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Interferometry for full temporal reconstruction of laser-plasma accelerator-based seeded free electron lasers

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Abstract

The spectacular development of Laser-Plasma Accelerators (LPA) appears very promising for a free electron laser application. The handling of the inherent properties of those LPA beams already allowed controlled production of LPA–based spontaneous undulator radiation. Stepping further, we here unveil that the forthcoming LPA–based seeded FELs will present distinctive spatio-spectral distributions. Relying on numerical simulations and simple analytical models, we show how those interferometric patterns can be exploited to retrieve, in single-shot, the spectro-temporal content and source point properties of the FEL pulses.

1. Introduction

Free Electron Lasers (FELs) [1, 2] deliver ultrashort, narrow-band and ultrabright pulses down to the hard x-ray range [3–5], enabling breakthrough experiments in chemical, physical, and biological sciences. These light sources rely on relativistic electron beams wiggling in the periodic magnetic field of an undulator as gain medium. Interacting with the spontaneous radiation of the undulator or an external seed, the electrons experience an energy modulation at the resonance wavelength which is further transformed into a density modulation by dispersive elements. After this 'lethargy' [6] phase, the beam density modulation allows the emission of a coherent radiation which can then be exponentially amplified. A saturation is reached when the electrons energy loss is such that the resonance condition is violated, causing a red spectral shift of the FEL line [7, 8].

Chicanes as dispersive elements are used to speed up the electron beam energy to density modulation conversion [9, 10]. Associated to electron beam energy chirping in the Radio-Frequency Accelerators (RFAs) structures producing those beams and then to frequency chirped seeds, they opened the door to FEL wavelength tunability [11], two-color operation [12], amplitude and phase control [13, 14], spectro-temporal shaping [15], chirped-pulse compression [16, 17] and temporal reconstruction [18].

In high-gain FELs [6], the initial radiation can be amplified by several orders of magnitudes up to saturation in one single pass. The gain of those systems is mainly driven by the electron beam properties. This is the reason why first FELs were developed using classical RFAs which presently deliver the requisite high-quality beams. However since those accelerators are typically several hundreds of meters long for a lasing in the x-ray range, the new Laser Plasma Accelerators (LPAs) [19] have been proposed as an alternative towards miniaturized FELs [20].

However from proposal to demonstration, the path is tortuous. In LPAs, the intrinsic large divergence (\approx few milliradians) and energy spread (\approx few percent) of the electron beam [21, 22] dramatically limit the FEL gain. Several techniques can be implemented to reach the mandatory Pierce parameter ρ_{fel} [6] qualifying the FEL gain: use plasma lenses [23–26] and/or high gradient quadrupoles in the immediate vicinity of the LPA source to handle the large divergence [27–29], use a chicane [30, 31] or a transverse gradient undulator [32, 33] to lower the slice energy spread, and/or use a chromatic matching [34] to synchronize the beam maximum density with the radiation pulse propagation inside the undulator. Few years after the first observation of LPA–based



Figure 1. Schematic of a LPA–based seeded FEL. From one main femtosecond infrared laser unit, two pulses centered at 800 nm are delivered with independent controls. The main laser branch is focused in a gas medium for electron beam generation. The second branch is used for seed generation. At the undulator exit, the radiation is observed using an imaging spectrometer while the electron beam is dumped.

undulator radiation [35–38], significative achievements in terms of electron beam transport and manipulation [39] turned a corner, allowing for control of the synchrotron radiation spectral properties [40]. LPA–based FEL amplification is now within reach at least in the visible or ultra-violet range. To ease this next step, a seeded configuration should be considered, since it enables to speed up the lethargy phase, i.e. requires shorter undulator lengths and relaxes the transport issues. Given the present LPAs performance, the first demonstration is more likely to be obtained in the low-gain [41] regime as proposed in [42].

