



HAL
open science

Addressing the cosmological H_0 tension by the Heisenberg uncertainty

Salvatore Capozziello, Micol Benetti, Alessandro D. A. M. Spallicci

► **To cite this version:**

Salvatore Capozziello, Micol Benetti, Alessandro D. A. M. Spallicci. Addressing the cosmological H_0 tension by the Heisenberg uncertainty. *Found.Phys.*, 2020, 50 (9), pp.893-899. <10.1007/s10701-020-00356-2>. <hal-02905943>

HAL Id: hal-02905943

<https://hal.science/hal-02905943v1>

Submitted on 24 Sep 2020

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.



HAL Authorization

Addressing the cosmological H_0 tension by the Heisenberg uncertainty

Salvatore Capozziello · Micol Benetti ·
Alessandro D.A.M Spallicci

Received: date / Accepted: date

Abstract The uncertainty on measurements, given by the Heisenberg principle, is a quantum concept usually not taken into account in General Relativity. From a cosmological point of view, several authors wonder how such a principle can be reconciled with the Big Bang singularity, but, generally, not whether it may affect the reliability of cosmological measurements. In this letter, we

S. Capozziello

- Università degli Studi di Napoli, Federico II (UNINA), Dipartimento di Fisica Ettore Pancini, Via Cinthia 9, 80126 Napoli, Italy,
- Istituto Nazionale di Fisica Nucleare (INFN), Sezione di Napoli, Via Cinthia 9, 80126 Napoli, Italy,
- Gran Sasso Science Institute (GSSI), Via Francesco Crispi 7, 67100 L'Aquila, Italy,
- Tomsk State University of Control Systems and Radioelectronics (TUSUR), Laboratory for Theoretical Cosmology, 40 prospect Lenina, 634050 Tomsk, Russia.
E-mail: capozziello@unina.it

M. Benetti

- Università degli Studi di Napoli, Federico II (UNINA), Dipartimento di Fisica Ettore Pancini, Via Cinthia 9, 80126 Napoli, Italy,
- Istituto Nazionale di Fisica Nucleare (INFN), Sezione di Napoli, Via Cinthia 9, 80126 Napoli, Italy.
E-mail: micol.benetti@infn.it

A.D.A.M Spallicci

- Université d'Orléans, Observatoire des Sciences de l'Univers en région Centre (OSUC) UMS 3116, 1A rue de la Férollerie, 45071 Orléans, France,
- Université d'Orléans, Collegium Sciences et Techniques (CoST), Pôle de Physique, Rue de Chartres, 45100 Orléans, France,
- Centre National de la Recherche Scientifique (CNRS), Laboratoire de Physique et Chimie de l'Environnement et de l'Espace (LPC2E) UMR 7328, Campus CNRS, 3A Avenue de la Recherche Scientifique, 45071 Orléans, France,
- Universidade do Estado do Rio de Janeiro (UERJ), Instituto de Física, Departamento de Física Teórica, Rua São Francisco Xavier 524, 20550-013 Maracanã, Rio de Janeiro, Brazil,
- Centro Brasileiro de Pesquisas Físicas (CBPF), Departamento de Astrofísica, Cosmologia e Interações Fundamentais (COSMO), Rua Xavier Sigaud 150, 22290-180 Urca, Rio de Janeiro, Brazil.
E-mail: spallicci@cnrs.orleans.fr

express the Compton mass as a function of the cosmological redshift. The cosmological application of the indetermination principle unveils the differences of the Hubble-Lemaître constant value, H_0 , as measured from the Cepheids estimates and from the Cosmic Microwave Background radiation constraints. In conclusion, the H_0 tension could be related to the effect of indetermination derived in comparing a kinematic with a dynamic measurement.

Keywords Heisenberg principle · observational cosmology · Hubble-Lemaître constant.

