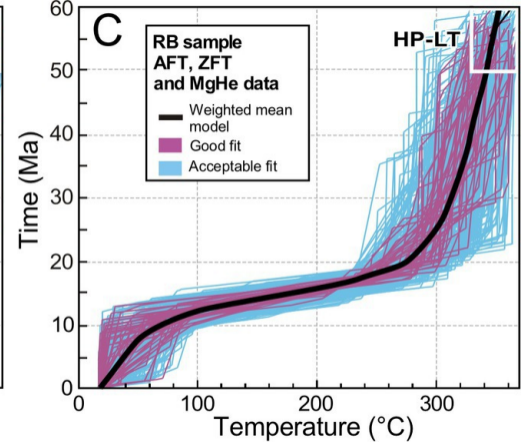
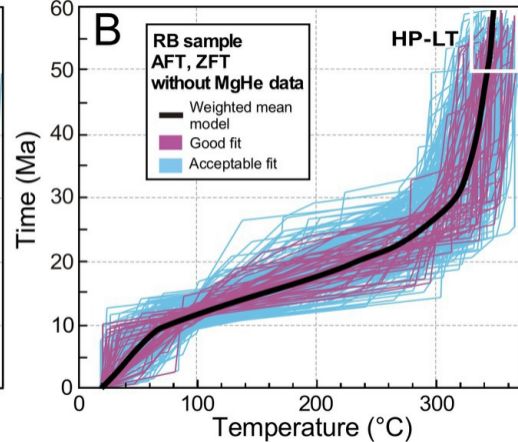
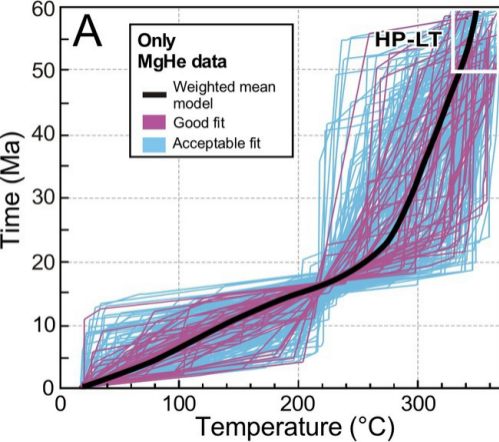
- ◆ Magnetite (low [Si])
- ◆ Magnetite (high [Si])

Zone 1

- Augite-Aegyrine
- Pl, Gl, Zo
- Chlorite

Zone 2

- Ilmenite
- Titanite



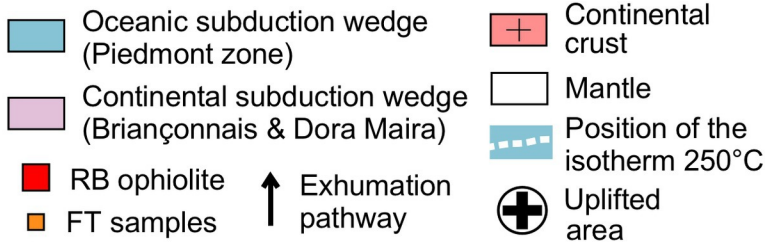
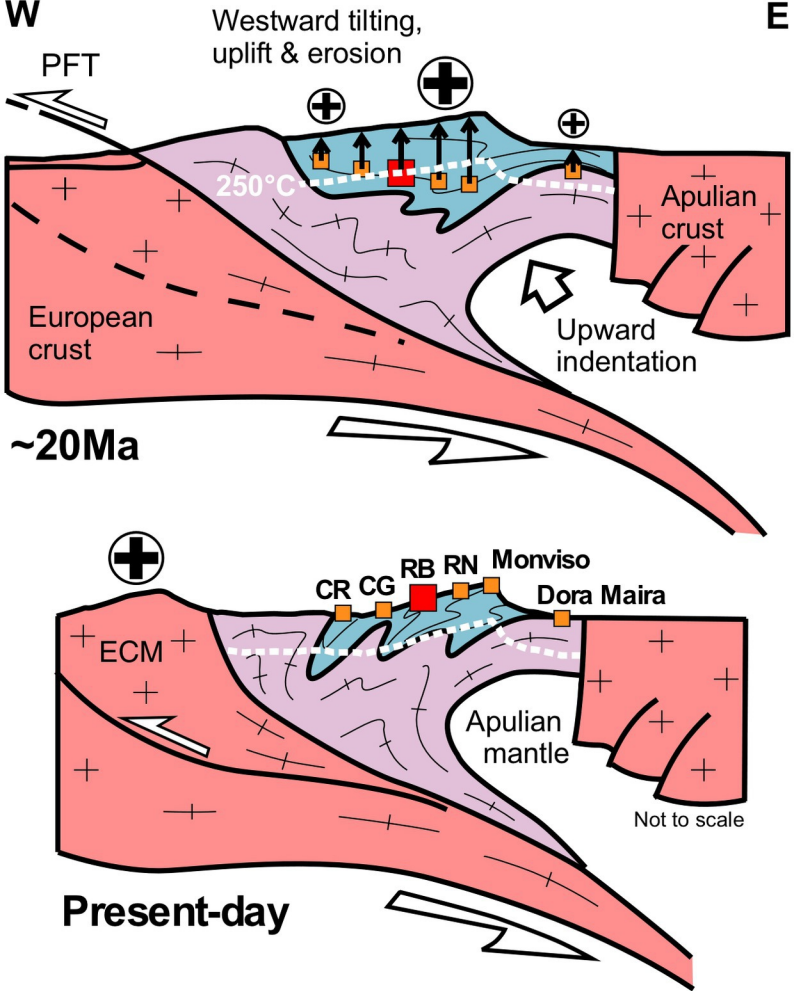