The aim of this work is to reveal that those forthcoming LPA–based seeded FELs can present original spatiospectral distributions that enable the full reconstruction of the temporal FEL pulse. Indeed, under simple conditions, the emitted radiation exhibits a clear red shift with respect to the input seed wavelength, leading to spectral interference fringes between this additional emitted radiation and the input seed. In the case of LPAs, the electron beam energy chirp, i.e. longitudinal sorting in energy, is not induced by the accelerating structure but by a dispersive transport of the high energy spread beam through a chicane. The resulting particles velocity gradient is so steep that the undulator dispersion drives the energy to density modulation conversion while simultaneously stretching the density modulation period (ahead traveling faster than tail particles). This effect is usually negligible on RFA based FELs, even though observed on a self-seeding configuration [43]. In the case of LPA beams however, the stretching is such that the final FEL wavelength can be red shifted by more than one seed linewidth. Due to the low-gain configuration, the red-shifted coherent emission is observed together with the seed, leading to strong interferences in the spectral domain. As in [18], this interferometric pattern can then be used for temporal reconstruction of the FEL pulses.

This paper is organized as follows. The typical schematic of an LPA–based seeded FEL is recalled in section 2 with the specific parameters considered in this study. The numerical simulation principle is also detailed. Section 3 describes the analytical model developed to explain the red-shift feature. The correlation between the FEL linewidth and the electron beam divergence is then investigated in section 4. To reproduce the interference pattern observed in the simulations, we present a double pulse model in section 5 and use it to achieve a full temporal reconstruction of the FEL pulse in section 6.

2. Experimental setup

A typical setup for a LPA–based seeded FEL is shown in figure 1. In state-of-the-art LPAs, nonlinear plasma waves are driven in a gas by femtosecond pulses of a multi-TW laser system, allowing for the trapping and further acceleration of electron bunches. The properties of the generated electron beams strongly depend on the LPA implementation. We choose a set of parameters (see table 1) compatible with the most recent achievements of the groups targeting both quality and reliability, mandatory for an FEL application [21, 42, 44–46]. With a short period undulator [47], these typical 200 MeV beams can be used to target the vacuum ultra-violet range. We therefore consider a seed at 160 nm (see table 1) generated in gas [48, 49] using a branch of the main laser [49]. The electron beam is transported along a dedicated line allowing for a three steps manipulation [34]: refocusing at the LPA exit using high gradient quadrupoles [50], longitudinal stretching and energy chirping using a chicane and chromatic matching using standard quadrupoles. Locally driving the beam off-axis, the chicane is also used to insert a mirror for the seed injection on-axis. Experimentally, an imaging spectrometer can provide the spatio-spectral distribution of the FEL radiation. Combined with adequate focusing optics, one can observe either the near- or far-field spectral distribution.

	Parameter	Value
Electron beam	Energy	200 MeV
	Relative energy spread	4 %
	Charge Q	20 pC
	Divergence $\sigma'_{x0,y0}$	1 mrad-rms
	Emittance $\epsilon_{x0,y0}$	0.1 mm mrad
	Chicane dispersion $R_{56,c}$	2.5 mm
Undulator	Period λ_u	18 mm
	Number of periods N_u	107
	Deflection parameter K_u	1.864
Seed laser	Central wavelength	160 nm
	Quadratic phase	$1.15 \times 10^{-4} {\rm fs}^{-2}$
	Duration	370 fs-FWHM
	Peak power	10 kW
	Waist position	0 m
	Rayleigh length	0.7 m

Table 1. Parameters for FEL simulation and modeling.

Using parameters of table 1, a 6D electron beam is generated with a flat top profile in energy and Gaussian profiles in other dimensions. The distribution is then transported to the undulator entrance using second order transport formalism [51]. The FEL amplification is simulated with GENESIS [52] using this 6D electron beam distribution together with an amplitude and phase temporal description of the seed while the undulator magnetic field is defined by its period λ_u , number of periods N_u and deflection parameter K_u (see table 1). The output radiation is then analyzed using Fourier optics to predict the observable near- and far-field spatio-spectral distributions. Typical simulation outputs are presented in figure 2. To identify the seed contribution, the charge Q is first forced to zero. As expected, Gaussian distributions are obtained in the near- and far-fields (figures 2(a) and (b)). The seed power P_{seed} is then forced to zero to distinguish the spontaneous emission contribution (figures 2(c) and (d)) expected at a resonance wavelength $\lambda_R = \lambda_u/(2\gamma^2)(1 + K^2/2 + \theta^2\gamma^2)$ for a mono-energetic beam with a Lorentz factor γ , θ being the observation angle. Hence the observed broad band emission results from the large energy spread (different γ) while the 'moon shapes' in the far-field result from the off-axis contribution in θ^2 [53]. When the seed overlaps the electron beam, i.e. in the FEL case, a significant red shift (≈ 1.5 nm) appears (figures 2(e) and (f)), together with high contrast fringes.