PACS PACS 98.80-k · 98.80.Es · 98.80.Jk · 03.65-w · 14.70.Bh

In the last decades, we are experiencing the so-called “precision cosmology”, which stands for the availability of numerous accurate cosmological measurements. Such a richness and precision allow confirmations of General Relativity (GR) and of the standard cosmological model, demonstrating that the universe can be our largest available laboratory. At the same time, several tensions and discrepancies emerged, indicating an increasing complexity and the need of adopting huge amounts of unknown and so far undetected dark entities. Observational data point out, substantially, a different evolution (expansion trend) of the universe at different scales. On the one hand, the increasingly accurate data coming from SNeIa allow a direct estimation of the expansion rate with high precision, deriving a value of the Hubble-Lemaître parameter $H_0 = 74.03 \pm 1.42$ [1]. On the other hand, detailed Cosmic Microwave Background (CMB) radiation maps, joined with Baryonic Acoustic Oscillations (BAO) data, point to the value $H_0 = 66.88 \pm 0.92$ [2]. This difference, named H_0 *tension*, of 4.4σ [1], is out of any conciliation and may suggest that the two different approaches, namely kinematic and dynamic, are not fully compatible. In principle, the H_0 tension can be addressed by examining alternative models to the cosmological standard one, or by trying to reduce data by further treating the systematics more and more accurately. These studies have highlighted additional model troubles, without leading to a definitive solution [3, 4, 5, 6, 7]. In conclusion, the debate consists whether we have to invoke new physics or a more detailed data analysis in view of solving the tension.

The discussion is open and it could lead to revise some fundamental assumptions as discussed, for example, in [8]. In this letter, we provide another reading of the issue by proposing, in a cosmological context, a discussion of the Heisenberg indetermination principle, possibly being the straightforward solution of the problem. In other words, the H_0 tension could be related to the concept of measurement in a cosmological context.

In order to start our discussion, it is worth recalling that the Compton wavelength expresses a fundamental limitation on measuring the position of a particle and it is inversely proportional to the particle rest mass, m_0 . In its reduced form, the Compton wavelength, λ_c is

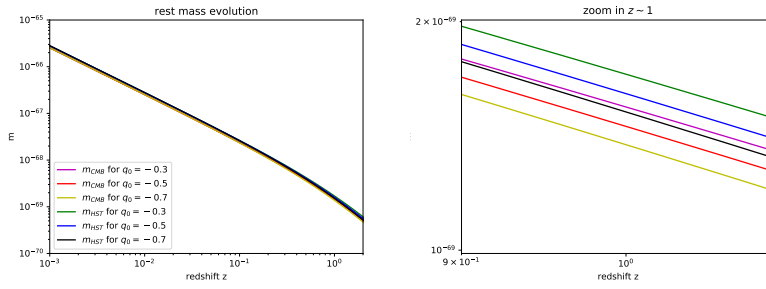


Fig. 1 Rest mass from Eq. (4) as function of the redshift, considering several values of the deceleration parameter and of the Hubble-Lemaître constant; m_{HST} refers to $H_0 = 74$ km/s/Mpc, while m_{CMB} to $H_0=67$ km/s/Mpc. In the right panel, the redshift range $[0.9 : 1.1]$ is zoomed-in.

$$\lambda_c = \frac{\hbar}{m_0 c}, \quad (1)$$

where \hbar is the reduced Planck constant and c is the speed of light. It can be interpreted as a fundamental limitation on the measure concept, since it can be related to the Heisenberg principle through

$$\Delta p \Delta x = \Delta E \Delta t = \Delta(m_0 c^2) \Delta t > \hbar/2, \quad (2)$$

considering the reduced wavelength as a length measurement, $\lambda_c = \Delta x$, and $m_0 c = \Delta p$ as a momentum measurement. Clearly, these considerations hold for any physical system where a “length” and a “momentum” can be defined.

For our purposes, let us stress that a large wavelength (or a large observation time) corresponds to low-energy photons, which means that the limit on measurability can be related to the energy (or, inversely, the wavelength) of the observed photons. This is intended for a single quantum particle. Instead, for an ensemble of particles, it is possible to independently draw mass upper limits, in presence of a macroscopic observable.

In this perspective, it is useful to recall the recent limit on the graviton mass of 2×10^{-58} kg given by LIGO, when detecting a coalescence of black holes, GW150914, for a signal duration of only 200 ms [9]. At its peak luminosity, the black hole binary emitted 10^{80} gravitons. If the graviton had an (effective) mass, then the speed of propagation would depend on the wavelength of the radiation. This would introduce dispersion in the signal, distorting the chirp waveform [10]. The distortion can be bounded using matched filtering. Conversely, a quantum estimate of the same upper limit on the graviton mass would impose an observation lasting more than 32 days.