Table 1: Elemental concentration in ppm obtained on representative magnetite crystals from zone 1 and zone 2 of the RB1 sample.

Element	Si	Ti	Cr	V	Mn	Mg
Detection Limit (ppm)	24	30	30	30	35	25
Euhedral magnetite						
RB1-Mag1 (n=76)						
Size 500 μm	680 \pm 474	138 \pm 102	55 \pm 11	2119 \pm 244	432 \pm 35	B.d.l.
Profile B – Fig. 3						
RB1-Mag3 (n=69)						
Size 650 μm	1019 \pm 638	72 \pm 48	48 \pm 10	1669 \pm 267	420 \pm 24	B.d.l.
RB1-Mag4 (n=61)						
Size 600 μm	744 \pm 349	60 \pm 38	45 \pm 10	1802 \pm 230	429 \pm 19	B.d.l.
Pseudo-euhedral magnetite						
RB1-Mag2 (n=53)						
Size 550 μm	993 \pm 650	59 \pm 31	43 \pm 11	1653 \pm 329	418 \pm 34	B.d.l.
RB1-Mag5 (n=42)						
480 μm	608 \pm 456	n.a.	n.a.	1863 \pm 257	n.a.	n.a.
Profile F – Fig. 3						
RB1 (n=301)	809\pm513	82\pm55	48\pm10	1821\pm265	425\pm28	B.d.l.

B.d.l.: Below detection limits

n.a. not analyzed

Table 2: In-situ U and Th concentrations obtained by LA-ICP-MS in sample RB1 (thick section).

Magnetite	M1	M2	M3	M4c[§]	M4r[§]	M5c	M6c	M6r
	Zone 1 - euhedral mag.					Zone 2 - pseudo euhedral mag.		
U \pm 2 σ (ppb)	15 \pm 2	5 \pm 0.4	17 \pm 4	8 \pm 2	21 \pm 2	7 \pm 1	7 \pm 3	24 \pm 16
Th \pm 2 σ (ppb)	3 \pm 3	bdl	6 \pm 3	1 \pm 0.7	6 \pm 9	3 \pm 0.6	4 \pm 1	bdl
Titanite	T1	T2	T3	T7	T8	T4*	T5*	T6*
	Zone 2							
U \pm 2 σ (ppb)	94 \pm 8	1569 \pm 141	704 \pm 56	217 \pm 72	111 \pm 9	27 \pm 5	194 \pm 16	38 \pm 6
Th \pm 2 σ (ppb)	28 \pm 22	628 \pm 546	109 \pm 9	163 \pm 117	4 \pm 1	4 \pm 1	36 \pm 3	6 \pm 1
Chlorite	C4	C5	C1	C2	C3			
	Zone 1		Zone 2					
U \pm 2 σ (ppb)	16 \pm 8	15 \pm 7	61 \pm 6	42 \pm 5	68 \pm 10			
Th \pm 2 σ (ppb)	5 \pm 1	13 \pm 5	7 \pm 1	13 \pm 3	17 \pm 12			

c[§] and r[§] mean magnetite core and rim

* Mn-rich titanite

bdl= below detection limit

Table 3: CT-scan and (U-Th-Sm)/He results of the magnetite-bearing chloritoschist RB1 sample