3. Analytic modeling of the FEL red shift

The standard FEL red spectral shift due to beam energy loss can only be responsible here for sub-nanometer spectral shifts (0.002 nm in the case of table 1). In order to understand the process underlying here, we reconsider the analytical treatment of the electron beam longitudinal motion along the undulator. For a relativistic electron of energy γmc^2 traveling in a purely vertical undulator magnetic field, the Lorentz equation in the horizontal direction can be integrated to derive the average longitudinal velocity [54]:

$$\tilde{v}_z \approx c \left[1 - \frac{1}{2\gamma^2} \left(1 + \frac{K_u^2}{2} \right) \right].$$

The time for an electron to travel the distance z in the undulator is simply $t = z/\tilde{v}_z$ so that the difference of travel time dt for electrons of different energies d γ close to γ is:

$$dt = z \frac{d}{d\gamma} \left(\frac{1}{\tilde{\nu}_z} \right) d\gamma \approx -\frac{z}{c} \left(1 + \frac{K_u^2}{2} \right) \frac{d\gamma}{\gamma^3}.$$
 (1)

Thus, electrons with different energies induce a chirp associated to a stretching $\Delta \lambda$ of the beam due to the longitudinal dispersion $R_{56,u}(z)$ of the undulator:

$$\Delta \lambda = c \Delta t = -z \left(1 + \frac{K_u^2}{2} \right) \frac{1}{\gamma^2} \frac{\Delta \gamma}{\gamma} = R_{56,u}(z) \frac{\Delta \gamma}{\gamma}.$$
 (2)



Seed only ($P_{seed} = 10 \text{ kW}$, Q = 0 pC), (c) and (d) Spontaneous emission only ($P_{seed} = 0 \text{ W}$, ($P_{seed} = 10 \text{ kW}$, Q = 20 pC). Other parameters from table 1.

In a first approximation, the longitudinal energy sorting of the electrons introduced in the chicane is given by:

$$\Delta s = R_{56,c} \frac{\Delta \gamma}{\gamma} \tag{3}$$

with $R_{56,c}$ the chicane strength or dispersion. Considering now a beam portion of one seed laser period, i.e. $\Delta s = \lambda_{\text{seed}}$, and inserting equation (3) into (2), we find that this beam portion is increased by:

$$\Delta\lambda(z) = \frac{R_{56,u}(z)}{R_{56,c}}\lambda_{\text{seed}} = -\frac{z\left(1 + \frac{K_u^2}{2}\right)}{\gamma_0^2 R_{56,c}}\lambda_{\text{seed}}.$$
(4)

The initial λ_{seed} density modulation is stretched proportionnally to the undulator and chicane dispersion ratio up to $\lambda_{seed} + \Delta\lambda(z)$. Since the coherent radiation is emitted at the final density modulation period, the FEL central wavelength is found shifted by $\Delta\lambda$. Because the density modulation is not present immediately at the undulator entrance, i.e. at z = 0, this model can be further refined replacing z in equation (4) by $z - z_i$, where z_i is the 'effective' starting point of the red shifting. Systematic analysis of the simulations reveals that z_i could be defined as the location for which the coherent emission intensity becomes ≈ 0.5 times the initial seed intensity.

Figure 3(a) shows a typical evolution of the FEL spectrum along the undulator. The FEL wavelength drifts away from the initial seed wavelength, following the analytical prediction of equation (4) (solid line). The horizontal dashed line corresponds to the starting point z_i in the model. Figure 3(b) summarizes the FEL central wavelengths obtained in simulation as a function of the chicane strength for various electron beam charges and divergences. The final wavelengths are found well within the theoretical predictions.