With these considerations in mind, let us develop our discussion for homogeneous and isotropic cosmological models with redshift $z \leq 1$. As a leading principle, we will use a cosmographic approach to describe the kinematics of

the observed universe. The characteristic Compton length is tantamount to the luminosity length derived from the observations.

Although the above redshift limit can be extended adopting appropriate cosmographic series, for our purposes we use a simple Taylor series.

It is worth noticing that $z \sim 1$ is, in some sense, the limit where the kinematic description of the pure phenomenological Hubble-Lemaître law is superseded by a dynamical description according to the Friedmann-Einstein equations. Beyond this value, the simple definition of redshift, coming from the Doppler formula $z \sim v/c$, is no longer valid and the pristine expression of the Hubble-Lemaître law, $v(z) = H_0 d(z)$, with $v(z)$ the recession velocity of the astrophysical objects at distance $d(z)$, is no longer guaranteed. For higher redshifts, we may opt between appropriate (model-independent) cosmographic series as, for example, considering more detailed approximants like Padé series [11], or a dynamical (model-dependent) description of the universe related to cosmological equations of motion. In any case, considerations on the cosmological concept of “measurement” can be addressed by simply adopting a cosmographic Taylor series truncated at the second order where kinematics and dynamics of Hubble-Lemaître flow have to be confronted. The luminosity distance $d_L(z)$ is then

$$d_L(z) = \frac{zc}{H_0} \left[1 + \frac{z}{2}(1 - q_0) \right], \quad (3)$$

where q_0 is the cosmographic deceleration parameter of the universe. This latter is constrained to negative values, namely either $q_0 = -0.64 \pm 0.22$ using the joint Pantheon data for SNeIa with BAO and time-delay measurements by H0LiCOW and angular diameter distances measured using water megamasers [12] or $q_0 = -0.28 \pm 0.49$ using JLA compilation SNeIa data with BAO and observational Hubble-Lemaître values [13].

Assuming the luminosity distance as the Compton wavelength of a particle at redshift z , we can substitute Eq. (3) into Eq. (1) obtaining the corresponding effective “rest mass” as

$$m = \frac{\hbar H_0}{zc^2 \left[1 + \frac{z}{2}(1 - q_0) \right]}, \quad (4)$$

which value of depends only on H_0 and q_0 and it is decreasing for increasing values of z . It is worth stressing that such an effective mass is derived assuming the Compton length of the observed universe given by the luminosity distance.

For the sake of simplicity in developing our arguments, we choose $z = 1$ and arbitrarily fix $q_0 = -1/2$. Setting the H_0 value of the order of the latest measurements of the Hubble Space Telescope (HST) [1], $H_0 \sim 74$ Km/s/Mpc, we find the corresponding mass $m_{\text{HST}} \simeq 1.61 \times 10^{-69}$ kg (we adopt the subscript HST to indicate that this value is obtained by this specific choice of H_0). Conversely, when we refer to the Hubble-Lemaître constant value, derived by the Planck measurements of the CMB anisotropies [2], for which $H_0 \sim 67$ Km/s/Mpc, we find $m_{\text{CMB}} = 1.46 \times 10^{-69}$ kg.

In both cases, we note an order of magnitude of 10^{-69} kg. Such a value is very much lower than the most stringent and official upper limit on the photon mass as reported by the Particle Data Group for a measurement obtained in the Solar System, $m_\gamma < 10^{-54}$ kg [14]; see [15] for a critical discussion. According to laboratory tests, the upper bound is instead $m_\gamma < 1.6 \times 10^{-50}$ kg [16].

A brief reminder on the impact of massive photons on physics is due. Although the first massive theory by de Broglie-Proca of a massive photon is not gauge invariant [17], later developments have got rid of this feature, considered a limit by some. Possibly the most popular approach is the quantizable theory by Stueckelberg [18] presenting a scalar compensating field. Boulware [19] showed the renormalizability and unitarity of the Quantum Electro-Dynamics with a de Broglie-Proca massive photon. Supposing that mass rises from the spontaneous symmetry $U(1)$ breaking, gauge invariance is insured also after breaking [20], possibly determined by the Higgs mechanism [21]. Another track was pursued by Guendelman [22] who showed quantization without spontaneous symmetry breaking. In the Standard-Model Extension, a gauge invariant effective photon mass emerges [23,24]. Its mass is proportional to the Lorentz-(Poincaré-)Symmetry Violation (LSV). Incidentally, the photon energy-momentum tensor is not conserved in presence of an external Electro-Magnetic (EM) field exchanging energy with the LSV vector or tensor and the EM field.