Aliquot	Volume (mm ³)	Surface area (mm ²)	Rs (μ m)	F _T	Mass (mg)	⁴ He (pmol)	²³⁸ U (ng)	²³² Th (ng)	¹⁴⁷ Sm (ng)	⁴ He \pm s (pmol/g)	²³⁸ U (ppb)	²³² Th (ppb)	¹⁴⁷ Sm (ppb)	eU (ppb)	Th/U	Raw date (Ma)	Corrected date \pm s (Ma)
Euhedral magnetite																	
RB1-3	0.29	2.72	444	0.98	1.54	2.33 \times 10 ⁻³	0.025	0.020	0.018	1.49 \pm 0.03	16.2	13.0	11.7	19	0.8	14.4 \pm 2.2	14.8 \pm 2.1
RB1-4	0.16	1.82	368	0.97	0.84	1.85 \times 10 ⁻³	0.019	0.007	0.021	2.18 \pm 0.04	22.7	8.4	25.1	25	0.4	16.9 \pm 2.5	17.4 \pm 2.6
RB1-5	0.21	2.28	410	0.97	1.12	1.16 \times 10 ⁻³	0.009	0.007	0.015	0.88 \pm 0.02	7.8	6.5	13.3	9	0.8	20.4 \pm 3.1	20.9 \pm 3.1
RB1-7	0.25	2.46	483	0.98	1.32	4.71 \times 10 ⁻³	0.038	0.027	0.075	1.85 \pm 0.04	28.7	20.2	56.9	34	0.7	19.5 \pm 2.9	19.9 \pm 2.9
RB1-10	0.16	1.82	362	0.97	0.85	2.03 \times 10 ⁻³	0.020	0.009	0.015	2.31 \pm 0.05	23.2	11.2	17.5	26	0.5	17.1 \pm 2.6	17.6 \pm 2.6
Pseudo-euhedral magnetite																	
RB1-8	0.18	2.12	477	0.98	0.94	3.23 \times 10 ⁻³	0.016	0.012	0.020	3.26 \pm 0.07	17.3	12.3	21.9	20	0.7	31.3 \pm 4.7	32.0 \pm 4.8
RB1-11	0.13	1.80	466	0.98	0.71	1.60 \times 10 ⁻³	0.008	0.007	0.004	2.39 \pm 0.05	11.7	9.7	6.2	14	0.8	29.7 \pm 4.5	30.4 \pm 4.6

Rs and F_T are the sphere equivalent radius and alpha ejection factor.

(U-Th-Sm)/He date has been corrected from alpha ejection factor (F_T) (Ketcham et al., 2011).

Table 4: Thermal History Inverse Modeling Parameters

1. Thermochronologic data	
Samples and data used	AFT and ZFT data from Schwartz et al. (2007) MgHe data from this paper: Table 2, samples RB1-3, 4, 5, 7, and 10
Data treatment: Dates, uncertainties, and other relevant constraints	
<u>AFT data</u>	Kinetics: Apatite assumed to have $Cl=0$ apfu Uncertainty model: Asymmetric 95% bounds
<u>ZFT data</u>	Initial mean track length: N/A; no track length data available Uncertainty model: Asymmetric 95% bounds
<u>MgHe data</u>	Uncertainty model: Uniform 15% (1SE) uncertainty
2. Additional geologic information and constraint placement	
Present-day temperature	20°C
Initial constraint	Assumed peak HP-LT conditions: $350\pm 20^\circ\text{C}$ at 58 ± 4 Ma
T-t path complexity	HeFTy code E3/50: Paths episodic, halved 3 times, with maximum cooling rate $50^\circ\text{C}/\text{m.y.}$
3. System- and model-specific parameters	
AFT annealing model: Ketcham et al. (1999)	
ZFT annealing model: Ketcham (2019)	
MgHe diffusion parameters: Blackburn et al. (2007)	
Modeling code: HeFTy v1.9.3	
Statistical fitting criteria: HeFTy default	
Ending condition: Models run until 50 good fits found	