Figure 3. (a) Normalized simulated spectrum along the undulator for a charge of 50 pC, with theoretical prediction of the peak wavelength evolution (solid white line) using a starting point z_i (dashed line) where the FEL intensity reaches 0.5 times the seed intensity. (b) Simulated spectral shift $\Delta\lambda$ versus the chicane dispersion $R_{56,c}$ for Q = 20 pC and $\sigma'_{x0,y0} = 1$ mrad (blue dot), Q = 40 pC and $\sigma'_{x0,y0} = 1$ mrad (red dot), Q = 20 pC and $\sigma'_{x0,y0} = 0.5$ mrad (blue triangle). Analytical prediction of the spectral shift for various starting points: $z_i = 0.25$ m (dashed line), $z_i = 0.5$ m (solid line), $z_i = 0.75$ m (dotted line). $R_{56,c}$ set to 1.5 mm corresponding to a value compatible with the insertion of a seed mirror in the chicane.



Figure 4. Influence of the initial energy spread on the FEL spectrum. Longitudinal energy distribution (a) at the LPA source and (b)–(c) at the undulator entrance, (c) longitudinal energy spread and current at the undulator entrance, FEL temporal profile (e) and on-axis spectrum (f) at the undulator exit. The initial energy spread is varied from $\pm 2\%$ (green), $\pm 1\%$ (blue), $\pm 0.5\%$ (red) and $\pm 0.3\%$ (black). Other parameters from table 1.

4. Analytic modeling of the FEL linewidth

The analysis of the FEL spectral content can then be pushed further. In the case of standard spontaneous emission, i.e. synchrotron radiation, the on-axis linewidth is inversely proportional to the number of undulator periods N_u : $\Delta\lambda/\lambda = 1/N_u$. An additional inhomogeneous broadening can then result from the electron beam relative energy spread σ_{γ} and/or divergence [55]. In the case of high-gain self-amplified FELs, the linewidth at saturation is essentially driven by ρ_{fel} according to $\Delta\lambda/\lambda \approx \rho_{\text{fel}}$. This linewidth can be further reduced down to the Fourier limit in the case of high-gain seeded FELs. In the low-gain case presented here, the linewidth dependencies are expected to be in an intermediate state which is worth being investigated. As shown in figure 4, varying the energy spread at the LPA source from ± 2 down to $\pm 0.5\%$ (figure 4(a)), the slice energy spread at the undulator entrance remains constant around 0.06% (figure 4(d)). This results from the strong stretching in the chicane (figure 4(b)) which cancels the contribution of the initial energy spread to the final slice energy spread. As a consequence, the final FEL linewidth is not sensitive to the initial energy spread for values above $\pm 0.5\%$ (figure 4(f)). For initial energy spread values below $\pm 0.5\%$ however, the FEL spectrum starts being affected



Figure 5. (a) Simulated on-axis spectrum of the seed at undulator entrance (dashed blue line) and exit (continuous blue line), and of the FEL at undulator exit (continuous black line). Dash lines: seed and FEL central wavelengths (maximum intensity location). (b) Simulated FEL linewidth versus beam initial divergence with $\epsilon_{x0,y0} = 0.1 \text{ mm} \text{ mrad}$ (full dots) and $\epsilon_{x0,y0} = 0.5 \text{ mm} \text{ mrad}$ (empty dots), using Q = 40 pC at $R_{56,c} = 2.5 \text{ mm}$. (c) Simulated electron beam slice energy spread inside the undulator versus beam initial divergence using same markers as for (b) with Q = 40 pC at $R_{56,c} = 2.5 \text{ mm}$ and calculated using equation (5) (black line) with $R_{522} = R_{544} = -0.23 \text{ m}$. (d) Simulated FEL linewidth versus chicane dispersion for $\sigma'_{x0,y0} = 0.5 \text{ mrad}$ -rms (blue dots), $\sigma'_{x0,y0} = 1.0 \text{ mrad}$ -rms (black dots), $\sigma'_{x0,y0} = 2.0 \text{ mrad}$ -rms (red dots), $\epsilon_{x0,y0} = 0.1 \text{ mm} \text{ mrad}$ (full dots) and $\epsilon_{x0,y0} = 0.5 \text{ mm}$ mrad (empty dots) with Q = 40 pC. Fits (dashed lines) of the linewidth with $f(a) = a/R_{56,c}$ where a = 1.7, 2.0 and 2.8 for respectively the 0.5, 1.0 and 2.0 mrad cases. Other parameters from table 1.