Concerning the electric charge conservation, we observe the following. In the de Broglie-Proca modified Gauss law, the coupling of the photon mass to the scalar potential implies a density of “pseudo-charge” proportional to the squared mass, added to the ordinary charges. Since the pseudo-charges represent somewhat the Higgs field which does not couple to the ordinary charges, the two kinds of charges are conserved separately. Generally speaking, if the electric flux out of a surface gives the total electric charge enclosed, then that charge must be locally conserved: the only way charge can change is by changing the flux at the same time, and for a distant surface that flux could not change instantaneously, if the charge changed. But we cannot impede a charge to change or decay. The mechanism behind charge change or decay should be included in the physical description [25].

From Eq. (2), referring to the Heisenberg uncertainty in the energy-time form, we can assume $\Delta t = 13.8$ Gy, the age of the universe, which is the largest possible observation time. Accordingly, we obtain that the smallest measurable mass is 1.35×10^{-69} kg, which is slightly below m_{HST} or m_{CMB} but of the same order of magnitude.

At the same time, we can represent the behaviour of the “rest mass” in Eq. (4), Fig. (1), assuming different values of q_0 . If, for larger redshift the “rest mass” decreases, looking at low redshifts, the value increases, nevertheless remaining much below the upper bound on the photon mass previously given, that is for $z = 0.1$ and still $q_0 = -1/2$, we get $m_{\text{HST}} = 2.61 \times 10^{-68}$ kg and $m_{\text{CMB}} = 2.37 \times 10^{-68}$ kg.

Finally, we have that

$$\Delta m|_{z=1} = m_{\text{HST}} - m_{\text{CMB}}|_{z=1} \sim 1.5 \times 10^{-70} \text{ kg}. \quad (5)$$

It means that a measurement of the photon mass only beyond the Heisenberg uncertainty can reconcile, in principle, the H_0 tension at $z = 1$. Specifically, a mass measurement of one order of magnitude below the indetermination limit can solve the current H_0 tension at 4.4σ [1]. This fact can be formulated also by writing

$$\frac{\Delta m}{m} = \frac{\Delta H_0}{H_0} \sim 0.1. \quad (6)$$

We can infer that the tension on the H_0 measurements can be the effect of the uncertainty of measurement itself at cosmological scales. In other words, the uncertainty on the photon mass, derived from a kinematic measurement and from dynamics, can be the reason of the H_0 tension.

Finally, assuming the photon mass rigorously equal to zero is not correct from the Heisenberg principle point of view and from the recent analysis of the Standard-Model Extension [23,24]. According to this consideration, the H_0 tension could not require any new physics but only a more accurate discussion of measurements.

Acknowledgements SC and MB acknowledge the Istituto Nazionale di Fisica Nucleare (INFN), sezione di Napoli, *iniziative specifiche* MOONLIGHT2 and QGSKY. ADAMS acknowledges the Erasmus+ programme for visiting the Università di Napoli, SC the Université d'Orléans and Campus France for the hospitality. The authors thank the referees and J. A. Helayël-Neto for precious suggestions which allowed to improve the paper.