(figure 4(f)). Since the slice energy spread is unchanged, it is not in cause. For low initial energy spread values, the bunch duration, even after stretching in the chicane, is significantly decreased (figure 4(c)). If this duration becomes of the order of—or smaller—than the slippage length $N_u \lambda$, the radiation slips ahead of the electronic medium before reaching the end of the undulator, aborting the FEL process. The final spectrum corresponds to an FEL process with less effective undulator periods, in good agreement with the smaller red shift observed (figure 4(f)). Finally, our FEL linewidth is not driven by the initial energy spread.

The FEL linewidth σ_{λ} (figure 5(a)) is then analyzed as a function of other initial beam parameters. It reveals strongly correlated to the initial beam divergence $\sigma_{x0,y0}$ while hardly sensitive to the initial beam emittance (figure 5(b)). Actually, it is the slice energy spread at the undulator entrance σ_{γ} which is directly driven by the initial beam divergence (figure 5(c)). This can be explained deriving analytically the electron beam slice energy spread [56] at the undulator entrance:

$$\sigma_{\gamma} \approx \frac{1}{R_{56,c}} \sqrt{\sigma_{z0}^2 + 2R_{522}^2 \sigma_{x0}^{\prime 4} + 2R_{544}^2 \sigma_{y0}^{\prime 4}}$$
(5)

with $R_{522,544}$ the transport matrix coefficients and σ_{z0} the length of the initial beam (at the LPA source). This simple formula properly fits the simulation results (figure 5(c) with black line). Hence, in the case of LPA beams, because the initial beam duration is short (few microns) while the initial divergence is large (\approx mrad), the dominant contribution after transport to the slice energy spread is the initial divergence. Following, the FEL linewidth is finally dominated by this initial divergence and turns out to be easily tunable with the chicane strength $R_{56,c}$ (figure 5(d)) as expected from (equation (5)). Such behaviors can be seen as a special feature of the LPA–based FELs.

5. Double pulse modeling of the interferometric fringes

The on-axis FEL spectra (figures 3(a) and 5(a)) present fringes that are also visible in the 2D spatio-spectral distribution (figures 2(e) and (f)). Most features of this interferometric pattern can be simply understood modeling the exit radiation as a superposition of two independent Gaussian light pulses: the seed laser E_{seed} and a coherent contribution E_{coh} . Each light pulse is described independently in 3D as a combination of a Gaussian beam in the transverse plane $E_{trans}(x, y, z)$ and a linearly chirped pulse in the temporal domain $E_{temp}(\Theta)$, where (x, y) stand for the transverse plane perpendicular to the propagation axis z and $\Theta = t - z/c$ stands for the retarded time with t the position in time and c the speed of light.

5.1. Temporal electric field

The electric field of a linearly chirped pulse is given by:

$$E_{\text{temp}}(\Theta) = E_0 \exp\left[-\frac{\Theta^2}{4\sigma_t^2}\right] \exp\left[-i(\phi + \omega_0\Theta + C_1\Theta^2)\right]$$
(6)

with σ_t the rms pulse duration, ϕ a phase offset, ω_0 the optical carrier frequency and C_1 the quadratic phase, i.e. the linear chirp, coefficient and E_0 the field amplitude.

5.2. Transverse electric field

Assuming a linear polarization and a propagation in the +z direction, the transverse electric field of a Gaussian beam is given by:

$$E_{\text{trans}}(x, y, z) = E_0 \frac{w_0}{w(z)} \exp\left(-\frac{r^2}{w(z)^2}\right) \times \exp\left[i\left(kz + k\frac{r^2}{2R(z)} - \psi(z)\right)\right],\tag{7}$$

where $r = \sqrt{x^2 + y^2}$ is the radial distance from the center axis of the beam, *z* is the propagation distance from the beam's focus (or 'waist'), $k = 2\pi/\lambda_0$ is the wavenumber, E_0 the electric field amplitude (same as in equation (6)), w(z) is the radius at the plane *z*, $w_0 = w(0)$ is the waist radius, R(z) is the radius of curvature of the beam's wavefront, and $\psi(z)$ is the Gouy phase. At a propagation distance *z*, the spot size parameter *w* is given by:

$$w(z) = w_0 \sqrt{1 + \left(\frac{z}{z_R}\right)^2},$$
 (8)

where $z_R = \pi w_0^2 / \lambda$ is called the Rayleigh length. The curvature of the wavefront is zero at the beam waist. It is equal to 1/R where R(z) is the radius of curvature as a function of position along the beam, given by:

$$R(z) = z \left[1 + \left(\frac{z_{\rm R}}{z}\right)^2 \right]. \tag{9}$$

The Gouy phase of the beam is given by: $\psi(z) = \arctan\left(\frac{z}{z_{\rm R}}\right)$.

5.3. Final 3D electric field

Using equations (6) and (7) to describe the separable longitudinal and transverse components, the 3D electric field of each pulse, seed E_{seed} and coherent contribution E_{coh} , can be expressed as:

$$E_{\text{seed/coh}}(x, y, z, \Theta) = \frac{E_{\text{temp}}(\Theta) \times E_{\text{trans}}(x, y, z)}{E_0}.$$
 (10)

The pulses can have different properties such as wavelengths, chirps, waists, etc... Assuming that the FEL results from the superposition of these two fields, its electric field is given by:

$$E_{\text{fel}}(x, y, z, \Theta) = E_{\text{seed}}(x, y, z, \Theta) + E_{\text{coh}}(x, y, z + \Delta z, \Theta - \tau)$$
(11)

with Δz the difference between the seed and coherent contribution waist positions and τ the temporal delay between the seed and coherent emission. The amplitude ratio $E_{\rm coh}/E_{\rm seed}$ corresponds to the FEL 'gain'.

5.4. Spatio-spectral distributions

The spatio-spectral distributions at the undulator exit, in the near- and far-field, are simply obtained using Fourier optics [57]. The near-field spatio-spectral distribution (in the vertical plane) is obtained by Fourier transforming the 3D electric field along the temporal axis for a given horizontal position $x = x_s$ while scanning the vertical position *y*:

$$\tilde{E}(x_s, y, z_0, \nu) = \int_{-\infty}^{+\infty} E_{\text{FEL}}(x_s, y, z_0, t) e^{-i2\pi\nu t} dt.$$
(12)

The far-field, i.e. angular, distribution is the Fourier transform of E_{FEL} in the transverse plane and is given by:

$$\widehat{E}(\nu_{x}, \nu_{y}, t) = \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} E_{\text{FEL}}(x, y, z, t) e^{-i2\pi(\nu_{x}x + \nu_{y}y)} dxdy,$$
(13)

while the angular distribution is defined as:

$$\widehat{E}_{ang}(\theta_x, \theta_y, t) = \widehat{E}(\nu_x, \nu_y, t)$$
(14)

with $\theta_{x,y} = \lambda \nu_{x,y}$. The far-field spatio-spectral distribution (in the vertical plane) is then obtained by Fourier transforming equations (13) and (14) along the temporal axis at a given horizontal angle $\theta_x = \theta_{x,s} = \lambda \nu_{x,s}$:



Figure 6. Example of double pulse model calculation. On-axis intensity (a) and spectrum (b) of the seed contribution (blue), coherent contribution (black) and total FEL field (red). Near-field (c) and far-field (d) intensity distributions of the FEL field. Double pulse model with $E_{0,\text{coh}}/E_{0,\text{seed}} = 3.1$, $\sigma_{t,\text{coh}} = 45$ fs, $\phi_{\text{coh}} = -8.238 \times 10^6$, $\omega_{0,\text{coh}} = 2\pi c/161.5$ nm, $C_{1,\text{coh}} = 4.0 \times 10^{-4}$ fs⁻², waist position 0.2 m before undulator exit, $z_{R,\text{coh}} = 0.25$ m and $\tau = 50$ fs.