References

1. A. G. Riess, S. Casertano, W. Yuan, L. M. Macri and D. Scolnic, *Large Magellanic Cloud Cepheid standards provide a 1% foundation for the determination of the Hubble constant and stronger evidence for physics beyond LambdaCDM*, *Astrophys. J.*, **876**, 85 (2019)
2. N. Aghanim et al. [Planck Collaboration], *Planck 2018 results. VI. Cosmological parameters*, arXiv:1807.06209v2 [astro-ph.CO] (2019).
3. L. Verde, T. Treu and A. G. Riess, *Tensions between the early and the late universe*, *Nat. Astron.*, **3**, 891 (2019)
4. M. Benetti, W. Miranda, H. A. Borges, C. Pigozzo, S. Carneiro and J. S. Alcaniz, *Looking for interactions in the cosmological dark sector*, *J. Cosm. Astropart. Phys.*, **1912**, 023 (2019)
5. L. L. Graef, M. Benetti and J. S. Alcaniz, *Primordial gravitational waves and the H_0 -tension problem*, *Phys. Rev. D*, **99**, 43519 (2019)
6. M. Benetti, L. L. Graef and J. S. Alcaniz, *The H_0 and σ_8 tensions and the scale invariant spectrum*, *J. Cosm. Astropart. Phys.*, **1807**, 066 (2018)
7. J. L. Bernal, L. Verde and A. G. Riess, *The trouble with H_0* , *J. Cosm. Astropart. Phys.*, **1610**, 019 (2016)
8. Y. Kang, Y. W. Lee, Y. L. Kim, C. Chung and C. H. Ree, *Early-type host galaxies of type Ia Supernovae. II. Evidence for luminosity evolution in Supernova cosmology*, *Astrophys. J.*, **889**, 8 (2020)
9. B. P. Abbott et al., *Observation of gravitational waves from a binary black hole merger*, *Phys. Rev. Lett.*, **116**, 061102 (2016); Errata Corrige: **121**, 122902(E) (2018)

10. C. M. Will, *Bounding the mass of the graviton using gravitational-wave observations of inspiralling compact binaries*, Phys. Rev. D, **57**, 2061 (1998)
11. A. Aviles, A. Bravetti, S. Capozziello and O. Luongo, *Precision cosmology with Padé rational approximations: Theoretical predictions versus observational limits*, Phys. Rev. D, **90**, 043531 (2014)
12. S. Capozziello, Ruchika and A. A. Sen, *Model independent constraints on dark energy evolution from low-redshift observations*, Mon. Not. Roy. Astron. Soc., **484**, 4484 (2019)
13. S. Capozziello, R. D'Agostino and O. Luongo, *Extended gravity cosmography*, Int. J. Mod. Phys. D, **28**, 1930016 (2019)
14. M. Tanabashi and the Particle Data Group, *Review of particle physics*, Phys. Rev. D, **98**, 030001 (2018)
15. A. Retinò, A. D. A. M. Spallicci and A. Vaivads, *Solar wind test of the de Broglie-Proca massive photon with Cluster multi-spacecraft data*, Astropart. Phys., **82**, 49 (2016)
16. E. R. Williams, J. E. Faller and H. A. Hill, *New experimental test of Coulomb's law: A laboratory upper limit on the photon rest mass*, Phys. Rev. Lett., **26**, 721 (1971)
17. L. de Broglie, *Nouvelles Recherches sur la Lumière*, Actualités Scientifiques et Industrielles (Hermann & C^{ie}, Paris) **411** (1936)
18. E. C. G. Stueckelberg, *Théorie de la radiation de photons de masse arbitrairement petite*, Helv. Phys. Acta, **30**, 209 (1957)
19. D. G. Boulware, *Renormalizability of massive non-Abelian gauge fields: a functional integral approach*, Ann. Phys., **56**, 140 (1970)
20. C. Itzykson and J.-B. Zuber, *Quantum Field Theory* (Dover Publications, Mineola, 2012)
21. A. Scharff Goldhaber and M. M. Nieto, *Photon and graviton mass limits*, Rev. Mod. Phys., **82**, 939 (2010)
22. E. Guedelman, *Gauge invariance and mass without spontaneous symmetry breaking*, Phys. Rev. Lett., **43**, 543 (1979)
23. L. Bonetti, L. R. dos Santos Filho, J. A. Helayël-Neto and A. D. A. M. Spallicci, *Effective photon mass from Super and Lorentz symmetry breaking*, Phys. Lett. B, **764**, 203 (2017)
24. L. Bonetti, L. R. dos Santos Filho, J. A. Helayël-Neto and A. D. A. M. Spallicci, *Photon sector analysis of Super and Lorentz symmetry breaking: effective photon mass, bi-refringence and dissipation*, Eur. Phys. J. C, **78**, 811 (2018)
25. S. Nussinov, *Charge-nonconserving decays*, Phys. Rev. Lett., **59**, 2401 (1987)