$$\widehat{E}(\nu_{y},\nu) = \widehat{E}(\nu_{x,s},\nu_{y},\nu) = \int_{-\infty}^{+\infty} \widehat{E}(\nu_{x,s},\nu_{y},t) e^{-i2\pi\nu t} dt.$$
(15)

Finally, the intensity distributions in the near-field I_{NF} and far-field I_{FF} for a finite slit aperture $2\delta x$ are obtained as follows:

$$I_{\rm NF}(y,\,\nu) = \int_{-\delta x}^{+\delta x} |\tilde{E}(x_s,\,y,\,\nu)|^2 \mathrm{d}x_s \tag{16}$$

$$I_{\rm FF}(\nu_{y},\,\nu) = \int_{-\delta\nu_{x}}^{+\delta\nu_{x}} |\widehat{E}(\nu_{x,s},\,\nu_{y},\,\nu)|^{2} \mathrm{d}\nu_{x,s}.$$
(17)

The double pulse model was first *tested* using a set of arbitrary input parameters. As shown for instance in figure 6, the fringes features in both the near- and far-fields are successfully generated, together with the curvature of the fringes along the vertical axis.

6. Interferometry for robust FEL pulse reconstruction

Assuming that the seed properties are known, the double pulse model (equation (11)) is used to fit the simulated near-field pattern (figures 7(a) and (b)) using a standard least mean square method. The result of the fit is shown in figures 7(c) and (d). Both near- and far-field distributions present a good agreement with the initial simulations. The few percent fit accuracy is illustrated in figure 7(e), comparing the near-field on-axis spectra (i.e. at y = 0 mm). Among the fitted parameters are found key features such as the FEL source point and Rayleigh length (respectively 0.23 m before the undulator exit and 0.25 m in the case of figures 7(c) and (d)). Finally, the reconstructed FEL intensity and phase in the time-domain are displayed in figure 7(f), again in very good agreement with the initial simulation.

Experimentally, the near-field intensity distribution is obtained placing a 2D imaging spectrometer in the image plane of a focussing lens at the undulator exit, while the far-field distribution is obtained moving the spectrometer to the focal plane of a focussing lens. Both distributions are therefore easy features to measure, meaning that combined with our double pulse model, an efficient tool for full FEL pulse reconstruction becomes available.

7. Conclusion

In conclusion, we presented simulation results of a feasible LPA–based seeded FEL. We showed that it presents original spatio-spectral distributions, with red shifted interference fringes. Those features were explained using simple considerations, allowing a deep understanding of the physical effects underlying. Moreover, we



 $E_{0,coh}/E_{0,seed} = 2.35$, $\sigma_{t,coh} = 17.3$ fs, $\phi_{coh} = -9.3 \times 10^6$, $\omega_{0,coh} = 2\pi c/161.5$ nm, $C_{1,coh} = 3.6 \times 10^{-4}$ fs⁻², waist position 0.23 m before undulator exit, $z_{R,coh} = 0.25$ m and $\tau = 50$ fs. Simulated (blue) and fitted (black) on-axis FEL spectral intensity (e) and field intensity and phase (f). Other parameters from table 1 with Q = 15 pC.

demonstrated that our double pulse model can provide with key features on the FEL source point as well as the full temporal reconstruction of the FEL pulse on a single-shot basis. We verified that this work can be extended to FELs in the visible or XUV range. Fringes can be indeed obtained as long as the red shift amplitude is of the order of the seed linewidth, and as the gain remains below two orders of magnitude.

In the framework of LPA–based FEL experiments, the double pulse model tool could first be used to characterize the FEL pulse obtained in the time domain, i.e. beyond the standard spectral characterizations. Thanks to its 2D features, it could also be used to optimize the seed spectral and transverse properties, to enhance the FEL amplification. One could then finally hope to control, using this diagnostic, the spectral and temporal properties of the FEL pulse via electron beam and seed tuning.

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